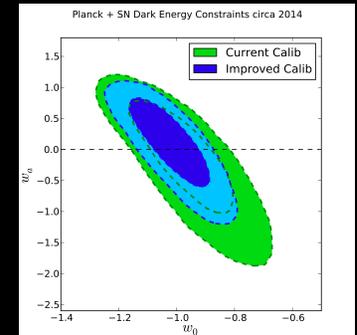
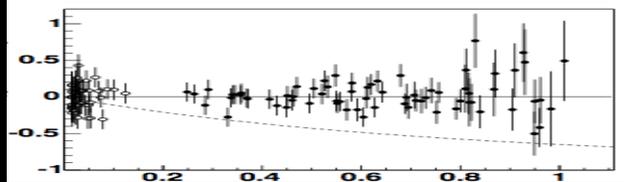
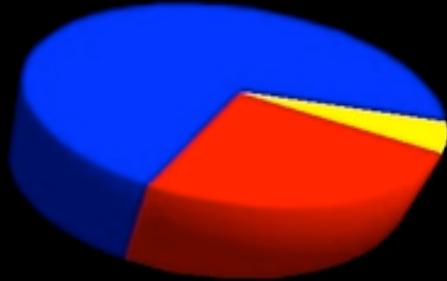
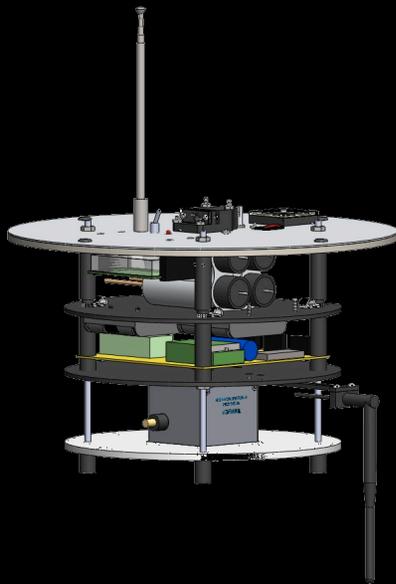


Precision Photometry via Man-Made Light Sources,



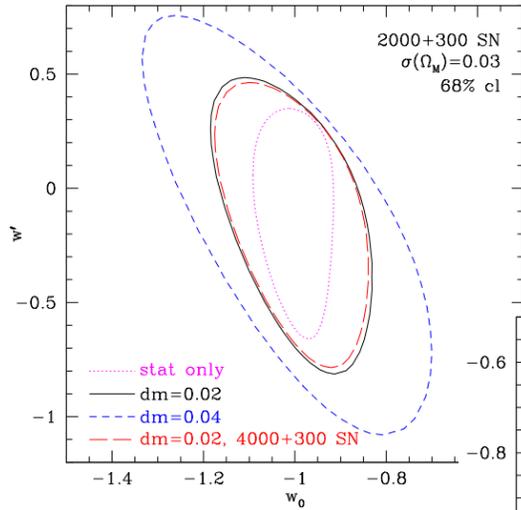
for Optical & Microwave Astronomy



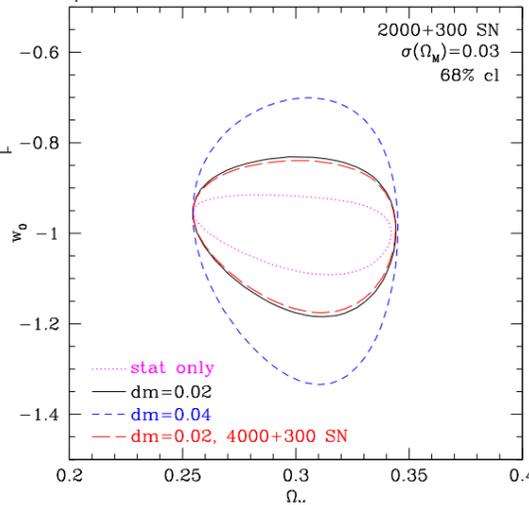
Justin Albert
Univ. of Victoria



Systematic Uncertainties Dominate Both Present & Future Dark Energy Measurements

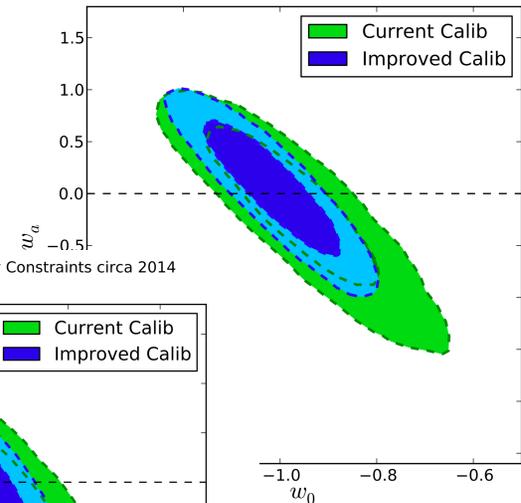


Kim, Linder, Miquel, & Mostek (MNRAS, 2004):

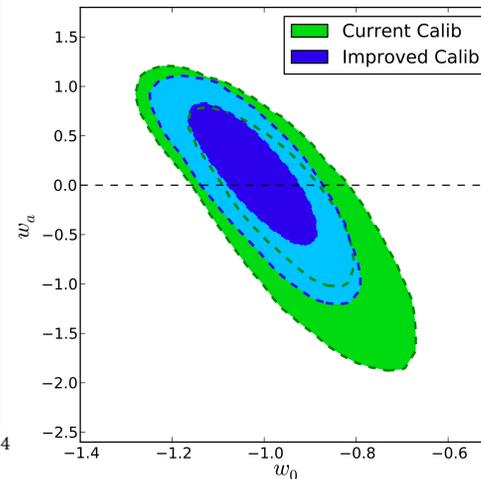


Conley (2011):

Planck + SN Dark Energy Constraints circa 2014 (flat Universe)



Planck + SN Dark Energy Constraints circa 2014

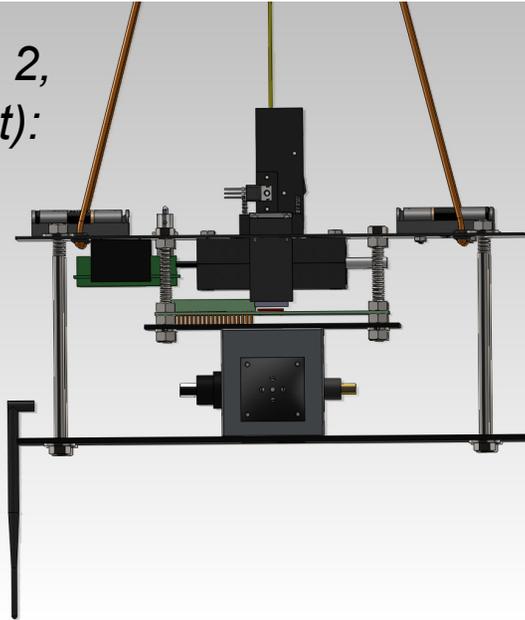


- 2% uncertainties in photometry fundamentally limit measurements of w_a to $> \pm 0.40$ and w_0 (when w_a is floated) to $> \pm 0.12$.
- Future SNIa surveys nearly **useless** for cosmology without significant photometry improvement.

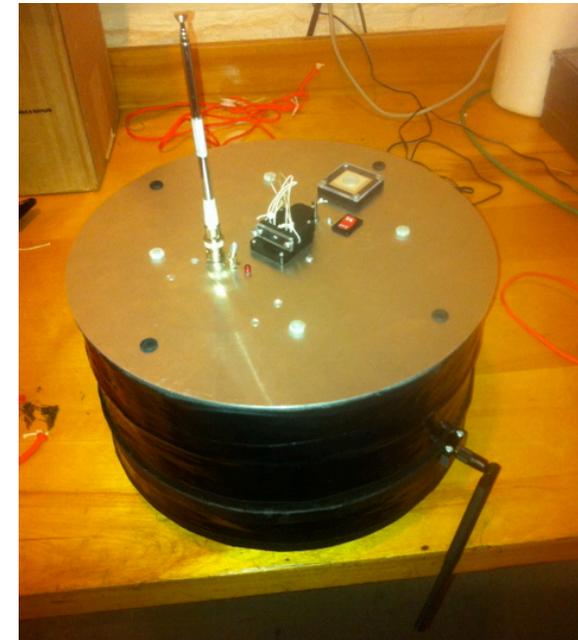
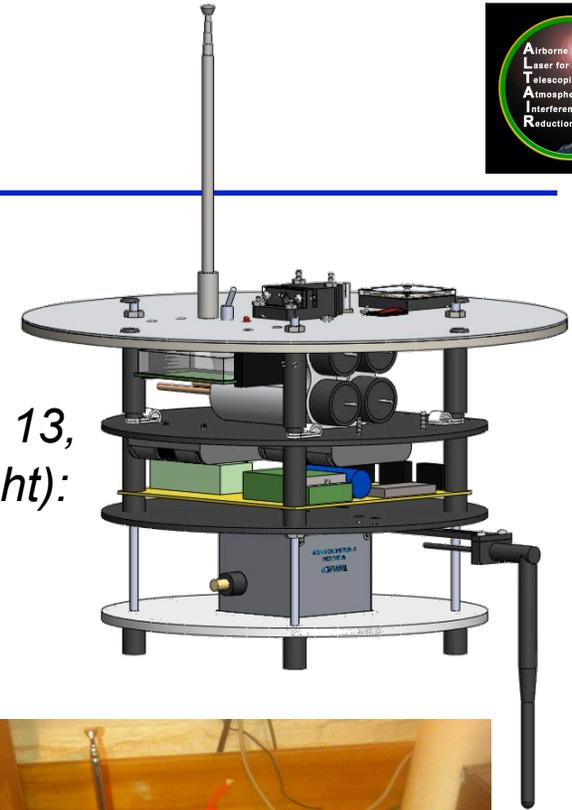


Our Payload

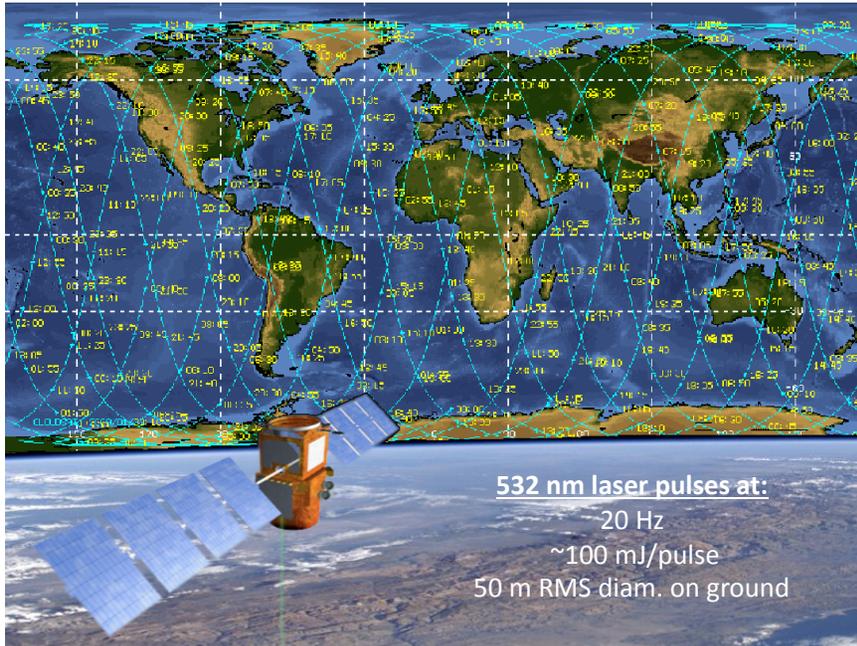
v.1 (Sept. 2, 2011 flight):



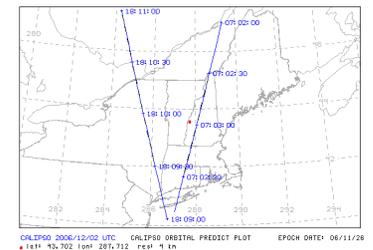
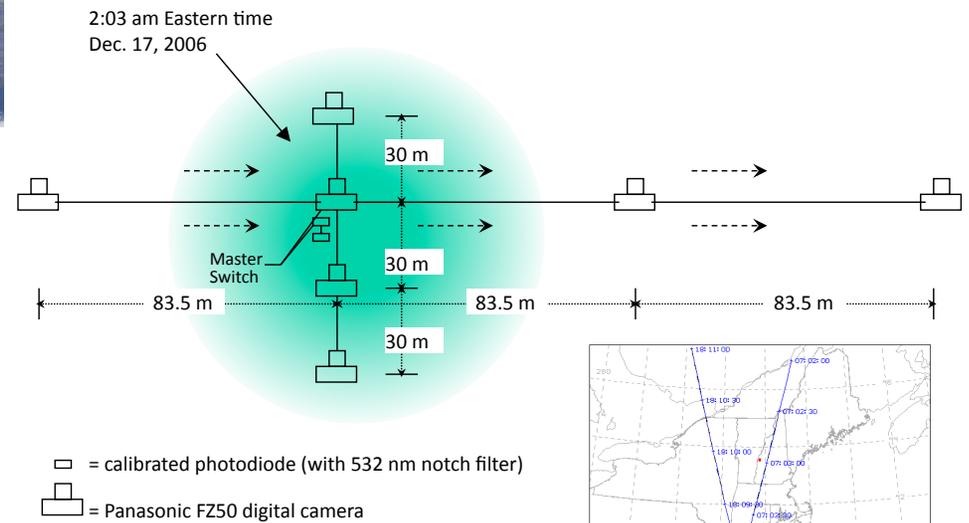
v.2 (Apr. 13, 2012 flight):



Some Pre-History, & How we Got Here



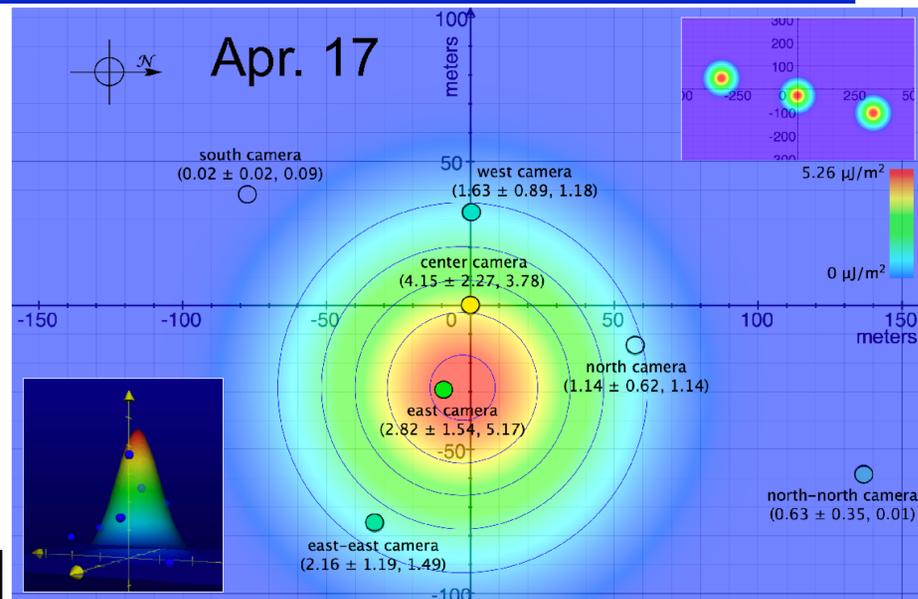
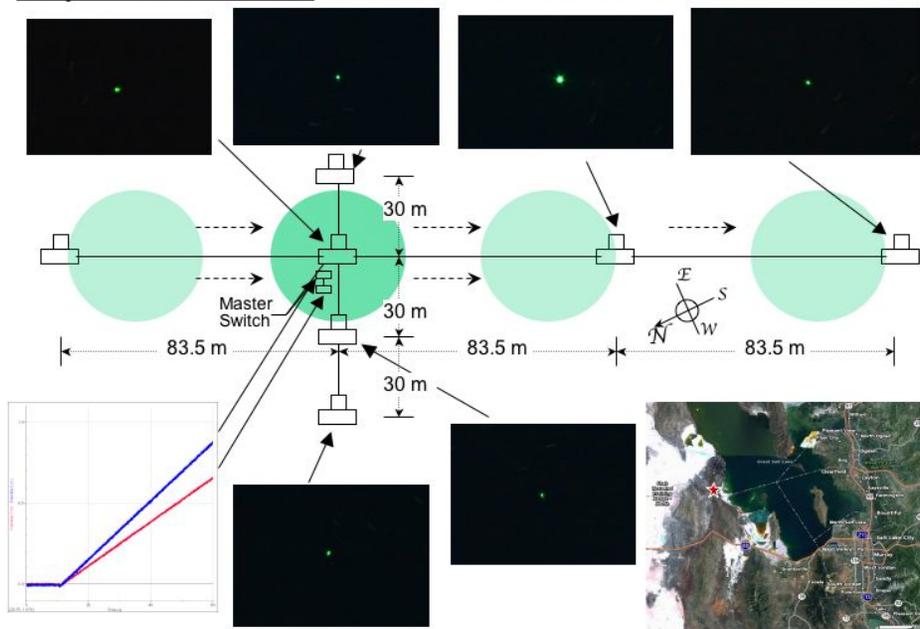
Setup for Calibrated CALIPSO Observations



Some Pre-History, & How we Got Here

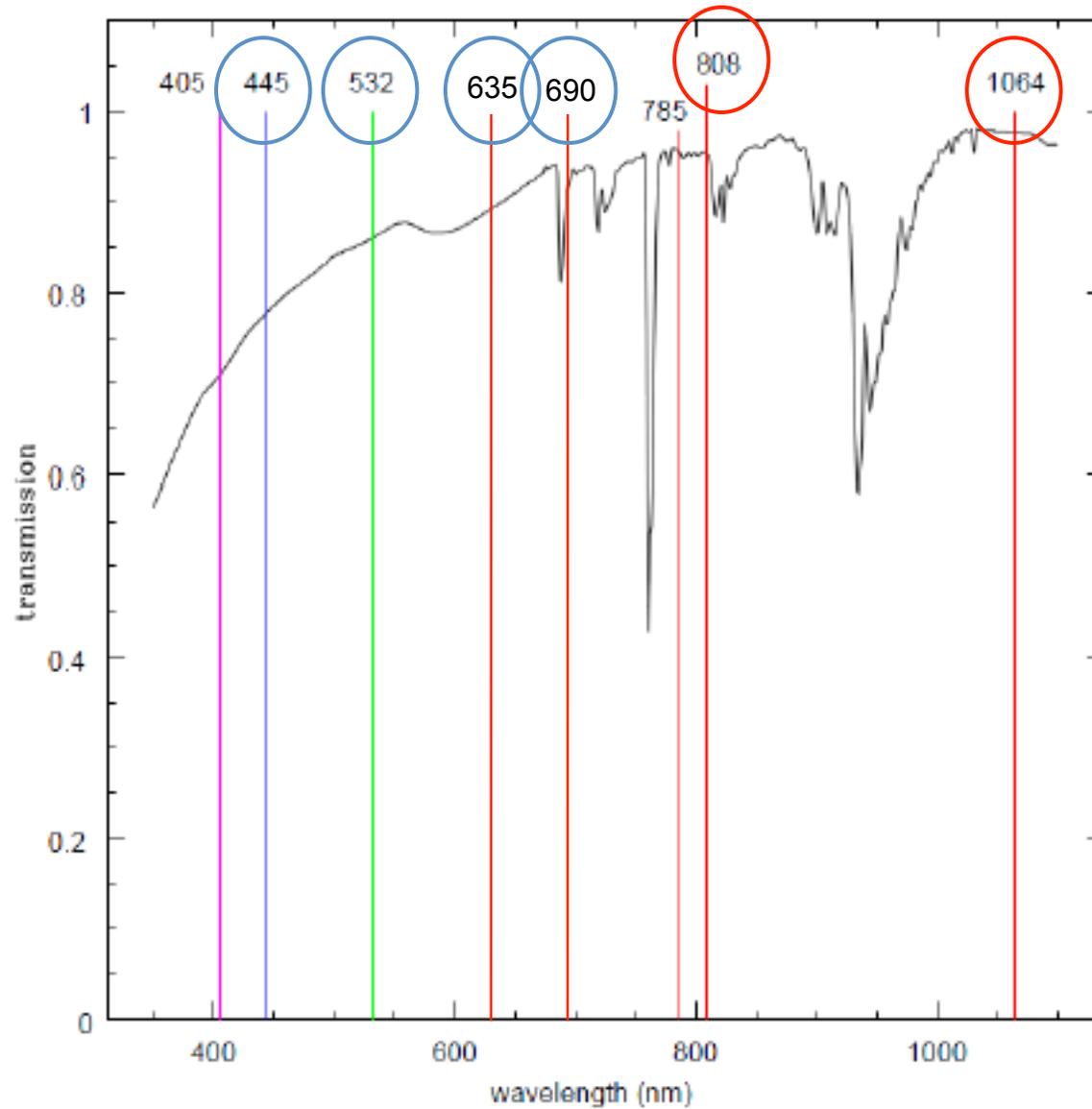
Astronomical Journal 143, 8 (2011)

May 1 observation: west of the Great Salt Lake, UT, 9:44:50 UTC



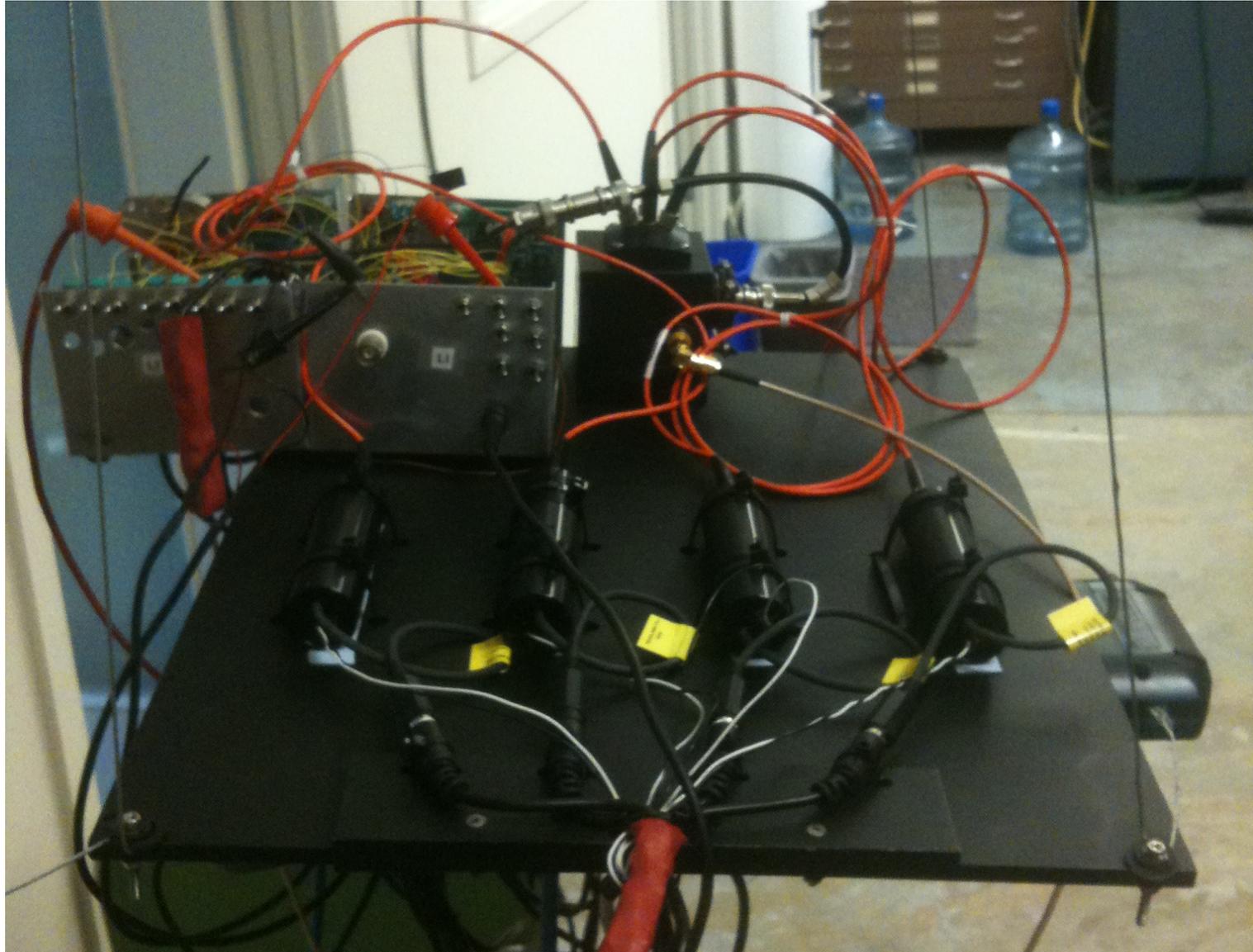


Optical Source Lines

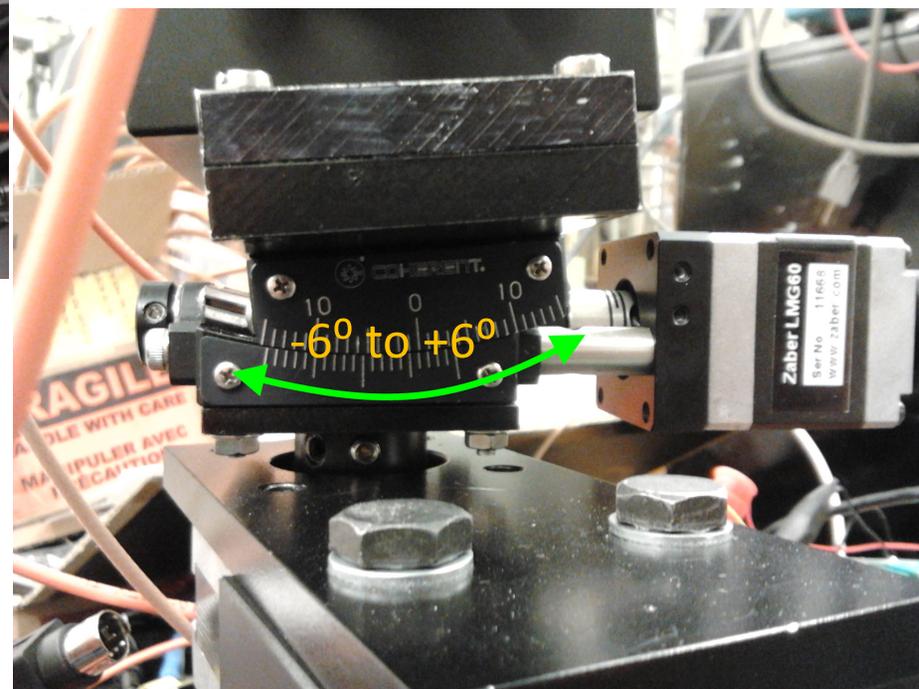
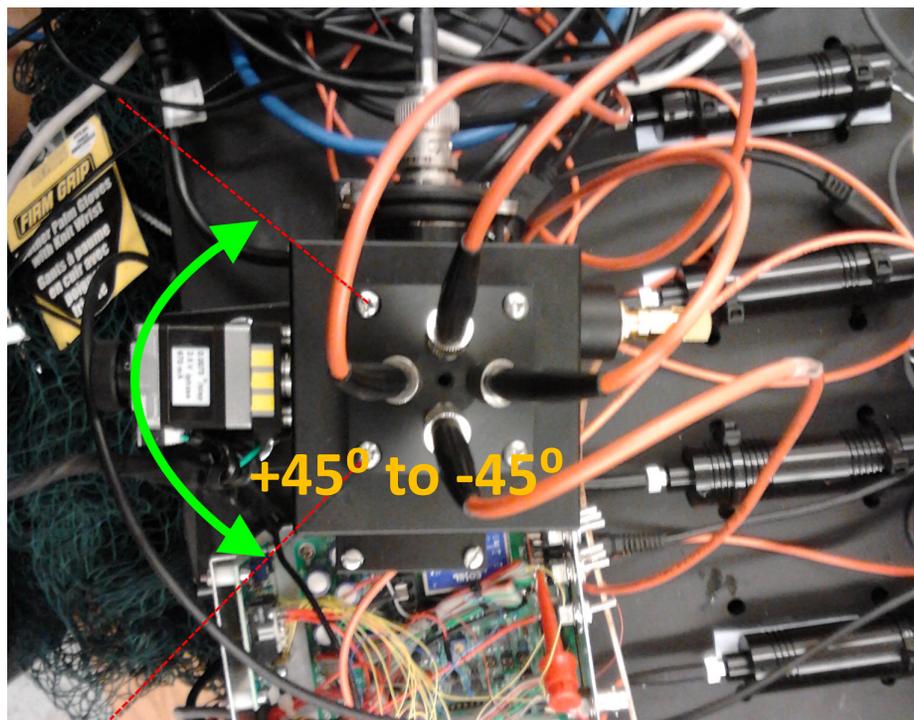




Optical Payload Test Platform



Testing the Optical Payload

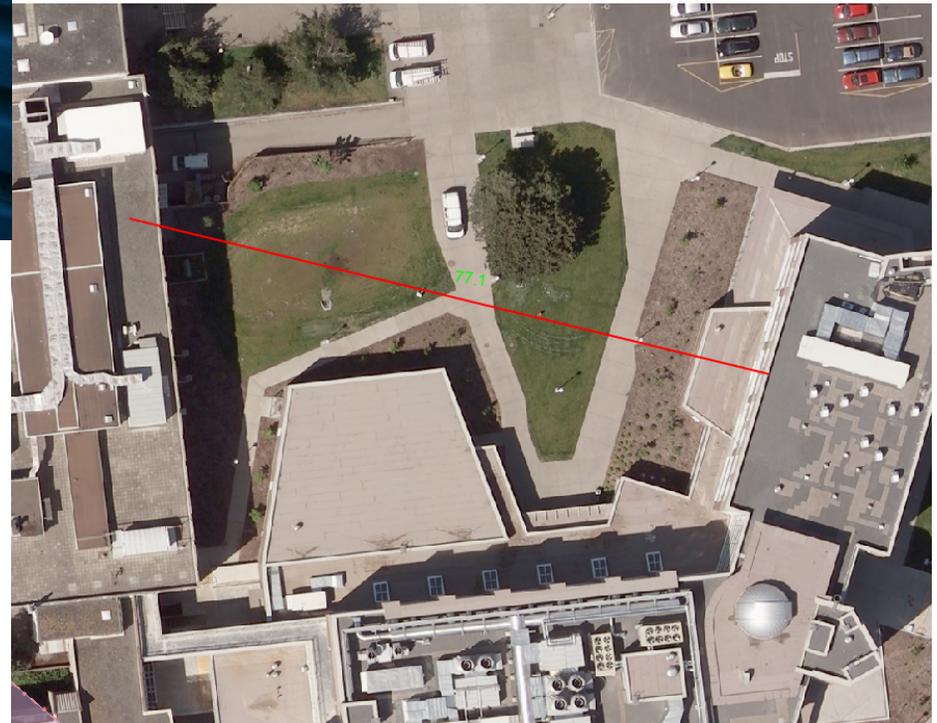


Testing the Optical Payload



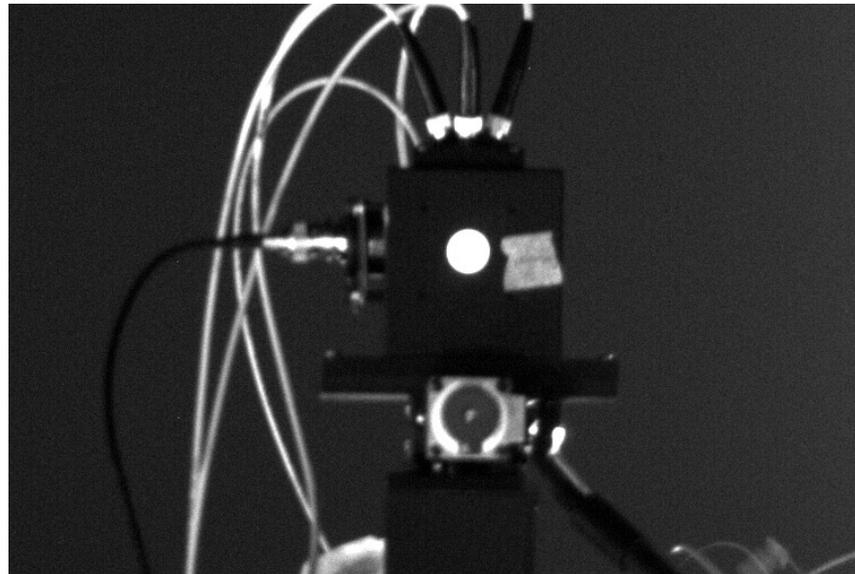
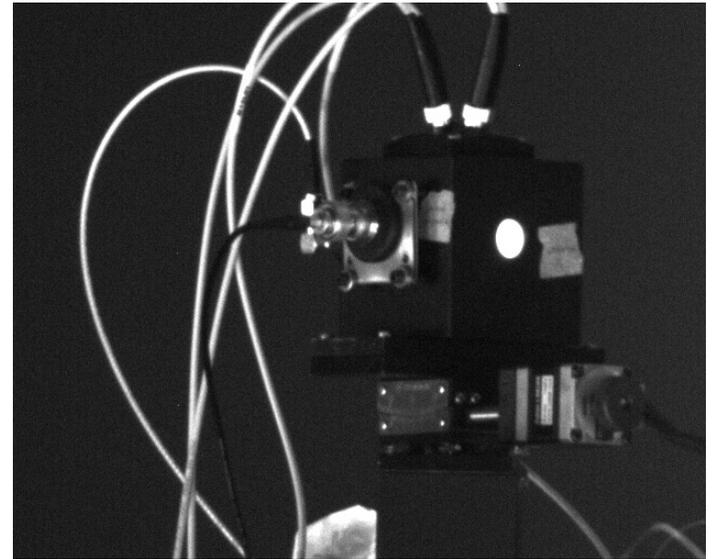
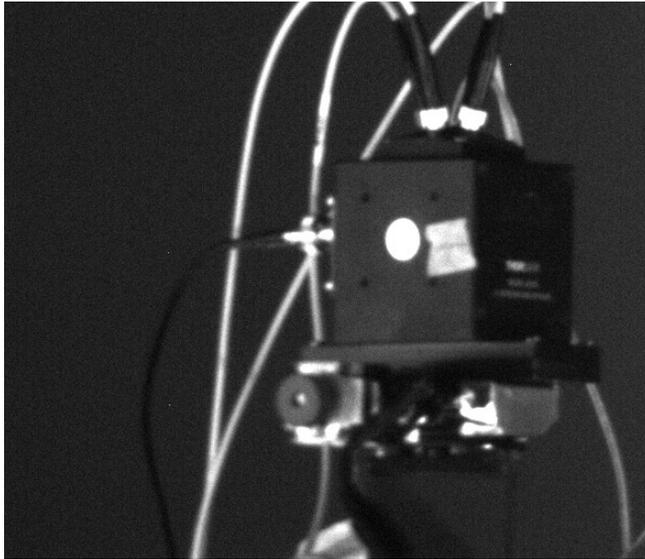


Testing the Optical Payload

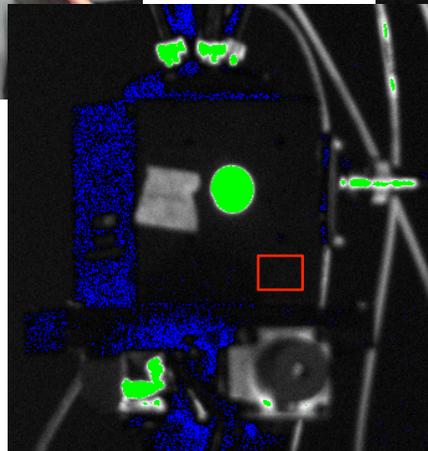
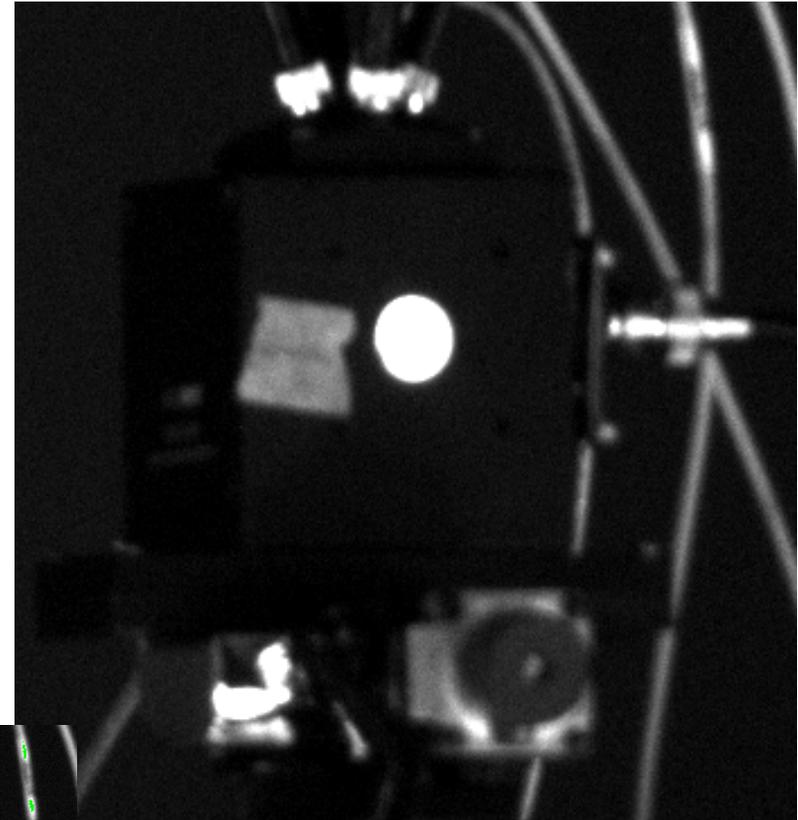
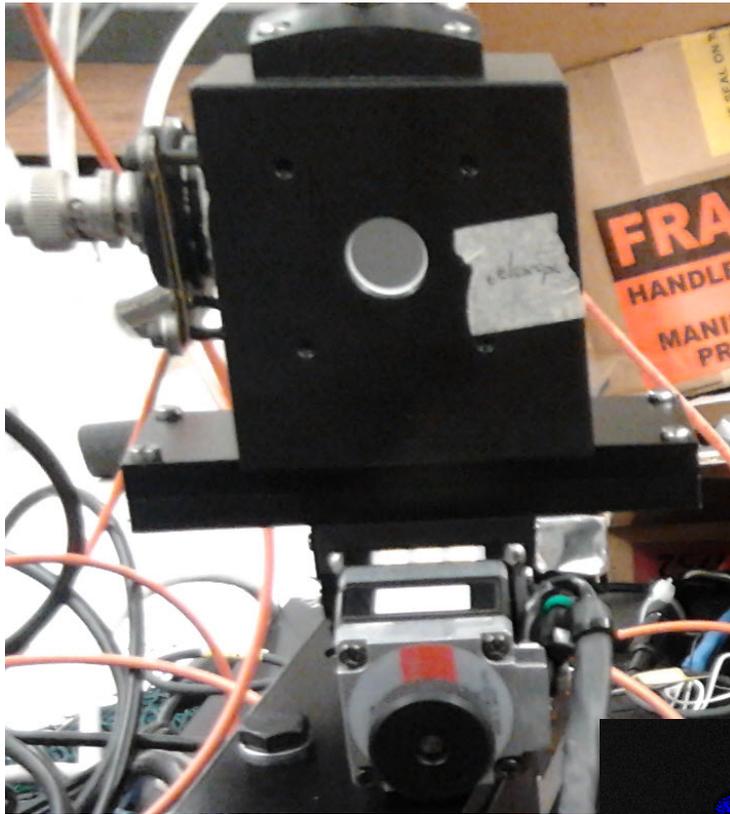




Some Near-Field (80 m) Images

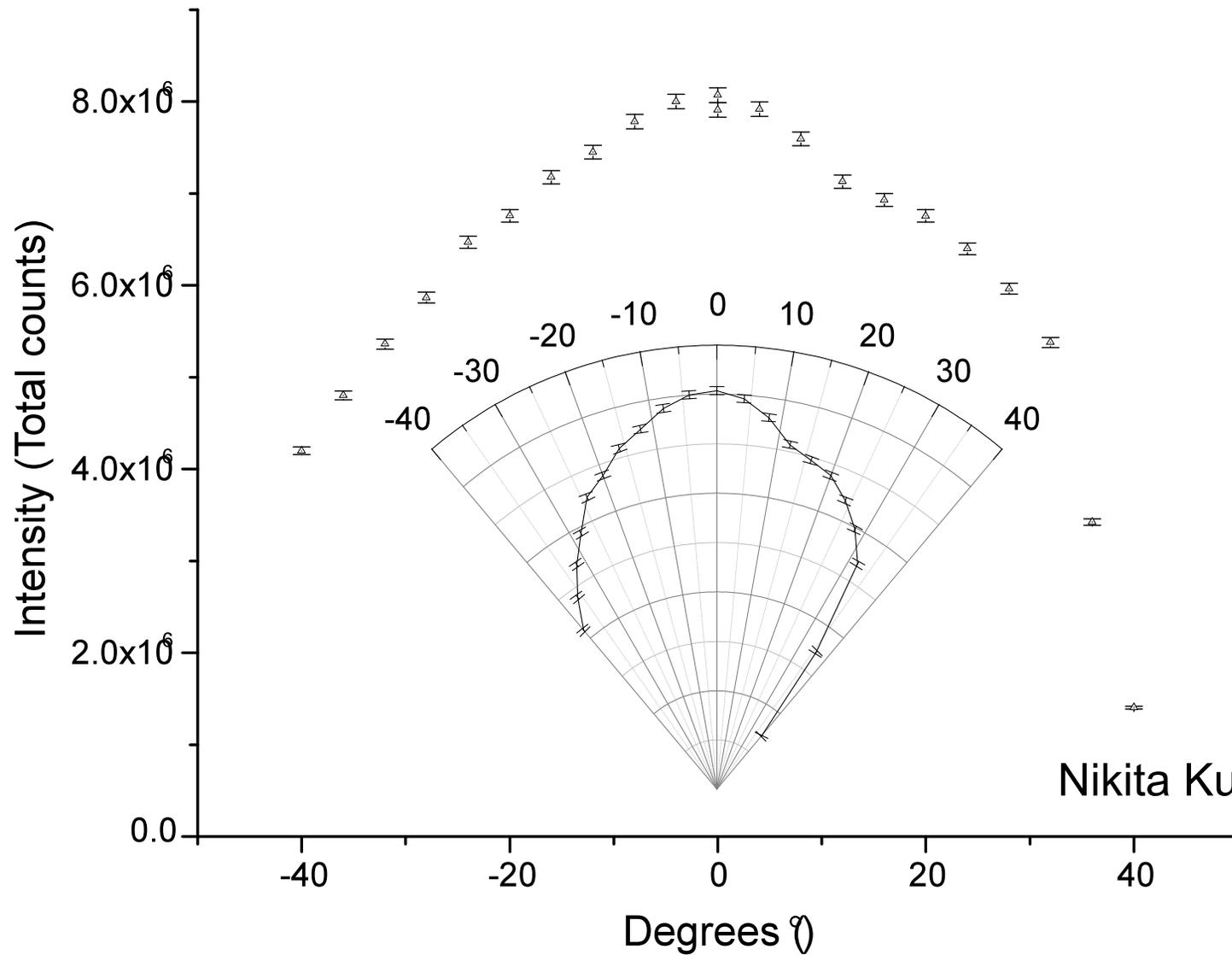


Photometry





Some Initial Results

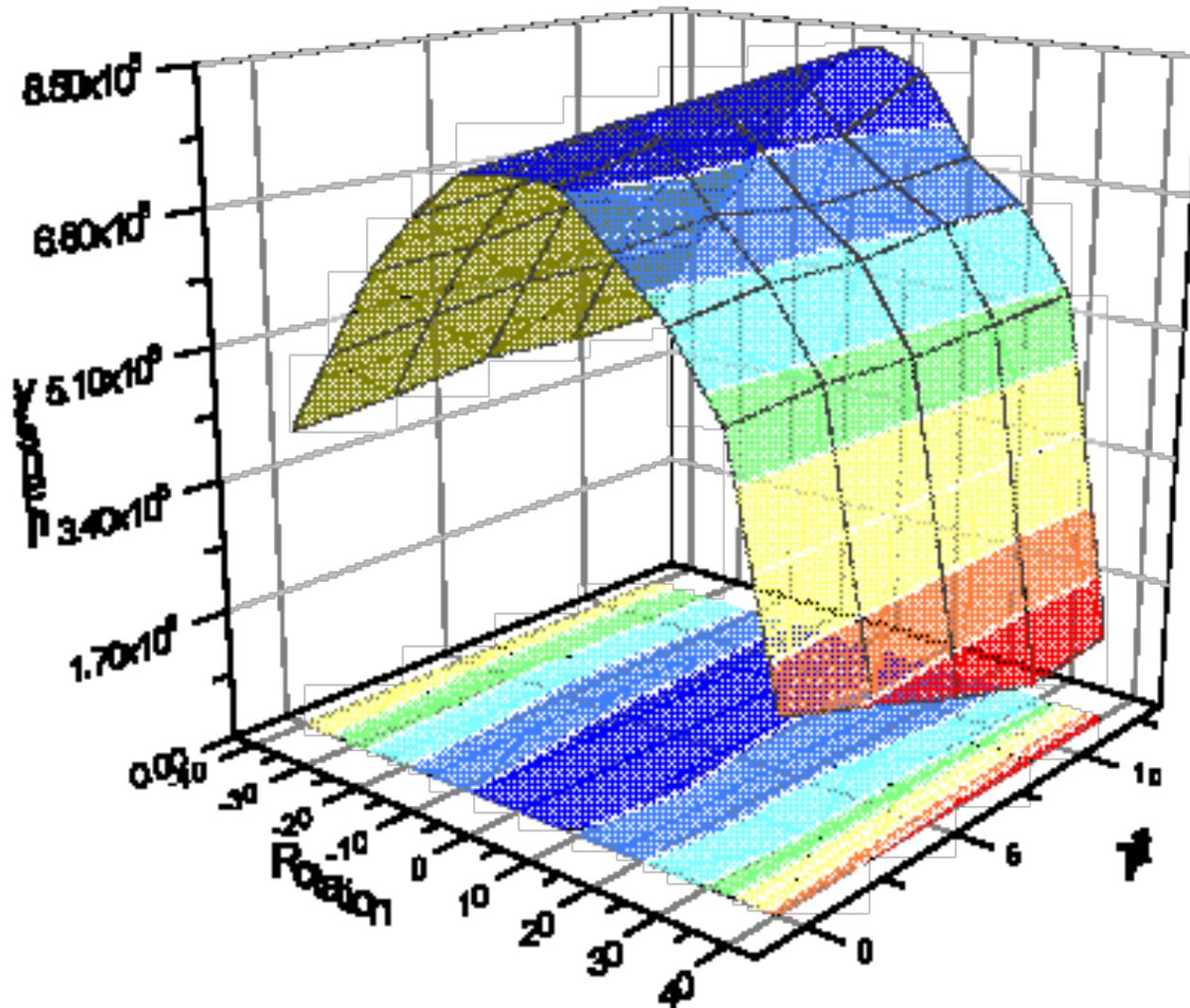


Nikita Kuklev



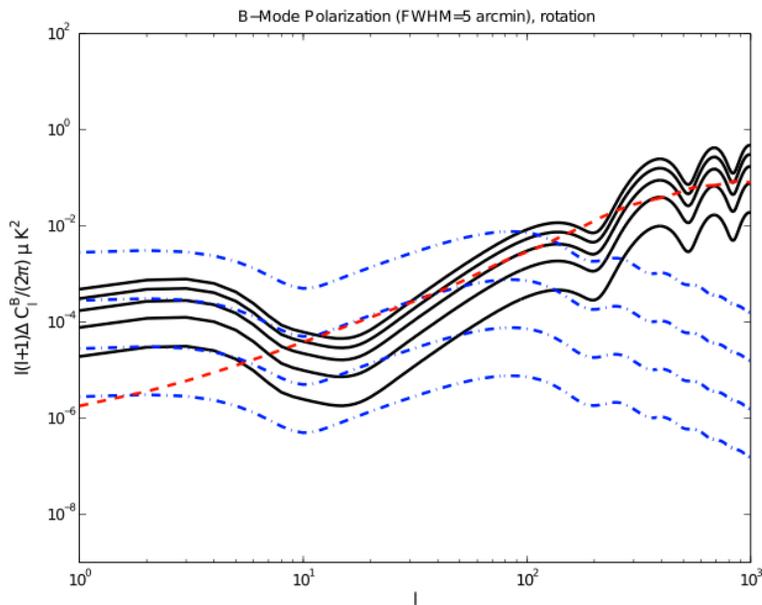


Some Initial Results

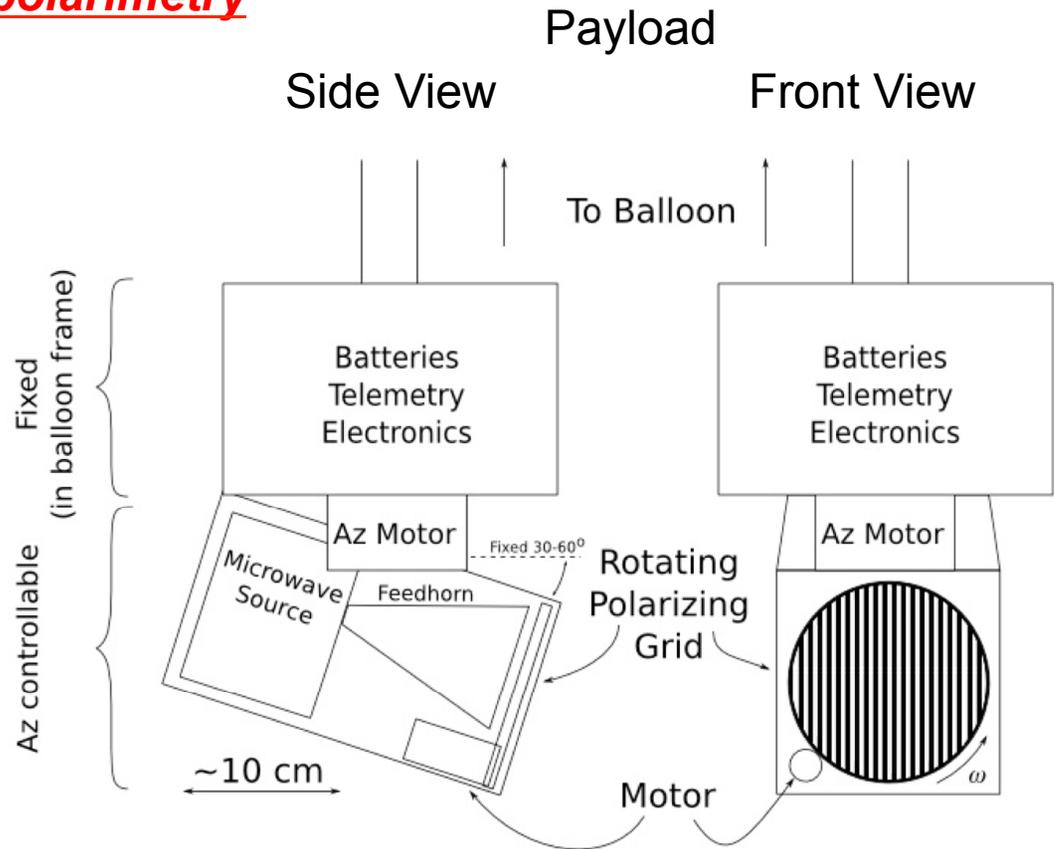


A Microwave Source ...

Calibrate polarimetry



Angular power spectrum of false CMB B-mode signals generated by errors in polarization sensitivity angles (from Shimon et al, PRD 77, 083003)



McGill University
(Matt Dobbs & Keith Vanderlinde)



Our Schedule & Plans

- Our next flight: New Hampshire night flight this spring – test observing strategy.
- Following that, flights over Mt. Hopkins NIST facility are planned.
- Canadian flights: over DAO (Vancouver Island) and Mont Megantic. Initial tests of microwave payload in ~ late 2013.
- Optical payload: work toward flights in Chile (and Hawaii) in 2014 and beyond.
- Microwave payload: flights over South Pole (SPT-Pol) and Cerro Chajnantor.



The More Distant Future: Wild Ideas

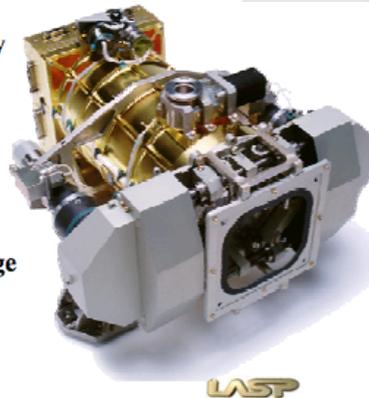
- We will achieve O(0.1%) calibration with simple NIST-calibrated photodiodes.
- Beyond that, an in-situ radiometer could achieve O(0.01%) or better ...

• Provide Total Solar Irradiance (TSI) measurements

- **Flight history:**
 - SORCE: January 25, 2003 to 2008
 - GLORY: 2008 -2013
 - NPOESS: 2013....
- **Extend the long-term TSI record.**
 - How does the Sun's output vary and what is its impact on terrestrial climate?
 - Address long-term climate change and natural variability

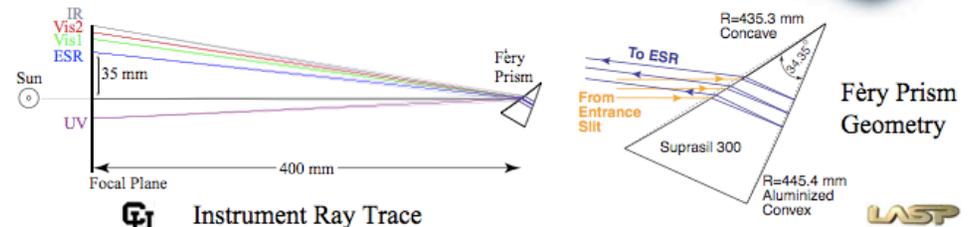
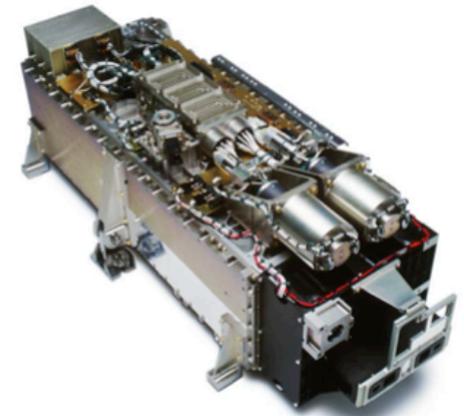
• Measurement Requirements

- **Performance**
 - Accuracy 300 ppm (0.5 W/m^2) (1σ)
 - Stability 10 ppm/yr (1σ)
 - Noise 10 ppm (1σ)
- **Report four 6-hourly averages and one daily average TSI measurement per day (Level 3 data products)**

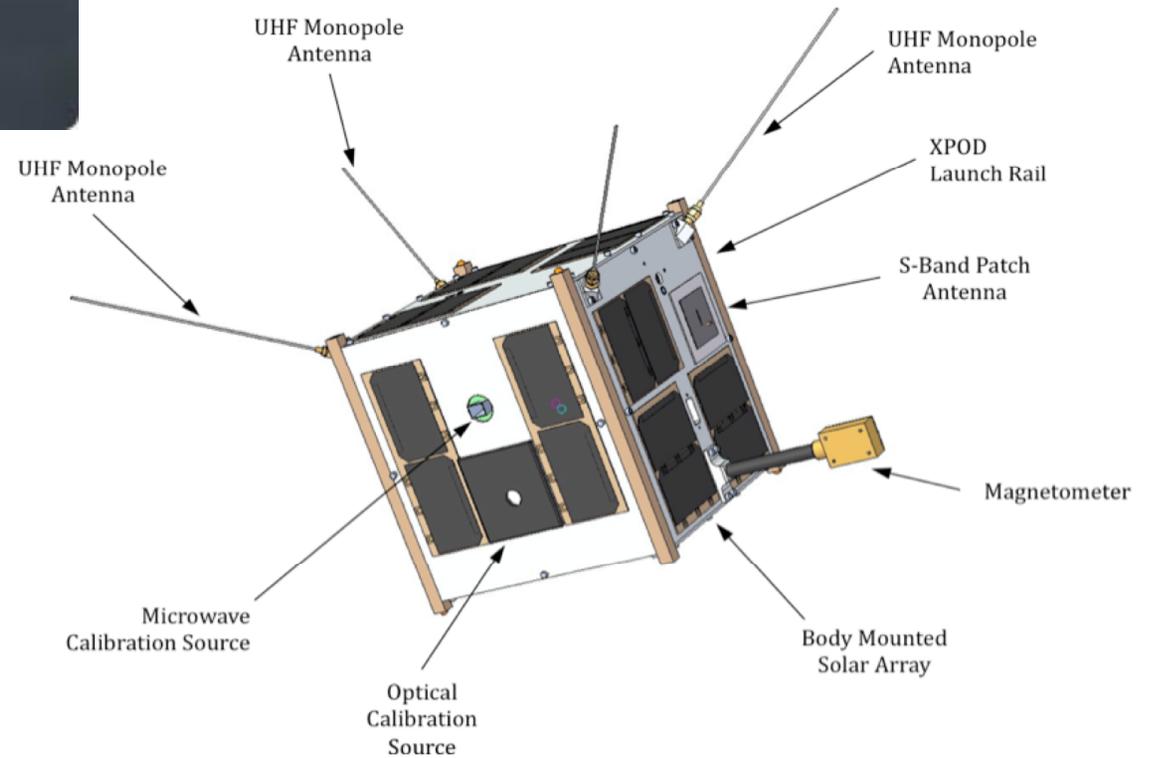


Spectral Irradiance Monitor SIM

- Measure 2 absolute solar irradiance spectra per day
- High measurement accuracy
 - Goal of 300 ppm ($\pm 1\sigma$)
 - Achieving about 2%
- High measurement precision
 - SNR ≈ 500 @ 300 nm
 - SNR ≈ 20000 @ 800 nm
- High wavelength accuracy
 - $1.3 \mu\text{m}$ knowledge in the focal plane
 - (or $\delta\lambda/\lambda < 150$ ppm)
- In-flight re-calibration
 - Prism transmission calibration
 - Duty cycling 2 independent spectrometers



The More Distant Future





Conclusions

- Man-made sources are, in principle, able to reach up to two orders of magnitude better photometric calibration precision than any natural light sources.
 - 1) Can *take them into the lab before and after use*, unlike stars.
 - 2) Can *monitor them in-situ*, in real time.
 - 3) Can be used to *calibrate white dwarfs* (and the Moon) very precisely, and on a detector-based standards scale.
 - 4) Small balloons are *inexpensive*.
 - 5) Your *choice of spectrum* & color on demand (including microwave! etc.), *brightness*, ... *location* in the sky, time of night (or day), ...
- arXiv:1101.5214 (astro-ph), AJ 143, 8 (2011).
- This is a core program for LSST: will be the primary photometry calibration method for LSST SNIa observations.

