

ESTIMATING SHAPE-LIKE SYSTEMATICS FOR THE LARGE EXTRA DIMENSIONS SEARCHES IN DUNE

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June 22nd, 2021



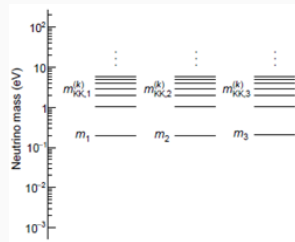
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Model signatures/consequences

LED model (Davoudiasl et. al 2002) :

- In this model, three bulk right-handed neutrinos coupled (via Yukawas's) to the three active brane neutrinos.
- After compactification of the effective extra dimension, from the four dimensional (brane) point of view, the right-handed neutrino appears as **an infinite tower of sterile neutrinos** or Kaluza-Klein modes.



Phenomenological consequences:

- The **sterile-active mixings** and the new oscillation frequencies modify the active 3ν -oscillations therefore **distorting the neutrino event energy spectrum**.
- Departures from the standard oscillations due to the existence of LED can then be probed at neutrino oscillation experiments (Long & Short baselines).

Vacuum probabilities

Three-active neutrino oscillation probability:

$$P_{\nu_\alpha \rightarrow \nu_\beta} = \left| \sum_{k=1}^3 U_{\alpha k}^* U_{\beta k} \exp\left(-i \frac{m_k^2}{2E}\right) \right|^2$$

LED oscillation probability, n -KK modes:

$$P_{\nu_\alpha \rightarrow \nu_\beta} = \left| \sum_{k=1}^3 \sum_{n=0}^{\infty} U_{\alpha k}^* U_{\beta k} (L_k^{0n})^2 \exp\left(-i \frac{(\lambda_k^{(n)})^2}{2ER^2}\right) \right|^2$$

& $\lambda_k^{(n)}$ is obtained from $\lambda_k^{(n)} - \pi(m_k^D R)^2 \cot(\pi \lambda_k^{(n)}) = 0$ with $\lambda_k^{(n)} \in [n, n + 1/2]$. We can then make the identification:

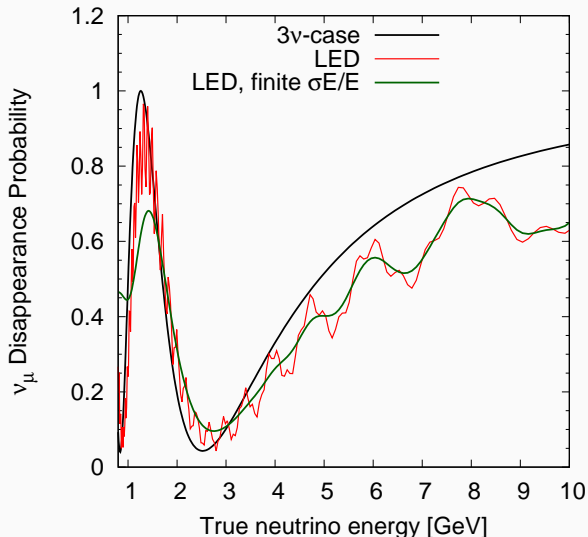
$$m_k^{(n)} = \frac{\lambda_k^{(n)}}{R} \xrightarrow{n \gg 1} \frac{n}{R}, \text{ and for the 'modified' mixing } U_{\alpha k} L_k^{0n}$$

Four free parameters m_1^D, m_2^D, m_3^D and R in the theory, reduced to two m_0 and R .

For ' $n = 0$ ' and ' $m^D R \ll 1$ ', 3ν -flavor phenomenology must be satisfied Davoudiasl et. al 2002.

Main features

DUNE FD; $R=1\mu\text{m}$, $m_0=0$



Most active (sterile) case corresponds to $n = 0$ ($n \ll 1$). The standard 3ν -neutrino oscillations are recovered in the limit $R \rightarrow 0$.

- Global reduction of survival probabilities, which is typically noticeable at high energies (Machado et. al 2011).
- Appearance of modulations and fast oscillations to Kaluza-Klein states.
- These **shape-like features** can be exploited at the analysis level. This have been done in MINOS (2016).
- Sensitivity analyses for several osc. Exps (Machado et. al 2011), IceCube (Esmaili et. al. 2014), DUNE (Berryman et. al 2016... “revamped” for DUNE FD TDR & ND CDR), and SBN (Stenico, DVF, Peres 2018).

Previous DUNE setup

$40kt \times (3.5yr(\nu) + 3.5yr(\bar{\nu})) \times 1.07MW = 300 kt MW \text{ years of exposure}$

Information considered in the analysis:

- Signal: CC, ν and $\bar{\nu}$, appearance and disappearance oscillation channels included in the analysis.
- Only FD information is considered, but ND fixes the flux normalization.

Systematics

T. Alion et. al. [arxiv:1606.09550](https://arxiv.org/abs/1606.09550) → **First GLoBES files release.**

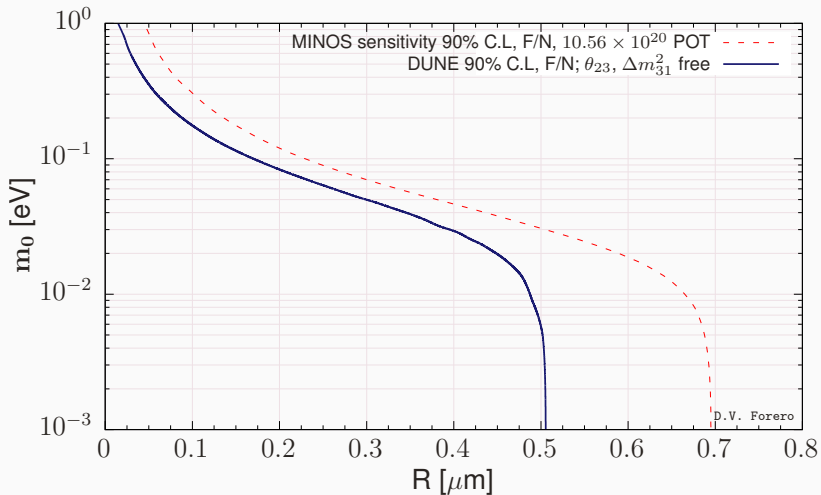
- Signal normalization systematical errors:
 $\sigma(\nu_e) = 0.02$, $\sigma(\bar{\nu}_e) = 0.02$, $\sigma(\nu_\mu) = 0.05$, $\sigma(\bar{\nu}_\mu) = 0.05$.
- Background normalization systematical errors:
 $\sigma(\nu_\mu) = 0.05$, $\sigma(\nu_e) = 0.05$, $\sigma(\nu_\tau) = 0.2$, $\sigma(\bar{\nu}_e) = 0.05$ & $\sigma(NC_{dis}) = 0.1$.
- At this point, **bin-to-bin uncorrelated systematics (or SHAPE syst.) not included!**

Fluxes

- The “Optimized Engineered Nov2017”.

DUNE Sensitivity to LED; 300 kt-MW-years of exposure

DUNE TDR arxiv:2002.03005



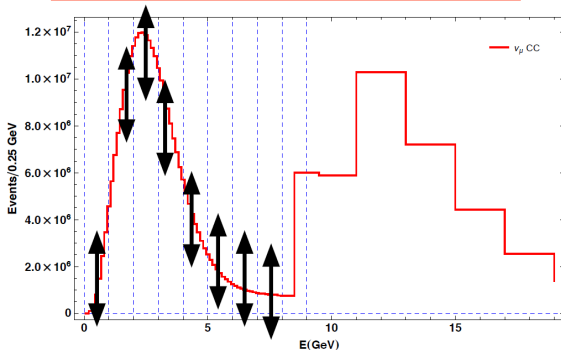
Thanks to [S. De Rijck](#) we can show MINOS sensitivity result (Asimov data).

Work in progress...

Shape systematics in GLoBES

Define energy points, p_E , and size of the systematic, s , at each p_E

E.g. One uncorrelated systematic every GeV



Current DUNE setup

$40kt \times (6.5yr(\nu) + 6.5yr(\bar{\nu})) \times 1.2MW = 624 \text{ kt-MW-yr}$ of exposure $\equiv 10 \text{ yrs(staged)}$

Information considered in the analysis:

- Signal: CC, ν and $\bar{\nu}$, appearance and disappearance oscillation channels included in the analysis.
- Only FD information is considered, but ND fixes the flux normalization.

Systematics

B. Abi, et. al, [arxiv:2103.04797](https://arxiv.org/abs/2103.04797) → Latest GLoBES files: $E_{\text{rec}} \text{ binwidth} = (\text{TDR binwidth})/2$.

- Signal normalization systematical errors:
 $\sigma(\nu_e) = 0.02$, $\sigma(\bar{\nu}_e) = 0.02$, $\sigma(\nu_\mu) = 0.05$, $\sigma(\bar{\nu}_\mu) = 0.05$.
- Background normalization systematical errors:
 $\sigma(\nu_\mu) = 0.05$, $\sigma(\nu_e) = 0.05$, $\sigma(\nu_\tau) = 0.2$, $\sigma(\bar{\nu}_e) = 0.05$ & $\sigma(NC_{dis}) = 0.1$.
- **bin-to-bin uncorrelated systematics (or SHAPE syst.) included**, as explained in slide 9.

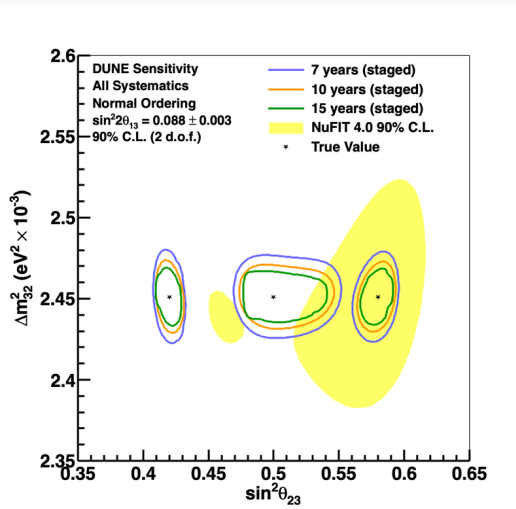
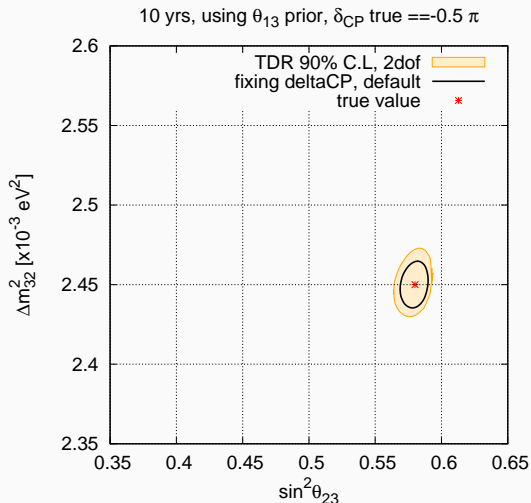
Fluxes

- The “Optimized Engineered Nov2017”.

Estimating the level of the 'shape' systematics

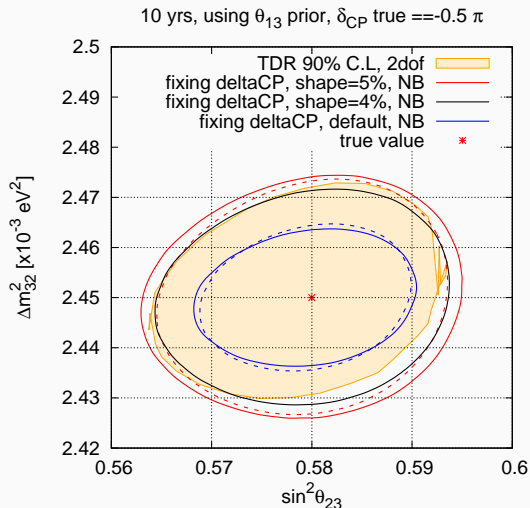
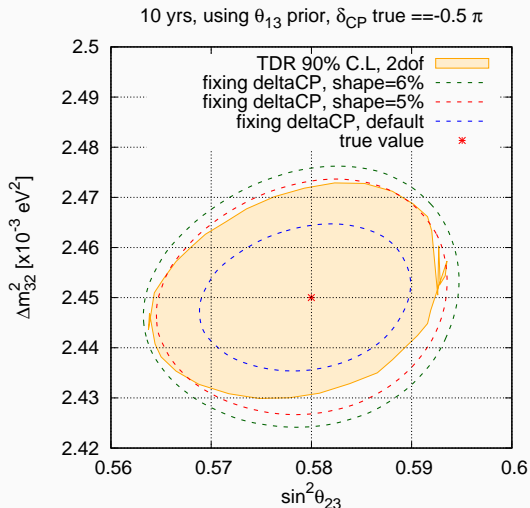
Atmospheric plane, the importance of the shape systematics

LBL phys. Potential of the DUNE Exp. arxiv:2006.16043, FIG. 26



Estimating the level of the 'shape' systematics

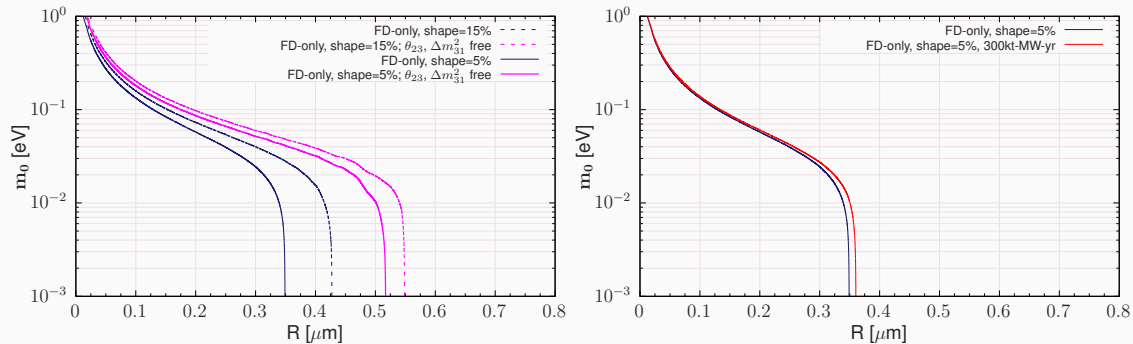
ZOOMING IN the Atmospheric plane, when using TDR binning result quoted as 'NB'



DUNE Sensitivity to LED, preliminary results

FD-Only, with TDR binning & 2 KK modes

624 kt-MW-years of exposure \equiv 10 yrs (staged). DUNE 90% of C.L for 2 d.o.f:



In coll. with A. Sousa, E. Fernandez-M, M. Blennow & S. Rosauero, as part of the **DUNE BSM Physics WG**

Present & Near Future Plans

Presented at the DUNE Coll. Meeting in May 2021

- We (A. Sousa, E. Fernandez-M, M. Blennow & S. Rosauero), and also Michael Wallbank, Jeremy Hewes and Callum Wilkinson had discussions on what it is needed to include a realistic set of systematics for the 2-detector analysis.
- Callum Wilkinson and Elizabeth Worcester provided us with the FULL systematical errors used in the 3-flavor neutrino analysis (for the TDR) in the form of covariance matrices (using CAFAna).
- A less than a week ago, Michael Wallbanck, Jeremy Hewes & Herilala Razafi 'translated' the covariance matrices into a format that can be included into our LED analysis in GLOBES.

Plans:

- To understand the departures from 3-flavor results in [arxiv:2006.16043](https://arxiv.org/abs/2006.16043), FIG. 26 obtained when estimating of the level of systematics, using GLOBES files (first procedure).
- To implement the covariance matrices for the 2-detector analysis including LED (second procedure).
- First procedure is considered as a cross check of the second one, at least for the FD-only sensitivity to LED.

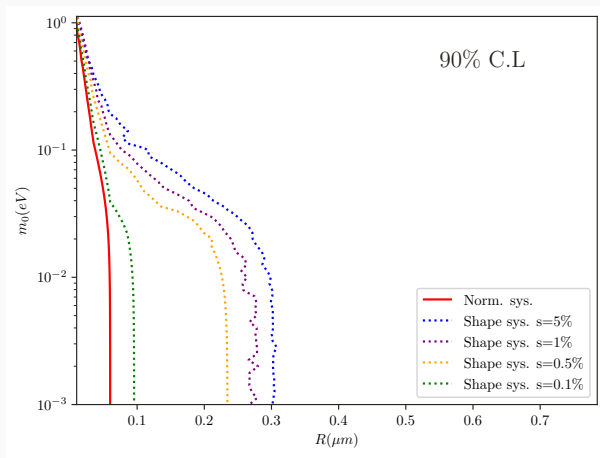
THANK YOU FOR YOUR ATTENTION!

Back up

Towards a two-detector fit

First results for 2 KK modes, with the old binning

Including a shape-like systematic error in the signal (uncorrelated between detectors) in the ND.



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