NuSTEC introducing: Expanding our palette

Improving the art of neutrino nuclei modelling with charged lepton scattering data

Implications & Motivation for DUNE

Alfons Weber for the DUNE Collaboration

NuSTEC Workshop on Electron Scattering

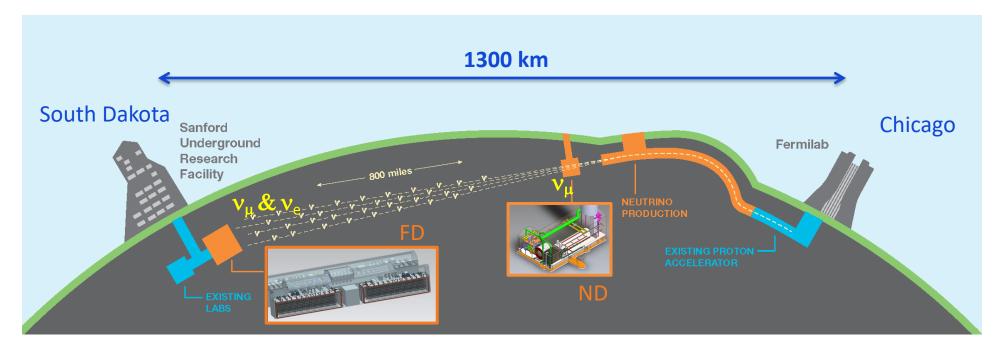






DUNE General Setup

- LBNF/DUNE will consist of
 - > An intense v-beam fired from Fermilab 1.2 MW upgradeable to 2.4 MW
 - Highly capable near detector complex
 - > A >20 kt fiducial mass LAr TPC far detector underground at SURF
 - A cavern for a full 40 kt detector system

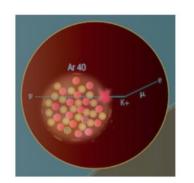




DUNE Physics Program

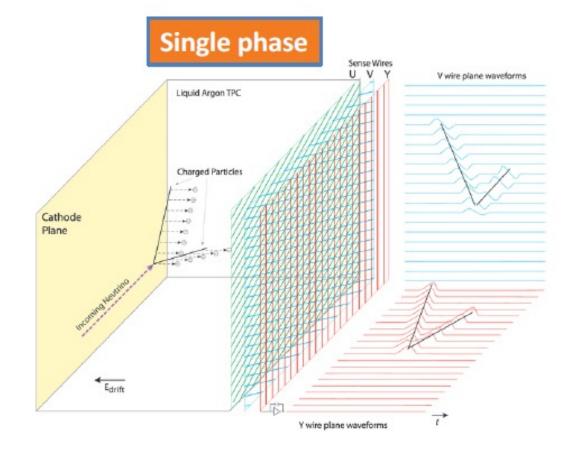
- Neutrino Oscillations
 - Search for leptonic CP violation
 - Determine neutrino mass ordering
 - Precision PMNS measurements
- Supernova Physics
 - Observation of time and flavour profile provides insight into collapse and evolution of supernova
 - Unique sensitivity to electron neutrinos
 - Baryon number violation
 - Predicted by many BSM theories
 - > LAr TPC technology well-suited to certain proton decay channels $(e.g., p \rightarrow K+\nu)$
 - > Δ (B-L) ≠ 0 channels accessible (*e.g.*, n→n)

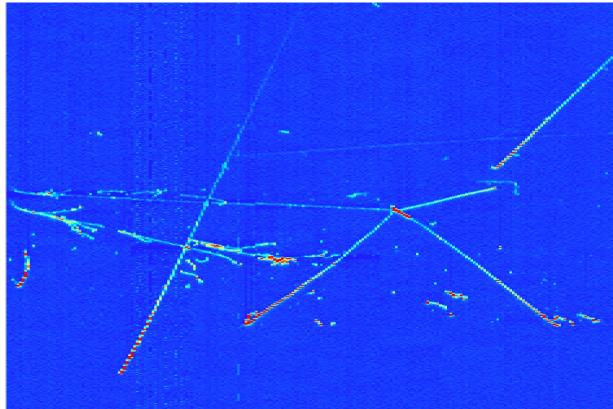




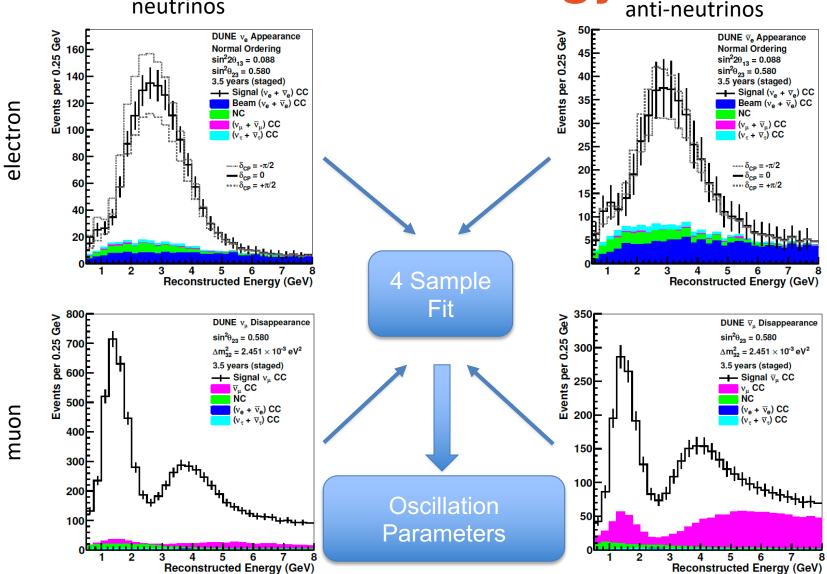
Far Detector (LArTPC)

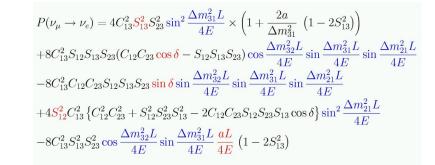
 $E_{\nu}^{rec} = E_{l}^{rec} + T_{p}^{rec} + T_{\pi^{+}}^{rec} + E_{\pi^{0}}^{rec} + T_{other}^{rec}$

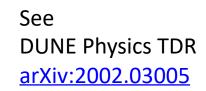














Measurement Strategy

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Measuring Oscillations

Oscillation probabilities

$$P_{\nu_{\mu} \to \nu_{e}}(E_{\nu}) = \frac{\phi_{\nu_{e}}^{far}(E_{\nu})}{\phi_{\nu_{\mu}}^{far,no-osc}(E_{\nu})} = \frac{\phi_{\nu_{e}}^{far}(E_{\nu})}{\phi_{\nu_{\mu}}^{near}(E_{\nu}) * F_{far/near}(E_{\nu})}$$
Well known (1-2%)



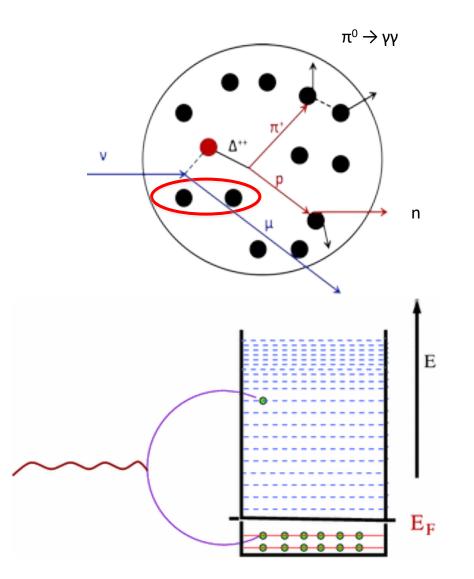
But in Reality

$$\frac{\frac{dN_{\nu_e}^{far}}{dE_{rec}}}{\frac{dN_{\nu_{\mu}}^{near}}{dE_{rec}}} = \underbrace{\int P_{\nu_{\mu} \to \nu_e}(E_{\nu}) * \phi_{\nu_{\mu}}^{near}(E_{\nu}) * F_{far/near}(E_{\nu}) * \sigma_{\nu_e}^{Ar}(E_{\nu}) * T_{\nu_e}^{far}(E_{\nu}, E_{rec}) dE_{\nu}}{\int \phi_{\nu_{\mu}}^{near}(E_{\nu}) * \sigma_{\nu_{\mu}}^{Ar}(E_{\nu}) * T_{\nu_{\mu}}^{near}(E_{\nu}, E_{rec}) dE_{\nu}}$$

- No cancellations
 - Unless you unfold
- Need to understand especially
 - Detector effects in near and far detector
 - Relation of visible to neutrino energy
 - Cross section ratios
 - Near to far flux extrapolation
- Flux normalisation cancels
 - > Shape is more important



Neutrino Interactions



- Tricky Problem
 - How to relate neutrino to measured energy?
 - Neutrino energy unknown
 - > Nuclear recoil not measurable
 - Nucleus will absorb some energy
- In a Near Detector measure charged and neutral particles
 - p from curvature in B-field
 - > Calorimetric energy for γ
 - $> \beta$ from ToF of neutron (recoil)



What does the ND need to measure?

ND Fluxes

 $\phi_{\nu_x}^{near}(E_{\nu})$

- Prior constrained 5-10%
- Total and differential cross sections on Argon

 $\frac{d^n \sigma_{\nu_{\chi}}^{Ar}}{da \ db \ dc \ \dots} (E_{\nu}) \quad \text{(Largely unknown)}$

True to reconstruction "matrix"

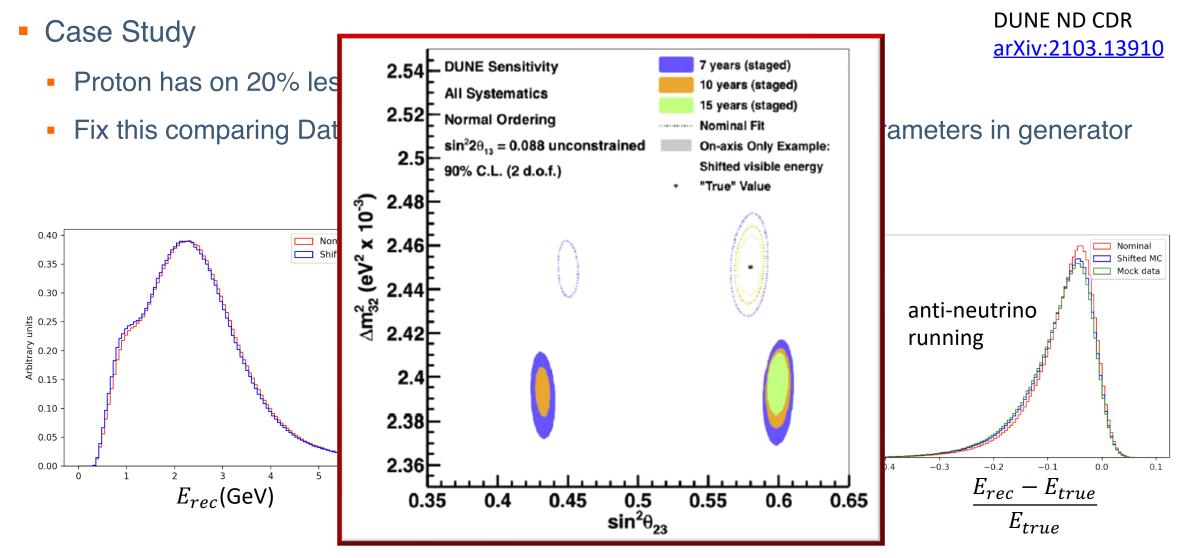
 $T_{\nu_x}^{far}(E_{\nu}, E_{rec})$ and $T_{\nu_x}^{near}(E_{\nu}, E_{rec})$

Depends on

- Detector effects
- differential cross sections



What Can Go Wrong?





Overarching ND Requirements

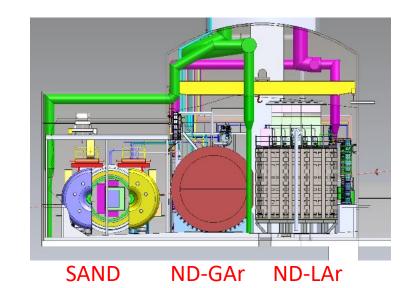
O0: Predict the neutrino spectrum at the FD: The Near Detector (ND) must measure neutrino events as a function of flavor and neutrino energy. This allows for neutrino cross-section measurements to be made and constrains the beam model and the extrapolation of neutrino energy event spectra from the ND to the FD.

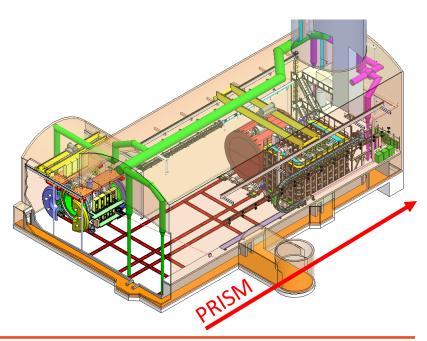
O0.1	Measure interactions on argon	Measure neutrino interactions on argon, determine the neutrino flavor, and measure the full kinematic range of the interactions that will be seen at the FD.
O0.2	Measure the neutrino energy	Reconstruct the neutrino energy in CC events and control for any biases in energy scale or resolution.
O0.3	Constrain the xsec model	Measure neutrino cross-sections in order to constrain the cross section model used in the oscillation analysis.
O0.4	Measure neutrino flux	Measure neutrino fluxes as a function of flavor and neutrino energy.
O0.5	Obtain data with different neutrino fluxes	Measure neutrino interactions in different beam fluxes in order to disentangle flux and cross sections and verify the beam model. (PRISM)
O0.6	Monitor the neutrino beam	Monitor the neutrino beam energy spectrum with sufficient statistics to be sensitive to intentional or accidental changes in the beam on short timescales.



The Near Detector Complex

- Four main components
 - System for on-Axis Neutrino Detection (SAND)
 - Liquid argon detector (ND-LAr)
 - Downstream tracker with gaseous argon target (ND-GAr)
 - ND-LAr and ND-GAr systems can move off-axis (PRISM concept)
- High statistics constrains
 - Cross section & neutrino flux
- Understand detector effects in LAr-TPC

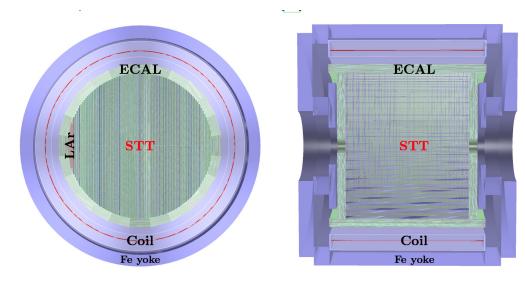


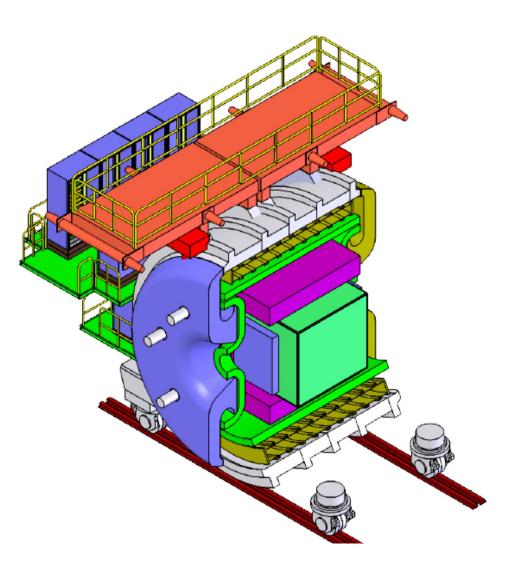




SAND

- On-Axis Beam monitor
 - ECAL & magnet (formerly KLOE)
 - Straw-Tube-Tracker (CH₂-Target)
 - Small LAr Target (GRAIN)
- Task
 - Measure neutrino flux (with H-measurement)
 - Monitor beam stability



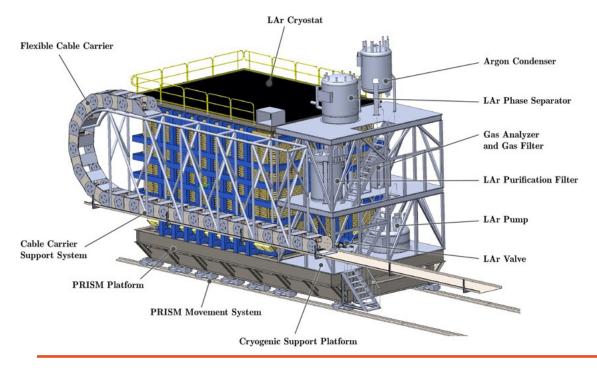


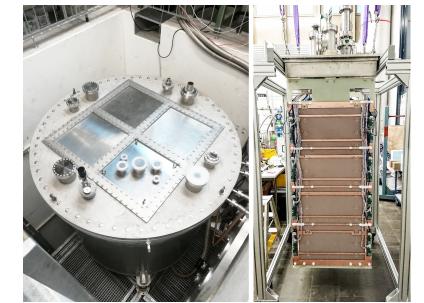


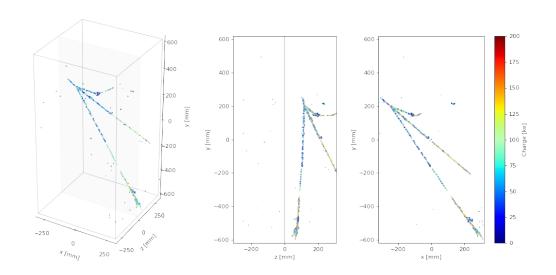
ND-LAr: Liquid Argon TPC

LAr TPC

- Study detector effects
- Similar to far detector technology
- High event rate
 - > Need compartmentalisation
 - Pixelized readout

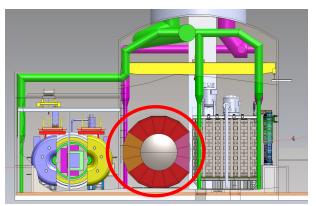


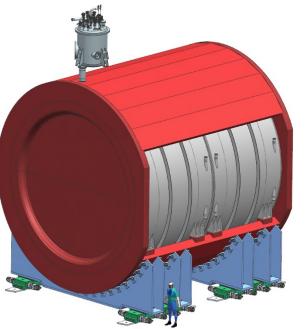




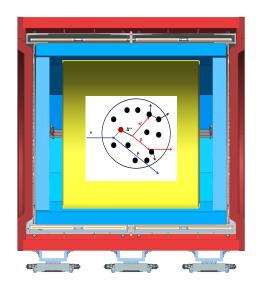


ND-GAr: Gaseous Argon TPC

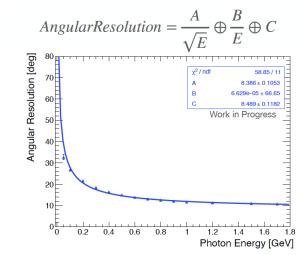


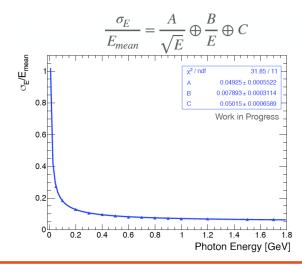


- Main detector components
 - High pressure (10 bar) gas TPC
 - ECAL & SC magnet
- Interactions on Ar gas
 - Low tracking threshold
 - Can detect neutrons
- Study nuclear effects



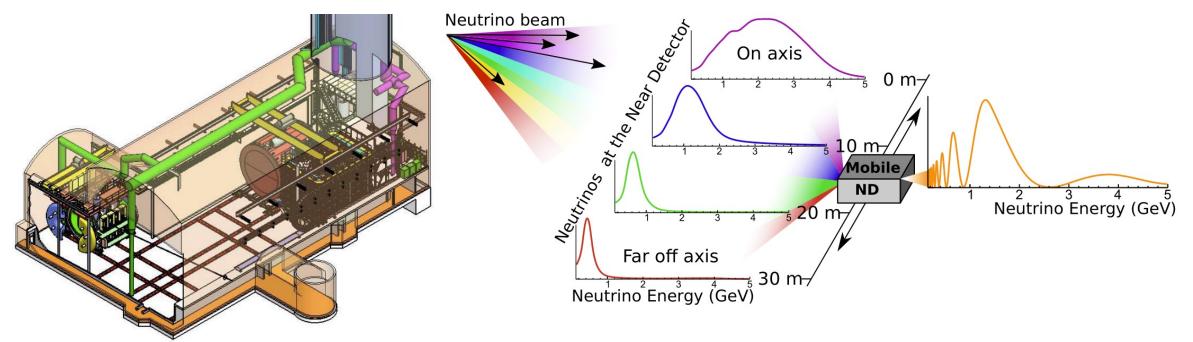
ECAL Performance







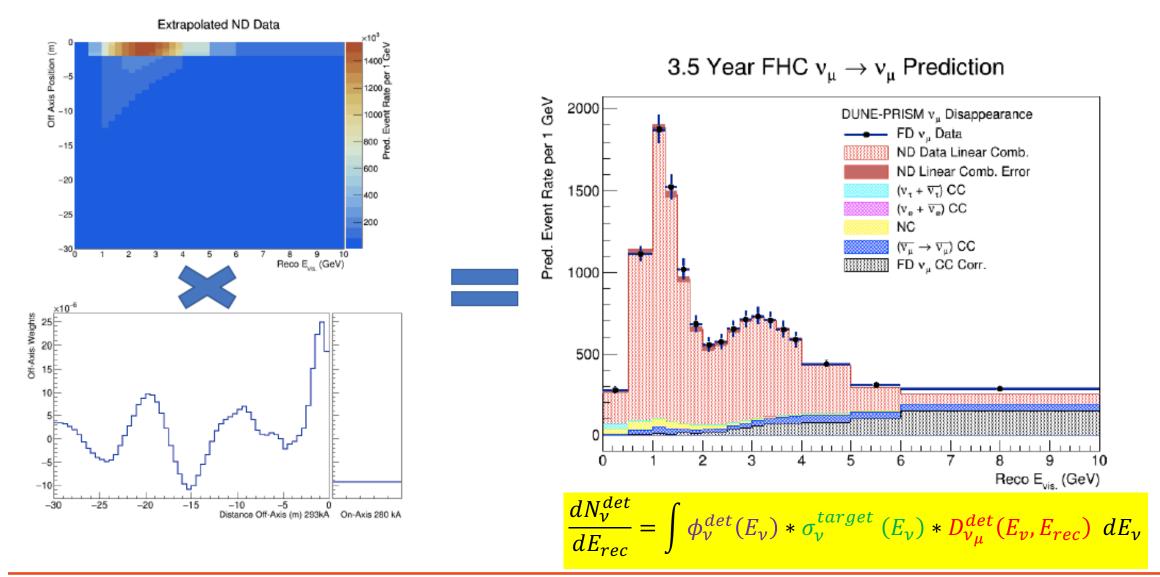
PRISM Concept



- Neutrino spectrum changes when going off-axis
- Can measure event rates with different fluxes
- Direct linear combination of date at different positions to construct FD oscillated spectrum
- Need to understand detector differences and flux only



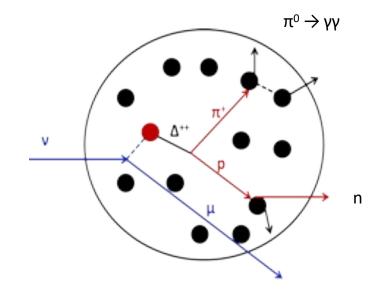
PRISM Analysis





Where are we?

- DUNE needs to make FD prediction, which have a 1-2% accuracy
 - Event rate and energy scales
- We don't really understand neutrino interactions
- Need to understand
 - Cross sections
 - Nuclear effects (FSI, IS)
 - All the particles emerging from the nucleus (full differential exclusive cross sections)
- Near Detector Design
 - Constrain uncertainties
 - > Become independent from uncertainties





Summary and Conclusion

- DUNE is an neutrino facility with an exciting physics program
- We have designed a capable Near Detector complex to constrain uncertainties
 - > ND-Lar

Liquid Argon TPC similar to FD

> ND-Gar

Gaseous Argon Target with low tracking threshold & neutron sensitivity

- PRISM Concept Measure at different off-axis angles/fluxes
- SAND

Measure flux and beam stability

- Will need to reach 1-2% systematic uncertainties
 - But this cannot be done without understanding the cross section



Backup

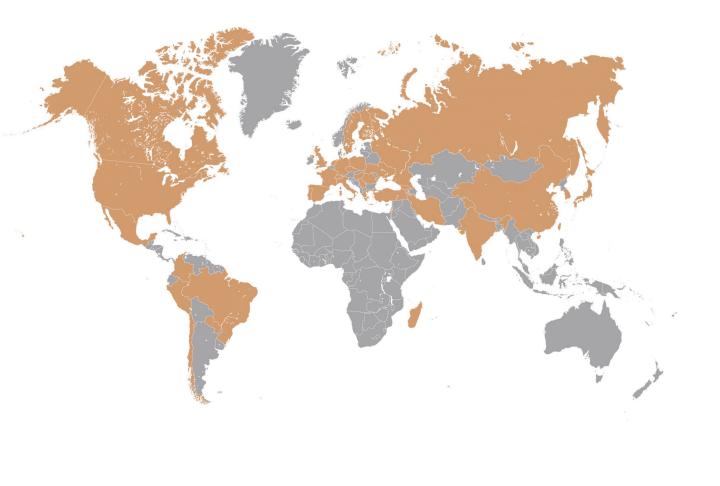




The DUNE Collaboration

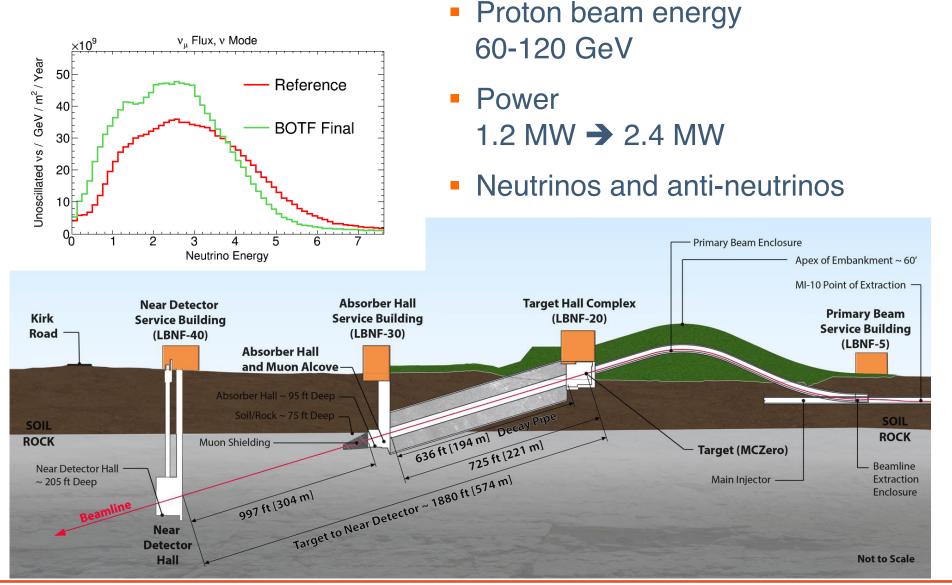
- International Collaboration
 - 1350 members
 - 200+ Institutions
 - 34 countries & CERN





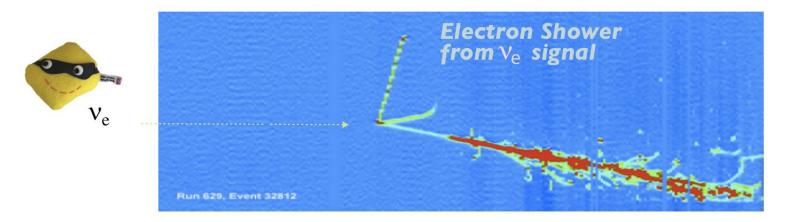


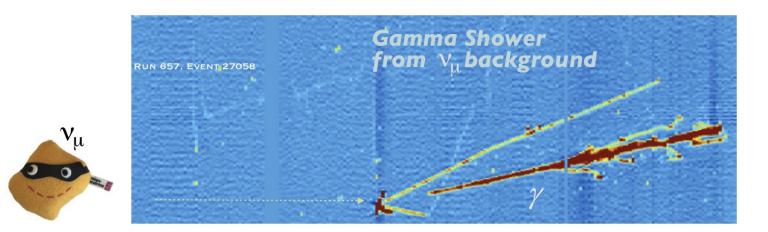




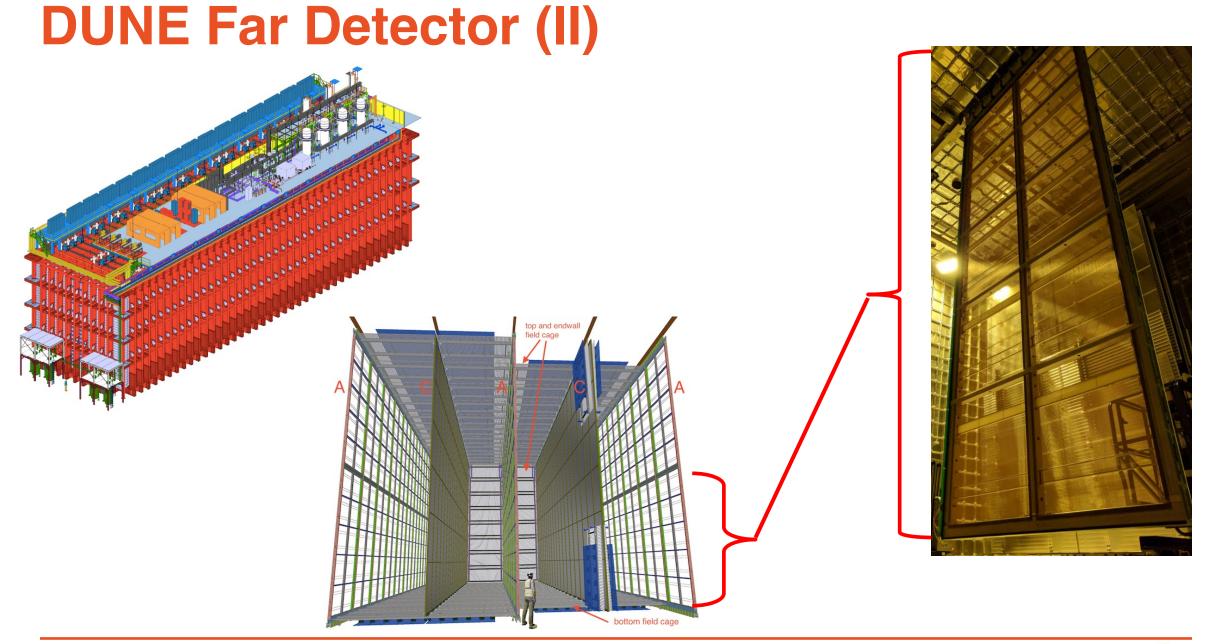


Event identification





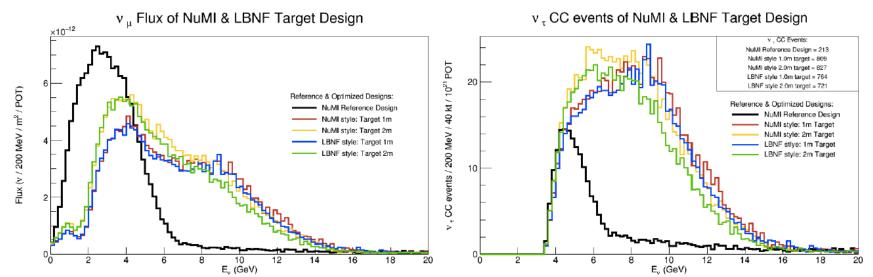






High Energy Tune

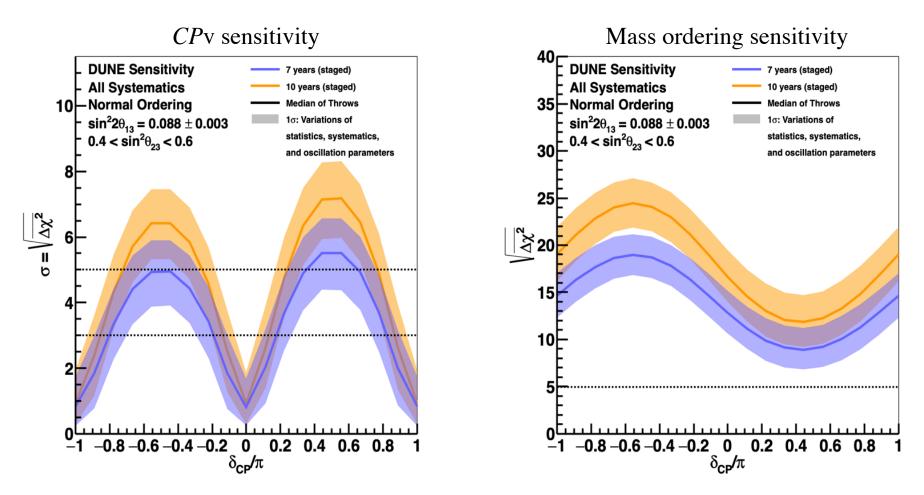
- Can change the flux by changing
 - Target positions
 - Horns (shape, position, current)



- Physics
 - Tau appearance
 - NSI



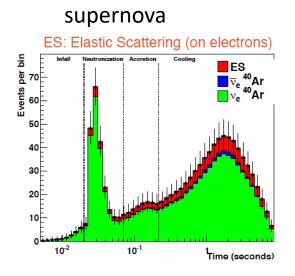
DUNE Sensitivity



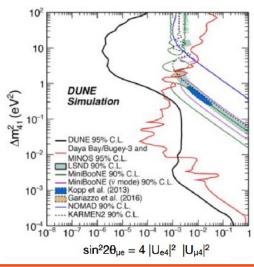
Move quickly to potential *CP* violation discovery Rapid, definitive mass ordering determination (>5 σ)



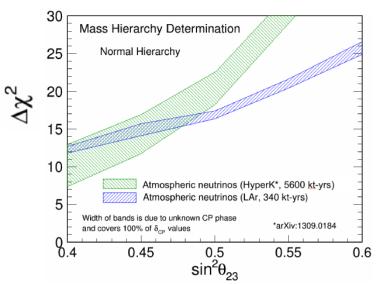
Other Physics



atmospherics



atmospherics



- Dark matter
- Large extra dimensions
- Dark photons
- NS interactions

