

# MATHUSA

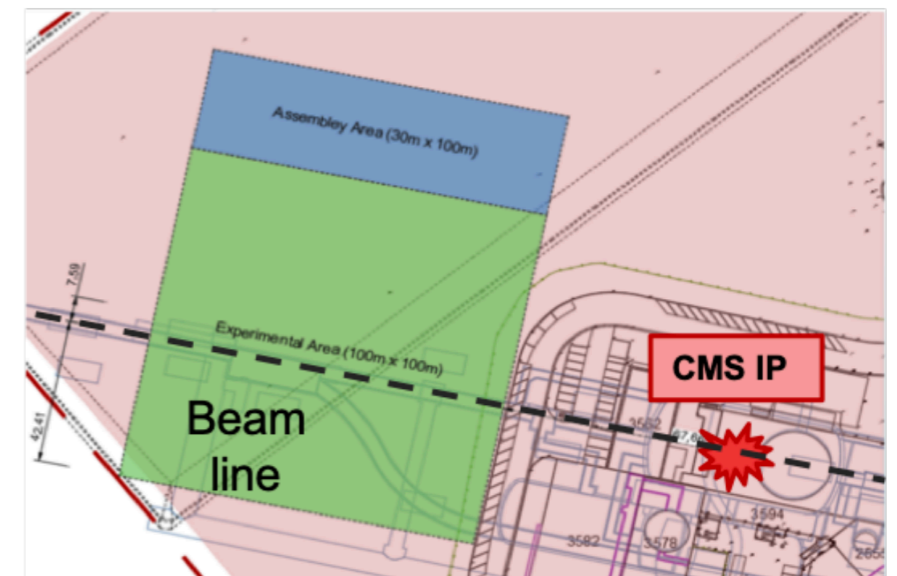
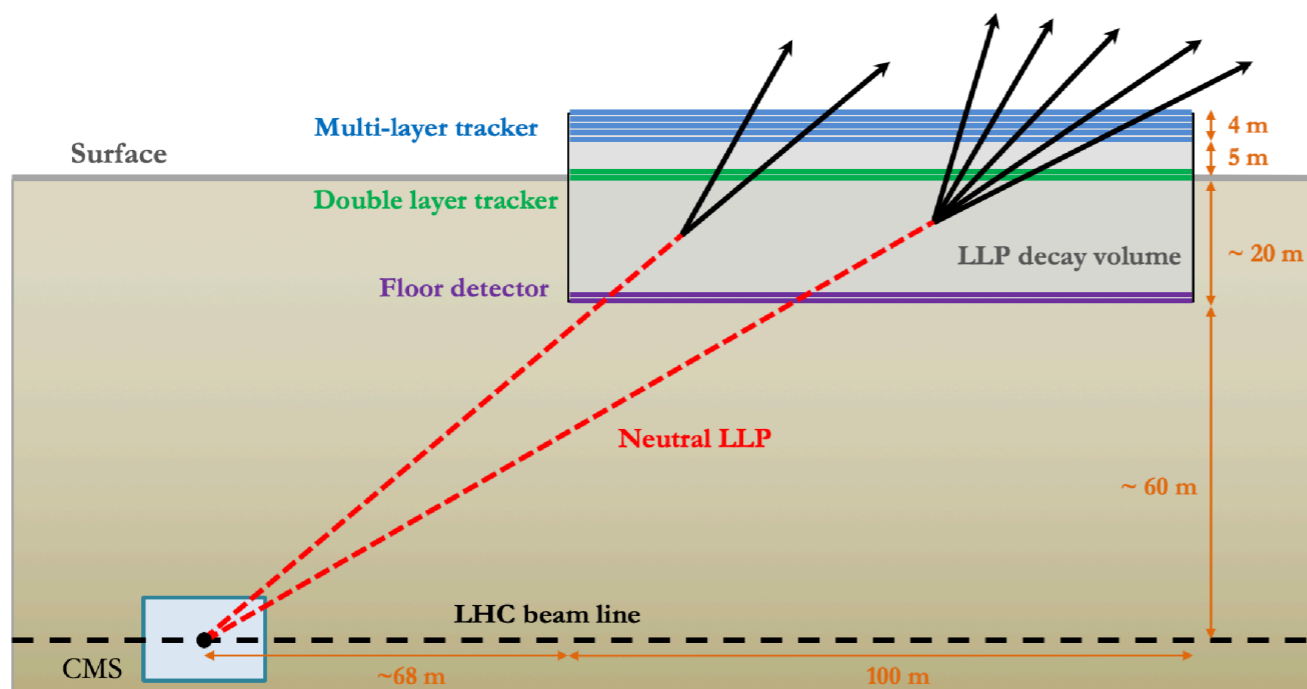
## *Status and goals*

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John Paul Chou  
Rutgers University

Friday, September 17, 2021

- MATHUSLA is a dedicated detector for long-lived particles
  - Designed to have applicability across a broad range of potential final states
  - Conceptually simple: build a big empty box with trackers on CERN-owned land near CMS
    - LLPs that decay inside will be reconstructed as displaced vertices
    - Backgrounds can be  $\sim O(1)$  because 80+ m rock shielding suppresses IP backgrounds and 4D tracking from  $\sim ns$  timing are distinct criteria for signal identification



# MATHUSLA COLLABORATION



- International collaboration including members & institutions from US, Canada, Chile, Bolivia, Mexico, Italy, Switzerland, ...
  - TDR in progress, aim publish early 2022.
  - Begin MATHUSLA operation with HL-LHC!
- Physics justification detailed in ~200 page report [1806.07396]

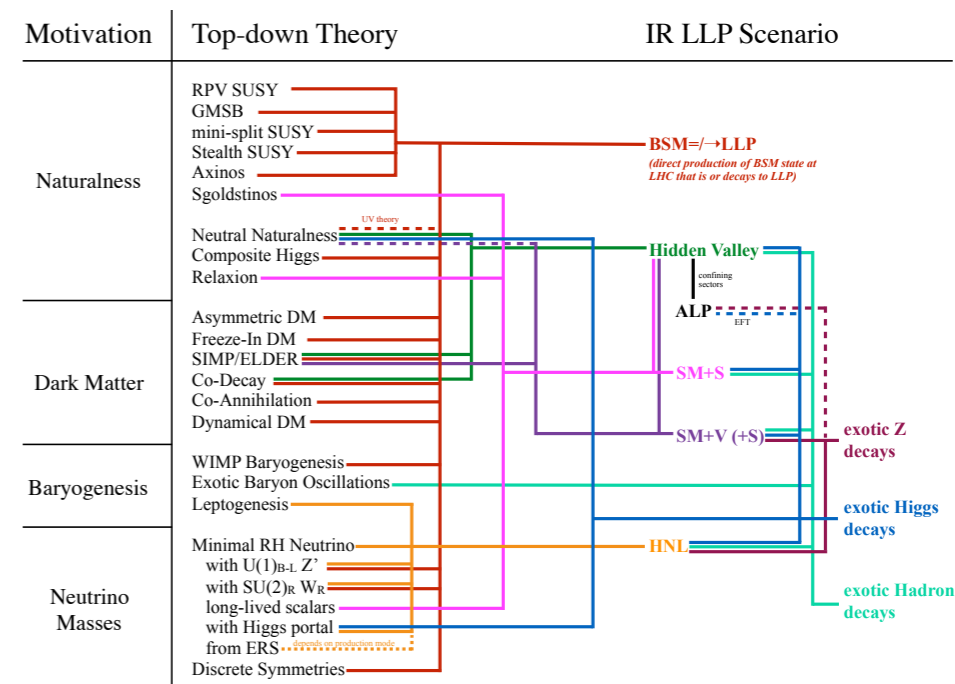
**A Letter of Intent for MATHUSLA: a dedicated displaced vertex detector above ATLAS or CMS**

1811.00927  
LHCC-I-031

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Long-Lived Particles at the Energy Frontier:  
The MATHUSLA Physics Case

1806.07396

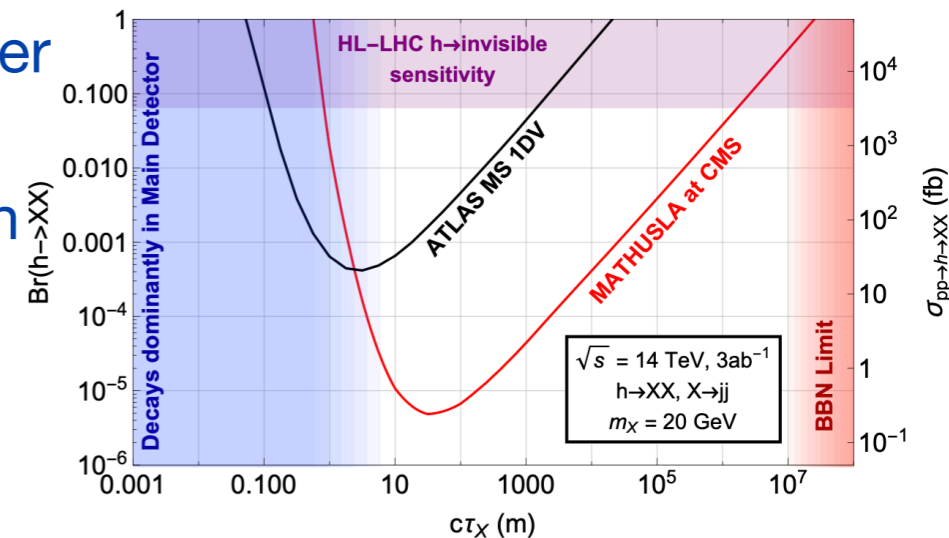


# DESIGN PRINCIPLES



- In the long lifetime regime  $> 100\text{m}$ , MATHUSLA has roughly same chance of “catching” an LLP decay in its decay volume as the main detectors

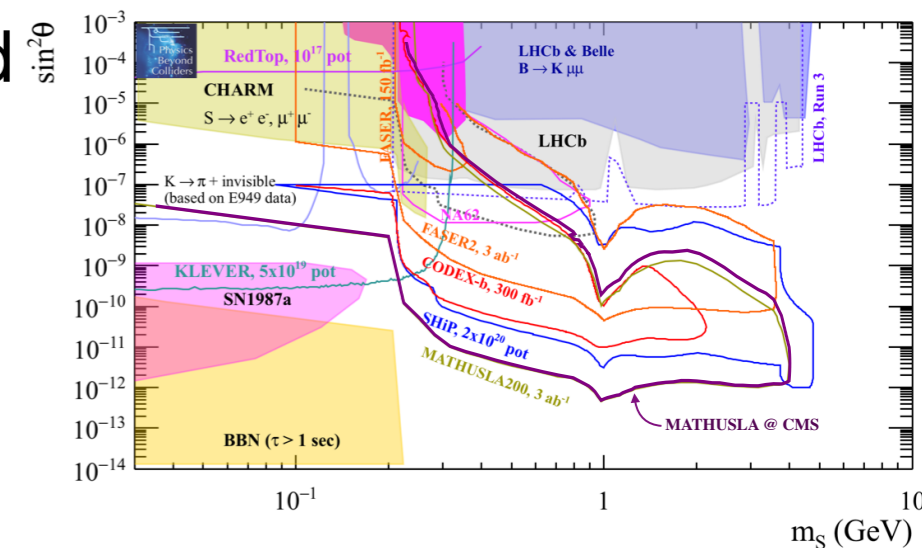
- Greater depth of decay volume compensates for smaller solid angle coverage
- For LLP searches, MATHUSLA’s greater sensitivity than ATLAS/CMS is due to
  - near-zero backgrounds
  - no trigger limitations



- Therefore, MATHUSLA will beat the main detectors for LLP signals where main detector searches are significantly impeded by background and trigger considerations (up to 1000x better reach)

- Targets (in order of priority)

- Hadronically decaying LLPs from few GeV to TeV
- LLPs with mass  $<$  few GeV (any decay mode)
- Cosmic Ray Physics

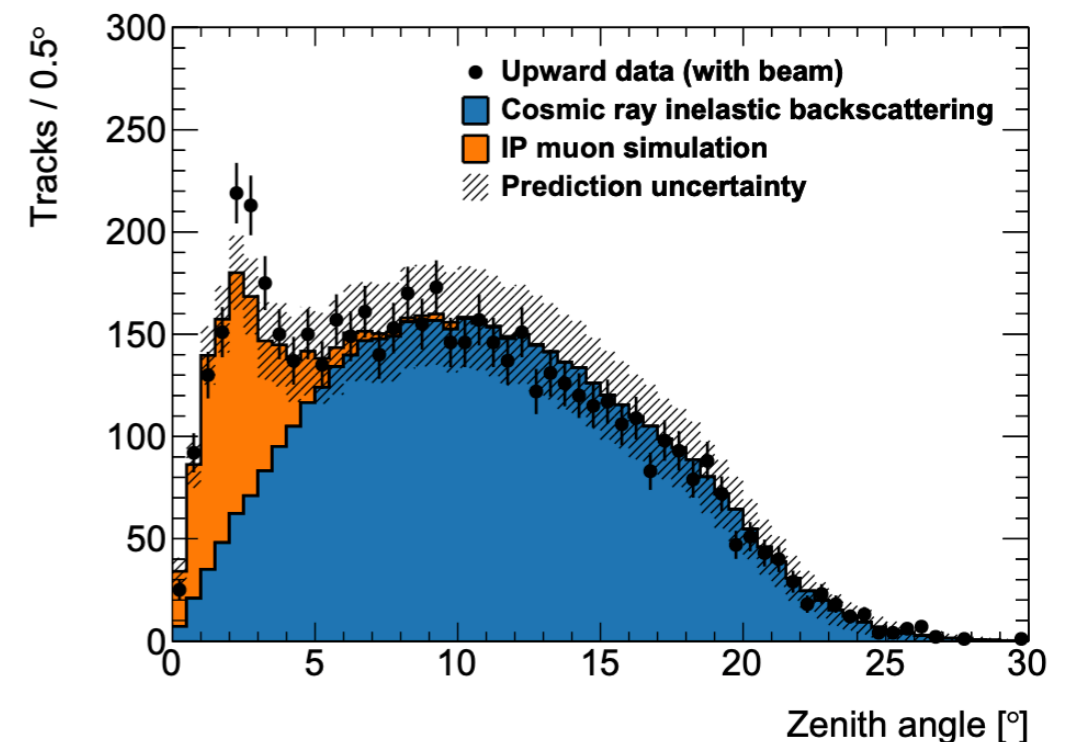
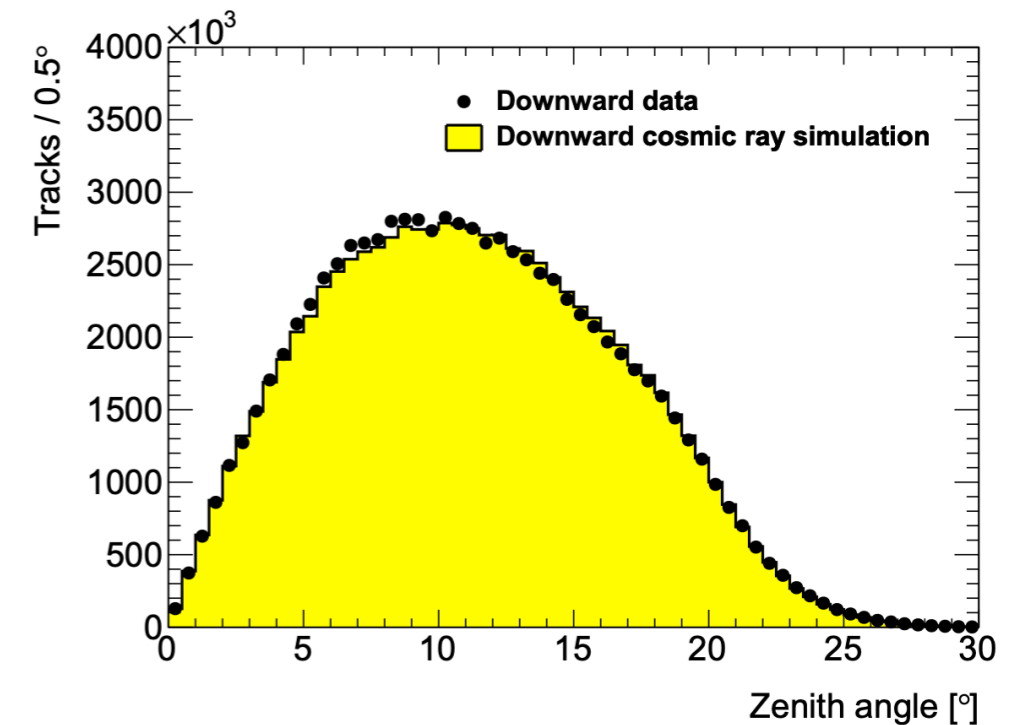
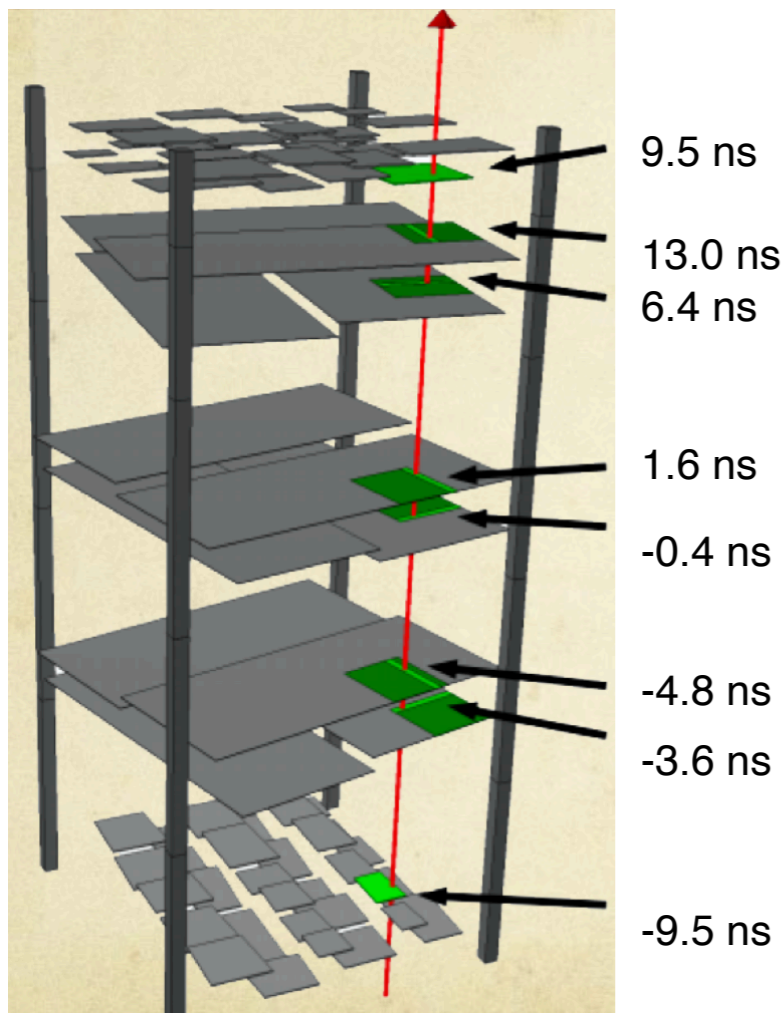


# TEST STAND RESULTS

[NIM A (2020) 164661]



- Test stand operated above ATLAS in 2018 – combination of plastic scintillator and RPCs
  - Both downward and upward rates/ angular distributions well predicted by simulation  $O(\sim 10\%)$

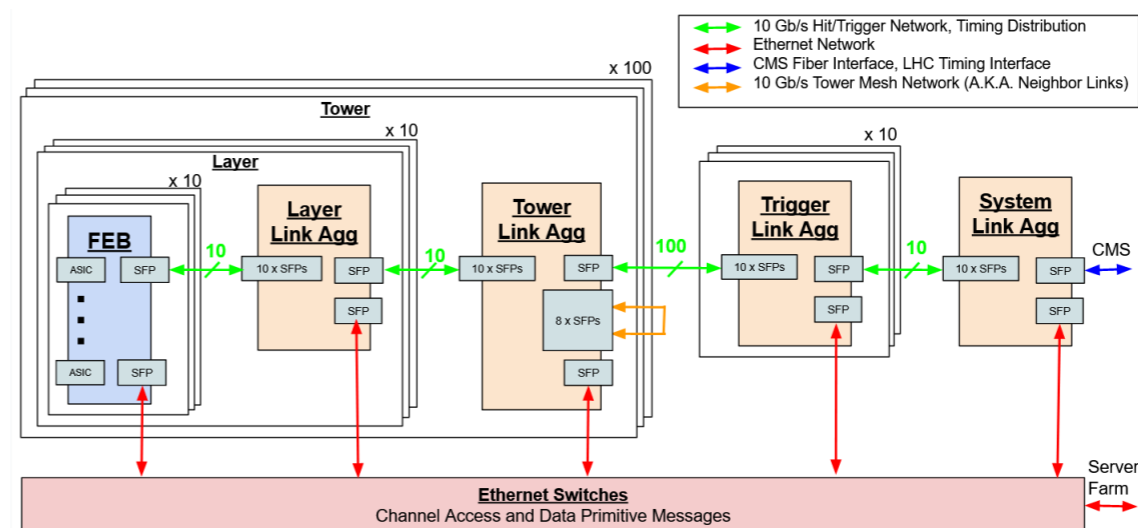


# ONGOING DEVELOPMENTS



- Collaboration is engaged in R&D for basic hardware/engineering of detector:
  - extruded scintillators, fibers, SiPMs, trigger & DAQ design
  - Simulation studies (LLP + Backgrounds)
- Implementing on custom tracking algorithms for MATHUSLA's unique environment to achieve high LLP reconstruction efficiency for low-multiplicity LLP final states
  - Low track multiplicity final states require careful optimization of detector geometry, tracking algorithm design, etc.
  - Backgrounds from IP muons and cosmic backscatter (e.g. neutral kaons) can create induce reconstructable vertices in the detector
- Cosmic Ray Studies
  - CR physics case white paper coming out this year 2021

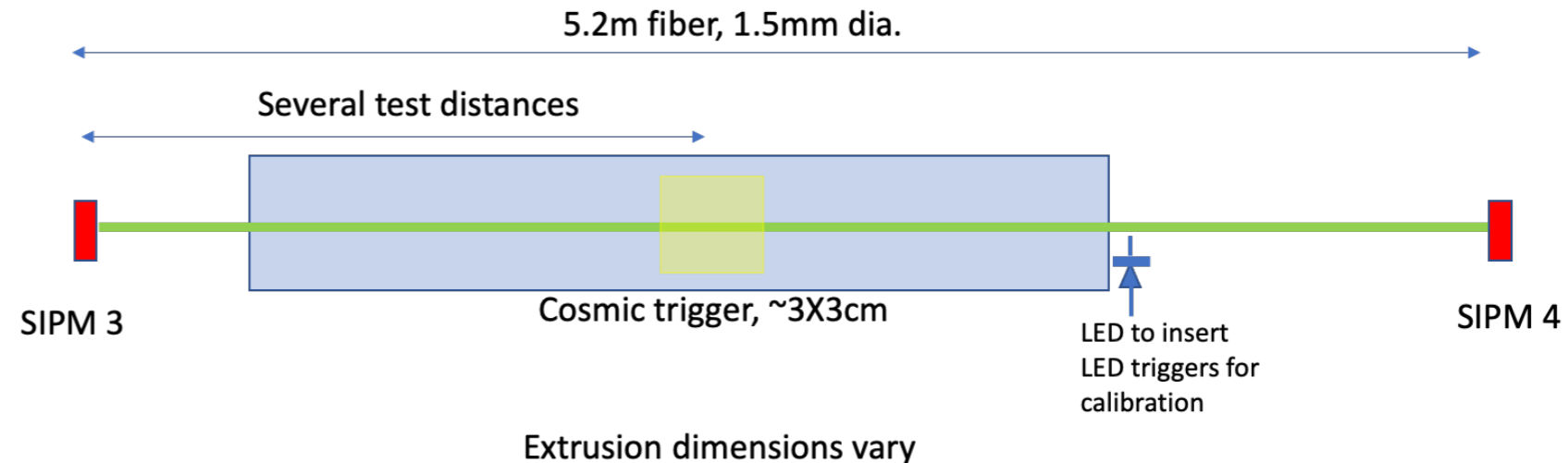
- DAQ
  - Modular design of the Front End Boards and link aggregation boards
  - All hits stored in buffer storage Data rate is well within COTS server
- Trigger
  - Tower aggregation module triggers on upward going tracks within 3 x 3 tower volumes
  - Selects data from buffer for permanent storage
- Trigger to CMS
  - Upward-going vertex forms trigger to CMS
  - MATHUSLA trigger latency estimates appear compatible with CMS L1 latency budget



# SCINTILLATOR TIMING AND TESTING



Cosmic setup: 2 small cosmic trigger counters define cosmic ray, light is measured at each end of fiber by SIPMs. Various algorithms to determine time resolution of difference.



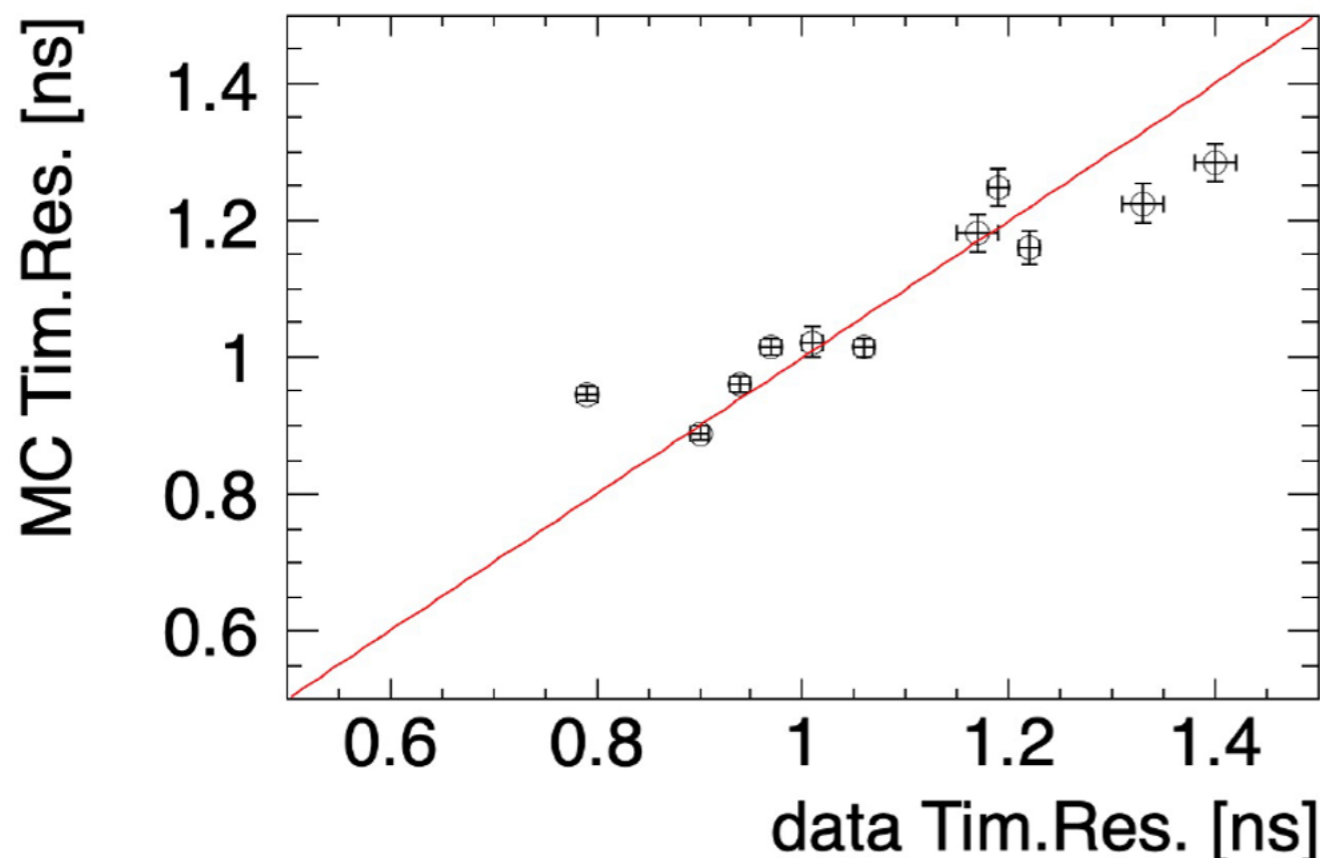
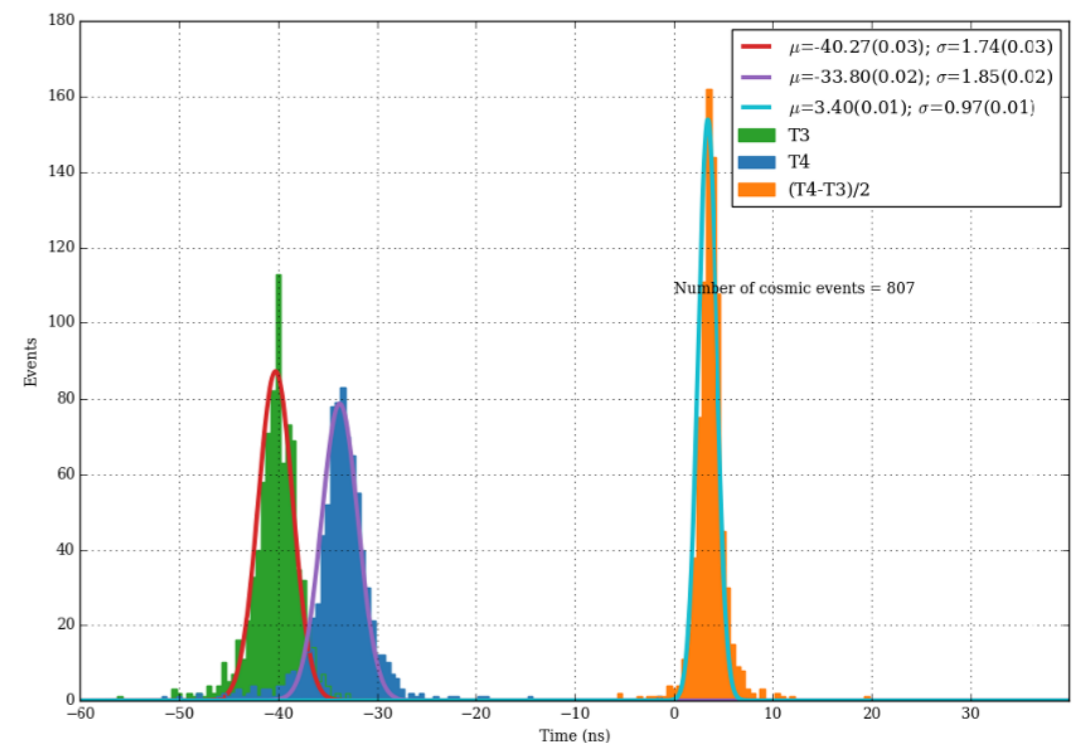
- Target timing resolution is  $\sim 1$  ns
  - Use difference in arrival time between separate measurements at two ends of extruded scintillator
  - Critical feature of the detector design
    - Separates downwards from upwards going tracks
    - Reject low beta particles from neutrino QIS
    - 4D tracking and vertexing reduces fakes/combinatorics



# SCINTILLATOR TIMING AND TESTING



- Use Geant to study extrusion and fiber choice to identify critical parameters
  - Goal of  $<1\text{ns}$  timing is achievable



DT result for cosmic run.

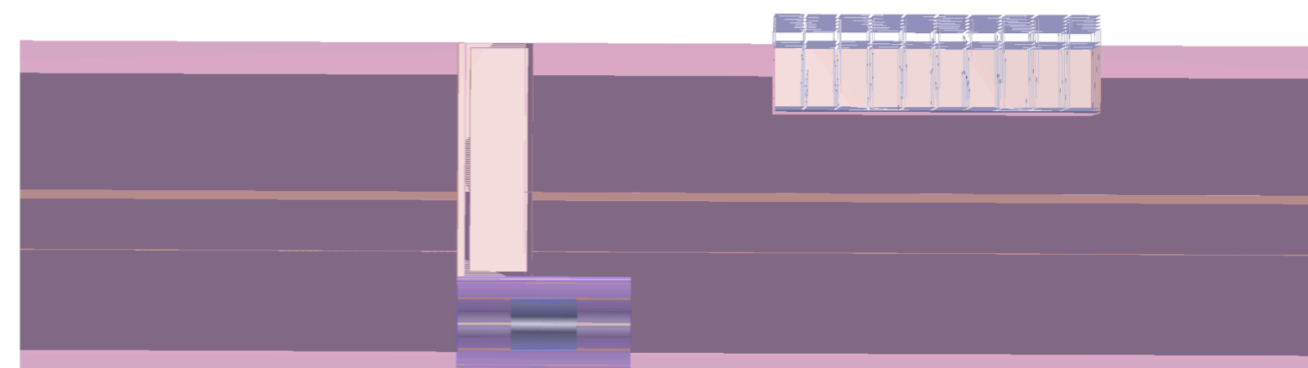
- 2X5 cm extrusion cross section
- 5m fiber, BCF92, 1.5mm diameter
- 1pe constant threshold

- Different extrusion thickness
- Different fiber diameters
- Different fiber lengths
- Different fiber vendors

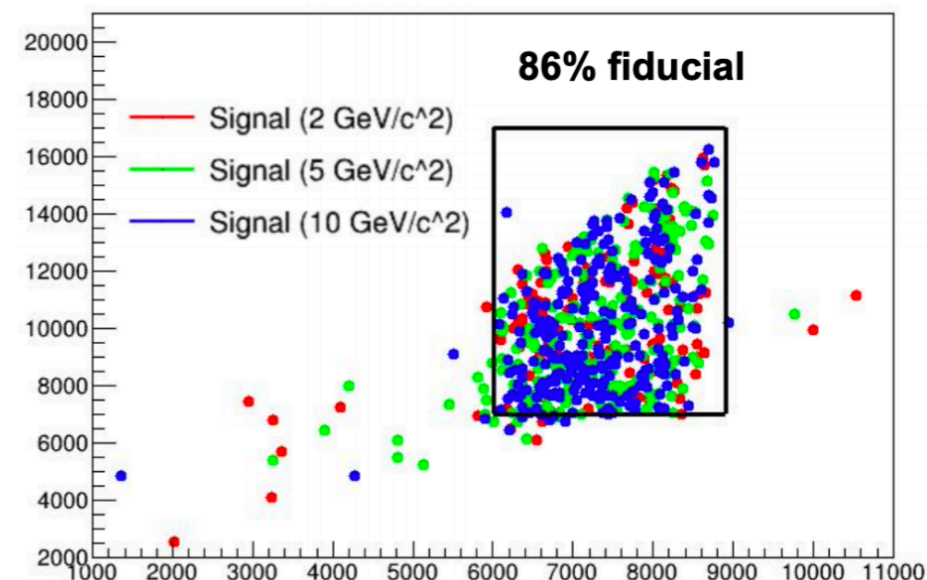
# SIMULATION OF BACKGROUNDS



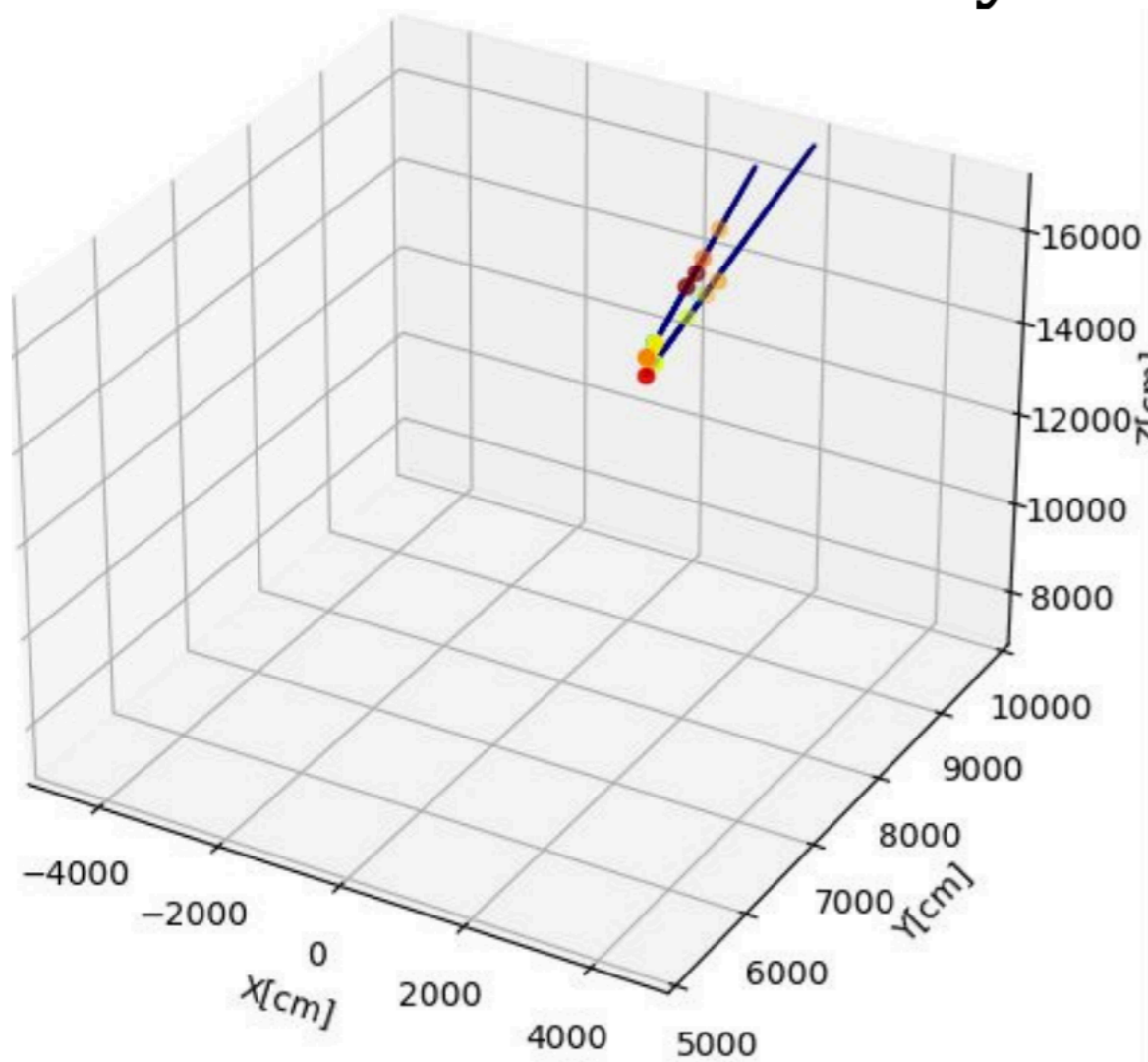
- Use Geant to model particle interactions in matter
- Backgrounds under study:
  - Upwards going muons from collisions (Pythia8)
  - Backscatter (to upwards going  $V^0$ ) from downwards going cosmic rays (Parma)
  - Neutrino interactions (Genie3)
- Cavern, access shaft, CMS, rock, and detector are all modeled
  - Rock is from a geological survey (same as for test stand)
- Analysis software uses Kalman Filtering to reconstruct tracks and form 4D vertices



Signal Vertex Location



# IP MUON BACKGROUNDS



- We expect  $\sim 10^{11}$  muons from  $W$  events over lifetime of HL-LHC
  - $\sim 10^9$  will reach MATHUSLA
- These muons can create vertices in a few different ways
  - Delta-rays
  - Induce EM Showers
  - 5-body decay in flight
- Backgrounds are suppressible with a high-coverage floor veto + topological constraints on the vertices

# CONCLUSIONS



- MATHUSLA has extensive reach and versatility to probe the LLP landscape
  - Significant progress is being achieved on multiple fronts: simulation of rare backgrounds, DAQ design, scintillator/fiber/SiPM properties, cosmic ray physics case
  - New member contributions always welcome!
- Hope to finish TDR by early 2022, followed by prototype module and full detector for HL-LHC