HIGH ENERGY AND ULTRA-HIGH ENERGY NEUTRINOS

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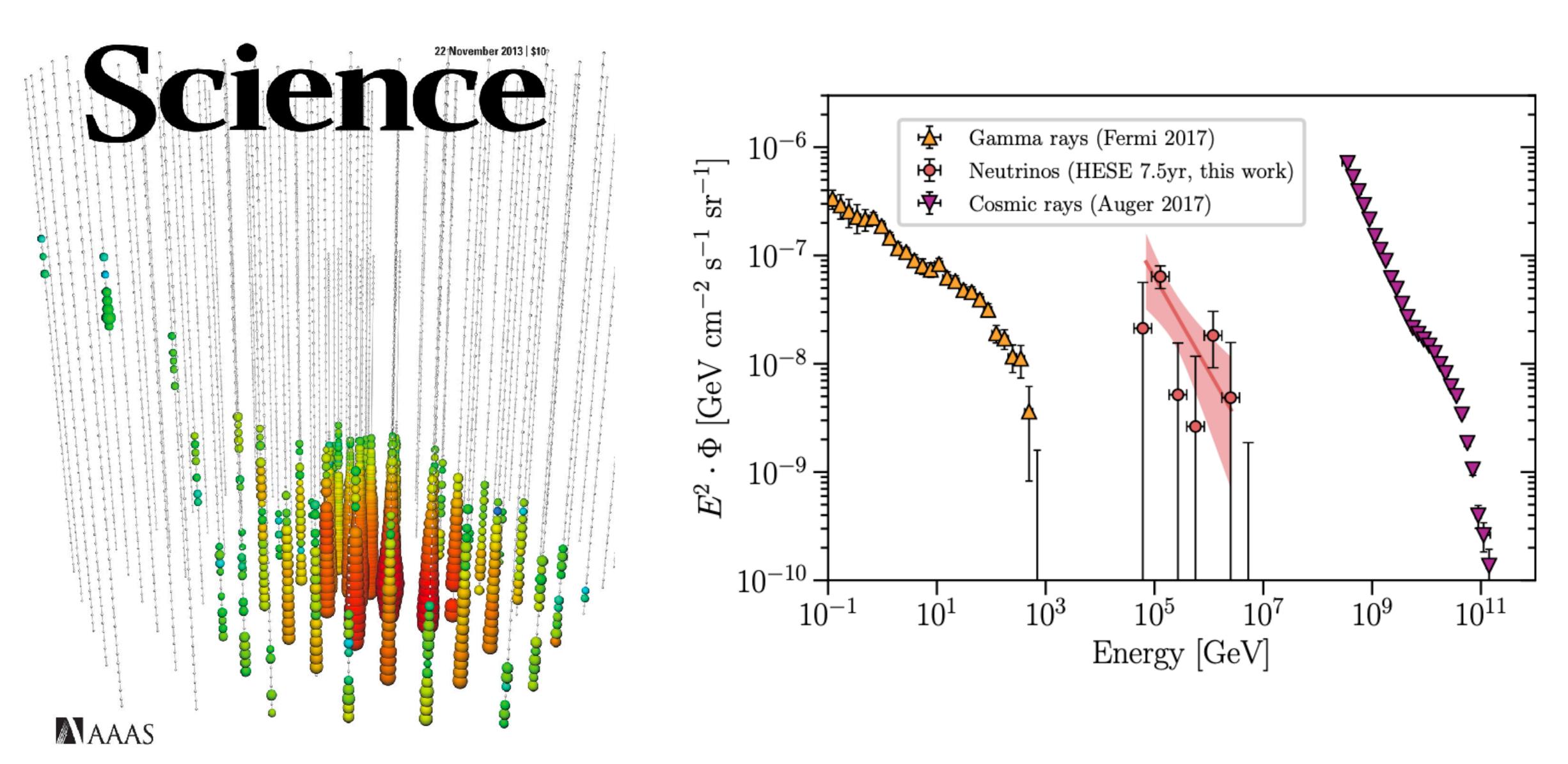
> CF-7 Day Workshop 18 November 2021

WHITE PAPER OUTLINE

- > Physics Opportunities with energetic cosmic neutrinos
 - > Starting point Lols: Snowmass Lols on HE/UHE Fundamental Physics and UHE neutrinos
- ➤ High-energy (TeV-PeV) and Ultra-High Energy (>100 PeV)

Contents 1 Introduction and overview 1.1 High-energy neutrinos (TeV-PeV) 1.2 Ultra-high-energy neutrinos (> 100 PeV) 1.3 HE and UHE cosmic neutrinos in particle physics HE and UHE cosmic neutrinos in astrophysics 1.5 Present and future experimental landscape 2 Current status and lessons learned 2.1 In particle physics 2.2 In astrophysics 3 Goals for the next decade 3.1 In particle physics 3.2 Simulation tools 3.3 In astrophysics 3.4 Detector requirements to achieve the science goals 4 Experimental landscape 4.1 Overview 4.2 TeV-PeV range 4.3 Multi-PeV range 4.4 > 100-PeV range 10 5 Summary and conclusions 11

THIS DECADE: DISCOVERY OF HIGH-ENERGY NEUTRINOS

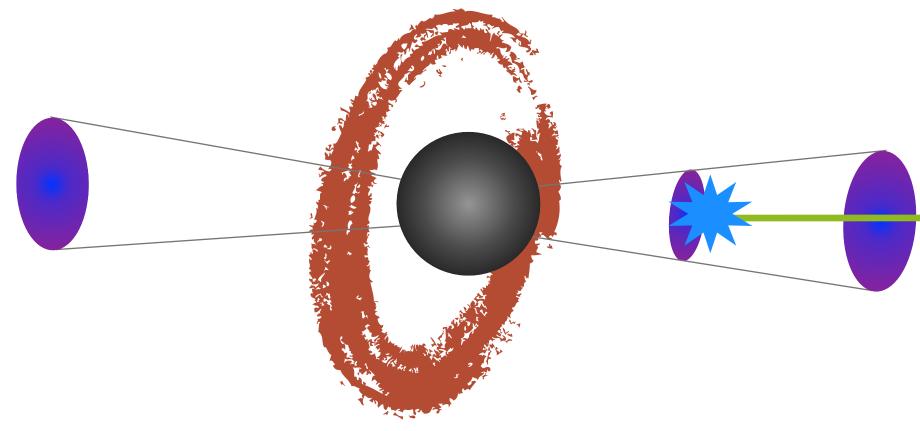


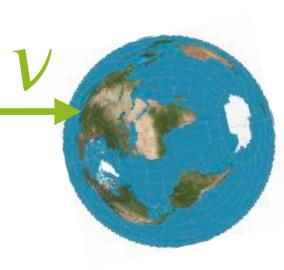
IceCube Science 2013

PRD 2021 arXiv:2011.03545

NEUTRINO SOURCES







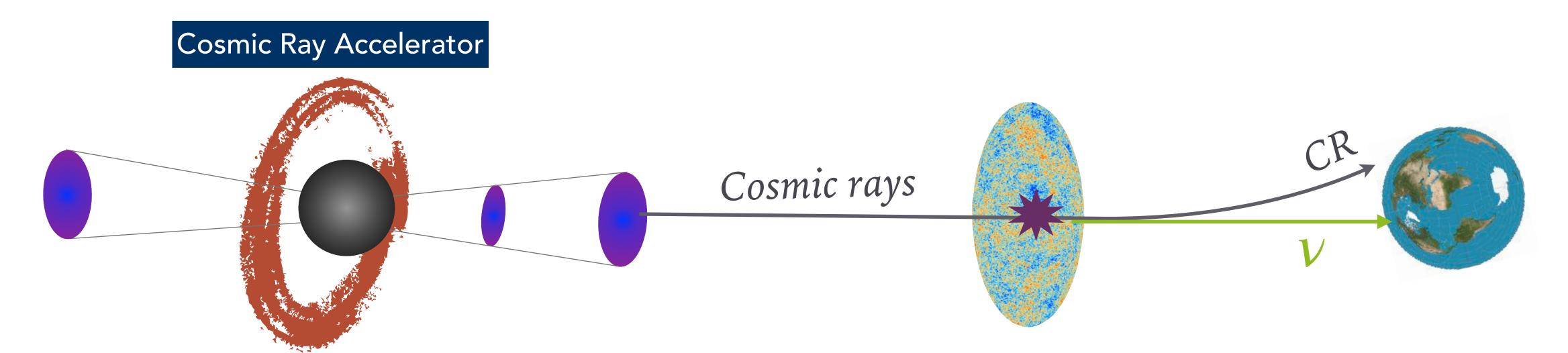


Cosmic ray interactions at the sources

Cosmogenic v's:

Origin of the most energetic cosmic rays

NEUTRINO SOURCES



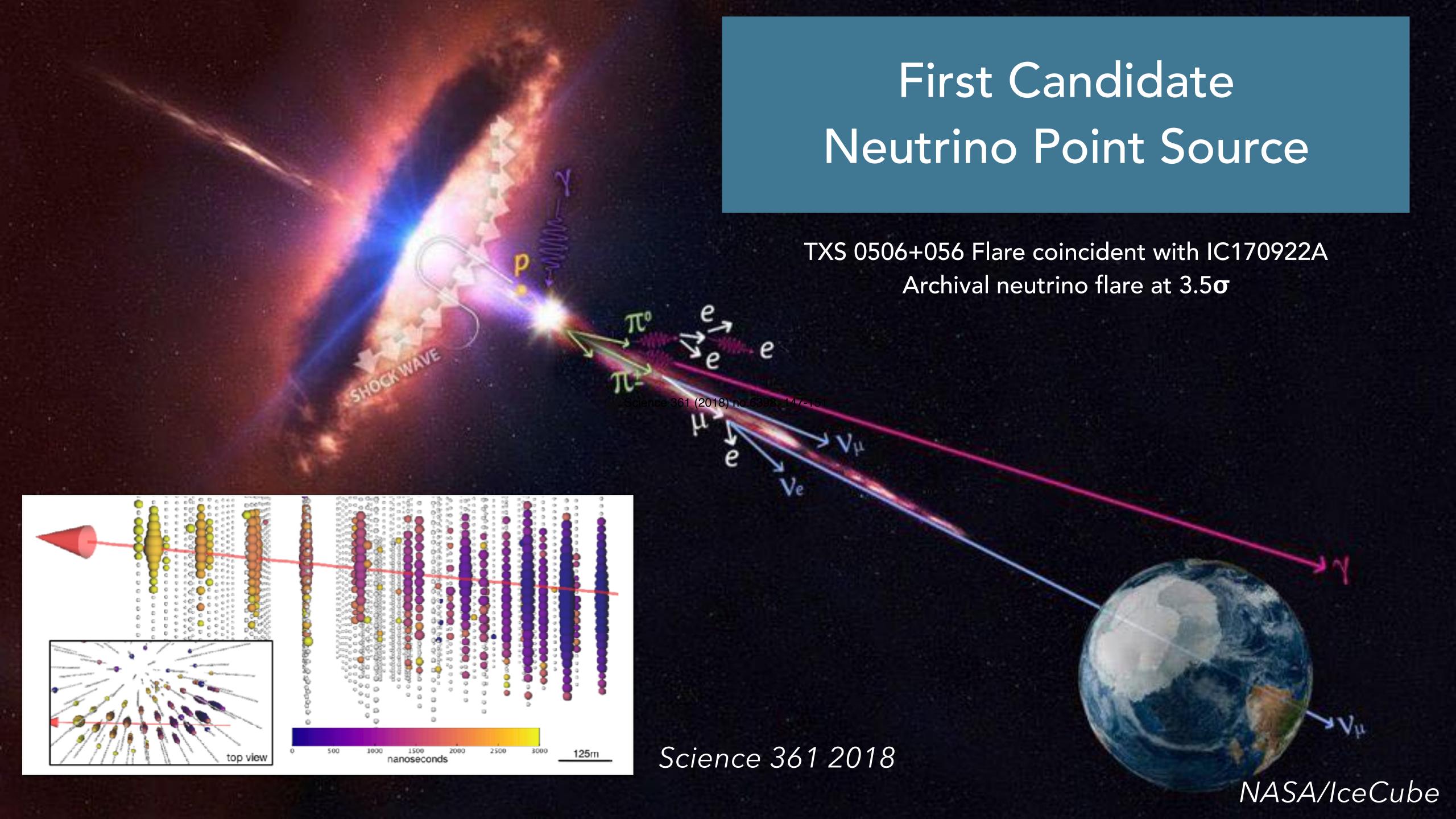


Cosmic ray interactions at the sources

Cosmogenic v's:

Origin of the most energetic cosmic rays

Cosmic Ray
$$p + \gamma \to \Delta^+ \to \pi^+ + n \\ \pi^+ \to \mu^+ + \nu_\mu \\ \mu^+ \to e^+ + \bar{\nu}_\mu + \nu_e$$



Science Targets in Astrophysics

Origin of cosmic rays

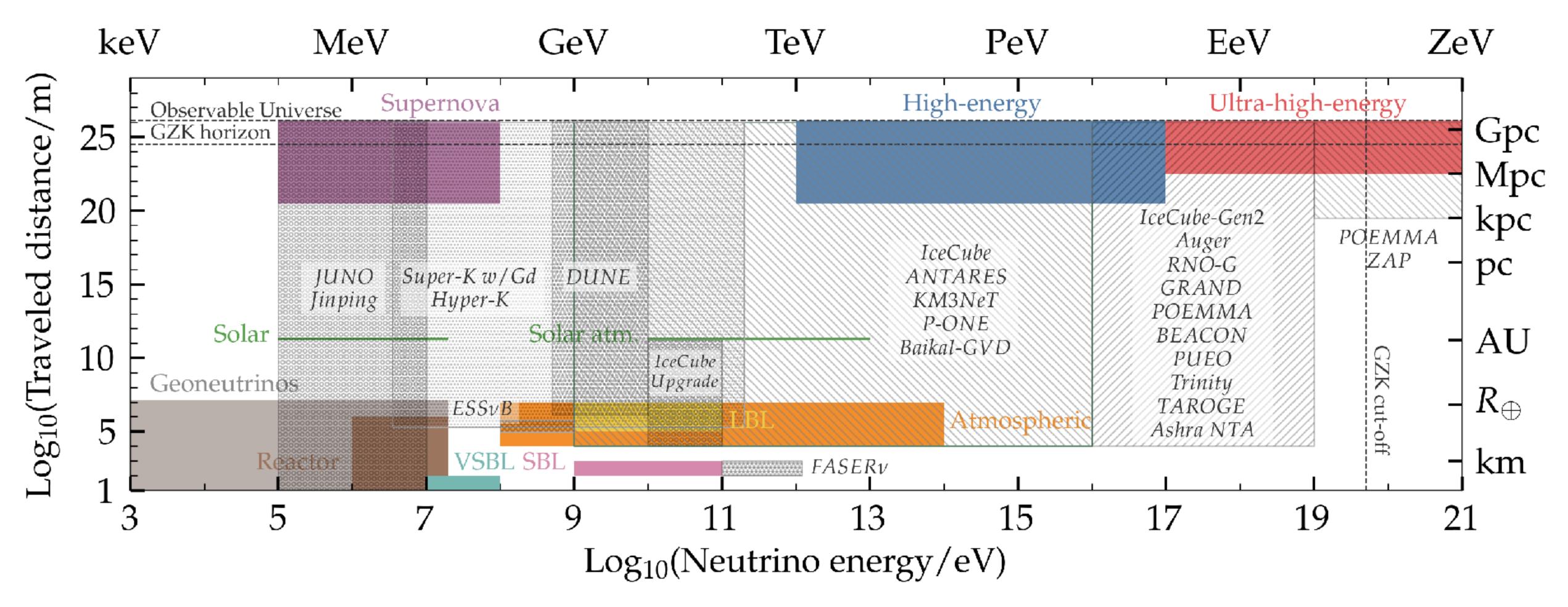
Revealing the sources of cosmic neutrinos

Neutrino point sources

New era of multi-messenger astrophysics

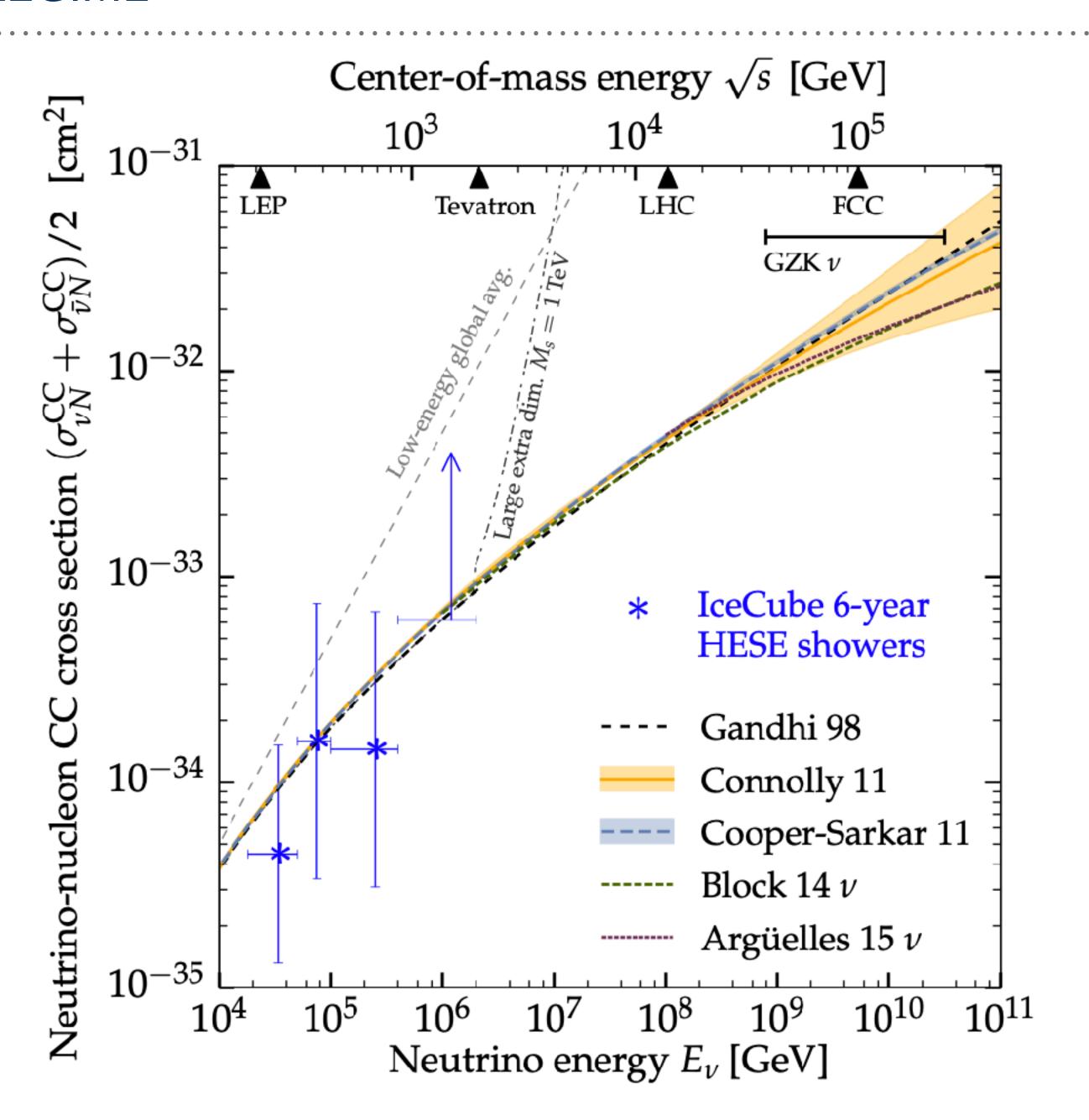
FUNDAMENTAL PHYSICS AT THE HIGHEST ENERGIES

- > Fundamental physics probe at the highest energy scales (TeV to ZeV)
- ➤ Longest baselines (~Gpc) allow even small effects to accumulate



CROSS SECTIONS IN A NEW ENERGY REGIME

- Probe neutrino interactions previously unmeasured
- ➤ Will improve with better statistics
- Push to ultra-high energies to probe new energy regime
 - ➤ EeV neutrinos $\rightarrow \sqrt{s} \sim 30 \text{ TeV}$

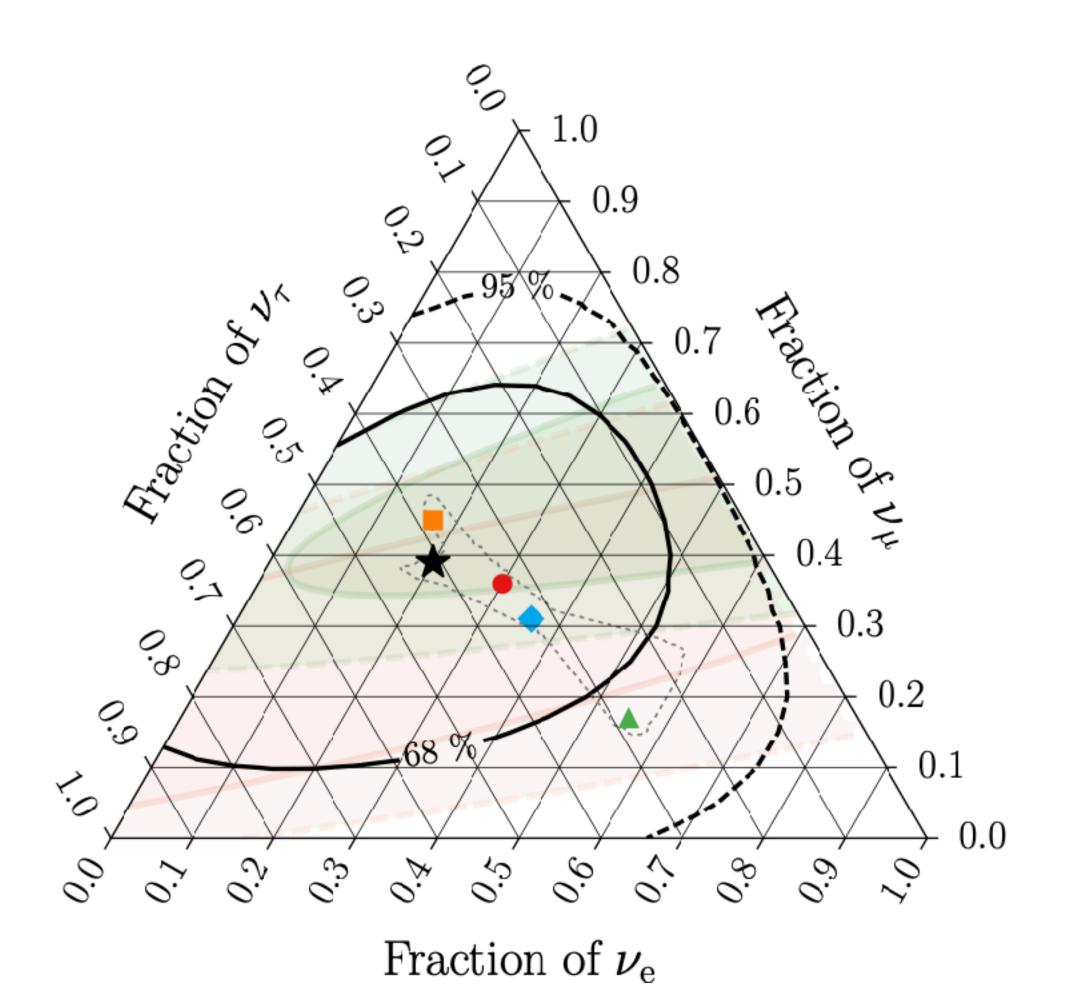


NEUTRINO FLAVOR

- ightharpoonup At the sources, we expect only $v_{\rm e}$ and/or $v_{\rm \mu}$
- ► Even Flavor ratios 1 v_e : 1 v_μ : 1 v_τ^* expected due to flavor oscillations over Gpc length scales

*Deviations from standard flavor ratios expected from Beyond-Standard-Model scenarios

$$p + \gamma \to \pi^{+} + n \begin{cases} \pi^{+} \to \mu^{+} + \nu_{\mu} \\ \mu^{+} \to e^{+} + \bar{\nu}_{\mu} + \nu_{e} \\ n \to p + e^{-} + + \bar{\nu}_{e} \end{cases}$$



Fundamental Physics Tests with TeV-PeV and EeV neutrinos

How do cross sections behave at higher energies?

How do flavors mix at high energies?

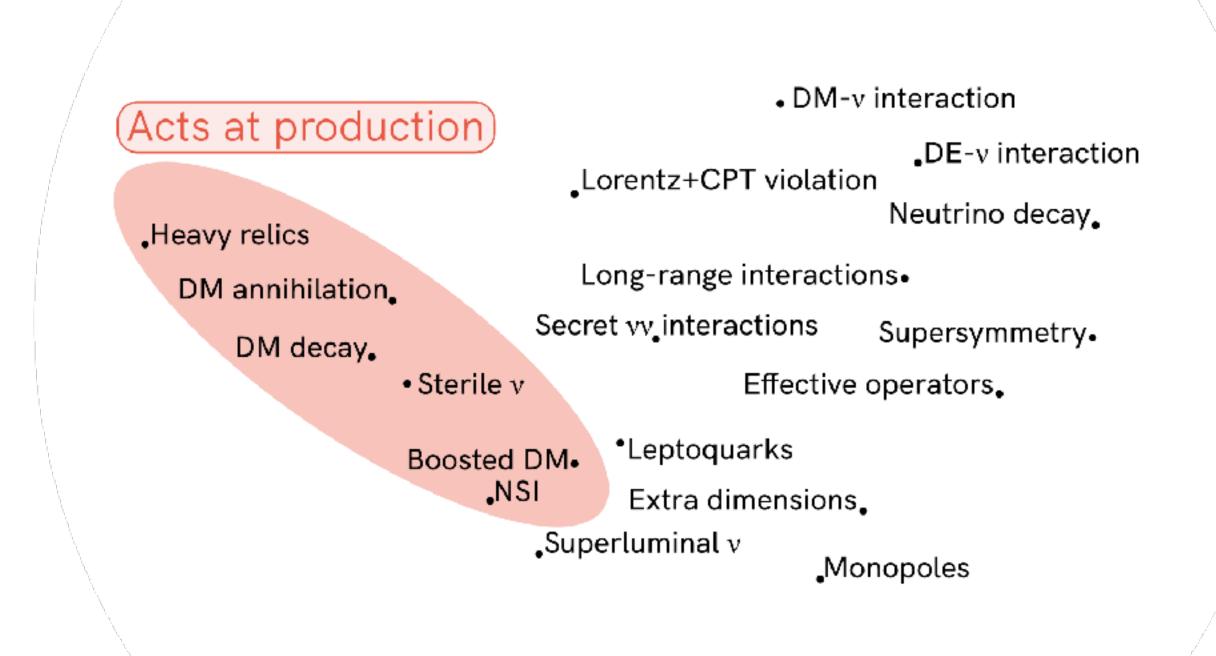
What are the fundamental symmetries of Nature?

Evidence of dark matter?

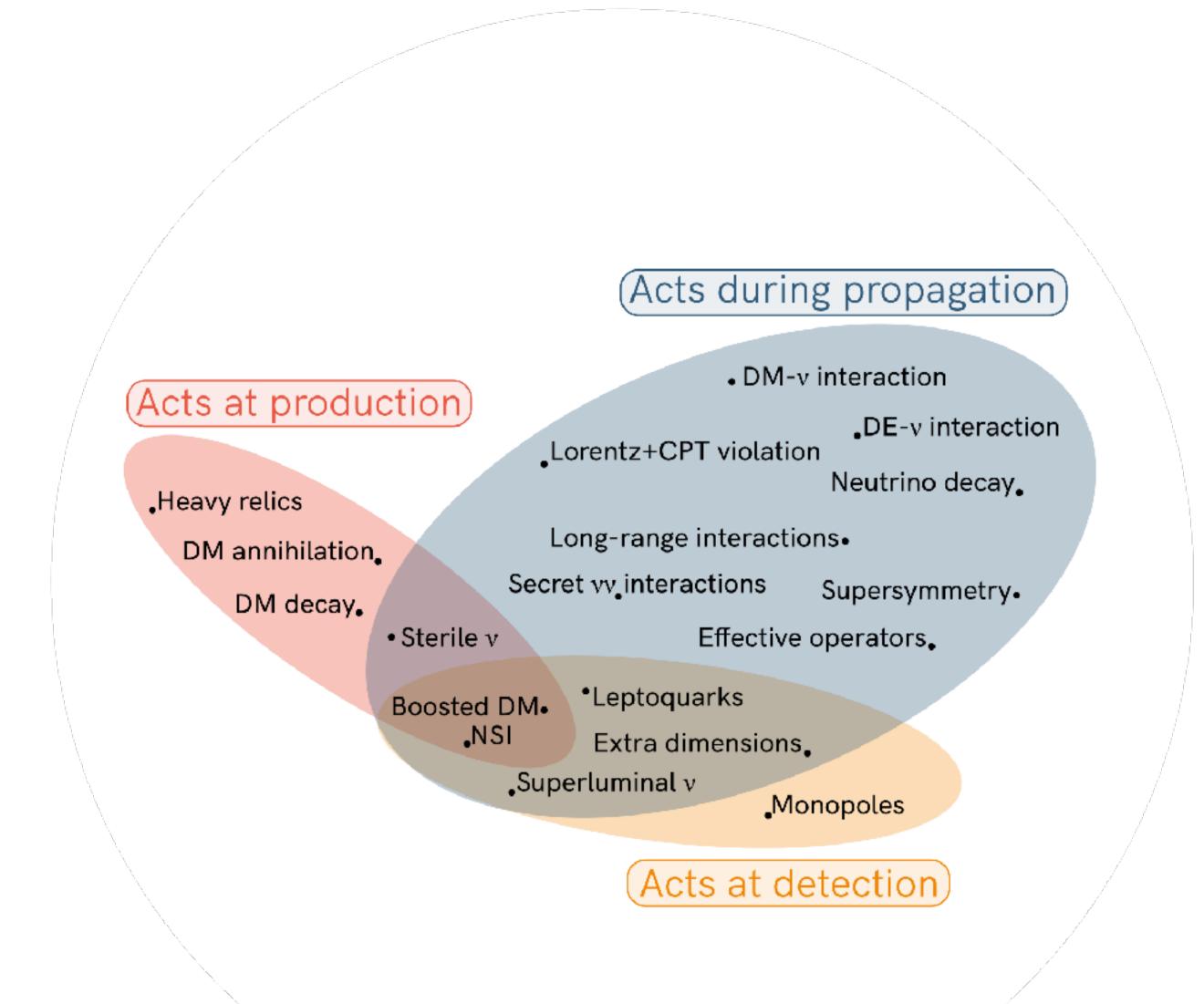
Hidden interactions with cosmic backgrounds?

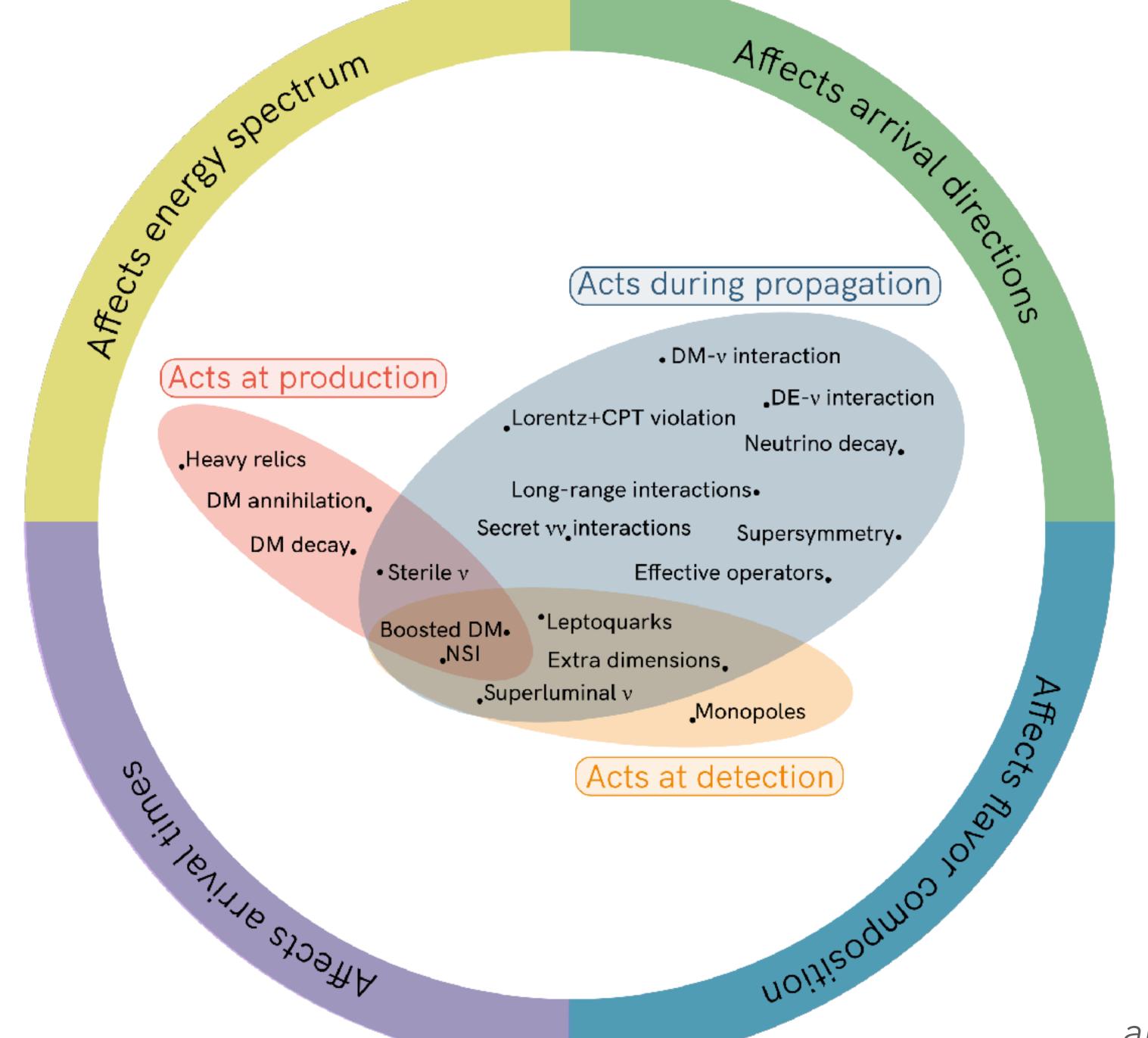
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    DM-v interaction

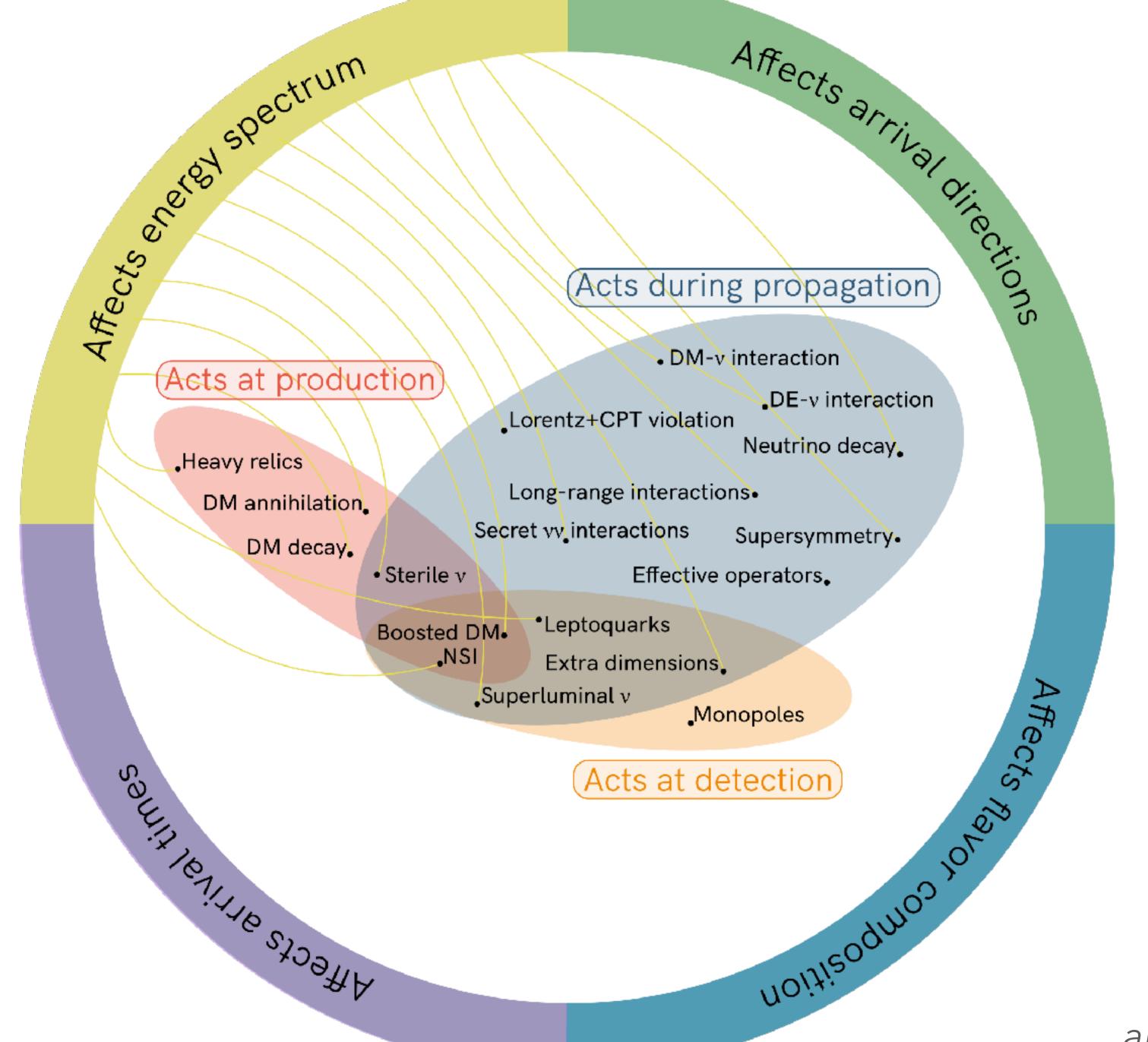
                                                      •DE-v interaction
                              Lorentz+CPT violation
                                                     Neutrino decay.
.Heavy relics
                               Long-range interactions.
  DM annihilation.
                           Secret vv.interactions
                                                    Supersymmetry.
      DM decay.
                                          Effective operators.
                  • Sterile v
                                 Leptoquarks
                   Boosted DM•
                        •NSI
                                  Extra dimensions.
                           •Superluminal v
                                                "Monopoles
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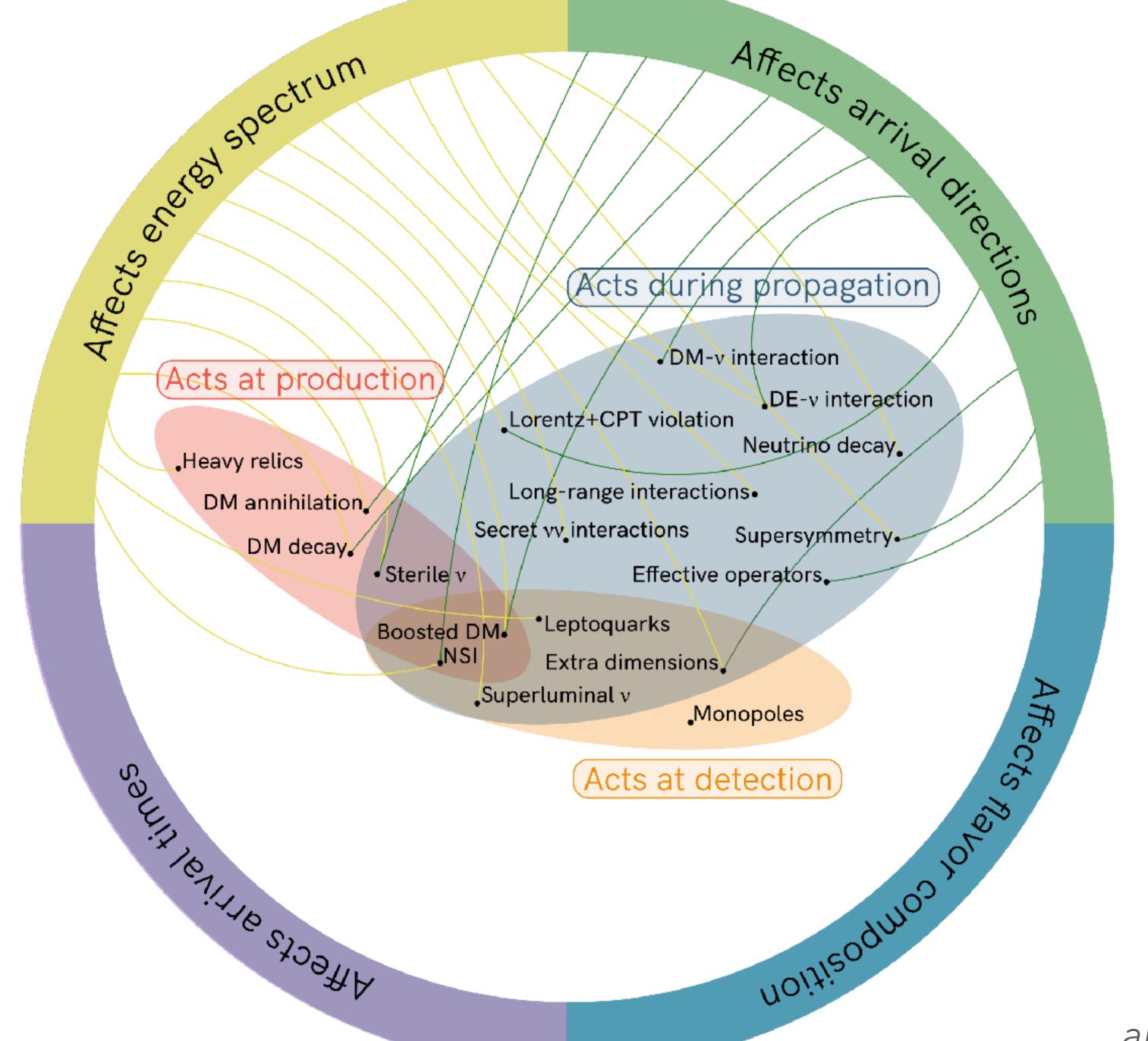


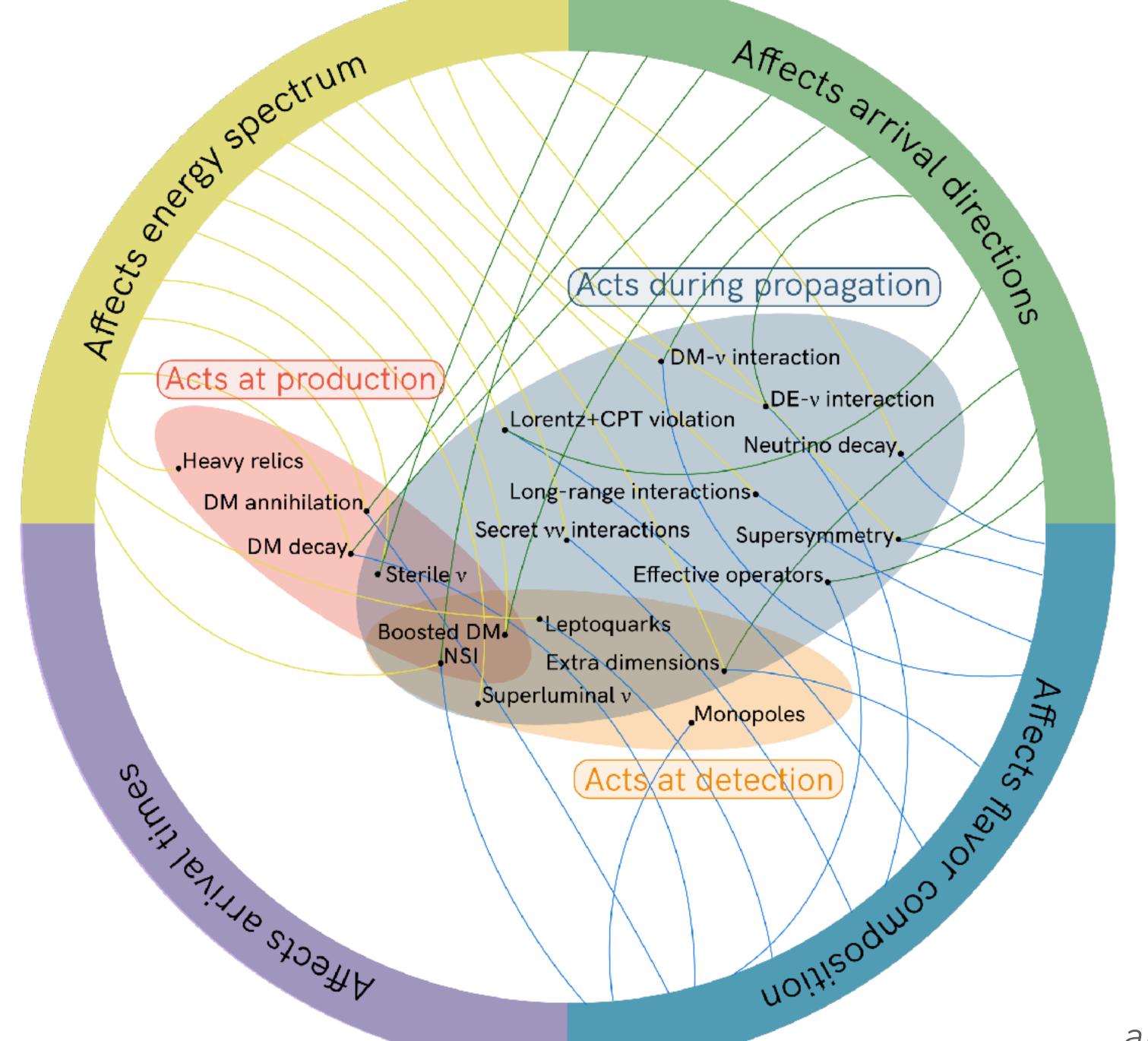
(Acts during propagation) • DM-v interaction (Acts at production) •DE-v interaction Lorentz+CPT violation Neutrino decay. .Heavy relics Long-range interactions. DM annihilation. Secret vv_interactions Supersymmetry. DM decay. Effective operators. • Sterile v *Leptoquarks Boosted DM• .NSI Extra dimensions. •Superluminal v Monopoles

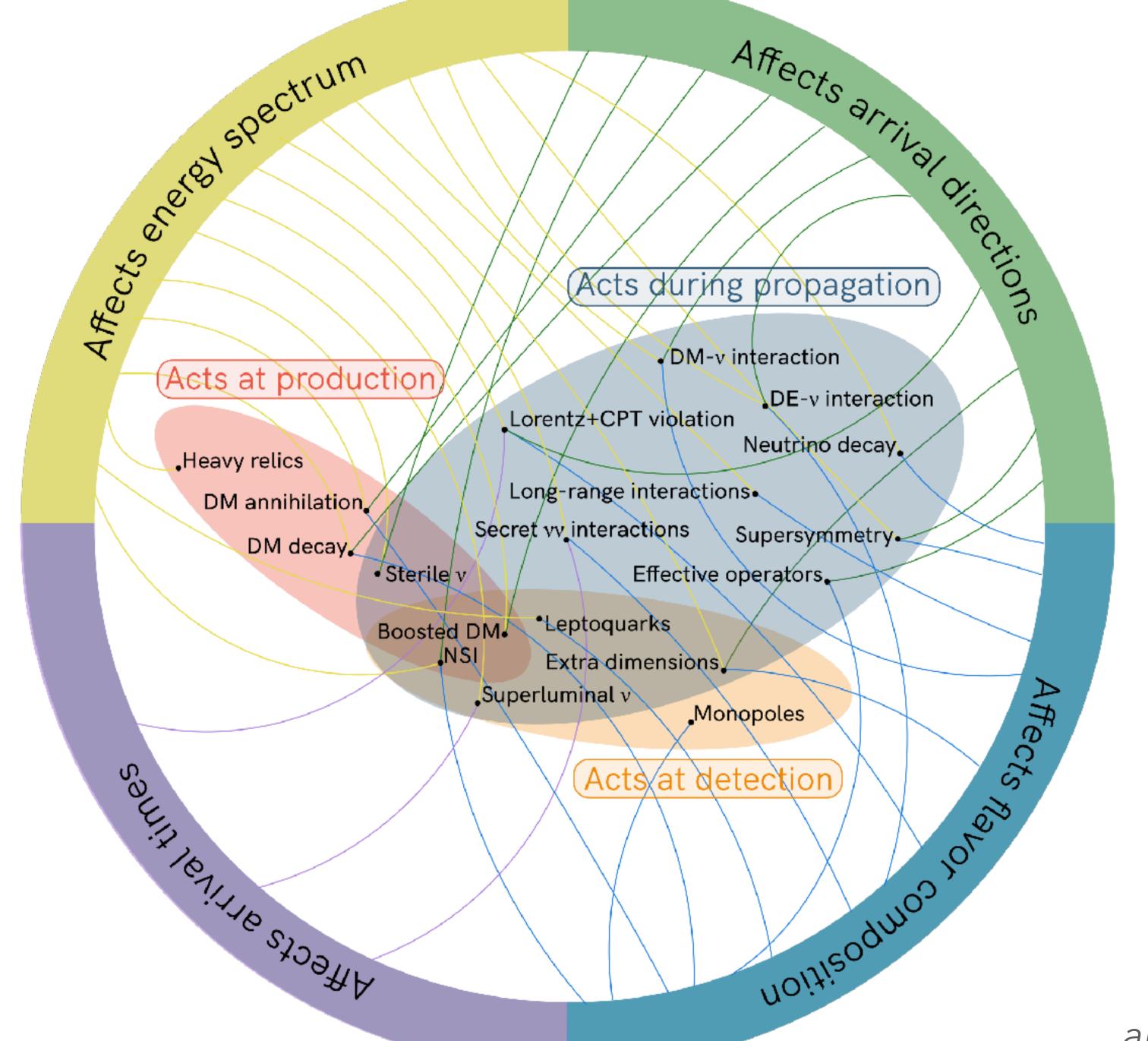






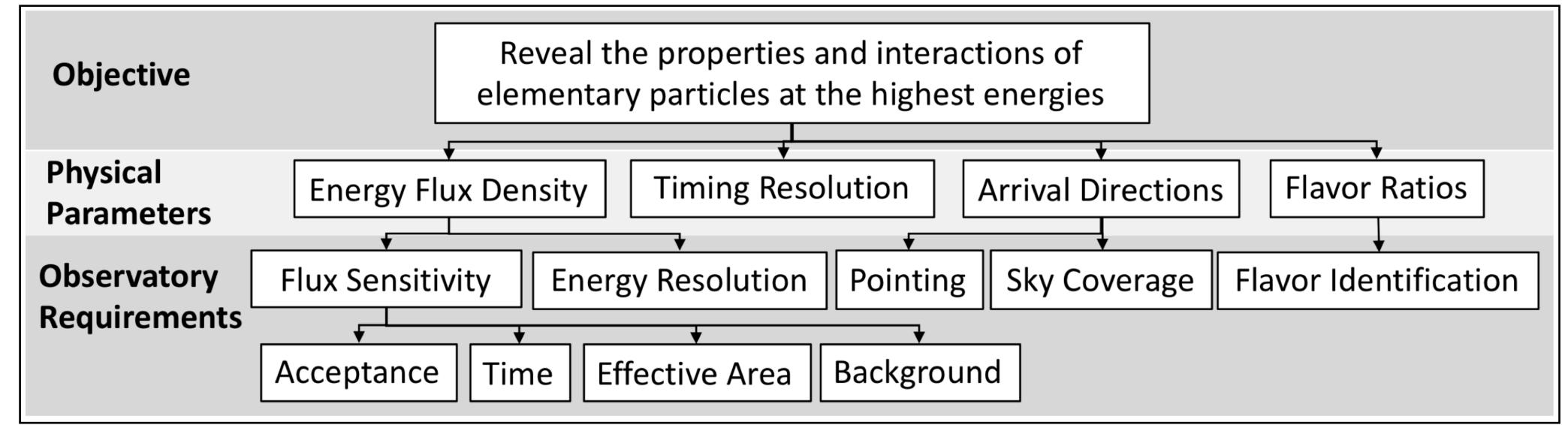


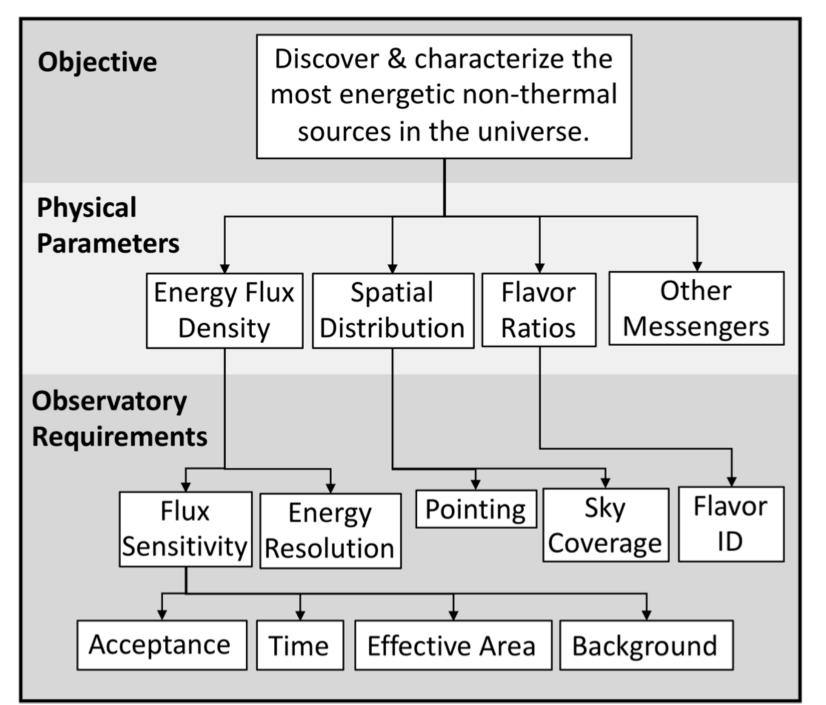




DETECTOR GOALS FOR THE NEXT DECADE

Update for Snowmass

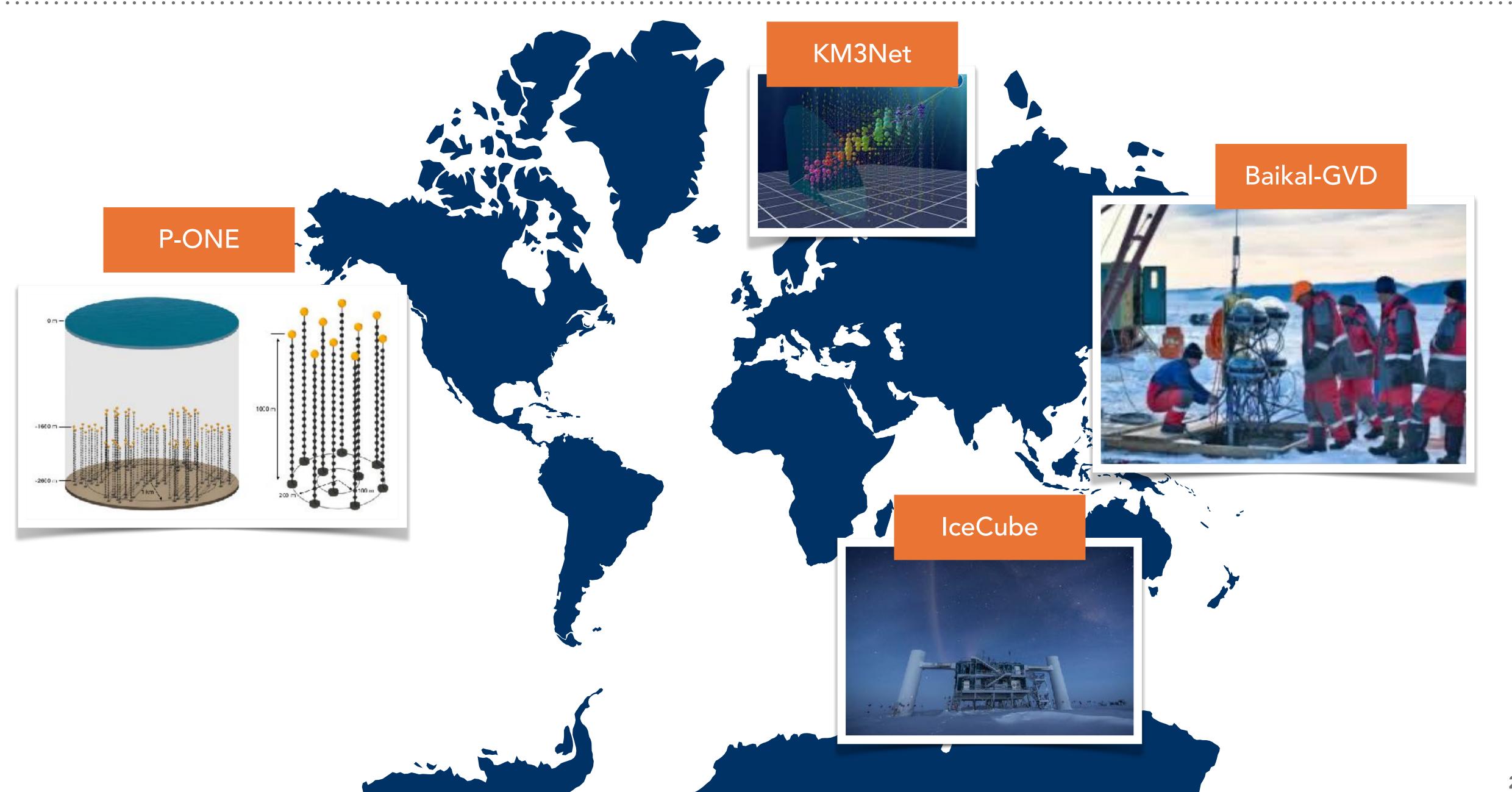




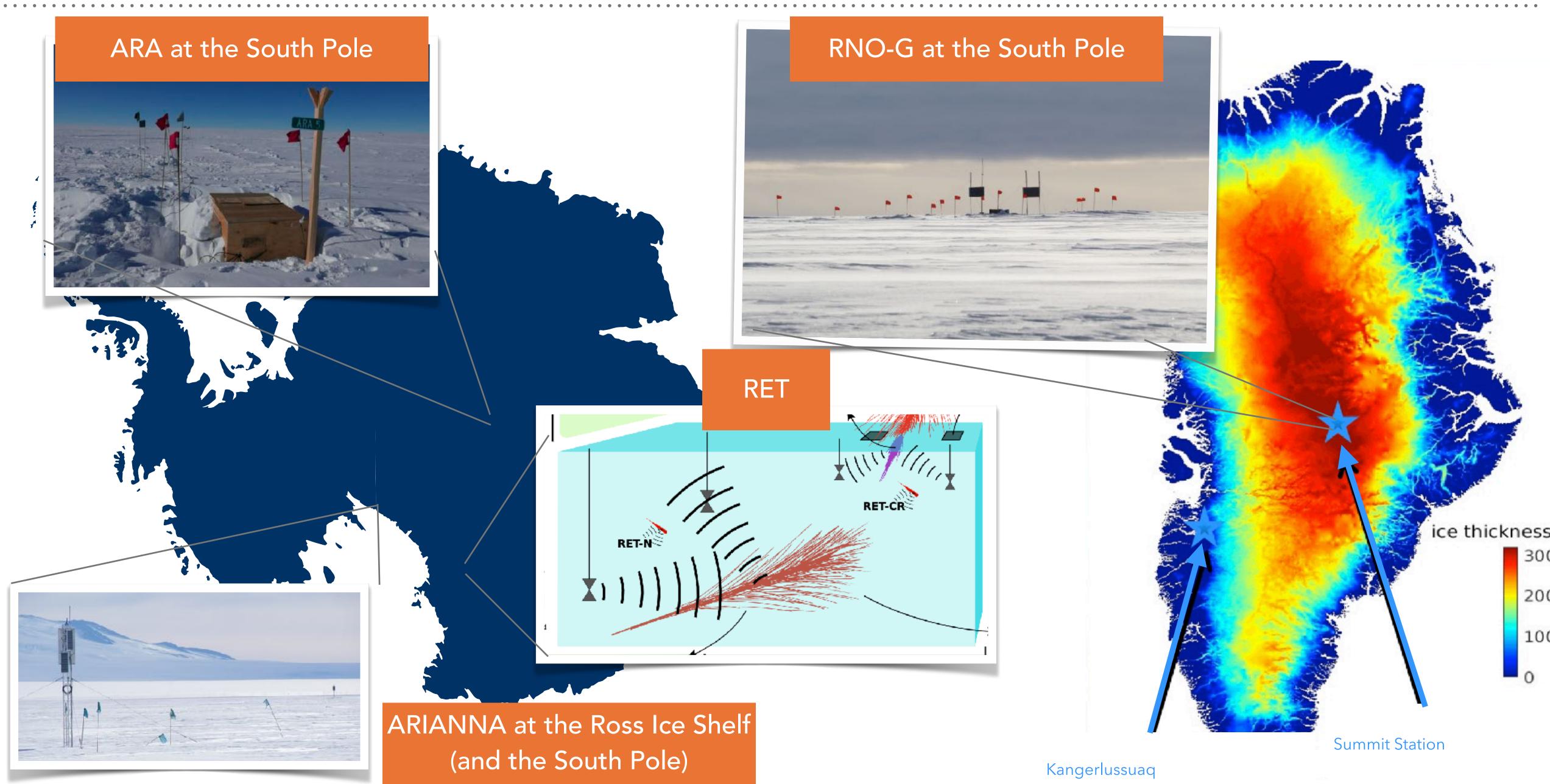
Astro2020

arXiv:1903.04334

EXPERIMENTAL LANDSCAPE: LARGE WATER CHERENKOV TELESCOPES



EXPERIMENTAL LANDSCAPE: IN-ICE RADIO EXPERIMENTS



EXPERIMENTAL LANDSCAPE: IN-ICE EXPERIMENTS



EXPERIMENTAL LANDSCAPE: EARTH-SKIMMING TAU NEUTRINOS





RELATIONSHIP WITH TAU NEUTRINO WHITEPAPER

- > Tau Neutrinos in the Next Decade: from GeV to EeV
 - Focus on tau neutrinos across a wide range of energies
 - > At HE/UHE energies, many science topics and experimental techniques overlap
 - ➤ <u>NuTau2021 Workshop</u> in September

- ➤ The HE/UHE whitepaper is broader than the tau paper in the following ways:
 - ➤ Astrophysics and particle physics topics
 - ➤ All flavor
-but more focused on >TeV energy regime

How Do I GET INVOLVED?

- ➤ Join our slack channel: #wp-cf07-he-uhe-nu on snowmass2021.slack.com
 - ➤ Find instructions to join here: https://snowmass21.org/#join_slack_workspace_and_snowmass_email_list

View-only Overleaf: https://www.overleaf.com/read/pwcjvtxjyrvy

- ➤ Contact the editor(s): Markus Ackermann, Mauricio Bustamante, Lu Lu, Nepomuk Otte, Mary Hall Reno, Stephanie Wissel
 - ➤ Indicate what you'd like to contribute. We will send you the editable Overleaf link

Schedule: First draft due by Jan. 7