NF05: Neutrino Properties

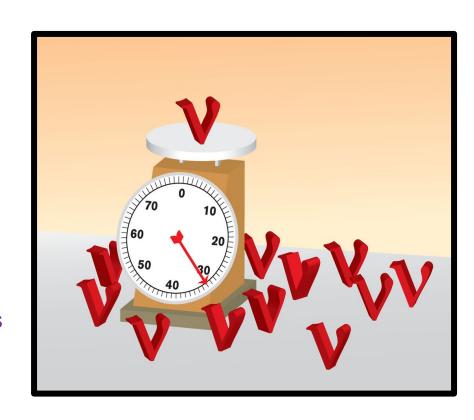
NF05 Conveners: Ben Jones, Lisa Kaufman, Carlo Giunti, Diana Parno

NF05 honorary co-conveners: Andrea Pocar (RF04), Julieta Gruszko

Group Scope

Neutrino Properties is a very broad term!

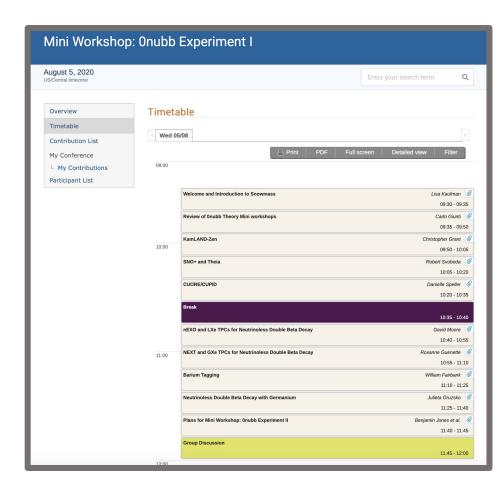
- Early in Snowmass we defined our groups scope as the following topics not covered elsewhere:
 - Direct neutrino mass measurements
 - Dirac vs Majorana nature
 - Electromagnetic properties



• Input for the NF05 report was sought via:

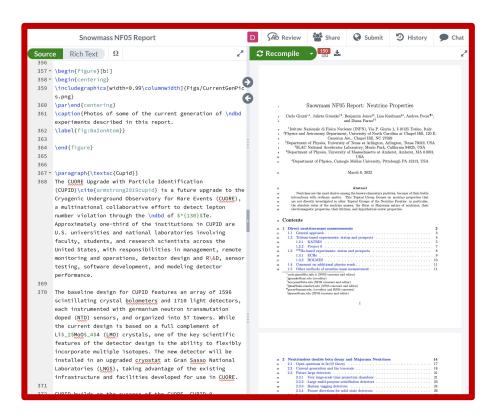
- 6 summer half-day workshops in Summer 2020
- A workshop on the ton-scale future of Onubb in Dec 2020
- Snowmass LOIs
- A fairly small number of whitepapers (not all of them yet in our hands)

 Everything presented in one of these formats <u>that we have</u> <u>received</u> is covered in our report draft.



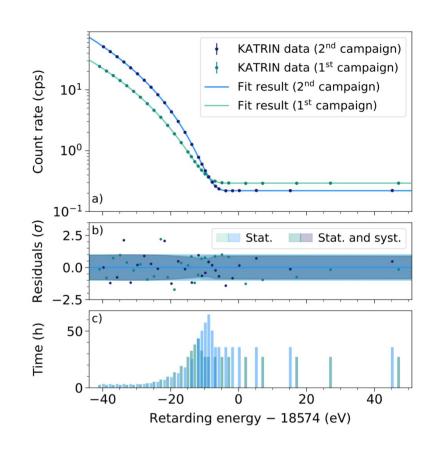
Report status

- Advanced draft in overleaf: 43 pages of text + 248 references.
- We have a few sections still to complete, expect it to top out around 46pages + 18pg of refs
- We have generally tried to be brief and where other white papers exist, or topic area is covered by another group, refer rather than rewrite.
- This means topics that have white papers take less space in our report than ones without white papers.
- We encourage readers not to equate length of text with importance of topic; instead consider it correlated to well-documented-elsewhere-ness.



Motivating considerations: Direct Neutrino Mass

- Community is fairly unified in its vision for direct neutrino mass.
- Two main experiments in tritium:
 - KATRIN (present gen)
 - Project8 (next gen)
- Two main experiments in Holmium:
 - ECHo
 - HOLMES
- Plus connections to other BSM physics searches
- And brief mention of other approaches to this question (cosmology, etc)



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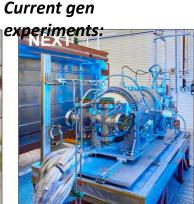
Motivating considerations: Onubb

 On the contrary, a rather wider array of perspectives exists in Onubb.

- We have tried to ensure they are all represented
 - Or at least, all the ones with US involvement!







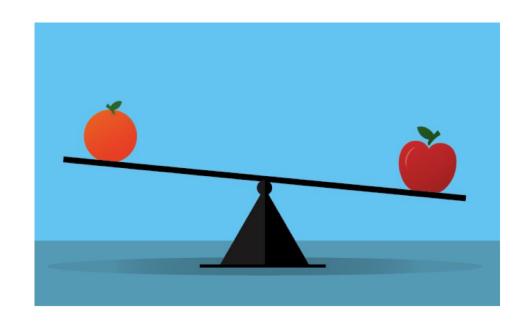




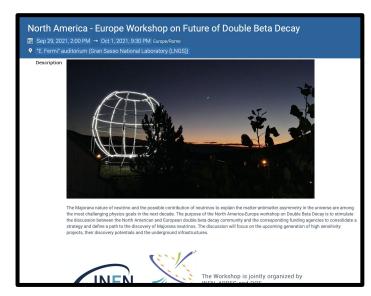


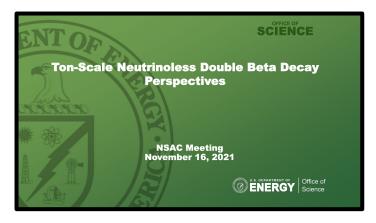
Motivating considerations: Onubb

- We have opted **not** to attempt to quantitatively compare different detectors yet to be built.
- While we can easily ask collaborations for the vital statistics of their future detector, we cannot correct for varying robustness of those projections.
- There seemed to us little value in an exercise comparing different levels of optimism in backgrounds of unbuilt detectors.



- There is an ongoing process within DOE NP concerning realization of at least one ton-scale Onubb experiment, and ideally more than one, within the next decade.
- This process has been proceeding with potentially quite consequential steps during this Snowmass process.
- We have endeavored not to interfere.
- We briefly review current-gen and ton-scale in our report, but primarily focus on the possible beyond-ton-scale future.
- The main message to convey is that:
- Searches for Onubb are a critical pillar of the world neutrino physics program; continued investment is needed to realize both this generation and the next generation of experiments.

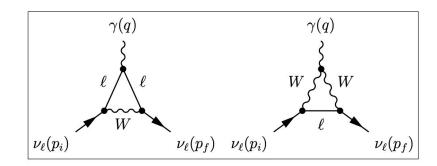




2	Neu	eutrinoless double beta decay and Majorana Neutrinos					
	2.1 Open questions in $0\nu\beta\beta$ theory						
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Motivating considerations: Electromagnetic Properties

- No major experiment in the US pursues electromagnetic properties as its primary goal.
- Nevertheless, they are important fundamental properties of neutrinos, accessible mainly through low energy neutrino scattering.
- A brief review of theoretical and experimental status is given.
- The most programmatically relevant aspect is probably connections to CEvNS, which is covered here in brief.



Method	Experiment	$\operatorname{Limit}\ [e]$	$_{\mathrm{CL}}$	Year	Ref.
Neutrality of matter	Bressi et al. [220]	$Q_{\nu_e} \in (-3.8, 2.6) \times 10^{-21}$	68%	2014	[177]
	TEXONO [221]	$ Q_{\nu_e} < 3.7 \times 10^{-12}$	90%	2006	[222]
Reactor $\bar{\nu}_e e^-$	GEMMA [195]	$ Q_{ u_e} < 1.5 imes 10^{-12}$	90%	2013	[223]
	TEXONO	$ Q_{ u_e} < 1.0 imes 10^{-12}$	90%	2014	[224]
Accelerator $(\nu_{\mu}, \bar{\nu}_{\mu}) e^{-}$	LSND [181]	$ Q_{\nu_{\mu}} < 3 \times 10^{-9}$	90%	2020	[225]
Beam Dump	BEBC [196]	$ Q_{\nu_{\tau}} < 4 \times 10^{-4}$	90%	1993	[226]
Accelerator $(\nu_{\tau}, \bar{\nu}_{\tau}) e^{-}$	DONUT [198]	$ Q_{\nu_{\mu}} < 4 \times 10^{-6}$	90%	2020	[225]
CEvNS	COHERENT [185, 186]	$\begin{aligned} Q_{\nu_e} &\in (-14, 34) \times 10^{-8} \\ Q_{\nu_e} &\in (-10, 12) \times 10^{-8} \\ Q_{\nu_{e\mu}} &< 17 \times 10^{-8} \\ Q_{\nu_{e\tau}} &< 27 \times 10^{-8} \\ Q_{\nu_{\mu\tau}} &< 20 \times 10^{-8} \\ Q_{\nu_e} &< 3.3 \times 10^{-12} \end{aligned}$	95% 90%	2020	[187]
	CONOS	$ Q_{\nu_e} < 3.3 \times 10^{-12}$ $ Q_{\nu_e} < 7.3 \times 10^{-12}$	3070	2022	[199]
Solar $\nu_e e^-$	XMASS-I	$ Q_{\nu_e} < 1.3 \times 10$ $ Q_{\nu_{\mu}} , Q_{\nu_{\tau}} < 1.1 \times 10^{-11}$	90%	2020	[204]
Table 5:	Table 5: Experimental limits for the neutrino electric charges.				

3	Neutrino Electromagnetic Properties					
	3.1	Charge Radius	37			
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Other.

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• Finally we make a very brief, passing reference to these two topics, though expect they mostly reside within the NFO3 group purview.

What did we miss?

- We will circulate report draft soon. If we missed something tell us!
- If you don't like the text about your experiment, reach out with a replacement!
- Our main goal was inclusivity and to ensure nothing that was contributed to Snowmass was omitted, while presenting an organized view of a busy and active area.
- Your feedback welcome!

