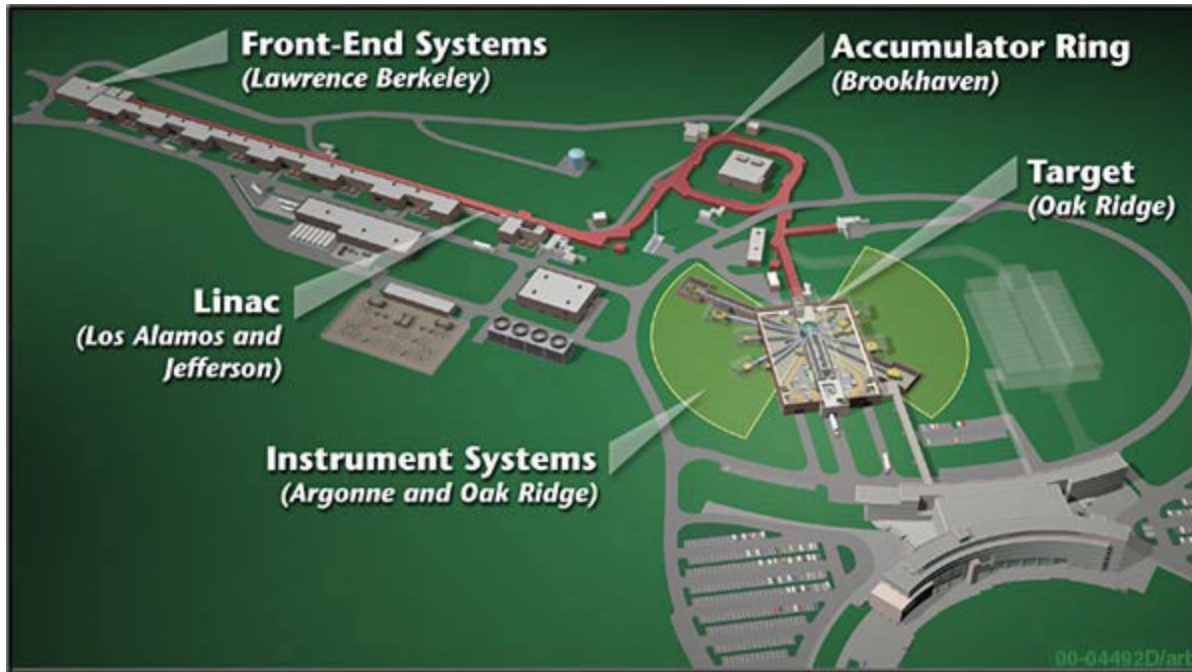


OscSNS: A Definitive Search for Sterile Neutrinos at the $\Delta m^2 \sim 1 \text{ eV}^2$ Scale

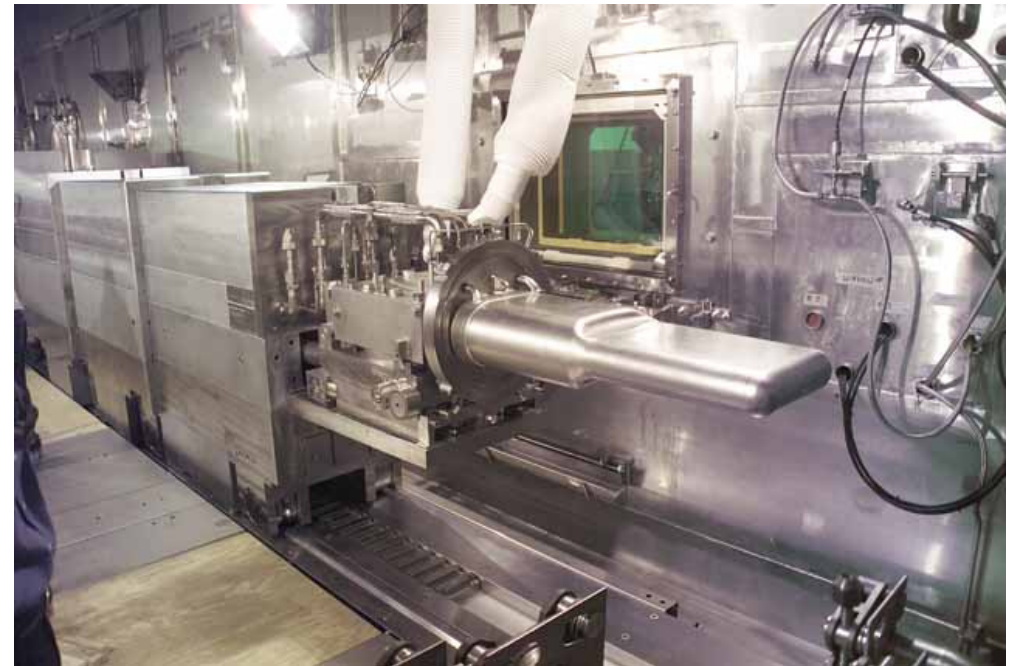
- Introduction & Motivation
- Spallation Neutron(ino) Source
- Physics Potential
- Stopped Pion source at FNAL
- Conclusions



Mercury target

Spallation Neutron Source Oak Ridge, Tennessee

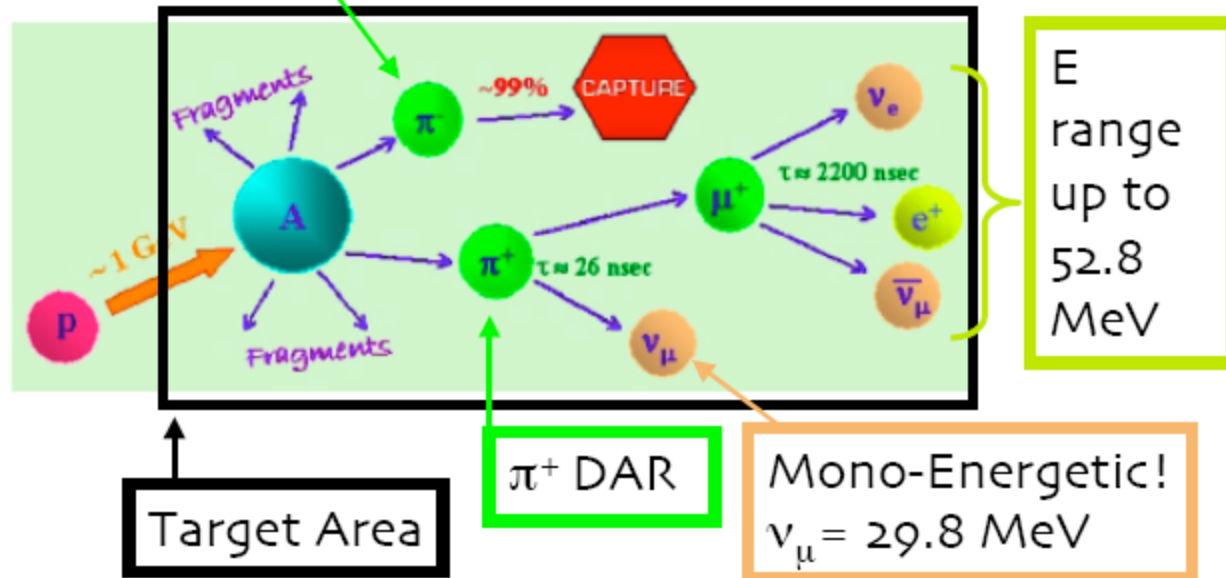
- 1.3 GeV, 1.2 MW proton beam on Hg target production of spallation neutrons
- 60 Hz with 695 ns pulse length
- Neutrinos are for free!



Neutrino production at SNS

Accelerator based Decay at Rest

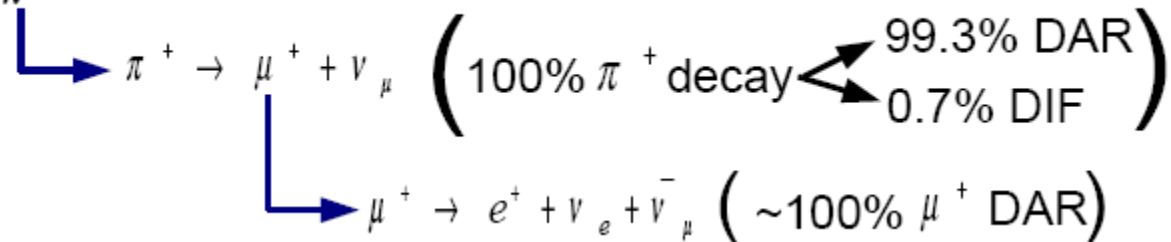
π^- absorbed by target



Target Area

Liquid Mercury (Hg+) target

11.6% p produce π^+



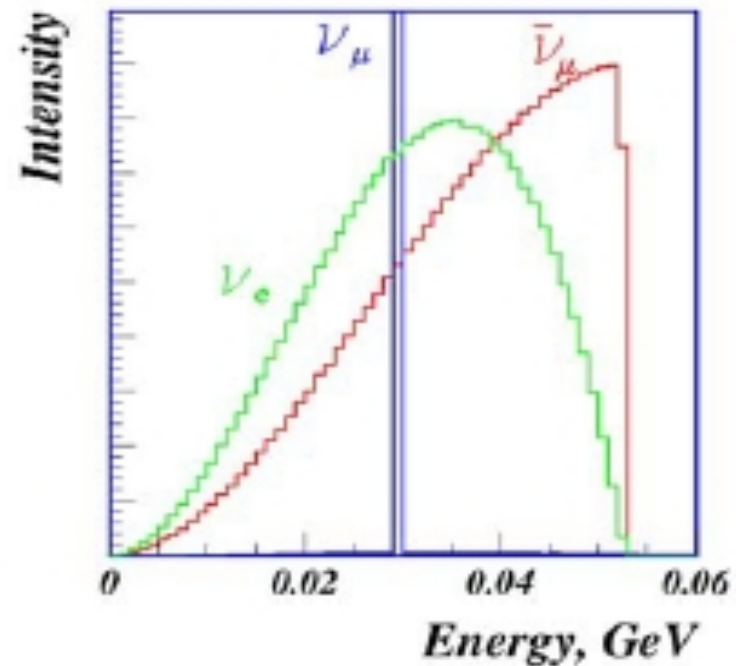
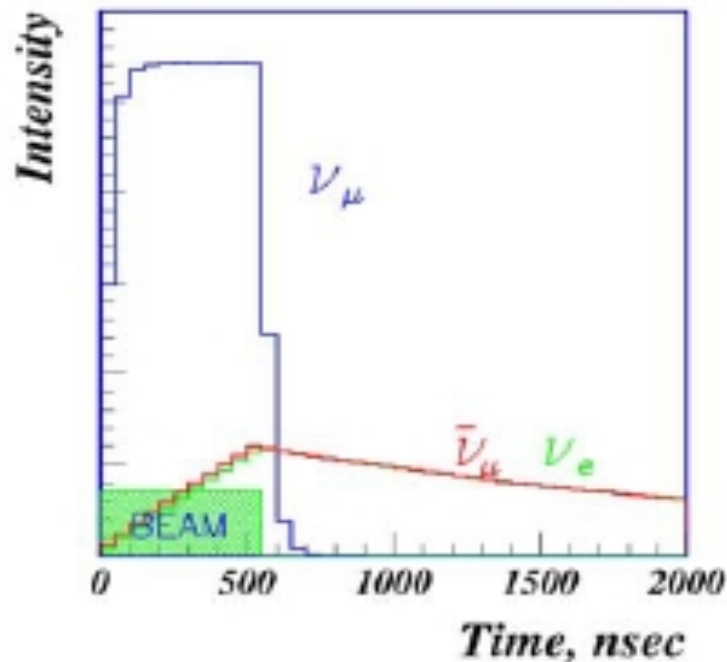
Stopped Pion/Muon Neutrino Source: Beam Time Structure and nu Flux Spectrum

→ Neutrino Flux from a Stopped Pion/Muon Source:

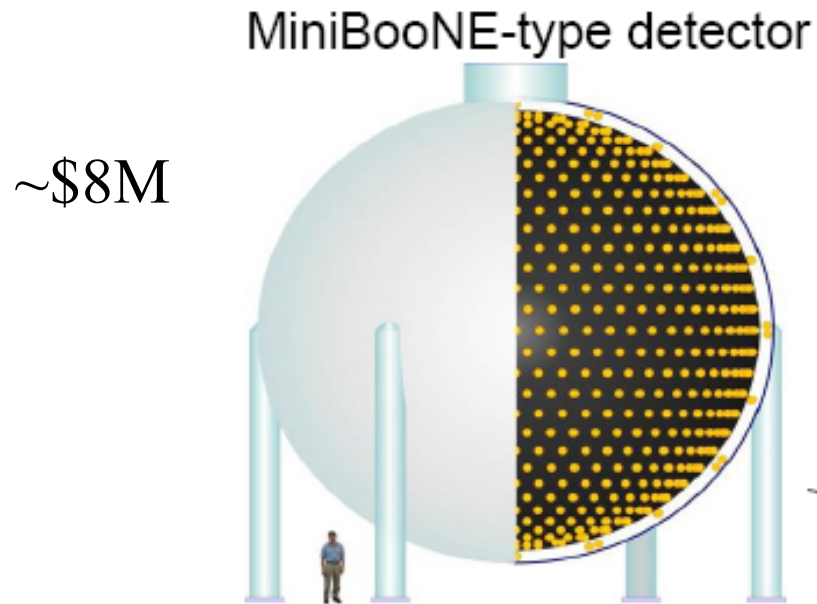
$$\pi^+ \rightarrow \mu^+ \nu_\mu, \quad \tau = 26 \text{ nsec}$$

$$\mu^+ \rightarrow e^+ \bar{\nu}_\mu \nu_e, \quad \tau = 2.2 \mu\text{sec.}$$

event types	0 → 0.695 μsec	0.695 → 5 μsec	> 5 μsec
π ⁺ DAR and π [±] DIF events	96.3%	3.7%	0
μ ⁺ DAR and oscillation candidates	14.3%	73.6%	12.1%



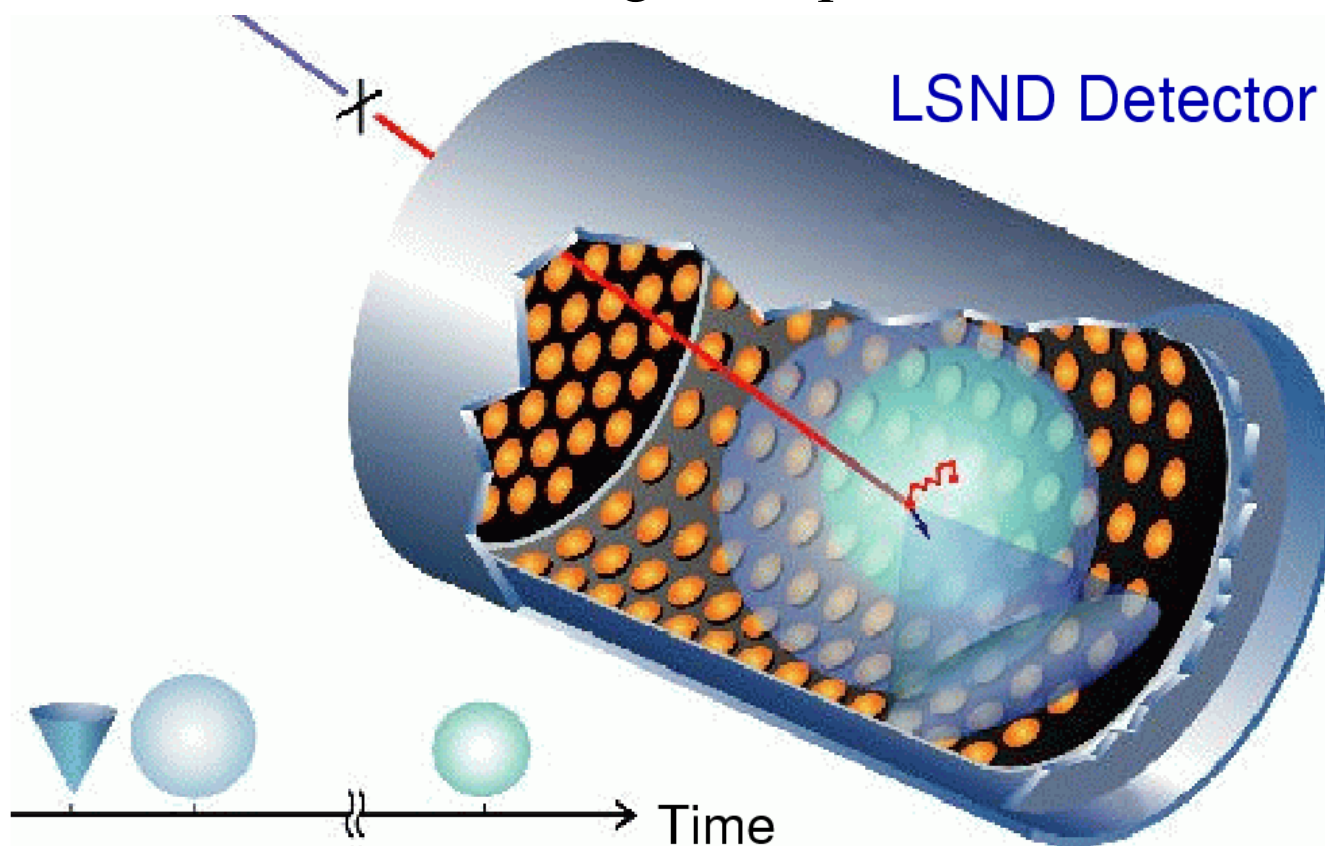
Neutrino Detection:



- 60 m upstream from the Hg target;
- 12-m diameter sphere (fiducial = 10 m);
- filled with 800 t of mineral oil CH_2 (fiducial = 450 t) + ~30 kg of butyl-PBD scintillator;
- 3502 phototubes: 3262 in detector (25% coverage) and 240 in veto region;
- buried under 10 ft of dirt – to suppress cosmic rays and beam-induced neutrons.
- Currently no ideal location for near detector due to space constraints, closest spot is ~20m

The Liquid Scintillator Neutrino Detector Concept

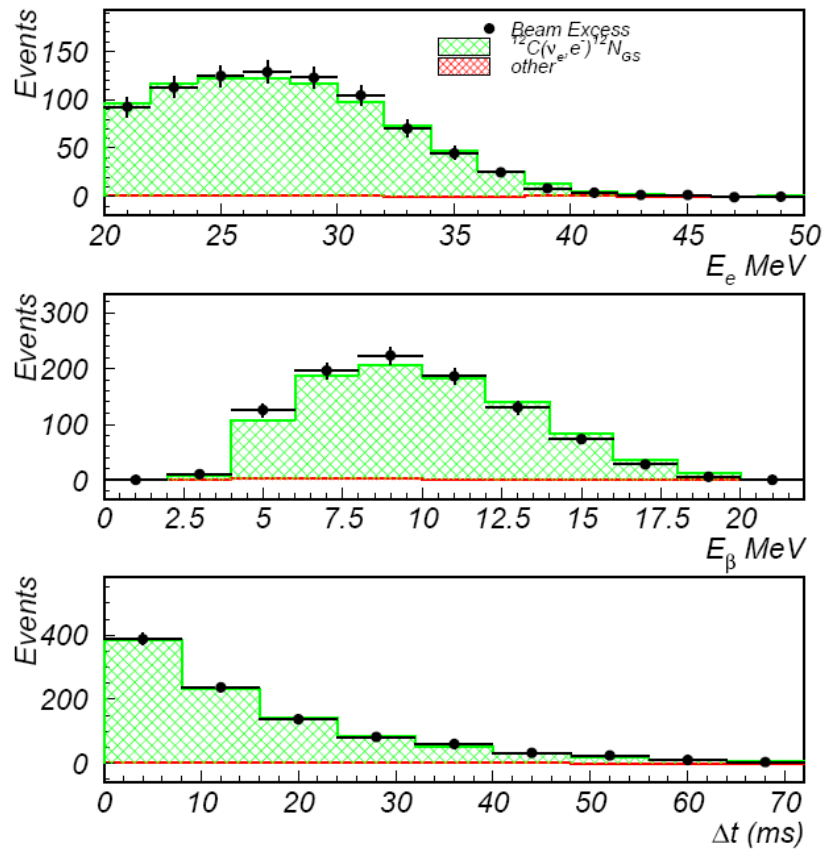
- Prompt Cherenkov light
- Delayed scintillation light
- Good electron ID, event timing, and spatial reconstruction



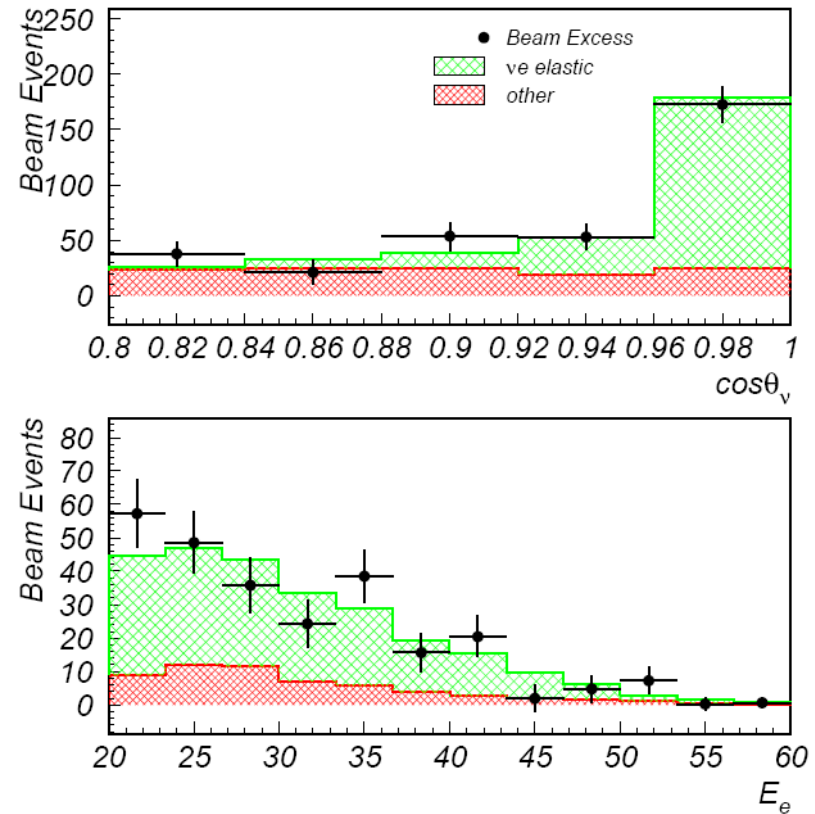
Detector calibration

- Many cosmic ray stopped muon decays
- ^{16}N source: [$^{16}\text{O}(n, p)^{16}\text{N}$] producing a γ -tagged 6.1 MeV gamma-ray.
- ^8Li source: electron energy spectrum up to 15 MeV.
- pT source: [$^3\text{H}(p, \gamma)^4\text{He}$] producing a 19.8 MeV gamma-ray.
- ^{252}Cf source: producing fission neutrons.

LSND Cross Sections



N12 ground state CC electrons



ν_e elastic electrons

OscSNS can repeat the LSND measurements with much higher precision and lower backgrounds plus additional measurements:

$\nu_\mu \rightarrow \nu_e$ appearance ($\nu_e \text{ }^{12}\text{C} \rightarrow e^- \text{ }^{12}\text{N}_{\text{gs}} + \beta$)

$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ appearance ($\bar{\nu}_e p \rightarrow e^+ n + \gamma$)

ν_μ disappearance & search for sterile ν

($\nu_\mu \text{ }^{12}\text{C} \rightarrow \nu_\mu \text{ C}^* + \gamma$) (~1300 events per year)

$\nu_e \rightarrow \nu_e$ elastic scattering (~1700 ev. per year)

νC cross sections (~4600 events per year)

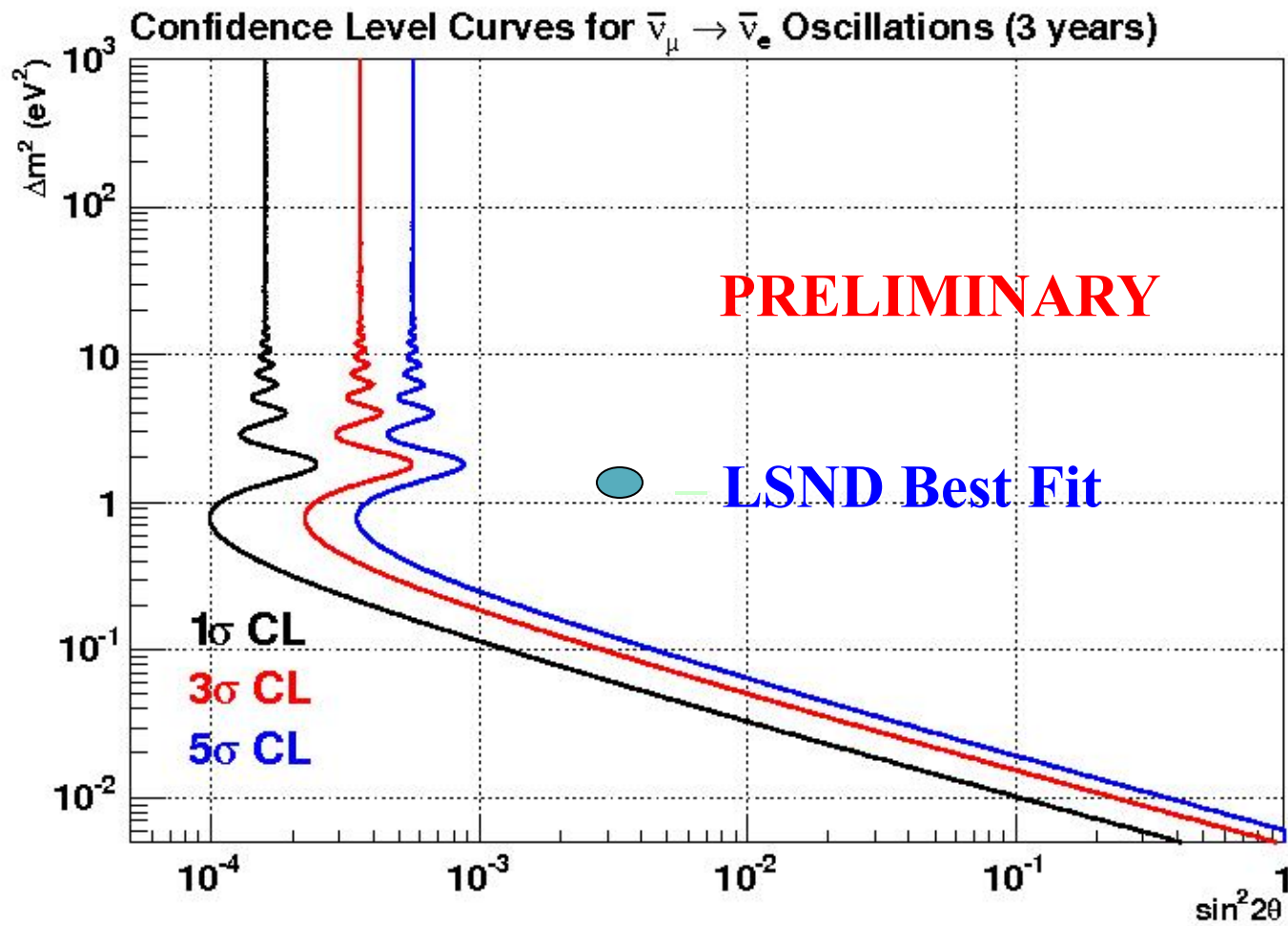
OscSNS Advantages Over Other Neutrino Oscillation Experiments

- Well understood ν flux ($\sim 7\%$, maybe better)
- Well understood ν flux spectrum
- Well understood ν cross sections (1-2%)
- Low duty factor and good timing
- Absence of nuclear effects that can affect energy reconstruction
- Very low backgrounds ($< 0.1\%$)
- Beam comes for free from the SNS
- SNS runs more than $\frac{1}{2}$ the year

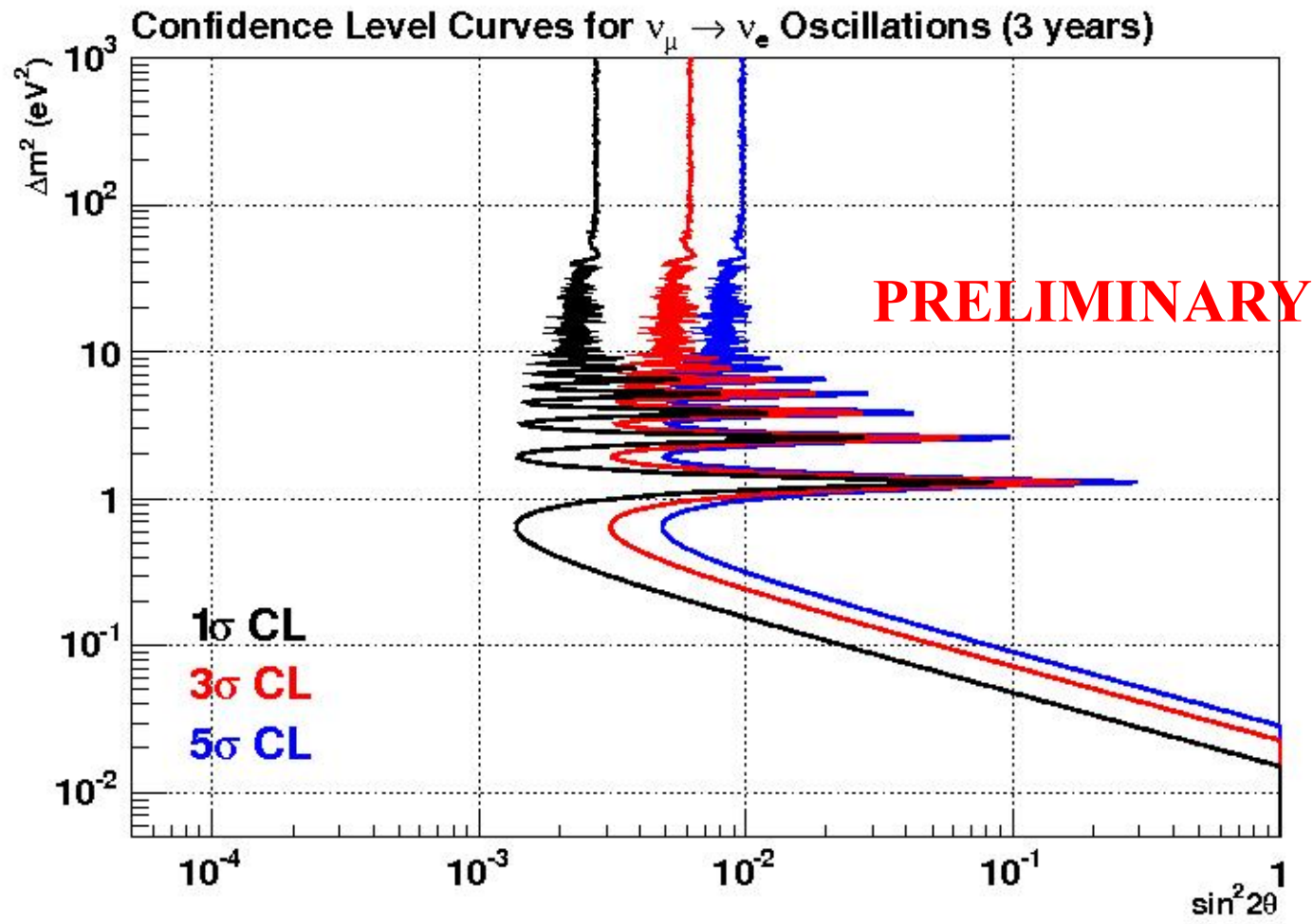
OscSNS $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ Experiment vs LSND (assuming $\Delta m^2 < 1 \text{ eV}^2$)

- More Detector Mass (x5)
- Higher Intensity Neutrino Source (x2)
- Lower Duty Factor (x100) (less cosmic background)
- No DIF Background (backward direction)
- Lower Neutrino Background (x4) (60m vs 30m)
- Better Signal/Background (x4)
- For LSND parameters, expect ~ 350 $\bar{\nu}_e$ oscillation events & < 50 background events per year!

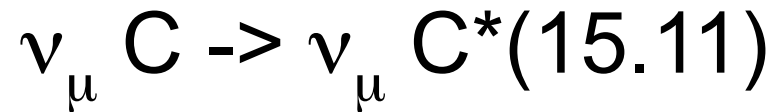
OscSNS Oscillation Sensitivities



OscSNS Oscillation Sensitivities

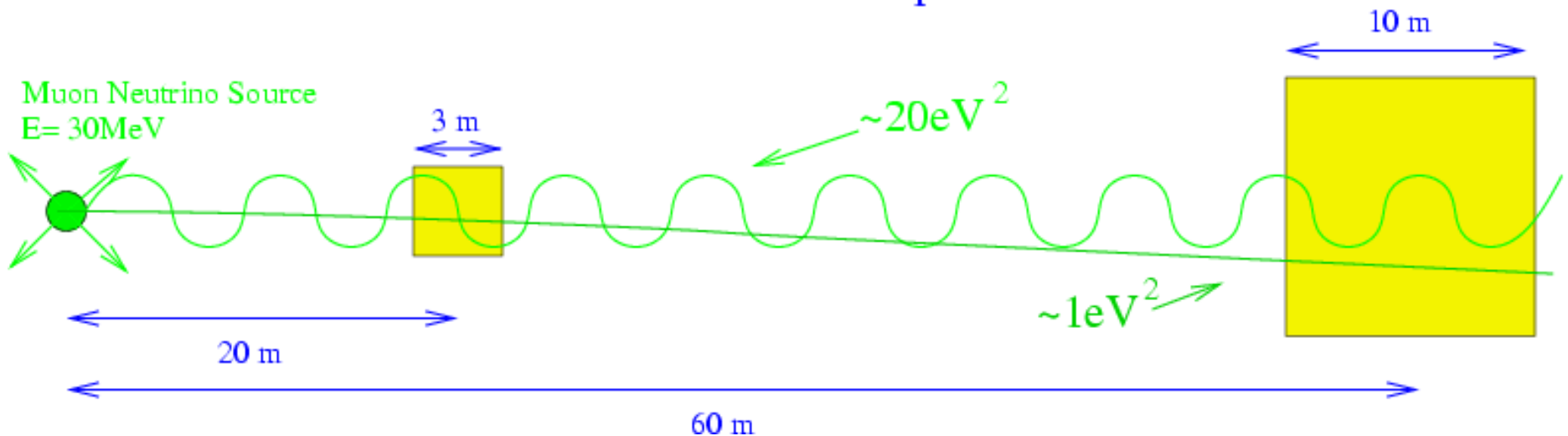


A Smoking Gun Search for Sterile Neutrinos Via Measurement of NC Reaction:



Garvey et al., Phys. Rev. D72 (2005) 092001

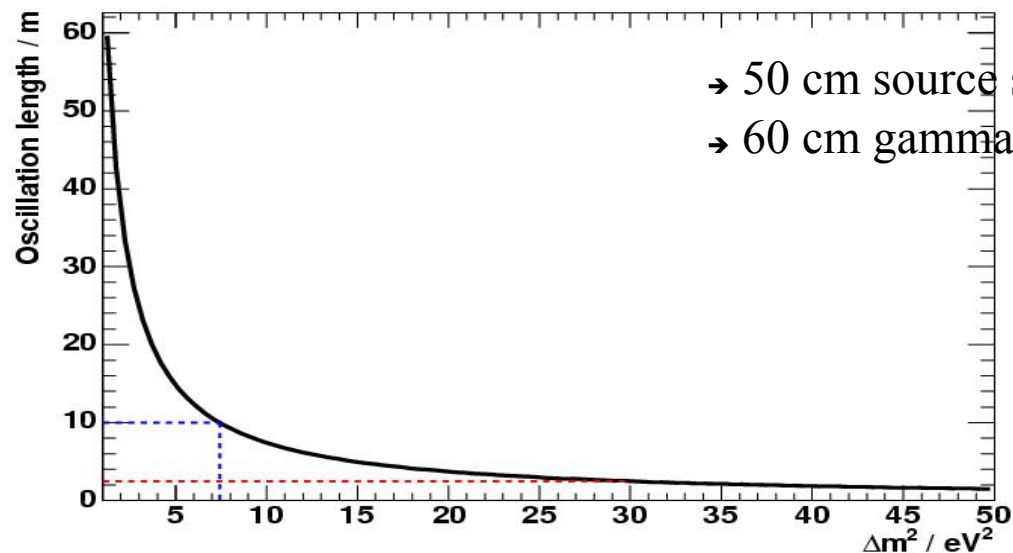
Neutral Current Disappearance Pattern
in a Two Detector Setup



Active-Sterile Neutrino Oscillations with Stopped Pions (hep-hp/0501013)

- If LSND oscillations is $\nu_\mu \rightarrow \nu_s \rightarrow \nu_e$, then we expect $P(\nu_\mu \rightarrow \nu_s) > 0.10$
- Can detect **all neutrinos** via NC reaction, $\nu_x {}^{12}\text{C} \rightarrow \nu_x {}^{12}\text{C}^*(15.11\text{MeV})$.
- Since we have **monoenergetic** ν_μ source, then look for NC rate distortion as a function of L.

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2(2\theta) \sin^2\left(\frac{1.27}{30} \delta m^2 L\right)$$

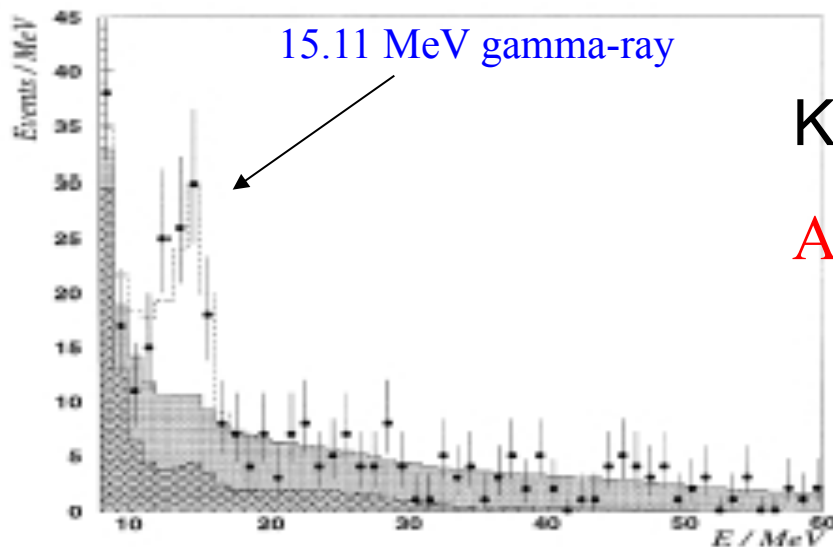


→ 50 cm source size

→ 60 cm gamma Compton scattering.

Sterile Neutrino Oscillations with $\nu_{\mu}^{12}\text{C} \rightarrow \nu_{\mu}^{12}\text{C}^*(15.11\text{MeV})$

- KARMEN measures $\sigma = (3.2 \pm 0.6) \times 10^{-42} \text{ cm}^2$



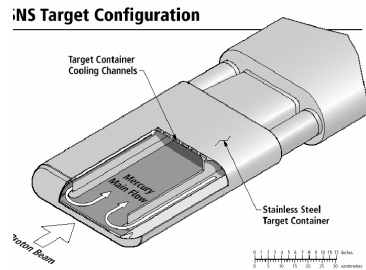
KARMEN gamma $\sigma(E) \sim 2\%$

At SNS design for $\sigma(E) \sim 4$ to 7%

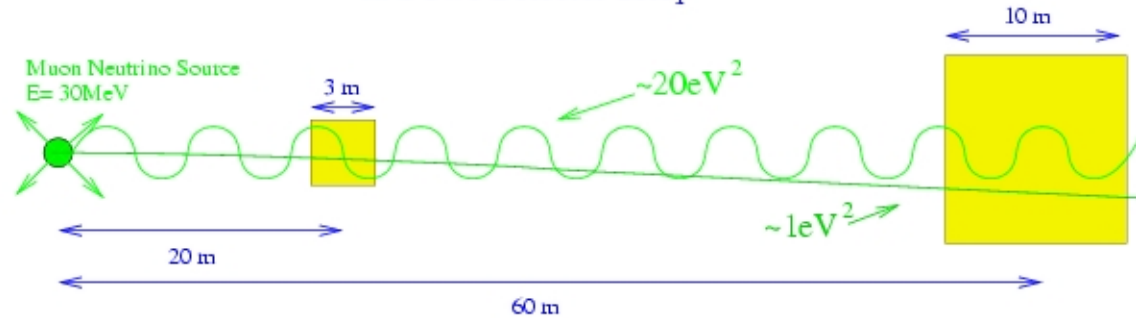
- Expected Backgrounds from NC $\bar{\nu}_{\mu}$ and CC ν_e , can be measured from beam off/out of time.

Sterile Neutrino Oscillations with $\nu_\mu^{12}\text{C} \rightarrow \nu_\nu^{12}\text{C}^*(15.11\text{MeV})$

SNS source L=50 cm

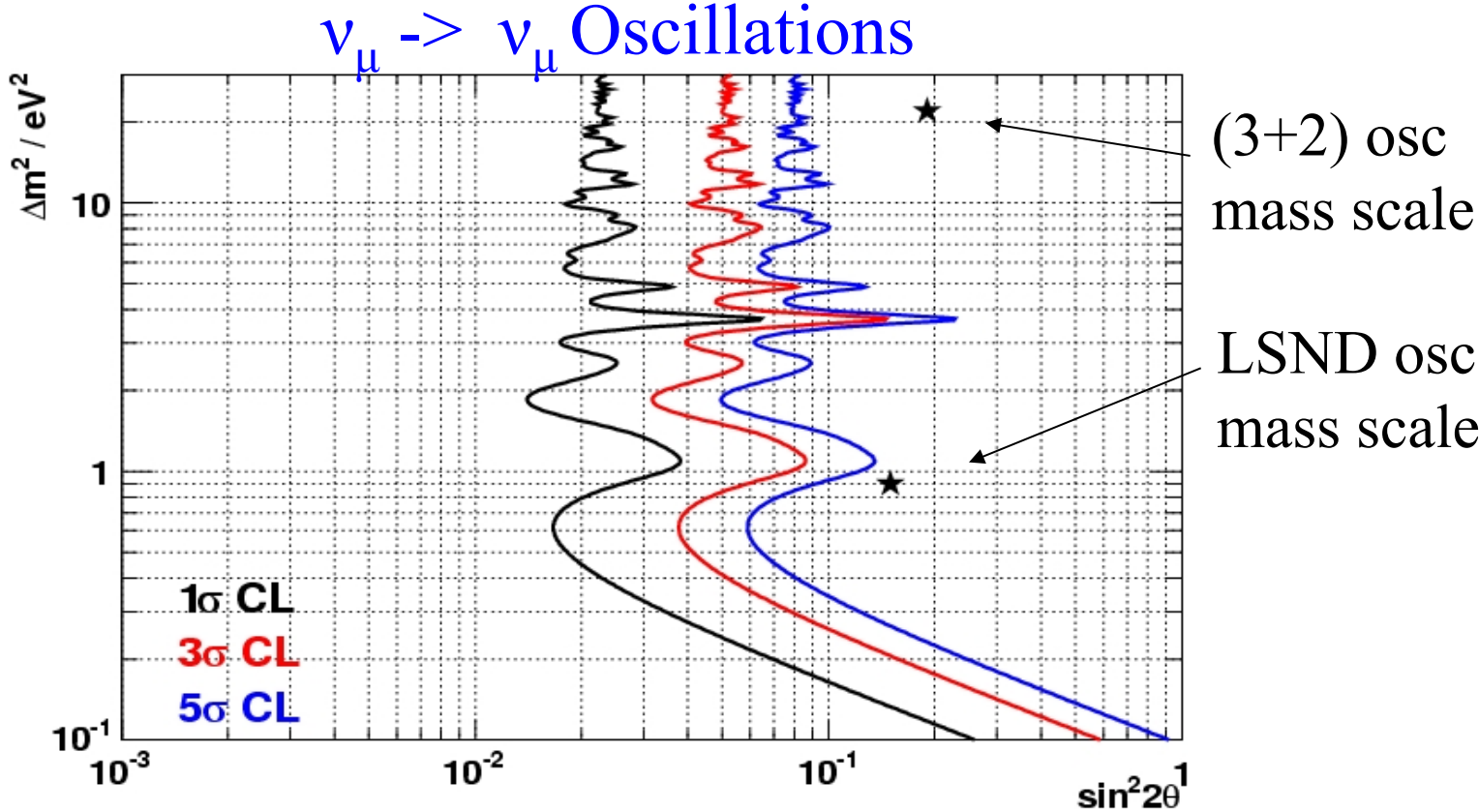


Neutral Current Disappearance Pattern in a Two Detector Setup



Detector	Source Dist. (m)	FD Size (tons)	FD Length (m)	$\nu_\mu^{12}\text{C} \rightarrow \nu_\nu^{12}\text{C}^*$ events/year
SNS Near	18	25	3	2056
SNS Far	60	500	10	3702

Sterile Neutrino Oscillations Sensitivity with SNS Source and Two Detectors (3 years), 5% flux+xsec systematic error.



Stopped Pion Source at FNAL

- MiniBooNE target area is a source of stopped pions! Calculated ν_μ fluxes (Steve Brice):
 - LSND at 30m: 1.2×10^6 $\nu/\text{cm}^2/\text{s}$ (detector 167 tons)
 - KARMEN at 17m: 1.6×10^6 $\nu/\text{cm}^2/\text{s}$ (detector 56 tons)
 - OscSNS at 60m: 1.2×10^6 $\nu/\text{cm}^2/\text{s}$ (proposed detector 1kton)
 - BNB at 15m: 0.7×10^6 $\nu/\text{cm}^2/\text{s}$
- However, MiniBooNE source very extended $\sim 5\text{m}$, and at higher energy there are Kaons (background or source?)
- Can build a dedicated target Hall on the BNB, take advantage of higher POT with Project X.

Conclusions

- A stopped pion beam provides a source of neutrinos with a well characterized flux (flavor, magnitude, and energy) and interaction cross sections.
- Using the stopped pion source at SNS could provide definitive ($>5\sigma$) evidence that LSND observed an excess of $\bar{\nu}_e$ events in a $\bar{\nu}_\mu$ beam which could be interpreted as oscillations at the $\sim 1\text{eV}^2$.
 - Two detector would prove if this was oscillations.
- A stopped pion source (SNS/BNB) with short duty cycle and two detectors could provide definitive evidence of sterile neutrino oscillations at the $\sim 1\text{eV}^2$ scale.