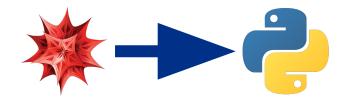




Case Study -- Python Implementation

Dylan J Temples MAGIS-100 Simulation & Analysis Workshop 26 January 2022

Case Study -- Python Implementation



Goal: Port all functionality of the Mathematica case study notebook to Python 3 Motivation:

- Mathematica is proprietary software, licenses are expensive for non-Universities
- Python is open source, widely used, well documented, and powerful
- Utilizing high-performance computing resources for simulations (e.g. FNAL grid)

All methods available in the Mathematica notebook are duplicated in Python, mostly with identical names, arguments, and returned quantities.



Git Repository and Installation

Currently available on personal github (private), will migrate to the MAGIS Redmine

https://github.com/dtemps123/MAGIS analysis

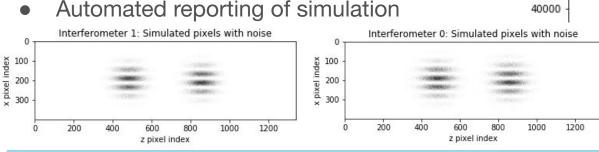
	E README.md
Clone with git	MAGIS-100 Case Study Python Tools
Install with pip	Dylan Temples October 22, 2021
Fully documented in the git Readme	This notebook details the usage of the backend MAGIS Case Study Python Tools, and should serve as some level of documentation. This goes through step-by-step usage of the main functions an average user would use to do an analysis. At some critical points, it compares the calculations in this notebook to the same calculations done in the "MAGIS-100 Camera Example Parameters R21.nb" Mathematica notebook available on the DocDB: https://nd-docdb.fnal.gov/cgi-bin/sso/ShowDocument?docid=615.
 Usage examples script-based interactive (Jupyter) 	Installation To download this repository: git clone git@github.com:dtemps123/MAGIS_analysis.git (this uses SSH so you'll need to add your public key to your github account). Once downloaded, move to that directory: cd MAGIS_analysis From here, the Python package can be installed which allows you to use this framework in code that lives outside this directory: pip install .



Current Functionality

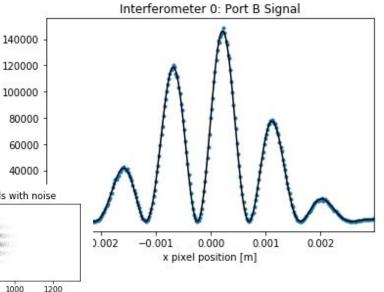
What does it do?

- Statistically generate atom positions from wavefunction
- Discretizes photon counts w/ noise across pixels to generate image using parameterized optics settings
- Simultaneous fits of up to 4x interference patterns to determine phase shift



What doesn't it do?

- Optical simulations (ray tracing)
- Quadratic wavefunction/fitting



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Performance vs Mathematica

Example code to generate one image per interferometer and extract the phase

Metric	Mathematica	Python
Time to simulate & fit images 100 times [sec]	5550	1487*
Mean Δφ [rad]	-0.000204	-0.003312
Mean Δφ error [rad]	0.000358	0.001577
Mean fit χ^2 /Ndof	16.6462 / 14	16.4153 / 14

* Most time is spent creating marginalized position distributions (only needs to be done once). Image generation and fitting is very fast.

```
1 import matplotlib.pyplot as plt
3
    from MAGIS_analysis import Parameters as sp
    from MAGIS_analysis import Interferometer as ic
    from MAGIS_analysis import Wavefunctions as wf
    from MAGIS_analysis import SimTools as st
    from MAGIS_analysis import FitTools as ft
9
    ## Show the fundamental constants
10 print(sp.hbar)
11 print(sp.m)
12 print(sp.in_to_m)
13 print(sp.Gamma)
14 print(sp.pi)
15
16
   ## Show the user options
17 sp.PrintUsrOpts()
18
    sp.PrintImgParams()
19
    sp.PrintSnsrParams()
20
21 ## Load up the wavefunctions
   wf.PrintParams()
22
24
   ## Create the interferometer objects
25 int0 = ic.Interferometer()
    int0.PrintParams()
26
27
28
    int1 = ic.Interferometer()
    int1.SetParam("atomsPerCloud", 1.1e6)
29
    int1.SetParam("cloudRadius" , 0.0018)
30
    int1.PrintParams()
31
32
33
    ## Calculate the marginal distributions of atom positions
34
    ## in both interferometers
    st.CalculateMarginalDistributions(int0)
35
    st.CalculateMarginalDistributions(int1)
36
37
38
    ## Generate a single image from each interferometer
    img0, x_marg0, z_marg0 = st.GenerateImage(int0, recalculate_pdfs=False)
    img1, x_marg1, z_marg1 = st.GenerateImage(int1, recalculate_pdfs=False)
40
41
42
    ## Extract the images from each port on both interferometers
43
    xvalsbox0, portAarray0, portBarray0, nzsummedA0, nzsummedB0 = ft.Extract2Ports(img0, x_marg0, z_marg0, 3, 3, int0
    xvalsbox1, portAarray1, portBarray1, nzsummedA1, nzsummedB1 = ft.Extract2Ports(img1, x_marg1, z_marg1, 3, 3, int1
44
45
46
    ## Fit the image distributions
47
    fr = ft.SimFits_2Port2Int(xvalsbox0, portAarray0, portBarray0, nzsummedA0, nzsummedB0
48
                             xvalsbox1, portAarrav1, portBarrav1, nzsummedA1, nzsummedB1
                             int0 , int1)
49
50
    fr, show()
51
52 plt.show()
                                                                            🚰 Fermilab
```

Next Steps & Interfaces

To do:

- Improve fitting to reduce uncertainty
- Implement quadratic wavefunction and fitting
- Implement single interferometer and single port fits
- Improve automated reporting
- Test all combinations of user options
- Optimization and parallelization

Potential interfaces:

- Optical simulation tool: use simulated atom positions to generate images
- Integration into data pipeline for quick online analysis / performance monitoring
- Integration with signal model choice: phase shift vs time







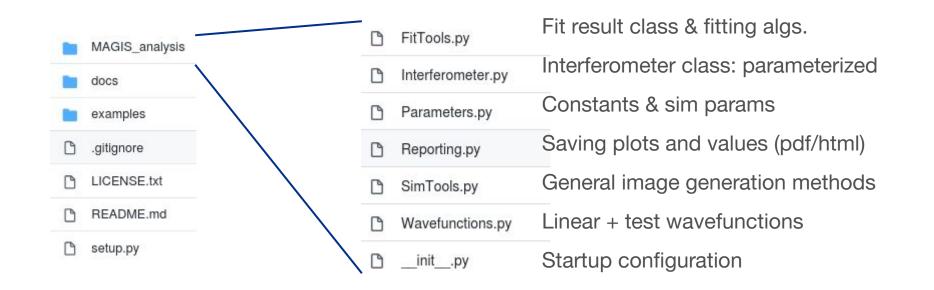
Thank You

Case Study -- Python Implementation

https://github.com/dtemps123/MAGIS analysis

Dylan J Temples dtemples@fnal.gov

Package Contents





Package Contents

FitTools.py	Interferometer.py	Wavefunctions.py
<pre> class FitResult(): </pre>	7 ## Create an instantiable class	
	8 class Interferometer:	8 tFinalBS = 3.0 ## [s]
def init (self,	$n param$): $\frac{9}{10}$ id_n = 0	9 tExtra = 0.1 ## [s] time be
	10 11 ## Define Default Interferon	10 portBvz = 0.15 ## [m/s] diffe
<pre>def show(self): ==</pre>	12 phiphi = $sp.pi / 2$.	11 2Sep = politov2 (LXLTa ##)
	13 atomsPerCloud = $1e6$	12 KITINGC - 170.0005
def show cov mat(s	<i>elf</i>): = 14 cloudRadius = 0.002 ##	[] 13 yFringe = 2.0 * sp.pi / kFringe
	15 xCenter = 0.000 ##	i 14 aQuad = 1e-12
def show cor mat(s	elf): = 16 zCenter = -0.0075 ##	
	17	16 ## Dictionary for Sensor Paramet
def params to dict	(self): 18 ## Derived Parameters	17 ⊳ params = { ==
	acceptedPhotons = atomsPerce	
def p errs to dict	(self): 20 signalPerCloud = acceptedPt	
	(Sell): 21 pixelAcceptance = (sp.snsr_r 22 signalPerPixel = pixelAccep	
<pre>> def GetFitQuality(data</pre>	x values, 23	32 ► def SetParam(key, val): ==
	24 ## Initialization Method	36
<pre>> def Z2Ports1Int_FitFun</pre>		
	44	43
<pre>> def FitZ2Ports1Int(z d</pre>	list hots 2 45 def CalculateDerivedParams(s	self): 44 ► def PrintParams():
		51
def X2PortsCombined Fi	tFunc(x, an 58 def SetParam(self, key, val)	JZ ## Delitie the actual wavefullette
	59► <i>def</i> GetParam(<i>self, key</i>):=	53 ⊳ <i>def</i> testfunction(x, z, xA, zA, µ
<pre>> def FitX2PortsCombined</pre>	(x dist hpt 65	65
	66 ► def PrintParams(self): =	66 ► def wavefunction(x, z, xA, zA, μ
def Extract2Ports(pxl_		
	74 ► def GetName(self): =	<pre>118 ► def EvaluateWavefunction(wf_obj, 120</pre>
def SimFits 2Port2Int(int0 x. int	138 130 - dof Evoluato(intfr x array z t
		139
		😴 Fern

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Package Contents

SimTools.py

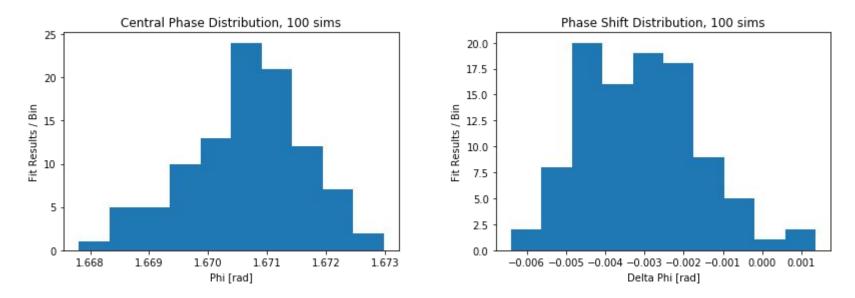
	def	<pre>PrintDerivedParameters(intfr0,intfr1): =</pre>
27 28 ► 78	def	CalculateMarginalDistributions(<i>intfr</i> , <i>n_samps</i> =2000): =
	def	<pre>SampleFromInvertedCDF(rvars, pvals, n_samps): =</pre>
97 ► 113	def	WaveFunctionPixelation(<i>intfr</i>): =
114 ► 134	def	<pre>PhotonPixelation(phtn_array): =</pre>
135 ► 160	def	<pre>DiffusedAtomPositions(N_atoms): =</pre>
161 ► 179	def	RandomSampleAtomPositions(<i>intfr, port</i> ="A"):
180 ► 209	def	<pre>ProjectToPixels(portA_atoms_x, portA_atoms_z, portB_ato</pre>
210 ►	def	GenerateImage(intfr, recalculate_pdfs=False): =

Reporting.py

7 8 9	<pre>data_path = "./" ## Path where this series_id = "2222-9999_0000" ## The uni</pre>
10 11	<pre>smry_prefix = "log" ## Appended to t smry_suffix = ".md" ## Appended to t</pre>
12 13 14	enable = False ## Do you
15 ► 23 24	<pre>options = {</pre>
25 ► 28	<pre>def SetOpt(key, val): =</pre>
35	<pre>def GetOpt(key): == def PrintOpts(): ==</pre>
43 44 ►	<pre>def Initialize(path_str="./",series=None):</pre>
92 93► 117	<pre>def LogTextLine(*message): =</pre>
118 ► 142	<pre>def LogOutputLine(*output_msg): =</pre>
163	<pre>def LogCodeBlock(*code): = def LogTmage(img_file_namefig_ebi): =</pre>
185	<pre>def LogImage(img_file_name, fig_obj): == def LogMarkdownDump(mkdn str): ==</pre>

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Phase Shift Statistics





1D Fit Result: Python vs Mathematica

