

Neutrino-electron scattering Project X

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Topics

- Neutrino magnetic moment
- Weak mixing angle $\sin^2(\theta_w)$ at low Q^2
- Neutrino Flux Calibration

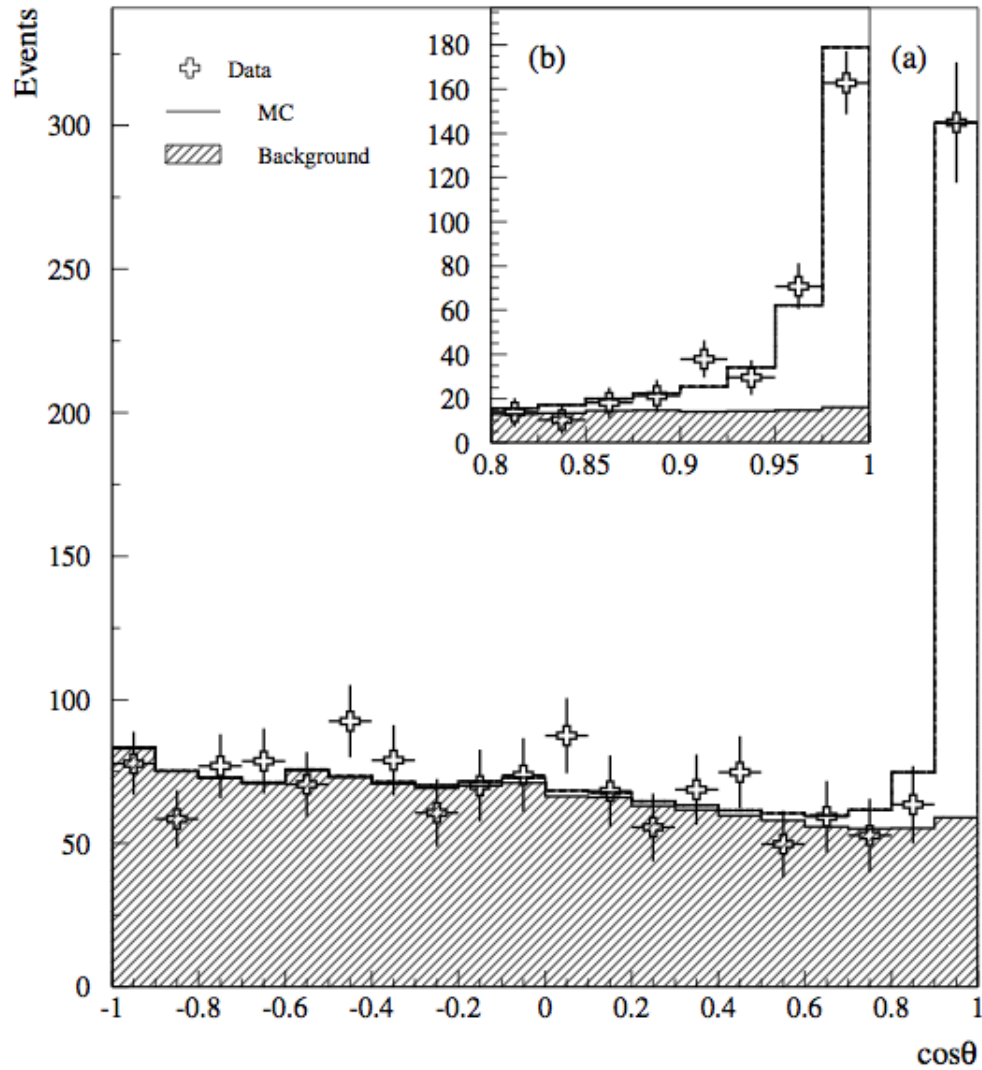
Neutrino Magnetic moment

- In minimally extended SM with Dirac Neutrinos expect less than 10^{-19} Bohr Magnetons.
- Majorana neutrino transition moments or right-handed weak currents could enhance to experimentally relevant ranges.

Neutrino Magnetic Moment

- Present limit for ν_μ is $6.8 \cdot 10^{-10}$ Bohr Magnetons. LSND PRD 63,112001
- Present limit for ν_e is $0.5 \cdot 10^{-10}$ Bohr Magnetons. Borexino PRL 101, 091302
- Neutrino mixing complicates interpretation of the limits. Could be Dirac or Majorana
- Because of mixing ν_μ limit may not be very useful if it is much worse than ν_e limit

LSND



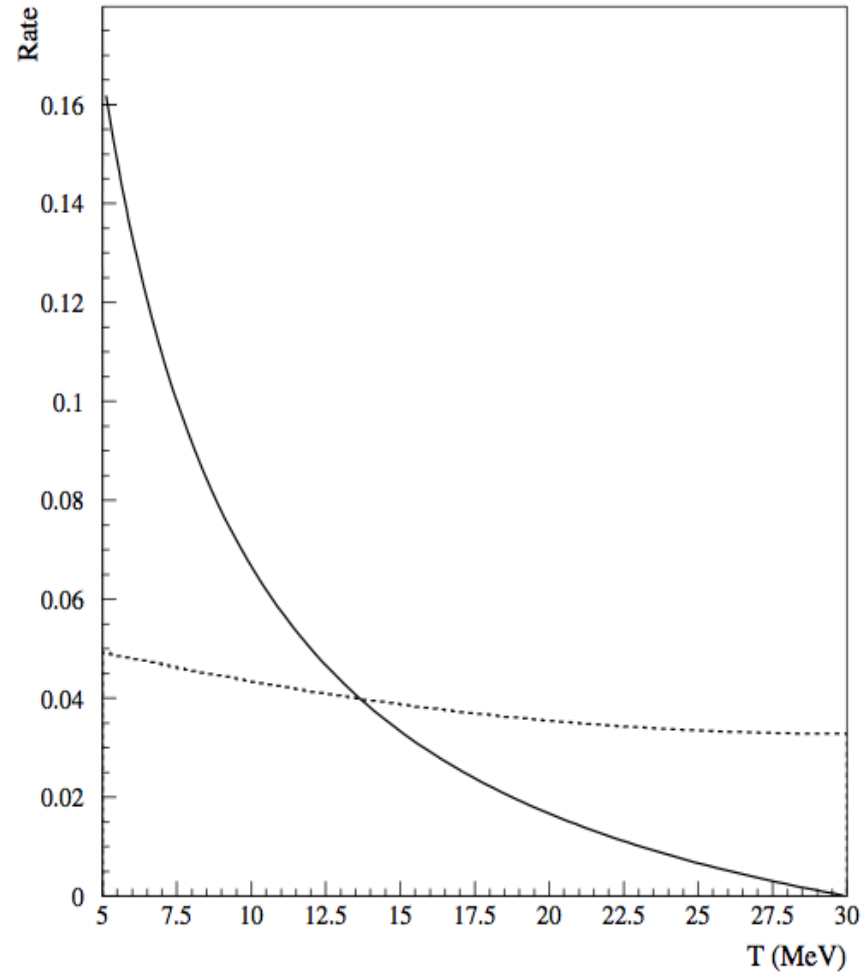
ν_{μ} magnetic moment

- At low electron energy, T_e , the magnetic moment cross section is given by
$$d\sigma^M/dT_e = f^2 * 2.5 * 10^{-25} \text{ cm}^2/T_e$$
 for $T_e \ll E_v$
where $f =$ magnetic moment in electron Bohr magnetons
- Because of $1/T_e$ factor want to look at as low an electron energy as possible

Standard model cross section

- For ν_μ
 $d\sigma^{\text{SM}}/dT_e = 22 \cdot 10^{-46} \text{cm}^2/\text{MeV}$ for $T_e \ll E_\nu$
- In the SM the number of elastic scattering events at low T_e measures the total flux for a beam like the LBNE ν_μ beam up to calculable corrections. This assumes no magnetic moment contribution.

Electron Energy



Neutrino Magnetic Moment

- Ratio of magnetic to SM cross section for $T_e \ll E_\nu$ is given by
$$d\sigma^M/dT_e / d\sigma^{SM}/dT_e = 1.1 * 10^{20} f^2 / T_e$$
- So event rates and ratio at low T_e depend on total neutrino flux. The shape is not important as long as $T_e \ll E_\nu$

Neutrino Magnetic Moment

- Experiments at high intensity proton sources might reach $f = 10^{-10}$.
- At $T_e = 15$ MeV and $f = 10^{-10}$ have $d\sigma^M/dT_e / d\sigma^{SM}/dT_e = 0.07$
- MiniBooNE set limit $< 10 \cdot 10^{-10}$ Bohr Magnetons for $5 \cdot 10^{20}$ POT. The number of electron scattering events in the final sample was low (~ 5 events) because ~ 500 m downstream of target.

Neutrino Magnetic Moment- miniBooNE

- Selected electrons $T_e > 15$ MeV.
- Cosmic ray background below 50 MeV was handled by rejecting past time muons.
- Below 15 MeV have large Boron 12 beta decay background in carbon target.
- Use angle cut and electron ID

Magnetic Moment

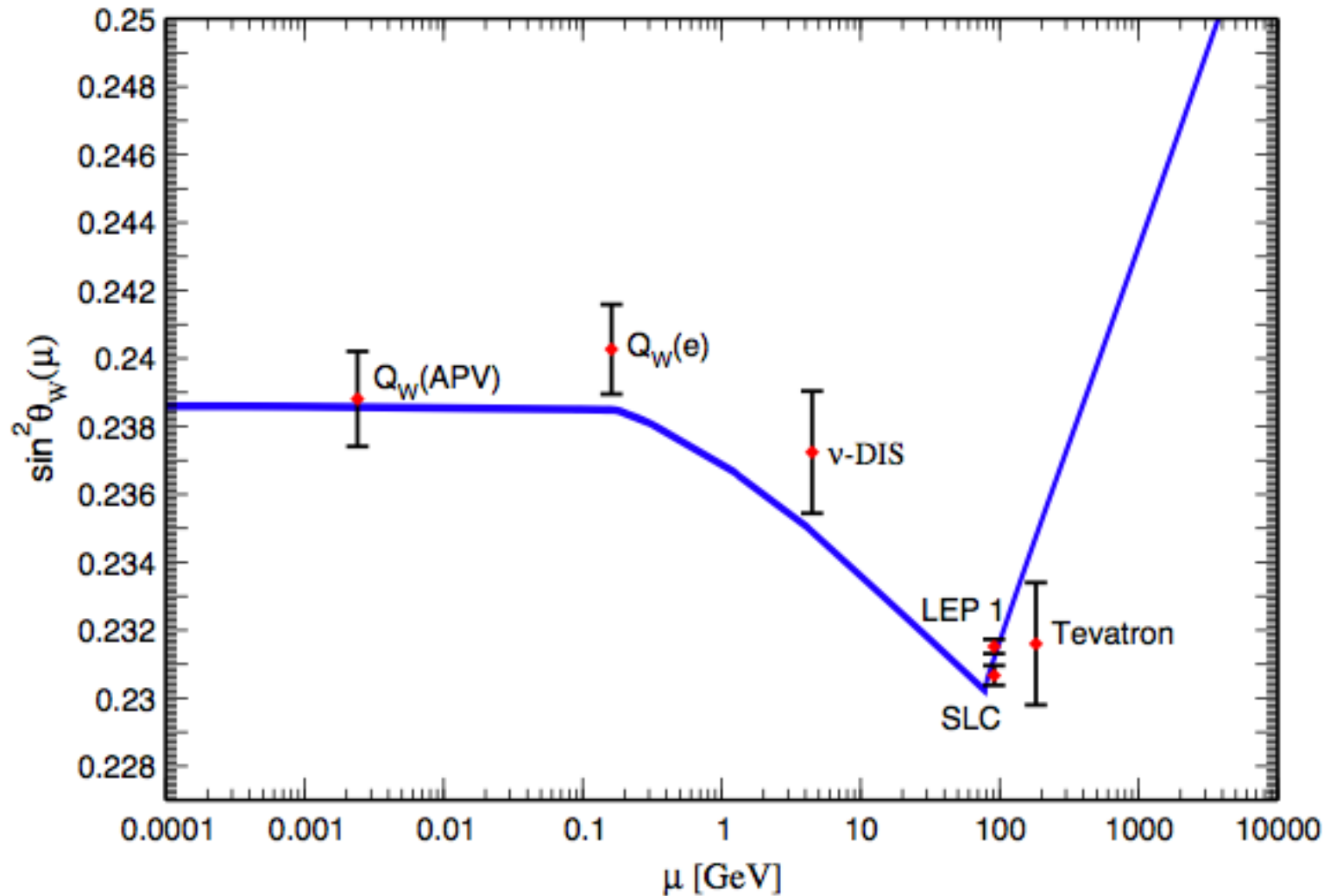
Future Experiments

- Need to do better than MiniBooNE.
- Cosmic ray background drops compared to signal with higher neutrino flux, better beam spill or location underground.
- Beam background is very detector specific. It may be limiting factor in sensitivity. Angular resolution very important.
- Could also reach $f = 10^{-10}$ with stopped π^+ neutrino beam

Weak mixing angle at low Q^2

- Running with Q^2 of $\sin^2\theta_w$.
- Present low Q^2 limit $\sim 0.5\%$ from parity violation asymmetry experiment.
- LBNE near detector report- get similarly sensitivity at fairly low Q^2 .

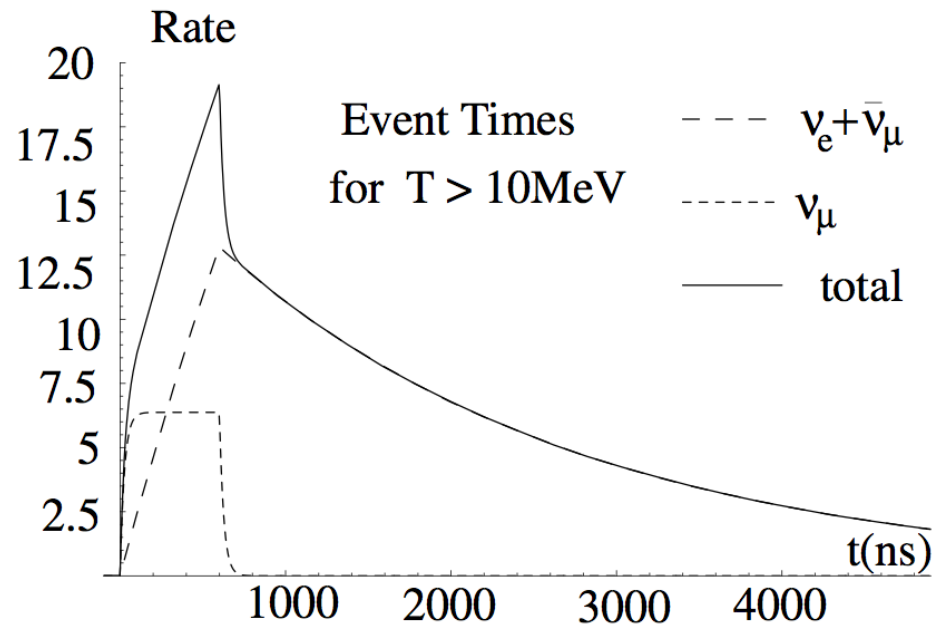
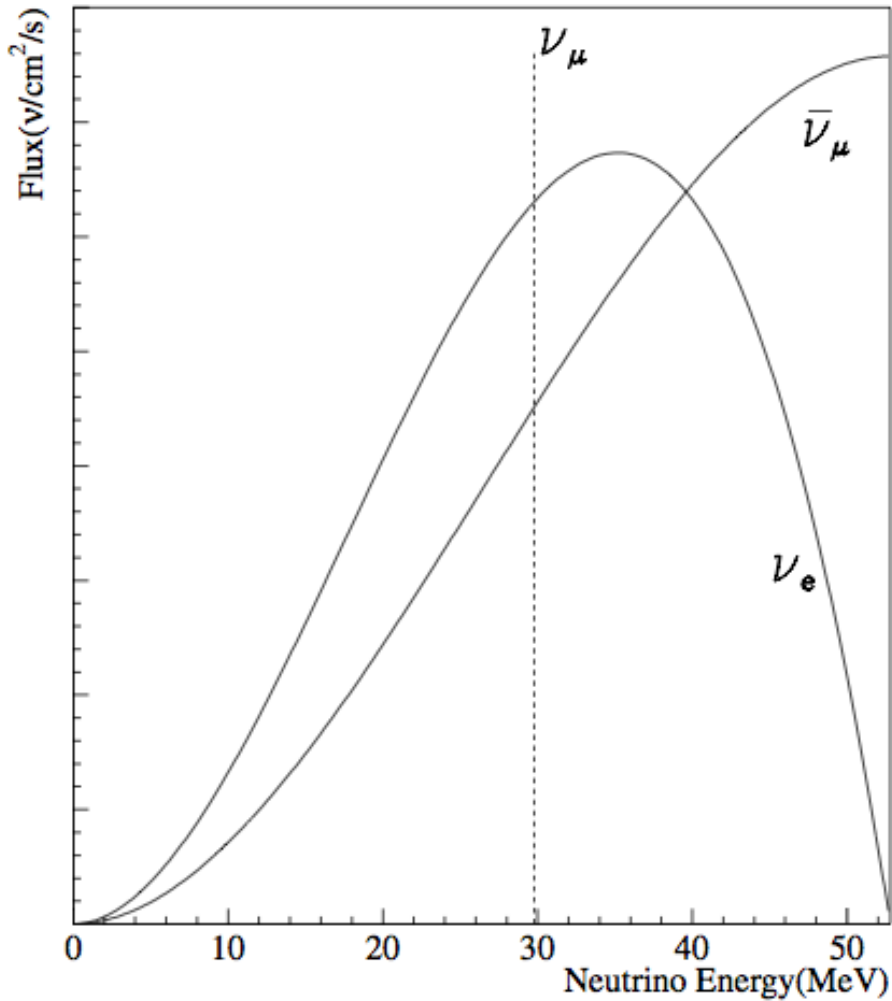
PDG 2010



Stopped Pion Source

- Could get another low Q^2 measurement with a stopped pion neutrino source.
- Use high Z , high density target to minimize π^- decay in flight background.
- Put detector ~ 90 degrees to minimize μ^- decay in flight background.

Stopped Pion Beam



Weak mixing angle at low Q^2

- Get mostly ν_e and $\bar{\nu}_\mu$ from μ^+ decay with 2.2 μs lifetime and ν_μ from π^+ decay with 26 ns lifetime.
- The ratio $R = \sigma(\nu_\mu e) / [\sigma(\nu_e e) + \sigma(\bar{\nu}_\mu e)]$
 $= (0.75 - 3\sin^2\theta_w + 4\sin^4\theta_w) / (1 + 2\sin^2\theta_w + 8\sin^4\theta_w)$

Weak mixing angle at low Q^2

- See J.Phys. G. Nucl. Part.Phys. 29(2003) 2647-2664 for design for SNS
- Measure R to 2-3% for 600 ns spill and $\sin^2\theta_w$ to 1 to 2% in oil or water detector.
- Might do factor of 2 better with 200 ns spill.

Weak mixing angle at low Q^2

- A stopped pion experiment is also of interest because:
- The elastic scattering of ν_e on e is one of very few reactions that has a large destructive interference between CC and NC channels
- Can set limits on the electron-neutrino charge radius.

Flux Calibration

- For ν_μ $d\sigma^{\text{SM}}/dT_e =$
 $12.8 \cdot 10^{-46} \text{cm}^2/\text{MeV} [1 + 0.73(1 - T_e/E_\nu)^2]$
- $\text{Cos}\theta = (1 + m_e/E_\nu) / (1 + 2m_e/T_e)^{0.5}$
- For a given neutrino flux shape the shape of the electron energy spectrum is completely determined.

Flux Calibration

- In principle, a perfectly known electron spectrum determines the neutrino spectrum.
- But in practice this does not work in a simple way.
- Electrons of a given energy have comparably contributions from all neutrinos with $E_\nu > T_e$.

Flux Calibration

- This is particularly bad at the lower end of the spectrum where the difference between the number of events in adjacent T_e bins provides information on the neutrino flux in that region but with huge statistical errors.
- Extreme case: the lowest energy electrons measure the total flux as discussed before. Have approximately equal contributions from the full spectrum.

Flux Calibration

- Do much better on the upper part of the neutrino spectrum.
- For example may be able to measure the high energy tail of LBNE fairly well if have clean elastic scattering sample.

Flux Calibration

- Need to include radiation corrections.
- Need to include information on ν_e contribution and wrong sign contribution.
- The number of ν_e from muon decay are tightly constrained by the number of ν_μ from π^+ decay. Need to include ν_e from kaon decay.
- For anti-neutrino running have a large wrong sign contribution.

Wrong Sign Contribution

- Want a good determination of wrong sign flux.
- For reference- MiniBooNE determined WS flux to 13% of its value using two methods: $CC\pi^+$ and μ^- capture.
- An experiment like LBNE could determine the WS flux in the near detector much better in several different ways.
- This is a discussion for another time.

Flux Shape

- Here is some guess work on how one might better determine the flux shape using the electron energy shape.
- Start with the simulated neutrino flux in energy bins with flux shape error matrix. The simulated spectrum will have strong correlations between adjacent energy bins.

Flux shape

- Do a fit to the electron energy distribution varying the flux shape as allowed by the flux shape correlated error matrix.
- The T_e distribution will help determine the relative flux for E_ν bins far apart.
- As discussed before the T_e distribution will help more with the higher part of the E_ν spectrum than the lower part.

Conclusions

- Can improve limit on ν_μ magnetic moment significantly. Unlikely to detect a signal and so this can't be the main motivation for an experiment.
- Can measure $\sin^2\theta_W$ at low Q^2 with comparably sensitivity to present Parity violating asymmetry experiments.
- Can measure flux normalization to 1-2% and in conjunction with the simulated flux and its errors improve on the flux shape.