





ESSnuSB from source to target and plans for the future

NuFact 2022

PRESENTED BY NATALIA MILAS

2022-07-14

The ESSnuSB Project





Goal: to measure the leptonic charge-parity violating phase **How:** by generating a muon (anti-)neutrino super beam and look at the electron neutrino appearance in a large detector.

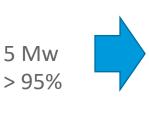
With 5 MW beam from the ESS linac we can use the second oscillation maximum where we are less limited by systematic uncertainties and thus have a higher sensitivity.



ESSnuSB High Level Parameters

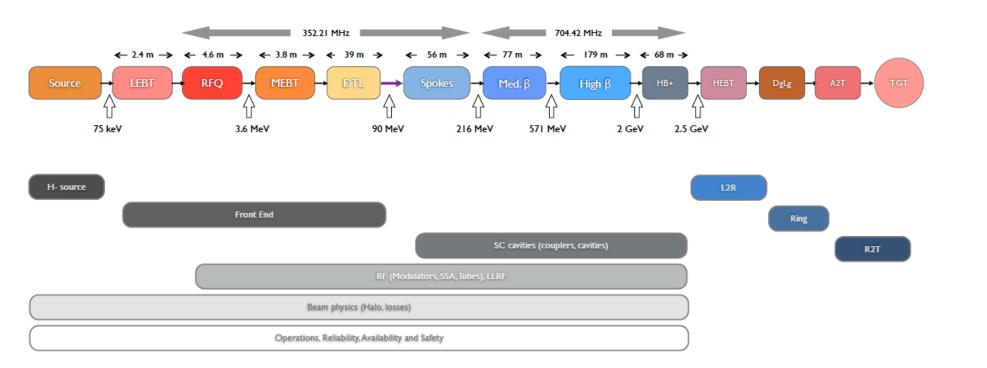


Drivers Design:High power5 MwHigh availability> 95%



Linac Parameters:Beam Energy2 GeVBeam Current62.5 mARep. Rate14 HzPulse length2.86 msLosses< 1 W/m</td>Ionsp

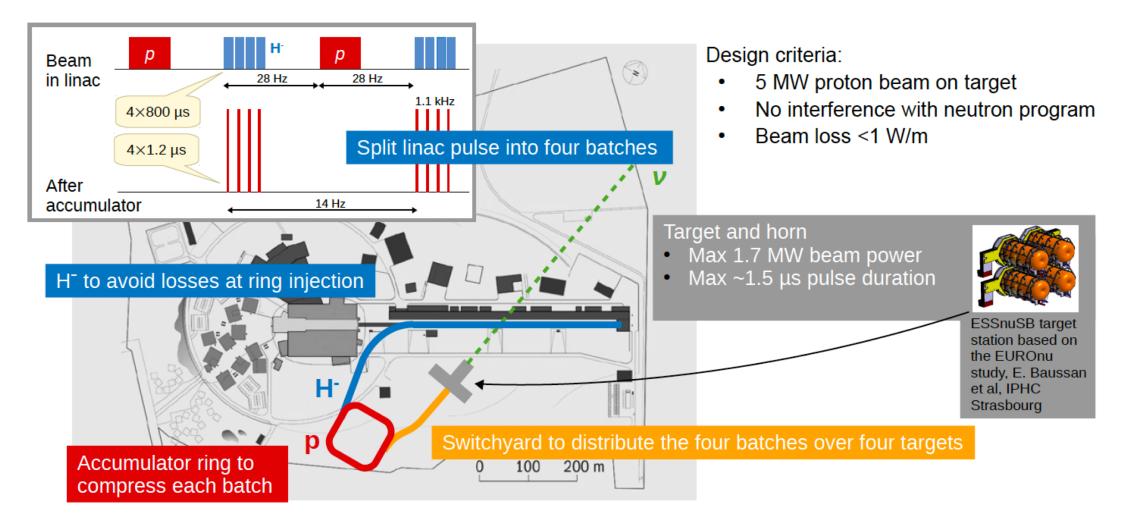
ESSnuSB beam:Beam Energy2.5 GeVBeam Current62 mA (50 mA)Rep. Rate14 HzPulse length< 3.5 ms</td>Losses< 1 W/m</td>IonsH⁻



ESSnuSB High Level Parameters



Constraints



EssnuSB Layout

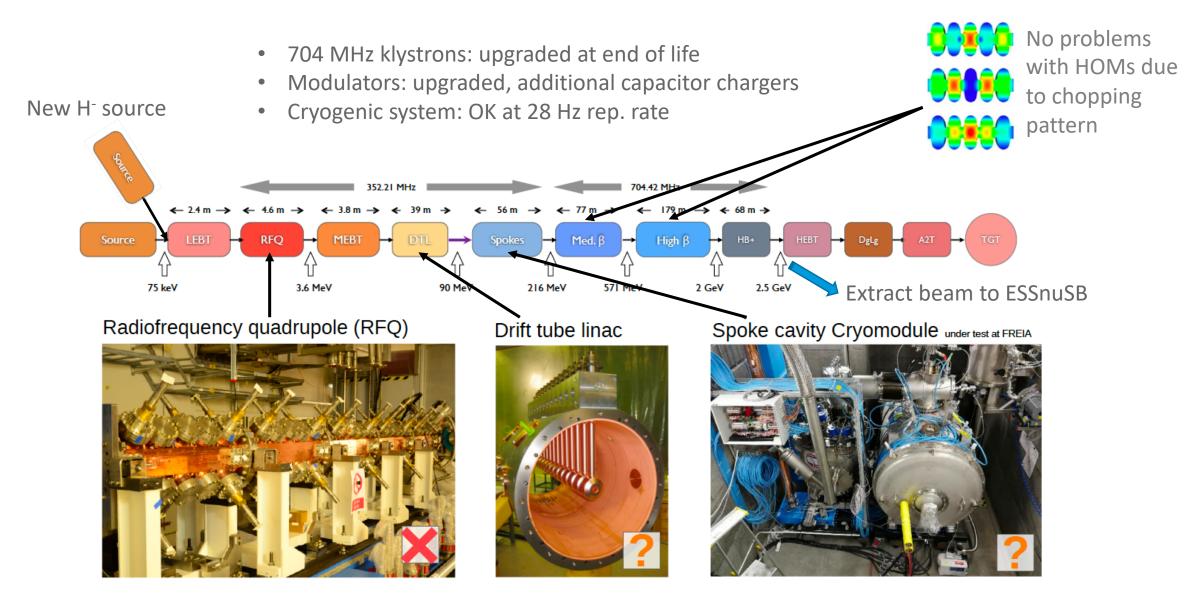




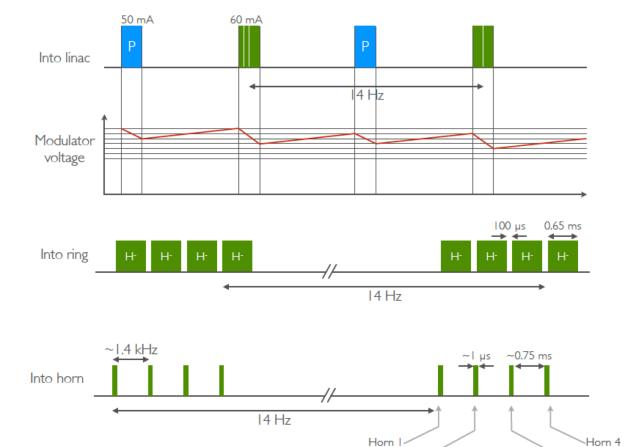
Rasmus Johansson and Nick Gazis

Linac Design





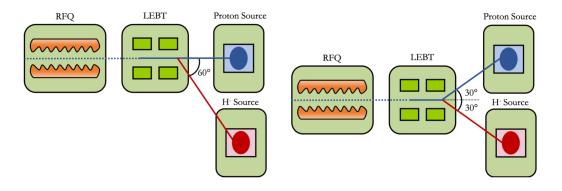
A New Front-End and Pulsing Scheme



Horn 2

Horn 3

Possibility of merging the two bean before the RFQ

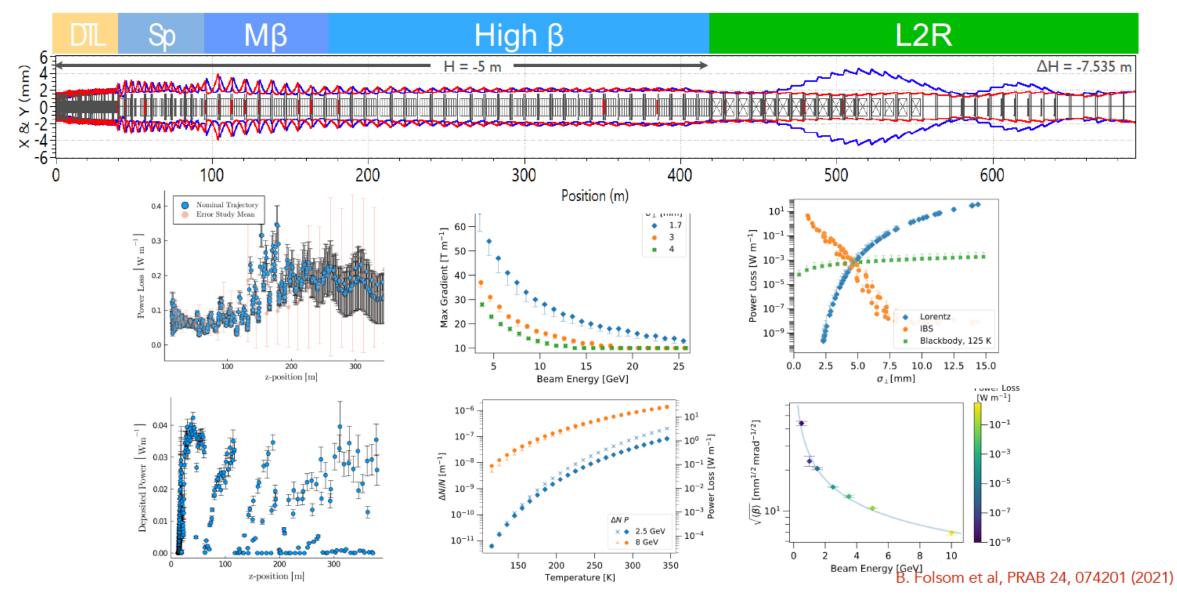


- Smaller footprint and easier to fit in the tune
- More difficult dynamics for both beam with the 30 deg option









2022-07-14 PRESENTATION TITLE/FOOTER

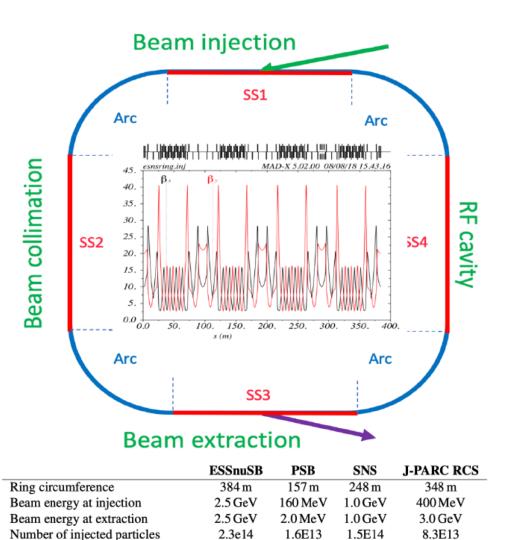
Accumulator Ring Design

Main challenges

- Beam loss control due to very high beam powe
- Space-charge tune shift due to very high beam intensity
- Instabilities (e-p instability)

Lattice design

- Developed by Horst Schönauer at CERN
- Circumference: 384 m
- 4-fold symmetry
- 4 straight sections (SS1~SS4) and 4 arc sections (Arc)
- Fixed injection chicane and fast programmable bump for injection painting







Accumulator Ring: Injection

Foil stripping and laser stripping

• Foil stripping: widely used in proton synchrotrons or accumulators, very challenging due to high power

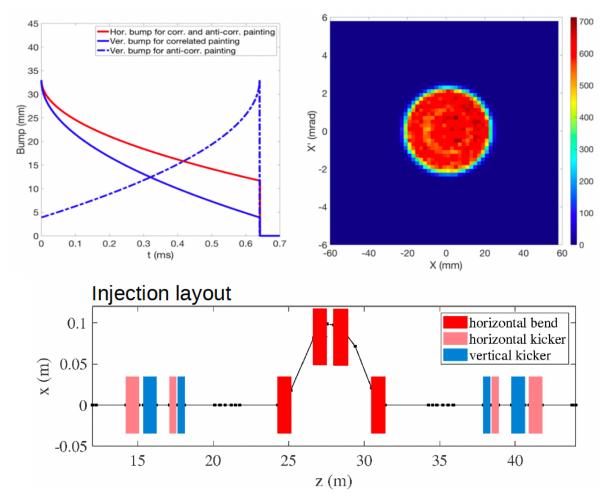
• Laser stripping: a promising alternative method

Painting

- Mitigate space charge issue
- Mitigate foil temperature issue

Foil temperature issue mitigation

- Mismatch injection
- Splitting the foil along beam direction
- Moving injection point



Y. Zou, Uppsala University



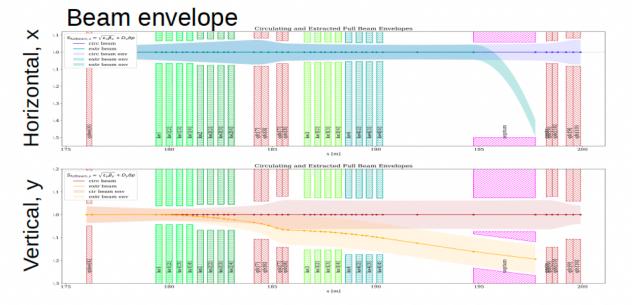
Accumulator Ring: Extraction

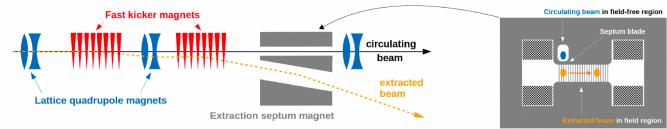
Beam extraction

- Single-turn extraction system is designed to extract the full beam in a single turn after accumulation
- Fast magnets (kickers) to extract the beam vertically out of the ring during the extraction gap of 100-130 ns
- Horizontal deflector (septum) to deflect the beam by 16.8 deg to the start of the extraction line

Challenges

- Loss-free extraction \Rightarrow optimize aperture sizes
- Rise-time of kickers ⇒ aperture size, B-fie technology

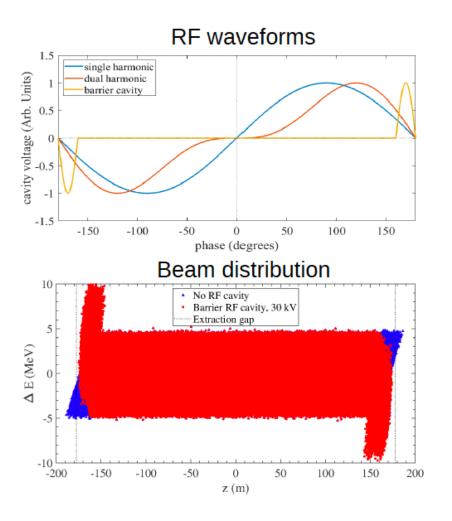




A. Alekou and I. Efthymiopoulos, CERN



Accumulator RF and Switchyard Design

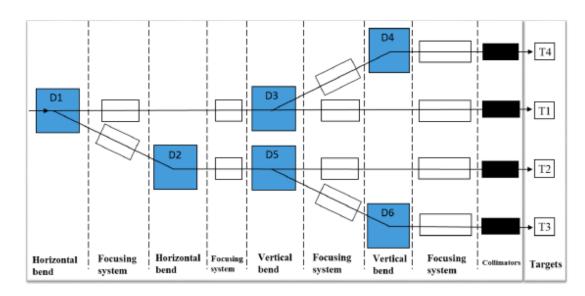


Y. Zou, Uppsala University

RF cavity

- RF cavity to keep the extraction gap clean while keep energy spread small
- Barrier bucket is chosen due to its very small leakage risk and small energy spread (±0.15%)

• Aperture optimized to make rise-time requirement easier to reach



E. Bouquerel, IPHC Strasbourg

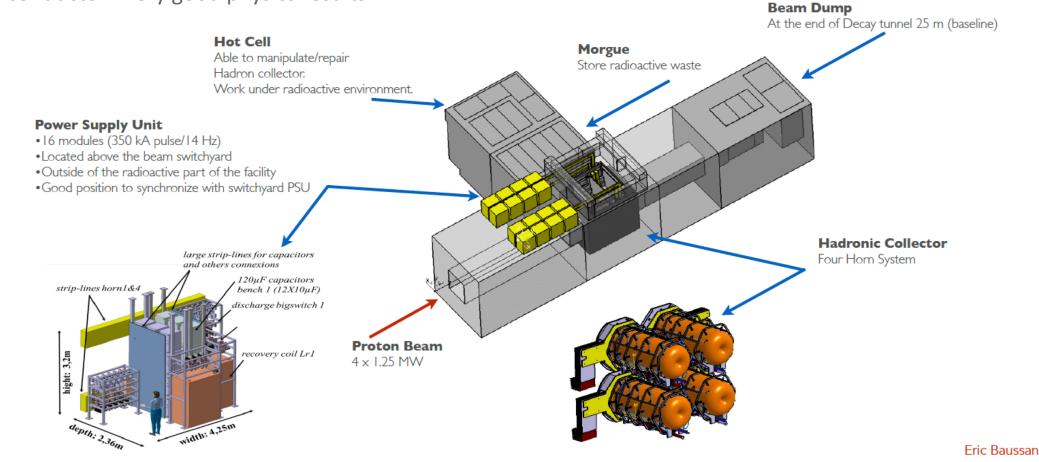
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Target Design

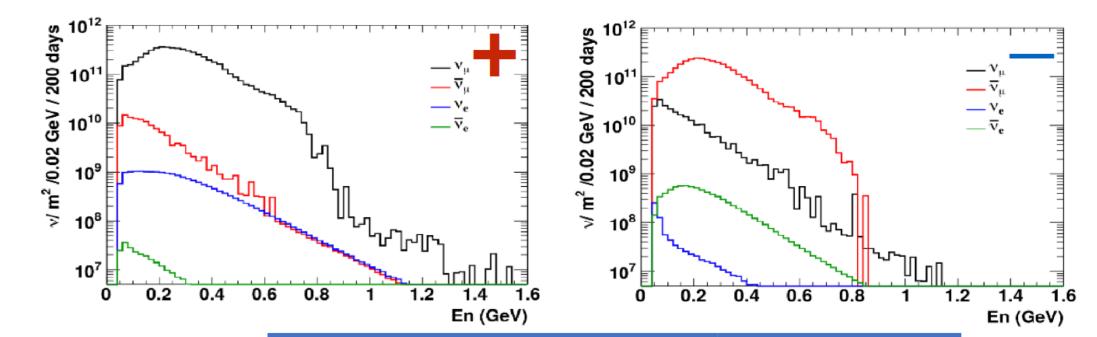


Packed Bed Target: Power 1.25 – 1.6 MW. Potential heat removal rates at the hundreds of kW level. Helium cooling (10 bars). Separated from the horn

Focusing System: 4-horn+4 target system to accommodate the MW power scale. Packed bed target integrated into the inner conductor : very good physics results.



Expected Neutrino Flux



	Positive Polarity		Negative Polarity	
	N (1E10 1/m ²)	%	N (1E10 1/m ²)	%
Muon neutrino	583	98	23.9	6.55
Muon anti neutrino	12.8	2.1	340	93.2
Electron neutrino	1.93	0.3	0.08	0.02
Electron anti neutrino	0.03	0.01	0.78	0.21

Eric Baussan

ESS NEUTRINO SUPER BEAM



Detectors

Two water Cherenkov detectors

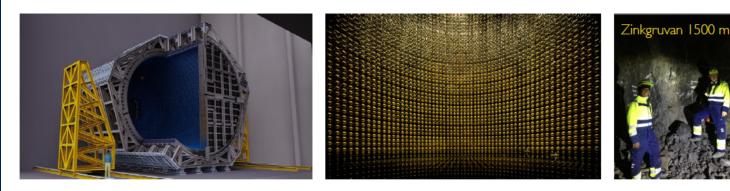
• Far detector (370 km from target)

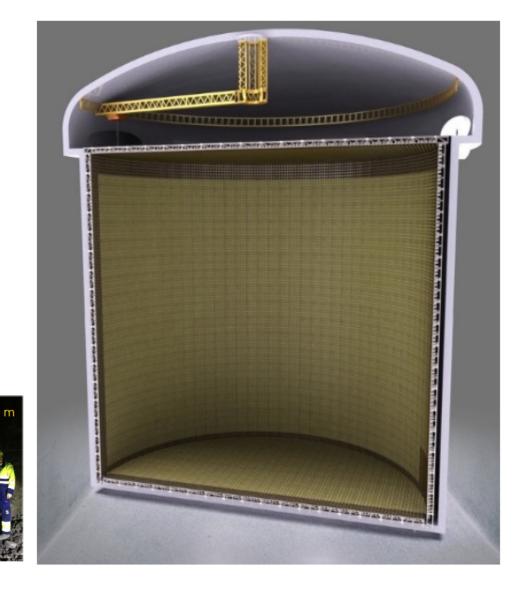
 – Pi x 372 x 74 m3 >0.5 Megaton of ultra pure water as the far detector . That is 10 times the volume of the super-Kamiokande.

- 540k m3 total fiducial volume
- 100k 20" PMTs gives 40% coverage
- 1500 meter deep underground
- To decrease the noise and background from cosmic

radiation

- Near detector (250 m from target)
 - Pi x 4.722 x 11 m3





Next Steps



Further important studies needed before start ESSnuSB construction.

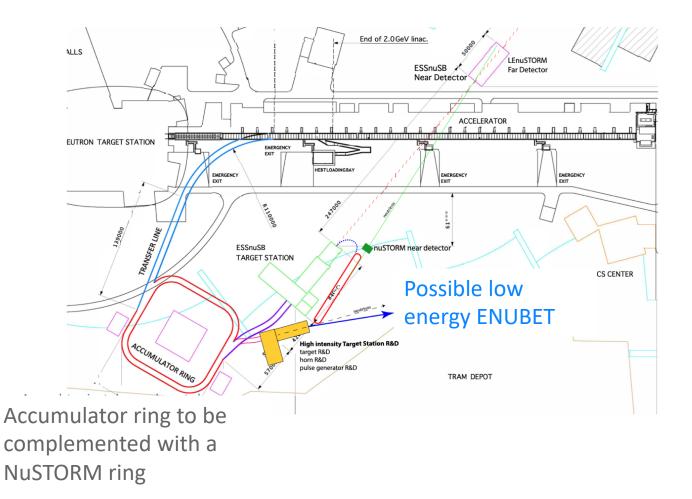
- supplementary studies required for the present proposal (according to present WP conclusions),
- site specific civil underground engineering: at ESS-site and far detector site,

Preparation of the R&D phase,

- detailed studies of the safety and licensing requirements needed for the later approval by the authorities as well as how to take environmental protective measures into account
- elaborated radiation protection issues, material and waste optimization,
- optimized safety operations (construction and maintenance) including the near and far detectors.
- provide energy and material saving/reuse solutions for the ESSnuSB installations on the ESS site and the Far Detector.

Next Steps

Proposed Scenario for the period 2022-25





Low Energy nu STORM

- ~0.5 GeV muons
- sterile neutrinos
- cross-section measurements

ENUBET

• cross-section measurements

Further studies on:

- Civil engeneering
- Target
- Accelerator improvements (fine tuning)

Final Goal:

 Multi purpose facility with great potential for material and particle physics studies!





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Thank you! Questions?