Current Understanding of the Earth's Core

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Outline

- Density Structure
- Core composition
 - Light elements
 - Radiogenic elements
 - Noble gases



Density structure of the Earth

- Derived from measurements of seismic wave speeds (v_p, v_s) and/or normal mode oscillation frequencies
- Adams-Williamson equation

$$\frac{d\rho}{dr} = -\frac{\rho(r)g(r)}{v_p^2 - \frac{4}{3}v_s^2}$$

Density Structure



Density Resolution



Nimmo (2022), Oxford Research Encyclopedia, https://doi.org/10.1093/acrefore/9780190647926.013.204



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Layers in the core?

- "F-layer" at the bottom of the outer core
 - 150 km thick, stably stratified?
 - Density change ~0.5 g/cc across layer
 - Could be a gradient in O (Gubbins et al. 2008)
- Stable layer at the top of the outer core?
 - Existence is disputed (Buffett et al. 2016, Helffrich & Kaneshima 2010, Irving et al. 2018)
 - -~300 km thick, light by ~0.6 g/cc
 - Could be a result of accumulation of light elements (e.g. O,Si) or be a non-convecting part of the outer core

Core Light Elements

- Based on cosmochemical models (McDonough & Sun 1995) the core contains Fe with about 5.5wt% Ni
- But the measured density of the core is too light by 7-10% to be a pure Fe-Ni alloy
- So the Earth contains one or more light elements
- Usual candidates are H,C,O,Si,S

Atomic Number vs Atomic Weight



Which light element is present?

- Two ways of addressing this question:
 - Compare densities and seismic velocities [nonunique]
 - Use partitioning models [assumption-dependent]

Density jump at the ICB

- The inner core is denser than the outer core by 0.36-0.62 g/cc after taking the phase change into account (Masters & Gubbins 2003)
- So the inner core contains fewer light elements than the outer core
- Oxygen is excluded from the solid core, whereas S and Si are not (Alfe et al. 2002, Ozawa et al. 2010)
- So the outer core must contain O

Densities and seismic velocities

- In principle, different light elements may produce different densities & velocities
- Measurements at correct P,T conditions are hard
- Sometimes *ab initio* simulations are used instead
- No single element can satisfy the observations
- There is little agreement, although oxygen is commonly found



Partitioning models

- "Partitioning" means the extent to which an element goes into the core rather than the mantle on melting
- Experiments at Earth-relevant conditions are difficult (e.g. Fischer et al. 2020 vs. Blanchard et al. 2022)
- "Partition coefficient" depends on P,T,fO₂
- Assumptions of initial fO₂ and composition have a big effect on the conclusions
- There is little agreement on how fO₂ has evolved with time (e.g. Badro et al. 2015, Rubie et al. 2015)
- But most models favour Si and O as the dominant core light elements

What about S,C and H?

- C core content ~0.1 wt% (Fischer et al. 2020, Blanchard et al. 2022)
- S is a possibility from velocity/density measurements; S partitioning experiments suggest
 <2 wt% in core (Suer et al. 2017)
- Both C and S are strongly depleted in most precursor Earth materials
- H could be as high as 0.6 wt% (Tagawa et al. 2021)
- But would require cores to form and acquire gas before the nebular dispersed (<5 Myr) – unlikely but not impossible

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Radiogenic Elements

- The Earth's heat engine is powered by long-lived radiogenic elements: K,U,Th
- U and Th are thought to reside exclusively in the mantle ("lithophile")
- But some K might be found in the core
- Radioactive K can significantly alter the thermal and magnetic history of the core (e.g. Nimmo et al. 2004)
- Partitioning data (Blanchard et al. 2017, Xiong et al. 2018) suggest only ~30ppm K in the core, too small to be important

Noble gases

- ~5x10⁻¹³ concentration for ³He (Olson & Sharp 2022)
- ~2x10⁻¹³ for ³He and ~7x10⁻¹² for ²²Ne (Bouhifd et al. 2020)
- ~ 3x10-9 for ⁴⁰Ar (from radioactive decay) (Wang et al. 2022)
- These are very uncertain numbers
- But the concentrations are very low!

Summary

- Density Structure
 - Well-known, except for possible layers at top and bottom of outer core
- Core composition
 - Light elements: *O, probably Si, possibly S*
 - Radiogenic elements: *K at the tens of ppm level*
 - Noble gases: Concentrations very low (~10⁻¹³-10⁻⁹)