

# Calibrating for Precision Calorimetry in LArTPCs at ICARUS and SBN

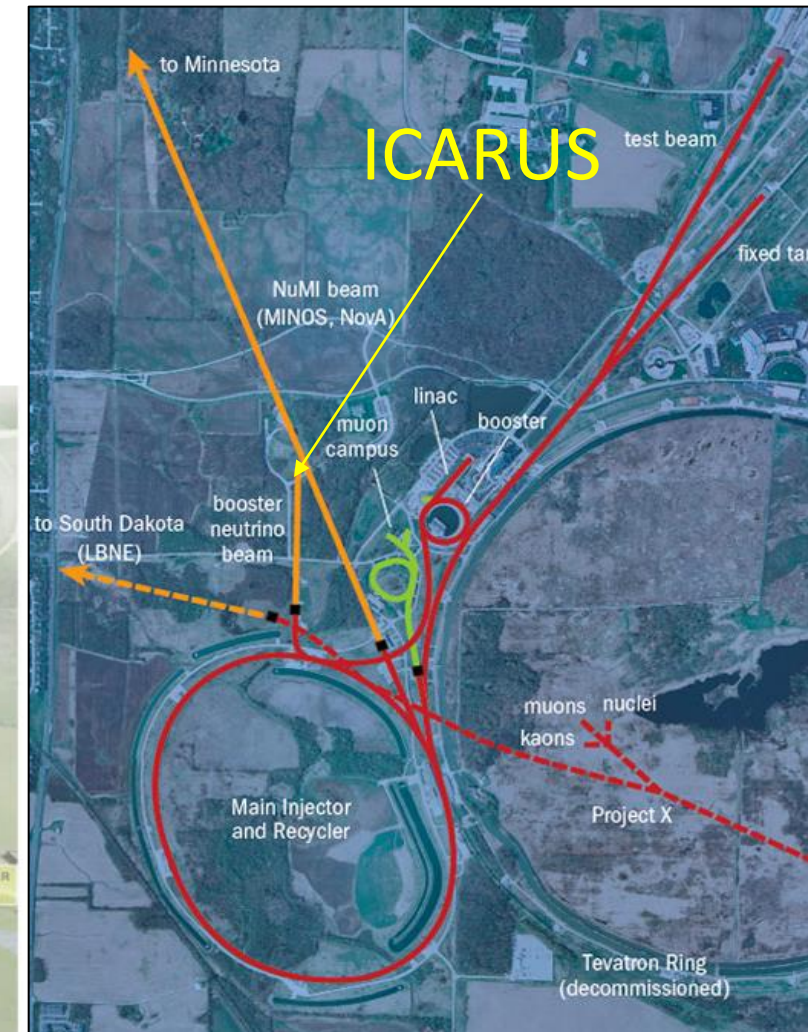
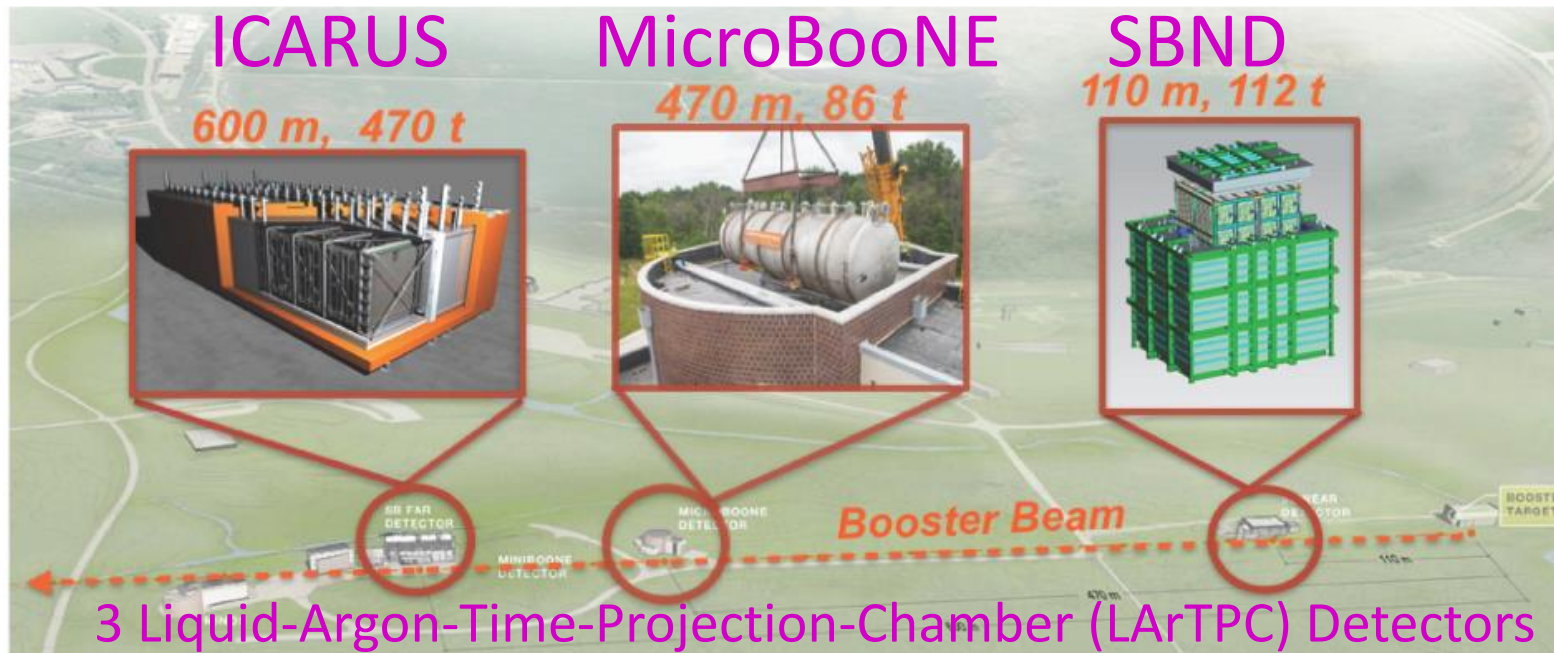
GRAY PUTNAM

UNIVERSITY OF CHICAGO



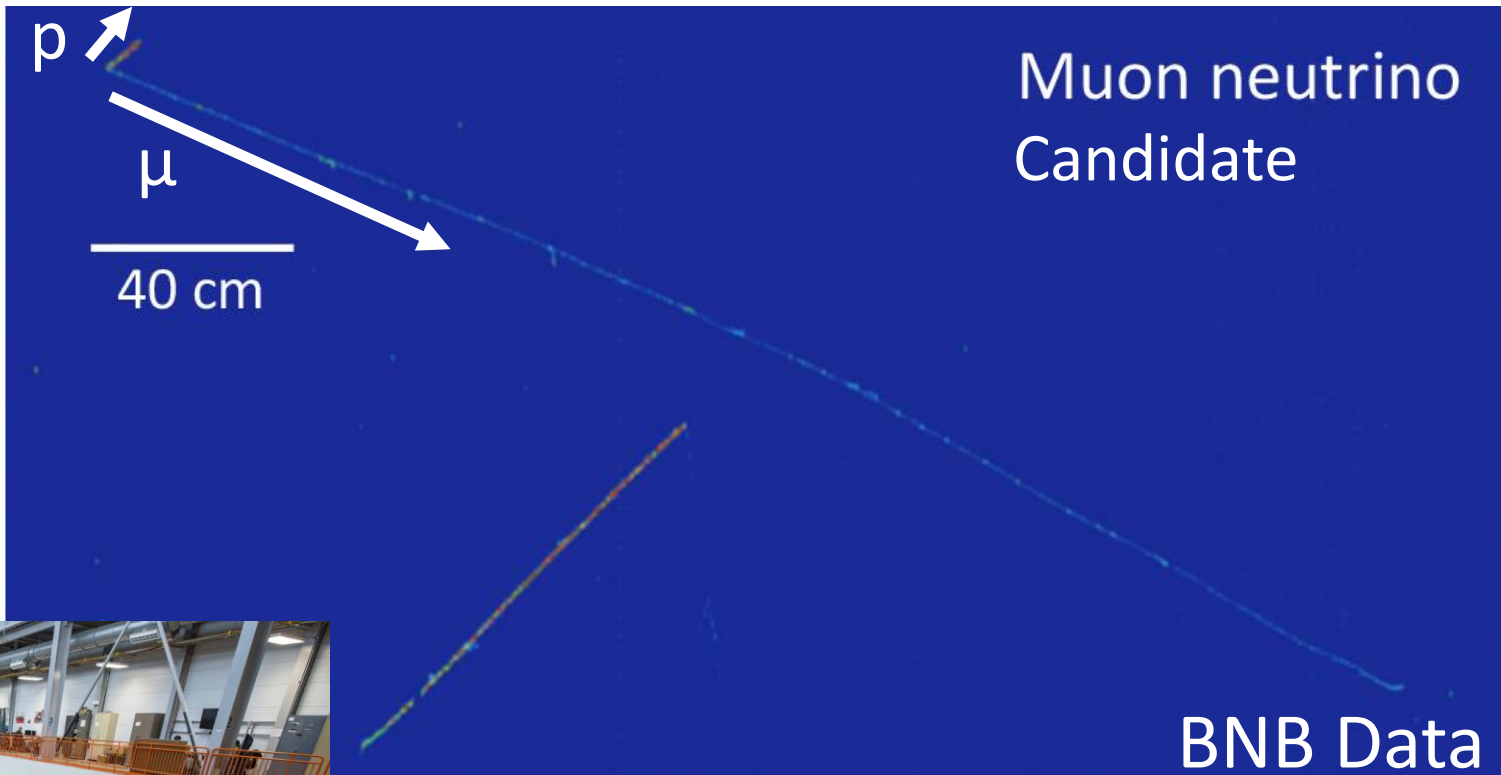
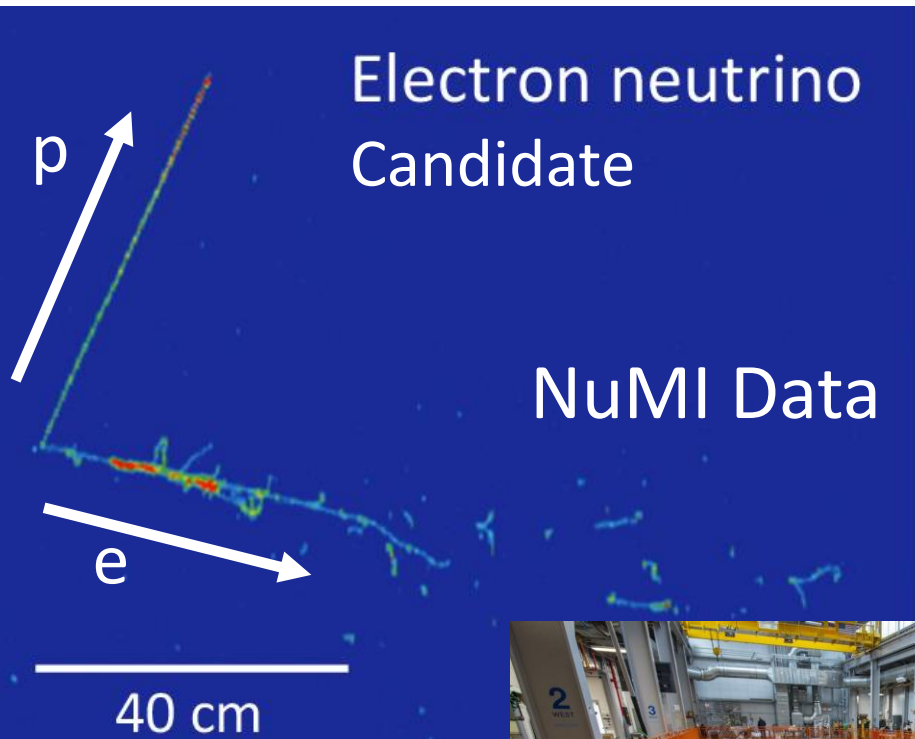
# ICARUS and SBN at Fermilab

- ICARUS is the Far Detector in the Short Baseline Neutrino (SBN) Program
- SBN program physics:
  - eV-scale sterile neutrino search
  - GeV-scale neutrino cross section measurements
  - Single Detector BSM physics searches





# Neutrino Images from the ICARUS LArTPC



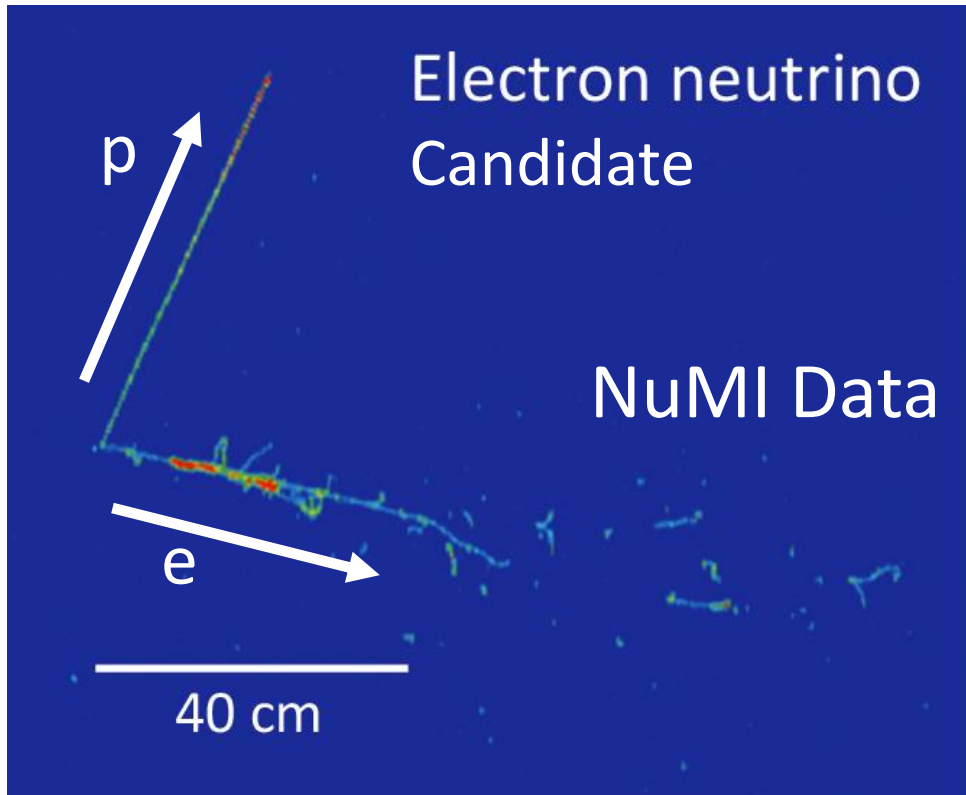
<https://news.fnal.gov/2021/05/icarus-gets-ready-to-fly>

ICARUS Detector  
at FNAL:



Each image is from one TPC  
inside each cryostat

# ICARUS: A Liquid Argon Time Projection Chamber (LArTPC)



A LArTPC is a calorimeter for measuring charged particles produced in  $\nu$  interactions

What measurements can you do with charge calorimetry in a LArTPC?

Easy

$\mu/p, e/\gamma$   
Separation

Electron  
showers

Low-energy  
particles from  $\nu$ s

Low-energy  
electron tracks

Very low energy  
electron blips

Hard

# Using Charge at DUNE: Low Energy Electrons

## Supernova neutrino burst detection with the deep underground neutrino experiment

DUNE Collaboration

### DUNE as the Next-Generation Solar Neutrino Experiment

Francesco Capozzi<sup>1,2,3,\*</sup> Shirley Weishi Li<sup>1,2,4,†</sup> Guanying Zhu<sup>1,2,‡</sup> and John F. Beacom<sup>1,2,5,§</sup>  
<sup>1</sup>*Center for Cosmology and AstroParticle Physics (CCAPP), Ohio State University, Columbus, Ohio 43210, USA*

<sup>2</sup>*Department of Physics, Ohio State University, Columbus, Ohio 43210, USA*

<sup>3</sup>*Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), 80805 München, Germany*

<sup>4</sup>*SLAC National Accelerator Laboratory, Menlo Park, California 94025, USA*

<sup>5</sup>*Department of Astronomy, Ohio State University, Columbus, Ohio 43210, USA*



(Received 4 September 2018; published 27 September 2019)

### Xenon-Doped Liquid Argon TPCs as a Neutrinoless Double Beta Decay Platform

A. Mastbaum,<sup>1</sup> F. Psihas,<sup>2</sup> and J. Zennaro<sup>2</sup>

<sup>1</sup>*Rutgers University, Piscataway, NJ, 08854, USA*

<sup>2</sup>*Fermi National Accelerator Laboratory (FNAL), Batavia, IL 60510, USA*

(Dated: March 29, 2022)

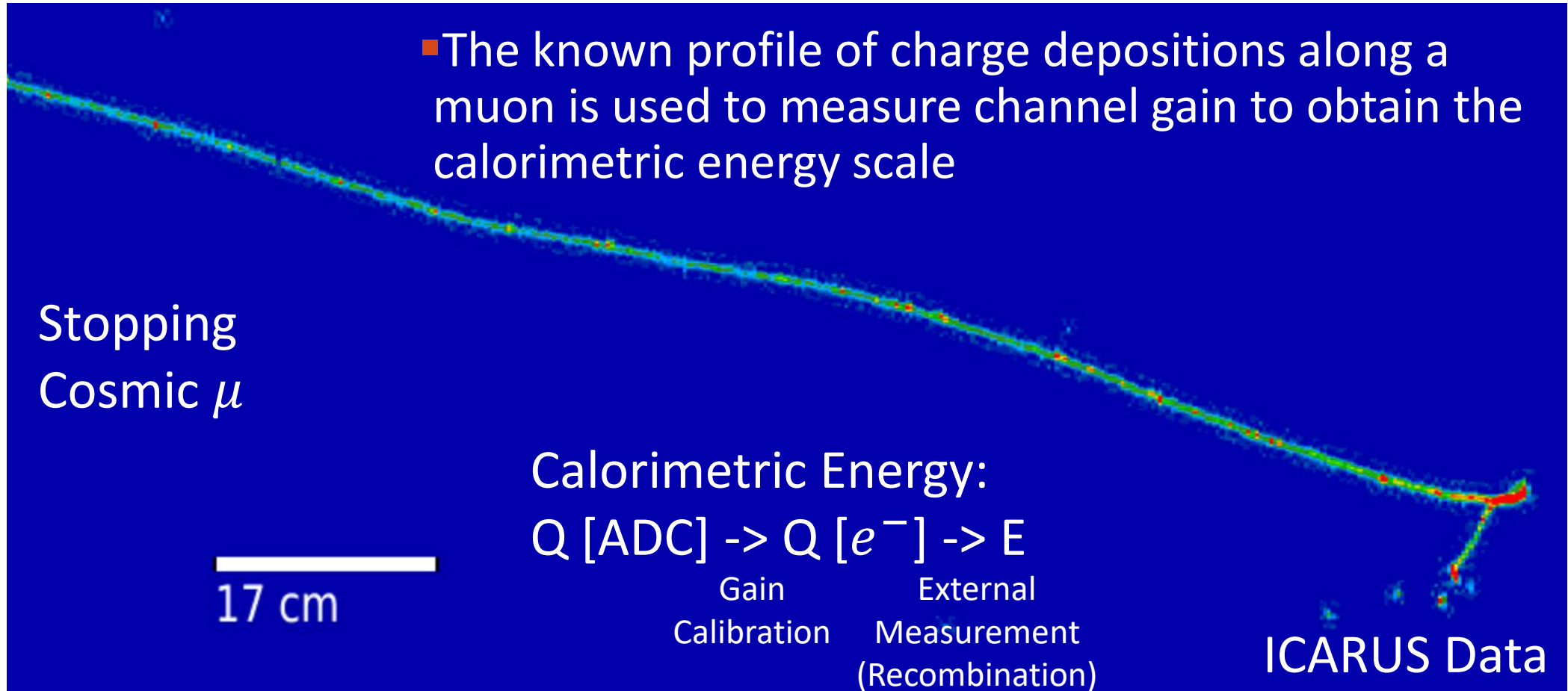
Physics goals for DUNE include a variety of signatures from low energy electrons, which would apply calorimetric energy measurements.

# How Well Can We Calibrate LArTPCs?

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# Cosmic Muons as a Standard Candle

- In LArTPC experiments, depositions from cosmic muons are used as a “standard candle” to calibrate the energy scale



# Accuracy of Energy Measurements

- In LArTPC experiments, depositions from cosmic muons are used as a “standard candle” to calibrate the energy scale
- The predicted ionization per length of a cosmic muon combines the Bethe-Bloch energy loss with a recombination model to map energy to charge

How well do we know the most-probable  $\frac{dQ}{dx}$  from a 1GeV muon deposition with 1ms of drift time?

Source	CV with Uncertainty	Percent Impact on $dQ/dx$
Recombination Modeling	$\alpha = 0.93 \pm 0.02,$ $\beta = 0.212 \pm 0.001$	3.8
Mean Excitation Energy ( $I_0$ )	$188 \pm 17$ eV	1.0
Transverse Diffusion ( $D_T$ )	$8.8 \pm 4.4$ cm <sup>2</sup> /s	1.0

ArgoNeuT Collab, JINST (2013)

Recombination model:  $\frac{dQ}{dx} = \frac{\ln \alpha + \frac{dE}{dx} \beta}{W_{ion} \beta}$

ICRU 37, plus uncertainty from GAr v. LAr

Extrapolation from longitudinal diffusion through Wannier relation



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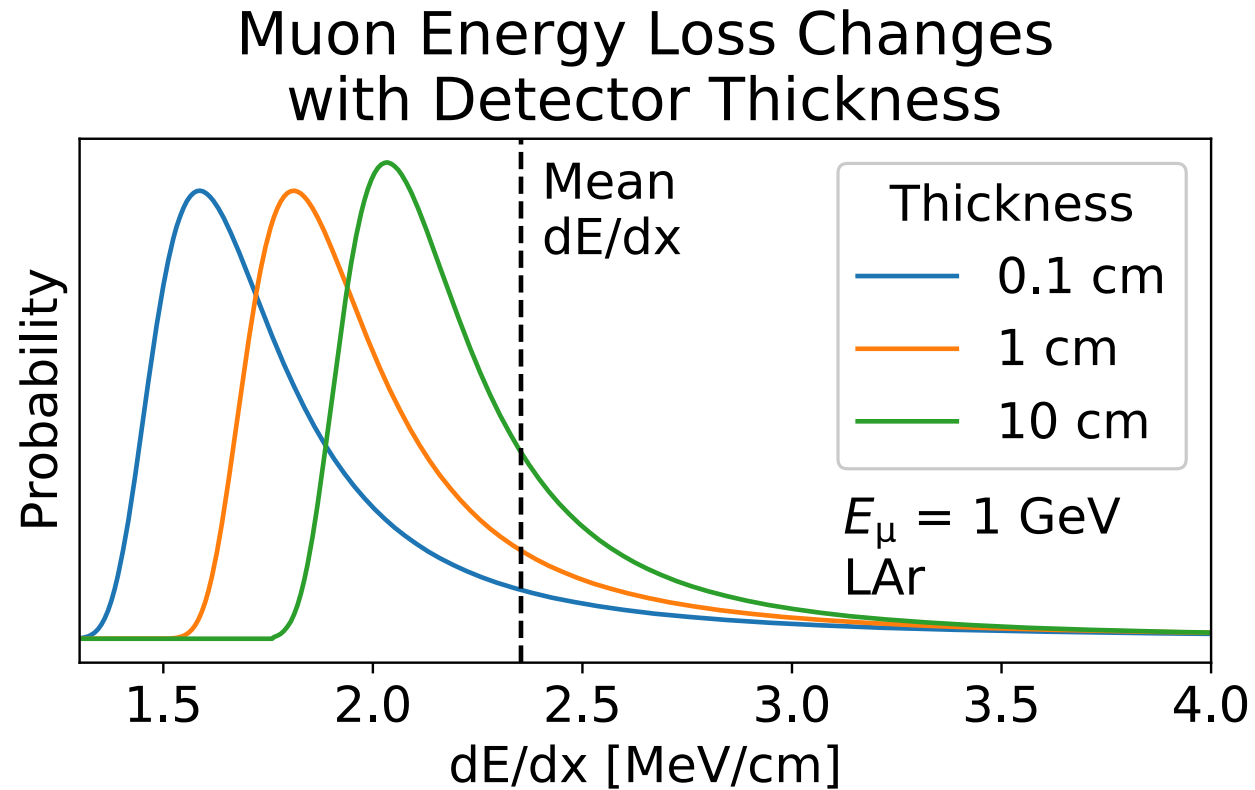
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ICRU 37, plus uncertainty from GAr v. LAr

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# The Landau Energy Loss Distribution Depends on Wire Thickness

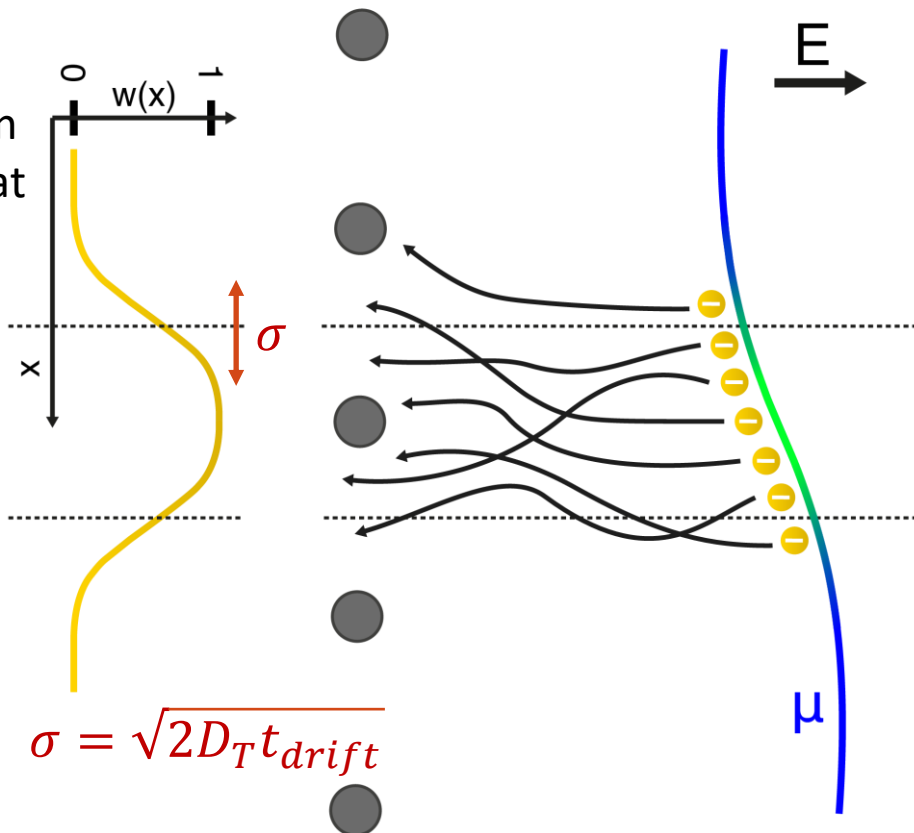
- The distribution of energy loss is a Landau distribution
- The peak of a Landau distribution has a dependence on the length of the particle observed by the wire
- As the thickness goes up, the most-probable-value (MPV) of energy loss goes up



# Diffusion Changes the Thickness!

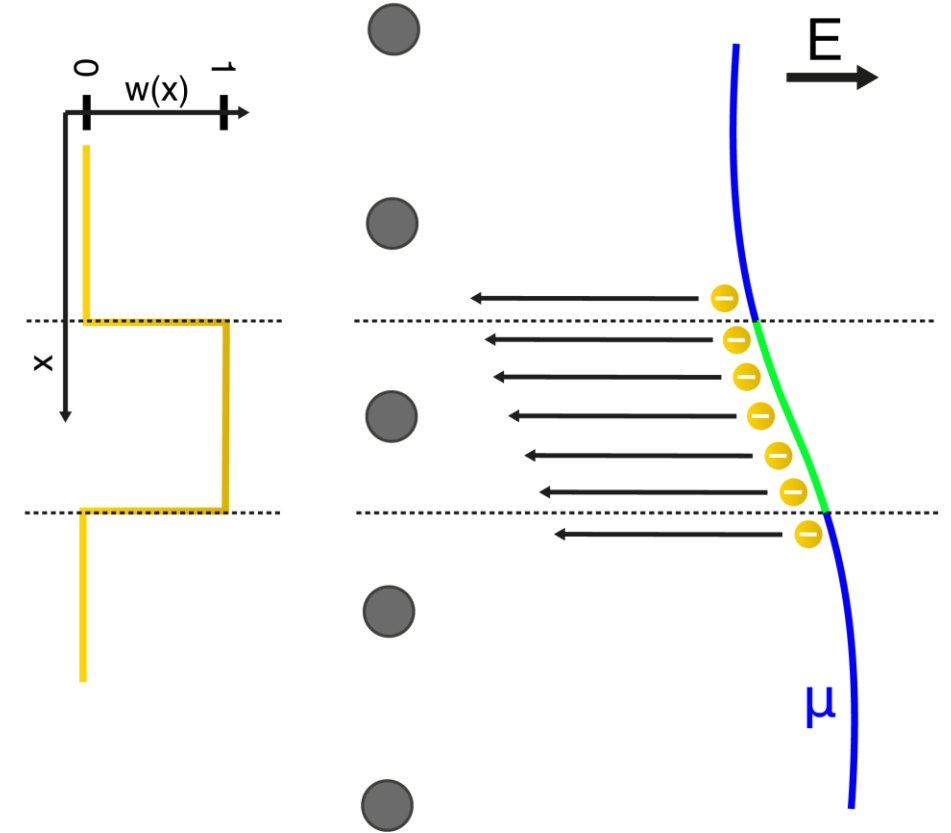
## WITH DIFFUSION

$w(x)$ : **weight function** which gives a weight to how much ionization charge a wire will see at each point along the muon trajectory.



$$\sigma = \sqrt{2D_T t_{drift}}$$

## WITHOUT DIFFUSION



- Diffusion transverse to the drift direction (and the wire direction) thickens the length of the muon that each wire is sensitive to – **this changes the MPV energy loss**

# Energy Scale Calibration at ICARUS

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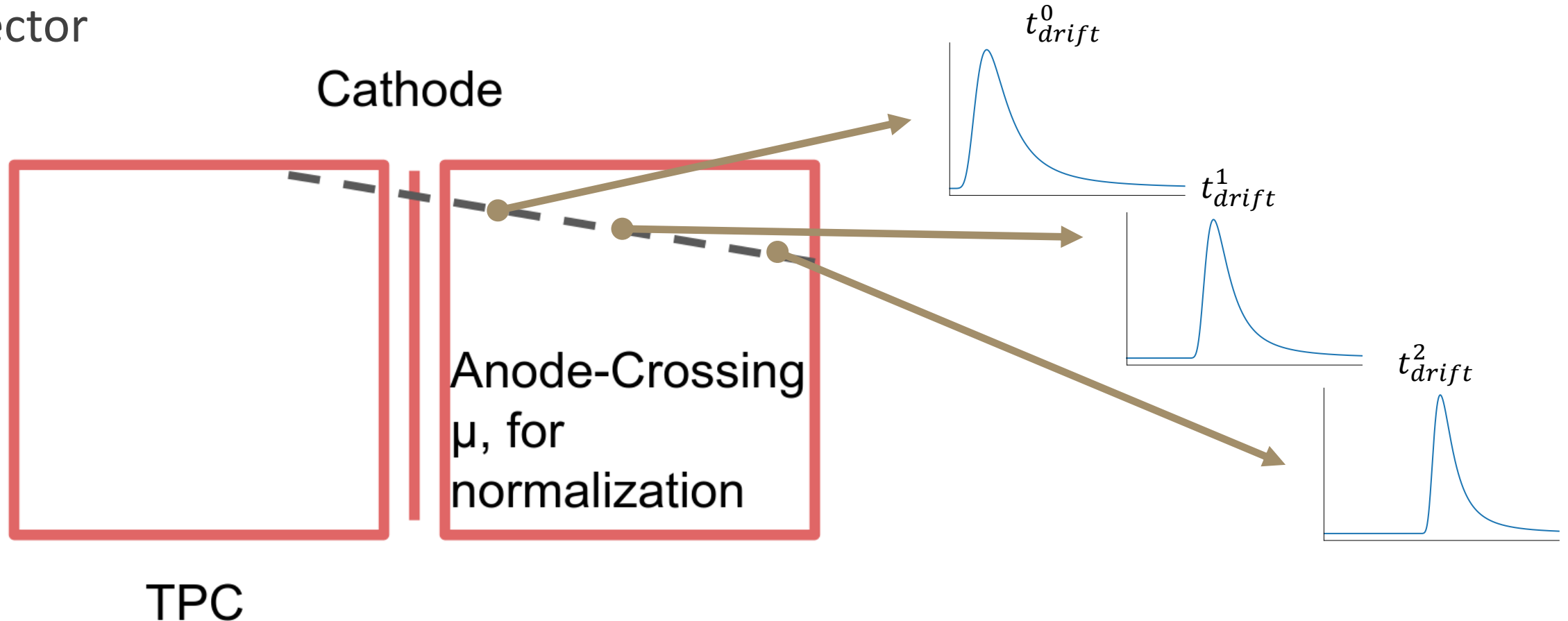
# Energy Scale Calibration Procedure

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- Step 1: **normalize** the detector response in the drift direction
  - This removes detector effects such as argon impurities which attenuate the signal
- Step 2: **calibrate** the energy scale
  - Examination of the systematic uncertainties and results at ICARUS
- For both steps, we have devised a procedure which addresses possible biases from diffusion

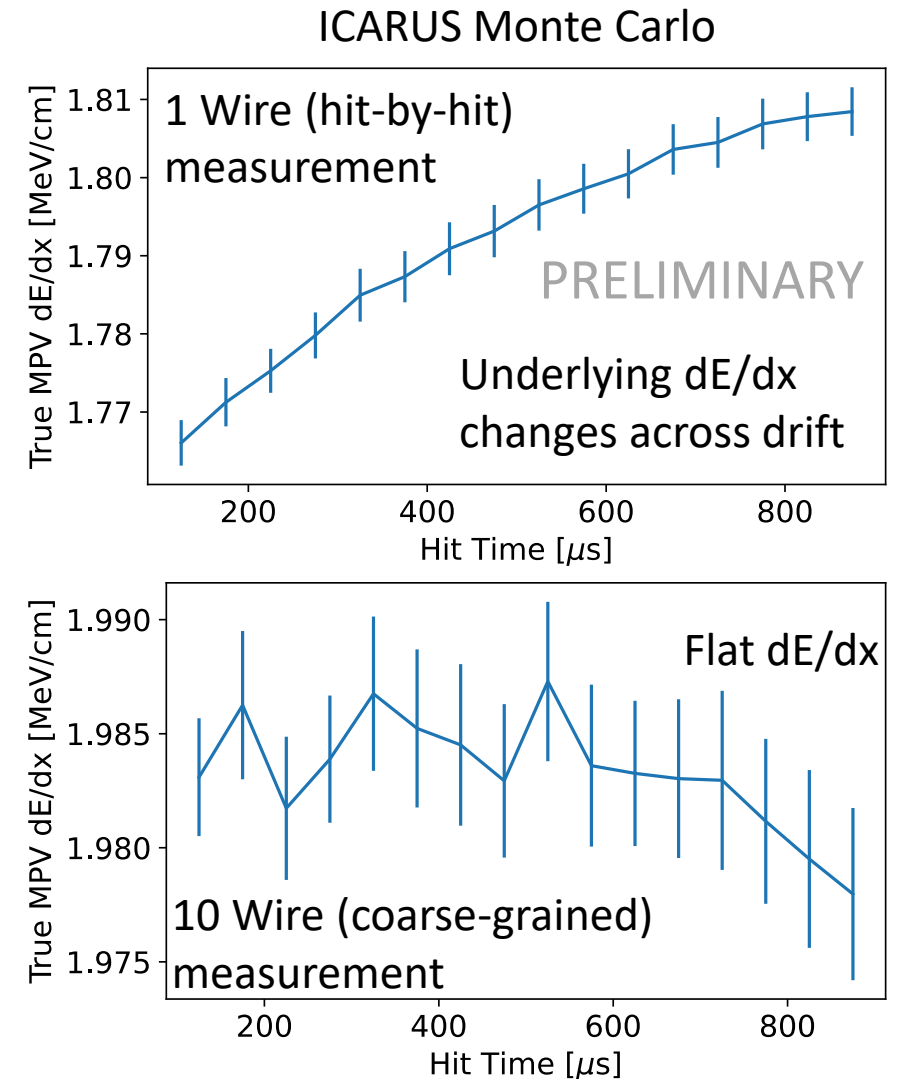
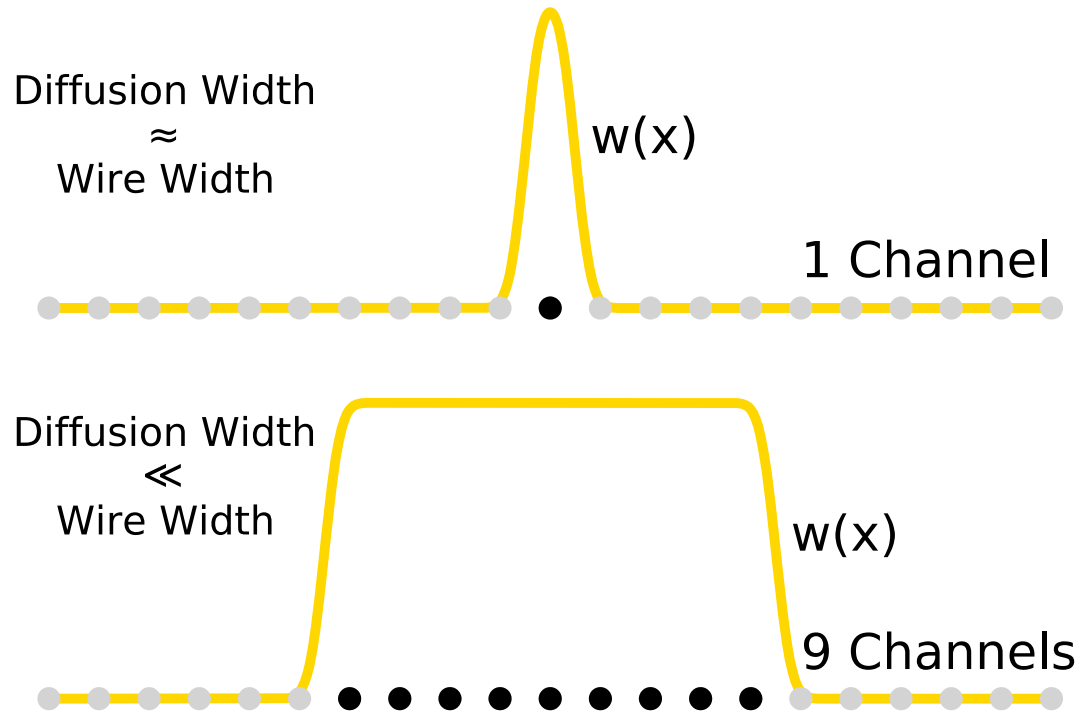
# Normalizing the Drift Direction Detector Response

- Impurities in the argon attenuate ionization electrons as a function of drift time
- To remove this effect: look at  $dQ/dx$  from cosmic muons, make it flat across the detector



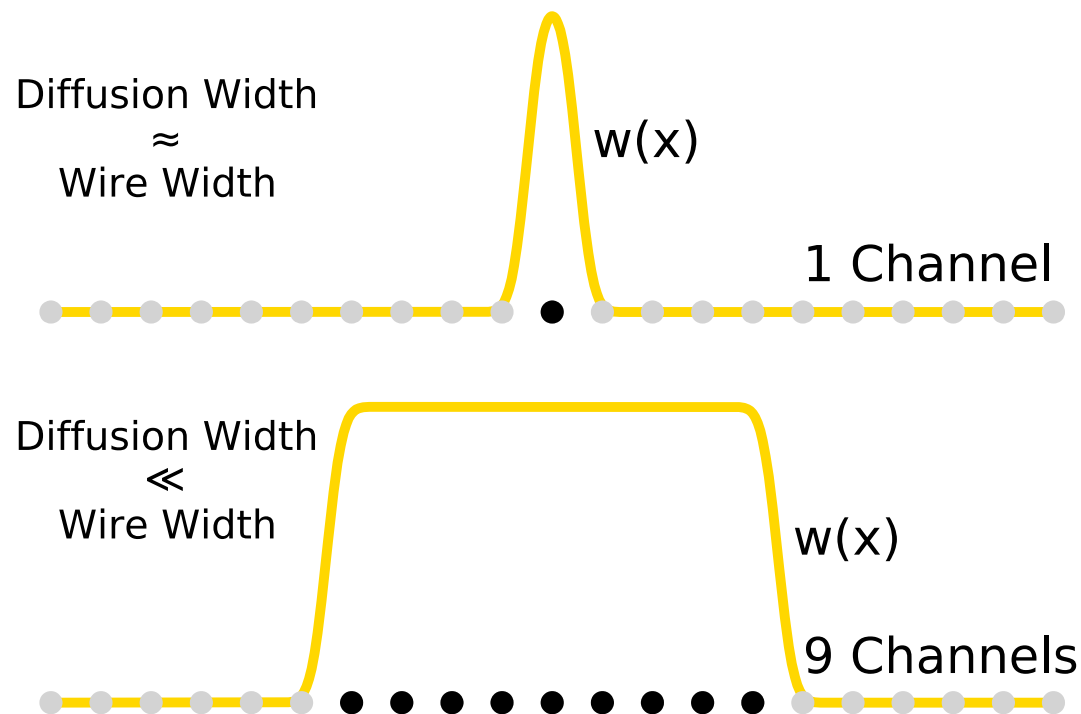
# Drift Direction Response Normalization with Diffusion

- Diffusion changes the underlying  $dE/dx$  of muon depositions across the drift direction
- We can remove this effect by coarse-graining the detector



# Drift Direction Response Normalization with Diffusion

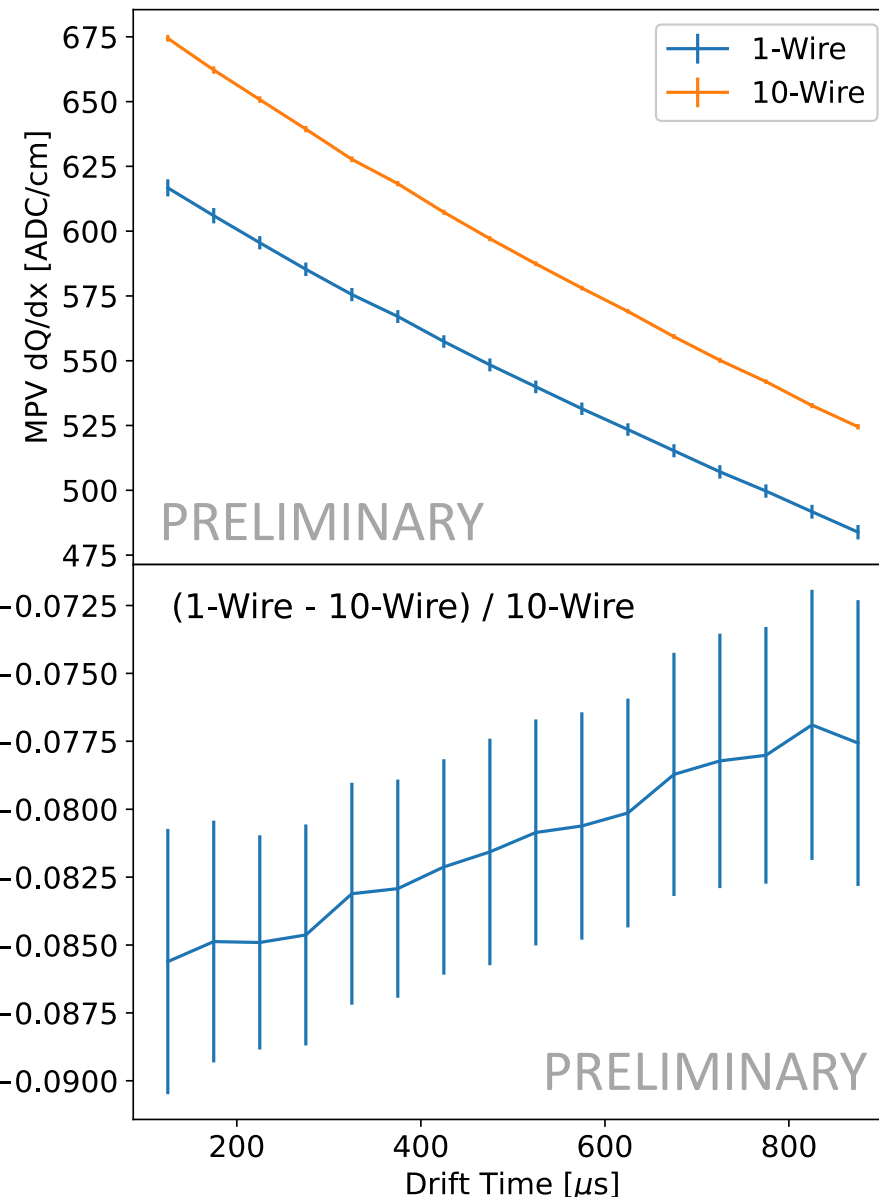
- Diffusion changes the underlying  $dE/dx$  of muon depositions across the drift direction
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The difference in thickness (and thus  $dE/dx$ ) narrows across the drift.

MPV  $dQ/dx$  Ratio

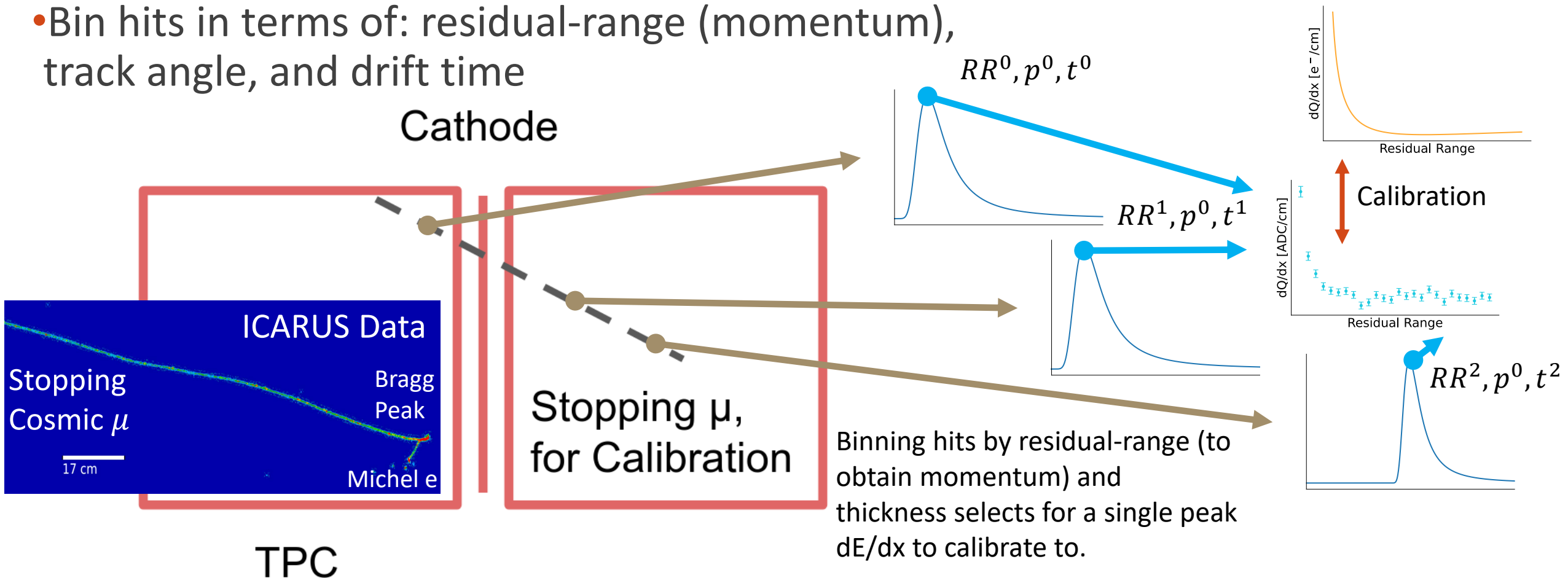
ICARUS Commissioning Data  
TPC WW Run 7897





# Energy Scale Calibration Procedure

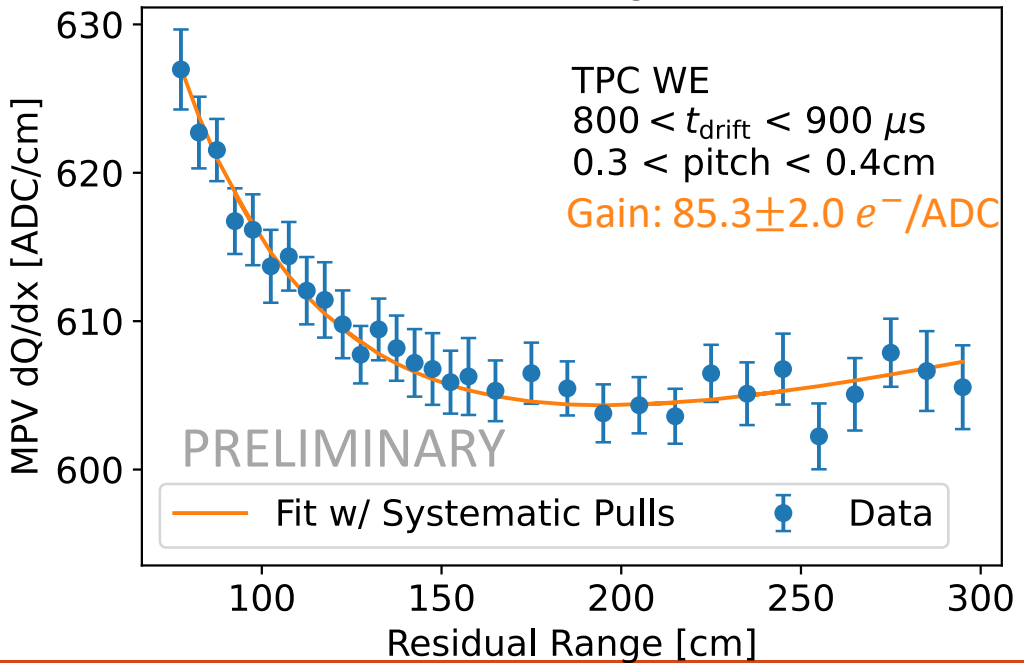
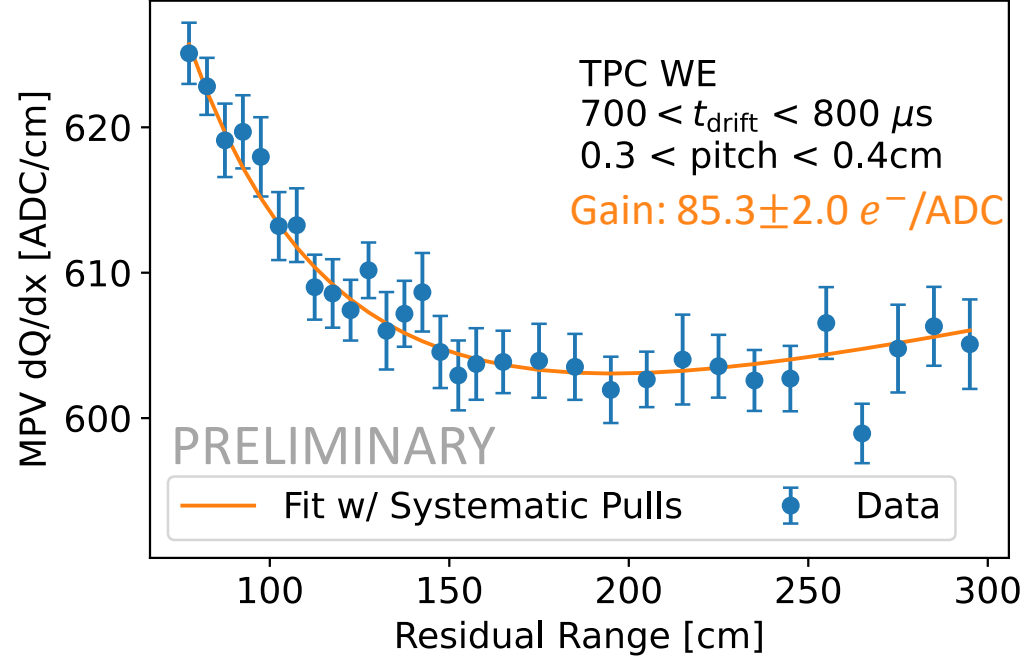
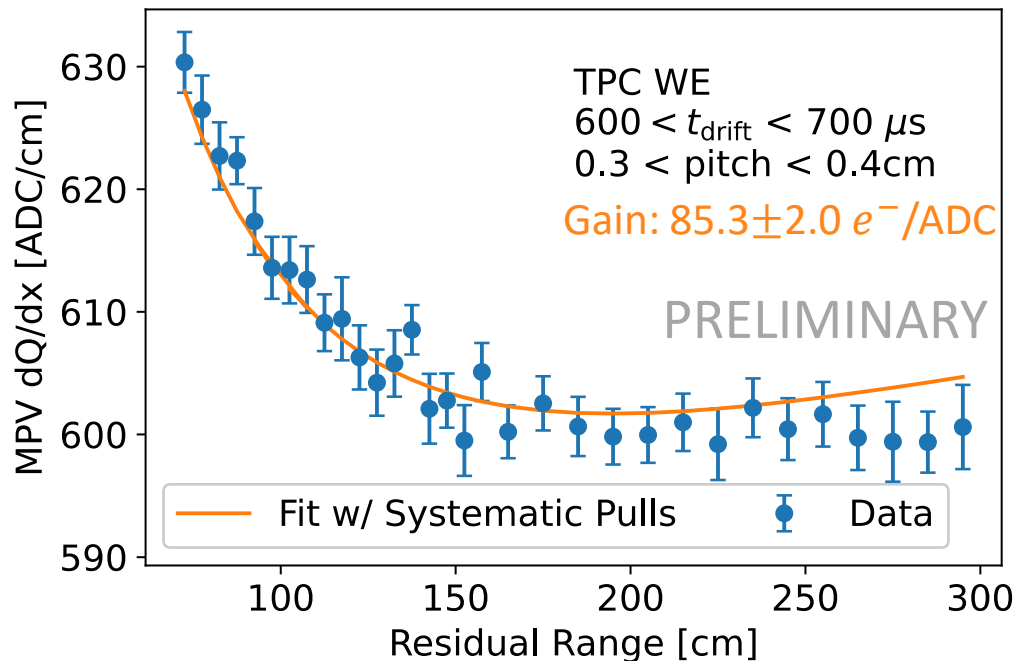
- After normalizing detector response, we calibrate the energy scale by fitting to the  $dQ/dx$  profile of stopping cosmic muons
- Bin hits in terms of: residual-range (momentum), track angle, and drift time



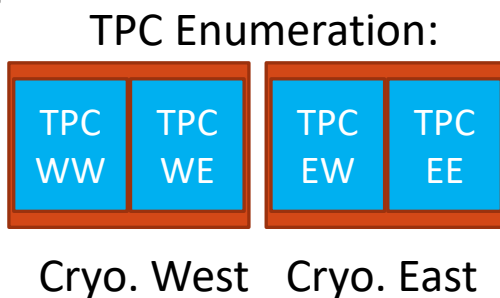
# Example dQ/dx Profile Data

ICARUS  
Commissioning  
Data

TPC WE



- Fit across all drift bins, with a separate gain in each TPC and “pulls” for systematics

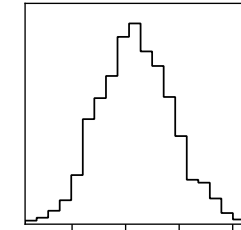


# Calibration Fit Results

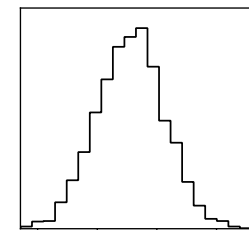
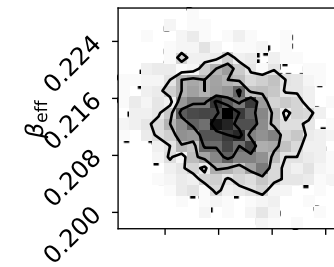
Systematic Parameters

Gain Measurement

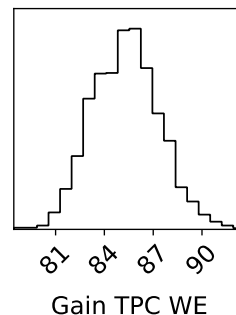
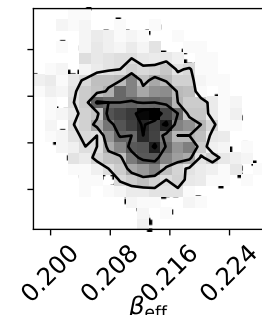
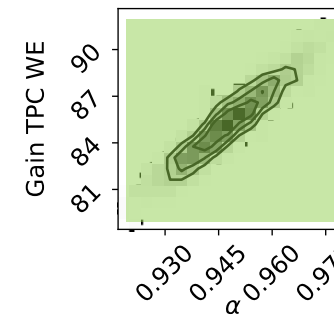
Parameter	Prior CV	Prior Unc.	Posterior CV	Posterior Unc.
Recombination $\alpha$	0.93	0.02	0.948	0.010
Recombination $\beta_{eff}$ [cm <sup>3</sup> /g][cm/kV]	0.212	0.005	0.212	0.005
Transverse Diffusion $D_T$ [cm <sup>2</sup> /s]	8.8	4.4	9.1	2.2
Mean Excitation Energy $I_0$ [eV]	188	17	194	15
Gain TPC EE [e <sup>-</sup> /ADC]			83.4	2.0
Gain TPC EW [e <sup>-</sup> /ADC]			81.8	2.0
Gain TPC WE [e <sup>-</sup> /ADC]			85.3	2.0
Gain TPC WW [e <sup>-</sup> /ADC]			84.3	2.0



ICARUS  
Commissioning  
Data



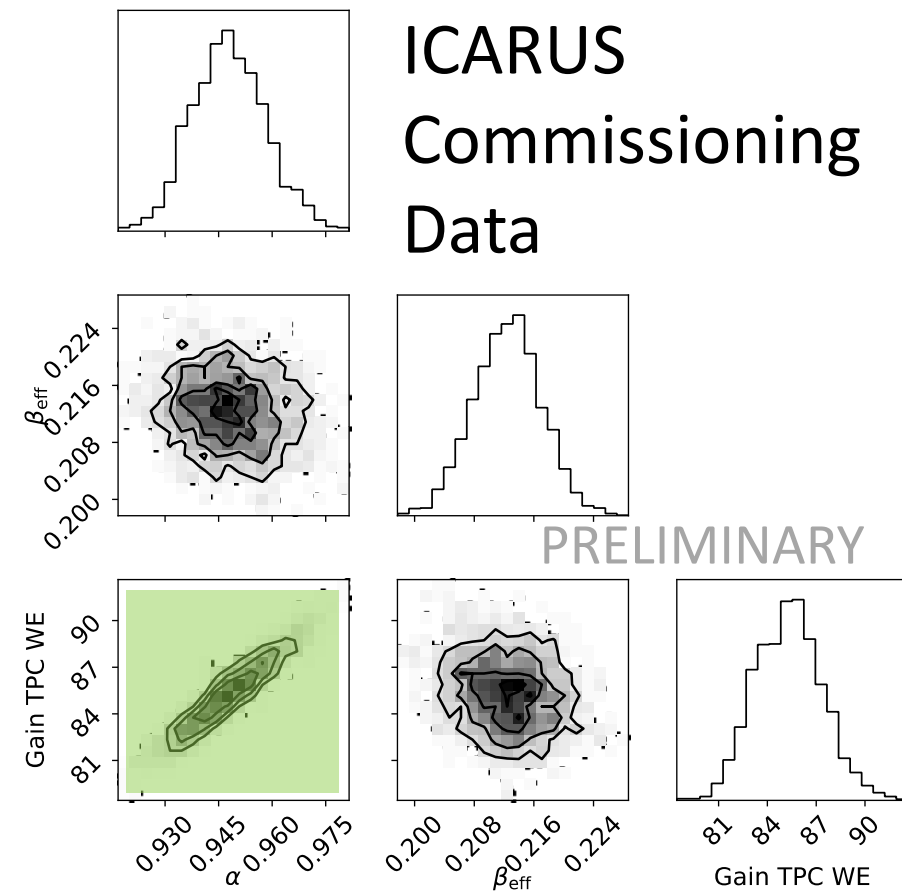
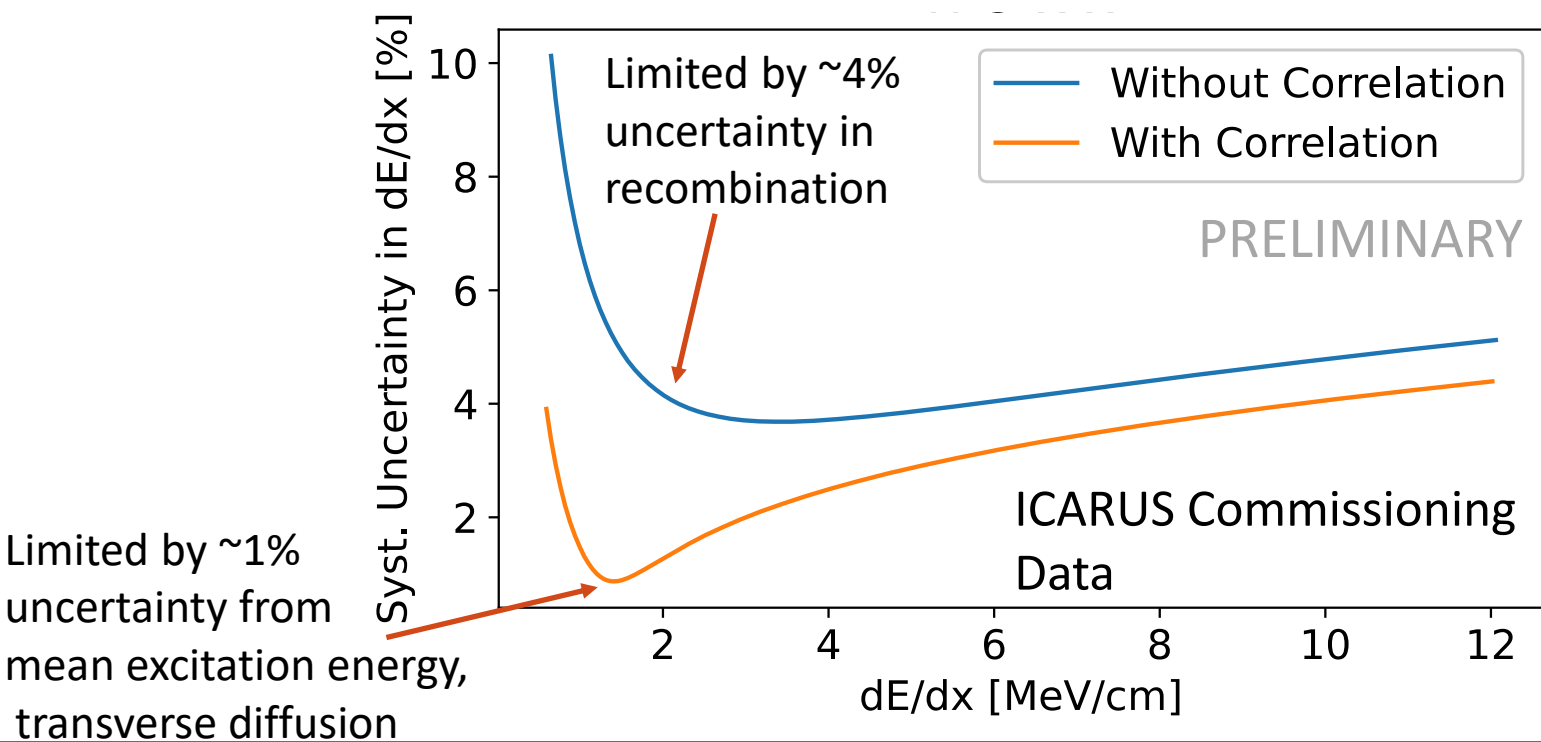
PRELIMINARY



Strong correlation between gain and recombination  $\alpha$  ( $\alpha$  determines behavior near the MIP dE/dx).

# Marginalize Over Recombination

- By leveraging the correlation between gain and recombination in our dataset, we can lower the systematic uncertainty in  $dE/dx$ 
  - i.e.: marginalize over gain and recombination together

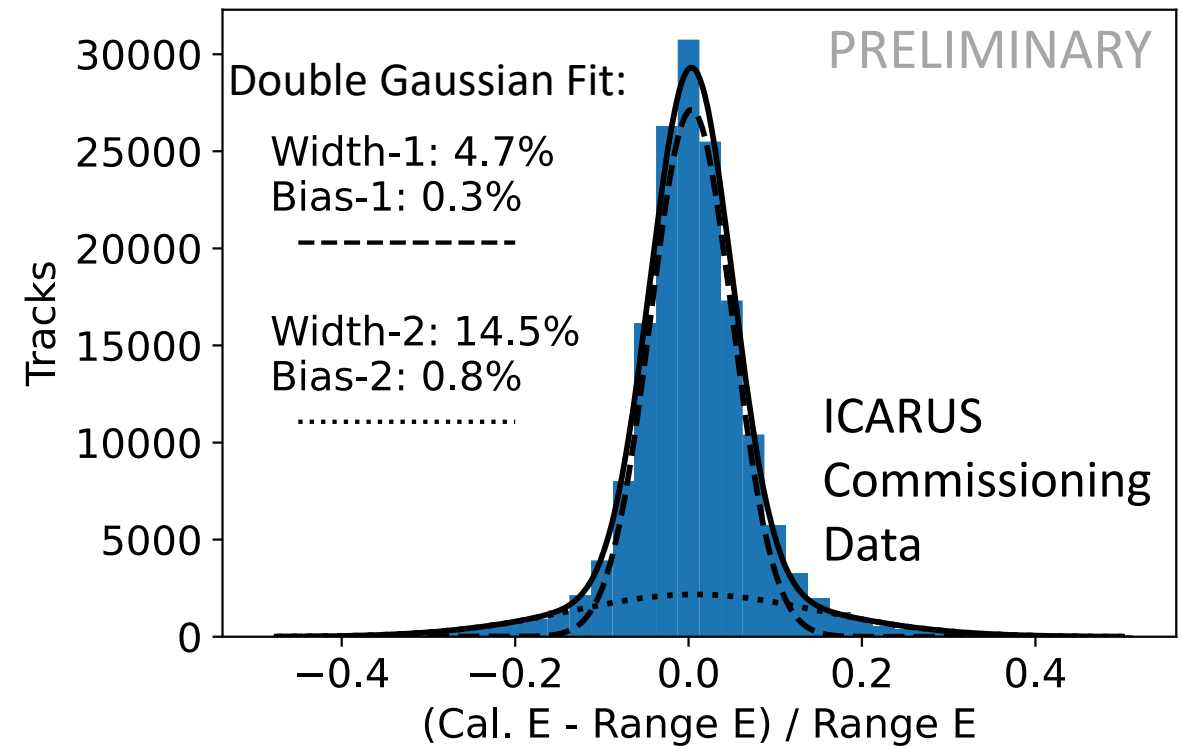
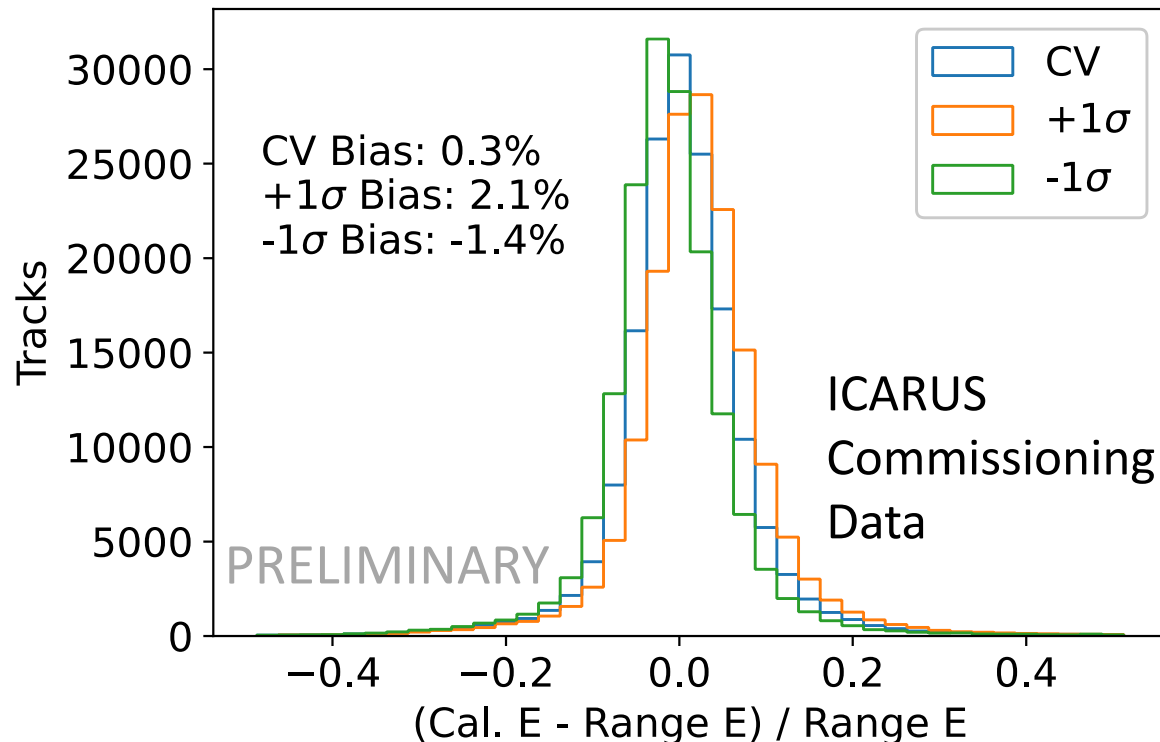


Strong correlation between gain and recombination  $\alpha$  ( $\alpha$  determines behavior near the MIP  $dE/dx$ ).



# Closure Test with Energy Reconstruction

- Compare calorimetric to range kinetic energy reconstruction for the stopping muon dataset

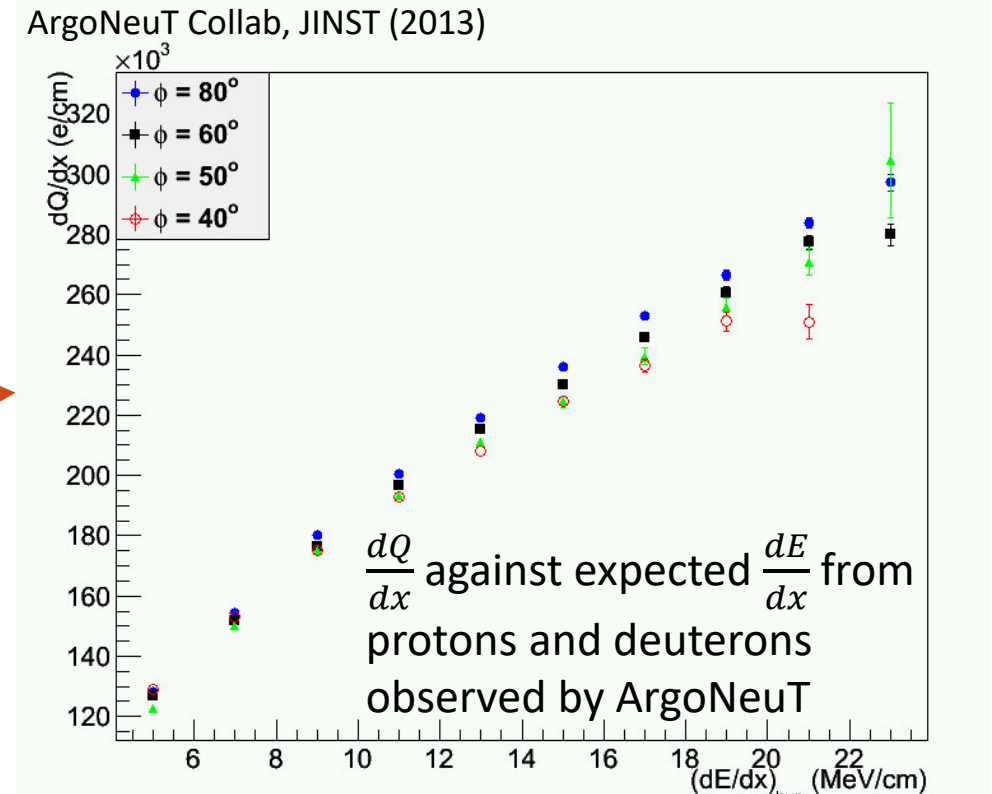


- 4.7% resolution in bulk of distribution (2% intrinsic resolution from range)
- Bias within 1 $\sigma$  range expected from systematic uncertainties

# Looking Forward: We Need More Measurements

- This procedure works well inside the (cosmic  $\mu$ ) dataset we have, but will it generalize to others?

- ArgoNeuT recombination found evidence for angular variations in recombination
- Could there be particle-type dependence in recombination (at the percent level)?



- To obtain the best possible energy measurements, we need to better understand the argon
  - At ICARUS we can measure recombination and diffusion

# Conclusion

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- Accurate and precise energy measurements are needed to unlock the physics potential of SBN and DUNE
- Systematic uncertainties in liquid argon properties limit the accuracy of calibrating the energy scale
  - One of these, transverse diffusion, plays a role not appreciated in LArTPC experiments that we are now accounting for in ICARUS
- In ICARUS, we've implemented a calibration procedure that limits the effects of these systematic uncertainties
- Looking forward, more measurements of LAr properties are needed to get the best possible energy measurements

# Backup Slides

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# Energy Loss by Elastic Scattering

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- Charged particles lose energy in elastic collisions with atomic electrons
- Above the mean excitation energy, this is described by the Rutherford formula:

$$\rho_e \frac{d\sigma}{dT} \propto \frac{1 - \beta^2 T/T_{max}}{T^2}$$

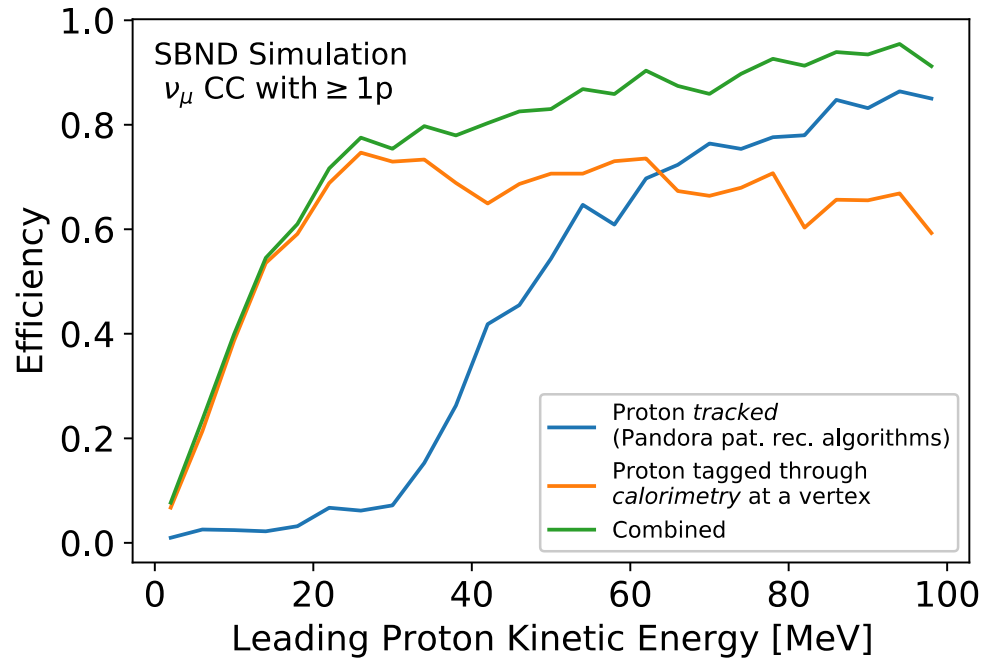
Due to the power-law behavior of Rutherford scattering, muons lose much of their energy in a small number of large energy-transfer collisions (delta rays)

Mean:  $\sim \int dT T \frac{d\sigma}{dT} \rho_e$  -- diverges at low  $T$  -> atomic effects important

Variance:  $\sim \int dT T^2 \frac{d\sigma}{dT} \rho_e$  -- converges at low  $T$  -> *delta rays determine variance*

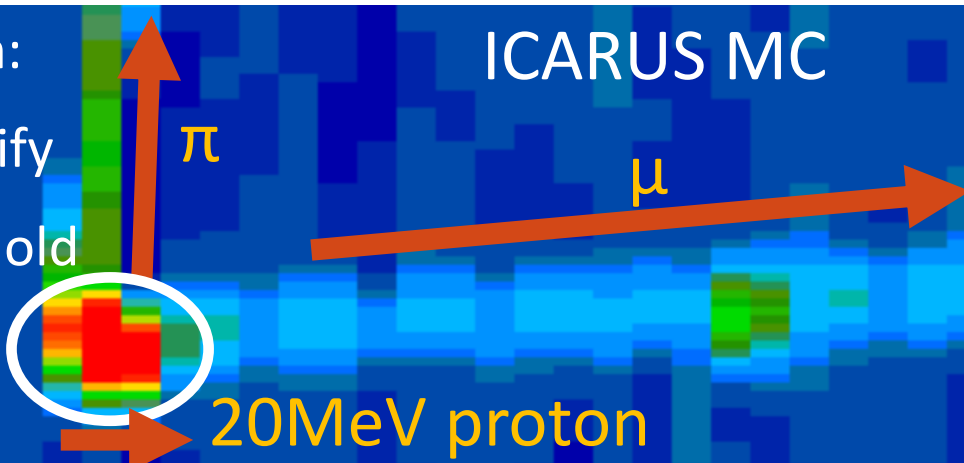
# Using Charge at SBN: Low Energy Protons

Charge-based reconstruction lowers the proton identification threshold, which can improve neutrino energy and topology reconstruction

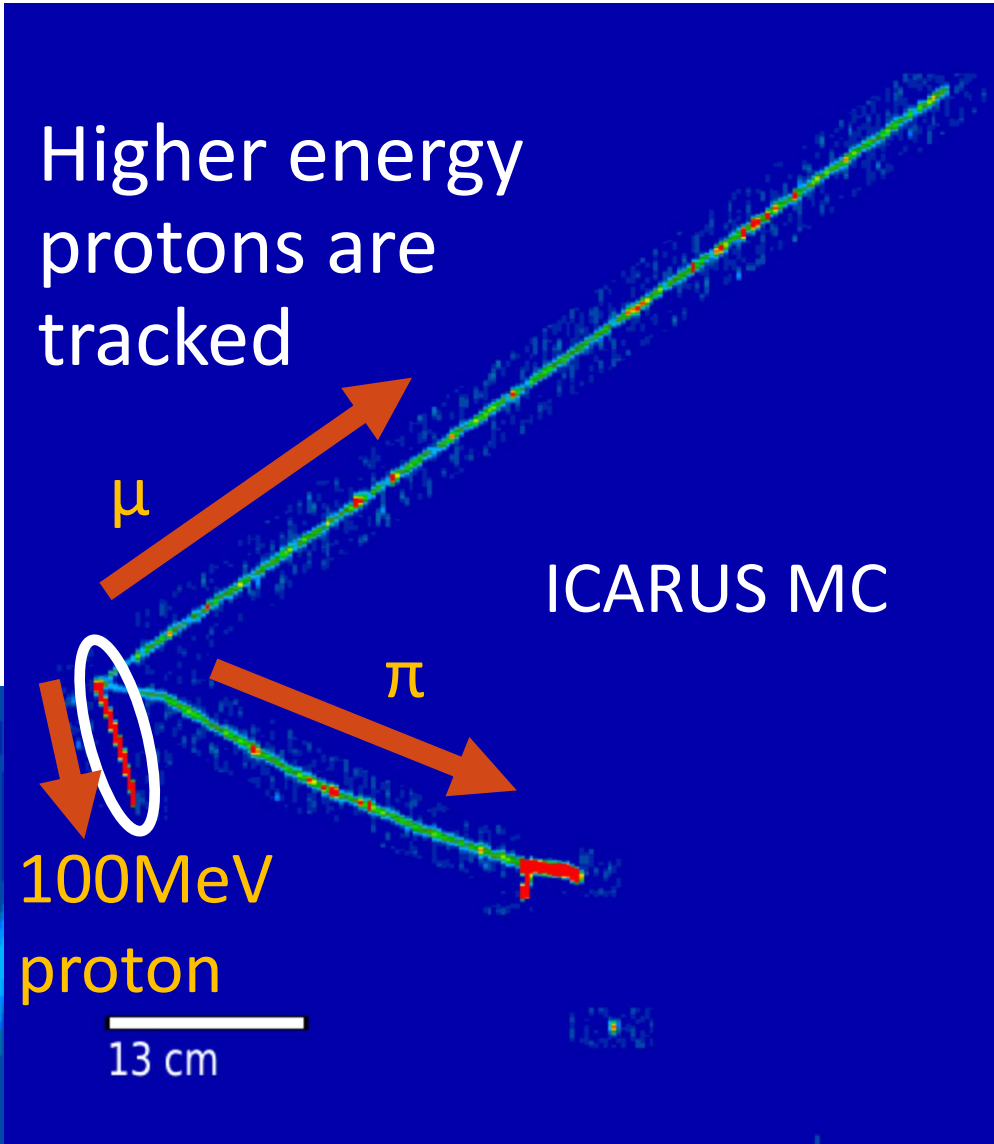


New Algorithm: use charge at vertex to identify protons below tracking threshold

1.5 cm

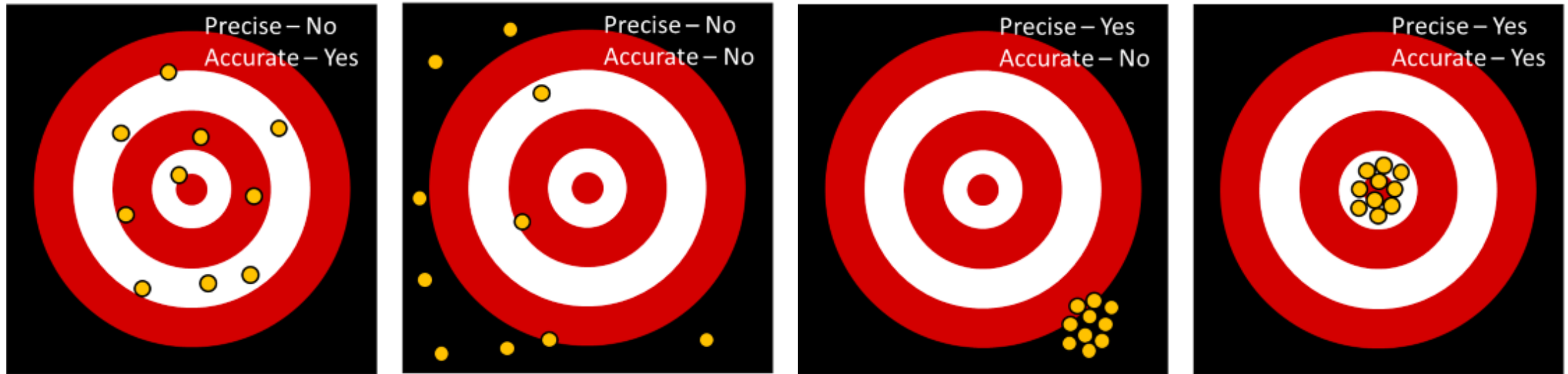


Higher energy protons are tracked



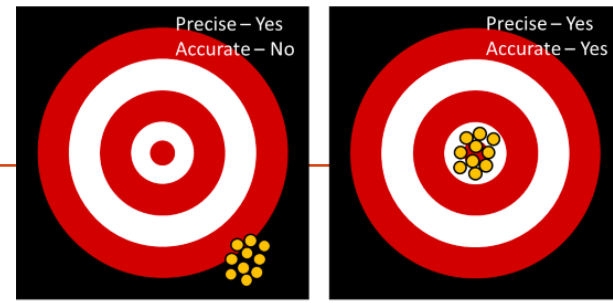
# How Well Can We Measure E using Q in LAr?

- Developments in LArTPC technology (cold electronics, high argon purity) increasingly provide excellent charge measurements
- How well can we turn those measurements of charge into energy?

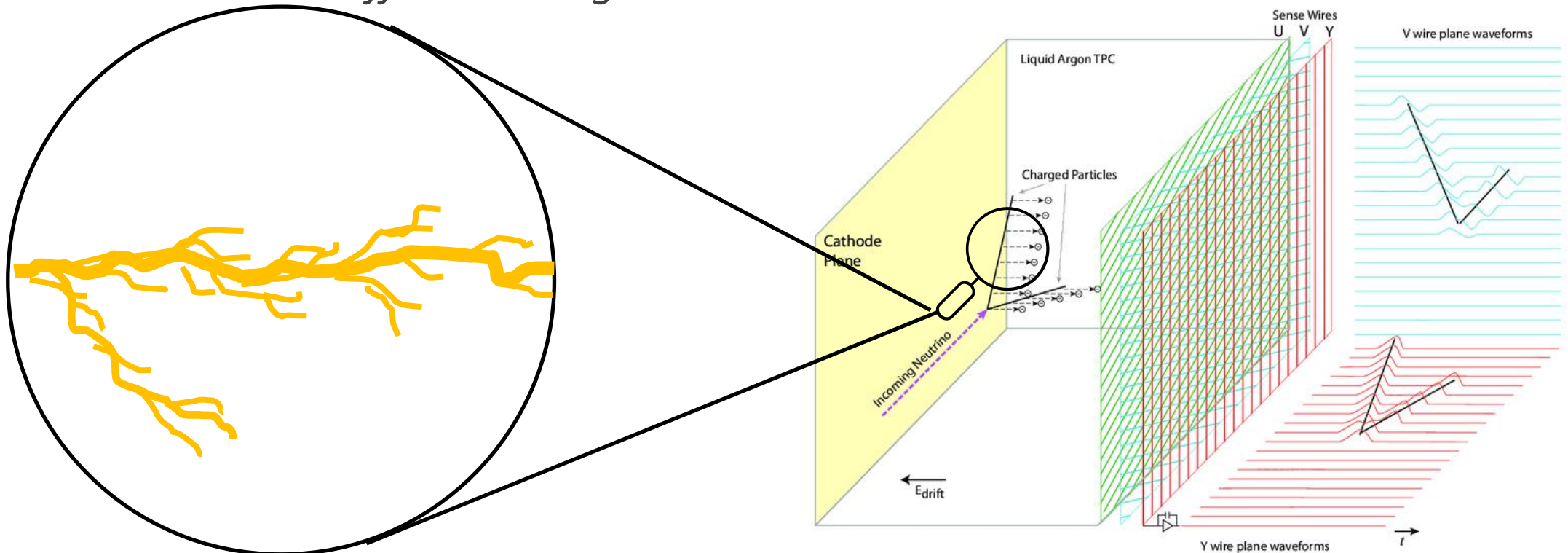


How *accurately* and *precisely* does charge measure energy in liquid argon?

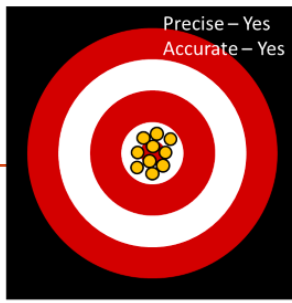
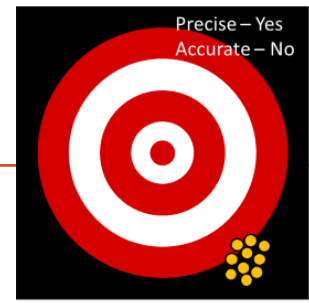
# Energy Resolution with Charge in LAr



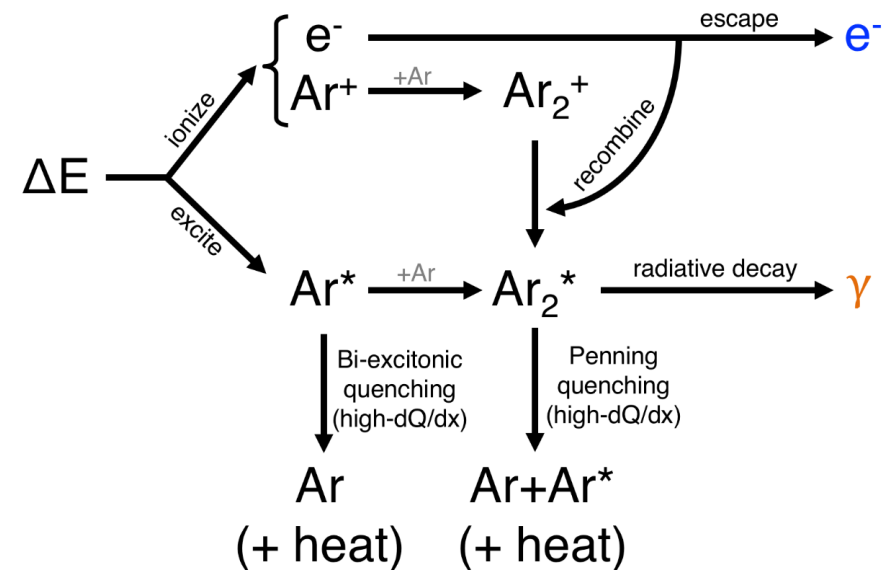
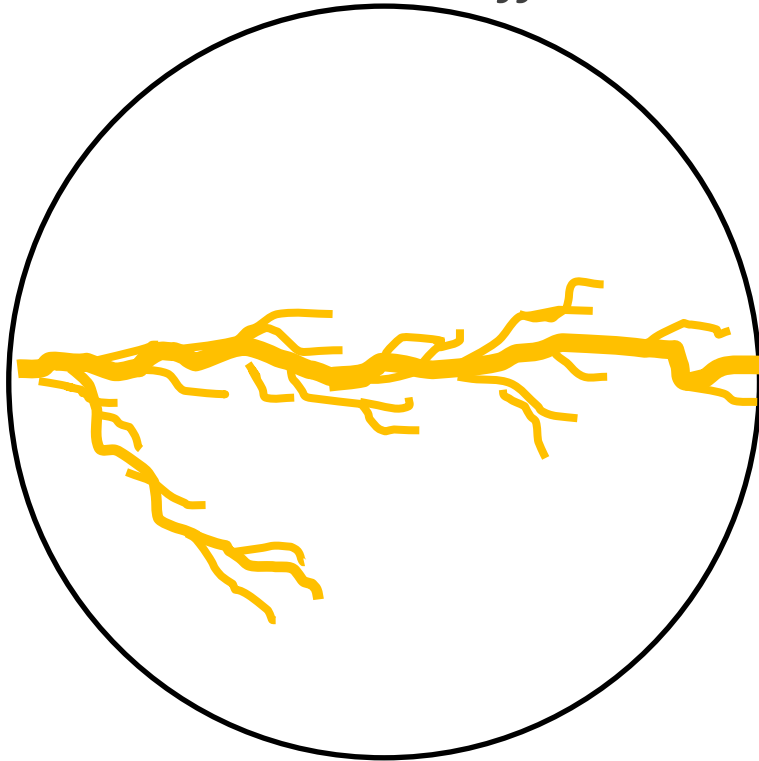
- Charged particles traversing argon create a cascade of delta rays
- The same energy loss across a wire can be produced by different spectra of delta rays, which recombine differently
  - *This produces a resolution effect in charge measurements!!*



# Energy Resolution with Charge in LAr

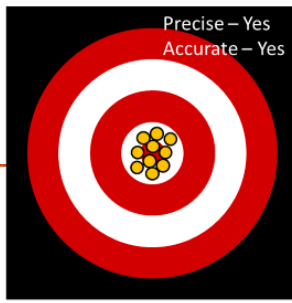
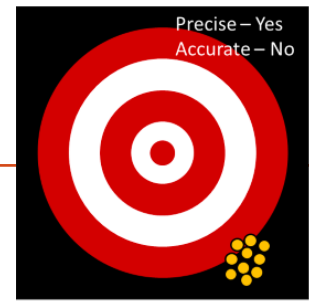


- Charged particles traversing argon create a cascade of delta rays
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Ionization electrons *recombine* into excited Ar states along each ionization column.

# Energy Resolution with Charge in LAr



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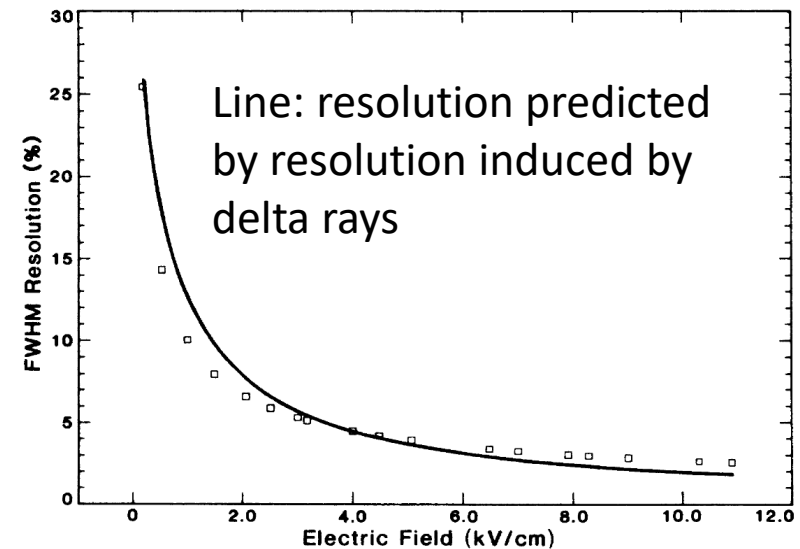
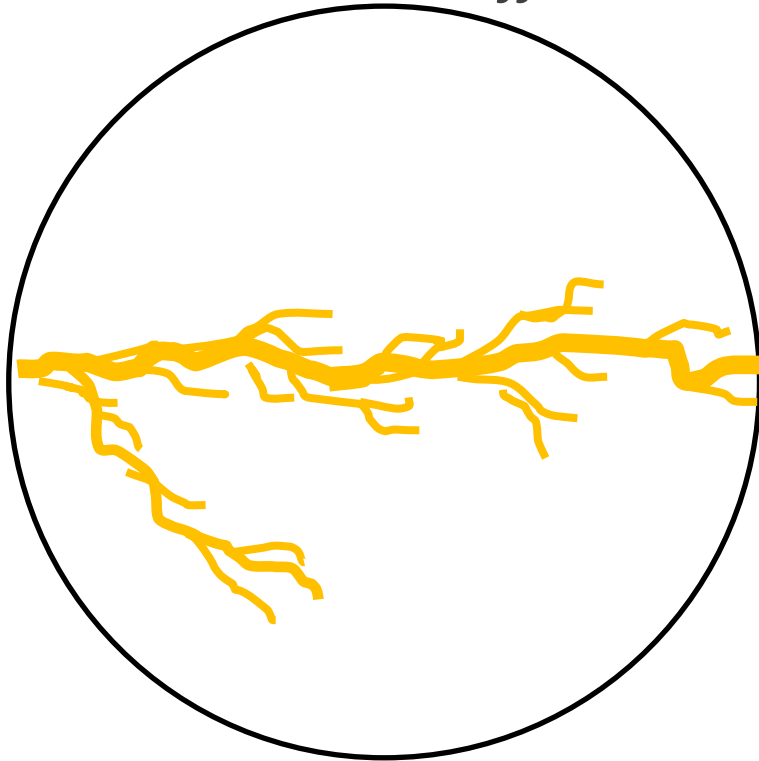
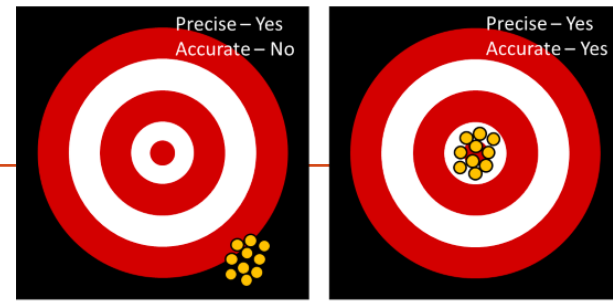


FIG. 3. A fit of Eq. (3) to the measured resolution of a 976-keV electron in liquid argon (Ref 12).

Thomas, Imel, and Biller PR A (1988)

# Can Resolution be Improved for DUNE?

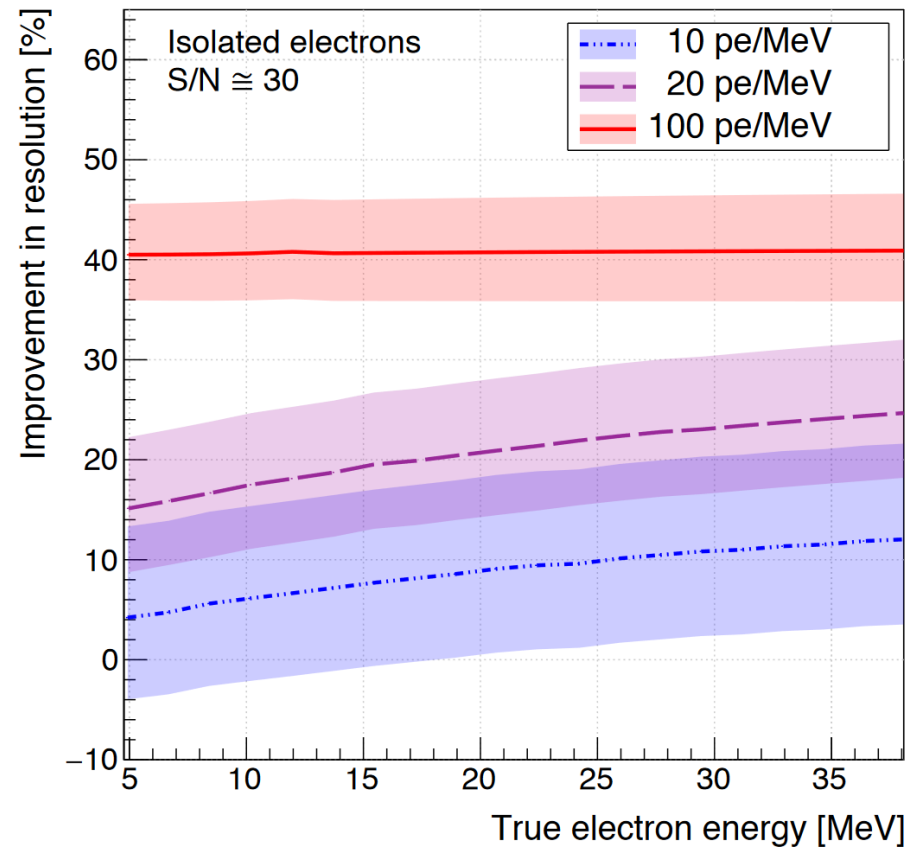


- Collecting both the charge and light from energy depositions could remove the resolution from recombination

Projected improvement in energy resolution using Q+L from LArIAT

There is not a significant improvement at the projected DUNE light yield (5 pe / MEV)

W Foreman et al., PR D (2020)

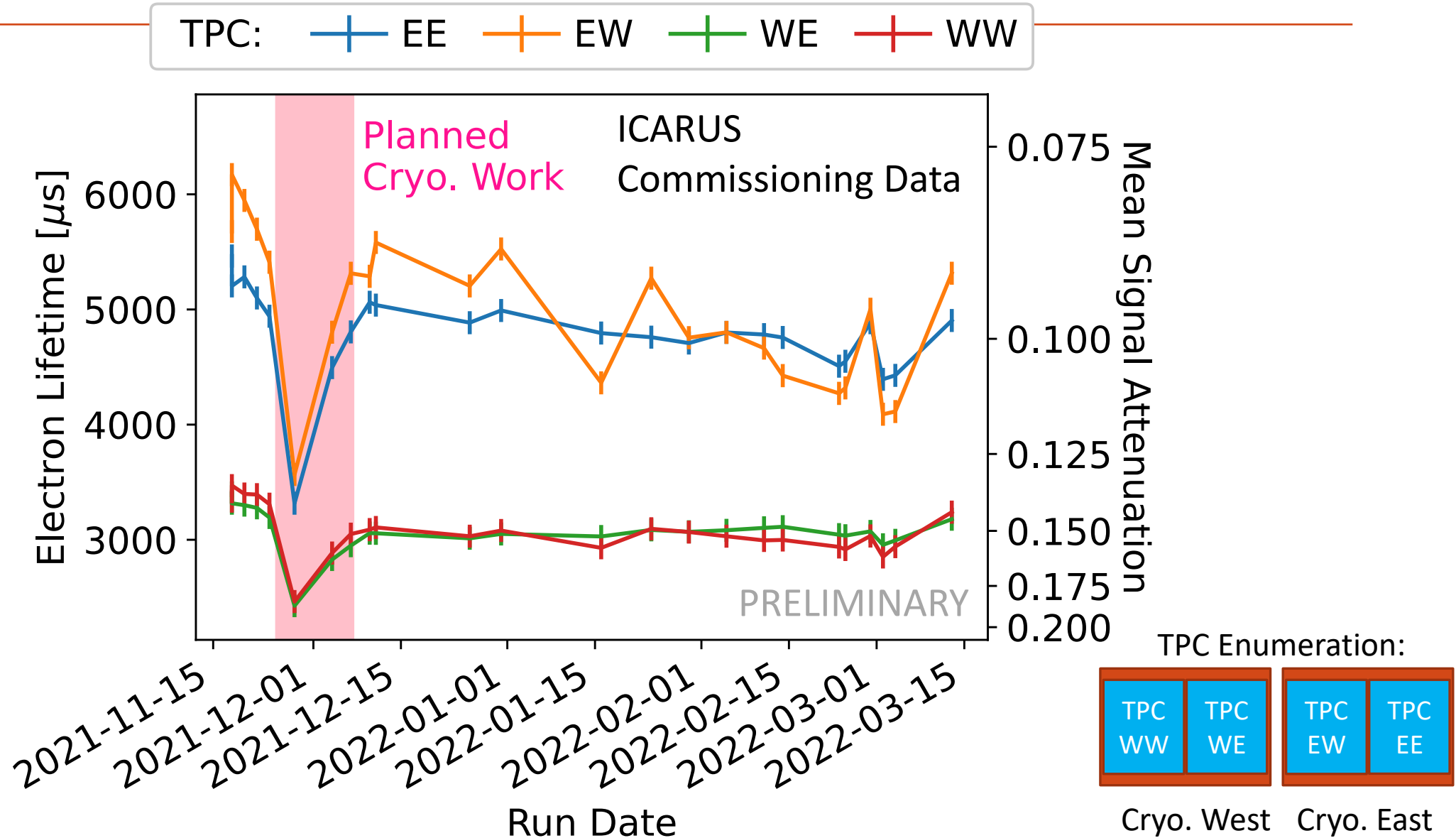


... however, adding a photosensitive dopant to the argon in DUNE would increase light-yield by turning it into charge collection.

Suggested by A Mastbaum, F Psihas, and J Zennaro (arxiv:2203.14700)



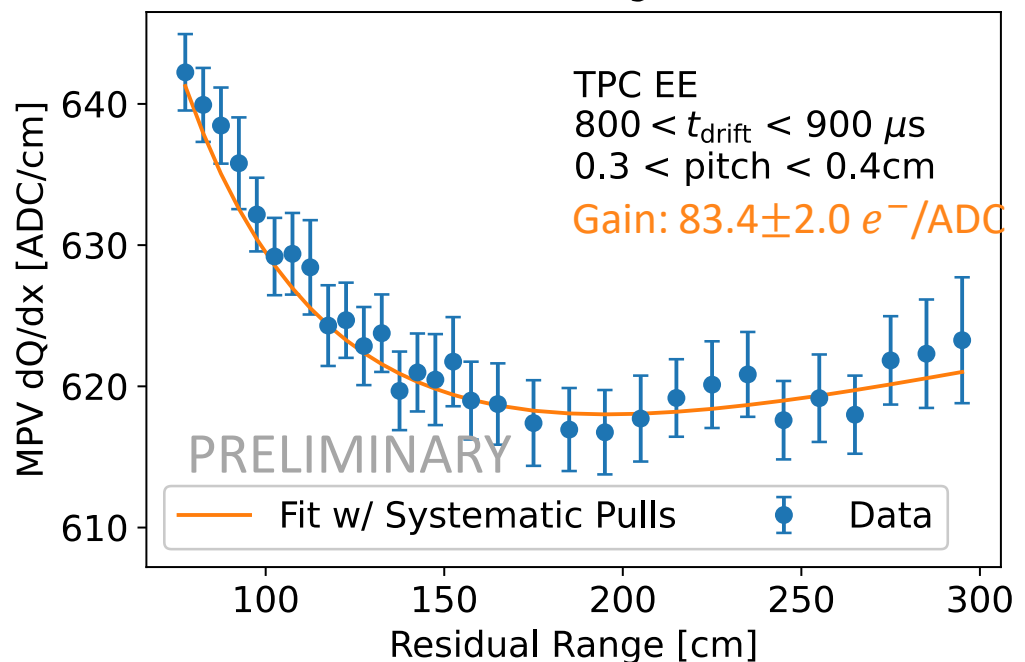
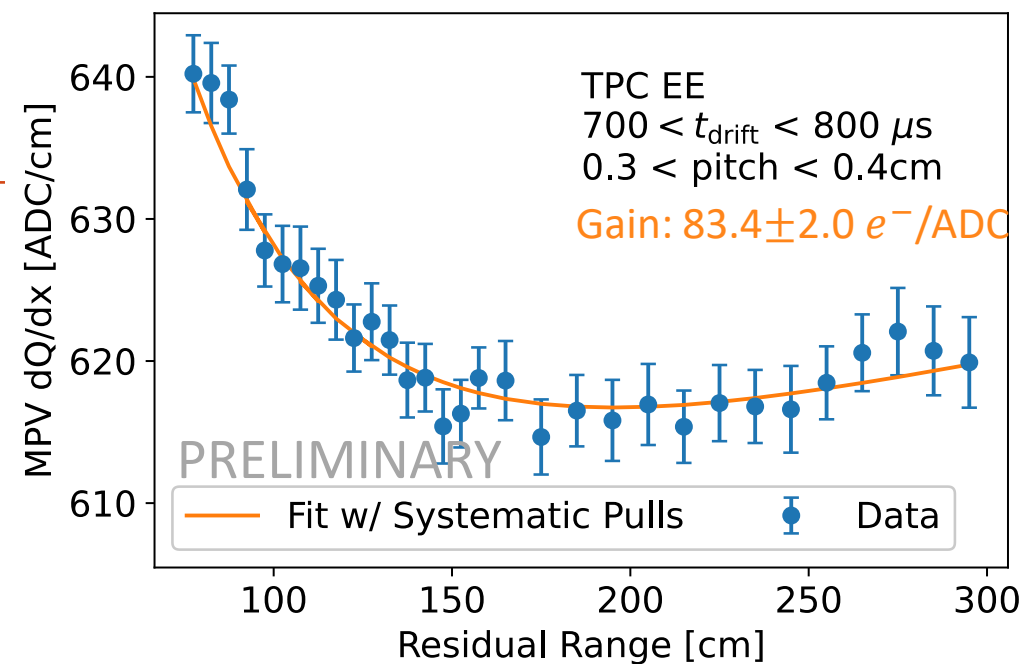
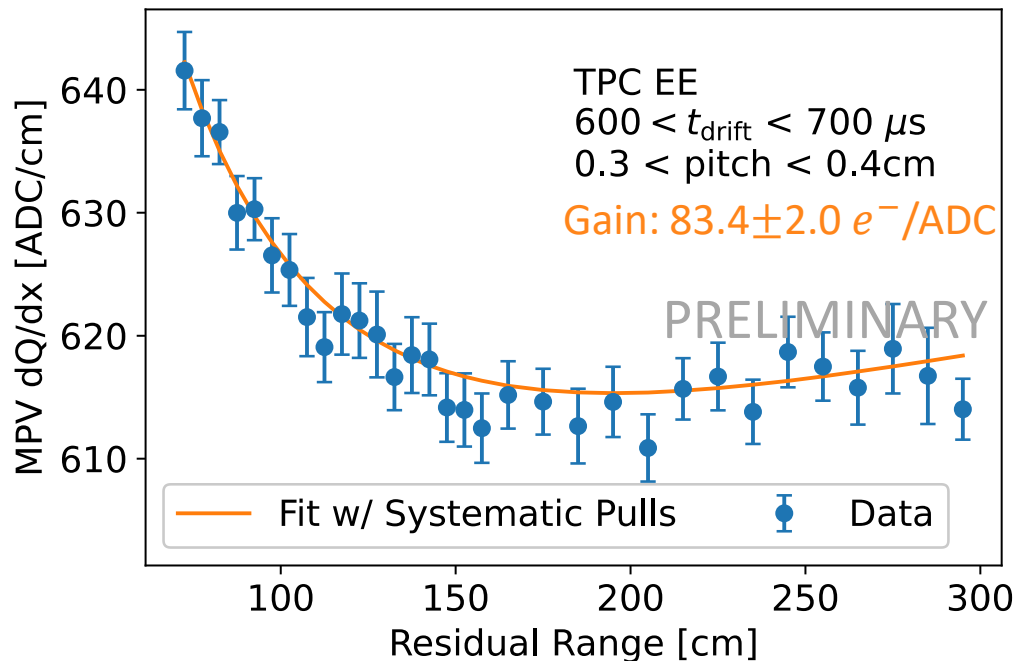
# Electron Lifetime Result on Calibration Dataset



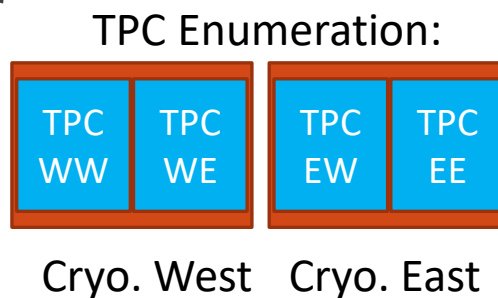
# Example dQ/dx Profile Data

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TPC WE



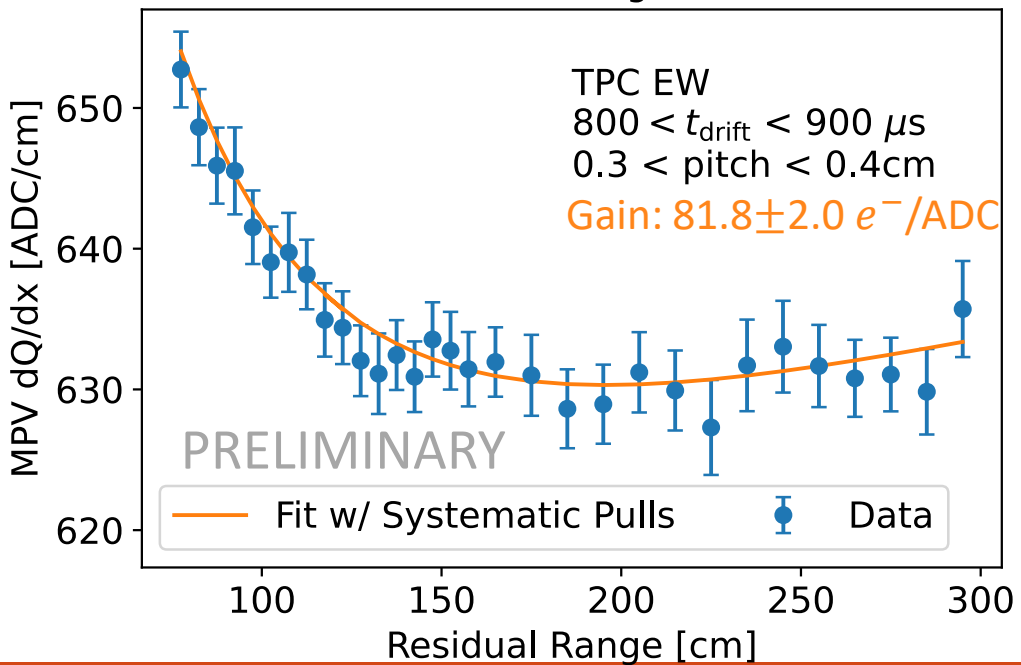
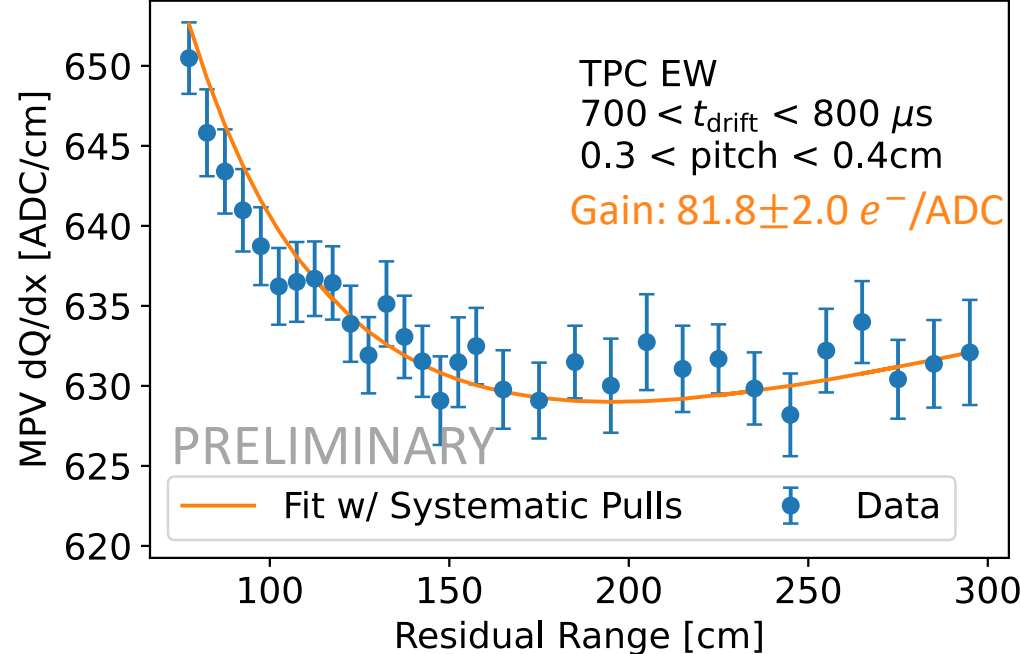
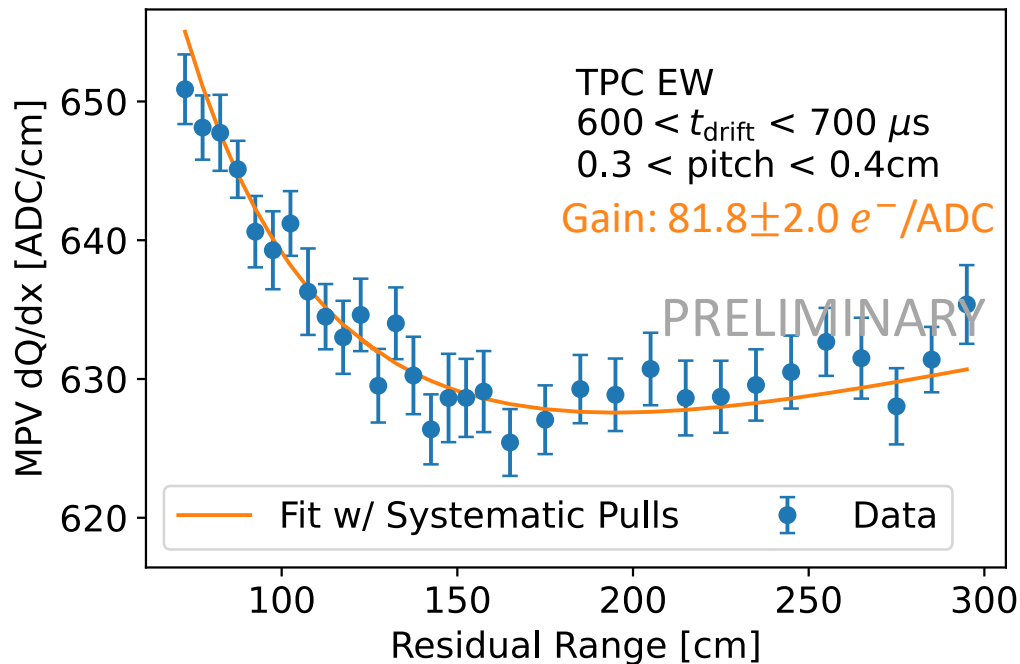
- Fit across all drift bins, with a separate gain in each TPC and “pulls” for systematics



# Example dQ/dx Profile Data

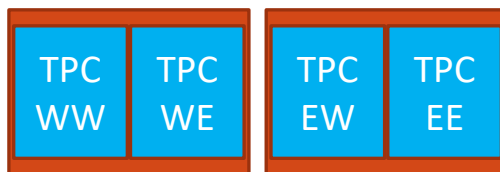
ICARUS  
Commissioning  
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- Fit across all drift bins, with a separate gain in each TPC and “pulls” for systematics

TPC Enumeration:

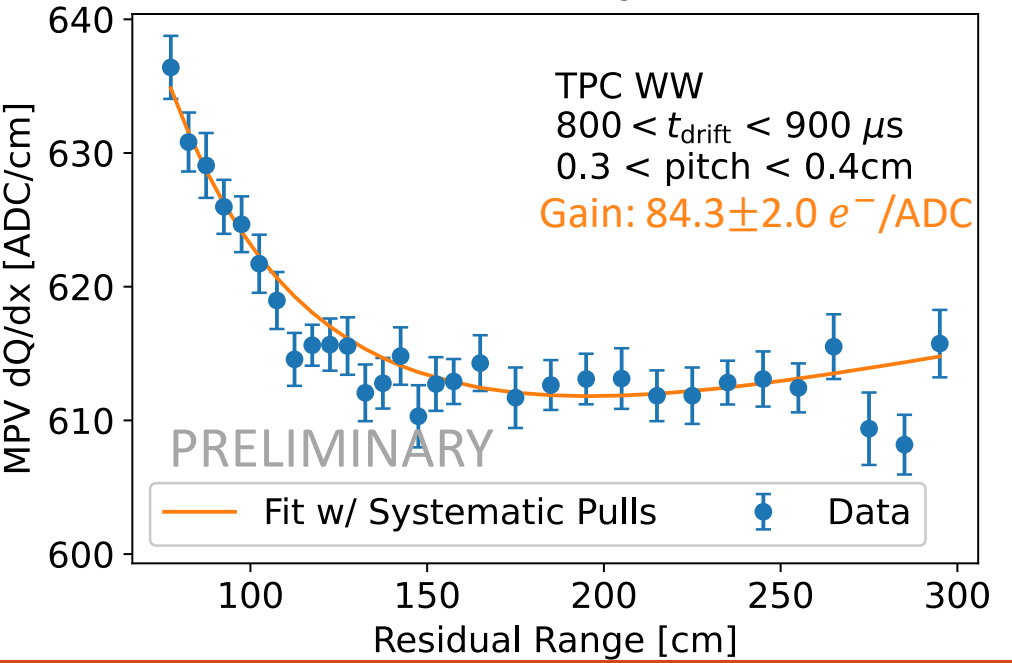
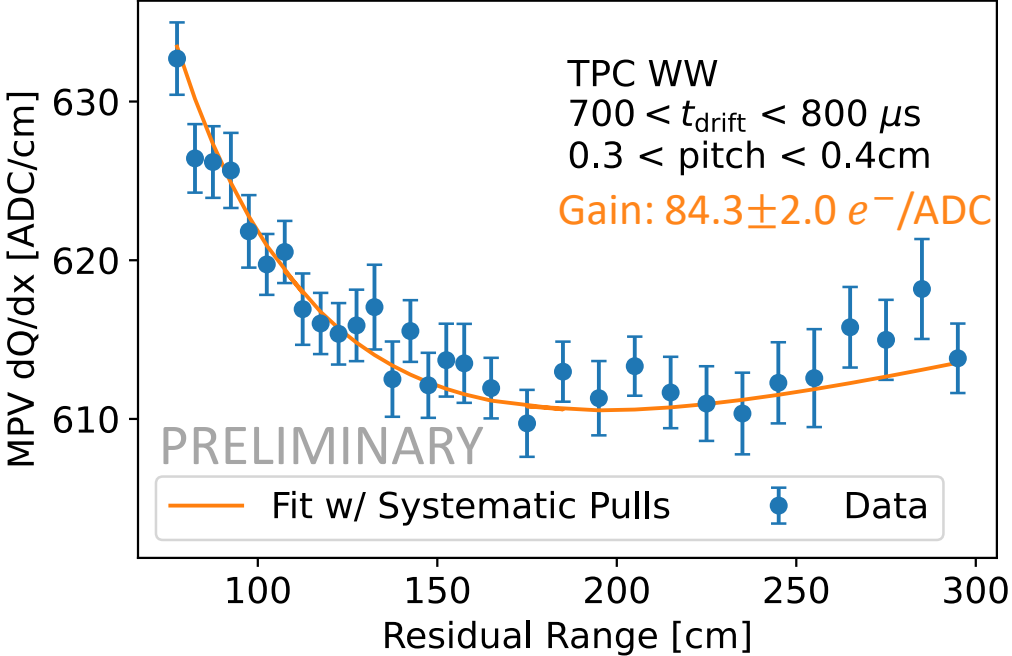
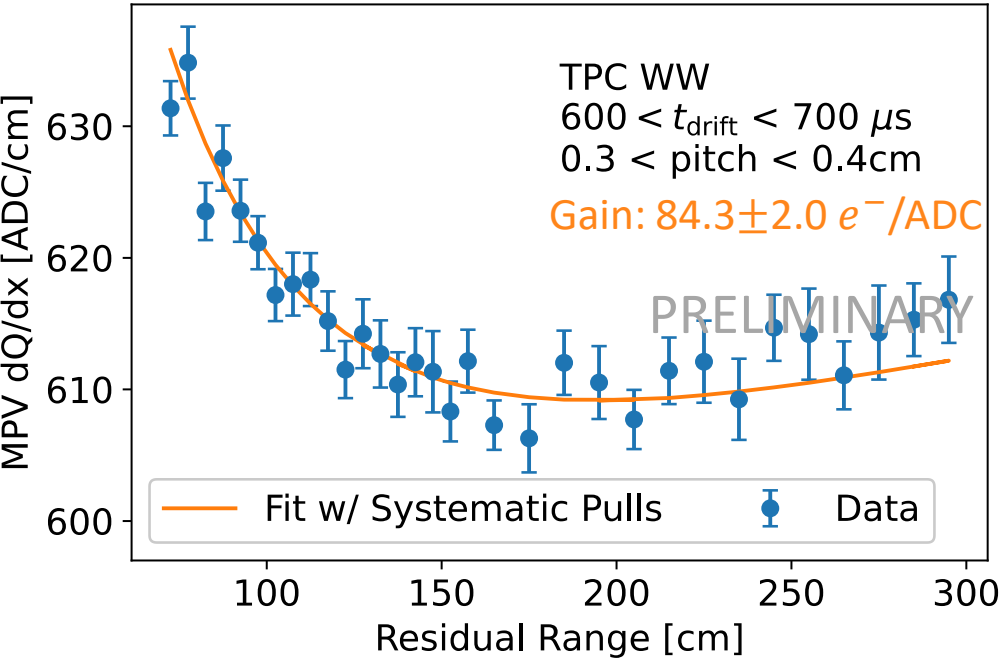


Cryo. West

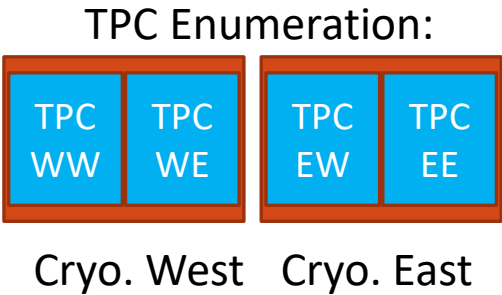
Cryo. East

# Example dQ/dx Profile Data

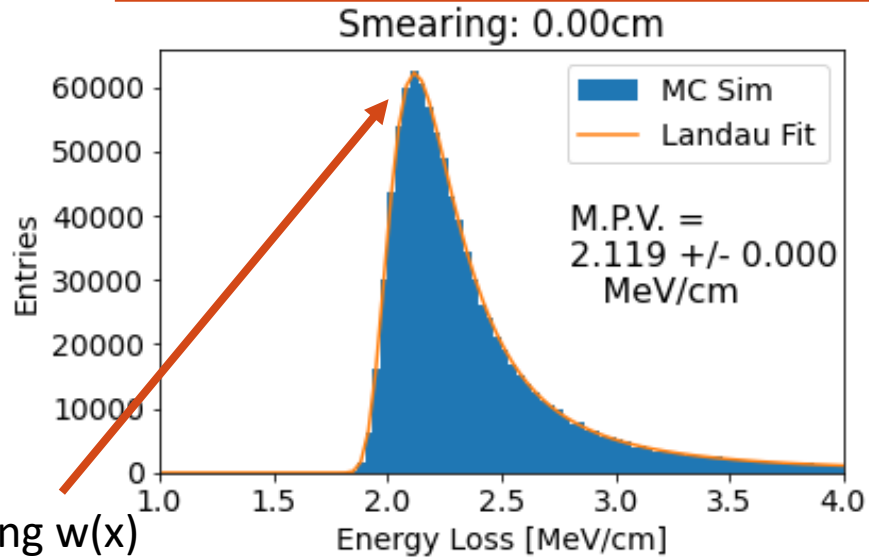
ICARUS  
Commissioning  
Data  
  
TPC WE



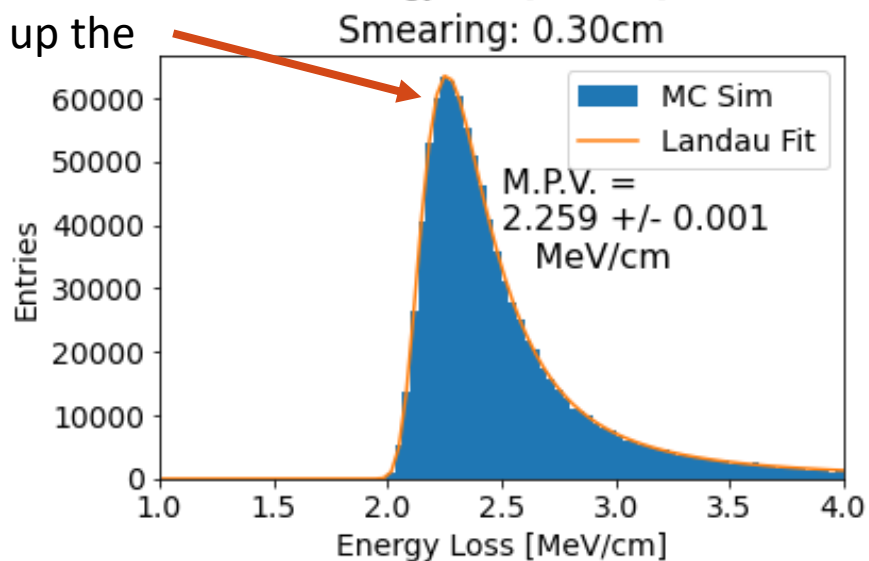
■ Fit across all drift bins, with a separate gain in each TPC and “pulls” for systematics



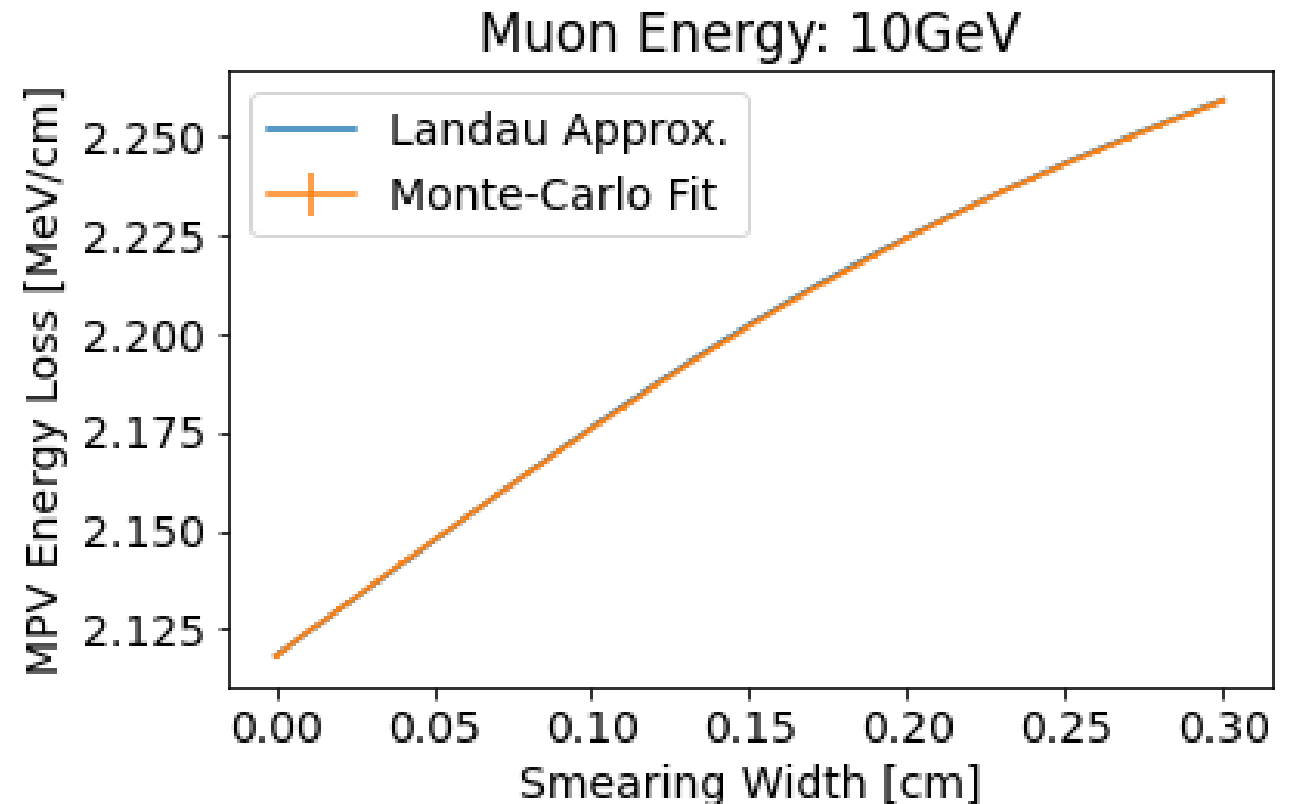
# Effect on the MPV: Toy MC



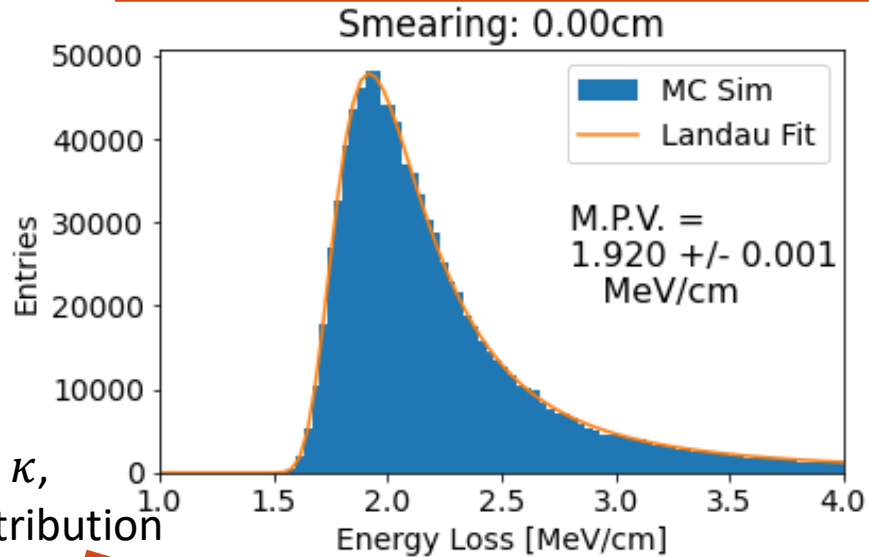
Smearing  $w(x)$   
Pushes up the  
MPV!



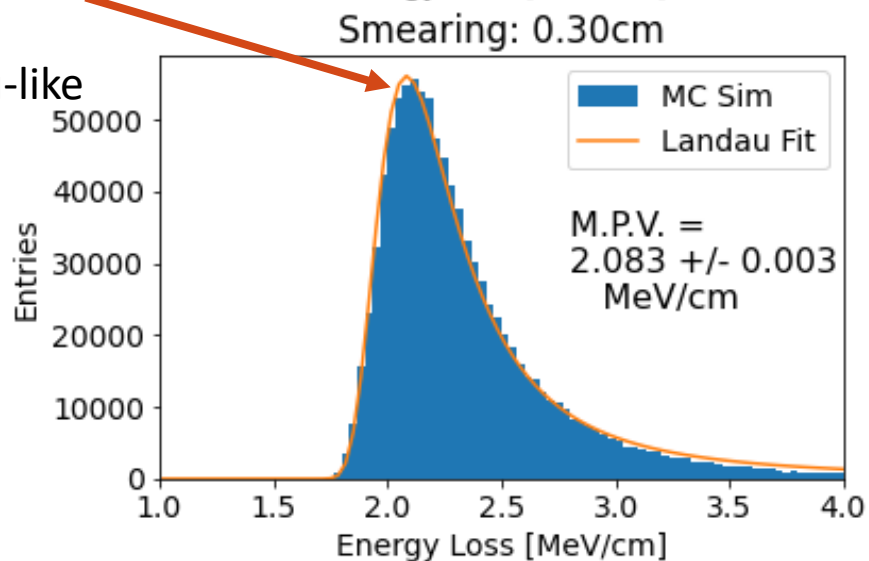
- Results from a toy MC of muon energy loss in LAr, for a wire spacing of 3mm



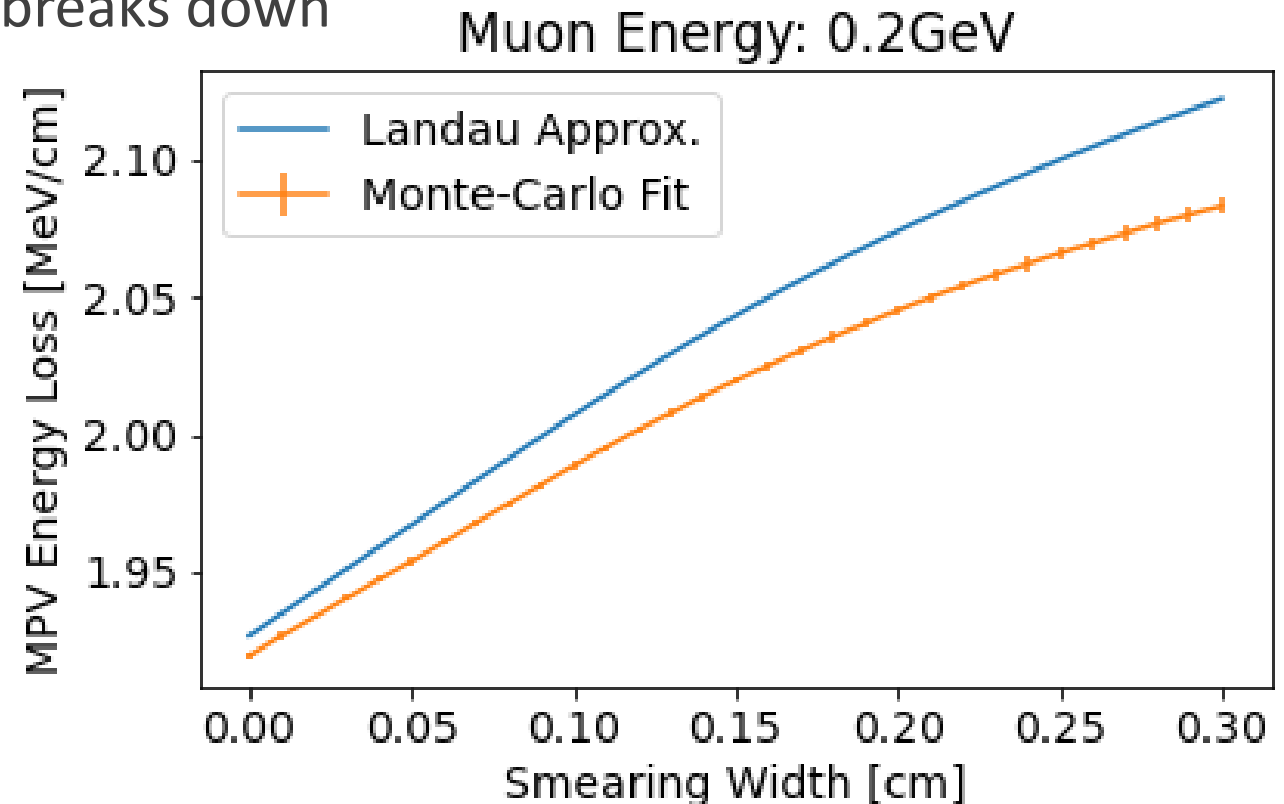
# Effect on the MPV at Large Thickness



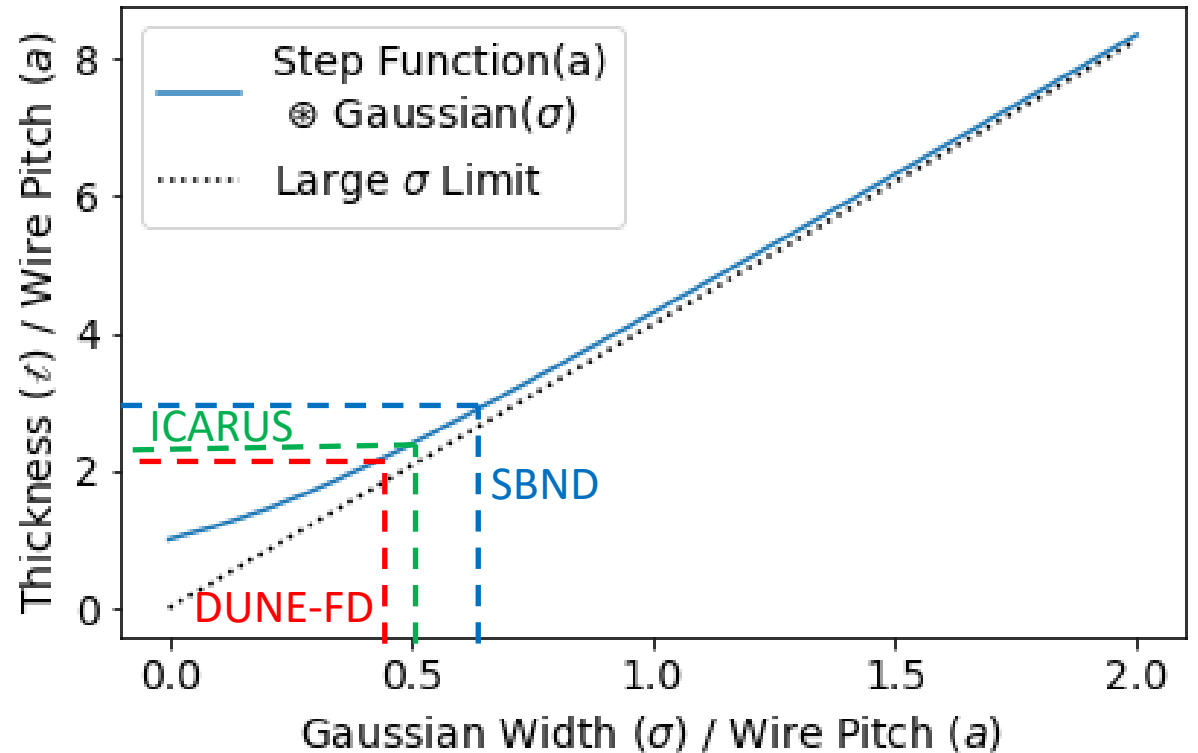
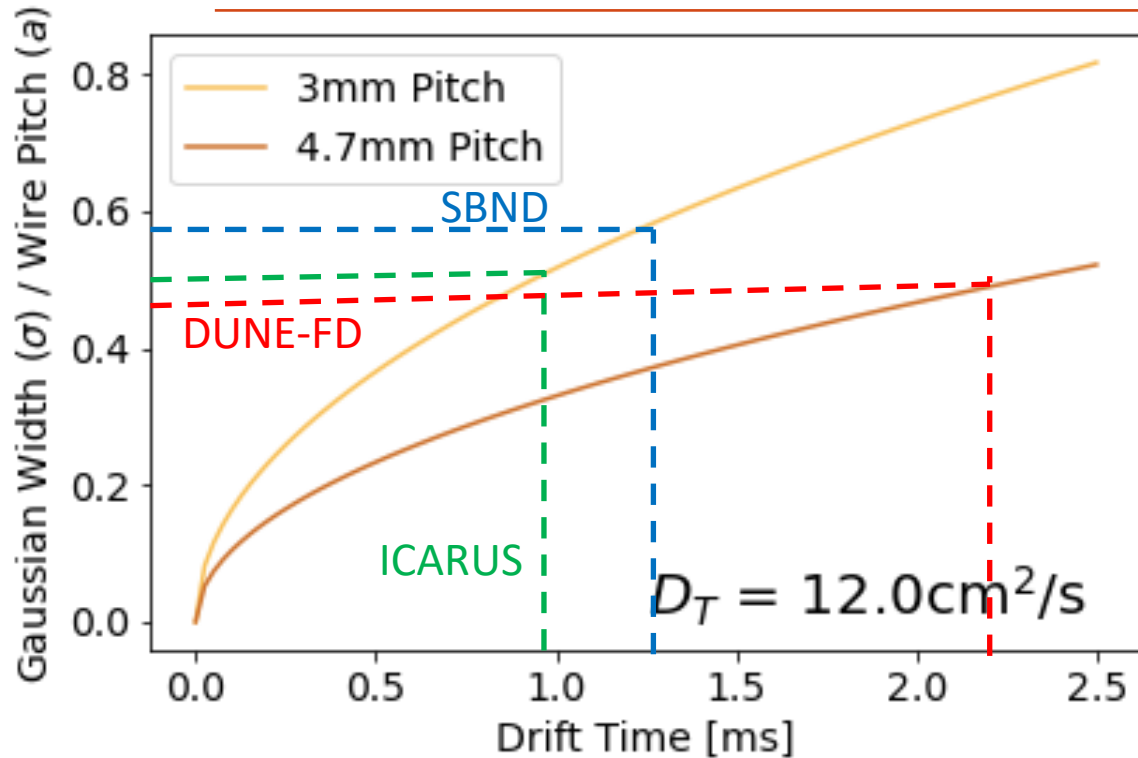
At high  $\kappa$ ,  
the distribution  
is less  
Landau-like



- Results from a toy MC of muon energy loss in LAr, for a wire spacing of 3mm
- At large thickness ( $\kappa$ ), the Landau approximation breaks down



# Impact of Diffusion on Thickness at Detectors



$$\sigma_T = \sqrt{2D_T t_{drift}}$$

$$p = \int w(x) dx$$

$$t = p e^{-\int w(x) \log w(x) dx} / p$$

$w(x)$ : convolution of the step-function wire and Gaussian diffusion

- At the cathode, the effect of diffusion about doubles the channel thickness relative to the wire pitch



# Impact on MPV in Relevant Detectors

## MPV Energy Loss for a 1 GeV Muon

Detector	Wire Pitch [mm]	Drift Time [ms]	Diffusion Const. $D_T$ [cm <sup>2</sup> /s]	MPV $dE/dx$ , No Diffusion [MeV/cm]	MPV $dE/dx$ at Cathode (Full Diff.) [MeV/cm]	Difference [%]
MicroBooNE [4]	3.00	2.33	5.85	1.69	1.79	5.9
ArgoNeuT [3]	4.00	0.295	12.0 (9.30)	1.72 (1.72)	1.76 (1.75)	2.3 (1.7)
ICARUS [5]	3.00	0.960	12.0 (9.30)	1.69 (1.69)	1.78 (1.77)	5.3 (4.7)
SBND [5]	3.00	1.28	12.0 (9.30)	1.69 (1.69)	1.79 (1.78)	5.9 (5.3)
DUNE-FD (SP) [7]	4.71	2.2	12.0 (9.30)	1.74 (1.74)	1.82 (1.81)	4.6 (4.0)

- This translates into a few percent change to the MPV  $dE/dx$  at the cathode