

# Neutrino Oscillation Measurement with KM3NeT/ORCA

on behalf of the KM3NeT Collaboration

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# The KM3NeT Collaboration

## Collaboration mainly European institutes

- also: Australia, Morocco, South Africa, U.A.E., Algeria and China

## Water Cherenkov detectors using sea water as target volume

- 2 detectors sharing 1 optical sensing technology

## Oscillation Research with Cosmics in the Abyss (ORCA)

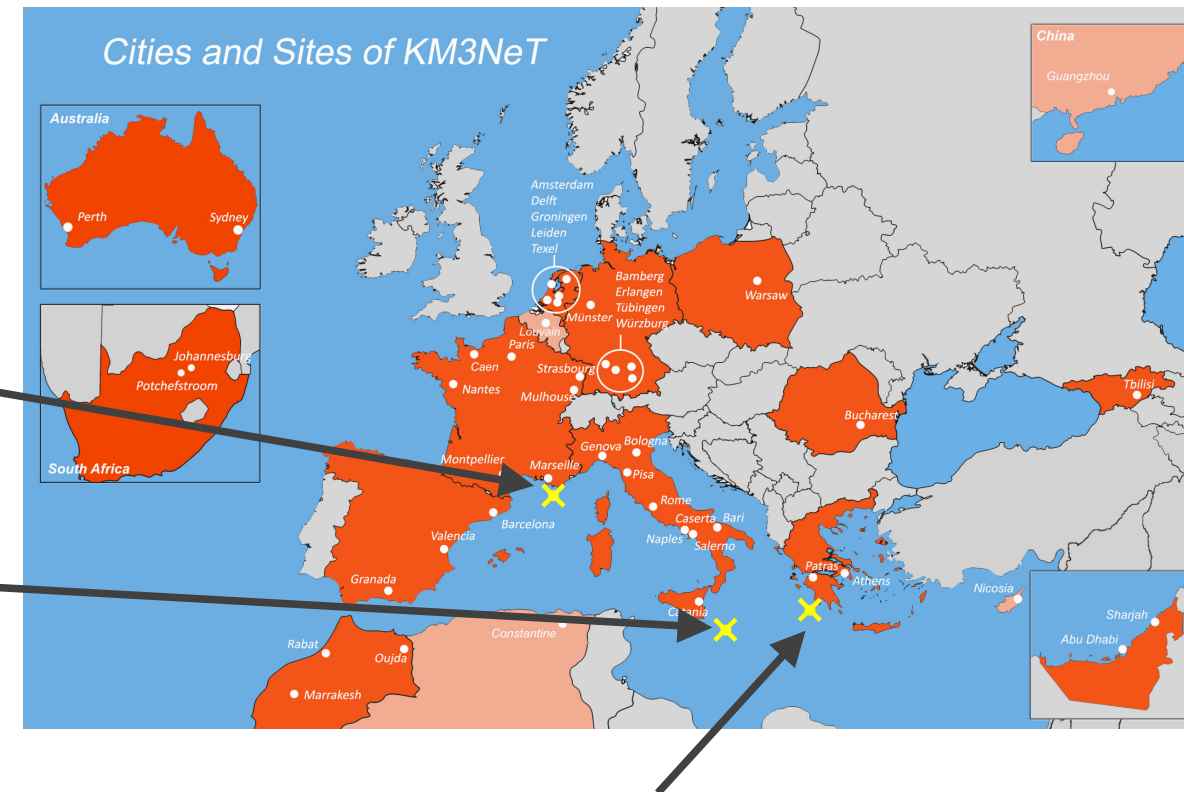
- Dense instrumentation for few GeV atmospheric neutrinos
- Determine neutrino mass hierarchy, oscillation parameters

## Astroparticle Research with Cosmics in the Abyss (ARCA)

- sparse instrumentation covering  $1\text{km}^3$  instrumented volume for TeV-PeV cosmic neutrinos
- High-energetic astrophysical neutrino sources, diffuse flux

## KM3NeT succeeding ANTARES

- data taking period: 2008 - 2022



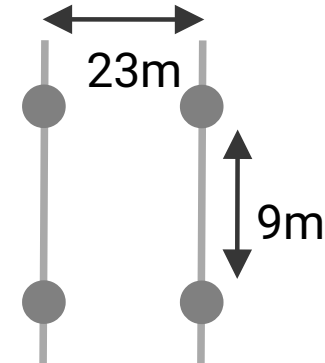
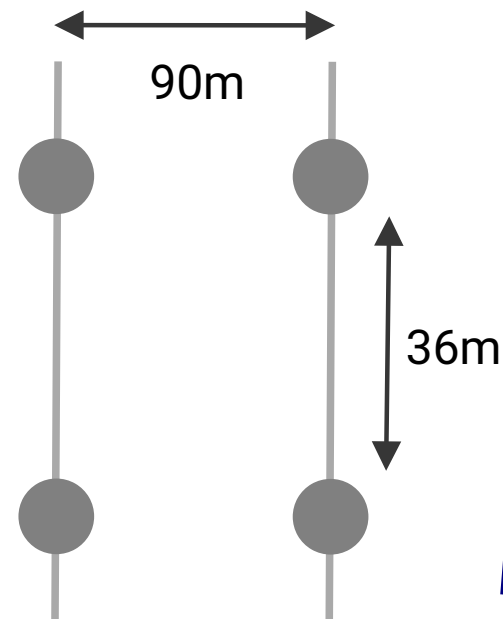
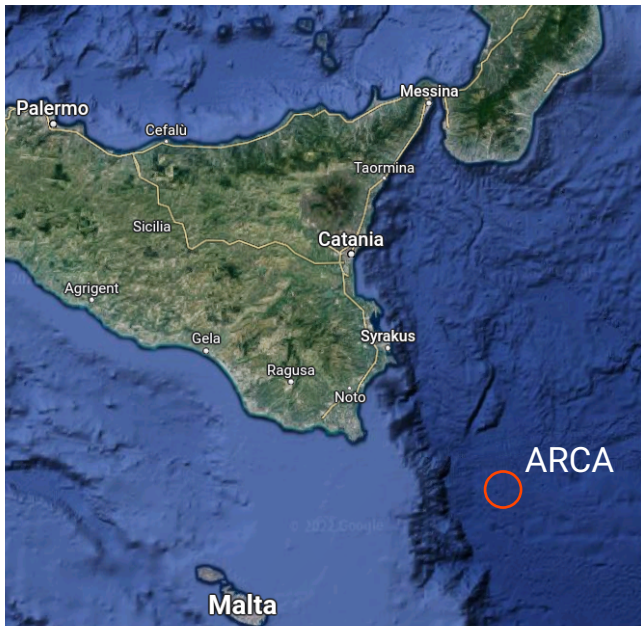
Possible location for 3<sup>rd</sup> detector site off the Greek coast

# The KM3NeT Detector Sites

Building Block =  
115 strings

## ARCA

- Final configuration: 2x Building Blocks, 128340 PMTs
- Positioned ~120km off Sicily at 3500m sea depth
- 19 strings deployed (6.3%)



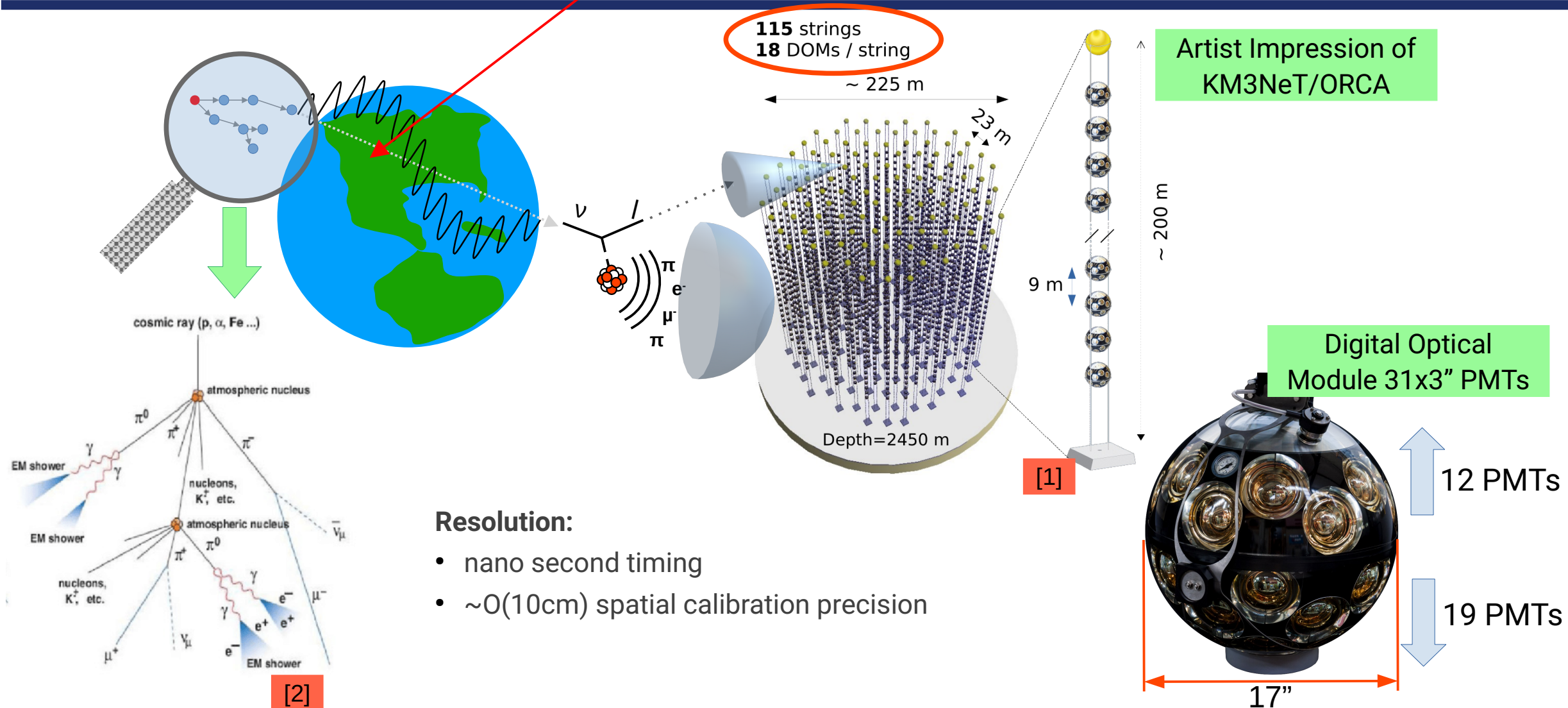
## ORCA

- Final configuration: 1x Building Block, 64170 PMTs
- 11 strings deployed (9.6%)
- Close to the former ANTARES site about 40km off southern french coast near Toulon
- Sea floor depth of the ORCA site ~2450m



# KM3NeT/ORCA in a nutshell

You are here



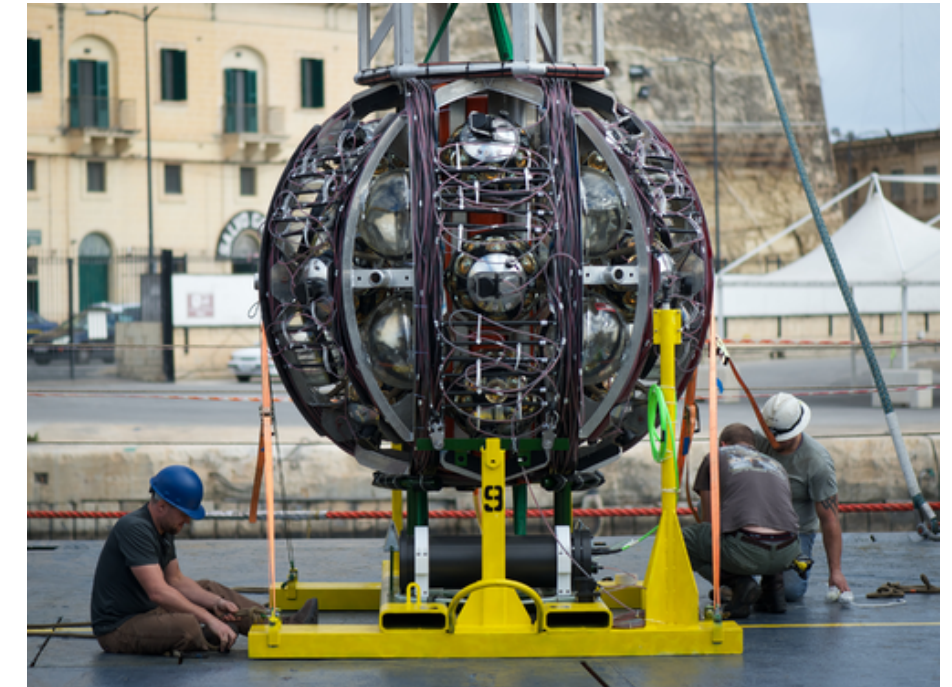
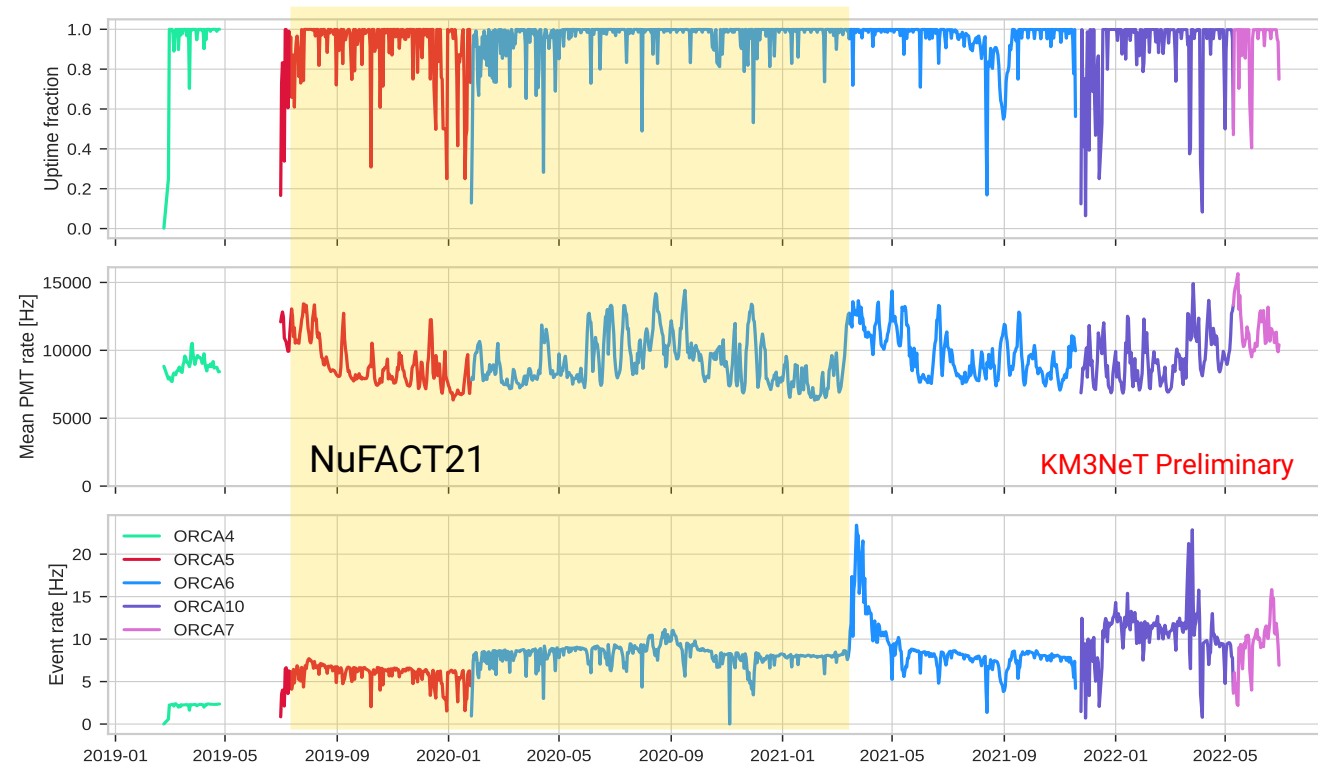


# KM3NeT/ORCA current status

## Detector Timeline (ORCA)

- Last NuFACT: 6 strings running
- Now: 11 strings deployed → currently 7 running

## Some numbers:

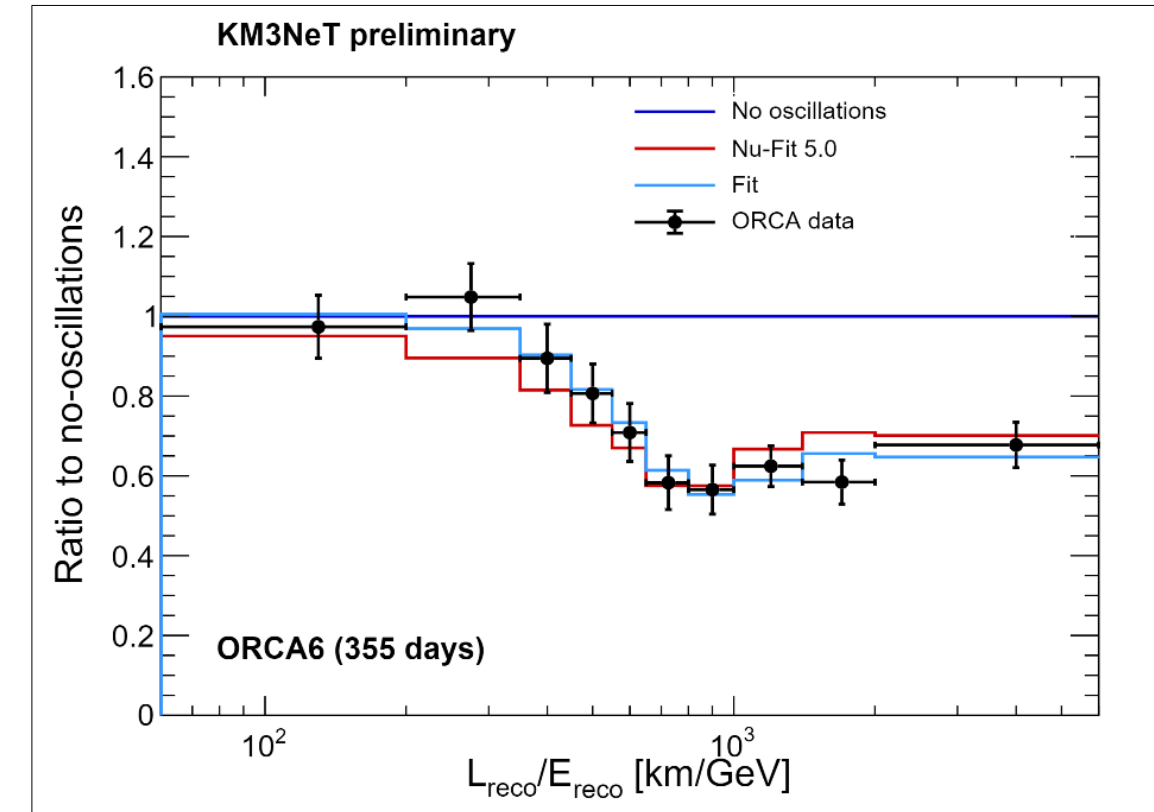


DU furled and ready for the deployment

## Oscillation measurement

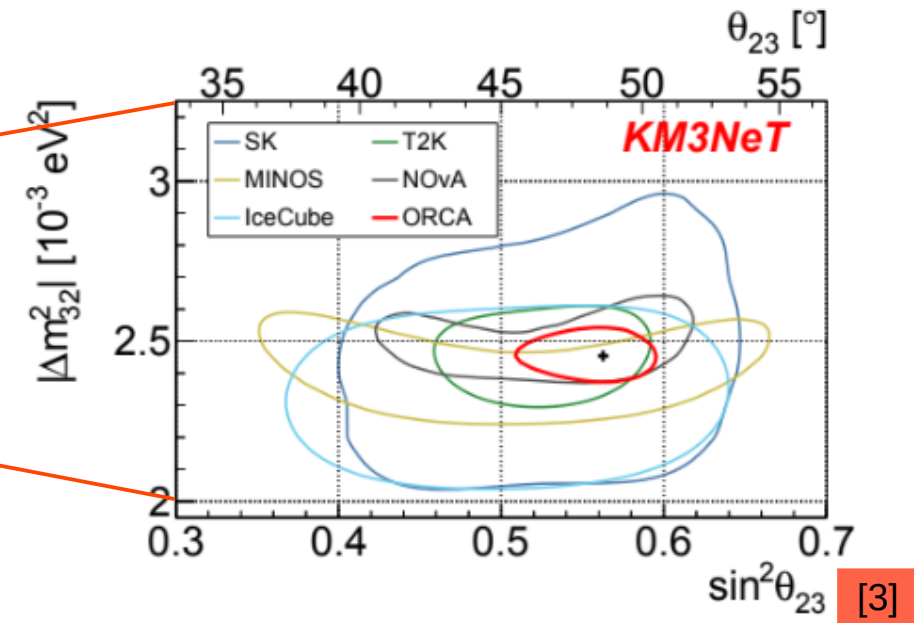
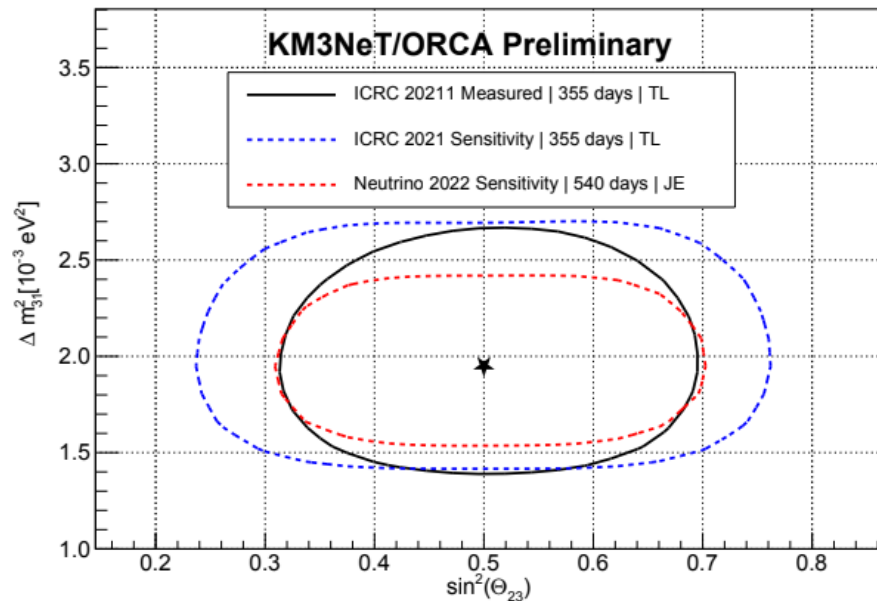
- Dataset contains 355 days with 6 ORCA strings
- Compared model:  
Atmospheric flux model → HKKM14  
Oscillation parameters: → NO & NuFit 5.0 values
- Oscillation pattern matches the prediction very well
- Analysis parameters:

Parameter	Treatment Free/Fixed/Prior
$\theta_{12}$ [deg]	Fixed
$\theta_{13}$ [deg]	Fixed
$\theta_{23}$ [deg]	Free
$\Delta m_{31}^2$ [ $10^{-3}$ GeV <sup>2</sup> ]	Free
$\Delta m_{21}^2$ [ $10^{-5}$ GeV <sup>2</sup> ]	Fixed
$\delta_{CP}$ [deg]	Fixed
Normalisation	Free
Spectral index	Prior: 10%
$n_{\nu_{\mu}}/n_{\nu_{\text{horiz}}}$	Prior: 7%
$n_{\nu_{\mu}}/n_{\nu_{\bar{\mu}}}$	Prior: 10%
$n_{\nu_e}/n_{\nu_{\bar{e}}}$	Prior: 10%
$n_{\nu_{\mu}}/n_{\nu_e}$	Prior: 3%
$n^{NC}$	Prior: 10%
$n_{\tau}^{CC}$	Prior: 20%
Energy scale	Prior: 10%



## Updated oscillation parameter measurement with 6 strings

- Livetime increased from 355 days to 540 days
- Improved selection & particle identification  
→ together: Sample increased by factor 4
- (Unblinding & Measurement update about to come)

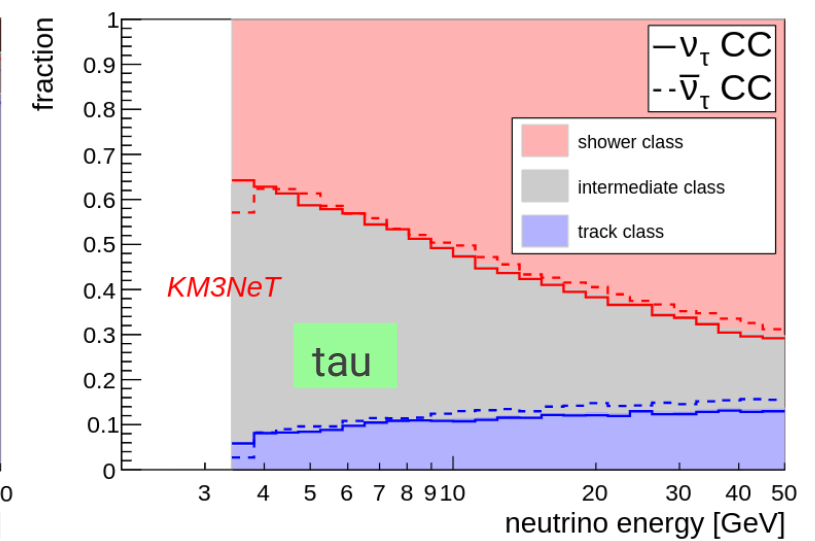
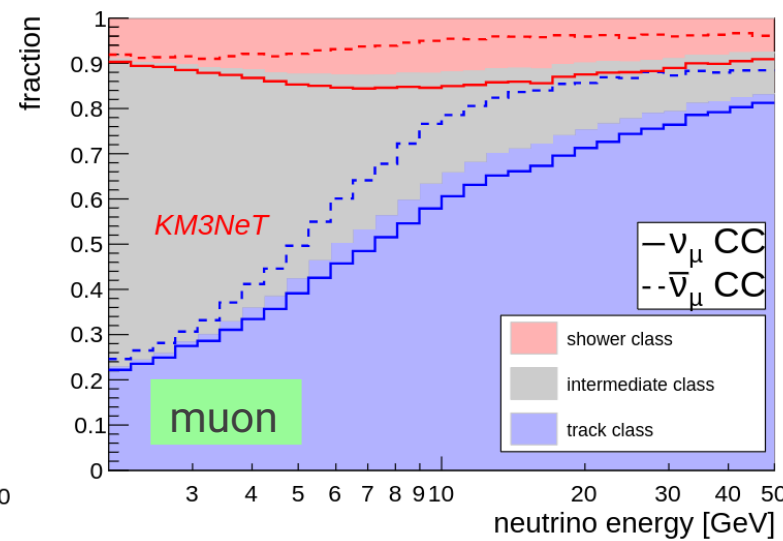
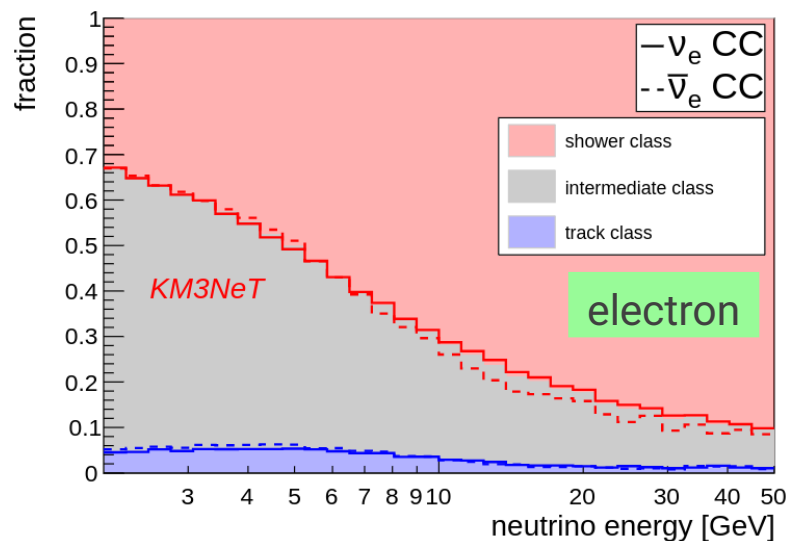


ORCA115  
Prediction



## Triggered events are reconstructed with different algorithms

- Reconstruction algorithms differ in the event geometry hypothesis, i.e. track and shower
- Events are classified by multiple features using random decision forests (RDF):  
**Features:** likelihood, reconstructed energy & direction, number of triggered hits, etc.
- $\nu / \mu$  classification  $\rightarrow$  muon score
- noise classification  $\rightarrow$  noise score
- track/shower classification  $\rightarrow$  track / intermediate / shower

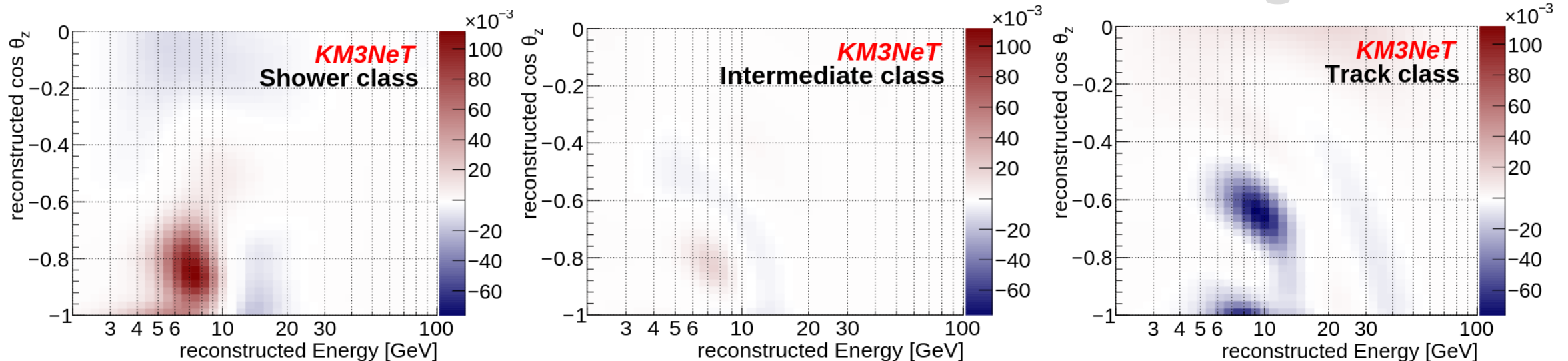


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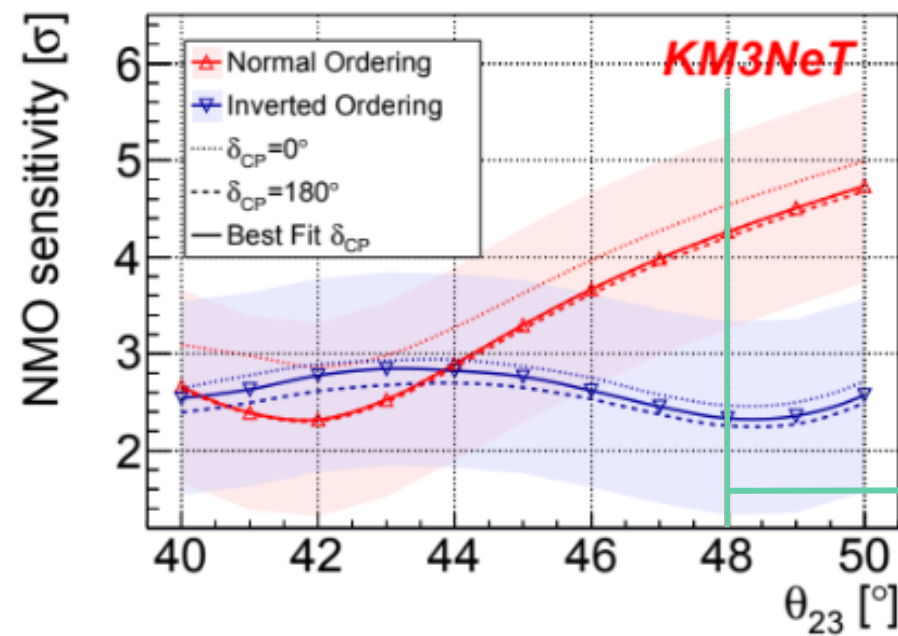
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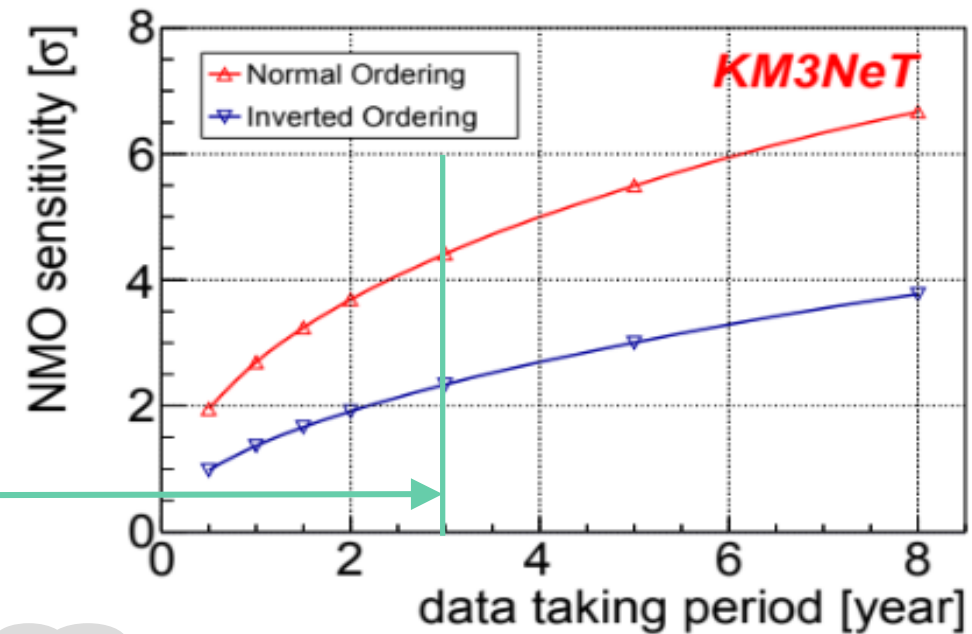
# KM3NeT/ORCA full detector estimated sensitivity

## ORCA Mass Ordering Sensitivity

- assumed data taking time: 3 years



3yrs



@  $\theta_{23} = 48^\circ$

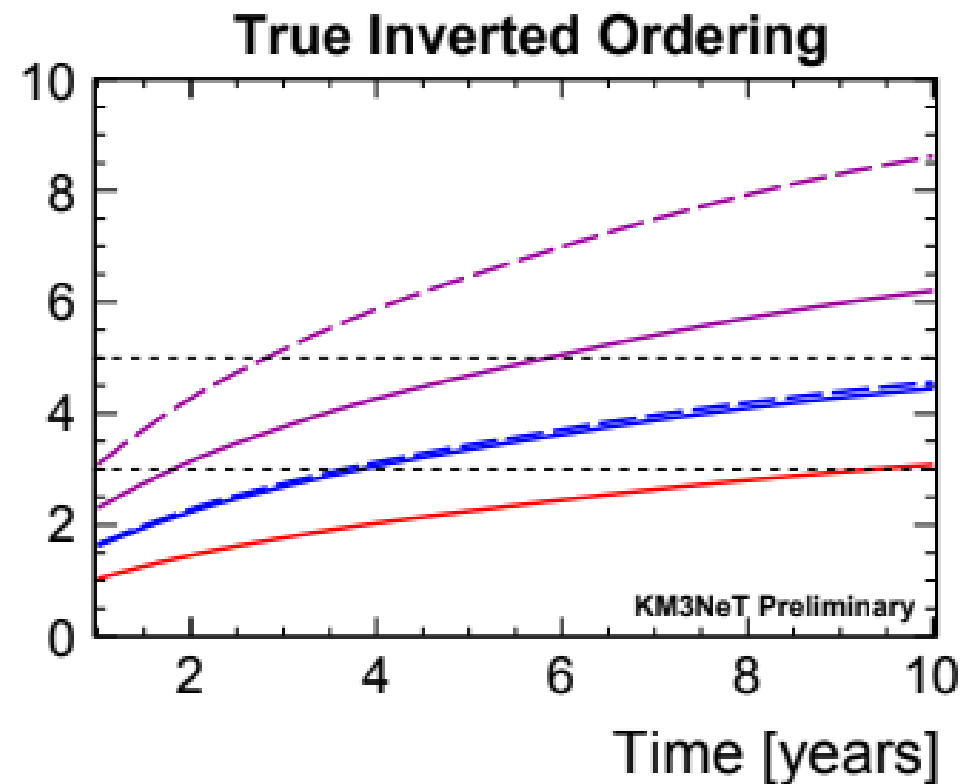
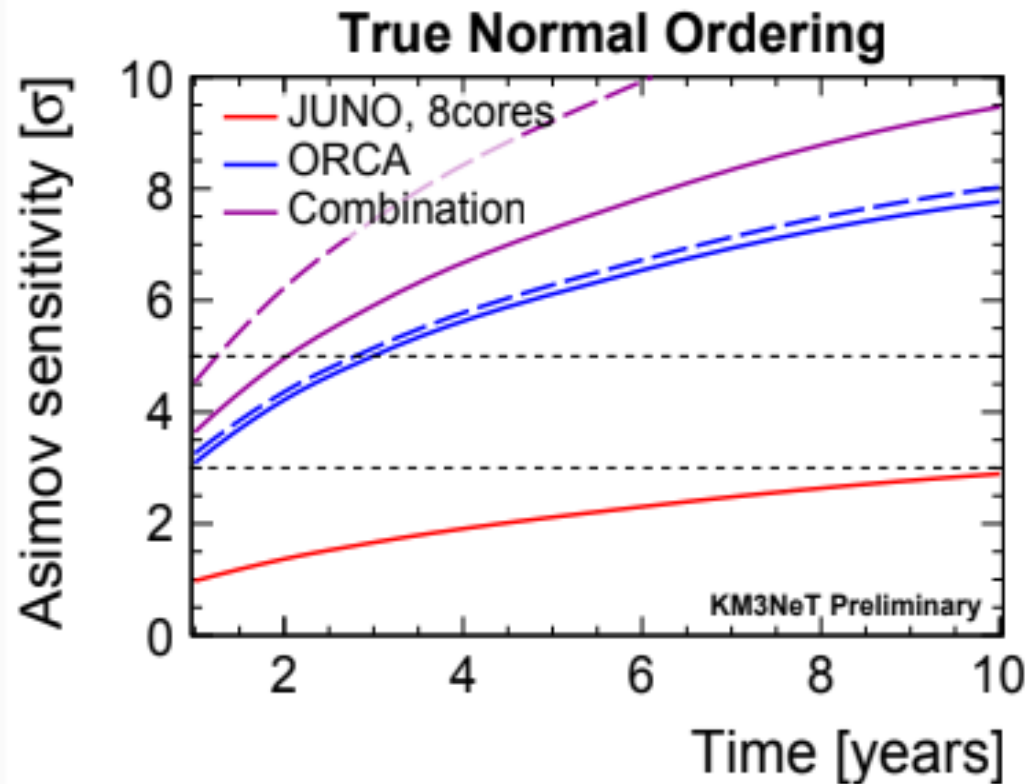
[3]



# NMH sensitivity combined with JUNO

## Massive increase of sensitivity by combining ORCA & JUNO

- Analysis approach:  $\Delta m_{31}$  disagreement on wrong NMO hypothesis
- Sensitivity boost compared to full ORCA only:  $4\sigma \rightarrow 6\sigma$  @3yrs, normal ordering and  $\theta_{23} = 48^\circ$



## 4<sup>th</sup> Sterile Flavour Properties

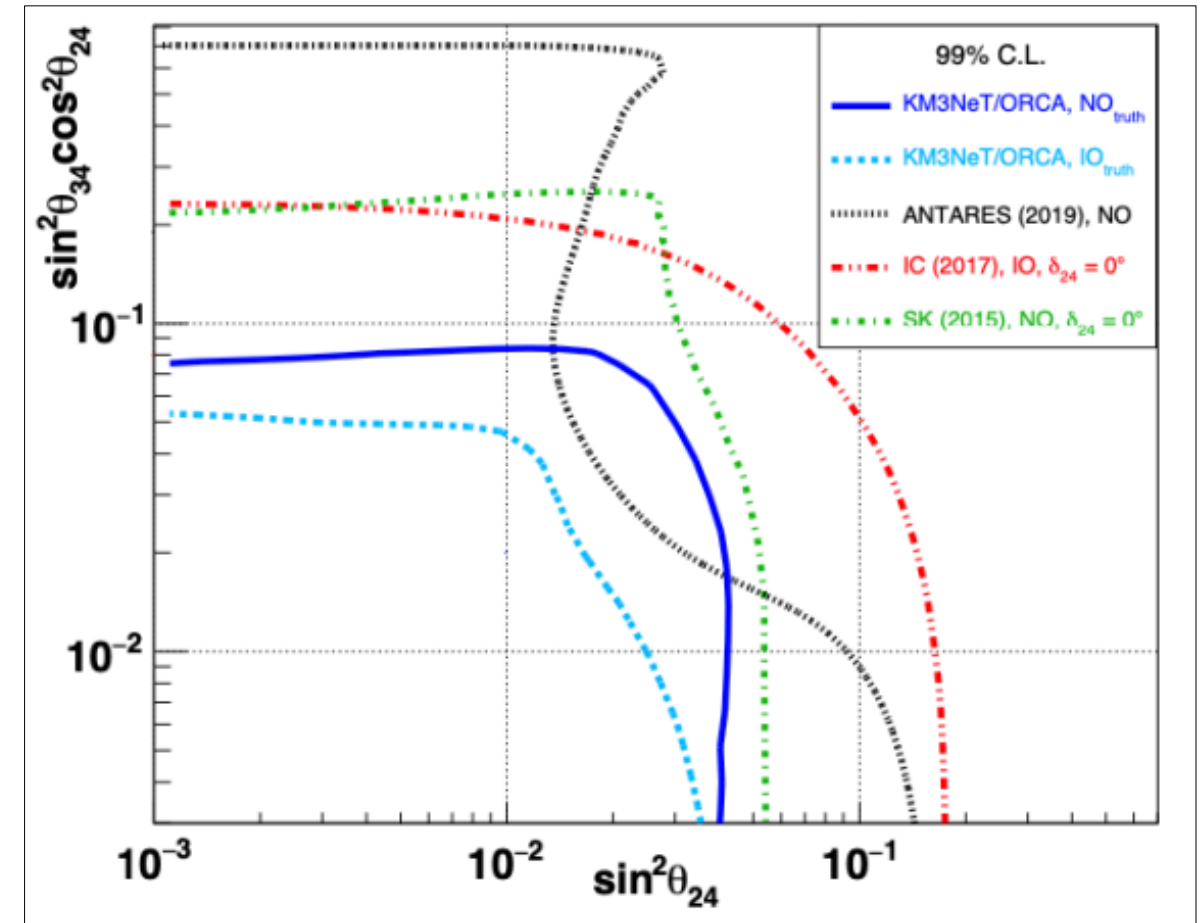
- 4 neutrino flavour yields:
  - 3 additional **mixing angles**
  - 2 additional **phases**
  - (1 additional **mass**)

**Given sensitivity for large  $\Delta m_{41}$  limit**

- assumed data taking time 3 years

**Free phase  $\delta_{24}$  worsens the sensitivity**

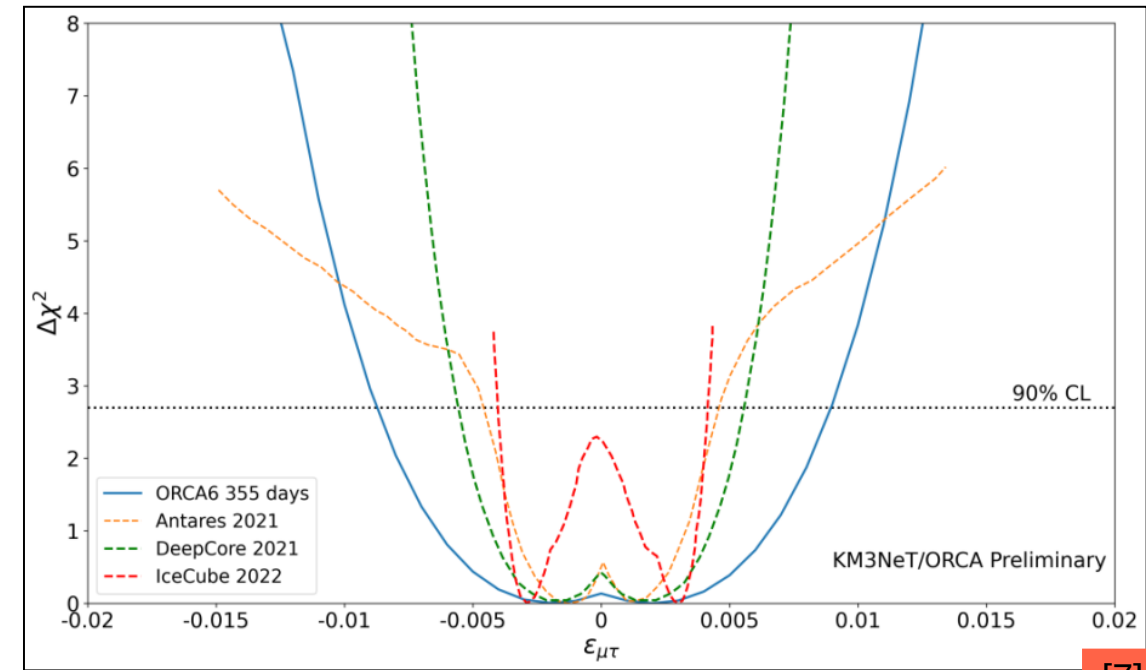
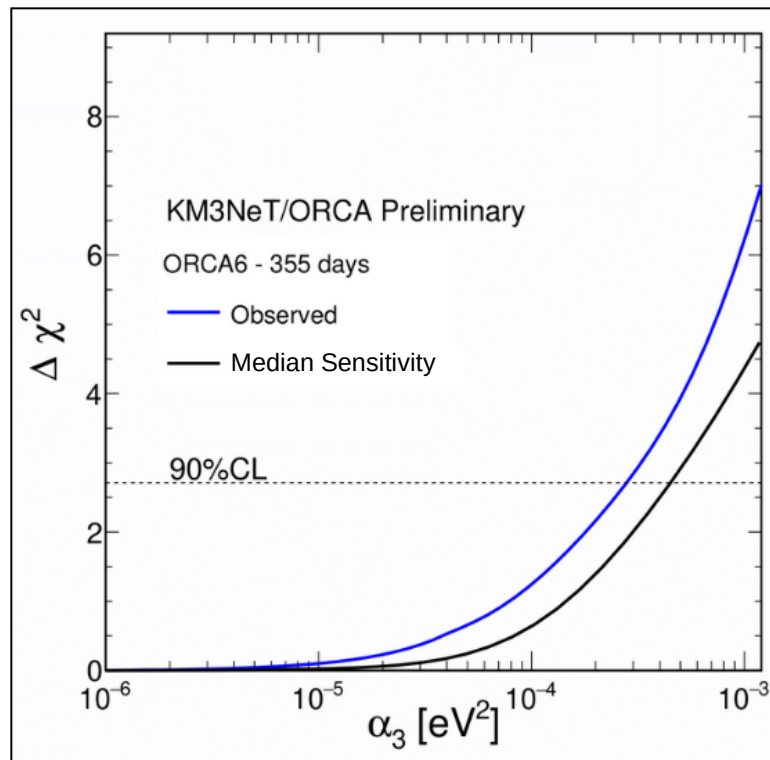
- here phase for IC and SK is assumed to be  $\delta_{24} = 0^\circ$



[6]

## Constraining non-standard interactions

- 1 year sensitivity about to be competitive to other experiments
- ORCA prediction for 355 days livetime in 6 string configuration



[7]

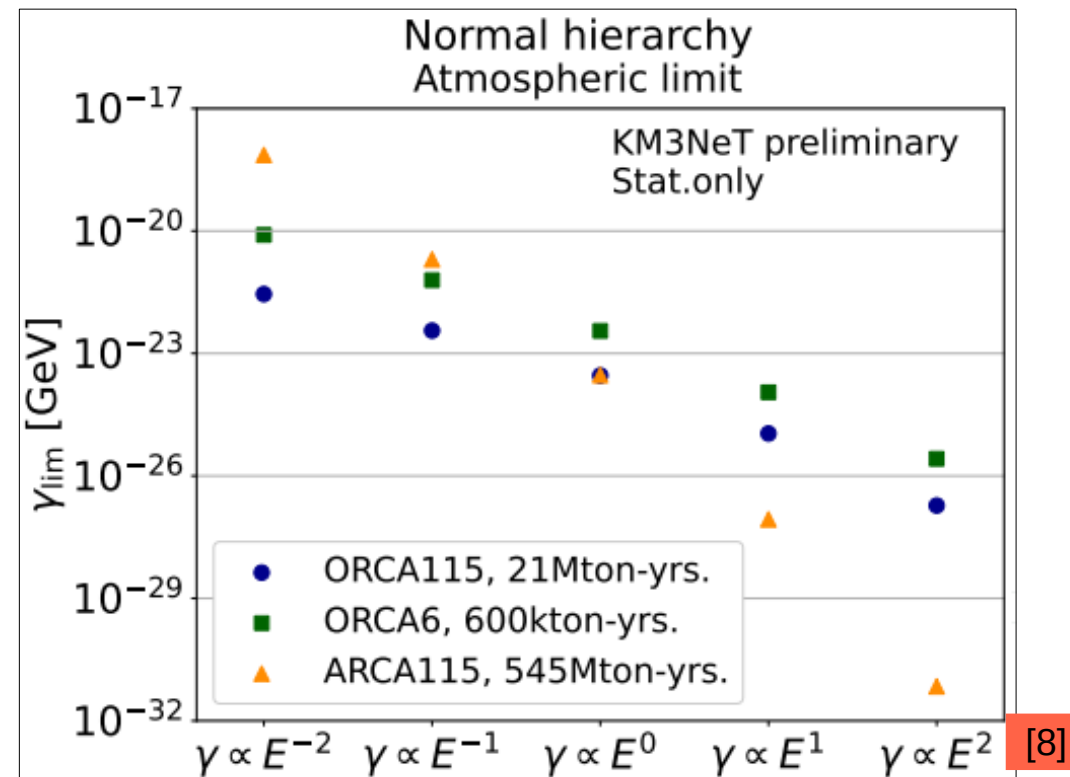
## Neutrino decay / livetime

- Proof of principle
- Full ORCA detector will be world-leading



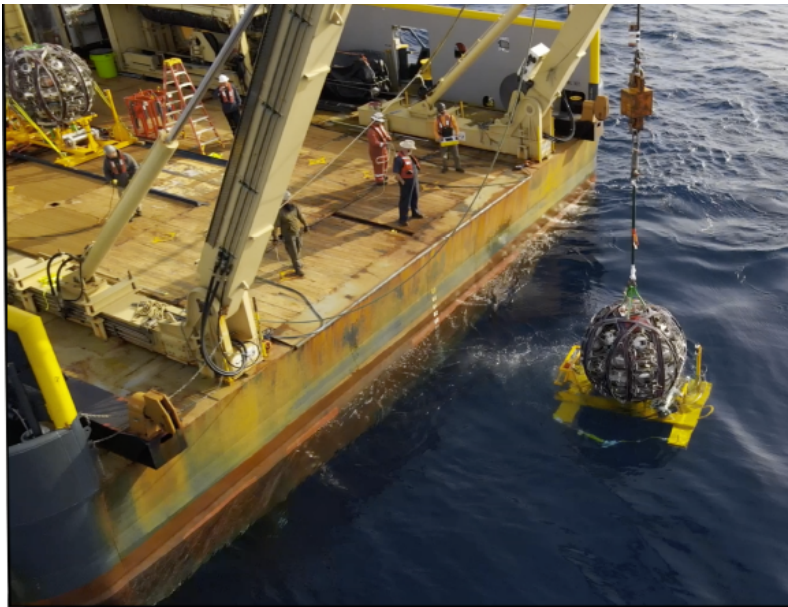
## Preliminary sensitivities

- Constraints on a wide range of power law indices, i.e.  $[-2, 2]$
- ORCA dominates negative indices  $\rightarrow$  lower energies more relevant
- ARCA dominates positive indices  $\rightarrow$  vice versa as higher energies are more relevant



## ORCA Detector Status

- First oscillation results already obtained  
→ e.g. NSI already competitive
- Data taking is on-going while detector is growing (currently 11 strings deployed)
- Refinement of the data analysis causes sample size growth larger than livetime growth



## Perspectives

- Upon completion KM3NeT/ORCA will be one of the leading experiments with respect to NMH & oscillation parameter measurement with atmospheric neutrinos
- Construction ramping up on the way to the full detector

# Thank You!

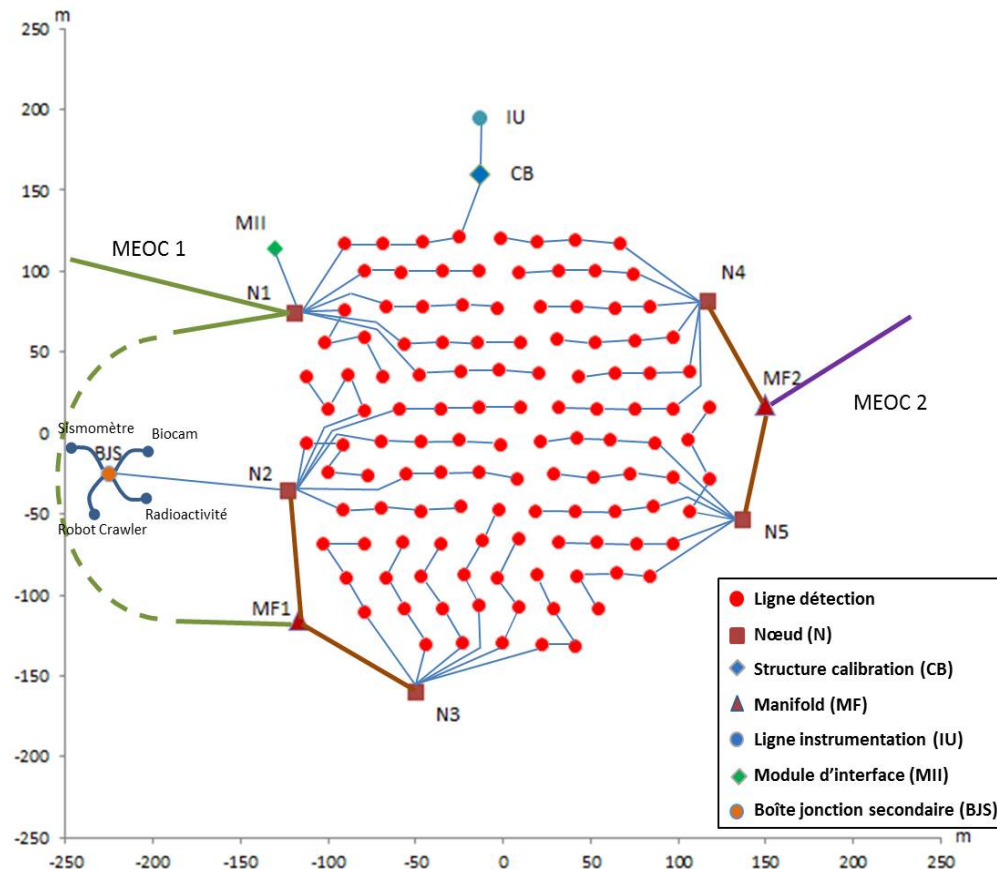
- [1] A. V. Akindinov and others, “Letter of Interest for a Neutrino Beam from Protvino to KM3NeT/ORCA”, Eur. Phys. J. C, vol. 79, no. 9, p. 758, 2019, doi: 10.1140/epjc/s10052-019-7259-5.
- [2] A. López-Oramas, Multi-year Campaign of the Gamma-Ray Binary LS I +61° 303 and Search for VHE Emission from Gamma-Ray Binary Candidates with the MAGIC Telescopes. 2015. doi: 10.13140/RG.2.1.4140.4969.
- [3] S. Aiello and others, “Determining the neutrino mass ordering and oscillation parameters with KM3NeT/ORCA”, Eur. Phys. J. C, vol. 82, no. 1, p. 26, 2022, doi: 10.1140/epjc/s10052-021-09893-0.
- [4] L. Nauta and others, “First neutrino oscillation measurement in KM3NeT/ORCA”, PoS, vol. ICRC2021, p. 1123, 2021, doi: 10.22323/1.395.1123.
- [5] N. Chau, J. P. Athayde Marcondes de André, V. Van Elewyck, A. Kouchner, L. Kalousis, and M. Dracos, “Neutrino mass ordering determination through combined analysis with JUNO and KM3NeT/ORCA”, JINST, vol. 16, no. 11, p. C11007, 2021, doi: 10.1088/1748-0221/16/11/C11007.
- [6] S. Aiello and others, “Sensitivity to light sterile neutrino mixing parameters with KM3NeT/ORCA”, JHEP, vol. 10, p. 180, 2021, doi: 10.1007/JHEP10(2021)180.
- [7] J. Manczak, “First limits on neutrino non-standard interactions with KM3NeT/ORCA6.” Zenodo, 2022. doi: 10.5281/zenodo.6785232.
- [8] Lessing, N. (2022). Sensitivity to quantum decoherence in neutrinooscillations with KM3NeT [Masterthesis]. Friedrich-Alexander-Universität Erlangen-Nürnberg.



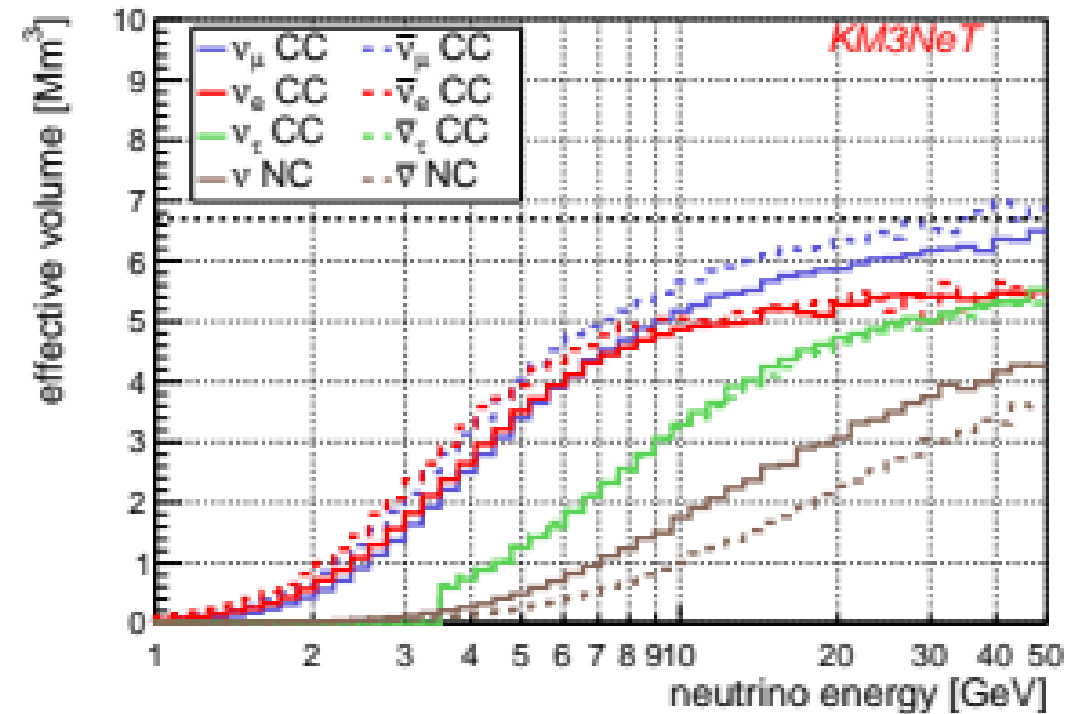


# Backup

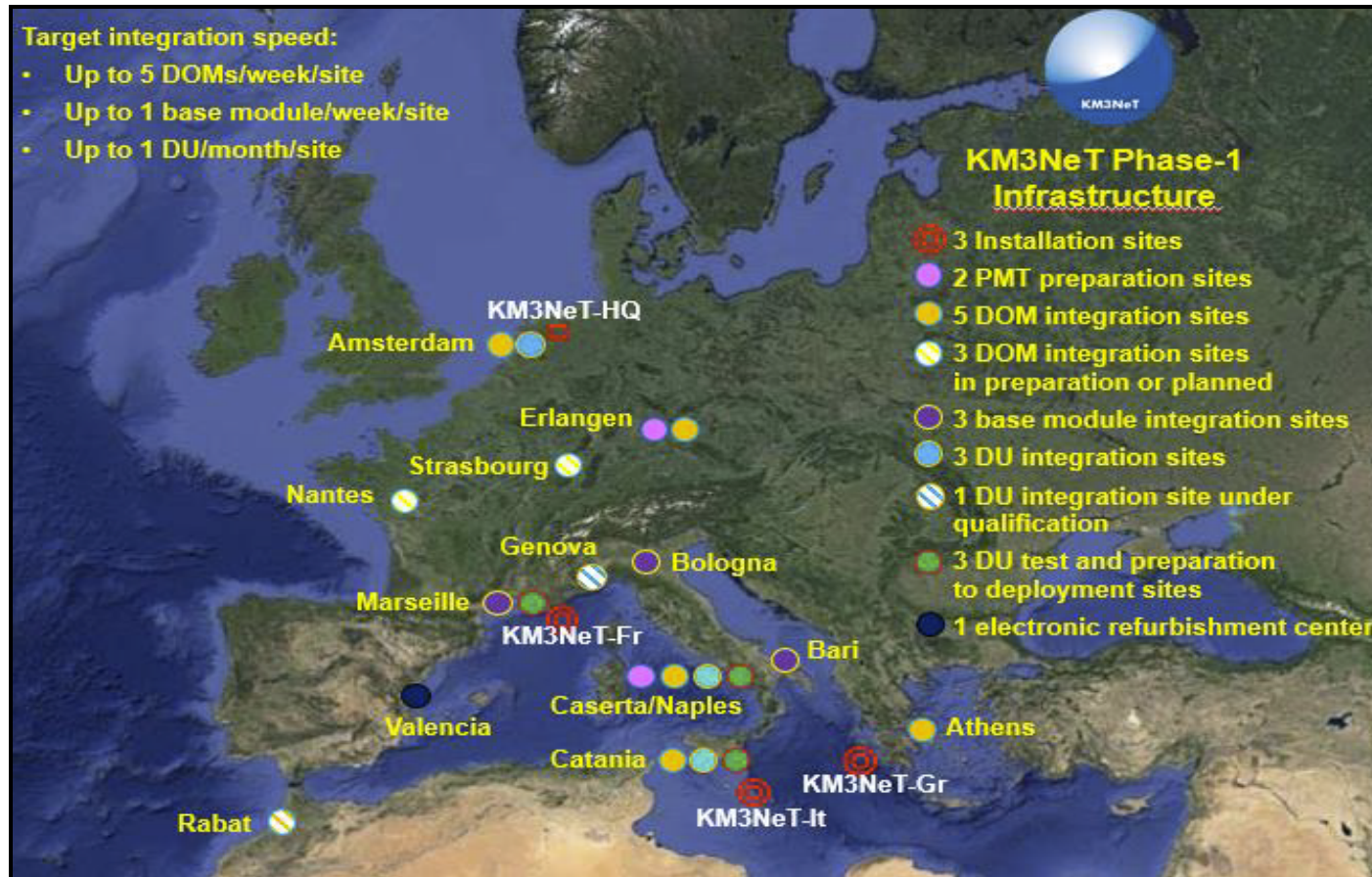
## Planned ORCA layout and components



## Estimated effective Volume

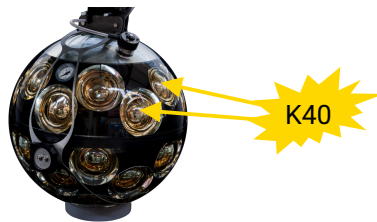


## Infrastructure and production sites

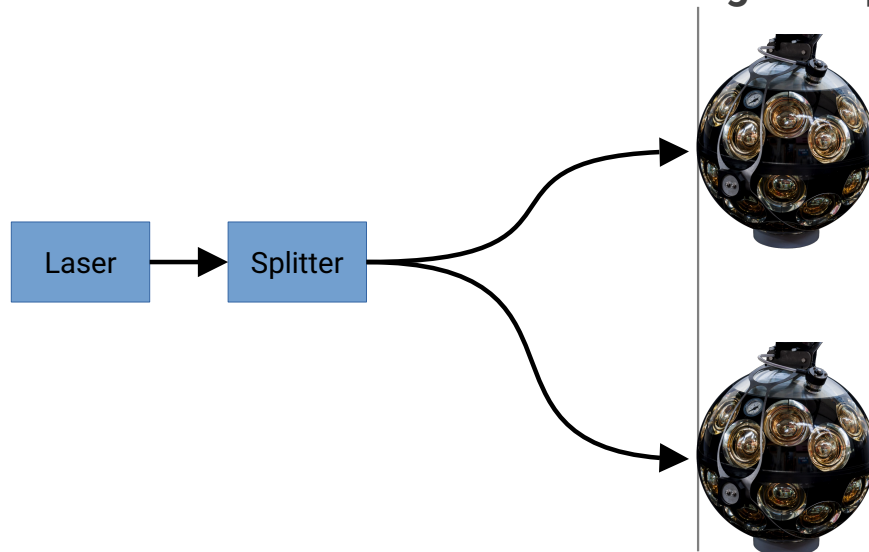


# KM3NeT/ORCA time calibration

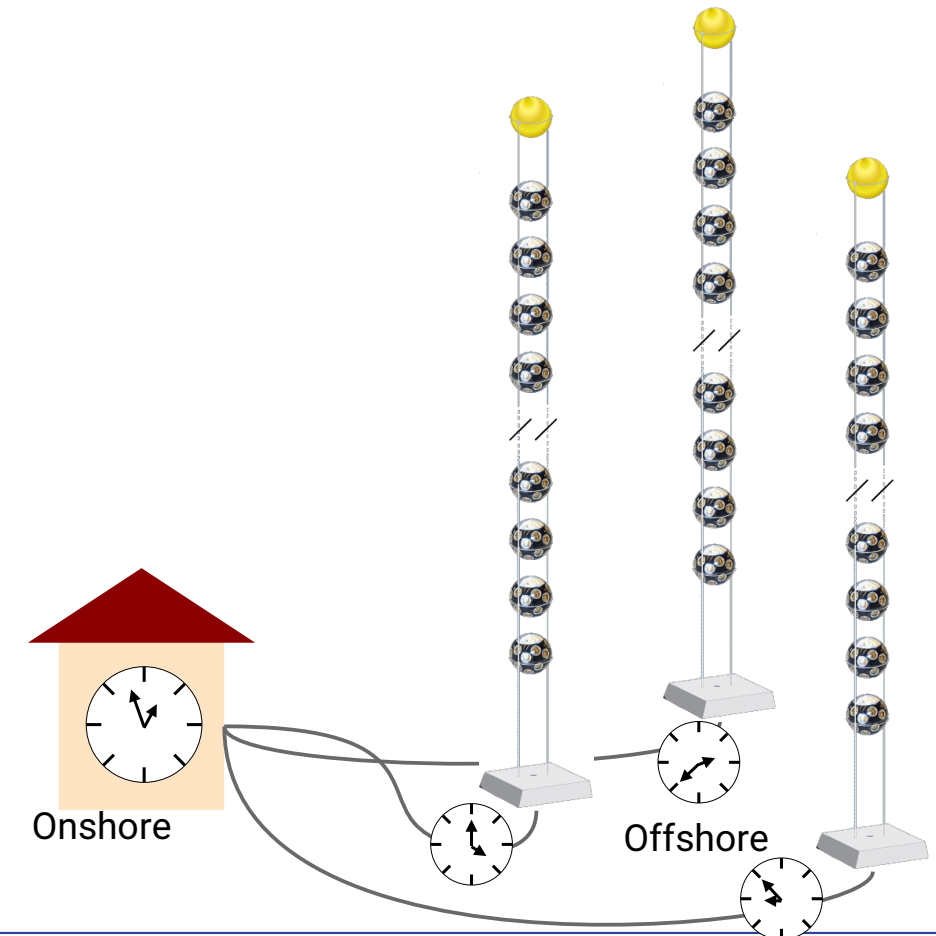
Inter PMT calibration → Insitu using Potassium 40



Inter DOM calibration → On-shore using laser pulses

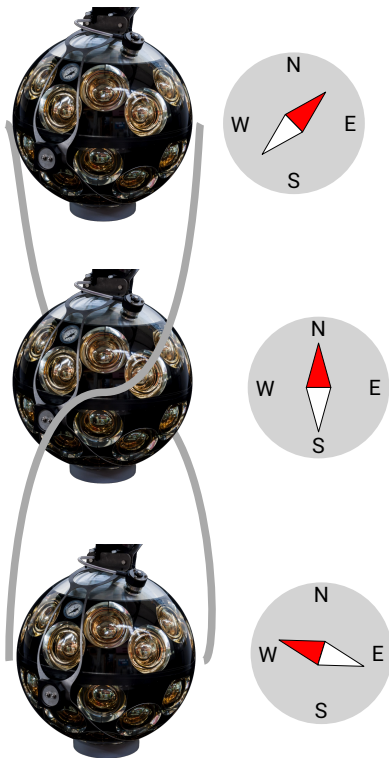


Inter string calibration → Synchronisation with network master clock

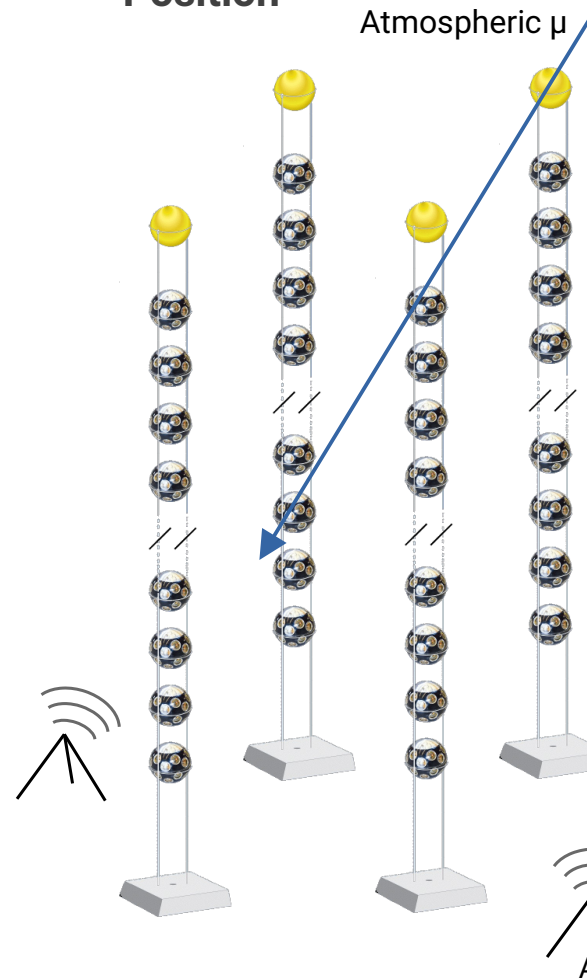


# KM3NeT/ORCA position calibration

## Orientation

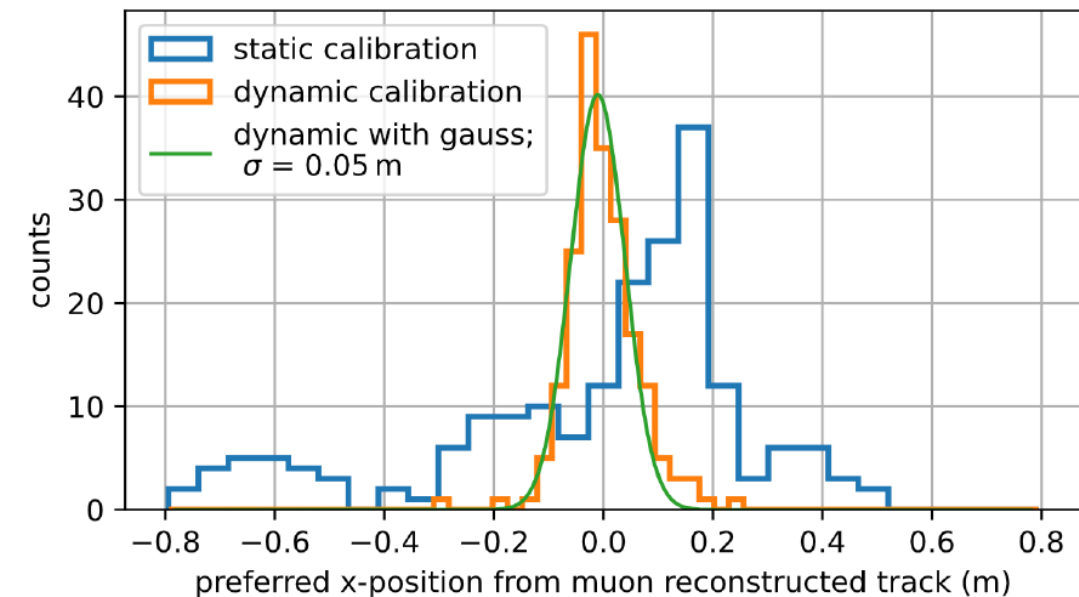


## Position



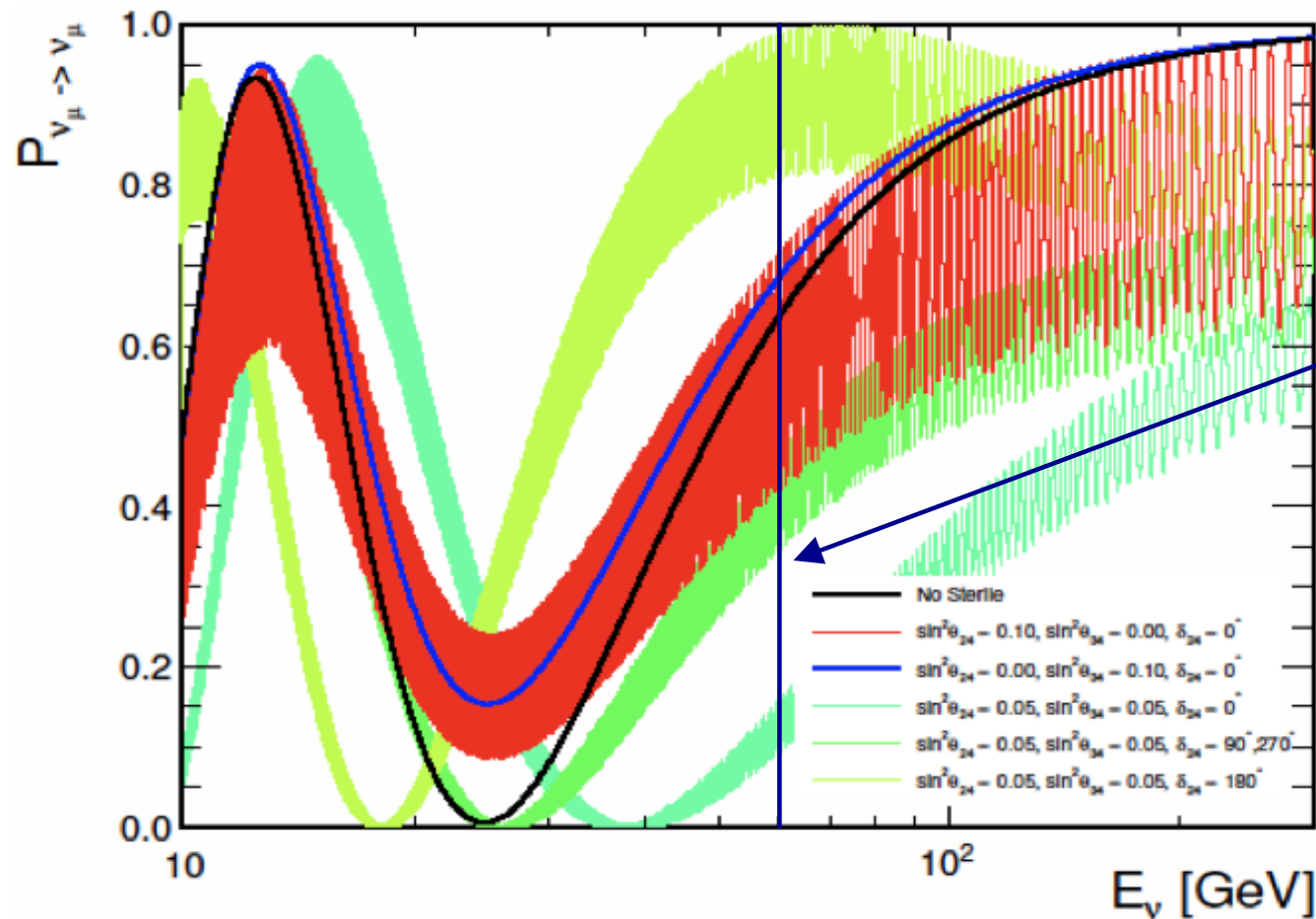
## Calibration result:

- Calibrate with a acoustic calibration and check muon reconstruction time residuals



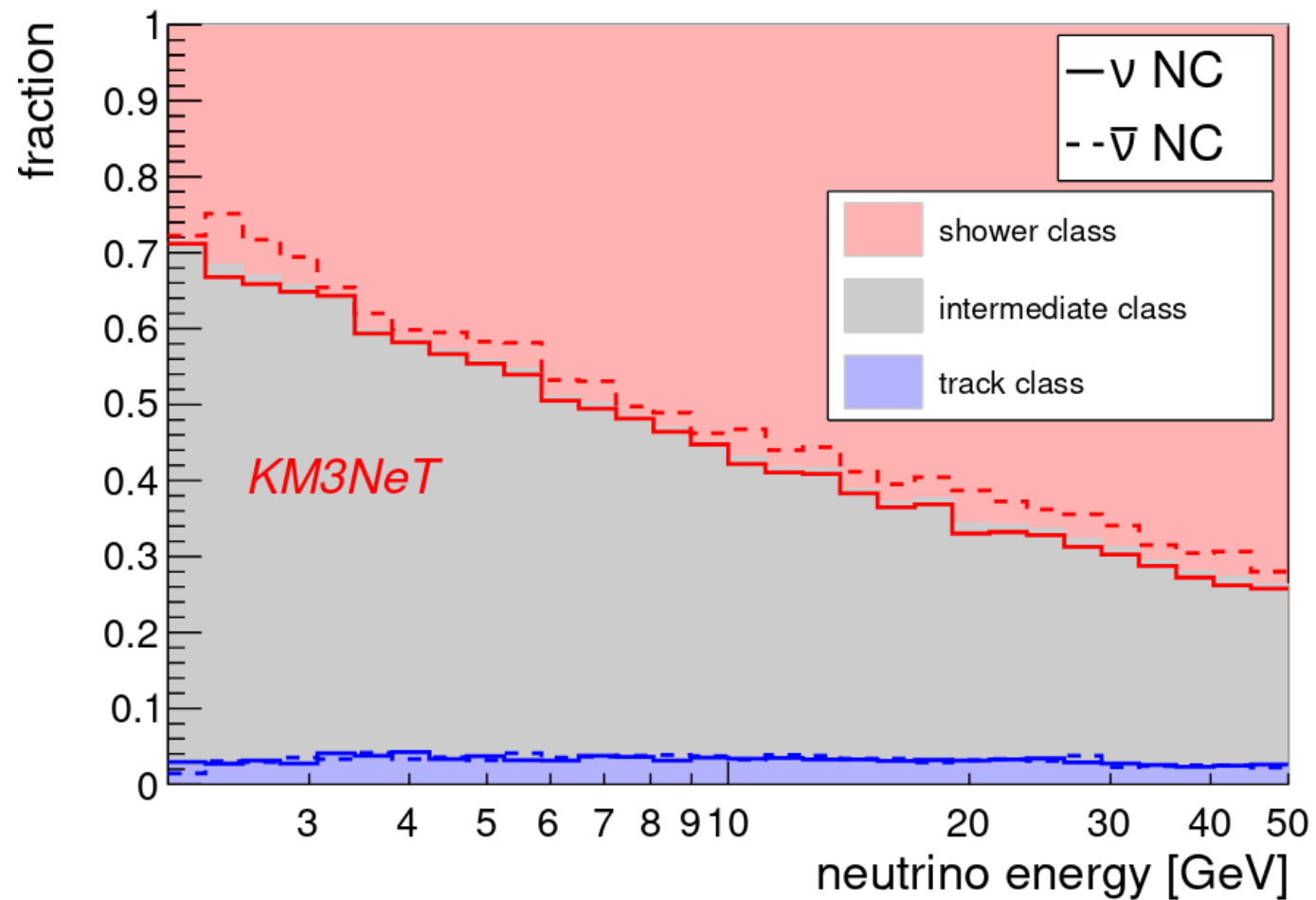


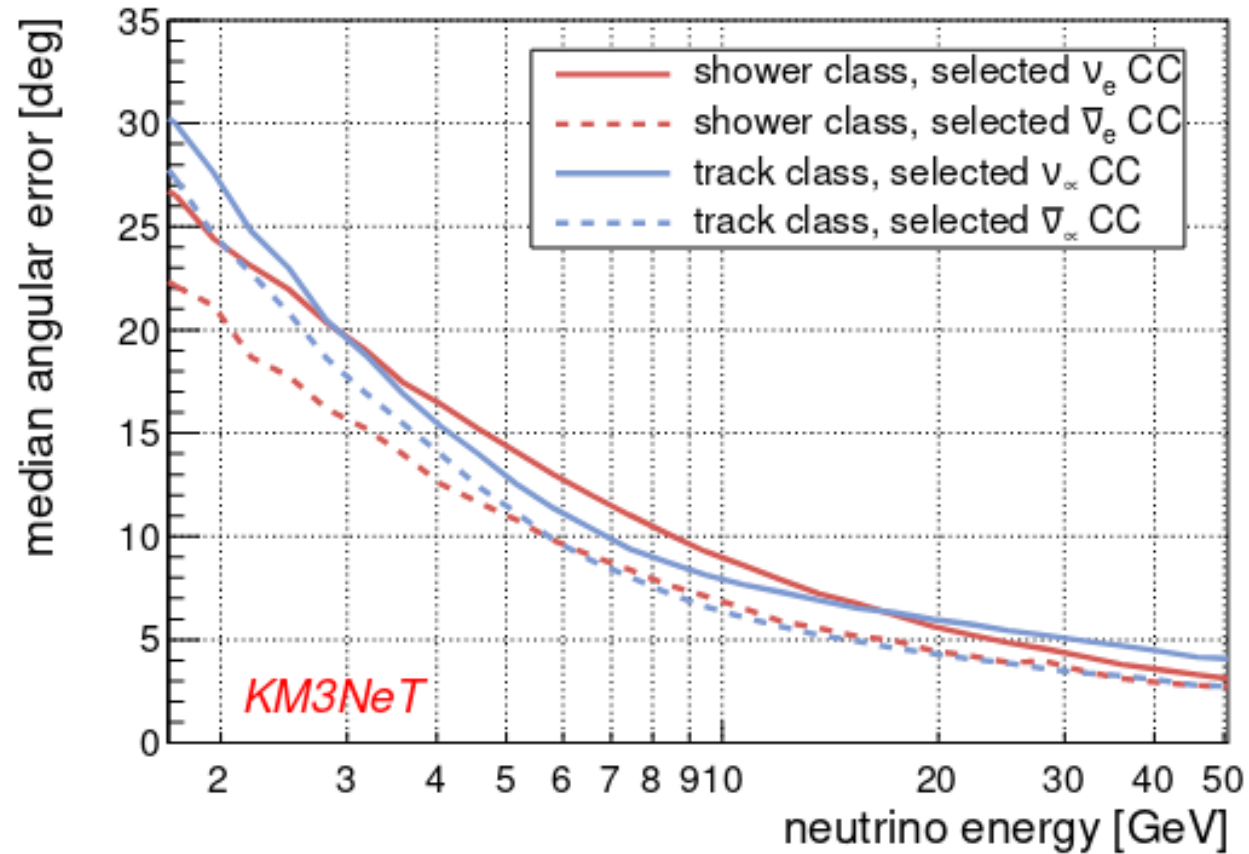
Oscillation probability  $P(\mu \rightarrow \mu)$ :



## ANTARES Sensitivity:

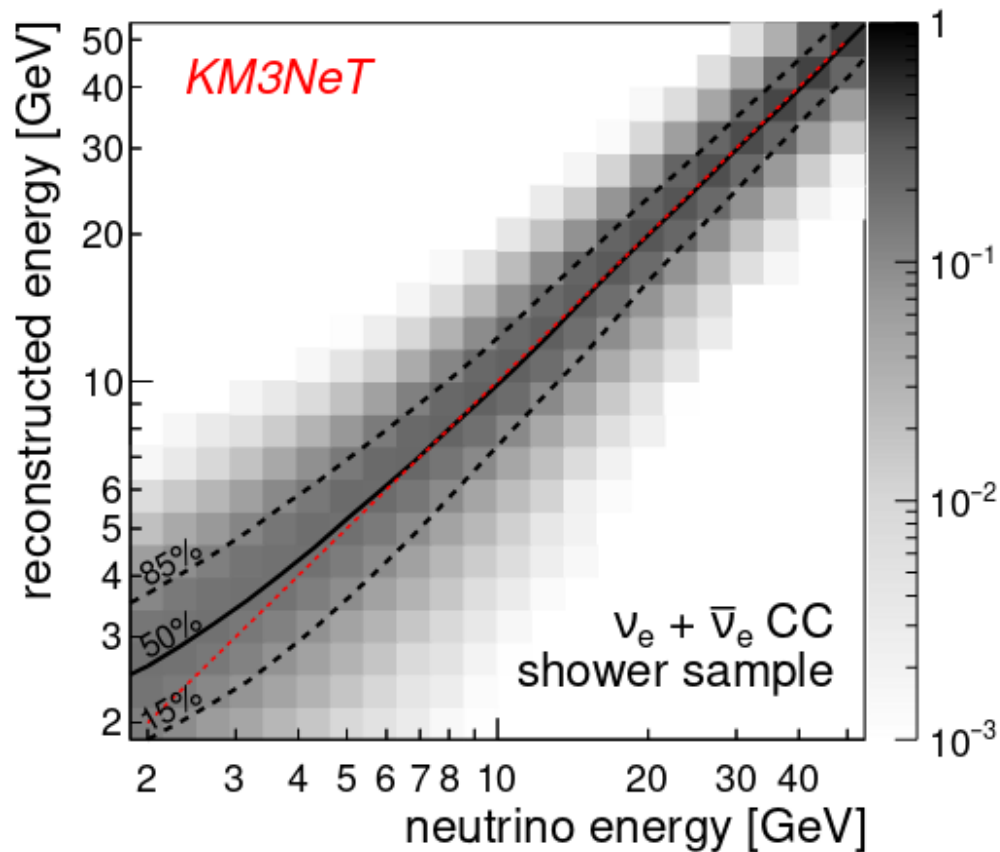
- IceCube dataset only up to  $\sim 60$  GeV while ANTARES dataset contains up to over 100 GeV
- ANTARES sensitivity decreases if dataset is limited to the IceCube used energy range





# Energy Reconstruction

Shower:



Track:

