

Phenomenology and search strategies

Stefania Gori
UC Santa Cruz



Accelerator-based Dark Sector Searches Agora
April 22, 2022

Outline of the talk

1. “Dark sectors at high-intensities”. Activities of the RF6 Snowmass topical group
2. **Search strategies** for dark sectors at or below the GeV scale
3. **Big idea solicited white papers**. Phenomenology and present / future experiments

Most of the plots are preliminary
Comments are welcome!

Apologies if I do not mention all the experiments

RF6: dark sectors at high intensities

Conveners: Stefania Gori (sgori@ucsc.edu), Mike Williams (mwill@mit.edu)

- * mailing list (SNOWMASS-RPF-06-DARK-SECTOR@fnal.gov)
- * SLACK channel ([#rpf-06-dark-sector](#))
- * webpage: <https://snowmass21.org/rare/dark>

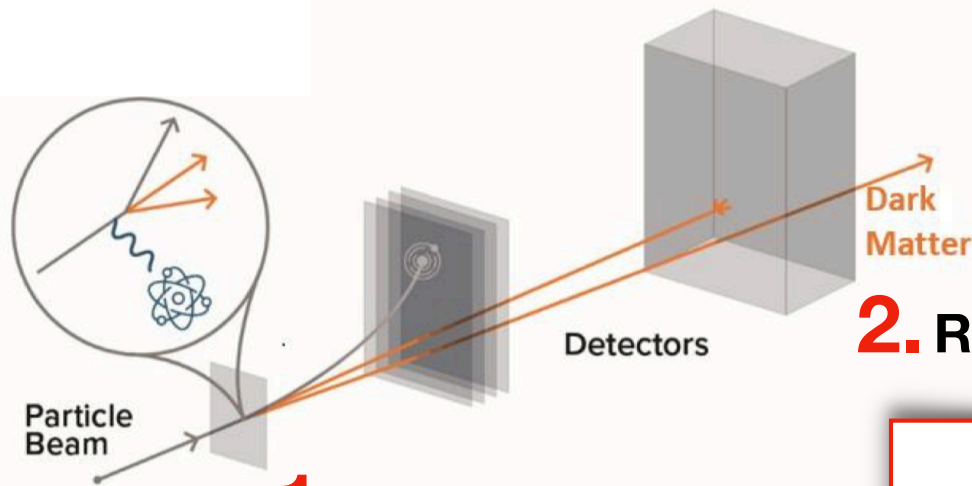
Key questions:

- * Why are dark sectors a compelling opportunity for the next decade?
see talk by Natalia
- * What are the needed theoretical developments to guarantee a broad exploration of the dark sector?
- * Which set of benchmark models comprehensively covers the interesting signatures we can look for?
- * How will experiments confront the dark sector landscape?
(RF6 involves experiments from small to large, both dedicated and multi-purpose).

Goal: to achieve a broad theoretical and experimental exploration of the physics of dark sectors at or below the GeV scale.

Search techniques

Production of dark matter

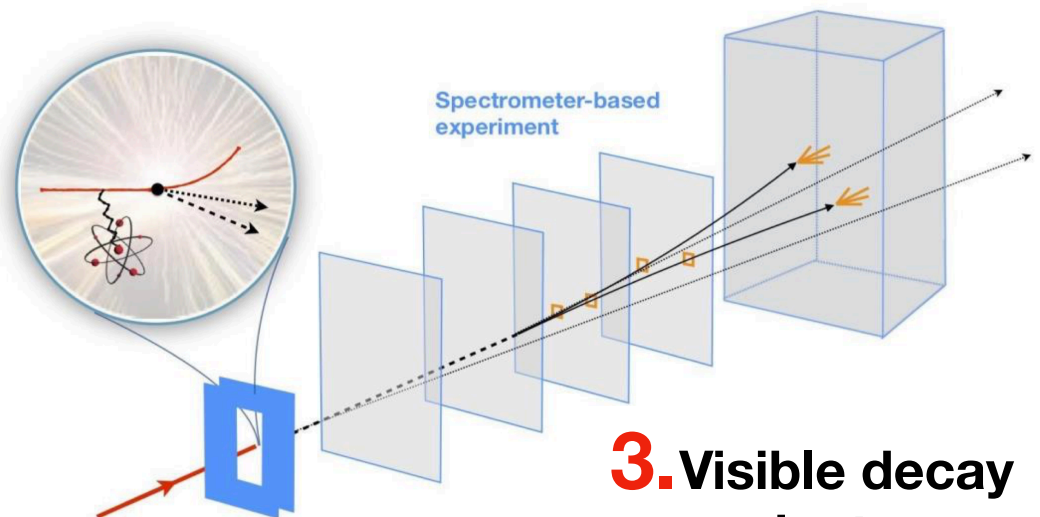


1. Missing energy/ momentum

2. Re-scattering

Several experiments, but
the experimental techniques
are only 3

Production of unstable dark sector particles

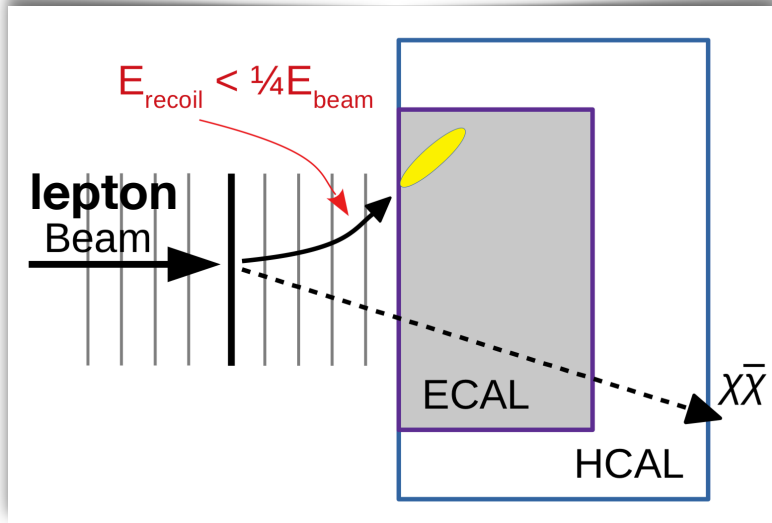


3. Visible decay products

Basic Research Needs for
DM Small projects New Initiatives (DMNI),
2018

[https://science.osti.gov/-/media/hep/pdf/Reports/
Dark_Matter_New_Initiatives_rpt.pdf](https://science.osti.gov/-/media/hep/pdf/Reports/Dark_Matter_New_Initiatives_rpt.pdf)

1. Missing energy/momentum

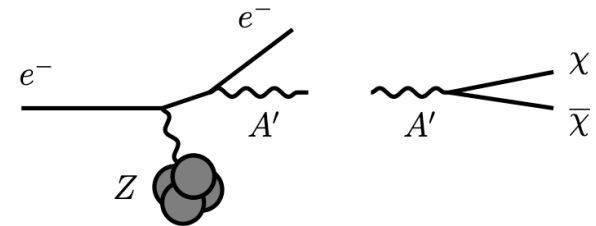


Dark matter events can be kinematically characterized by the calorimetric “disappearance” of a sizable fraction of the beam energy.

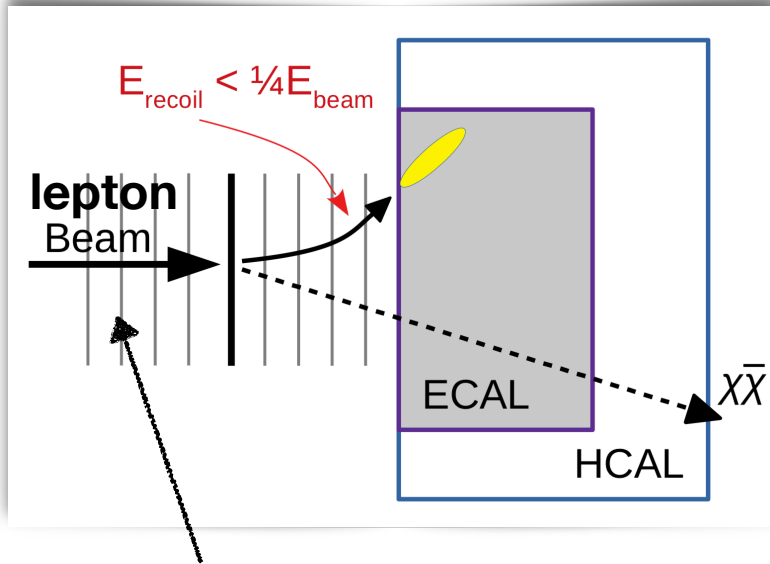
Detection strategy

Dark Matter can be produced through the mediation of a on-shell or off-shell mediator.

For example,



1. Missing energy/momentum



e⁻ beam for the **NA64** experiment,
[Andreas et al., 1312.3309](#) Running
 at CERN

e⁻ beam for the **LDMX** experiment,
[Akesson et al., 1808.05219](#)

e⁺ beam for the **POKER** experiment,
[Andreev et al., 2108.04195](#)

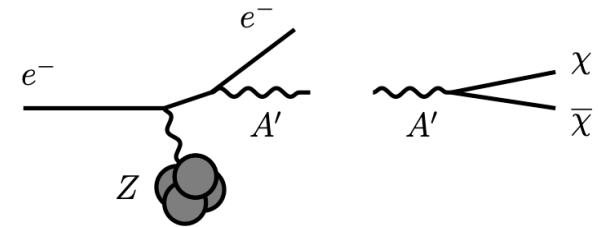
μ⁻ beam for the **M³** experiment,
[Kahn et al., 1804.03144](#) Future
 experiments

Dark matter events can be kinematically characterized by the calorimetric “disappearance” of a sizable fraction of the beam energy.

Detection strategy

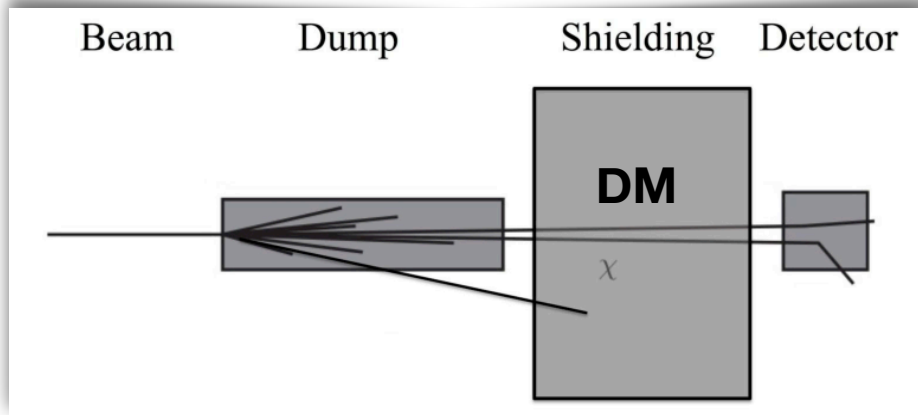
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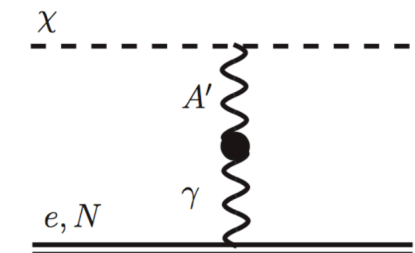
DMNI funding

2. Re-scattering

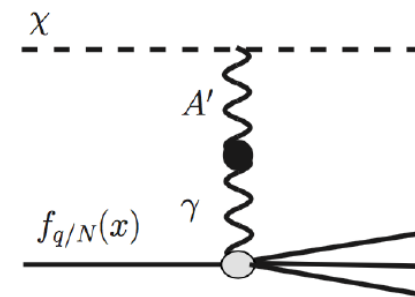


Production of dark matter in the dump and detection of its scattering in forward detectors.

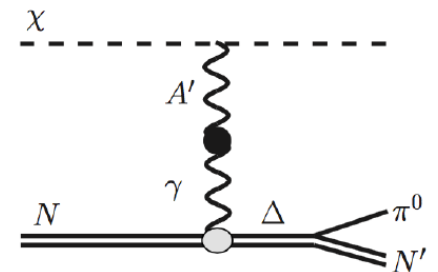
Detection strategy



Elastic electron, nucleon ,
nucleus (coherent) scattering

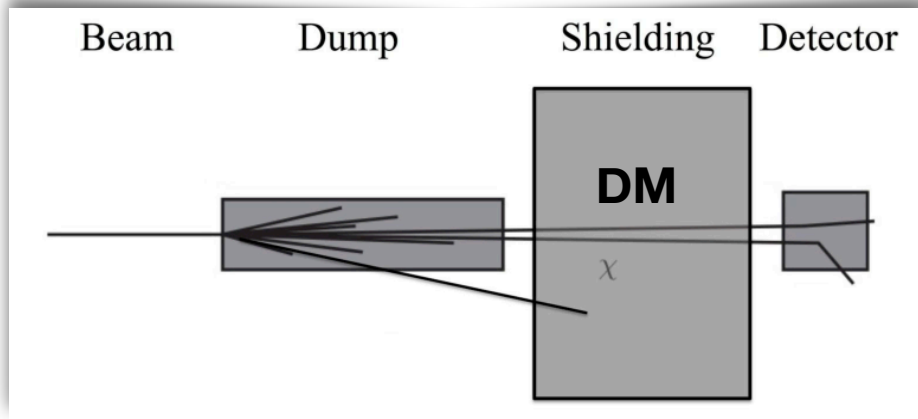


Deep inelastic scattering



Inelastic neutral pion
production

2. Re-scattering



Proton-beam experiments are highly synergistic with the accelerator-based neutrino physics program. They use the same beamlines and detectors:

LSND, MiniBooNE, COHERENT, CCM

(see Brian's talk)

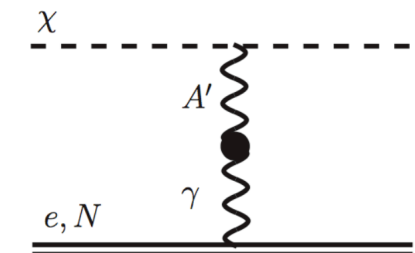
DMNI funding

Electron-beam experiments have the advantage of a more compact secondary DM beam (BDX experiment)

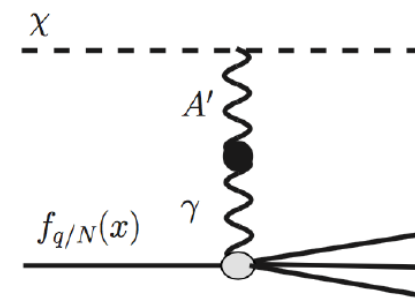
Synergy with beam dump-experiments that utilize high energy beams (forward facility, future colliders (see Jonathan's talk))

Production of dark matter in the dump and detection of its scattering in forward detectors.

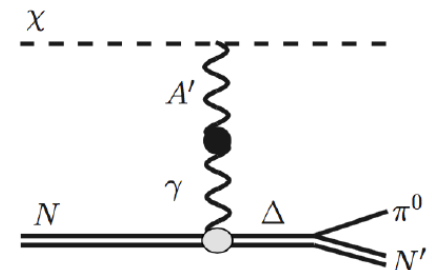
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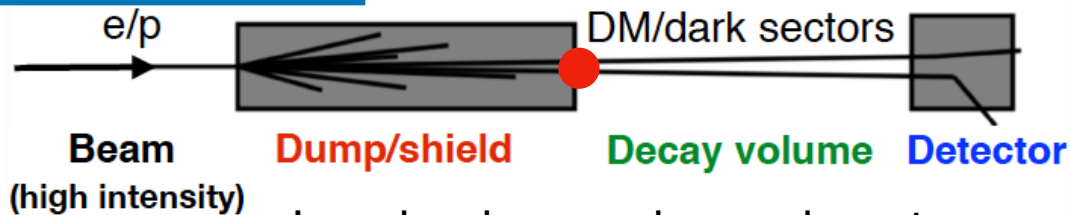
Deep inelastic scattering



Inelastic neutral pion production

3. Visible signatures

DISPLACED



Low background experiments
(depending on the size of the dump)

p beam for the SeaQuest/DarkQuest experiment at Fermilab

p beam for the NA62, KLEVER experiments at CERN

e- beam for the HPS experiment at JLAB

e- beam for the DarkLight experiment at TRIUMF

Production of an unstable dark sector particle in the dump and detection of its SM decay products in forward detectors.

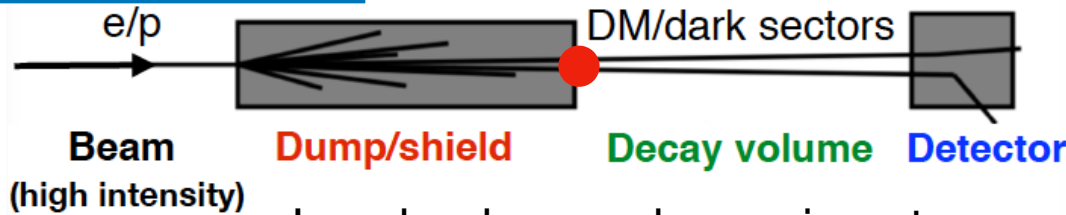
Detection strategy

Running
experiments

Future
experiments

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Detection strategy

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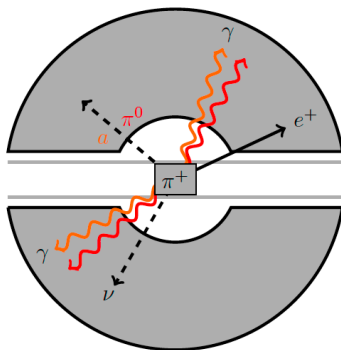
p beam for the NA62, KLEVER experiments at CERN

e- beam for the HPS experiment at JLAB

e- beam for the DarkLight experiment at TRIUMF

Running
experiments

Future
experiments



PROMPT

Production of an unstable dark sector particle from meson decay and detection of its SM decay products. **Detection strategy**

Pion decaying at rest (PIONEER experiment)

Eta/eta' decaying (almost) at rest (REDTOP experiment)

Enormous synergy with collider experiments! Belle II, LHCb, ...

A broad set of benchmarks

initiated by N.Toro

Benchmarks in Final State x Portal Organization

| | DM Production | Mediator Decay Via Portal | Structure of Dark Sector |
|----------|---|---|---|
| Vector | m_χ vs. y [$m_A/m_\chi=3, \alpha_D=.5$] $m_{A'}$ vs. y [$\alpha_D=0.5, 3$ m_χ values] m_χ vs. α_D [$m_A/m_\chi=3, y=y_{fo}$] m_χ vs. m_A [$\alpha_D=0.5, y=y_{fo}$] <i>Millicharge m vs. q</i> | $m_{A'}$ vs. ϵ [decay-mode agnostic] $m_{A'}$ vs. ϵ [decays] | iDM m_χ vs. y [$m_A/m_\chi=3, \alpha_D=.5$] (anom connection) SIMP-motivated cascades [slices TBD] $U(1)_{B-L} / \mu-\tau / B-3\tau$ (DM or SM decays) |
| Scalar | m_χ vs. $\sin\theta$ [$\lambda=0$, fix $m_S/m_\chi, g_D$] (thermal target excluded 1512.04119, should still include) Note secluded DM relevance of $S \rightarrow SM$ of mediator searches | m_S vs. $\sin\theta$ [$\lambda=0$] m_S vs. $\sin\theta$ [$\lambda=s.t. \text{Br}(H \rightarrow ss \sim 10^{-2})$]? | Dark Higgsstrahlung (w/vector) scalar SIMP models Leptophilic/leptophobic dark Higgs |
| Neutrino | $e/\mu/\tau$ at 1709.07001 Batell, Han, McKeen, Es Haghi | m_N vs. U_e m_N vs. U_μ m_N vs. U_τ Think more about reasonable flavor structures | Sterile neutrinos with new forces |
| ALP | m_χ vs. f_q/l [$\lambda=0$, fix $m_a/m_\chi, g_D$] (thermal target excluded) What about f_y, f_G ? | m_a vs. f_y m_a vs. f_G m_a vs. $f_q=f_l$ m_a vs. f_w | FV axion couplings |

Focus: both the DM and the **mediator** are light ($O(\text{GeV})$ or less)

Bold = BRN benchmark, italic = PBC benchmark, others are new suggestions

BRN = basic research needs for DM small projects

RF6 “Big ideas” solicited papers

- Organization around science goals/questions.
- Arrange the breadth of RF6 science so that all the main techniques have a chance to shine.
- Span $\geq 95\%$ of white-paper interests

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1. Detect dark matter particle production (production reaction or through subsequent DM scattering), with a focus on exploring sensitivity to thermal DM interaction strengths.

Editors: Gordan Krnjaic, Natalia Toro

2. Explore the structure of the dark sector by producing and detecting unstable dark particles: Minimal Portal Interactions.

Editors: Brian Batell, Chris Hearty

3. New Flavors and Rich Structures in Dark Sectors.

Editors: Phil Harris, Philip Schuster, Jure Zupan

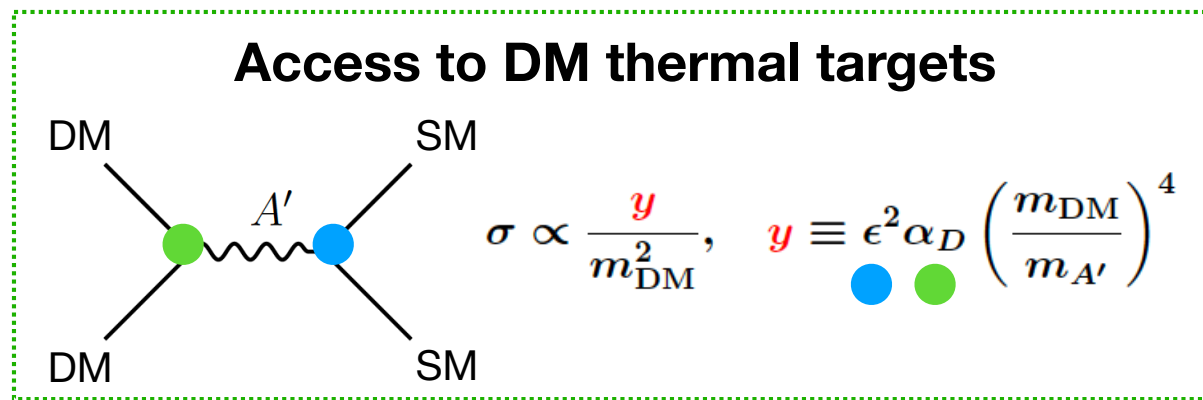
4. Experiments/facilities

Editors: Phil Ilten, Nhan Tran

More details: <https://docs.google.com/document/d/1R0O23wjGLxRzsc93a4pJIFn17yW9TCTq>
(in our google drive folder, https://drive.google.com/drive/folders/1sMn1cWI2ddqzu46Yi4TcMIX7Cm2GUxO_)

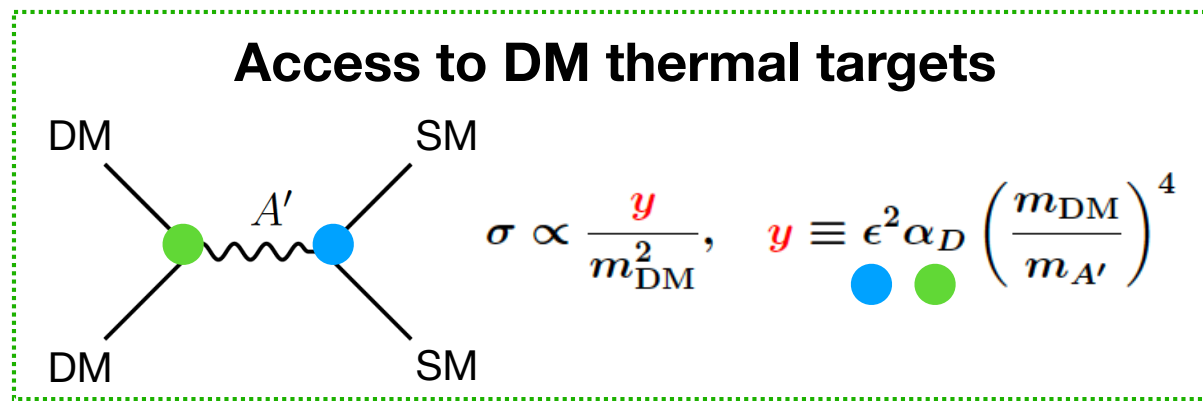
Big idea 1 solicited white paper

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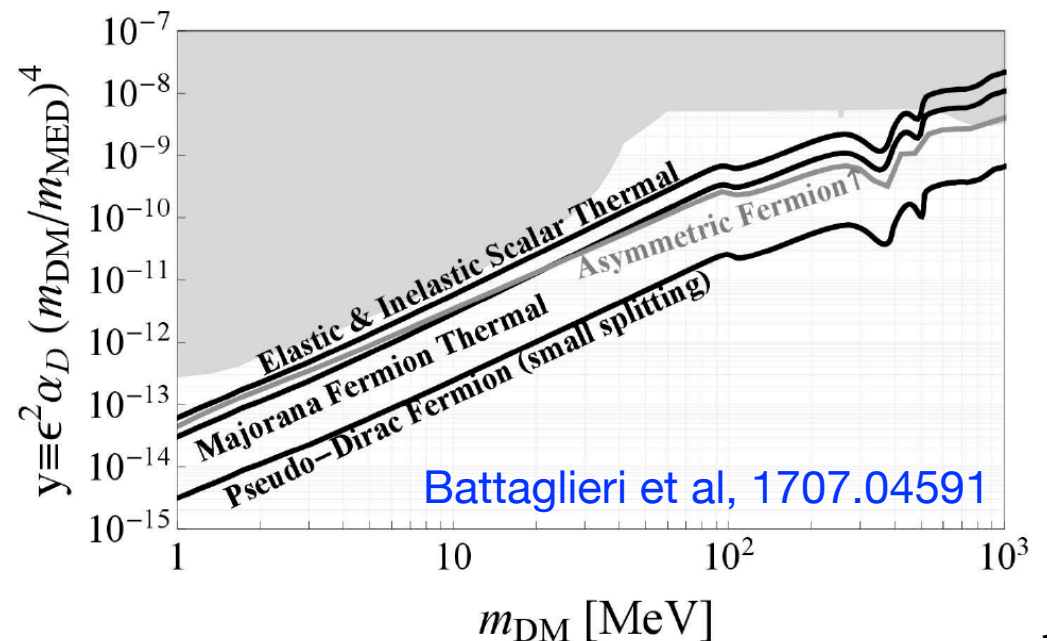


In the case of a dark photon mediator:

$$\text{Majorana} : J_D^\mu = \frac{1}{2} \bar{\chi} \gamma^\mu \gamma_5 \chi$$

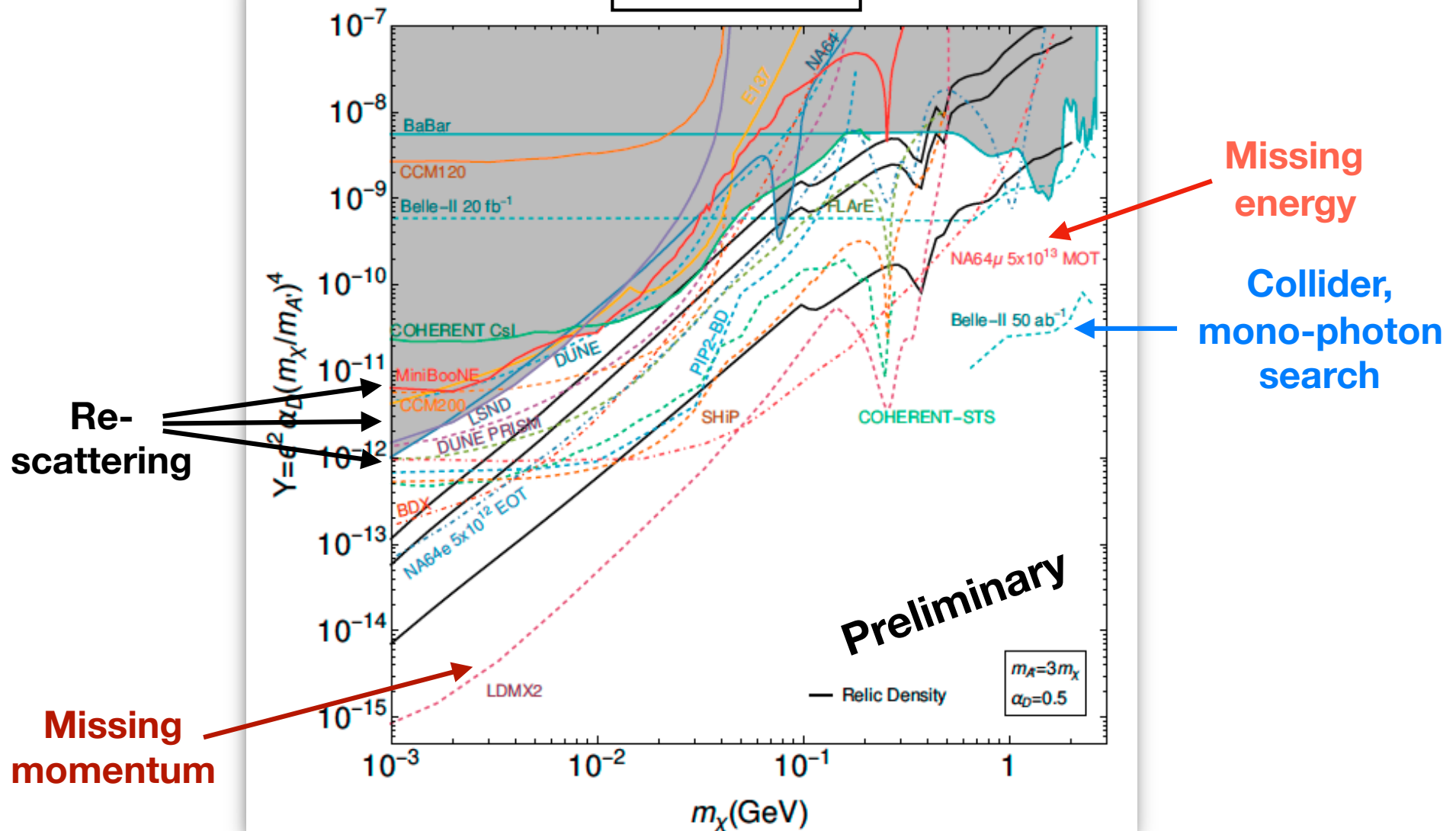
$$\text{PseudoDirac} : J_D^\mu = i \bar{\chi}_1 \gamma^\mu \chi_2$$

$$\text{Scalar} : J_D^\mu = i(\chi^* \partial^\mu \chi - \chi \partial^\mu \chi^*)$$



Big idea 1, invisible dark photon

$$A' \rightarrow \chi\chi$$



Big idea 2 solicited white paper

Explore the structure of the dark sector by producing and detecting unstable dark particles: Minimal Portal Interactions.

Editors: Brian Batell, Chris Hearty

| | | |
|--------------------|---|---|
| * dark photon | $\epsilon B^{\mu\nu} A'_{\mu\nu}$ | $A' \rightarrow \ell^+ \ell^-, \dots$ |
| * dark scalar | $\kappa H ^2 S ^2$ | $S \rightarrow \mu^+ \mu^-, \pi^+ \pi^-, KK, \dots$ |
| * sterile neutrino | $y H L N$ | $N \rightarrow \ell \pi, \dots$ |
| * ALP | $g_{a\gamma} a \tilde{F}_{\mu\nu} F^{\mu\nu}$ | $a \rightarrow \gamma\gamma, \dots$ |

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How to test these couplings?

Sizable coupling \rightarrow prompt decay (generically larger backgrounds)

Small coupling \rightarrow displaced decay (generically small backgrounds)

Experimental targets:

Secluded DM scenarios

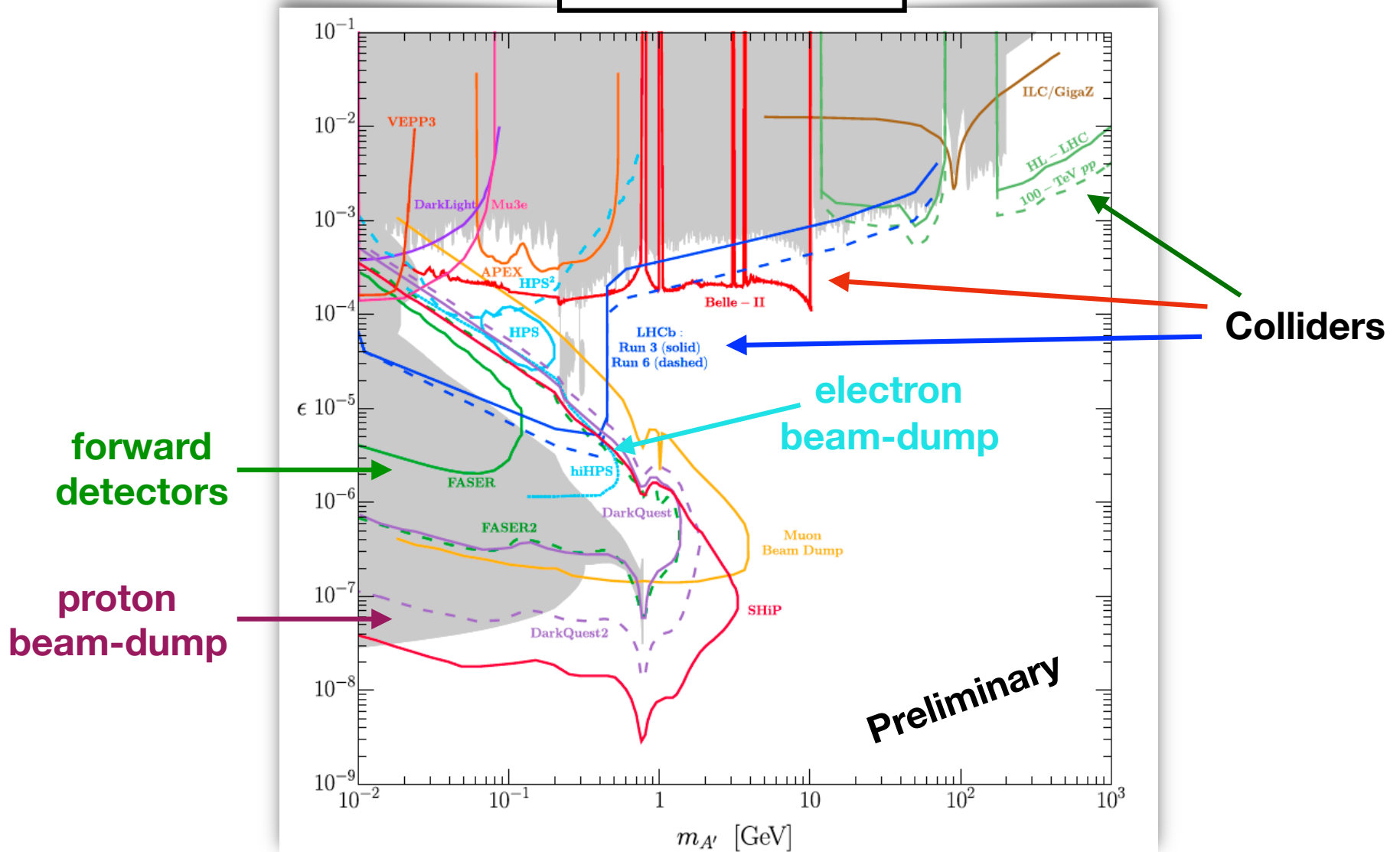
(Pospelov, Ritz, Voloshin, 0711.4866)

Forbidden DM scenarios

(D'Agnolo, Ruderman, 1505.07107)

Big idea 2, visible dark photon

$$A' \rightarrow \text{visible}$$



Big idea 3 solicited white paper

Beyond Minimal Models: New Flavors and Rich Structures in Dark Sectors.

Editors: Phil Harris, Philip Schuster, Jure Zupan

Richer phenomenology  rethinking of experimental strategies for achieving optimized sensitivities

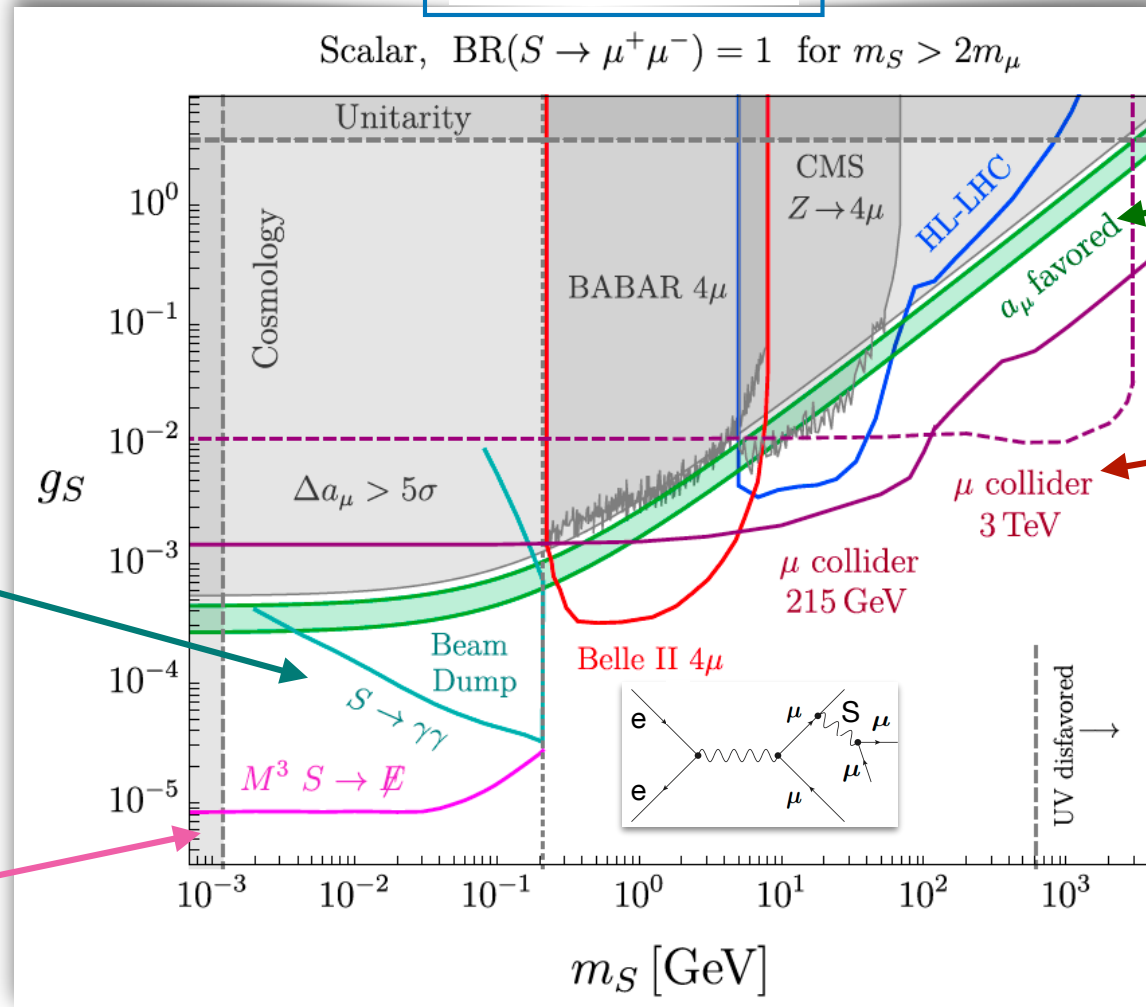
2 themes:

- * Dark sector benchmarks for anomalies in data (e.g. $(g - 2)_\mu$, flavor anomalies, Xenon 1T excess, MiniBoone excess, ...)
- * Commonly used benchmarks going beyond the assumption of minimality (e.g. flavor violating ALPs, strongly interactive massive particle DM, inelastic DM, ...)

Big idea 3, anomaly motivated scenarios

“flavor-specific” dark sectors

$$\mathcal{L} = -g_S S \bar{\mu}\mu$$



$(g - 2)_\mu$ favored region

high energy muon colliders

muon beam-dump
Chen, Pospelov, Zhong, 1701.07437

muon missing momentum
Kahn, Krnjaic, Tran, Whitbeck, 1804.03144

Capdevilla, Curtin, Kahn, Krnjaic, 2112.08377

Big idea 3, strongly interacting massive particle DM

Realized in a QCD-like theory $SU(N_c)$ with

$SU(N_f) \times SU(N_f) \rightarrow SU(N_f)$ Light pions (some of them can be DM)

$$\mathcal{L}_{\text{WZW}} = \frac{2N_c}{15\pi^2 f_\pi^5} \epsilon^{\mu\nu\rho\sigma} \text{Tr}(\pi \partial_\mu \pi \partial_\nu \pi \partial_\rho \pi \partial_\sigma \pi)$$

| | |
|-----------------------------|--|
| Relic abundance: | $3 \rightarrow 2$ annihilation |
|-----------------------------|--|

Big idea 3, strongly interacting massive particle DM

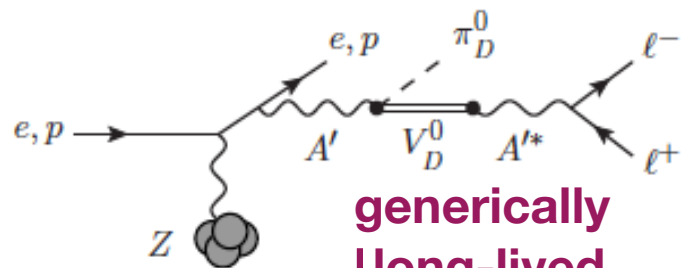
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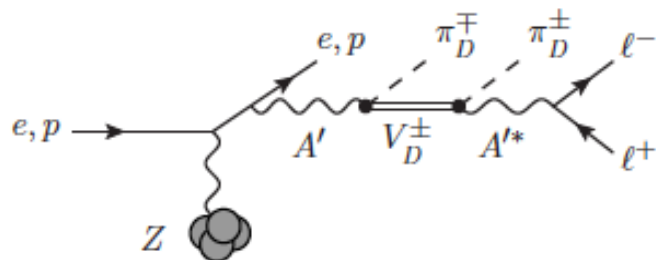
Relic abundance: $3 \rightarrow 2$ annihilation

Signatures



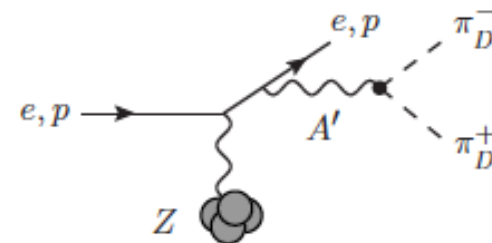
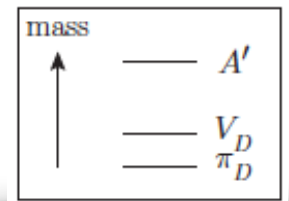
generically
long-lived

$$A' \rightarrow V_D \pi_D, V_D \rightarrow A'^* \rightarrow e^+ e^-$$



$$A' \rightarrow V_D \pi_D, V_D \rightarrow A'^* \pi_D \rightarrow e^+ e^- \pi_D$$

(Semi) Visible
 A' decay



Invisible
 A' decay

$$A' \rightarrow \pi_D \pi_D$$

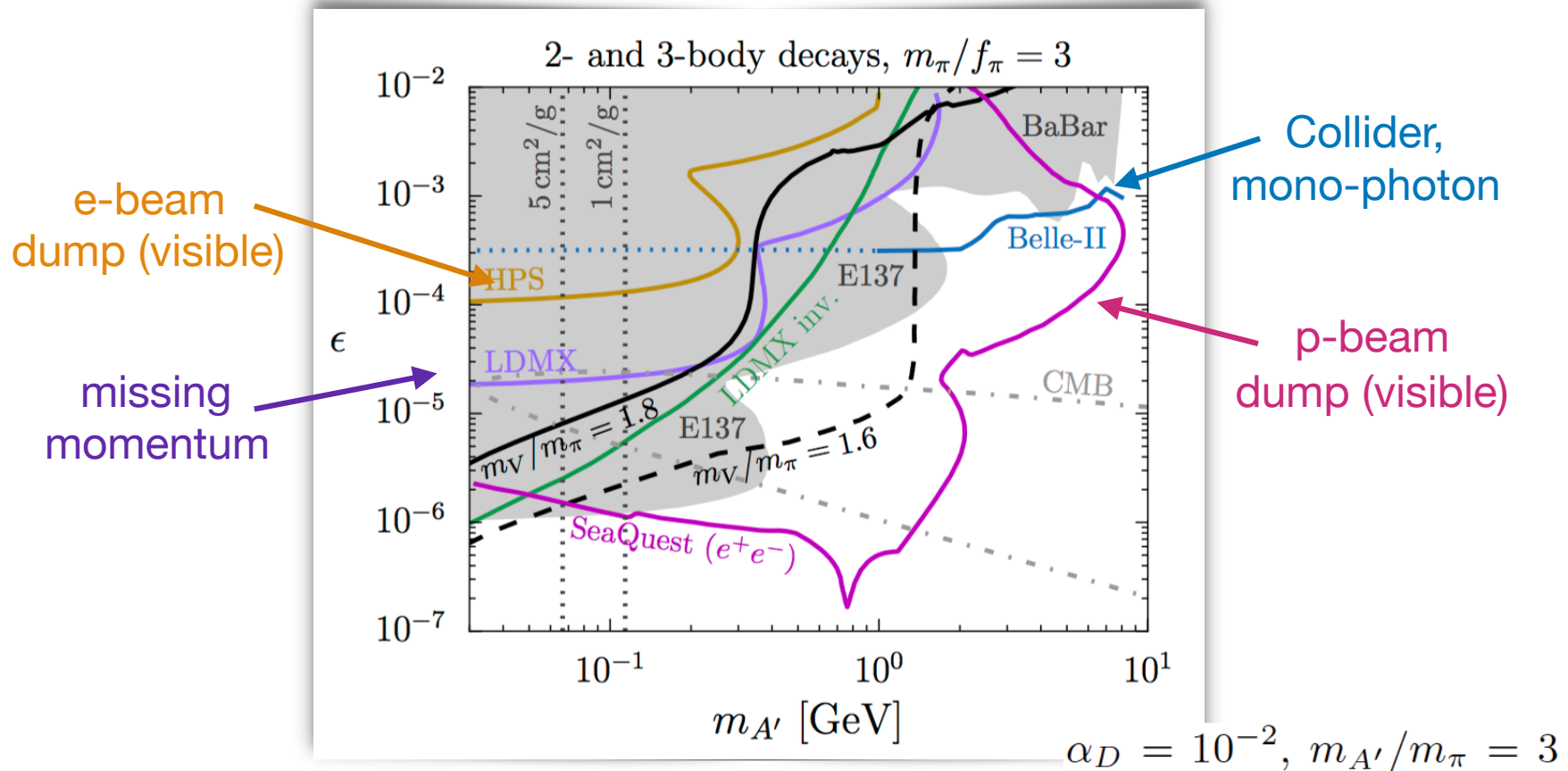
Big idea 3, strongly interacting massive particle DM

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$SU(N_f) \times SU(N_f) \rightarrow SU(N_f)$ Light pions (some of them can be DM)

$$\mathcal{L}_{\text{WZW}} = \frac{2N_c}{15\pi^2 f_\pi^5} \epsilon^{\mu\nu\rho\sigma} \text{Tr}(\pi \partial_\mu \pi \partial_\nu \pi \partial_\rho \pi \partial_\sigma \pi)$$

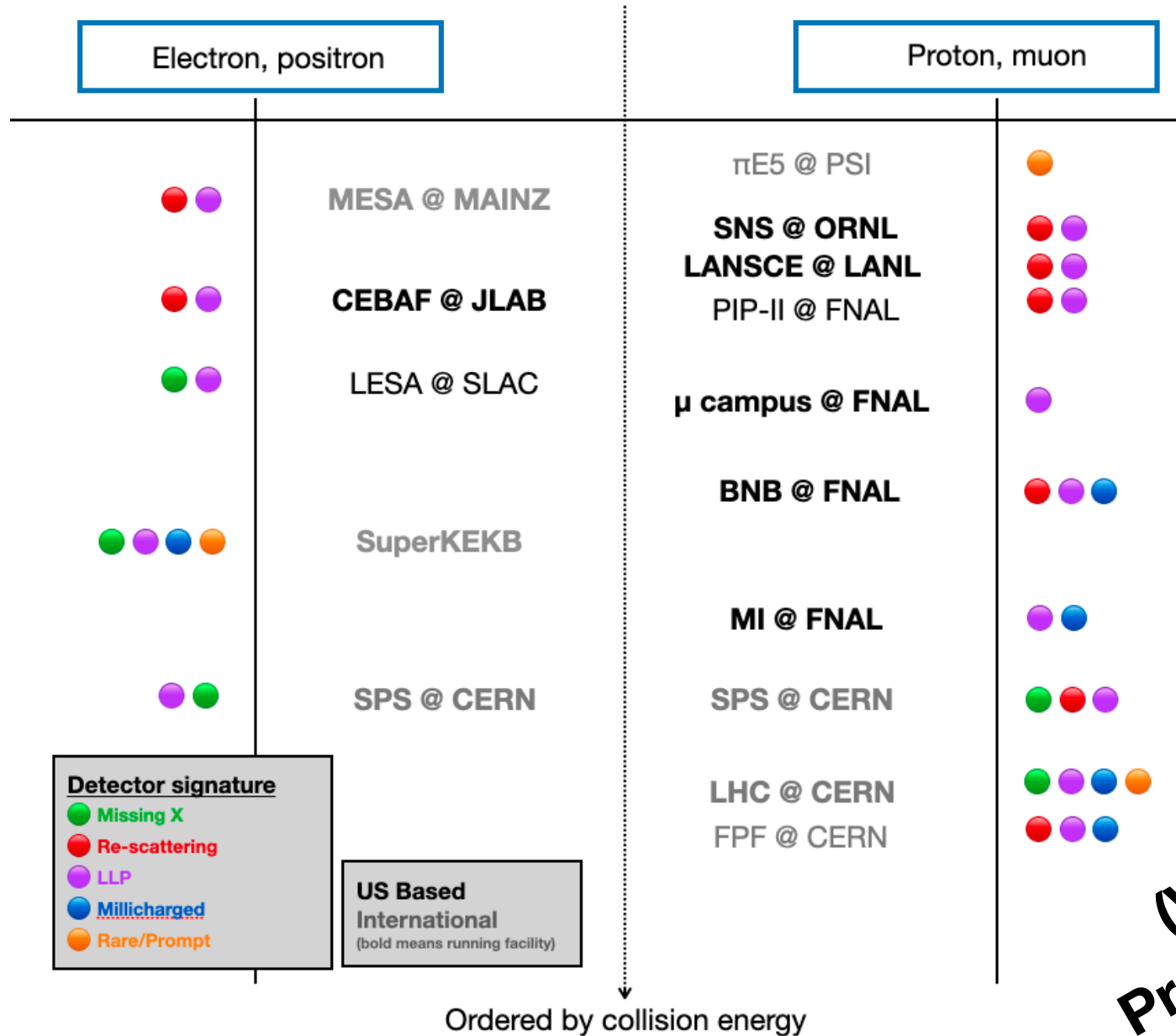
Relic abundance: $3 \rightarrow 2$ annihilation



Berlin, Blinov, SG, Schuster, Toro, 1801.05805

Experiments and facilities solicited white paper

Editors:
Phil Ilten, Nhan Tran



(Very)
Preliminary



Conclusions & Outlook

Dark sector particles naturally appear in many well-motivated extensions of the Standard Model.

High-intensity experiments based on 3 different techniques:

- missing energy/momentum
- re-scattering
- searches for visible decay products

will offer an **unprecedented opportunity to test dark sectors at or below the GeV scale**

Full exploration of many benchmarks is expected in the coming few years.

Crucial complementarity with

- (high energy) colliders
- auxiliary detectors at colliders
- direct detection experiments

Big idea 3, inelastic dark

$$-\mathcal{L} \supset m_D \eta \xi + \frac{1}{2} \delta_\eta \eta^2 + \frac{1}{2} \delta_\xi \xi^2 + \text{h.c.}$$

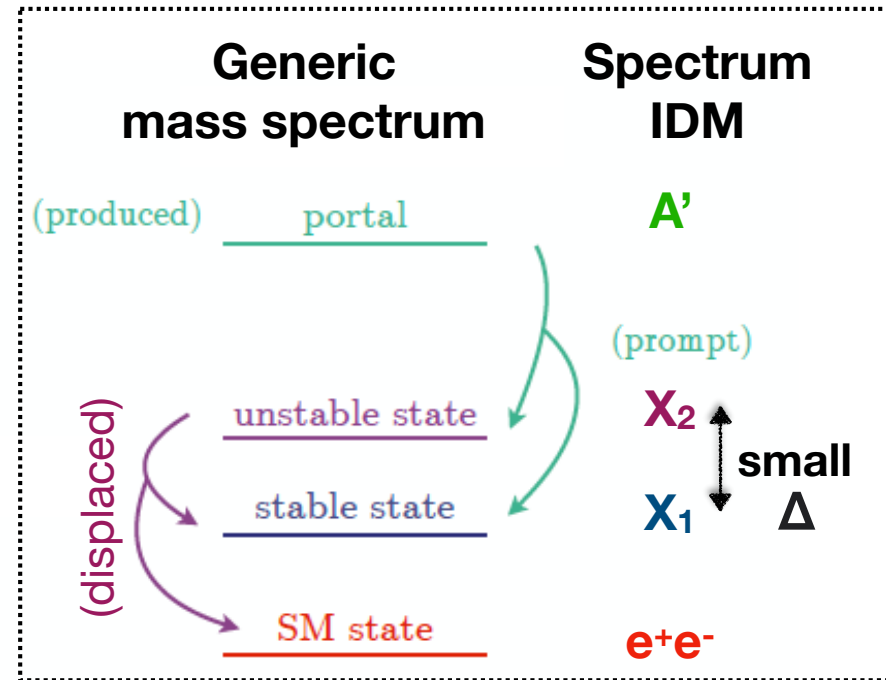
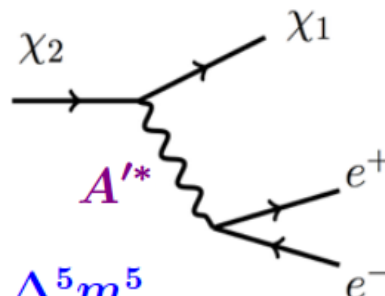
$$\mathcal{L} \supset \frac{ie_D m_D}{\sqrt{m_D^2 + (\delta_\xi - \delta_\eta)^2/4}} A'_\mu (\bar{\chi}_1 \gamma^\mu \chi_2 - \bar{\chi}_2 \gamma^\mu \chi_1)$$

2-component Weyl spinors
with opposite charge under U(1)'

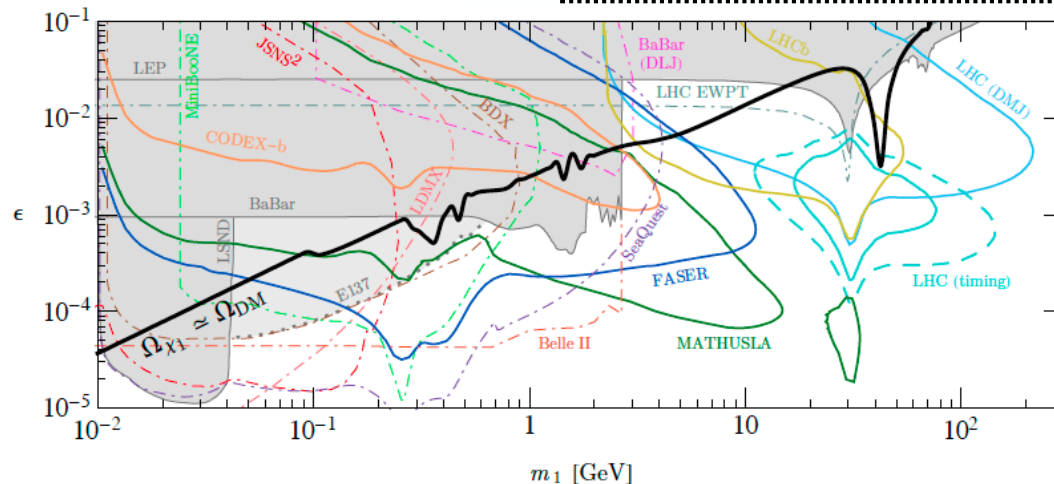
* Freeze-out: $\chi_1 \chi_2 \rightarrow \text{SM}$

* Signatures in our labs:

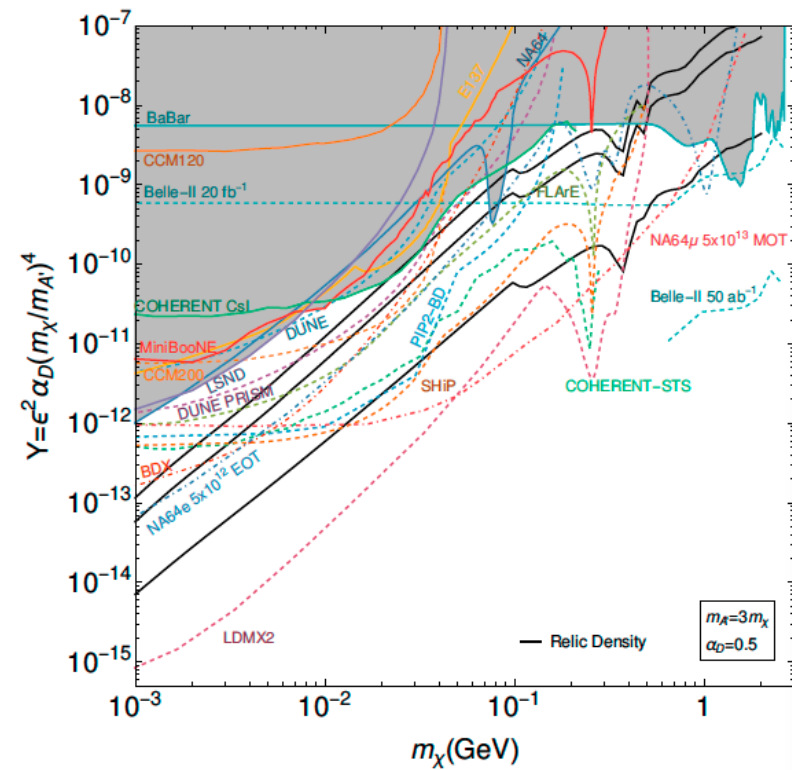
$$\Gamma(\chi_2 \rightarrow \chi_1 e^+ e^-) \simeq \frac{4\epsilon^2 \alpha_{\text{em}} \alpha_D \Delta^5 m_1^5}{15\pi m_{A'}^4}$$



Visible decays if large
enough mass splitting.
Invisible component as
well



Invisible dark photon parameter space



Preliminary

$\epsilon, \alpha_D, m_\chi, m_{A'}$

