



# Cluster Cosmology

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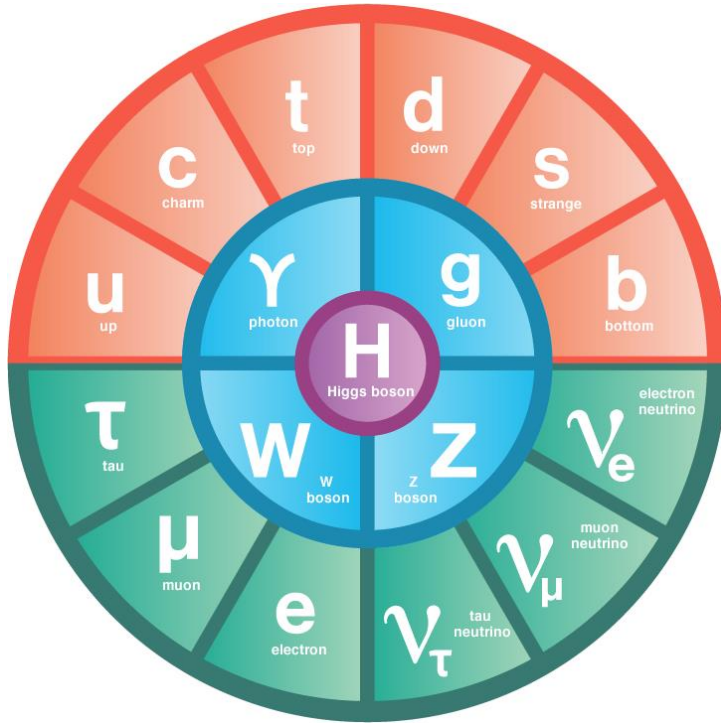
SIST 5 minutes, 5 slides

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**DARK ENERGY  
SURVEY**

- Standard model of cosmology

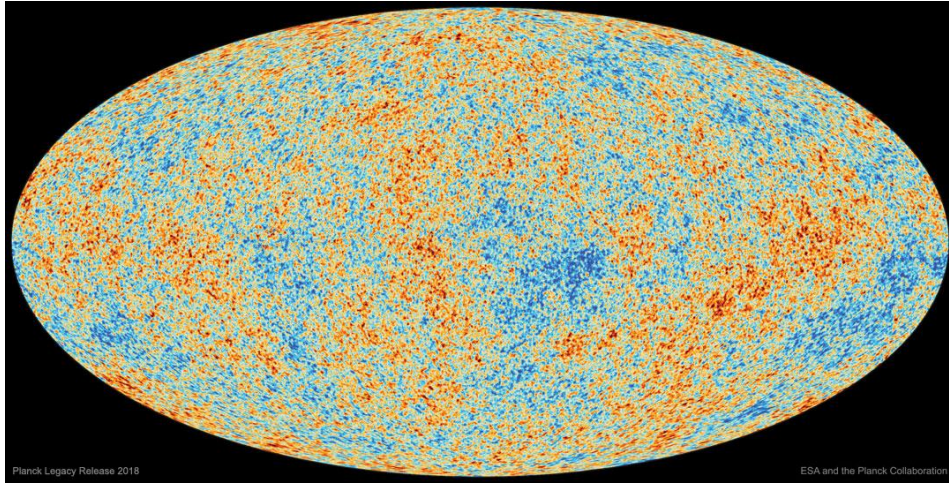


Description	Symbol	Value
Physical baryon density parameter <sup>[a]</sup>	$\Omega_b h^2$	$0.022\ 30 \pm 0.000\ 14$
Physical dark matter density parameter <sup>[a]</sup>	$\Omega_c h^2$	$0.1188 \pm 0.0010$
Age of the universe	$t_0$	$13.799 \pm 0.021 \times 10^9$ years
Scalar spectral index	$n_s$	$0.9667 \pm 0.0040$
Curvature fluctuation amplitude, $k_0 = 0.002\ \text{Mpc}^{-1}$	$\Delta_R^2$	$2.441^{+0.088}_{-0.092} \times 10^{-9}$ <sup>[88]</sup>
Reionization optical depth	$\tau$	$0.066 \pm 0.012$

Total density parameter <sup>[b]</sup>	$\Omega_{\text{tot}}$	1
Equation of state of dark energy	$w$	-1
Tensor/scalar ratio	$r$	0
Running of spectral index	$dn_s/d \ln k$	0
Sum of three neutrino masses	$\sum m_\nu$	$0.06\ \text{eV}/c^2$ <sup>[c][84]:40</sup>
Effective number of relativistic degrees of freedom	$N_{\text{eff}}$	$3.046$ <sup>[d][84]:47</sup>
Hubble constant	$H_0$	$67.74 \pm 0.46\ \text{km s}^{-1}\ \text{Mpc}^{-1}$
Baryon density parameter <sup>[b]</sup>	$\Omega_b$	$0.0486 \pm 0.0010$ <sup>[e]</sup>
Dark matter density parameter <sup>[b]</sup>	$\Omega_c$	$0.2589 \pm 0.0057$ <sup>[f]</sup>
Matter density parameter <sup>[b]</sup>	$\Omega_m$	$0.3089 \pm 0.0062$
Dark energy density parameter <sup>[b]</sup>	$\Omega_\Lambda$	$0.6911 \pm 0.0062$
Critical density	$\rho_{\text{crit}}$	$(8.62 \pm 0.12) \times 10^{-27}\ \text{kg/m}^3$ <sup>[g]</sup>
The present root-mean-square matter fluctuation averaged over a sphere of radius $8h^{-1}\ \text{Mpc}$	$\sigma_8$	$0.8159 \pm 0.0086$
Redshift at decoupling	$z_*$	$1\ 089.90 \pm 0.23$
Age at decoupling	$t_*$	$377\ 700 \pm 3200$ years <sup>[88]</sup>
Redshift of reionization (with uniform prior)	$z_{\text{re}}$	$8.5^{+1.0}_{-1.1}$ <sup>[89]</sup>



# CMB



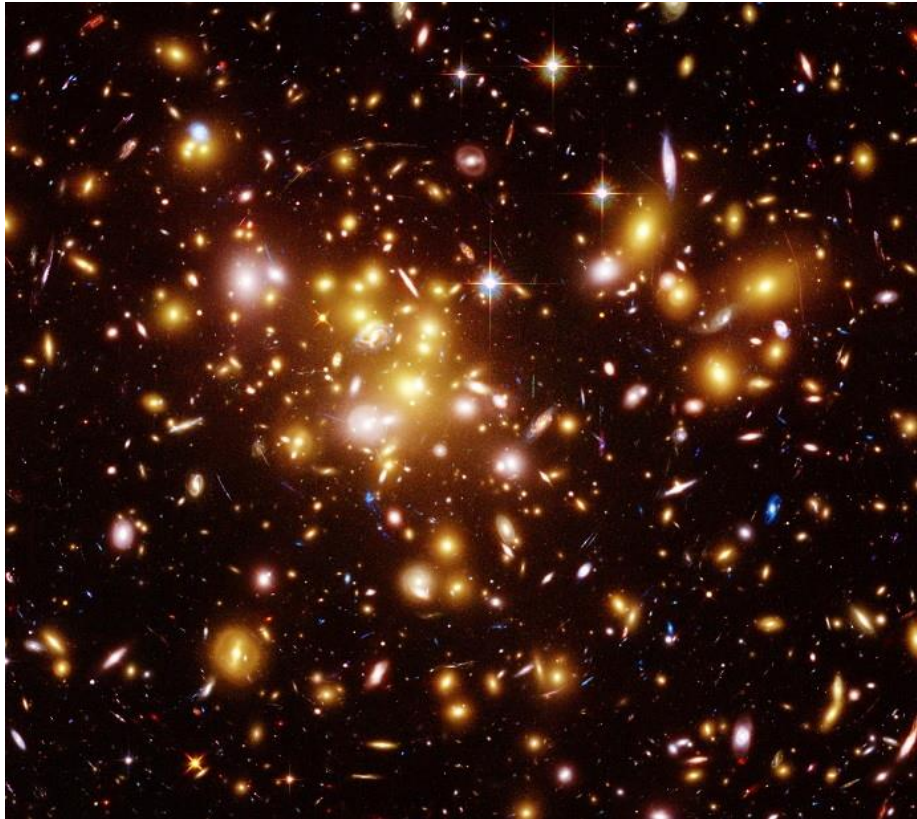
(Picture from Planck collaboration)

(Picture 2MASS Large Scale Survey)



# Galaxy clusters and weak lensing

- Structures with the largest scale in the Universe
- Their mass can be found using weak lensing





# Cluster Cosmology

Matter density parameter <sup>[b]</sup>	$\Omega_m$	$0.3089 \pm 0.0062$
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- In 2020, the DES cluster group measured  $\Omega_m = 0.2$ .

## References

- Abbott, T. M. C., and others. "Dark Energy Survey Year 1 Results: Cosmological constraints from cluster abundances and weak lensing". *Phys. Rev. D* 102, no.2 (2020): 023509.
- Frieman, Joshua, Michael, Turner, and Dragan, Huterer. "Dark Energy and the Accelerating Universe". *Ann. Rev. Astron. Astrophys.* 46 (2008): 385–432.
- Costanzi, Matteo, Francisco, Villaescusa-Navarro, Matteo, Viel, Jun-Qing, Xia, Stefano, Borgani, Emanuele, Castorina, and Emiliano, Sefusatti. "Cosmology with massive neutrinos III: the halo mass function and an application to galaxy clusters". *JCAP* 12 (2013): 012.

