# VD PHOTON DETECTION SYSTEM -REQUIREMENTS

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### **Basic Considerations**

- We believe that the PDs of HD and VD modules of the far detector should have the same minimal physics requirements;
- This means that the VD module should be able to do at least the same physics as the HD one, but without limitations on possible scope expansions.

# PD HD Physics and Detector Requirements

Label	Description	Specification (Goal)	Rationale	Validation
SP-FD-3	Light yield	> 20 PE/MeV (avg), $>$ 0.5 PE/MeV (min)	Gives PDS energy resolution comparable to that of the TPC for 5-7 MeV SN $\nu$ s, and allows tagging of > 99% of nucleon decay backgrounds with light at all points in de- tector.	Supernova and nu- cleon decay events in the FD with full simulation and re- construction.
SP-FD-4	Time resolution	$< 1  \mu s$ (< 100 ns)	Enables 1 mm position reso- lution for 10 MeV SNB can- didate events for instanta- neous rate $< 1 \mathrm{m^{-3}ms^{-1}}$ .	

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The Physics requirements have been translated into detector requirements through detailed (and very time consuming) full simulations of the SP HD far detector and of large samples of Supernova and nucleon decay events

# Requirements

- The detector requirements of an average LY of 20 phel/MeV inside the TPC comes directly from the physics requirement of being able to reach, with the HD PD, a similar energy resolution for Supernova neutrinos as that achievable with the detection of the charge with the TPC;
- The detector requirement of a minimum LY of 0.5 phel/MeV comes from the physics requirement of being able to measure the T<sub>0</sub> for nucleon decay background events everywhere inside the active volume.



# Path towards PD VD Detector Requirements

- The work done for the HD represents a solid starting point and can allow us to make some educated guesses fabout VD requirements;
- It is reasonable to think that the requirement on the average LY, related to the *average energy resolution* of the PD, will be similar to the HD one - around 20 phel/MeV. This requirement is assumed also for the VD;
- The requirement *on the minimum light yield could be slightly different*, because the TPCs have different geometrical configurations;
- Conservatively, it looks appropriate to assume the minimum LY requirement of 0.5 phel/MeV everywhere
  8 in the TPC also for the VD

# Path towards PD VD Detector Requirements

- Assuming for the VD the same minimal physics requirements of the HD, the derivation of the detector requirements will follow the same process:
  - Simulation and reconstruction of large samples of Supernova neutrinos and nucleon decay events
- The Physics and Simulation WG of the PD Consortium has developed the Monte Carlo of the VD module. Production of large samples of SN neutrino and nucleon decay events will start in the next weeks. A refinement of the VD requirements is expected for mid-June

### **TB held Requirements**

• TB held required of FD1 are assumed also for FD2, for the moment.

SP-FD-15	LAr nitrogen con- tamination	< <sup>:</sup> 10 ppm	Maintain 0.5 PE/MeV PDS sensitivity required for trig- gering proton decay near cathode.	In situ measur- ment
SP-PDS-1	Clean assembly area	Class 100,000 clean assembly area	Demonstrated as satisfac- tory in ProtoDUNE-SP, and is the Deep Under- ground Neutrino Experi- ment (DUNE) assembly area standard.	<u>ProtoDUNE-SP</u> and in Fermilab materials test stand
SP-PDS-2	Spatial localization in y-z plane	$< 2.5 \mathrm{m}$	Enables accurate matching of photon detector (PD) and TPC signals.	SNB     neutrino and       nucleon     decay       (NDK)     simulation       in the FD
SP-PDS-3	Environmental light exposure	No exposure to Sunlight. All other unfiltered sources: < 30 minutes integrated across all exposures	Shown to prevent damage to wavelength-shifting (WLS) coatings due to UV.	Studies in ProtoDUNE-SP, and at IU
SP-PDS-4	Environmental humidity limit	<~50 % RH at 70 °F	Demonstrated to prevent damage to WLS coatings due to humidity.	PD optical coating studies

### **TB held Requirements**

SP-PDS-14	Signal-to-noise i SP-PD	in	> 4	Keep data rate within elec- tronics bandwidth limits.
SP-PDS-15	Dark noise rate i SP-PD	in	$< 1  \rm kHz$	Keep data rate within elec- tronics bandwidth limits.
SP-PDS-16	Dynamic Range i SP-PD	in	< 20 %	Keep the rate of saturating channels low enough for ef- fective mitigation.

### **Additional Considerations**

- The VD design offers much more flexibility for the design of the PD than the HD and has much less mechanical constraints. It offers the possibility of expanding the Physics reach of the detector;
- An increased and more uniform Light Yield translates into:
  - Improved resolution for low energy events (SNB and solar neutrinos) → better identification of SNB spectral features, which can allow to discriminate different dynamical models;
  - <sup>~</sup> Possibility of a light based and high efficiency SNB trigger;
  - Improved background rejection for low energy events by increasing the flash-track matching;
  - <sup>~</sup> Improve the overall resolution for beam events (through
- 12 the combination of charge and light signals).