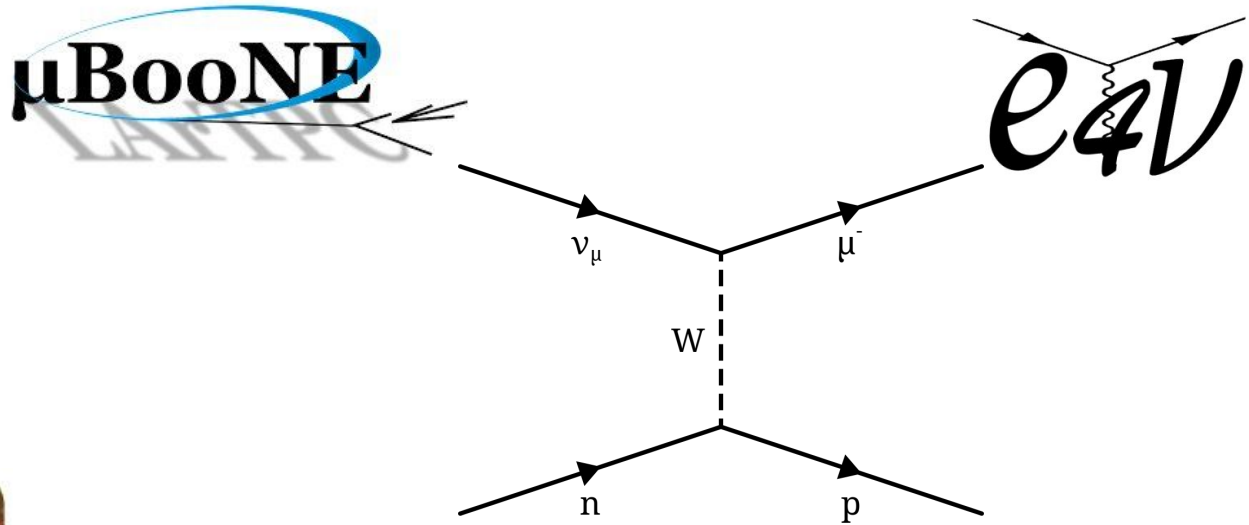
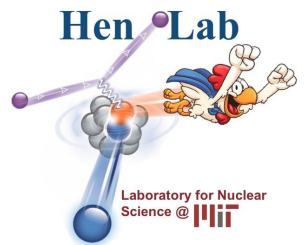


# Exclusive Quasielastic-like $\nu_{\mu}$ - $^{40}\text{Ar}$ Interactions in MicroBooNE and Connections with Electron Scattering

Wine & Cheese Seminar, Oct 30 2020



A. Papadopoulou (apapadop@mit.edu)  
on behalf of the MicroBooNE & e4ν collaborations

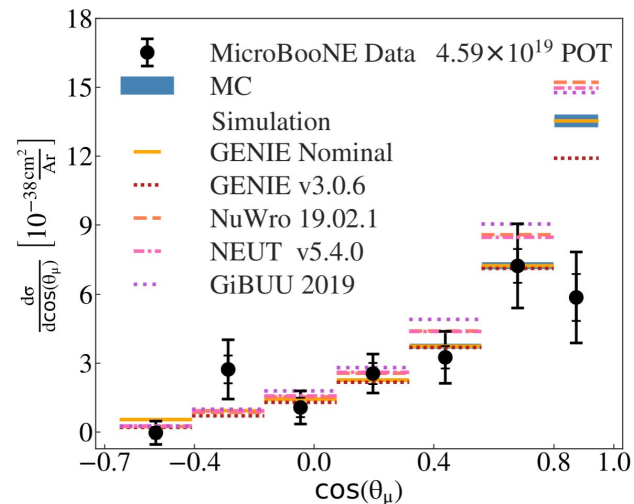


# Overview

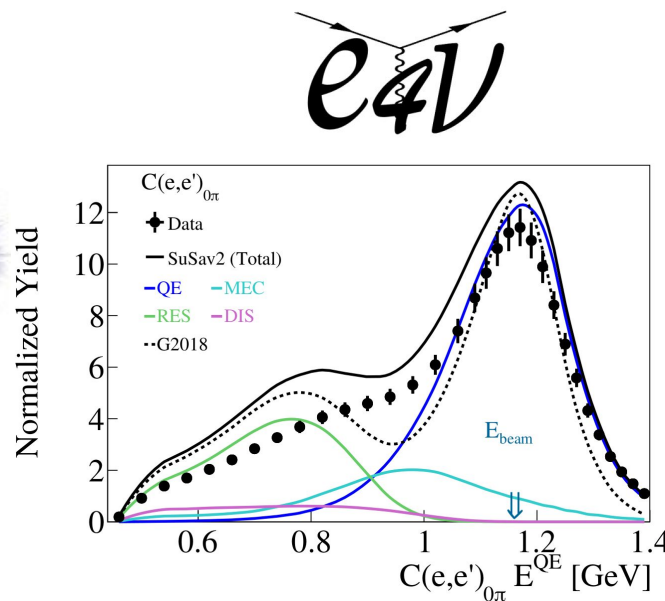
- Era of high precision neutrino oscillation measurements
- Liquid Argon Time Projection Chambers (LArTPCs) are state-of-the-art detectors in neutrino physics
- Demand for highly precise  $\nu$  cross section measurements

# Outline

 [Phys. Rev. Lett. 125, 201803 \(2020\)](#)

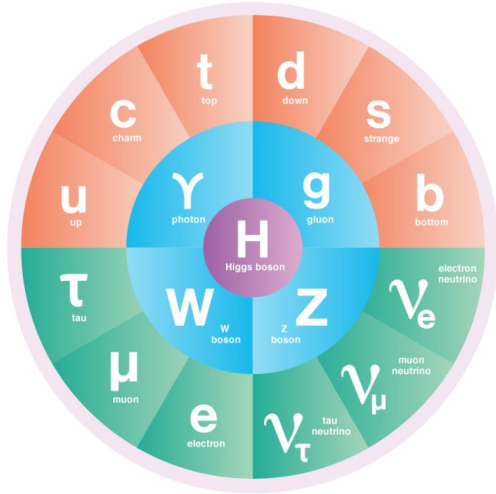


## Modeling Input



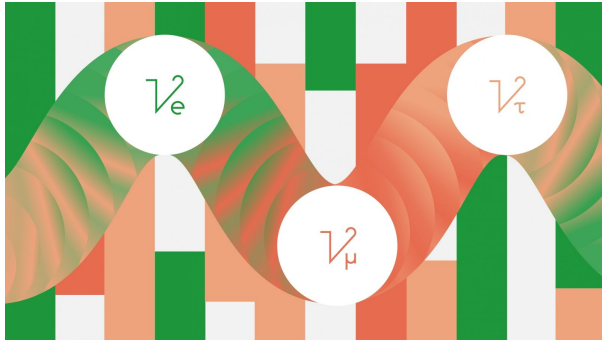
- First exclusive measurement of Quasielastic-like interactions using the MicroBooNE detector
- Connections with electron scattering

# Neutrinos



## Standard Model

- 3 (anti-)neutrinos of 3 flavours
- Massless
- Interact weakly (W / Z exchange)



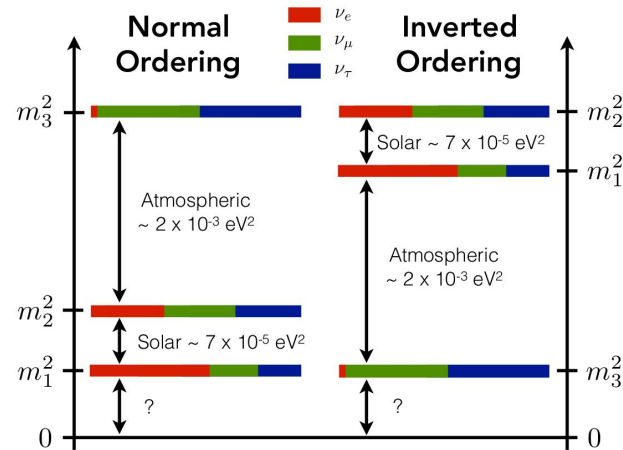
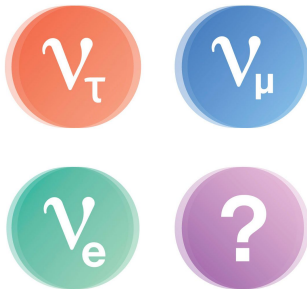
## Beyond Standard Model

- Neutrino oscillate → have mass

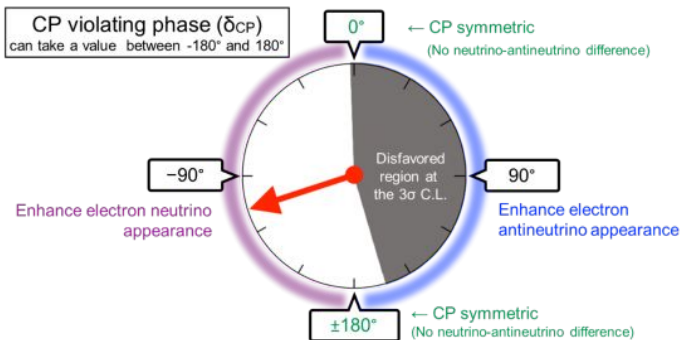
**Yet a lot of open questions ...**

# Open Questions

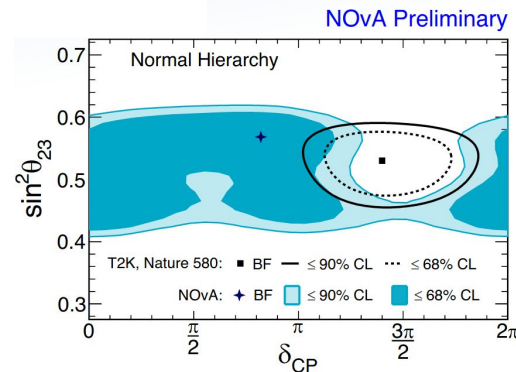
- Absolute mass
- Mass ordering
- CP violation
- Nature of neutrinos (Dirac vs Majorana)
- Sterile neutrinos



T2K [Nature 580, 339–344 \(2020\)](#)



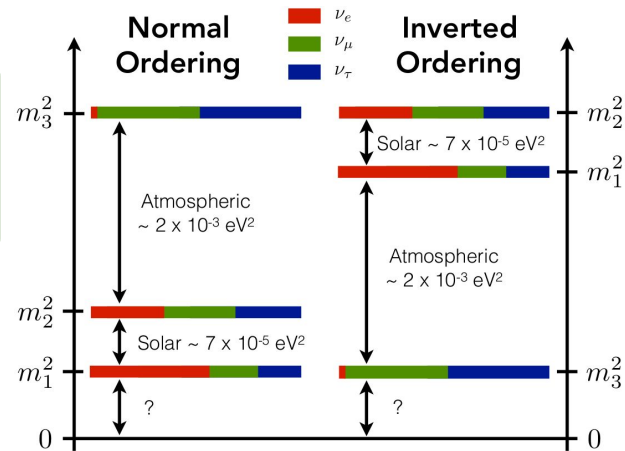
[A.Himmel, Neutrino 2020](#)



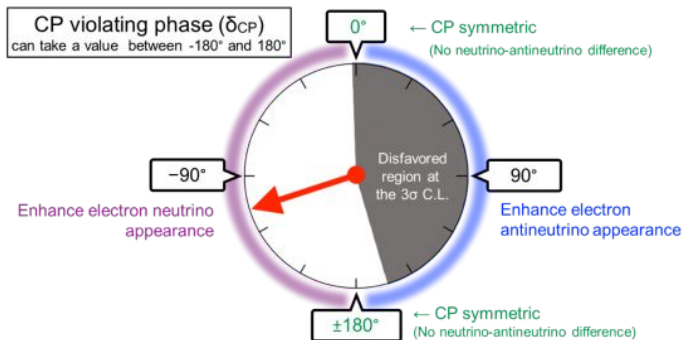
# Open Questions

- Absolute masses
- **Mass ordering**
- **CP violation**
- Nature of neutrinos (Dirac vs Majorana)
- **Sterile neutrinos**

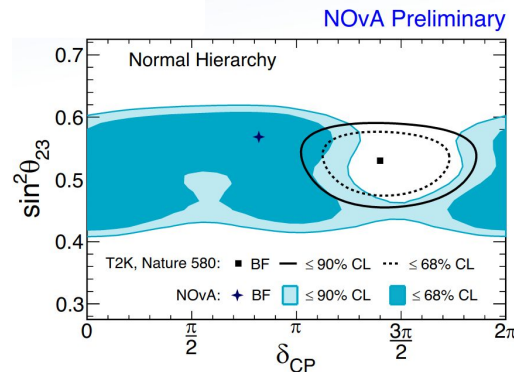
Investigated with short- & long-baseline oscillation experiments



T2K *Nature* 580, 339–344 (2020)



[A.Himmel, Neutrino 2020](#)



# Neutrino Oscillation Experiments

High precision appearance / disappearance measurements



Near Detector

Detect  $\nu_\alpha$

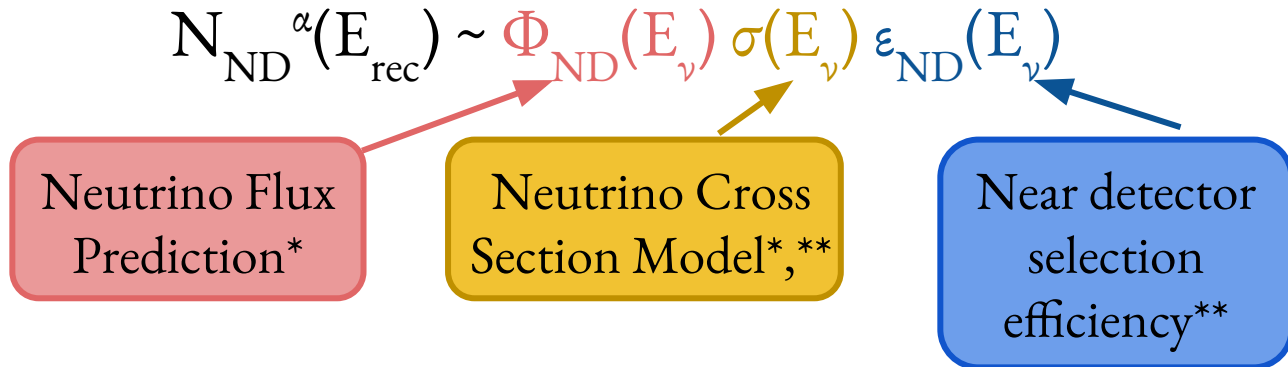
$N_{\text{ND}}^\alpha(E_{\text{rec}})$   
Number of events  
of flavor  $\alpha$

Appearance  
Measurement

Far Detector

$E_{\text{rec}}$  = Reconstructed  $\nu$  energy

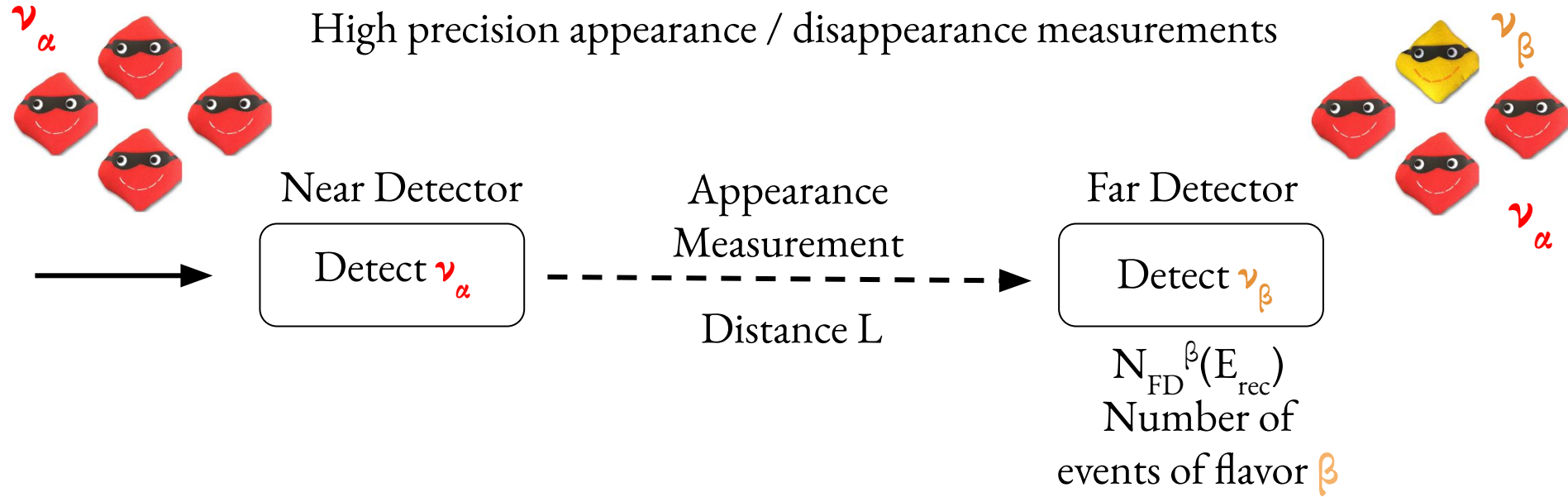
# Near Detector



\* $E_{\nu}$  = True  $\nu$  energy    \*\* $\sigma$  also depends on observables like muon momentum

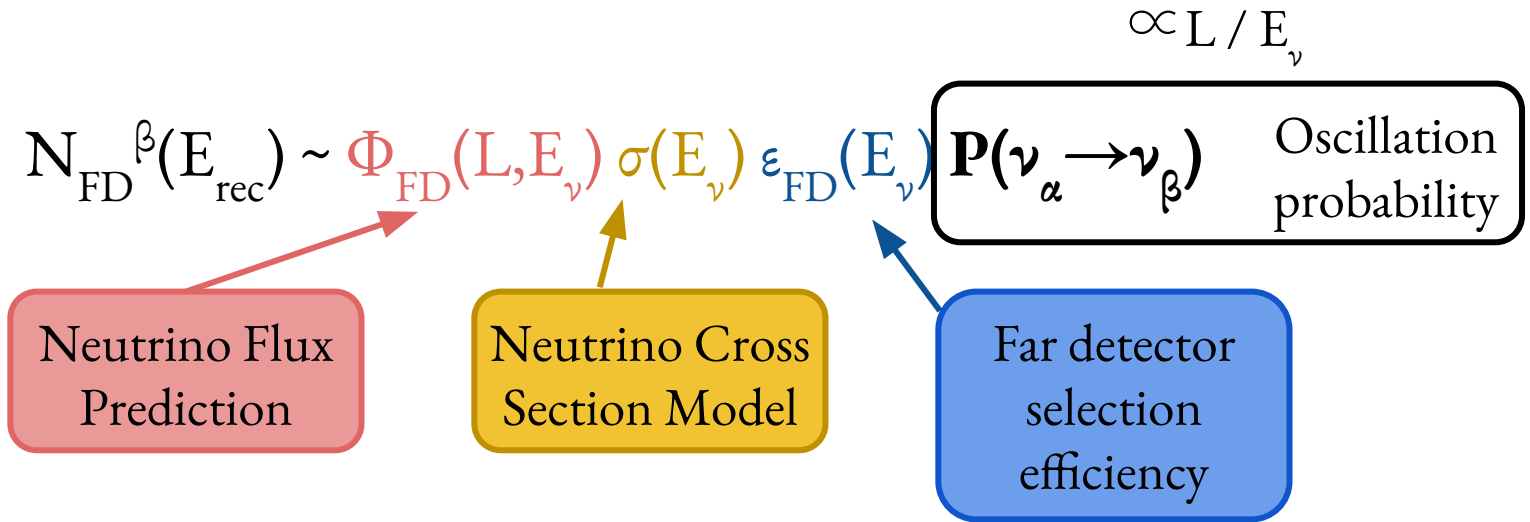


# Neutrino Oscillation Experiments



$E_{\text{rec}}$  = Reconstructed  $\nu$  energy

# Far Detector



# Far Detector

$$N_{\text{FD}}^{\beta}(E_{\text{rec}}) \sim \Phi_{\text{FD}}(L, E_{\nu}) \sigma(E_{\nu}) \varepsilon_{\text{FD}}(E_{\nu}) \mathbf{P}(\nu_{\alpha} \rightarrow \nu_{\beta})$$

Oscillation probability

Modeling Input

- Smearing relating  $E_{\nu}/E_{\text{rec}}$
- Signal / background topologies
- ...

# Far Detector

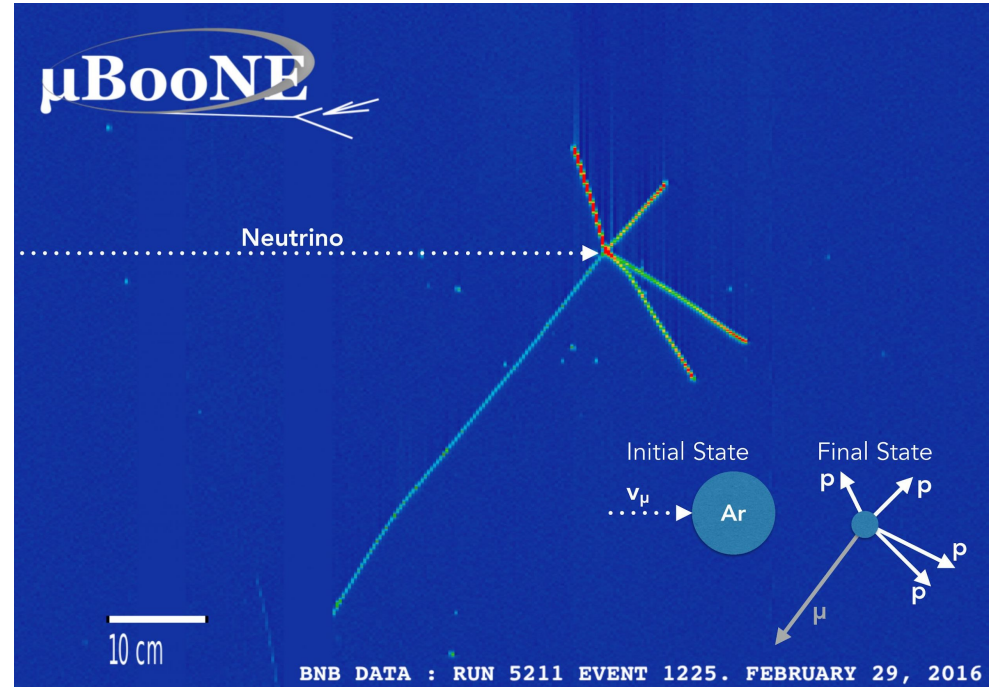
$$N_{\text{FD}}^{\beta}(E_{\text{rec}}) \sim \Phi_{\text{FD}}(L, E_{\nu}) \sigma(E_{\nu}) \epsilon_{\text{FD}}(E_{\nu}) \mathbf{P}(\nu_{\alpha} \rightarrow \nu_{\beta})$$

Oscillation probability

Need for high precision  $\nu$  cross section measurements  
& for state-of-the-art detectors

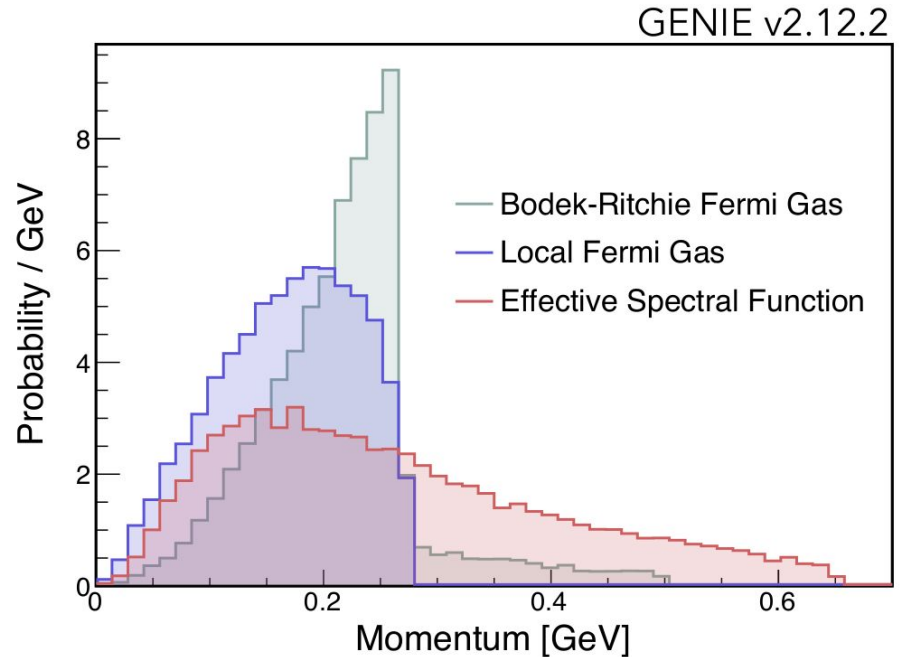
# Neutrino Interactions

**Nuclear effects** not well understood especially on heavy elements



# Fermi Motion

Neutrino scattering off nucleon in complex nuclear system



Credit: M. Del Tutto

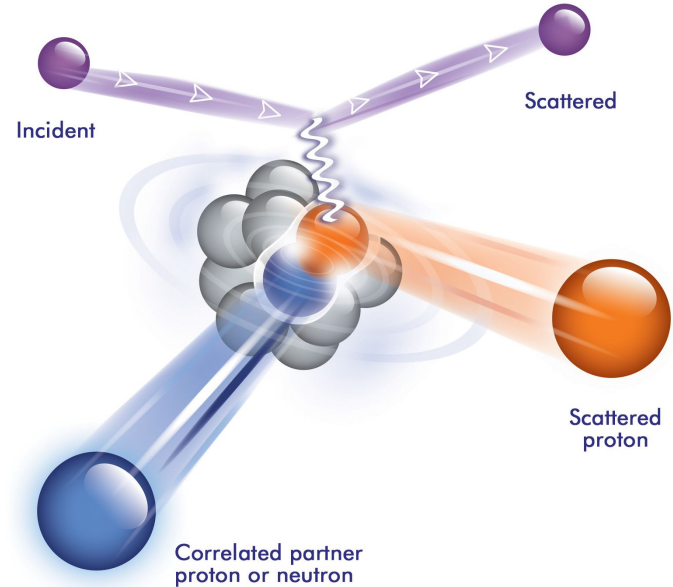
# Nucleon-nucleon Correlations

- Meson Exchange Current (MEC)
- Short Range Correlations

*Rev. Mod. Phys.* 89, 045002 (2017)

Need to measure outgoing particles

**Possible in LArTPCs!**



# RPA Effects

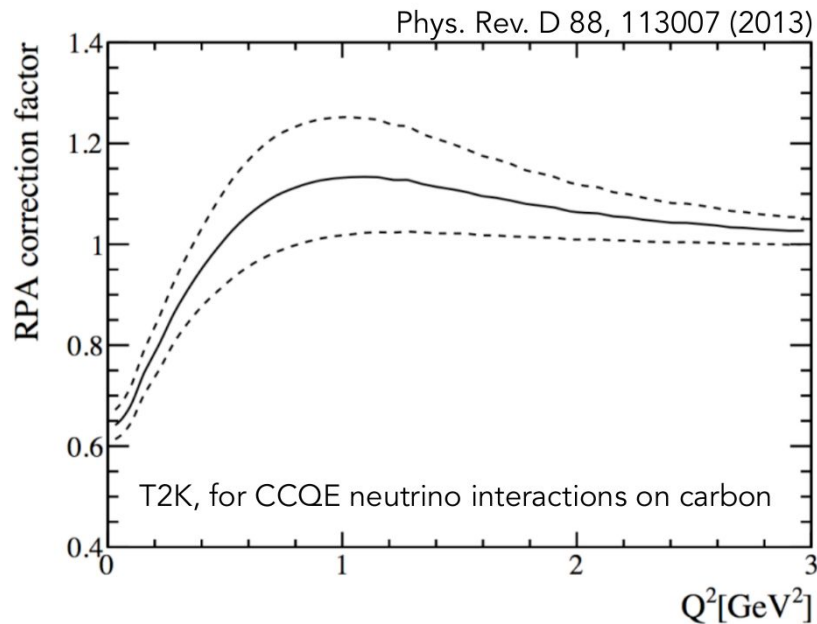
Use random phase approximation method (RPA)

Suppression at low  $Q^2$

Calculations on carbon & extrapolation to argon

[Phys. Rev. C 97 044616 \(2018\)](#)

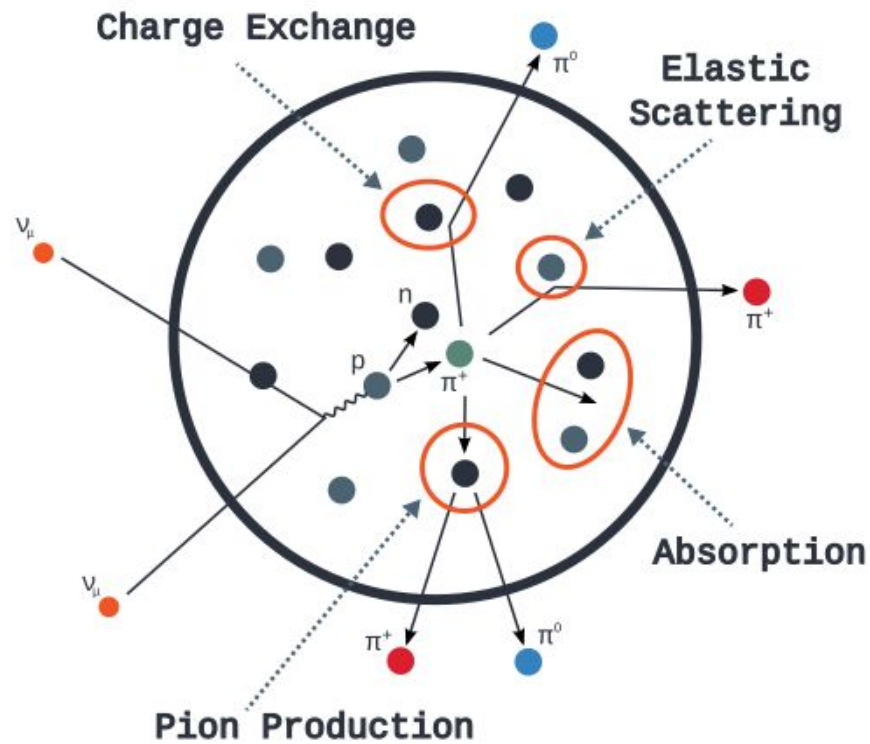
[Phys. Rev. C 92 024606 \(2015\)](#)





# Final State Interactions

Re-absorption or multiple hadron production in final state



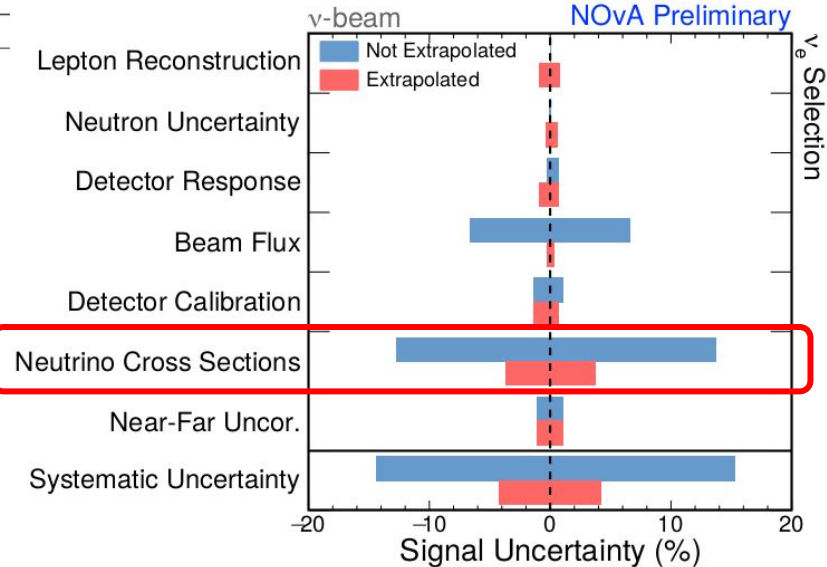
Credit: T. Golan

# Cross Section Uncertainty Dominate

T2K *Nature* **580**, 339–344 (2020)

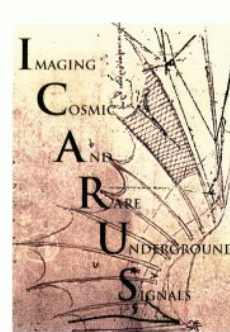
Type of Uncertainty	$\nu_e/\bar{\nu}_e$ Candidate Relative Uncertainty (%)
Super-K Detector Model	1.5
Pion Final State Interaction and Rescattering Model	1.6
Neutrino Production and Interaction Model Constrained by ND280 Data	2.7
Electron Neutrino and Antineutrino Interaction Model	3.0
Nucleon Removal Energy in Interaction Model	3.7
Modeling of Neutral Current Interactions with Single $\gamma$ Production	1.5
Modeling of Other Neutral Current Interactions	0.2
<b>Total Systematic Uncertainty</b>	<b>6.0</b>

**5% out of 6%!**

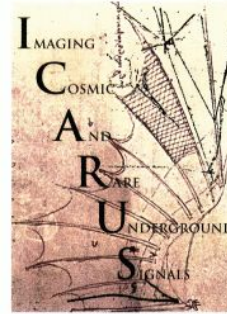
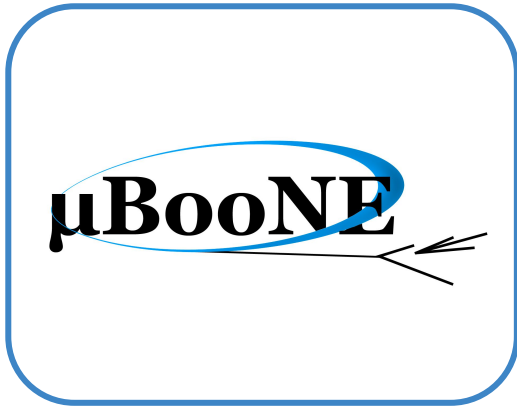


# Future Experiments

- Oscillation parameter extraction with percent level uncertainties
- Demand for precise  $\nu$  cross section modelling
- Head start with Short-Baseline Neutrino (SBN) Program



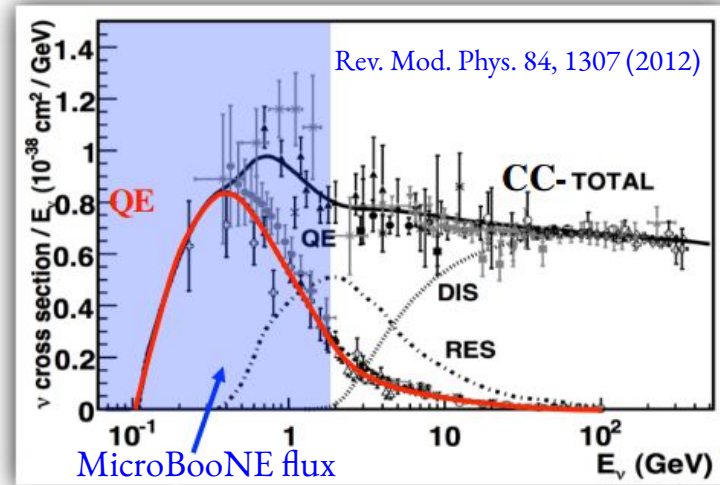
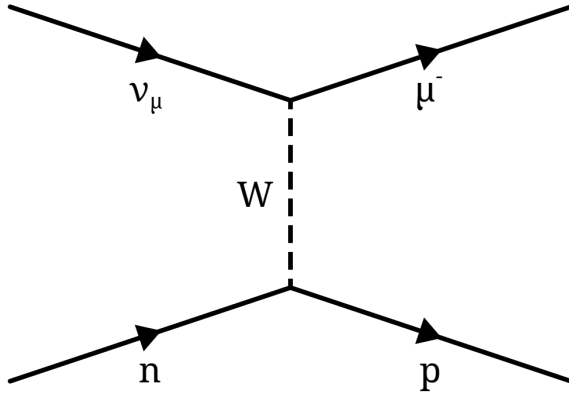
# LArTPC Experiments



Ideal for more detailed topology measurements

**MicroBooNE** has already recorded  $\sim 500\text{k}$   $\nu$  scattering events over the past 5 years  
(an initial  $\sim 5\%$  data set used in this analysis)

# Charged Current Quasi-elastic Interactions



Simple  $1p0\pi$  topology

Dominant at energies relevant for MicroBooNE

# Previous Measurements

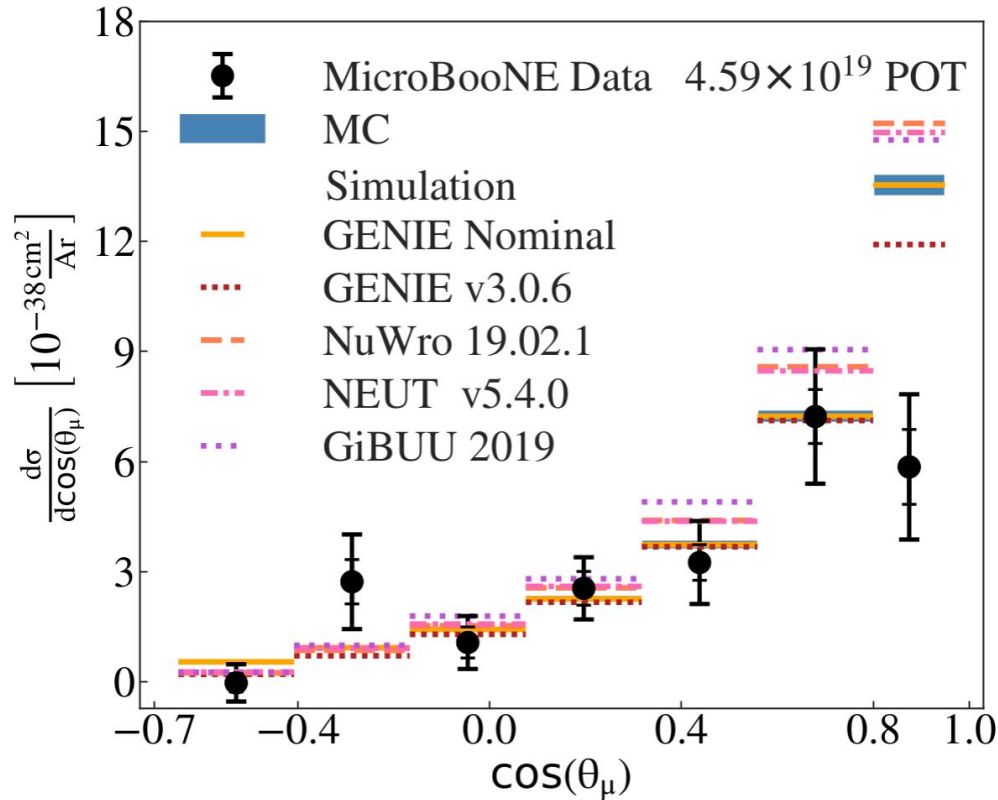
Experiment	Target
ArgoNeuT	Ar
MINERvA	CH, C/CH, Fe/CH, Pb/CH
MINOS	Fe
NOMAD	C
SciBooNE	CH
T2K	CH, H <sub>2</sub> O
NOvA	CH
MiniBooNE	CH
MicroBooNE	Ar

No CCQE-like measurement

Different nuclei

Large data sets for exclusive measurements

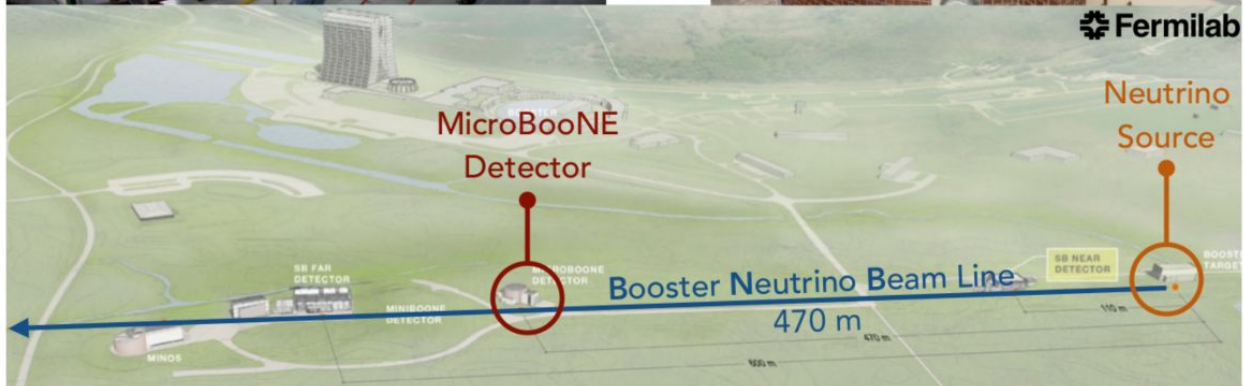
# This talk



First measurement of differential charged current quasielastic-like  $\nu_\mu - {}^{40}\text{Ar}$  scattering cross sections using the MicroBooNE detector

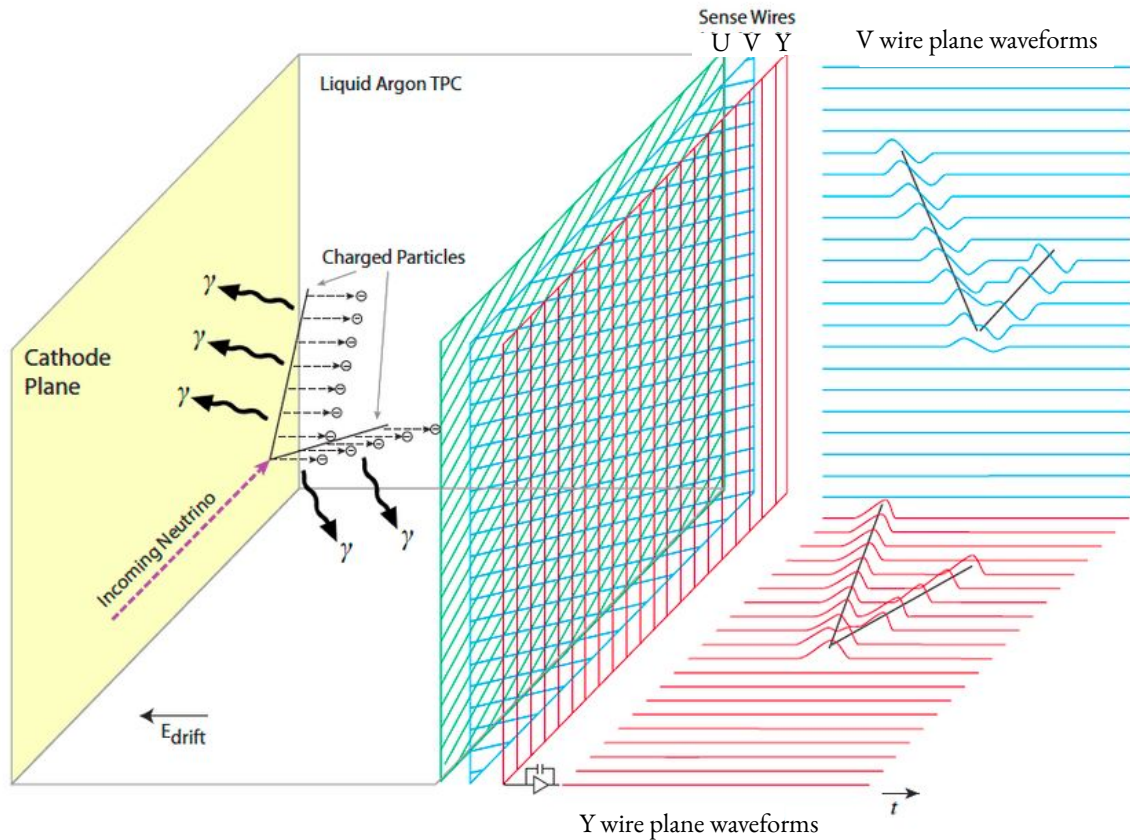
[Phys. Rev. Lett. 125, 201803 \(2020\)](#)

# MicroBooNE



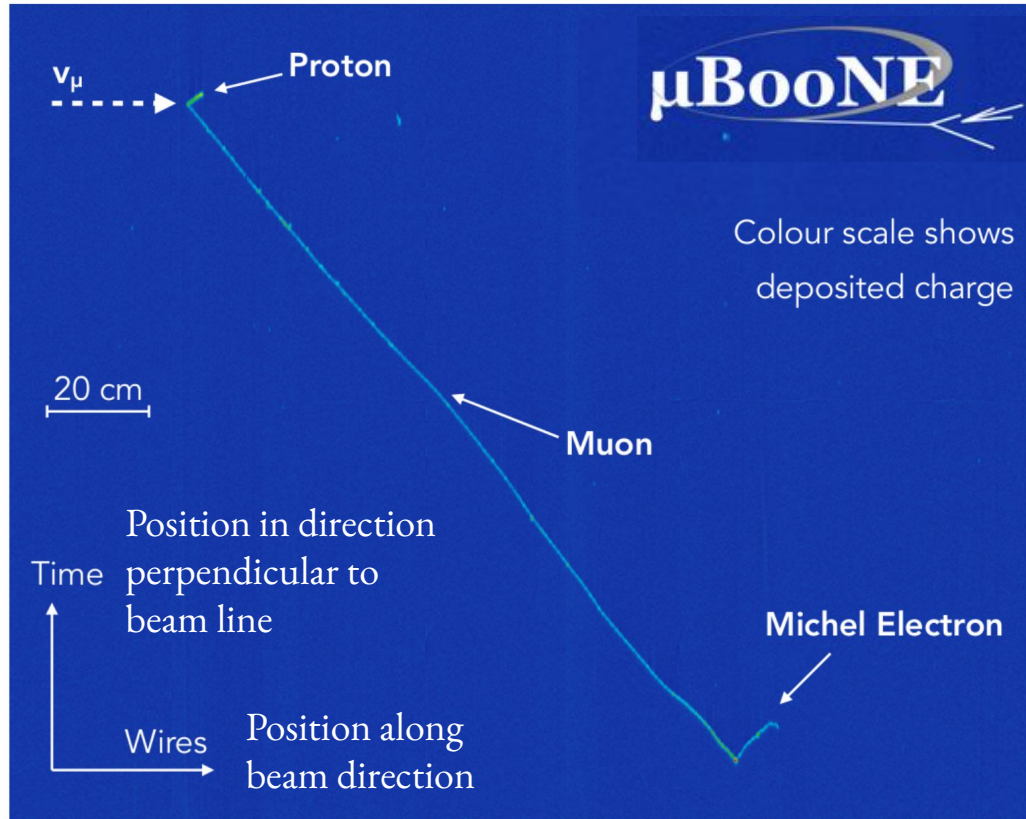


# Time Projection Chambers



- 3 wire planes
- 8192 gold coated wires
- 3 mm wire spacing
- 32 PMTs

# Event Displays



# Event Reconstruction

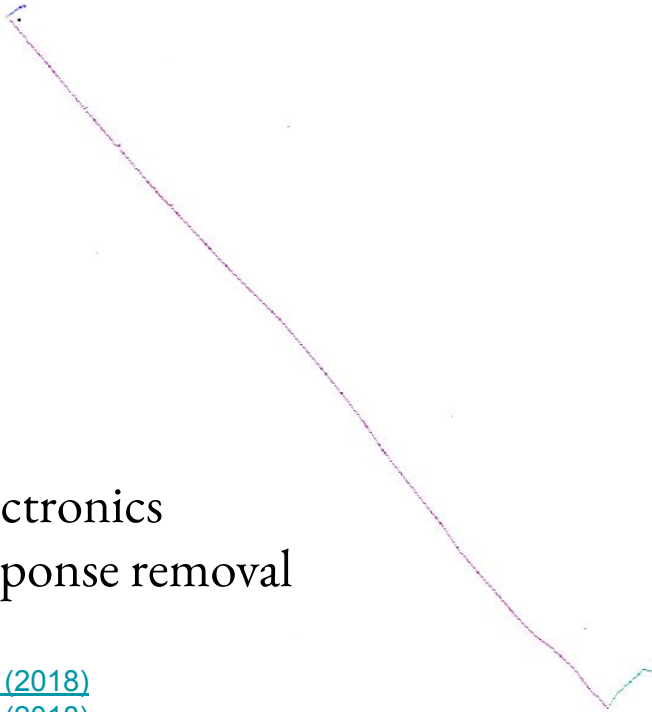
From raw hits to  
particle reconstruction

## **Pandora Pattern Recognition**

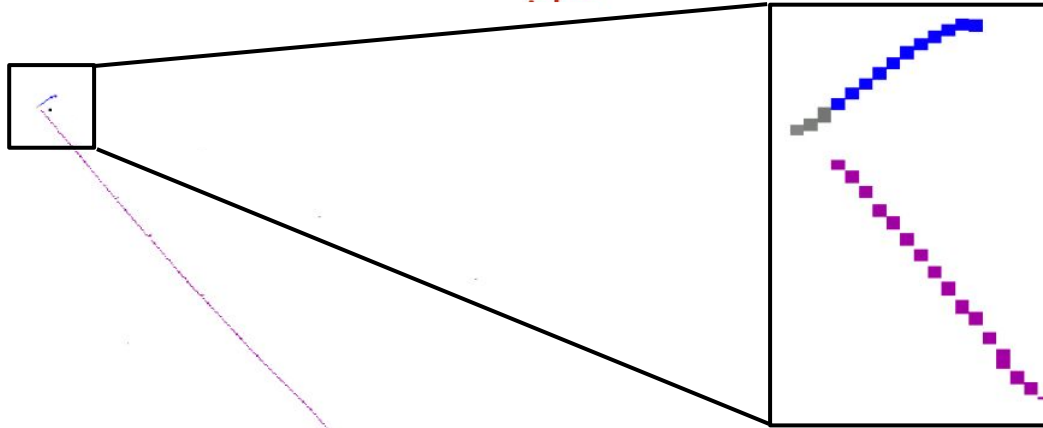
- [Eur. Phys. J. C78, 1, 82 \(2018\)](#)

Readout electronics  
and field response removal

[JINST 13, P07006 \(2018\)](#)  
[JINST 13, P07007 \(2018\)](#)



# Event Reconstruction

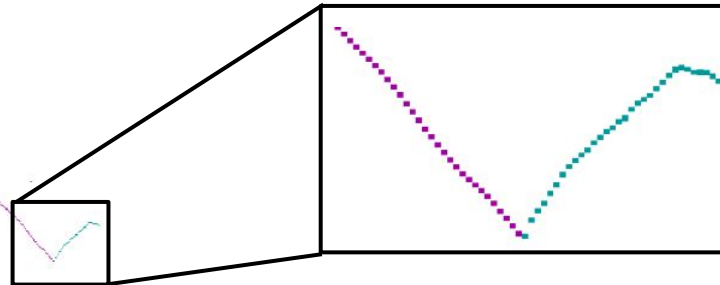


From raw hits to  
particle reconstruction

**Pandora Pattern  
Recognition**

[Eur. Phys. J. C78, 1, 82 \(2018\)](#)

Readout electronics  
and field response removal

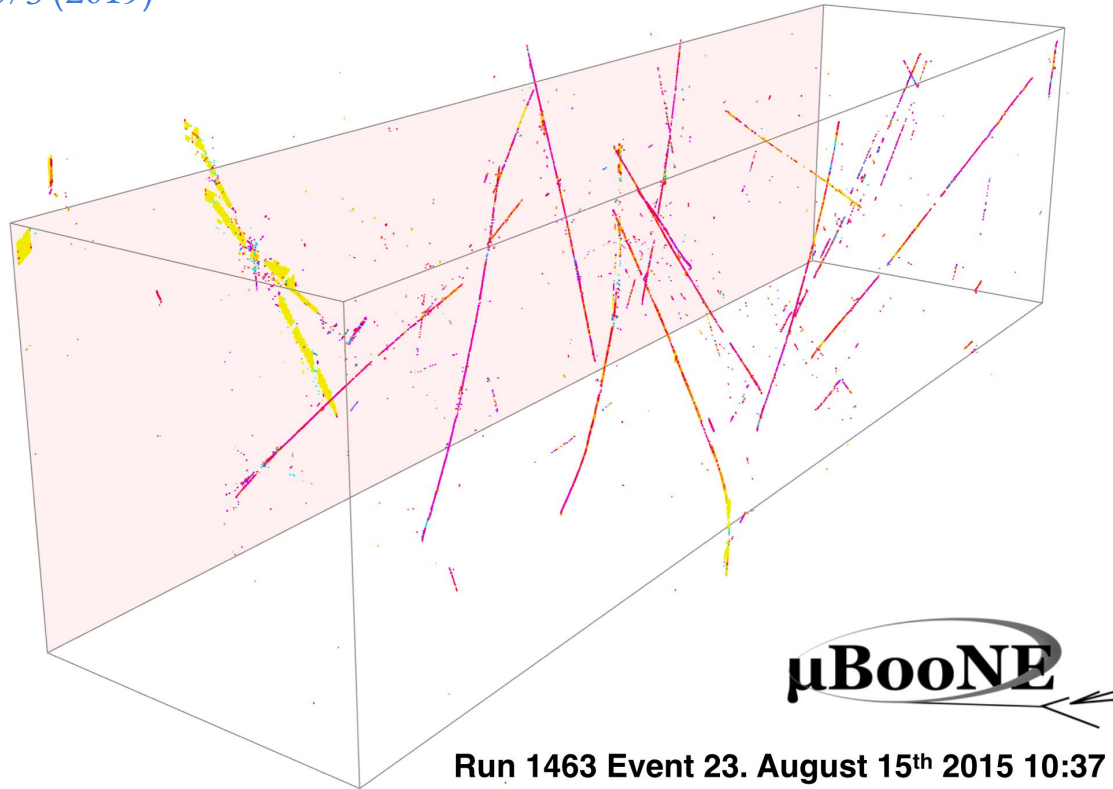


[JINST 13, P07006 \(2018\)](#)

[JINST 13, P07007 \(2018\)](#)

# Cosmic Background Dominance

Eur. Phys. J. C 79 673 (2019)

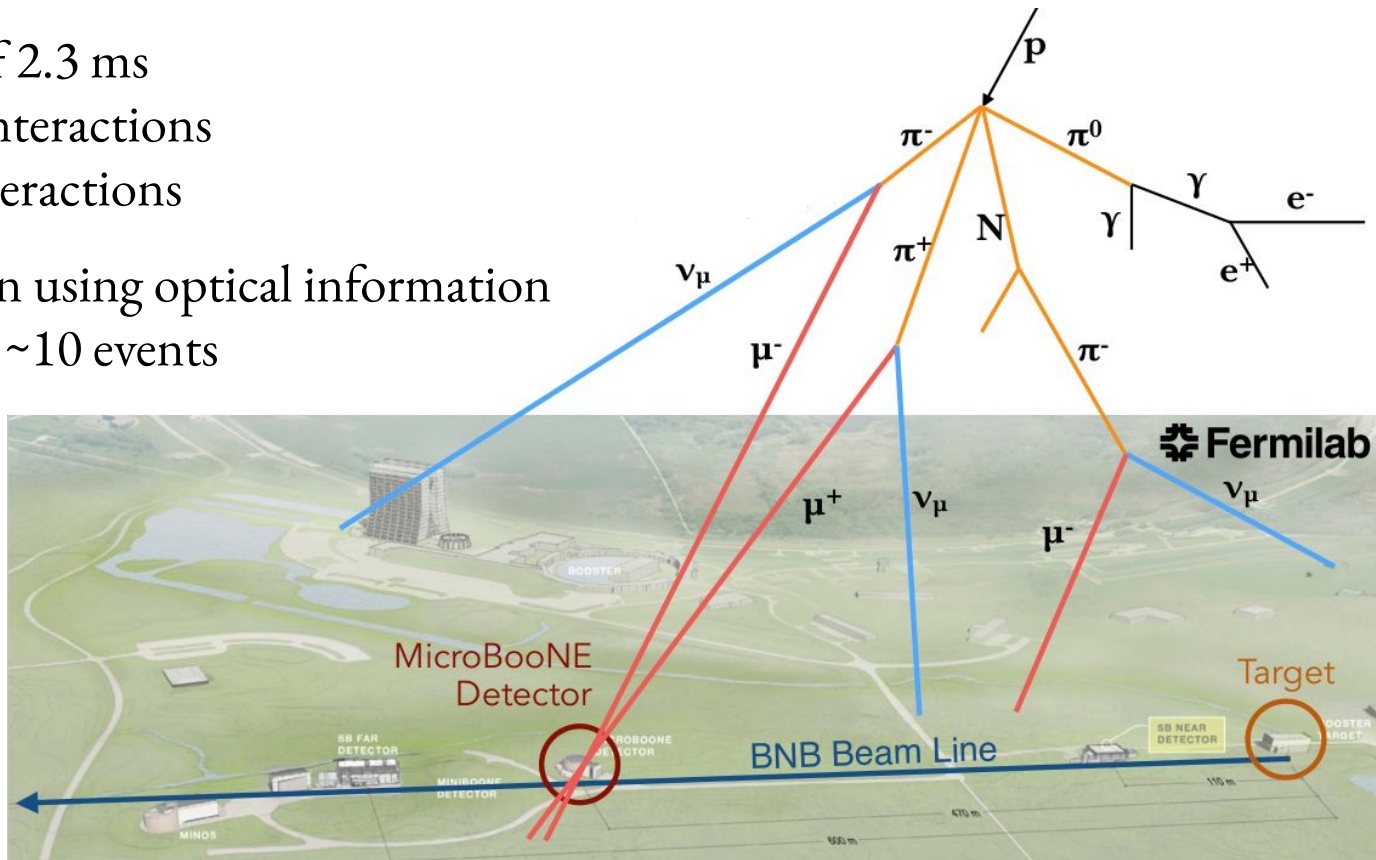


# Cosmic Background Dominance

Readout window of 2.3 ms

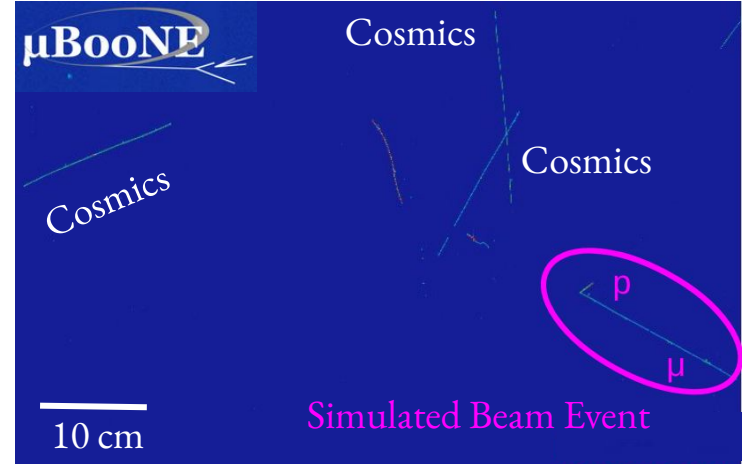
- ~20 cosmic interactions
- ~0.0017  $\nu$  interactions

Significant reduction using optical information  
to 1  $\nu$  interaction in ~10 events



# Overlay Samples

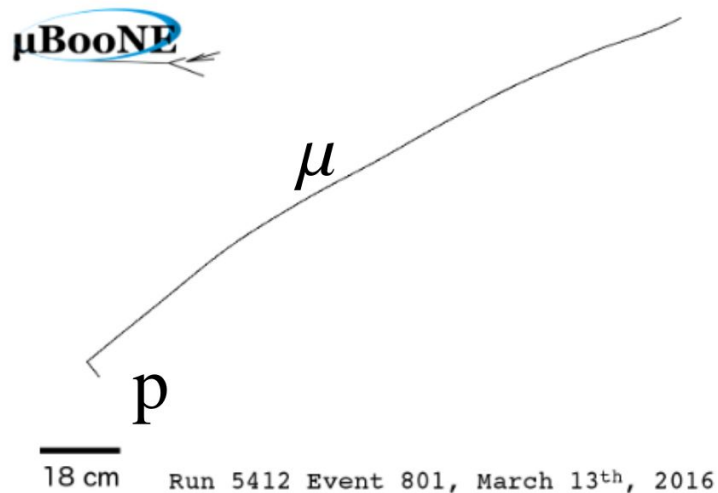
- First MicroBooNE analysis using novel technique in the LArTPC community
- Simulated events on top of real cosmic events
- Will be used as default simulation technique by MicroBooNE



# Signal Definition

Vertex of two tracks

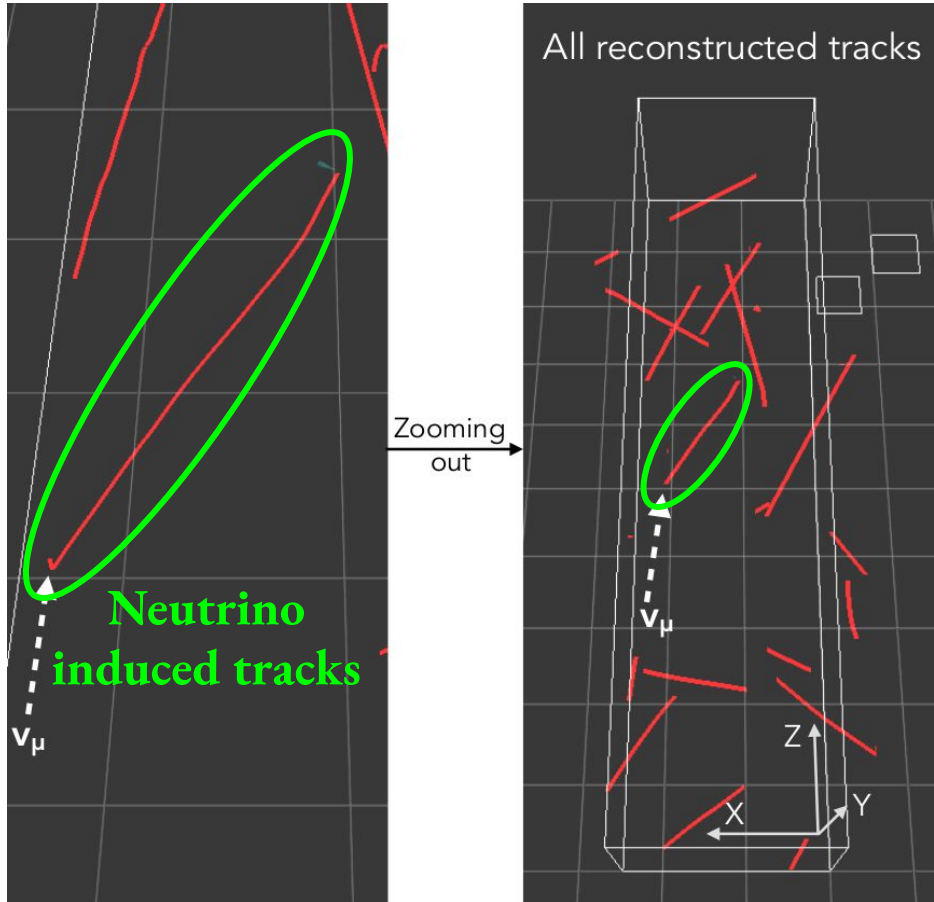
- 1 muon ( $> 100 \text{ MeV}/c$ )
- 1 proton ( $> 300 \text{ MeV}/c$ )
- No  $\pi^\pm$  ( $> 70 \text{ MeV}/c$ )



CC1p0 $\pi$  topology



# CC1p0 $\pi$ Event Selection



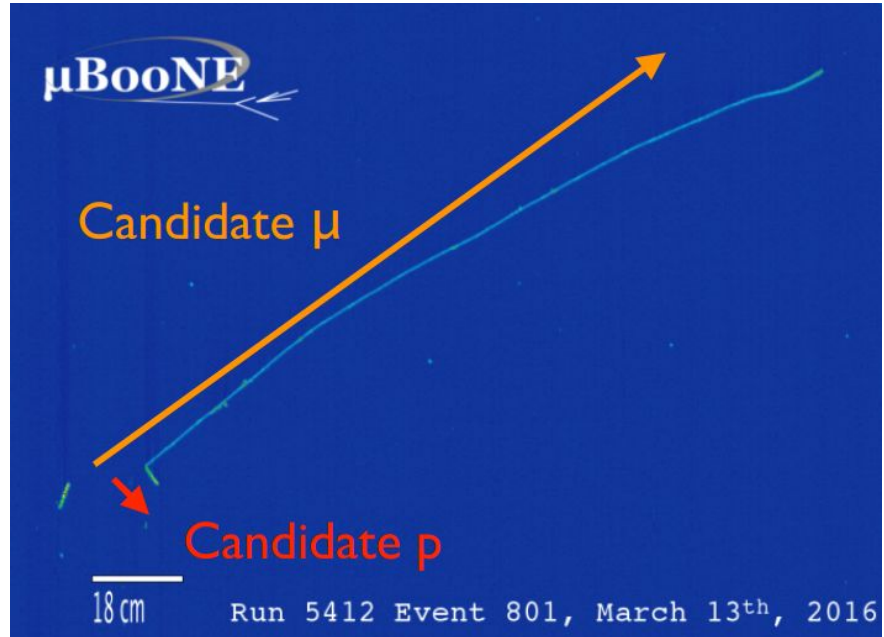
Detector and kinematics-based cuts  
to reject cosmics & to enhance signal

[Eur. Phys. J. C 79 673 \(2019\)](#)

Credit: M. Del Tutto

# Cosmic Rejection

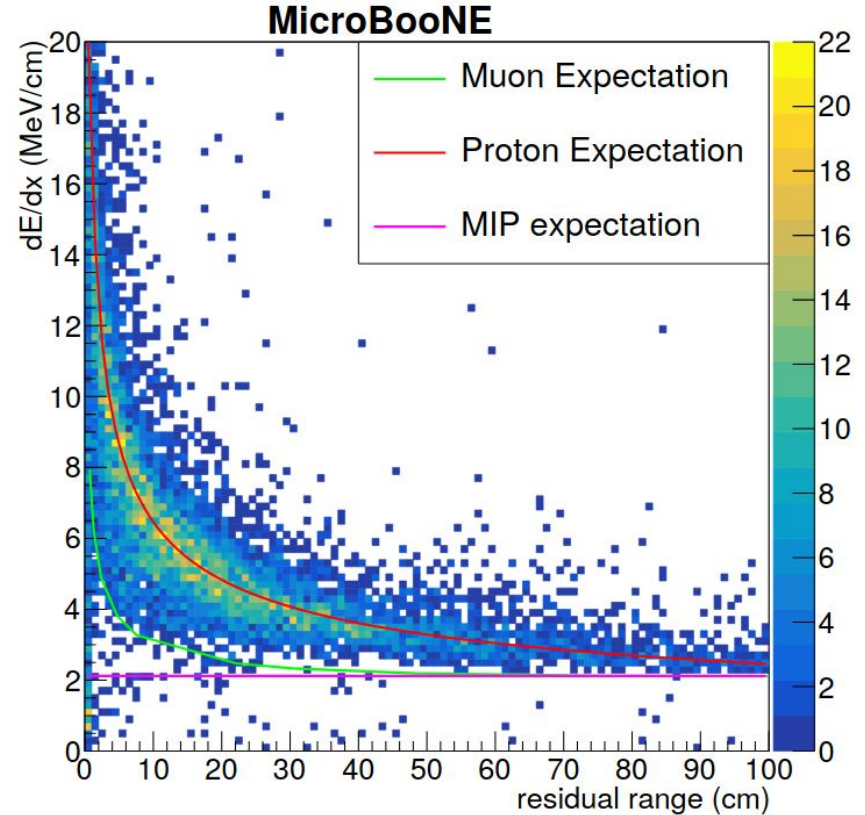
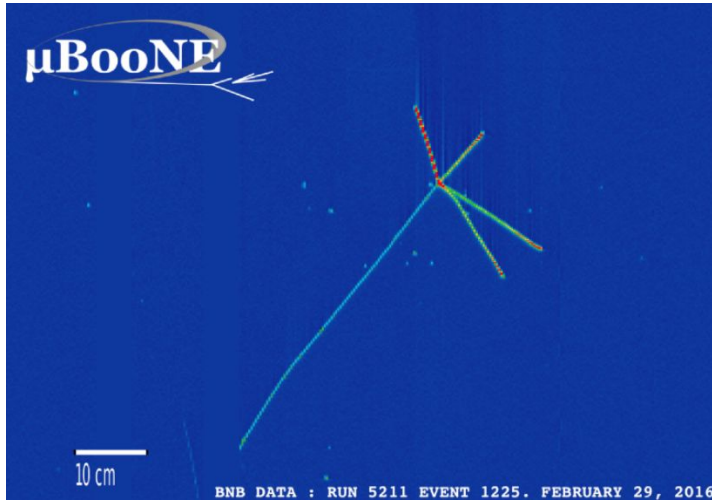
- Energy deposition profile (particle identification)
- Track length
- Scintillation light at vertex
- Non-collinearity  
 $|\Delta\theta_{\mu p} - 90^\circ| < 55^\circ$
- Charge deposition at vertex



# Energy Deposition Profile

Protons identified by Bragg peak  
in last 30 cm of track

Low proton threshold @ **300 MeV/c**



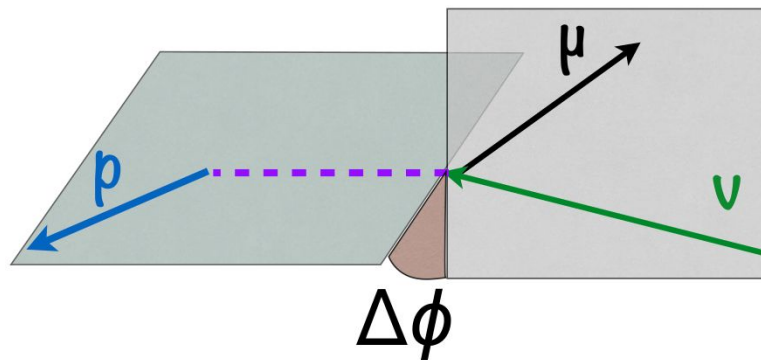
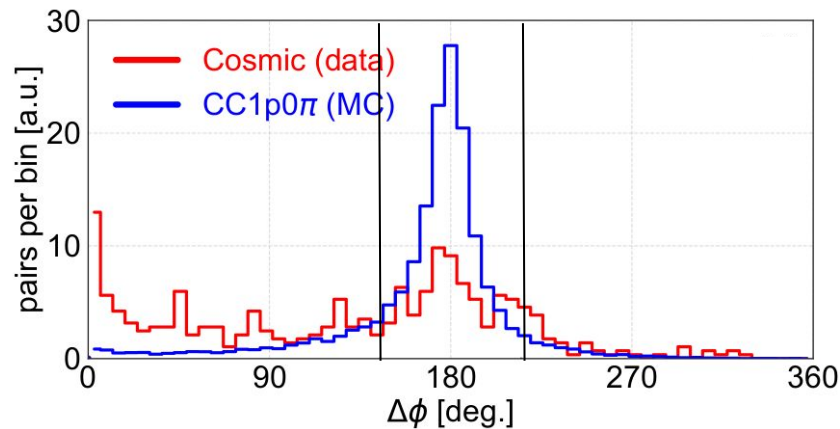
# CCQE Enhancement

## Coplanarity

$$|\Delta\phi_{\mu p} - 180^\circ| < 35^\circ$$

Transverse imbalance

$$P_T = |\vec{P}_\mu + \vec{P}_p|_T < 0.35 \text{ GeV}/c$$



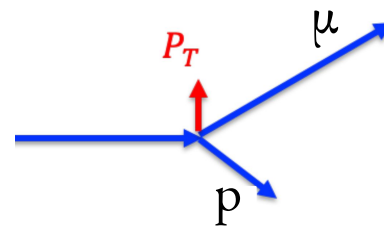
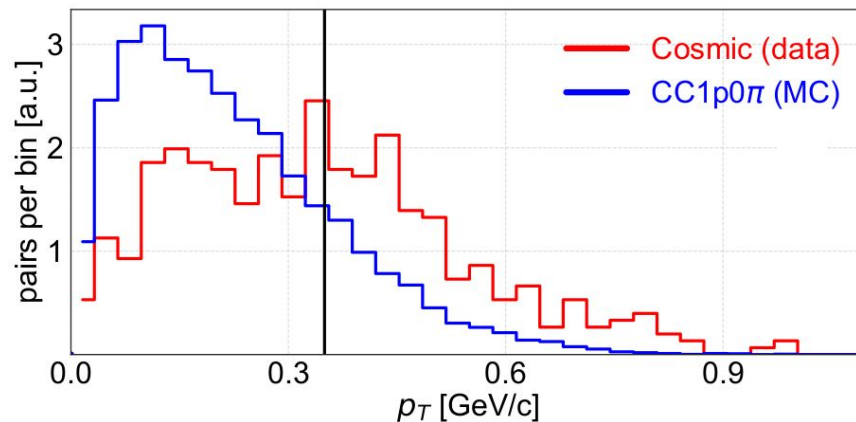
# CCQE Enhancement

Coplanarity

$$|\Delta\phi_{\mu p} - 180^\circ| < 35^\circ$$

**Transverse imbalance**

$$P_T = |\vec{P}_\mu + \vec{P}_p|_T < 0.35 \text{ GeV}/c$$



# Statistics

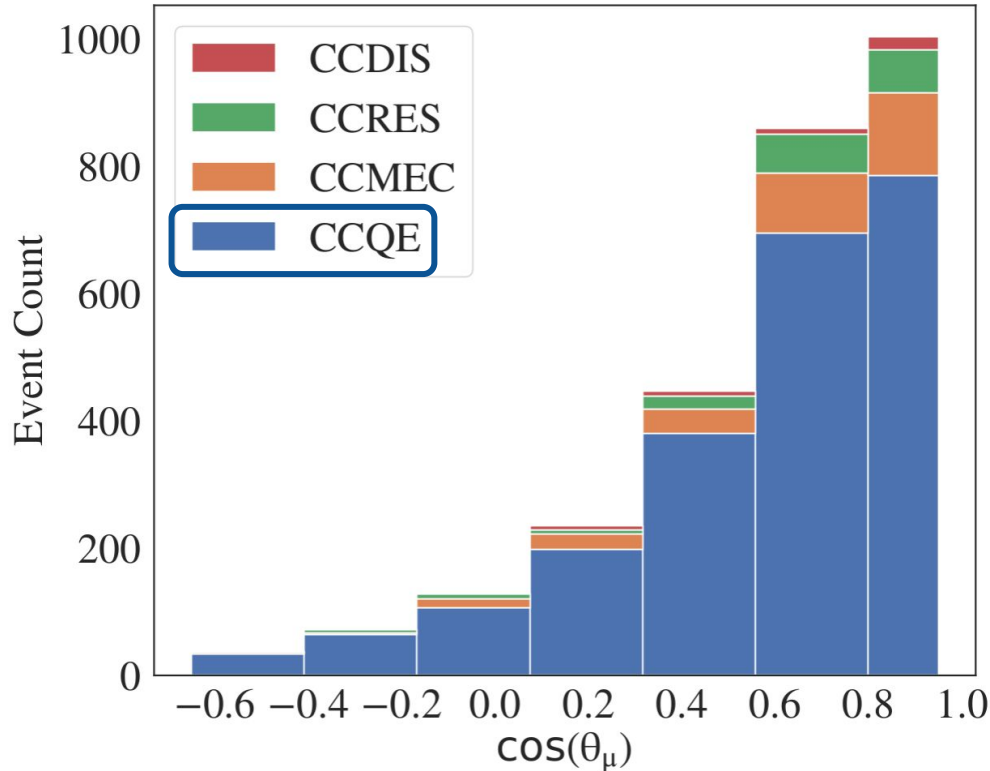
After selection cuts

# data events: 410

Using our MC, we report  
a  $CC1p0\pi$  topological signal purity  $\sim 84\%$

# CCQE Dominance

## MicroBooNE Simulation



Interaction Mode	Fractional Contribution (%)
CCQE	81.1
CCMEC	10.9
CCRES	6.6
CCDIS	1.4



GENIE v2.12.2

# Cross Section Extraction

- Event selection in data sample
- Cosmic background subtraction
- Beam related MC background subtraction
- Detection efficiency correction

$$\frac{d\sigma}{dX_n} = \frac{N_n^{\text{on}} - N_n^{\text{off}} - B_n}{\epsilon_n \cdot \Phi_\nu \cdot N_{\text{target}} \cdot \Delta_n}$$

X = kinematic variable of interest



# Cross Section Extraction

$$\frac{d\sigma}{dX_n} = \frac{N_n^{\text{on}} - N_n^{\text{off}} - B_n}{\epsilon_n \cdot \Phi_\nu \cdot N_{\text{target}} \cdot \Delta_n}$$

$N^{\text{on}}$  → # events in beam-on data  
( $\nu$  interactions in beam spill time)

$N^{\text{off}}$  → # events in beam-off data  
(cosmic activity)

**Data**

**Constants**

$\Phi_\nu$  → integrated neutrino flux

$N_{\text{targets}}$  → number of target nuclei

$\Delta$  → bin width

# Cross Section Extraction

$$\frac{d\sigma}{dX_n} = \frac{N_n^{\text{on}} - N_n^{\text{off}} - \mathbf{B}_n}{\epsilon_n \cdot \Phi_\nu \cdot N_{\text{target}} \cdot \Delta_n}$$

B → beam related MC background

(mimic  $1\mu 1p$  signal but not true  $CC1p0\pi$ )

$\epsilon$  → effective detection efficiency

**MC**

# Effective Efficiency

$$\epsilon_n = \frac{\text{true CC1p0}\pi \text{ reconstructed in bin n}}{\text{true CC1p0}\pi \text{ generated in bin n}}$$

Reconstruction  
effects

Resolution  
effects

Bin migration  
effects

Unfolding technique

Results reported as a function of **true variables**

# Systematic Uncertainties

$$\text{Syst} = \text{Flux} \oplus \text{XSec} \oplus \text{Detector}$$


Framework from  
MiniBooNE collaboration  
[PRD 79, 072002, 2009](#)

More than 40 reweightable  
GENIE parameters  
[MICROBOONE-NOTE-1074-PUB](#)

- Longitudinal & transverse diffusion
- Light yield outside the TPC
- Electron lifetime
- Recombination model
- Space charge map
- Electronic response

[MICROBOONE-NOTE-1045-PUB](#)

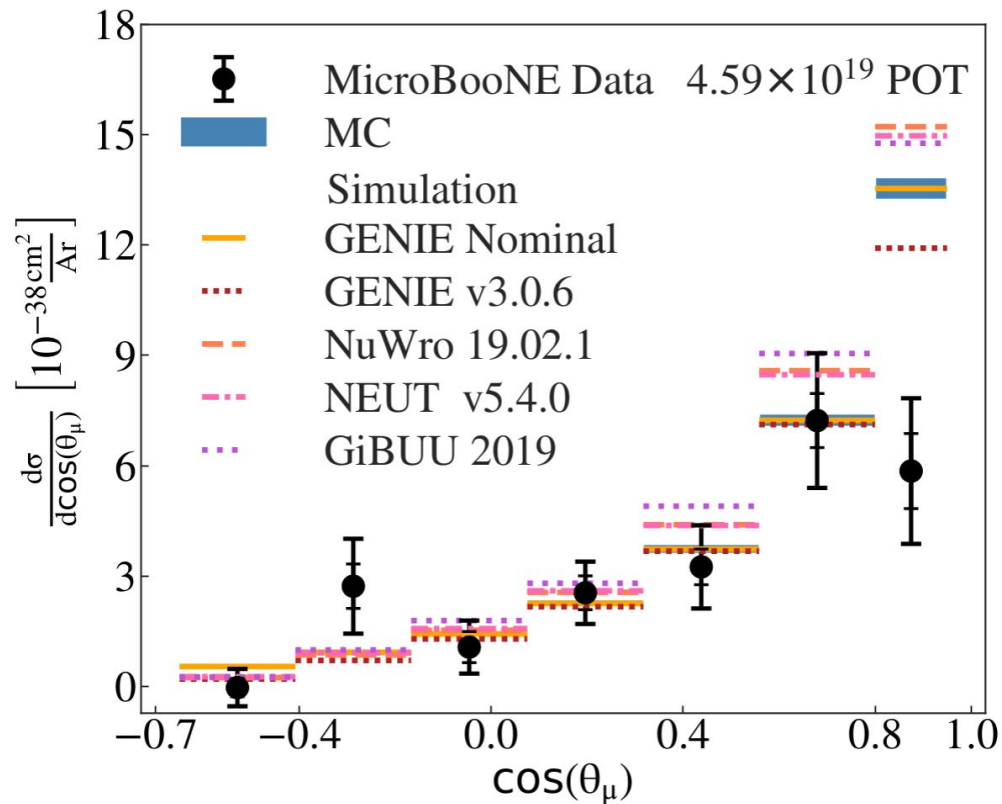
[MICROBOONE-NOTE-1032-PUB](#)

# Integrated Cross Section Uncertainty Breakdown

Source Of Uncertainty	Relative Uncertainty (%)
Beam Flux & POT	19
Cross Section Modelling	7
Detector Response	18
Efficiency $\mu$ -p decoupling	6
Statistical	16
<b>Total</b>	<b>32</b>

Already significant improvements for future analyses on MicroBooNE to reduce detector & cross section systematic uncertainties

# Muon Angular Results




Good agreement with models,  
except at very forward  
muon scattering angles

GENIE Nominal = GENIE v2.12.2

GENIE v3.0.6 = G18\_10a\_02\_11a

# Model Comparisons

Model Component	 <b>GENIE Nominal</b> <b>v2.12.2</b> <b>This work</b>	<b>GENIE v3.0.6</b> <b>G18_10a_02_11a</b> <b>Future analyses</b>	<b>NuWro</b> <b>(19.02.1)</b>	<b>NEUT</b> <b>(v5.4.0)</b>	<b>GiBUU</b> <b>(2019)</b>
<b>Nuclear Model</b>	Bodek-Ritchie Fermi Gas [1]	Local Fermi Gas [2, 3]	Local Fermi Gas [2, 3]	Local Fermi Gas [2, 3]	Consistent nuclear medium corrections
<b>Quasi-elastic</b>	Llewellyn-Smith [4]	Nieves [2, 3]	Nieves [2, 3]	Nieves [2, 3]	LFG model for nucleon momenta
<b>MEC</b>	Empirical [5]	Nieves [2, 3]	Nieves [2, 3]	Nieves [2, 3]	Separate MEC model [11]
<b>Resonant</b>	Rein-Seghal [6]	Berger-Seghal [7]	Berger-Seghal [7]	Berger-Seghal [7]	
<b>Coherent</b>	Rein-Seghal [6]	Berger-Seghal [7]	Berger-Seghal [7]	Rein-Seghal [6]	Propagates final state particles according to the BUU equations [11]
<b>FSI</b>	hA [8]	hA2018 [8]	Oset [10]	Oset [10]	

Significant differences in model components used  
in GENIE Nominal and other generators

# Model Comparisons

Model Component



**GENIE Nominal**  
v2.12.2  
This work

GENIE v3.0.6  
G18\_10a\_02\_11a  
Future analyses

NuWro  
(19.02.1)

NEUT  
(v5.4.0)

GiBUU  
(2019)

Nuclear Model

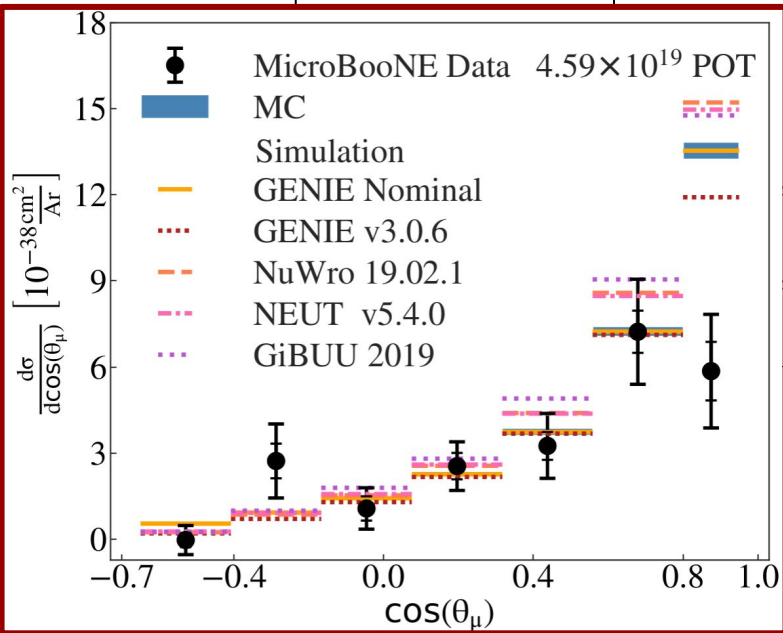
Bodek-Ritchie Fermi  
Gas [1]

Local Fermi  
Gas [2, 3]

Local Fermi  
Gas [2, 3]

Local Fermi  
Gas [2, 3]

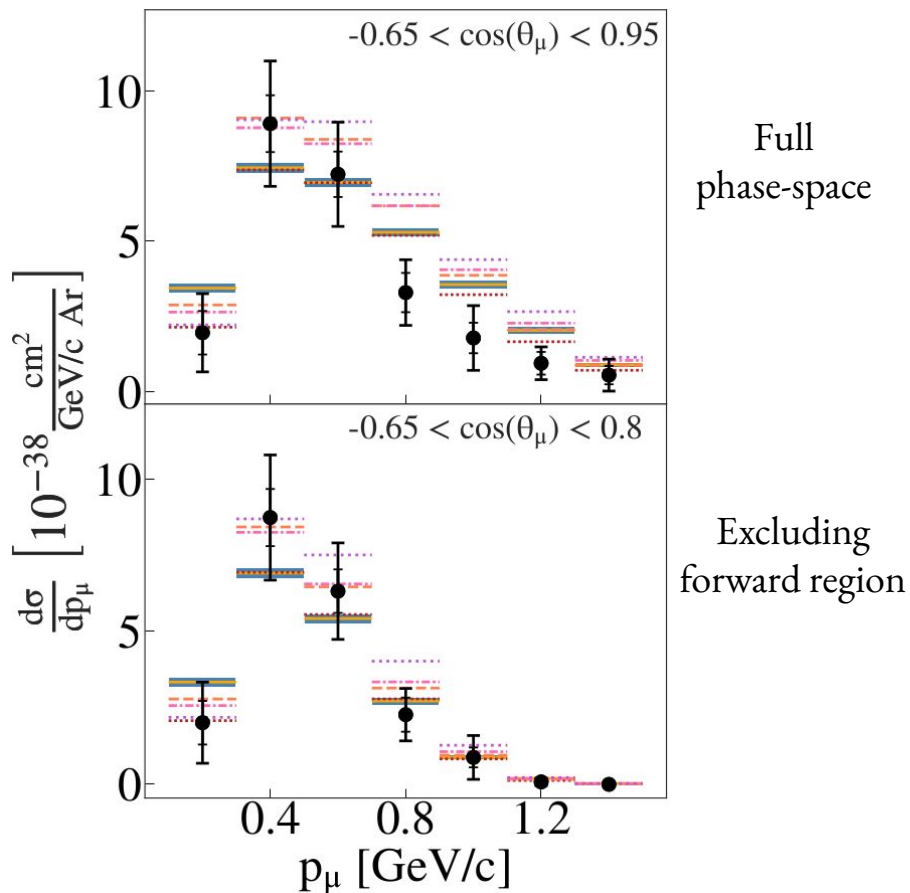
Consistent nuclear  
medium corrections



**None of them reproduces deficit  
in forward direction**

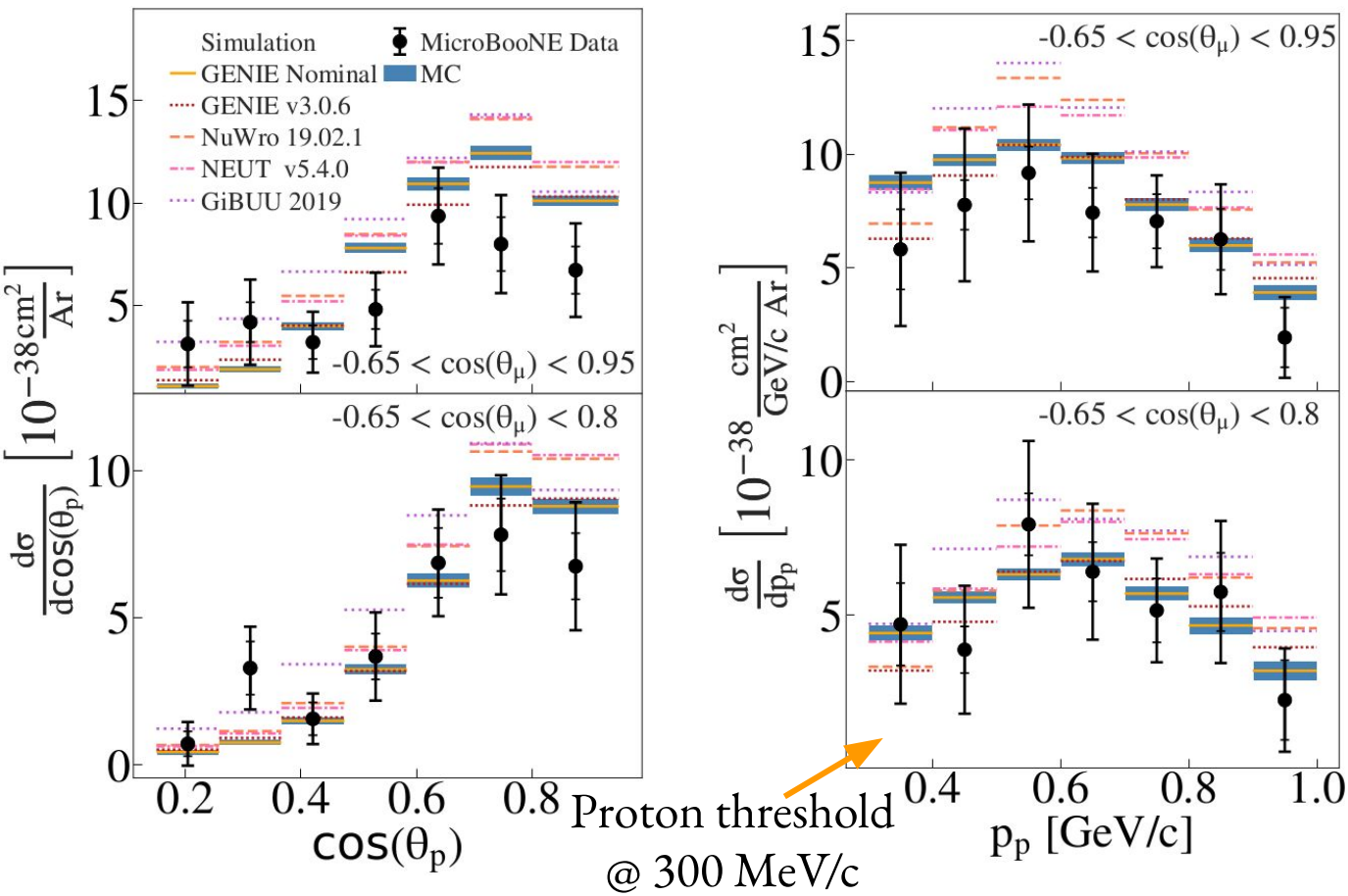


# Muon Momentum Results



Improved agreement  
if forward muon angles  
are excluded

# Proton Results



Full  
phase-space

Improved agreement  
if forward muon angles  
are excluded

Excluding  
forward region

# $\chi^2$ Comparisons

Linear sum of 1D  $\chi^2$ 's

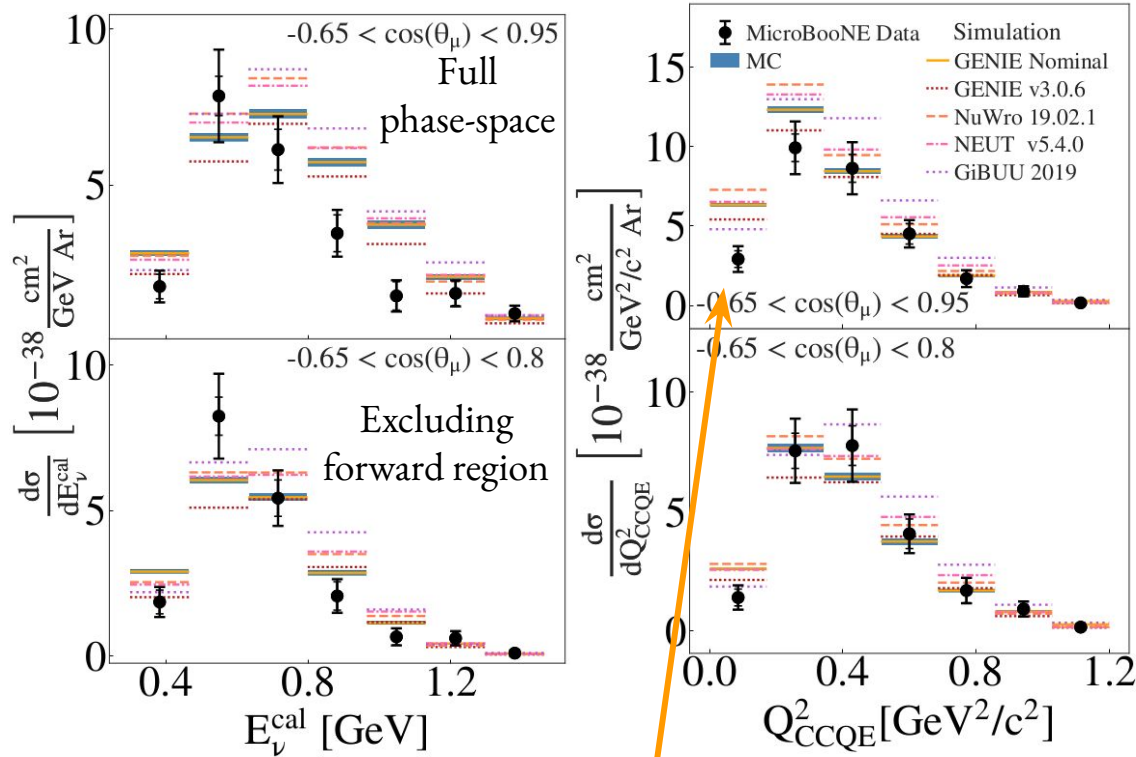
Not accounting for correlations between distributions

		Differential Cross Section $\chi^2$ /d.o.f	
		$-0.65 < \cos(\theta_\mu) < 0.95$	$-0.65 < \cos(\theta_\mu) < 0.8$
Generators	GENIE Nominal	63.2/28	30.1/27
	GENIE v3.0.6	34.6/28	21.4/27
	NuWro 19.02.1	76.7/28	29.9/27
	NEUT v5.4.0	78.5/28	32.2/27
	GiBUU 2019	82.2./28	40.0/27

GENIE v3.0.6 results in the lowest  $\chi^2$  /d.o.f.

Same conclusion as in inclusive analysis [Phys. Rev. Lett. 123, 131801 \(2019\)](#)

# Derived Quantities



$$E_\nu^{\text{cal}} = E_\mu + T_p + 40 \text{ MeV}$$

$$\vec{p}_\nu = (0, 0, E_\nu^{\text{cal}})$$

$$Q_{\text{CCQE}}^2 = (E_\nu^{\text{cal}} - E_\mu)^2 - (\vec{p}_\nu - \vec{p}_\mu)^2$$

Expected low  $Q^2$  deficit given  $\cos\theta_\mu$  result

# Model Comparisons

Model Component



GENIE Nominal  
v2.12.2  
This work

GENIE v3.0.6  
G18\_10a\_02\_11a  
Future analyses

NuWro  
(19.02.1)

NEUT  
(v5.4.0)

GiBUU  
(2019)

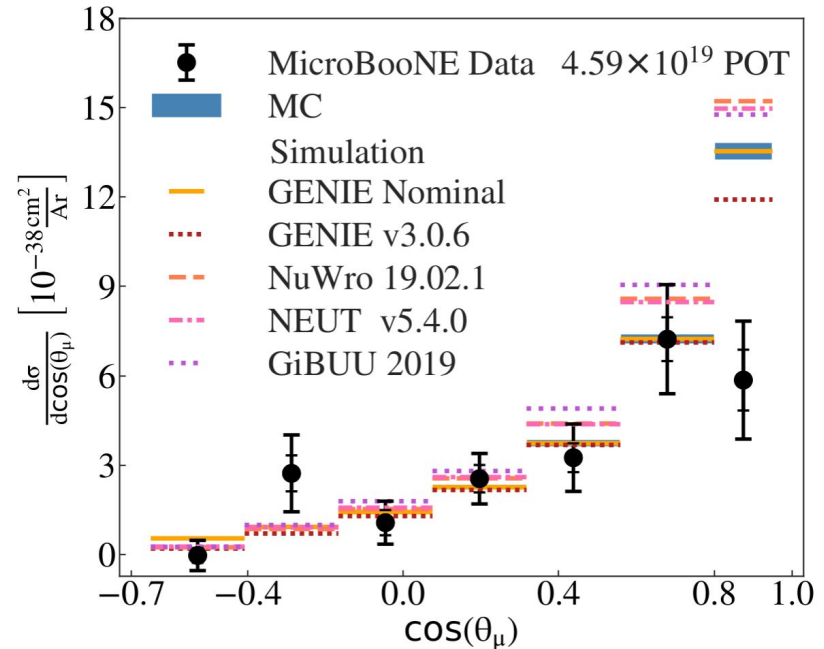
Model Component	GENIE Nominal v2.12.2 This work	GENIE v3.0.6 G18_10a_02_11a Future analyses	NuWro (19.02.1)	NEUT (v5.4.0)	GiBUU (2019)
<b>Nuclear Model</b>	Bodek-Ritchie Fermi Gas [1]	Local Fermi Gas [2, 3]	Local Fermi Gas [2, 3]	Local Fermi Gas [2, 3]	Consistent nuclear medium corrections
<b>Quasi-elastic</b>	Llewellyn-Smith [4]	Nieves [2, 3]	Nieves [2, 3]	Nieves [2, 3]	LFG model for nucleon momenta
<b>MEC</b>	Empirical [5]	Nieves [2, 3]	Nieves [2, 3]	Nieves [2, 3]	Separate MEC model [11]
<b>Resonant</b>	Rein-Seghal [6]	Berger-Seghal [7]	Berger-Seghal [7]	Berger-Seghal [7]	
<b>Coherent</b>	Rein-Seghal [6]	Berger-Seghal [7]	Berger-Seghal [7]	Rein-Seghal [6]	Propagates final state particles according to the BUU equations [11]
<b>FSI</b>	hA [8]	hA2018 [8]	Oset [10]	Oset [10]	

Differences in coulomb corrections, parameter tuning and implementation of RPA correction

**Our data indicates that these seemingly small differences can have a highly significant impact**

# MicroBooNE Wrap Up

- First measurement of  $\nu_{\mu} - {}^{40}\text{Ar}$  CCQE-like cross sections
- Powerful cosmic background rejection and high topological signal purity
- Improvement of theoretical models is required in forward region



[Phys. Rev. Lett. 125, 201803 \(2020\)](#)

# Far Detector

$$N_{\text{FD}}^{\beta}(E_{\text{rec}}) \sim \Phi_{\text{FD}}(E_{\nu}) \sigma(E_{\nu}) \varepsilon_{\text{FD}}(E_{\nu}) \mathbf{P}(\nu_{\alpha} \rightarrow \nu_{\beta})$$

Oscillation probability

Modeling Input

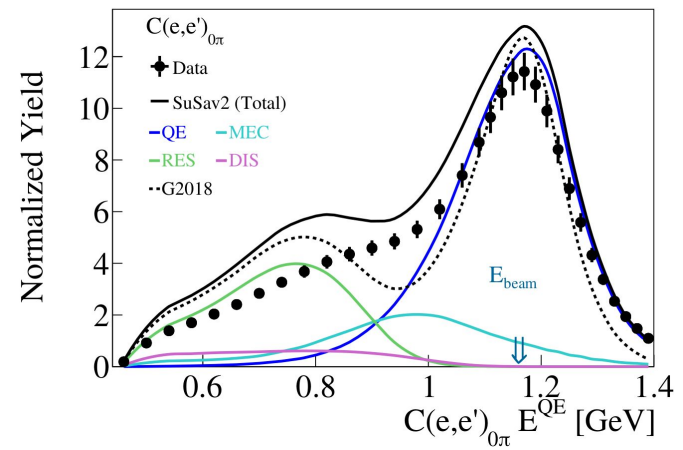
# How To Improve Modeling?

$\nu$  near detector constraints

Improved theory  
in event generators



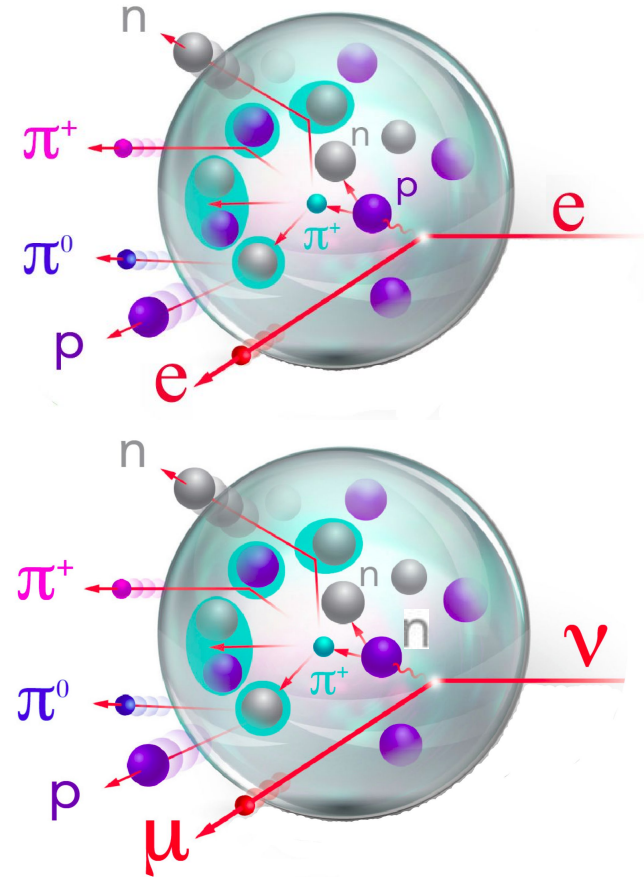
## External Data





# Why electrons?

- Very similar interactions
- Nuclear effects practically identical
- Known electron beam energy
- Benchmark  $\nu$  event generators



# Exclusive $1p0\pi$ Analysis With CLAS

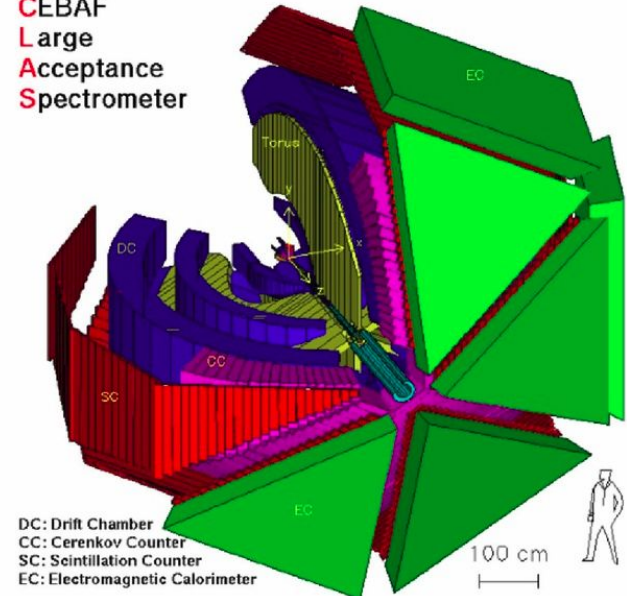
- **Data Mining**
- Complementary to the one performed on



- Large acceptance with  $\theta_e > 15^\circ$ , investigation of forward deficit
- Charged particle detection threshold comparable to neutrino experiments

**Jefferson Lab**  
Thomas Jefferson National Accelerator Facility

CEBAF  
Large  
Acceptance  
Spectrometer



DC: Drift Chamber  
CC: Cerenkov Counter  
SC: Scintillation Counter  
EC: Electromagnetic Calorimeter



# Playing The Neutrino Game



- 1 muon ( $> 100 \text{ MeV}/c$ )
- 1 proton ( $> 300 \text{ MeV}/c$ )
- No  $\pi^\pm$  ( $> 70 \text{ MeV}/c$ )



- Select "clean" (e,e'p) events  
1 proton ( $> 300 \text{ MeV}/c$ )  
No  $\pi^\pm$  ( $> 150 \text{ MeV}/c$ )
- Scale electron data by  $Q^4$

- Study energy reconstruction
- Benchmark event generators

# CLAS Available Data Sets

- Targets

$^4\text{He}$ ,  $^{12}\text{C}$ ,  $^{56}\text{Fe}$



$\text{H}_2\text{O}$



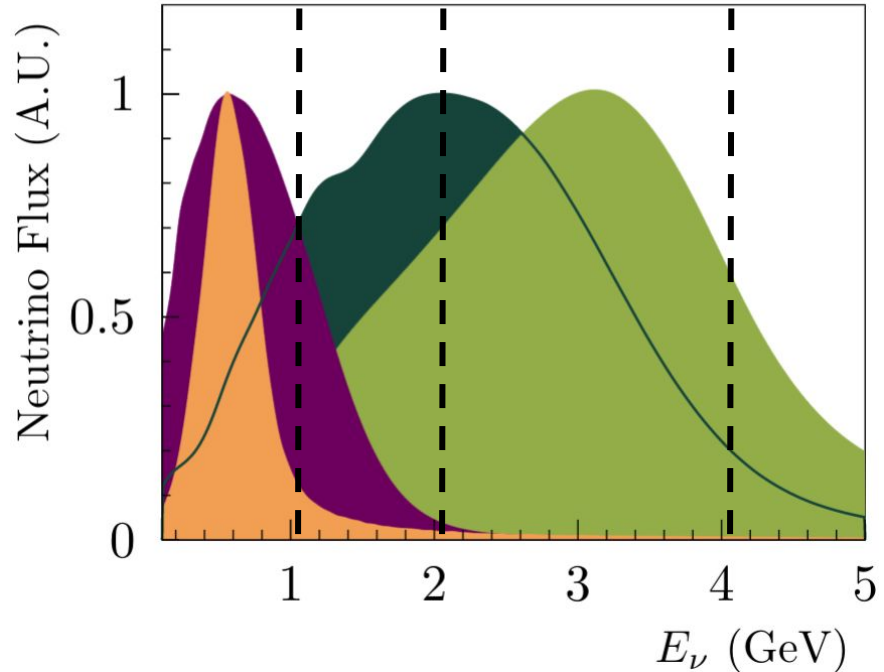
CH



Ar

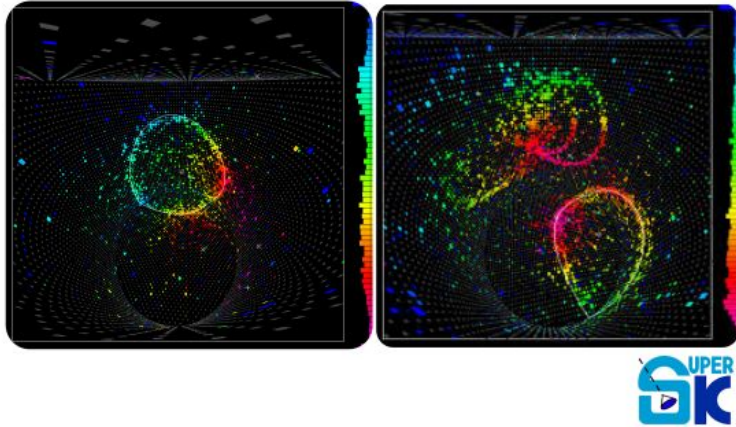
- Energies

1, 2 & 4 GeV



Credit: L. Pickering

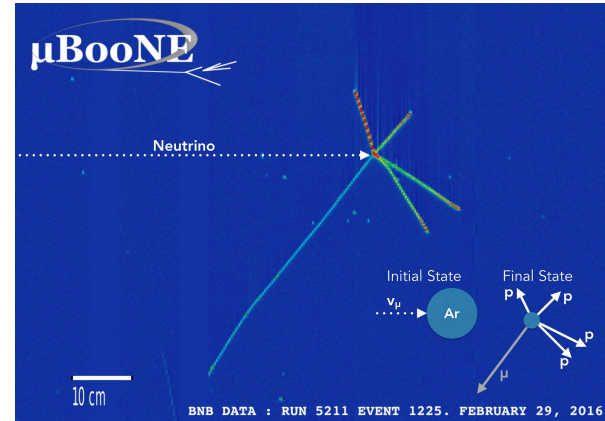
# Energy Reconstruction



Cherenkov detectors  
Assuming QE interaction  
Using lepton kinematics

$$E_{QE} = \frac{2M\epsilon + 2ME_l - m_l^2}{2(M - E_l + |k_l|\cos\theta_l)}$$

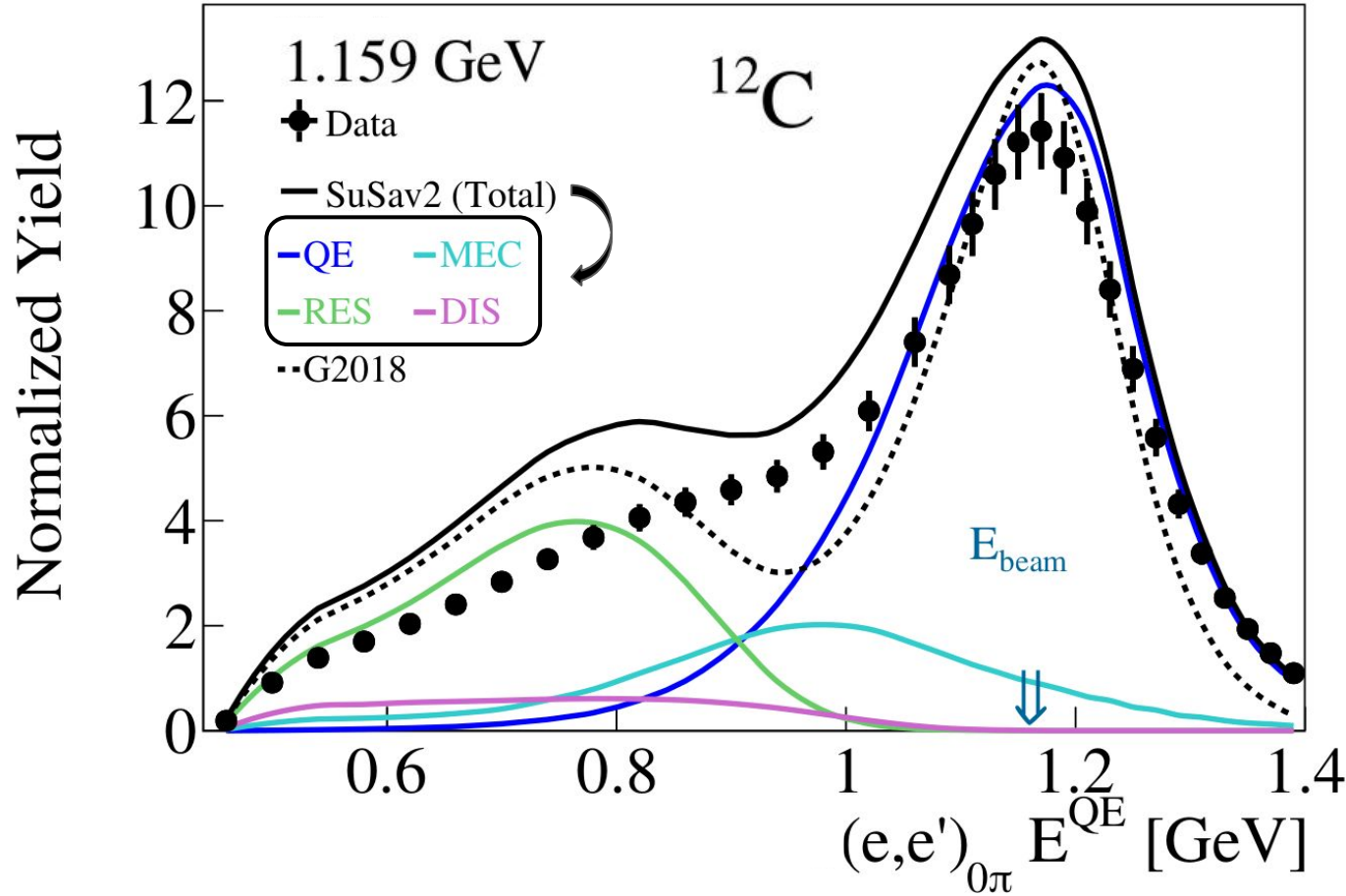
nucleon separation energy  $\epsilon \sim 20$  MeV



Tracking detectors  
Calorimetric sum  
Using all detected particles

$$E_{cal} = E_l + T_p + \epsilon$$

# Inclusive $0\pi$ $E^{QE}$ Reconstruction



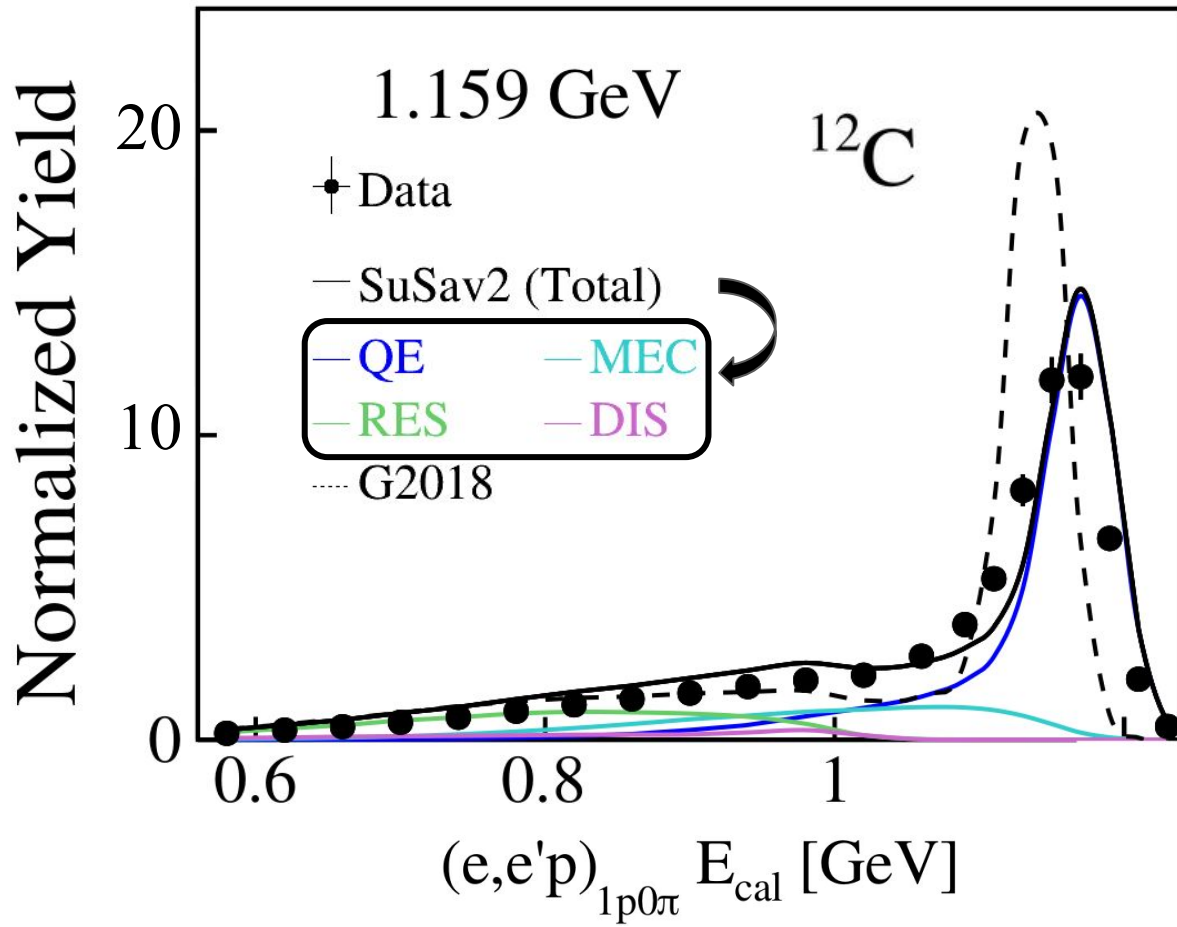
Simulation fails  
to capture details



G2018 = GENIE v3 G18\_10a\_02\_11a

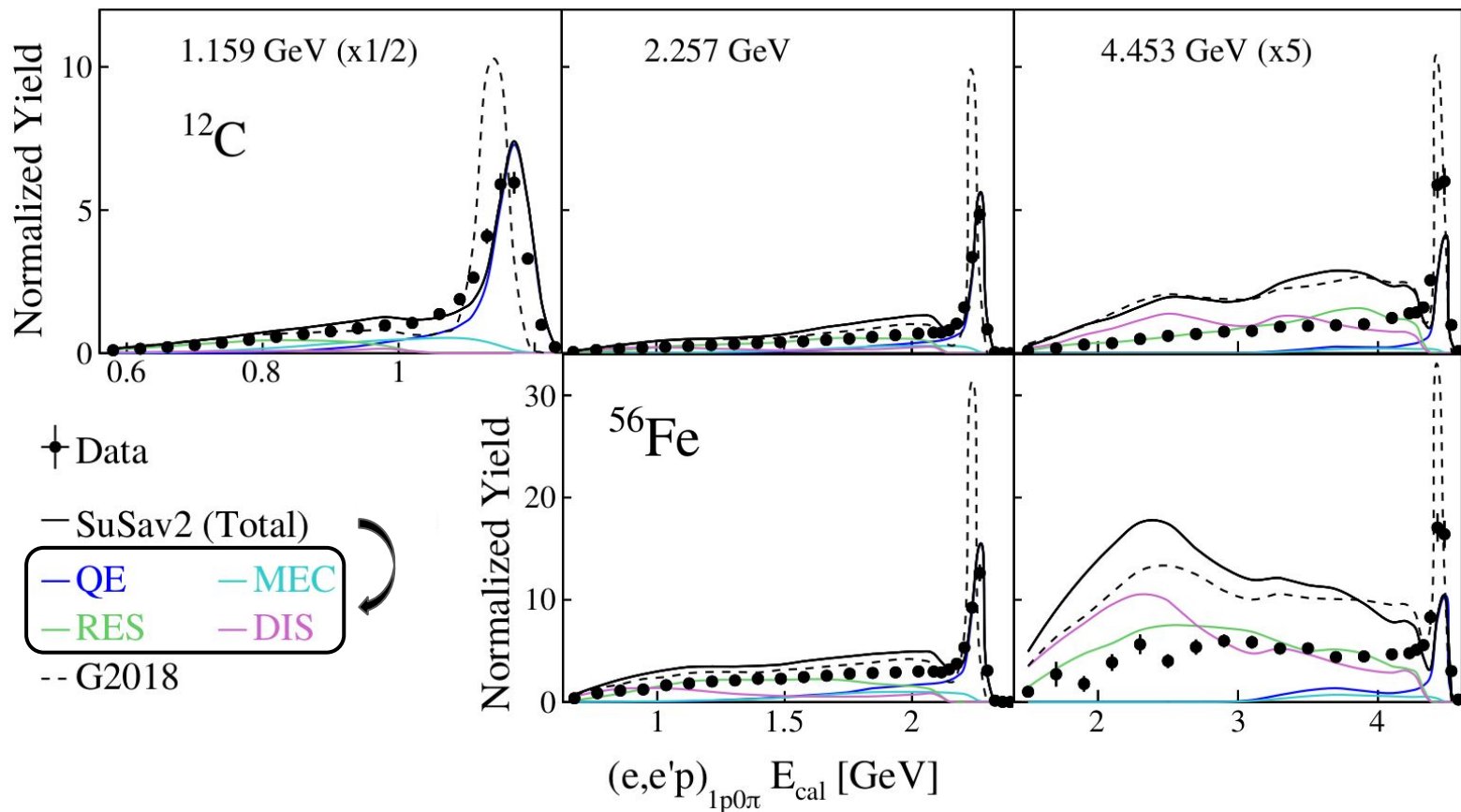
SuSav2 = Superscaling model  
[Phys. Rev. D 94, 013012 \(2016\)](#) 62

# Exclusive $1p0\pi$ $E_{cal}$ Reconstruction



SuSav2 captures  
QE peak

# Beam Energy & Nucleus Dependence

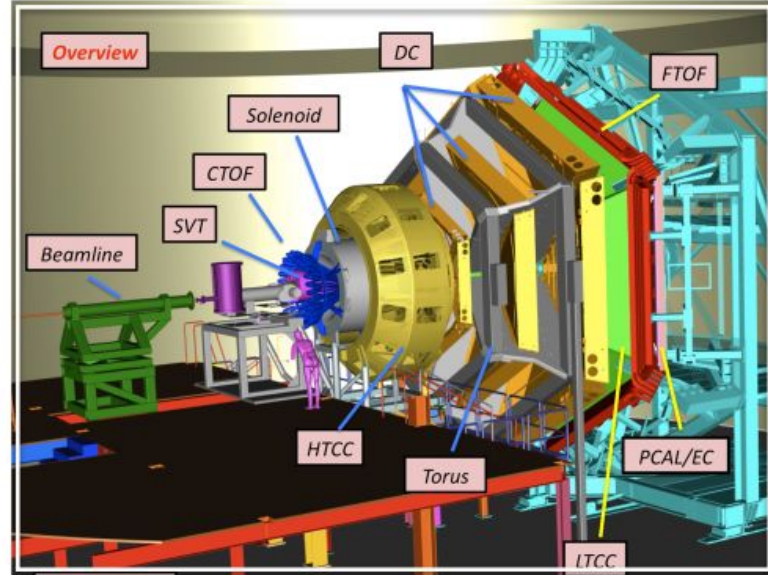


Simulation fails at high energies & on heavier nuclei



# Future Plans @ CLAS12

- Acceptance down to  $5^\circ$
- x10 luminosity [ $10^{35} \text{ cm}^{-2} \text{ s}^{-1}$ ]
- Targets  
 $^2\text{D}$ ,  $^4\text{He}$ ,  $^{12}\text{C}$ ,  $^{16}\text{O}$ ,  $^{40}\text{Ar}$ ,  $^{120}\text{Sn}$
- 1 - 7 GeV beam energies
- Data taking in 2021



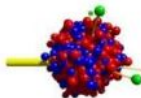
ICECUBE  
SOUTH POLE NEUTRINO OBSERVATORY



Hyper-Kamiokande



MINERvA



GiBUU

The Giessen Boltzmann-Uehling-Uhlenbeck Project

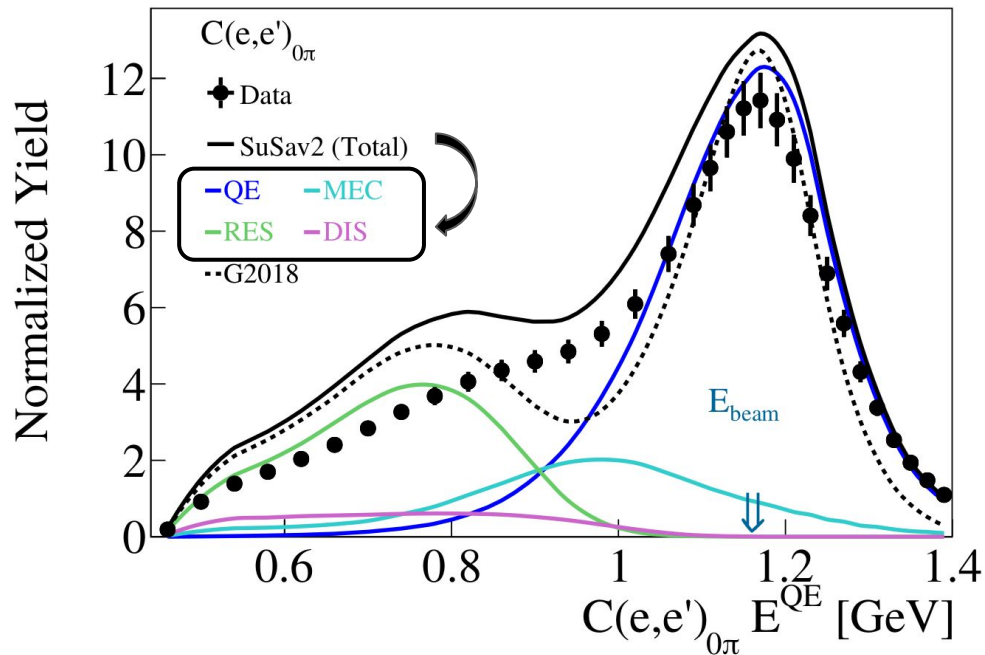


Genie



# $e4\nu$ Wrap Up

- Benchmarking  $\nu$  models against wide phase-space electron data sets
- Data/MC disagreements even for QE-like topologies
- Need for more electron scattering data sets to constrain neutrino models

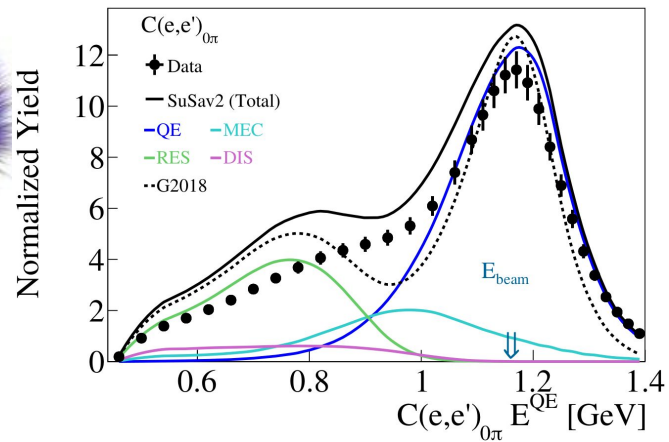
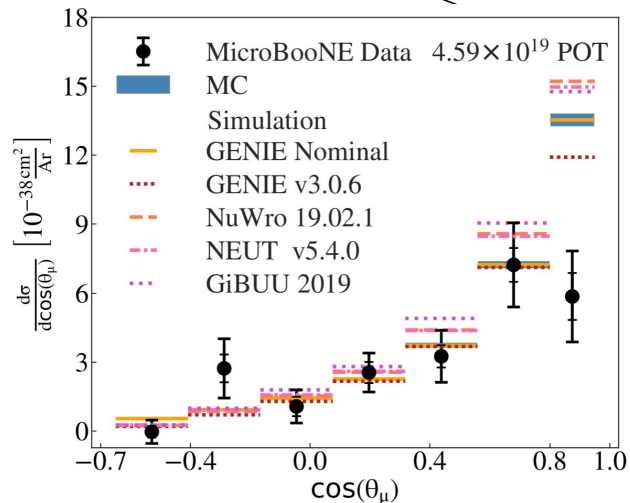


**Potential impact on DUNE**

# Summary

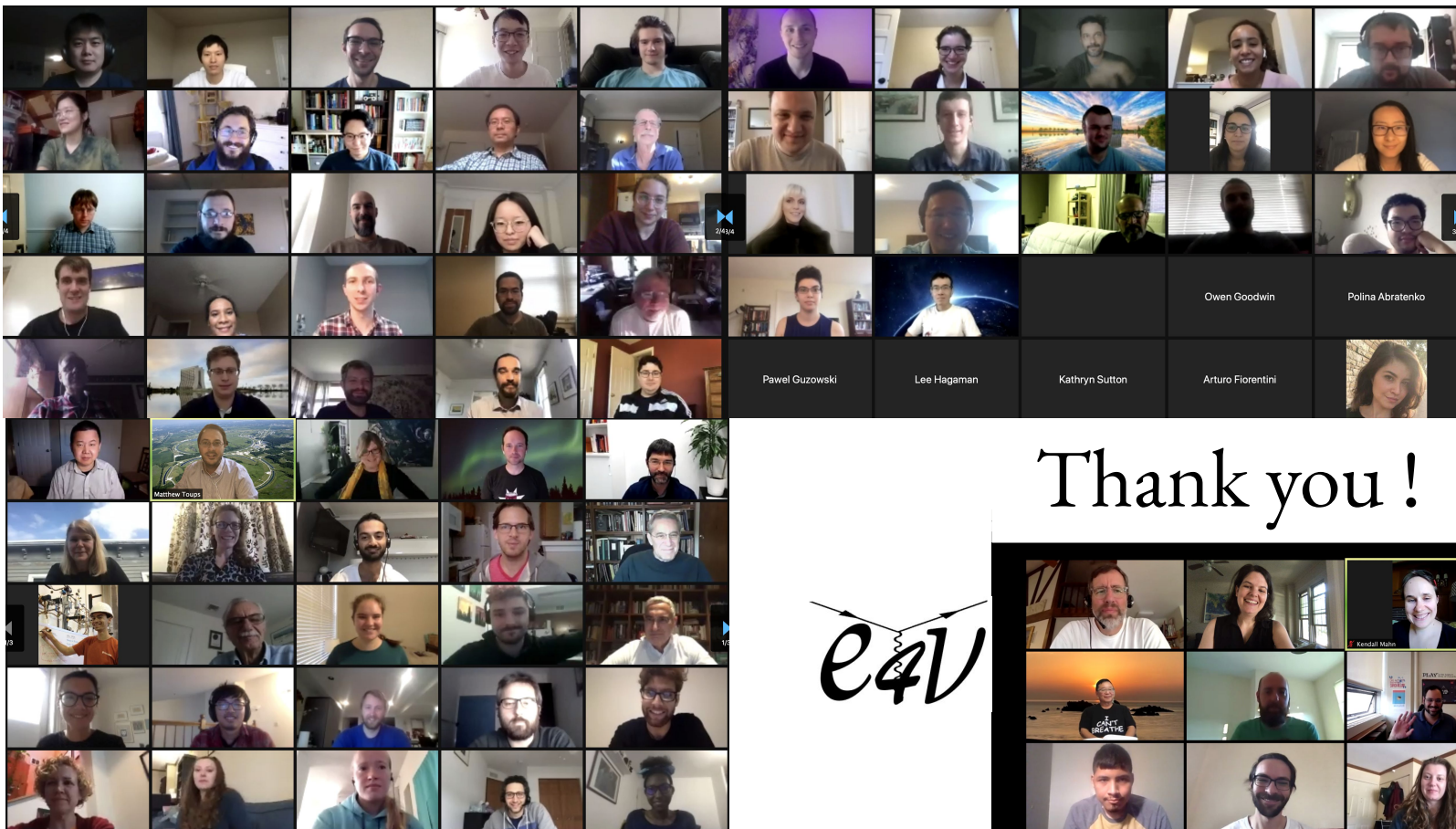


## Modeling Input

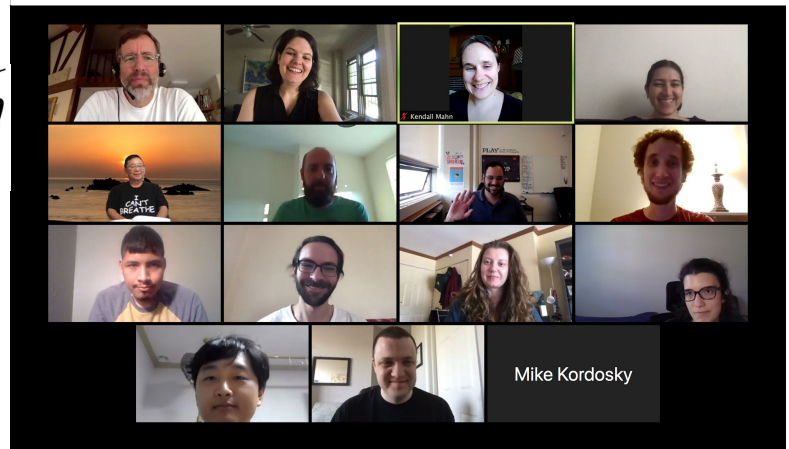


[Phys. Rev. Lett. 125, 201803 \(2020\)](#)

- First exclusive measurement of Quasielastic-like interactions using the MicroBooNE detector
- First head-to-head comparison between neutrino & electron scattering



# Thank you !





Questions ?





## Backup Slides

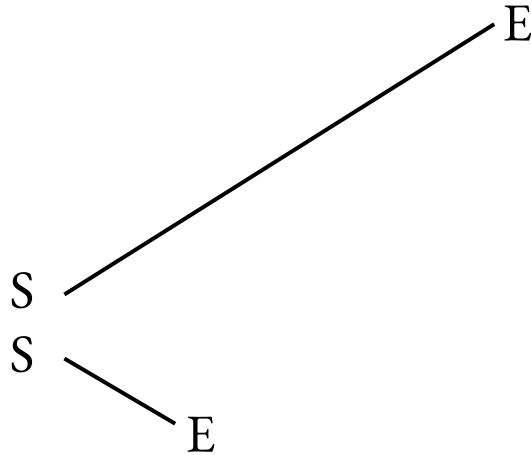


# References

- [1] Phys. Rev. D.24 (1981), p. 1400
- [2] Phys. Rev. C.83 (2011), p. 045501
- [3] Phys. Rev. D.88 (2018), p. 113007
- [4] Phys. Rept. 3 (1972), p. 261
- [5] AIP Conf. Proc. 1663 (2015), p. 030001
- [6] Ann. Phys. 133 (1981), p. 79
- [7] Phys. Rev. D.79 (2009), p. 053003
- [8] arXiv:1510.05494
- [9] Phys. Rev. D 80, 093001 (2009)
- [10] Nuclear Physics A 484, 557 (1988)
- [11] Physics Reports 512, 1 (2012)

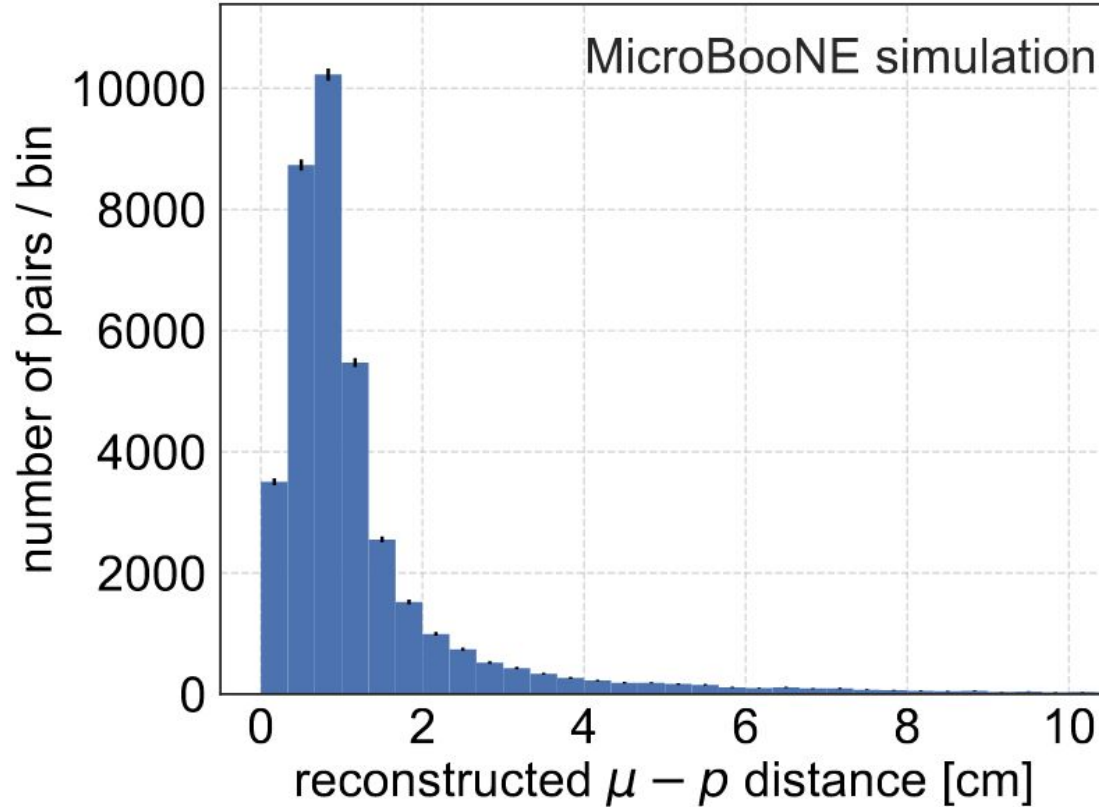
# Preselection

Pairs of tracks at close proximity with a distance  $< 11$  cm between any two edges  
(start-start, start-end, end-start, end-end)



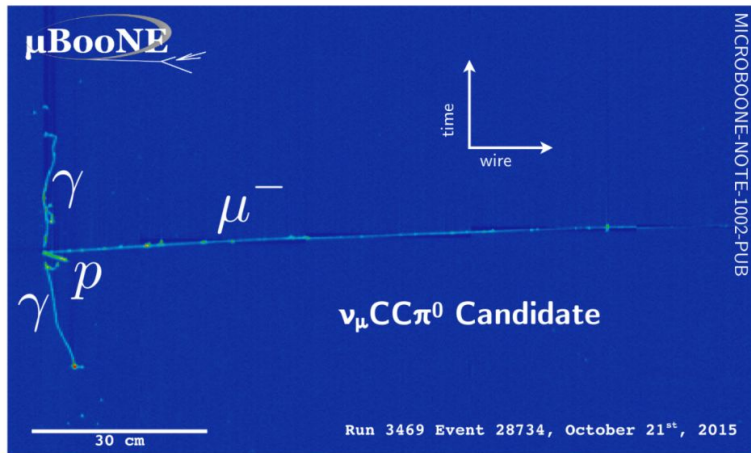


# Preselection

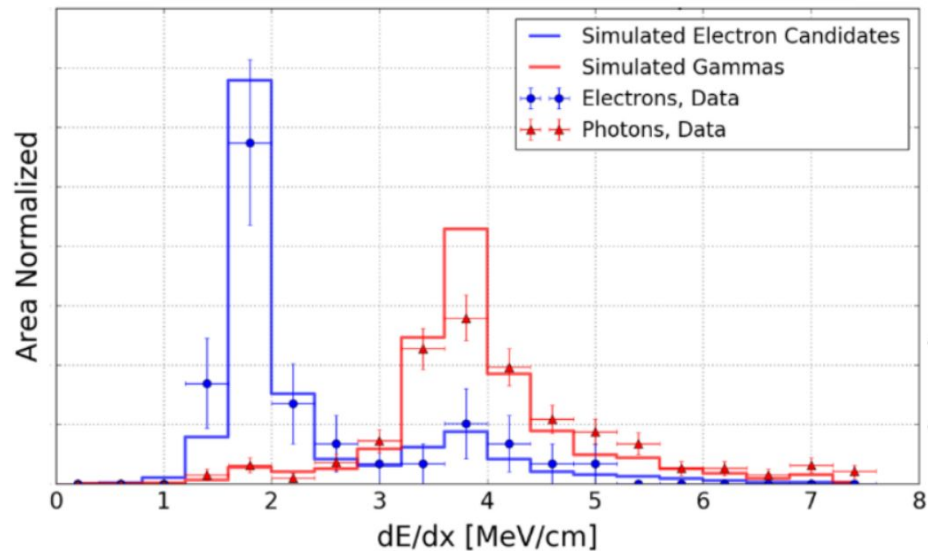


# Particle Identification In LArTPCs

- Detailed 3D imaging of  $\nu$  interactions
- Electron / photon discrimination



## ArgoNEUT



PRD 95, 072005 (2017)

# CC1p0 $\pi$ Topological Signal Definition

Vertex of two tracks

- 1 muon ( $> 100$  MeV/c)

Either semi-contained or fully contained

[Multiple Coulomb Scattering]

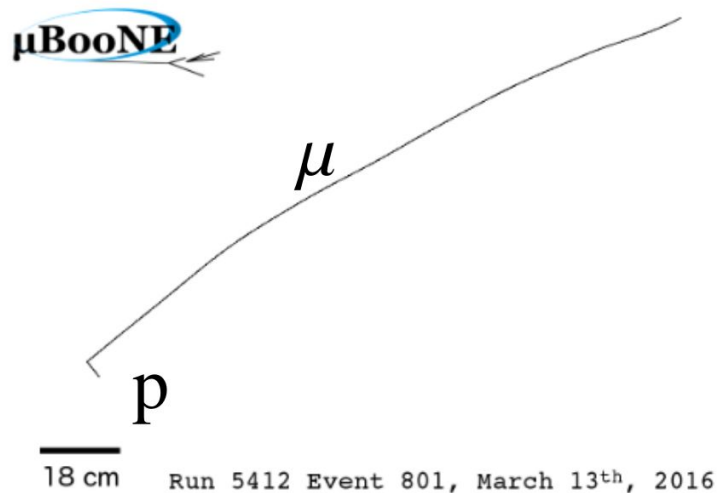
[JINST 12 P10010 \(2017\)](#)

- 1 proton ( $> 300$  MeV/c)

Fully contained

[Momentum from range\*]

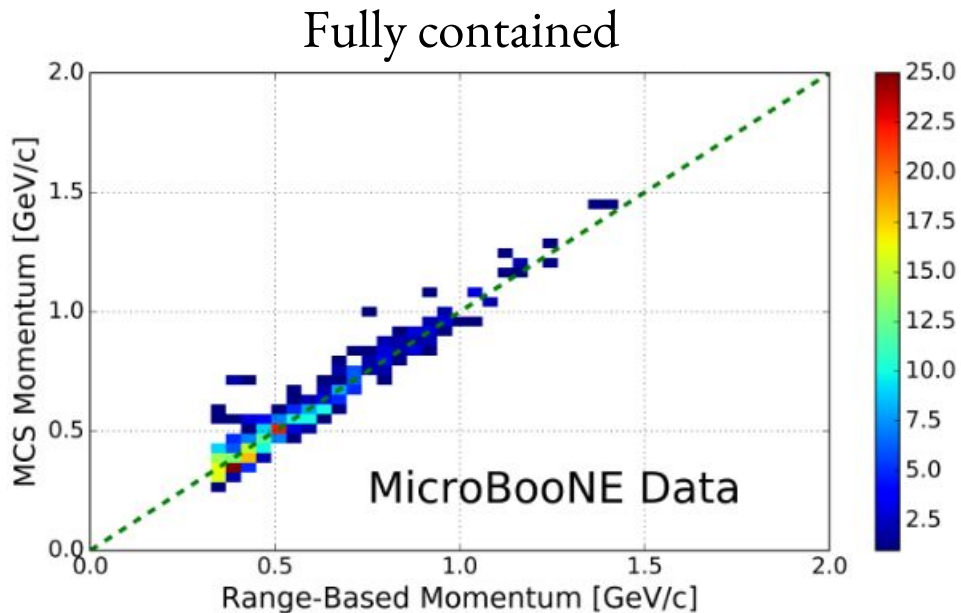
- No  $\pi^\pm$  ( $> 70$  MeV/c)



\* Table 289: Muons in Liquid argon (Ar)

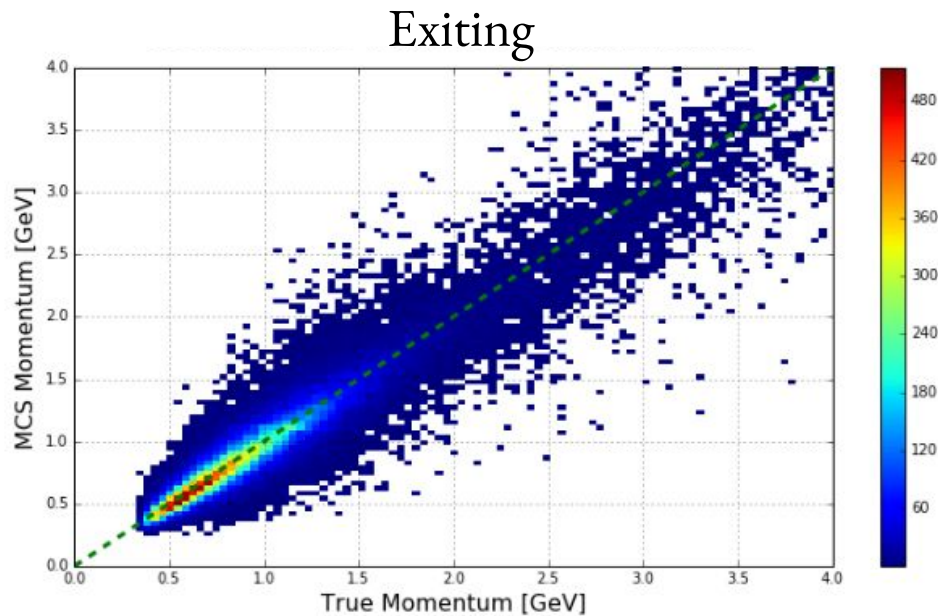
[http://pdg.lbl.gov/2012/AtomicNuclearProperties/MUON\\_ELOSS\\_TABLES/muonloss\\_289.pdf](http://pdg.lbl.gov/2012/AtomicNuclearProperties/MUON_ELOSS_TABLES/muonloss_289.pdf)

# Multiple Coulomb Scattering (MCS)



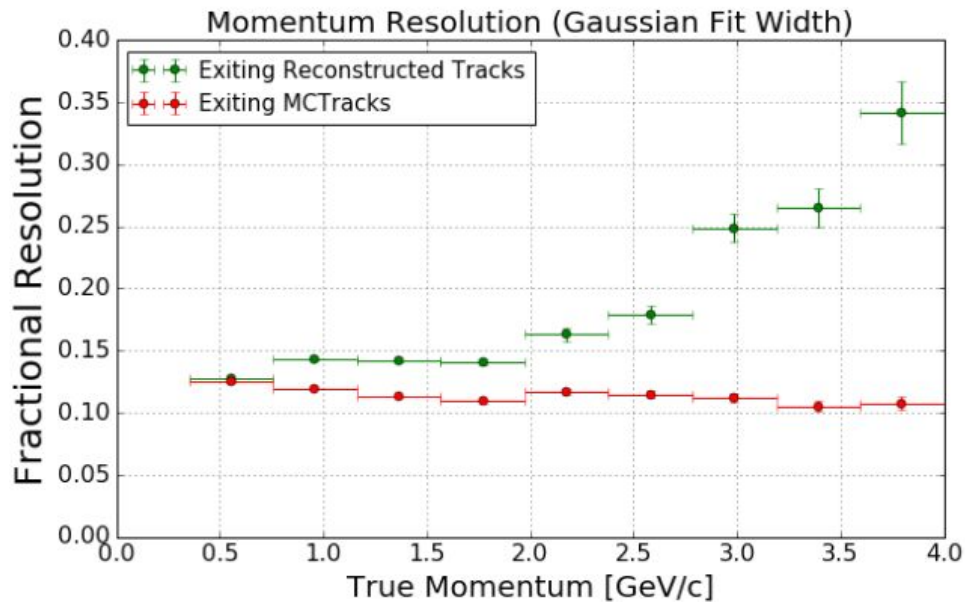
**Figure 7.** MCS-computed momentum versus range momentum for the automatically selected beam neutrino-induced fully contained muon sample in MicroBooNE data after hand scanning to remove poorly reconstructed tracks and obvious mis-identification topologies. The color (z) scale indicates number of tracks.

# Multiple Coulomb Scattering (MCS)



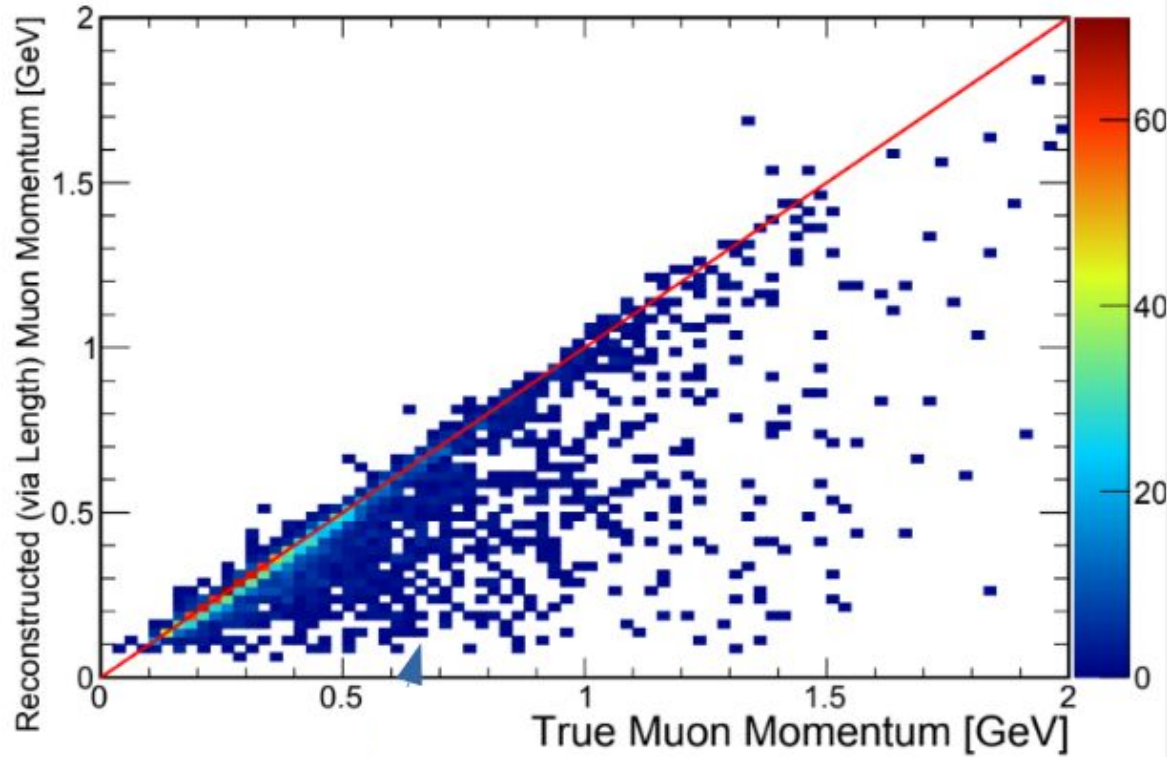
**Figure 10.** MCS-computed momentum versus true momentum for the MCTRACKS (truth-based) sample of simulated exiting muons from BNB  $\nu_{\mu}$ CC interactions in MicroBooNE with at least one meter of track contained within the TPC. The color (z) scale indicates number of tracks.

# Multiple Coulomb Scattering (MCS)



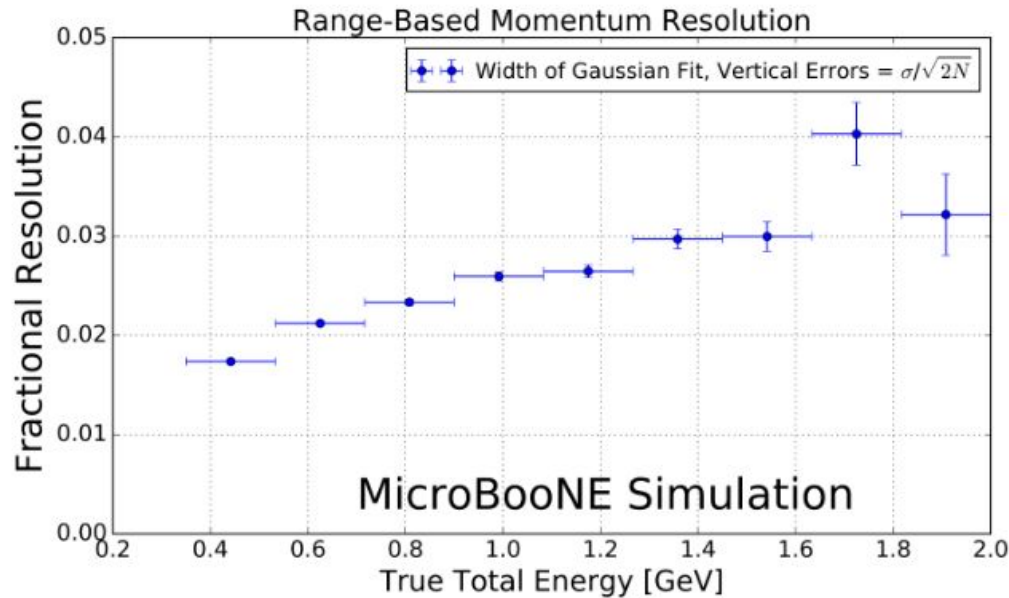
**Figure 12.** MCS momentum fractional bias (top) and resolution (bottom) from Gaussian fits to the reconstructed momentum as a function of true momentum for the exiting muon using the MCTRACKS (truth-based) sample and using the fully-simulated sample.

# Momentum From Range



[JINST 12 P10010 \(2017\)](#)

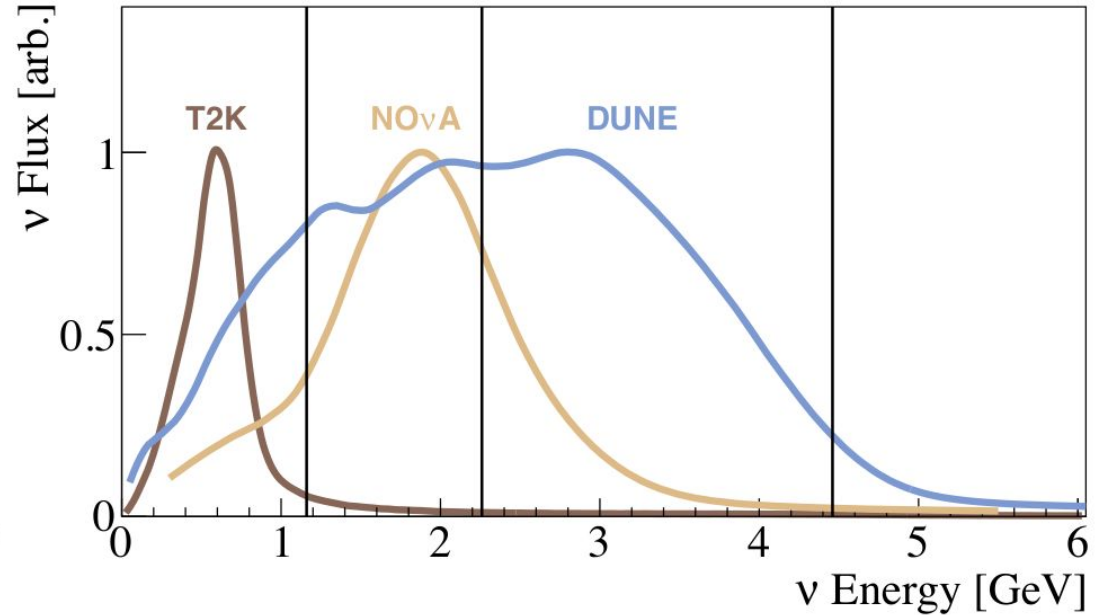
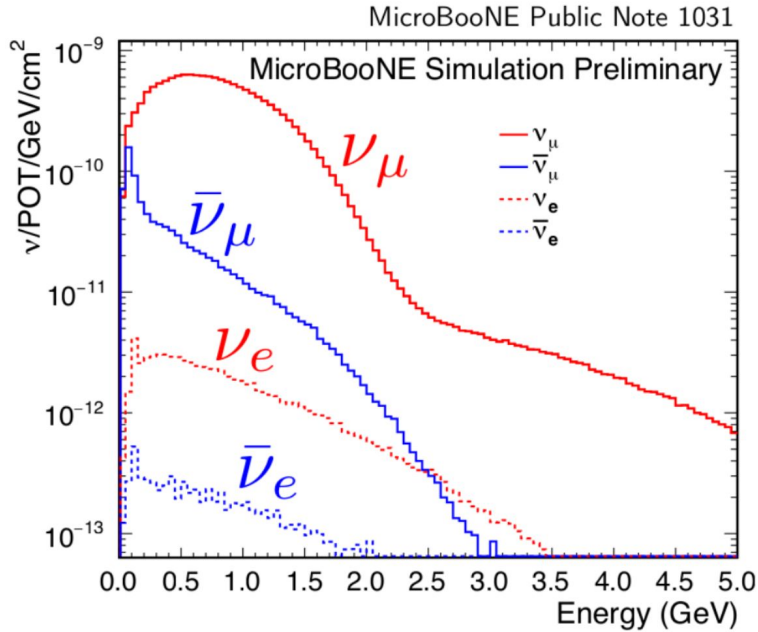
# Momentum From Range



**Figure 5.** Range-based energy fractional bias (top) and resolution (bottom) from a sample of simulated fully contained BNB neutrino-induced muons using true starting and stopping positions of the track. The bias is less than 1% and the resolution is better than 4%.

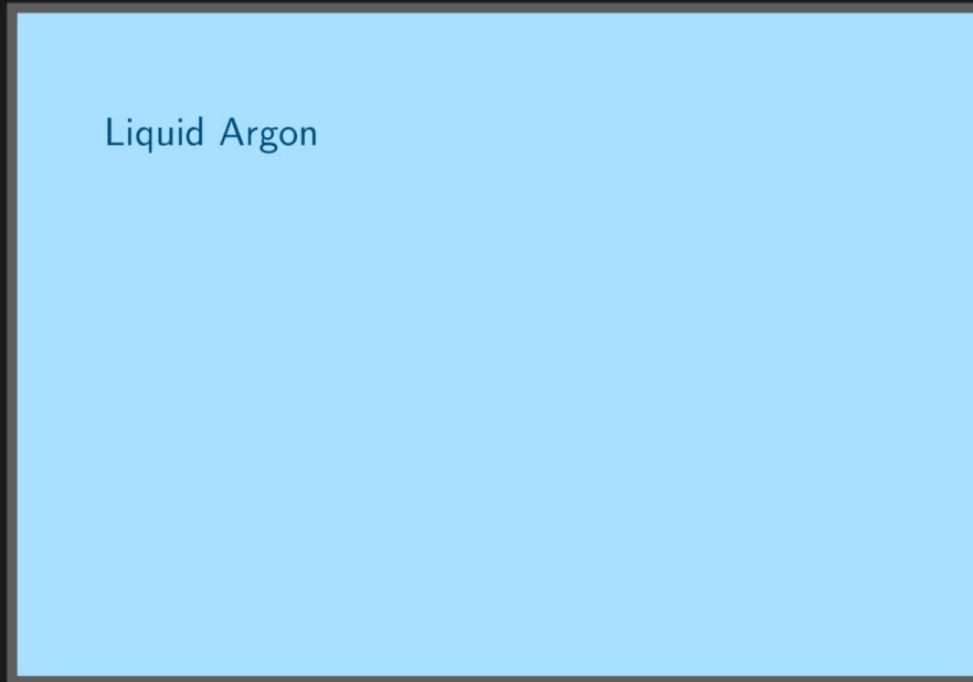


# Booster Neutrino Beamline

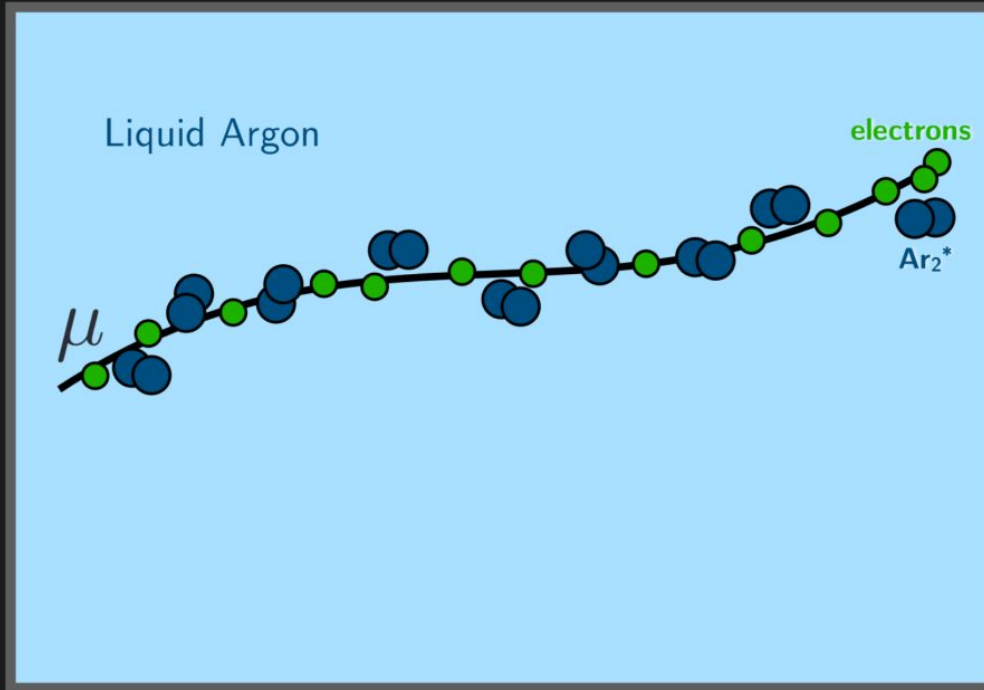


- Low energy & wide spectrum

# Liquid Argon TPCs



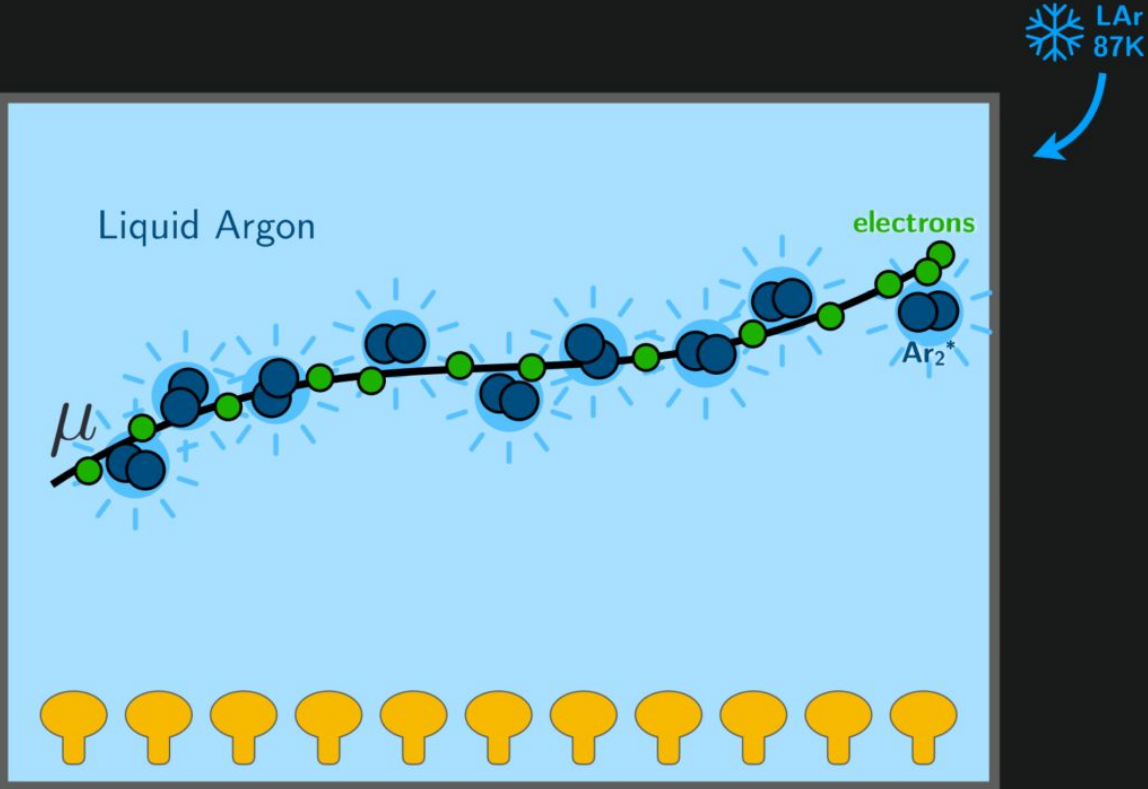
# Liquid Argon TPCs



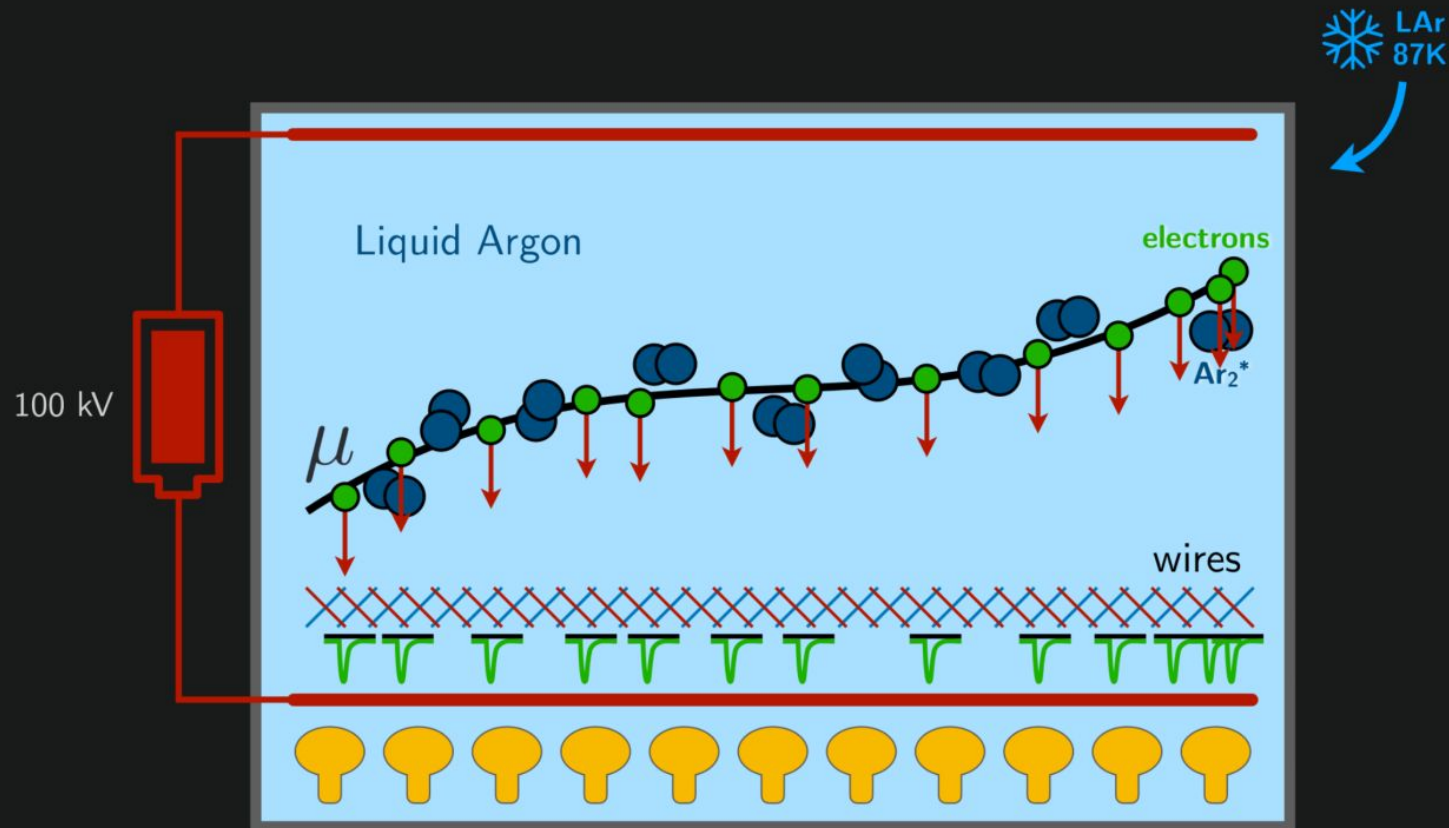
 LAr  
87K

A blue arrow points from the text and snowflake icon down towards the diagram, indicating the temperature of the liquid argon.

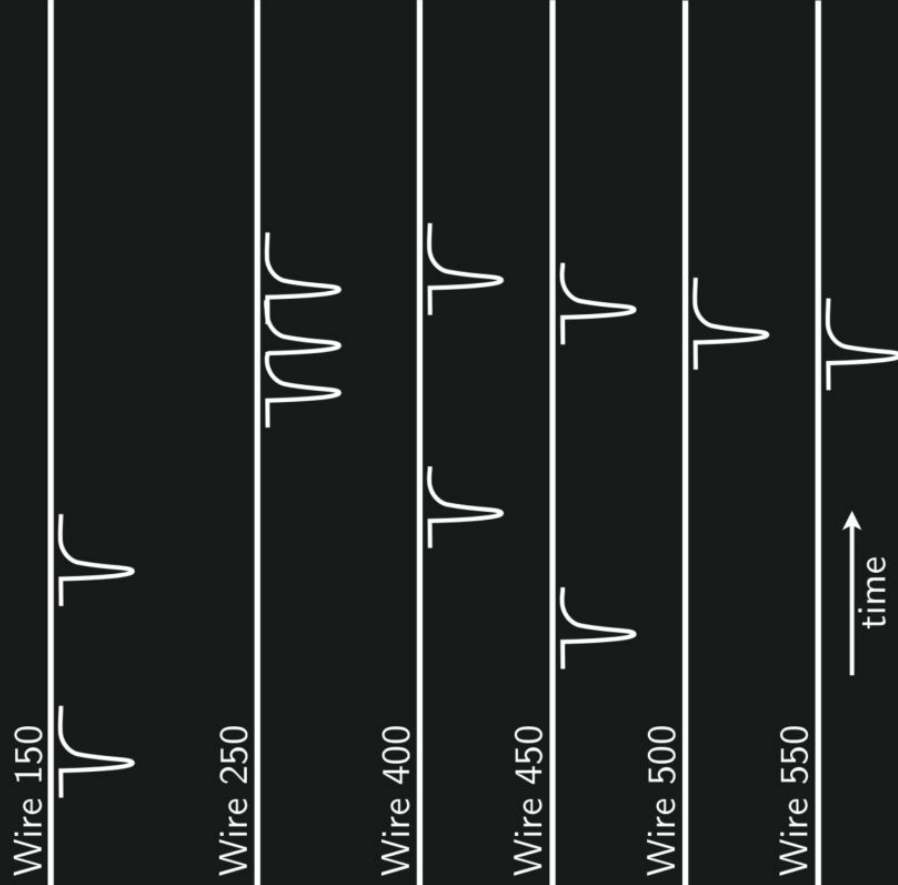
# Liquid Argon TPCs



# Liquid Argon TPCs



# Liquid Argon TPCs



$\mu$ BooNE

13 cm

Wire 150



Wire 250



Wire 400



Wire 450



Wire 500



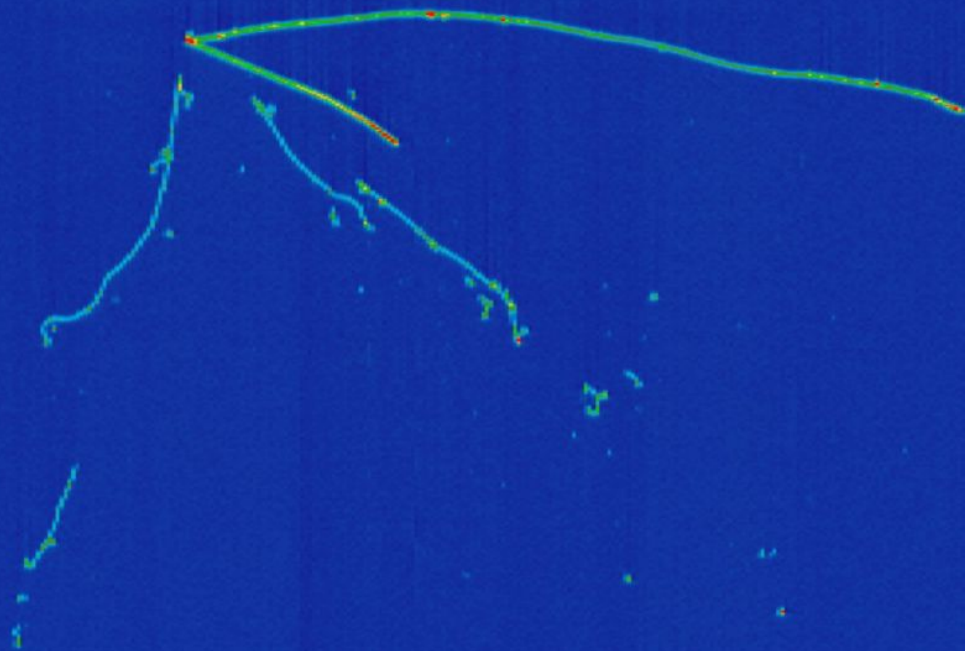
Wire 550



time

BNB DATA : RUN 5536 EVENT 1612. MARCH 22, 2016

$\mu$ BooNE

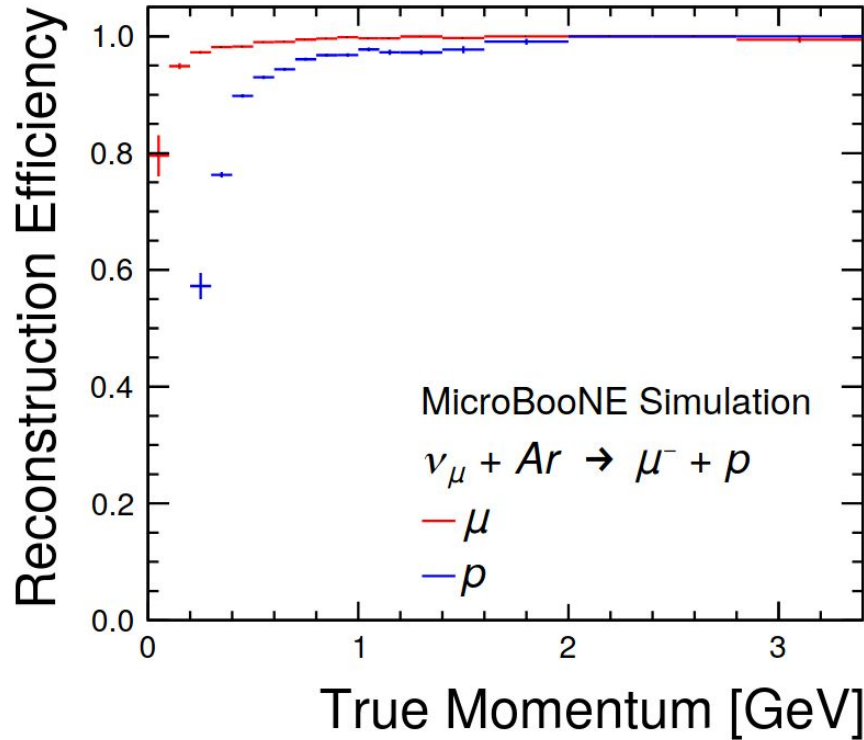


13 cm

BNB DATA : RUN 5536 EVENT 1612. MARCH 22, 2016

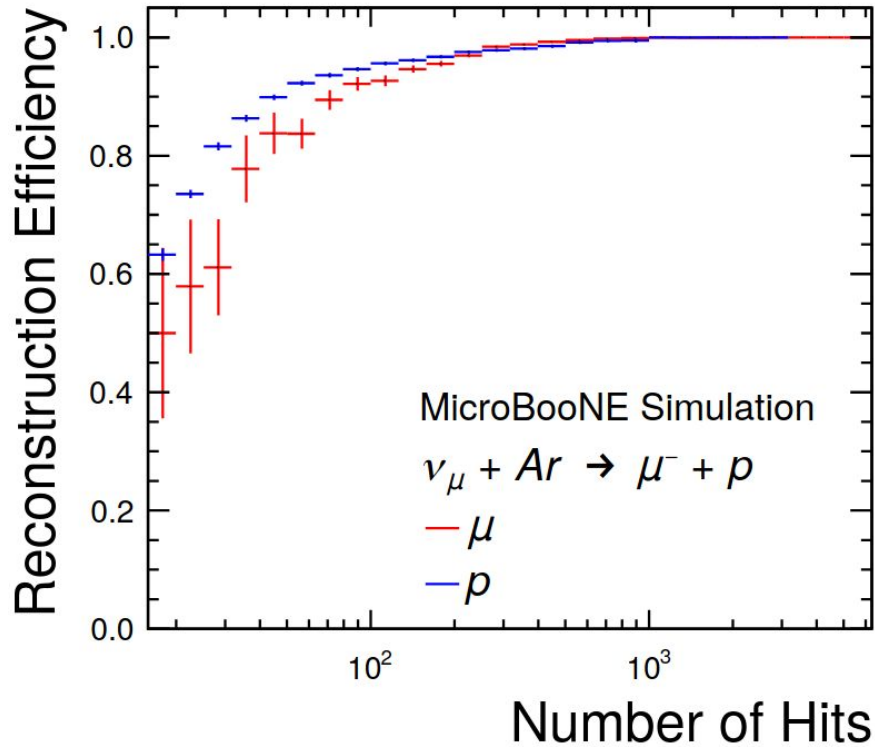


# Pandora Reconstruction



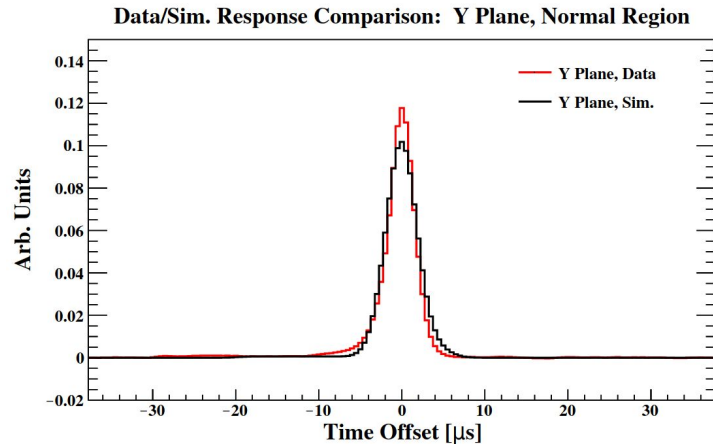
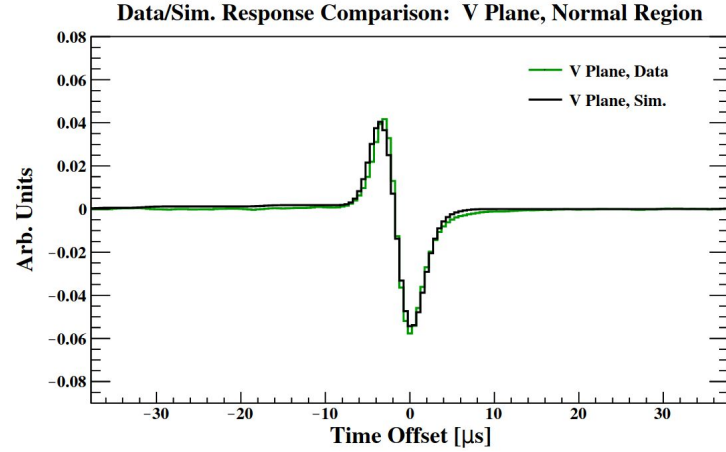
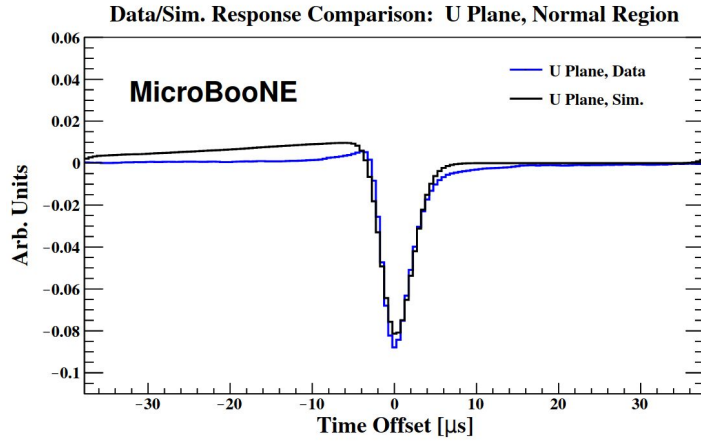
- High efficiency & accuracy
- Spatial resolution  $\sim 1$  cm

# Pandora Reconstruction



- High efficiency & accuracy
- Spatial resolution  $\sim 1$  cm

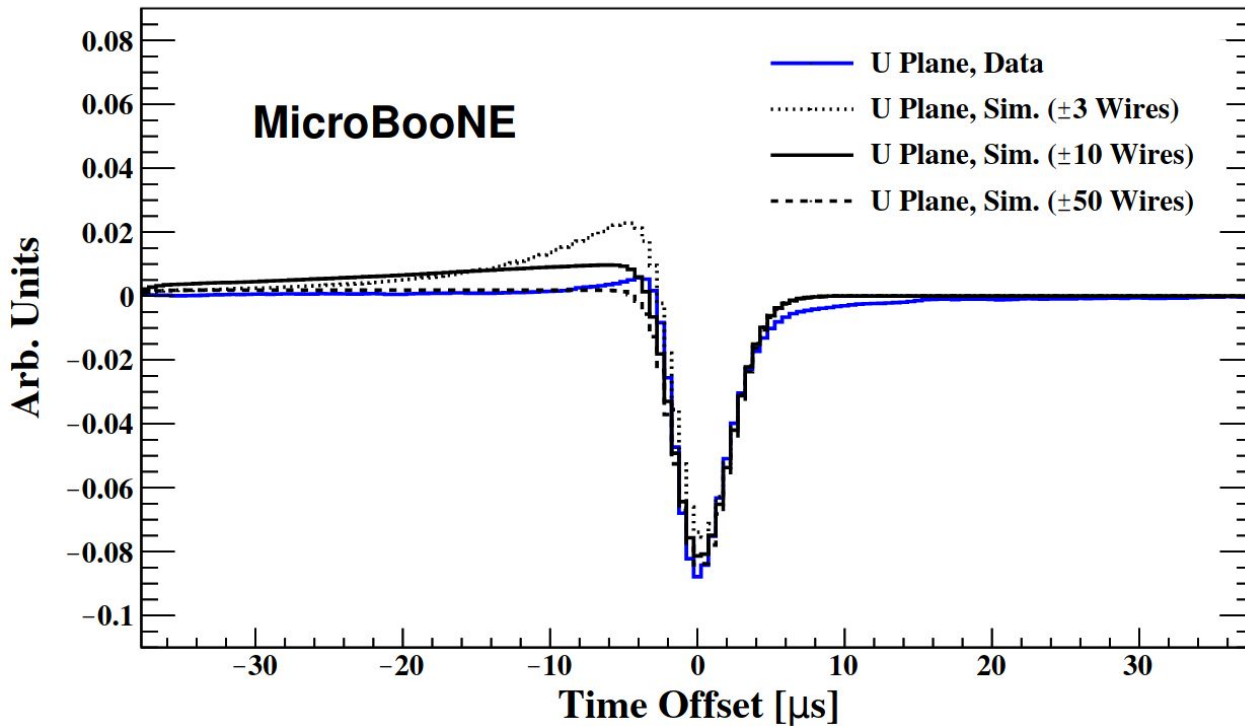
# Signal Processing



[JINST 13, P07007 \(2018\)](#)

# Signal Processing

Data/Sim. Response Comparison: U Plane, Normal Region

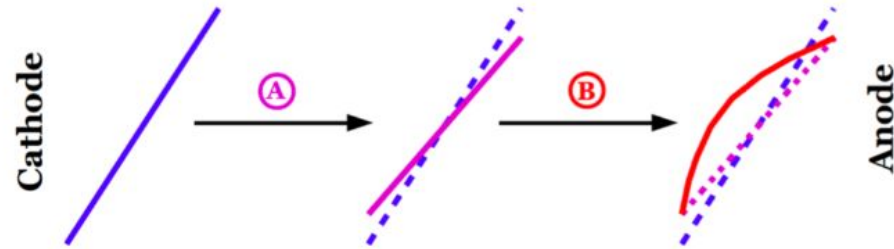


# Signal Processing

- Wires further away also see a weaker induction signal
- Different points in the charge distribution give signals at different times
- Correct simulation needs to sum **all** the contributions of all signals with time offsets
- Improved signal processing over past 3 years
- **For this analysis**, we are still using the “simple” nearest-wire treatment & the collection plane

# Space Charge Effects (SCE)

cumulative effect of the space charge in track reconstruction leads to squeezing of the reconstructed track in transverse directions and bending towards the cathode. See Figure 3.

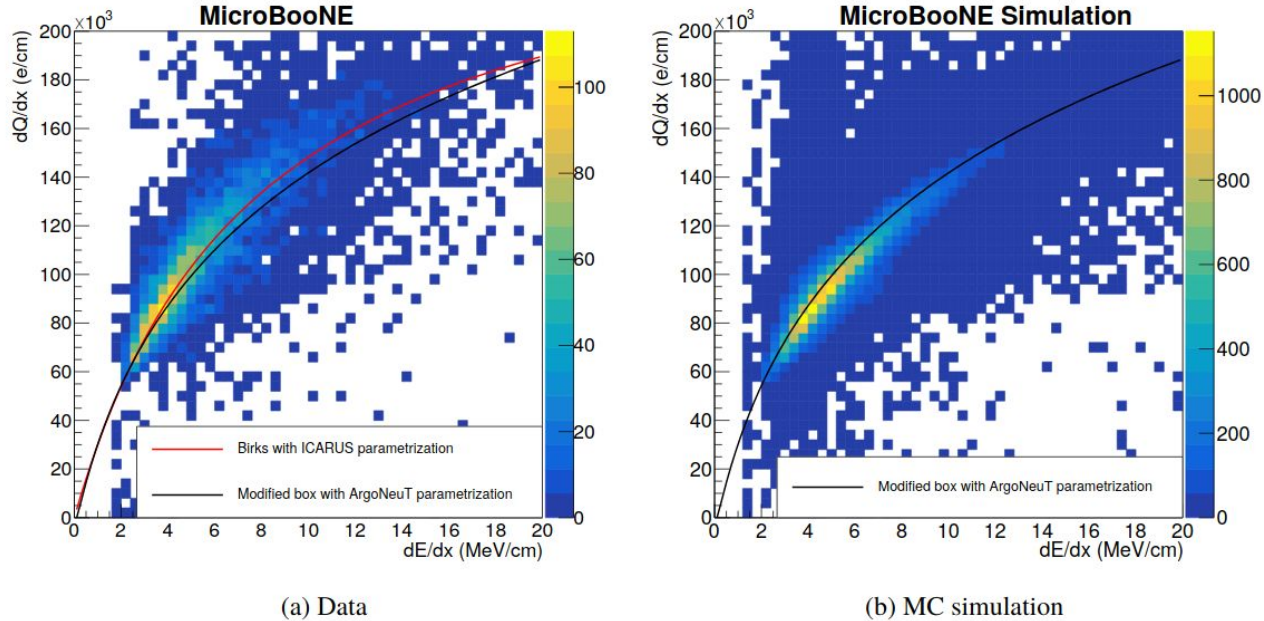


**Figure 3.** Diagram showing the effects of space charge on track reconstruction. The impact is two-fold: the reconstructed track could be squeezed by two extremes in the transverse directions of the TPC, as indicated in the rotation A, and bent towards cathode, as indicated by transformation B. (Image from Ref. [6] used by permission of its creator.)

# Recombination

- Ionisation electrons can “recombine” with argon ions
- The rate at which they do this depends on the local density of argon ions
- Non-trivial conversion from observed charge to deposited energy

# Recombination



**Figure 21.** (a) Comparison of the measured  $dQ/dx$  vs.  $(dE/dx)_{\text{range}}$  distribution with recombination models for selected proton tracks in data; (b) same comparison for selected protons in the MC simulation. The red curve corresponds to Birks' law and the black curve corresponds to the modified box model, using the parameters measured by the ICARUS and ArgoNeuT collaborations. Both model predictions are calculated at the nominal electric field of 0.273 kV/cm for MicroBooNE.



# Recombination

- Modified box model

$$\left(\frac{dE}{dx}\right)_{\text{calibrated}} = \frac{\exp\left(\frac{(\frac{dQ}{dx})_{\text{calibrated}} \beta' W_{\text{ion}}}{C_{\text{cal}} \rho \mathcal{E}}\right) - \alpha}{\frac{\beta'}{\rho \mathcal{E}}}, \quad (3.1)$$

with

$C_{\text{cal}}$  is a calibration constant used to convert ADC values to number of electrons,

$W_{\text{ion}} = 23.6 \times 10^{-6}$  MeV/electron (work function of argon),

$\mathcal{E} = 0.273$  kV/cm (MicroBooNE drift electric field),

$\rho = 1.38$  g/cm<sup>3</sup> (liquid argon density at a pressure 124.106 kPa),

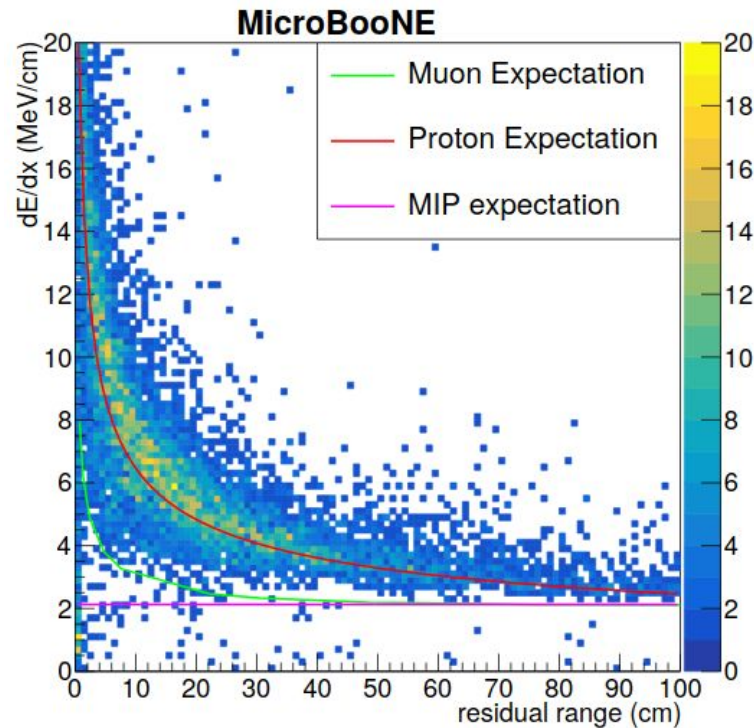
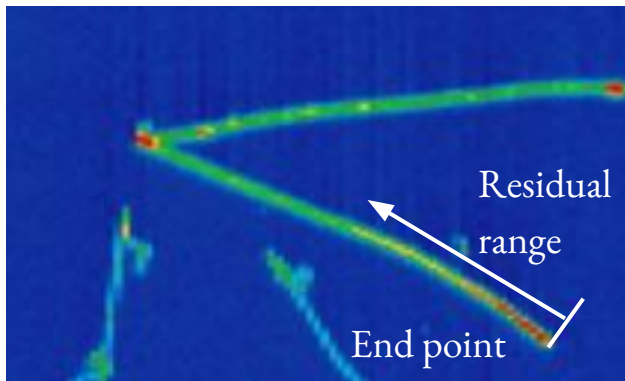
$\beta' = 0.212$  (kV/cm)(g/cm<sup>2</sup>)/MeV, and

$\alpha = 0.93$ .

The last two parameters were measured by the ArgoNeuT experiment [13] at an electric field of 0.481 kV/cm. The modified box model is applied at MicroBooNE's electric field of 0.273 kV/cm.

# PID In MicroBooNE

- For track-like particles, we rely on the Bragg peak
- Requires particles come to a stop in the detector
- PID all based off  $dE/dx$  vs Residual Range
- Pions and muons are functionally indistinguishable
- So PID is basically – is it a proton or not?



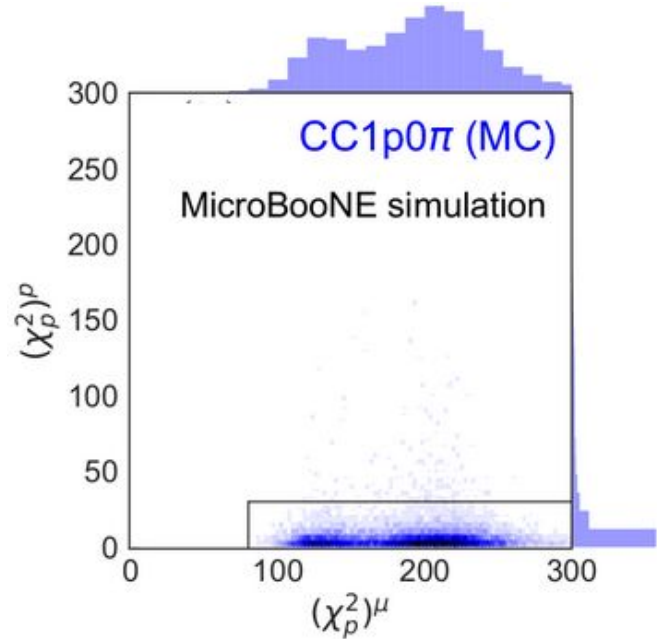
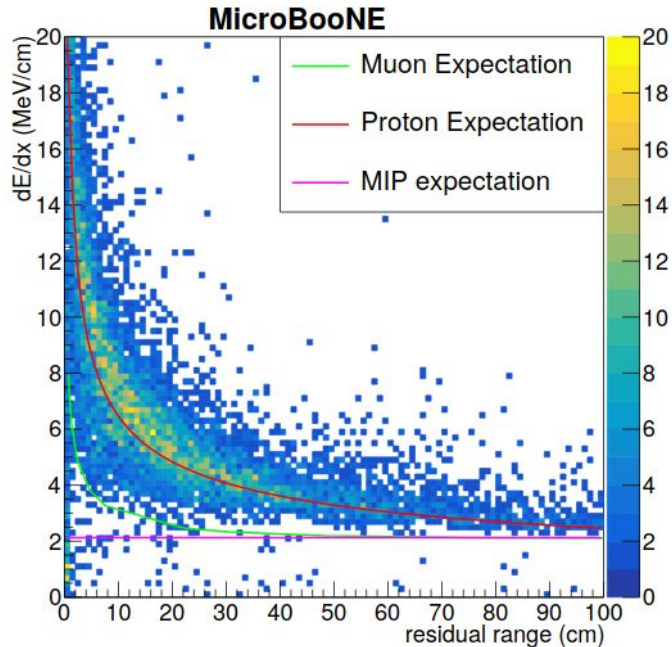
[JINST 15, P03022 \(2020\)](#)

# PID In MicroBooNE

- We convert the dE/dx vs residual range into a single number, a summed average distance from proton expectation
- Low is proton-like, high is not proton-like

$$PID = \chi^2_{\text{proton}} / d.o.f. = \sum_{\text{Track hits}} \left( \frac{(dE/dx_{\text{measured}} - dE/dx_{\text{theory}})^2}{\sigma_{dE/dx}} \right) / d.o.f.$$

JINST 15, P03022 (2020)

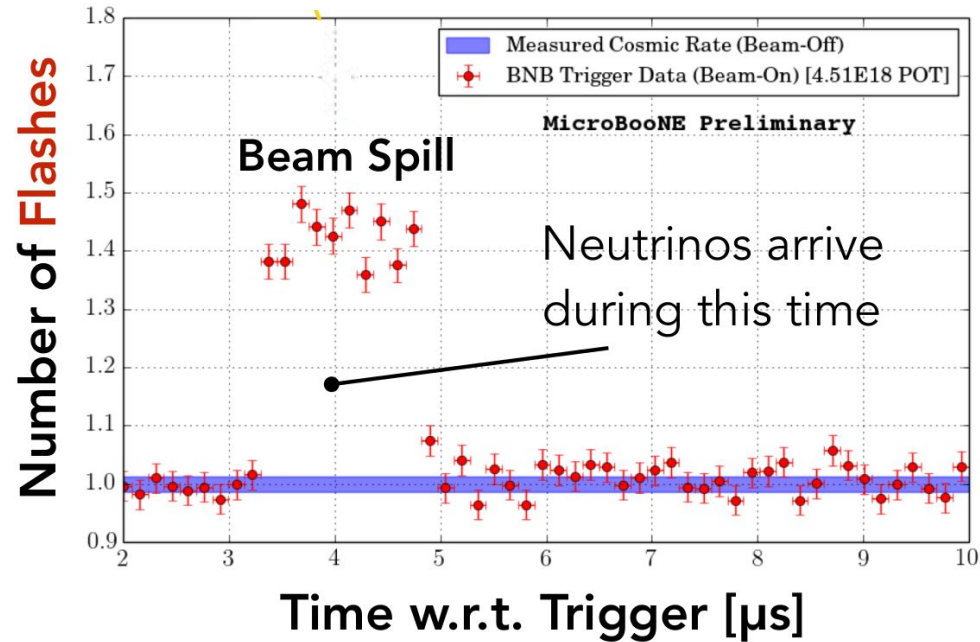


# Optical Information

Reduction of cosmic background  
is achieved using the Beam Spill Time

Neutrinos are delivered in spills of **1.6  $\mu\text{s}$**

The light information (prompt light  $\text{O}(\text{ns})$ )  
can be used to identify neutrino interactions  
during a spill



# Detector Response

$$\chi^2_{\mu} > \chi^2_p$$
$$\chi^2_{\mu} > 80 \ \& \ \chi^2_p < 30$$

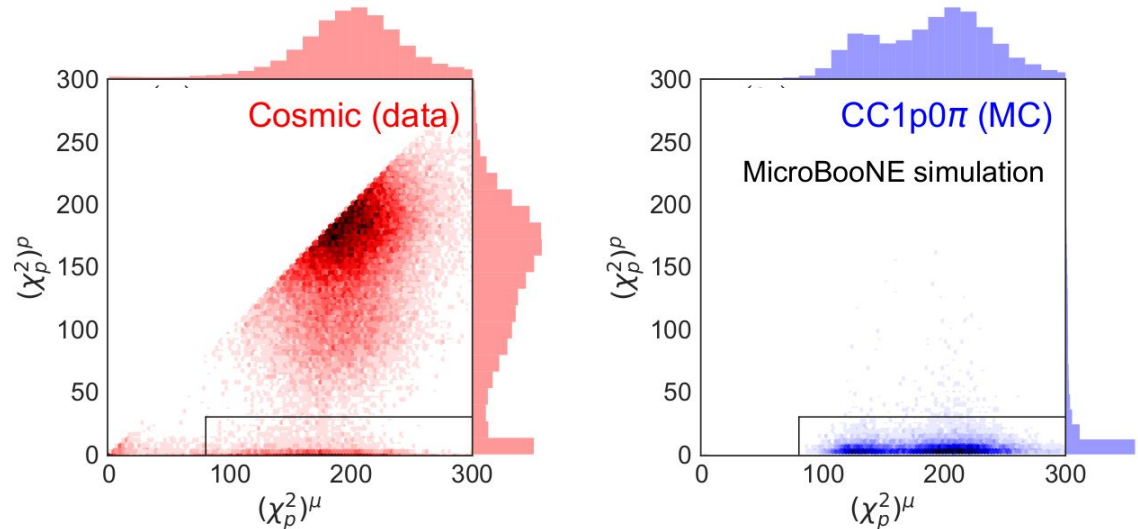
- **Energy deposition profile**

- Track length

- Scintillation light

- Non-collinearity

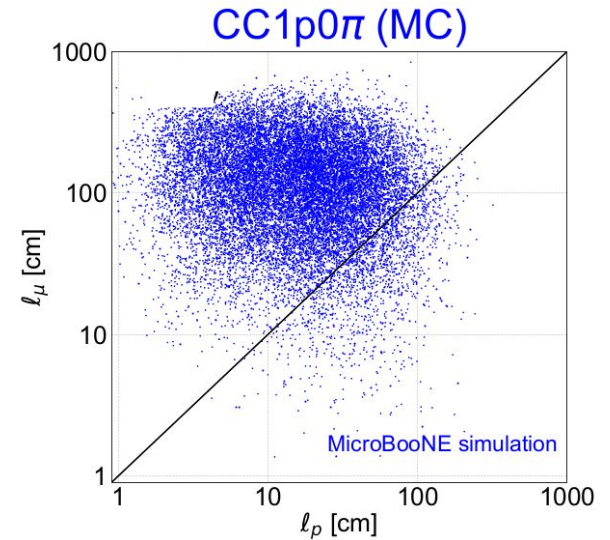
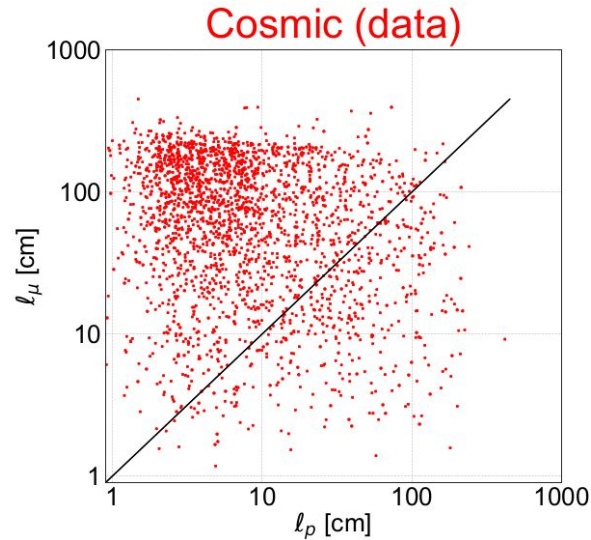
$$|\Delta\theta_{\mu p} - 90^\circ| < 55^\circ$$



# Detector Response

- Energy deposition profile
- **Track length**
- Scintillation light
- Non-collinearity  
 $|\Delta\theta_{\mu p} - 90^\circ| < 55^\circ$

$$l_\mu > l_p$$

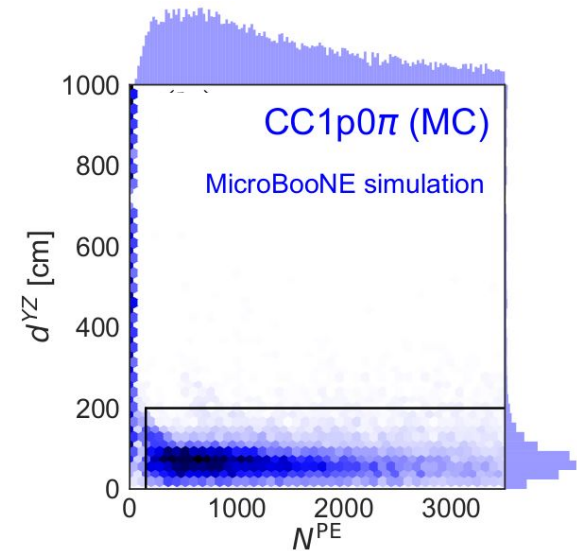
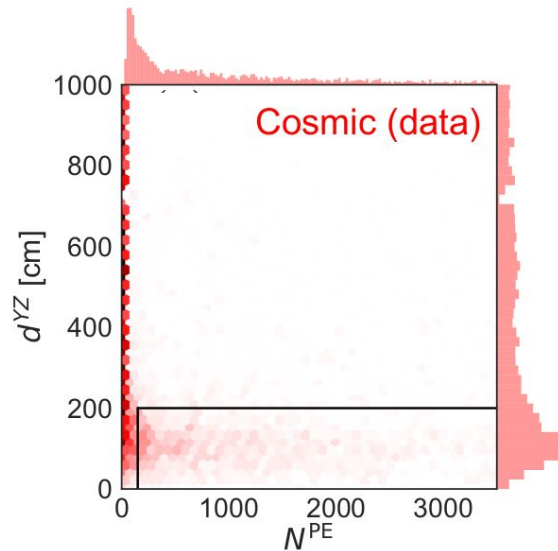


# Detector Response

# photoelectrons > 200  
YZ vertex-flash distance < 2 m

- Energy deposition profile
- Track length
- **Scintillation light**
- Non-collinearity

$$|\Delta\theta_{\mu p} - 90^\circ| < 55^\circ$$

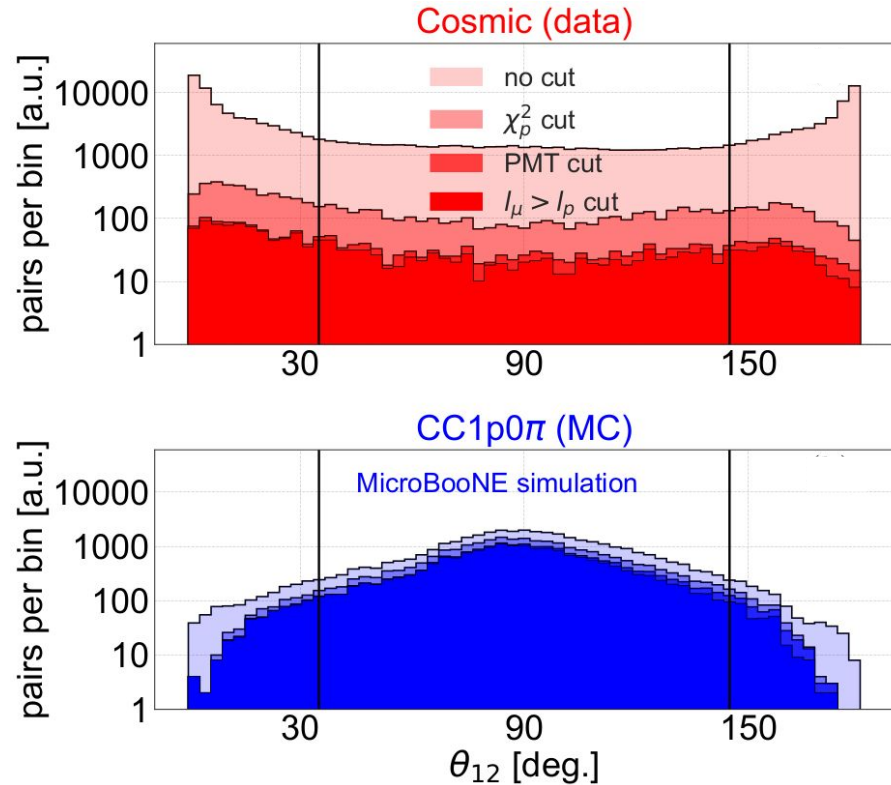
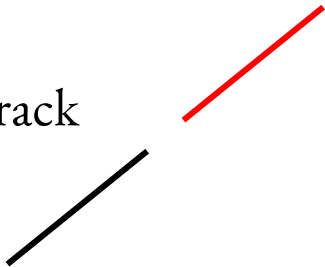


# Detector Response

- Energy deposition profile
- Track length
- Scintillation light
- **Non-collinearity**

$$|\Delta\theta_{\mu p} - 90^\circ| < 55^\circ$$

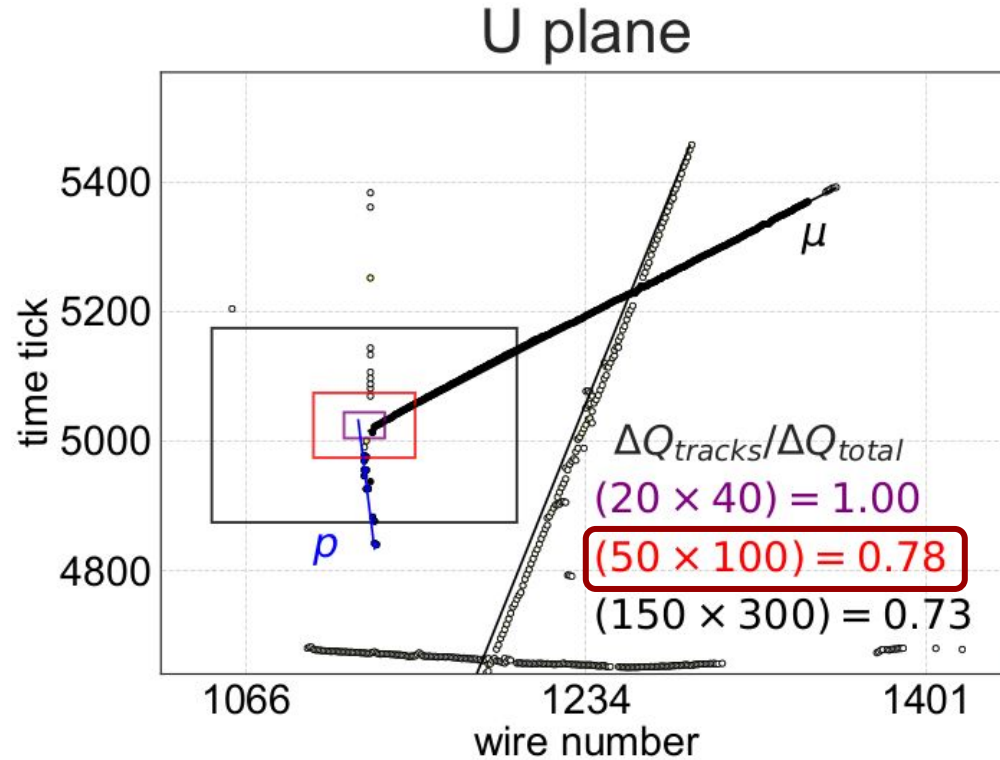
Broken track





# Kinematics

- Vertex activity

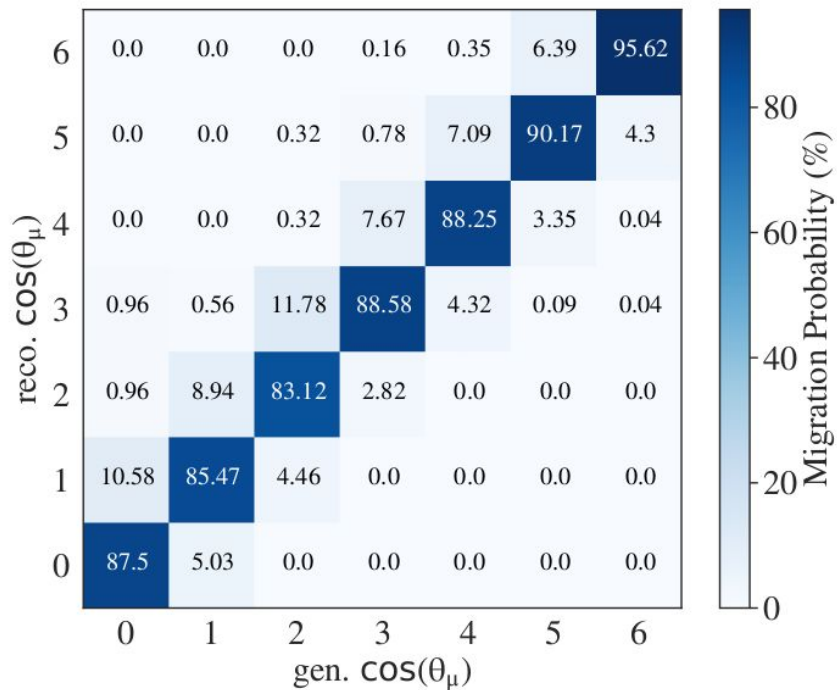


# Selection Cuts

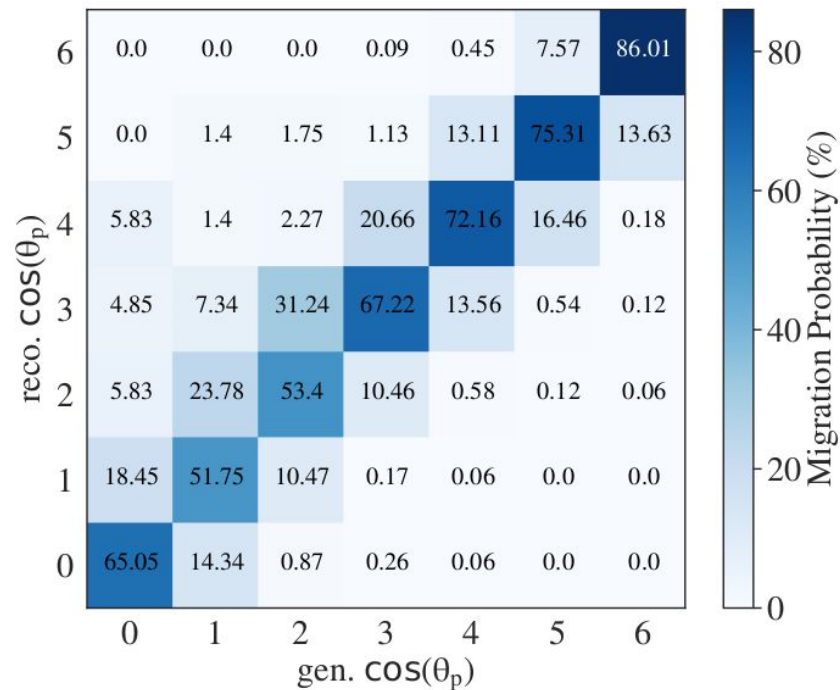
	DATA	simulated signal	
criterion	cosmic	CC1p0 $\pi$	purity
preselection			
preselection	155416(100.0%)	37228(100.0%)	13.1%
detector-response requirements			
$dE/dx$ profile	8327(5.4%)	25016(67.2%)	38.8%
optical filter	2256(1.5%)	19208(51.6%)	43.6%
track lengths	1874(1.2%)	17623(47.3%)	46.5%
collinearity	839(0.54%)	16796(45.1%)	50.5%
kinematical requirements			
vertex activity	467(0.30%)	15034(40.4%)	62.1%
coplanarity only	189(0.12%)	11824(31.8%)	75.2%
$p_T$ imbalance only	256(0.16%)	12261(32.9%)	69.3%
$\Delta\phi$ & $p_T$	104(0.07%)	10020(26.9%)	78.4%

# Migration Matrices

## MicroBooNE Simulation

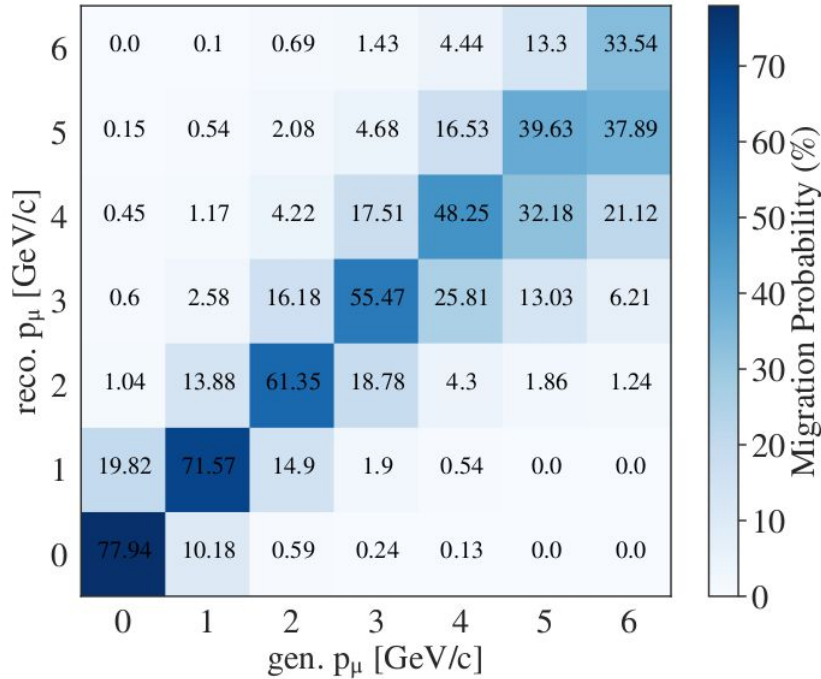


## MicroBooNE Simulation

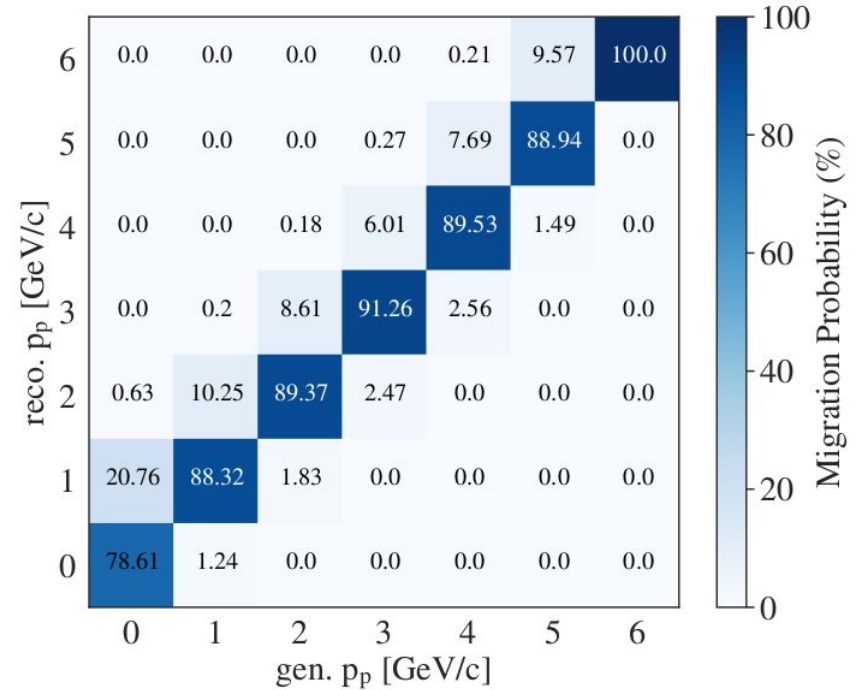


# Migration Matrices

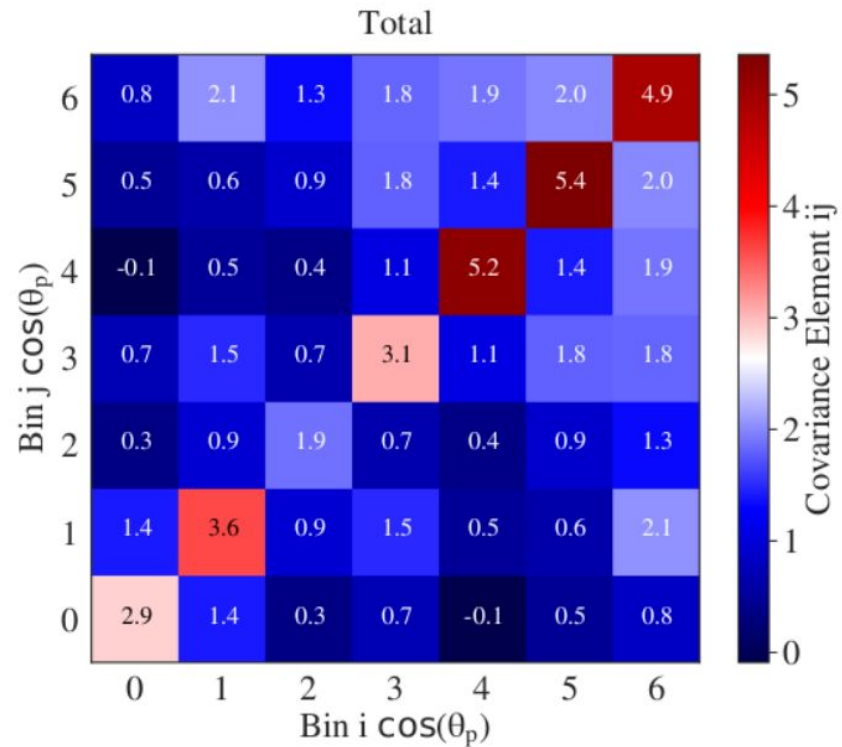
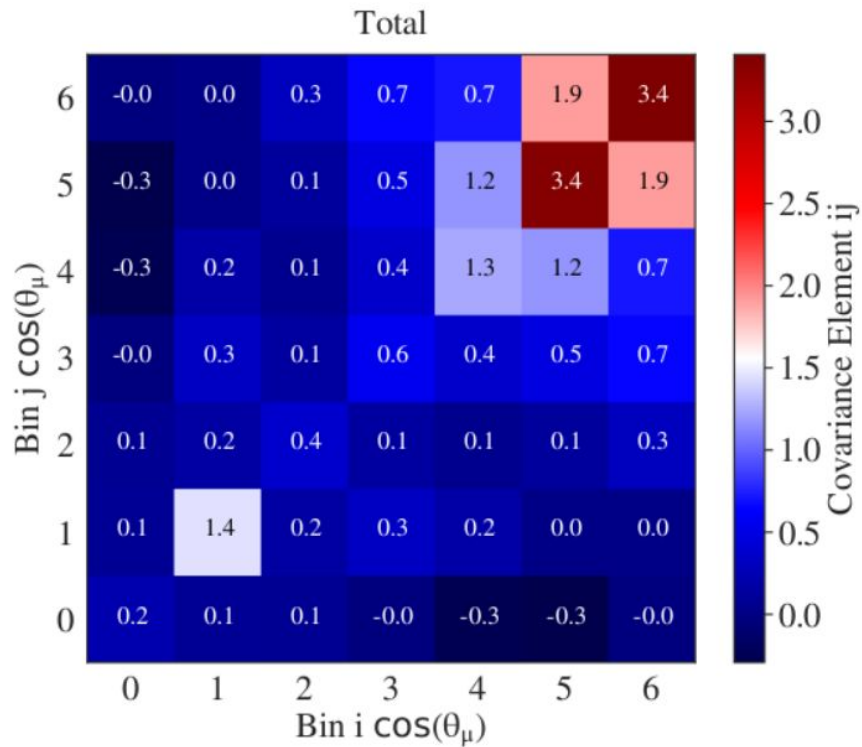
## MicroBooNE Simulation



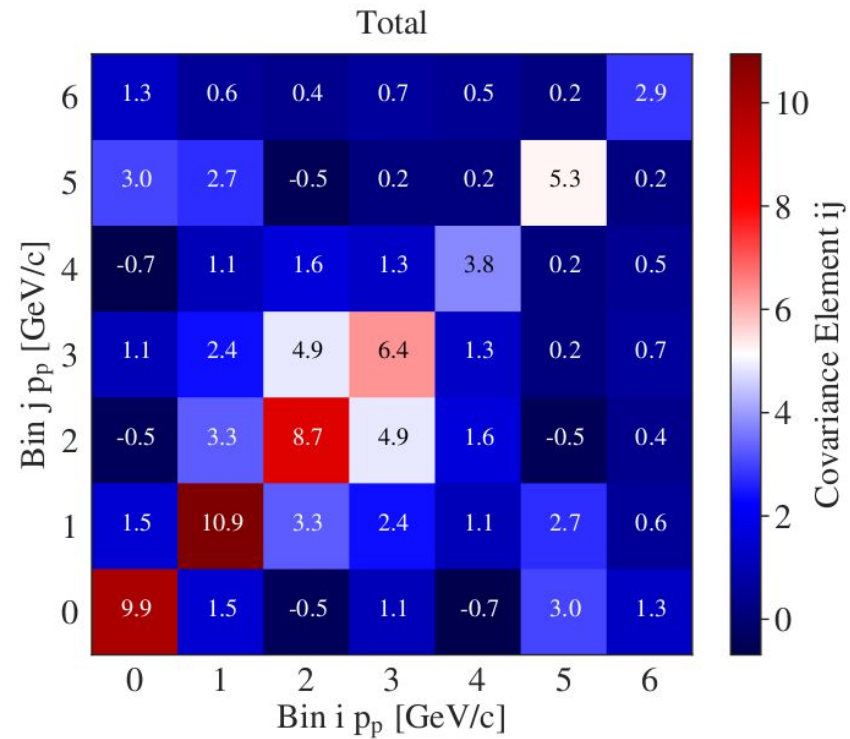
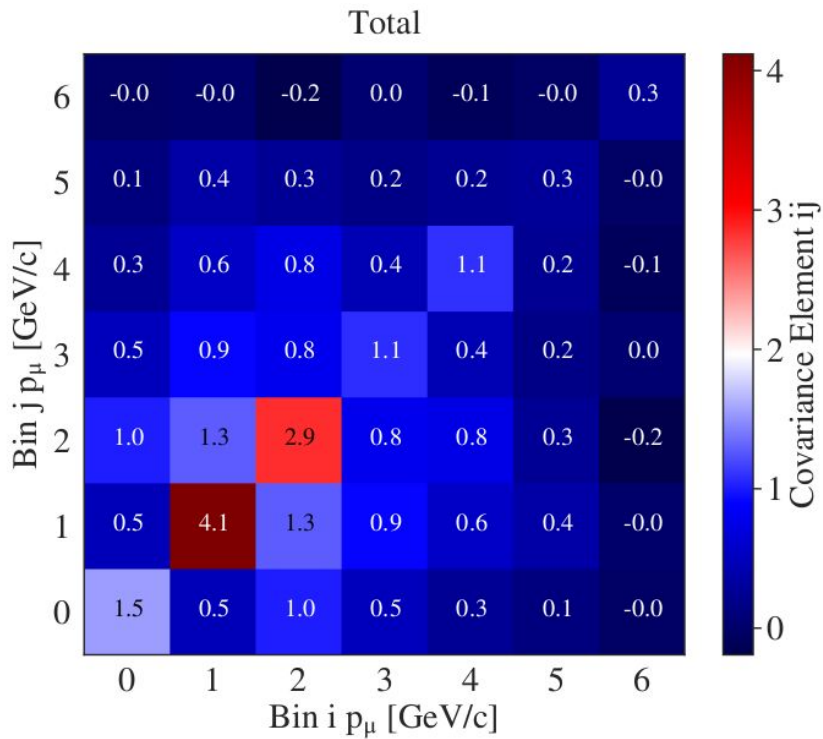
## MicroBooNE Simulation



# Covariance Matrices



# Covariance Matrices



# $\chi^2$ Comparisons For Each Kinematic Variable

Data - GENIE Nominal comparisons

Kinematic Variable	$-0.65 < \cos(\theta_\mu) < 0.95$	$-0.65 < \cos(\theta_\mu) < 0.8$
$P_\mu$	14.2/7	8.4/7
$\cos(\theta_\mu)$	33.8/7	7.3/6
$P_p$	2.8/7	5.1/7
$\cos(\theta_p)$	12.4/7	9.2/7

Agreement significantly improved if forward muon angles are excluded

# GENIE Reweightable Cross Section Parameters

$x_P$	Description of $P$	$\delta P/P$
$x_{M_A^{NC EL}}$	Axial mass for NC elastic	$\pm 25\%$
$x_{\eta^{NC EL}}$	Strange axial form factor $\eta$ for NC elastic	$\pm 30\%$
$x_{M_A^{CCQE}}$	Axial mass for CC quasi-elastic	$-15\% +25\%$
$x_{CCQE-Norm}$	Normalization factor for CCQE	
$x_{CCQE-PauliSup}$	CCQE Pauli suppression (via changes in Fermi level $k_F$ )	$\pm 35\%$
$x_{CCQE-VecFF}$	Choice of CCQE vector form factors (BBA05 $\leftrightarrow$ Dipole)	-
$x_{CCRES-Norm}$	Normalization factor for CC resonance neutrino production	
$x_{NCRES-Norm}$	Normalization factor for NC resonance neutrino production	
$x_{M_A^{CCRES}}$	Axial mass for CC resonance neutrino production	$\pm 20\%$
$x_{M_V^{CCRES}}$	Vector mass for CC resonance neutrino production	$\pm 10\%$
$x_{M_A^{NCRES}}$	Axial mass for NC resonance neutrino production	$\pm 20\%$
$x_{M_V^{NCRES}}$	Vector mass for NC resonance neutrino production	$\pm 10\%$
$x_{M_A^{COH\pi}}$	Axial mass for CC and NC coherent pion production	$\pm 50\%$
$x_{R_0^{COH\pi}}$	Nuclear size param. controlling $\pi$ absorption in RS model	$\pm 10\%$
$x_{R_{bkg}^{\nu p, CC1\pi}}$	Non-resonance bkg in $\nu p$ $CC1\pi$ reactions	$\pm 50\%$
$x_{R_{bkg}^{\nu p, CC2\pi}}$	Non-resonance bkg in $\nu p$ $CC2\pi$ reactions	$\pm 50\%$
$x_{R_{bkg}^{\nu n, CC1\pi}}$	Non-resonance bkg in $\nu n$ $CC1\pi$ reactions	$\pm 50\%$
$x_{R_{bkg}^{\nu n, CC2\pi}}$	Non-resonance bkg in $\nu n$ $CC2\pi$ reactions	$\pm 50\%$
$x_{R_{bkg}^{\nu p, NC1\pi}}$	Non-resonance bkg in $\nu p$ $NC1\pi$ reactions	$\pm 50\%$
$x_{R_{bkg}^{\nu p, NC2\pi}}$	Non-resonance bkg in $\nu p$ $NC2\pi$ reactions	$\pm 50\%$
$x_{R_{bkg}^{\nu n, NC1\pi}}$	Non-resonance bkg in $\nu n$ $NC1\pi$ reactions	$\pm 50\%$
$x_{R_{bkg}^{\nu n, NC2\pi}}$	Non-resonance bkg in $\nu n$ $NC2\pi$ reactions	$\pm 50\%$
$x_{A_{HT}^{BY}}$	$A_{HT}$ higher-twist param in BY model scaling variable $\xi_w$	$\pm 25\%$

$x_{C_{V1u}^{BY}}$

$C_{V1u}$  u valence GRV98 PDF correction param in BY model  $\pm 30\%$

$x_{C_{V2u}^{BY}}$

$C_{V2u}$  u valence GRV98 PDF correction param in BY model  $\pm 40\%$

$x_{CCDIS}$

Inclusive CC cross-section normalization factor

$x_{CC\bar{\nu}/\nu}$

$\bar{\nu}/\nu$  CC ratio

$x_{DIS-NuclMod}$

DIS nuclear modification (shadowing, anti-shadowing, EMC)

Also included comparison to Nieves

CCQE & MEC models



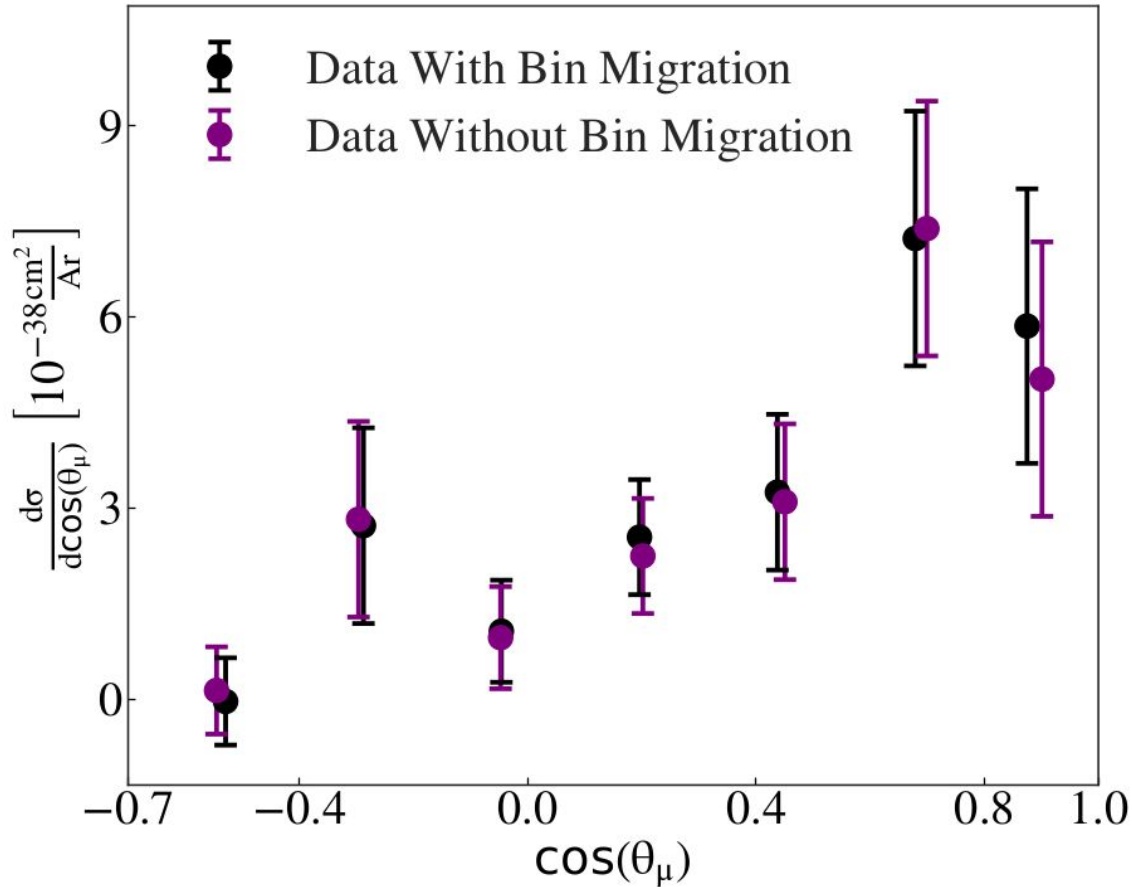
# GENIE Reweightable Hadronization & Resonant Parameters

$x_P$	Description of $P$	$\delta P/P$
$x_{AGKY}^{pT1\pi}$	Pion transverse momentum ( $p_T$ ) for $N\pi$ states in AGKY	-
$x_{AGKY}^{xF1\pi}$	Pion Feynman x ( $x_F$ ) for $N\pi$ states in AGKY	-
$x_{fz}$	Hadron formation zone	$\pm 50\%$
$x_{\theta_\pi}^{\Delta \rightarrow \pi N}$	Pion angular distribution in $\Delta \rightarrow \pi N$ (isotropic $\leftrightarrow$ RS)	-
$x_{BR}^{R \rightarrow X+1\gamma}$	Branching ratio for radiative resonance decays	$\pm 50\%$
$x_{BR}^{R \rightarrow X+1\eta}$	Branching ratio for single- $\eta$ resonance decays	$\pm 50\%$

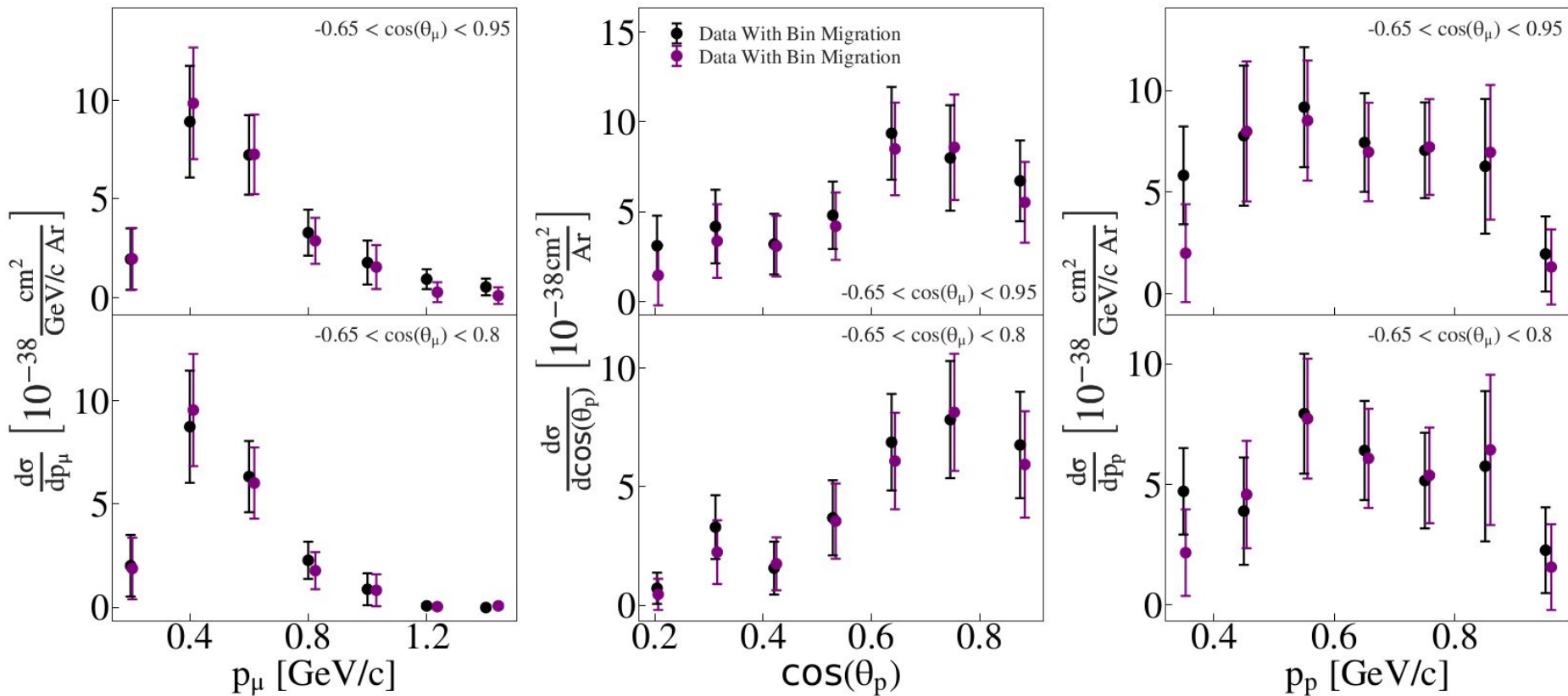
# GENIE Reweightable Hadron Transport Parameters

$x_P$	Description of $P$	$\delta P/P$
$x_{mfp}^N$	Nucleon mean free path (total rescattering probability)	$\pm 20\%$
$x_{ccx}^N$	Nucleon charge exchange probability	$\pm 50\%$
$x_{cl}^N$	Nucleon elastic reaction probability	$\pm 30\%$
$x_{inel}^N$	Nucleon inelastic reaction probability	$\pm 40\%$
$x_{abs}^N$	Nucleon absorption probability	$\pm 20\%$
$x_{\pi}^N$	Nucleon $\pi$ -production probability	$\pm 20\%$
$x_{mfp}^{\pi}$	$\pi$ mean free path (total rescattering probability)	$\pm 20\%$
$x_{ccx}^{\pi}$	$\pi$ charge exchange probability	$\pm 50\%$
$x_{cl}^{\pi}$	$\pi$ elastic reaction probability	$\pm 10\%$
$x_{inel}^{\pi}$	$\pi$ inelastic reaction probability	$\pm 40\%$
$x_{abs}^{\pi}$	$\pi$ absorption probability	$\pm 20\%$
$x_{\pi}^{\pi}$	$\pi$ $\pi$ -production probability	$\pm 20\%$

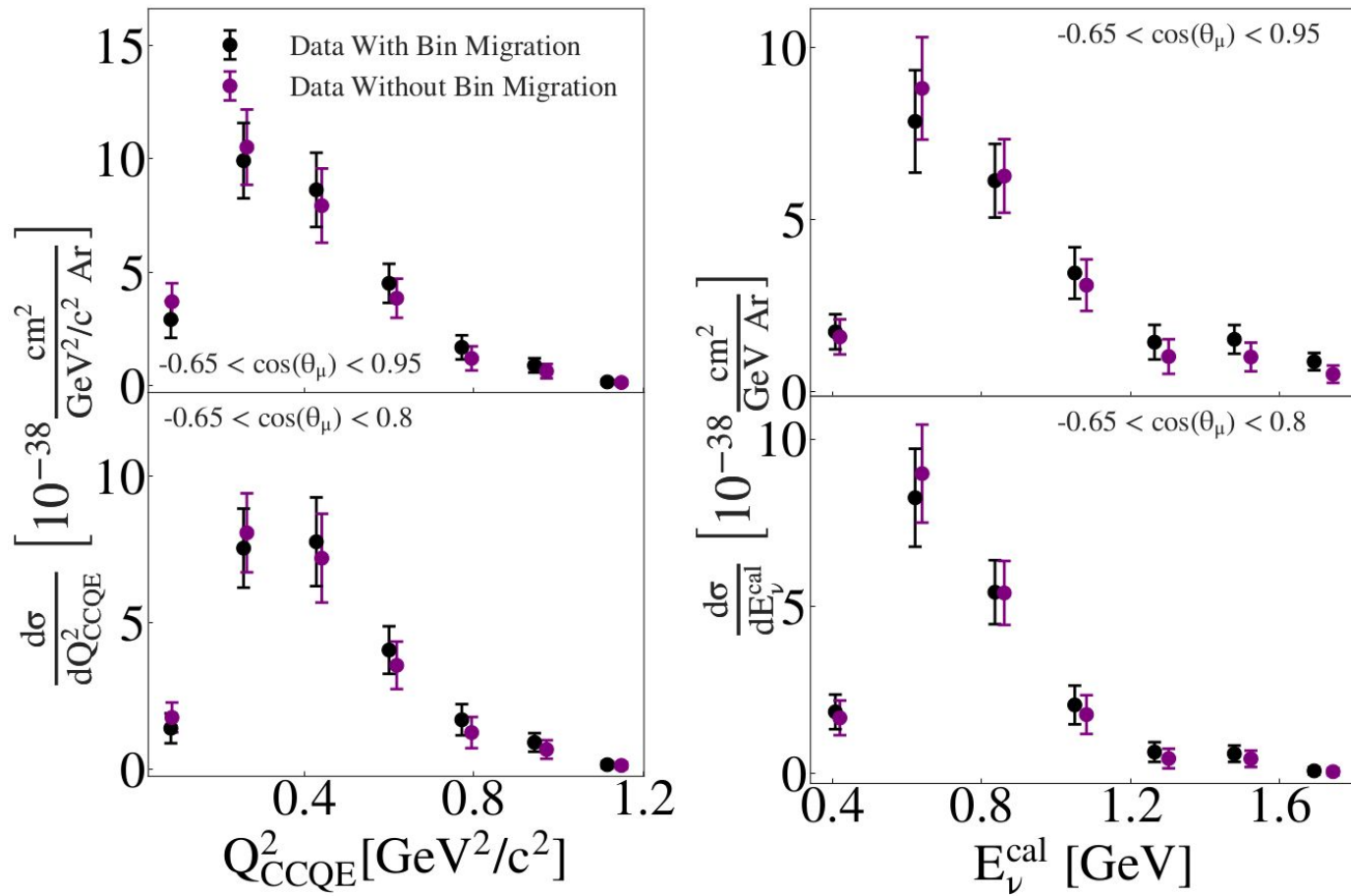
# Bin Migration Effects



# Bin Migration Effects

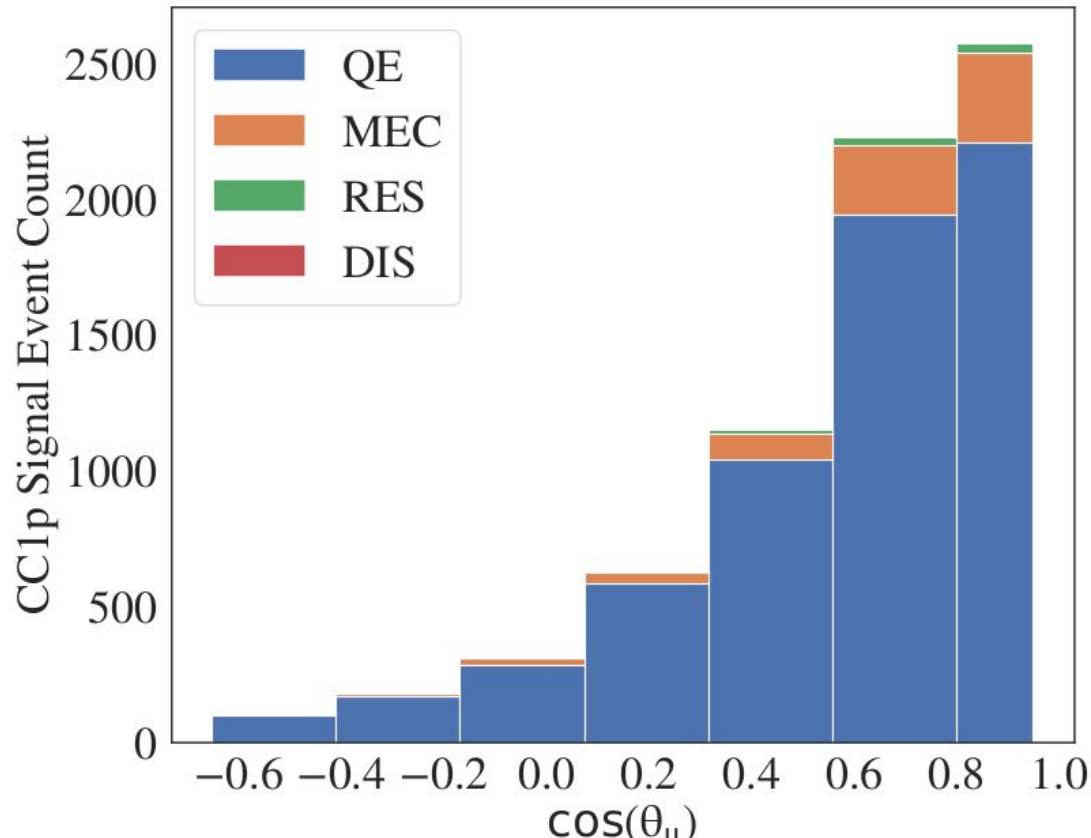


# Bin Migration Effects



# CC1p Signal Breakdown

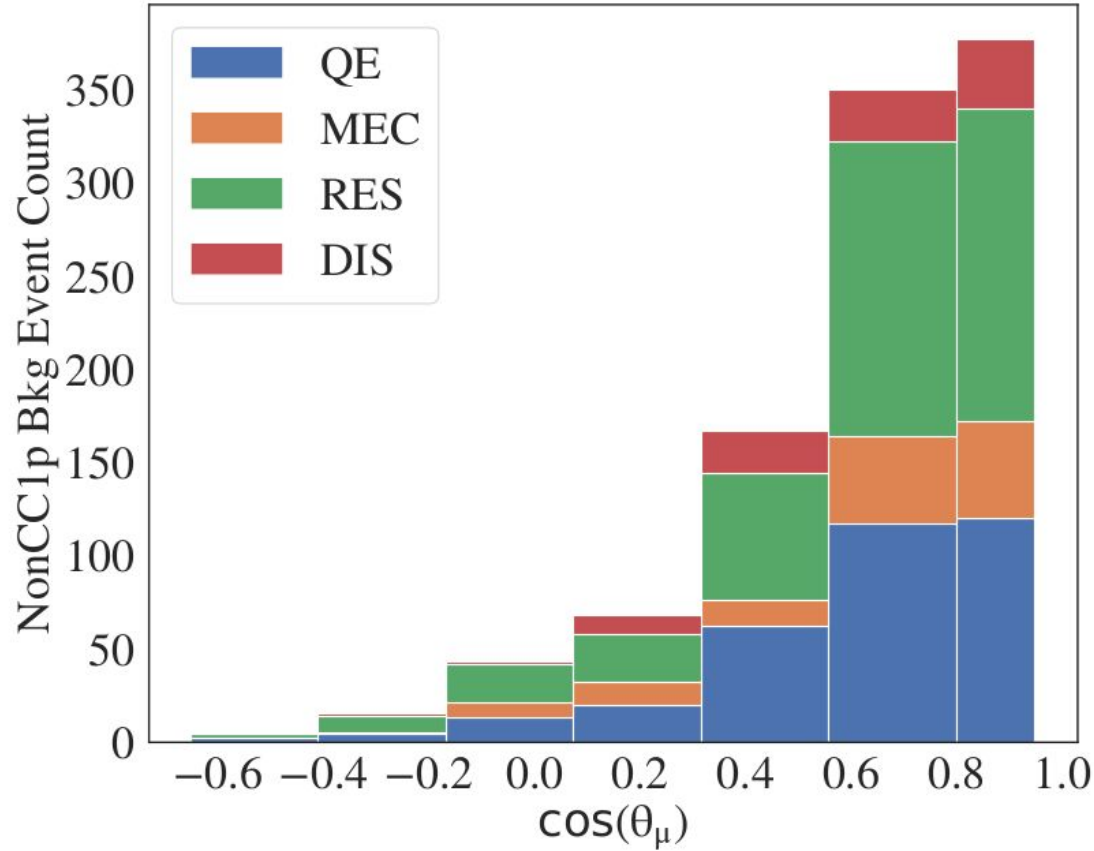
MicroBooNE Simulation



Interaction Mode	Fractional Contribution (%)
QE	87.9
MEC	10.6
RES	1.2
DIS	0.2

# Simulated Beam Related Background Breakdown

MicroBooNE Simulation



Interaction Mode	Fractional Contribution (%)
QE	33.0
MEC	13.1
RES	44.1
DIS	9.9

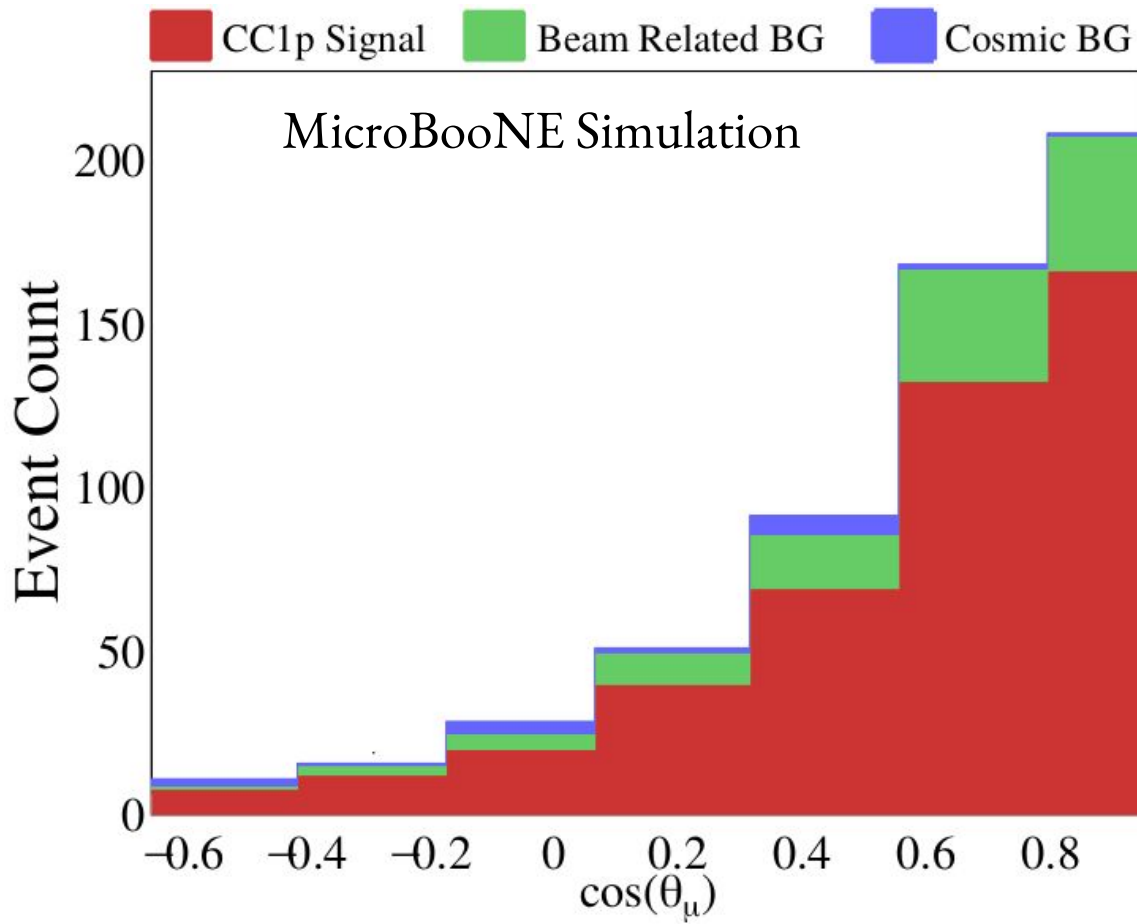
# Simulated Beam Related Background Breakdown

TABLE XIV: The number of simulated beam related background events from the other non  $\mu$ - $p$  pairs passing the analysis event selection cuts. The “beam-on” equivalent column lists the number of events scaled to the beam-on available POT.

background pairs	number of events	beam-on equivalent	fractional contribution to beam-on
broken tracks	$13 \pm 3.6$	$0.9 \pm 0.2$	$0.22 \pm 0.06\%$
$\mu^+p$	$11 \pm 3.3$	$0.8 \pm 0.2$	$0.18 \pm 0.06\%$
$pp$	$18 \pm 4.2$	$1.2 \pm 0.3$	$0.30 \pm 0.07\%$
$\pi p$	$109 \pm 10.4$	$7.5 \pm 0.7$	$1.82 \pm 0.20\%$
$ep$	$19 \pm 4.4$	$1.3 \pm 0.3$	$0.32 \pm 0.07\%$
$\mu D$	$1.0 \pm 1.0$	$0.1 \pm 0.1$	$0.02 \pm 0.02\%$
others	$1.0 \pm 1.0$	$0.1 \pm 0.1$	$0.02 \pm 0.02\%$



# $\text{Cos}(\theta_\mu)$ Topological Breakdown



# Effective Efficiency & Unfolding

$$\epsilon_n = \frac{N_{\text{reconstructed in bin } n}^{\text{true CC1p0}\pi} \text{ (generated in any bin)}}{N_{\text{generated in bin } n}^{\text{true CC1p0}\pi}}$$

$$= \frac{N_{\text{reconstructed in any bin}}^{\text{true CC1p0}\pi} \text{ (generated in bin } n\text{)}}{N_{\text{generated in bin } n}^{\text{true CC1p0}\pi}}$$

Standard efficiency

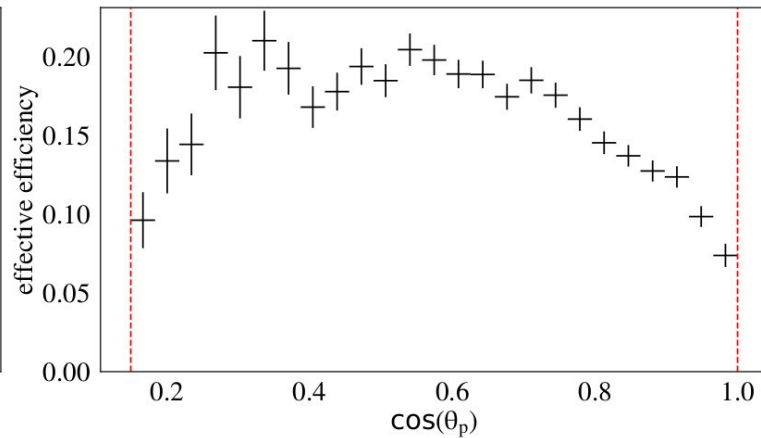
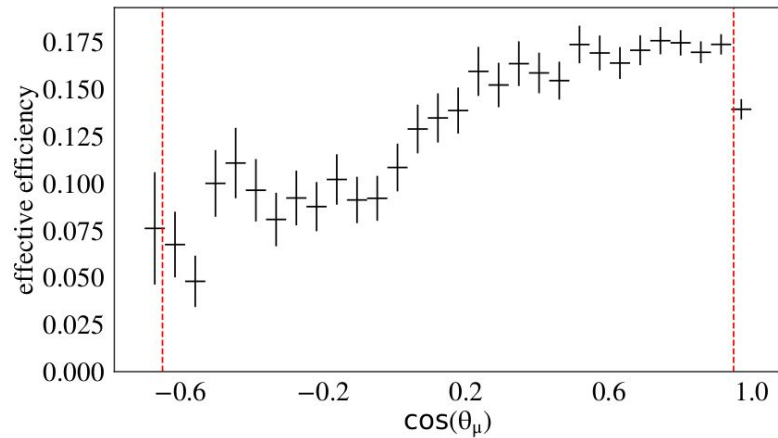
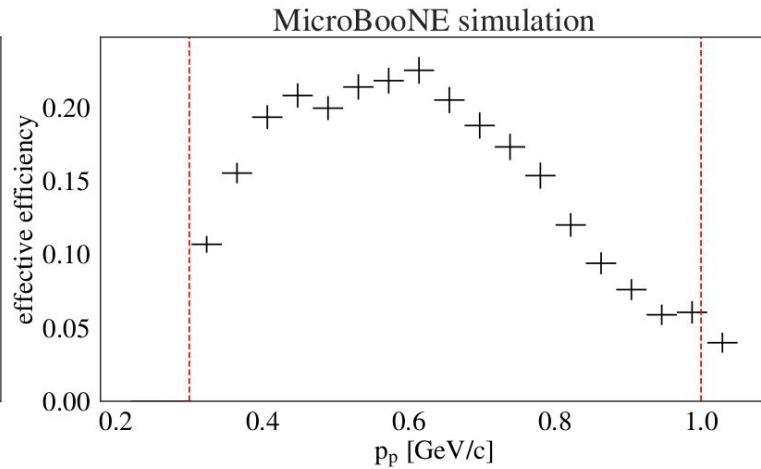
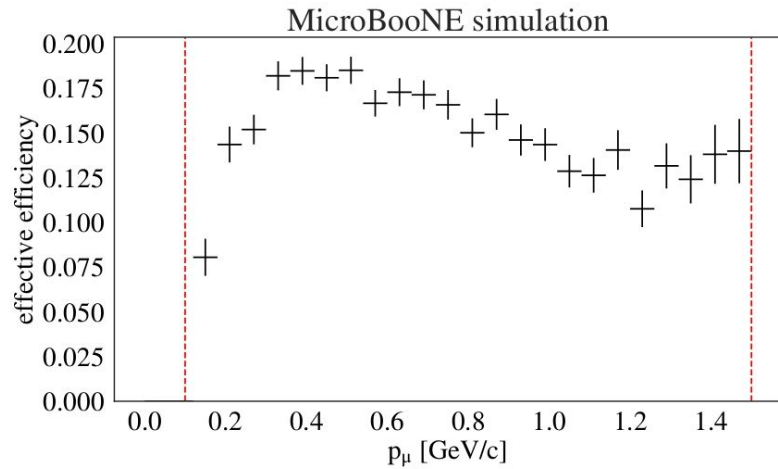
$$+ \frac{\sum_{j=1}^{N_{\text{bins}}} N_{\text{reconstructed in bin } n}^{\text{true CC1p0}\pi} \text{ (generated in bin } j \neq n\text{)}}{N_{\text{generated in bin } n}^{\text{true CC1p0}\pi}}$$

$$- \frac{\sum_{k=1}^{N_{\text{bins}}} N_{\text{reconstructed in bin } k \neq n}^{\text{true CC1p0}\pi} \text{ (generated in bin } n\text{)}}{N_{\text{generated in bin } n}^{\text{true CC1p0}\pi}}$$

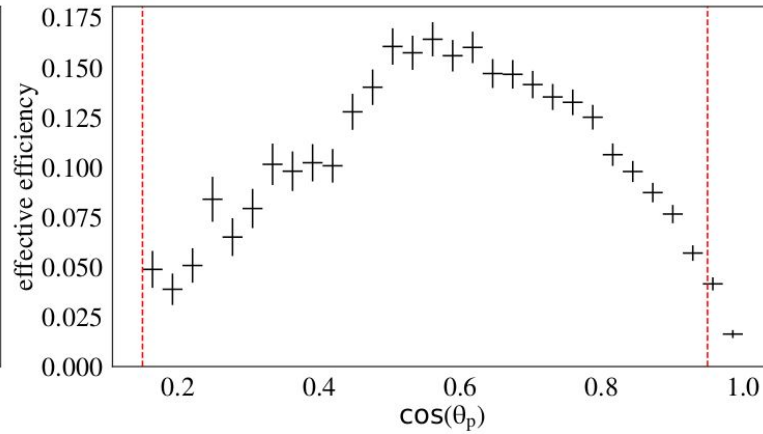
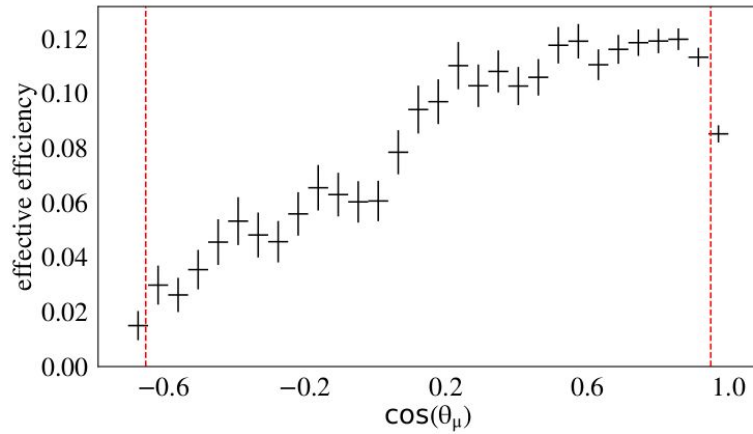
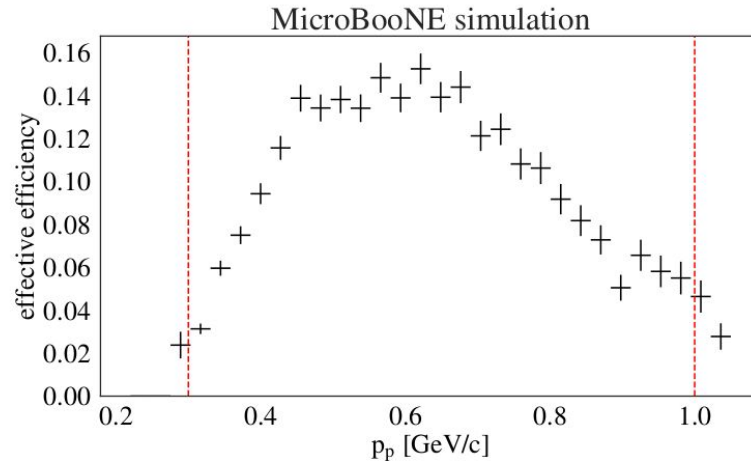
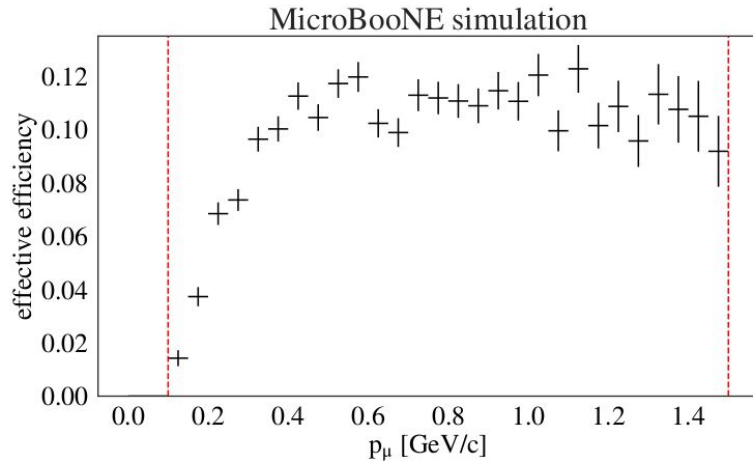
Bin migration effects

Results reported as a function of **true variables**

# Effective Efficiencies



# Effective Efficiencies



# Vertex Distribution

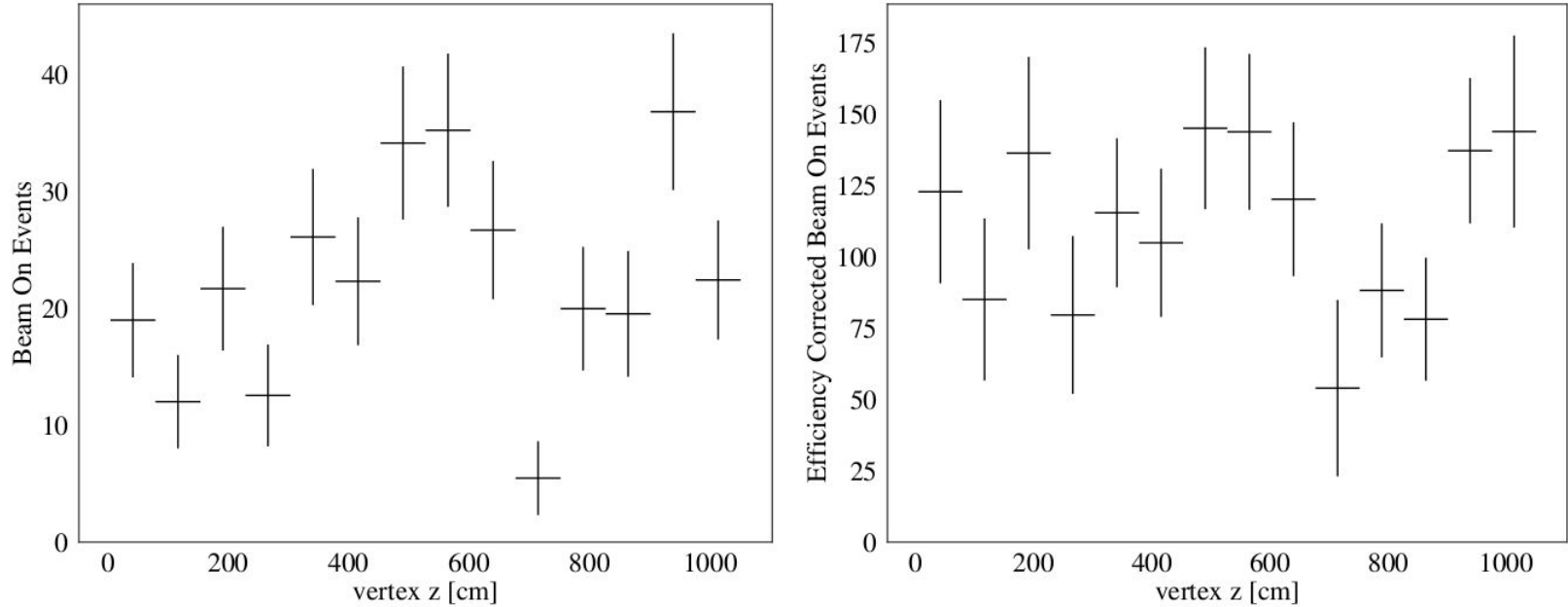


FIG. 19. Vertex  $z$  distribution for the measured events, after the beam related MC background has been subtracted, before (left) and after (right) detection efficiency corrections. No small- $z$  enhancement is observed and, with efficiency corrections, the measured distribution is consistent with that of a uniform neutrino interaction vertex.

# Kinematical Correlations

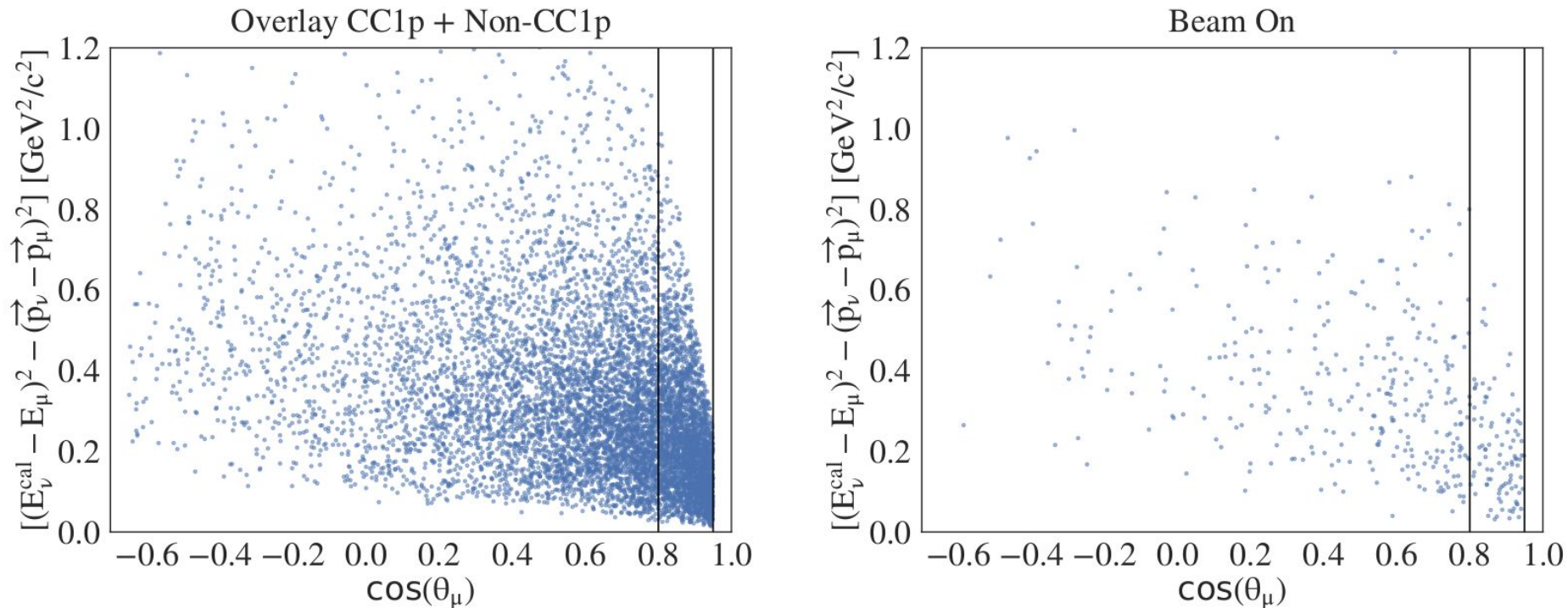


FIG. 5. Kinematic correlations between the high-level variable  $Q_{CQE}^2 = (E_\nu^{\text{cal}} - E_\mu)^2 - (\vec{p}_\nu - \vec{p}_\mu)^2$  and  $\text{cos}(\theta_\mu)$  for the overlay sample (left) and beam on (right). The black vertical lines indicate the limits of the last  $\text{cos}(\theta_\mu)$  bin.

# Systematic Uncertainties

$$E^{\text{syst}} = E^{\text{flux}} + E^{\text{xsec}} + E^{\text{detector}}$$

multisim

unisim

Generation of several MC replicas  
called “universes”

Changing one detector parameter at a time  
using the relevant uncertainty

Parameter variation within their  
uncertainties

Each variation requires a  
new MC generation

Performed via  
event reweighting

Take the difference between the central  
value cross section and the variation as the  
 $1\sigma$  uncertainty on the cross section

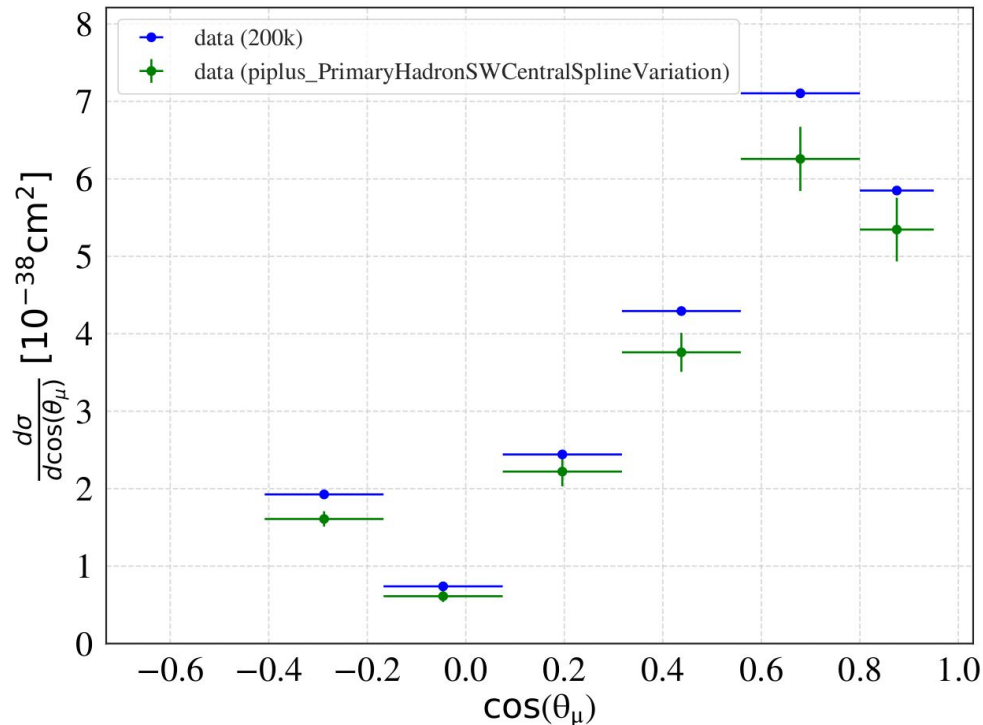
One MC run required

# Flux Systematic Uncertainties

1. uncertainties related to hadron production ( $\pi^\pm, K^\pm, K^0$ ) that arise due to uncertainties in the production of secondary particles when protons collide with the beryllium target.
2. all other related uncertainties, which are here called “non-hadron”. These uncertainties arise from errors in estimating the current that runs in the horn conductor, the depth by which such current penetrates the conductor ("skin effect") and the pion and nucleon cross sections (total, inelastic, and quasi-elastic) on aluminum and beryllium.



# Flux Systematic Uncertainties



$$\text{Flux Systematic Uncertainty}_{bin\ i} = \sqrt{\frac{(\text{nominal} - \mu)_{bin\ i}^2}{12} + \sigma_{bin\ i}^2}$$

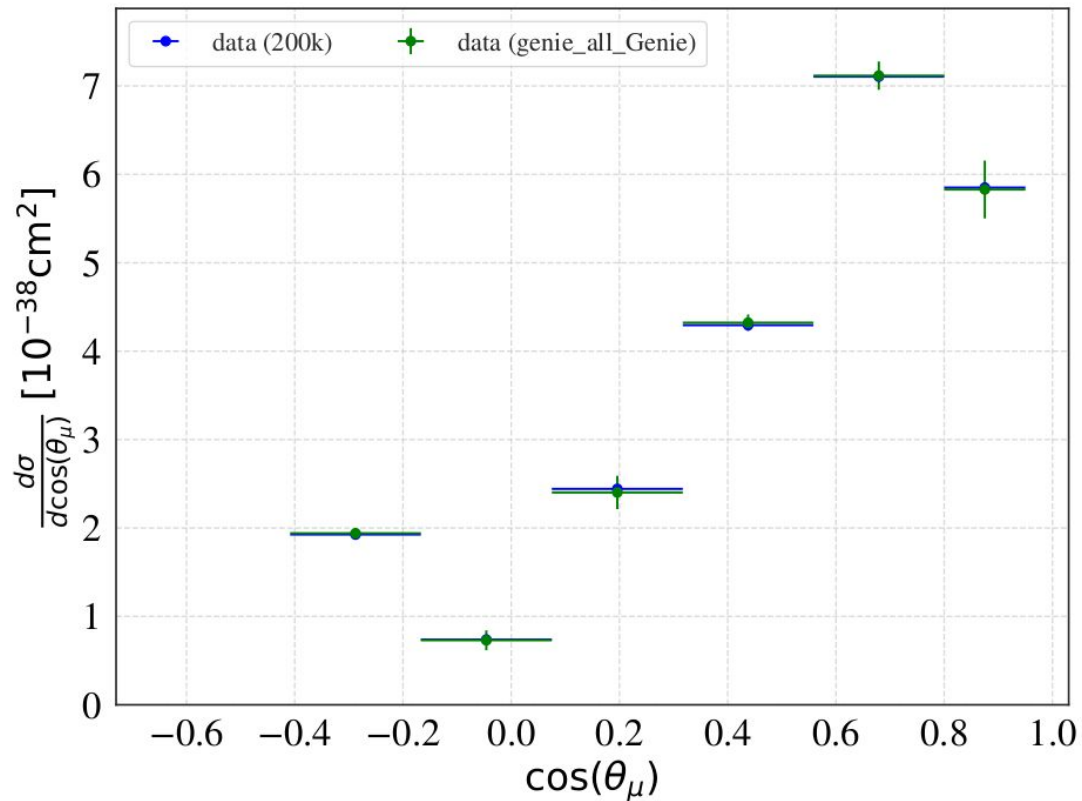
# Cross Section Systematic Uncertainties

## 1.6 Genie Cross Section Uncertainty

The study of the uncertainties related to the genie cross sections followed. The EventWeight label that is relevant for this part of the analysis is “genie\_all” and the knobs that tweaked within the context of this label are shown below starting from 200k events.

- genie\_qema
- genie\_ncelAxial
- genie\_ncelEta
- genie\_qevc
- genie\_ccresAxial
- genie\_ccresVector
- genie\_ncresAxial
- genie\_ncresVector
- genie\_cohMA
- genie\_cohR0
- genie\_NonResRvp1pi
- genie\_NonResRvbarp1pi
- genie\_NonResRvp2pi
- genie\_NonResRvbarp2pi
- genie\_ResDecayGamma
- genie\_ResDecayEta
- genie\_ResDecayTheta
- genie\_NC
- genie\_DISAth
- genie\_DISBth
- genie\_DISCv1u
- genie\_DISCv2u
- genie\_AGKYxF
- genie\_AGKYpT
- genie\_FormZone
- genie\_FermiGasModelKf
- genie\_FermiGasModelSf
- genie\_IntraNukeNmfp
- genie\_IntraNukeNcex
- genie\_IntraNukeNel
- genie\_IntraNukeNinel
- genie\_IntraNukeNabs
- genie\_IntraNukeNpi
- genie\_IntraNukePImfp
- genie\_IntraNukePIcex
- genie\_IntraNukePIel
- genie\_IntraNukePIinel
- genie\_IntraNukePIabs
- genie\_IntraNukePIpi

# Cross Section Systematic Uncertainties

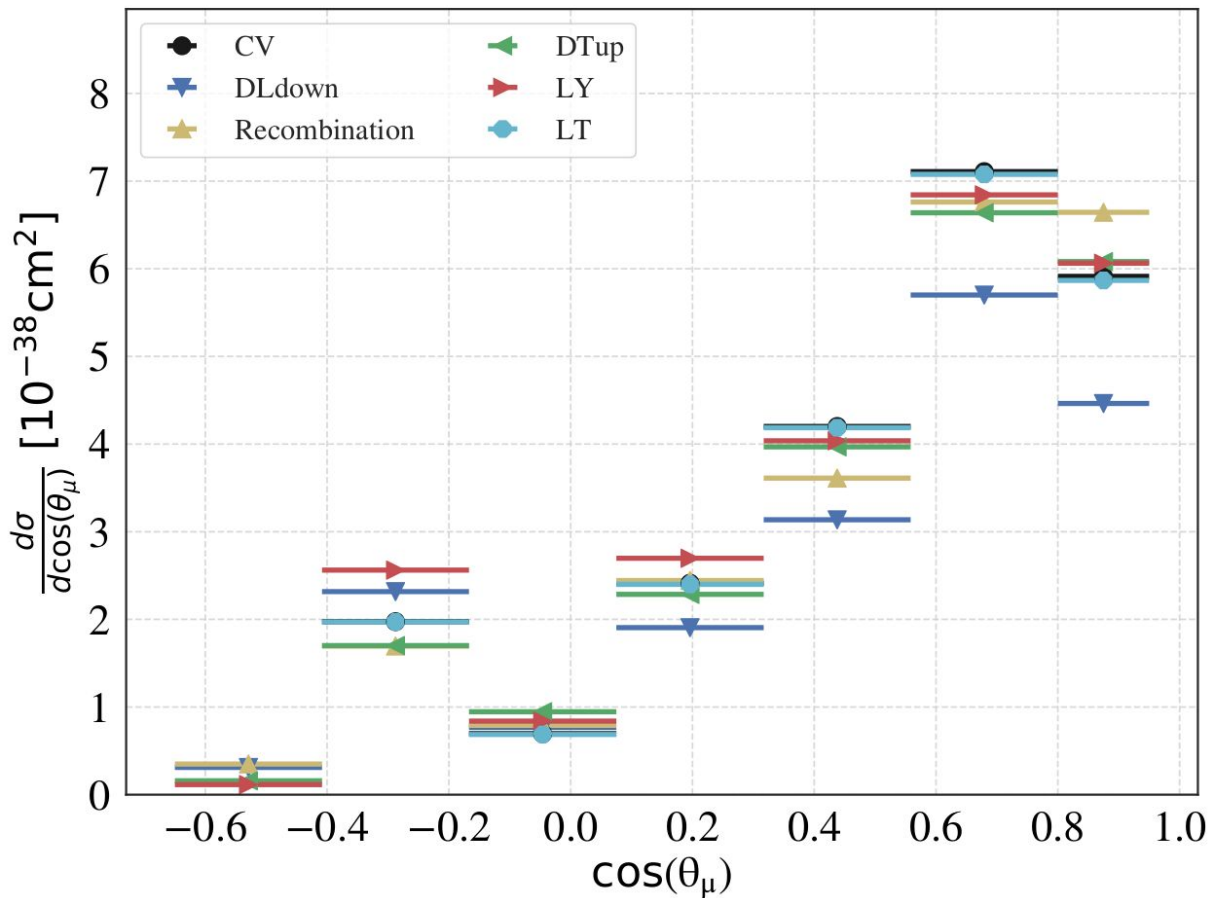


$$\text{Genie Systematic Uncertainty}_{bin\ i} = \sqrt{\frac{(\text{nominal} - \mu)_{bin\ i}^2}{12} + \sigma_{bin\ i}^2}$$

# Detector Systematic Uncertainties

In order to assess the detector systematic uncertainties, several samples are generated, where one detector parameter is varied at a time, and the same neutrino events are generated on top of the same cosmic background. Each of the following detector parameters are varied by  $1\sigma$ : longitudinal diffusion value, light yield outside the TPC, transverse diffusion, electron lifetime, recombination model and space charge map. The electronic response variation uses simulated neutrino events on top of a simulated cosmic background with CORSIKA [1]. The cross section is extracted using each one of those samples instead of the nominal simulation, where the relevant efficiencies, beam related background and data driven correction are rederived using the same procedures described in the paper. The uncertainty for each sample is obtained as the difference between the nominal cross section and the variation sample.

# Detector Systematic Uncertainties



# Dynamically Induced Charge

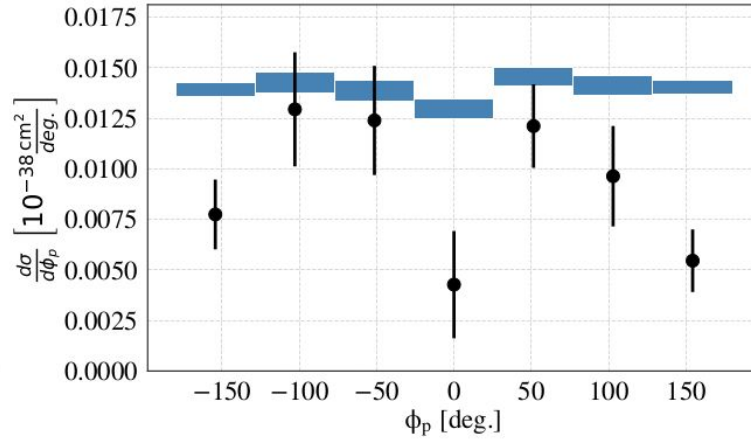
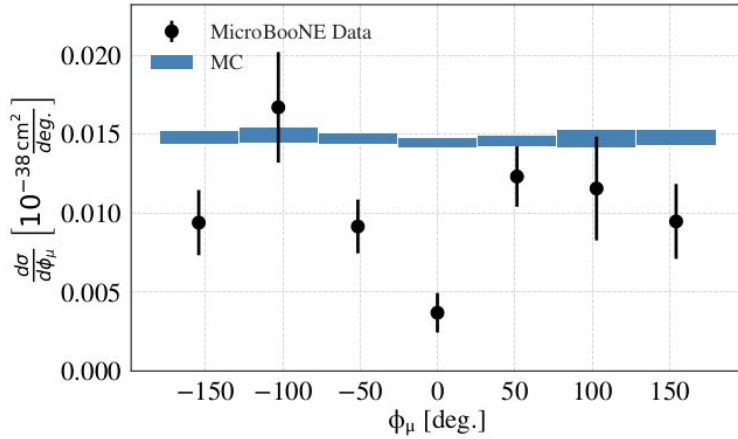
The extracted cross sections are expected to be independent of the azimuthal angle  $\phi$ . However, the simple model used to simulate the effect of induced charge on neighboring TPC wires leads to a low reconstruction efficiency of tracks perpendicular to the wire planes ( $\phi \approx 0$  and  $\phi \approx \pm\pi$ ) that created an artificial  $\phi$  dependence to the cross section. We correct for this effect using an iterative procedure. We first reweight events with a muon track falling in the  $\phi \approx 0$  bin and  $|\sin\theta| > 0.3$

to the weighted average of the cross sections in all other bins of  $\phi_\mu$  where  $|\sin\theta| > 0.3$ . Due to the coplanarity requirement, this reweighting affects the distribution of  $\phi_p \approx \pm\pi$ . We repeat the process starting from a proton track with  $\phi_p \approx 0$  until the cross section change is less than 0.01%, typically after 5 iterations.

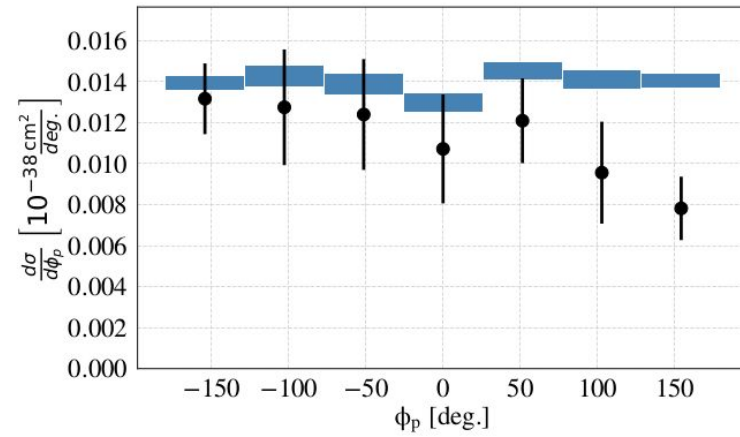
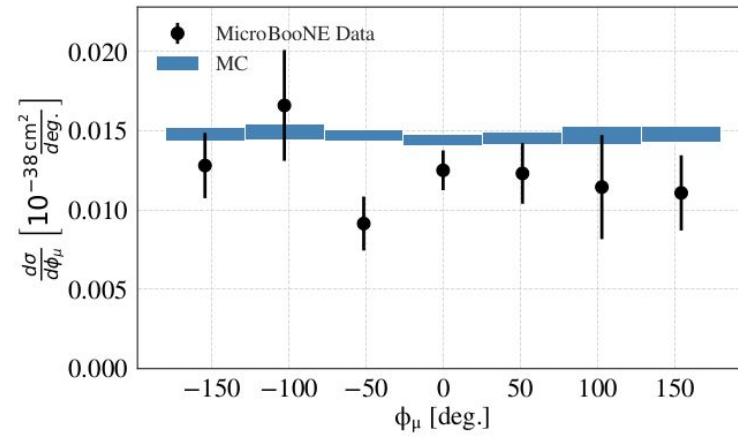
# Dynamically Induced Charge

In order to address the deficit in the central  $\phi$  bin due to the missing dynamically induced charge effects, we apply a data driven correction, as described in the paper. We estimate the uncertainty due to this procedure by producing a dedicated dynamically induced charge variation sample and by treating the sample as data. The difference between the cross section extracted using this variation as data and the nominal MC cross section results is assigned to be the uncertainty due to our data driven correction.

# Data Driven Correction



Before



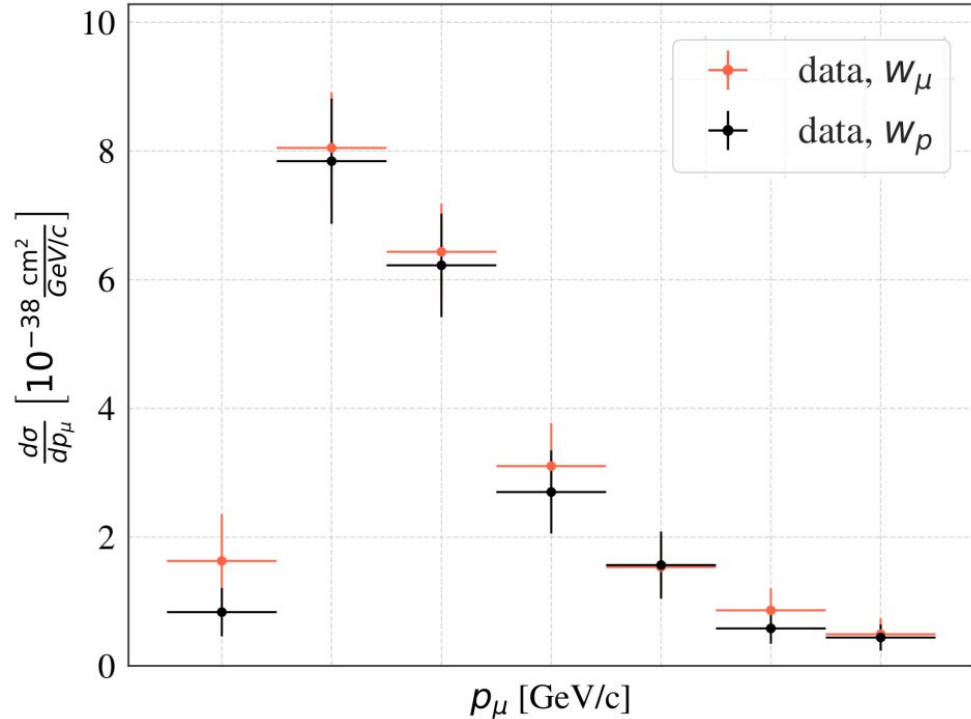
After



# Decoupling $\mu$ - $p$ Uncertainties

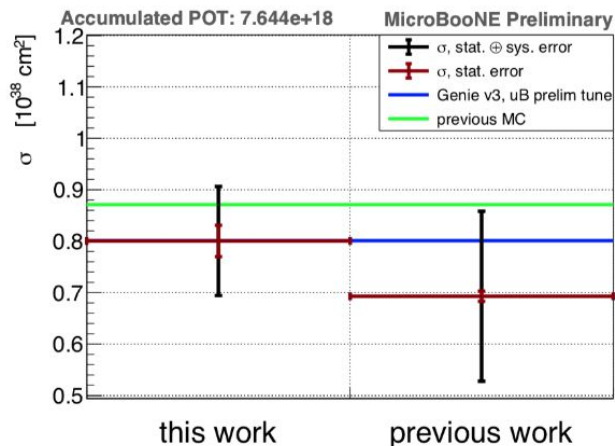
Due to the limited MC and data statistics, we divide the data into two three-dimensional (3D) bins and extract the differential cross-sections in each of the decoupled 3D spaces. This is justified in the limit that the muon and proton efficiencies are decoupled. This is not the case in our exclusive measurement.

$$\Delta_{\text{sys}}^{2 \times 3D} = \frac{1}{2} \left| \frac{d\sigma}{dp_\mu}(w_\mu) - \frac{d\sigma}{dp_\mu}(w_p) \right|$$



# Drastically Reduced Systematic Uncertainties

MICROBOONE-NOTE-1075-PUB · MICROBOONE-NOTE-1069-PUB



Flux-integrated cross section  
**consistent with previous measurement**

**Drastically reduced** systematic uncertainties

Instead of cosmic ray simulation, now use overlay: simulated neutrino interactions overlaid on real cosmic data  $\rightarrow$  no uncertainty in cosmic ray model

Source	Uncertainty	
	Previous Analysis	This Analysis
Detector response	16.2%	3.3%
Cross section	3.9%	2.7%
Flux	12.4%	10.5%
Dirt background	10.9%	3.3%
Cosmic ray background	4.2%	N/A
POT counting	2.0%	2.0%
CRT	N/A	1.7%
Total Sys. Error	23.8%	12.1%
Statistics	1.4%	3.8%
<b>Total (Quadratic Sum)</b>	<b>23.8%</b>	<b>12.7%</b>

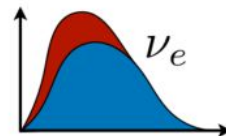
PRL 123, 131801 (2019)

# The Ingredients

## Leveraging Correlations to Minimize Systematics

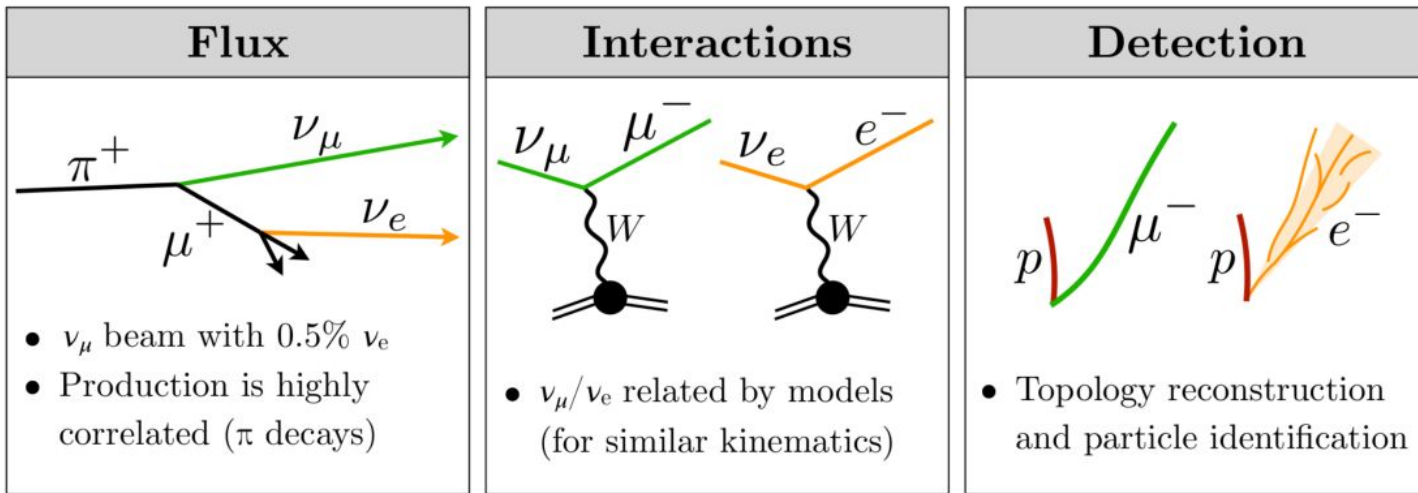


Given the small appearance probability of  $\sim 0.15\%$ , we are looking for **an excess** of  $\nu_e$  interactions on top of the  $0.5\%$  **intrinsic to the beam**



A key to MicroBooNE's single-detector appearance measurement is using  $\nu_\mu$  events to constrain the small  $\nu_e$  flux intrinsic to the beam

Credit A.Mastbaum



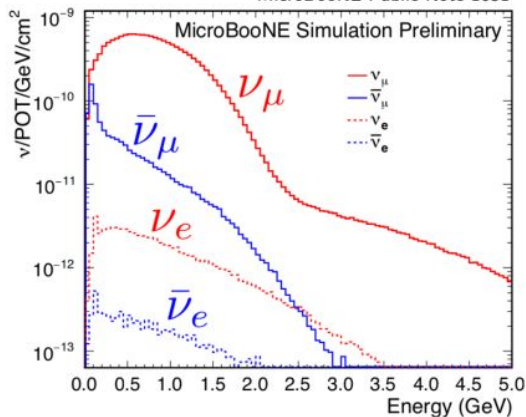
# Neutrino Flux

## Flux Modeling & Uncertainties

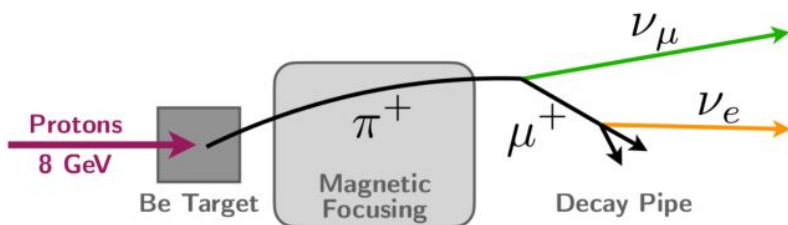


MicroBooNE Public Note 1031

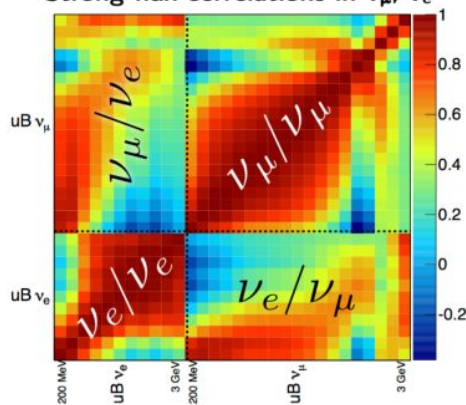
- **Fermilab Booster Neutrino Beam**
  - ▶ 8 GeV protons on a Be target
  - ▶ Primarily a  $\nu_\mu$  beam, with 0.5%  $\nu_e$
- **MiniBooNE-based flux uncertainties**
  - ▶ Beamline modeling in Geant4
  - ▶  $\pi$ ,  $K$  production (HARP, SciBooNE)
  - ▶  $\pi$ , nucleon interactions
  - ▶ ~12% integral flux uncertainty
- **Leverage strong  $\nu_\mu/\nu_e$  correlations** →



Credit A.Mastbaum

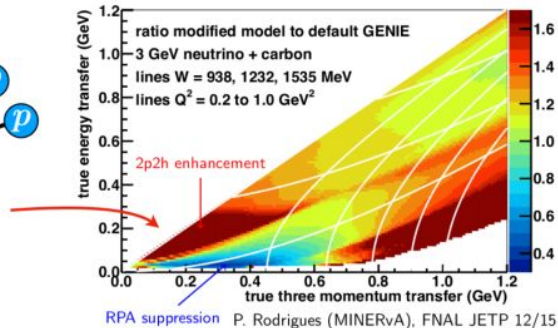
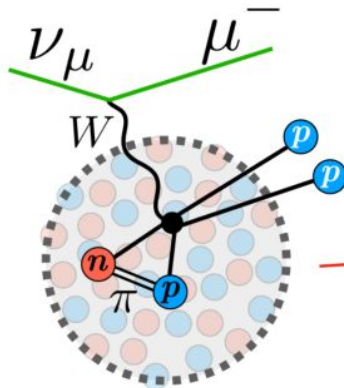
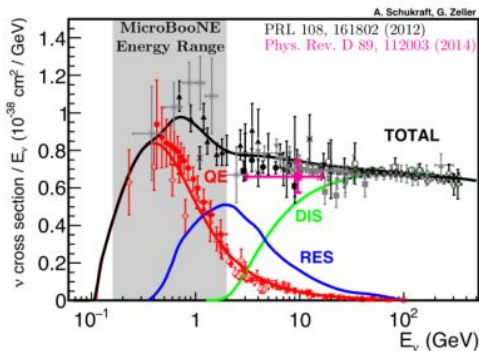


Strong flux correlations in  $\nu_\mu, \nu_e$



# Neutrino Interactions

## Interaction Modeling & Uncertainties



Credit A.Mastbaum

Searching for possible new physics in a regime with poor *a priori* constraints

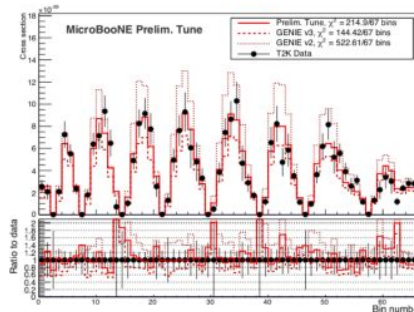
( $\nu$ -Ar interactions at 200 MeV)

Nuclear physics effects

e.g. interactions with correlated nucleon pairs ( $2p2h$ , or MEC)



- Latest **theory-driven models**
- New **tunes** including T2K CC0 $\pi$  data
- MICROBOONE-NOTE-1074-PUB
- Integrating Ar targets, ( $e, e'$ ) data

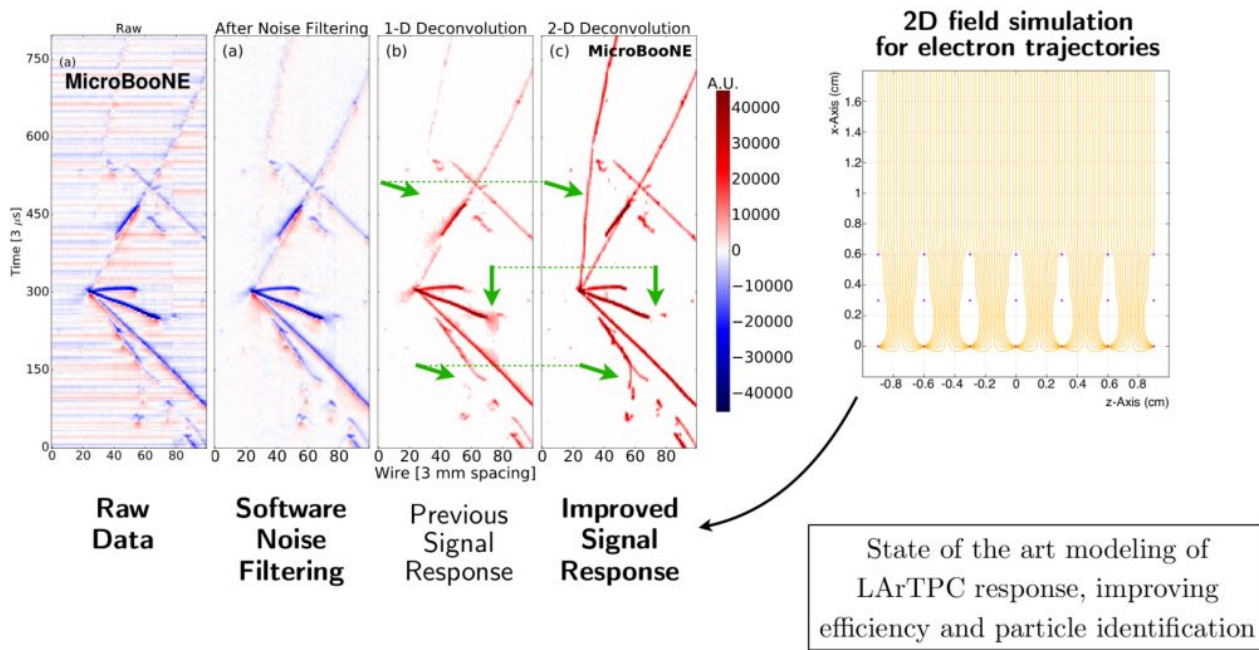


# Detector Modeling



Detailed modeling of **particle propagation**, **electron drift** & **detector response**, and **photon propagation**

## 1. Signal processing (MicroBooNE Collaboration, JINST 13 P07007, 2018 and JINST 13 P07007, 2018)



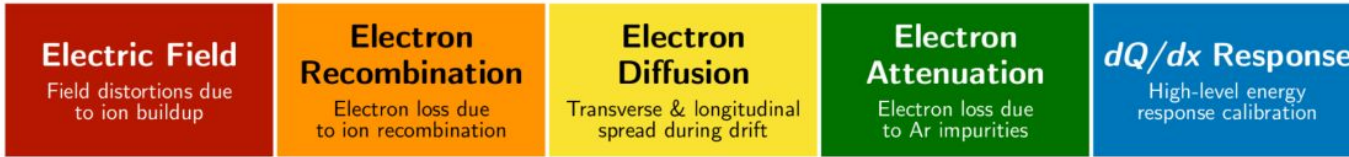
Credit A.Mastbaum

# Detector Modeling

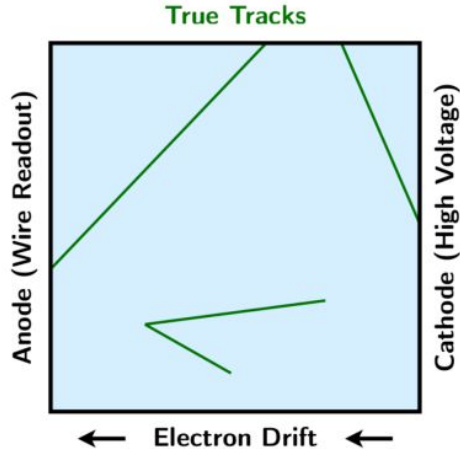


Detailed modeling of **particle propagation**, **electron drift & detector response**, and **photon propagation**

1. **Signal processing** (MicroBooNE Collaboration, JINST 13 P07007, 2018 and JINST 13 P07007, 2018)
2. **Response Calibration** or How Things Go Wrong



Credit A.Mastbaum



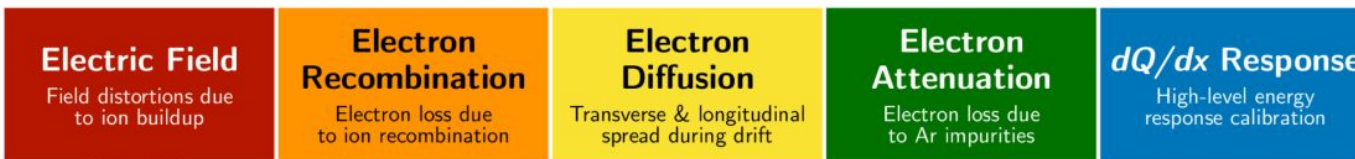
# Detector Modeling



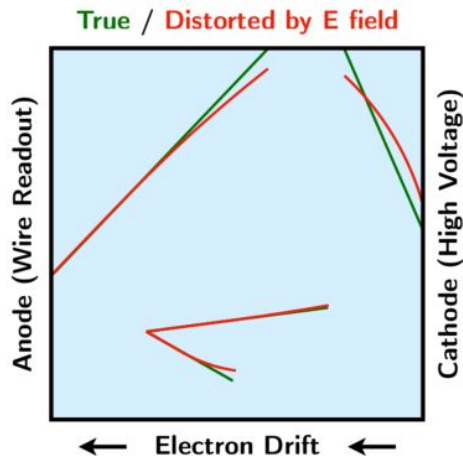
Detailed modeling of **particle propagation**, **electron drift & detector response**, and **photon propagation**

1. **Signal processing** (MicroBooNE Collaboration, JINST 13 P07007, 2018 and JINST 13 P07007, 2018)

2. **Response Calibration** or How Things Go Wrong



Credit A.Mastbaum



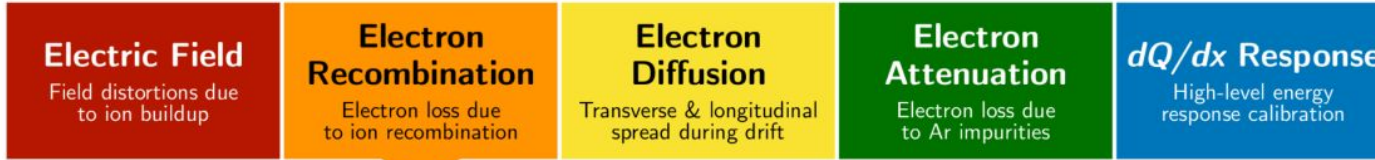


# Detector Modeling

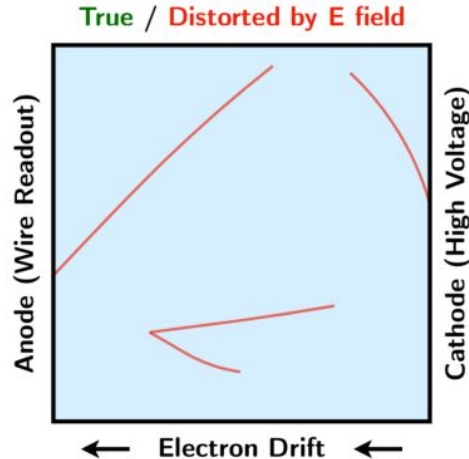


Detailed modeling of **particle propagation**, **electron drift** & **detector response**, and **photon propagation**

1. **Signal processing** (MicroBooNE Collaboration, JINST 13 P07007, 2018 and JINST 13 P07007, 2018)
2. **Response Calibration** or How Things Go Wrong



Credit A.Mastbaum

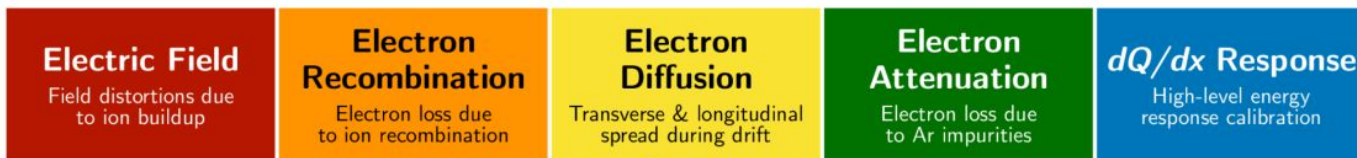


# Detector Modeling

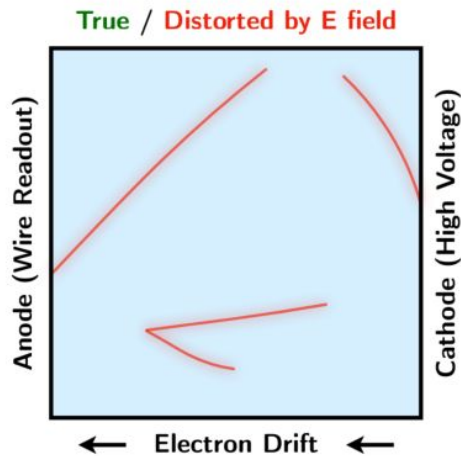


Detailed modeling of **particle propagation**, **electron drift** & **detector response**, and **photon propagation**

1. **Signal processing** (MicroBooNE Collaboration, JINST 13 P07007, 2018 and JINST 13 P07007, 2018)
2. **Response Calibration** or How Things Go Wrong



Credit A.Mastbaum

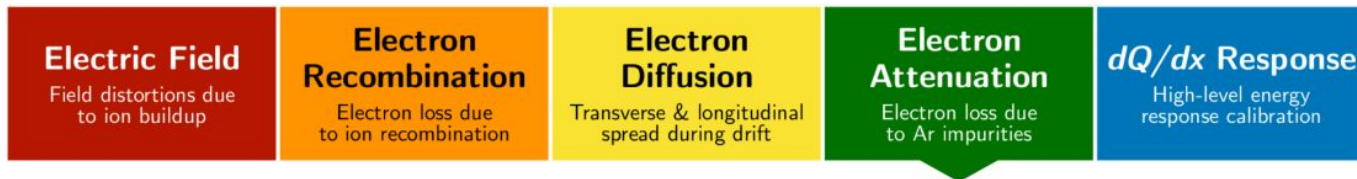


# Detector Modeling

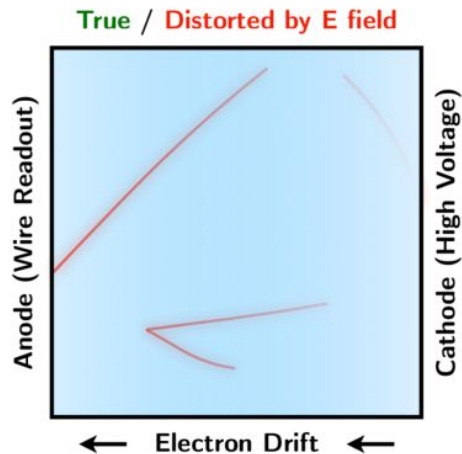


Detailed modeling of **particle propagation**, **electron drift** & **detector response**, and **photon propagation**

1. **Signal processing** (MicroBooNE Collaboration, JINST 13 P07007, 2018 and JINST 13 P07007, 2018)
2. **Response Calibration** *or* How Things Go Wrong



Credit A.Mastbaum



# Detector Modeling



Detailed modeling of **particle propagation**, **electron drift & detector response**, and **photon propagation**

1. **Signal processing** (MicroBooNE Collaboration, JINST 13 P07007, 2018 and JINST 13 P07007, 2018)
2. **Response Calibration** *or* How Things Go Wrong

## Electric Field

Field distortions due to ion buildup

## Electron Recombination

Electron loss due to ion recombination

## Electron Diffusion

Transverse & longitudinal spread during drift

## Electron Attenuation

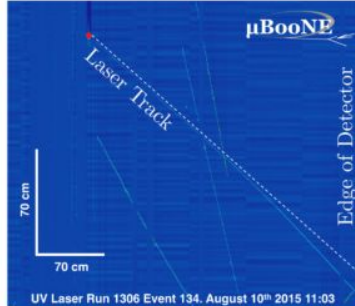
Electron loss due to Ar impurities

## $dQ/dx$ Response

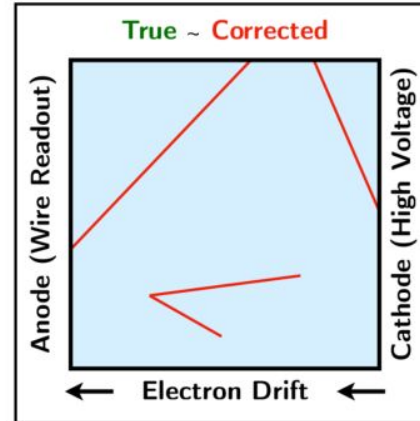
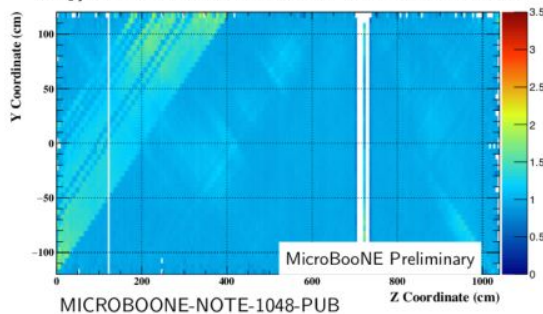
High-level energy response calibration

Credit A.Mastbaum

### In Situ Laser System



### $dQ/dx$ Calibration with Cosmic Muons



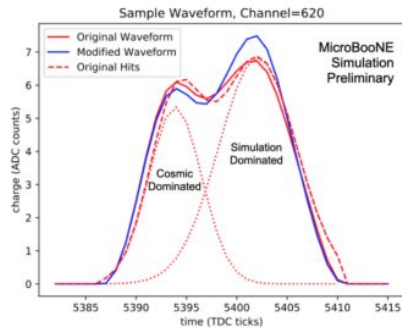
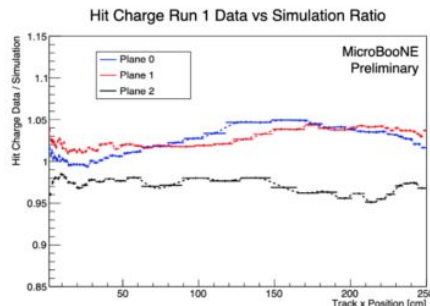
# Detector Modeling



Detailed modeling of **particle propagation**, **electron drift & detector response**, and **photon propagation**

1. **Signal processing** (MicroBooNE Collaboration, JINST 13 P07007, 2018 and JINST 13 P07007, 2018)
2. **Response Calibration** or How Things Go Wrong
3. **Systematic Uncertainties**

- Highest precision measurements yet performed with LArTPC technology
- Many subtle and correlated effects in the detector response model
- **First iteration:** Vary parameters according to *in situ* measurements and re-run full MC and analysis
- **Latest approach:** Capture waveform-level data/ MC differences in response as a function of  $x$ ,  $yz$ , angle, etc. as a correction and residual detector modeling systematic (*MICROBOONE-NOTE-1075-PUB*)



Credit A.Mastbaum

# Coverage of MCC8 Detector Variations



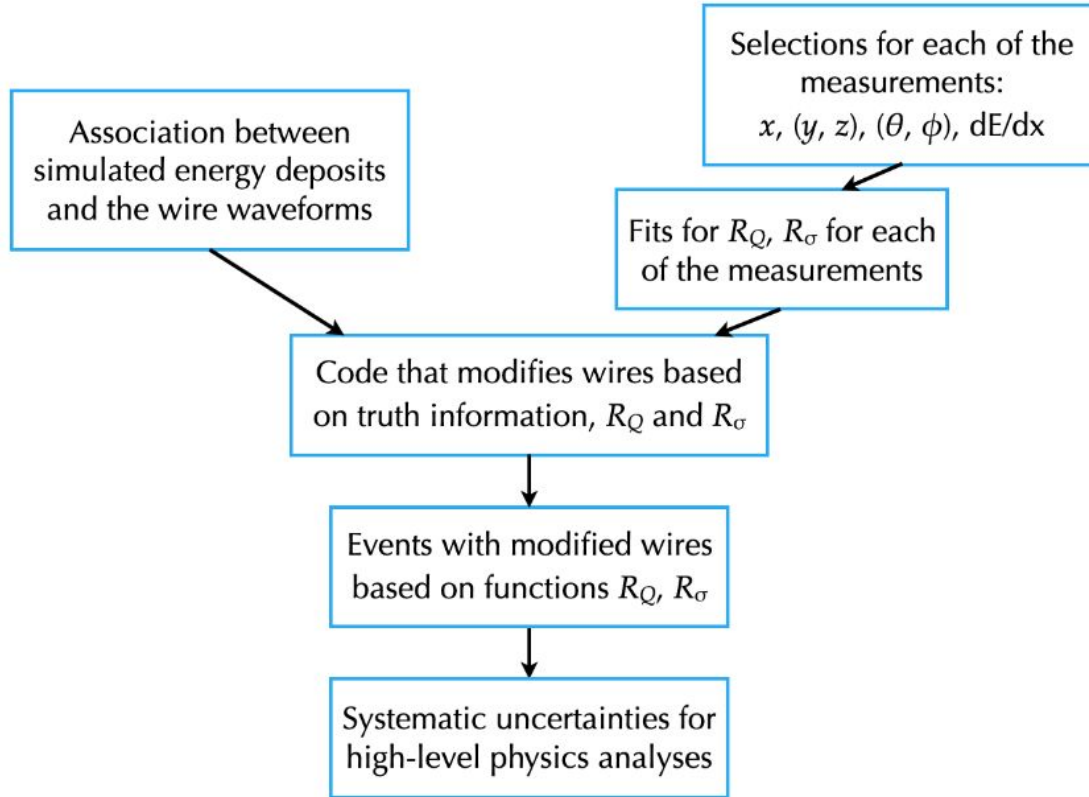
Want to make sure that MCC8 detector variations will be covered by the approach to simulation and detector systematics in MCC9

Some are covered by the move to cosmic data overlay for the MC

The rest are covered by this method, SCE variations, and light

Credit L.Yates

Covered by...	Detector Variations
x	Longitudinal & transverse diffusion, electron lifetime, SCE
(y, z)	Wire response, SCE
angles	Induced charge
dE/dx	Recombination
SCE variation	SCE (esp. spatial distortions)
light variations	Light model, light yield
overlay MC	Wire noise, PMT noise, saturated channels, dead channels



Credit L.Yates

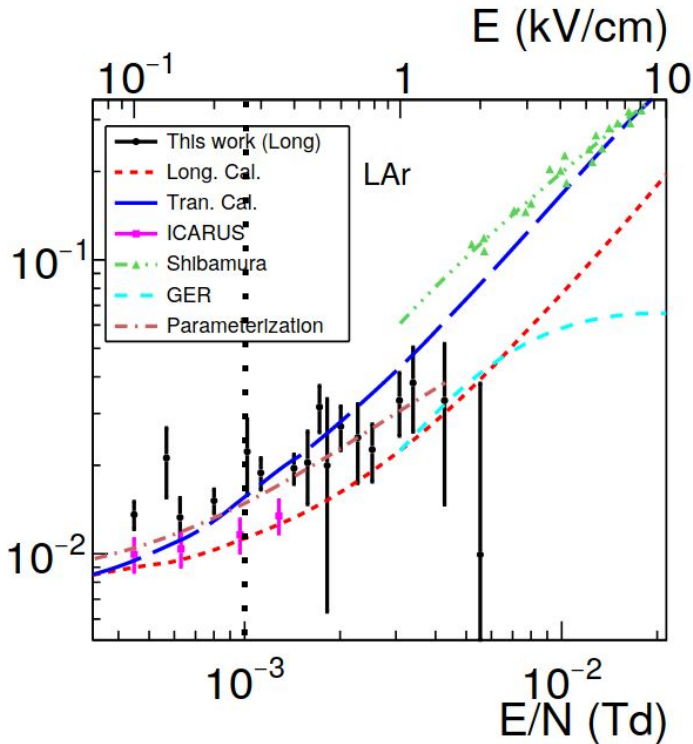
## MCC8 Detector Variations



# SCE

In the variation sample, a simple data-driven correction is applied to the central value simulation of these effects in order to better match what is measured in data. The central value space charge simulation includes spatial migration of charge and electron-ion recombination variations due to electric field distortions in the detector. The variation is defined by a set of corrections to the spatial migration and the electric field distortion. The spatial migration is altered by applying a scale factor,  $F(x)$ , which is dependent on the drift direction only, such that the spatial distortions observed at the top and bottom of the detector match between data and MC. This scale factor is applied to the distortions throughout the detector in MC. Additionally, the electric field distortions are reduced in the unisim to 70% of the nominal value in order to reflect the fact that space charge effects are less severe in data than they are in the nominal MC samples.

# Longitudinal Diffusion

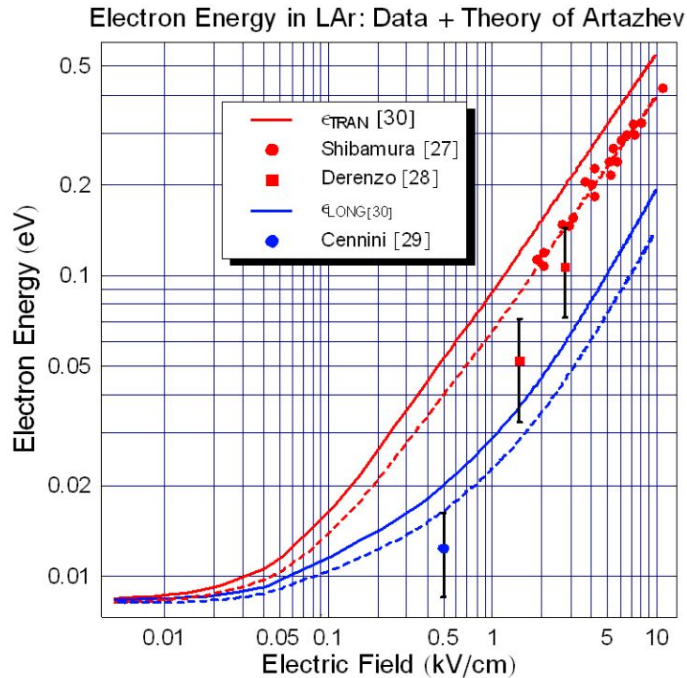


In the variation samples, the longitudinal diffusion constant is increased ( $DL_{up}$ ) or decreased ( $DL_{dn}$ ) relative to the MCC8 central value. We chose these variations by to span the range of previous measurements of longitudinal diffusion in liquid argon, using Fig. 11 (plot on the right side) in [this reference](#).<sup>3</sup> We looked at the data bands at the MicroBooNE field strength, 0.273 kV/cm. We used the upper extent of those bands for the  $DL_{up}$  variation and the lower extent of those bands for the  $DL_{dn}$  variation. The MCC8 central value for this constant is between these values.

# Transverse Diffusion

These variations were defined by extrapolation from the measurements shown on p. 9 in [DocDB 412](#). Unlike in the case of longitudinal diffusion, the available world data points do not cover the MicroBooNE field strength. Instead, we looked at a theoretical extrapolations (red curves on p. 9 of previously referenced document).

For the DTdn variation, we used the transverse diffusion constant coming from the data-scaled theory (red dashed curve on that plot) at 0.273 kV/cm. For the DTup variation, we used the constant coming from the unscaled theory (red solid curve) at 0.273 kV/cm; this was a relatively arbitrary choice intended to give a reasonable value above our best guess of the CV. The central value MCC8 uses a transverse diffusion constant corresponding to a nominal MicroBooNE electric field magnitude of 0.5 kV/cm, not the actual field of 0.273 kV/cm. This means that both the DTdn and DTup variations are actually below the CV.

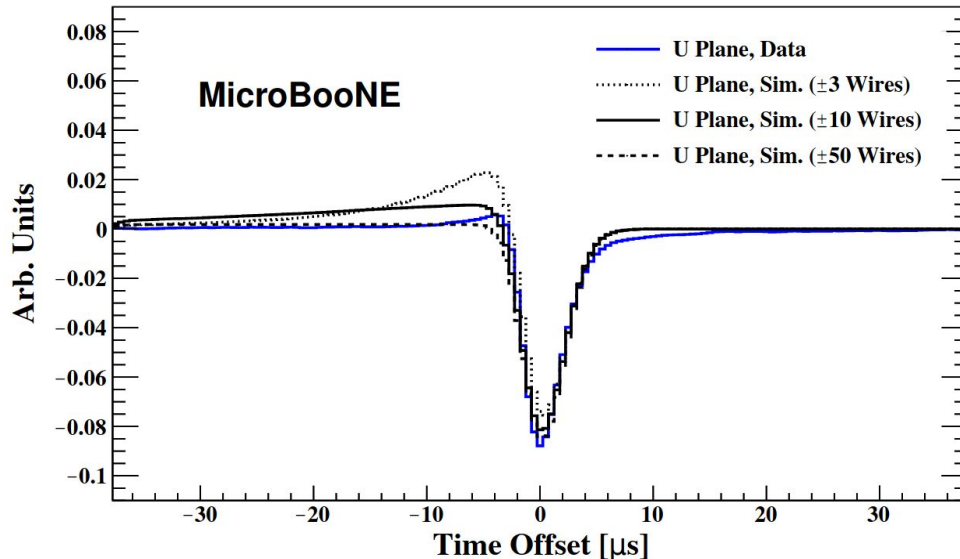


# Dynamically Induced Charge

**Description of variation:** In central value MCC8, induction is simulated only on the wire closest to the drifting charge.

In this variation, the central value MCC8 model for charge induction is replaced by a model that includes longer-range charge induction. This model simulations induction on the closest wire, plus the 10 neighboring wires on each side. The new model is more realistic than the central value MCC8 model, but it is still imperfect in, for example, lacking smooth inter-wire treatment. An improved model developed by the WireCell team will be included in MCC9.

**Data/Sim. Response Comparison: U Plane, Normal Region**



[JINST 13, P07007 \(2018\)](#)

# Wire Response Function

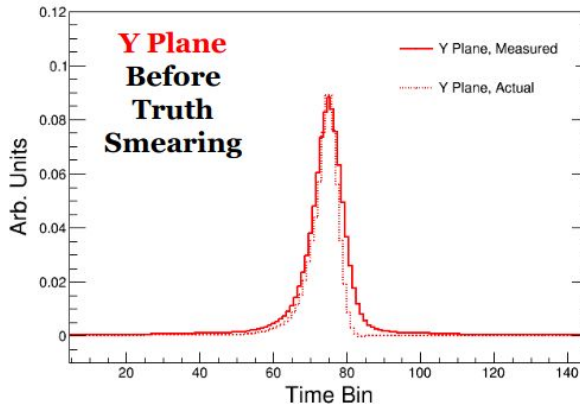
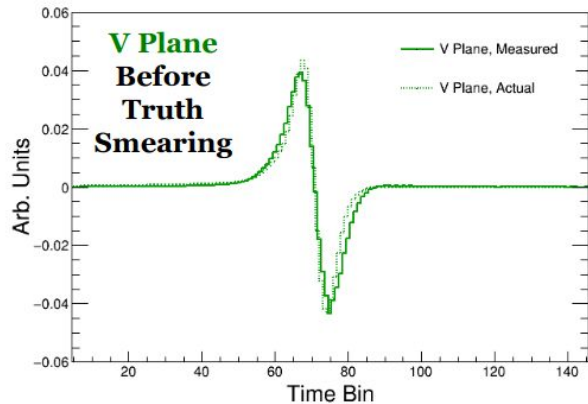
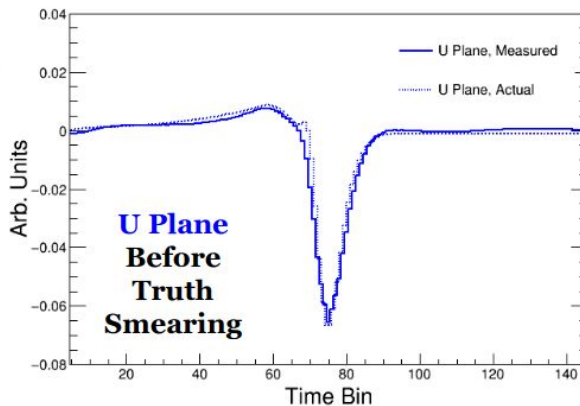
**Description of variation:** Charge induction on the TPC wires is simulated using a set of response functions, bipolar on the U and V planes and unipolar on the Y plane. The same functions are later used to deconvolve the wire signals. The derivation of these functions is described in full in the [MicroBooNE Signal Processing II paper](#).<sup>5</sup>

In this variation, the wire response function of each plane is squeezed by 20%. This represents the approximate uncertainty on the functions, as shown on slides 23–24 of [DocDB 9419](#), according to detector expert Mike Mooney. These squeezed functions are used to simulate the charge induction; the central value functions are still used to deconvolute the signals. We squeezed each plane's response functions for the nominal, shorted-U, and shorted-Y regions of the detector and properly time-aligned all of the squeezed functions.

Credit M.Mooney

# Wire Response Function

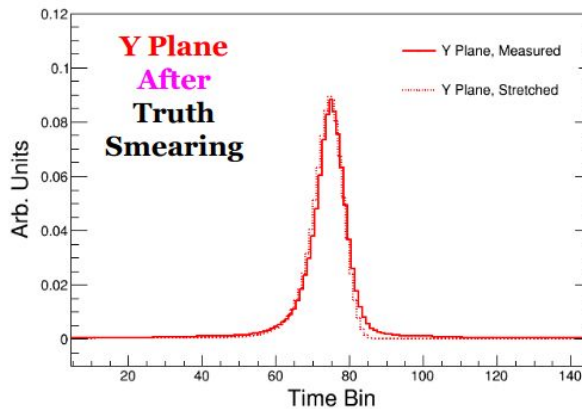
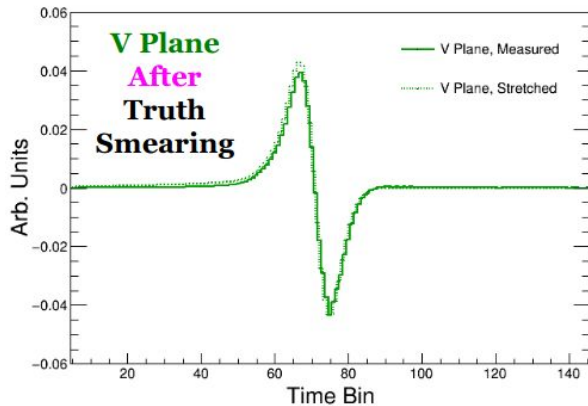
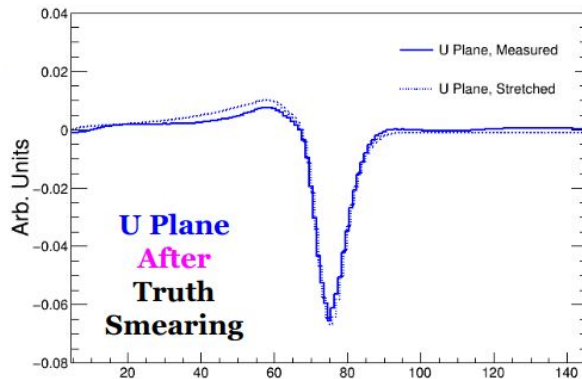
- ◆ Check accuracy/precision of technique by comparing response estimated from MC to the true simulated response
  - Using the old MCC7 (“v0.1”) responses for this check
- ◆ Find:  $\sim 1 \mu\text{s}$  residual smearing



Credit M.Mooney

# Wire Response Function

- ◆ Check accuracy/precision of technique by comparing response estimated from MC to the true simulated response
  - Using the old MCC7 (“vo.1”) responses for this check
- ◆ Find:  $\sim 1 \mu\text{s}$  residual smearing



Credit M.Mooney

# Light Out Of TPC

**Description of variation:** Light production outside the TPC may be imprecisely modeled. In this variation, we increase the light yield outside the TPC by 50%.

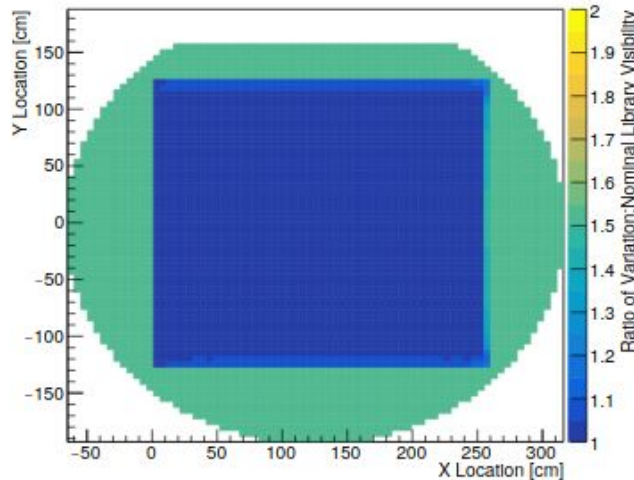
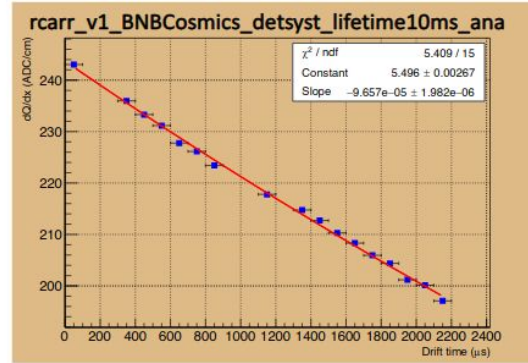
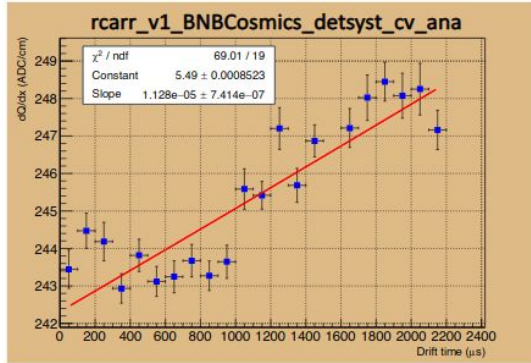


Figure 4: The ratio of systematically varied photon visibility library over our nominal photon visibility library projected onto the  $xy$  plane. The core of the TPC is fixed at 1 while the outer volume shifts to 1.5. The blank region at the top of the cryostat is the ullage where no scintillation light is simulated. Plot from Joseph Zennamo.



# Electron Lifetime

**Description of variation:** The central value MCC8 models the electron lifetime in liquid argon as infinite. In this variation, we reduce the lifetime to 10 ms, representative of the electron lifetime in the lowest-purity runs included in the good run list.



Electron lifetime : -88 619 +/-5.822  
QA/QC = 1.025 +/- 0.003 (amount charge left after @ 2.2 ms/amount of initial charge @ 0 ms)

Electron lifetime : 10.355 +/-0.212  
QA/QC = 0.809 +/- 0.006 (amount charge left after @ 2.2 ms/amount of initial charge @ 0 ms)

Unphysical e lifetime so QA/QC is more ideal

Figure 3: Left, from central value MC:  $dQ/dx$  of anode- or cathode-piercing tracks, as a function of drift distance. An estimate of the electron lifetime can be extracted from this plot, following the method in [DocDB 6318](#). The estimate for this sample comes out to be an unphysical value, indicating that the lifetime is effectively infinite. Right, from the sample with an electron lifetime of 10 ms: The same type of plot, showing that for this sample, the extracted lifetime is the expected 10 ms. Plots from Varuna Meddage.

# Recombination

**Description of variation:** The central value MC models electron-ion recombination using the modified box model with parameters fit to ArgoNeuT data. This variation substitutes the Birks model with parameters tuned to ICARUS data.

# Future Prospects

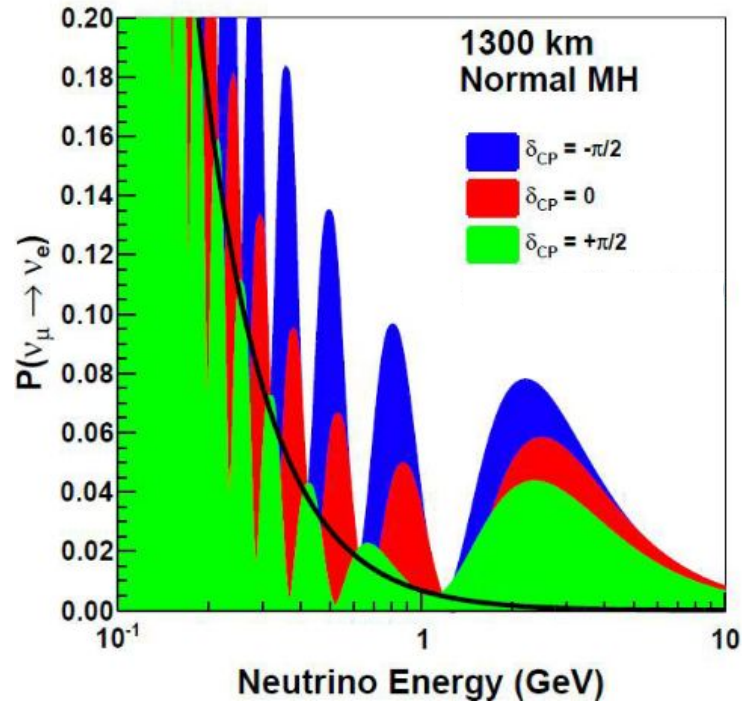
- Higher statistics
- New simulation
- New signal processing
- Completely new detector uncertainties
- 3-plane PID
- Working on various derived variables

A Feynman diagram showing an electron-positron annihilation process. Two lines, one solid and one dashed, meet at a vertex. From this vertex, a wavy line representing a photon extends downwards. Below the vertex, the text  $e^+e^-$  is written in a stylized, handwritten font.

# Future Experiments

Long-Baseline Neutrino Facility (LBNF) and  
Deep Underground Neutrino Experiment (DUNE)

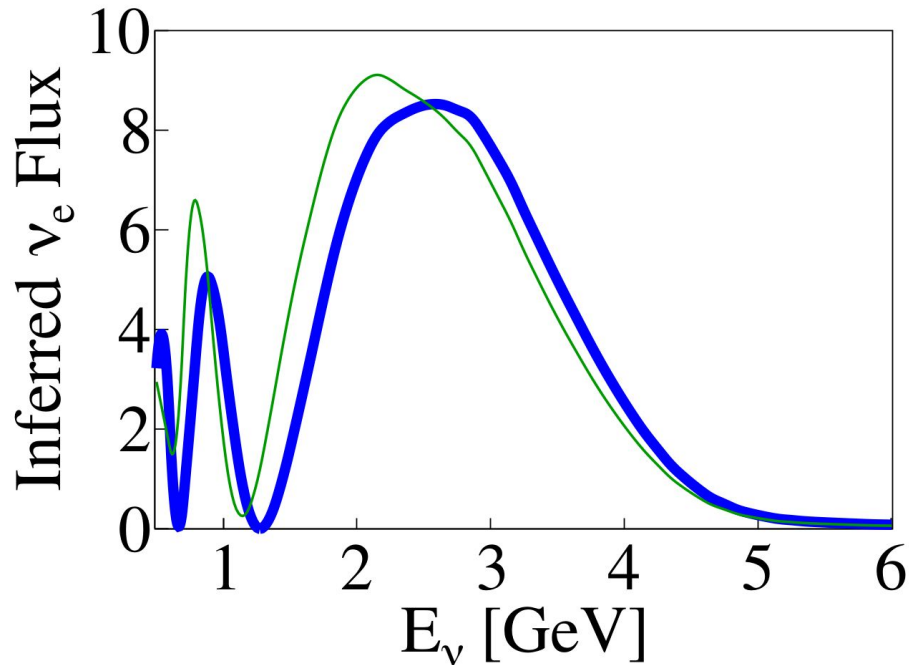
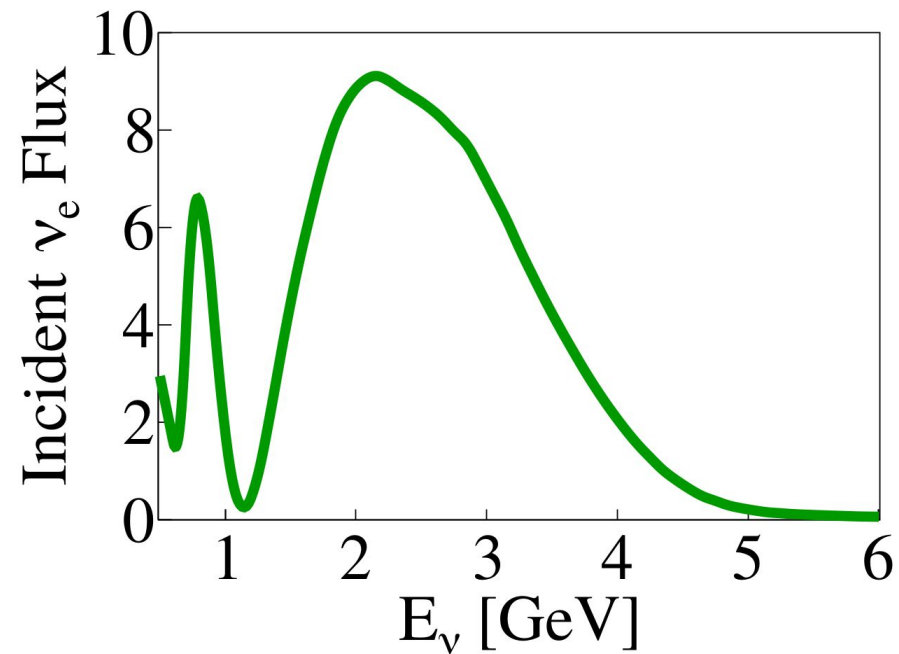
Conceptual Design Report



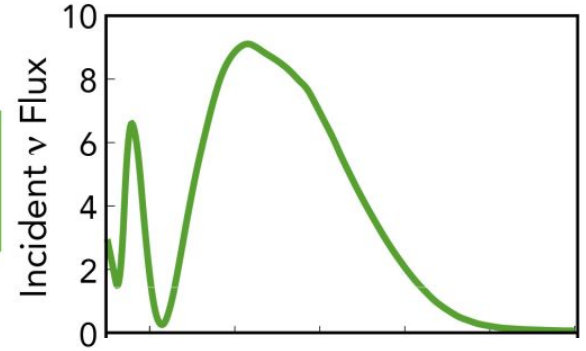
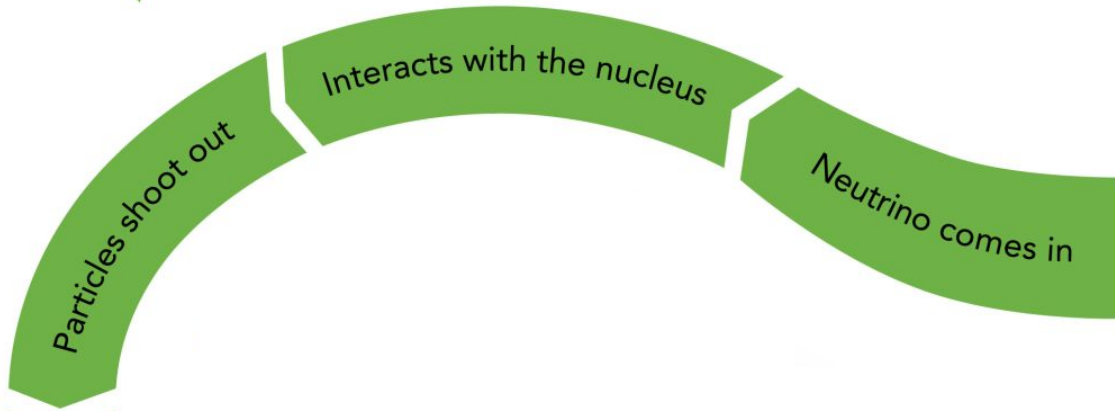
“Uncertainties on the oscillation analysis indicate that uncertainties exceeding 1% for signal and 5% for backgrounds may result in substantial degradation of the sensitivity to CP violation and mass hierarchy”



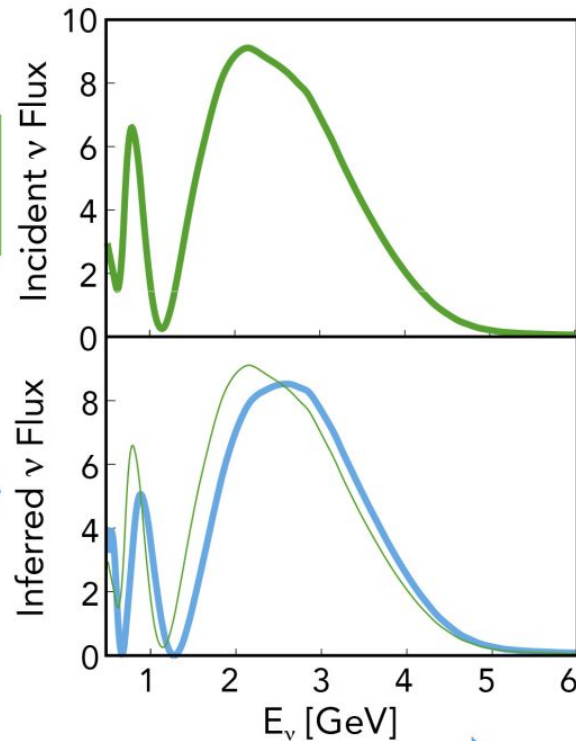
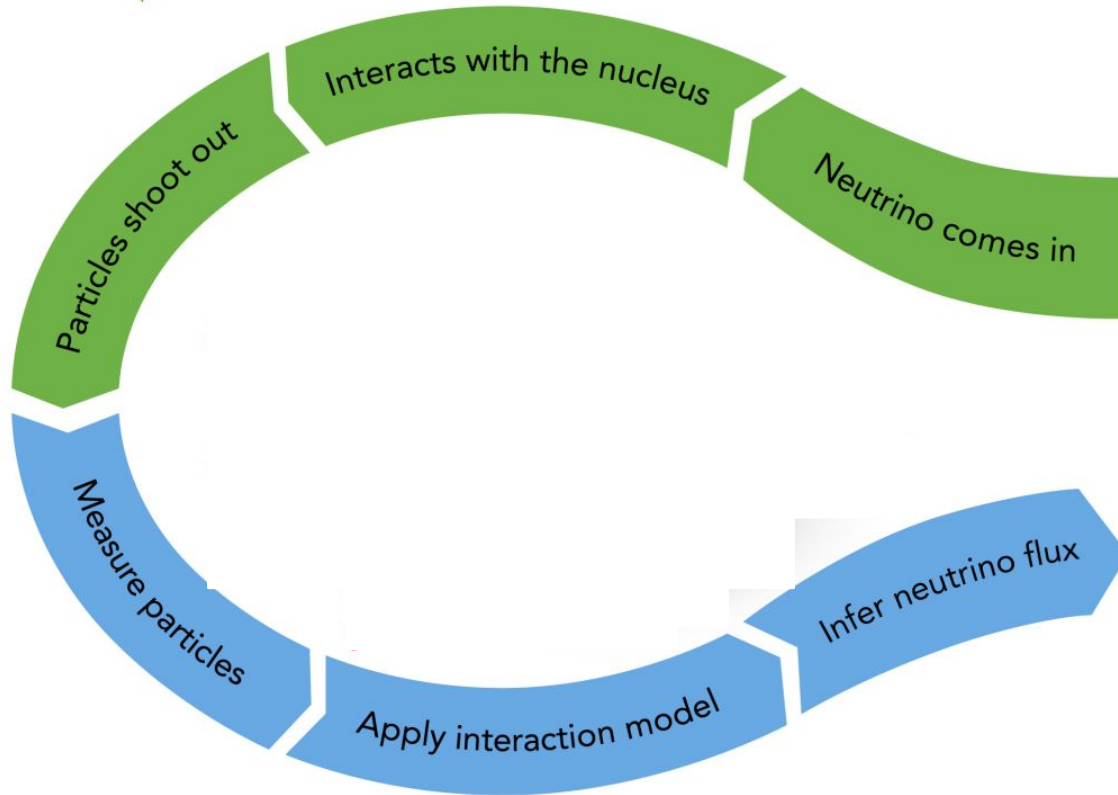
# Incident & Inferred Fluxes



# PHYSICS PROCESS



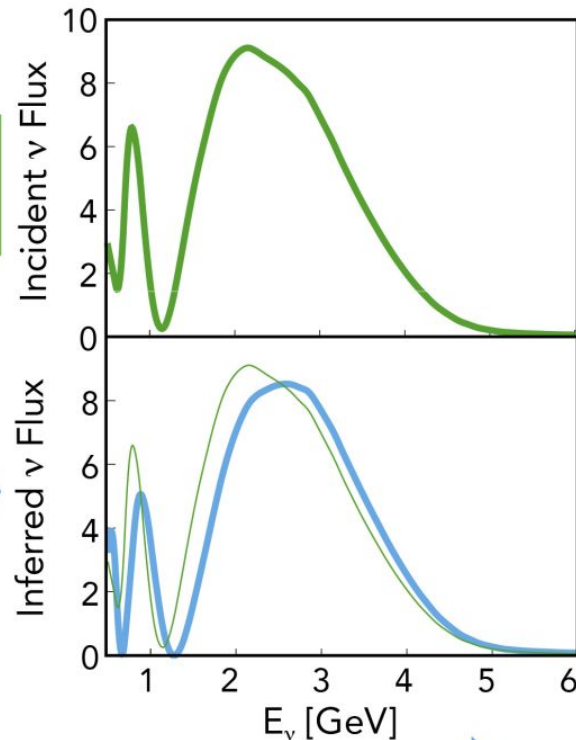
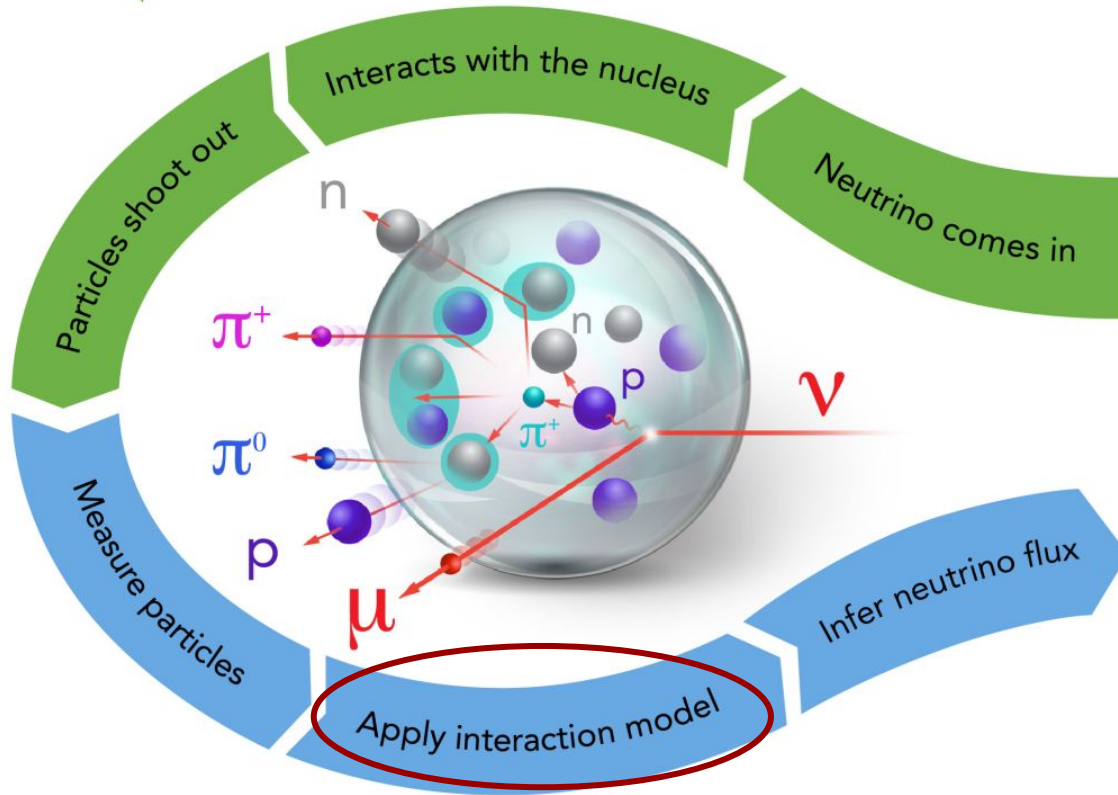
# PHYSICS PROCESS



# EXPERIMENTAL ANALYSIS



# PHYSICS PROCESS



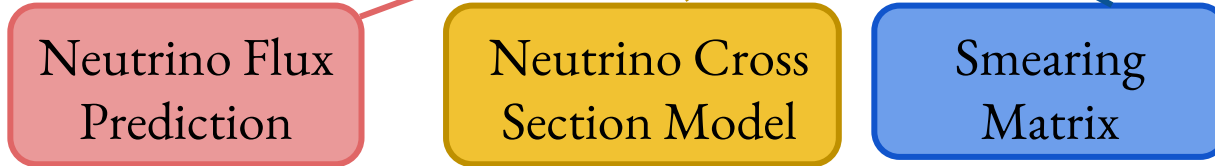
Apply interaction model

Infer neutrino flux

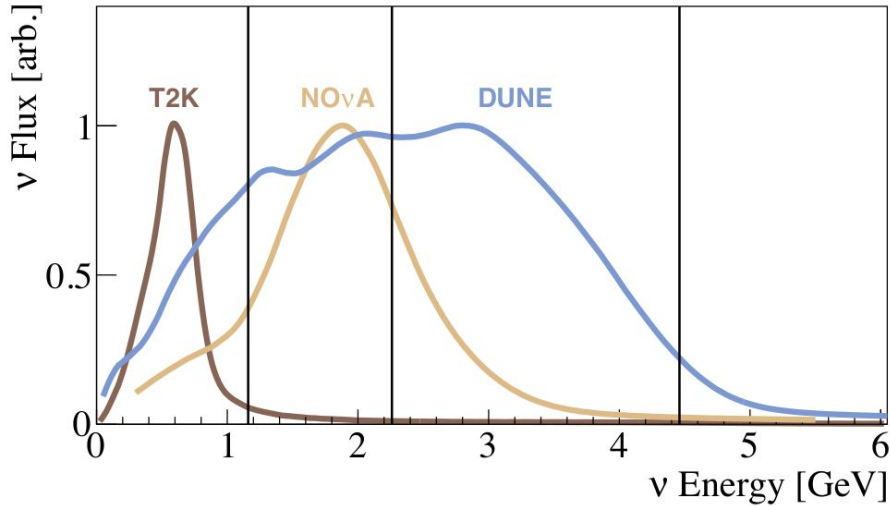
# EXPERIMENTAL ANALYSIS

# Event Rate

$$N(E_{\text{rec}}, L) \sim \sum_i \int \Phi(E_\nu, L) \sigma_i(E_\nu) f_{\sigma_i}(E_{\text{rec}}, E_\nu) dE$$

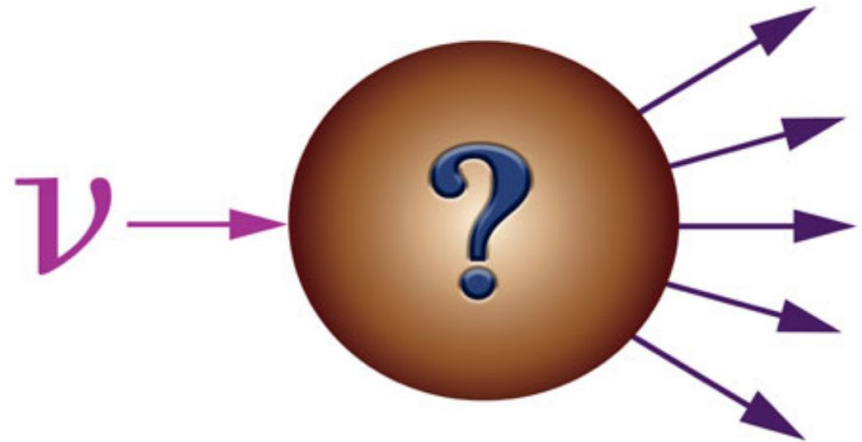


# Wide Energy Spectra

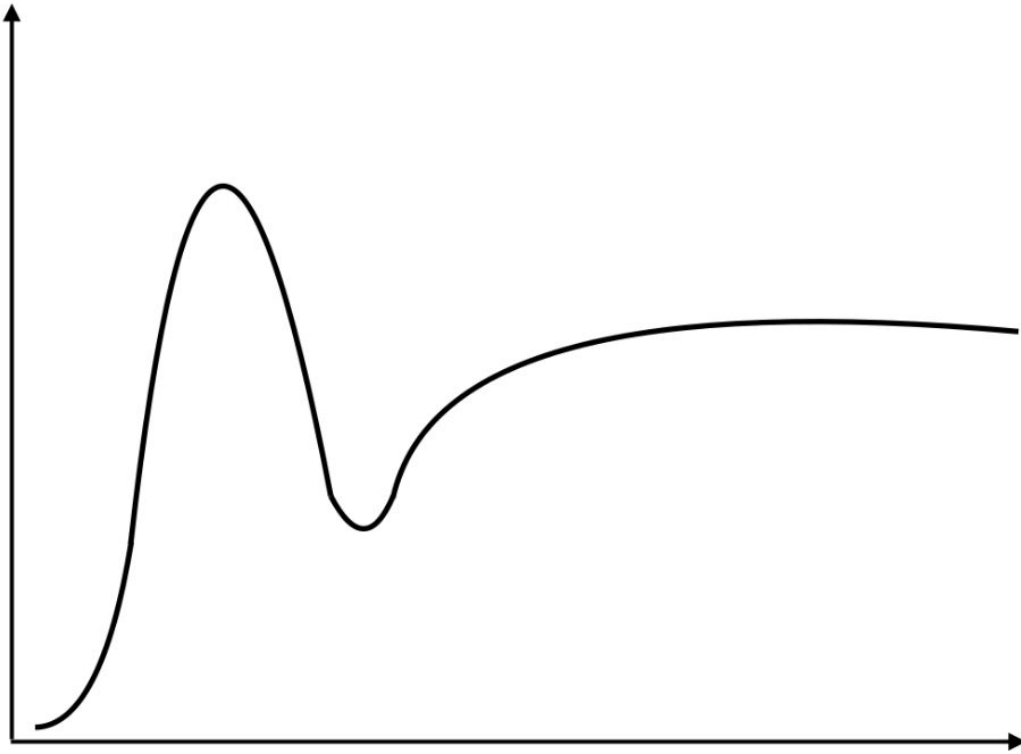


- Very challenging due to nuclear effects & final state interactions

- Use final state particles to reconstruct  $\nu$  energy & direction

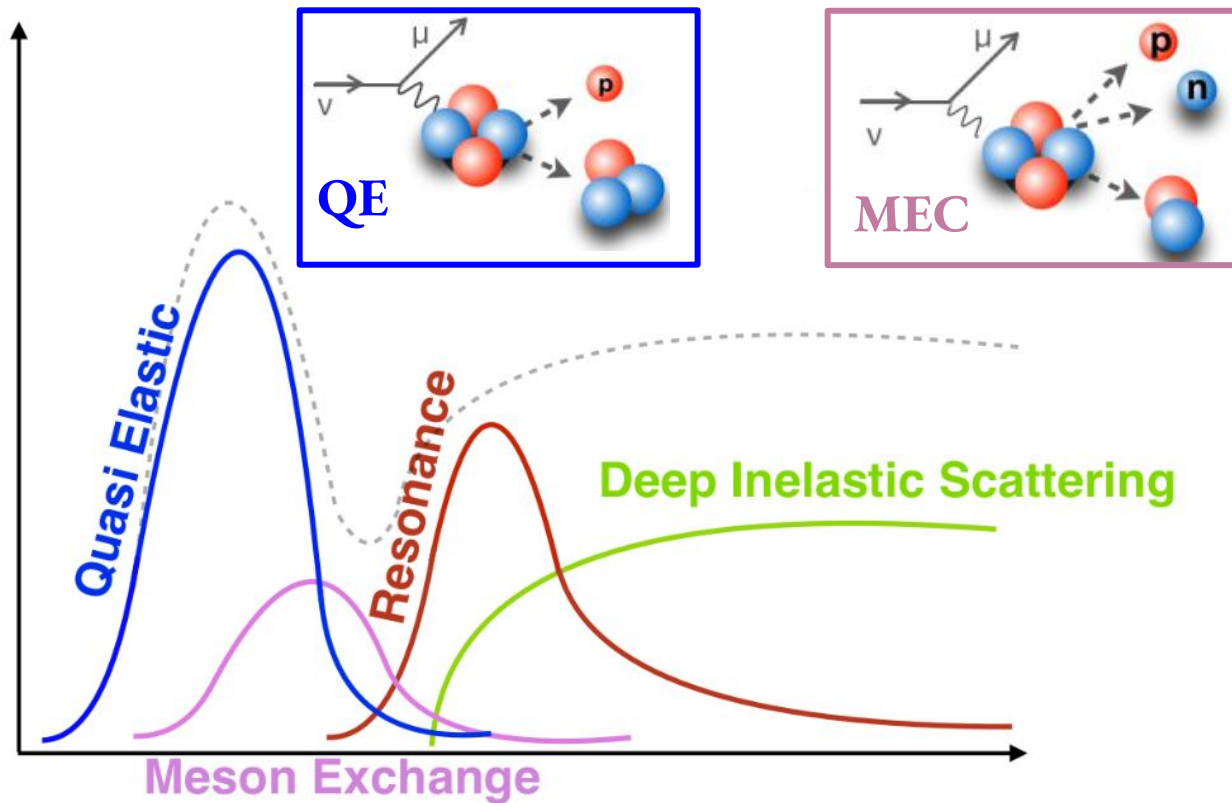


# Event Generators & Interaction Breakdown

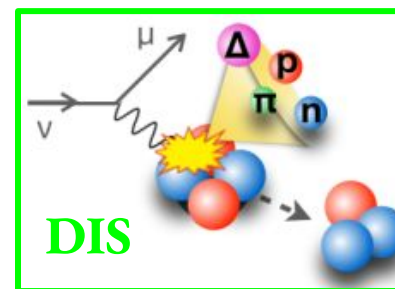
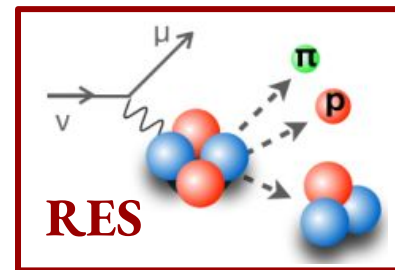


Energy Transfer =  $E_1 - E_p$

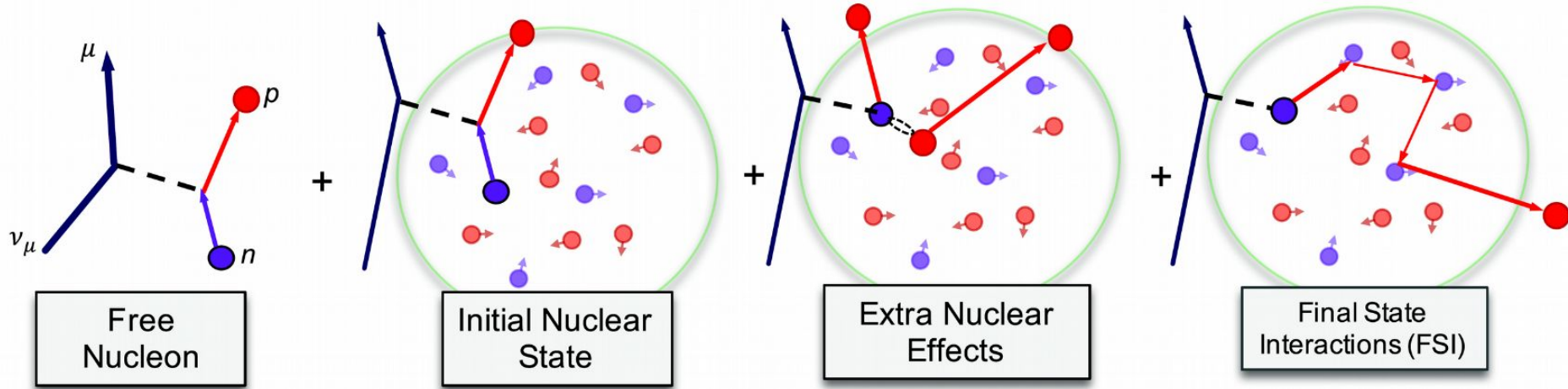
# Event Generators & Interaction Breakdown



$$\text{Energy Transfer} = E_1 - E_f$$



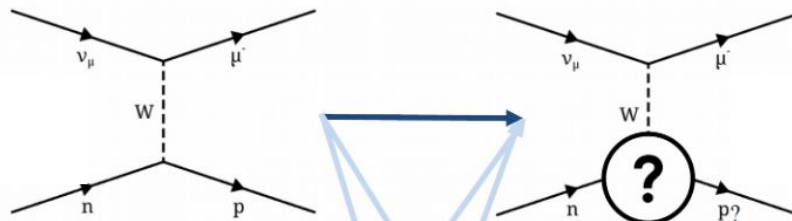
# Nature & Topological Breakdown



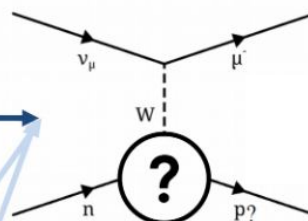
Final state particle content does not isolate initial interaction type!

# Topological Breakdown

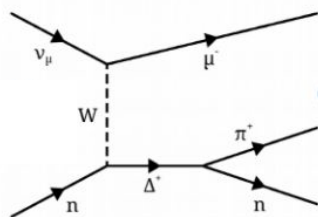
CCQE



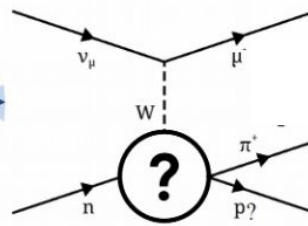
$CC0\pi$   
(CCQE-like)



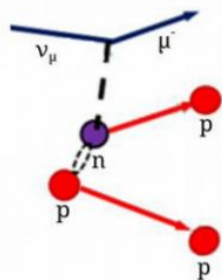
CCRES



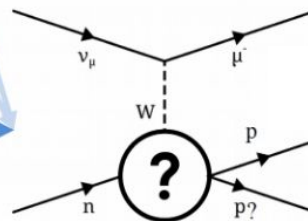
$CC1\pi$   
(CCRES-like)



2p2h

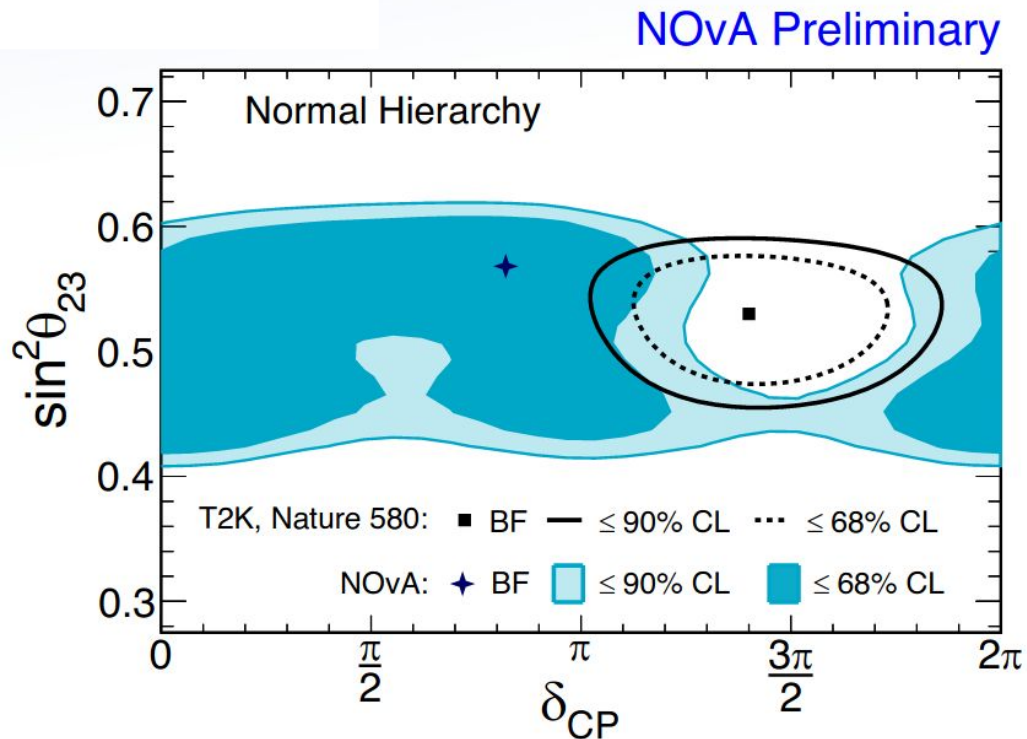
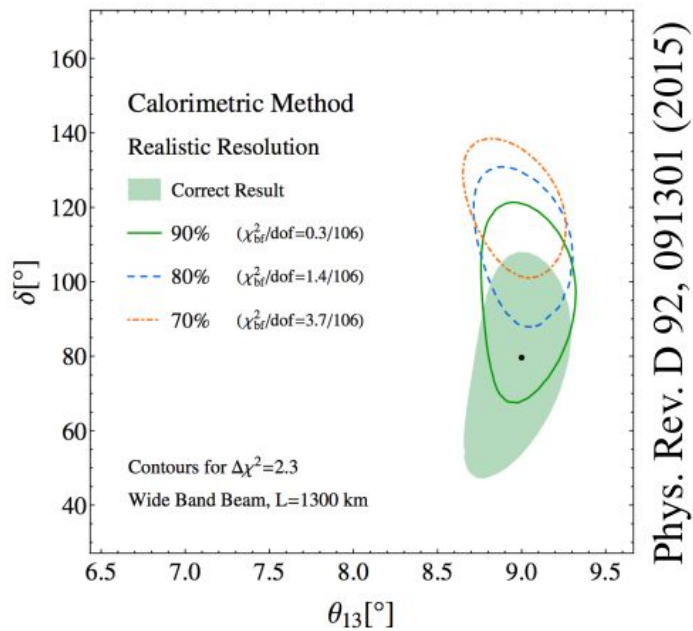


$CC0\pi+Np$   
( $N > 0$ )



# Miss-modelling might impact mixing parameters

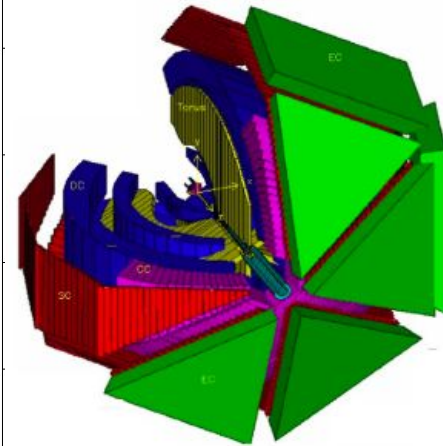
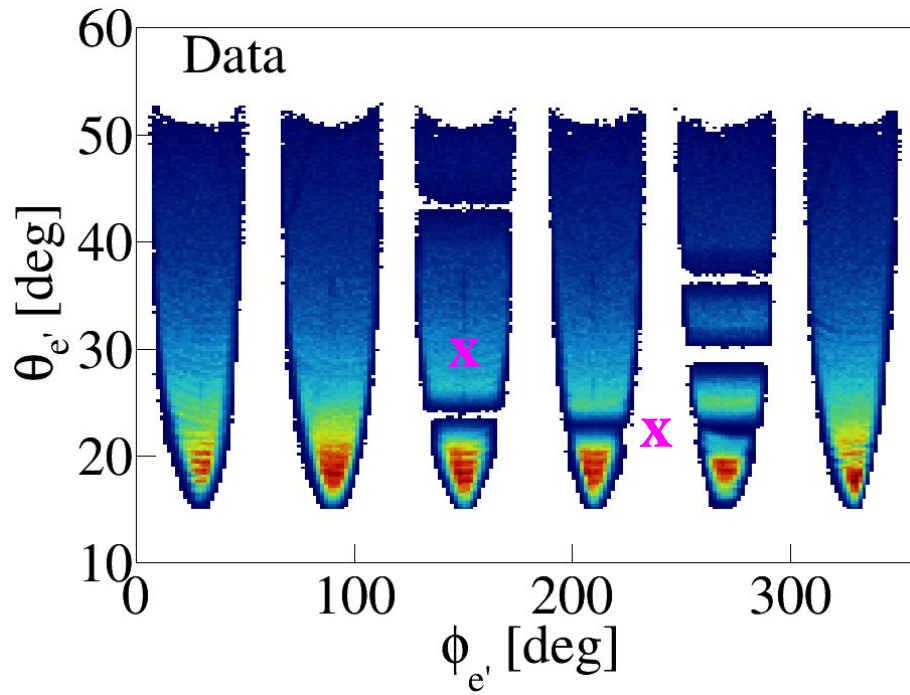
Extract Oscillations





# Background Subtraction

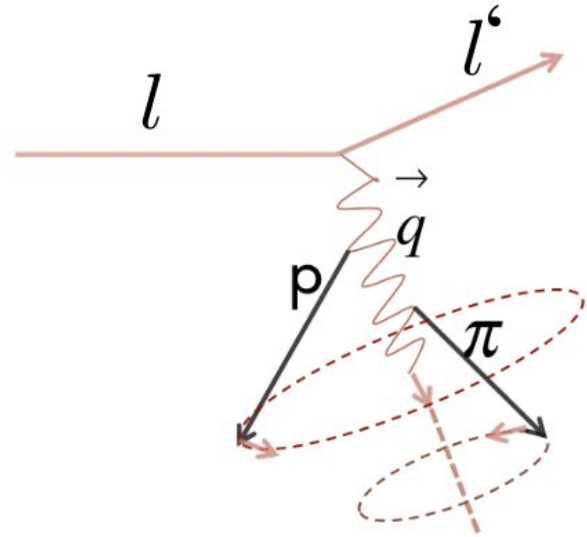
Non-QE interactions lead to multi-hadron final states  
Gaps make them look like  $(e,e'p)$  events



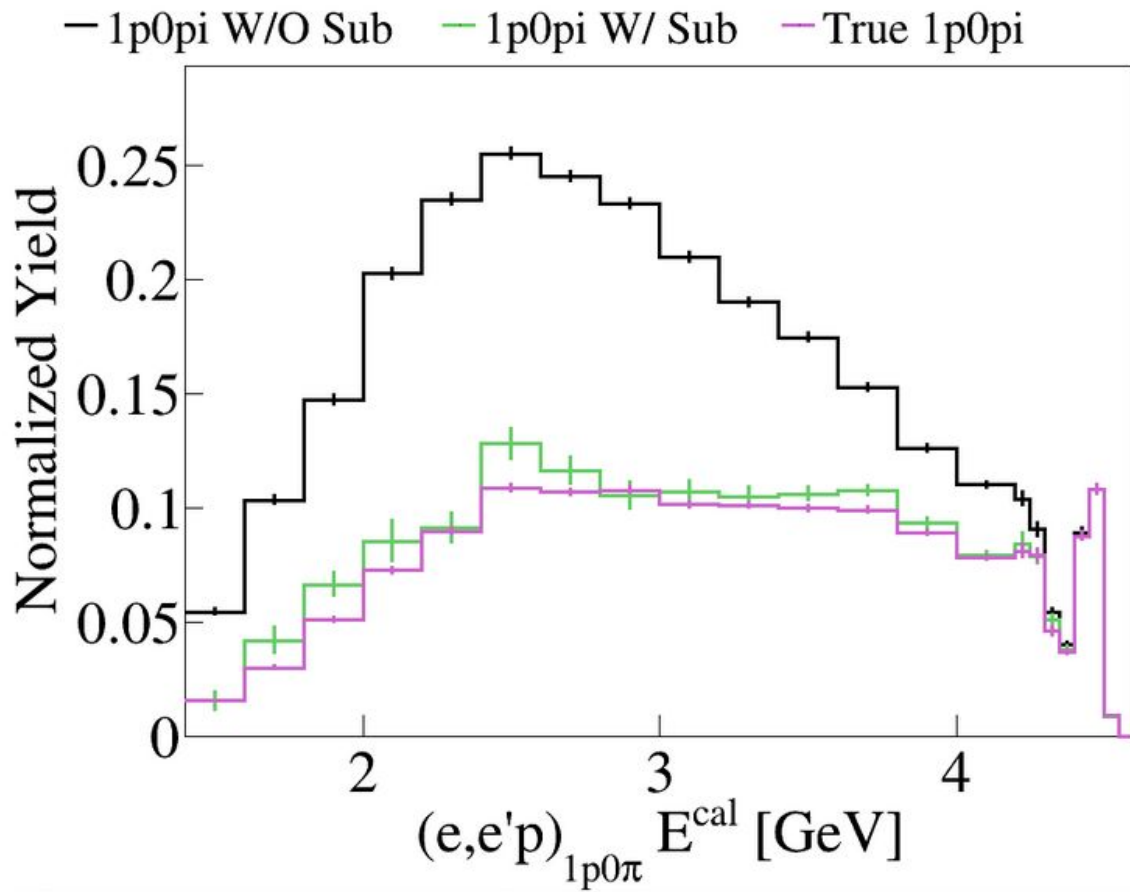
# Data Driven Correction

Non-QE interactions lead to multi-hadron final states  
Gaps make them look like  $(e,e'p)$  events

- Use measured  $(e,e'p\pi)$  events
- Rotate  $p, \pi$  around  $q$  to determine  $\pi$  detection efficiency
- Subtract undetected  $(e,e'p\pi)$
- Repeat for higher hadron multiplicities



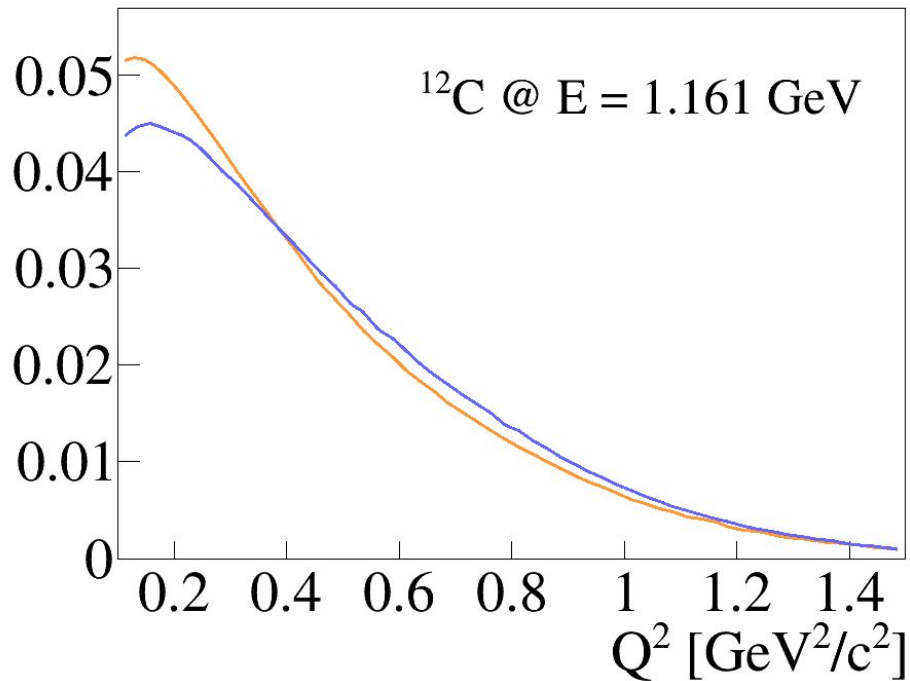
# Closure Test



# e & $\nu$ Similarities (Inclusive Channel)

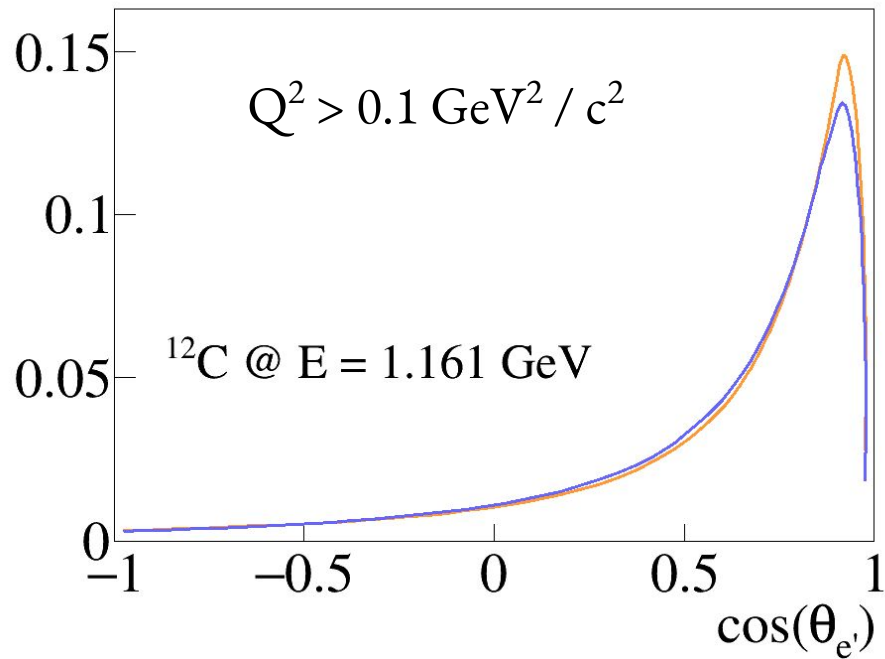
—  $\nu$  SuSav2

— e SuSav2



—  $\nu$  SuSav2

— e SuSav2

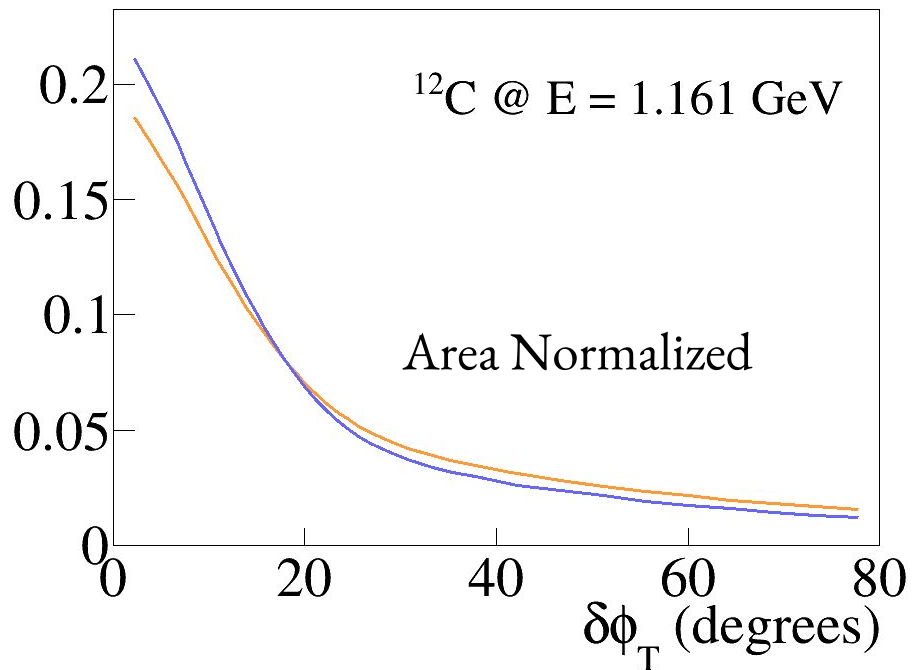


Area Normalized

# e & $\nu$ Similarities (Exclusive Channel)

—  $\nu$  SuSav2

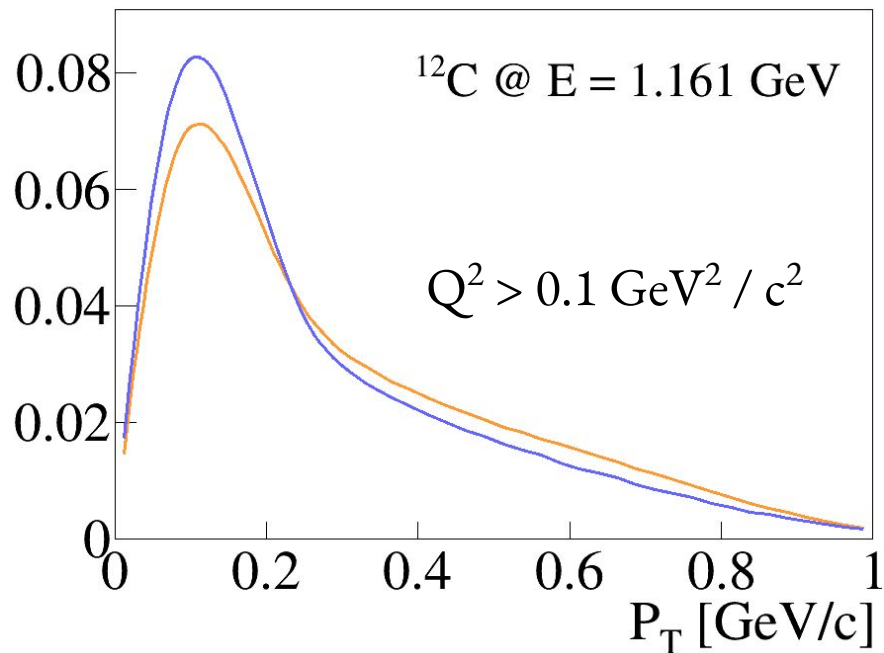
— e SuSav2



- 1 proton  $> 300$  MeV / c
- 0  $\pi^{+/-} > 70$  MeV / c

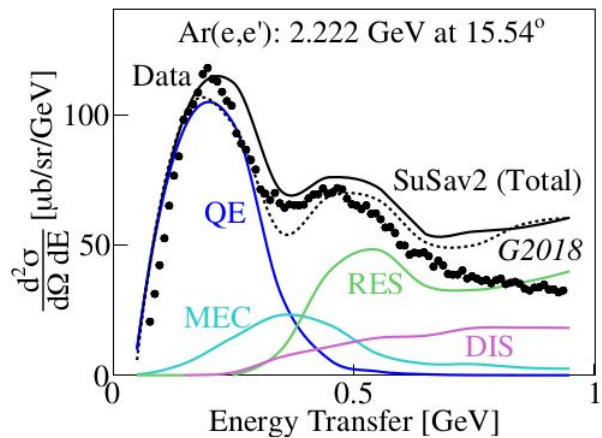
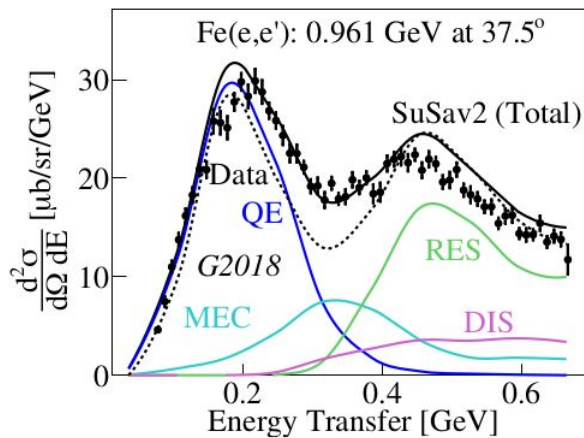
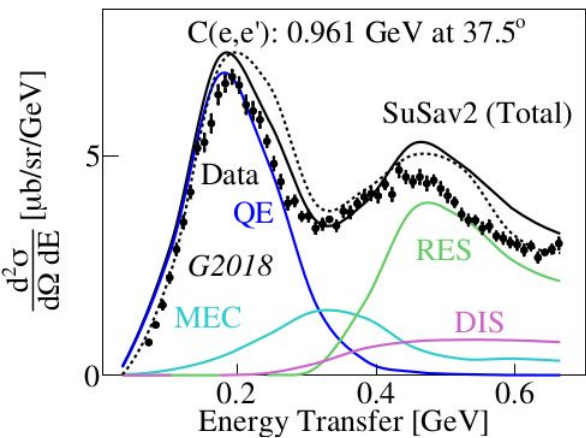
—  $\nu$  SuSav2

— e SuSav2



- 0  $\pi^0$  or  $\gamma$
- any number of n

# GENIE vs Inclusive Electron Data



G2018 = G18\_10a\_02\_11a  
(default MicroBooNE GENIE tune)

SuSav2 = Superscaling model

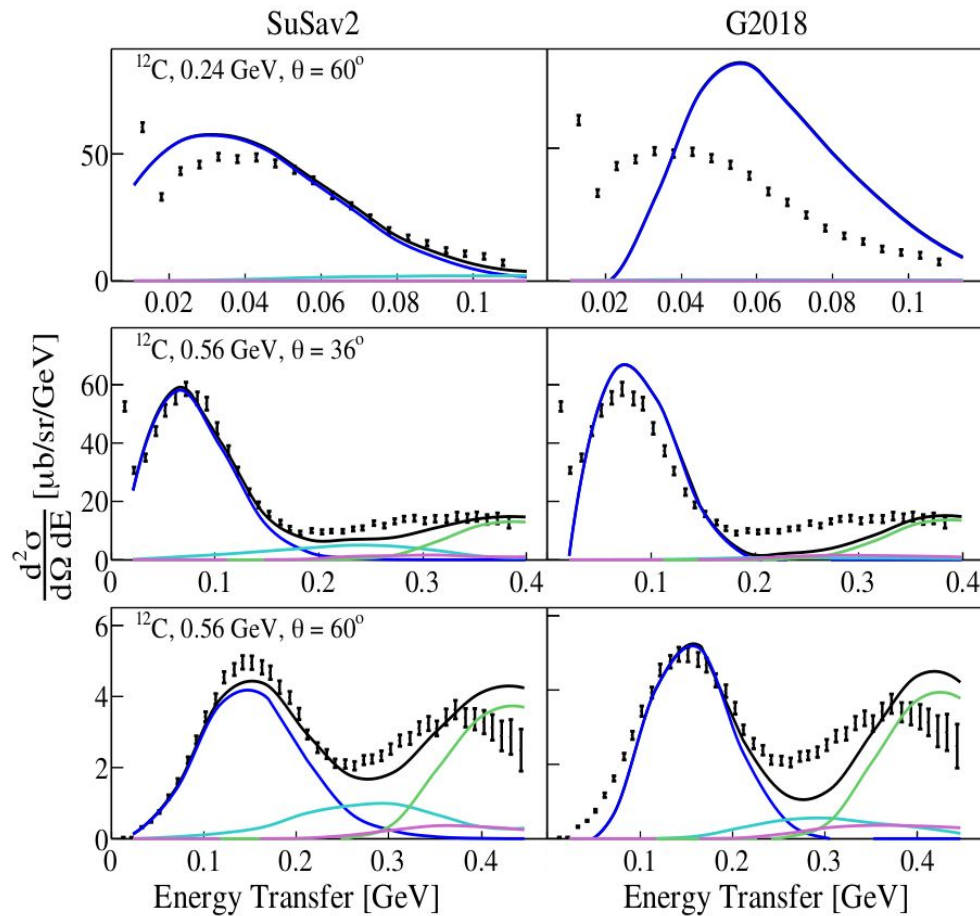
[Phys. Rev. D 94, 013012 \(2016\)](#)

(First time that QE & MEC events use the same model for e &  $\nu$  scattering in GENIE)

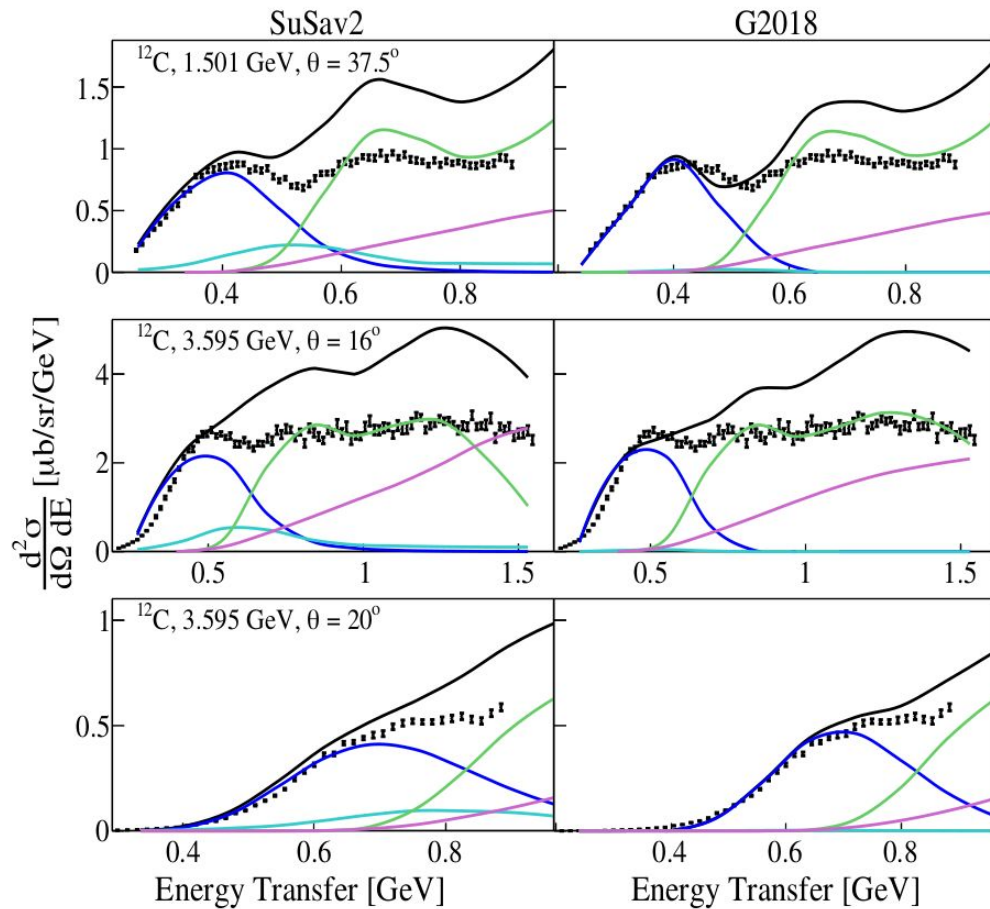
More inclusive comparisons [arXiv:2009.07228 \[nucl-th\]](#)



# GENIE vs Inclusive Electron Data



# GENIE vs Inclusive Electron Data

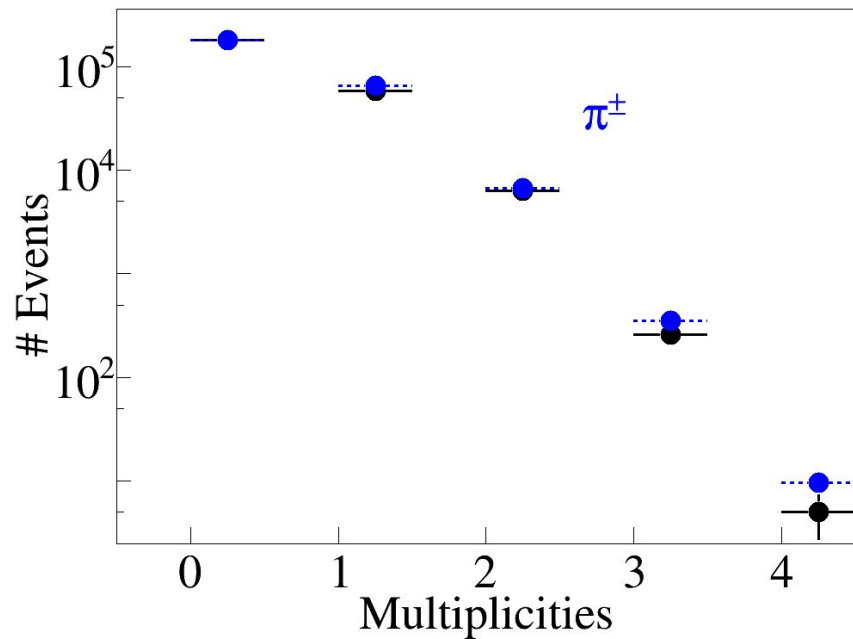
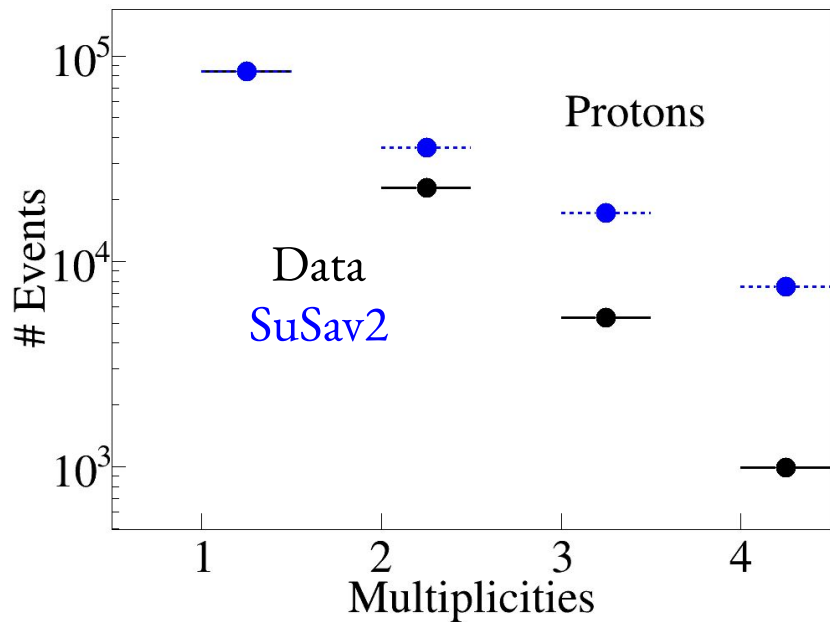




# Systematic Uncertainty

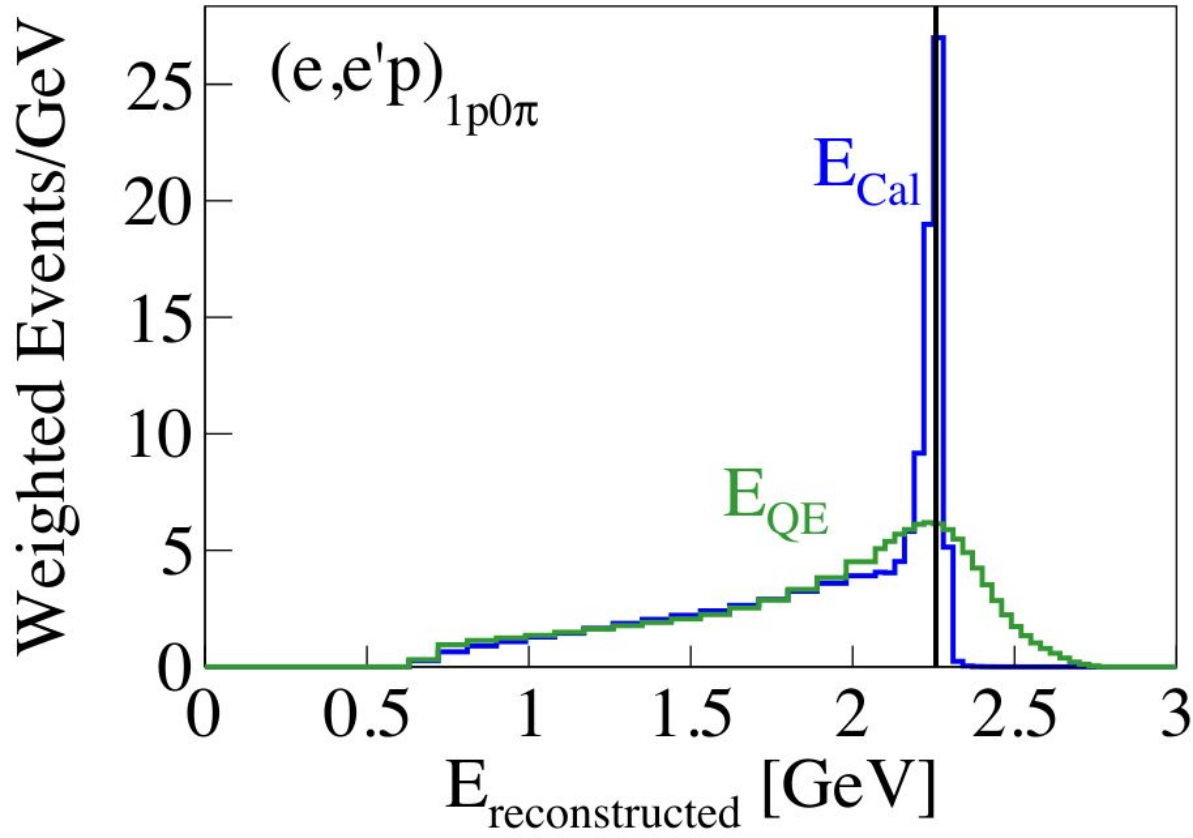
- Cross section sector variation (6%)
- Overall normalization (20% ?????)
- Acceptance correction (G2018 vs SuSav2) (20% ?????)
- 
-

# Multiplicities



# Energy Reconstruction Accuracy

2.257 GeV  $^{12}\text{C}$

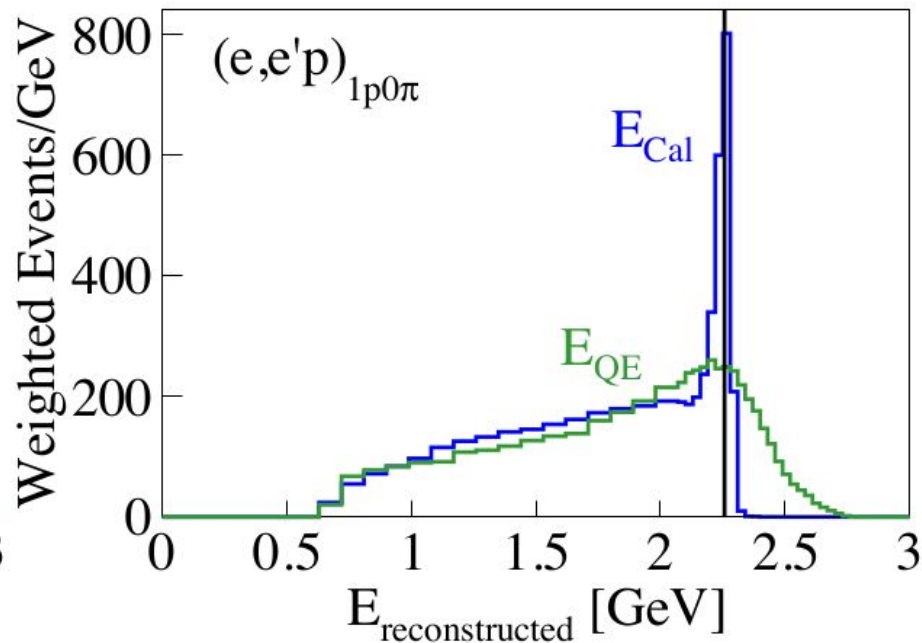
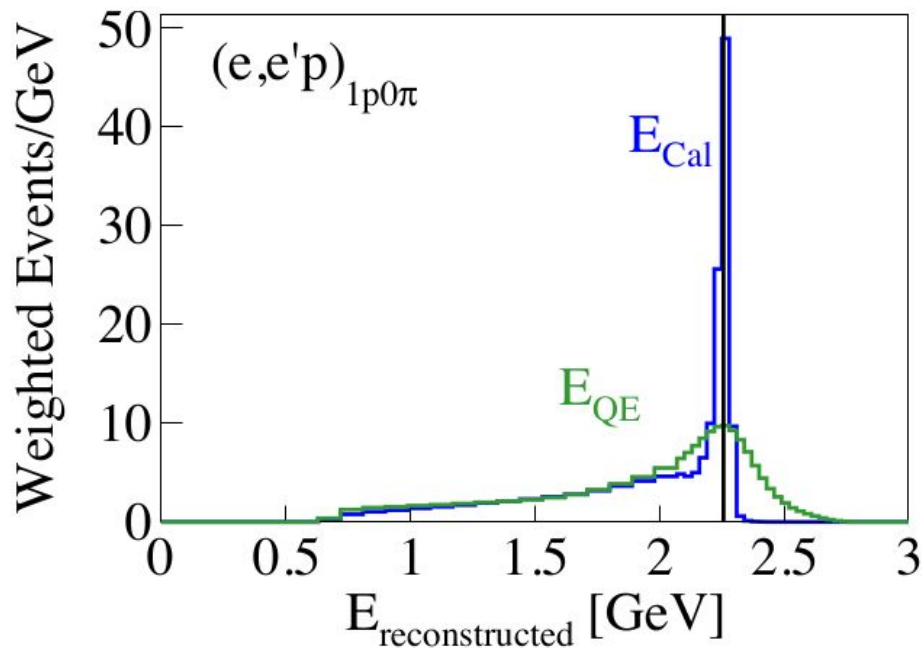


# Worse Reconstruction With Higher Mass

2.257 GeV

${}^4\text{He}$

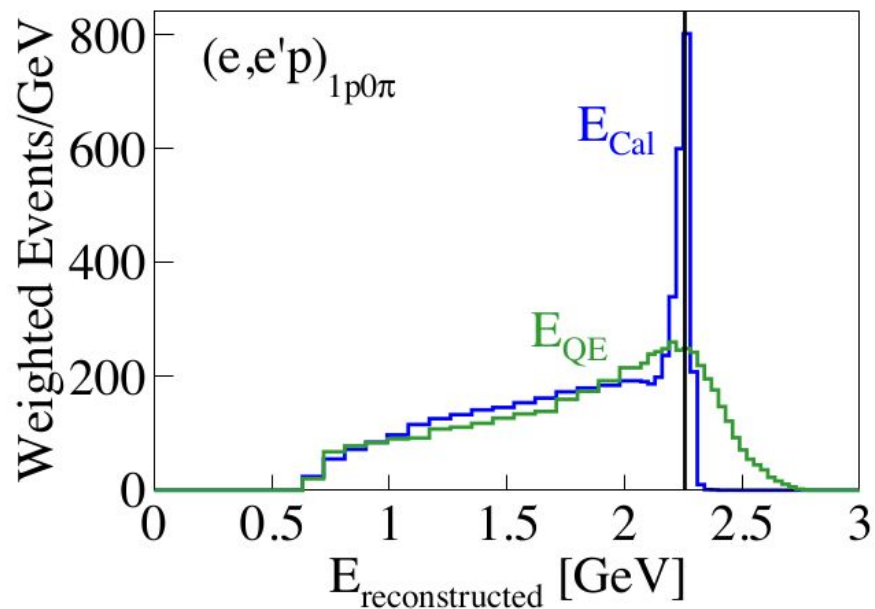
${}^{56}\text{Fe}$



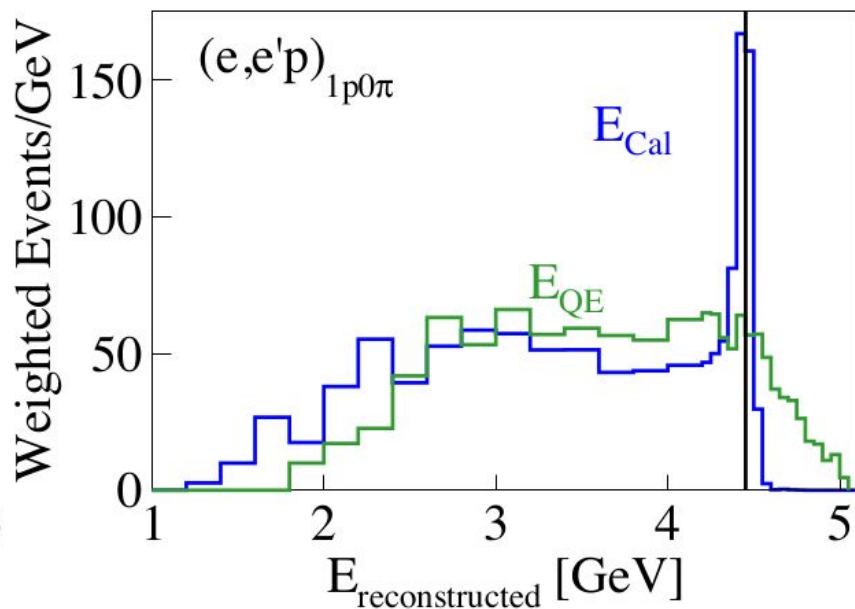
# Worse Reconstruction With Higher Energy

$^{56}\text{Fe}$

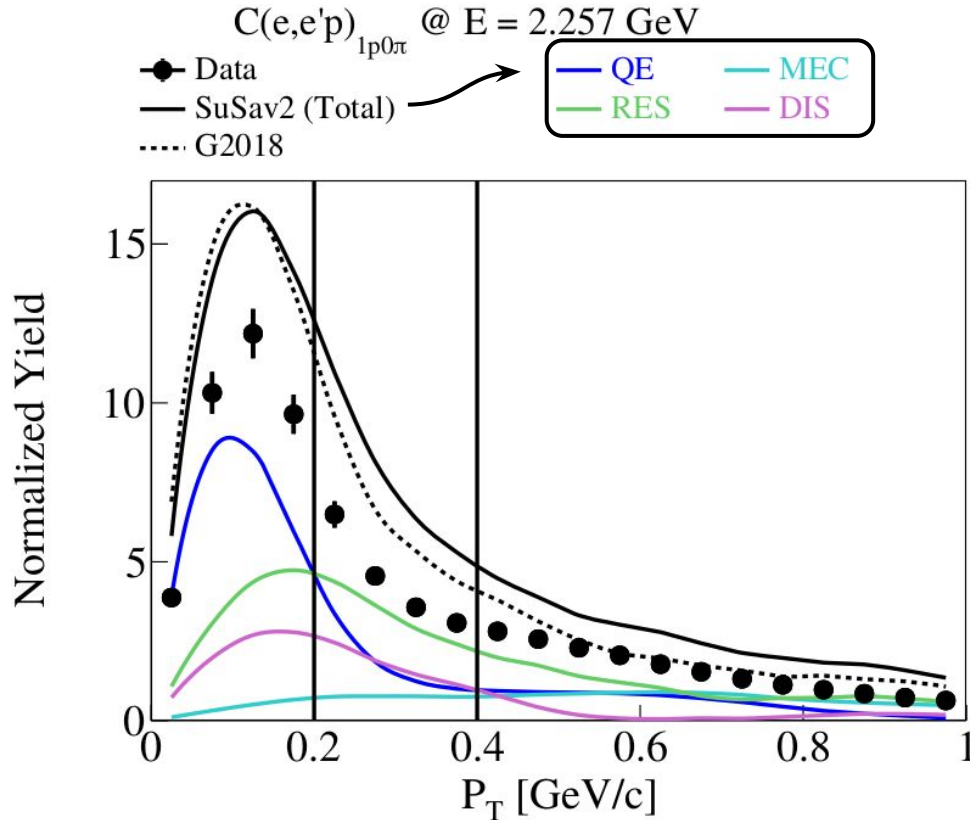
2.257 GeV



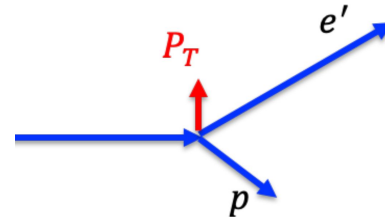
4.453 GeV



# Transverse Missing Momentum



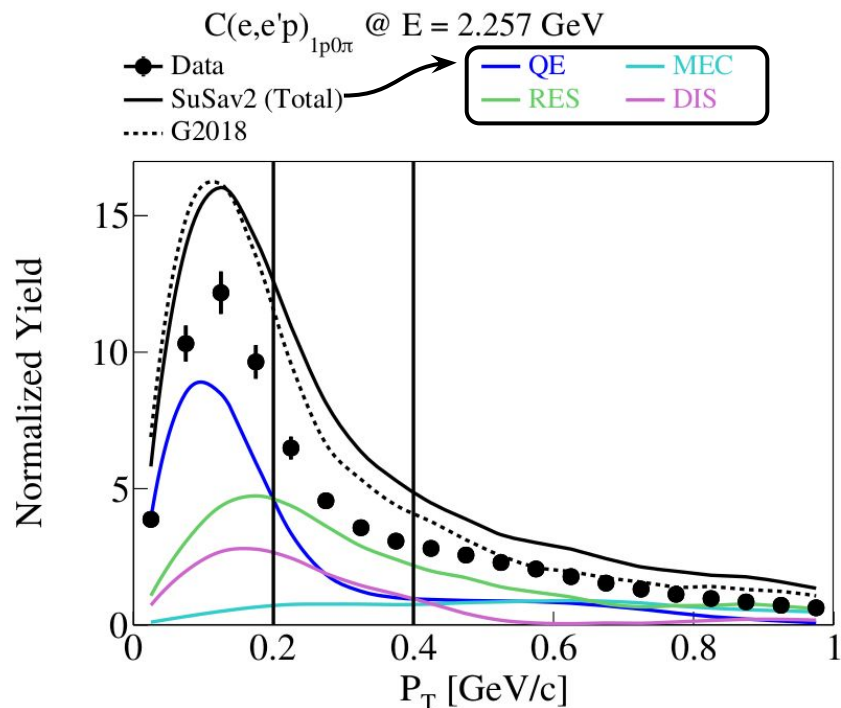
$$\vec{P}_T = \vec{P}_T^{e'} + \vec{P}_T^p$$



Simulation overpredicts strength  
However, overall good shape agreement

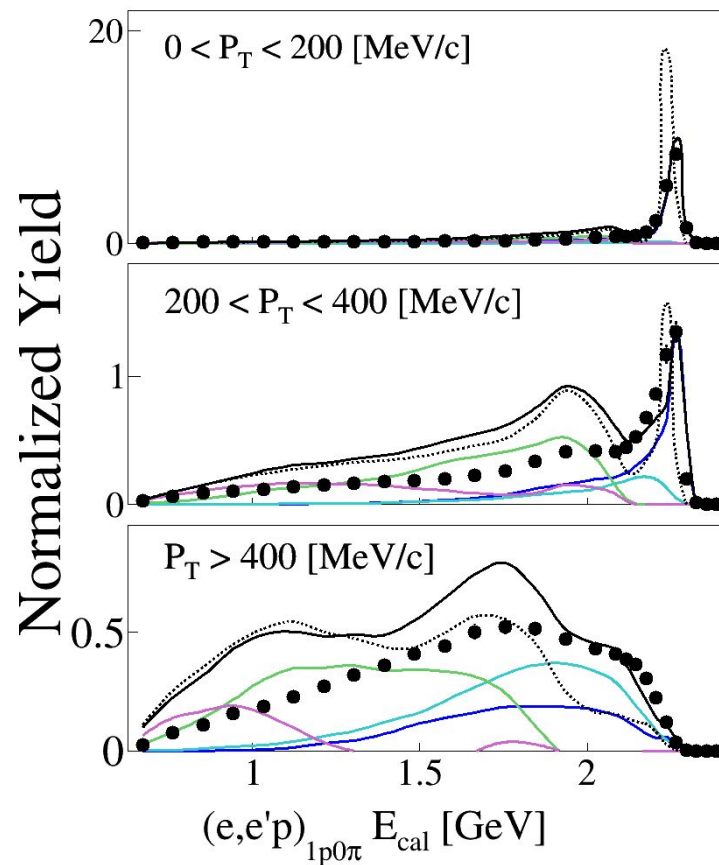
Equivalent MicroBooNE Single  
Transverse Variable analysis will follow

# Transverse Missing Momentum



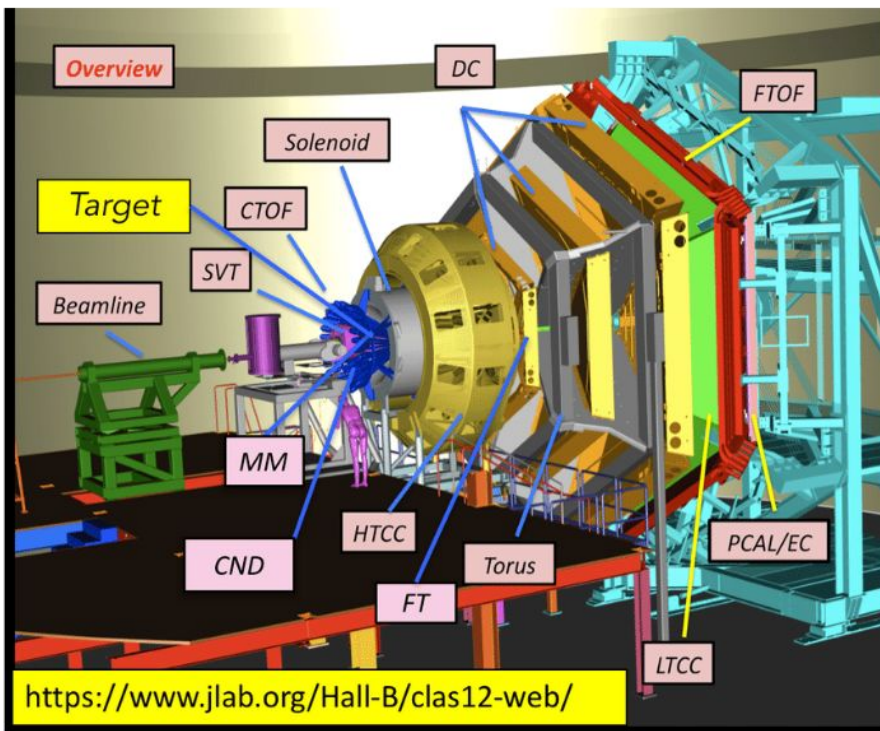
Splitting phase space in  $P_T$  slices

Data/MC disagreement in  $E_{cal}$  for higher  $P_T$



## CLAS 12

Jefferson Lab



### Forward Detector

- Torus Magnet
- Drift Chamber (DC)
- Forward Time of Flight (FTOF)
- High-threshold Cherenkov Counter (HTCC)
- Low-threshold Cherenkov Counter (LTCC)
- Ring Imaging Cherenkov Detector (RICH)
- Preshower + Electromagnetic Calorimeter (PCAL/EC)
- Forward Tagger (FT)

### Longitudinally Polarized Electron Beam

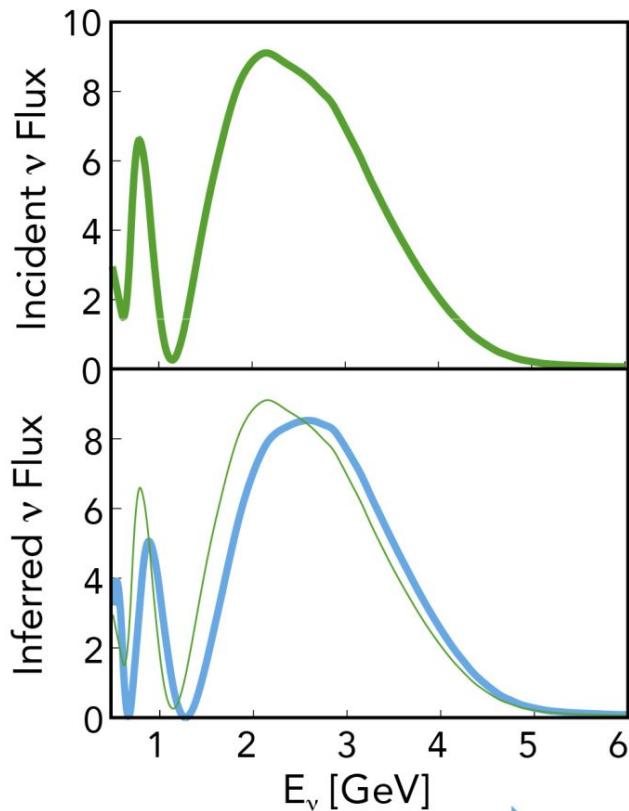
- $E = 10.6 \text{ GeV}$
- $P = 86\text{--}89\%$

### Unpolarized Liquid $\text{H}_2$ Fixed Target

- Torus magnet  $\rightarrow$  electrons inbending



# Reconstructed Energy



$$N(E_{\text{rec}}, L) \sim \sum_i \int \Phi(E_\nu, L) \sigma_i(E_\nu) f_{\sigma_i}(E_{\text{rec}}, E_\nu) dE$$

**Oscillated incident energy spectrum**

Smearing  
Matrix

Smear with GENIE-derived feed-down matrix

**Reconstruct with data-derived feed-dow matrix**

Improvement of event generators needed !

# Smearing Matrix