

# A MEASUREMENT OF THE COSMIC MICROWAVE BACKGROUND DAMPING TAIL FROM THE 2500-SQUARE-DEGREE SPT-SZ SURVEY

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## ABSTRACT

We present a measurement of the cosmic microwave background (CMB) temperature power spectrum using data from the recently completed South Pole Telescope Sunyaev-Zel'dovich (SPT-SZ) survey. This measurement is made from observations of 2540 deg<sup>2</sup> of sky with arcminute resolution at 150 GHz, and improves upon previous measurements using the SPT by tripling the sky area. We report CMB temperature anisotropy power over the multipole range  $650 < \ell < 3000$ . We fit the SPT bandpowers, combined with the results from the seven-year Wilkinson Microwave Anisotropy Probe (*WMAP7*) data release, with a six-parameter  $\Lambda$ CDM cosmological model and find that the two datasets are consistent and well fit by the model. Adding SPT measurements significantly improves  $\Lambda$ CDM parameter constraints, and in particular tightens the constraint on the angular sound horizon  $\theta_s$  by a factor of 2.7. The impact of gravitational lensing on the CMB power spectrum is detected with  $8.1\sigma$ , the most significant detection to date. The inferred amplitude of the lensing spectrum is consistent with the  $\Lambda$ CDM prediction. This sensitivity of the SPT+*WMAP7* data to lensing by large-scale structure at low redshifts allows us to constrain the mean curvature of the observable universe with CMB data alone to be  $\Omega_k = -0.003^{+0.014}_{-0.018}$ . Using the SPT+*WMAP7* data, we measure the spectral index of scalar fluctuations to be  $n_s = 0.9623 \pm 0.0097$  in the  $\Lambda$ CDM model, a  $3.9\sigma$  preference for a scale-dependent spectrum with  $n_s < 1$ . The SPT measurement of the CMB damping tail helps break the degeneracy that exists between the tensor-to-scalar ratio  $r$  and  $n_s$  in large-scale CMB measurements, leading to an upper limit of  $r < 0.18$  (95% C.L.) in the  $\Lambda$ CDM+ $r$  model. Adding low-redshift measurements of the Hubble constant ( $H_0$ ) and the baryon acoustic oscillation (BAO) feature to the SPT+*WMAP7* data leads to further improvements. The combination of SPT+*WMAP7*+ $H_0$ +BAO constrains  $n_s = 0.9538 \pm 0.0081$  in the  $\Lambda$ CDM model, a  $5.7\sigma$  detection of  $n_s < 1$ , and places an upper limit of  $r < 0.11$  (95% C.L.) in the  $\Lambda$ CDM+ $r$  model. These new constraints on  $n_s$  and  $r$  have significant implications for our understanding of inflation, which we discuss in the context of selected single-field inflation models.

*Subject headings:* cosmology – cosmology:cosmic microwave background – cosmology: observations – large-scale structure of universe

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