

Experience Implementing GiBUU in SBND Simulation

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Introduction



GiBUU Implementation for SBND



Motivation

» The GiBUU project provides a unified theoretical and simulation framework for particle interactions with nucleus in a wide energy range (MeV-GeV)

- Based on first-principles interactions
- » GiBUU propagates particles from the initial interactions using a transport model (the BUU equation) which describes the time evolution of the Wigner transform of the real-time Green's function

GiBUU showed good agreement (shape and normalization) with results from many neutrino experiments, including Ar targets (MicroBooNE)







Motivation

- Not implemented as an event generator yet **>>**
 - 0 Larsoft.
- Our first application is the SBN physics program **>>**

Advantages

- 0 efficiency, neutrino energy reconstruction, etc.
- 0





Our plan is to implement GiBUU as an alternative generator, event by event, in

An independent account for corrections coming from simulations: purity (signal-to-bkg migrations),

Crucial for extrapolations in oscillation analysis (model dependent analysis) and important for model independent measurements (such as neutrino-Ar cross sections) and to understand in BSM searches





GiBUU weighted events

» The GiBUU's output is a list events

These weights

- Each event has a weight
- can have negative values (if destructive interference terms, such as some 0 resonances which are relevant for higher energies)
- have to be added to obtain the cross section for any variable (neutrino, FS, etc)

These cross sections have to be converted into a generated event-by-event interaction and propagated through the detector



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General approach

- - For negative weights we take the absolute value to produce events. 0
- results can be calculated by subtracting both.
- » We use the flux and geometry GENIE infrastructure to propagate those GiBUU events in the detector geometry (SBND).



» The first step is to generate GiBUU events with (reasonable) high statistics and store them in a standard format we can use. We produce two libraries, one for positive and one for negative weights

» We made two productions (one corresponding to a negative and one for positive weights). The final







GiBUU weights

- » GiBUU weights have a wide distribution that span over several orders of magnitude
 - Not only depend on the neutrino energy but the different processes.



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GiBUU weights

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 - Not only depend on the neutrino energy but the different processes.

- » In a first approach, we tried to use the event Library Interface (EvtLib) from GENIE (available in version >= 3.2)
- » Unfortunately the EvtLib does not handles weighed library inputs and does not propagate the interaction mode of external generators)









Libraries

- » A standalone code to generate GiBUU events and create the libraries
 - We use the most recent GiBUU version (2021)
 - 0 completeness
 - The library contains all neutrino interaction flavors (ν_{μ} , ∇_{μ} , ν_{e} , ∇_{e} , and NC and CC).



Only interactions in Ar40 are produced for now. We need to expand to other materials for

Each entry has:

- Neutrino energy, weight, process ID (QE, 2p2h, etc)
- Array of final state particles: ID and kinematics •





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- - library
 - 0 vs GiBUU cross section. For instance: $w = \sigma^{GiBUU}_{v\mu CC} / \sigma^{GENIE}_{v\mu CC}$



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Each entry has:

- Neutrino energy, weight, process ID (QE, 2p2h, etc)
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» A code (ART module based on **GENIEGen**) generates and replaces GENIE events with GiBUU events • We keep the interaction vertices and use the neutrino information to find an event in the GiBUU

We calculate and propagate a weight to account for the difference between the inclusive GENIE









Choosing the GiBUU entry

- » The GiBUU entry is chosen given the distributions of weights
 - This distribution is made using a narrow neutrino energy: $\Delta = 0.001$ GeV
 - An entry from the library from a random selection. 0
 - Given the wide span of the weights, the challenge is not 0 repeating multiple times entries with very large weights



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 - We add a random rotation around the 0 neutrino axis for the final particles
 - This introduces some 0 randomization and reduce the cost of repeating events



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Prof-of-principle: Full implementation for SBND







Short-Baseline Near Detector





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Short-Baseline Near Detector



neutrino analysis:

- **Near Detector: SBND >>**
- **Far Detector : ICARUS** \rightarrow

Same neutrino beam, nuclear target and detector technology (LAr TPC detectors) to reduce systematic uncertainties to the % level.

v-Ar interactions: with an order of magnitude more data than is currently available (5000 v events/per day)

In addition to the sterile and cross-section programs, the SBND large detector mass and proximity to intense beams enable a broad physics program such BSM searches



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Short-Baseline Near Detector

Large-mass Liquid Argon Time Projection Chamber (LArTPC)

- 3D reconstruction with a mm position resolution **>>**
- Fine-granularity calorimetry $\rangle\rangle$
- Excellent particle identification with dE/dx information **>>**
- Low energy thresholds: few MeV **>>**

Photon Detection System (PDS)

- Novel technology of PMTs and X-Arapucas. **>>**
- Scintillation & reflected light => high and uniform \rightarrow light yield and excellent timing resolution

Cosmic Ray Tagger (CRT)

Timing and position resolution allows for triggering \rightarrow on entering/exiting particles

Cold commissioning - Summer 2023

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Implementation

- » We have implemented GiBUU into LarftSoft
 - analysis framework (CAFAna)



• We generate GiBUU events and propagate them through the full simulation, including the final







Interaction modes

» We also propagate the GiBUU process ID through the simulation

SBND

• This information is crucial to calculate the neutrino interaction systematics.







Other checks: enhanced QE and 2p2h



1p, 0n and 0pi: enhanced QE

• 2p, 0n and 0pi: enhanced 2p2h





Other checks: hadron production

» Average multiplicity per neutrino interaction and the average kinetic energy of hadrons in NuMu-CC interactions



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Neutrino Energy

Distributions for the neutrino **>>** energy per current and neutrino type







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Interaction vertex distributions

Transverse view





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Interaction vertex distributions

Longitudinal view





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Plan for uncertainties



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Principle

» Split the SBND sample into 2 regions in the de ICARUS' projection (sample A) and an externa





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etector al ring	r: one circle of g (sample B) .		
	0.34 ⁰	Sample A	
SBND		Sample B	ICARUS







Principle

Split the SBND sample into 2 regions in the detector: one circle of **>> ICARUS' projection (sample A)** and an **external ring (sample B)**.

- Use **sample A** for the oscillation fit (need study stats. limitations on topology $\rangle\rangle$ sampling). Likely cancellation of flux uncertainties (residual uncertainties expected)
- Uncertainties will come from two sources: \rightarrow

>> enhancing the sensitivity to the parameters we want.



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0.34 ⁰	Sample A	
SBND	Sample B	ICARUS

External e- and hadron scattering

Neutrino scattering in sample B

Event selections in sample B will be done with our own detector efficiencies, signal definitions, and





Principle

Pros

- » it can provide a robust model extrapolation and let us target desired sample for oscillation and error estimation.
- » Beam uncertainties when comparing sample A and B should be small, so we expect better error precision
- » Will reduce to its minimum flux uncertainties in the osc measurement.

Outgoing studies on selection efficiency, statistics, etc. Final answer till we look at the data.





Conclusions

- coming from simulations.
 - GiBUU has shown good agreement with neutrino-nucleus data, even in Ar.
- We have implemented GiBUU as an alternative generator, event by event, in LarSoft. **>>**
 - This includes the propagation through the full chain till the analysis stage.
- » We are focused to incorporate realistic systematics coming from our own neutrino-Ar data.
 - We are collaborating with Ulrich Mosel and Kai Gallmeinster (GiBUU authors) to be loyal to the physics.
 - experience of Costas Andreopoulos and Marco Roda.



» Having an alternative generator has the advantage of an independent account for all corrections

• We are exploring using GENIE-Reweight to propagate systematics. We are learning from the





Backup



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GiBUU interaction modes

GiBUU modes:

- 1: nucleon (QE)
- 2-31: non-strange baryon resonance
- 32: pi neutron-background (e.g. $v + n \rightarrow \mu + \pi^+ + n$)
- 33: pi proton-background (e.g. $v + n \rightarrow \mu + \pi^0 + p$)
- 34: DIS
- 35: 2p2h QE
- 36: 2p2h Delta
- 37: two pion background

Link here



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Comparison to GENIE

Neutrino energy for muon- and electron-neutrinos





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