# Gravitational effects in quantum systems

MAGIS group meeting December 7, 2022

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- 1. Spacetime curvature across a quantum state
- 2. Gravitational Aharonov-Bohm effect
- 3. Implications for gravitational field superposition
- 4. Future directions and prospects for MAGIS



# A neutron interferometer in an approximately uniform gravitational field



Phase:  $\Phi = 4\pi \lambda g h^{-2} M^2 d (d + a \cos \theta) \tan \theta \sin \phi$  $= k g T^2 \cos \theta \sin \phi$ 

R. Collela, A. W. Overhauser, and S. A. Werner, Phys. Rev. Lett. 34, 23 (1975)



Gravity is a purely nonlocal phenomenon

Gravity cannot be observed in any local measurement



Implication: uniform gravitational fields are not observable



-EP says that there are no local gravitational effects

-To observe gravity, need a nonlocal interferometer

-Resolvable relative acceleration between interferometer arms:

$$\left|\frac{\partial^2 V}{\partial x^2} \cdot \Delta x\right| > \delta g$$



## Gravity as spacetime curvature

#### -Tidal forces across wavefunction



indicates "genuine gravitational effects in quantum interference as opposed to a mere test of the equivalence principle." J. Anandan, *Phys. Rev. D* **30**, 8 (1984)



#### Apparatus







### Nonlocal interferometer: implementation







P. Asenbaum et al., Phys. Rev. Lett. 118, 183602 (2017)

#### Observation of the tidal phase shift





P. Asenbaum et al., Phys. Rev. Lett. 118, 183602 (2017)

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Beyond the midpoint theorem

$$\phi = \phi_{MP} \qquad \phi = \phi_{MP} + \phi_{\Delta S}$$

$$\int \int \phi_{\Delta S} = \frac{\Delta S}{\hbar}$$

$$S = \int \mathcal{L}(x, \dot{x}, t) dt$$

$$C. \text{ Overstreet et al., Am. J. Phys. 89, 324 (2021)}$$

#### The Aharonov-Bohm effect



FIG. 1. Schematic experiment to demonstrate interference with time-dependent scalar potential. A, B, C, D, E: suitable devices to separate and divert beams.  $W_1, W_2$ : wave packets.  $M_1, M_2$ : cylindrical metal tubes. F: interference region.

Aharonov and Bohm, Phys. Rev. (1959)

Previous gravitational experiments:

Aharonov-Bohm experiments:

Gravitational Aharonov-Bohm?



FIG. 2. Schematic experiment to demonstrate interference with time-independent vector potential.

$$\phi_{\Delta S} = 0, \ \phi = \phi_{\rm MP}$$

$$\phi_{\rm MP} = 0, \ \phi = \phi_{\Delta S}$$

J. Audretsch and C. Lämmerzahl, J. Phys. A (1983) M. A. Hohensee et al., Phys. Rev. Lett. (2012)



#### Gravitational Aharonov-Bohm effect: ingredients



#### Gravitational Aharonov-Bohm effect: results



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The success of quantum field theory in describing all other interactions suggests that gravity is quantum.

Gravity is well-described in the low-energy (linearized) limit as a spin-2 quantum field.

BUT:

-GR QFT is not perturbatively renormalizable. -Our best theory of gravity (GR) is classical. -Semiclassical gravity is mathematically well-defined.

Problem: no experimental input!



Classical physics: potential energy of a particle = energy of the field

$$U_{\rm EM} = qV_{\rm EM} = \frac{1}{2}\epsilon_0 \int |\mathbf{E}|^2 dV$$

-Also true in quantum mechanics

Field energy principle: potential energy is located in the field



C. Overstreet et al., arXiv:2209.02214 (2022)

Relativity principle (RP): laws of physics take the same form in every coordinate system

Quantum relativity principle (QRP): extends RP to quantum reference frames



C. Overstreet et al., arXiv:2209.02214 (2022)

-Semiclassical gravitational models must reject field energy principle and quantum relativity principle

-Schrödinger-Newton, Oppenheim models not ruled out

-QRP test: effect of lower-mass particle on a higher-mass superposition





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What we've learned: the trajectory and the action of a quantum system in a Newtonian gravitational field are correctly predicted by quantum mechanics.

Next steps: beyond the Newtonian approximation (GR phase shifts) -Nonlinearity of gravity -Gravitation of kinetic energy -Others?

Simulation/design work could be useful here!

S. Dimopoulos et al., Phys. Rev. D 78, 042003 (2008)



Direction 2: how do quantum systems source gravity?

To test semiclassical gravity, need to create a high-mass superposition state (~10<sup>12</sup> amu)

State of the art: 2.5 x 10<sup>4</sup> amu (Fein 2019)

Idea: use atoms! The effective mass of *N* entangled atoms of mass *m* is *Nm*.

Goal: > 10<sup>10</sup> highly entangled atoms in a spatial superposition

-Note that 10<sup>6</sup> atoms with 20 dB squeezing would improve the state of the art!



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