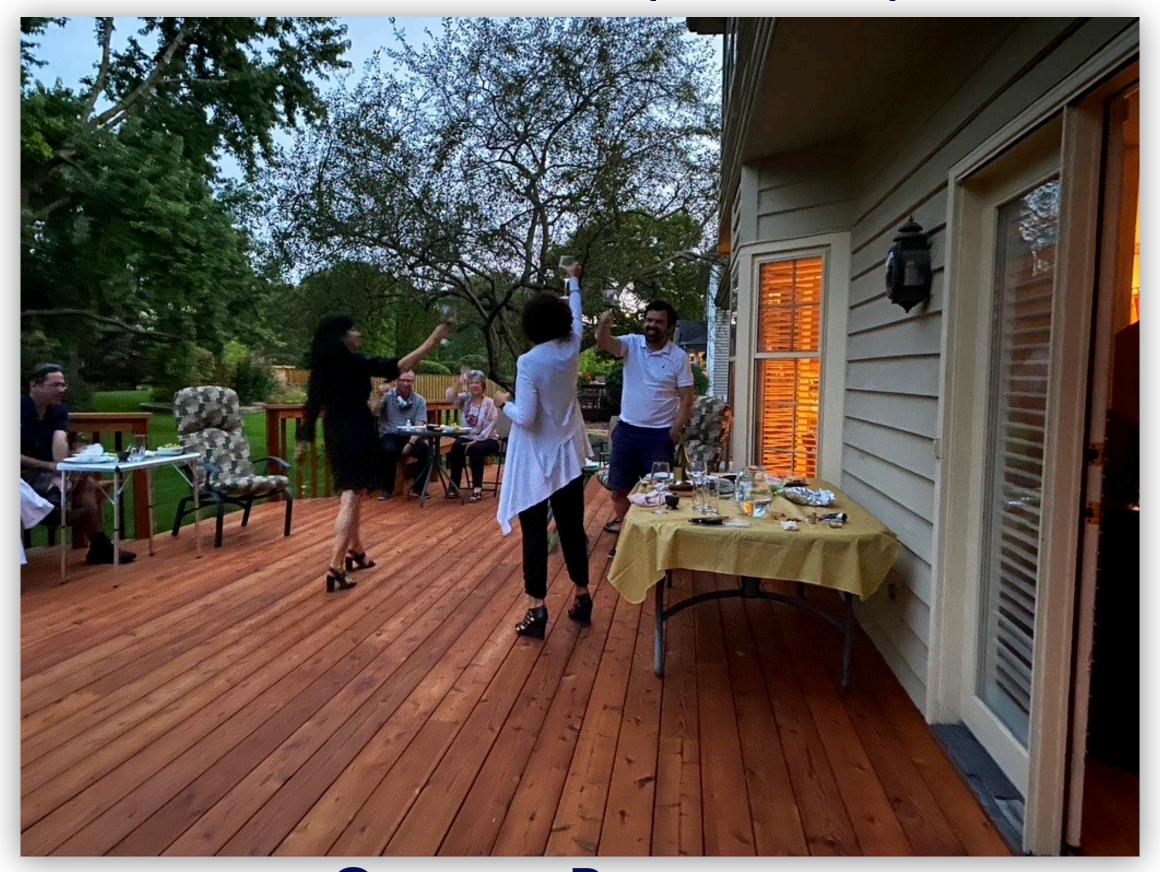
# NEW OPPORTUNITIES IN NEUTRINO PHYSICS AND BEYOND

BEYOND THE SM FROM COLLIDER TO EARLY UNIVERSE

Symposium in honor of Marcela carena and Carlos Wagner 60th birthday

UNIVERSITY OF CHICAGO, MAY 29<sup>TH</sup>, 2023



ORNELLA PALAMARA
FERMILAB

# TO BEGIN

We're delighted to celebrate our friends Marcela and Carlos and congratulate them for their brilliant and impactful scientific careers.

We appreciate that they have, directly and indirectly, fostered so much influential work and have helped cultivate the careers of so many people in our field.



As a neutrino experimentalist, I want to thank you for exposing to and trapping me into a variety of **New Physics ideas!** 

# OUTLINE

Beyond the Standard Model with Neutrinos

A successful history of experiment-theory joint projects\*

Some examples, including physics from "empty" events...

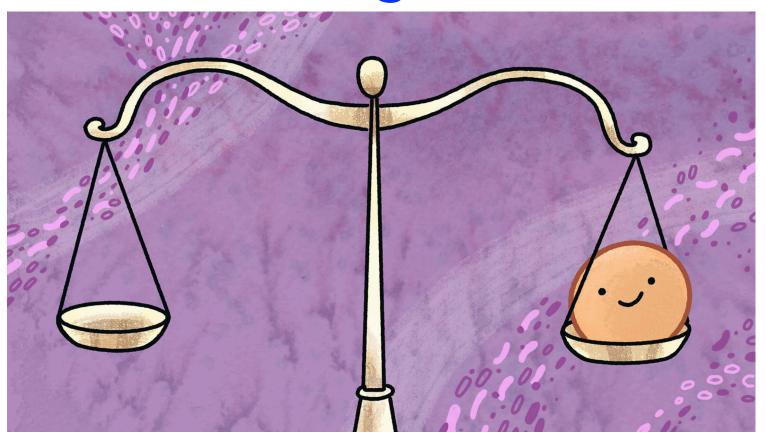
ArgoNeuT: constraints on New Physics in unexplored parameter space regions

Short-Baseline Neutrino program: eV-scale sterile neutrinos & other BSM explorations.

<sup>\*</sup>Examples are from personal experience! Focus on Short-Baseline Liquid Argon Time Projection Chamber neutrino detectors.

The existence of non-zero neutrino masses, inferred from neutrino oscillation measurements, is the only laboratory-based evidence of physics beyond the standard model!

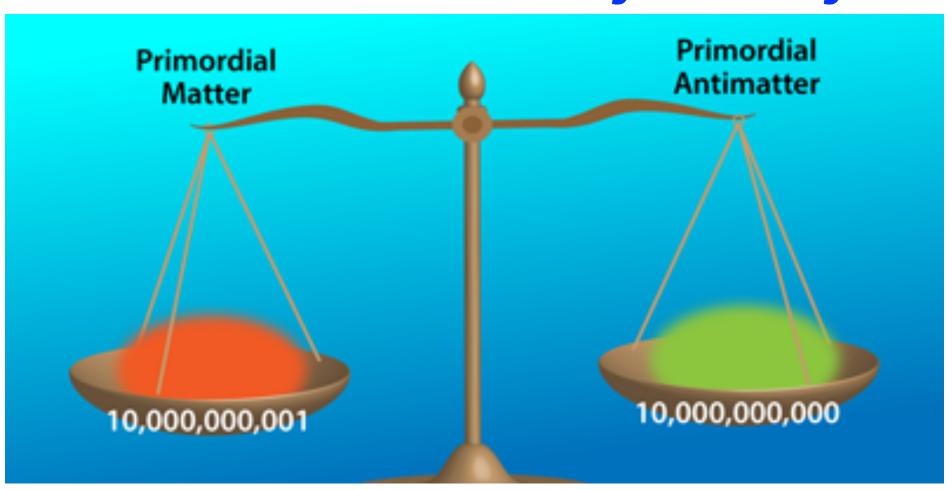
#### How do neutrino get their masses?



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How do neutrino get their masses?

Are neutrinos the key to the matter-antimatter asymmetry?

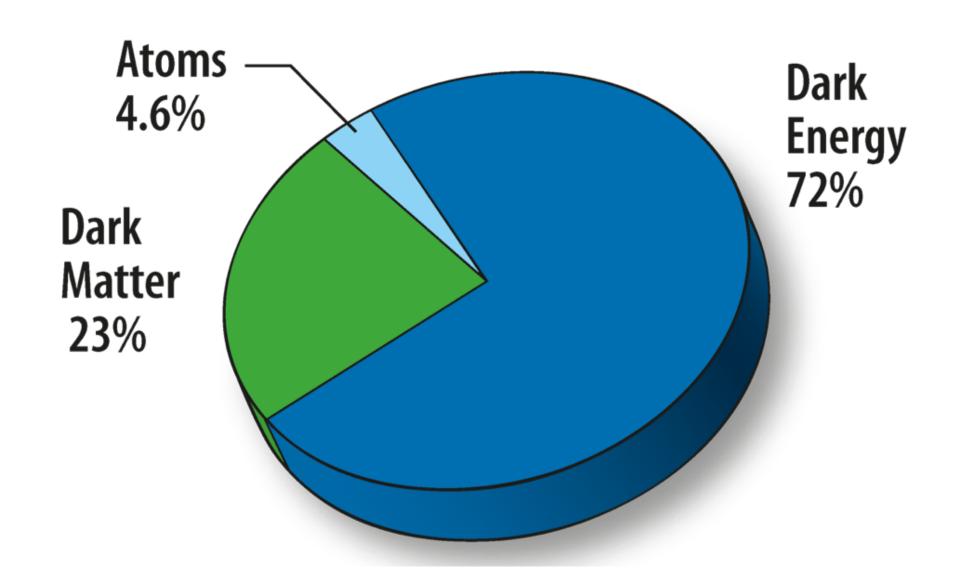


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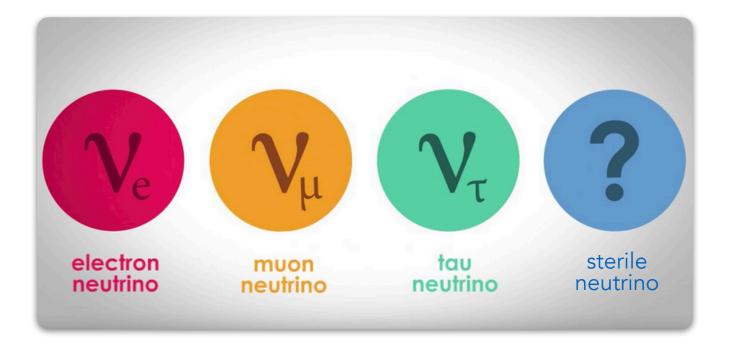
How do neutrino get their masses?

Are neutrinos the key to the matter-antimatter asymmetry?

Do neutrinos communicate with DM?

Are neutrinos their own antiparticle?

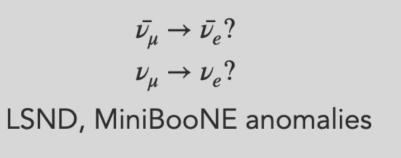
Are there more than three neutrinos?



### SHORT-BASELINE NEUTRINO ANOMALIES -BEYOND THREE NEUTRINO MIXING?

Three neutrino mixing is well established (data from solar, atmospheric, reactor and accelerator neutrino experiments)! Picture consistent with the mixing of 3 neutrino flavors with 3 mass eigenstates - with relatively small mass differences

It is a long (more than 30 years!) standing mystery if other types of neutrinos might exist.



 $\bar{\nu_e} \rightarrow \bar{\nu_e}$ ? Reactor, Gallium anomaly



**LSND** Liquid Scintillator Neutrino Detector

anti- $v_e$  appearance in anti- $\nu_{\mu}$  beam

3.8 σ excess



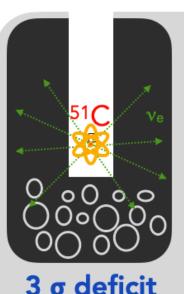
4.5 (2.8)  $\sigma$  excess



**Reactor Anomaly** reactor flux calculations

anti-ve disappearance in anti- $v_e$  beam

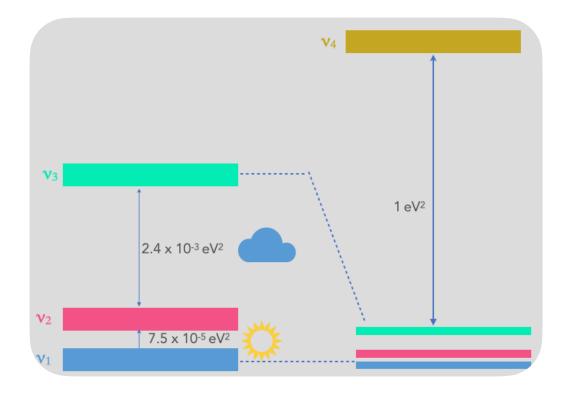
3 σ deficit



**Gallium Anomaly** Solar neutrino expt. calibrations

anti- $v_e$  disappearance anti- $v_e$  source

3 σ deficit



Could point at new physics in the neutrino sector: additional neutrino state(s) with large mass differences (eV<sup>2</sup> scale), driving neutrino oscillation at short distances\*

Any additional neutrino doesn't participate in weak interactions ⇒

"sterile" neutrino

\*Short-Baseline: L~100-1000 m,  $\Delta m^2 \sim 1eV^2$ cfr.: Long-Baseline: L~100-1000 km,  $\Delta m^2 \sim 10^{-3} eV^2$ 

 $\nu_{\mu} \rightarrow \nu_{\mu}$ not Observed Strong appearance-disappearance tension!

But...

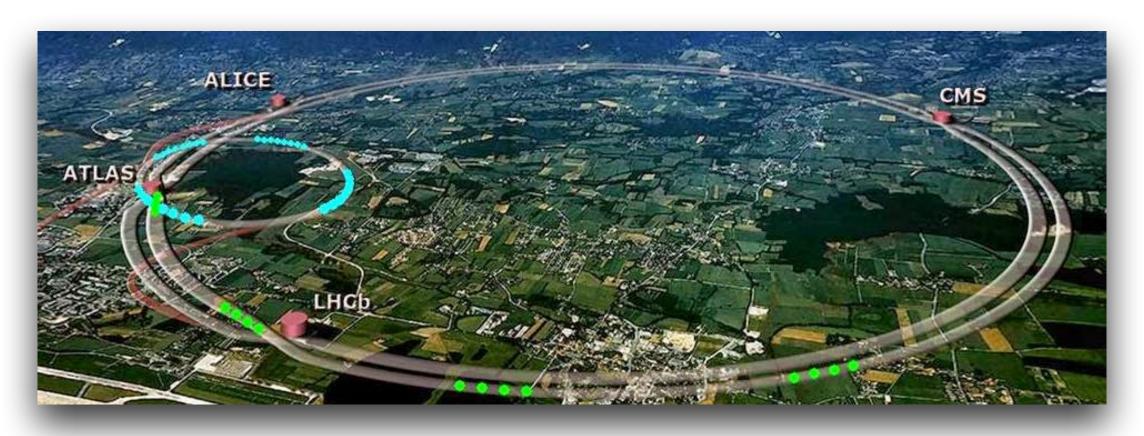
# WHERE TO STUDY NEW PHYSICS?

We know there needs to be physics beyond the standard model. We have no idea of what and where that is.

A reach science program to explore the unknown!

For a given experiment, search in many diverse signal regions for new models.

New physics searches have been the domain of high-energy colliders for decades.



Increasing interest for similar/complementary opportunities to be explored in **neutrino experiments**!

The BSM models accessible are specific to the neutrino source used.

Here we will focus on Short-Baseline Liquid Argon detectors on accelerator neutrino beams at Fermilab

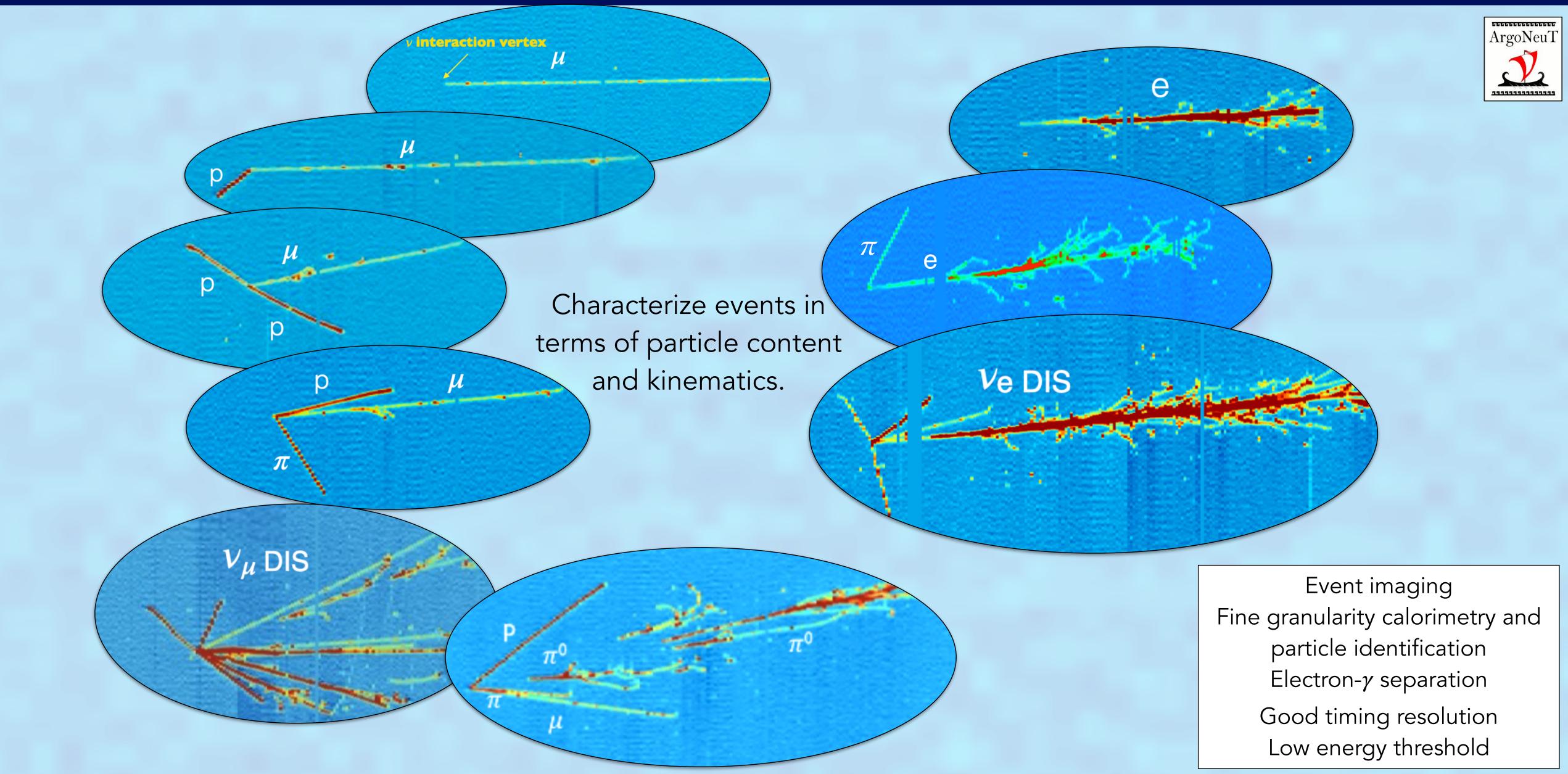




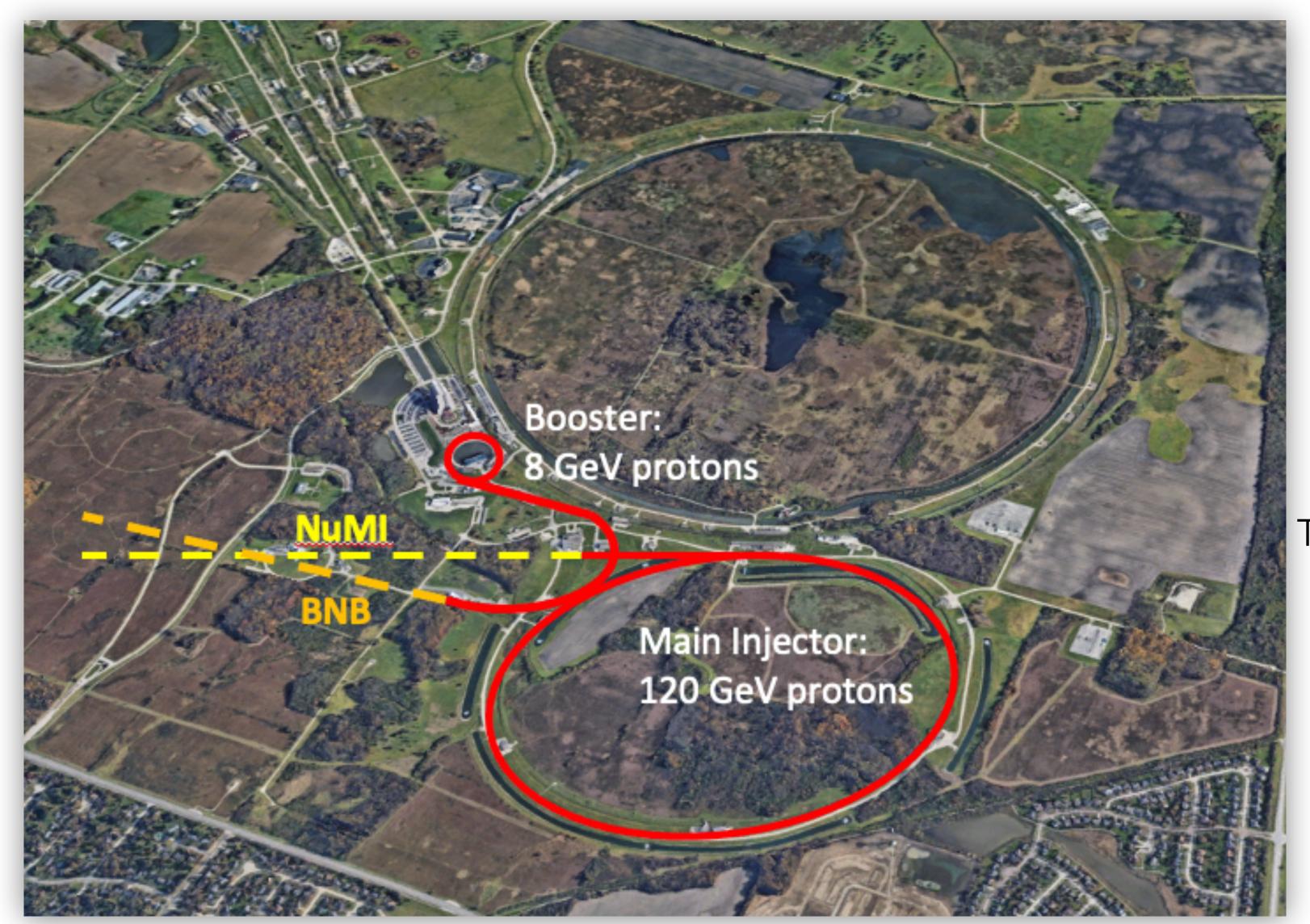




# Why Liquid Argon Neutrino Detectors?



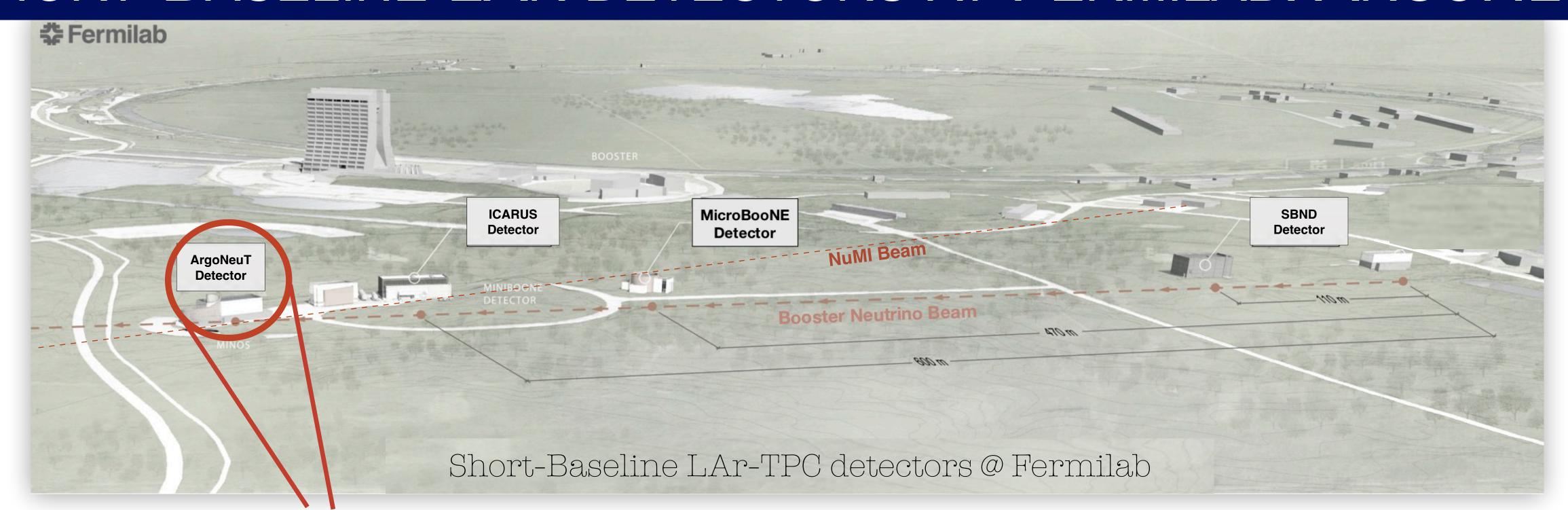
# FERMILAB - NEUTRINO BEAMS



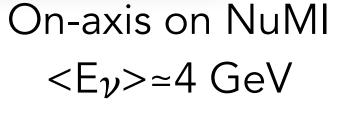
#### Two neutrino beams:

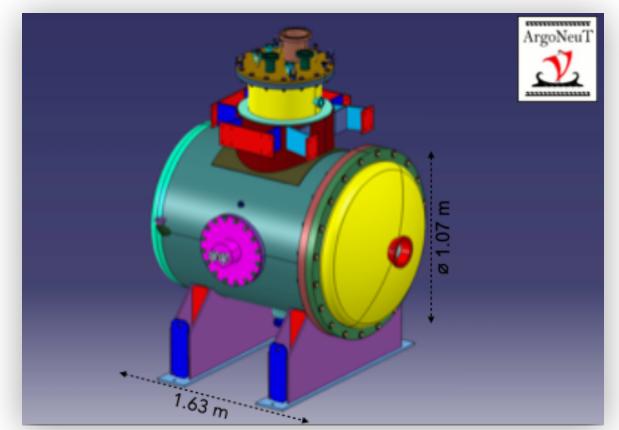
- Neutrinos at the Main Injector (NuMI)
- Booster Neutrino Beam (BNB)

### SHORT-BASELINE LAR DETECTORS AT FERMILAB: ARGONEUT

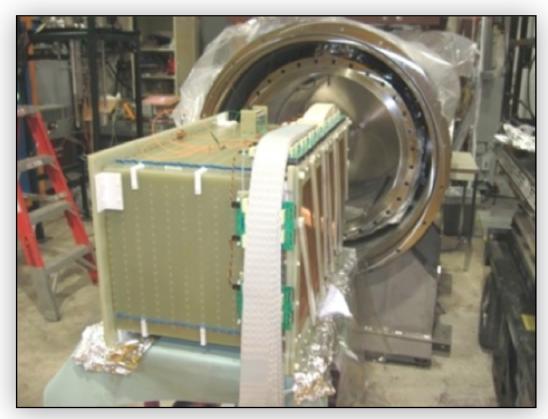


First LAr TPC detector at Fermilab. 5 months neutrino data collected in 2009-2010

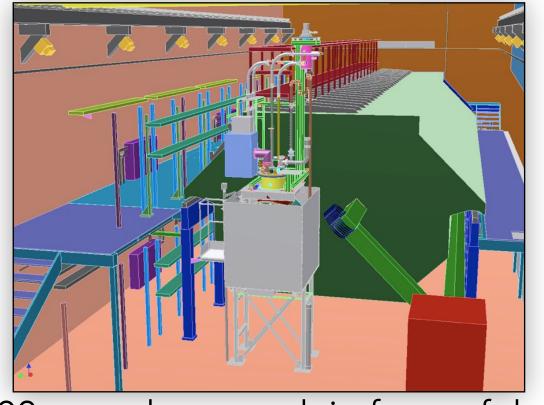




ArgoNeuT Collaboration, JINST 7 (2012) P10019



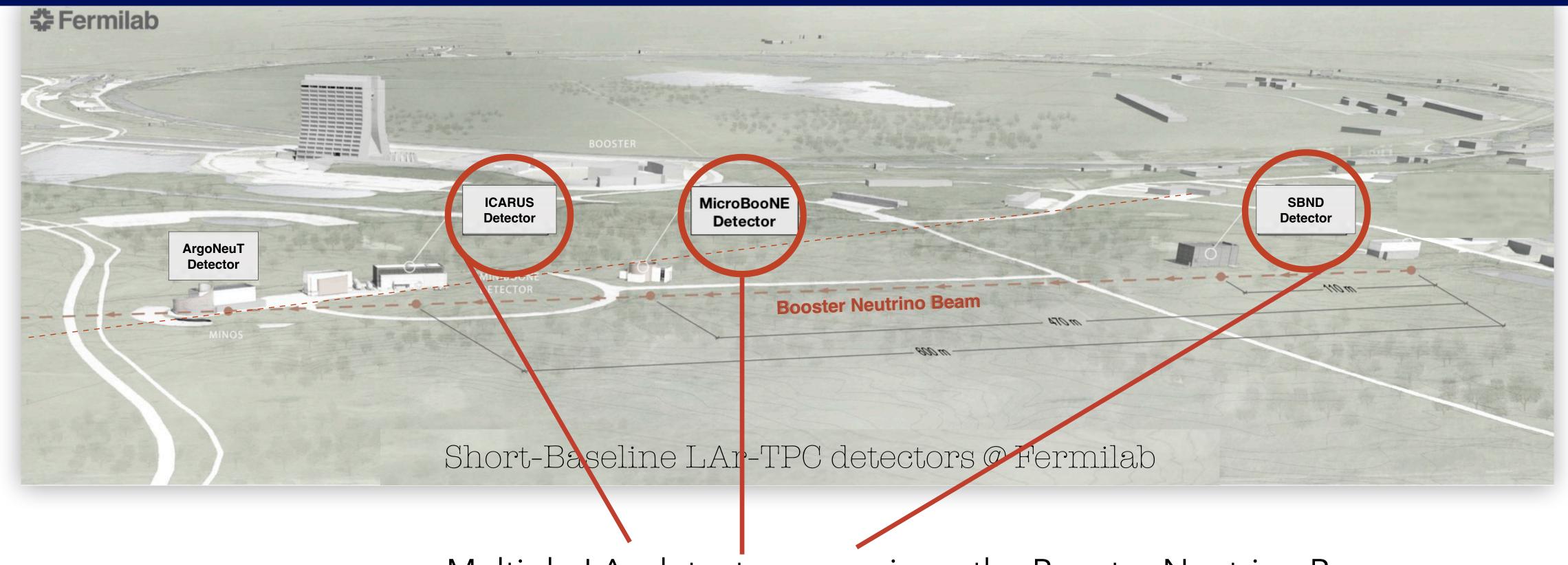
0.24 tons active volume LAr TPC



100 m underground, in front of the MINOS ND, ~ **1km from target** 

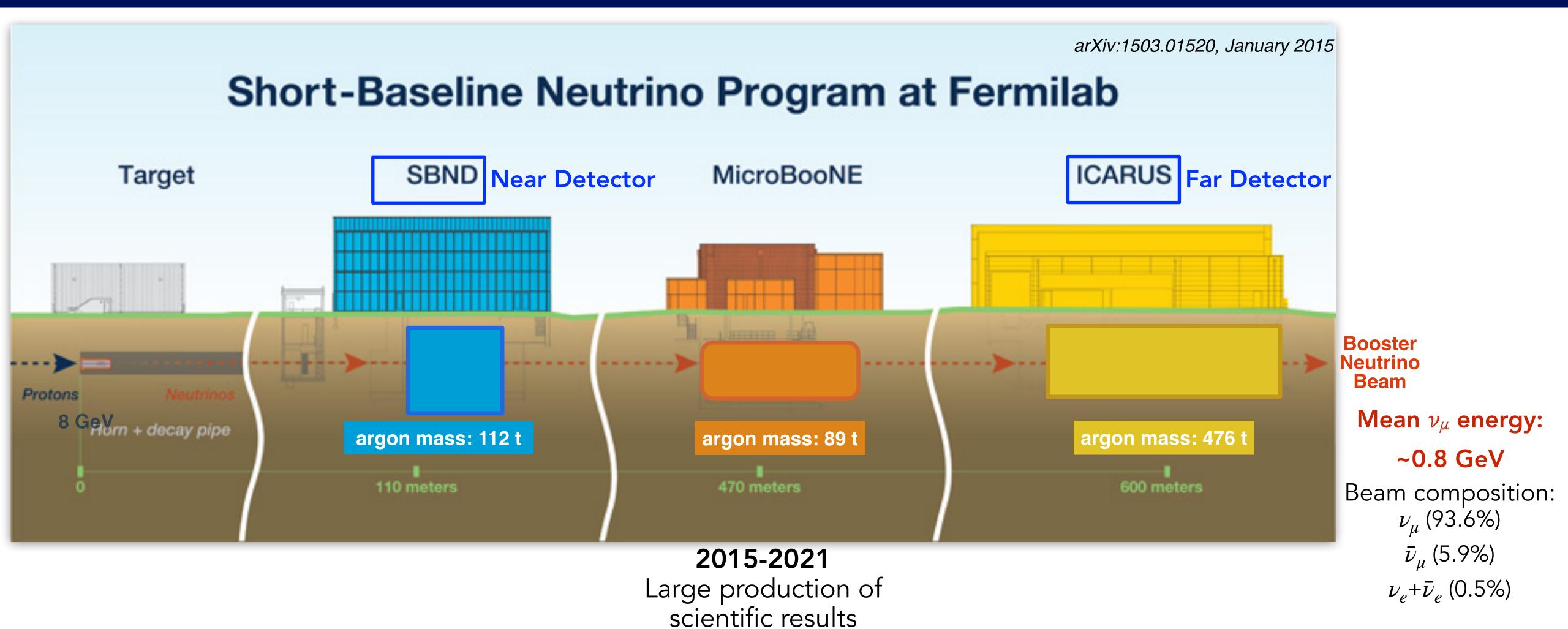
Table-top size, built as a test experiment... but still producing physics results!

### SHORT-BASELINE LAR DETECTORS AT FERMILAB: SBN DETECTORS



Multiple LAr detectors on-axis on the Booster Neutrino Beam

# SHORT BASELINE NEUTRINO PROGRAM



A program designed for **Sterile Neutrino** searches: same neutrino beam, nuclear target and detector technology to reduce systematic uncertainties to the % level.

But large mass LAr detectors and proximity to intense beams enables a broad physics program.

See later...

# SBN FAR DETECTOR: ICARUS



From Gran Sasso Laboratory to Fermilab via CERN



Four Time Projection Chambers in two modules

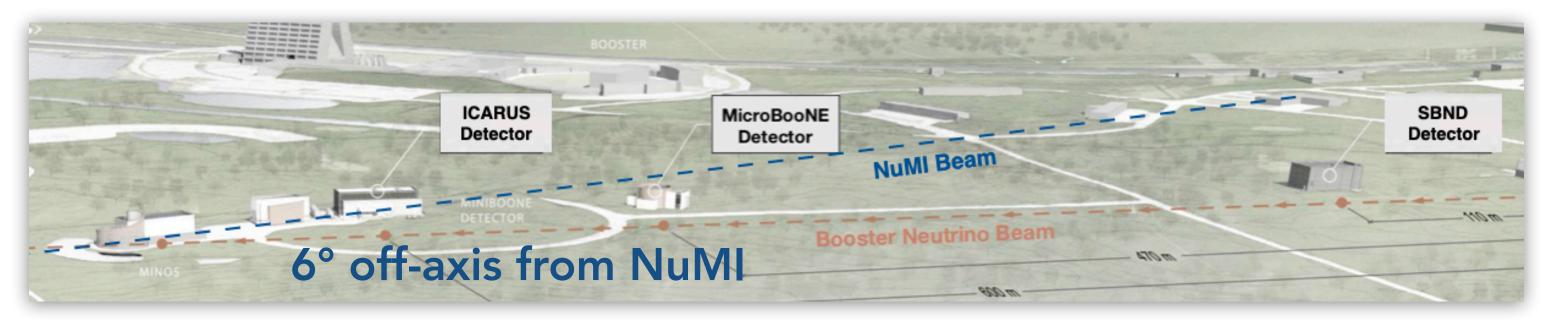
Each module dimension:

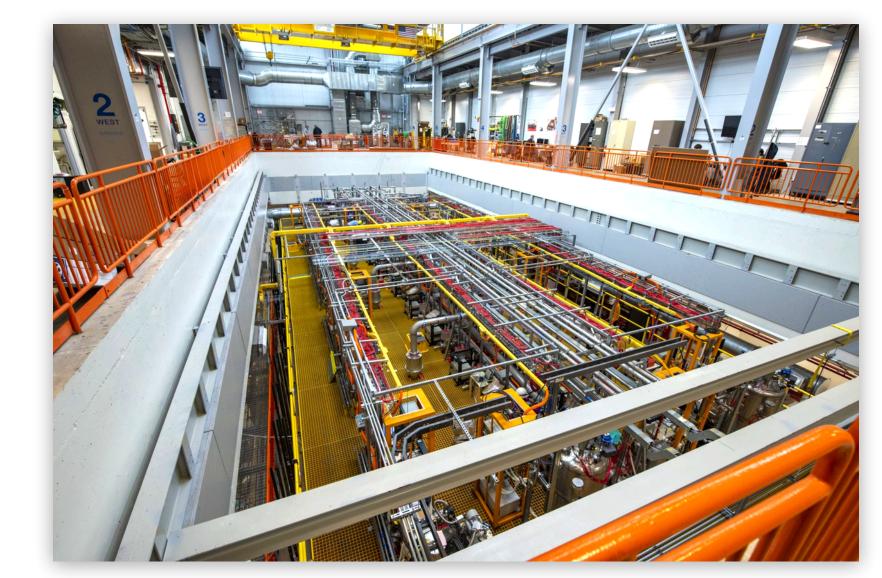
3m x 4m x 18m

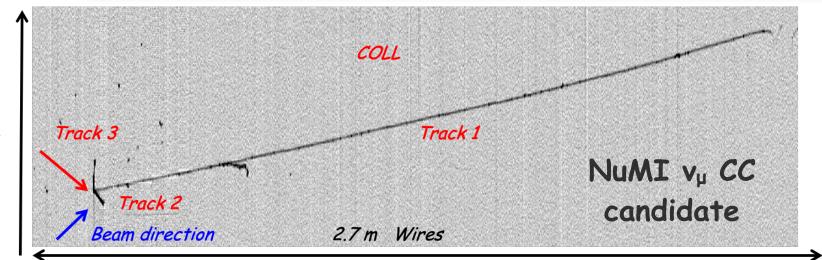
Installation of ICARUS cryostats

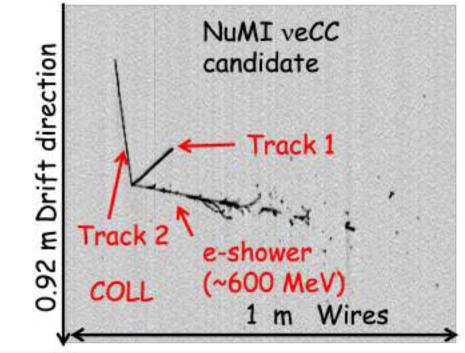
Collecting data in final configuration since **June 2022** 





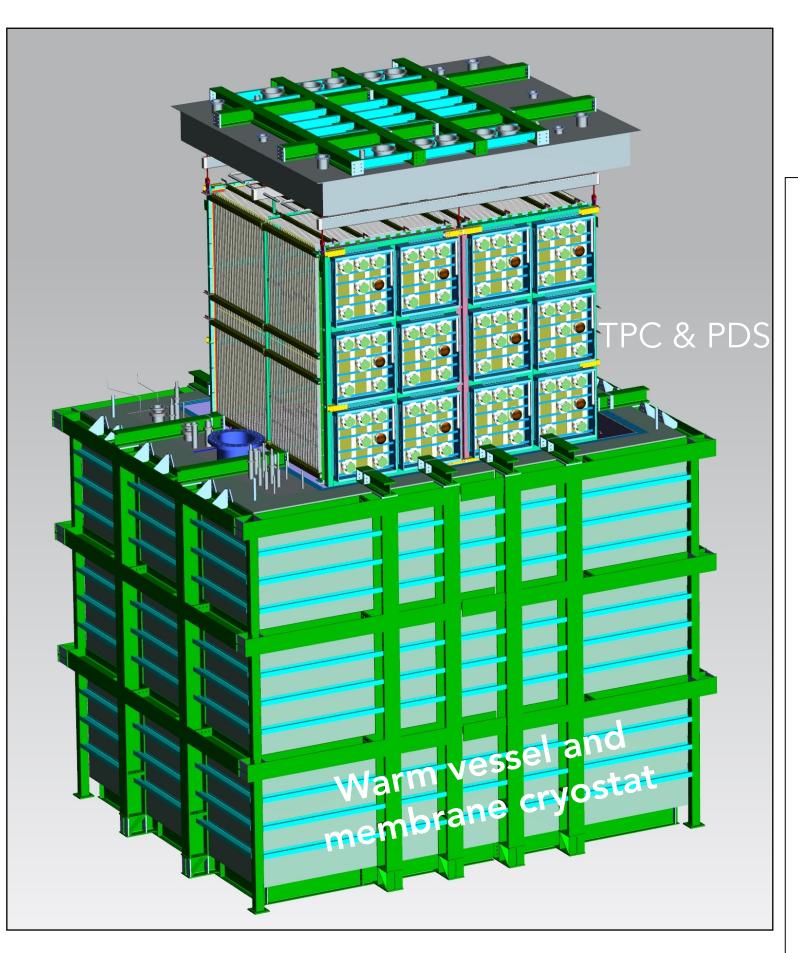


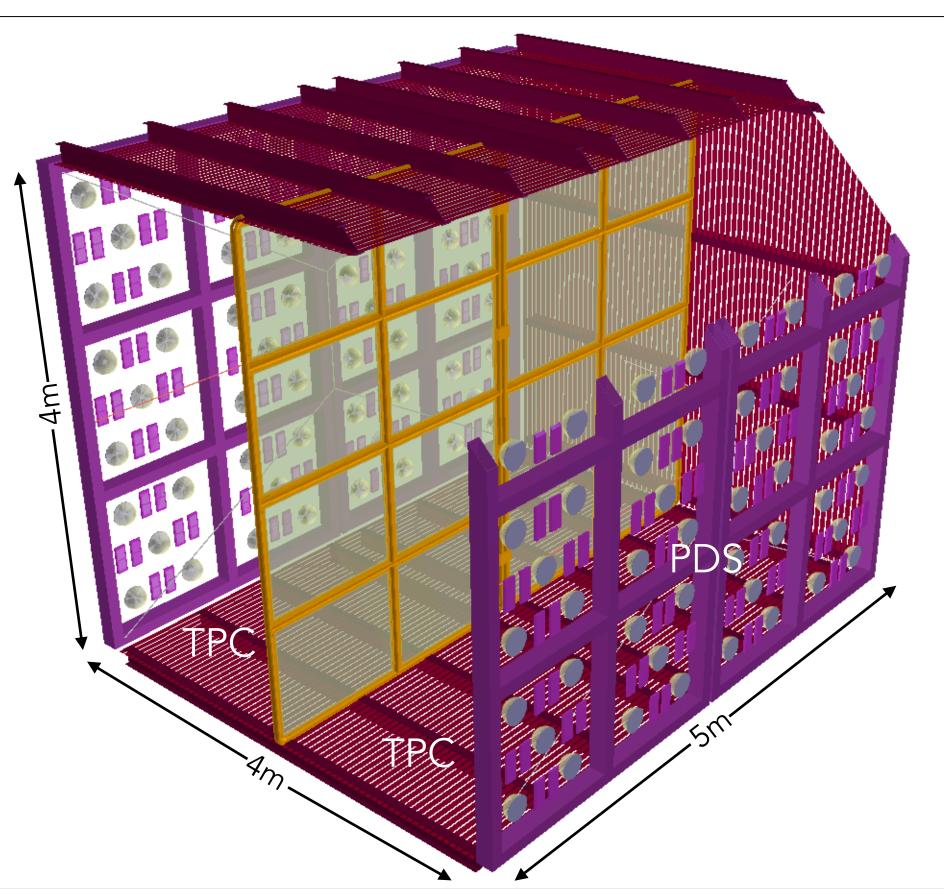




# SBN NEAR DETECTOR: SBND

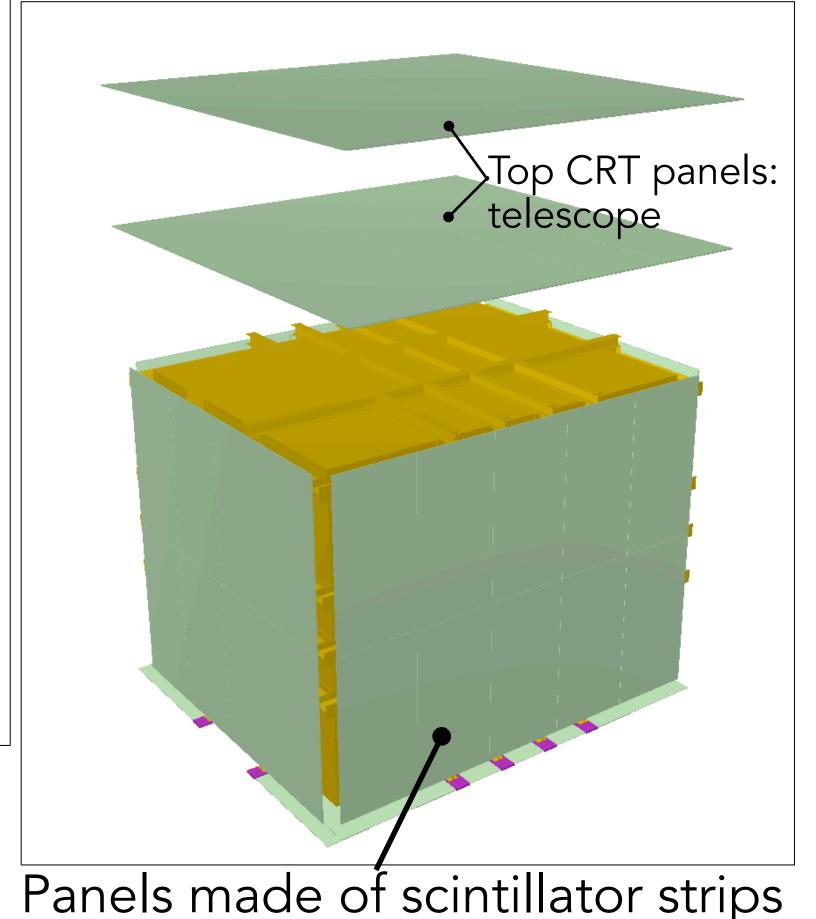






Two Time Projection Chambers and Photon Detection systems

Cryostat surrounded by a Cosmic Ray Tagger system for cosmic ray rejection



# SBND DETECTOR: TPC AND PDS



TPC Cold electronics

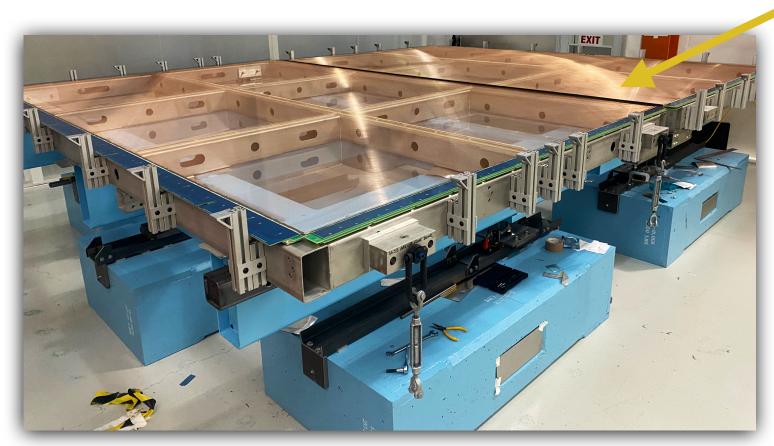


Two Time Projection Chambers **Total dimension**: 4m x 4m x 5m

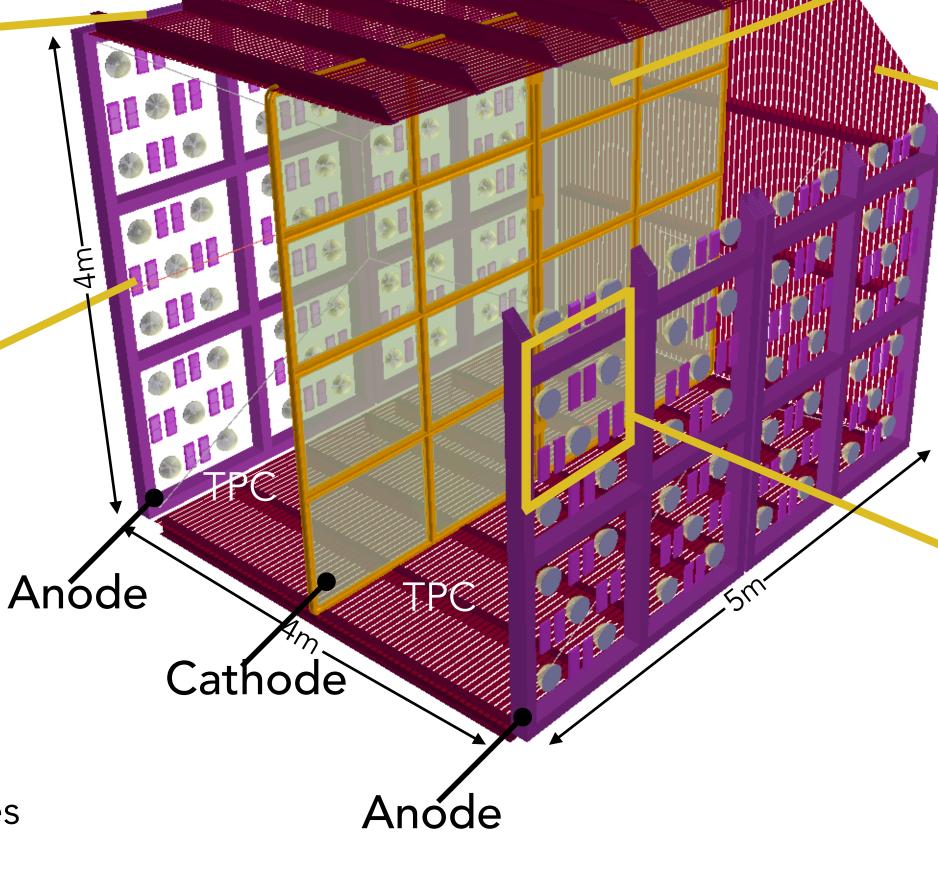


CPA-Cathode covered with TPB coated reflectors

Field Cage



APA-Wire Planes- 3 planes, ~11000 wires



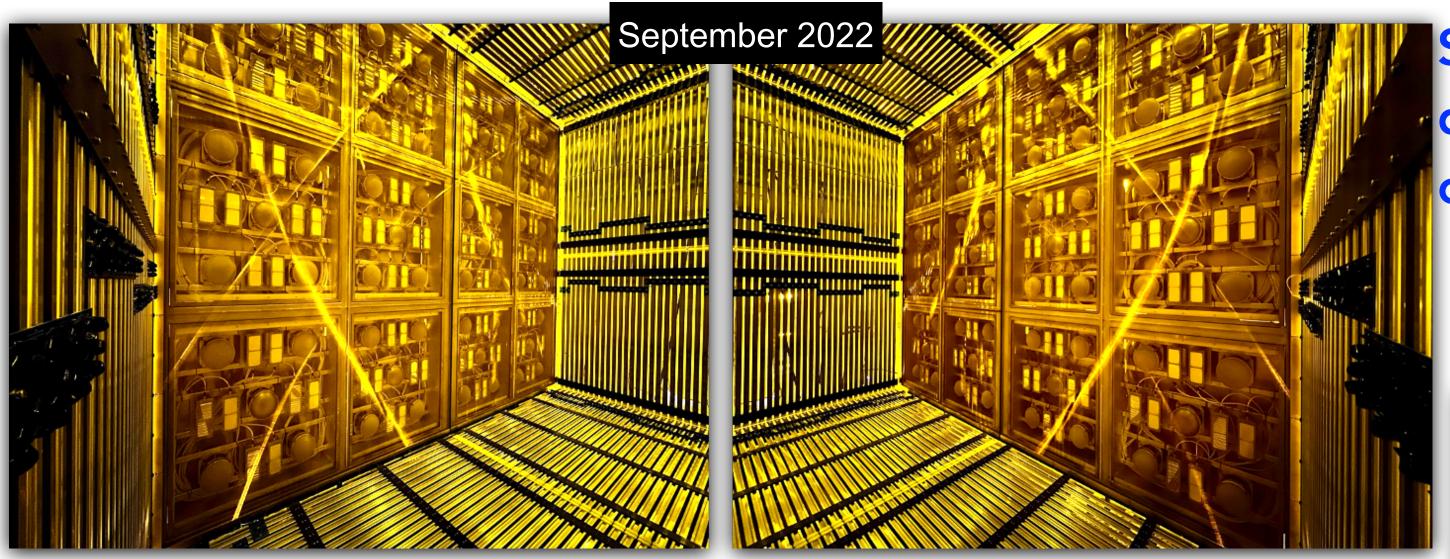




Photon Detection Systems: 120 PMTs, 192 X-Arapucas

# SBND STATUS





SBND detector completed

Membrane cryostat SBND cryostat completed





SBND detector move



SBND Detector
Building or Baseline Comer National Community Community Comer National Community Commun

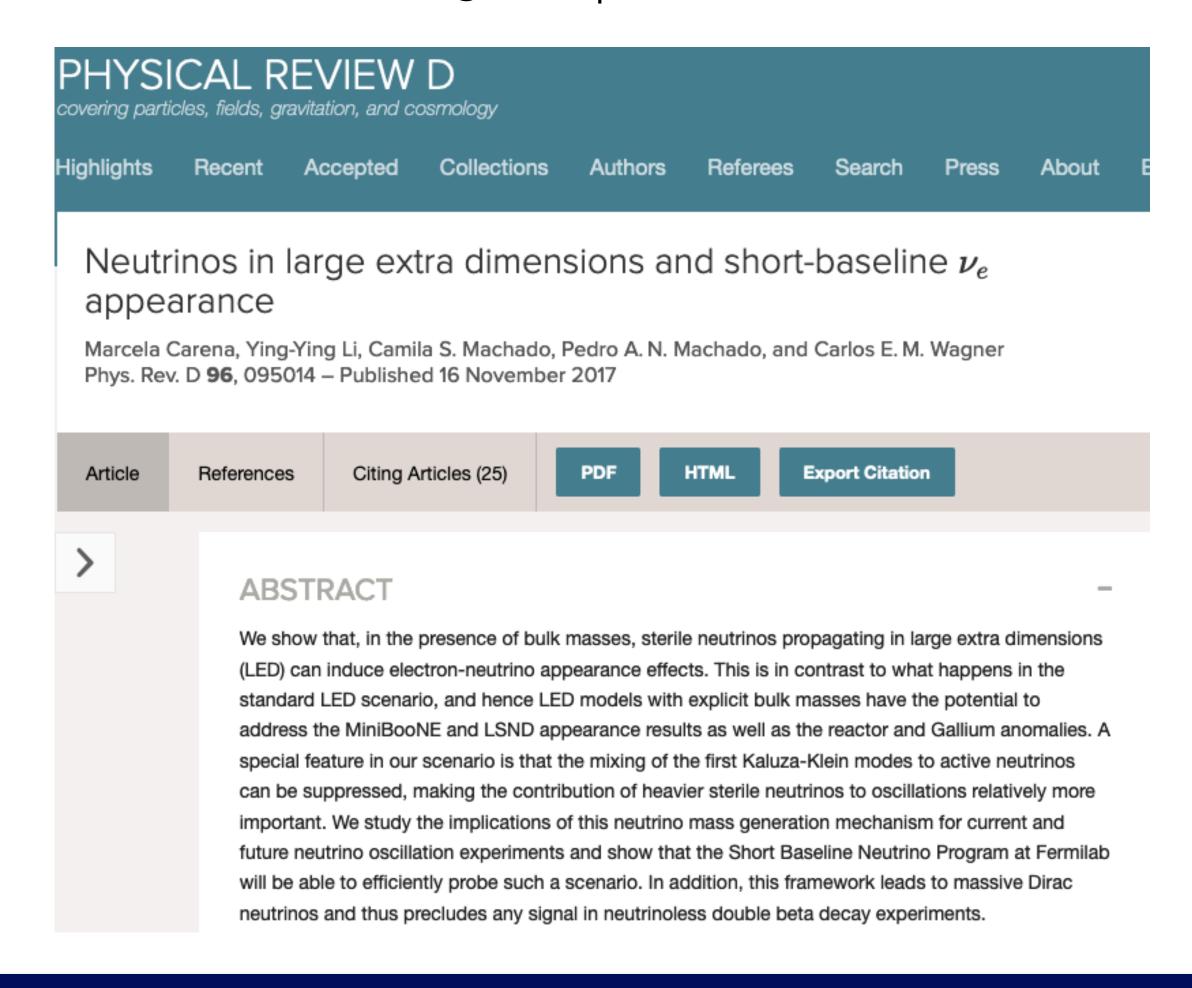
SBND detector into the cryostat

Expected to begin operations in **Fall 2023** 

### JOINT EXPERIMENT-THEORY PROJECTS - IT ALL BEGINS IN 2017...

Soon after I moved to Fermilab in 2014, I was invited to be part of a couple of theory search committees and this gave me the opportunity to start interacting with the Fermilab Theory Division, lead by Marcela.

In **2017**, **Marcela** contacted me to discuss about a neutrino model that, with **Carlos** and Collaborators, they wanted to test against possible Short-Baseline Neutrino (SBN) program sensitivity.



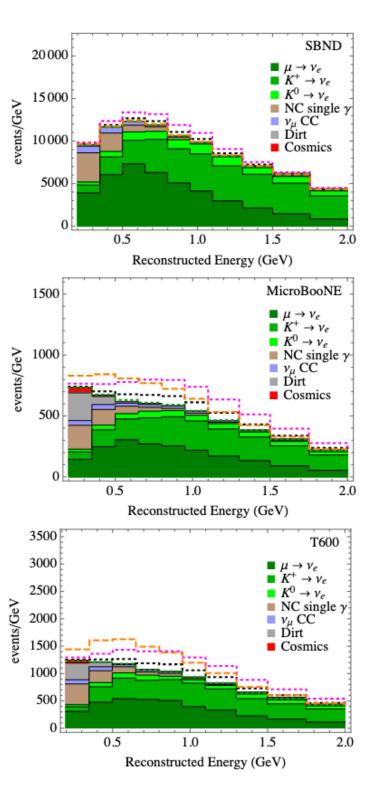
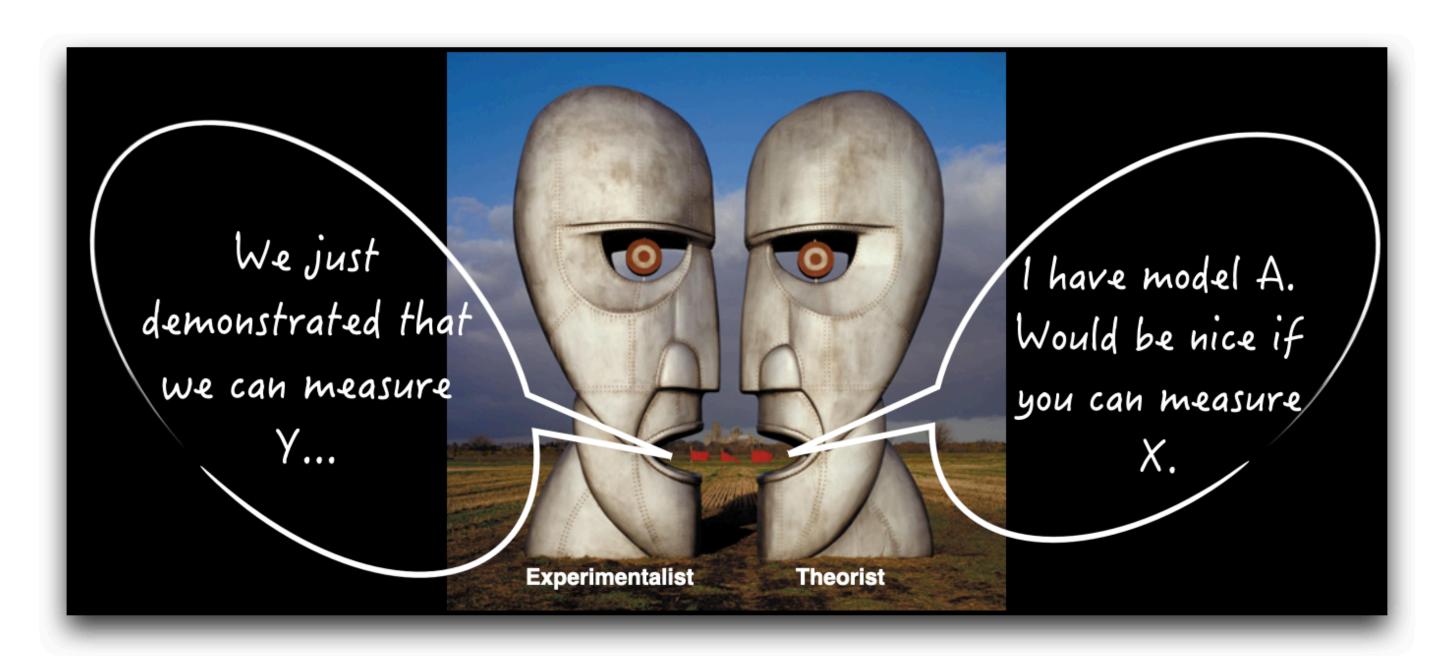


FIG. 10:  $\nu_e$  appearance spectrum at SBN detectors: (top) LAr1-ND, (middle) MicroBooNE, and (bottom) ICARUS-T600. Full data set at these three detectors are assumed (see text). The shaded histograms are the different background components as indicated in the legend (taken from Ref. [32]). Black, green, magenta and orange lines are for points 1, 2, 3 in Table [I] and the best-fit 3+1 sterile neutrino model, respectively.

Appearance spectrum at SBN detectors

### A BRIDGE BETWEEN THEORY AND NEUTRINO EXPERIMENTAL PROGRAM



Fermilab is an ideal place to develop this program

Some informal discussion between neutrino experimentalists and theorists turned into the **SBN-Theory working group** (started in 2018 - a joint experiment+theory effort).

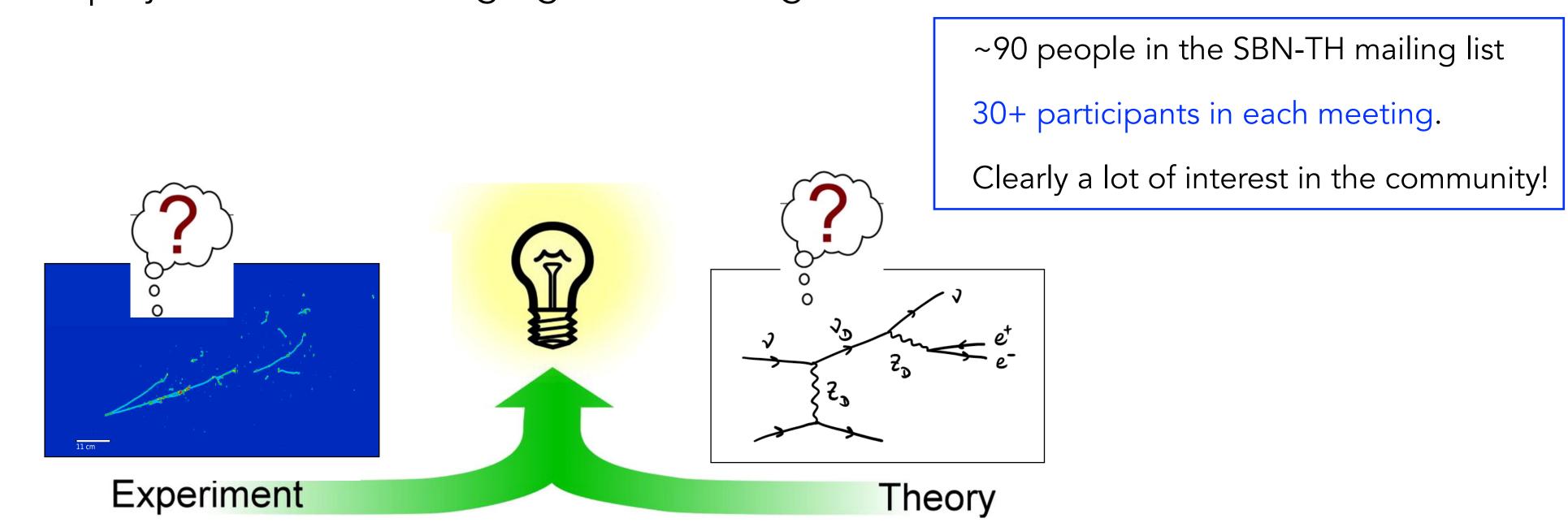
Goal: to exploit possibilities to enrich the physics program of LAr TPC neutrino experiments with **New Physics scenarios**.

Direct, regular and continued collaboration turned in several **successful projects**[a few examples in the following]

# SBN-THEORY WORKING GROUP MEETINGS

<u>Informal meetings (slides+backboard!!)</u>
Organizer: Pedro Machado

- Discuss experimental capabilities of LAr TPC neutrino detectors, current status of event reconstruction, thresholds and resolution.
- Discuss New Physics models.
- Theorists to better know how the LAr TPC technology works and experimentalists to know what to look for.
- Work side-by-side on various projects, understanding signal and background.



Marcela's and Carlos' influential impact on a new way to make physics today!

# SBN PROGRAM: NOT ONLY OSCILLATION PHYSICS!

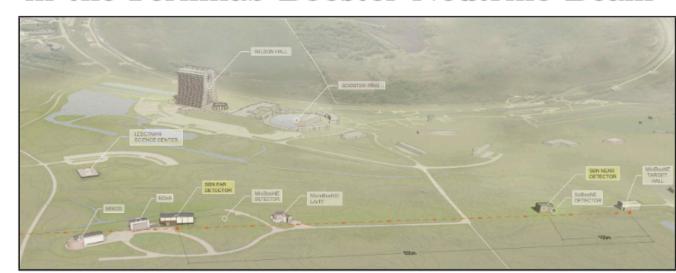
The Short-Baseline Neutrino program proposal (**2014**) only mentioned about another possible search for "New Physics":

"SBN will also be able to search for sub-GeV dark matter (mass less than a few hundred MeV) by running in beam-dump mode, where the 8 GeV proton beam is steered above the beryllium target and into the 50 m (or 25 m) downstream absorber.

It is an evolving landscape!

Alternative potential explanations to the SBL neutrino anomalies from more recently emerging new physics scenarios from theory.

A Proposal for a Three Detector Short-Baseline Neutrino Oscillation Program in the Fermilab Booster Neutrino Beam



The ICARUS, SBND and MicroBooNE collaborations arXiv:1503.01520, January 2014

Designed for **Sterile Neutrino** searches
Same neutrino beam, nuclear target and detector technology:
reducing systematic uncertainties to the % level



Joint experiment+theory effort evolved into a review paper (2019), where a range of ideas [now under active development] for using the SBN detectors and beam to search for signatures of **New Physics** are discussed!

Annual Review of Nuclear and Particle Science

The Short-Baseline Neutrino Program at Fermilab

Pedro A.N. Machado,<sup>1</sup> Ornella Palamara,<sup>1</sup> and David W. Schmitz<sup>2</sup>

<sup>1</sup>Fermi National Accelerator Laboratory, Batavia, Illinois 60510, USA
 <sup>2</sup>Enrico Fermi Institute and Department of Physics, University of Chicago, Chicago, Illinois 60637, USA; email: dwschmitz@uchicago.edu

Annu. Rev. Nucl. Part. Sci. 69 363-387 (2019)

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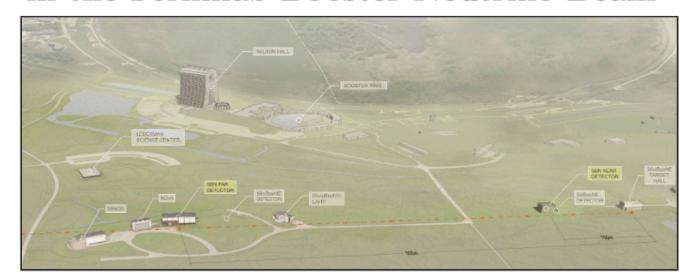
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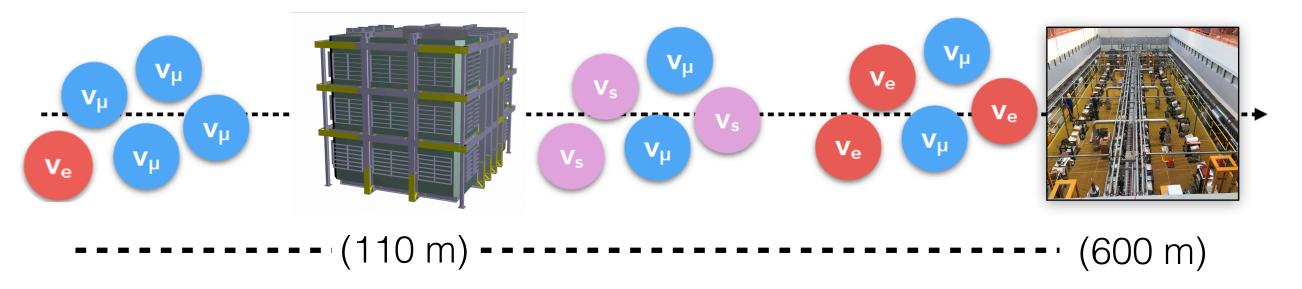
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Same neutrino beam, nuclear target and detector technology:
reducing systematic uncertainties to the % level

Pedro Machado & OP - Neutrino University (2019)

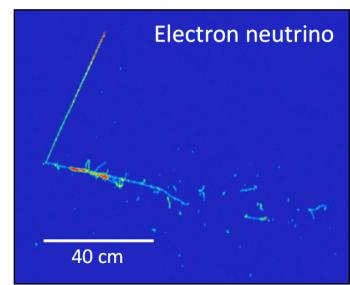


### SBN - A BROAD PHYSICS PROGRAM

eV-scale sterile neutrinos: searches for physics beyond the threeneutrino mixing with multiple-detectors at different baselines.



Neutrino-argon interactions: with an order of magnitude more data than is currently available.

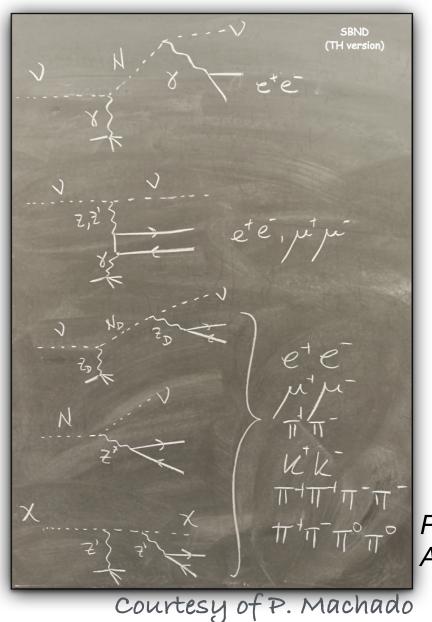


Up to 7000  $\nu$  events/per day in SBND!

#### We have expanded the physics program!

New physics scenarios: study alternative explanations of the short-baseline anomalies and other Beyond Standard Model scenarios.

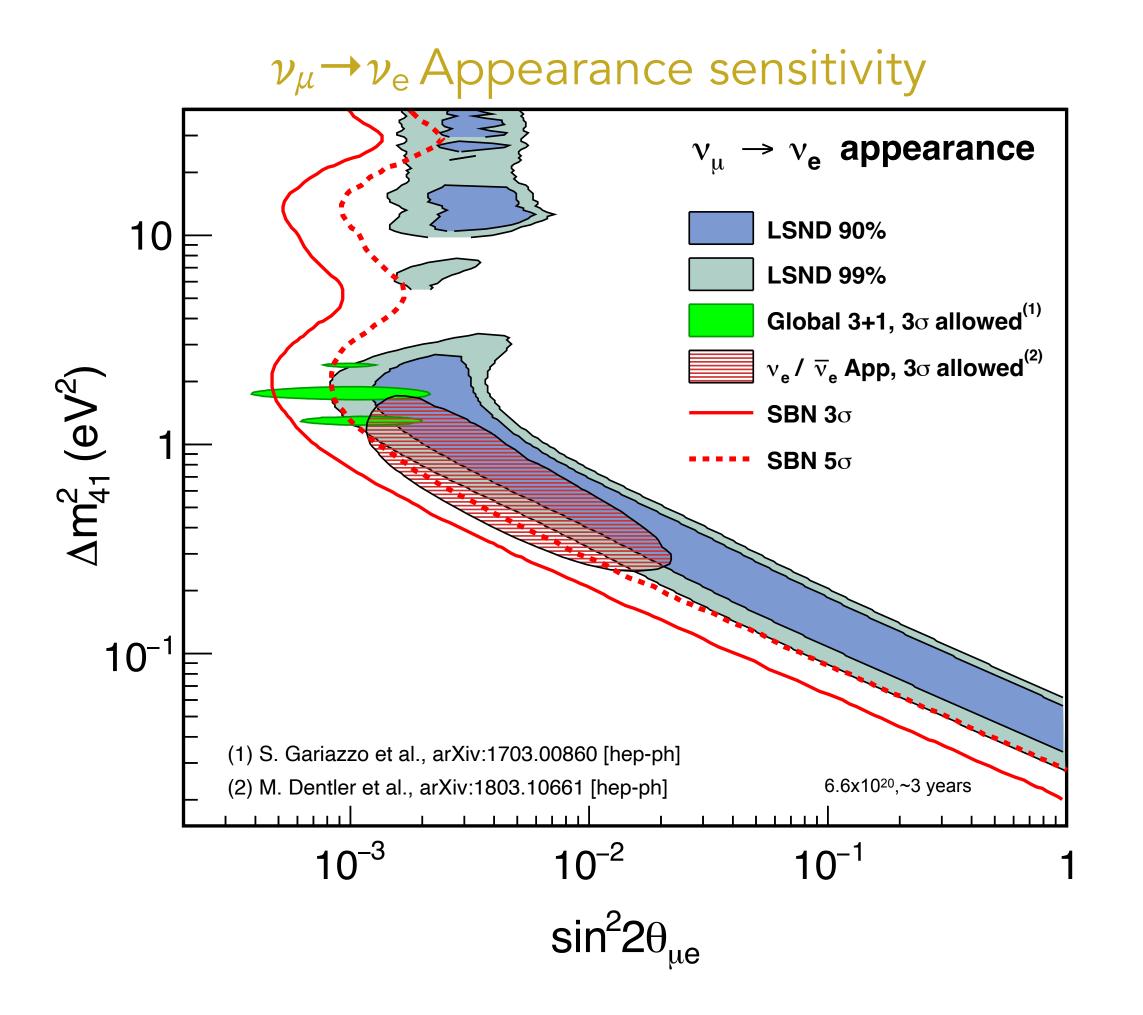
Many ideas for new searches emerging from collaboration with theory colleagues.

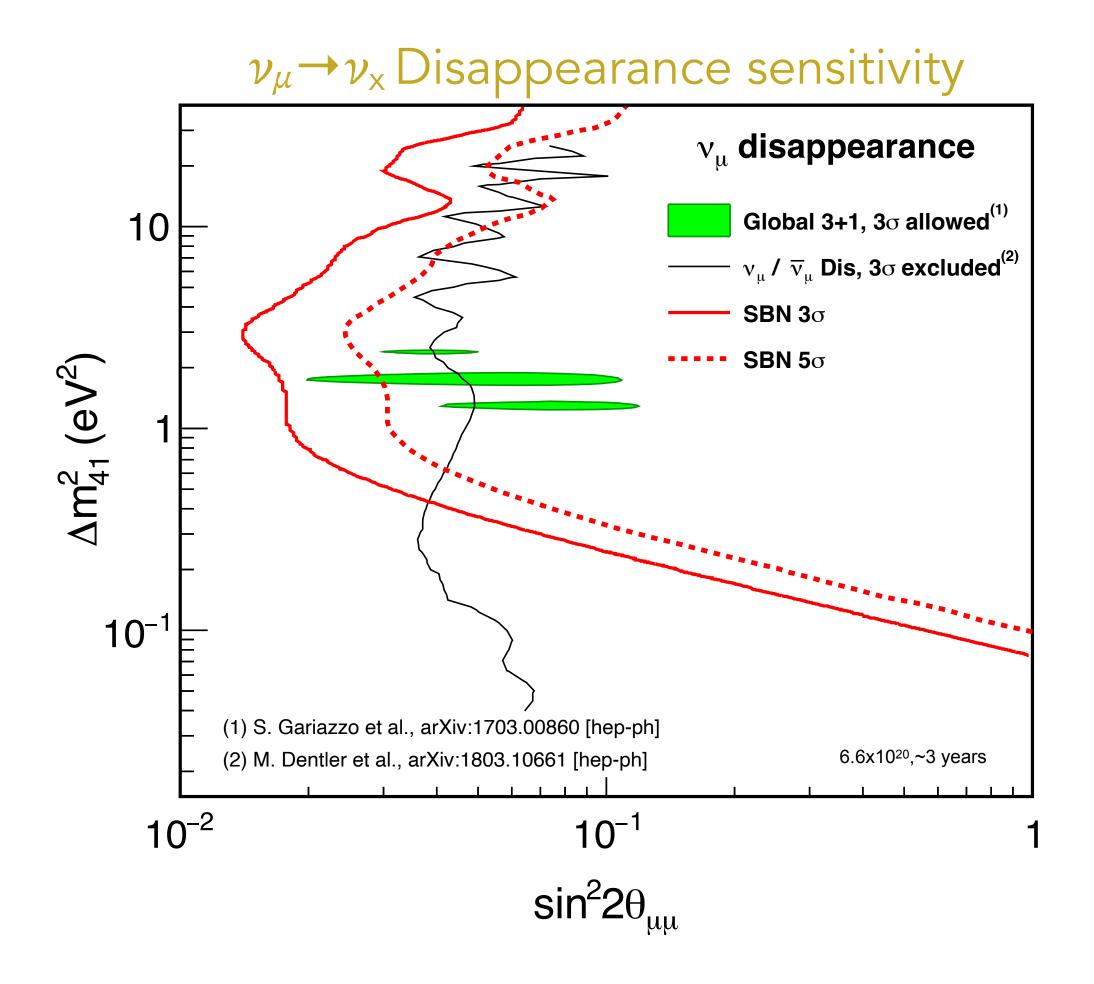


P.Machado, O.P., D. Schmitz, Annu. Rev. Nucl. Part. Sci. 69 363-387 (2019)

# SBN STERILE NEUTRINO SENSITIVITIES

P. Machado, O.P., D. Schmitz: Annual Rev. Nucl. Part. Sci. 69 363-387 (2019)



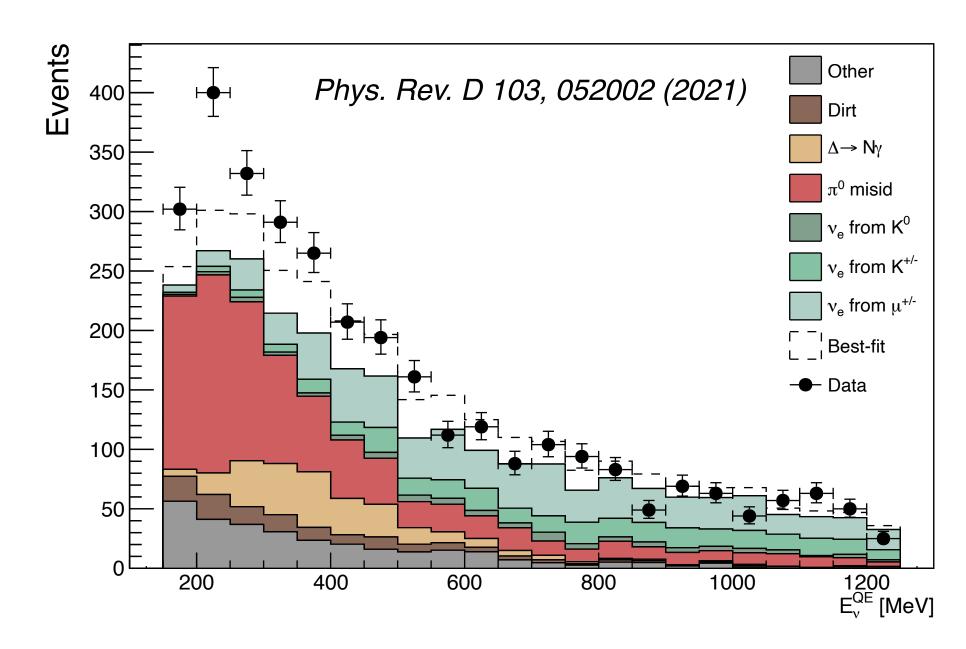


SBN can cover the parameters allowed by past anomalies at  $5\sigma$  significance

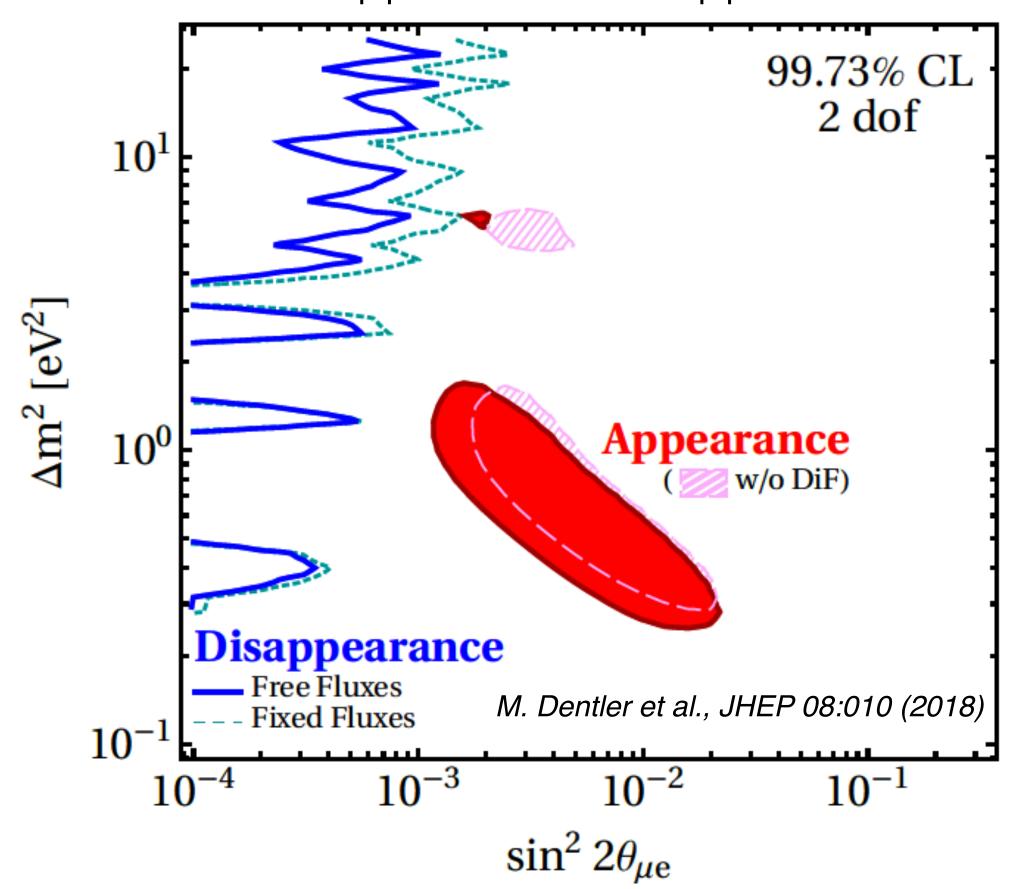
### THE LIGHT STERILE NEUTRINO EXPERIMENTAL LANDSCAPE

A 4.7  $\sigma$  tension arises when combining  $\nu_{\rm e}$  appearance and  $\nu_{\mu}$  disappearance data sets.



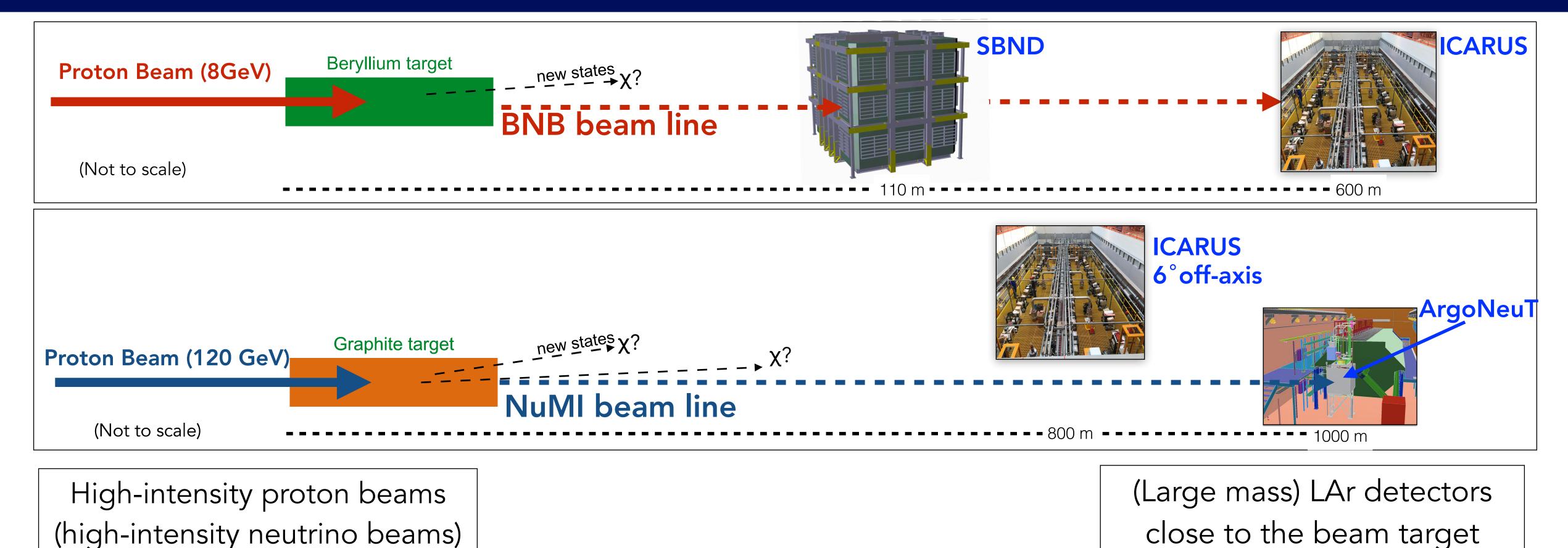


Limits from disappearance and appearance allowed region



Alternative (Beyond Standard Model) explanations exist that could explain the MiniBooNE anomaly

# SIGNATURES FOR NEW PHYSICS IN SBN AND ARGONEUT

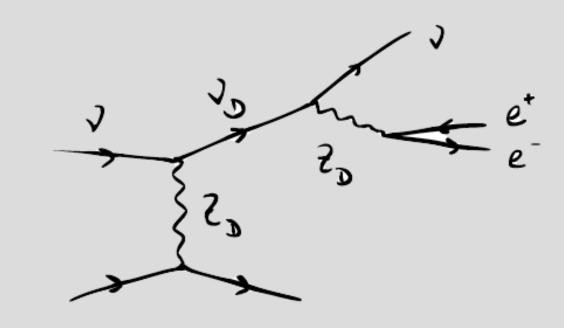


Opportunities to **probe signatures** for **new physics scenarios** in the **neutrino sector** (modifications to the neutrino oscillation paradigm) and **beyond** (novel experimental signatures produced in the beam target )

### JOINT EXPERIMENT-THEORY PROJECTS - NEW PHYSICS SEARCHES

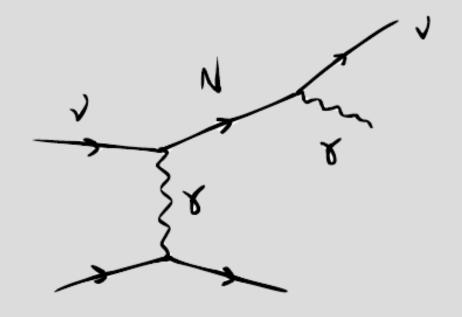
Alternative explanations
of the MiniBooNE excess
and other Beyond
Standard Model scenarios

#### Dark Neutrinos



Bertuzzo Jana Machado Zukanovich PRL 2018, PLB 2019 Arguelles Hostert Tsai PRL 2019 Ballett Pascoli Ross-Lonergan PRD 2019 Ballett Hostert Pascoli PRD 2020

#### Transition Magnetic Moment



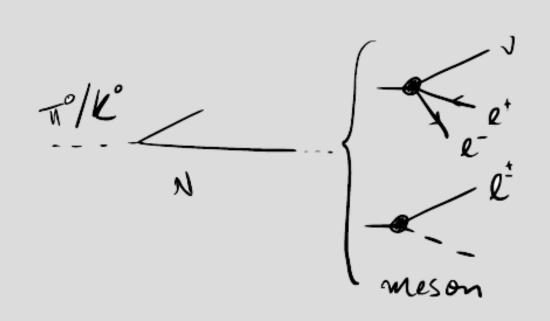
Gninenko PRL 2009 Coloma Machado Soler Shoemaker PRL 2017 Atkinson et al 2021 Vergani et al 2021

#### Axion-like Particles



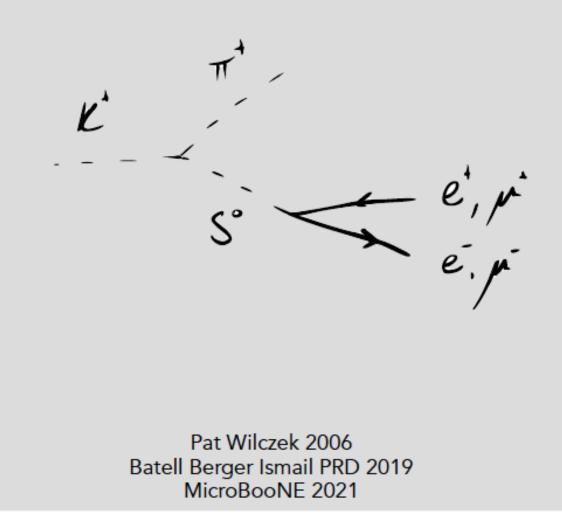
Kelly Kumar Liu PRD 2021 Brdar et al PRL 2021

#### Heavy Neutral Leptons

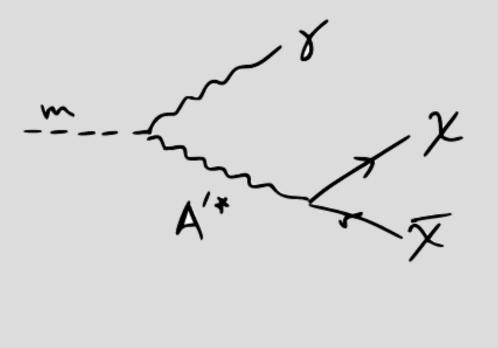


Ballett Pascoli Ross-Lonergan JHEP 2017 Kelly Machado PRD 2021

#### Higgs Portal Scalar

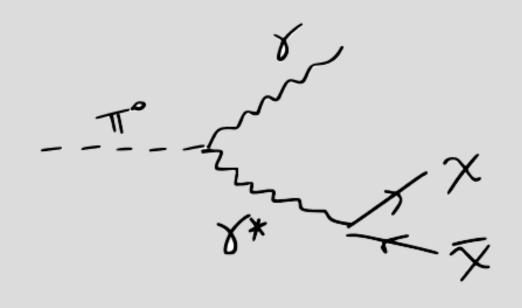


#### Light Dark Matter



Romeri Kelley Machado PRD 2019

#### Millicharged Particles



Magill, Plestid, Pospelov, Tsai, PRL 2019 Harnik Liu Palamara, JHEP 2019

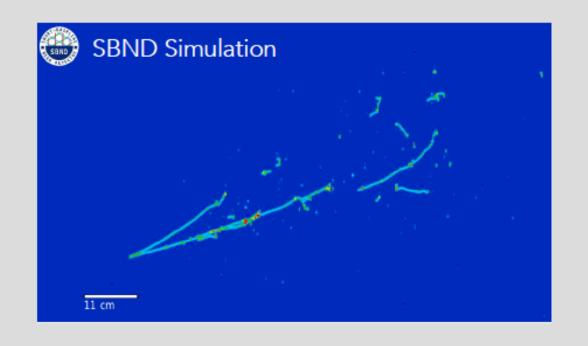
Note: not an exhaustive list!

### Joint Experiment-Theory projects - New Physics searches in SBND



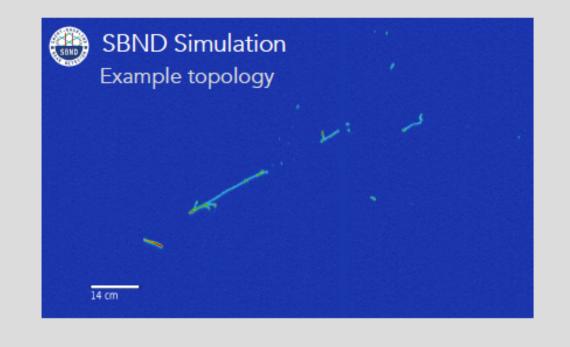
SBN(D) experiment(s) will explore the landscape and test not only the sterile neutrino hypothesis, but also other New Physics models

#### Dark Neutrinos



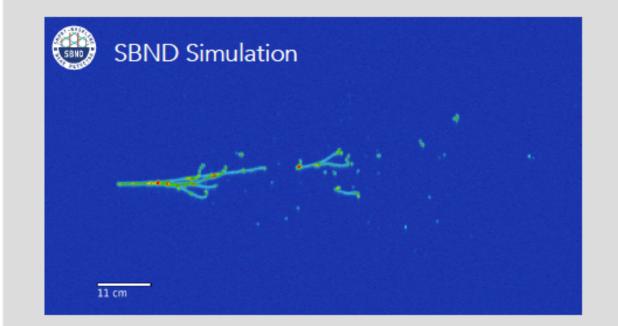
e+e- pair w/ or w/o hadronic activity

#### Transition Magnetic Moment



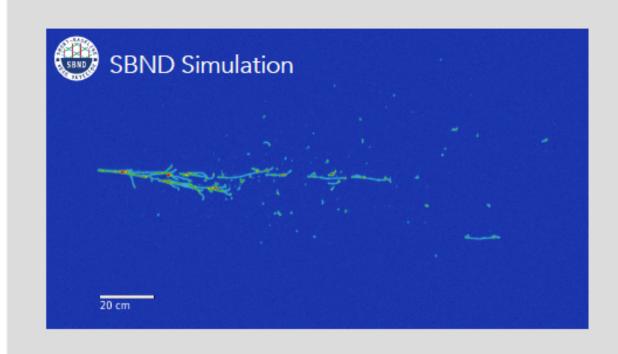
photon shower and hadronic activity

#### Axion-like Particles



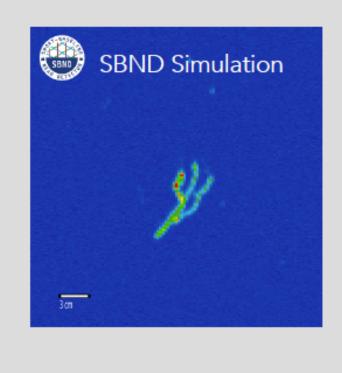
high-energy e+e-, μ+μ-

#### Heavy Neutral Leptons



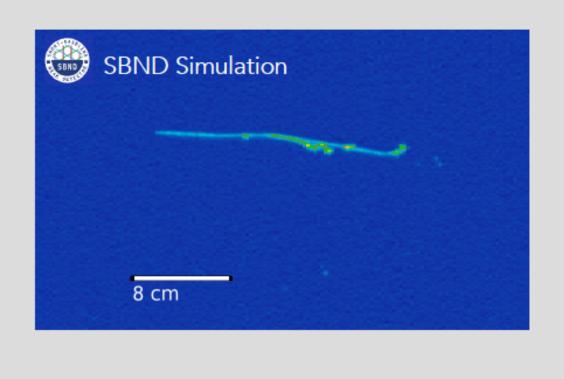
e+e-, μ+μ-, μπ

#### Higgs Portal Scalar



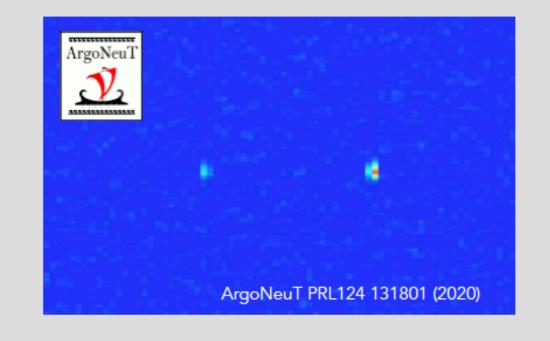
e+e-, μ+μ-, no hadronic activity

#### Light Dark Matter



electron scattering

#### Millicharged Particles

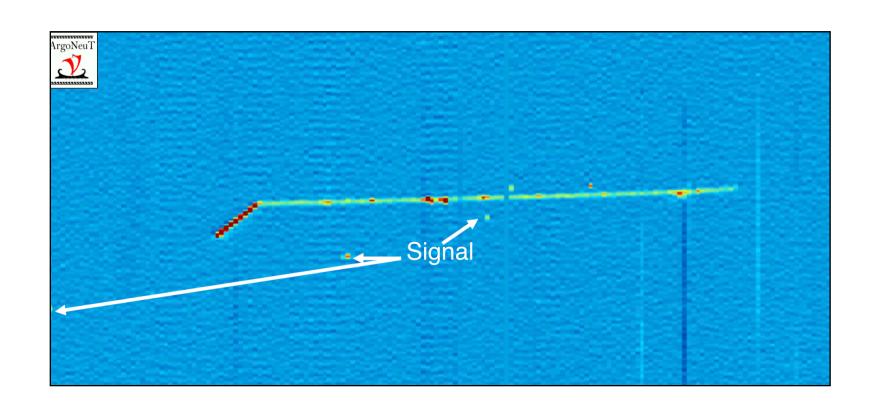


blips/faint tracks

Note: not an exhaustive list!

### Joint Experiment-Theory project - Search for Millicharged Particles

Motivated by the fact that LAr TPC's have demonstrated to be able to detect and reconstruct

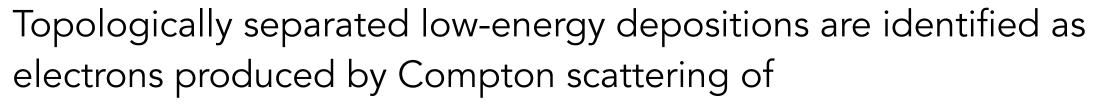


(sub-)MeV energy depositions

300 KeV threshold In ArgoNeuT



R. Acciarri et al., PRD 99, 012002 (2019)



- de-excitation photons from the target nucleus and
- photon produced by neutron inelastic interactions

We started a common project, which evolved into two papers:

Roni Harnik, Zhen Liu, and O. P., "Millicharged particles in liquid argon neutrino experiments", JHEP 07, 170 (2019)

ArgoNeuT Collaboration + 2 theorists (R. Harnik and Z. Liu)
R. Acciarri et al., PRL124 131801 (2020)

Total Reconstructed Energy in an Event

Data
FLUKA
FLUKA Stat. Error
FLUKA Total Error

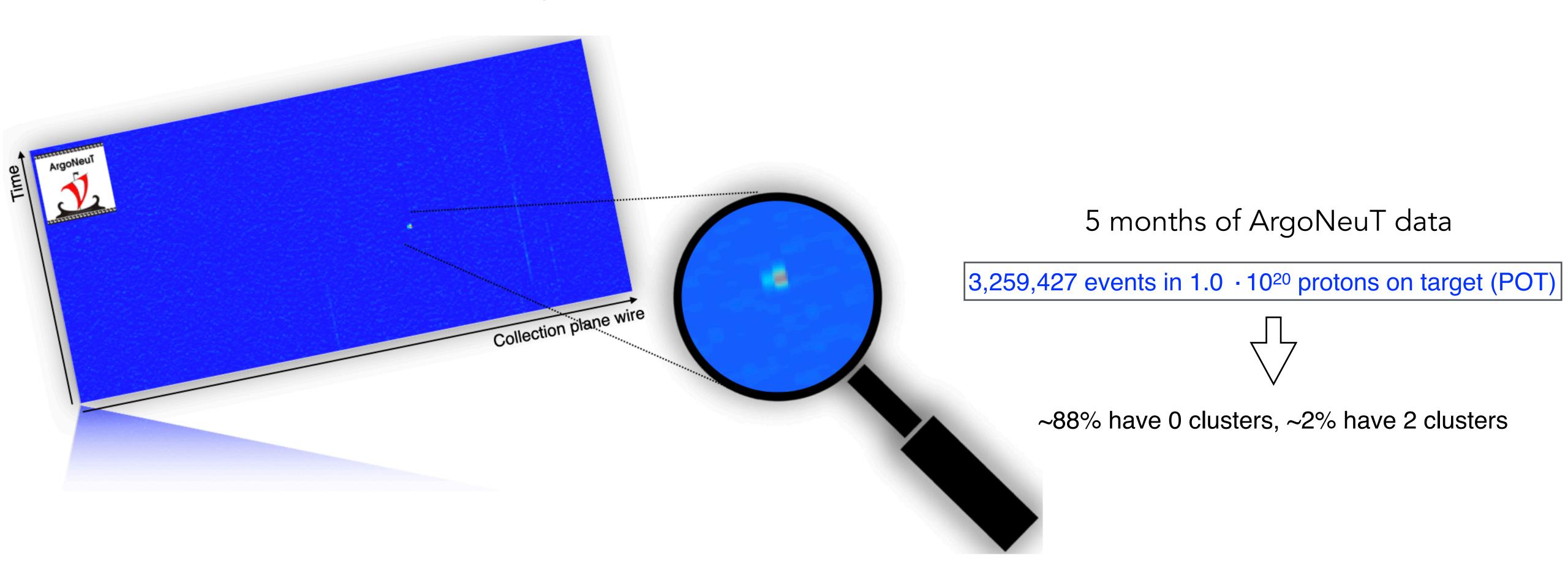
Total Energy (MeV)

Experimental paper

### JOINT EXPERIMENT+THEORY PROJECT: USE OF "EMPTY" EVENTS!



The vast majority of NuMI beam spills delivered did not produce a neutrino interaction within the ArgoNeuT TPC (due to the limited size of the detector)



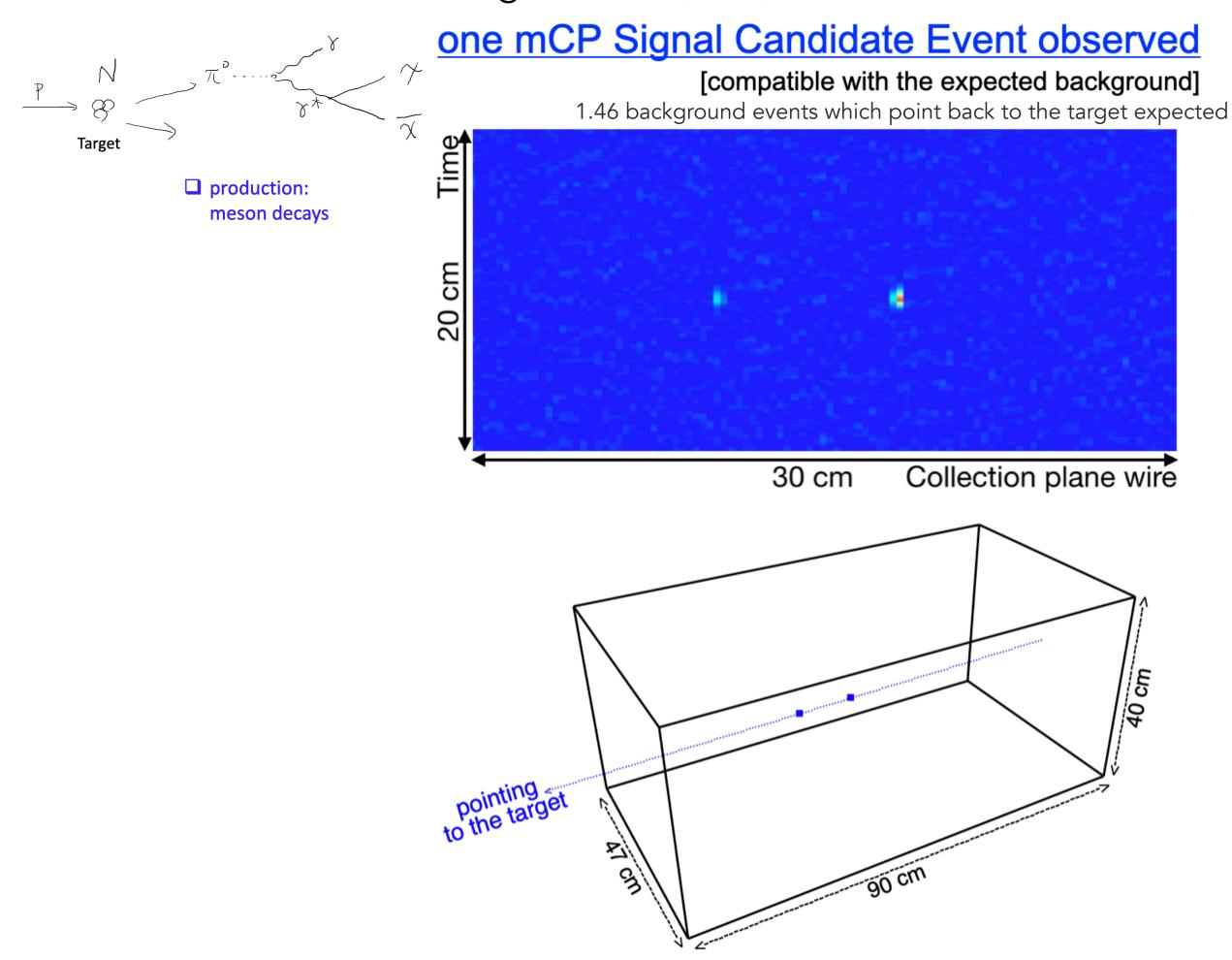
We searched for the possible presence of mCPs in these empty events.

# JOINT EXPERIMENT-THEORY PROJECT



### First search for Millicharged Particles in LAr TPC

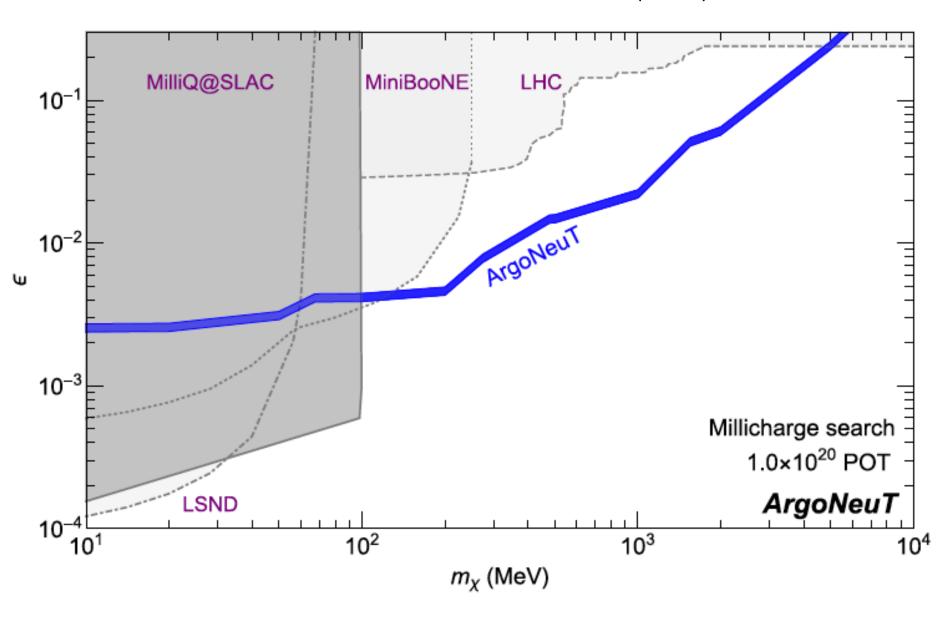
mCP have an electric charge  $Q = \varepsilon \cdot e \ (\varepsilon \ll 1)$ 



Low energy threshold (300 KeV) is the key!

Roni Harnik, Zhen Liu, and O. P., "Millicharged particles in liquid argon neutrino experiments", JHEP 07, 170 (2019)

ArgoNeuT Collaboration + 2 theorists (R. Harnik and Z. Liu)
R. Acciarri et al., PRL124 131801 (2020)



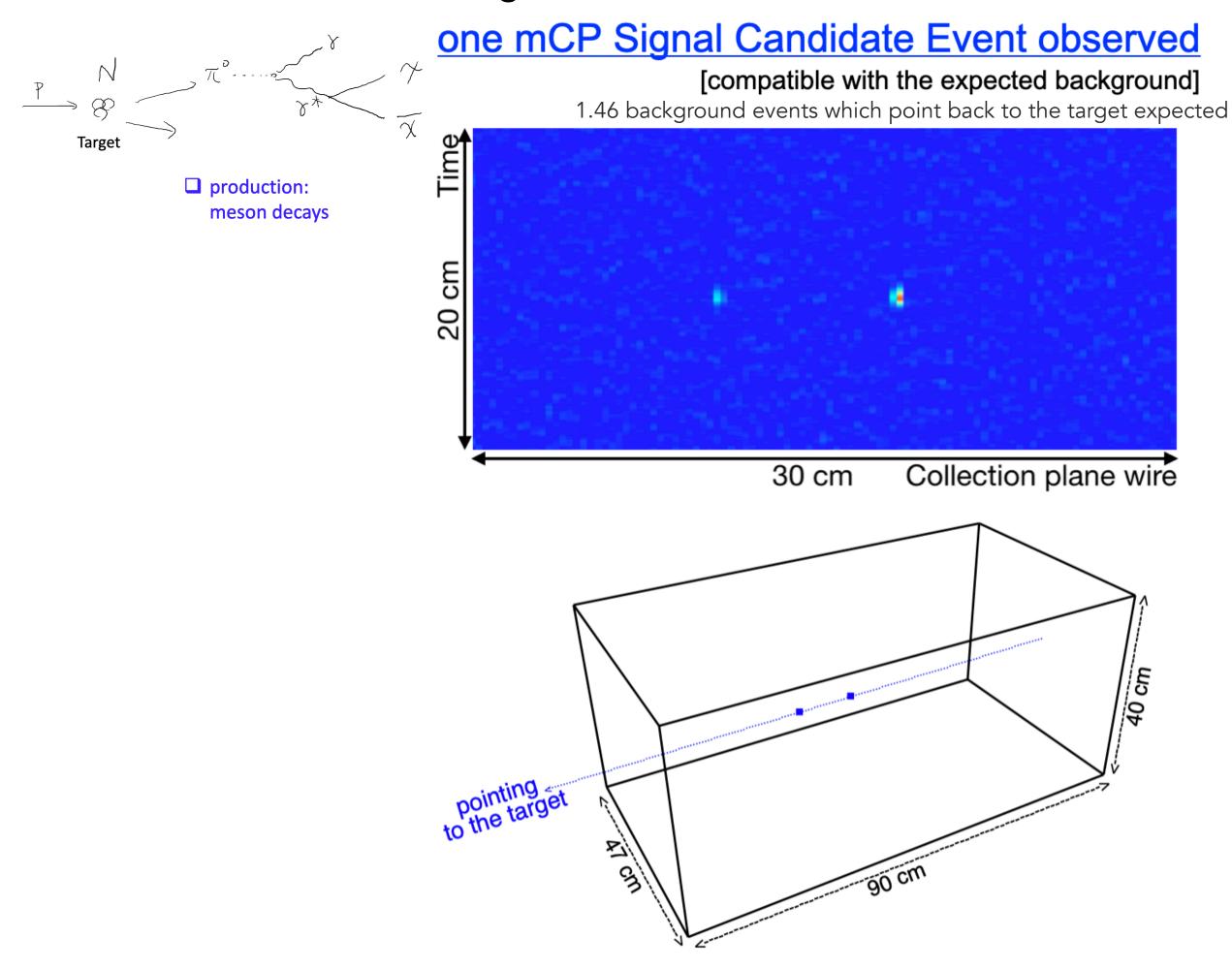
Leading constraints in unexplored parameter region!

# JOINT EXPERIMENT-THEORY PROJECT



### First search for Millicharged Particles in LAr TPC

mCP have an electric charge  $Q = \varepsilon \cdot e \ (\varepsilon \ll 1)$ 

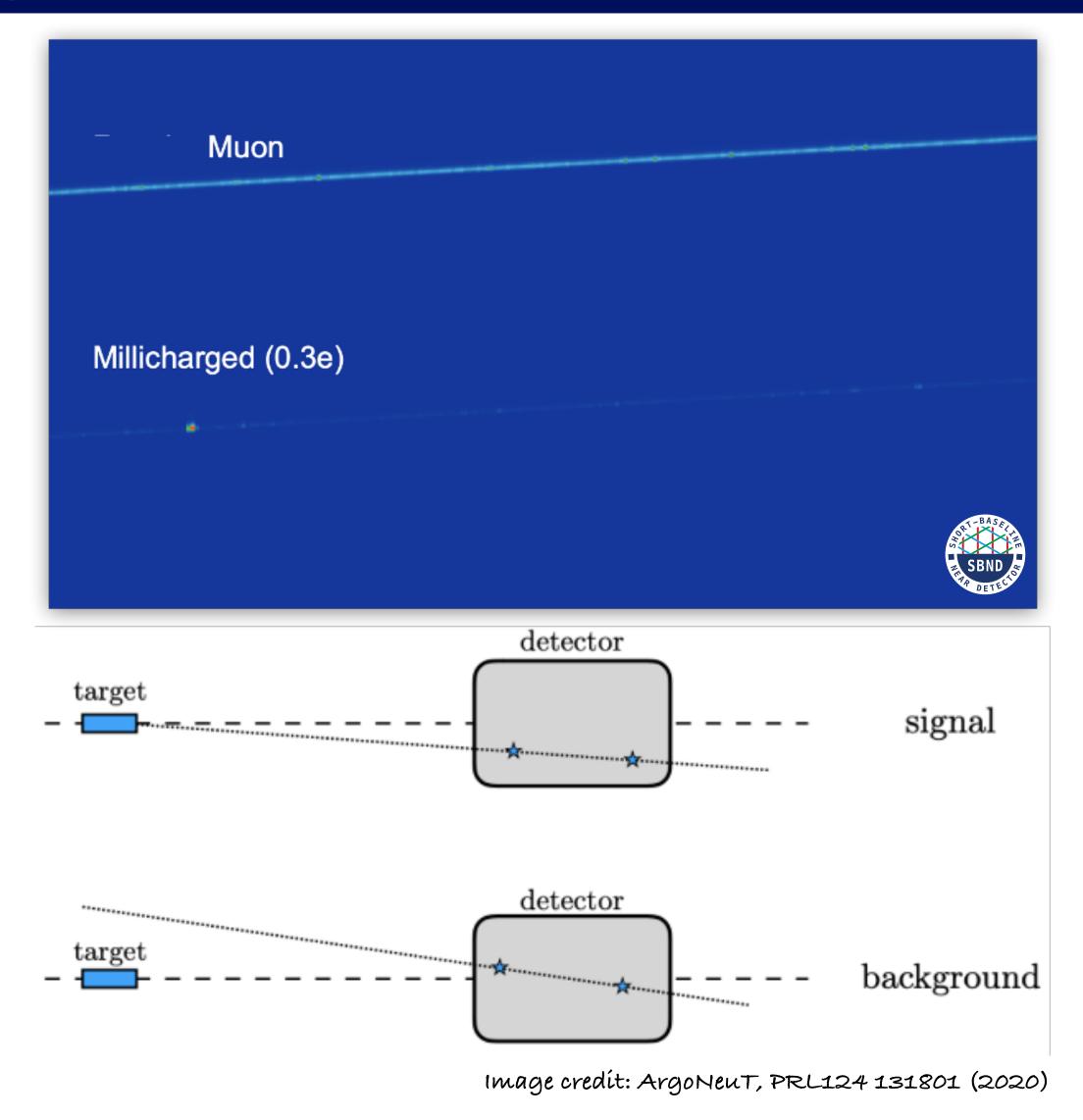


Low energy threshold (300 KeV) is the key!

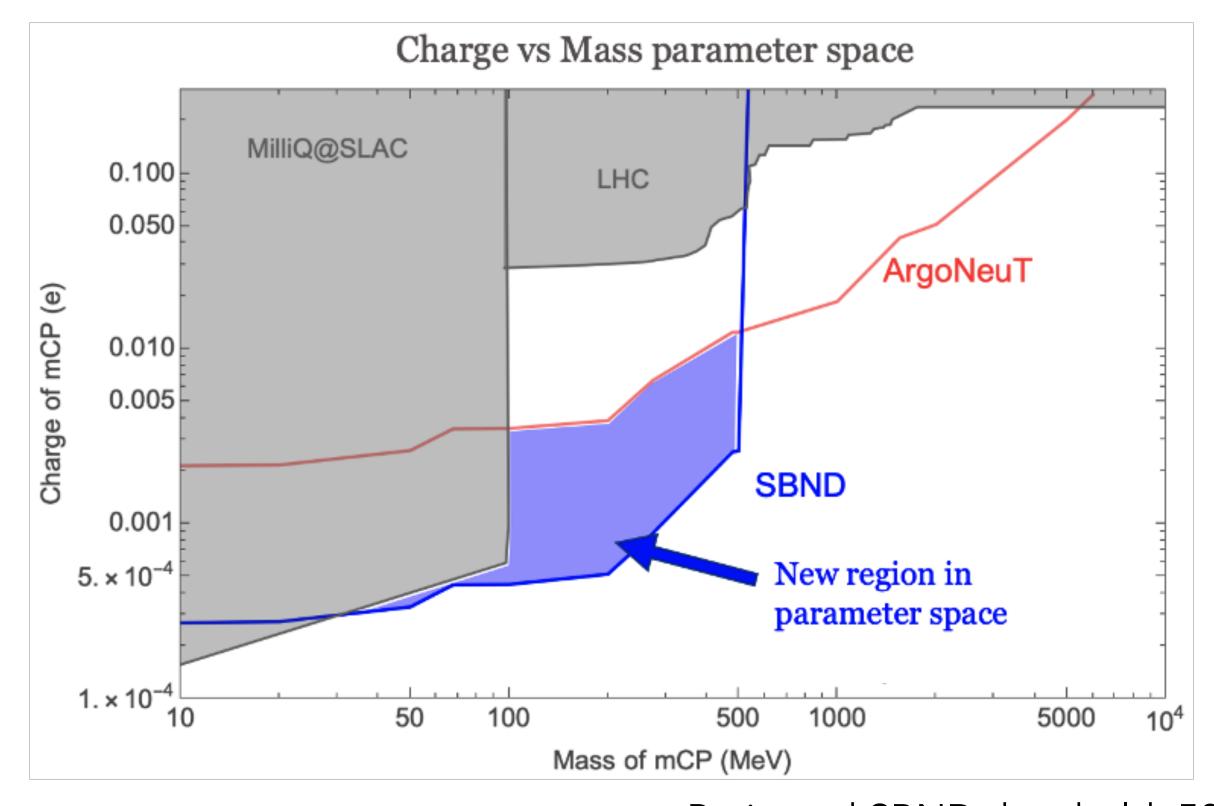
Roni Harnik & OP - Wine and Cheese seminar (2020)



# SEARCH FOR MILLICHARGED PARTICLES IN SBND



Millicharged particles would appear in SBND as **blips** or **faint tracks** pointing back to the target.



To reduce background: search for double hits events aligned with the target

Projected SBND threshold: 50 keV

# JOINT EXPERIMENT-THEORY PROJECT

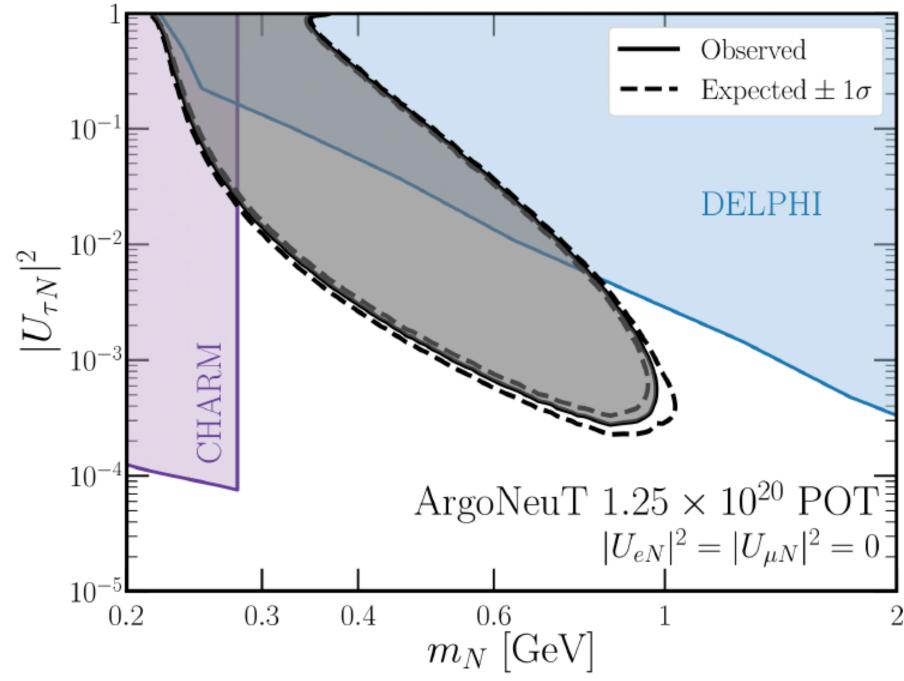


### First search for Heavy neutral Leptons $N \rightarrow \nu \mu^+ \mu^-$ in LAr TPC

Consider tau-coupled scenario, i.e.  $|U_{\tau N}|^2 \neq 0$  and  $|U_{eN}|^2 = |U_{\mu N}|^2 = 0$ Assuming HNL production predominately from  $\tau^{\pm}$  decay\*:  $D/D_s$  decay to  $\tau$ , that subsequently decay to HNLs  $\tau^{\pm} \rightarrow N X^{\pm}$  (X<sup>±</sup> is a SM particle e.g.  $\pi^{\pm}$ ) HNL decay  $N \rightarrow \nu \mu^+ \mu^-$ *Two-track* Event MINOS-ND ArgoNeuT Ν Double-MIP Event MINOS-ND ArgoNeuT ArgoNeuT Ν Minos-ND

**0 events observed in the data,** consistent with background expectation of 0.4±0.2 event

ArgoNeuT Collaboration + 2 theorists (K. Kelly and A. de Gouvêa)
R. Acciarri et al., PRL 127 121801 (2021)



Significant increase in the parameter space exclusion region! New exclusion limits for tau-coupled HNLs with  $m_N=280-970$  MeV.

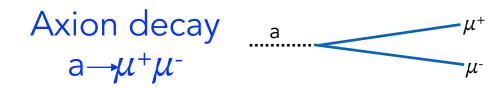
\*For details see: P. Coloma et al. Eur. Phys. J. C, 81(1):78, 2021

# JOINT EXPERIMENT-THEORY PROJECT

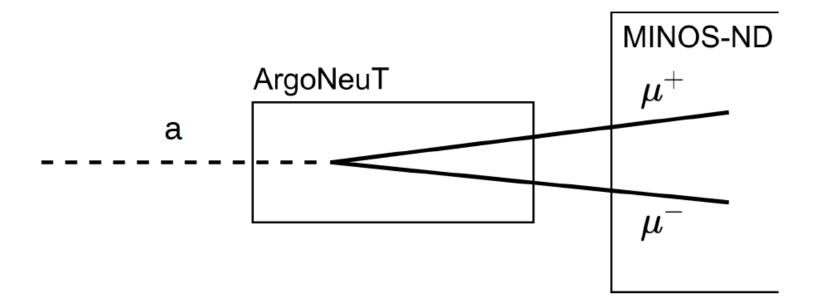


### First search for Heavy QCD Axions in LAr TPC

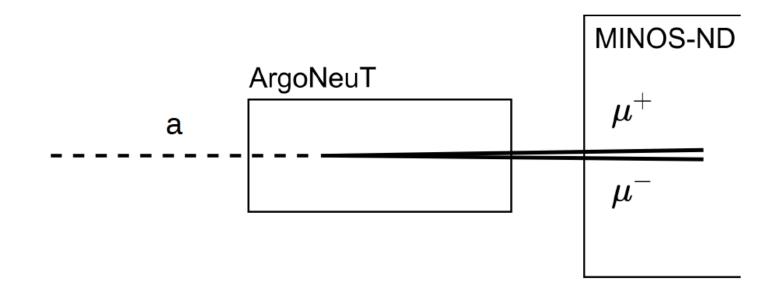
Heavy QCD axions production from  $\pi^0$ ,  $\eta$  and  $\eta'$  mesons\* and decay to ee,  $\mu\mu$ ,  $\gamma\gamma$  + hadronic modes. Contributions depend on axion-lepton coupling, cl: two benchmark scenarios cl = 1/36 and cl = 1/100.



*Two-track* Event



Double-MIP Event

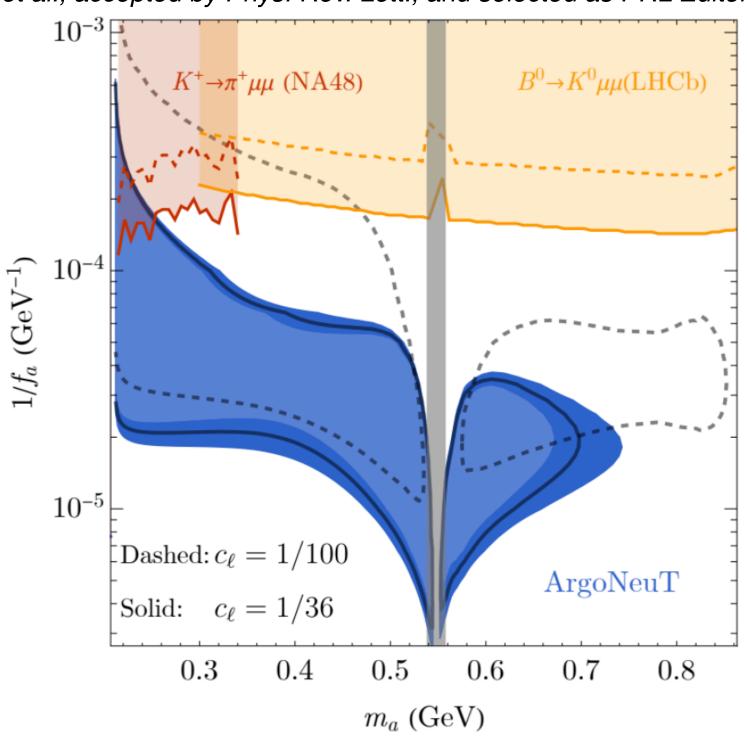


\*For details see: K. Kelly, S. Kumar and Z. Liu Phys. Rev. D 103 (2021) 9, 095002

**0 events observed in the data,** consistent with background expectation of 0.1±0.1 event

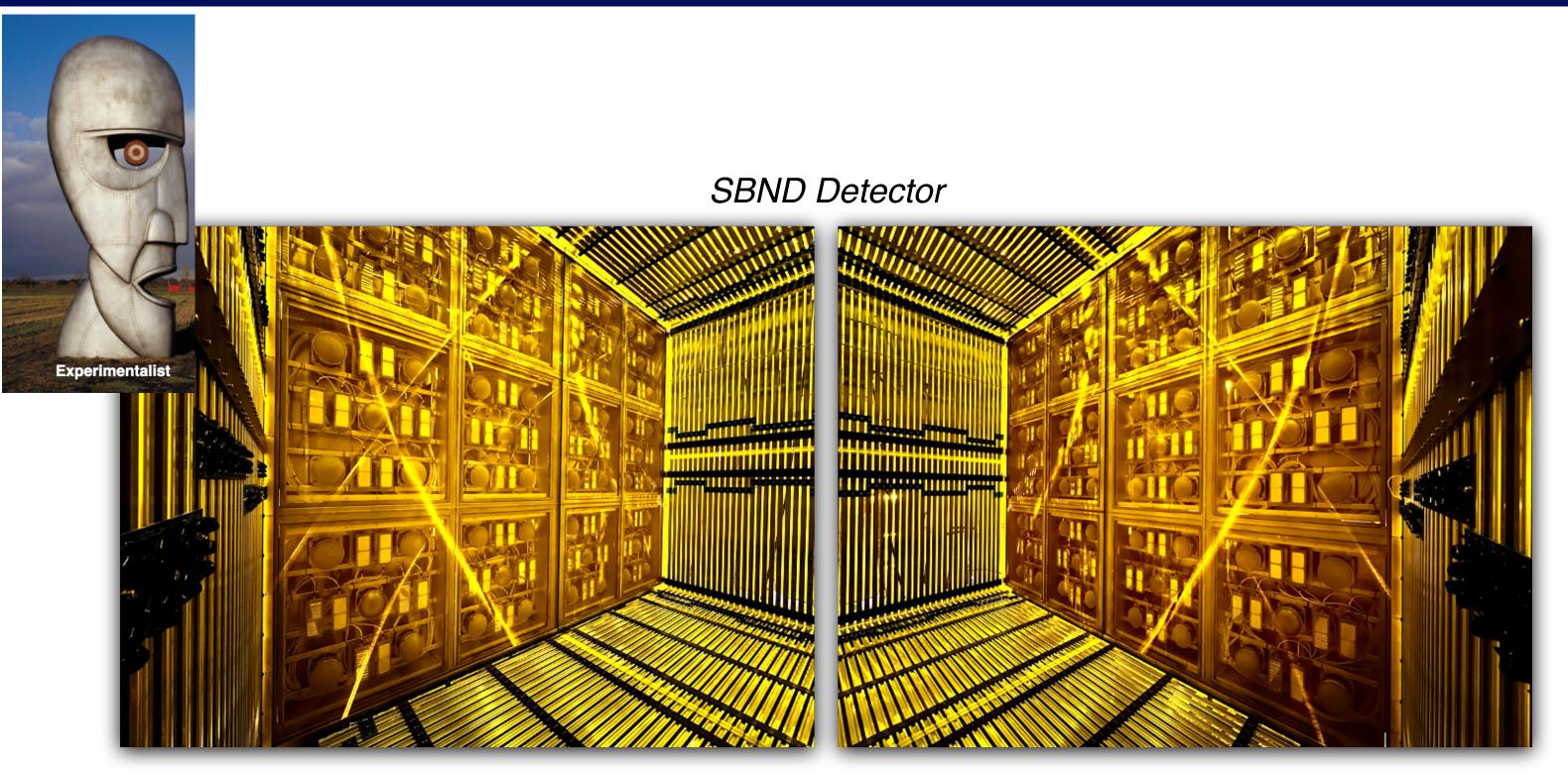
ArgoNeuT Collaboration + 6 theorists (R. Co, R. Harnik, K. Kelly, S. Kumar, Z. Liu. K Lyu)

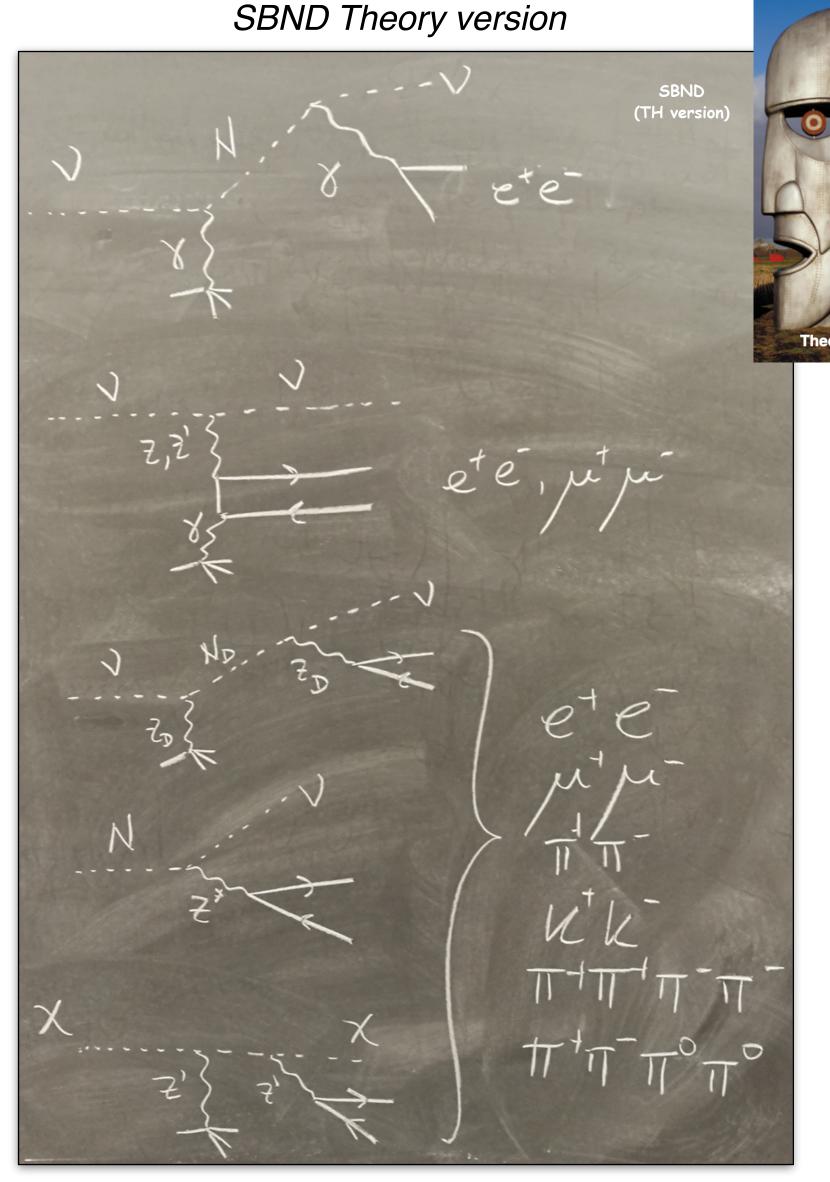
R. Acciarri et al., accepted by Phys. Rev. Lett., and selected as PRL Editors' Suggestion



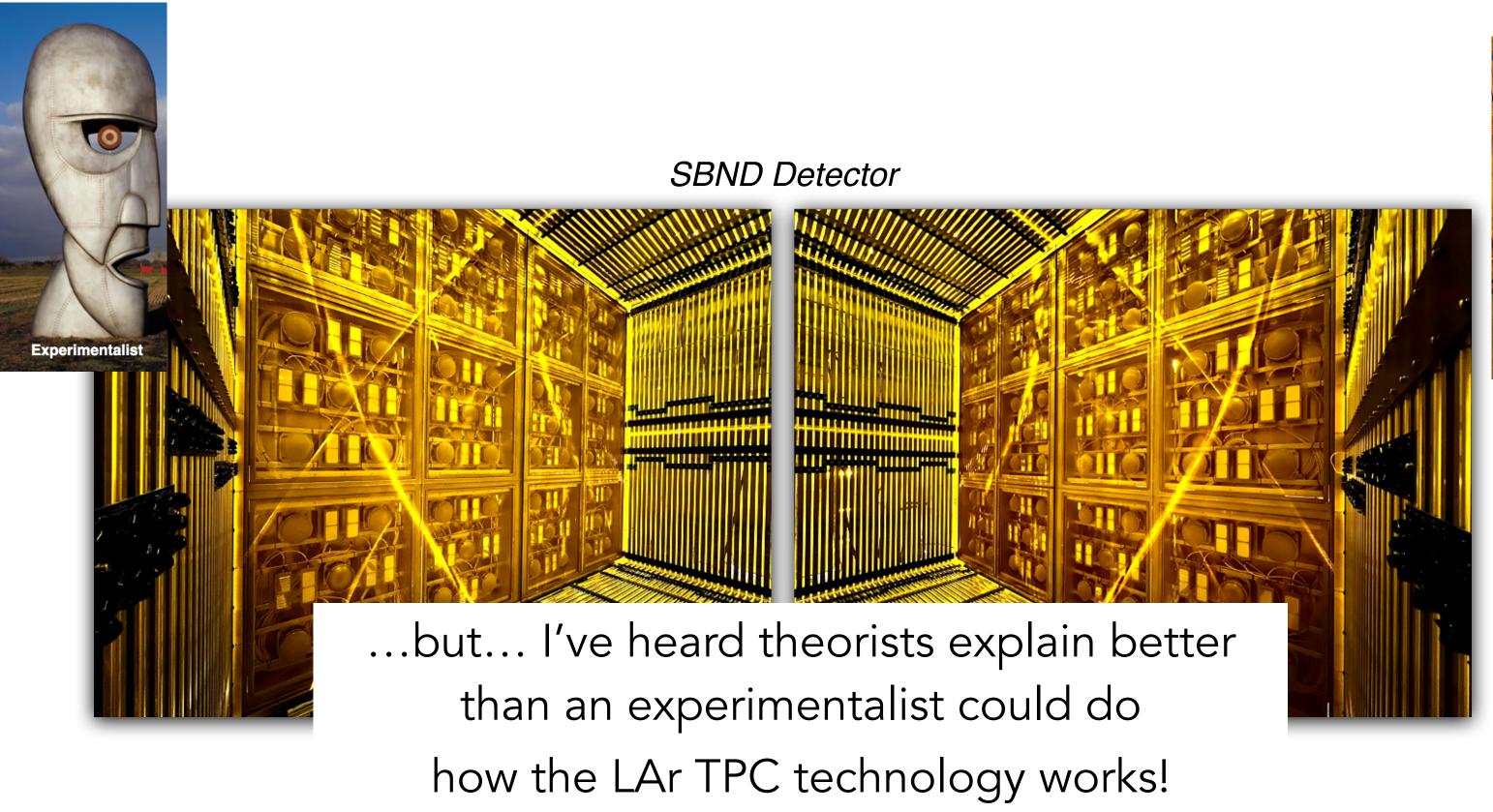
New exclusion constraints for heavy QCD axions with  $m_a \sim 0.2-0.9$  GeV and axion decay constant  $f_a \sim 10$  TeV.

# SAME PERSPECTIVE FROM TWO DIFFERENT ANGLES





# SAME PERSPECTIVE FROM TWO DIFFERENT ANGLES





Pedro Machado



SBND detector move



Roni Harnik, Pedro Machado, Olivia Vizcarra

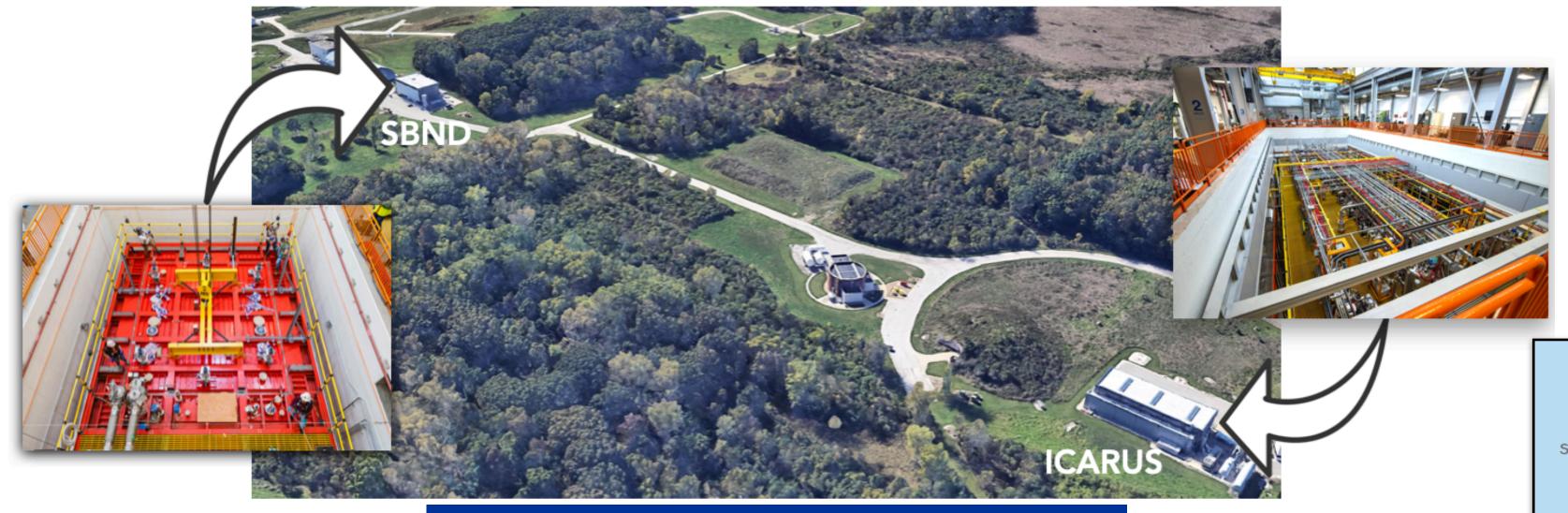
...and theorists support experimentalist...

Theorists are now members of the SBND Collaboration! ...and I'm invited to attend phenomenology conferences...

# SUMMARY

ArgoNeuT, a small LAr TPC running for 5 months on the NuMI beam produced world leading BSM searches.

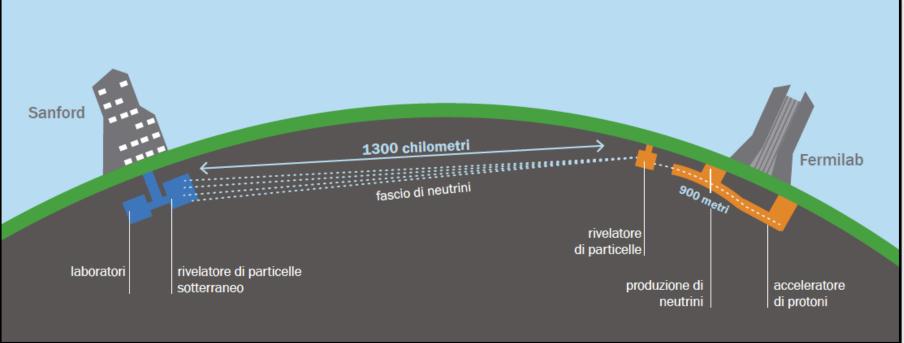
Current and next- generation LAr neutrino experiments are fantastic tools to look for **New Physics in the neutrino sector and beyond**.



Short-Baseline Neutrino Program

coming online soon...

#### in the next decade





Collaborations between experimentalists and theorists are of great relevance to the neutrino physics community, that needs the theory support to strengthen the neutrino experimental measurements, particularly when searching for New Physics.

## ...Now...on a personal level...

Marcela and Carlos are amazing friends!

I have so many memories of beautiful moments spent with them.

Here are some examples...

#### Staying 3 feet apart but together during the pandemic!





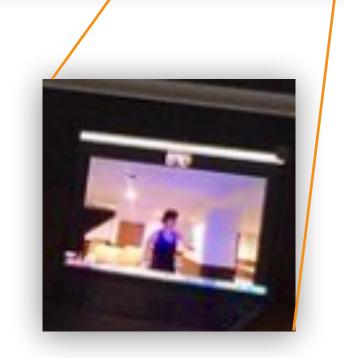




at Marcela & Carlos house

Marcela and Ornella Zoom workout during the pandemic!





## ...Now...on a personal level...

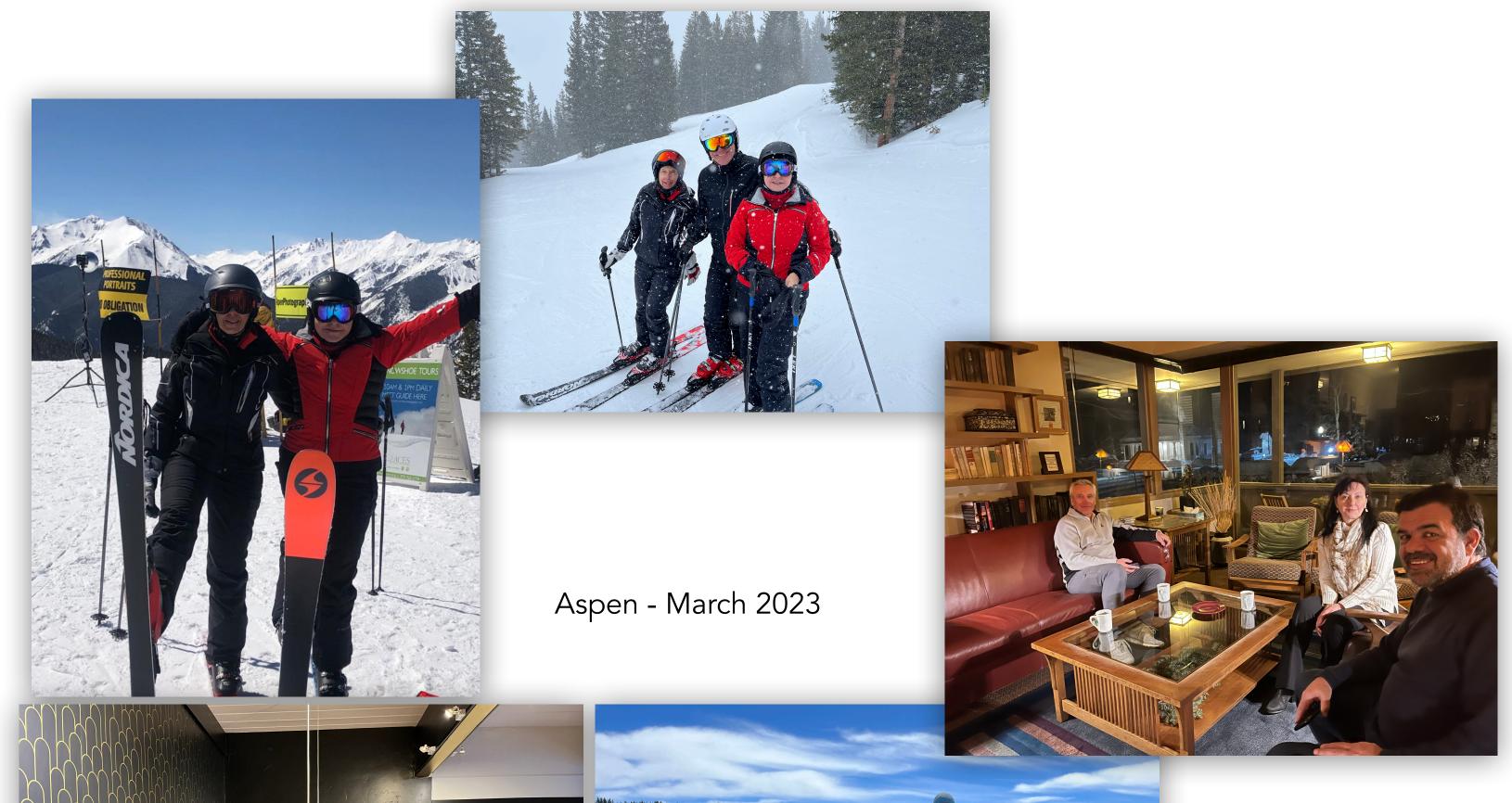


Sailing - Lake Geneva (WI) - Sept. 11 2021 Carlo's birthday!





## ...Now...on a personal level...









Marcela and Ornella workout now!



Marcela's influential impact on a new way to workout today!

## To Conclude

Marcela & Carlos, we feel extremely fortunate to have you as friends and honored to be here!

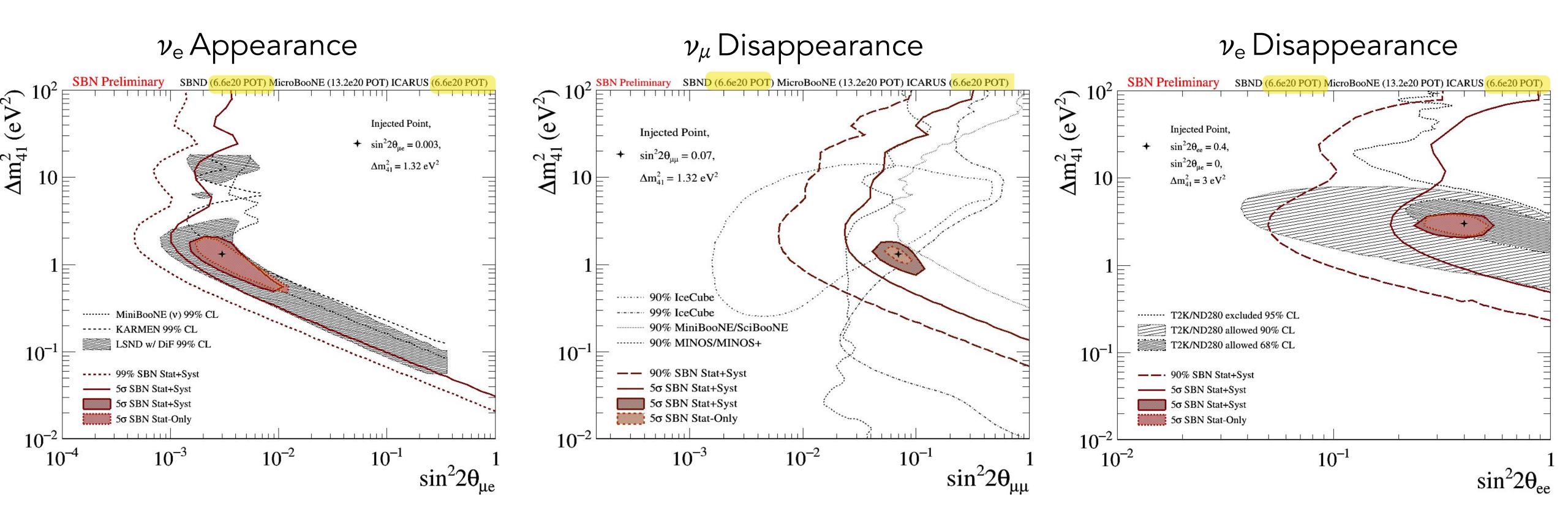
Congratulations on your 60th birthday!



To many more years filled with New Physics questions and lovely friendship to come!

# EXTRAS

## SBN STERILE NEUTRINO SENSITIVITIES



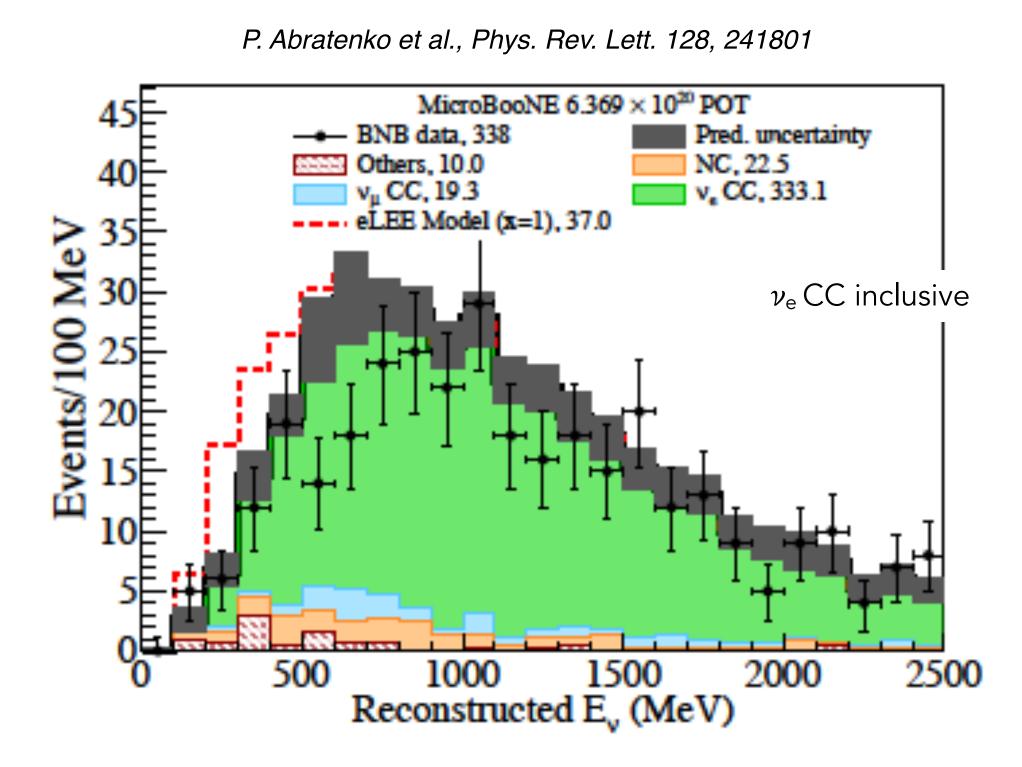
The SBN program tests the sterile neutrino hypothesis by covering the parameter regions allowed by past anomalies at  $5\sigma$  significance.

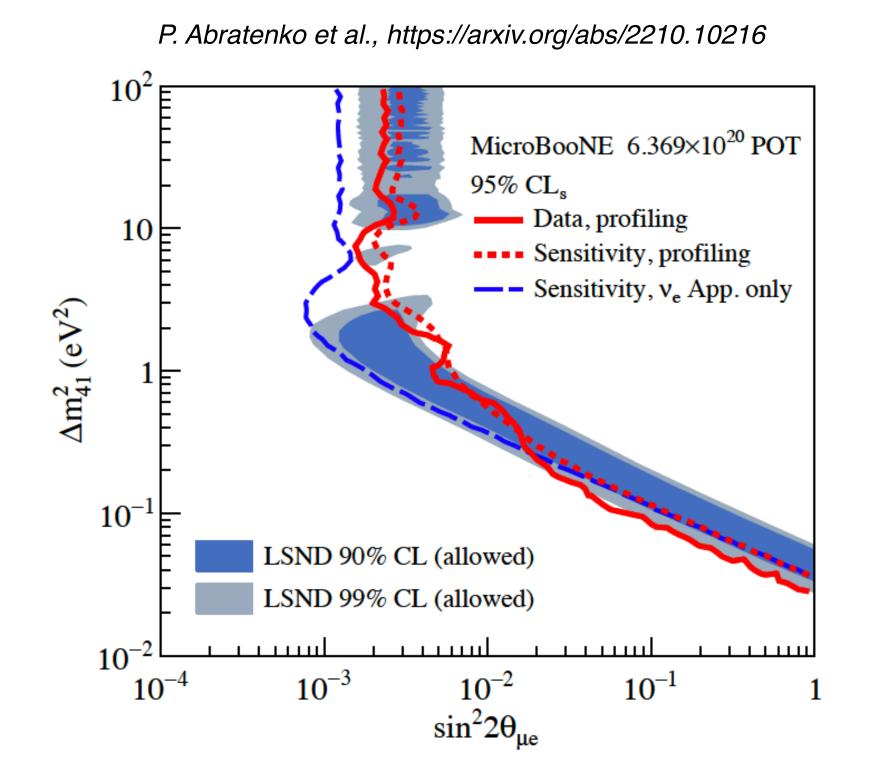
Complementary measurements in different modes: important for interpretation in terms of sterile neutrino oscillation.

## Sterile Neutrino Searches Beyond MicroboonE

The MicroBooNE experiment presented very interesting results of first analyses searching for an excess of low-energy electromagnetic events:

no hints of an electromagnetic event excess, but results do not rule out existence of sterile neutrinos.





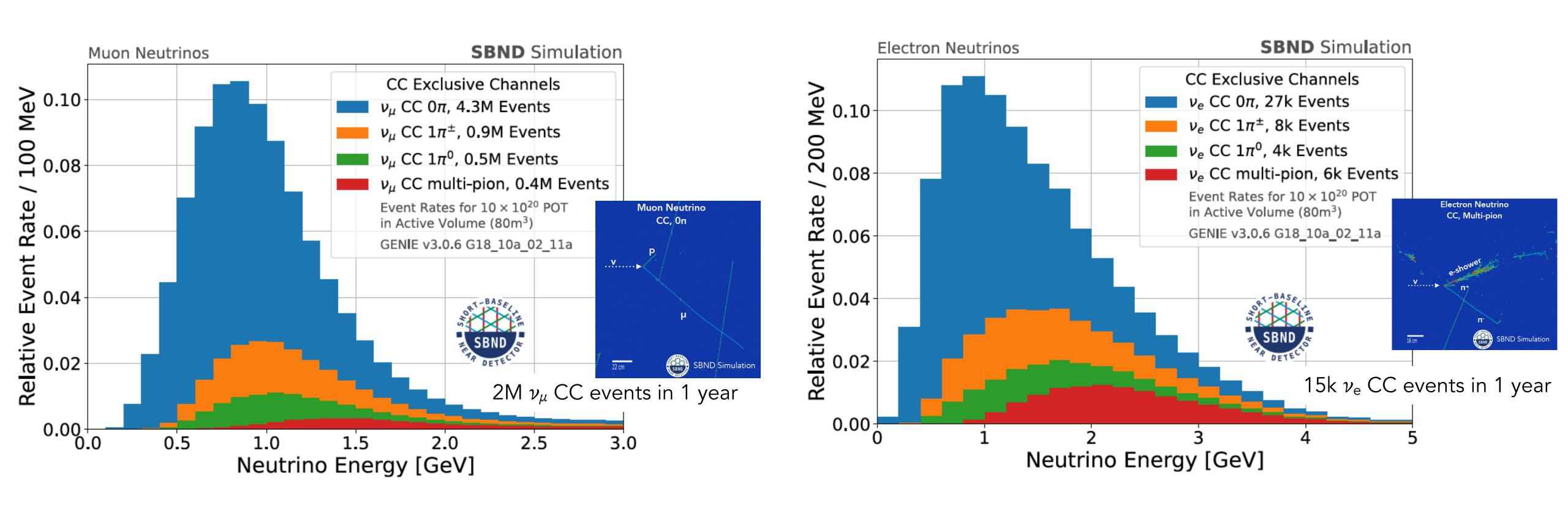
Entering the next phase now, with a Near detector and a larger Far detector.

## PRECISION STUDIES OF NEUTRINO-ARGON INTERACTIONS IN SBND



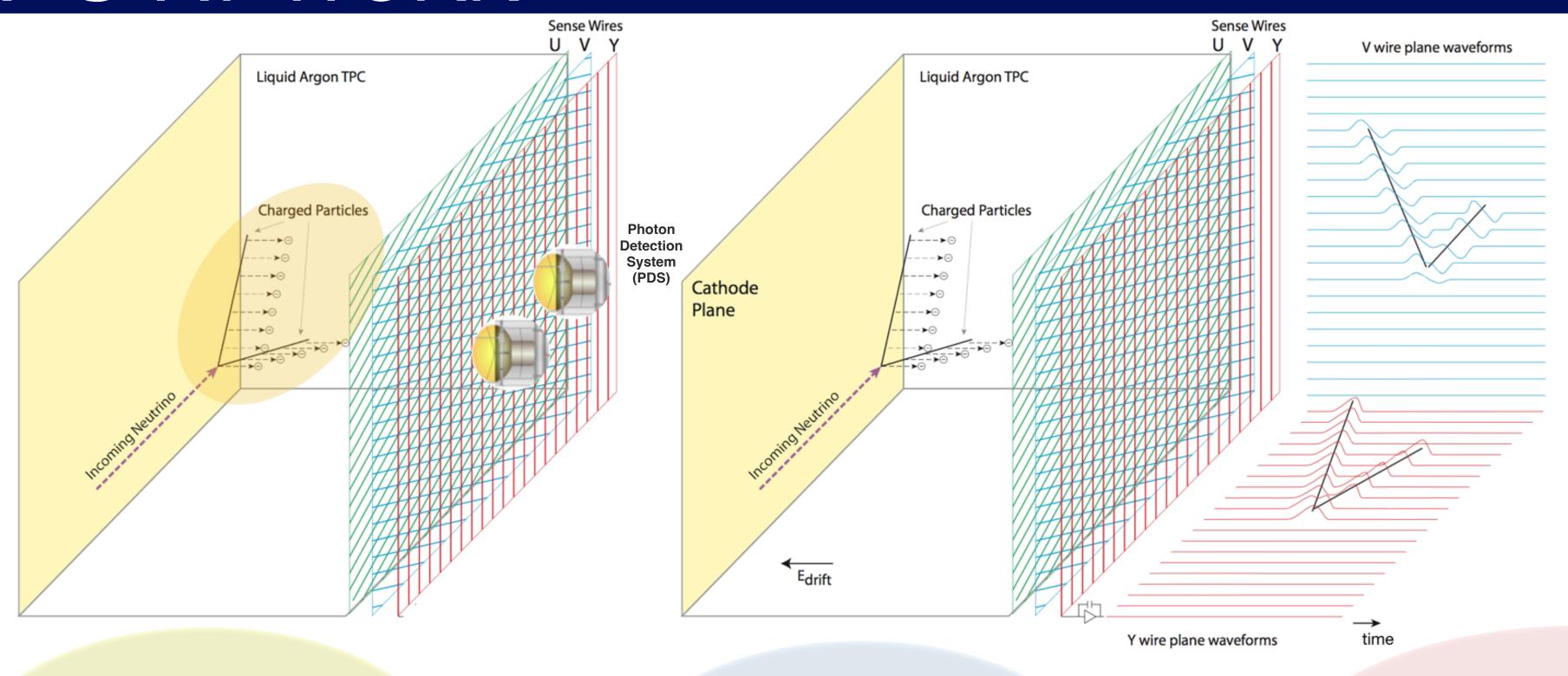
With its proximity to the neutrino source, **SBND** will compile neutrino data with unprecedented high event rate and will enable a generational advance in the study of neutrino-argon interactions in the GeV energy range.

### Up to 7000 $\nu$ events/per day in SBND!



SBND will record 20-30x more neutrino-argon interactions than is currently available (10 million total events (CC+NC), including around 50,000  $\nu_{\rm u}$  CC events above 2 GeV, and 50,000  $\nu_{\rm e}$  CC events).

# LAR TPC AT WORK



Charged particles in LAr produce free ionization electrons and scintillation light

m.i.p. at 500 V/cm: ~ 60,000 e/cm ~ 50,000 photons/cm

VUV photons propagate and are shifted into VIS photons

Ionization charge <u>drifts</u>
in a <u>uniform electric</u> field
towards the readout
wire-planes

Electron drift time ~ ms

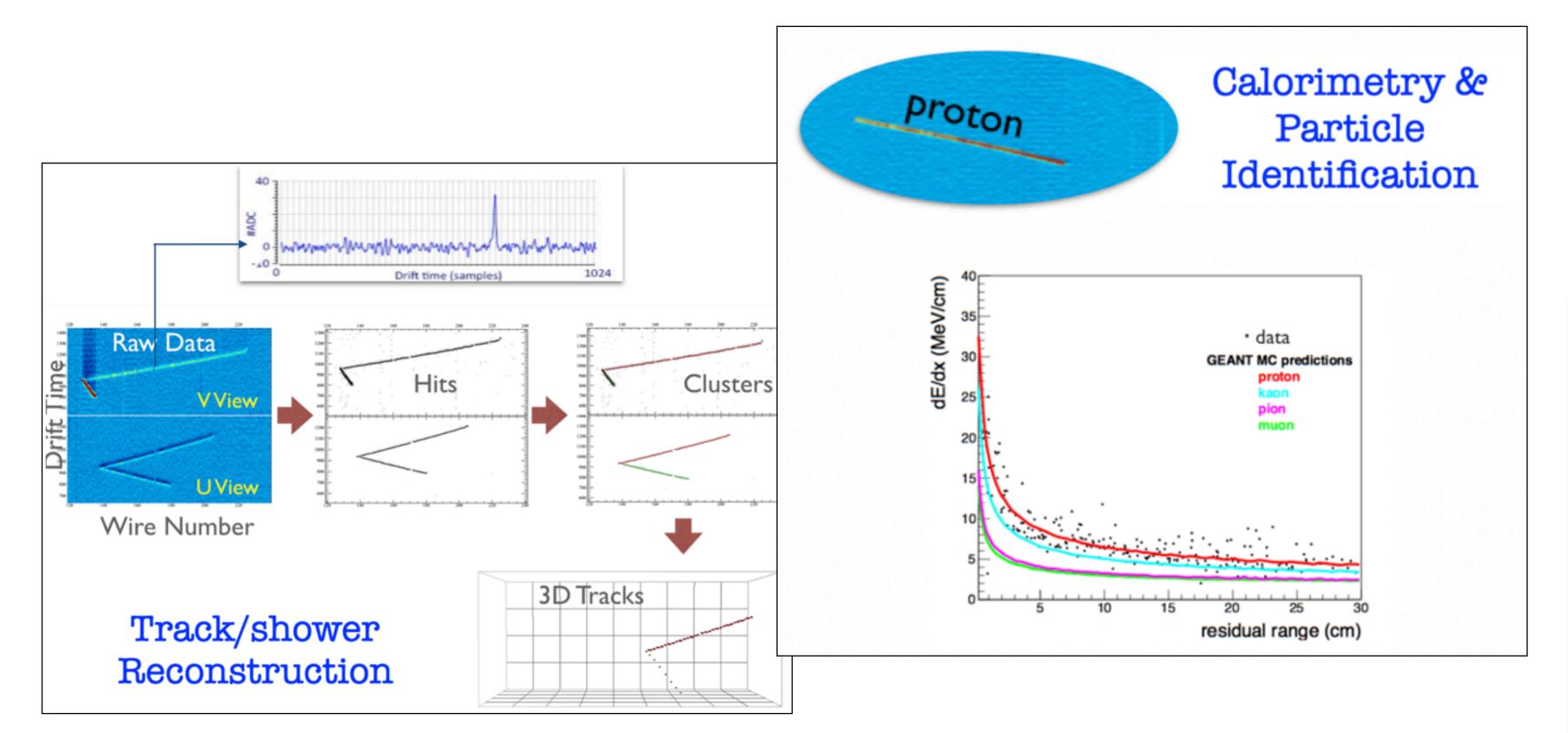
Digitized signals from the wires are collected [time of the wire pulses gives the drift coordinate of the track and amplitude gives the deposited charge]

Scintillation light fast signals from PDSs give event timing

# THE LAR TECHNOLOGY

Measure neutrino interactions with millimeter position resolution. Excellent capability for energy depositions from sub-MeV to few GeV, far beyond that offered by any other neutrino detector.

LArTPC at work: imagining, energy and timing



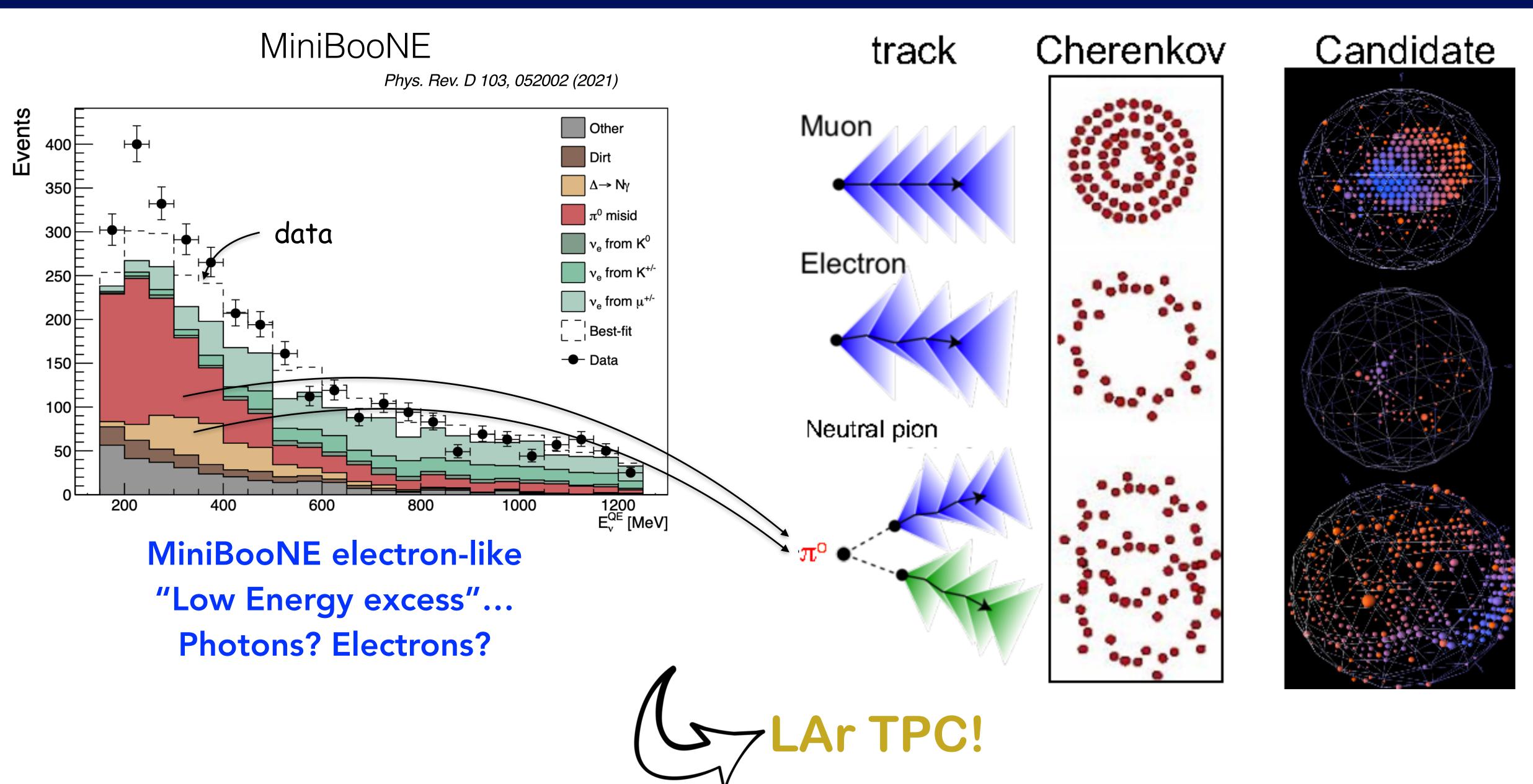
- Multiple 2D and the 3D reconstruction of charged particles ⇒ Imaging
- Total charge proportional to the deposited energy ⇒

#### **Calorimetry**

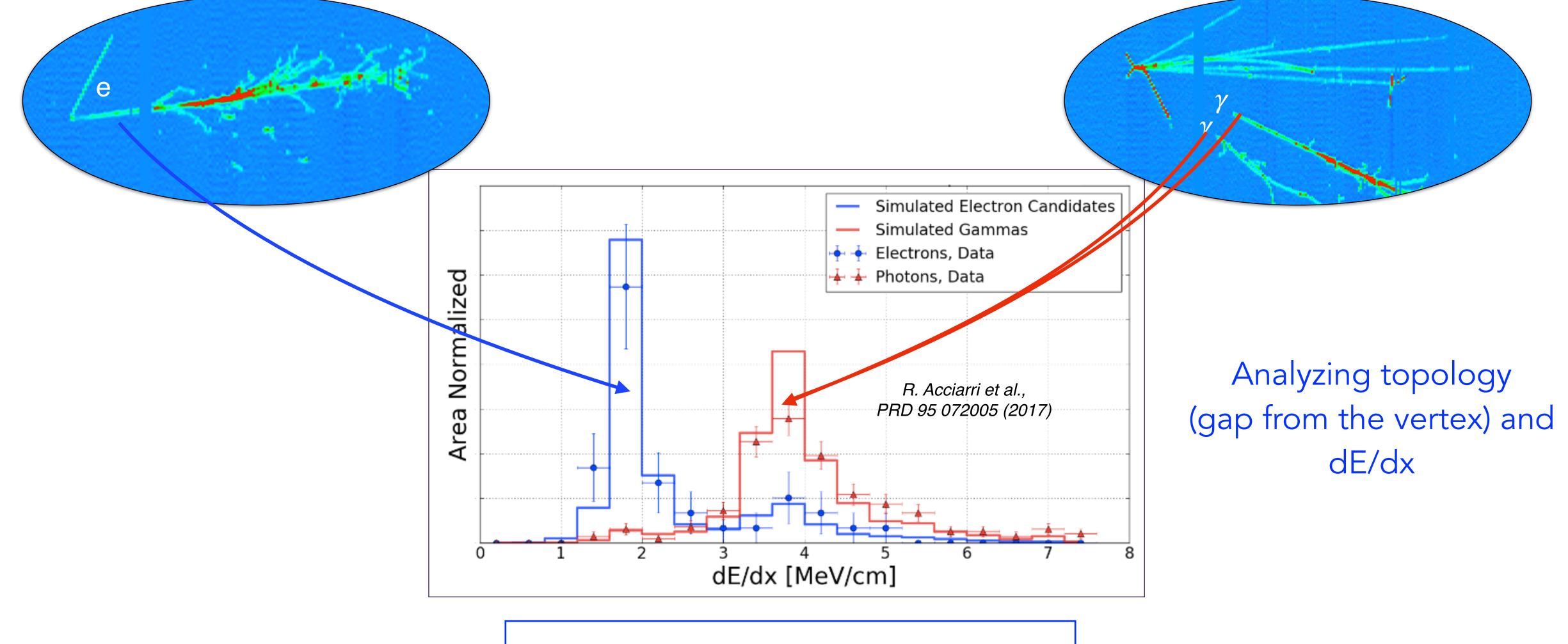
dE/dx along the track ⇒
 Particle Identification

and few ns timing resolution from photon detection system(s).

# ELECTRONS OR PHOTONS?

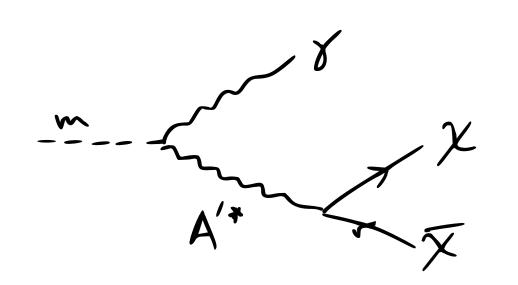


# ELECTRON-γ DISCRIMINATION IN LAR TPC



e-γ discrimination capability of LAr is crucial to disentangle the signal/background nature of the electron-like excess observed by MiniBooNE

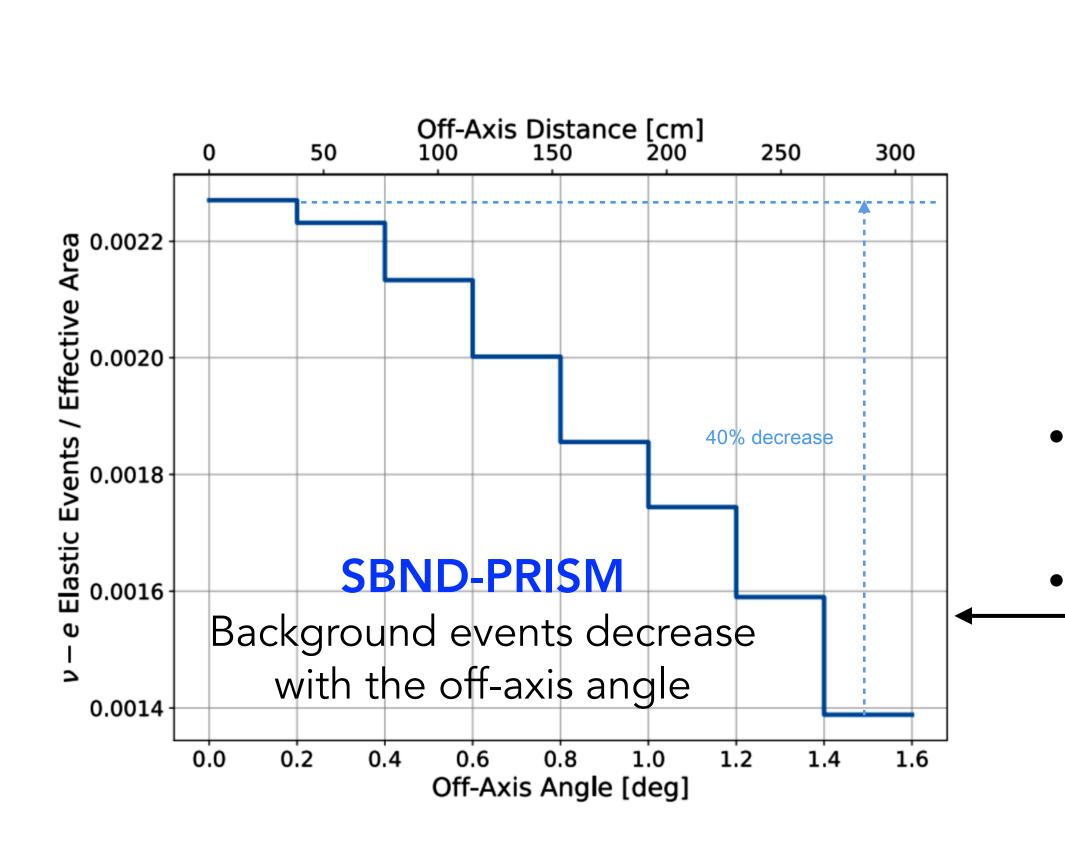
# SEARCHES FOR LIGHT DARK MATTER IN SBND

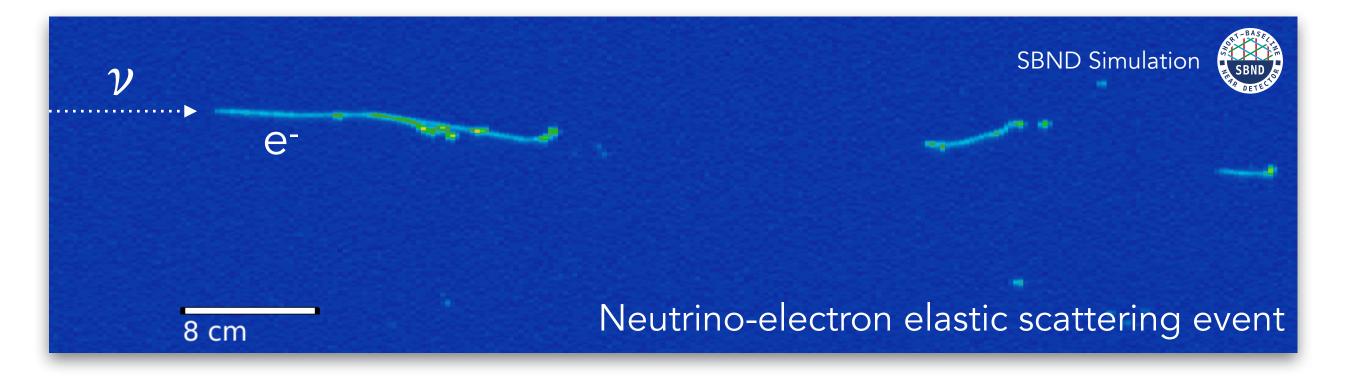


An example: **light** (sub-GeV) **dark matter**, coupled to the Standard Model via a dark photon. Dark photons can be produced by the decay of neutral meson (pions, etas) in the target, and

decay into dark matter.

The dark matter, through the dark photon, scatter off electrons in the detector.



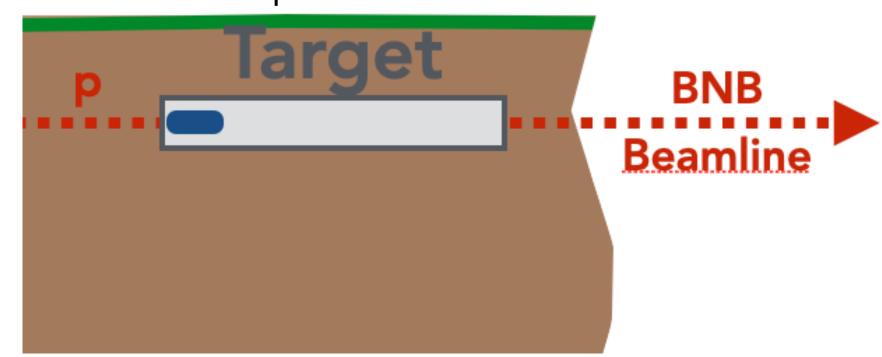


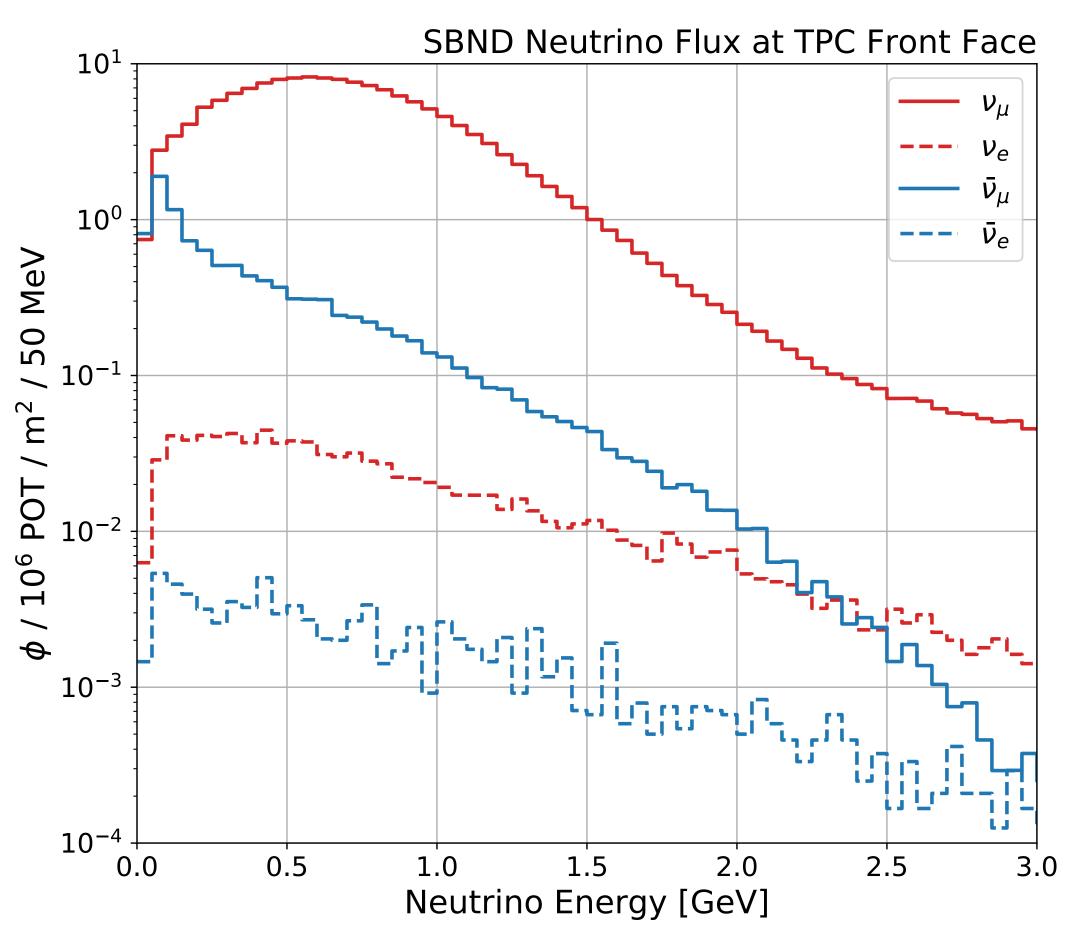
- **Signal**: DM elastic scattering electron events. DM comes from neutral (unfocused) mesons.
  - **Background**: neutrino-electron elastic scattering. Neutrinos come from charged (focused) mesons.
  - Neutrino flux drops off more sharply as a function of radius!

## BOOSTER NEUTRINO BEAM

#### High-intensity neutrino beam

from 8 GeV proton beam.





Neutrino flux at the SBND front face.

Mean muon-neutrino energy: ~0.8 GeV

Beam composition:

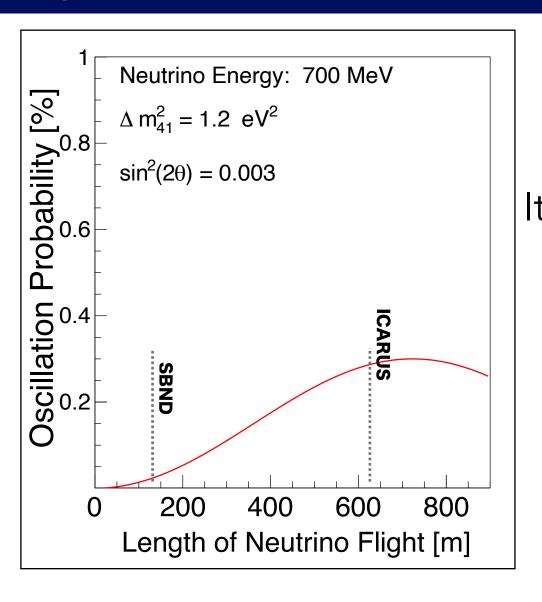
$$\nu_{\mu}$$
 (93.6%)

$$\bar{\nu}_{\mu}$$
 (5.9%)

$$\nu_e + \bar{\nu}_e$$
 (0.5%)

## WHAT MAKES THE SBN PROGRAM UNIQUE?

LAr Technology



#### Near detector SBND

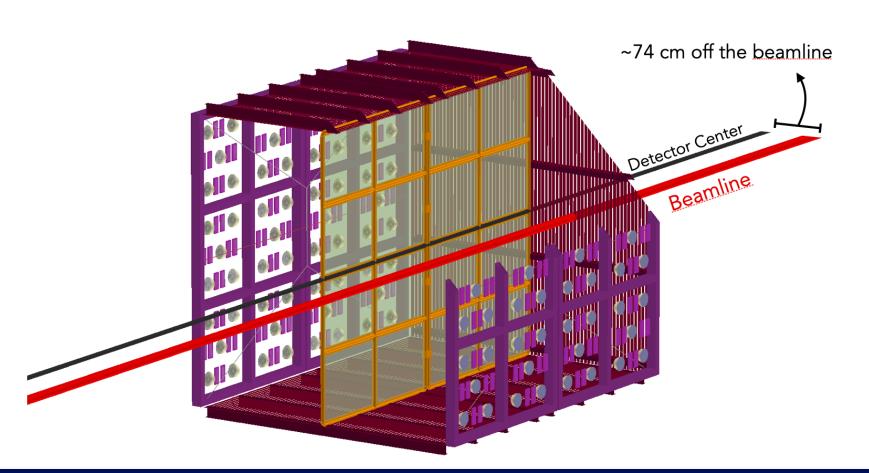
Crucial for oscillation searches.

Sitting close to the neutrino source, SBND plays a **unique role.**It sits before oscillations turn on @eV-scale → it characterizes the beam and **addresses the dominant systematic uncertainties** 

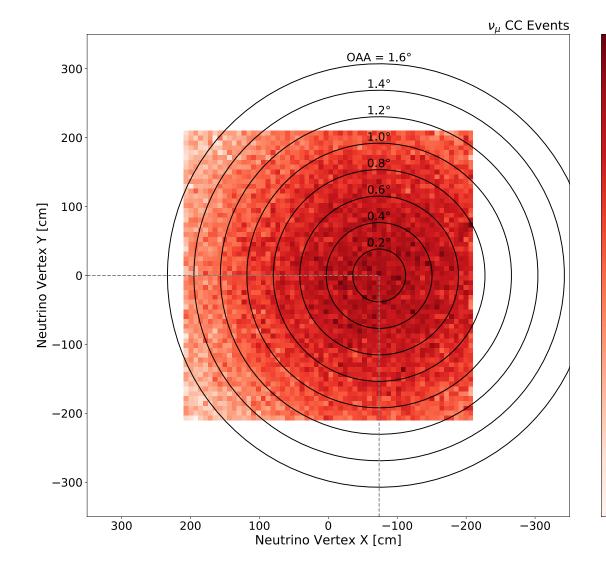
#### Far detector ICARUS

Given its far location and large mass provides big exposure to oscillated neutrinos, allowing for a **high sensitivity oscillation search** 

SBND is close (110 m) to the neutrino source and intentionally positioned offset relative to the beam center.



#### **SBND-PRISM**



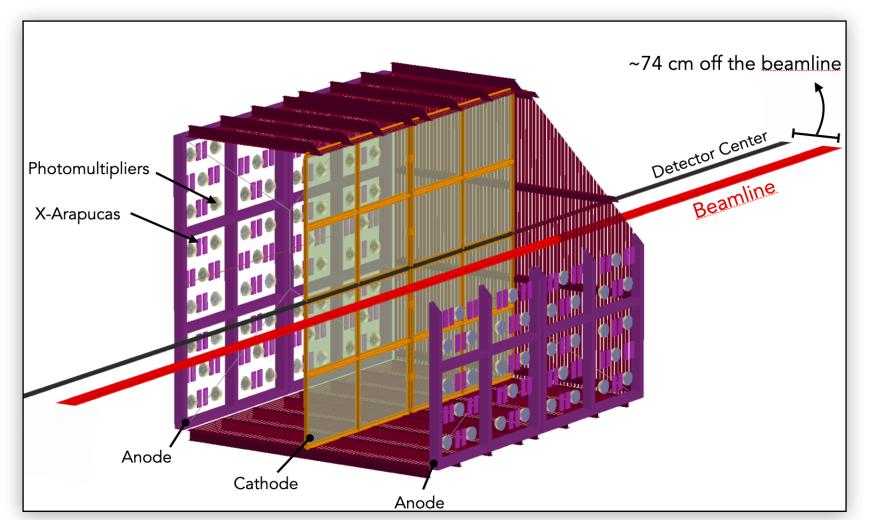
This "PRISM"\* feature of SBND allows sampling multiple neutrino fluxes in the detector

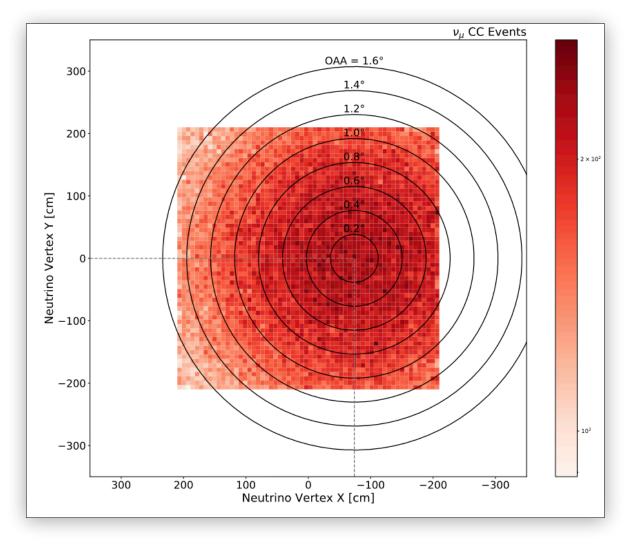
SBND-PRISM provides unique constraints of systematic uncertainties, helps mitigate backgrounds, and expands the SBN(D) physics potentials

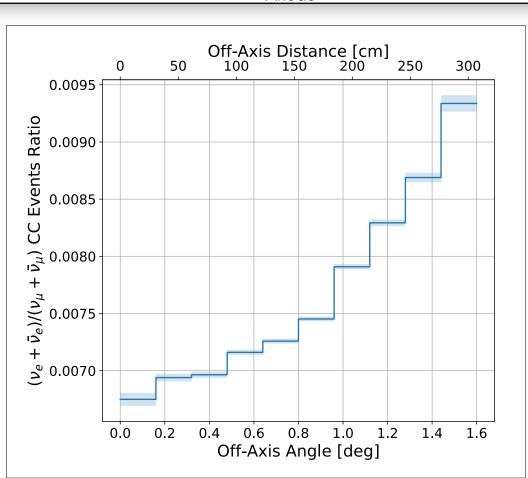
<sup>\*</sup>Similar to the nu-PRISM and DUNE-PRISM concepts, but with a fixed detector.

## SBND-PRISM: SAMPLING MULTIPLE OFF-AXIS FLUXES WITH THE SAME DETECTIONS

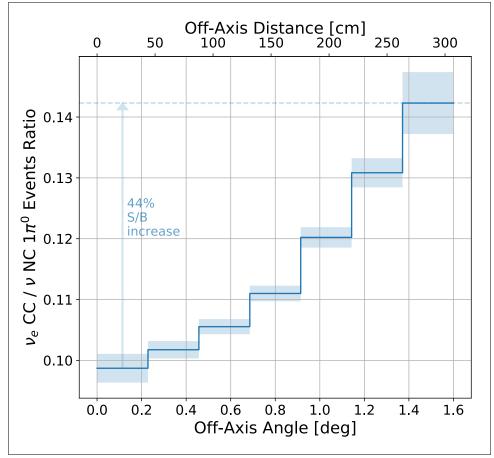
#### A Slightly Off-Axis Detector close to the neutrino source



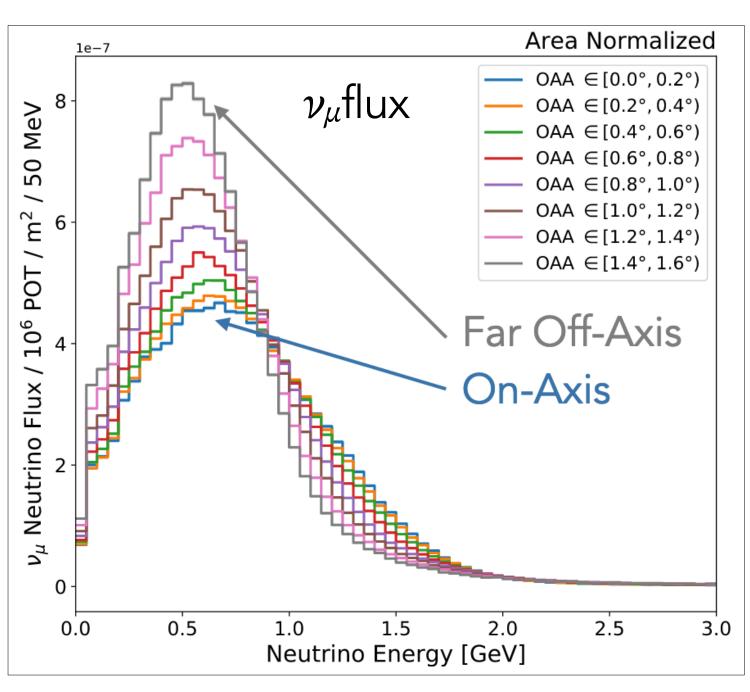




 $V_e$  to  $V_\mu$  ratio changes moving off-axis



NC  $\pi^0$  decreases moving off-axis



#### Additional physics potentials from SBND-PRISM

- Further constrain neutrino interactions in oscillation physics.
- Perform targeted neutrino interaction measurements and disentangle nuclear effects.
- Background reduction moving off-axis.



# SBND-PRISM - NEUTRINO FLUXES

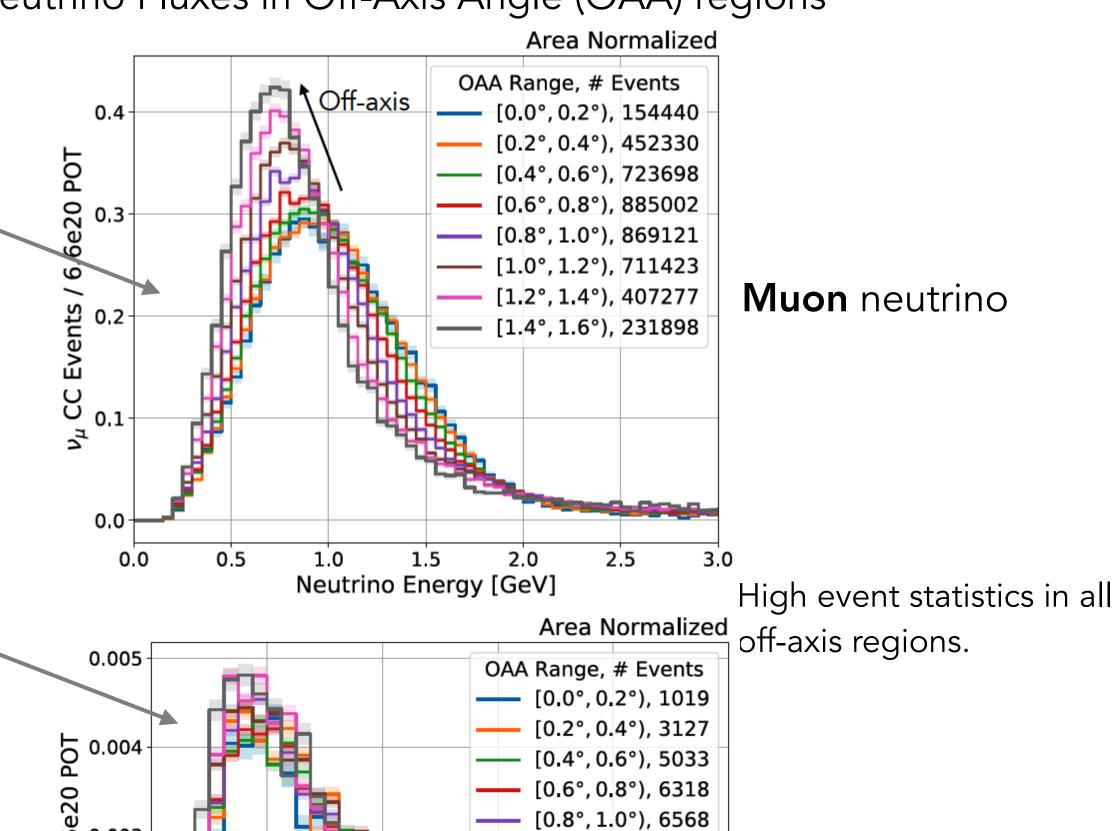
The **Muon** neutrino energy distributions are affected by the off-axis position  $[\nu_{\mu}]$  come predominantly from two-body decay]. Larger off-axis angle  $\rightarrow$  lower mean energy.

The **Electron** neutrino energy distributions also change, but they are less affected by off-axis position [ $\nu_e$  come from three-body decay].

Muon and electron neutrino spectra change in a different way!

Leveraging the different behavior of muon and electron neutrinos in the OAA regions, we can improve sensitivity for sterile neutrino searches.

Neutrino Fluxes in Off-Axis Angle (OAA) regions



Neutrino Energy [GeV]

[1.0°, 1.2°), 5509

[1.2°, 1.4°), 3370

- [1.4°, 1.6°), 2033

**Electron** neutrino

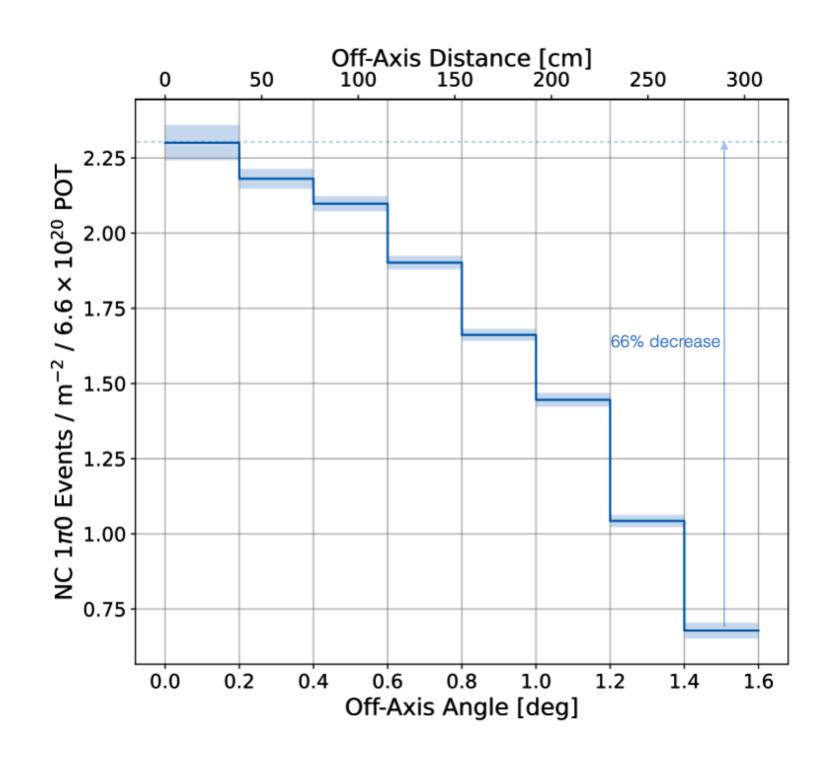
Events / (

0.000

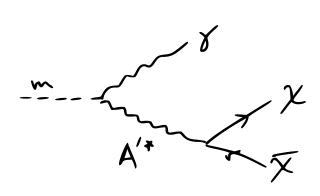
## SBND-PRISM TO MITIGATE BACKGROUNDS

#### An example: electron neutrino measurements

Main background for electron neutrino: NC 1  $\pi^0$  events.

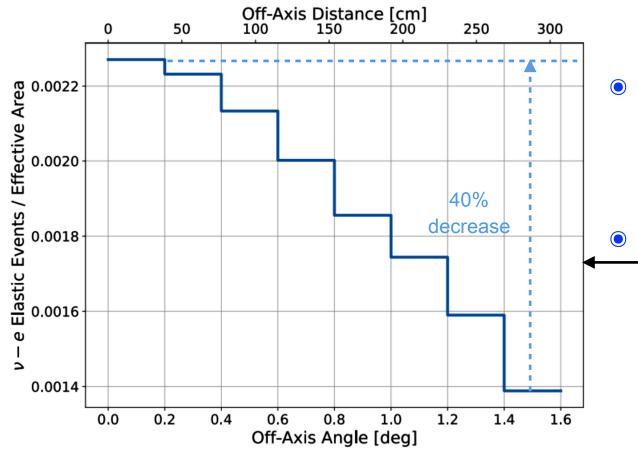


#### Another example: search for Light (sub-GeV) Dark Matter

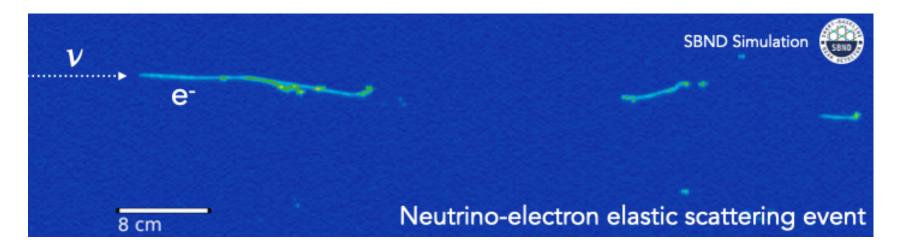


Dark photons, produced by the decay of neutral meson (pions, etas) in the target and decay into dark matter.

The dark matter, through the dark photon, scatter off electrons in the detector.



- **Signal**: DM elastic scattering electron events. DM comes from neutral (unfocused) mesons.
- Background: neutrino-electron elastic scattering. Neutrinos come from two-body decays of charged (focused) mesons.



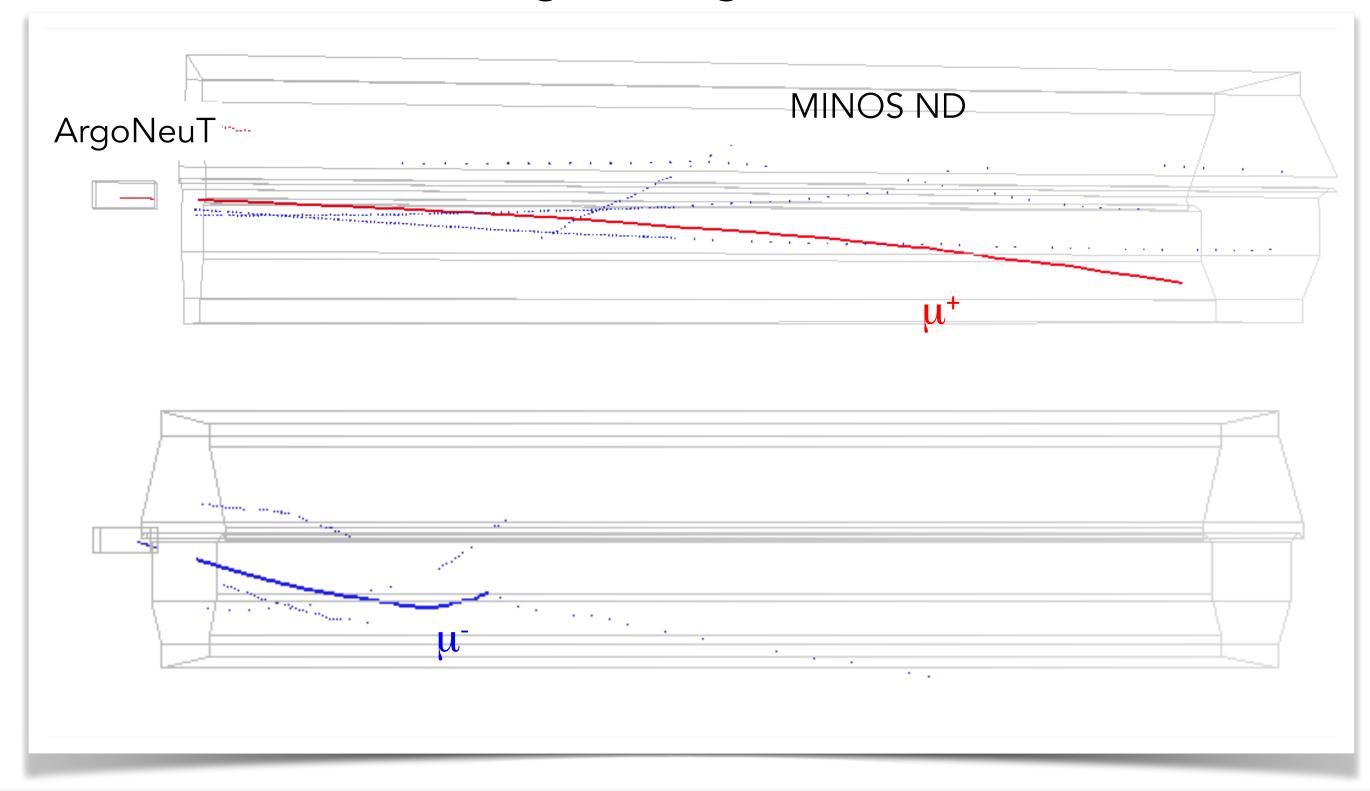
SBND-PRISM provides a natural way to reduce backgrounds by looking off-axis.

# ADVANTAGES OF ARGONEUT

Despite being a small LAr TPC and taking data for a short time, Argoneut has some advantages:

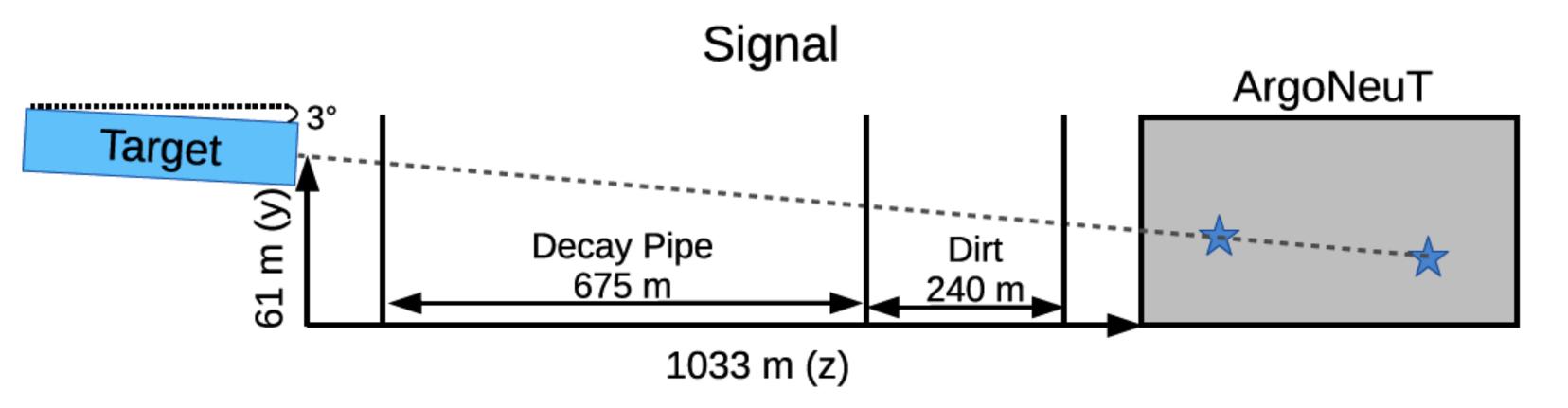
- 100 m underground → no cosmics
- Well understood/calibrated data set
- MINOS ND as spectrometer
  - Magnetic field allows muon momentum measurement and charge recognition
  - Allows to distinguish pion from muon, typically difficult in LAr TPCs

MINOS ND used to identify muons - tracks exiting ArgoNeuT are reconstructed in MINOS ND

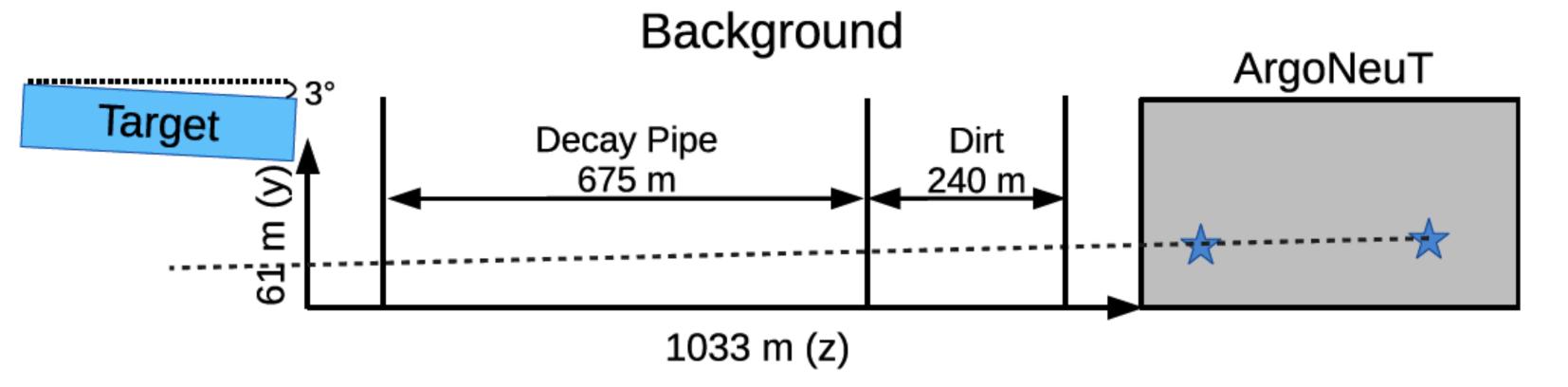


# MCP SEARCH - ARGONEUT ANALYSIS TECHNIQUE

En route to the detector, mCPs travel through hundreds of meters of dirt, energy loss is negligible and angular deflections are small  $\rightarrow$  mCPs point back to the target.



Signal is a double-hit event with a line defined by the two hits pointing to the target



A background double-hit event doesn't point to the target

To reduce background: search for double hits events aligned with the target

# HIGHLY FORWARD-GOING DI-MUON SIGNATURES

HNL and Heavy QCD axions are very different models... but can produce similar decay signatures in ArgoNeuT

**HNLs** decaying to muon pair N  $\rightarrow$  v  $\mu$ +  $\mu$ -Muons highly boosted:

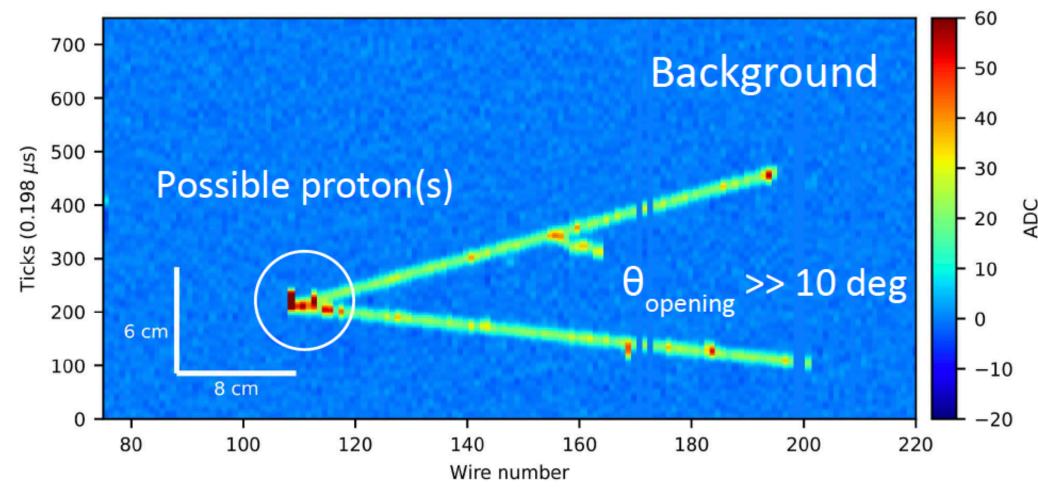
- average energy ~ 7 GeV
- average angle with respect to beam direction ~ 1.5 deg
- average opening angle ~ 3 deg

**Axions** decaying to muon pair a  $\rightarrow \mu + \mu$ -Muons highly boosted:

- average energy ~ 20 GeV
- average angle with respect to beam direction ~ 1.5 deg
- average opening angle ~ 2.5 deg

ArgoNeuT + MINOS ND, ideal to search for  $\mu^+\mu^-$  signatures

- ArgoNeuT LArTPC: vertex identification and reconstruction of low energy particles - allows identification of background
- MINOS ND: muon charge reconstruction and pion rejection



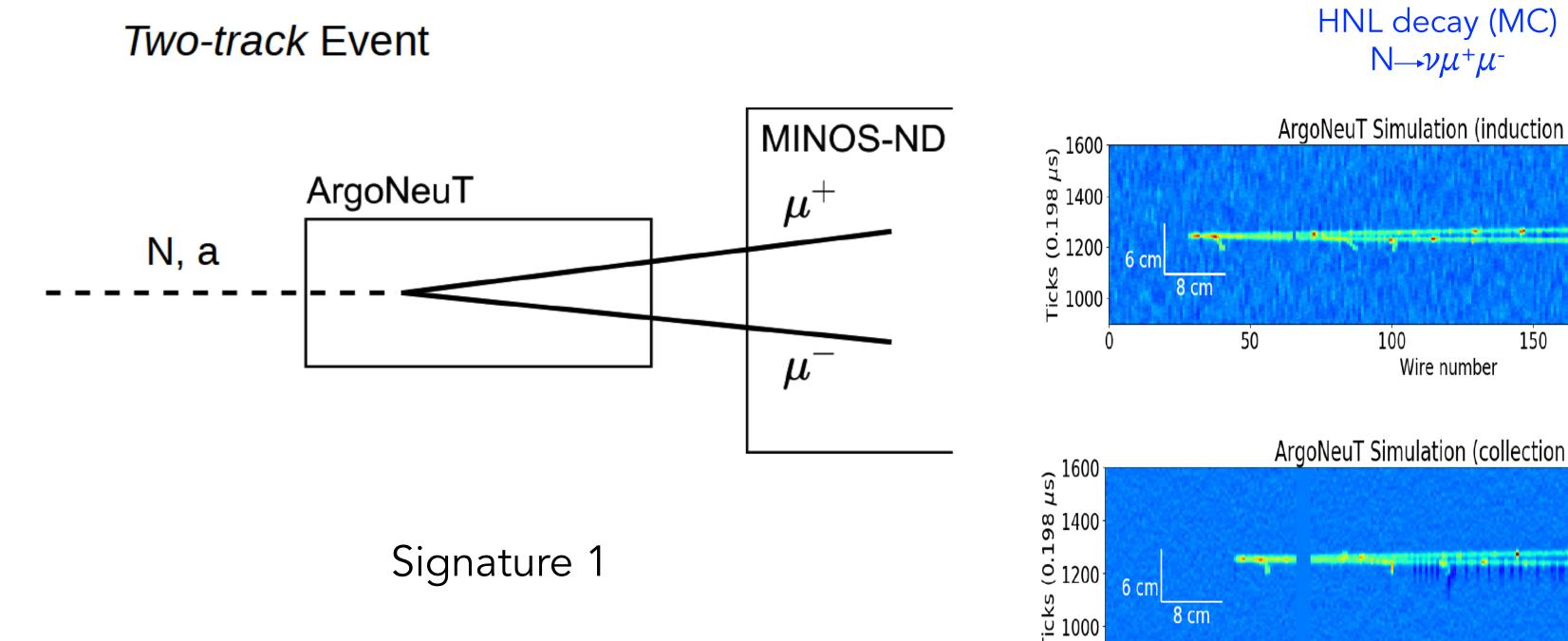
Identify the presence of hadronic activity at the vertex

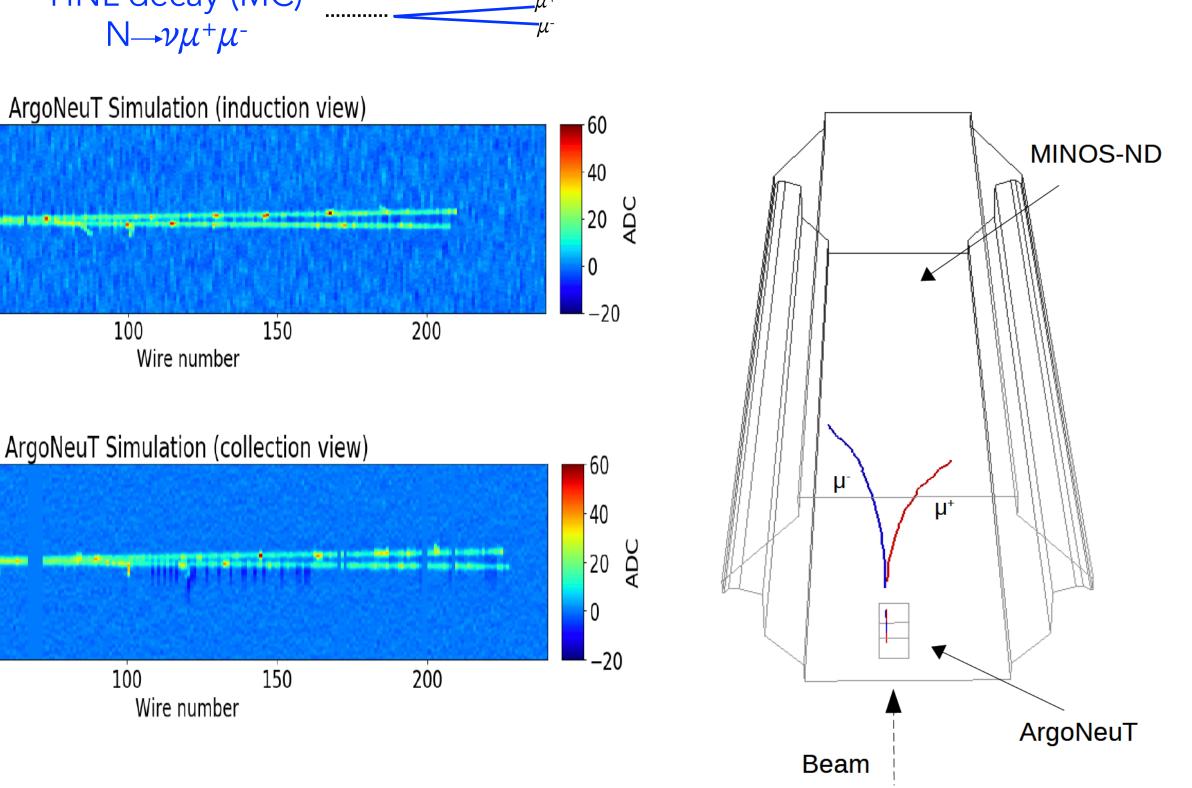
# DI-MUON SIGNATURES IN ARGONEUT

Two different signatures, depending on how forward going the muons are:

Signature 1: two MIP tracks in ArgoNeuT, match to two tracks in MINOS-ND

Signature 2: single double-MIP dE/dx track in ArgoNeuT, matches to two tracks in MINOS-ND



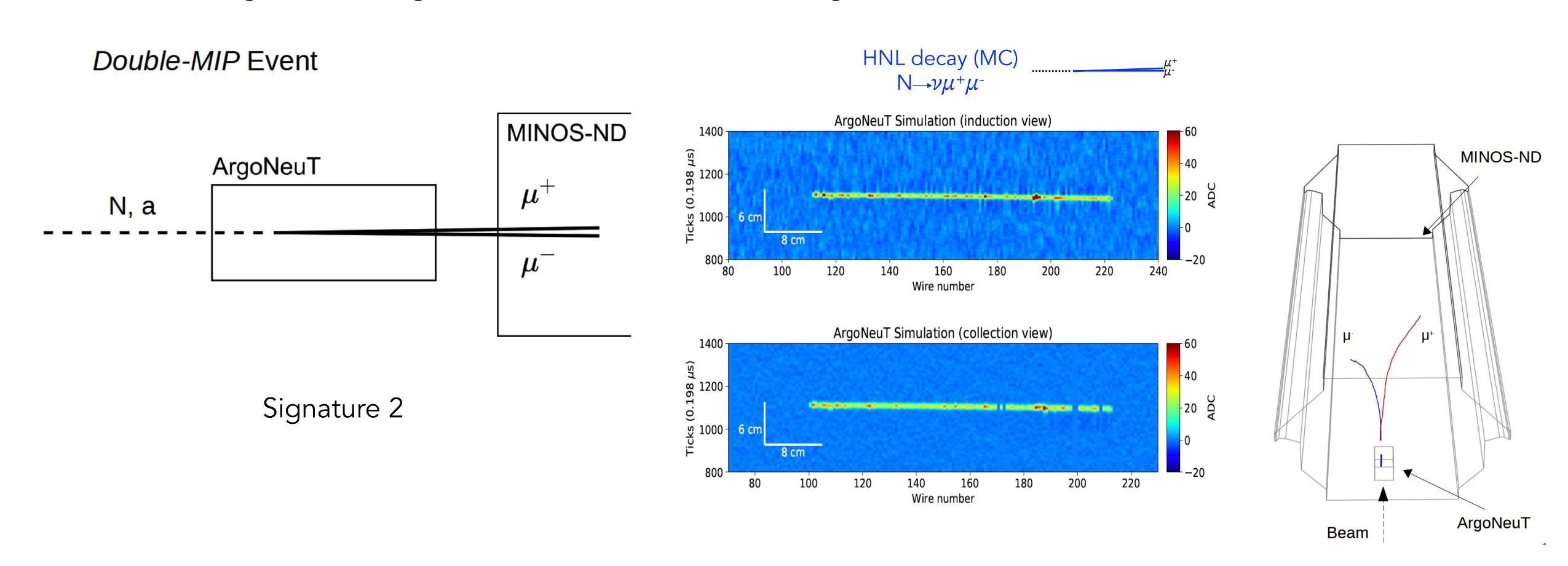


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In most forward-going cases, muon pair may be reconstructed as overlapping in ArgoNeuT.

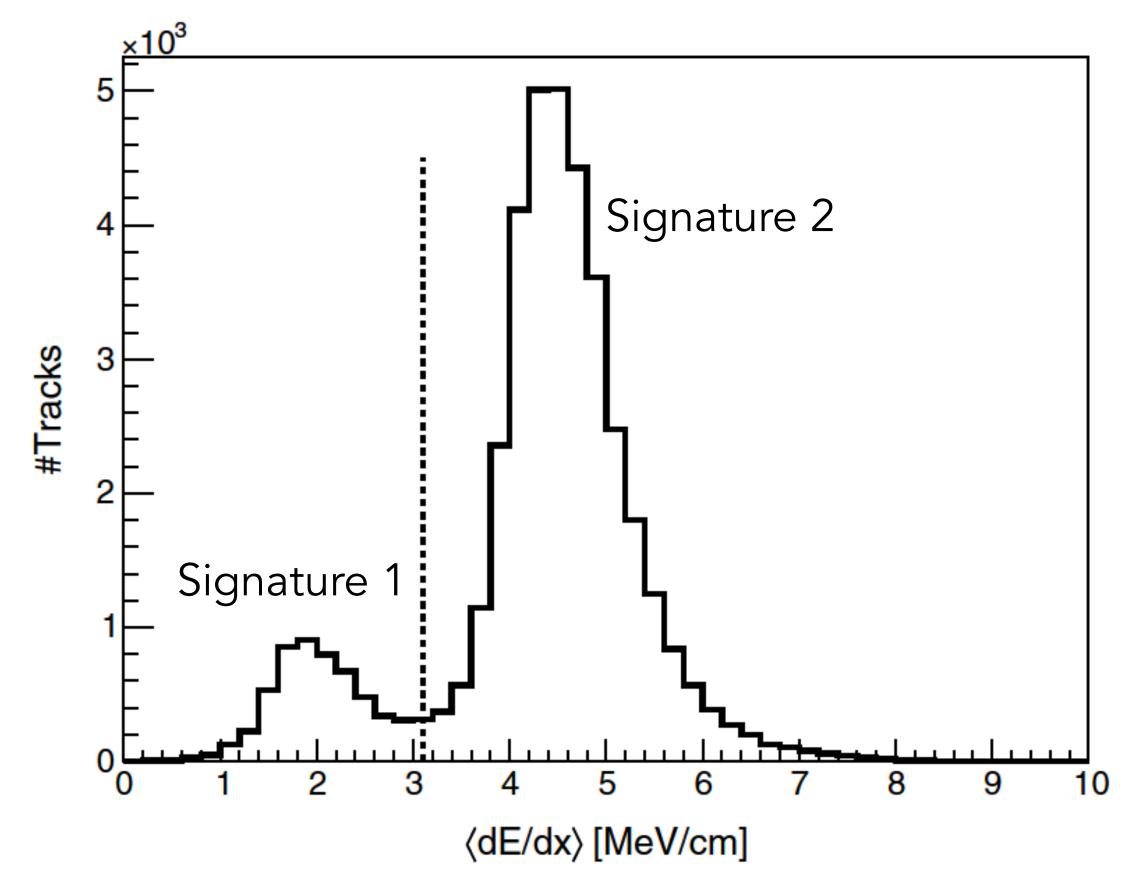
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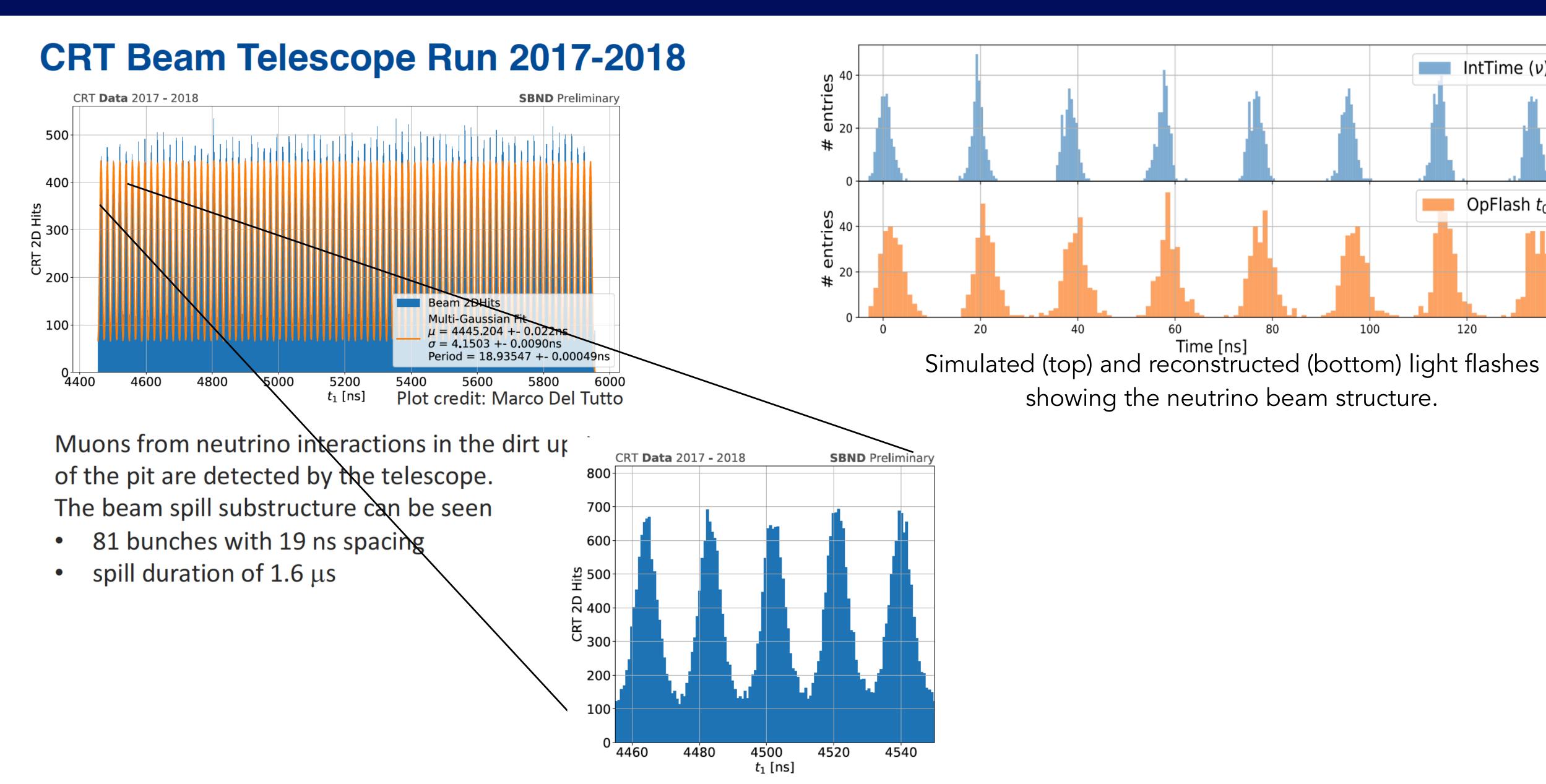
Developed new techniques to identify highly-forward-going muon pairs, applicable to future searches in LArTPC detectors, e.g. the DUNE near detector.



Identifying overlapping muon pairs:

Average reconstructed dE/dx over the first 5 cm of tracks resulting from simulated HNL decays.

## BNB TIME STRUCTURE



IntTime (v)

OpFlash  $t_0$ 

120