

# Characterization of a PbWO<sub>4</sub> crystal using Geant4 simulations for single particle collisions

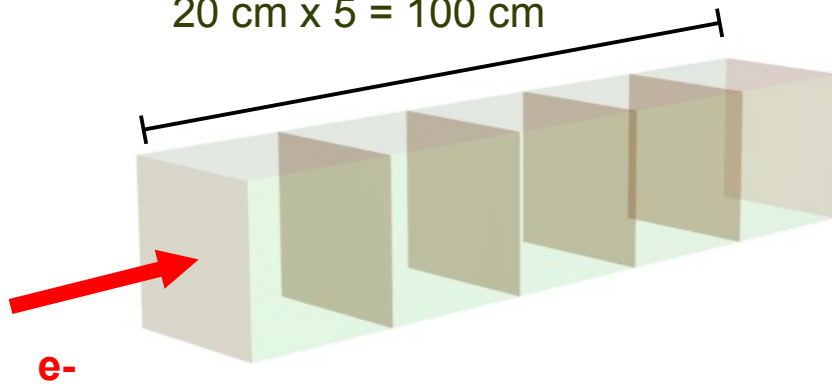
S.Chekanov + S.Eno

CalVision meeting

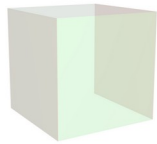
Feb 9, 2022

# Simulation of a simple HCAL tower using PbWO<sub>4</sub>

20 cm x 5 = 100 cm



e<sup>-</sup>



20 cm x 20 cm ~ 21 X<sub>0</sub> ~ 1 λ<sub>I</sub>

Simulate hits, N(scintillation), N(cherenov) for particles of different types between 0.5 – 20 GeV

Challenging! 20 GeV particles produces 4 million photons (on average)

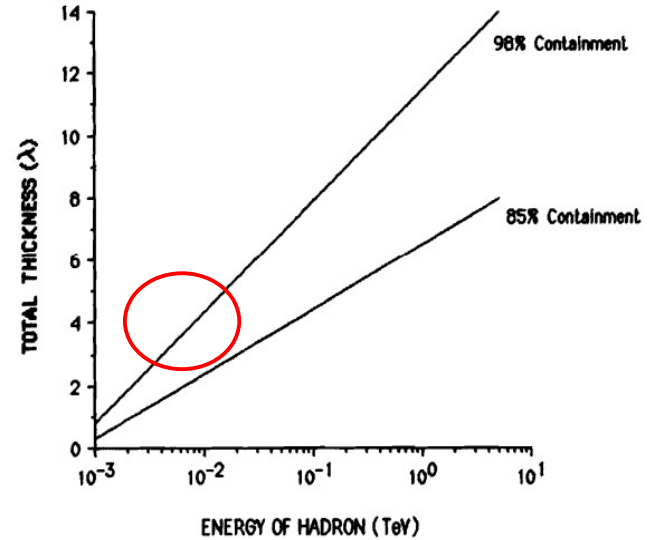
21 X<sub>0</sub> (typical width for ECAL FCCee)

5 interaction length

~98% containment of hadronic shower for 1 GeV particles

<http://lss.fnal.gov/conf/C860623/p355.pdf>

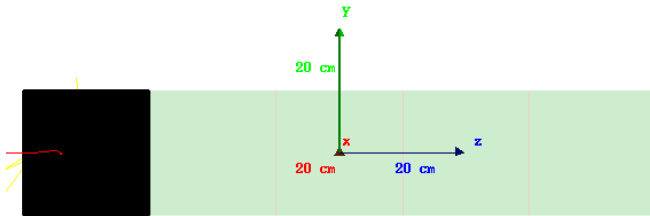
Containment of hadron showers



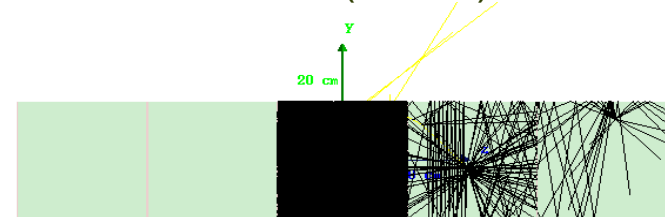
# Examples of collisions

- Simulation have been created for 8 particles. 1000 events per particle type
- (e, gamma, pi, pi0, p, proton, neutron, mu, K-)
- Keep optical photons with wavelength 300 nm – 1000 nm
- All files are stored in HepSim: <https://atlaswww.hep.anl.gov/hepsim/info.php?item=362>

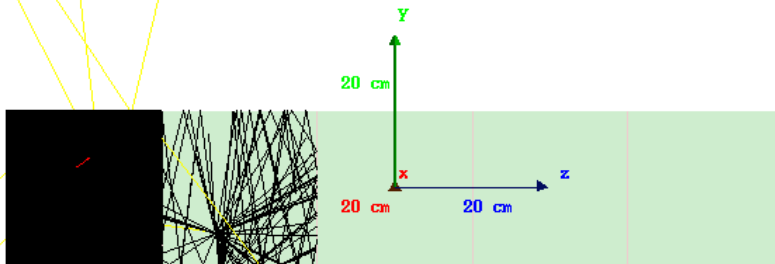
Electron (1 GeV)



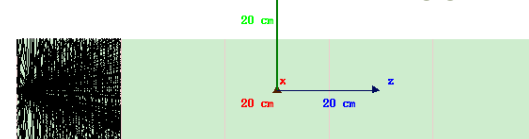
Neutron (5 GeV)



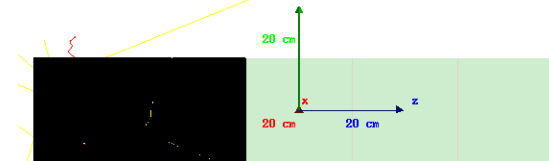
K- 5 GeV



Proton 0.5 GeV (Bragg effect)

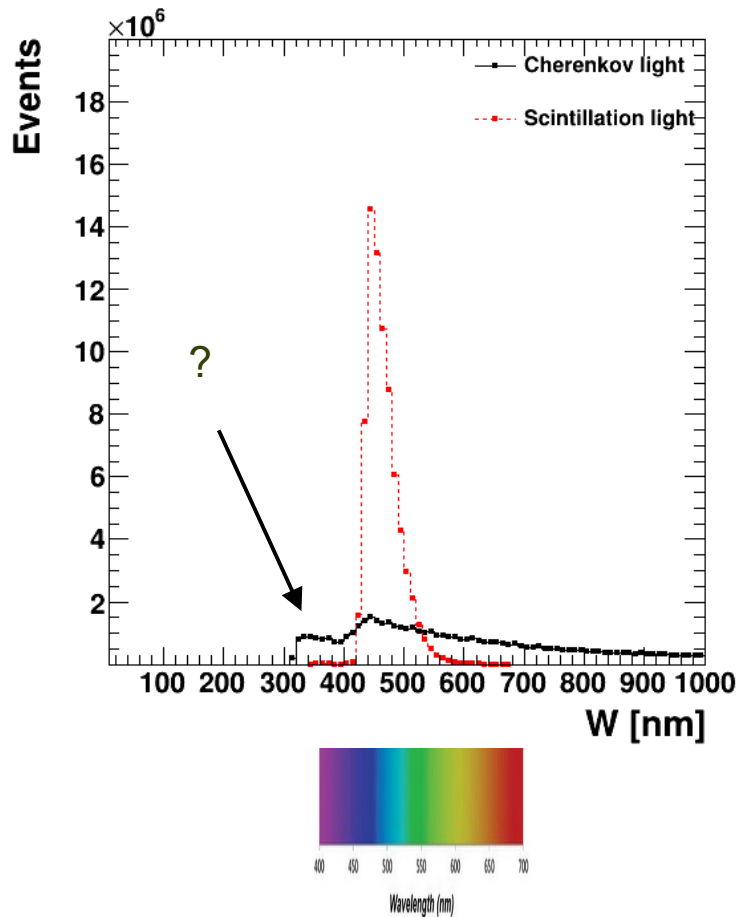


Proton 5 GeV



Photons are shown with black lines  
(dark area – too many to show!)

# Wavelength for Cherenkov and Scintillation light



**Scintillation light peak: ~460 nm**

Wavelength for 1000 injected electrons (1 GeV)

Counting of photons in [300-1000] nm range

No any instrumental filter, efficiency of SiPMT etc.

“**Luminosity per MeV**” calculated as the average number of photons per MeV (for e-):

Scintillation light:

**193 + -1 per MeV**

Agrees with expectations for PbWO<sub>4</sub>

see <http://scintillator.lbl.gov/>

Cherenkov light:

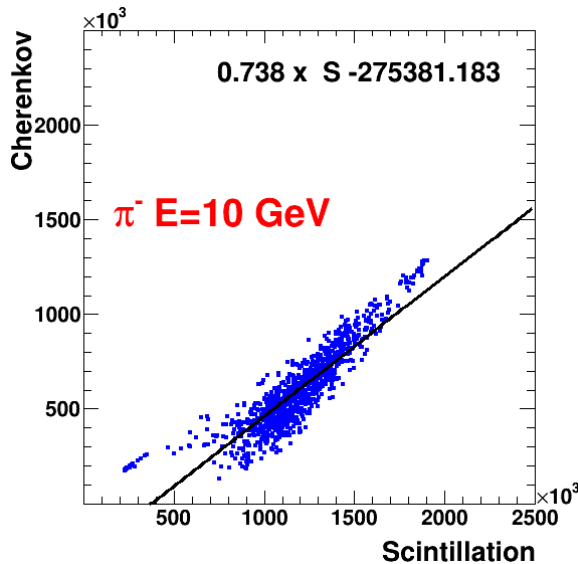
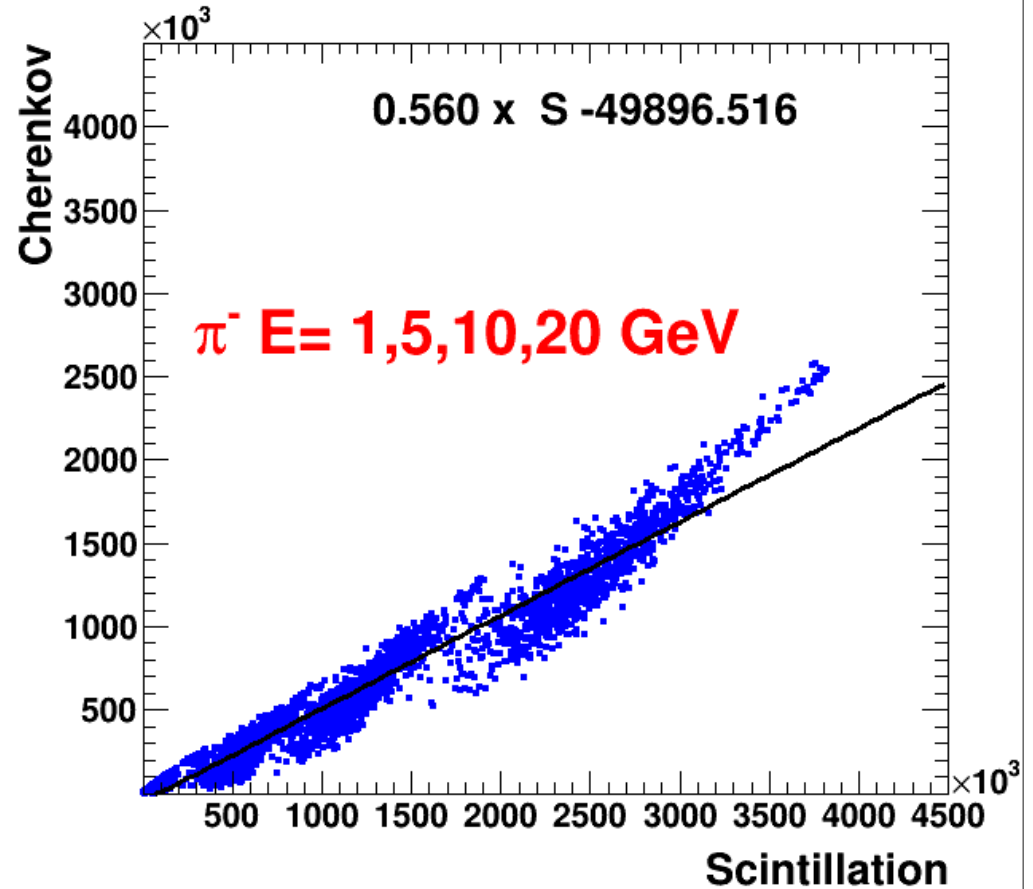
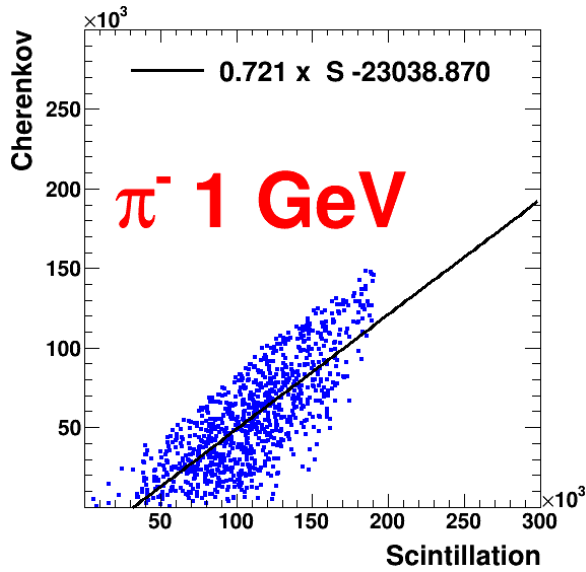
**127 + - 3 per MeV**

Problem:

Expect Cherenkov peak below 300 nm.

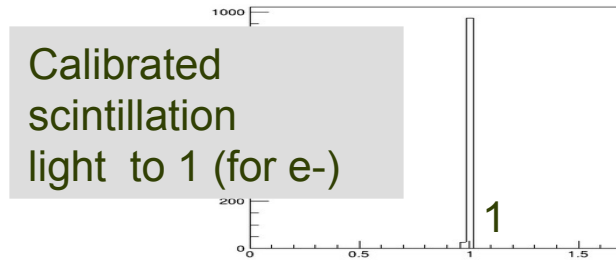
The code has 200 nm cut, but optical properties are tabulated starting from 300 nm

# Scintillation light vs Cherenkov

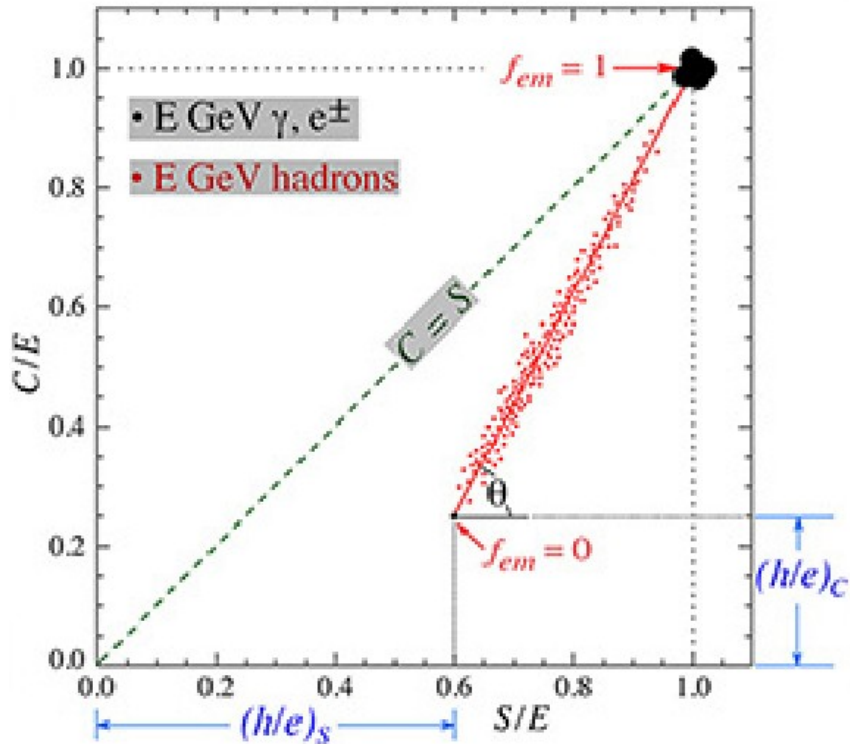


# Calibrated Scintillation vs Cherenkov

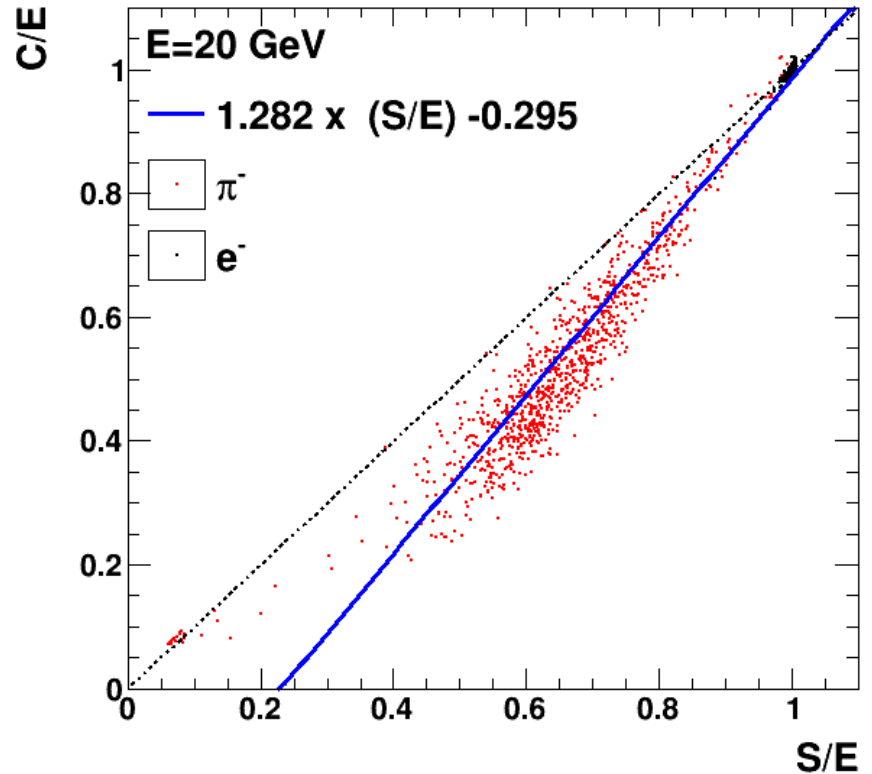
- calibrate Scintillation and Cherenkov light to electrons:  
 $N(\text{optical photons})/E(\text{beam}) = 1$  using 20 GeV electrons



DREAM: 200 GeV



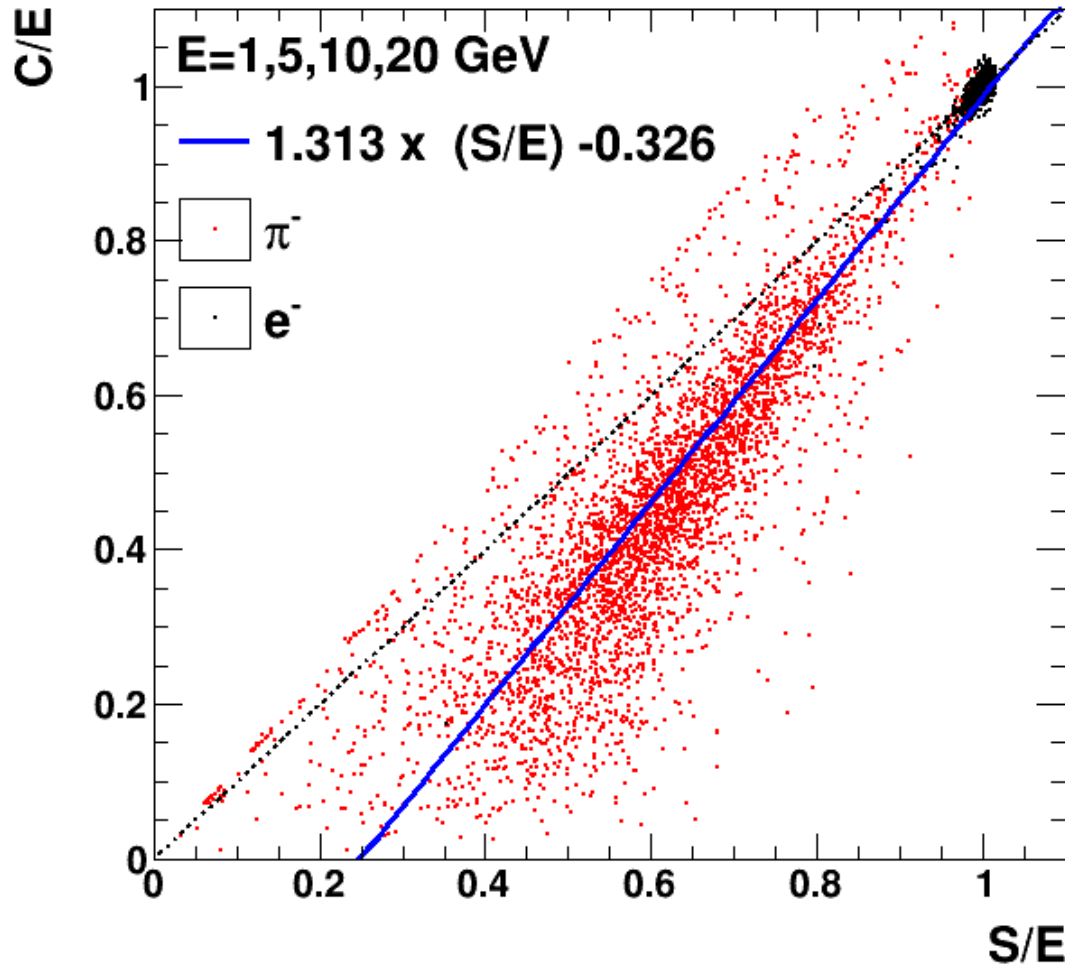
20 GeV simulations



Dual-Readout Calorimetry  
 Sehwook Lee, Michele Livan, Richard Wigmans  
 Fig. 8: <https://arxiv.org/pdf/1712.05494.pdf>

Fit includes electrons to force the fit to cross (1,1)

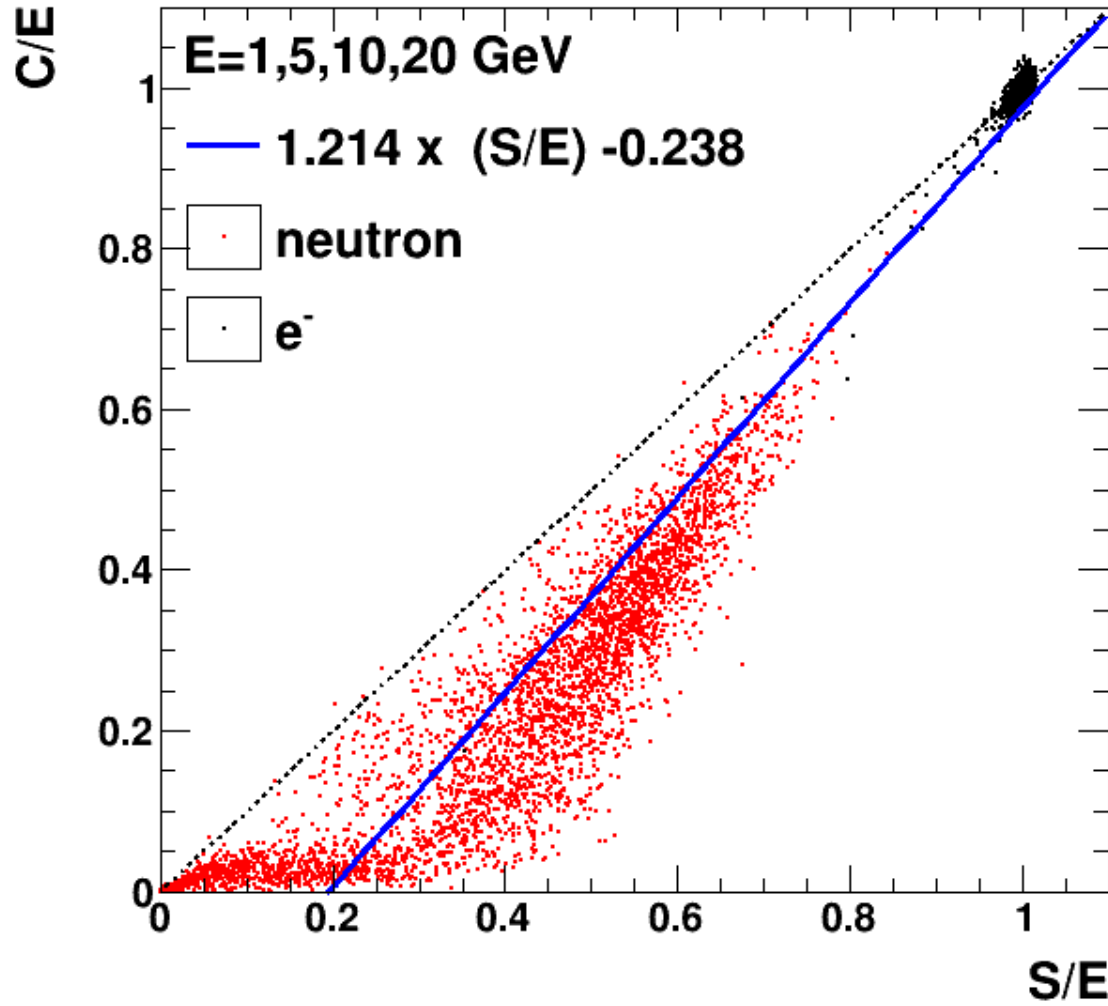
# Pions: Scintillation light vs Cherenkov after calibration to electrons (all energies)



All energies from 1 to 20 GeV

1000 events per energy

# Neutrons: Scintillation light vs Cherenkov after calibration



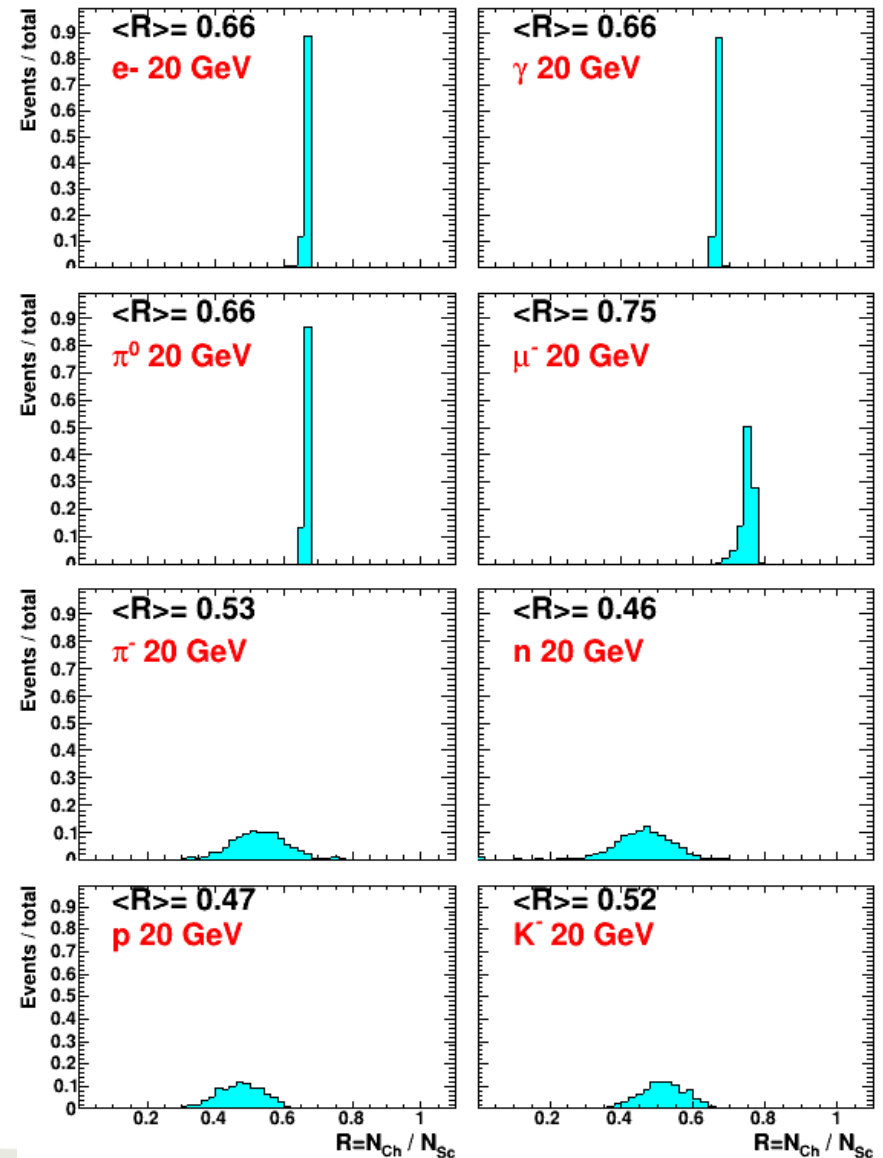
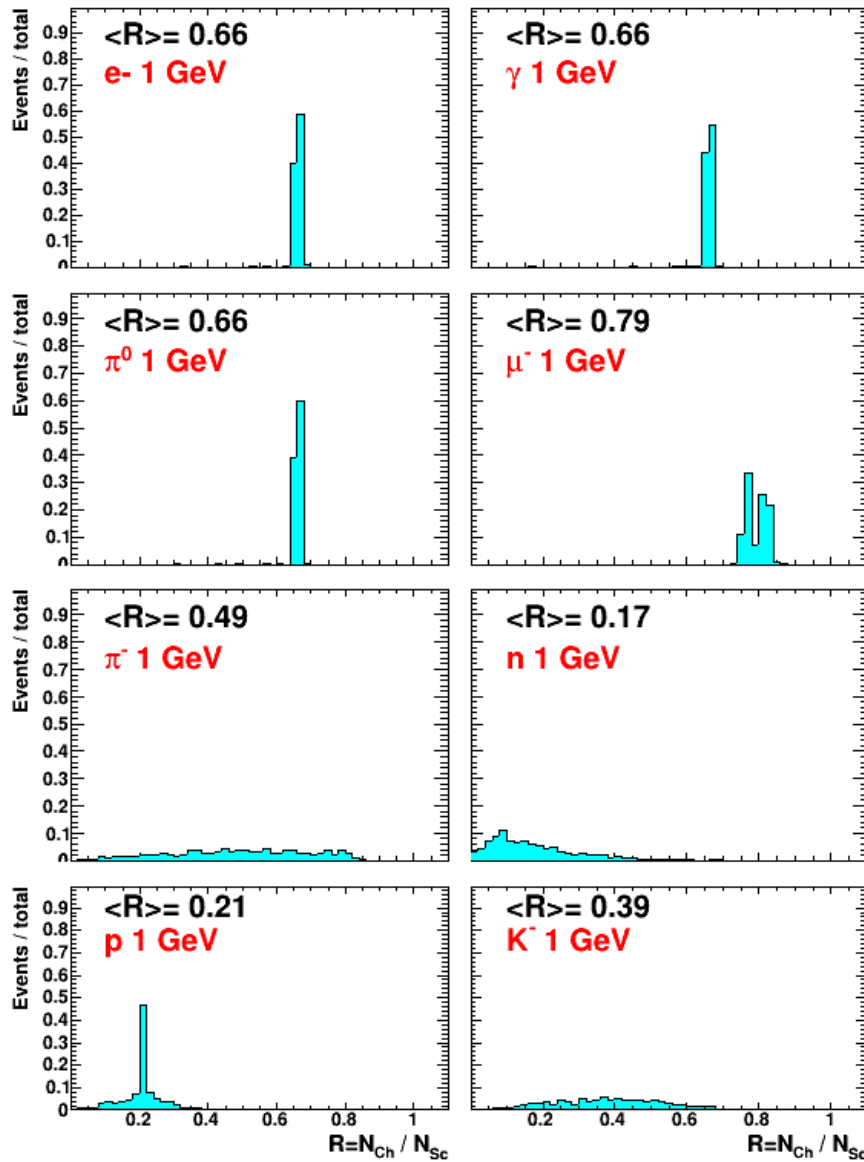
Neutrons have many events with almost no Cherenkov light



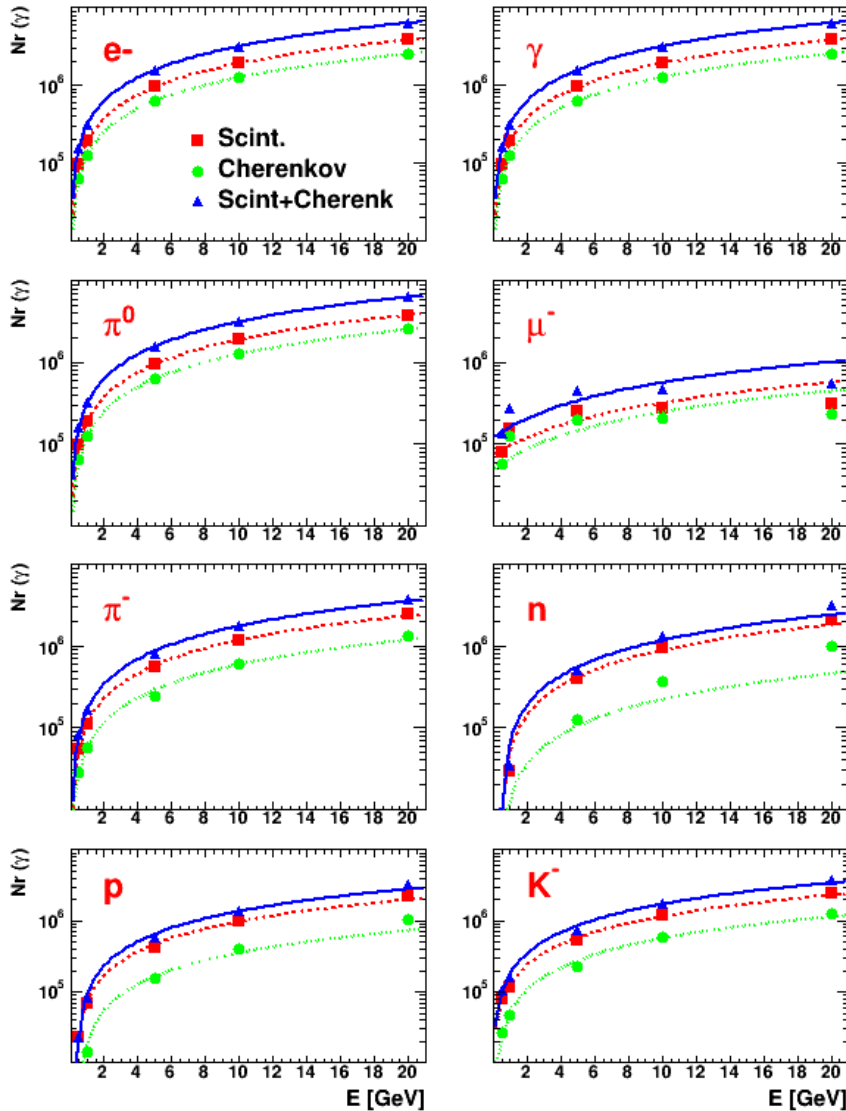
# Ratio of Cherenkov / Scintillation photons

1 GeV

20 GeV



# Rate of optical photons



Rate of Scintillation+Cherenkov  
Fit with straight lines

**e-: 191049 x E - 385**

$h/e = 0.58$  for Scintillation light  
 $h/e = 0.39$  for Cherenkov light  
(calculated using e- and **pi-** at 5 GeV)

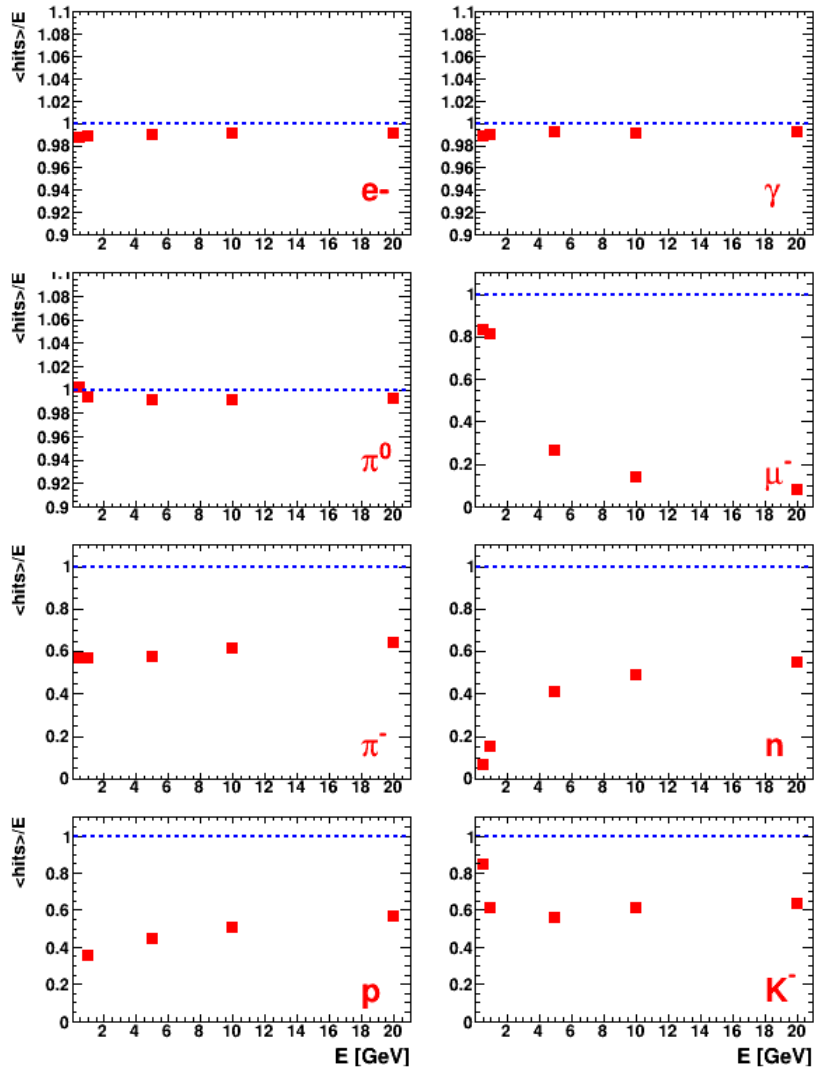
$h/e = 0.45$  for Scintillation light  
 $h/e = 0.25$  for Cherenkov light  
(calculated using e- and **protons** at 5 GeV)

$$\chi = \frac{1 - \eta_S}{1 - \eta_C} = 0.68$$

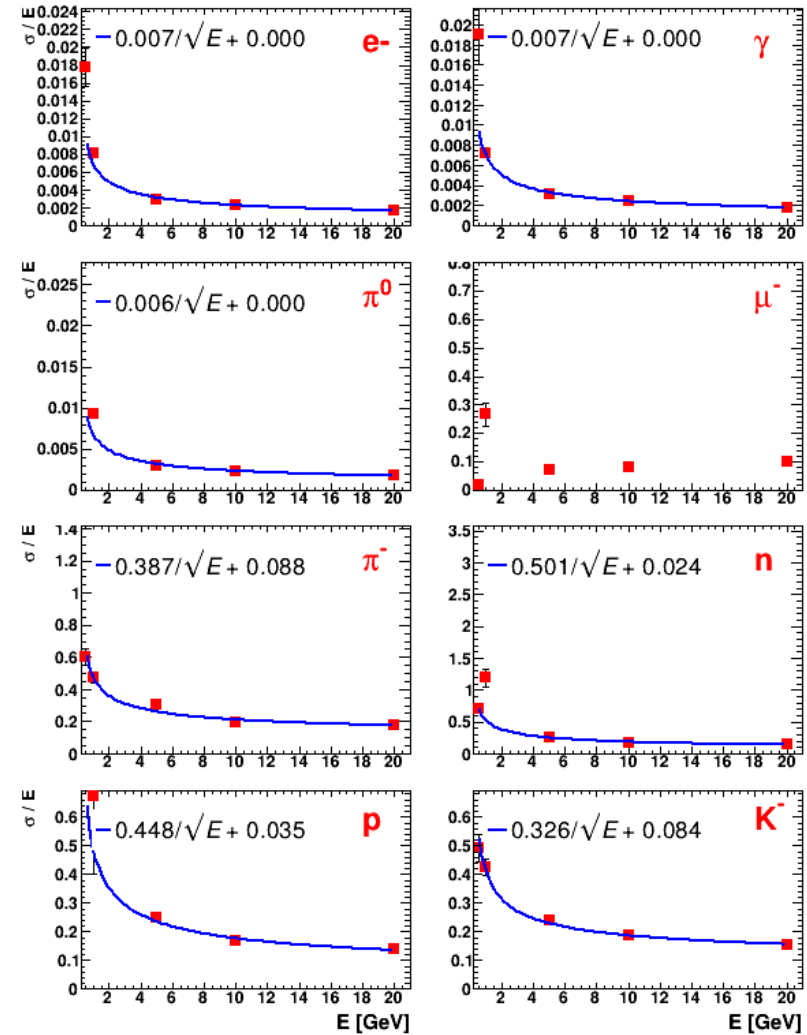
$$E = \frac{S - \chi C}{1 - \chi} \quad \leftarrow \text{correction formula}$$

<https://arxiv.org/pdf/1712.05494.pdf>

# Response and energy resolution of hits



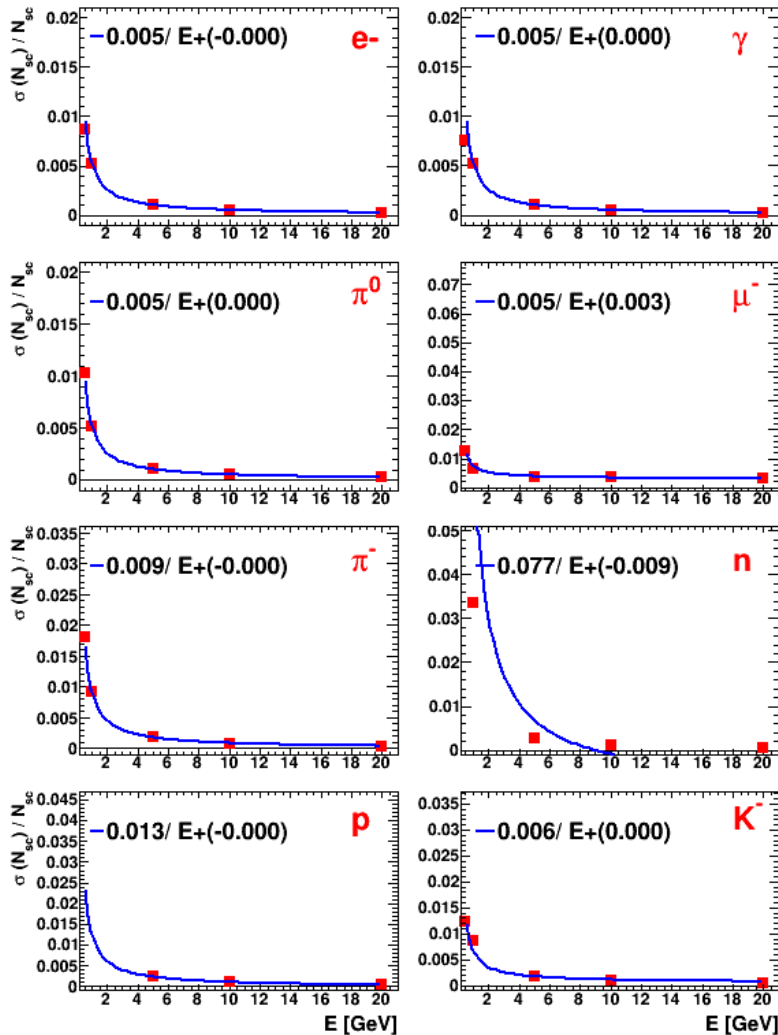
Response



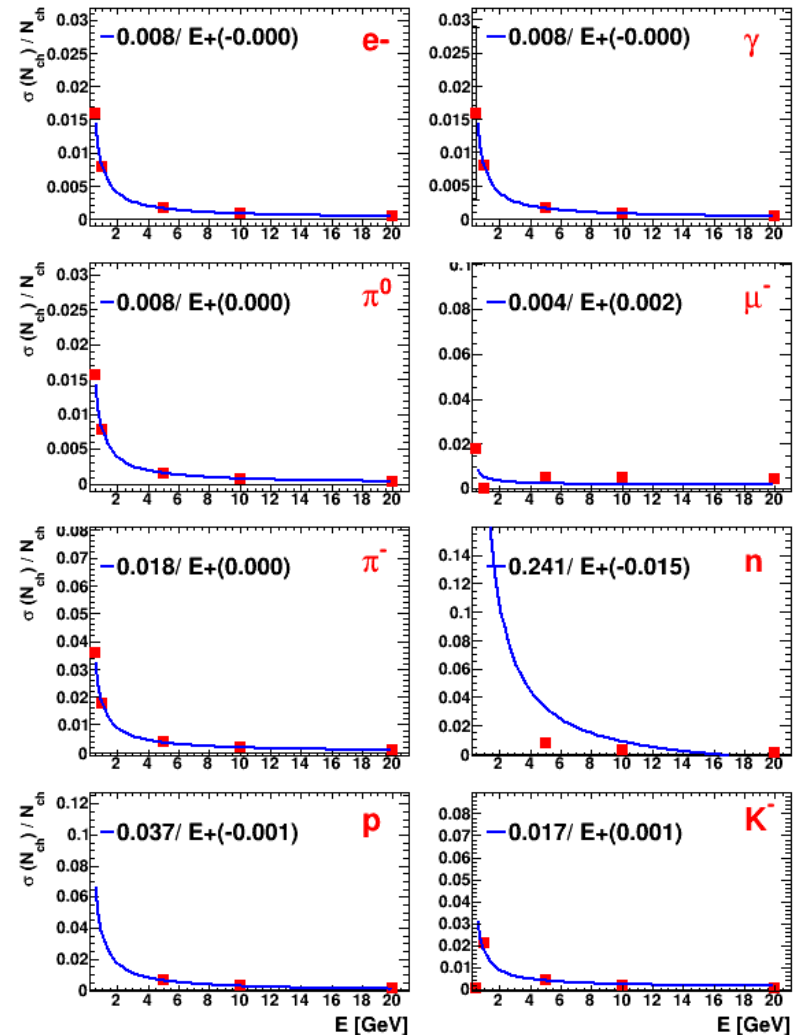
Resolution

# Resolution of Scintillation and Cherenkov $\gamma$

## Scintillation

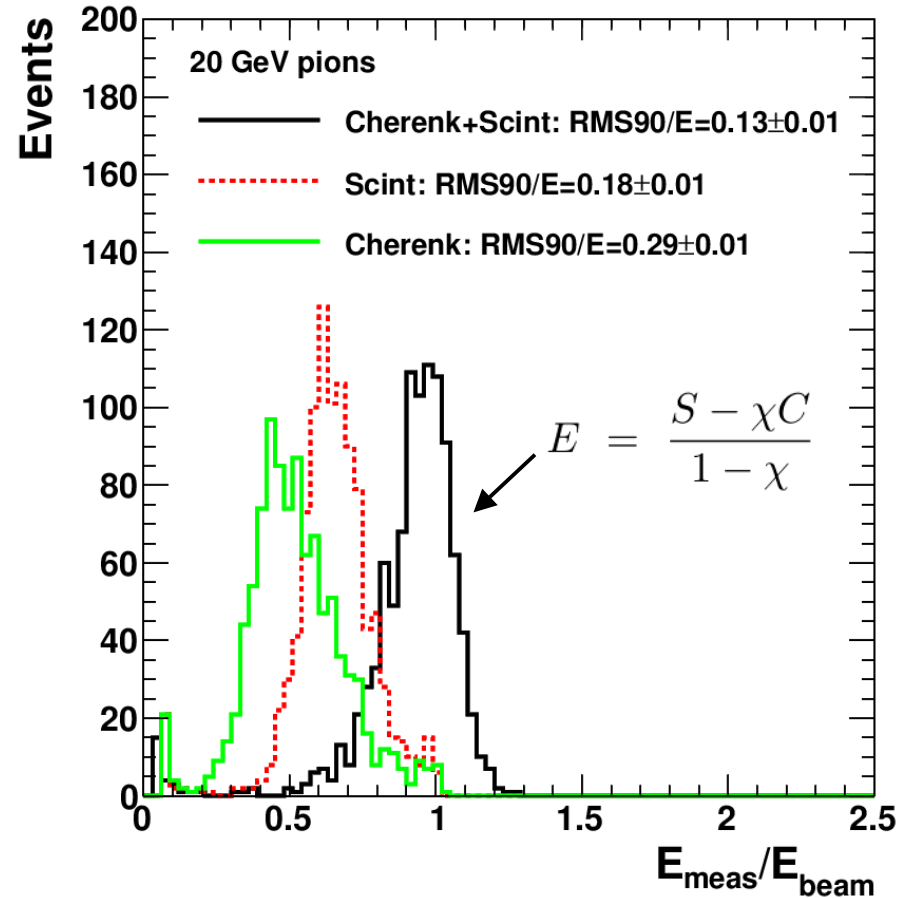
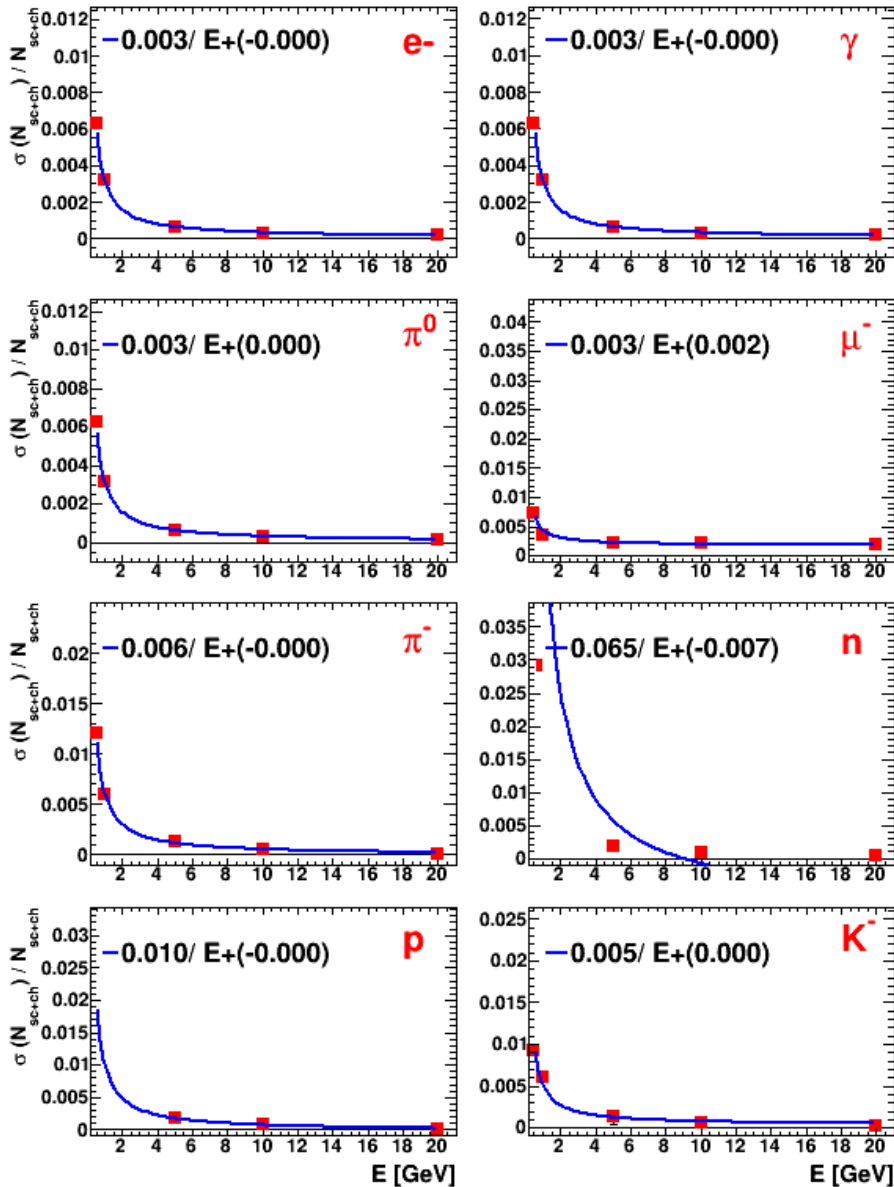


## Cherenkov



Cherenkov light has larger relative resolution (especially for hadrons)  
 $1/E$  describes the light better than  $1/\sqrt{E}$

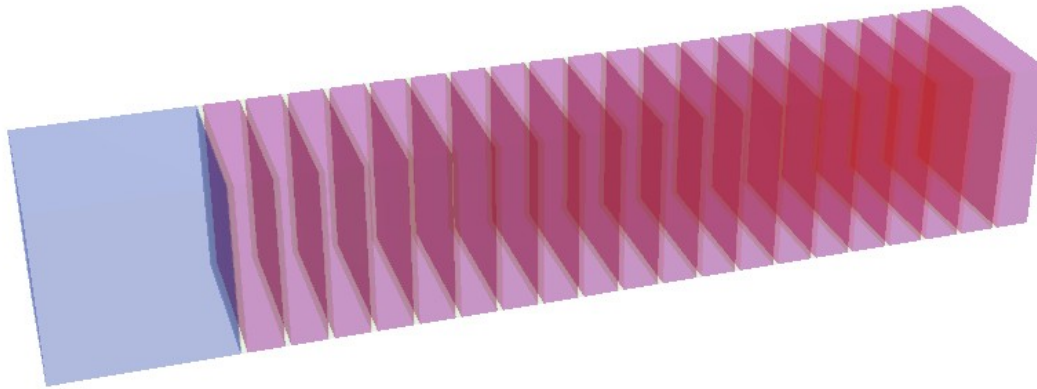
# Resolution of Scintillation + Cherenkov photons



Some improvements in the resolution

# Simplified sampling HCAL

What about sampling option?



↑  
Place for ECAL  
("vacuum")

**20 layers**

→ 5.2 interaction lengths

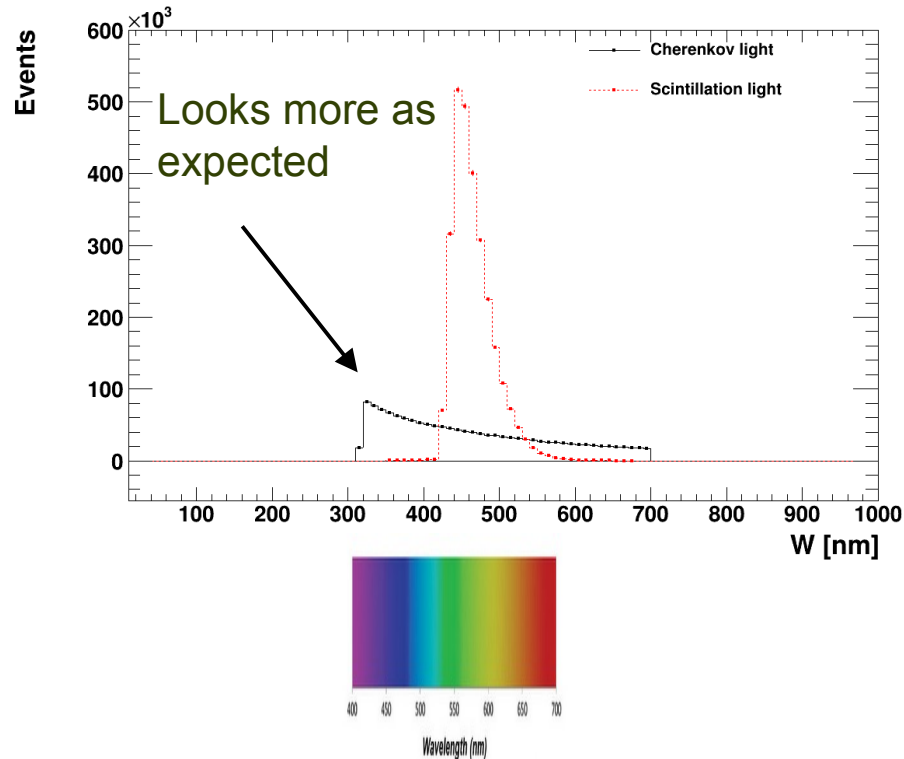
**Each layer has**

- *4 cm steel (red color)*
- *0.5 cm of PbWO<sub>4</sub>*  
*surrounded by kill media*
- *Sampling fraction ~0.11*

- Geant4 is about a factor 10 faster than for the solid crystals since photons are created in the thin active layers
- "killing" optical photons after simulation may not be needed



# Wavelength for Cherenkov and Scintillation light



Counting of photons in [300-700] nm range

No any instrumental filter, efficiency of SiPMT etc.

“**Luminosity per MeV**” calculated as the average number of photons per MeV (for e-):

Scintillation light:

**193 + -1 per MeV**

Agrees with expectations for PbWO<sub>4</sub>

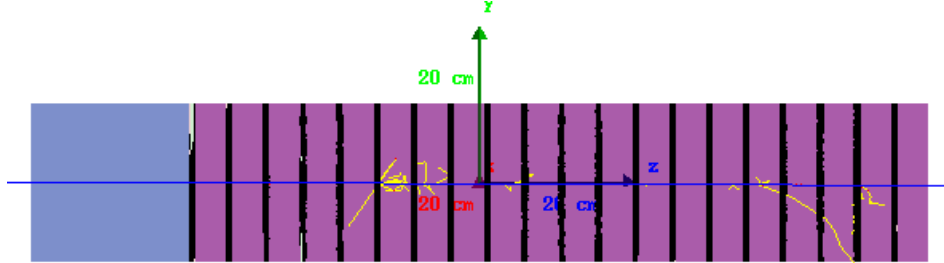
see <http://scintillator.lbl.gov/>

**Scintillation light peak: ~460 nm**

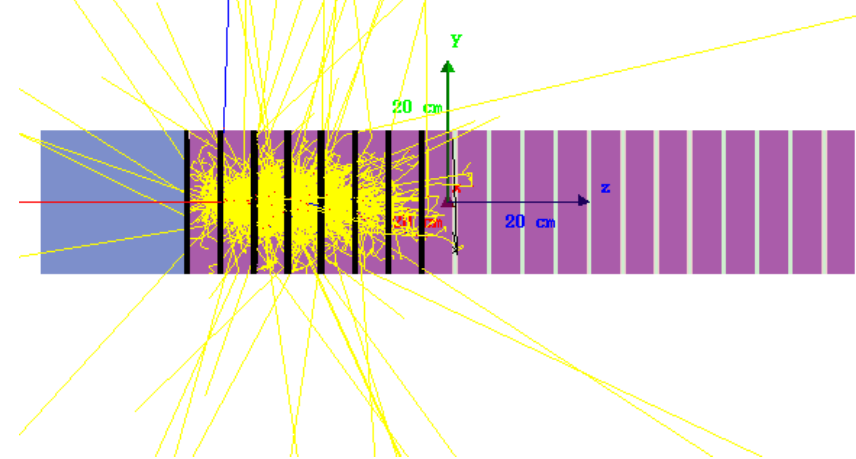
Problem: Cherenkov cannot be generated below 300 since optical properties are not defined

# Simplified sampling HCAL: Examples

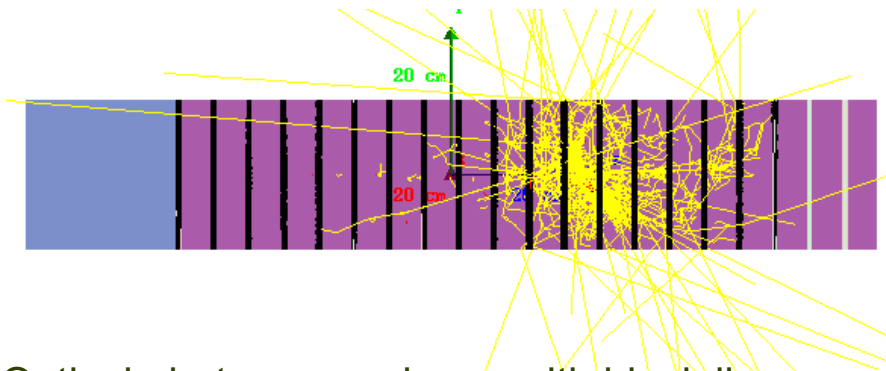
Muons 10 GeV



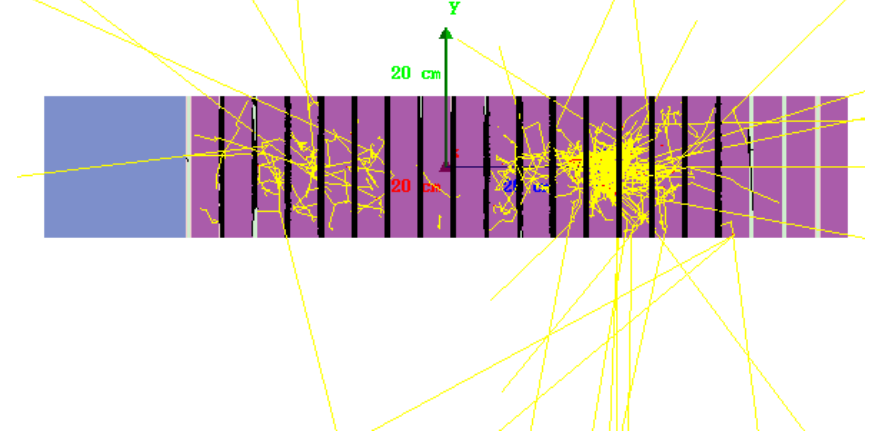
Electron 10 GeV



Pion 10 GeV



Neutron 10 GeV



Optical photons are shown with black lines  
(dark area – too many to show!)

Yellow: Photons

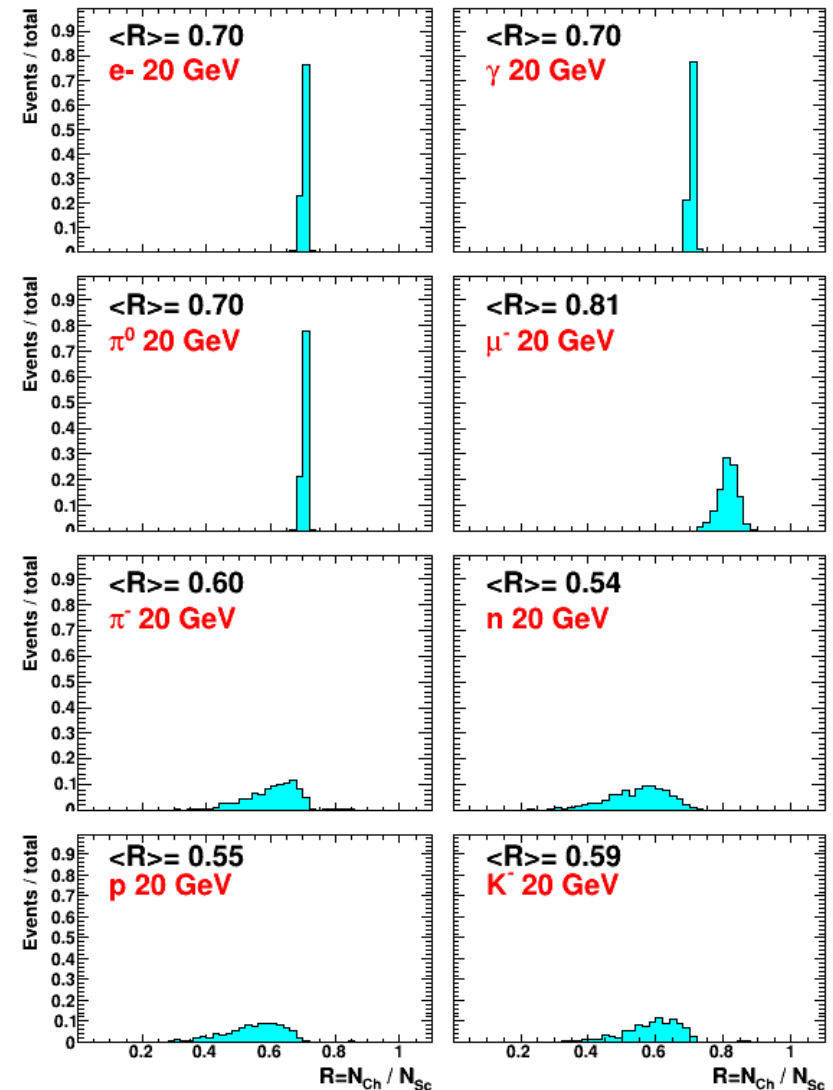
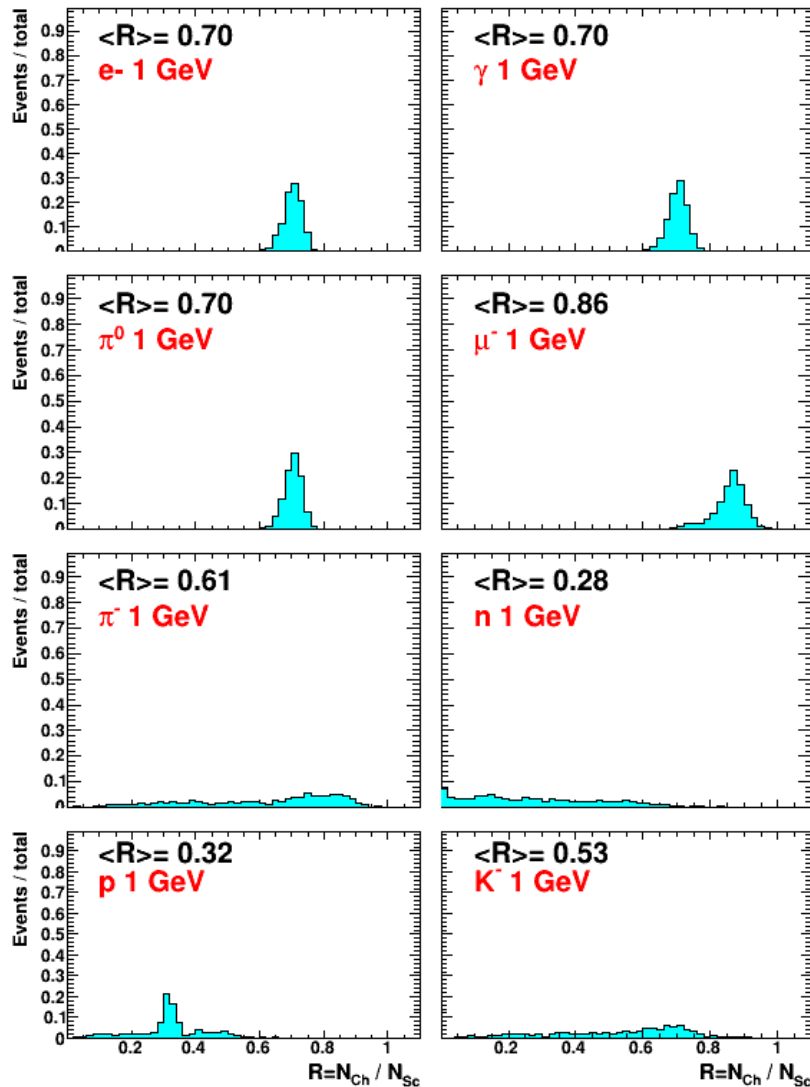




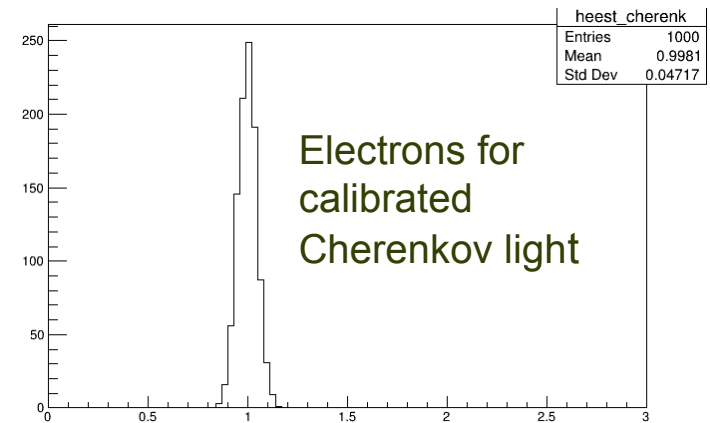
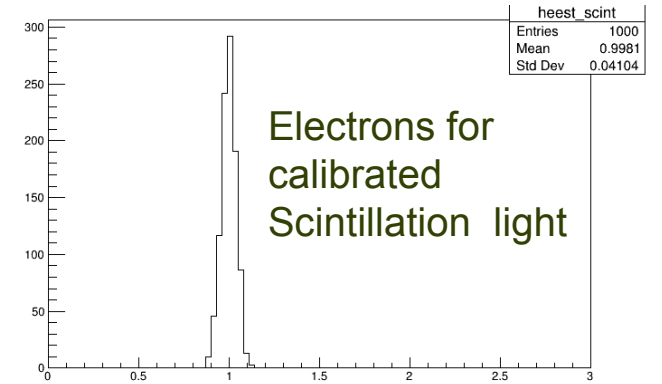
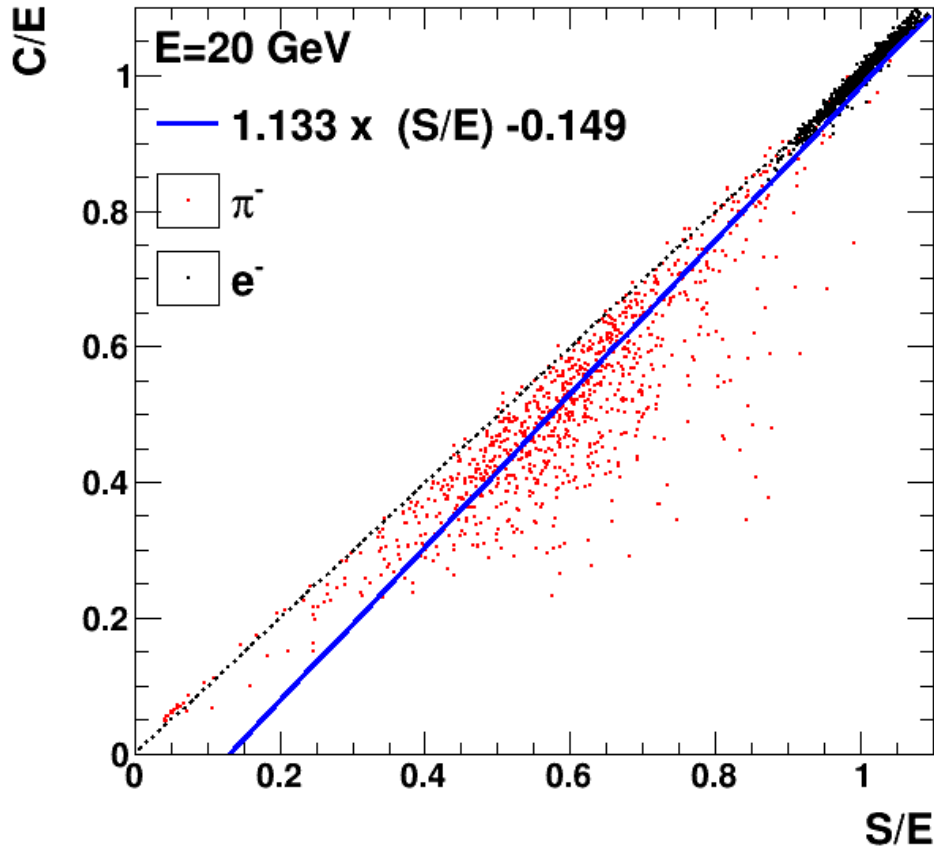
# Simplified sampling HCAL: cherenkov/scintillation

1 GeV

20 GeV

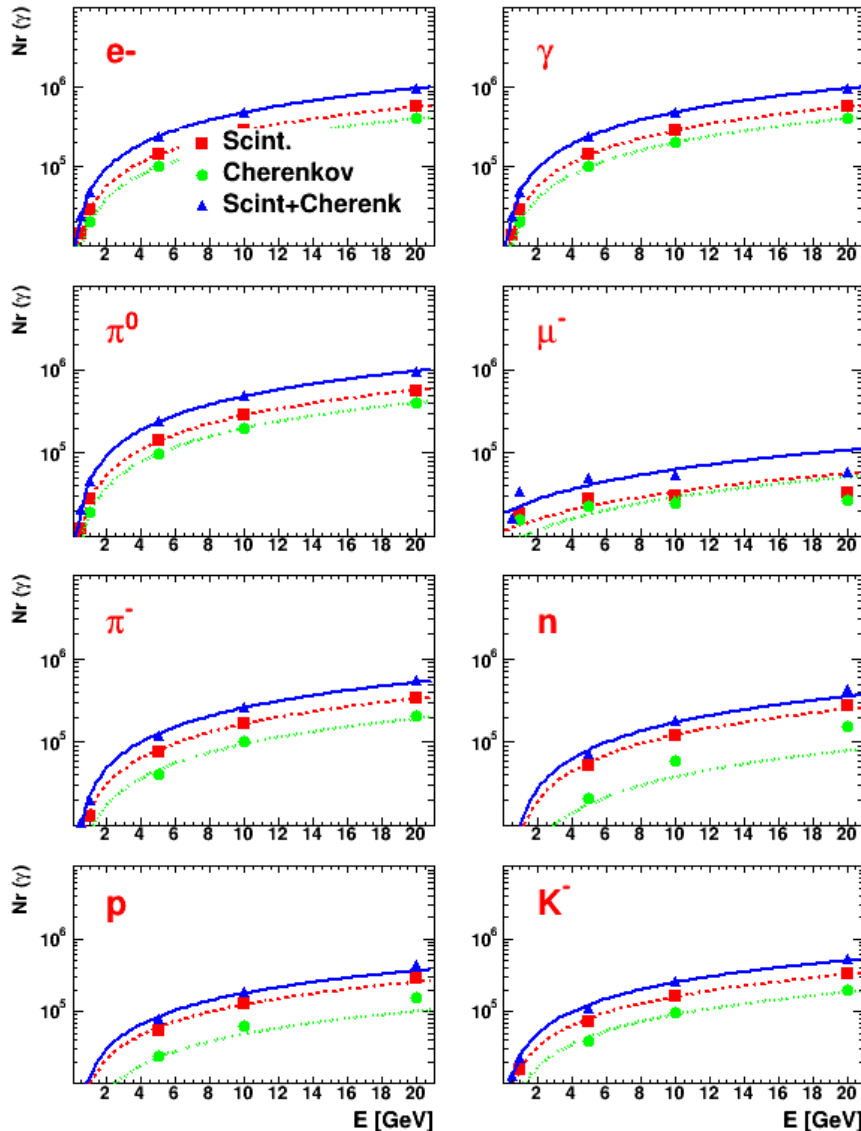


# Scintillation light vs Cherenkov after calibration



- Larger spread for electrons (expected)
- Slope for pions is weaker than for pure crystal “HCAL”

# Rate of optical photons



Rate of Scintillation+Cherenkov  
Fit with straight line:

e-:  $28236 \times E^{-2.27}$

(6.7 times lower than for pure crystals)

$h/e = 0.53$  for Scintillation light

$h/e = 0.40$  for Cherenkov light

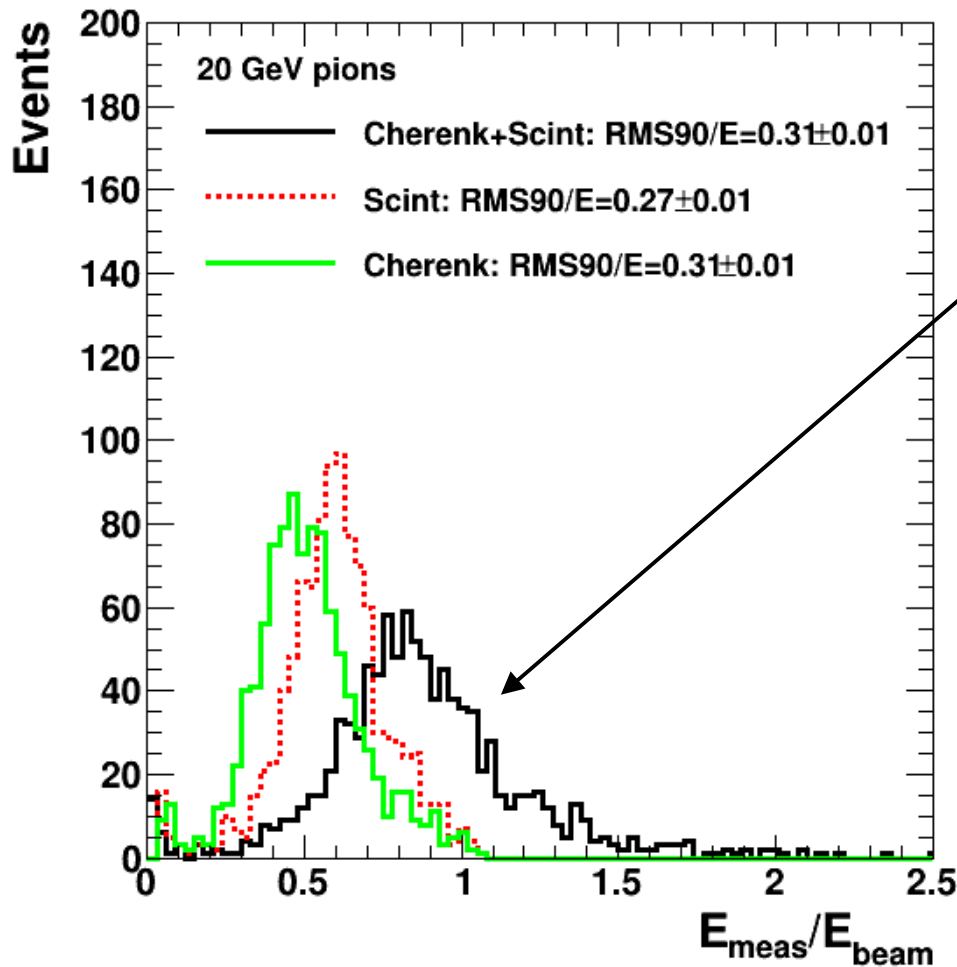
(calculated using e- and  $\pi^-$  at 5 GeV)

$$\chi = \frac{1 - \eta_S}{1 - \eta_C} = 0.79$$

$$E = \frac{S - \chi C}{1 - \chi} \quad \leftarrow \text{Corrected energy}$$

<https://arxiv.org/pdf/1712.05494.pdf>

# 20 GeV pions for calibrated response



$$E = \frac{S - \chi C}{1 - \chi}$$

No improvement in resolution

Similar conclusions for 50 GeV, protons, neutrons...

Will try to make a few variations in the simulations..