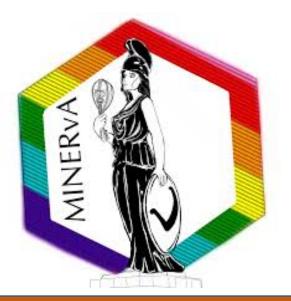
#### High Statistics Quasielastic AntiNeutrino Scattering at MINERvA

#### Amit Bashyal For the MINERvA Collaboration

10th March, 2023



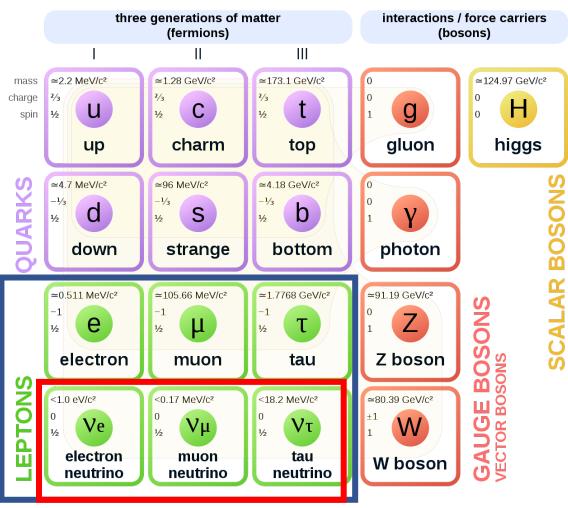
arxiv.org/abs/2211.10402



\*Now at Argonne National Lab

#### **Neutrinos In Standard Model**

#### **Standard Model of Elementary Particles**



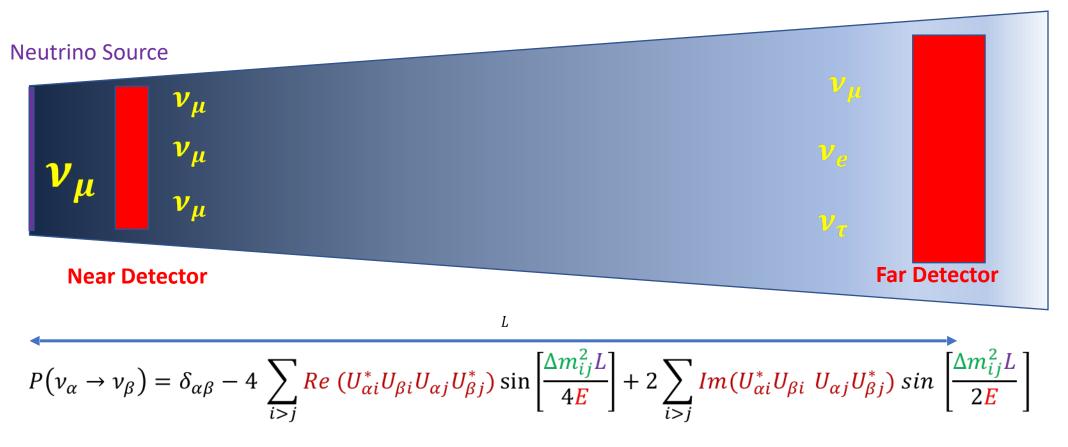
Neutrinos are electrically neutral leptons.

3 Generations of charged leptons —>3 generations of neutral massless neutrinos

$W^+ \rightarrow e^+ + \nu_e$	$W^- \rightarrow e^- + \bar{\nu}_e$
$W^+ \rightarrow \mu^+ + \nu_\mu$	$W^-  o \mu^- + \bar{\nu}_\mu$
$W^+ \rightarrow \tau^+ + \nu_{\tau}$	$W^- \to \tau^- + \bar{\nu}_\tau$

•  $v_e, v_\mu, v_\tau$  are only interact through weak force.

#### **Neutrino Oscillation Experiments**



Decreasing flux intensity

 $L \rightarrow$  Distance Between neutrino source and far detector E  $\rightarrow$  Energy of the neutrino

 $U_{\alpha i}, U_{\beta j} \rightarrow \text{Mixing matrix elements } (\alpha, \beta \rightarrow \nu_e, \nu_\mu, \nu_\tau \& i, j \rightarrow \nu_1, \nu_2, \nu_3)$  $\Delta m_{ij}^2 = m_i^2 - m_j^2 \rightarrow \text{Difference of mass of eigenstates } \nu_i, \nu_j$ 

#### **Neutrino Oscillation: Measurement**

In a  $\nu_{\mu} \rightarrow \nu_{e}$  oscillation experiment, quantity we want to measure is:

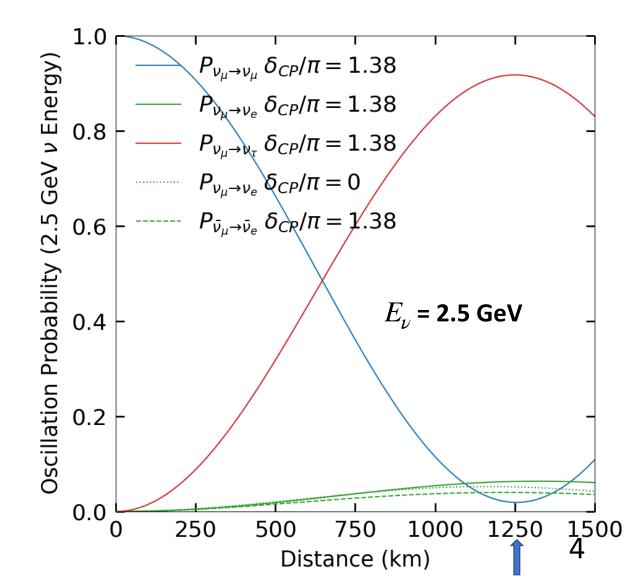
$$P_{\nu_{\mu} \to \nu_{e}}(E,L) \propto \frac{\Phi_{\nu_{e}}(E,L)}{\Phi_{\nu_{\mu}}(E,0)}$$

 $\Phi_{\nu_e}(E,L) \rightarrow$  Electron neutrino flux measured at distance L from the source

 $\Phi_{\nu_{\mu}}(E,0) \rightarrow$  Muon neutrino flux measured at the source

One of the goals of current oscillation experiments is to measure how much CP violation happens in lepton sector  $\delta_{CP} = 0 \rightarrow \text{No CP Violation} \quad P(\nu_{\mu} \rightarrow \nu_{e}) = P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e})$  $\delta_{CP} \neq 0 \rightarrow \text{CP Violation} \quad P(\nu_{\mu} \rightarrow \nu_{e}) \neq P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e})$ 

Neutrino experiments don't measure flux directly. They measure the number of events N as a function of reconstructed neutrino energy  $E_{rec}$ .



#### **Neutrino Oscillation: Measurement**

 $N_e(E_{rec},L) \propto \sum \Phi_e(E,L)\sigma_i(E)f_{\sigma i}(E,Erec)dEdM$ 

Estimation of neutrino flux: MINERvA employs various methods to accurately estimate neutrino flux and constrain flux related uncertainties.

Neutrino Interaction Cross section: Cross section measurements improve our neutrino interaction models (Accelerator neutrino) oscillation experiments like NOvA, DUNE, T2K,HK need accurate estimations of neutrino flux ( $\Phi(E)$ ) and cross section ( $\sigma(E)$ ) to extract the oscillation parameters accurately.

**Cross section is convoluted by** 

 $\left[\sigma(E_{\nu}) * \theta(E_{\nu})\right] | \text{nuclear effects} \left(\theta(E_{\nu})\right)$ 

Nuclear effects complicate the cross-section measurements.

Smearing of true neutrino energy (E): Limitation on how accurately neutrino energy can be reconstructed ( $E_{rec}$ ) Smearing of true (E) to reconstructed ( $E_{rec}$ ) depends upon detector properties.

#### **Neutrino Oscillation: Measurement**

 $N_e(E_{rec},L) \propto \sum \Phi_e(E,L)\sigma_i(E)f_{\sigma i}(E,Erec)dEdM$ 

*i* : neutrino-nucleus interaction channel

Estimation of neutrino flux: MINERvA employs various methods to accurately estimate neutrino flux and constrain flux related uncertainties.

Neutrino Interaction Cross section: Cross section measurements improve our neutrino interaction models

MINERvA was designed to study the poorly understood neutrino-heavy nucleus interactions

- MINERvA has developed a comprehensive flux strategy to constrain the neutrino flux uncertainties.
- MINERvA has provided cross section measurements of various neutrino scattering processes over a wide neutrino energy range.
- Measurements in various target materials has helped us to understand nuclear effects in neutrino-heavy nucleus interactions.
- Neutrino oscillation experiments often use heavy target material detectors.

Understanding nuclear effects is crucial to reconstruct neutrino energy in near and far detectors.

#### **Neutrino Production in NuMI Beamline**

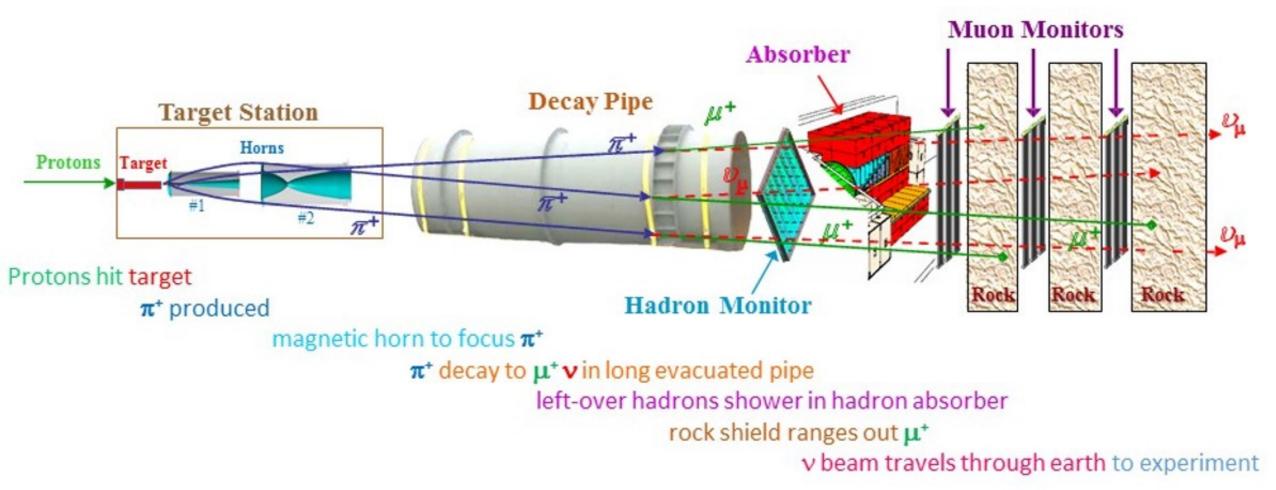
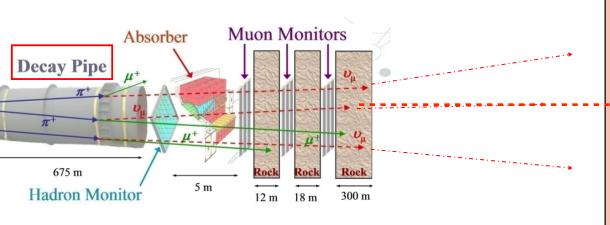
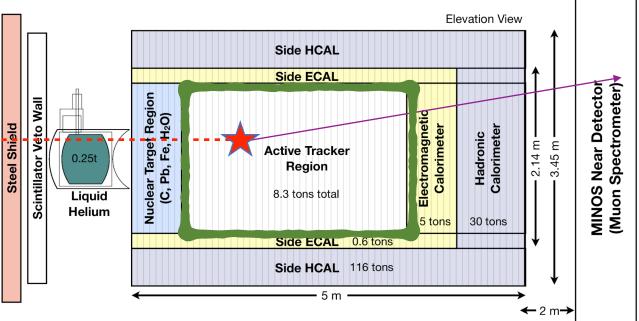


Image Credit: Fermilab

Many thanks to Accelerator Division for the awesome neutrino beam!

## **MINERvA Experiment**





- Main Injector Neutrino Nucleon (A) experiment
- MINERvA is a neutrino cross-section measurement experiment
- Downstream of the NuMI beam line

- MINERvA detector  $\rightarrow$  1.032 km from the NuMI target
- MINOS near detector  $\rightarrow$  1.04 km from the NuMI target
  - MINOS near detector is magnetized.
  - Identify charge of muon from interaction

#### **MINERvA Experiment**

#### The MINERvA Collaboration



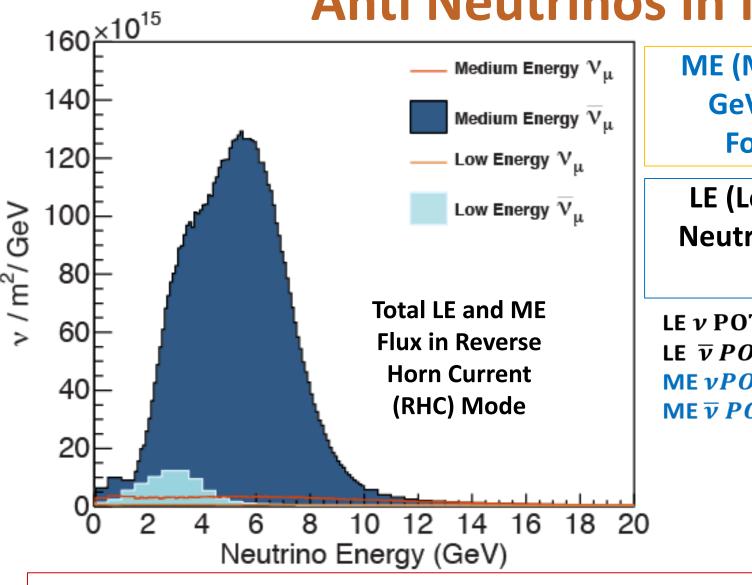
Detector shutting down for the last time.

#### Front face of the MINERvA detector



Celebrating for collecting 30e20 POT

## Anti Neutrinos in MINERvA



decay to give  $\bar{\nu}_{\mu}$ .

**ME (Medium Energy) = 6 GeV Neutrino Energy Focusing Peak Run** 

LE (Low Energy) = 3 GeV **Neutrino energy Focusing** Peak Run

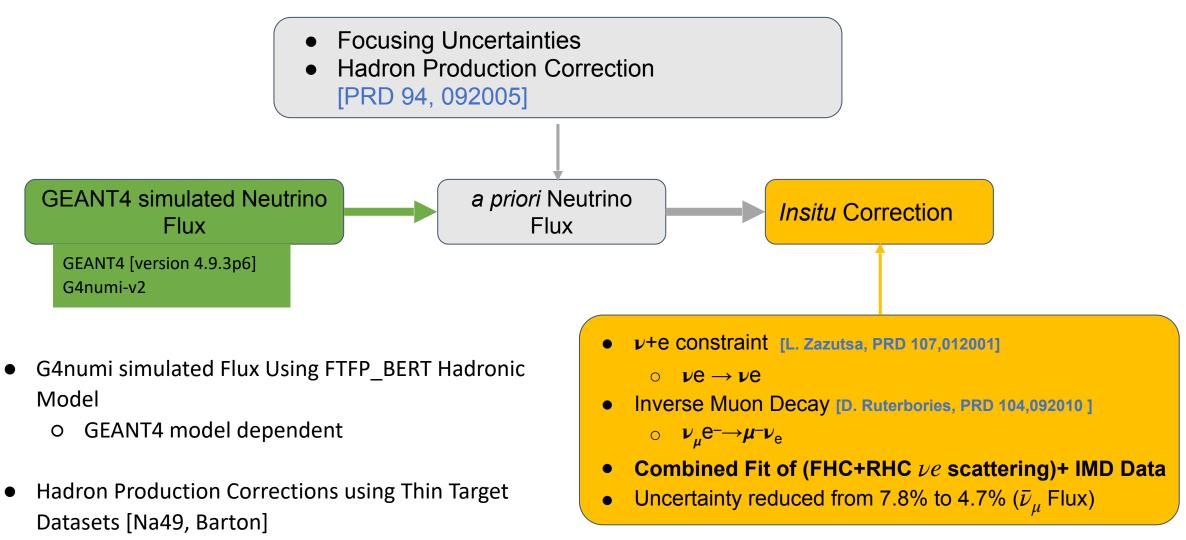
LE  $\nu$  POT: 4. 0×10<sup>20</sup> LE  $\overline{\nu} POT$ : 1. 7×10<sup>20</sup> ME  $\nu POT \sim 3 \times LE \nu POT$  $\mathsf{ME}\,\overline{\boldsymbol{\nu}}\,\boldsymbol{POT}\sim\boldsymbol{7}\times\boldsymbol{LE}\,\overline{\boldsymbol{\nu}}\,\boldsymbol{POT}$ 

 $LE \rightarrow 2$  years run ME  $\rightarrow$  5.5 years run

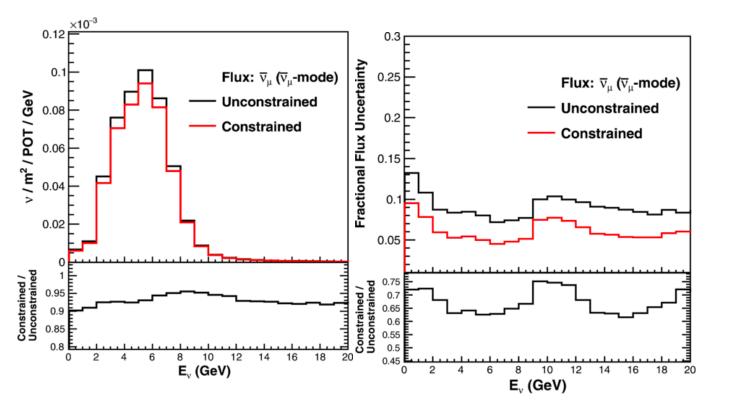
**ME data:** 

**High Statistics+ High Energy** 

# **MINERvA Flux Strategy**



## Insitu correction on the Flux

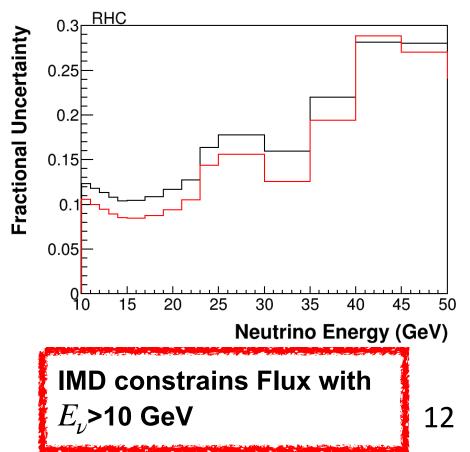


*a priori* anti neutrino Flux before and after constrained by *v*e scattering data Uncertainty on predicted anti neutrino flux before and after constrained by ve scattering data

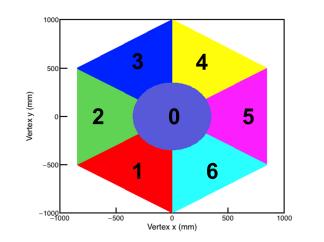
Combined Fit of (FHC+RHC  $\nu e$  scattering) + IMD Data Uncertainty reduced from 7.8% to 4.7% ( $\bar{\nu}_u$  Flux)  ν+e constraint [L. Zazueta, PRD 107,012001]

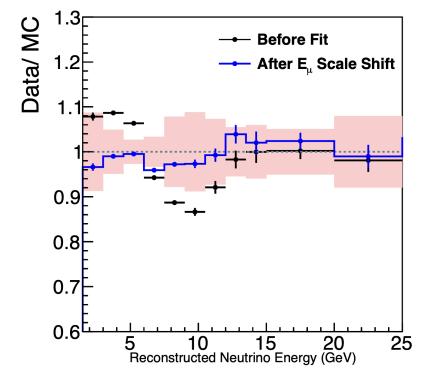
 $\circ \nu e \rightarrow \nu e$ 

 Inverse Muon Decay [D. Ruterbories, PRD 104,092010]



#### Low nu Fit to resolve discrepancy





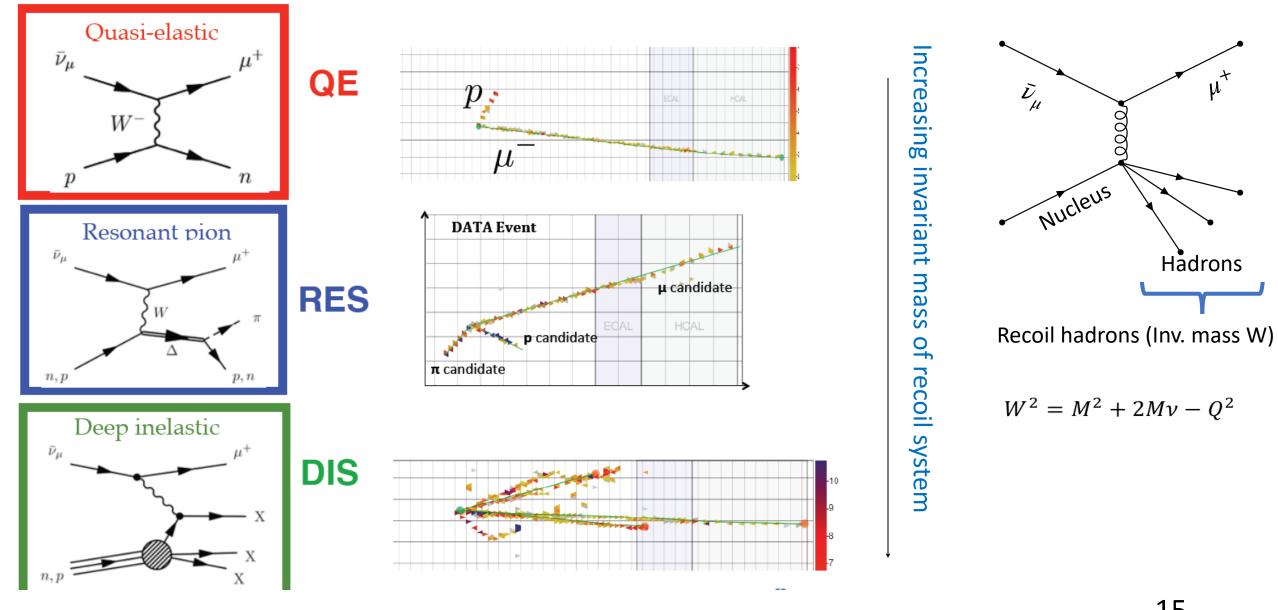
- MINERvA saw discrepancy between its data and simulated sample.
- Discrepancy due to shift in energy spectrum
- A multi parameter fit
  - Fit done with Low recoil sample
    - Cross-section independent of  $E_{\nu}$ —> Shape depends on Flux
  - Focusing parameters and MINOS muon energy scale as fit parameters
- Shift of MINOS muon energy scale by 1.8  $\sigma$  to resolve discrepancy

#### A. Bashyal,arXiv:2104.05769

Parameter	Nominal	Best Fit (No Prior)	Best Fit (Prior)
Beam Position (X)	0.0 mm	$-0.3 \pm 0.3 \pm 0.1 \text{ mm}$	$-0.3 \pm 0.2 \pm 0.1 \text{ mm}$
Beam Position (Y)	0.0 mm	$0.8 \pm 0.3 \pm 0.3$ mm	$0.7 \pm 0.2 \pm 0.2 \text{ mm}$
Target Position (X)	0.0 mm	$-0.8 \pm 0.3 \pm 0.1 \text{ mm}$	$-0.8 \pm 0.3 \pm 0.1 \text{ mm}$
Target Position (Y)	0.0 mm	$2.3 \pm 0.7 \pm 1.2 \text{ mm}$	$1.7 \pm 0.6 \pm 0.8 \text{ mm}$
Target Position (Z)	-1433 mm	$-1432.4 \pm 2.4 \pm 0.3$ mm	$-1431 \pm 1.8 \pm 0.3$ mm
Horn 1 Position (X)	0.0 mm	$-0.3 \pm 0.4 \pm 0.5 \text{ mm}$	$-0.1 \pm 0.3 \pm 0.1 \text{ mm}$
Horn 1 Position (Y)	0.0 mm	$0.1 \pm 0.5 \pm 0.5 \text{ mm}$	$0.0 \pm 0.3 \pm 0.3$ mm
Beam Spot Size	1.5 mm	$1.41 \pm 0.09 \pm 0.03$ mm	$1.32 \pm 0.09 \pm 0.03$ mm
Horn Water Layer	1.0 mm	$1.2 \pm 0.3 \pm 0.05 \text{ mm}$	$1.3 \pm 0.25 \pm 0.1 \text{ mm}$
Horn Current	200 kA	$198.0 \pm 1.4 \pm 1.4$ kA	$199.1 \pm 0.7 \pm 0.5 \text{ kA}$
Muon Energy Scale	1.0	$1.032 \pm 0.004 \pm 0.008$	$1.036 \pm 0.004 \pm 0.006$

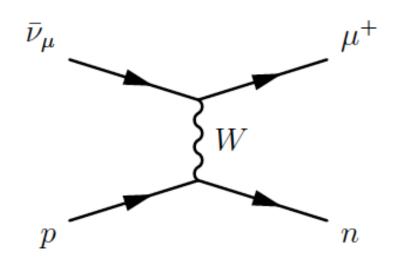
# $\bar{\nu}_{\mu} \text{ Charged Current Quasi} \\ \text{ Elastic Cross-section} \\ \text{ Measurement in CH Target} \\ \end{array}$

#### **General Neutrino Nucleon Interactions in MINERvA**



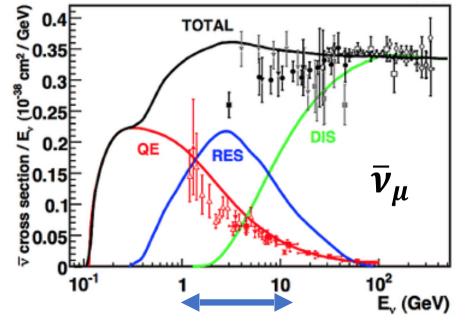
Hadrons

## **CCQE Interactions**



Feynman diagram of an anti muon neutrino CCQE interaction

CCQE Processes  $\nu_l + n \rightarrow l^- + p$   $\bar{\nu}_l + p \rightarrow l^+ + n$  $l \rightarrow e, \mu, \tau \quad n \rightarrow neutron \quad p \rightarrow proton$ 



**MINERVA, DUNE, NOVA** 

Oscillation experiments like NOvA and DUNE (will) see CCQE interactions as one of the major interactions.

**CCQE cross-section** measured by MINERvA will help oscillation experiment to understand their data.

*Z.A Formaggio and Z.P. Zeller, Rev. Mod. Phys. 84, 1307-1341, 2012* 

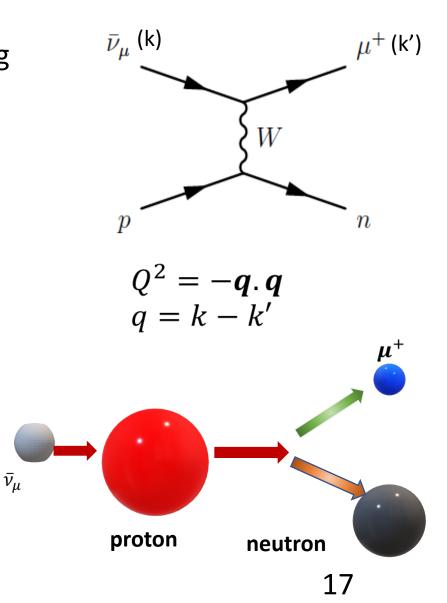
Anti-neutrino nucleon cross-section (per nucleon) for different interaction channels. Red: QE cross-section

#### **Neutrino Nucleon QE Scattering**

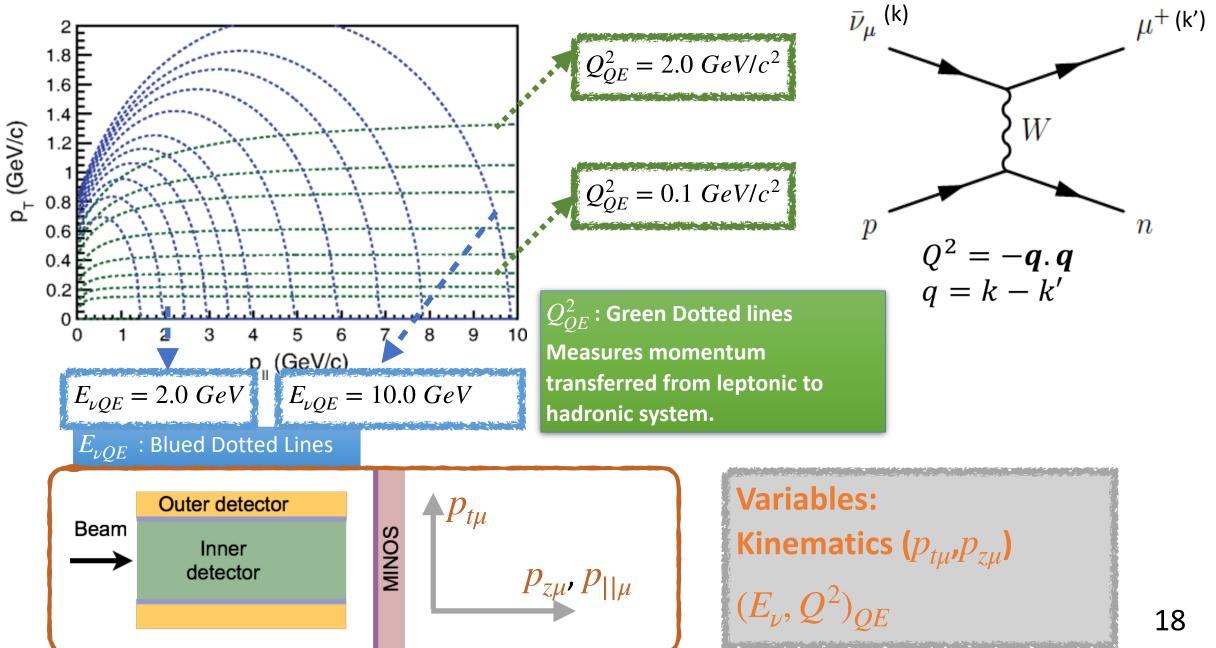
- CCQE : Relatively clean process
- Assuming the nucleon is at rest, the energy of the incoming neutrino can be reconstructed using the kinematics of outgoing muon.

$$E_{\nu}^{QE} = \frac{m_n^2 - (m_p - E_b)^2 - m_{\mu}^2 + 2(m_p - E_b)E_{\mu}}{2(m_p - E_b - E_{\mu} + p_{\mu}\cos\theta_{\mu})}$$
$$Q_{QE}^2 = 2E_{\nu}^{QE}(E_{\mu} - p_{\mu}\cos\theta_{\mu}) - m_{\mu}^2$$

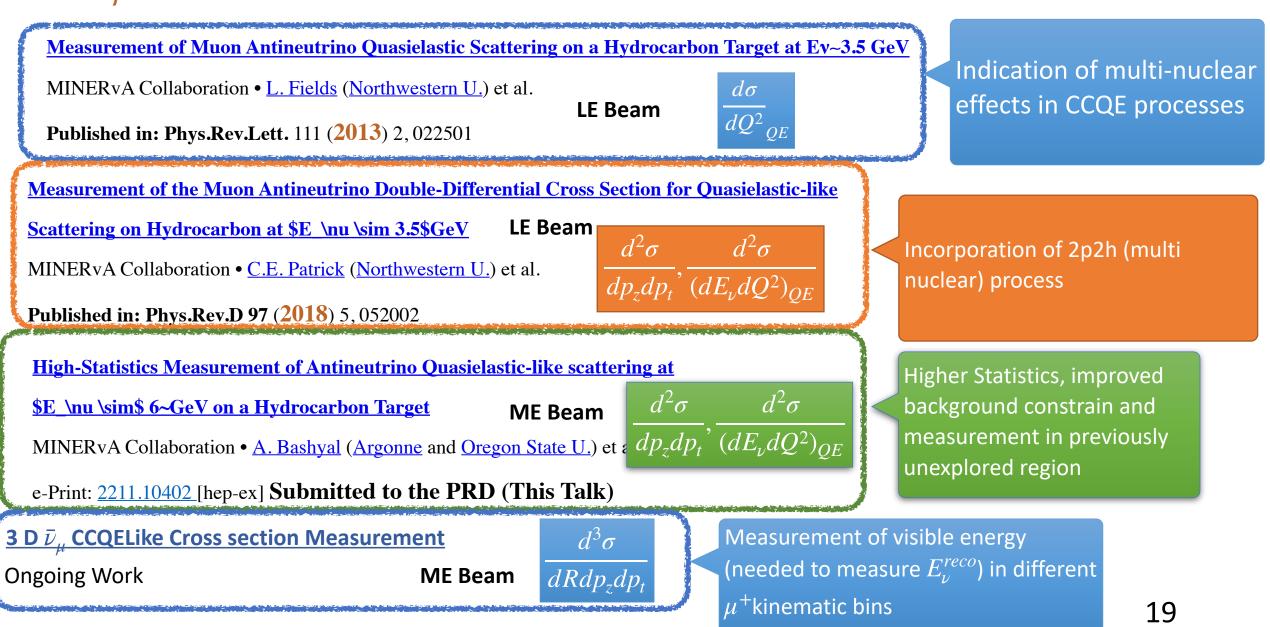
$$\begin{split} m_p &
ightarrow mass \ of \ proton \ m_n &
ightarrow mass \ of \ neutron \ E_b &
ightarrow Binding \ Energy \ of \ the \ neutron \ in \ nucleus \ E_\mu &
ightarrow Energy \ of \ the \ outgoing \ muon \ p_\mu &
ightarrow Momentum \ of \ the \ outgoing \ muon \ Q^2 &
ightarrow Four \ momentum \ transferred \ squared \end{split}$$



#### **Cross section Measurement Variables**



## $\bar{\nu}_{\mu}$ CCQELike (CH Target) Measurement in MINERvA

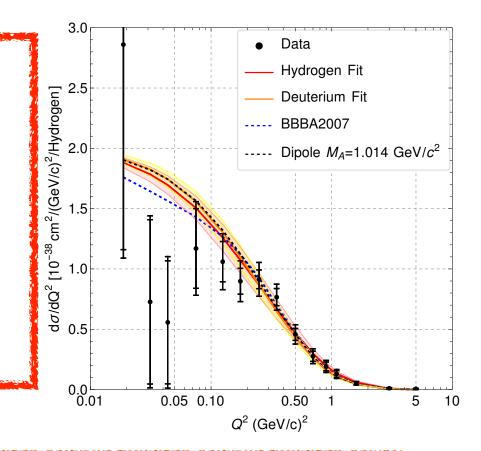


# Measurement of the axial vector form factor from antineutrino-proton scattering in $MINER\nu A$

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Joint Experimental-Theoretical Physics Seminar, Feb 1, 2023



Article | Open Access | Published: 01 February 2023

#### Measurement of the axial vector form factor from antineutrino-proton scattering

T. Cai <sup>C</sup>, M. L. Moore, A. Olivier, S. Akhter, Z. Ahmad Dar, V. Ansari, M. V. Ascencio, A. Bashyal, A. Bercellie, M. Betancourt, A. Bodek, J. L. Bonilla, A. Bravar, H. Budd, G. Caceres, M. F. Carneiro, G. A. Díaz, H. da Motta, J. Felix, L. Fields, A. Filkins, R. Fine, A. M. Gago, H. Gallagher, ... L. Zazueta + Show authors

 Nature
 614, 48–53 (2023)
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 6757
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 158
 Altmetric
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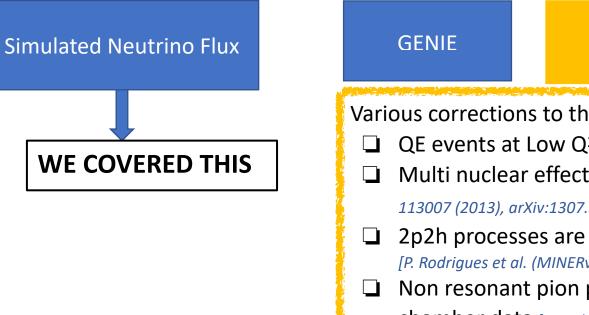
Signal of above analysis is subset of this analysis' Signal.

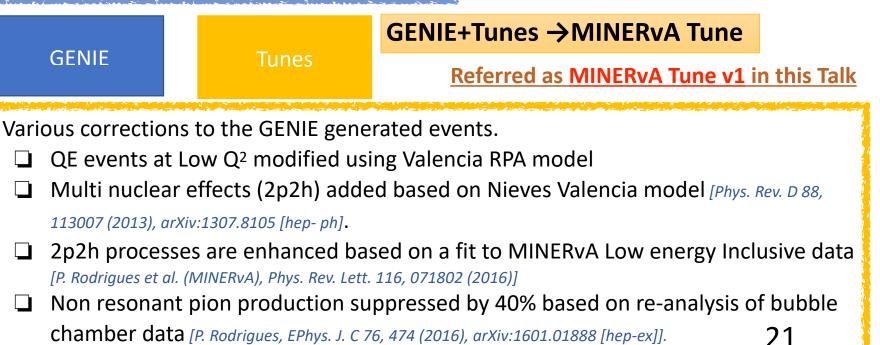
#### Simulation of Neutrino Nucleon Interaction in MINERvA

**GENIE 2.12.6** [*arXiv:1510.05494* [*hep-ph*]] is our neutrino MC generator:

- □ Nuclear Model (initial state) → Relativistic Fermi Gas (RFG) Model for initial nuclear state with an additional high energy tail as prescribed by Bodek and Ritchie [A. Bodek and J. Ritchie, Phys. Rev. D 23, 1070 (1981)].
- □ Final State Interaction of hadrons →INTRANUKE h-A model
- **QE** Process  $\rightarrow$  Llewellyn-Smith formalism + BBA05 with M<sub>A</sub> = 0.99 GeV C.H. Llewellyn Smith, Phys. Rept. 3C, 261 (1972)
- □ Resonance  $\pi$  production  $\rightarrow$  Rein Sehgal model with M<sub>A</sub>= 1.12 GeV
- □ Deep Inelastic Scattering → Bodek- Yang Model [A. Bodek,arXiv:hep-ph/0411202 [hep-ph].]

Full Detector Simulation (GEANT4) to simulate the response of the detector for particles that interact with the detector.



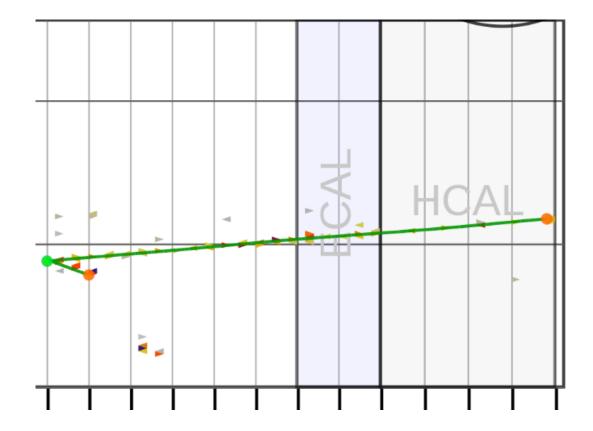


#### **MINERvA Detector Simulation**

GENIE Generated Neutrino Interactions



Full Detector Simulation (GEANT4) to simulate the response of the detector for particles that interact with the detector.



- Data Overlay is done on the detector simulated events.
- Simulate pileup effect from other interactions (upstream and in the target) in the simulated interactions.

#### **CCQE Cross-section**

 CCQE Cross-section is generally expressed in Llewellyn Smith formalism:

$$\frac{d\sigma}{dQ^2} \begin{pmatrix} \nu n \to l^- p \\ \bar{\nu}p \to l^+ n \end{pmatrix} = \frac{M^2 G^2 \cos^2 \theta_c}{8\pi E_{\nu}^2} \left[ A(Q^2) \mp B(Q^2) \frac{(s-u)}{M^2} + \frac{C(Q^2)(s-u)^2}{M^4} \right]$$
(1)

$$\begin{split} M &= \frac{m_p + m_n}{2} \\ s - u &= 4ME - Q^2 - m_l^2 \\ G_F^2 &= 1.18 \times 10^{-5} GeV^{-2} \ (Fermi \ Coupling \ constant) \\ \theta_C &= 0.974 \end{split}$$

#### **CCQE Cross-section**

24

$$A(Q^{2}) = 4 \frac{Q^{2}}{4M^{2}} [F_{A}]^{2} (1 + \frac{Q^{2}}{4M^{2}}) - |F_{V}^{1}|^{2} (1 - \frac{Q^{2}}{4M^{2}}) + (1) \quad \text{Valid for } \nu_{\mu}, \nu_{e}$$

$$|\xi F_{V}^{2}|^{2} \frac{Q^{2}}{4M^{2}} (1 - \frac{Q^{2}}{4M^{2}}) + 4F_{V}^{1}\xi F_{V}^{2} \frac{Q^{2}}{4M^{2}}] \quad (1) \quad \text{Valid for } \nu_{\mu}, \nu_{e}$$

$$B(Q^{2}) = 4 \frac{Q^{2}}{4M^{2}} [F_{A}(F_{V}^{1} + \xi F_{V}^{2})] \quad (2)$$

$$C(Q^{2}) = \frac{1}{4} [F_{A}]] + |F_{V}^{1}|^{2} + \frac{Q^{2}}{4M^{2}} |\xi F_{V}^{2}|^{2}] \quad (3)$$

$$\cdot M_{A} \sim 0.99 \text{ GeV} (\text{Axial Mass})$$

$$\cdot \text{Based on bubble chamber (hydrogen targets) measurements}$$

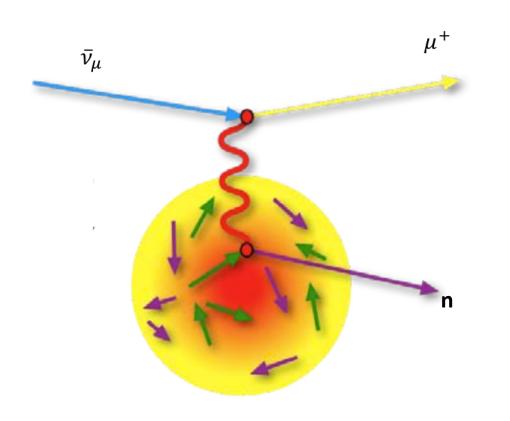
$$\cdot \text{Measurements in heavier target report slightly higher axial mass}$$

$$\cdot \text{Nuclear Effects}$$

$$\cdot \text{Dipole Form Factor approximation breaks at high } Q^{2}$$

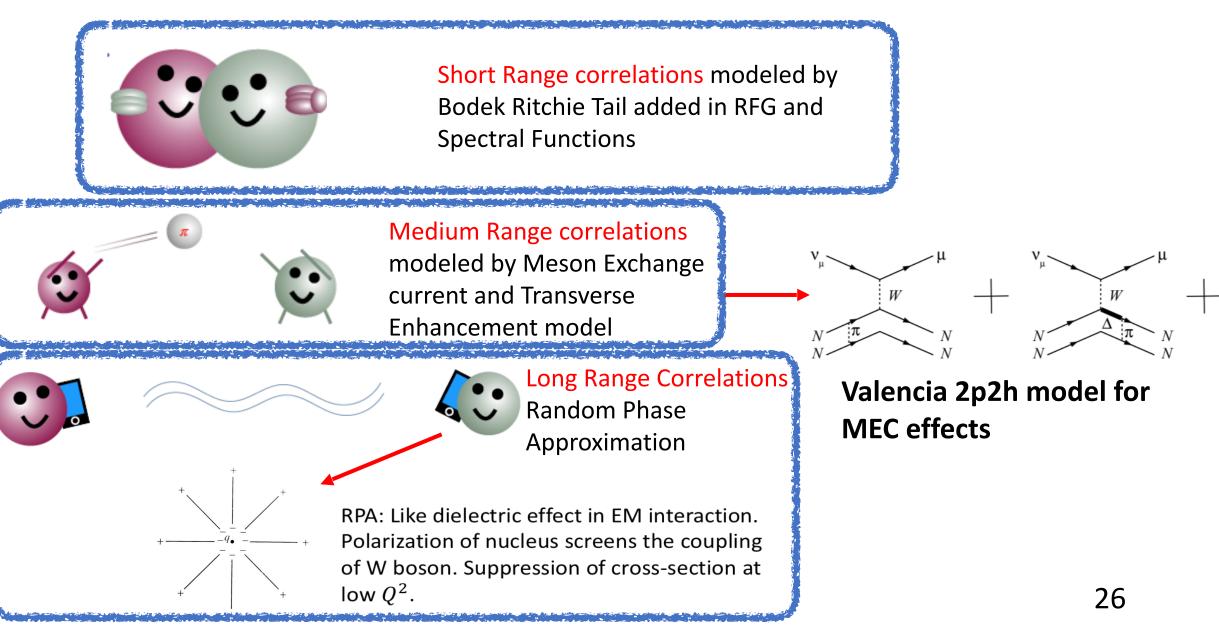
## **Neutrino Nucleon Cross-section Modeling**

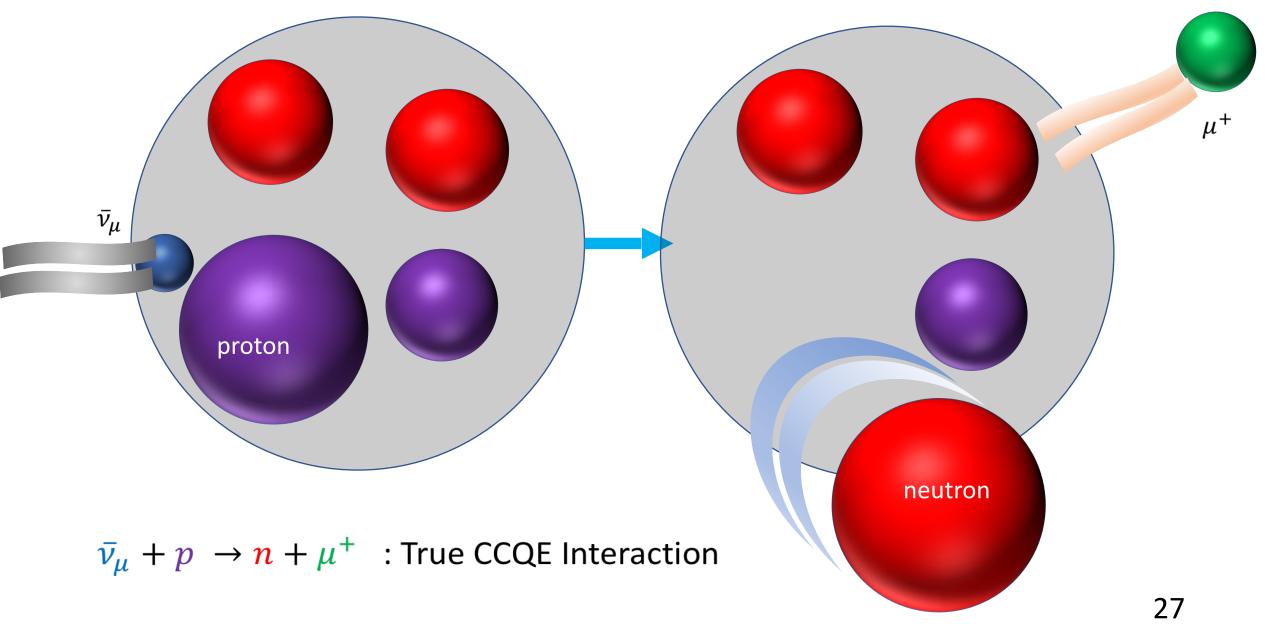
CCQE cross-section model based on the Llewellyn Smith formalism.

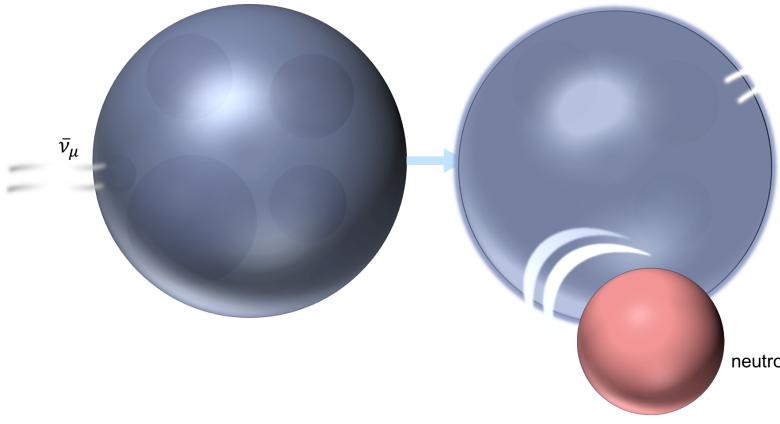


- Cross-section models are based on neutrino interaction with the free nucleons.
- In heavy targets like carbon (this analysis), argon, etc., the final state particle (that exits the nucleus) can be changed due to nuclear effects that are not modeled by the Fermi Gas Model.

## **Nuclear Effects: nucleon-nucleon correlation**





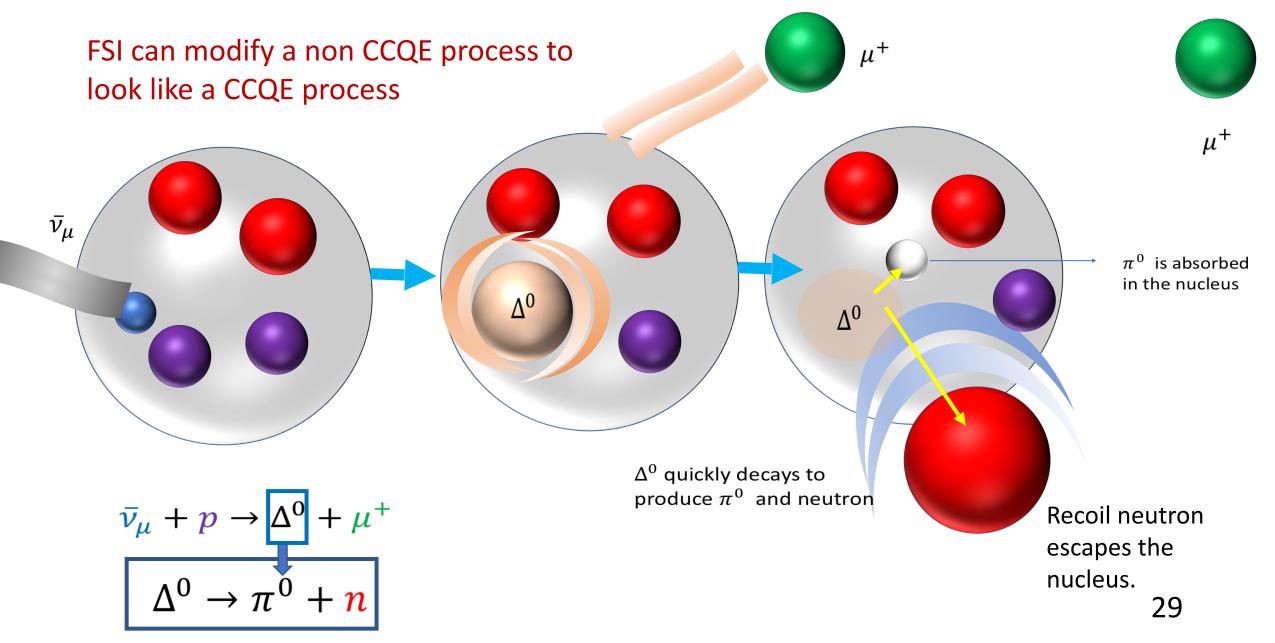


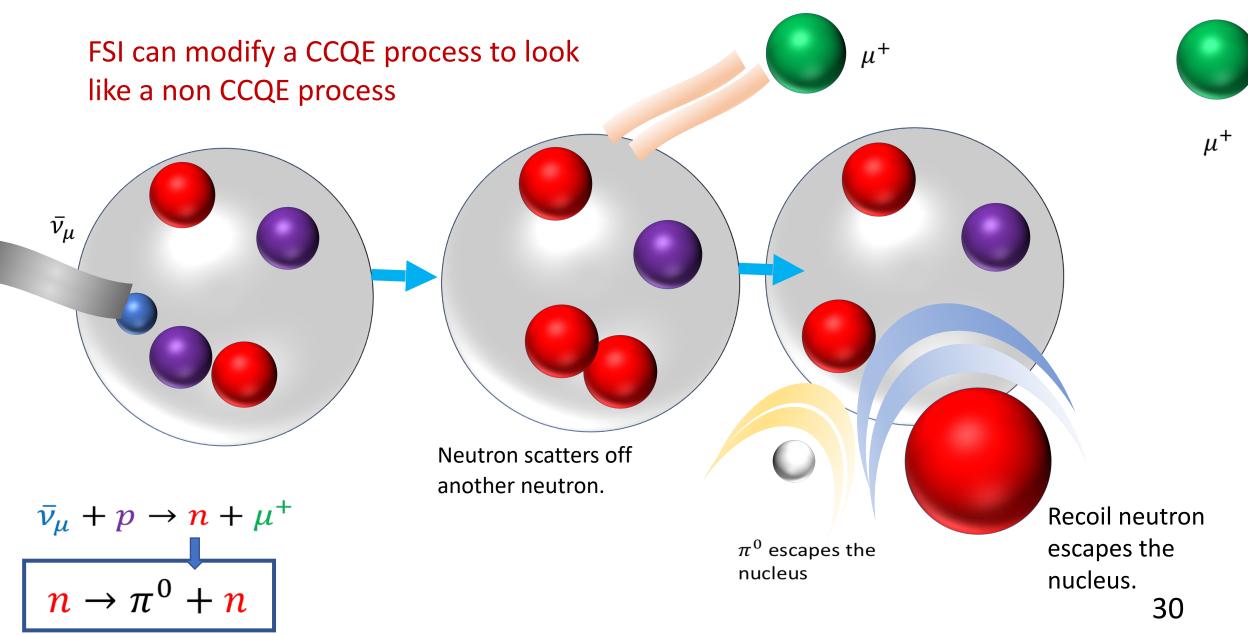
But we don't know what happens inside the nucleus.

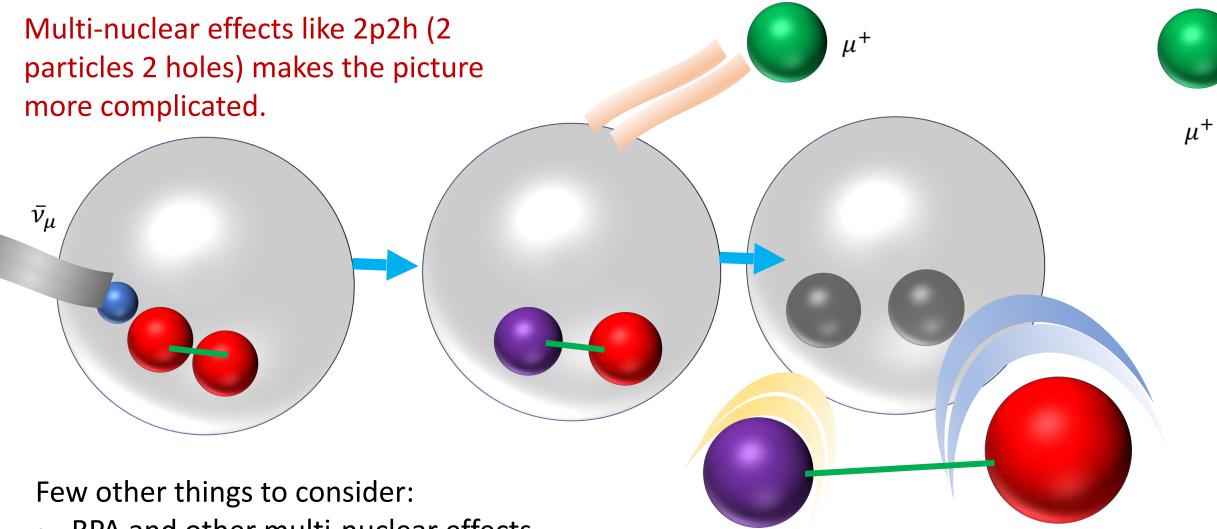
We can only see the final state particles in the detector.

neutron

 $\bar{\nu}_{\mu} + p \rightarrow n + \mu^+$ : True CCQE Interaction





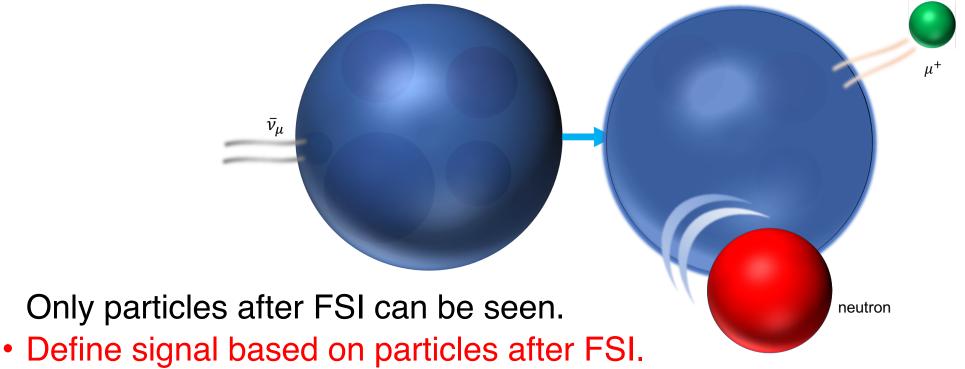


- RPA and other multi-nuclear effects
- Nucleons are not at rest

Multi Nuclear effect ejects the nucleon with its correlated partner

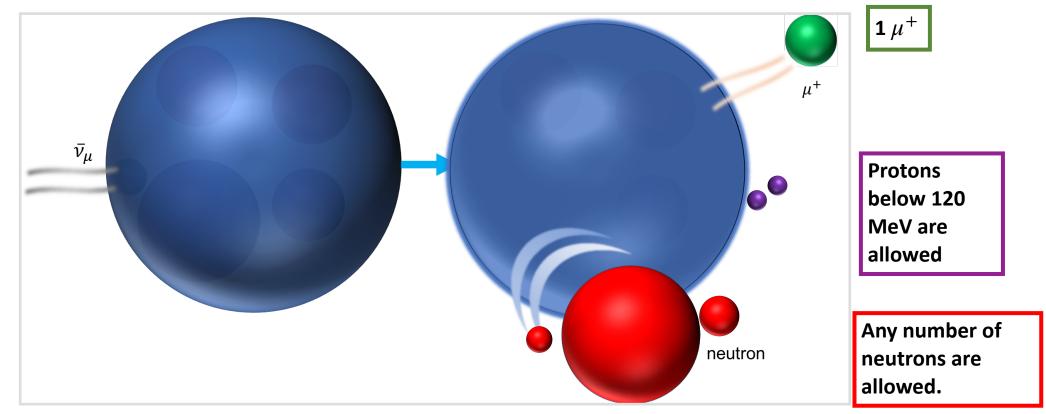
#### **CCQE-Like Event**

- FSI can fake a CCQE process to look like a non CCQE process.
- FSI can fake a non CCQE process to look like a CCQE process.



CCQELike processes

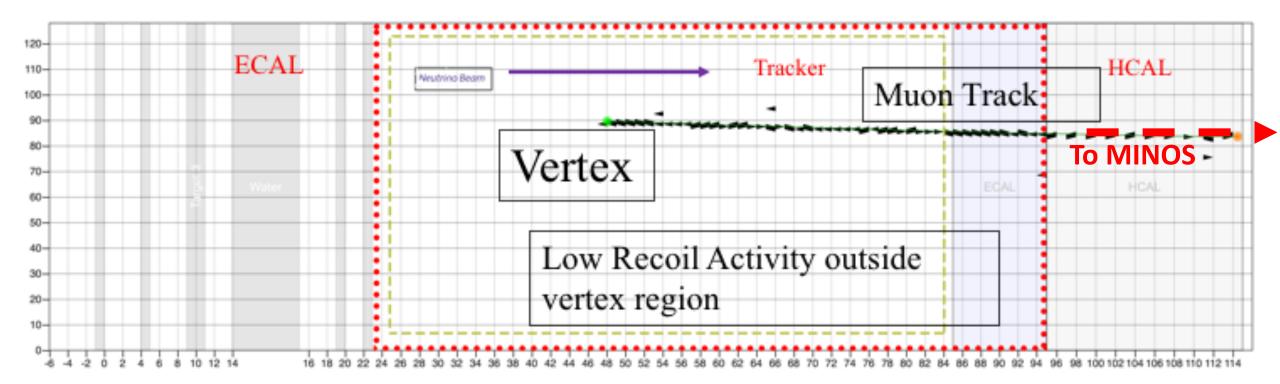
## **Signal Definition: CCQELike Process**



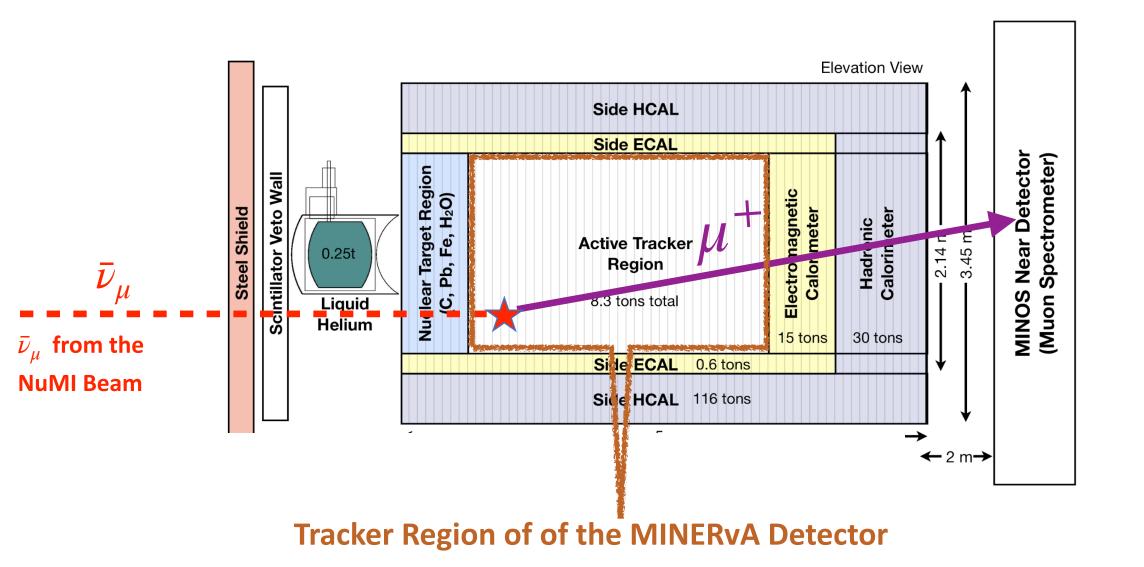
Signal Definition based on Final State Particles

- 1 positive muon ( $\mu^+$ )
- Any number of neutrons
- Any number of Protons below 120 MeV Kinetic Energy
- No mesons (particles like  $\pi^{\pm}$  which are produced in Resonance processes)

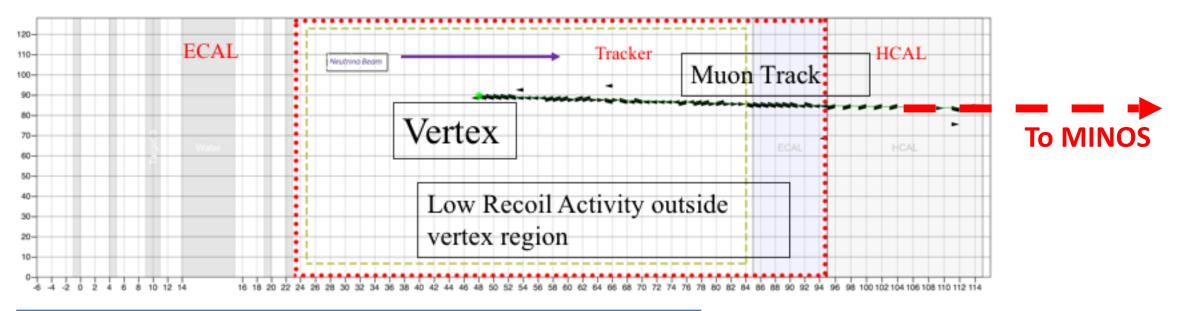
#### **Event Reconstruction:CCQELike Process**



Event should be in the **Tracker (CH)** region of the detector. We are interested in the cross-section in CH target.



#### **Event Reconstruction:CCQELike Process**



#### **Signal Definition**

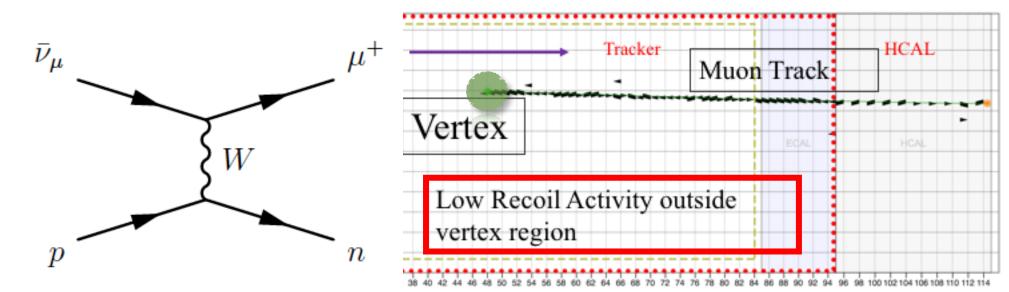
# Event should have 1 positive muon in the final state.

• 1 MINOS matched muon track (to

identify the charge of the muon)

- Apply a  $20^o$  angle cut on muon track (with respect to the  $\bar{\nu}_\mu$  beam)
- No Additional tracks (in next few slides)

#### **Event Reconstruction:CCQELike Process**



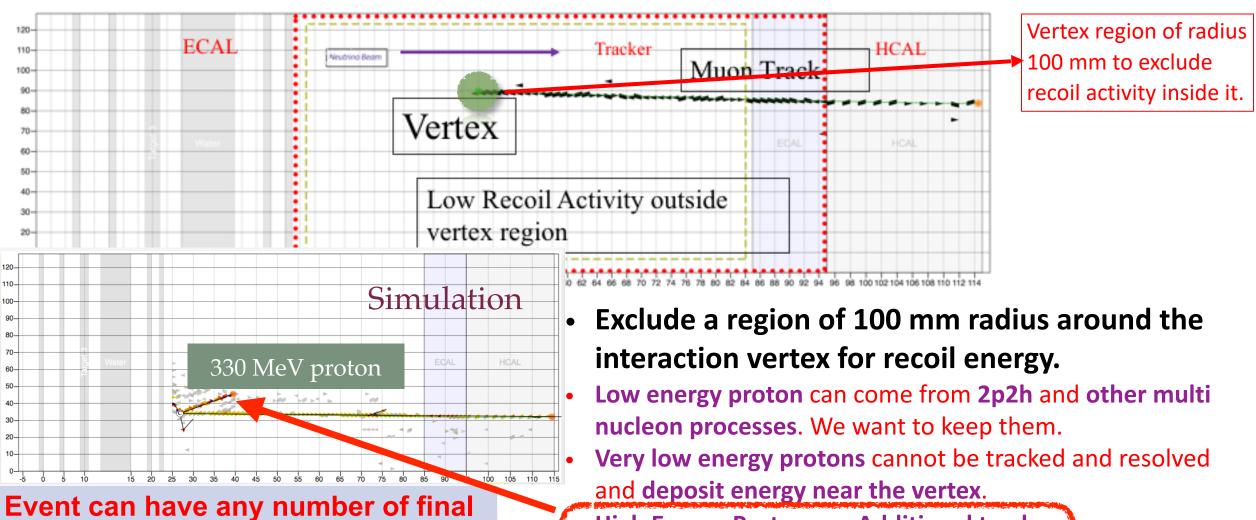
#### **Signal Definition**

Event should have 1 positive muon in the final state.

Event can have any number of neutrons in the final state.

- Low Recoil Activity outside the vertex region
- Most of the energy from the interaction carried away by the muon
- Recoil Activity → Isolated clusters outside the vertex region
- High recoil activity events are dominated by resonance and Deep Inelastic events.

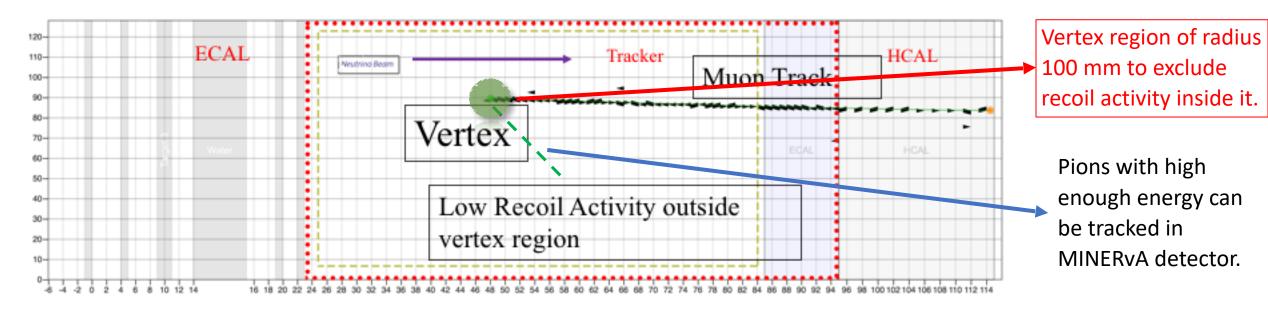
#### **Event Reconstruction:CCQELike Process**



state protons less than 120 MeV.

High Energy Protons  $\rightarrow$  Additional tracks.

#### **Event Reconstruction:CCQELike Process**



#### **Signal Definition**

Event should have 1 positive muon in the final state.

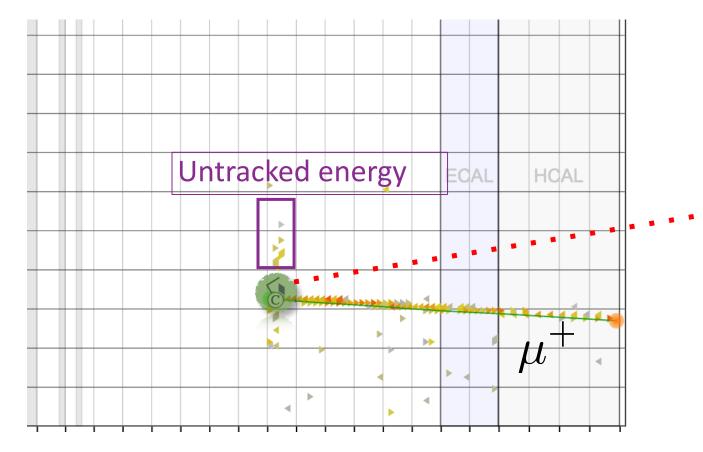
Event can have any number of neutrons in the final state.

Event can have any number of final state protons less than 120 MeV.

# Event cannot have any mesons in the final state.

- Only 1 track events (track being muon track) are selected.
- Additional tracks  $\rightarrow$  charged pions are rejected.
  - Remember protons with high Kinetic Energy also make tracks [and (>120 MeV) rejected]

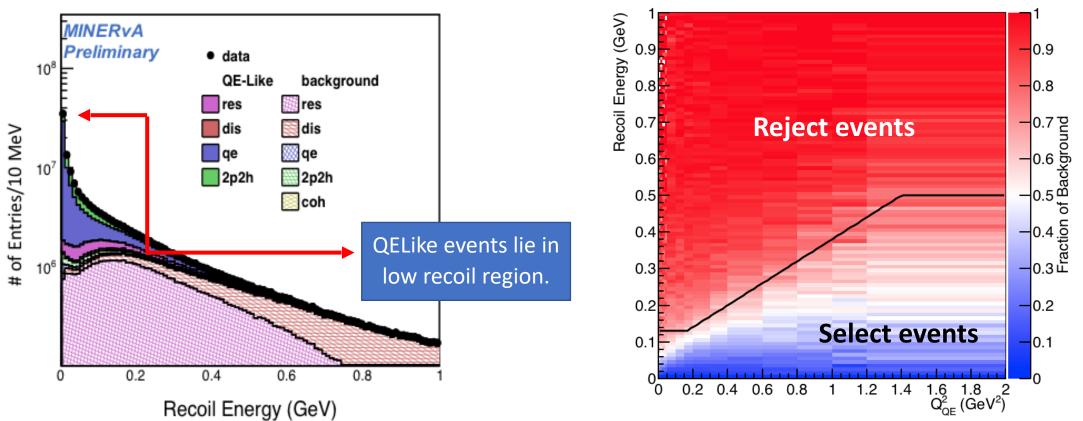
## **Recoil Energy Definition**



All the untracked energy outside the spherical blob of radius 100 mm outside the interaction vertex.

Untracked energy is the energy of the clusters that are not part of the  $\mu^+$  track.

#### **Recoil Energy**



- Distribution of Recoil Energy with Data and Various MC components.
- Recoil Energy cut based on the previous iteration of this analysis [Phys.Rev.D 97 (2018) 5, 052002]
- **Optimized** for signal selection efficiency+purity of selected sample.
- Loose cut at high  $Q_{QE}^2$  region
  - Keep 2p2h events in this region that is not well-understood.

#### **Extraction of Cross-section**

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

- This analysis measures cross section ( $\sigma$ ) as a function of 2 variables x and y.
- Event reconstruction in  $(\alpha, \beta)$  bins.
- Want to measure in true bins (i, j) bins.
- More on this later.

•  $(\frac{d^2\sigma}{dp_z dp_t}) \rightarrow$  Cross section as a function of muon kinematics  $E_{\nu}^{QE} = \frac{m_n^2 - (m_p - E_b)^2 - m_{\mu}^2 + 2(m_p - E_b)E_{\mu}}{2(m_p - E_b - E_{\mu} + p_{\mu}\cos\theta_{\mu})}$ •  $(\frac{d^2\sigma}{dE_{\nu}dQ^2})_{QE} \rightarrow$  Cross section as a function of  $E_{\nu}$  and four  $Q_{QE}^2 = 2E_{\nu}^{QE}(E_{\mu} - p_{\mu}\cos\theta_{\mu}) - m_{\mu}^2$ momentum transferred ( $Q^2$ ) based on QE hypothesis

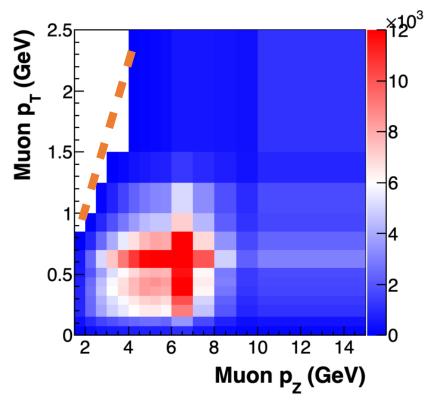
#### **Event Reconstruction**

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} \left(N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg}\right)}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

#### **Raw Event Reconstruction**

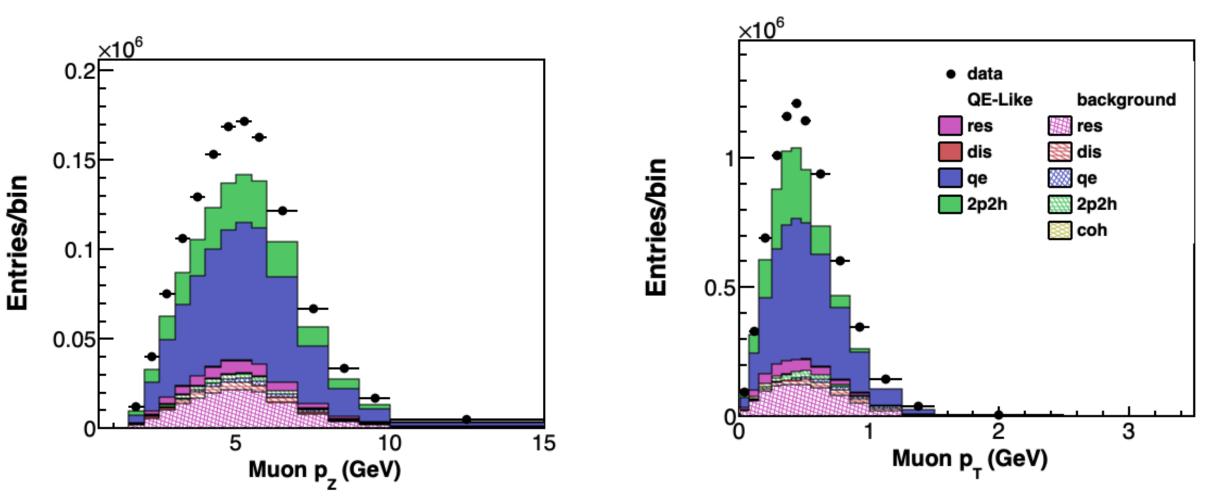
$$(\frac{d^2\sigma}{dxdy})_{ij} =$$

$$\frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$



- Apply the selection conditions on the data and simulated sample to select the CCQE Like candidate events.
- Requiring muon track angle to be less than 20 degrees rejects events at high p<sub>T</sub> and low p<sub>Z</sub> phase space.

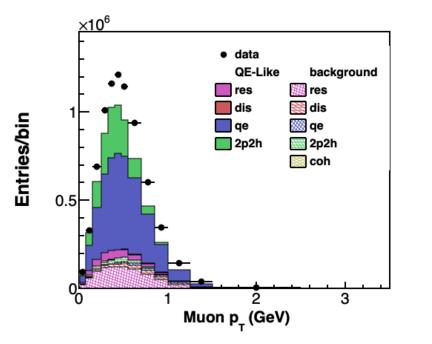
#### **Raw Event Selection**



We look at our simulated sample to estimate the types of events we select in our data. \*1 D are the projections from 2 D distributions.

Muon  $p_z$  || Neutrino Beam

## **Selected Events (Signal)**



Signal Components	Fraction of Total MC Events
QE	0.54
RES	0.05
DIS	0.003
2p2h	0.193
Total	0.786

#### Signal (QE-Like)

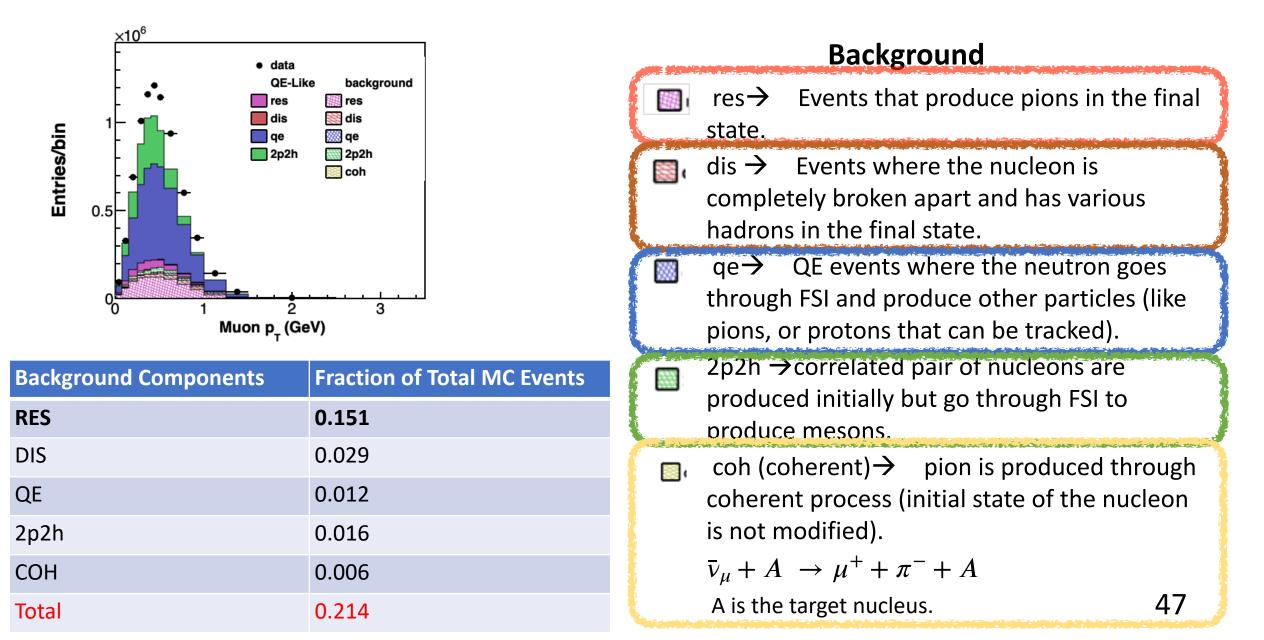
■ res→ Events that produced pions initially but go through FSI to produce neutron in the final state.

dis→ Events where the neutrino completely breaks the nucleon and creates high recoil activity but only neutron is the Final State Particle (very rare)

 $qe \rightarrow$  True QE events. Recoil neutrons do not go through FSI and escape the nucleus.

■: 2p2h → Events in which the neutrino interacts with a correlated pair of nucleons and both nucleons exit the nucleus.

## Selected Events (Background)



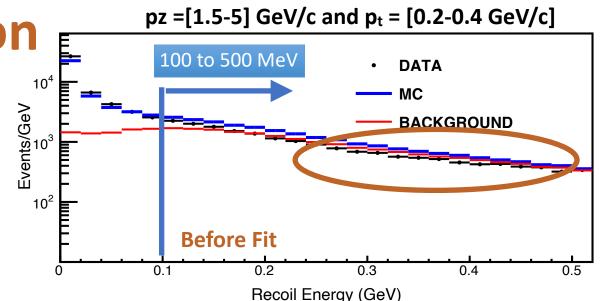
#### Background

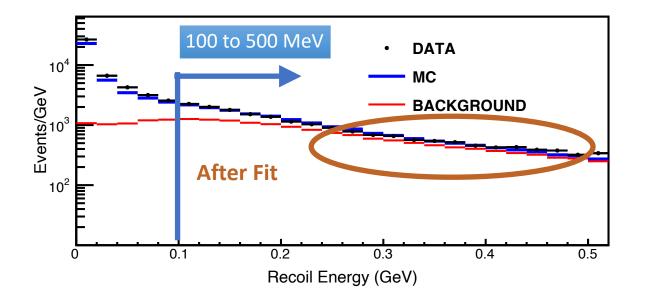
$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

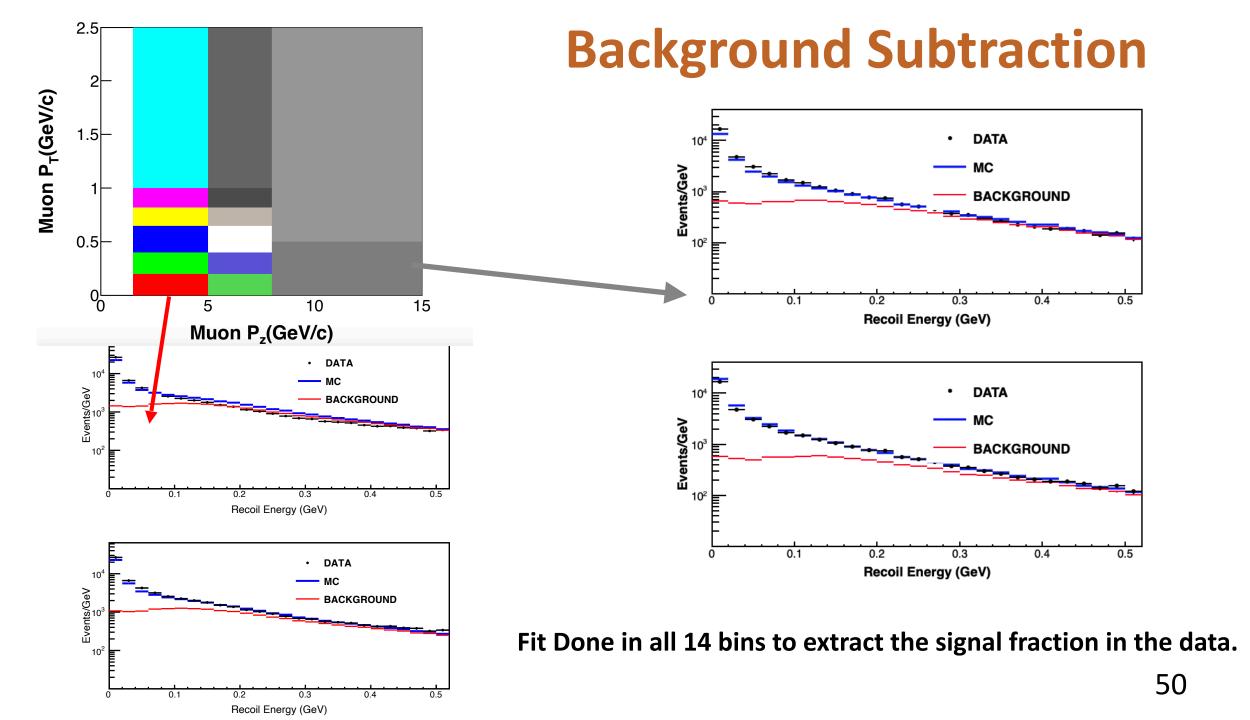
- Our selected data has both signal and background contribution.
- We need to subtract the background from our selected sample.
- Cannot rely on MC completely to estimate the background in the data.
  - Pion production may be over predicted by our simulation.
  - Nuclear effects have significant uncertainties.
  - We want to improve simulation models from our data after all.
- Use Data driven method to subtract the background from the data.

## **Background Subtraction**

- We look at our recoil energy distribution in 14 different bins of muon p<sub>T</sub> and p<sub>z</sub>.
- Fraction fit of data with simulated signal and background recoil energy distribution shapes.
  - ROOT::TFractionFitter [root.cern.ch]
  - Fit done between 100 to 500 MeV recoil energy region.
    - Background rich region
- Fit gives the best estimation of signal and background fraction in our data.
  - Efficiency correction to get the signal fraction in each fit region.







#### $635,592\pm1,251$ (stat.) $\pm 13,850$ (syst.) events after background subtraction.

$p_T ~ \text{GeV/c}$	$p_z \text{ Gev/c}$	Signal Before	Signal After	Signal Efficiency	Efficiency Corrected Signal	$\chi^2/NDF$
0.0 - 0.2	1.5 - 5.0	0.56	$0.610\pm0.027$	0.901	0.78	18.12/19
0.2 - 0.4	1.5 - 5.0	0.62	$0.700\pm0.011$	0.905	0.86	26.82/19
0.4 - 0.65	1.5 - 5.0	0.61	$0.673 \pm 0.008$	0.901	0.84	16.08/19
0.65 - 0.82	1.5 - 5.0	0.59	$0.622\pm0.012$	0.922	0.75	5.96/19
0.82 - 1.0	1.5 - 5.0	0.59	$0.582 \pm 0.0188$	0.951	0.66	14.74/19
1.0 - 2.5	1.5 - 5.0	0.0.59	$0.597 \pm 0.0373$	0.988	0.61	16.44/19
0.0 - 0.2	5.0 - 8.0	0.62	$0.788 \pm 0.0367$	0.923	0.89	11.43/19
0.2 - 0.4	5.0 - 8.0	0.67	$0.772\pm0.015$	0.922	0.89	16.64/19
0.4 - 0.65	5.0 - 8.0	0.67	$0.719 \pm 0.010$	0.918	0.85	26.45/19
0.65 - 0.82	5.0 - 8.0	0.65	$0.700 \pm 0.0119$	0.930	0.80	21.69/19
0.82 - 1.0	5.0 - 8.0	0.64	$0.638 \pm 0.0166$	0.952	0.71	11.33/19
1.0 - 2.5	5.0 - 8.0	0.62	$0.615\pm0.026$	0.983	0.63	19.48/19
0.0 - 0.5	8.0 - 15.0	0.69	$0.778 \pm 0.026$	0.926	0.89	18.86/19
0.0 - 0.5	8.0 - 15.0	0.66	$0.69\pm0.0156$	0.950	0.77	11.87/19
	•			•		
Signal fraction in the simulated Signal fraction predicted by the fit (100 to Fraction of signal in the data sample						ata sample t

sample (100 to 500 MeV recoil energy)

Signal fraction predicted by the fit (100 toFraction of signal in the data sample that500 MeV recoil energy)51

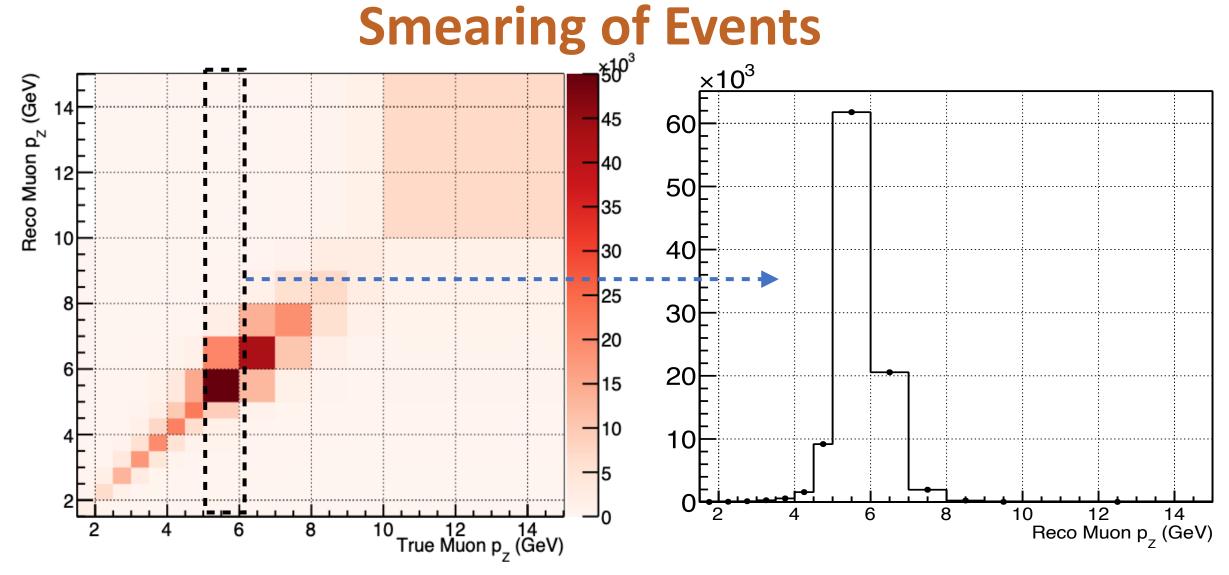
#### **Unsmearing the reconstructed events**

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

- We look in our data as a function of reconstructed quantities  $[(p_z, p_t)_{\mu}, (E_{\nu}, Q^2)_{QE}]$ .
- Cannot measure these quantities perfectly
  - Limitation of our detector resolution
  - Reconstruction algorithms

• Example

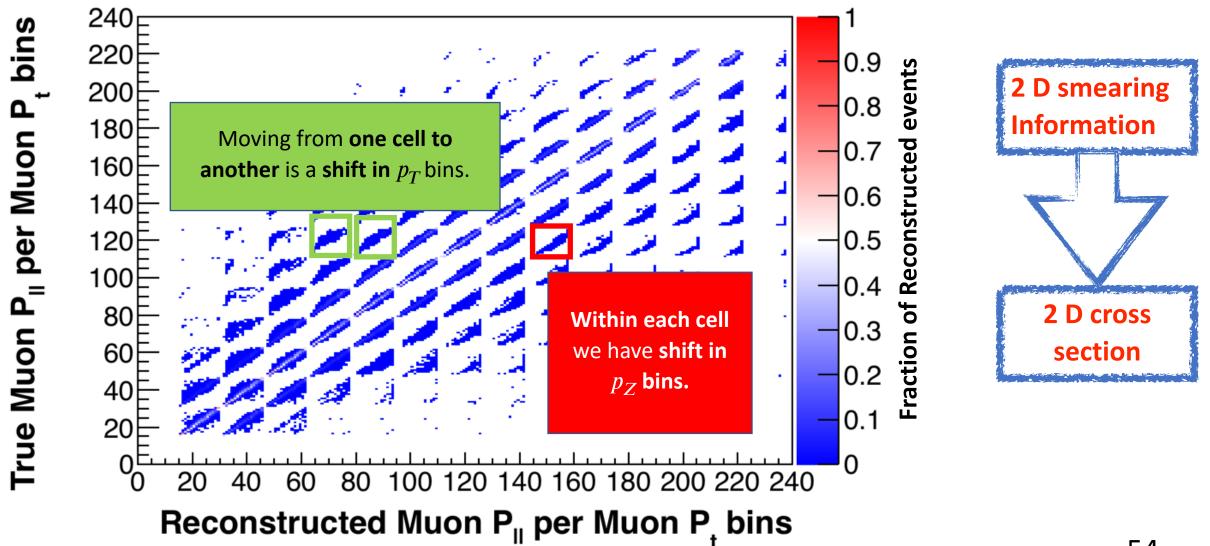
- A  $\mu^+$  whose actual (true) momentum is 5 GeV could be reconstructed as 4 GeV sometimes or 3 GeV sometimes.
- Called the **smearing of the events** from true bin to the reconstructed bins.
- We want to **test models against our measurements** 
  - Models are based on True Variables
- Need to correct our reconstructed events to their true phase space.
- $U_{\alpha\beta ij} \rightarrow$  Matrix that contain the smearing information of events in true bins *ij* to reconstructed bins  $\alpha\beta$
- We use Iterative Bayesian Unfolding method [cite]to unfold our reconstructed data into the true bins.



Smearing of events from their true  $p_z$  bins to reconstructed  $p_z$  bins.

Smearing of 5 to 6 GeV (true) events in different reconstructed bins

#### **Unfolding Matrix**



## **Unfolding of the Event**

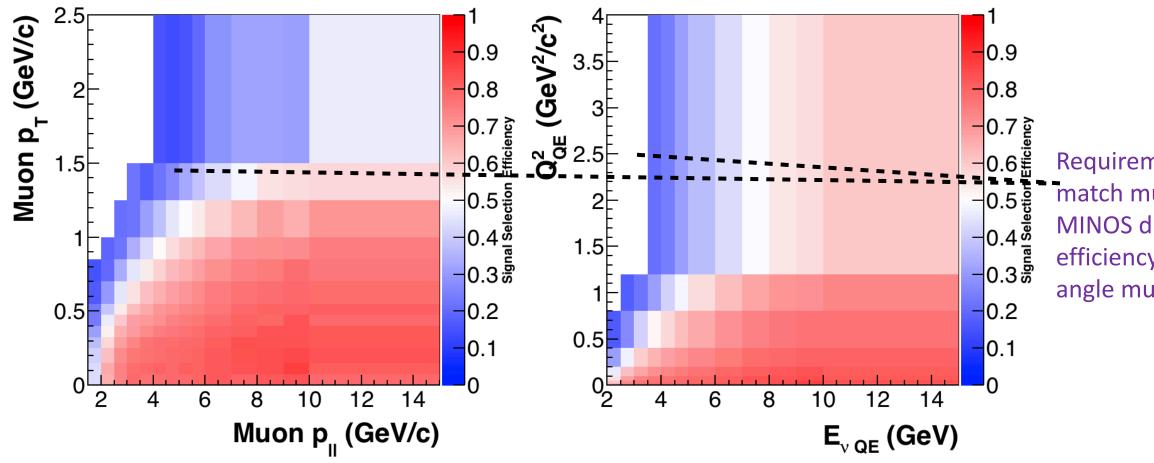
- Background subtracted sample is unfolded to the the true kinematic variables.
  - Iterative Bayesian Unfolding [G. D' Agostini, Nuclear Instrument Method]
  - Based on RooUnfold Algorithm.
- Unfolding studies done with various model predictions.
  - Find Optimum number of iteration needed to unfold
  - Test the stability of the unfolding matrix
- $(p_z, p_t)_{\mu}$  = 4 iterations
- $(E_{\nu}, Q^2)_{QE}$  = 8 iterations

### **Signal Selection Efficiency Correction**

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

$$\epsilon = \frac{N_{QELIKE}^{Pass Cut}}{N_{QELIKE}^{Total Events}}$$
  
→ Outer detector
  
Beam
  
Inner
  
detector

- We fail to reconstruct some signal events due to:
  - Detector acceptance
    - Remember MINOS acceptance requirement?
  - Reconstruction Efficiency
    - Our algorithms are not 100% perfect.
- Correct for the fraction of events that we failed to reconstruct.



#### Signal Selection Efficiency $\approx$ 70% to 80%

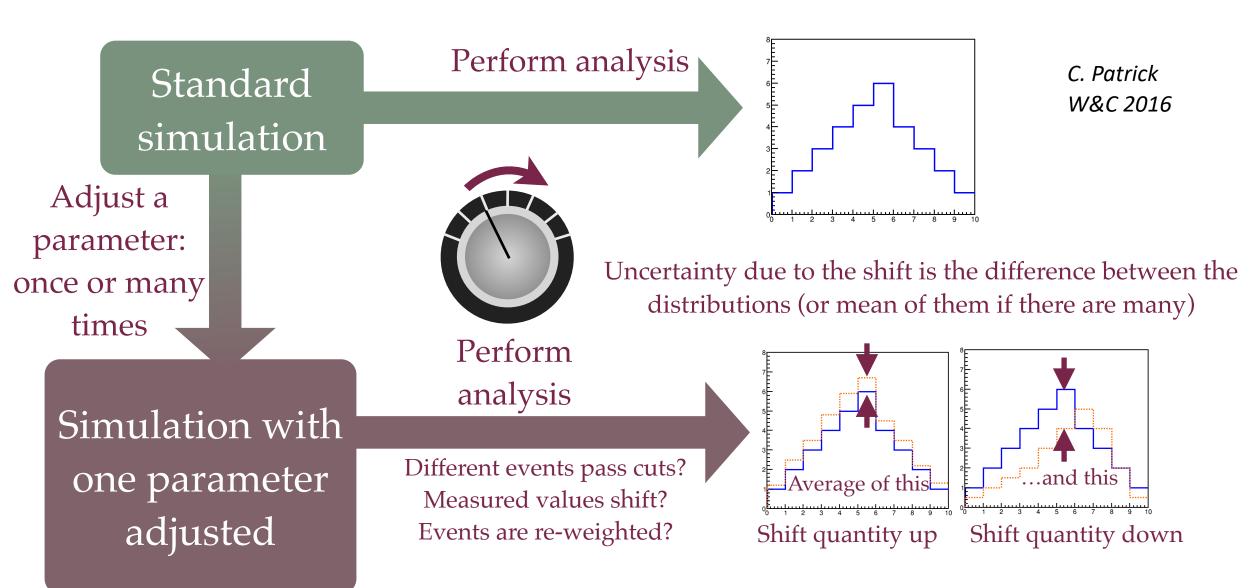
Requirement to match muons at MINOS drops our efficiency for high angle muons.

#### **Flux And Target Normalization**

$$\left(\frac{d^2\sigma}{dxdy}\right)_{ij} = \frac{\sum_{\alpha\beta ij} U_{\alpha\beta ij} (N_{data,\alpha\beta} - N_{data,\alpha\beta}^{bkg})}{\epsilon_{ij} \Phi T \Delta x_i \Delta y_j}$$

- Muon  $P_z vs$ .  $P_T$ : We use the integrated neutrino flux for normalization.
- $E_v vs$ .  $Q^2$ : distribution is normalized by neutrino flux of corresponding  $E_v$  to get the neutrino cross-section independent of the shape of the flux.
- $T \rightarrow$  Total number of nucleons in the tracker region.
- $3.23 \times 10^{30}$  nucleons (protons+neutrons) in the tracker region.
- $\Phi$  (Flux) integrated from 0 to 120 GeV to obtain the differential cross section

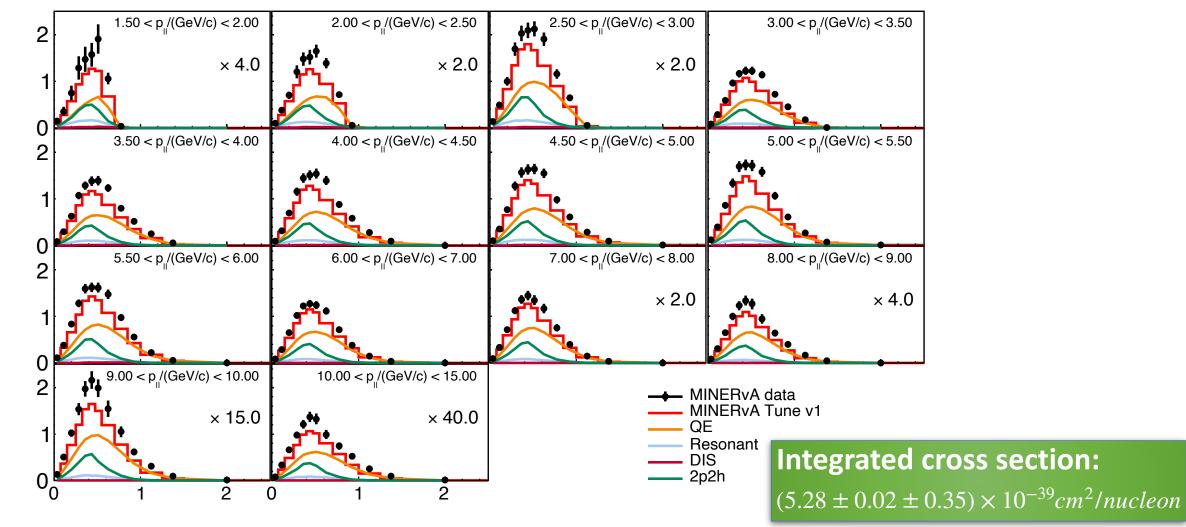
#### **Systematic Uncertainties**



Examples: increase resonant cross section by 10% OR 500 "universes" of flux changes

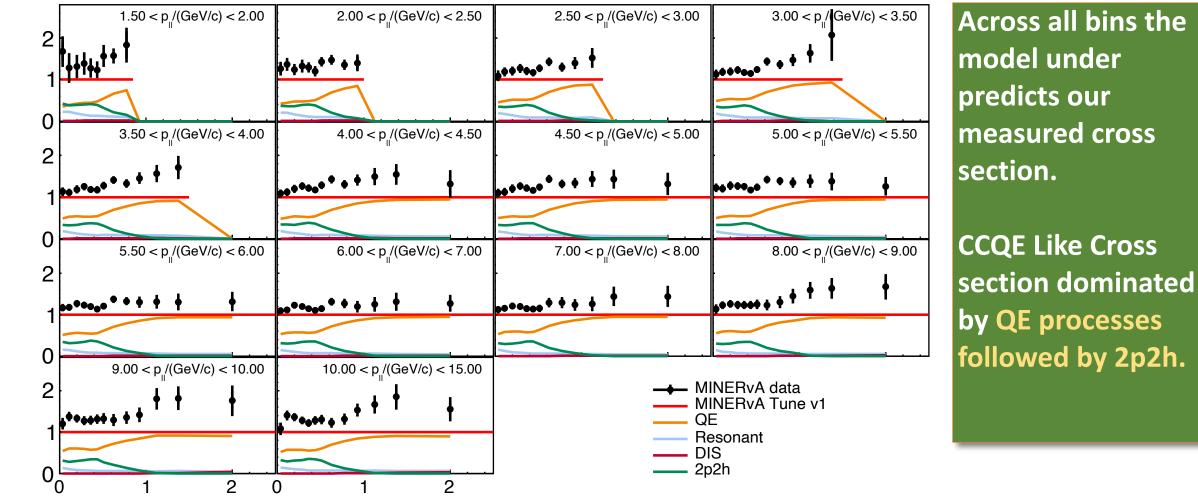
# **Cross section Results**

# $p_t \operatorname{cross} \operatorname{section}$ in the bins of $p_z$



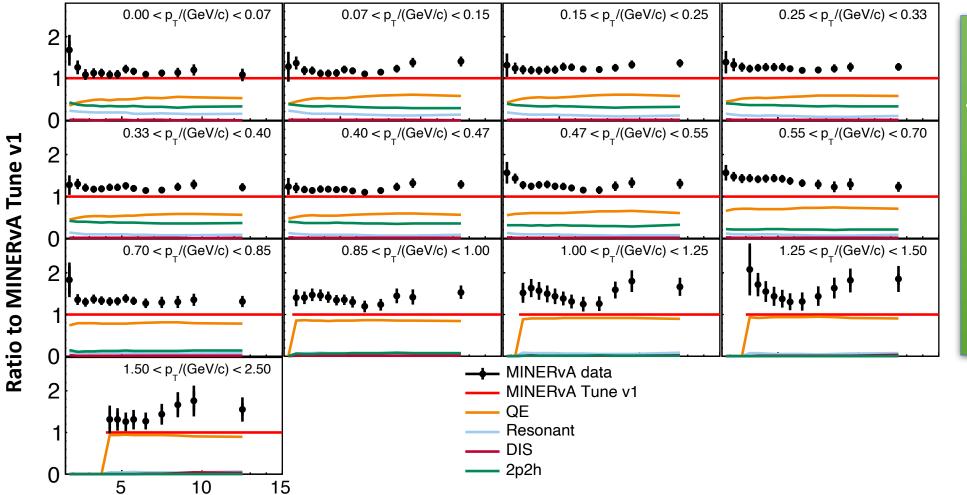
Muon Transverse Momentum (GeV/c)

# Cross section Ratio ( $p_t$ in the bins of $p_z$ )



Muon Transverse Momentum (GeV/c)

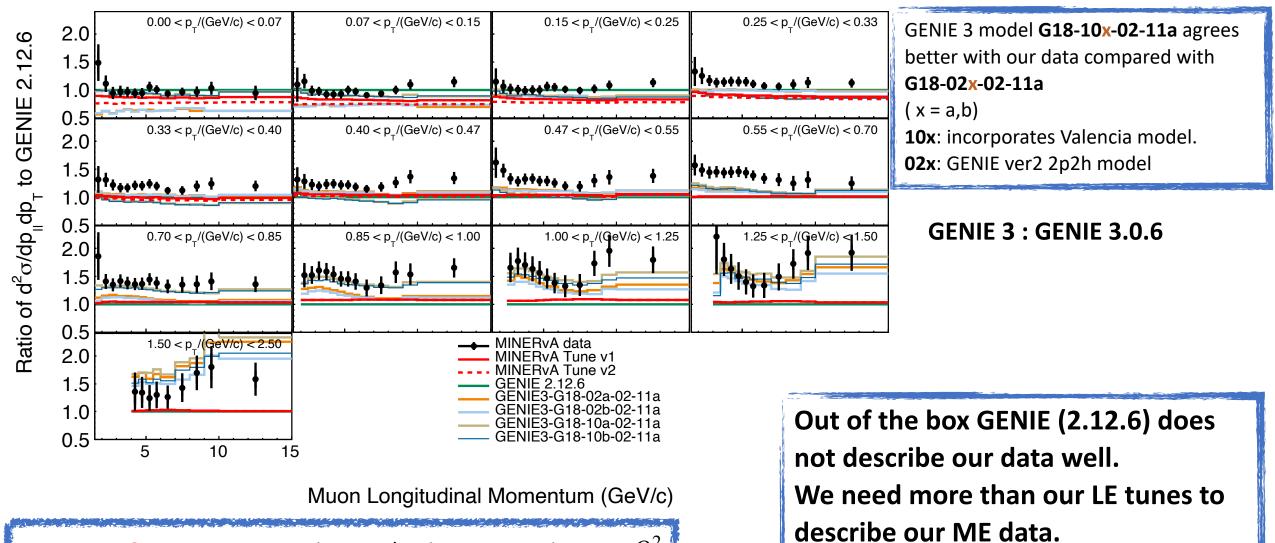
# Cross section ratio ( $p_z$ in the bins of $p_t$ )



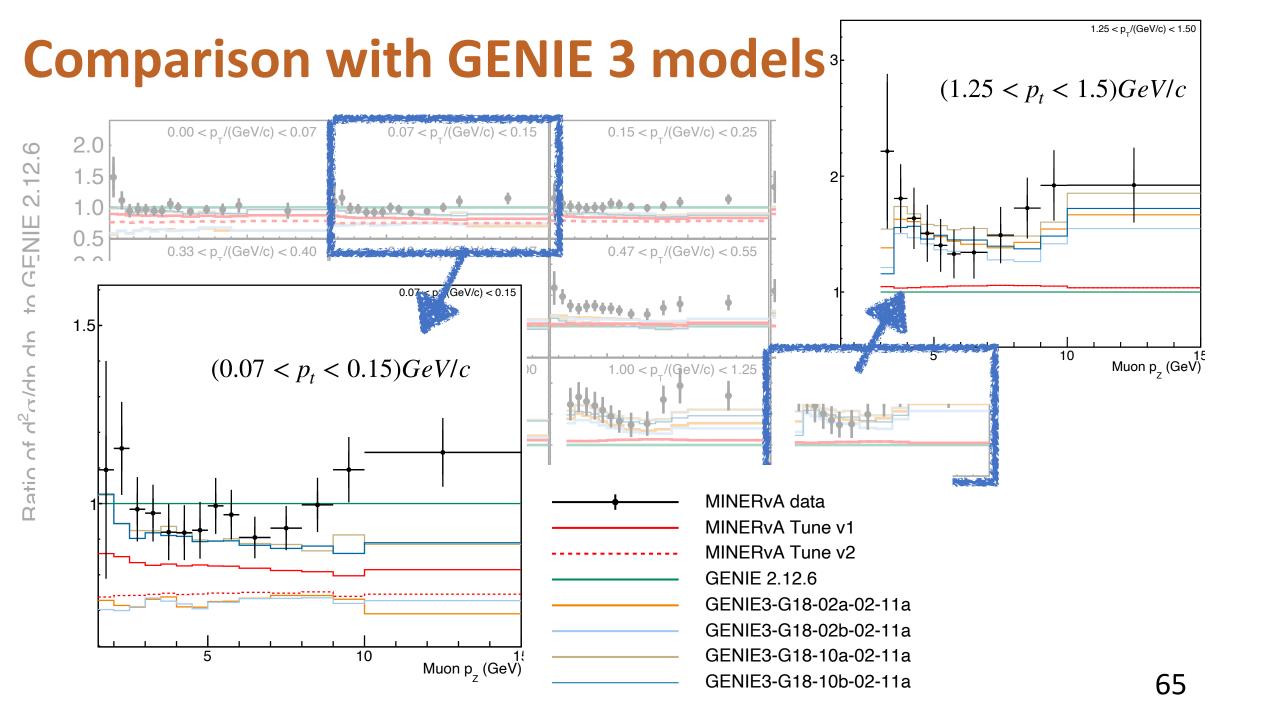
QE (1p1h) and 2p2h processes dominate the low transverse region whereas the higher transverse region is dominated by QE processes entirely.

Muon Longitudinal Momentum (GeV/c)

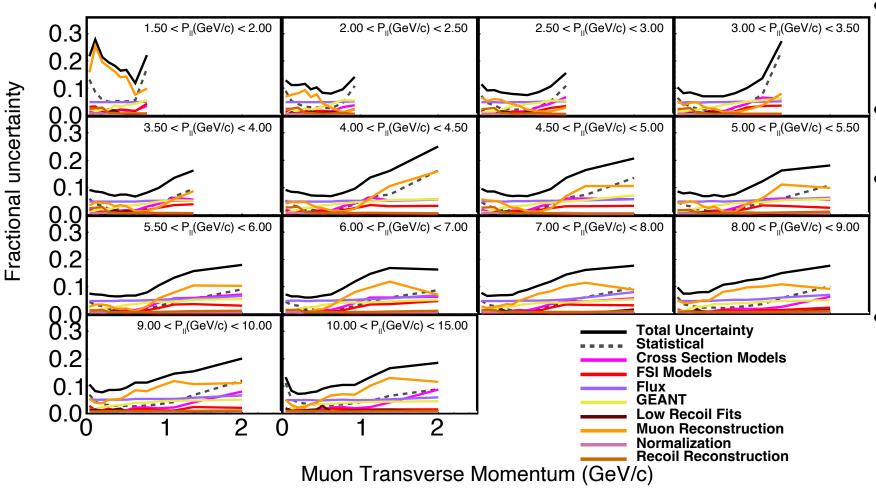
#### **Comparison with GENIE 3 models**



**MnvTune v2** = **MnvTune v1**+ Pion Production suppression Low  $Q^2$ [*P. Stowell*, arXiv:1903.01558.]



## **Error Summary on the data**



- Uncertainty is dominated by the muon reconstruction.
- Flux contribution is around 5%.
- Neutron related error dominates the GEANT category and contribute unto 5% in some regions.
   GENIE cross section models and FSI contribute contribute less than 5%

overall.



Model	$\chi^2$ - linear $\chi^2$ - log	g
GENIE 2.12.6 Tunes		
MINERvA Tune v1	362.6 580.4	4
MINERVA Tune v2	364.4 601.4	4
GENIE w/o 2p2h	226.5 $473.2$	2
GENIE (Default)	346.4 550.6	6
GENIE∓πtune	354.3 568.	5
GENIE+RPA	230.0 406.7	7
GENIE+RPA+ $\pi$ tune	231.7 414.0	D
GENIE+Low Recoil Tune	755.4 1059.4	4
GENIE+Low Recoil Tune+RPA	361.2 570.0	0
GENIE+Low Recoil Tune+ $\pi$ tune	e 760.6 1081.8	8
GENIE 3.0.6 Tunes		
GENIE 3.0.6 G18_02a_02_11a	602.9 $865.0$	0
GENIE 3.0.6 G18_02b_02_11a	586.9 $878.3$	3
GENIE 3.0.6 G18_10a_02_11a	353.1  447.3	5

TABLE II.  $p_{\parallel} - p_{\perp} \chi^2$  between data and model variants derived from GENIE. The number of degrees of freedom is 171. Both the  $\chi^2$  between the values and between the logs of the values are listed.

312.8

421.7

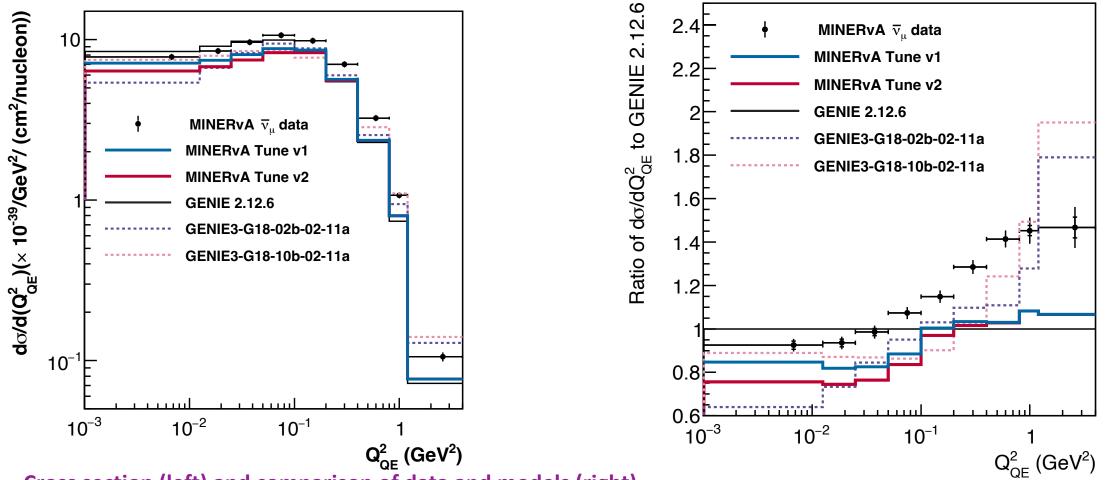
GENIE 3.0.6 G18\_10b\_02\_11a

- Full systematic and bin to bin correlations are treated to calculate the  $\chi^2$  between data and models.
- χ<sup>2</sup> tells us that GENIE 2 variations that were based on low energy data are not sufficient to describe our Medium Energy data.
- GENIE 3 models with Nieves 2p2h implementation performs better than GENIE 2 like 2p2h models.

# 1 D projections from $(E_{\nu}, Q^2)_{QE}$ measurements

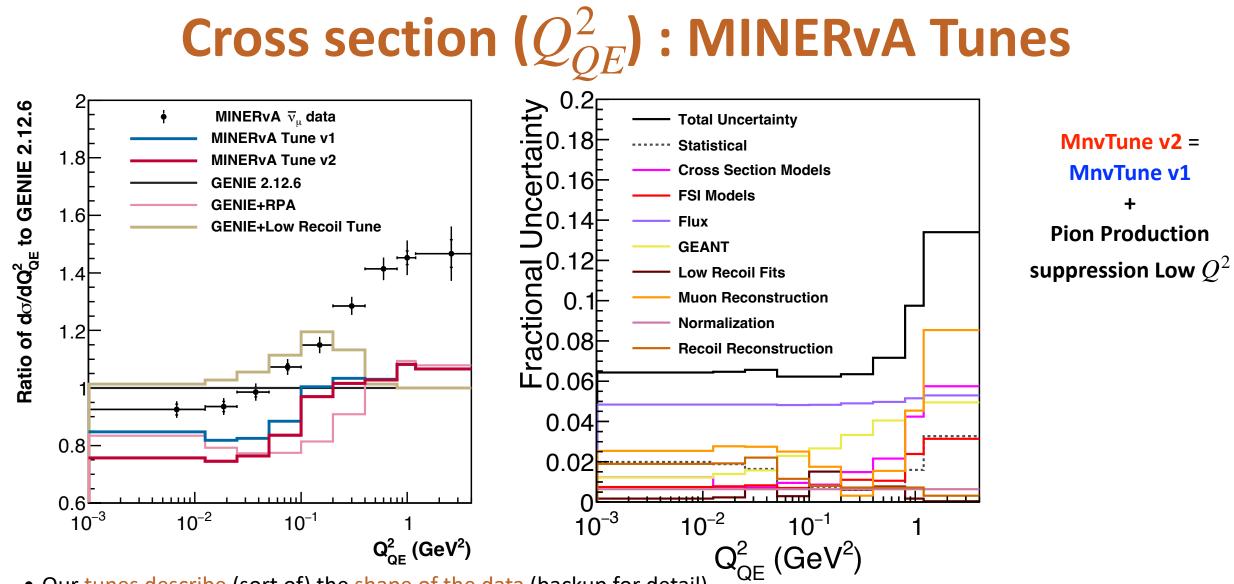
# **Cross section (** $Q_{QE}^2$ **) : MINERvA Tune and GENIE 3**

**MINERvA Tune v2** = MnvTune v1+Pion Production suppression Low  $Q^2$  based on P. Stowell, arXiv:1903.01558.



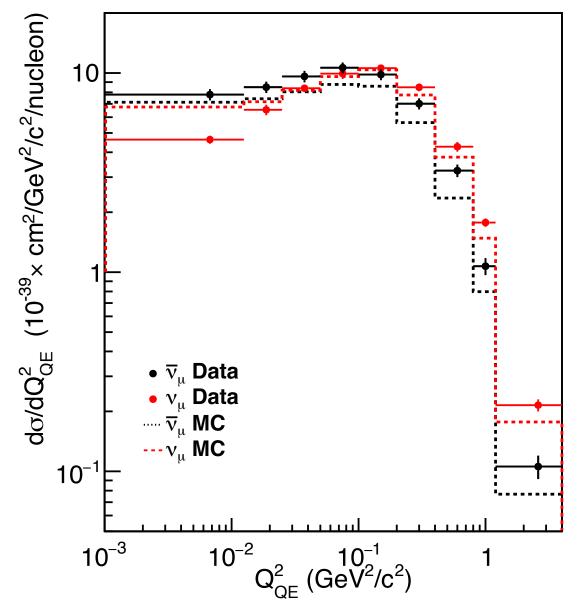
Cross section (left) and comparison of data and models (right).

GENIE 3 model with Nieves and hN (G18-10b series) describes data better at high  $Q_{QE}^2$  than older models. None of the models agree with the data at highest  $Q_{OE}^2$  bin



- Our tunes describe (sort of) the shape of the data (backup for detail) .
  - But needs more than LE tunes to describe our ME data
- Uncertainties on  $Q_{QE}^2$  cross section (right) are dominated by Flux, Muon reconstruction.
- GEANT4 uncertainties are dominated by neutron related interactions.

#### Comparison with MINERvA ME $\nu_{\mu}$ CC0pi measurement



 $u_{\mu} + n \rightarrow \mu^{-} + p$  [*M. Carniero*, PRL 124,121801]  $\bar{\nu}_{\mu} + p \rightarrow \mu^{+} + n$  [This Talk]

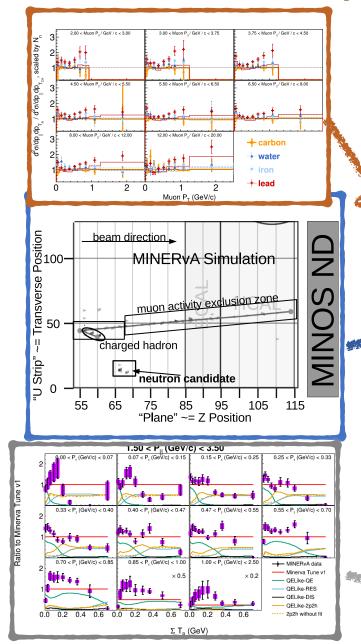
 Not one to one comparison but can show the agreement with data and our model for CCQELike cross section measurements for both neutrinos (black) and anti neutrinos (red)

# Conclusions

#### Conclusions

- This analysis provides high statistics cross section measurement of anti neutrinos CCQE like process.
- Extends measurements to previously unexplored kinematic regions.
- More than LE tunes are needed to describe our ME data.
- Models under predict the data
  - Similar to the LE era analysis
- Higher Statistics, better constrained flux systematics
  - Valuable information for upcoming oscillation experiments.

**CCQELike LandScape** • High Statistics results  $\nu_{\mu}$  [*M. Carneiro*, PhysRevLett.124.121801]



and  $\bar{\nu}_{\mu}$  CCQELike are published.

 $(\frac{o_{\nu_{\mu}}}{\sigma_{\bar{\nu}_{\mu}}})_{CH}$  analysis with full treatment of systematics and

correlations is ongoing.

•  $(\sigma_{\nu_{\mu}})_{A}, (\sigma_{\nu_{\mu}})_{CH}$  and ratios where A is  $C, H_{2}0, Fe, Pb$  [J,

*Kleykamp,* arxiv: 2301.02272]

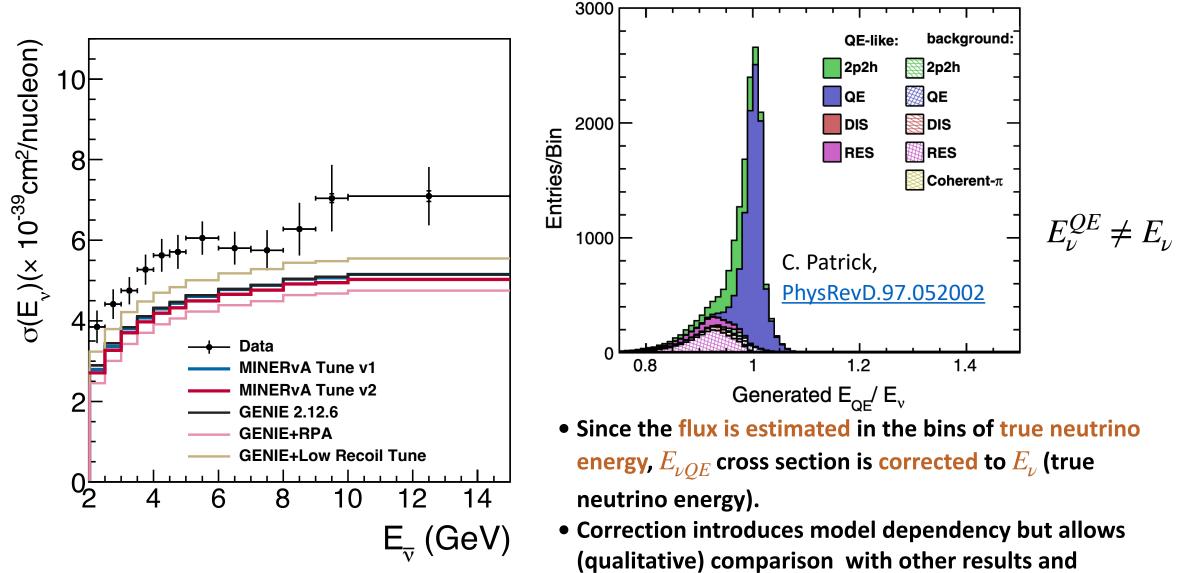
- Neutrons are the Final State particles of this analysis.
- Paper being prepared on cross section with 2 or more neutrons in final state.
- $\bar{\nu}_{\mu}$  CCQELike cross section measurement in heavier target (ongoing)
- $\bar{\nu}_{\mu}$  CCQELike cross section as a function of recoil energy in

muon kinematic bins (ongoing)

- $\nu_{\mu}$  CCQELike version of this analysis published [D.
  - Ruterbories, arxiv 2203.08022]

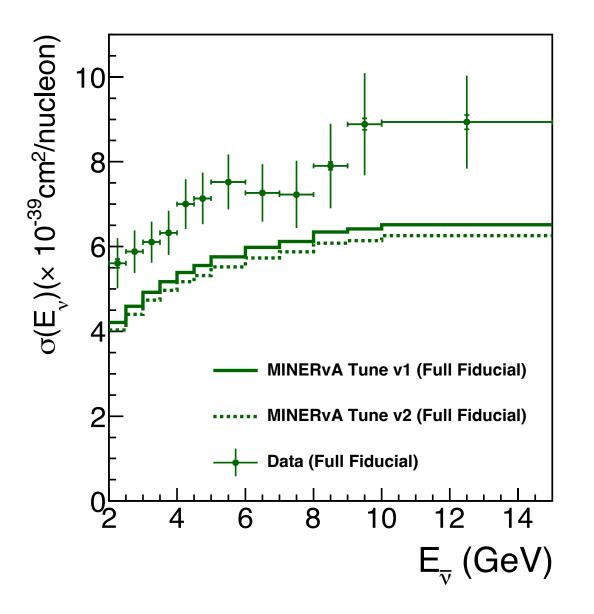
# Back up

## Cross section ( $E_{\nu}$ )



theoretical models.

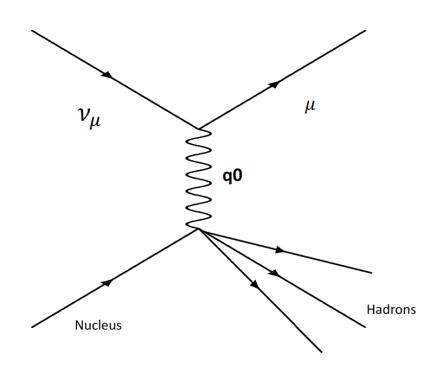
### Full Fiducial Cross section ( $E_{\nu}$ )



Signal Definition includes events with proton KE less than 120 MeV. Remember that our selection criteria requires a 20% angle cut.

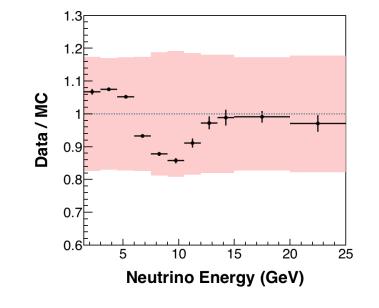
Removing the proton KE threshold and correcting for the angle cut gives the full fiducial cross section. Allows closer comparison with other results.

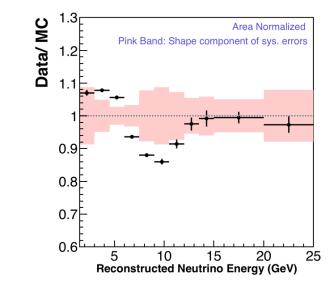
#### Low nu Fit to resolve Data/MC discrepancy



Charged Current Inclusive Event  $u_{\mu} + N \rightarrow \mu^{+} + Hadrons$ 

- $\nu$  (nu)  $\rightarrow$  Energy transferred (q<sub>o</sub>) to the recoil system
- Low nu Events  $\rightarrow q_0 \ll E_{\nu}$
- Cross section Independent of incoming neutrino energy →Shape of Low nu distribution depends on flux shape only.

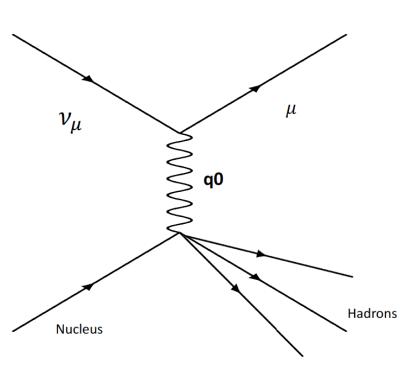




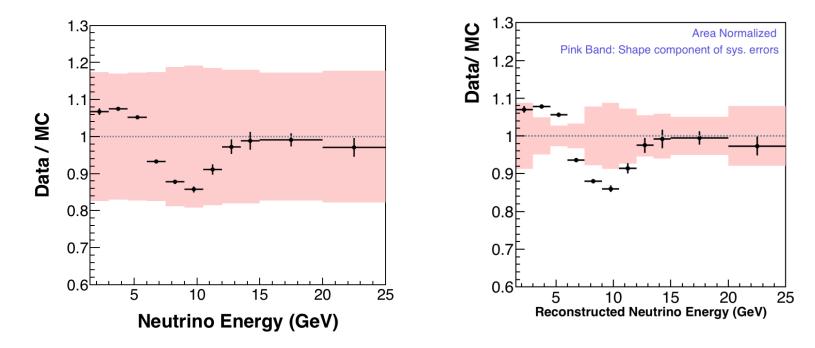
Ratio between MINERvA low nu data and simulated sample. The pink shade shows shape+normalization coverage (left) and shape coverage (right) by the systematic errors. This excludes cross section mismodeling as a candidate cause of discrepancy.

$$\frac{d\sigma}{d\nu} = \frac{G_F^2 M}{\pi} \int_0^1 (F_2 - \frac{\nu}{E_\nu} [F_2 \pm xF_3] + \frac{\nu}{2E_\nu^2} [\frac{Mx(1 - R_L}{1 + R_L} F_2] + \frac{\nu^2}{2E_\nu^2} [\frac{F_2}{1 + R_L} \pm xF_3]) dx$$
$$\frac{d\sigma}{d\nu} \approx \frac{G_F^2 M}{\pi} \int_0^1 (F_2) dx$$

#### Low nu Fit to resolve Data/MC discrepancy



Charged Current Inclusive Event  $u_{\mu} + N \rightarrow \mu^{+} + Hadrons$ 



- Ratio between MINERvA low nu data and simulated sample.
- Left: Pink shade shows shape+normalization coverage by systematic errors.
- Right: Pink shade shows shape coverage by systematic errors.

- Low nu Events  $\rightarrow q_o \ll E_{\nu}$
- Cross section Independent of incoming neutrino energy →Shape of Low nu distribution depends on flux shape only.

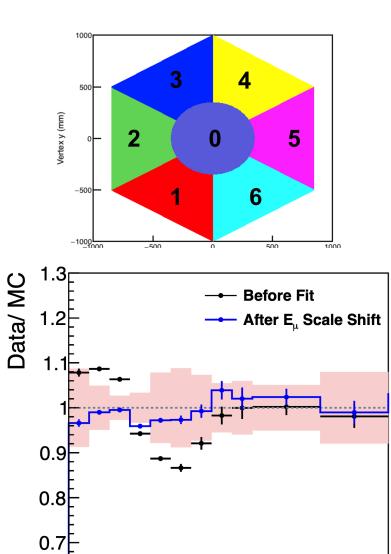
 $\approx \frac{G_F^2 M}{-1} \int_0^1 (F_2) dx$  $rac{d\sigma}{d
u}$ 79

#### **CCQE Cross-section**

$$\frac{d\sigma}{dQ^{2}} \begin{pmatrix} \nu n \to l^{-}p \\ \bar{\nu}p \to l^{+}n \end{pmatrix} = \frac{M^{2}G^{2}\cos^{2}\theta_{c}}{8\pi E_{\nu}^{2}} [A(Q^{2}) + B(Q^{2}) \frac{(s-u)}{M^{2}} + \frac{C(Q^{2})(s-u)^{2}}{M^{4}}]$$
(1)
$$A(Q^{2}) = 4\frac{Q^{2}}{4M^{2}} [|F_{A}|^{2}(1 + \frac{Q^{2}}{4M^{2}}) - |F_{V}^{1}|^{2}(1 - \frac{Q^{2}}{4M^{2}}) + \frac{(1)}{[\xi F_{V}^{2}]^{2}} \frac{Q^{2}}{4M^{2}} (1 - \frac{Q^{2}}{4M^{2}}) + 4F_{V}^{1}\xi F_{V}^{2}\frac{Q^{2}}{4M^{2}}]$$

$$B(Q^{2}) = 4\frac{Q^{2}}{4M^{2}} [F_{A}(F_{V}^{1} + \xi F_{V}^{2})]$$
(2)
$$C(Q^{2}) = \frac{1}{4} [|F_{A}|^{2} + |F_{V}^{1}|^{2} + \frac{Q^{2}}{4M^{2}} [\xi F_{V}^{2}]^{2}]$$
(3)

#### Low nu Fit to resolve discrepancy



10

15

Reconstructed Neutrino Energy (GeV)

20

25

0.6

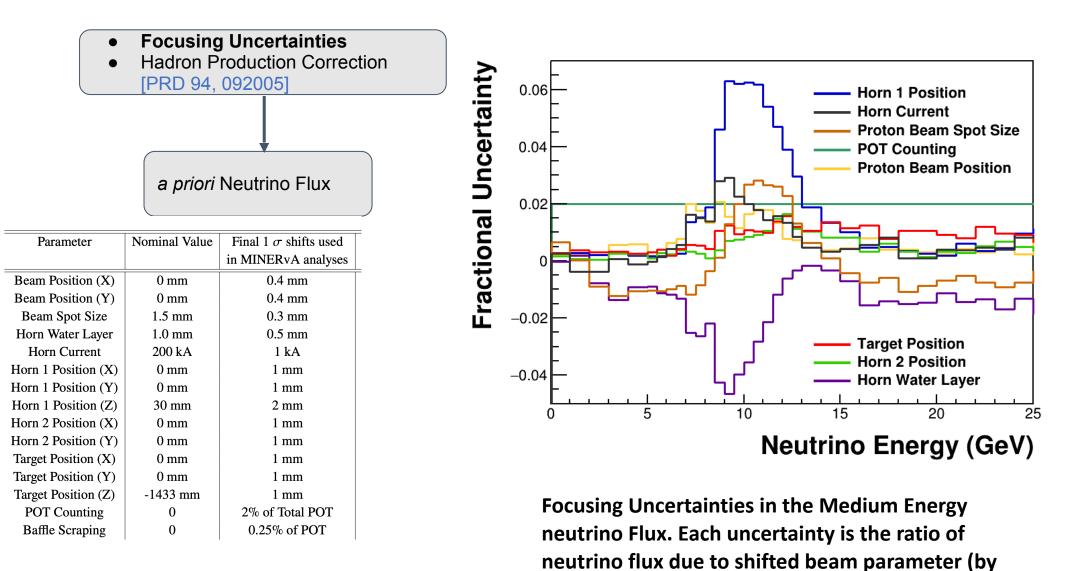
- Transverse face of the MINERvA detector was divided into 7 different areas and a multi parameter fit with the low nu sample was performed.
- Focusing + MINOS Muon energy scale as fit parameters
- Fit with and without prior uncertainty taken into account on each parameter
- Fit with (out) preferred a 3.2 (3.6)% shift in muon energy scale.
- MINERvA shifted muon energy scale by 3.6% for all sample (bottom left plot)

Parameter	Nominal	Best Fit (No Prior)	Best Fit (Prior)
Beam Position (X)	0.0 mm	$-0.3 \pm 0.3 \pm 0.1 \text{ mm}$	$-0.3 \pm 0.2 \pm 0.1 \text{ mm}$
Beam Position (Y)	0.0 mm	$0.8 \pm 0.3 \pm 0.3 \text{ mm}$	$0.7 \pm 0.2 \pm 0.2 \text{ mm}$
Target Position (X)	0.0 mm	$-0.8 \pm 0.3 \pm 0.1 \text{ mm}$	$-0.8 \pm 0.3 \pm 0.1 \text{ mm}$
Target Position (Y)	0.0 mm	$2.3 \pm 0.7 \pm 1.2 \text{ mm}$	$1.7 \pm 0.6 \pm 0.8 \text{ mm}$
Target Position (Z)	-1433 mm	$-1432.4 \pm 2.4 \pm 0.3$ mm	$-1431 \pm 1.8 \pm 0.3$ mm
Horn 1 Position (X)	0.0 mm	$-0.3 \pm 0.4 \pm 0.5 \text{ mm}$	$-0.1 \pm 0.3 \pm 0.1 \text{ mm}$
Horn 1 Position (Y)	0.0 mm	$0.1 \pm 0.5 \pm 0.5 \text{ mm}$	$0.0 \pm 0.3 \pm 0.3$ mm
Beam Spot Size	1.5 mm	$1.41 \pm 0.09 \pm 0.03$ mm	$1.32 \pm 0.09 \pm 0.03$ mm
Horn Water Layer	1.0 mm	$1.2 \pm 0.3 \pm 0.05 \text{ mm}$	$1.3 \pm 0.25 \pm 0.1 \text{ mm}$
Horn Current	200 kA	$198.0 \pm 1.4 \pm 1.4$ kA	$199.1 \pm 0.7 \pm 0.5 \text{ kA}$
Muon Energy Scale	1.0	$1.032 \pm 0.004 \pm 0.008$	$1.036 \pm 0.004 \pm 0.006$
			01

#### Low-nu Fit with Focusing Parameters only

Parameter	Nominal Value	New Value
Beam Position (X)	0 mm	$-0.2 \pm 0.12 \text{ mm}$
Beam Position (Y)	0 mm	$-0.53 \pm 0.14$
Beam Spot Size	1.5 mm	$1.22 \pm 0.14 \text{ mm}$
Horn Water Layer	1 mm	$0.895 \pm 0.16 \text{ mm}$
Horn Current	200 kA	197.41 ± 0.76 kA
Horn 1 Position (X)	0 mm	0. ± 0.17 mm
Horn 1 Position (Y)	0 mm	-0.39 ± 0.17 mm
Target Position (X)	0 mm	$-0.32 \pm 0.17 \text{ mm}$
Target Position (Y)	0 mm	$1.65 \pm 0.5 \text{ mm}$
Target Position (Z)	-1433 mm	-1419.44 ± 1.83 mm

#### **MINERvA Flux Strategy**

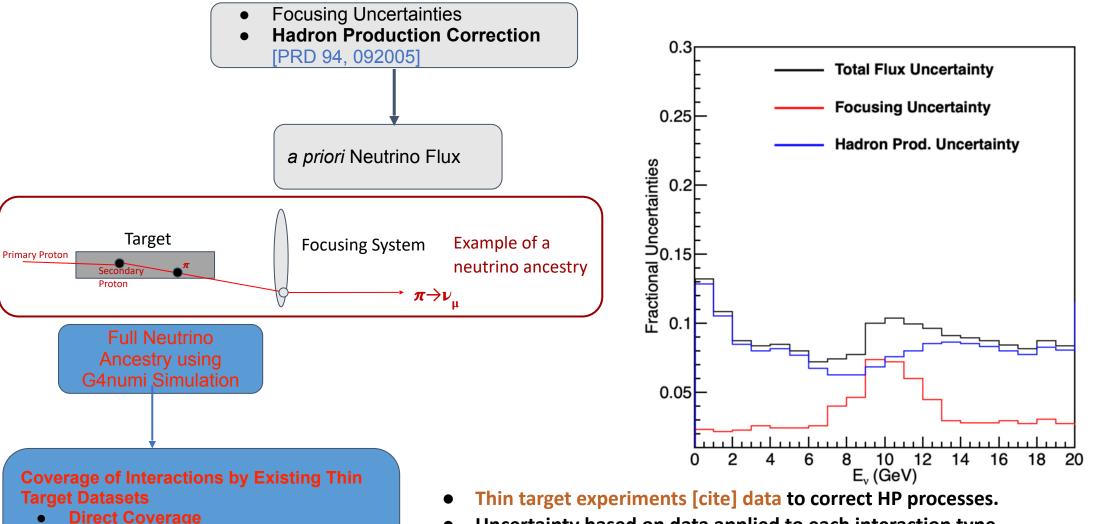


+1  $\sigma$ ) to the nominal neutrino flux.

Table of Beam Parameters and their values at their nominal andshifted position.

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### **MINERvA Flux Strategy**

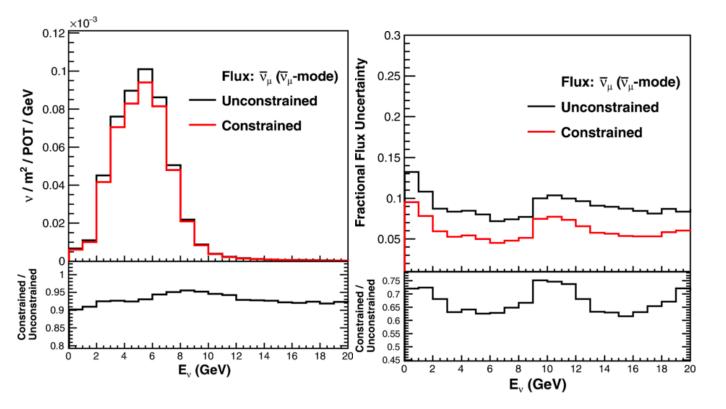


**Coverage by Extension of data sets** 

No coverage at all

- Uncertainty based on data applied to each interaction type.
  - \*Might need plot approval on lower plot (unc from gen2thin (-14) minervame6A) 0

### **MINERvA Flux Strategy**



*a priori* anti neutrino Flux before and after constrained by *v*e scattering data

Uncertainty on predicted anti neutrino flux before and after constrained by  $\nu$ e scattering data • v+e constraint [PRD 107,012001]

 $\circ \nu e \rightarrow \nu e$ 

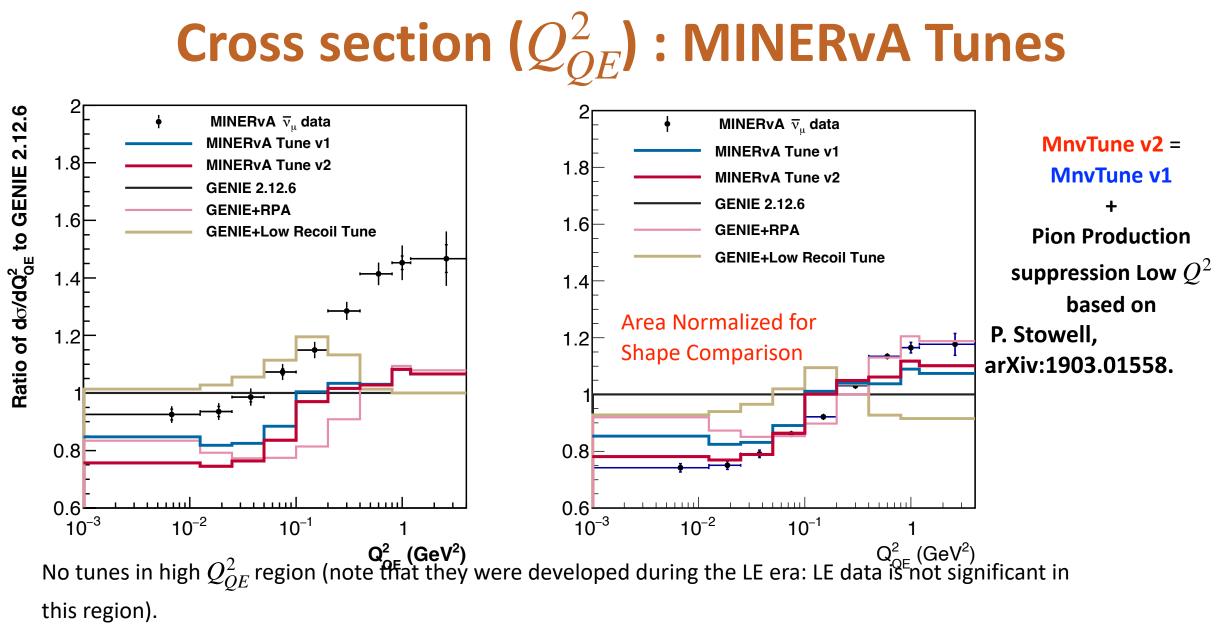
• Inverse Muon Decay [PRD 104,092010]

$$\circ \nu_{\mu}e^{-} \rightarrow \mu^{-}\nu_{e}$$

Neutrino Flux (left) and uncertainty before and after IMD constraint.

Because of  $E_{\nu} \cong 11$  GeV, only constrain high energy

region.



Plots show that our tunes predict the shape of the distribution relatively well (right) compared to absolute distribution.

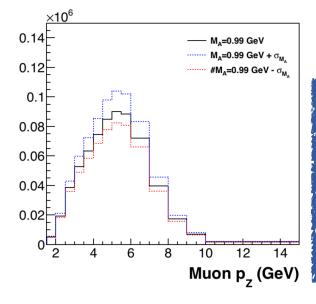
### **Systematic Uncertainties**

**Reconstruction Related Uncertainties:** Uncertainties related to reconstruction of blobs, tracks, PID etc

#### **GENIE Model Related Uncertainties**

Uncertainties coming from **GENIE** models to model neutrino interactions.

They get propagated to our data during background subtraction, efficiency correction and unfolding.



#### A GENIE $M_A$ related systematic



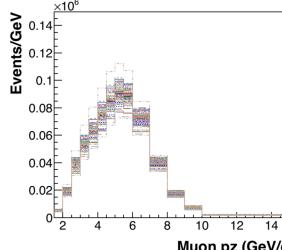
#### • Shift the parameter by $\pm 1\sigma$

- Get the alternate distributions (universe).
- Calculate uncertainty from the spread relative to central value (CV) universe.

Flux Uncertainties (Already covered):

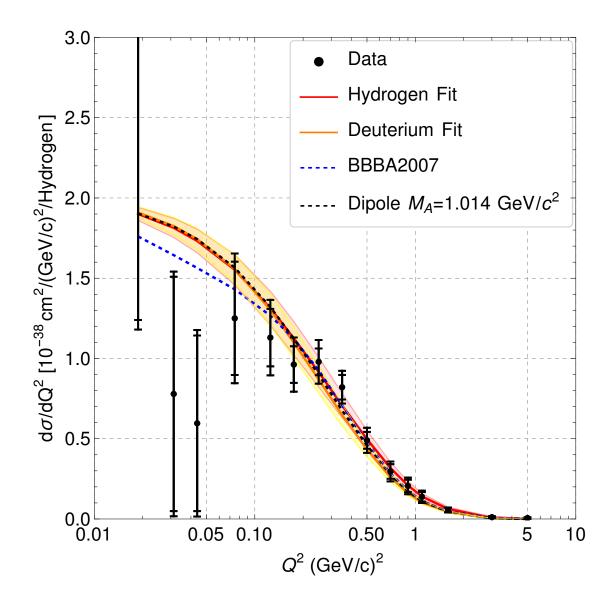
- Focusing Uncertainties
- Hadron Production Uncertainties

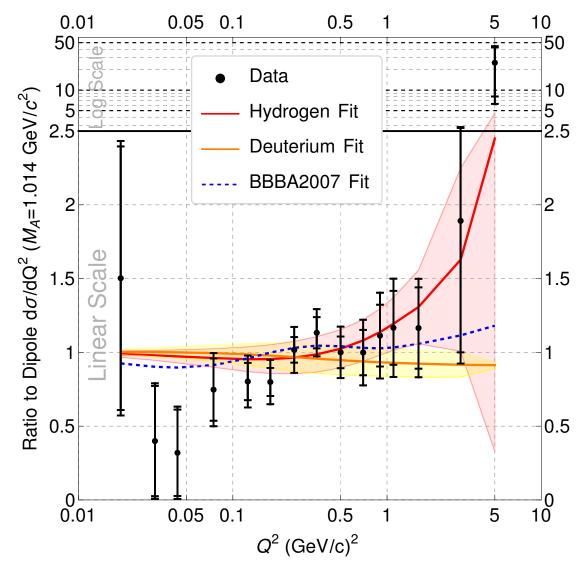
#### Multi Universe Method



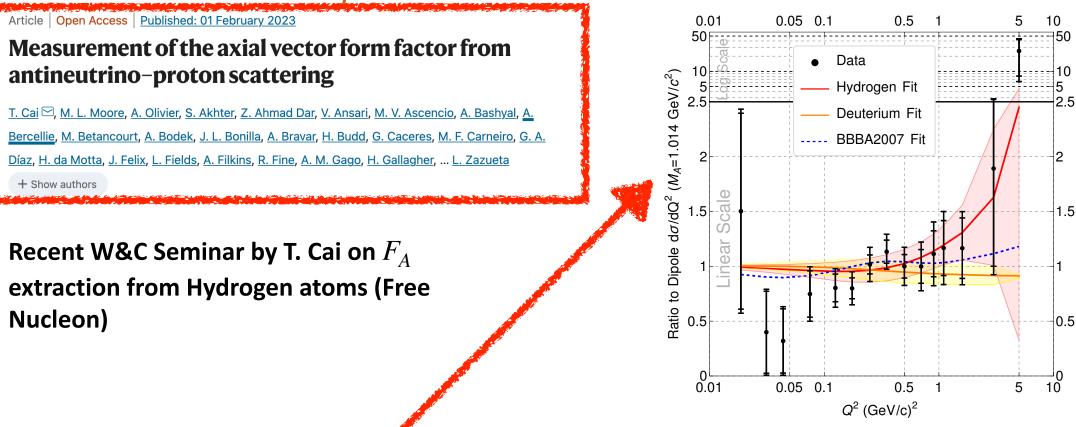
- Many parameters (correlated)
- Shift parameters within  $\pm 1\sigma$
- Generate N universes (500 in this analysis)
- Calculate uncertainty from the spread of the distribution.

Muon pz (GeV/c

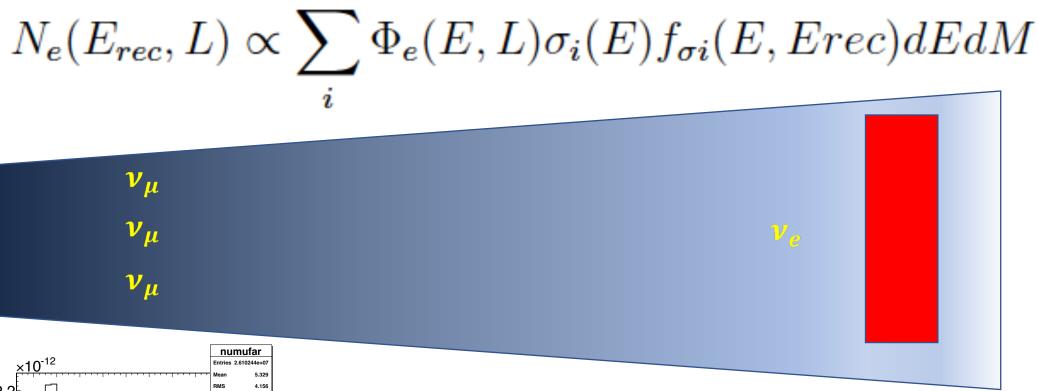


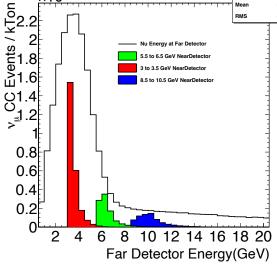


# $\bar{\nu}_{\mu}$ CCQE (H) Cross-section



- $F_A \rightarrow$  Axial Form Factor  $F_A(Q^2) = -\frac{g_A}{(1+rac{Q^2}{M_A^2})^2}$
- Based on bubble chamber (hydrogen targets) measurements
- Measurements in heavier target report slightly higher axial mass
  - Nuclear Effects
- Dipole Form Factor approximation breaks at high  $Q^2$

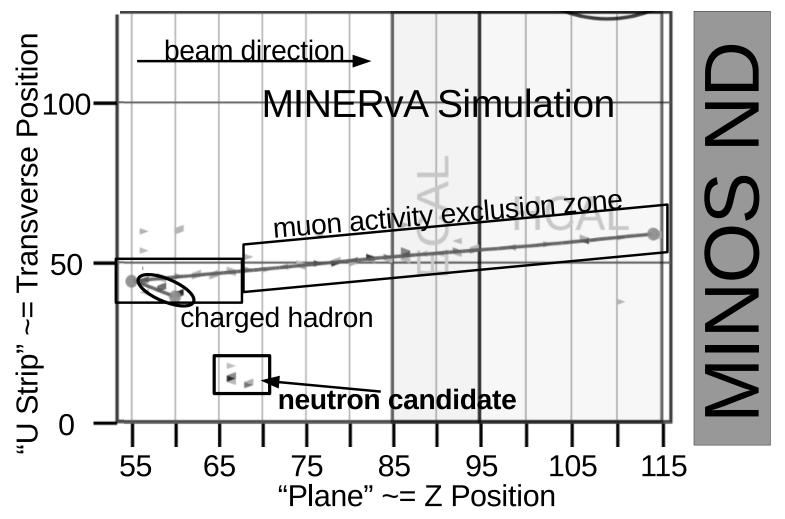




- Near and Far Detector will see different neutrino energy spectra.
- Different **nuclear effects** for same  $E_{rec}$  at Near and Far Detector.
- $\sigma_{\!Near} 
  eq \sigma_{\!Far}$  for the same  $E_{\!rec}$ 
  - Need to understand the nuclear effects in heavy nucleus.

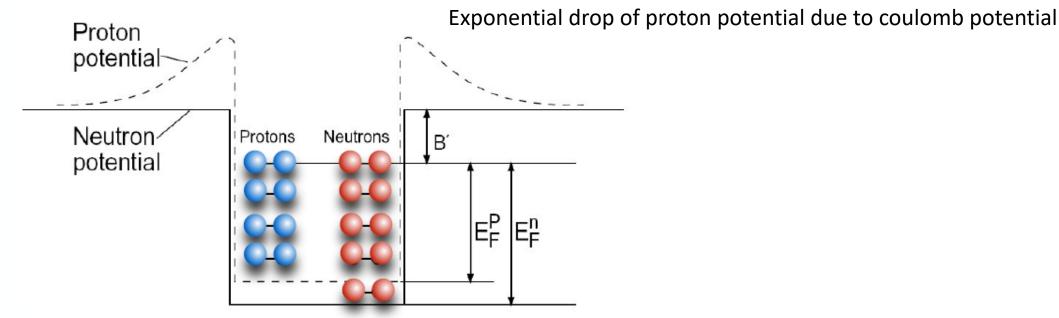
MINERvA was designed to study the poorly understood neutrino-heavy nucleus interactions

#### **Neutron Reconstruction in MINERvA**



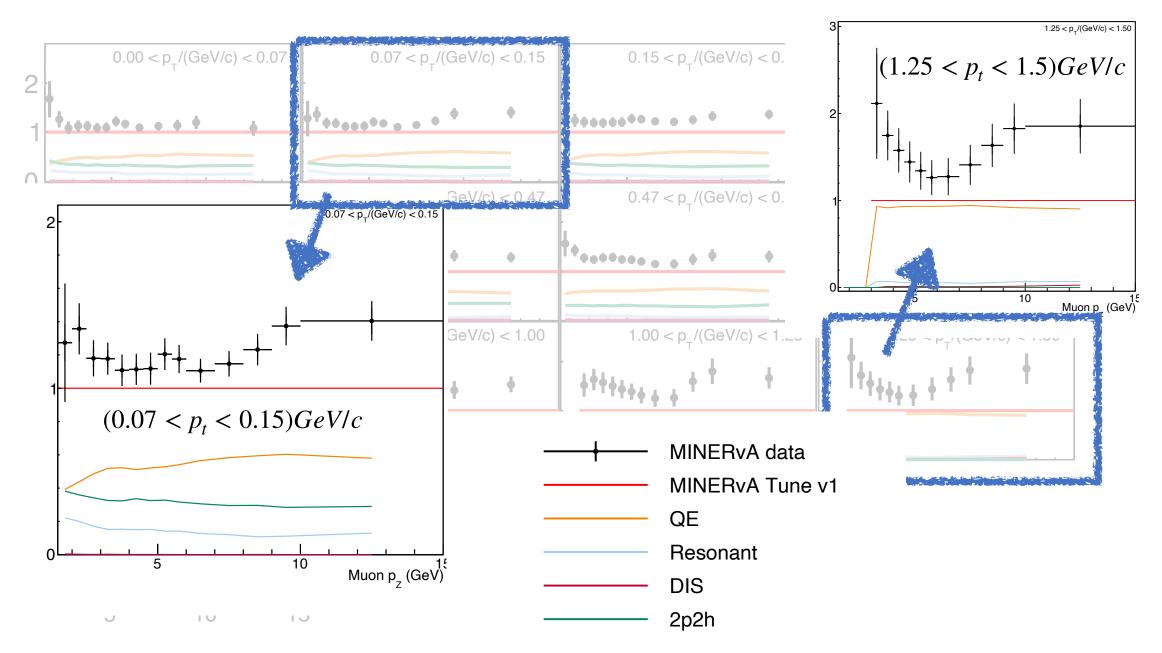
Paper in Preparation to submit to the PRD.

#### Nucleus and Nucleons: Relativistic Fermi Gas Model



#### **Relativistic Fermi Gas:**

- Nucleons as independent particles with some fermi momentum in a mean field generated by the rest of the nucleus.
- Nucleons can interact with other nucleons.



Muon Longitudinal Momentum (GeV/c)