

Baryogenesis and Dark Matter in the Mirror Twin Higgs

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Theory-driven models before

1. **Electroweak scale stability: SUSY**
2. **Dark Matter = WIMPs**
3. **Gauge Unification + Baryogenesis**

Theory-driven models before

1. Electroweak scale stability: SUSY

→ Challenged by LHC data

2. Dark Matter = WIMPs

→ Challenged by DM detection

3. Gauge Unification + Baryogenesis

→ Challenged by
proton decay
experiments

→ High energies
make difficult
to look for...

This talk: Another Option

Mirror Twin Higgs Model

1. **EW Stability: Neutral Naturalness**
2. **DM: Asymmetric Dark Matter**
3. **Baryogenesis \leftrightarrow Dark Matter**

Mirror Twin Higgs Model

Higgs = pNGB

$$SU(4) \times U(1) \rightarrow SU(3) \times U(1) \subset [SU(2)_L \times U(1)_Y]_{\text{Local}}$$

Gauge/Yukawa generate a potential for the Higgs.

\mathbb{Z}_2 parity



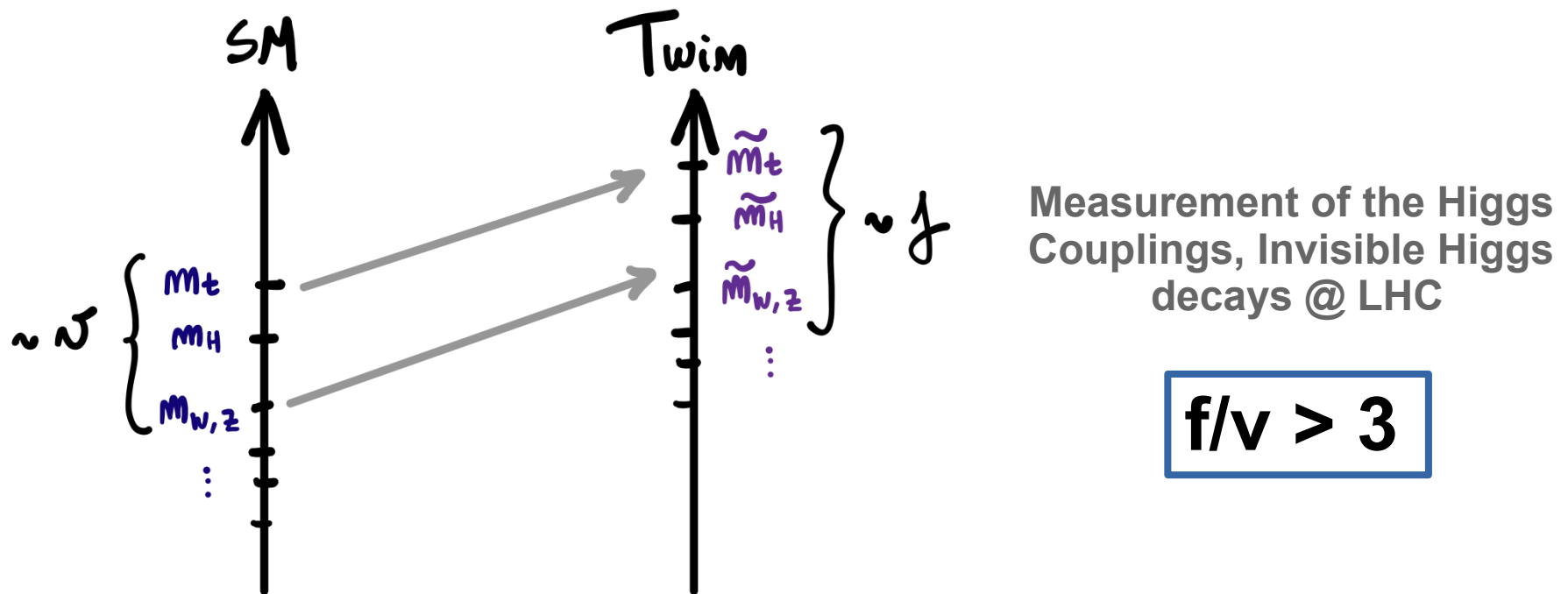
**Restores the full SU(4)
symmetry of the potential**

Consequences of the \mathbb{Z}_2 symmetry

- **Mirror copy of the SM**
- **Hierarchy Problem \rightarrow Neutral Naturalness**
- **Twin Sector \rightarrow Dark Sector**
- **Cosmology,**

**Sandbox to
connect different
BSM problems**

The Twin Higgs Model



- **Soft Z_2 breaking** → **Safe for the HP** (Masses)
- **Hard Z_2 breaking** → **Reintroduces the HP** (Couplings)

Dark Matter in the Twin Higgs

Z_2 + Dark B \rightarrow **Built-in candidates for DM = Twin Nucleons**

 **Asymmetric
Dark Matter**

$$\Omega_{DM} \simeq 5\Omega_{Baryons}$$

 **Common origin of
visible and dark matter**

Twin Dark Matter is naturally ADM

Dark Matter in the Twin Higgs

However ...

$$\frac{\tilde{\Lambda}_{QCD}}{\Lambda_{QCD}} = \left(\frac{f}{v}\right)^{2/9}$$

$$\boxed{f/v \sim 3}$$



$$\Lambda_{QCD_B} \sim 1.4 \Lambda_{QCD_A}$$

$$m_{p,\text{twin}} \simeq 1.4 m_p$$

**Twin Dark Matter cannot
be much heavier than this...**

$$* \frac{\Omega_{DM}}{\Omega_{Baryons}} = \left(\frac{n_{DM}}{n_{VM}}\right) \frac{m_{DM}}{m_p} \neq 5$$

Whereas...

$$\boxed{m_{DM} \simeq 5 m_N}$$

Dark Matter in the Twin Higgs

- Our approach:
1. Choose a baryogenesis model for the visible sector.
 2. Use the Z_2 to copy the mechanism to the twin sector.

- [Literature]
- Farina, JCAP 11 (2015) 017.
 - Farina, Phys.Rev.D 94 (2016) 3, 035017.
 - Garcia Garcia, Phys.Rev.Lett. 115 (2015) 12, 121801.

- * Hard Z_2 breakings in QCD/Yukawa couplings
- * Connect the two sectors with portals beyond the Higgs

It is possible to increase the twin DM mass by brute force.

 But this requires hard Z_2 breaking!

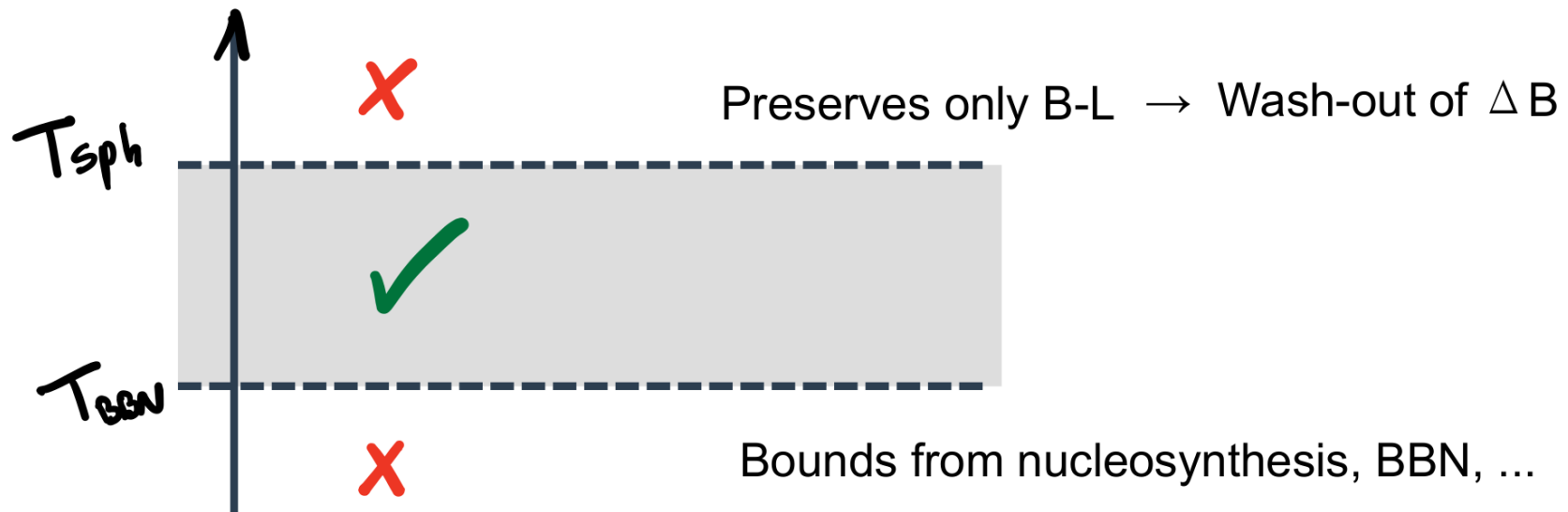
We do not want that.

Baryogenesis Model

Example: Low-temperature Baryogenesis

Generate the B asymmetry directly through out-of-equilibrium decays

Window: $T_{\text{BBN}} < T_{\text{baryogenesis}} < T_{\text{Sphalerons}}$

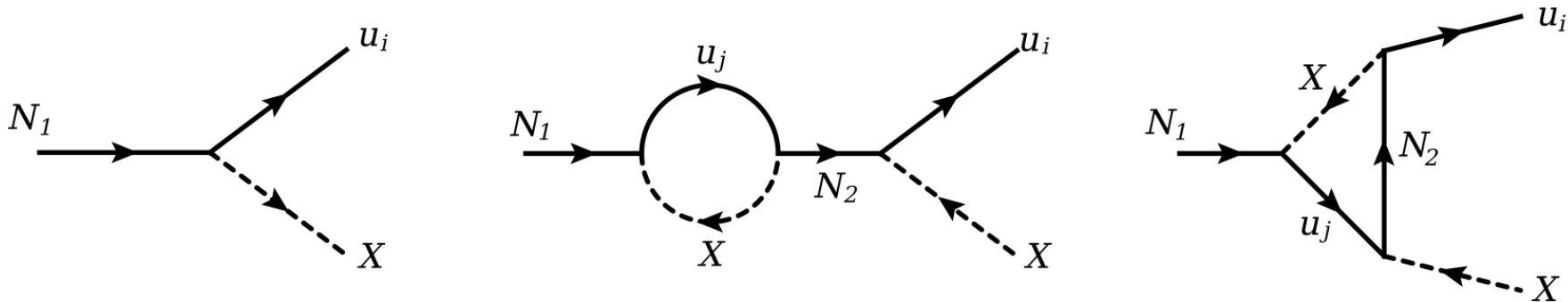


Baryogenesis Model

- Need:
- 2 flavors of N_α $\alpha=1,2.$
 - 1 colored X

$$\Delta\mathcal{L}_{\text{Baryogenesis}} = \lambda_{i\alpha} N_\alpha u_i^c X + \lambda'_{ij} d_i^c d_j^c \bar{X} - M_\alpha N_\alpha N_\alpha + M_X^2 X \bar{X}$$

Decay of N_1 generates the baryon asymmetry



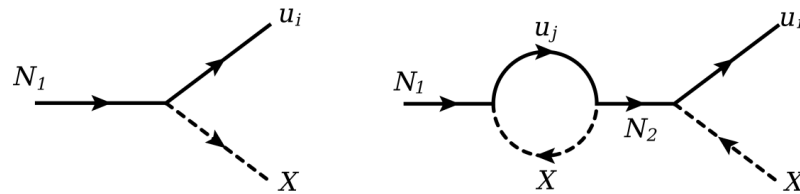
Baryogenesis Model

- Assume that N_1 was efficiently produced in the early universe.
- Calculate the **Baryon asymmetry** from N_1 's **Yield** and the **CP asymmetry**.

$$Y_{\Delta B} = \frac{n_{N_1}}{s} \left(\frac{\Gamma(N_1 \rightarrow B) - \Gamma(N_1 \rightarrow \bar{B})}{\Gamma(N_1 \rightarrow \text{tot})} \right) \equiv Y_{N_1} \epsilon_{CP}^{N_1}$$

$$\frac{Y_{\Delta B}}{8.7 \times 10^{-11}} = \left(\frac{213.5}{g_{*,S}(T_{FO})} \right) \left(\frac{3\mathcal{F}_S(1.5) + \mathcal{F}_V(1.5)}{15.3} \right) \left(\frac{|\lambda_{u2}| \sin^{1/2} \phi}{2.3 \times 10^{-4}} \right)^2$$

$\Phi \rightarrow$ Phases of the couplings in the interference of:



Asymmetric Dark Matter

What about dark matter?

Same story: Copy $\Delta\mathcal{L}_{\text{Baryogenesis}}$ with the the Z_2 symmetry

(We do not want hard Z_2 breakings)

Allowed soft breakings:

- **Soft f/v breaking:** Twin N masses get lifted

↷ CP asymmetry depends on the ratio of twin N_1 and N_2 masses

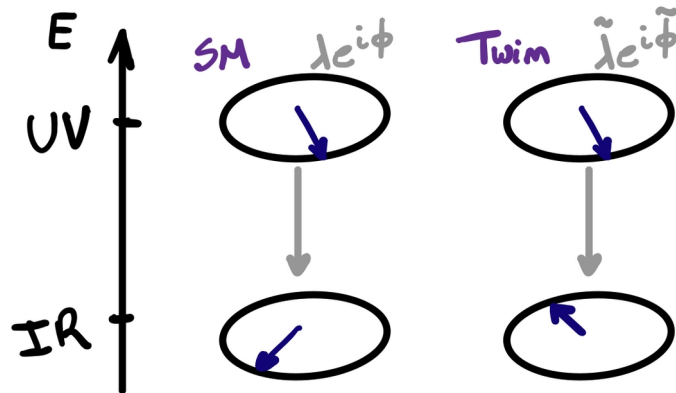
↷ Not very effective in changing the baryon asymmetry in the twin sector

- **Soft phase breaking:** CP violating phase is not fixed for the twin sector

Twin Baryogenesis

Different phases between the SM and the Twin sectors

- Phases do not reintroduce the Hierarchy Problem
- A priori, no reason for equal phases since we do not provide any mechanism that generate them



Analogies:

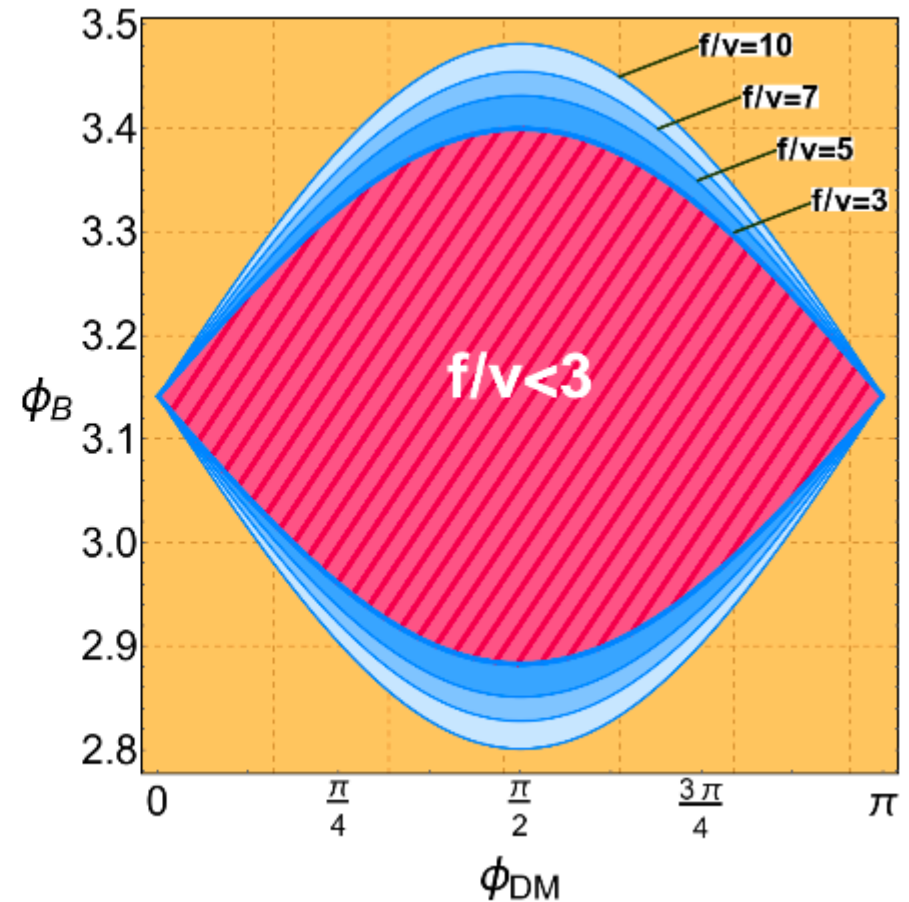
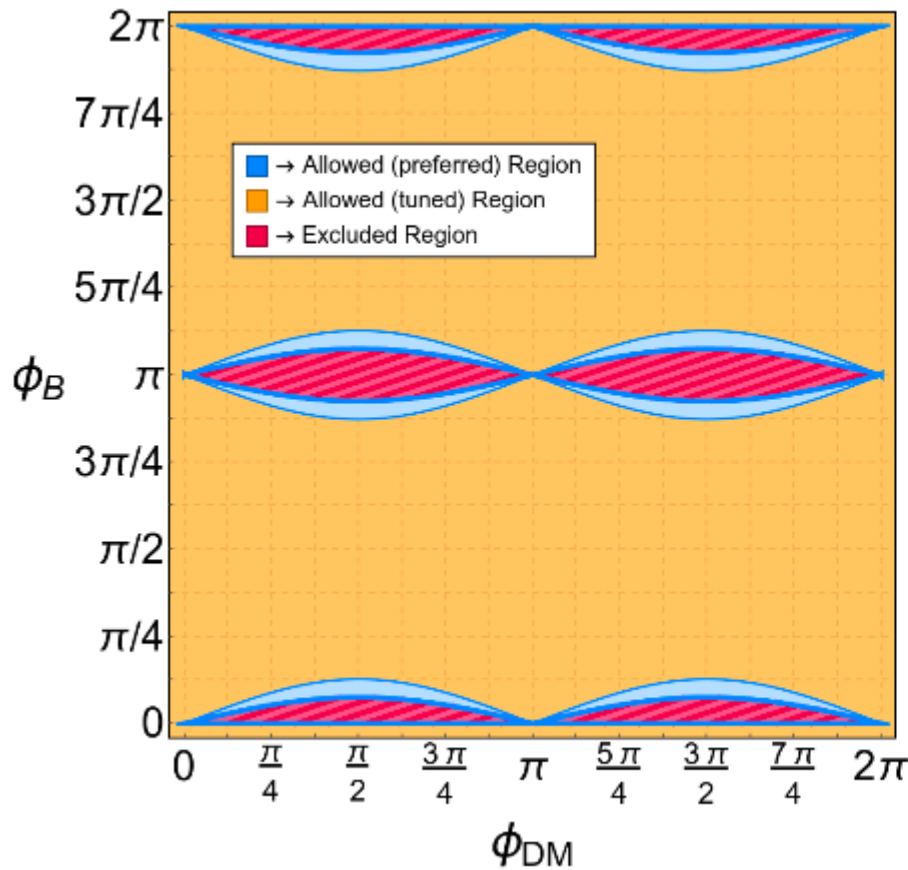
- “Vacuum misalignment”
- “Ferromagnets”
- “Twin CKM phase”

The same reasoning apply to different baryogenesis models:

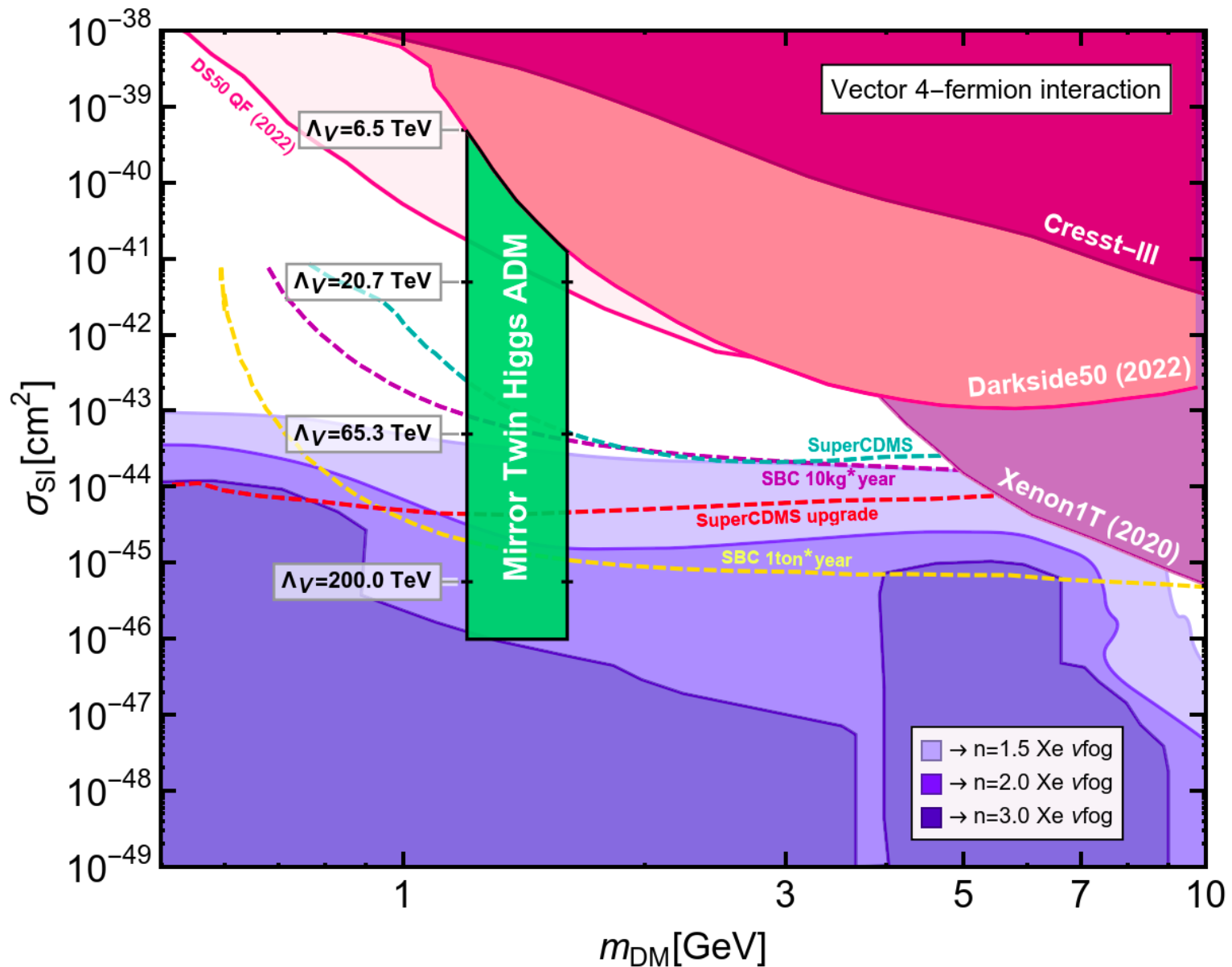
i.e. High temperature Leptogenesis

Twin Baryogenesis

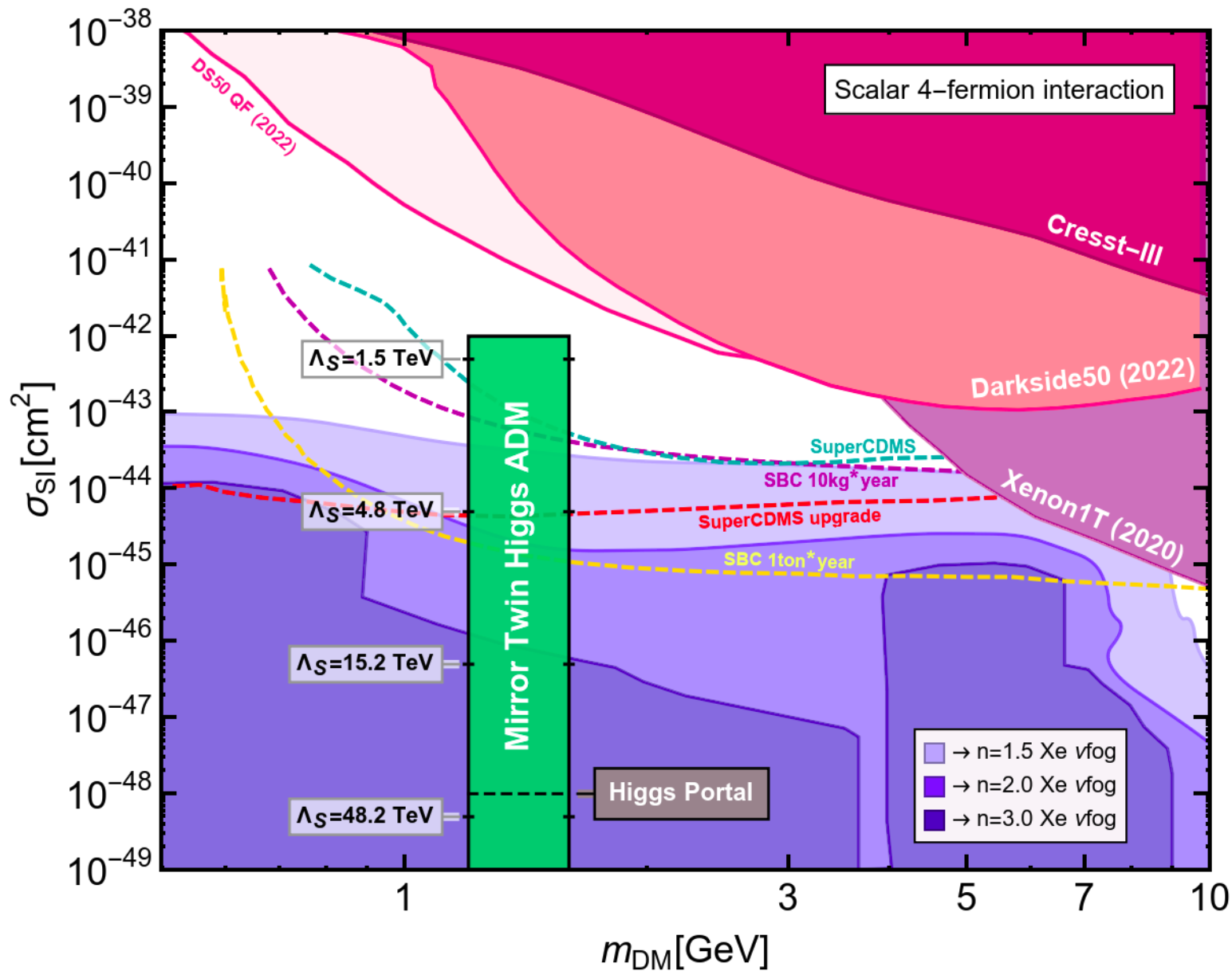
Different phases between the SM and the Twin sectors



Direct Detection: Vector interaction



Direct Detection: Scalar interaction



What about the LHC?

Low temperature Baryogenesis ↔ New colored particles at TeV

- Theory-motivated models of GeV DM might start receiving attention once we reach the neutrino fog.

→ **Low scale baryogenesis is specially interesting for the Twin Higgs.**

- Scalar di-quarks/Leptoquarks
- Hadronic displaced vertices

From the Twin Higgs side → New physics is naturally hidden

- Higgs couplings modifications
- Higgs invisible decays

Theory motivation for experiment

Summary

1. Twin Higgs as a sandbox for connecting different BSM phenomena
2. Baryogenesis model $\xrightarrow{Z_2 \text{ Copy}}$ DM genesis model
3. Phase misalignment: CP violating phases can be different in the mirrored sector
4. Look forward for direct detection around ~ 1 GeV!



Thank you!