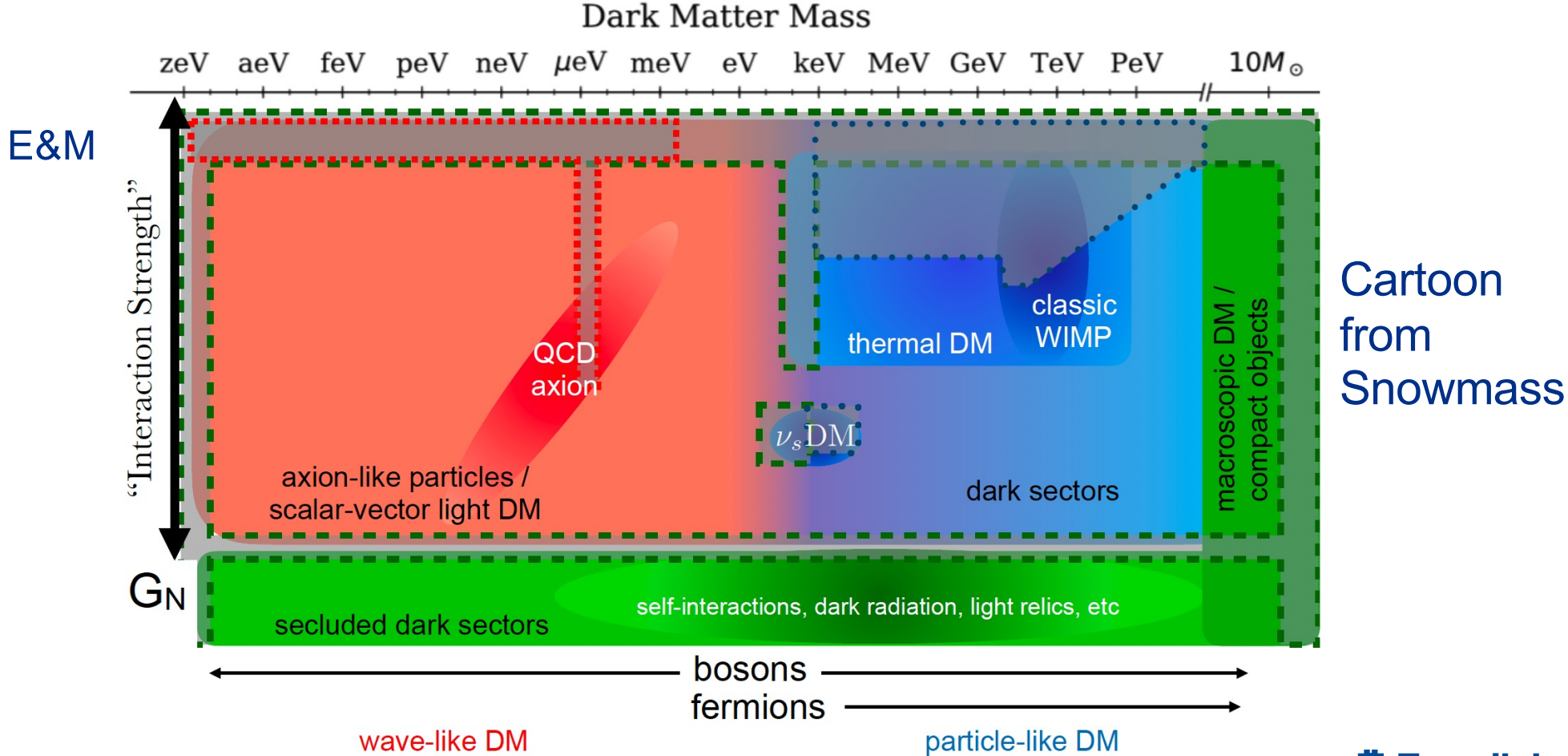
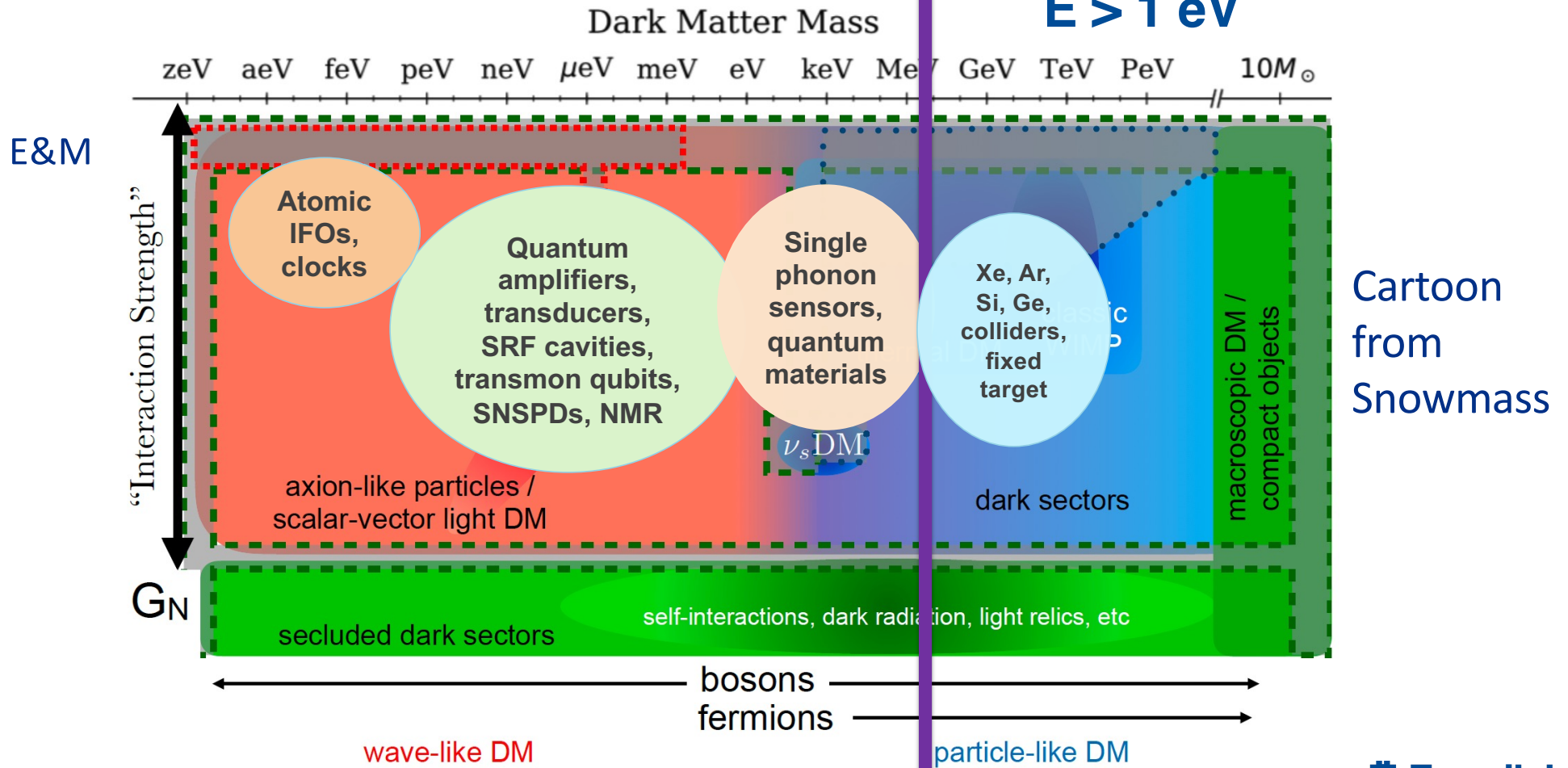


# Dark Matter Strategy: Delve Deep, Search Wide ... but how?

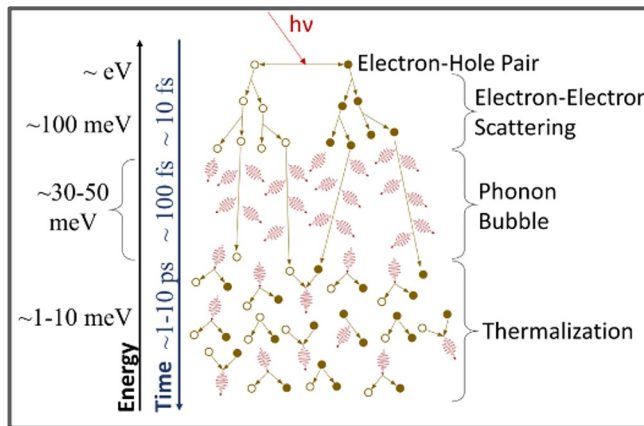
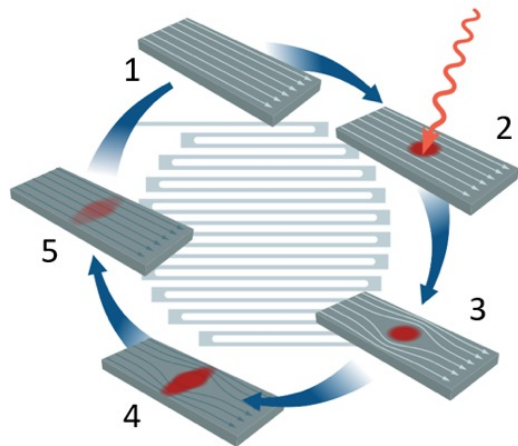


Cartoon from Snowmass

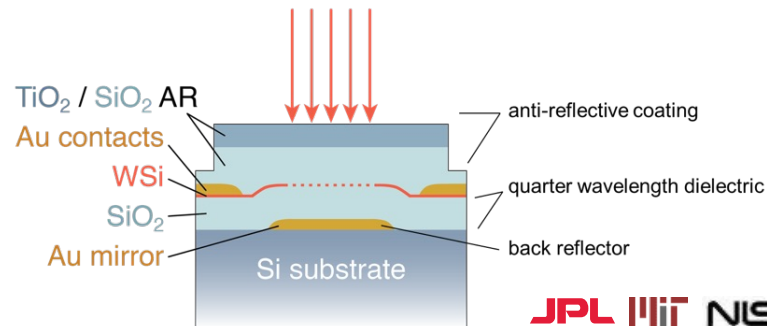
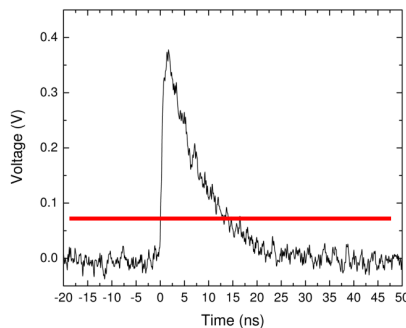
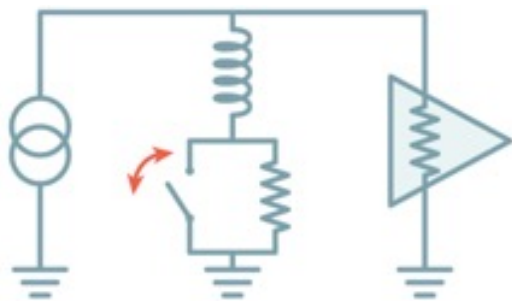
Quantum sensors:  $E < 1 \text{ eV}$  ← → Conventional tech:  $E > 1 \text{ eV}$



# Superconducting Nanowire Single Photon Detectors Matt Shaw (JPL)

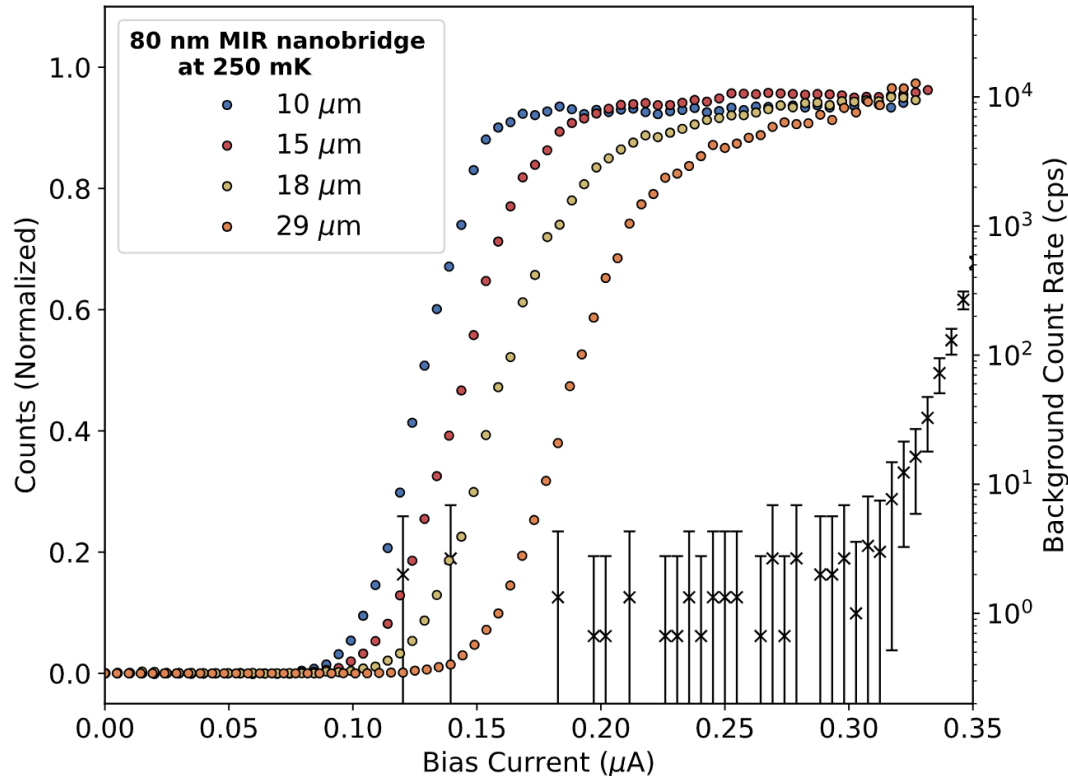


- 98% System Detection Efficiency at 1550 nm (NIST)
- 0.27 – 29  $\mu\text{m}$  operating wavelength (JPL / NIST / MIT)
- $<1\text{e-}5$  cps dark count rate (MIT / NIST)
- $\text{mm}^2$  active area (JPL / NIST / MIT)
- 400 kpix camera format (NIST)
- 1.5 Gcps maximum count rate (JPL)
- $<3$  ps FWHM timing jitter (MIT / JPL / NIST)
- High radiation tolerance



# New Result: Single Photon Detection at 29 $\mu\text{m}$

Matt Shaw (JPL)

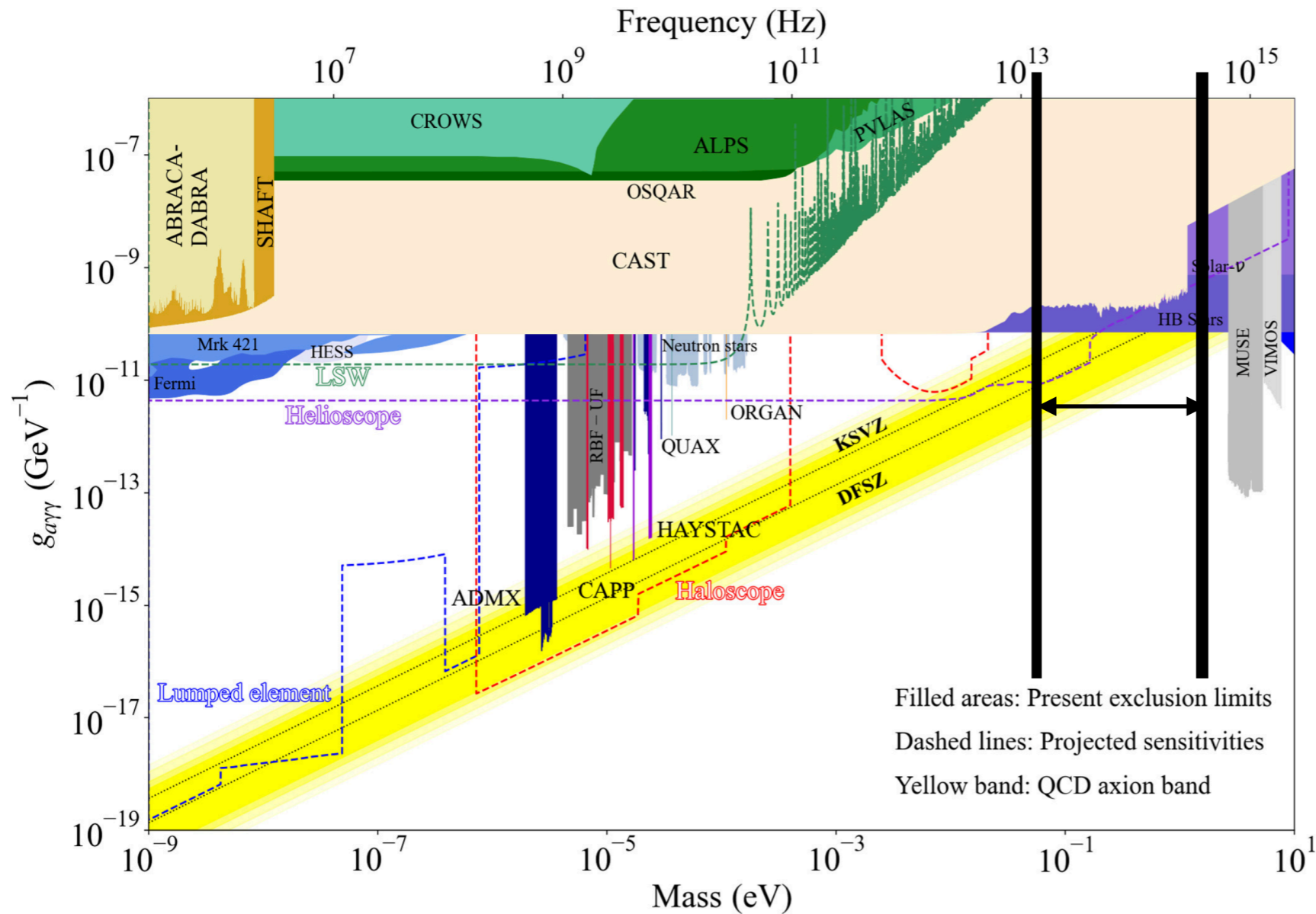


- SNSPD nanobridge with Si-rich WSi for reduced superconducting gap energy (1.3 K  $T_c$ , 80 nm nanobridge)
- First demonstration of single-photon counting at far-infrared wavelengths (43 meV, 10 THz)

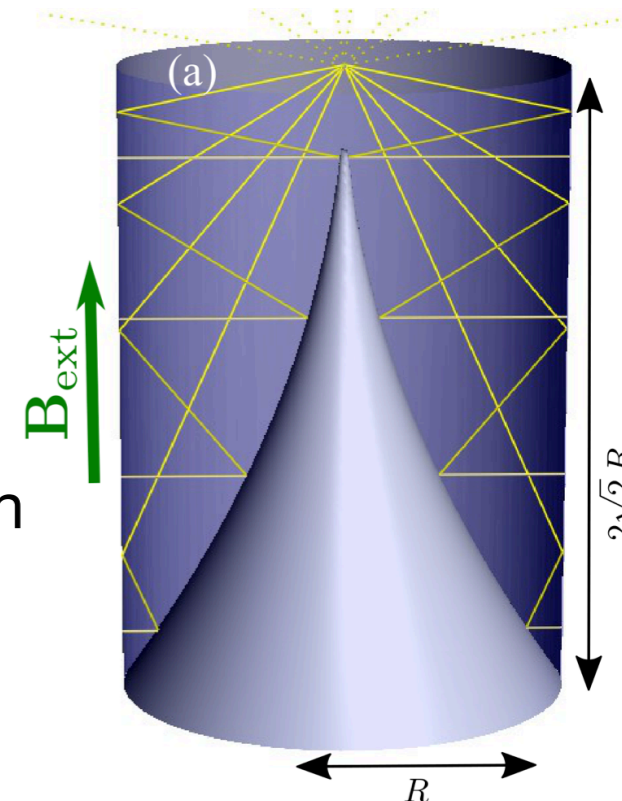
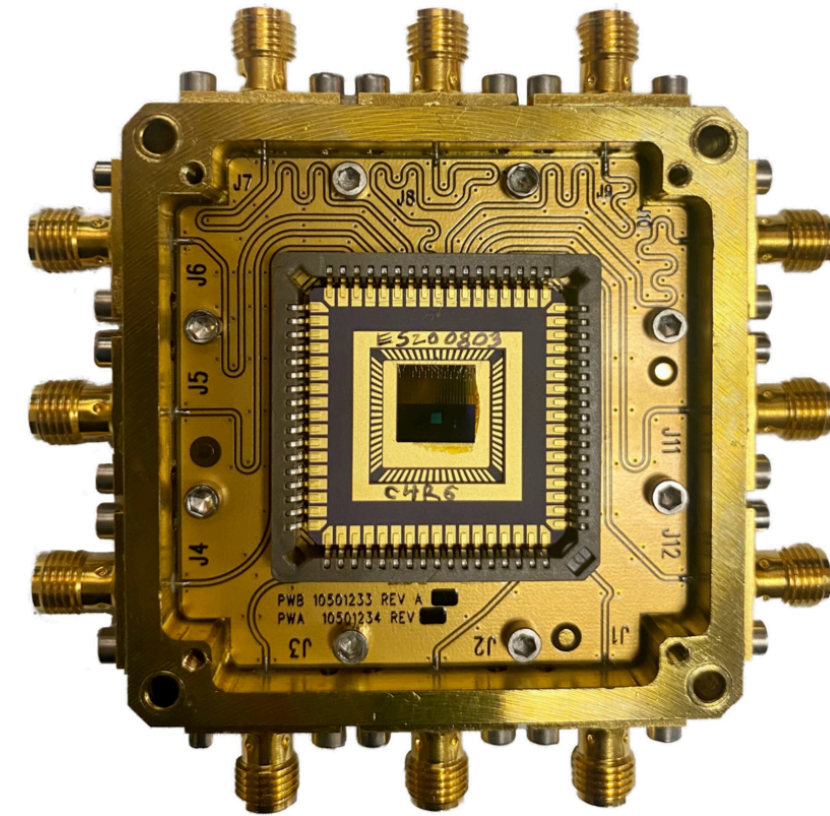


# Axions with low-threshold SNSPD

Low-threshold photon detectors will play a key role in broadband axion detection



SNSPD overview:  
see M. Shaw's slides



SNSPD demonstrated nanowire sensitivity to 0.04 - 1 eV  $\rightarrow$  mass reach

Best DCR  $< 1\text{e-}5$  c.p.s  $\rightarrow$  axions coupling sensitivity

Final SNSPD to realize axion detection require challenging R&D

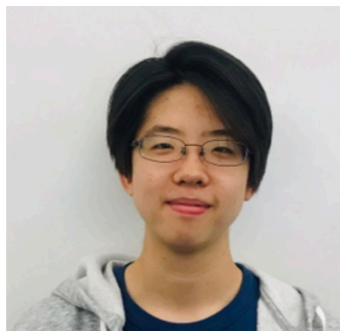


# Axions with low-threshold SNSPD

**BREAD**  
COLLABORATION



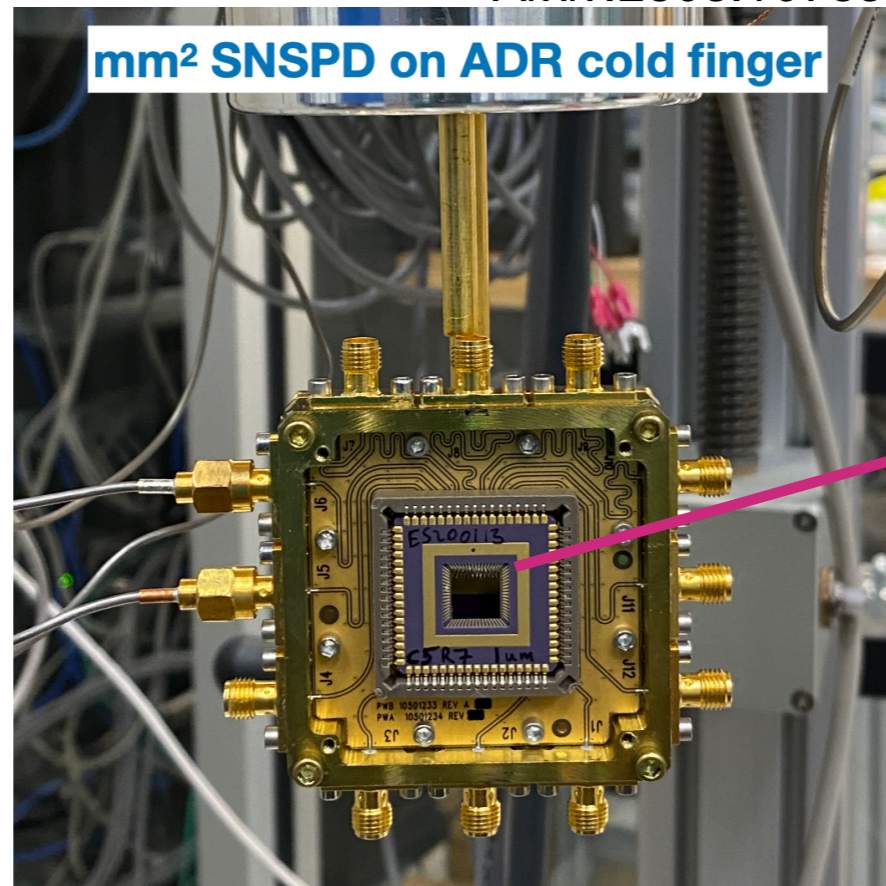
Jamie Luskin



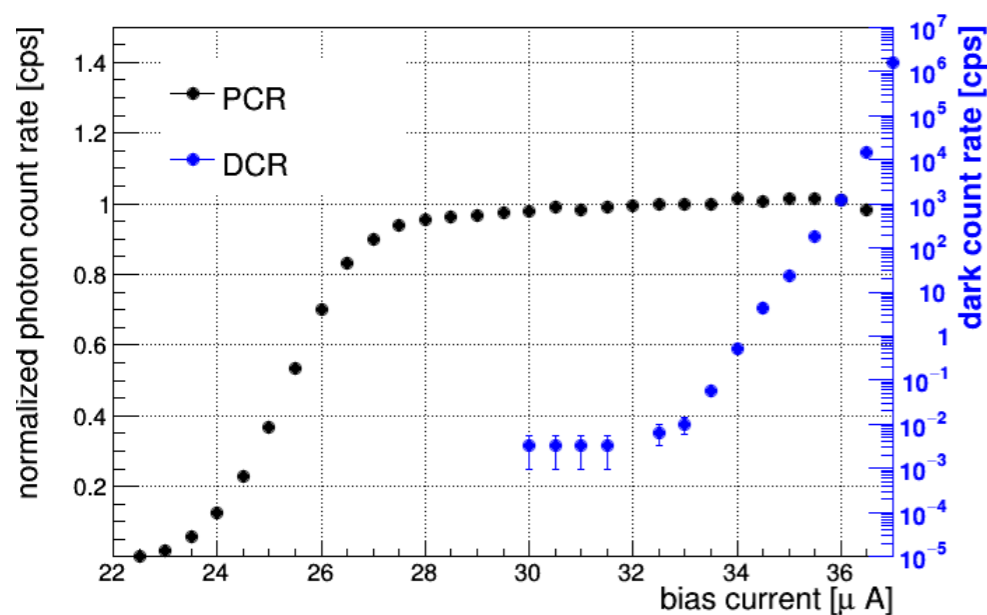
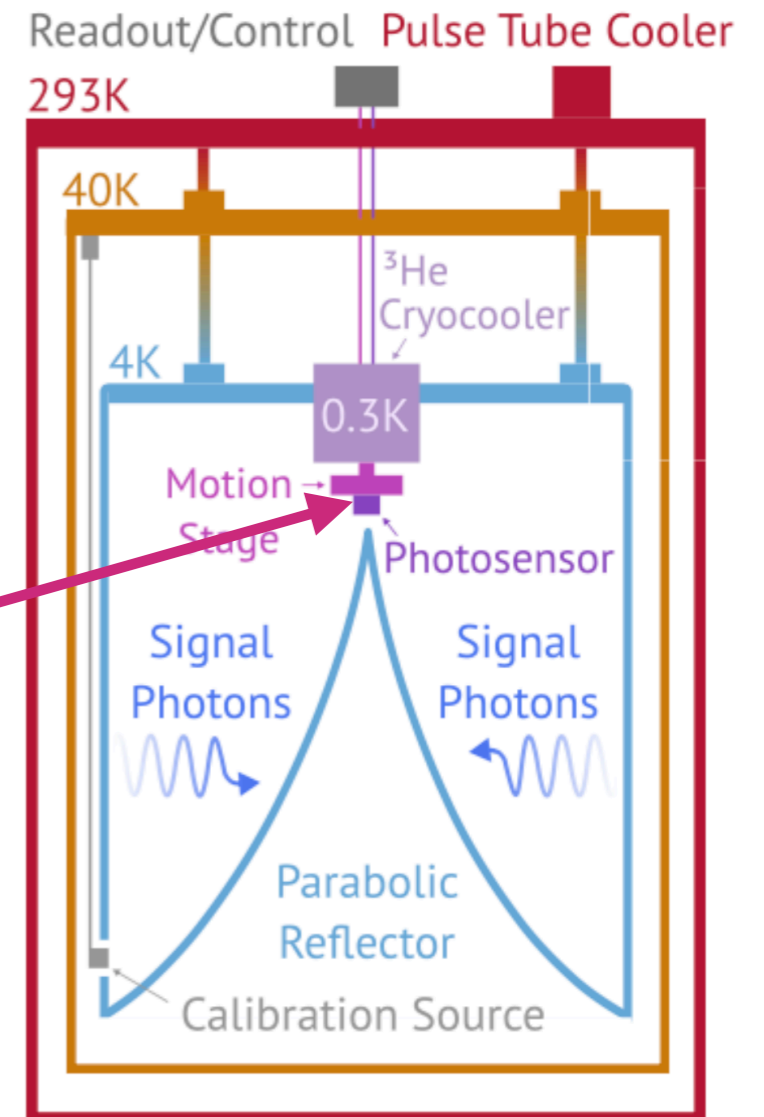
Christina Wang

ongoing R&D to meet  
sensor parameters

Arxiv:2303.10739



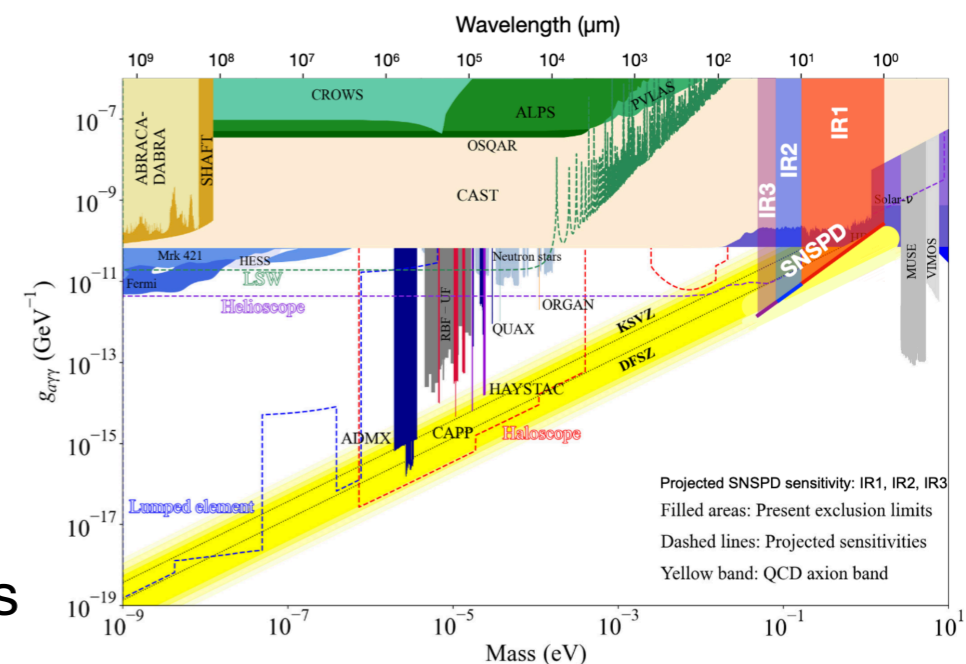
mm<sup>2</sup> SNSPD on ADR cold finger



Towards new  
axion sensitivity



SNSPD area/channel  
scaling: see S. Xie's slides

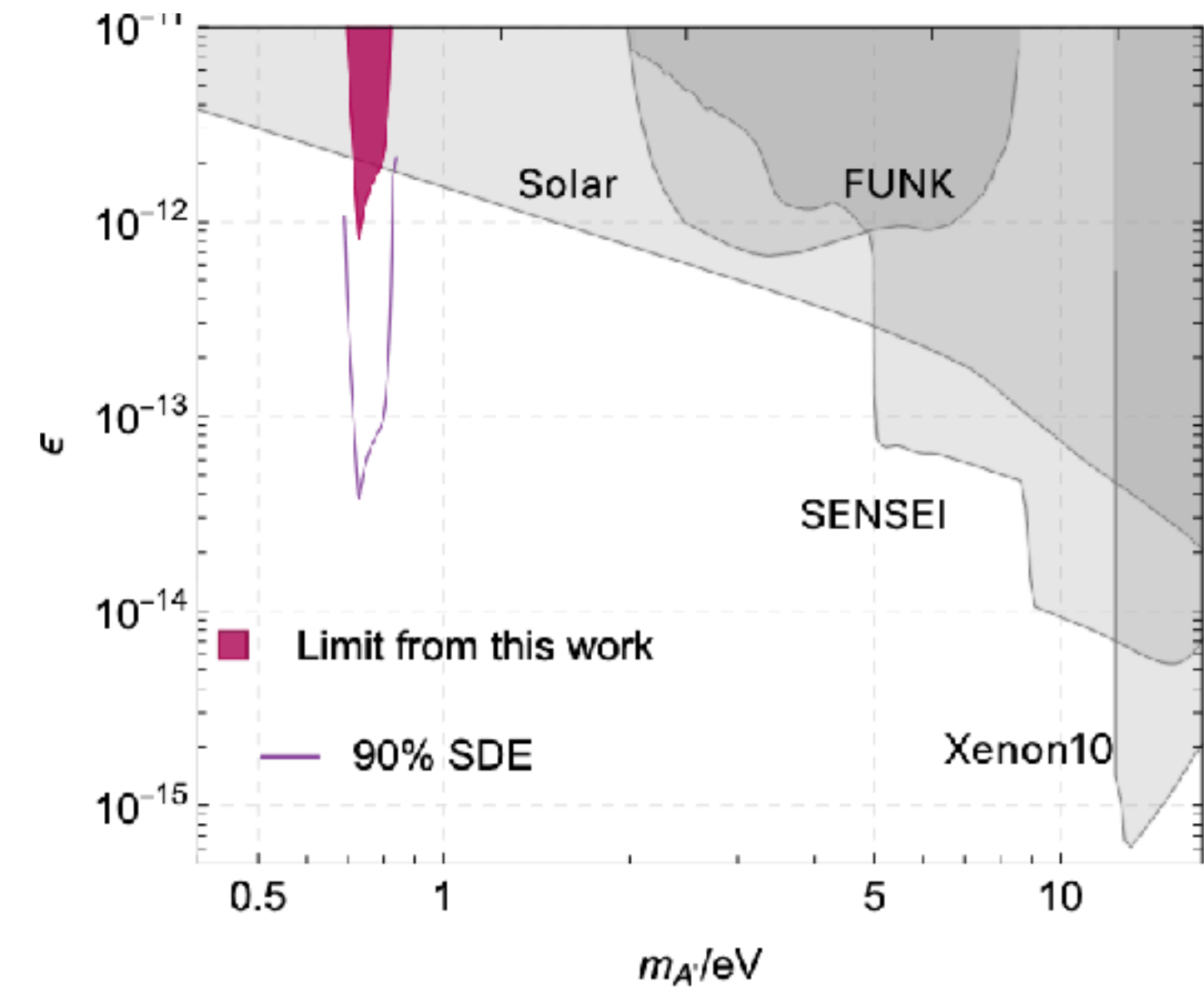
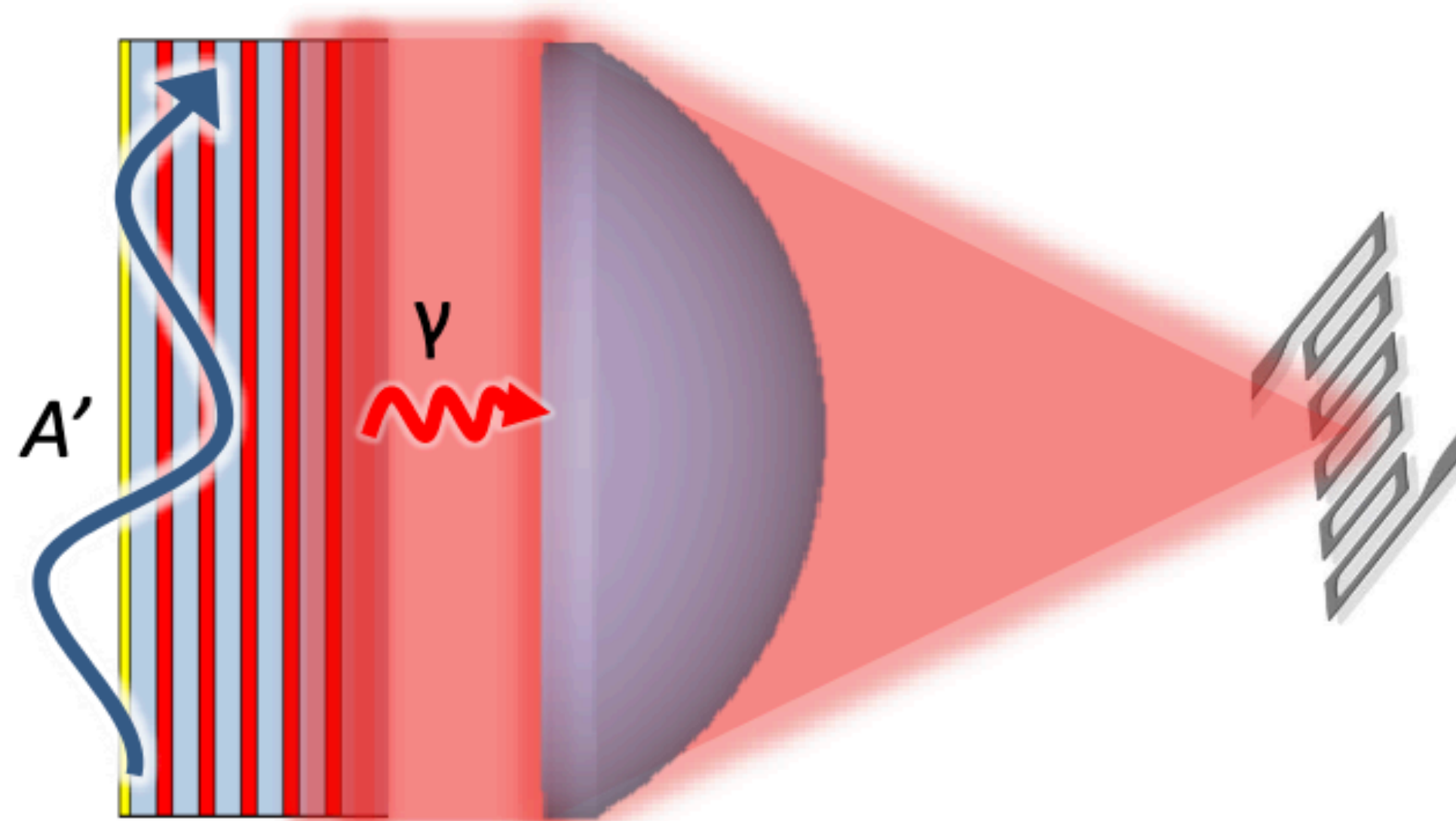
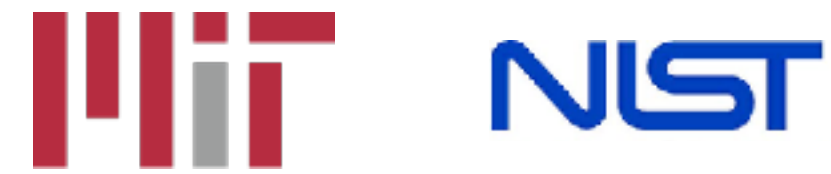




# LAMPOST: Dielectric Haloscope with an SNSPD

Stewart Koppell<sup>1</sup>, Dip Joti Paul<sup>1</sup>, Jeffrey Chiles<sup>2</sup>, Junwu Huang<sup>3</sup>, Masha Baryakhtar<sup>4</sup>, Sae Woo Nam<sup>2</sup>, Karl Berggren<sup>1</sup> (1. MIT, 2. NIST Boulder, 3. Perimeter Institute for Theoretical Physics, 4. University of Washington)

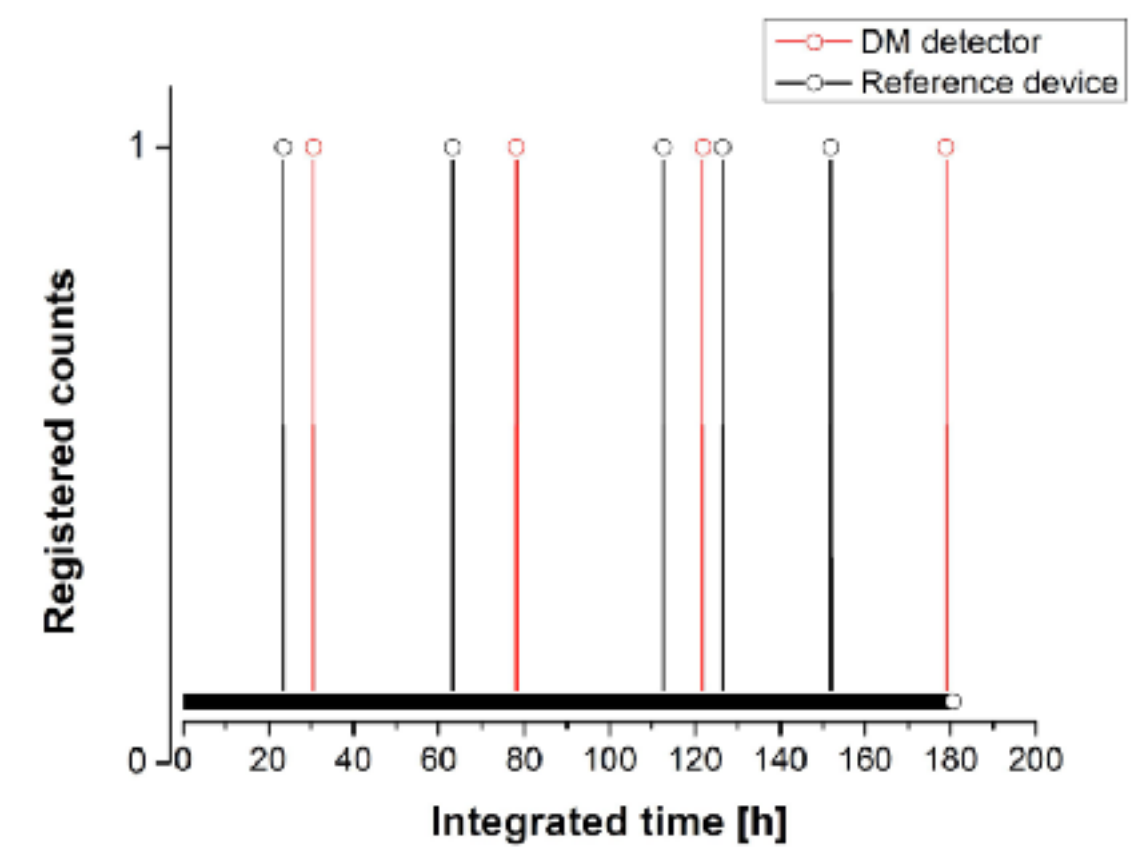
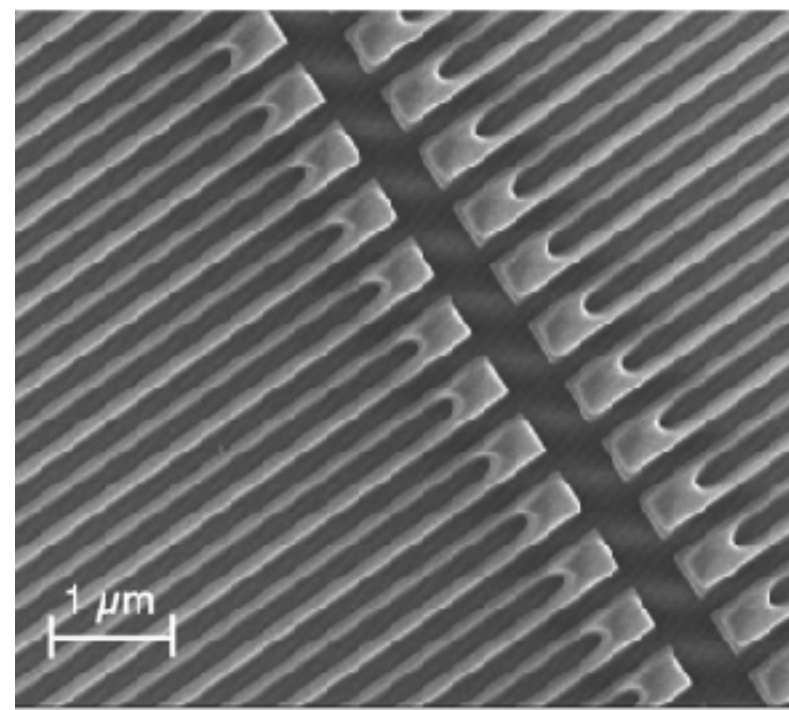
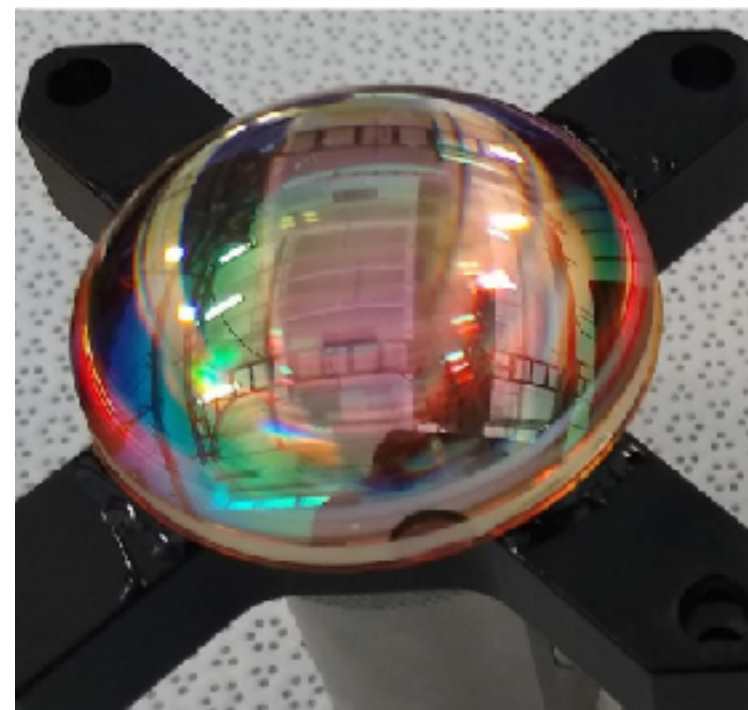
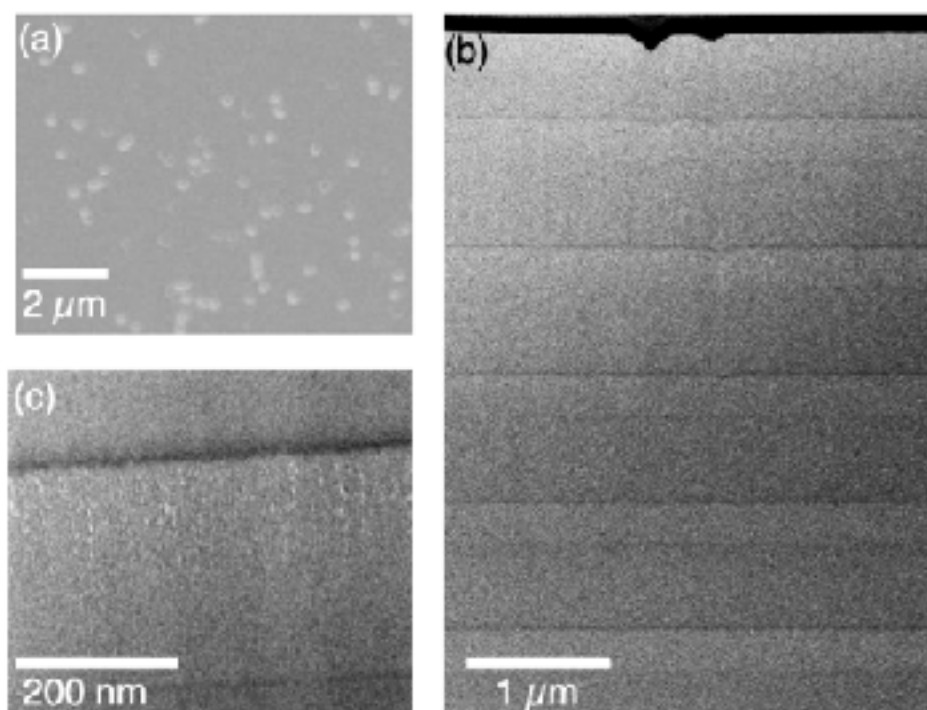
Ilya Charaev Jeffrey Chiles



Dark Matter Absorber

Lens

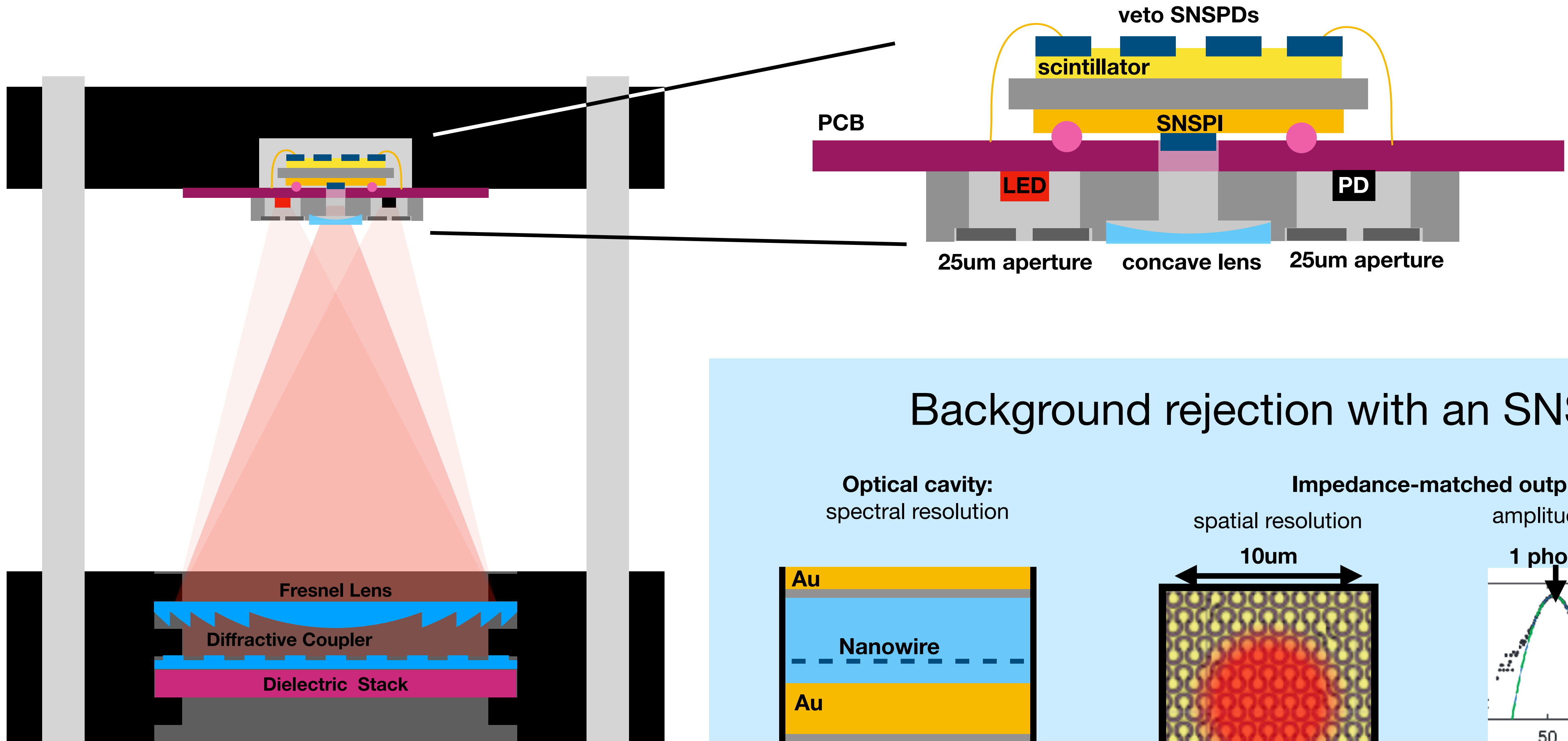
SNSPD



J. Chiles et al, Phys. Rev. Lett. 128 (2022)

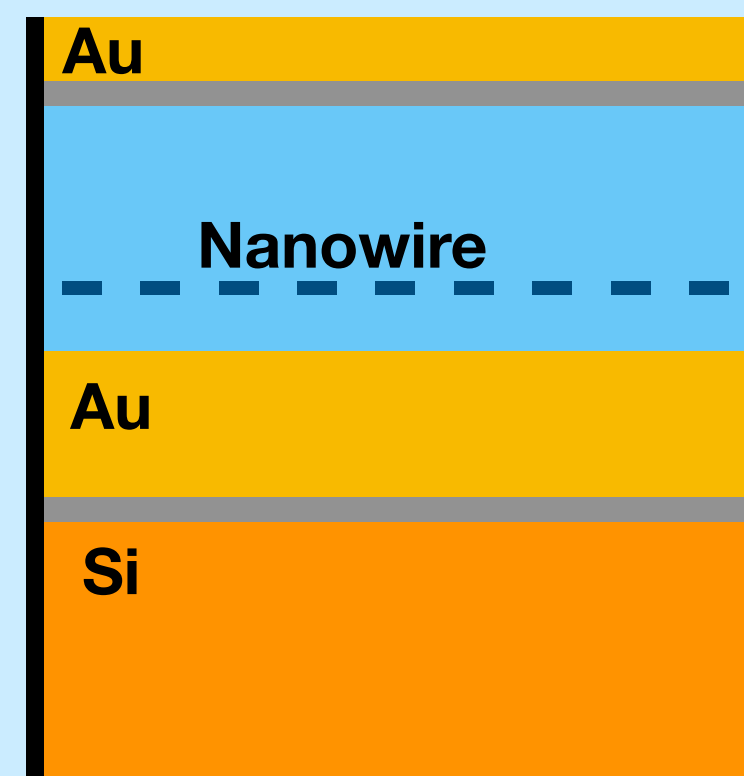
**DCR  $\sim 1/40$ hrs  $\sim 1/(730$ years  $\cdot$  micron<sup>2</sup>)**

# LAMPOST 2: Dielectric Haloscope with an SNSPI



## Background rejection with an SNSPI

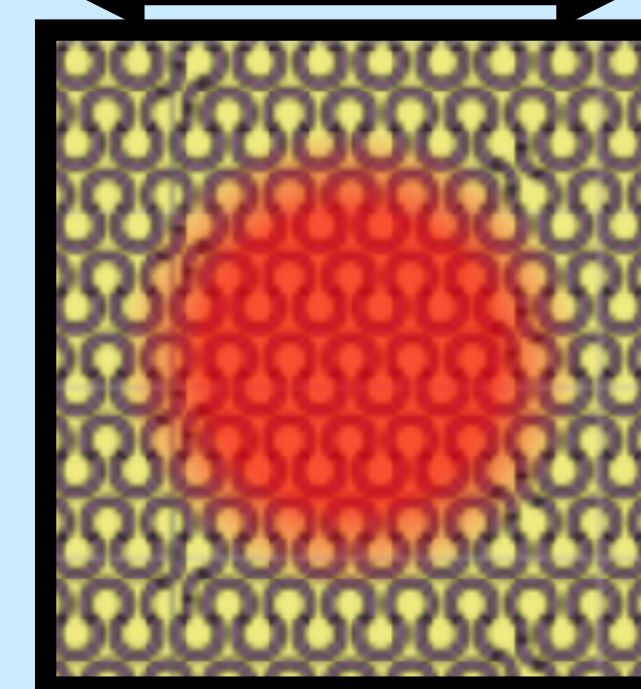
Optical cavity:  
spectral resolution



Impedance-matched output:  
amplitude resolution

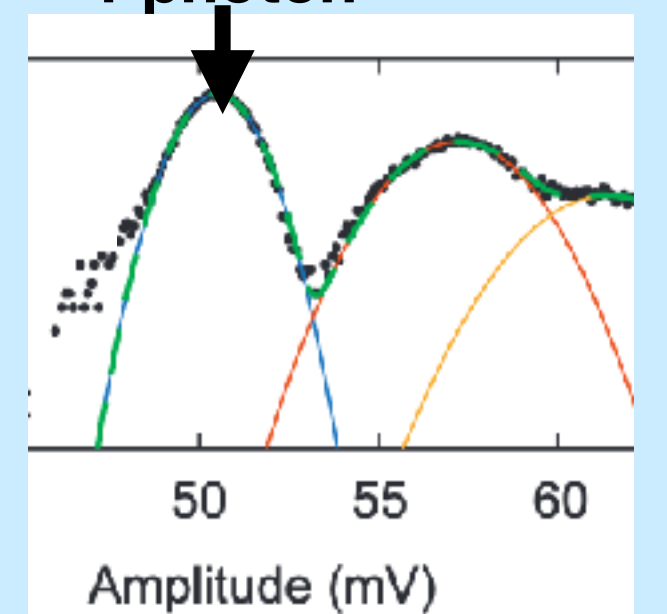
spatial resolution

10µm



amplitude resolution

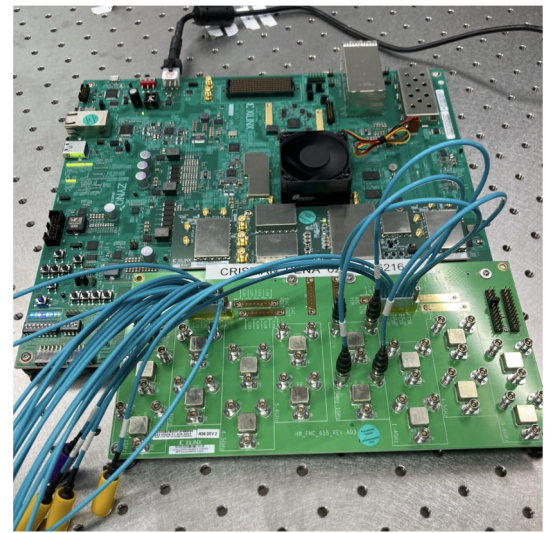
1 photon





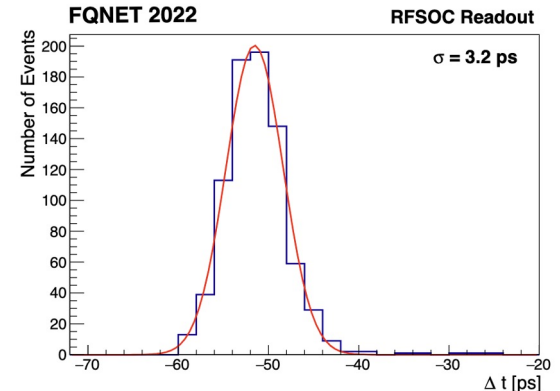
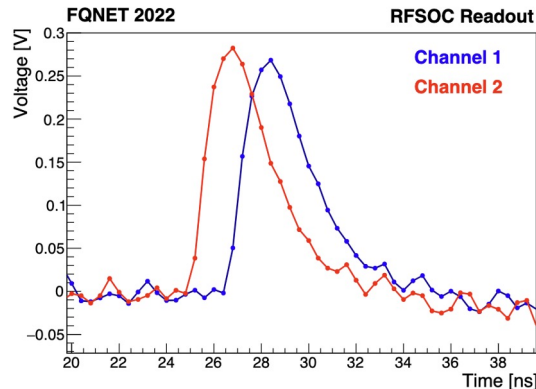
# QICK SNSPD Readout Demonstrator

- SNSPDs are ideal low-threshold photodetectors for axion detection experiments (see talk by C.Pena)
- **Quantum Instrumentation Control Kit** System offers integrated solution to FPGA-based detector control and readout
- First demo of time-domain readout of SNSPDs sampled at 2.5Gsp/s, measured sync time resolution of 3ps

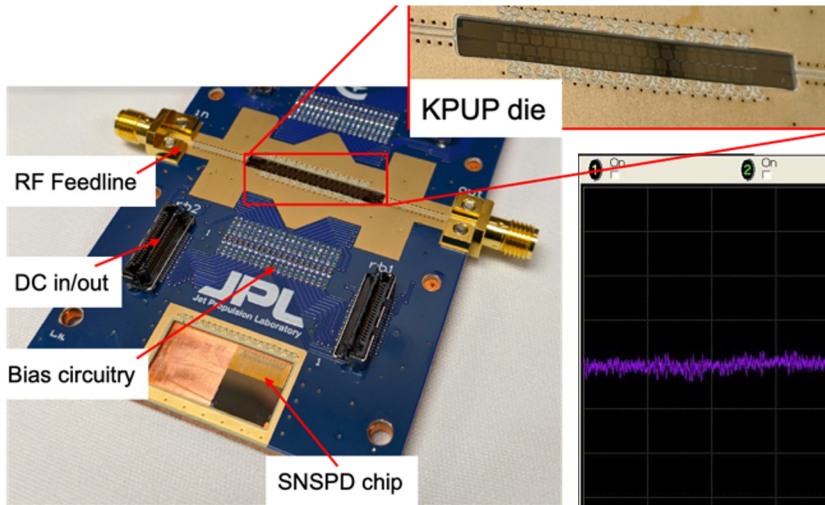


Entangled Photon Pair Source Demonstrator using the Quantum Instrumentation Control Kit System, S. Xie et al, arxiv:2304.01190, submitted to IEEE JQE.

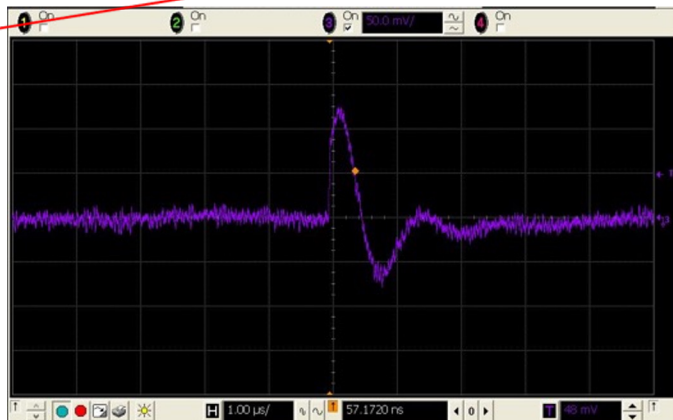
- **Paving the way towards fast and scalable readout !**



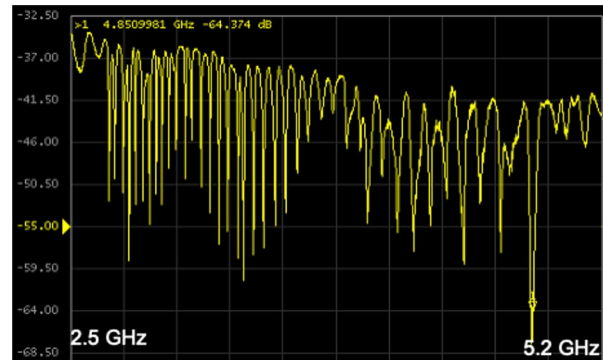
# Frequency-Domain SNSPD Readout with QICK



SNSPD array interfaced to resonator chip

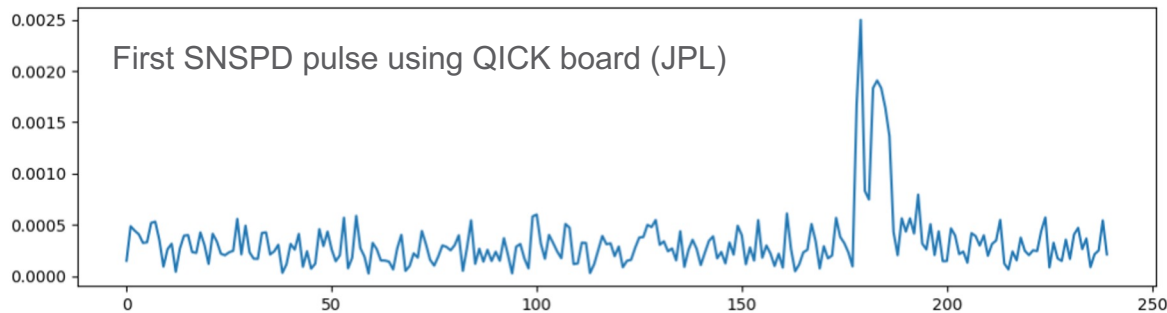


Frequency domain SNSPD readout with analog demux



40 resonators on one feedline

- Frequency-domain multiplexing is the best approach to scale to large arrays of long-wavelength ( $>10 \mu\text{m}$ ) SNSPDs due to high current sensitivity
- Fermilab / JPL / Caltech / ASU collaboration is has recently measured first SNSPD pulses using QICK hardware and kinetic inductance parametric upconverter (KPUP)

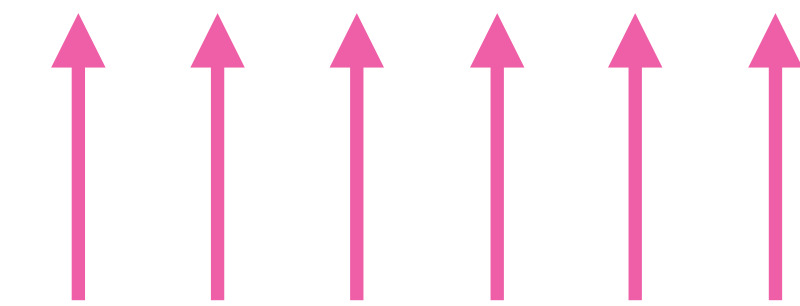
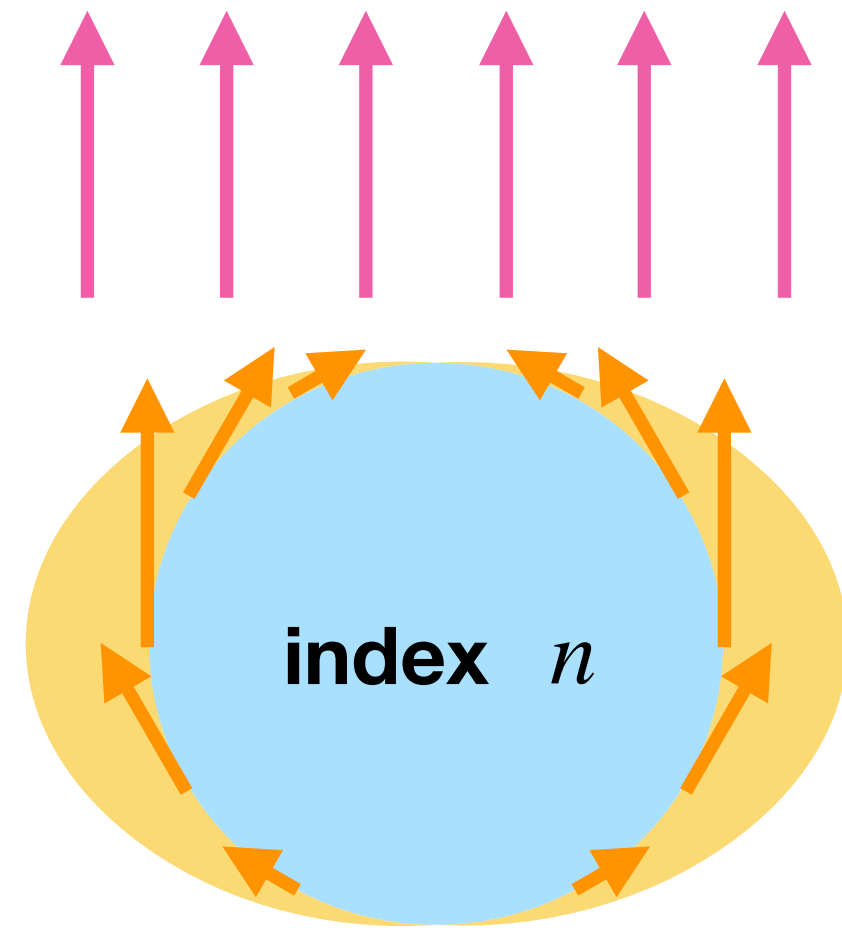


Sasha Sypkens



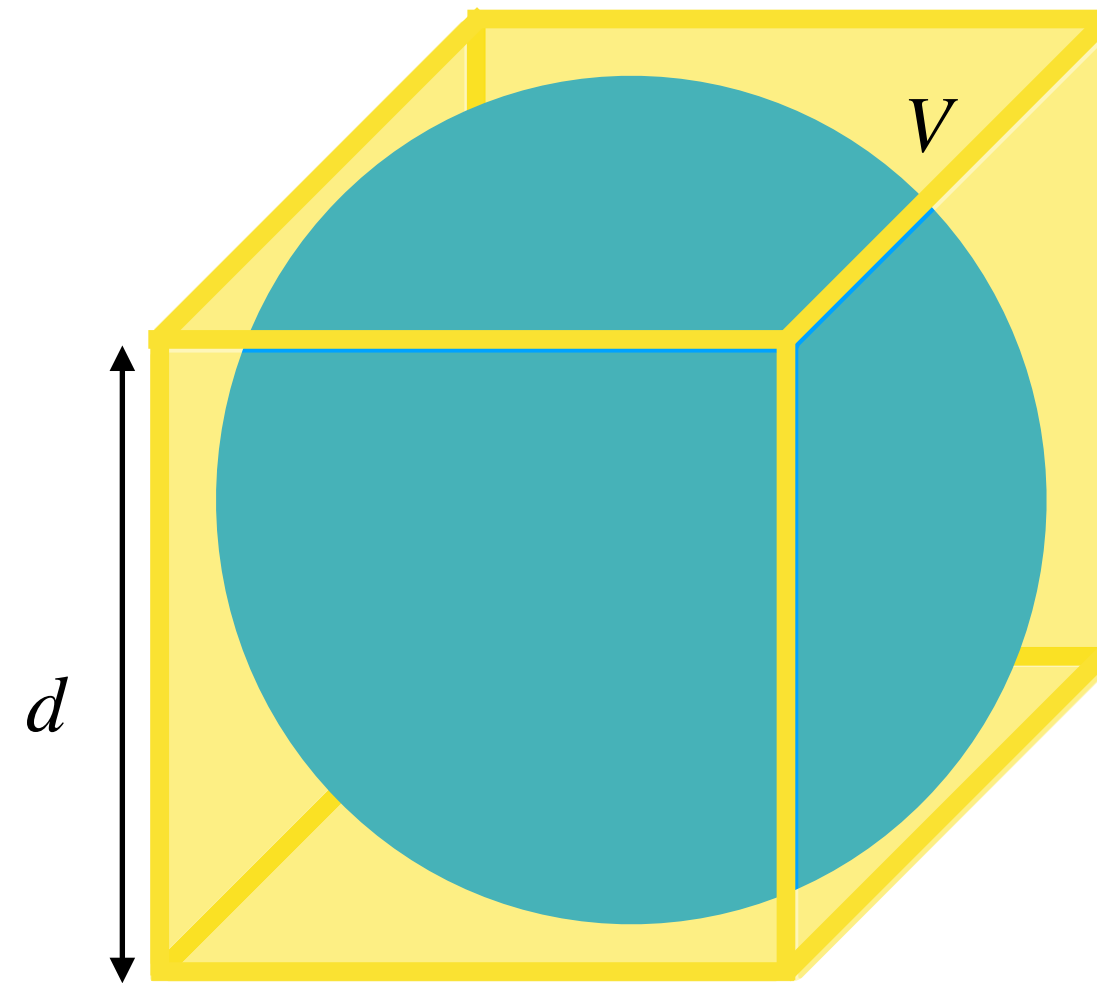
# Dark Photon Absorption by a 3D Microstructure

Stewart Koppell<sup>1</sup>, Dip Joti Paul<sup>1</sup>, Junwu Huang<sup>2</sup>, Karl Berggren<sup>1</sup> (1. Massachusetts Institute of Technology, 2. Perimeter Institute for Theoretical Physics)



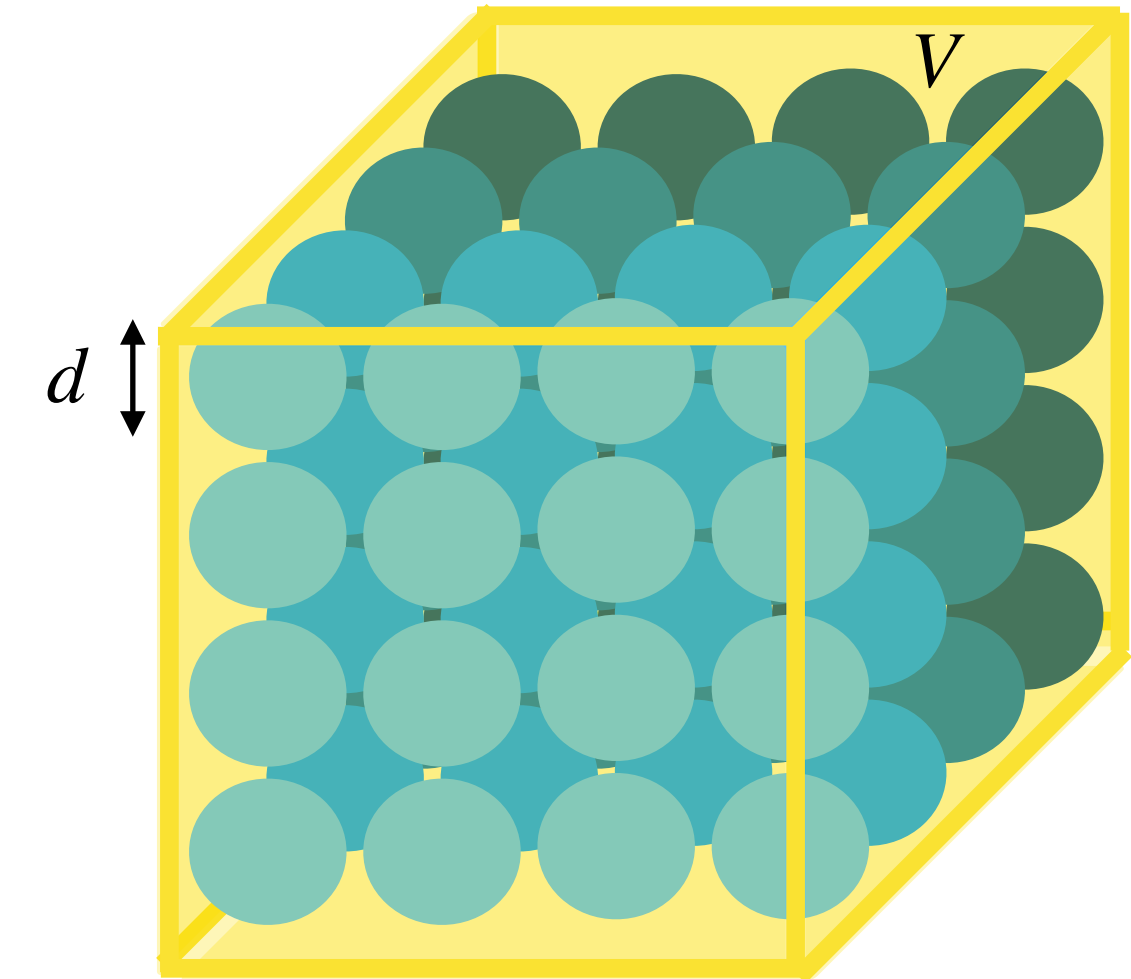
$$E_{\parallel} = E^a \cos(\theta)$$

$$P = \mathbf{SA} E_{\parallel}^2 \left(1 + \frac{1}{n}\right) \left(1 - \frac{1}{n}\right)^2$$



Above diffraction limit:  $d \gg \lambda/2\pi$

$$\frac{P}{V} \equiv \rho \propto \frac{\mathbf{SA}}{V} \propto \frac{1}{d}$$



Below diffraction limit:  $d \ll \lambda/2\pi$

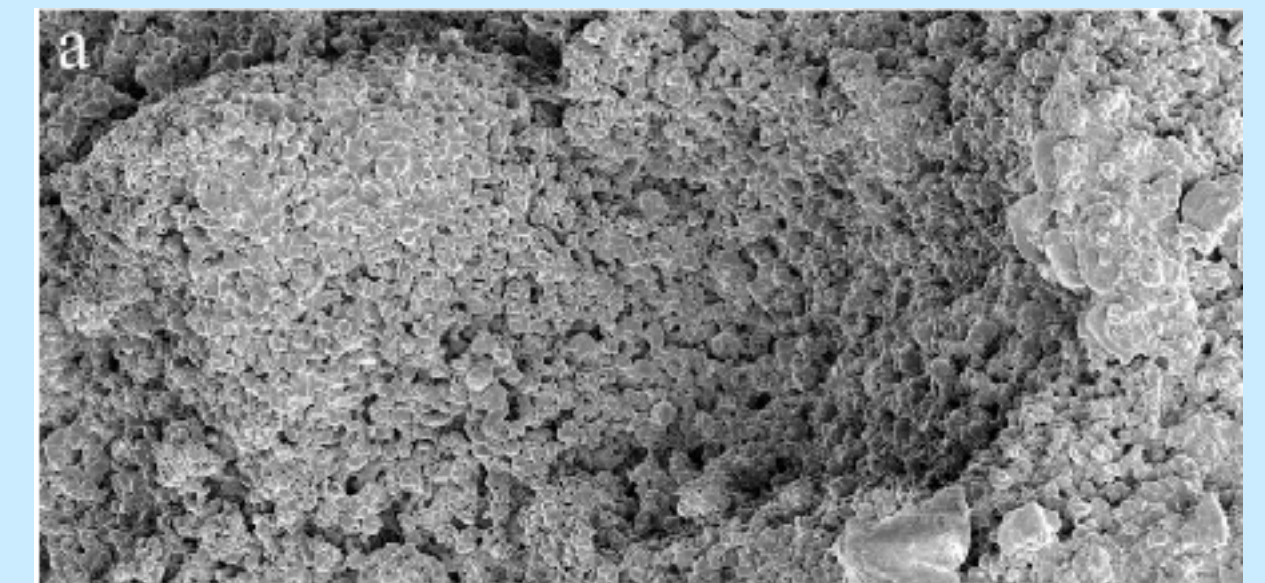
$$\rho \propto \frac{d^3}{\lambda^4}$$

## Porous Alumina

Absorption  $\alpha \sim 0.1 m^{-1}$

Scattering  $\beta \sim 0.2 cm^{-1}$

Diffusion  $\gamma \sim \sqrt{\alpha\beta}$



Near-infrared optical properties of a porous alumina ceramics produced by hydrothermal oxidation of aluminum

Aleksey V. Lisitsyn<sup>3</sup>, Leonid A. Dombrovsky<sup>3,\*</sup>, Vladimir Ya. Mendeleyev<sup>3</sup>, Anatoly V. Grigorenko<sup>3</sup>, Mikhail S. Vlaskin<sup>4</sup>, Andrey Z. Zhuk<sup>4</sup>

<sup>3</sup>Joint Institute for High Temperatures, NCHIME, Krasnokazarmennaya 17A, Moscow 111116, Russia  
<sup>4</sup>Joint Institute for High Temperatures, Izhor'skaya 13-2, Moscow 525003, Russia

# Collecting photons

0.7 $\mu$ m Dielectric Powder

Diffusion Length  $1/\gamma = 3mm?$

Specific Surface area  $3m^2/cm^3$

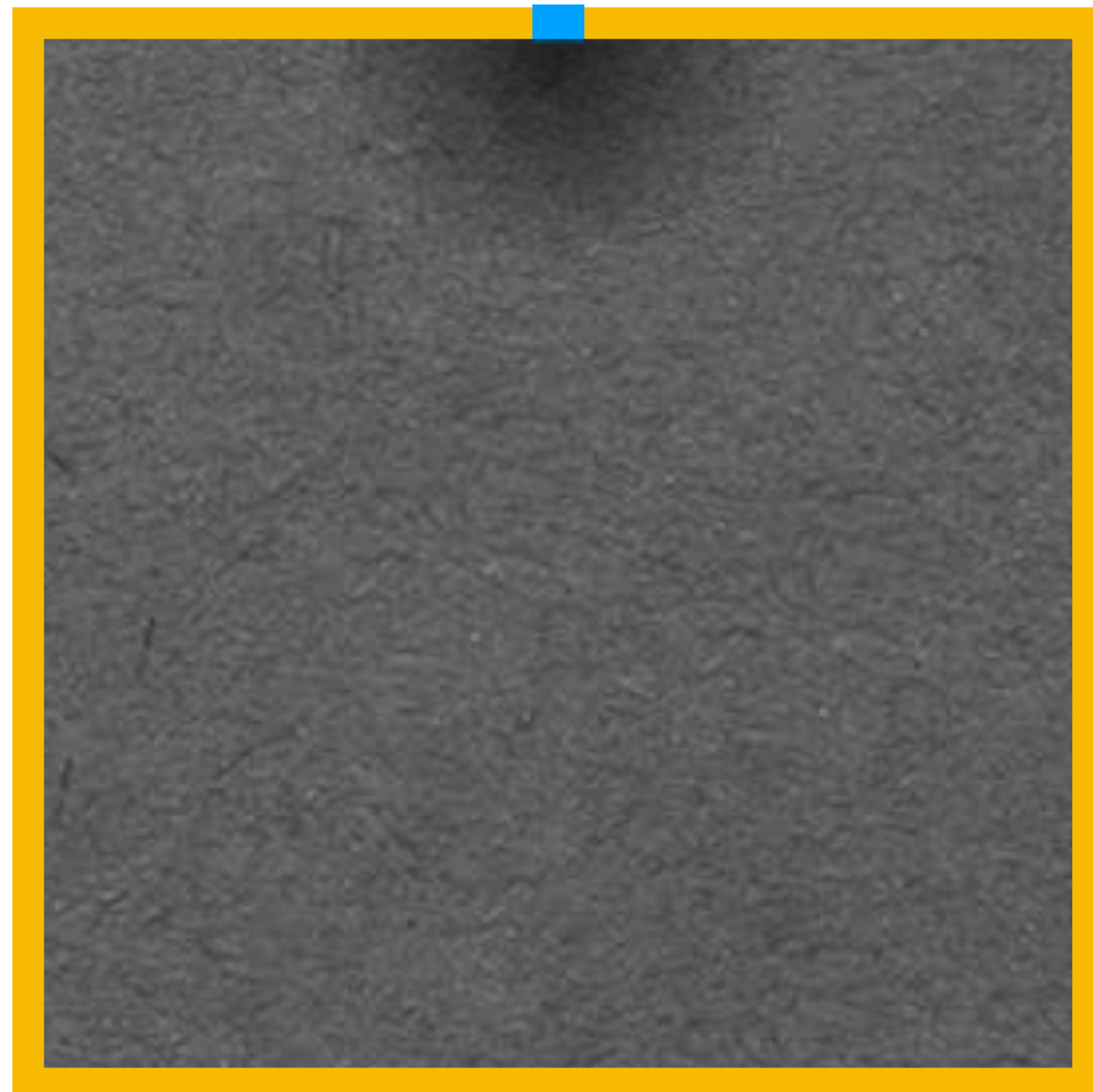
Effective volume:

$$1/\gamma^3 = (3mm)^3$$

Effective area:

$$0.1m^2$$

SNSPD



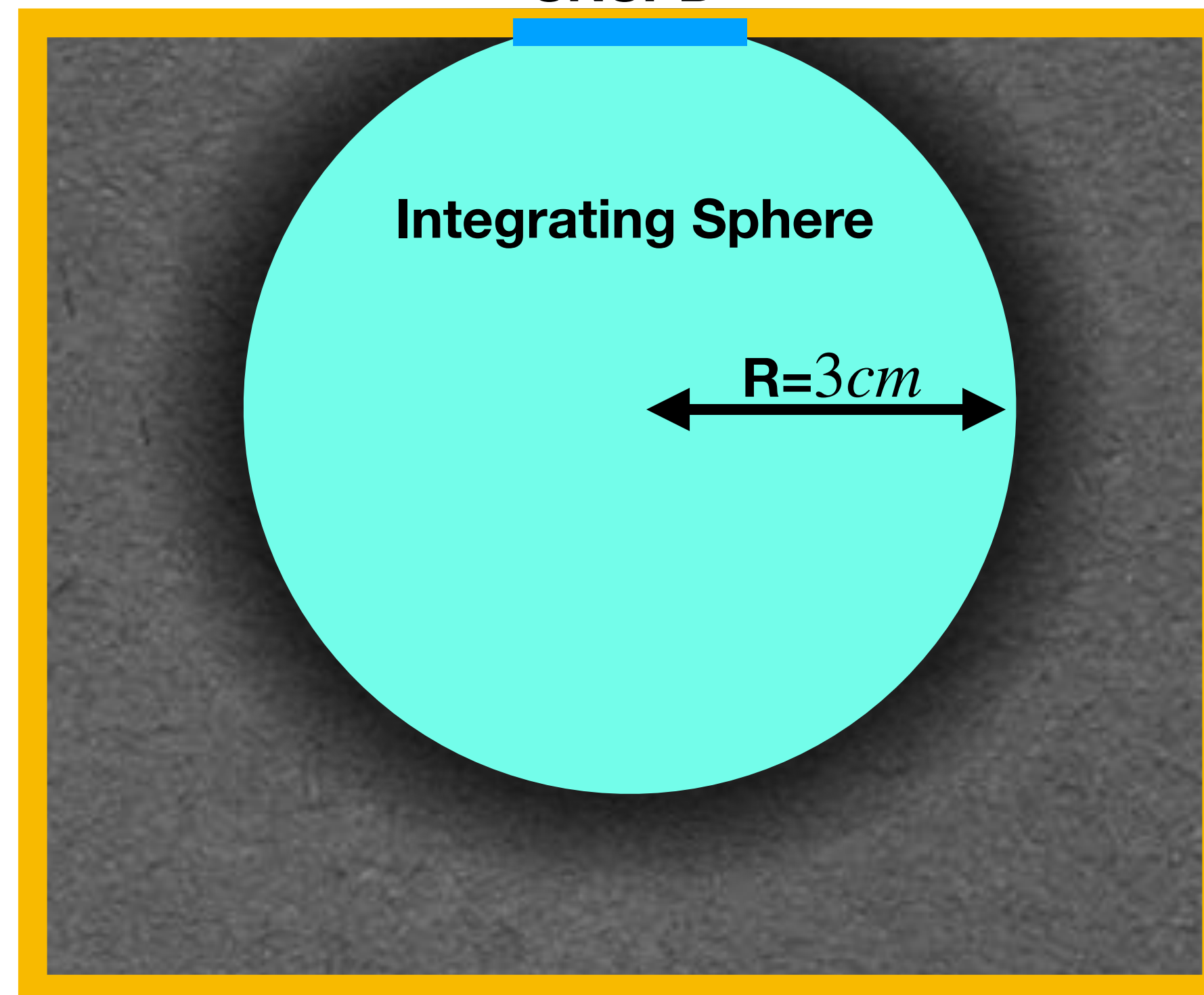
Effective volume:

$$4\pi R^2/\gamma = 34cm^3$$

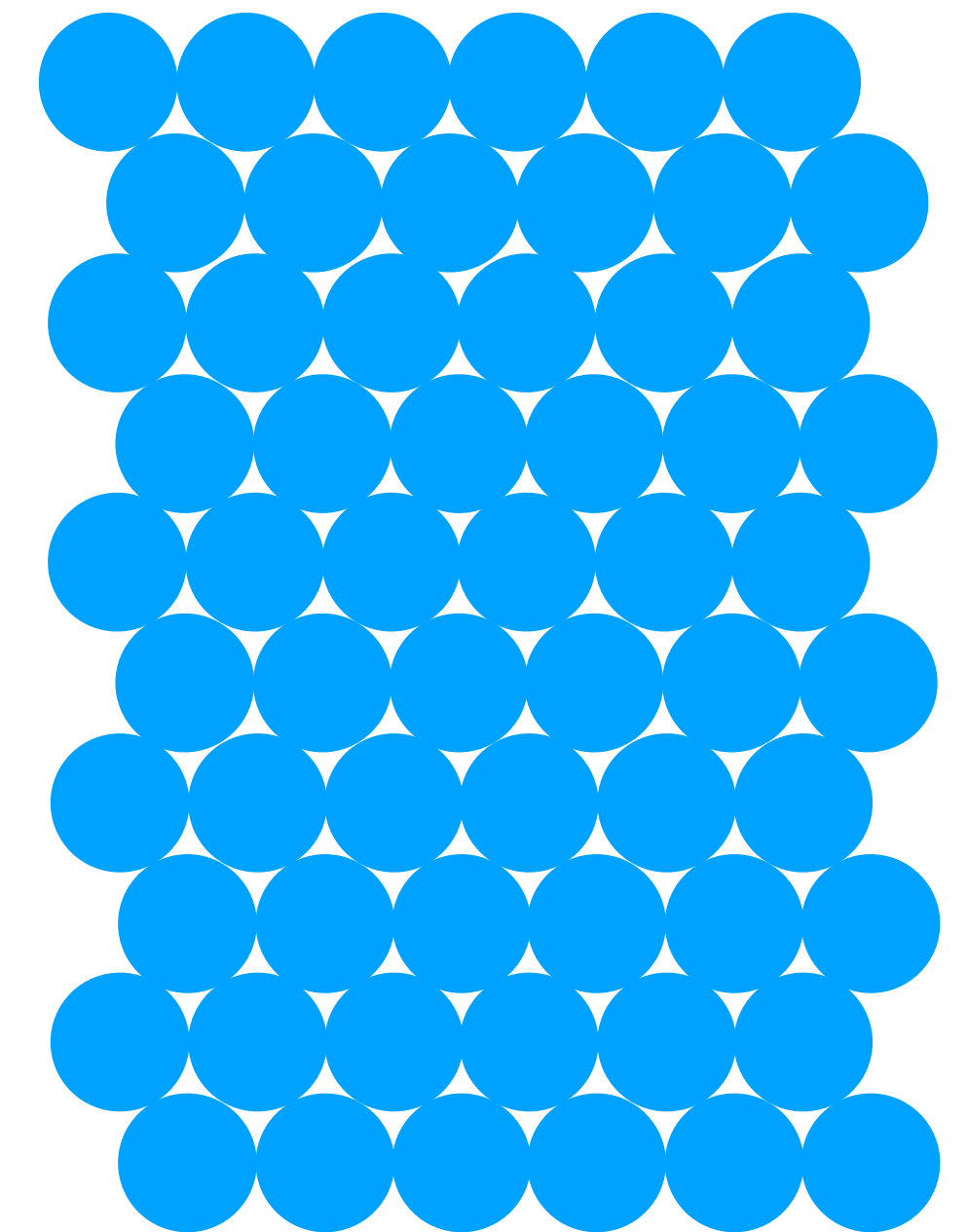
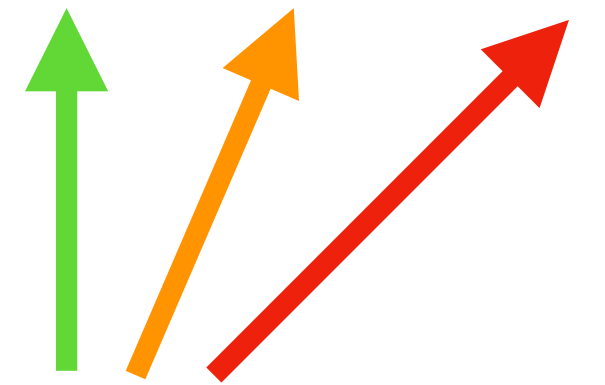
Effective area:

$$100m^2$$

SNSPD



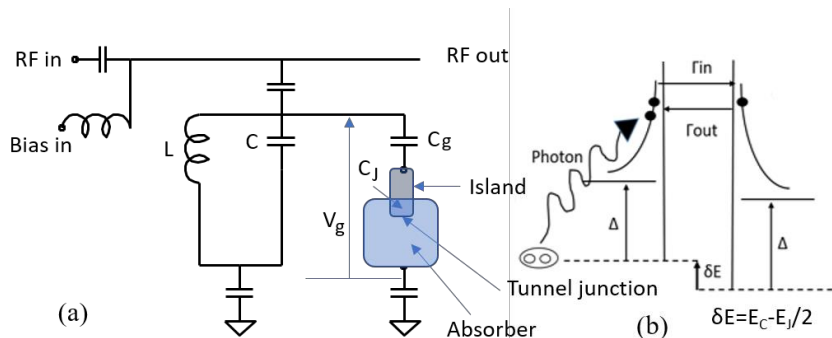
Photonic Crystal



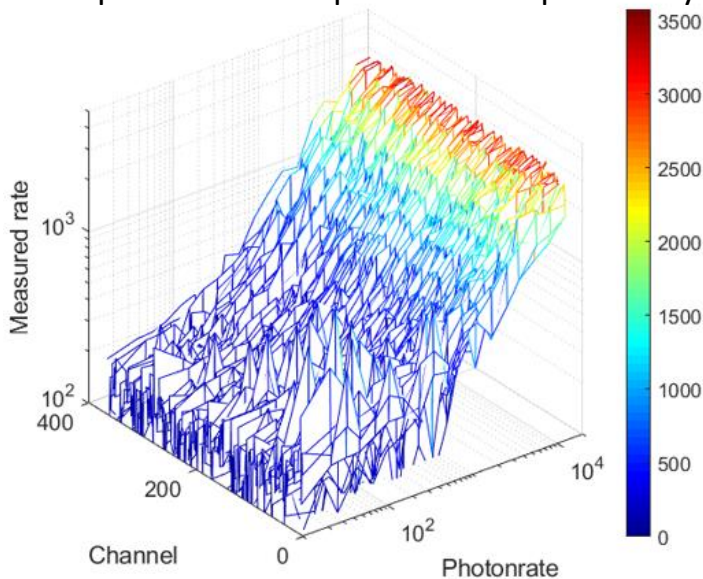




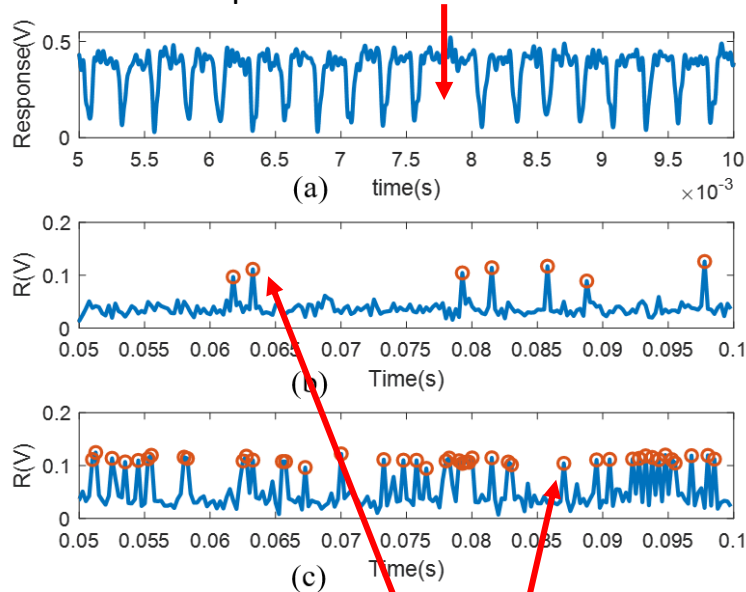
QCDs exploit the extreme susceptibility of a charge qubit to external influences to make an extremely sensitive far IR detector



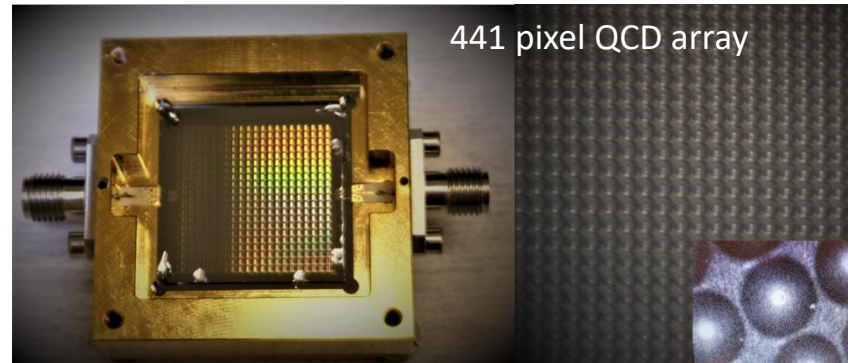
Measured photon rate as function  
Of optical radiation power in 441 pixel array

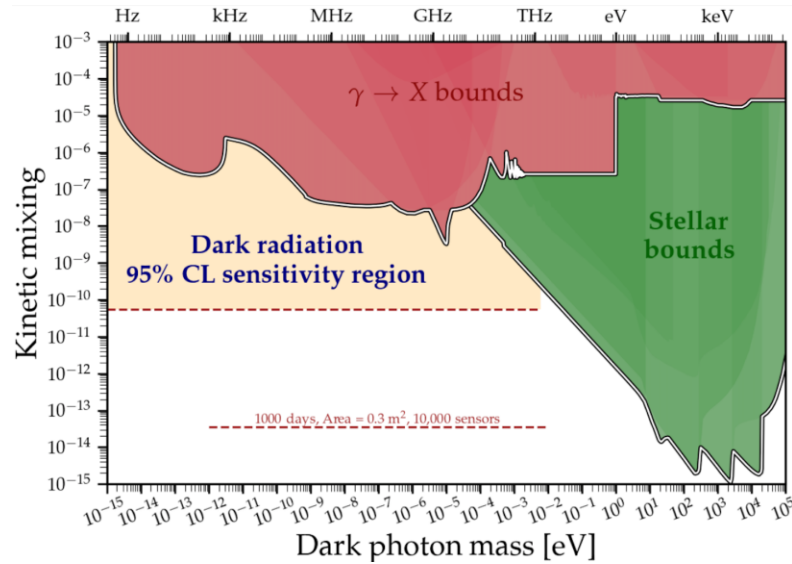


Missing quantum capacitance peak  
indicates photon detection event

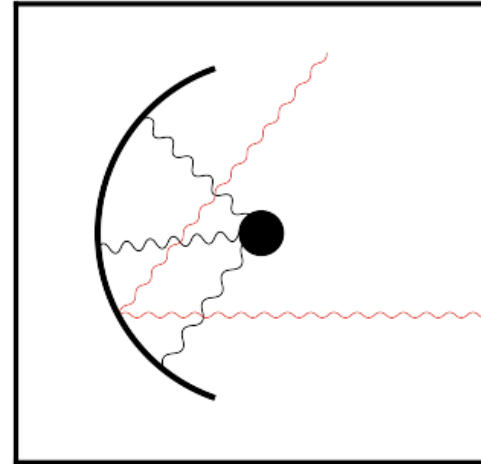


Photon events for cold and hot black body

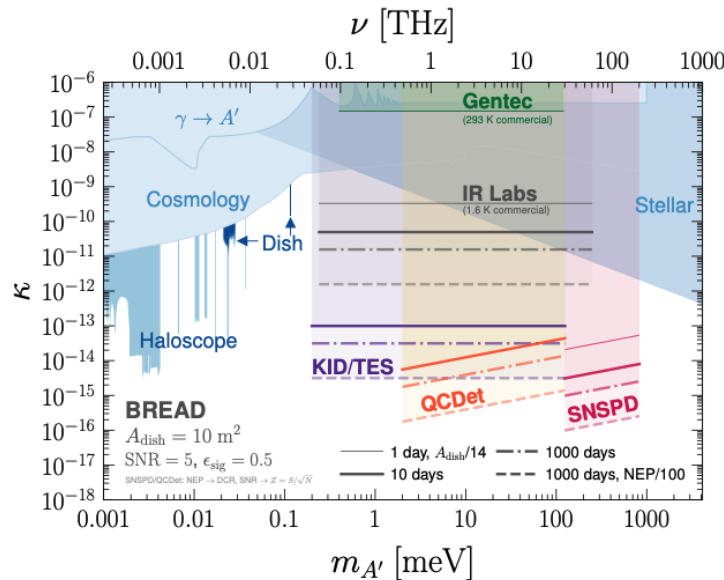




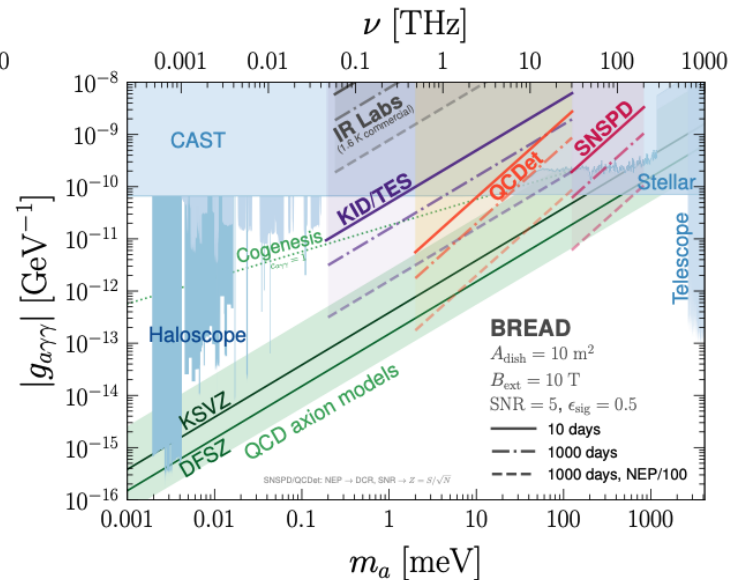
Dark photons convert to visible photons upon scattering on impedance mismatched materials



Dark Photon sensitivity



Axion sensitivity



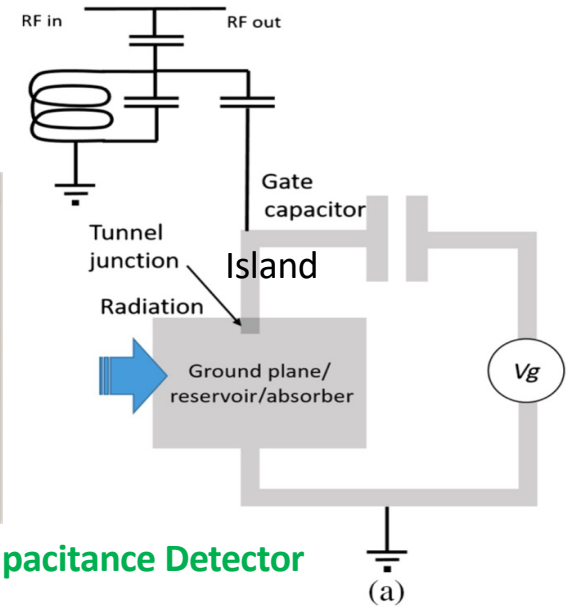
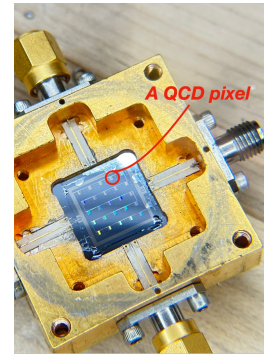
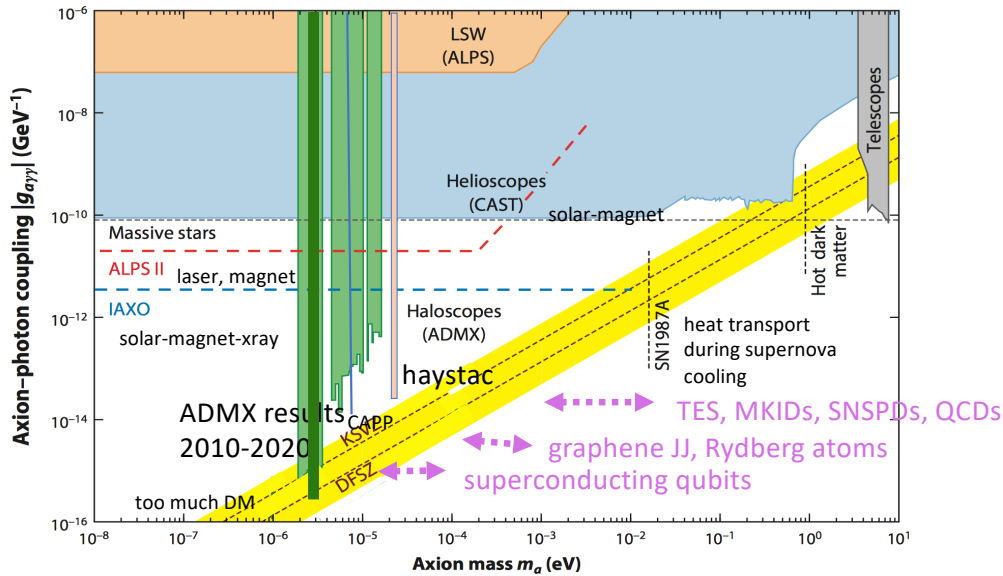
# Dark Matter detection with Quantum Capacitance Detectors

Rakshya Khatiwada

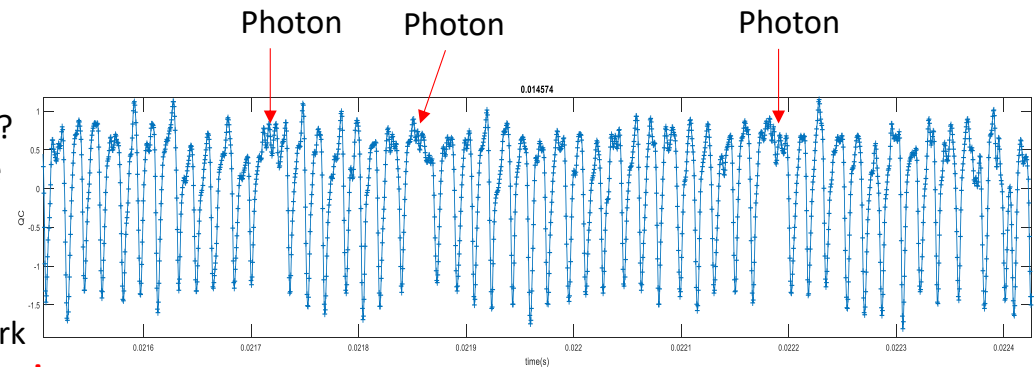
Illinois Institute of Technology

Fermilab

# Motivation



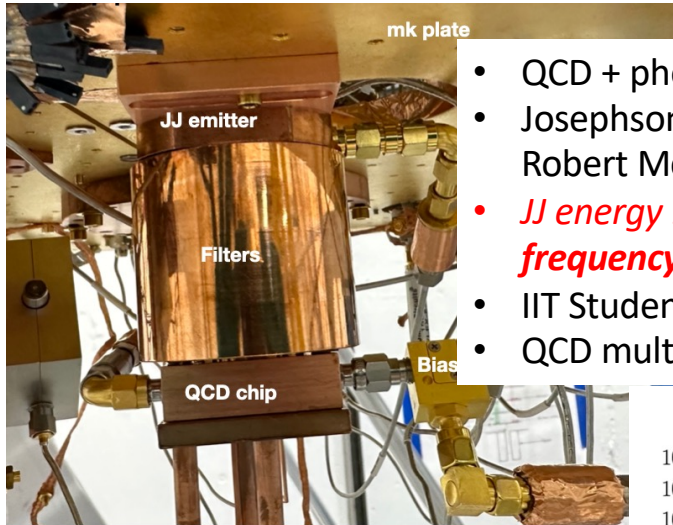
Quantum Capacitance Detector



- How do we bridge the gap between ultralight to GeV Dark Matter?
- Most successful Haloscope experiments limited by quantum noise
- Lack of detection technology limiting the progress
- Qubits: great potential for low energy radiation detection
- Proven single photon counting with Qubits  $\sim$  few GHz
- Can we expand it further to probe even higher mass, wave like Dark Matter? **JPL (Echternack), Fermilab (Chou, Khatiwada, Yu), Wisconsin (McDermott, Harrison) collaboration**



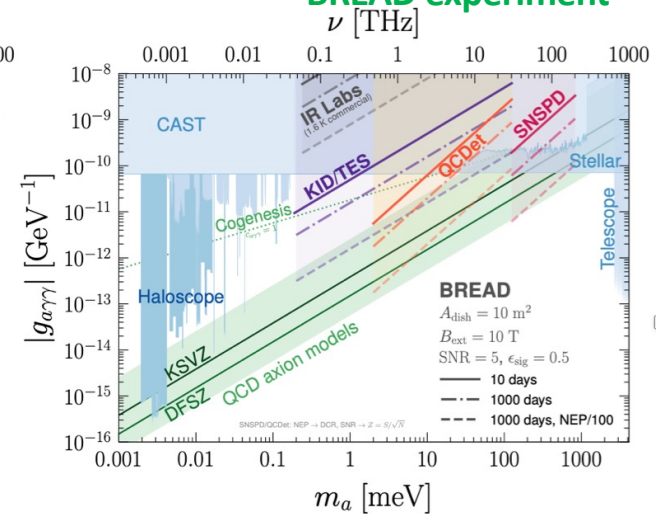
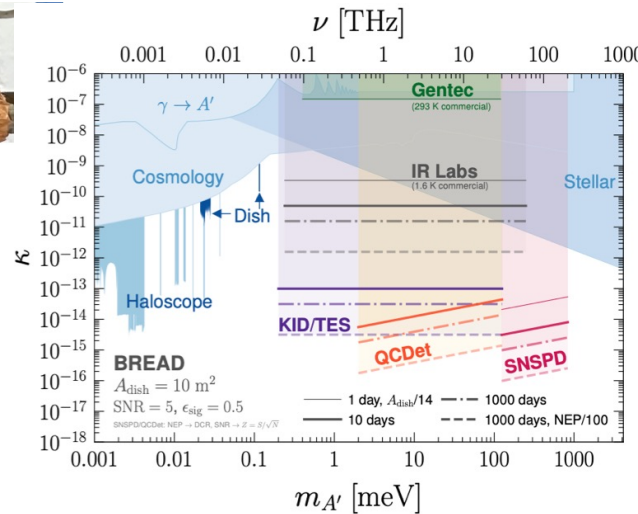
# QCD single photon counter for Dark Matter



- QCD + photon emitter source integration in progress at FNAL/IIT
- Josephson Junction based photon emitter source for fab. by Robert McDermott's group
- ***JJ energy  $E_j = 2eV_{DC} \Rightarrow f = 2eV_{DC}/h$  produces photons at this frequency***
- IIT Student supported by FNAL LDRD
- QCD multiple array readout and characterization in progress

**Projected sensitivity for Broadband Antenna BREAD experiment**

- QCD: great potential for THz photon detection
- **Map out: Energy deposited from DM photon in Al  $\rightarrow$  Broken cooper pairs  $\rightarrow$  quasiparticles tunneling rate higher  $\rightarrow$  qubit charge parity 'signal' readout**



(BREAD Collaboration)  
Phys. Rev. Lett. **128**, 131801 – Published 28 March 2022

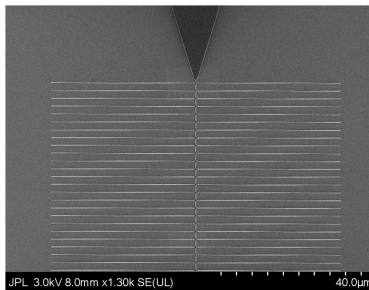
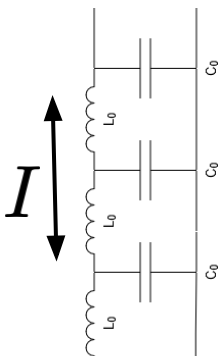
R. Basu Thakur<sup>a,b</sup>, P. K. Day<sup>a</sup>, D. Cunnane<sup>a</sup>, K. Ramanathan<sup>b</sup>, J. Zumidzinas<sup>b</sup>

## Nonlinear Kinetic Inductance (NLKI)

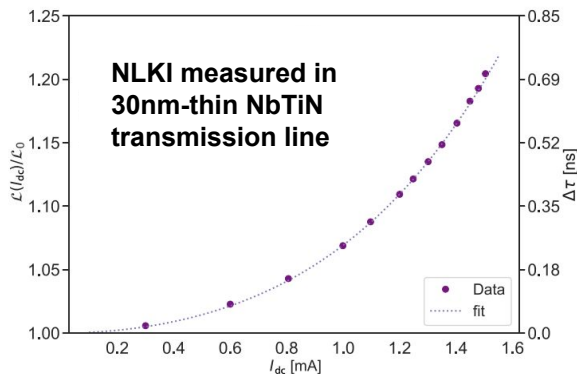
Kinetic inductance in thin-film superconductors, NbTiN, MgB2 etc. varies non-linearly with supercurrent

$$\mathcal{L}(I) \approx \mathcal{L}_0 \left( 1 + \left( \frac{I}{I_*} \right)^2 + \left( \frac{I}{I_*} \right)^4 + \dots \right)$$

## NLKI in transmission lines



**NbTiN transmission line  
 250nm-wide central conductor**



NLKI measured with resonators & transmission lines in 1-100 GHz

THz devices under R&D

*Devices fabricated with phase velocities  $\sim 0.1\%$  &  $\sim 10\%$  NLKI*

- In dispersion-engineered transmission lines, NLKI can enable frequency mixing processes (akin to nonlinear crystals)

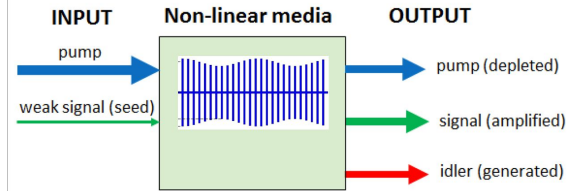
$$\frac{1}{\mathcal{L}_0 C} \frac{\partial^2 I}{\partial z^2} = \frac{\partial^2 I}{\partial t^2} + \frac{1}{I_*^2} \frac{\partial}{\partial t} \left[ I^2 \frac{\partial I}{\partial t} \right]$$

*Parametric amplification @ SQL & frequency conversion demonstrated*

- In non-dispersive transmission lines, NLKI can enable field-tunable delay of the RF signal by applying DC current

$$\Delta\tau(I) = \sqrt{\mathcal{L}(I)C} \quad \sim 1 \text{ ns achieved, targeting micro-seconds}$$

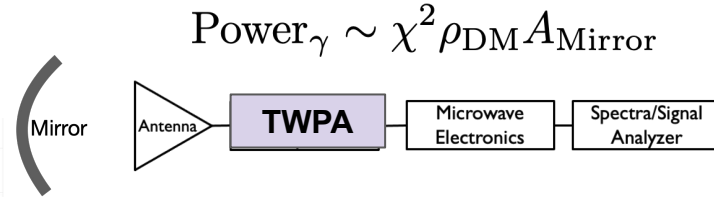
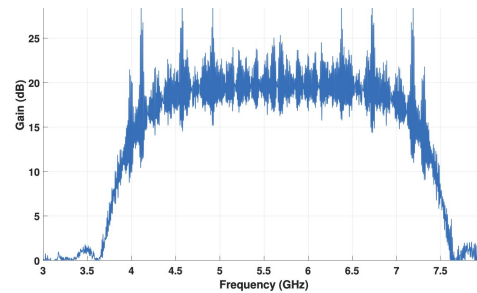
## Traveling Wave Parametric Amplifiers (TWPA)



Transmission line with phase-matching & dispersion engineering, thus amplification

### TWPAs used for dark matter searches

~17 dB gain = 3 x standard quantum limit noise, actively pushing to SQL

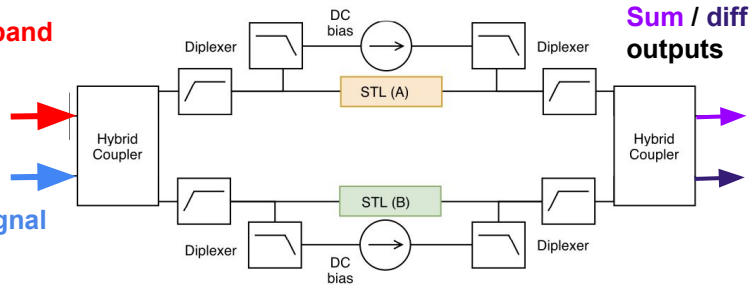


$$\text{Power}_\gamma \sim \chi^2 \rho_{\text{DM}} A_{\text{Mirror}}$$

**QUANTUM LIMITED PHOTONS IN THE DARK**  
 Experiment set leading limits on hidden photon dark matter in 20-30 ueV ...  
 planned expansion to 500 ueV

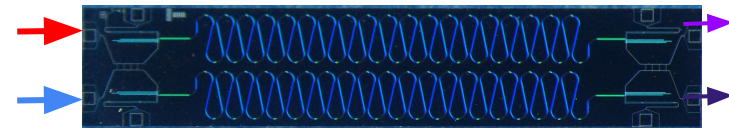
## Superconducting On-chip Fourier Transform Spectrometer (SOFTS)

Broadband input



4-port Mach Zehnder form with two transmission lines  
 SOFTS does calibrated lossless broadband spectroscopy

Tunable *electromagnetic* delay (no mirrors) via NLKI  
 Ultracompact as  $v_{\text{phase}} \sim 0.1\%c$ , broadband and lossless  
 Integrable with TWPAs/ QCDs/ KIDs for photonic circuits

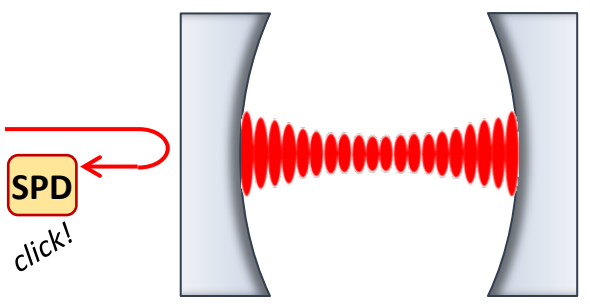


← 6mm prototype device for 20-40 GHz signals with 1GHz resolution →

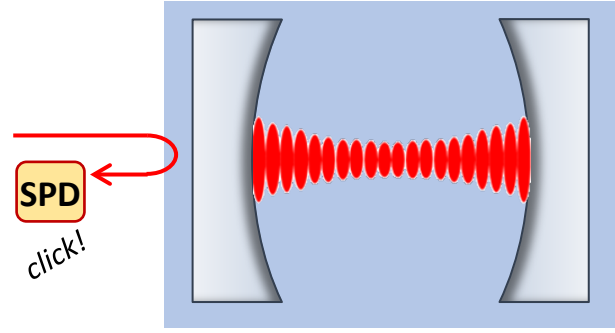
Line intensity, galaxy-cluster spectro-imaging, CMB spectral distortions, dark-sector searches, quantum-optics experiments

• Almost everything you can say about EM cavities can also be said about acoustic cavities. *Quality factor, volume, geometry, readout, etc.*

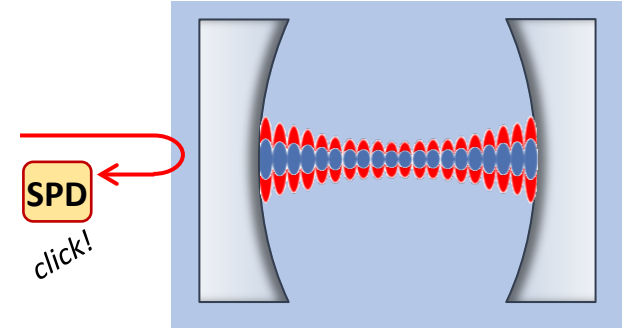
• One nice recipe for this:  
Take any electromagnetic cavity...



...and fill it with superfluid helium.  
The EM modes are unchanged...



...and the superfluid has acoustic modes!



• The electromagnetic modes can be used to readout and control the acoustic modes – and vice versa.

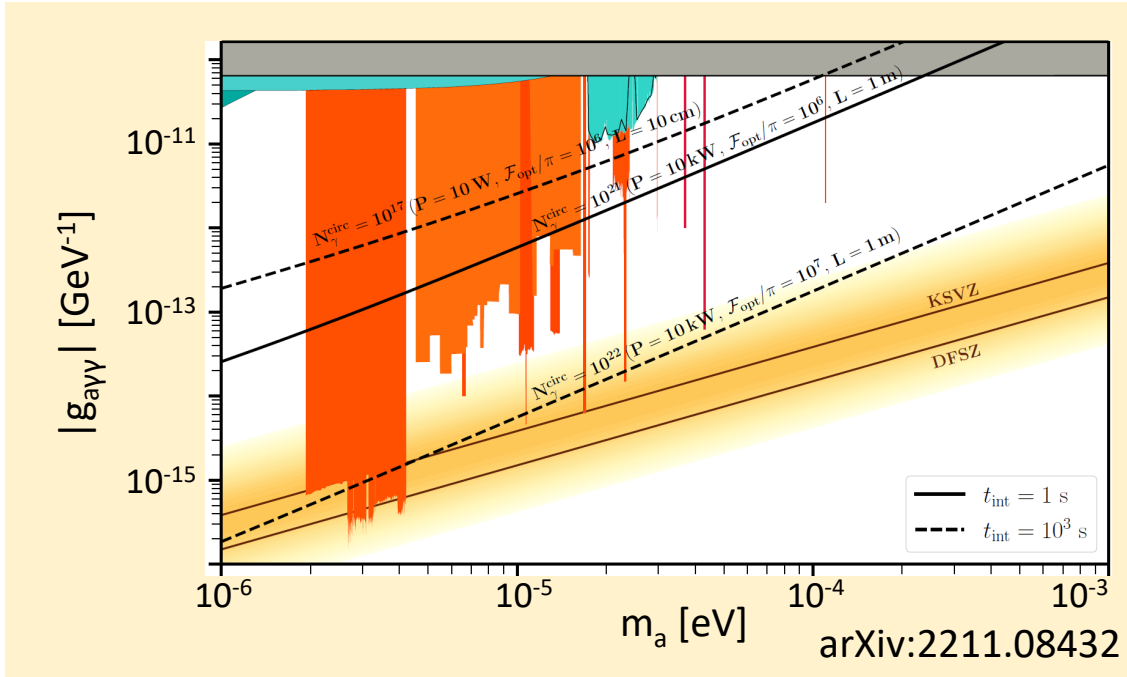
• A nice consequence of this interaction:  
detection of a single photon can serve as detection of a single phonon  
 $\approx 1 \text{ eV}$   $\approx (v_{\text{He}}/c) \times 1 \text{ eV} \approx 1 \mu\text{eV}$

• An important fact: Only phonons in a single mode of the cavity are detected

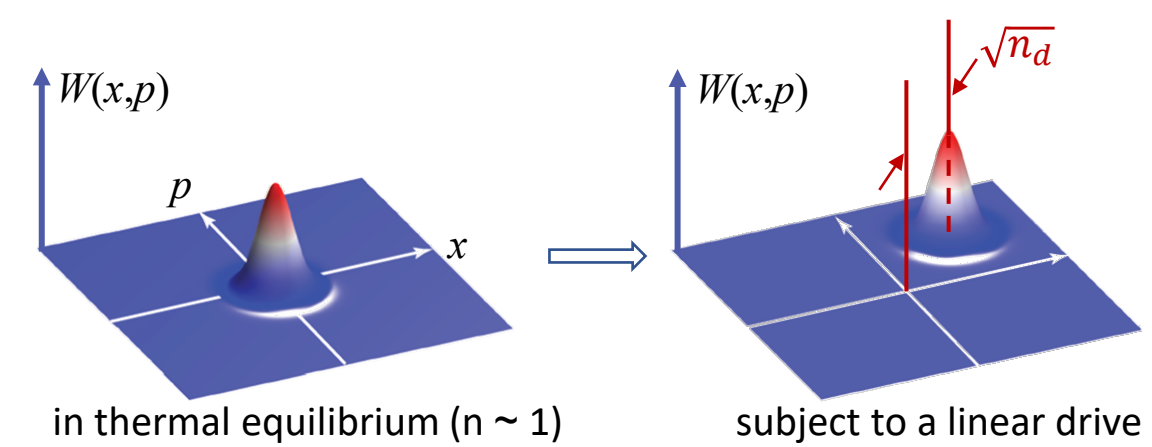
**Bad for typical DM searches** →

*Ongoing collaboration with Kathryn Zurek, Clara Murgui, Yikun Wang. Stay tuned!*

**Good for typical cavity quantum dynamics** (next slide, please!)



Testing quantum linearity with single-mode optomechanics



If the oscillator is linear, the state is simply displaced.

Phonon statistics are well-known:  $g^{(2)}(0) = 1 + \frac{2n_{th}}{n_d}$

For weak nonlinearity, there will be deviations for large  $n_d$ .

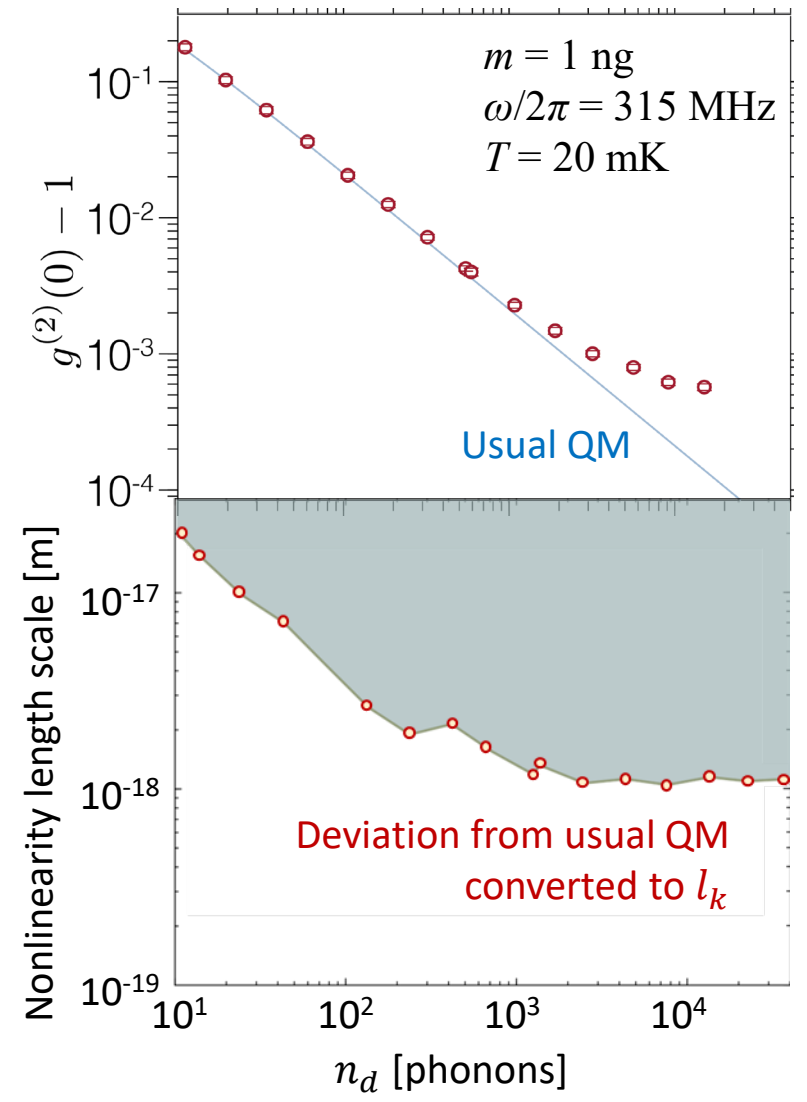
**A proposal\* from quantum gravity phenomenology:**

- To reconcile LI & discreteness at Planck scale, the box operator must become nonlinear.
- In non-relativistic limit, Schrödinger operator  $\mathcal{S}$  inherits this nonlinearity:

$$\mathcal{S} \longrightarrow \mathcal{S} + \sum_{n=2}^{\infty} b_n \left( \frac{-2m}{\hbar^2} \right)^{n-1} \underbrace{l_k^{2n-2}}_{\text{empirical length scale, not necessarily Planck}} \mathcal{S}^n$$

- This would change the phonon statistics:  $g^{(2)}(0) = \underbrace{1 + \frac{2n_{th}}{n_d}}_{\text{Usual QM}} + \underbrace{\frac{18n_d m \omega}{\hbar} \left(1 + \frac{n_{th}}{2}\right) b_2 l_k^2}_{\text{Nonlinearity}}$

Measured phonon statistics:



\*Belenchia, et al., PRD 95, 026012 (2017);  
Belenchia, et al., PRL 116, 161303 (2016);



# HEP Interferometry via photon counting

Lee McCuller **Caltech**

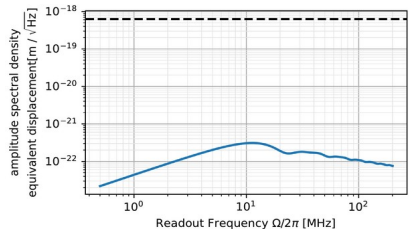
Interferometry & optomechanics:

- Profoundly sensitive to gravitational physics
- macroscopic quantum mechanics
- applications in DM detection

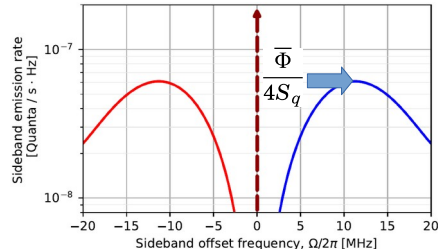
“wavelike” detectors limited by “quantum noise” from homodyne readout or parametric amp readout

Make interferometry more like rare-process HEP  
Signal power statistics → linear in time detection/exclusion  
– no background “counts” from vacuum fluctuations  
**requires suitable search statistics**

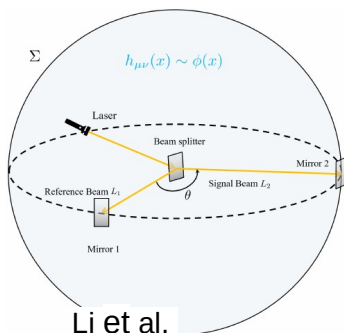
McCuller 2211.04016



Small, wideband stochastic metric fluctuation signal  
Orders of magnitude below shot noise  
~3 months for  $1\sigma$  by incoherent search stat.



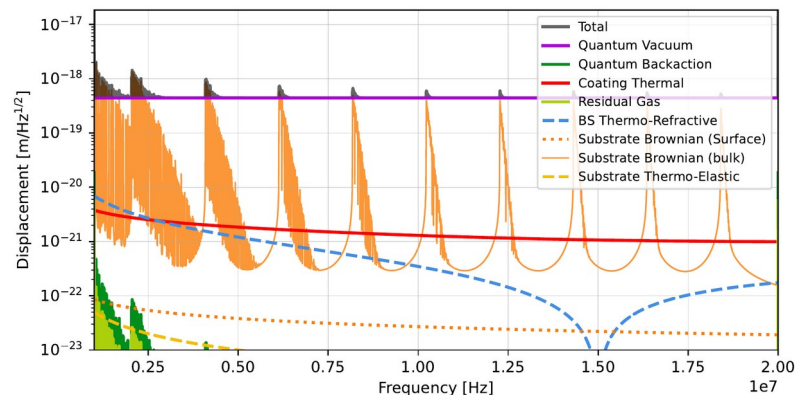
Equivalent sideband photopower  
Emits 1 photon/second  
Vastly accelerated search



Li et al,  
[arXiv:2209.07543](https://arxiv.org/abs/2209.07543)

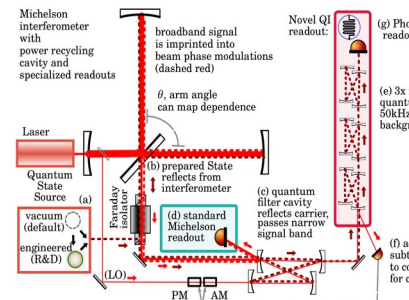
# GQuEST:

Gravity from Quantum Entanglement of Space-Time



Pathfinder for high-contrast photon counting  
10kW interferometer → mHz count rate

High-frequency signal amenable to first attempts at counting but requires new sophistication → unique design elements



Finds an entanglement-entropy basis for gravitation - via Metric fluctuation signature:

*Banks, KZ 2108.04806*  
*E. Verlinde, KZ 1902.08207*  
*E. Verlinde, KZ 1911.02018*



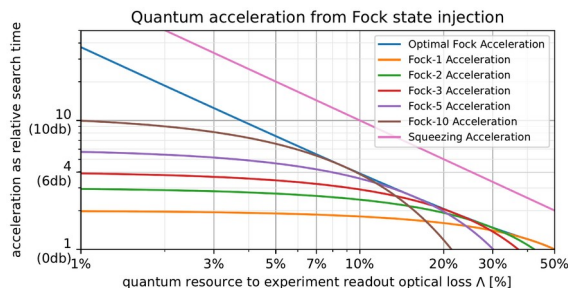
# Novel Quantum Enhancements

LIGO's performance speaks volumes.  
we've learned to saturate the benefits from squeezing.  
Loss-engineering (in optical) will be  
*incremental* and *trade* with higher power

Developing photon counting is a prerequisite to demonstrate *any* non-Gaussian observable at high contrast

Squeezing adds background counts → better quantum enhancements exist.

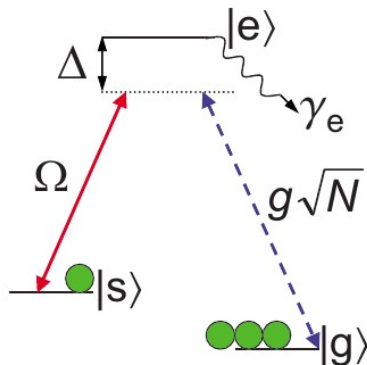
research quantum-enhanced Non-quadrature, non-Gaussian observables.



# Generalized Counting

Employ quantum memories  
implement matched-template search.  
Generalize temporal-mode basis  
beyond Lorentzian signal wavelets

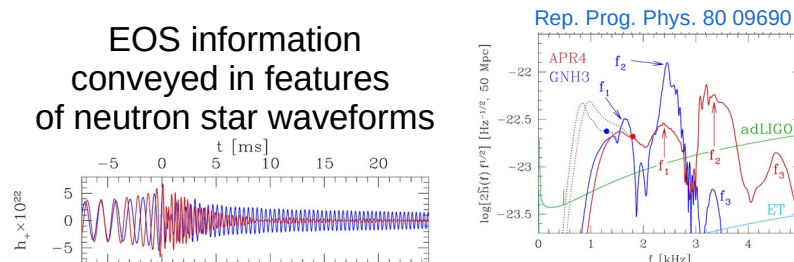
[PhysRevA.76.033804](https://arxiv.org/abs/1603.033804)



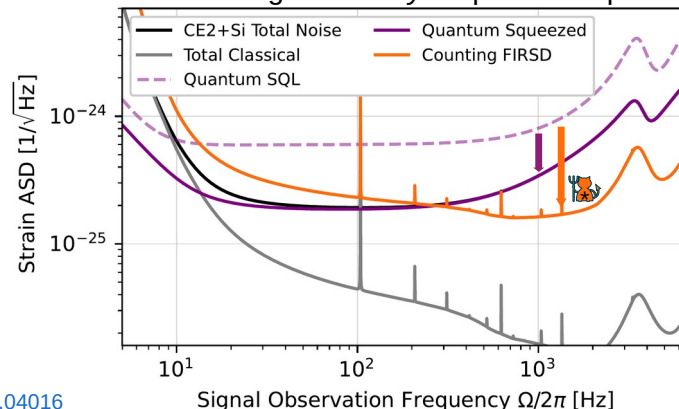
# Science Goal

Binary Neutron Star inspiral “Equation of State”  
strong-field nuclear matter in strong-field GR  
Neutrino energy transport highly influences

EOS information conveyed in features of neutron star waveforms



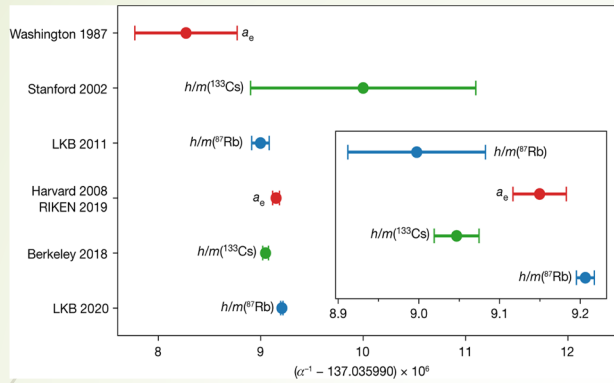
Proposed future detectors, e.g. Cosmic Explorer, to detect 1e5 neutron stars/yr  
Majority of total signal power near shot-noise limit  
Photon counting can vastly outperform squeezing\*



[arXiv:2211.04016](https://arxiv.org/abs/2211.04016)

CE2, 2um Cryo cSI Tech. W/ Photon counting

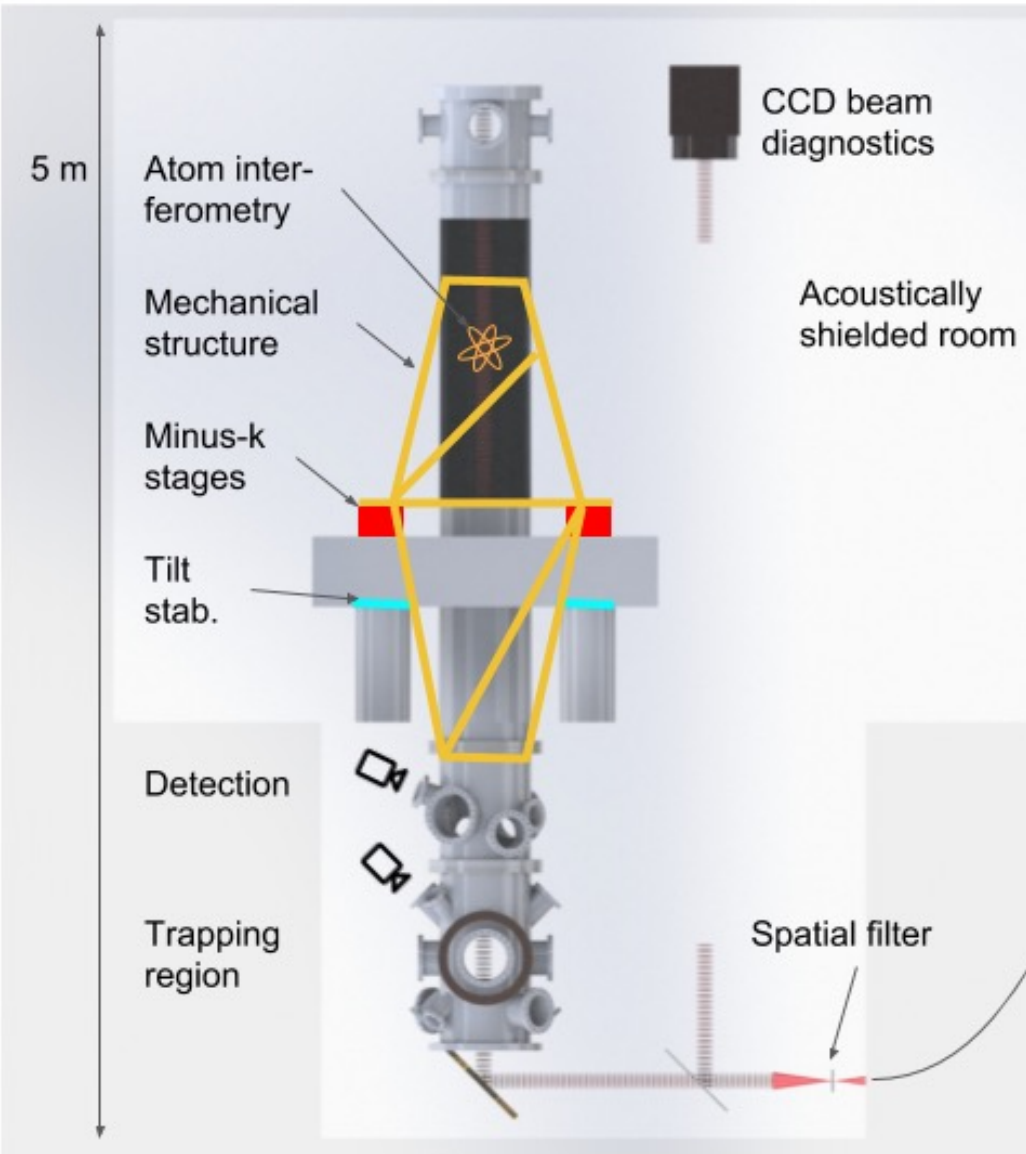
\*statistic/informatic quantum tradeoffs inspire further study



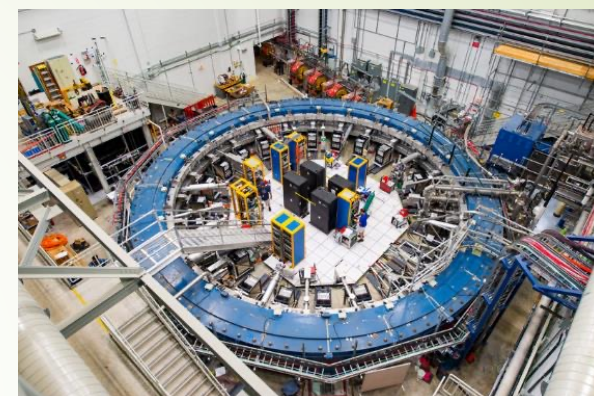
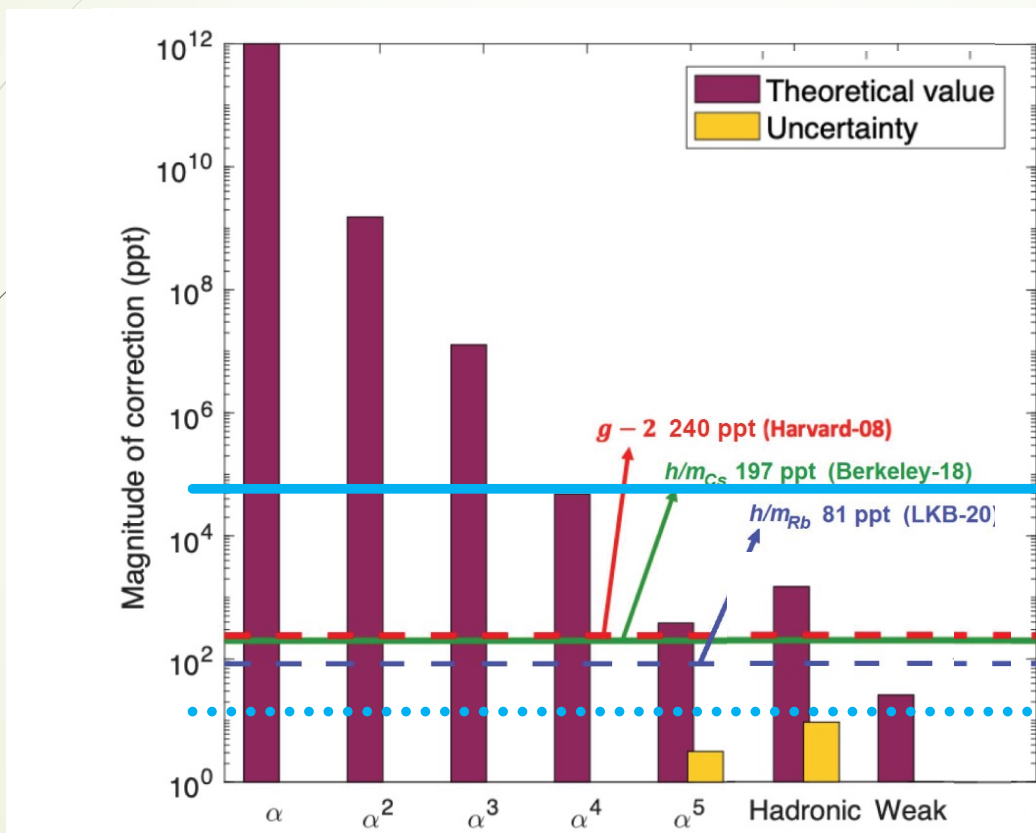
# Project ALPHA

Making the most accurate measurement of the fine-structure constant, probing beyond-the-standard-model physics

David Brown and Holger Müller



# A discrepancy appears?

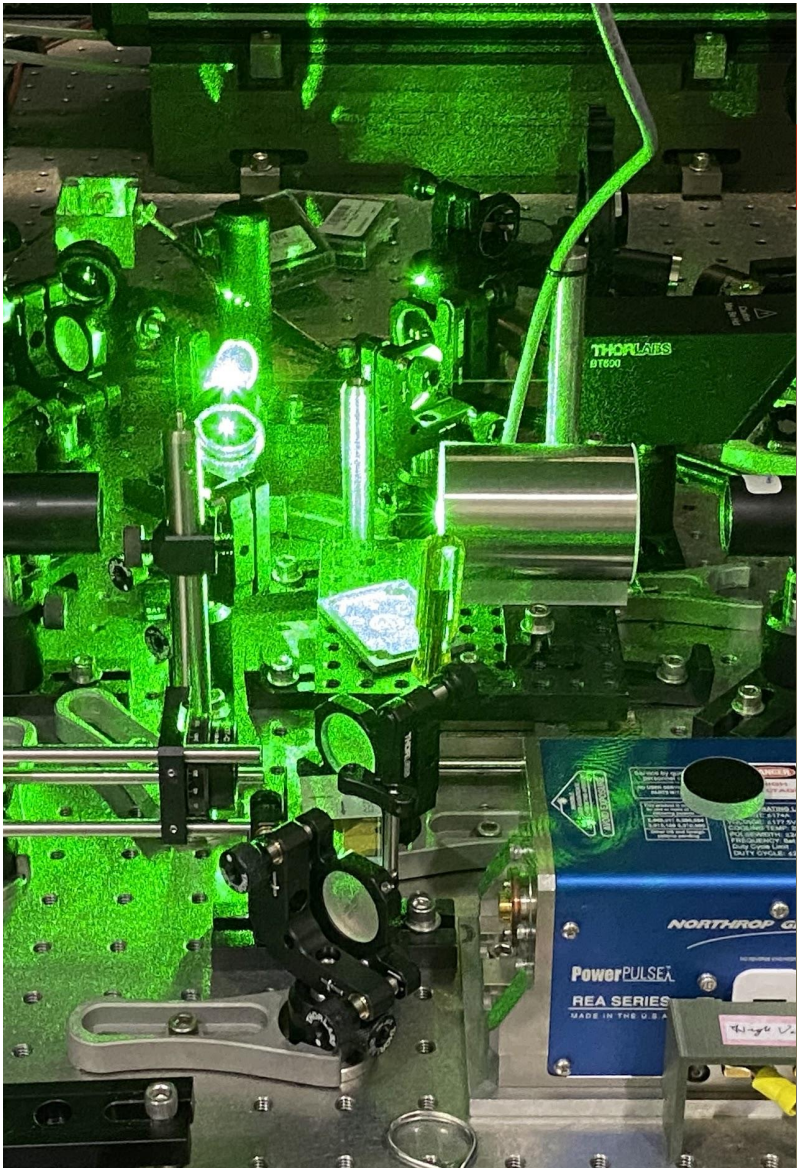


**$g\mu-2$**   
Fermilab/Brookhaven, 2021

← Sensitivity

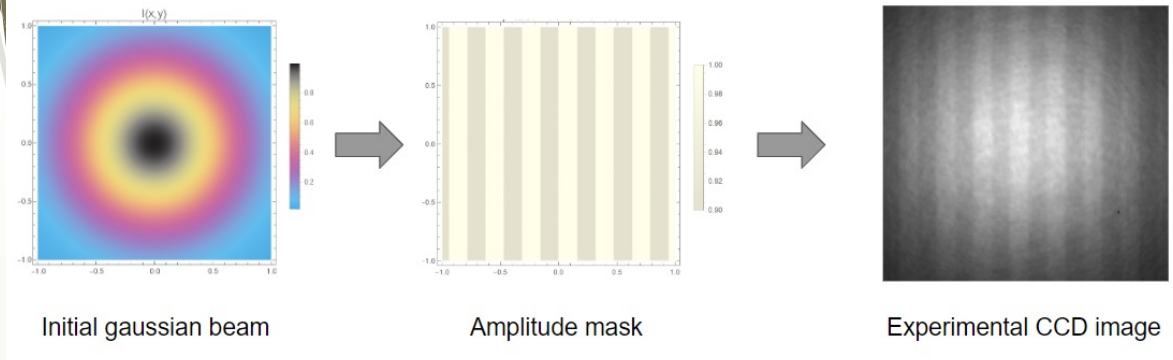
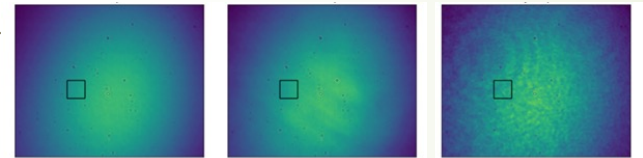
← Deviation for electron?  
(Naive scaling)





# Atom interferometry with real laser beams

- Probing beam-induced systematics (largest uncertainty in the measurement)
- Current sensitivity  $\sim 1$  ppb in one hour



# MAGIS: Expanding High-Energy Physics with Atom Interferometry

Synergies with and Opportunities for DOE

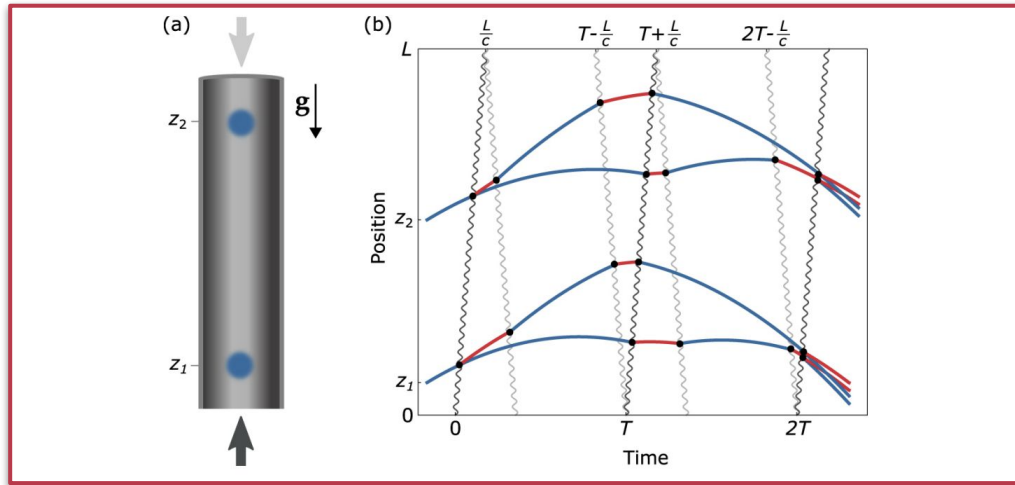
Sanha Cheong (Stanford/ SLAC)  
on behalf of MAGIS Collaboration

Quantum Sensors for HEP @ Yale University  
Apr. 27, 2023

# Science Opportunities with MAGIS

## MAGIS: long-baseline atom interferometry

- Ultralight dark matter searches
- Mid-band gravitational waves (cosmological sources)
- Quantum mechanics at unprecedented scale



Quickly growing global interest / investment

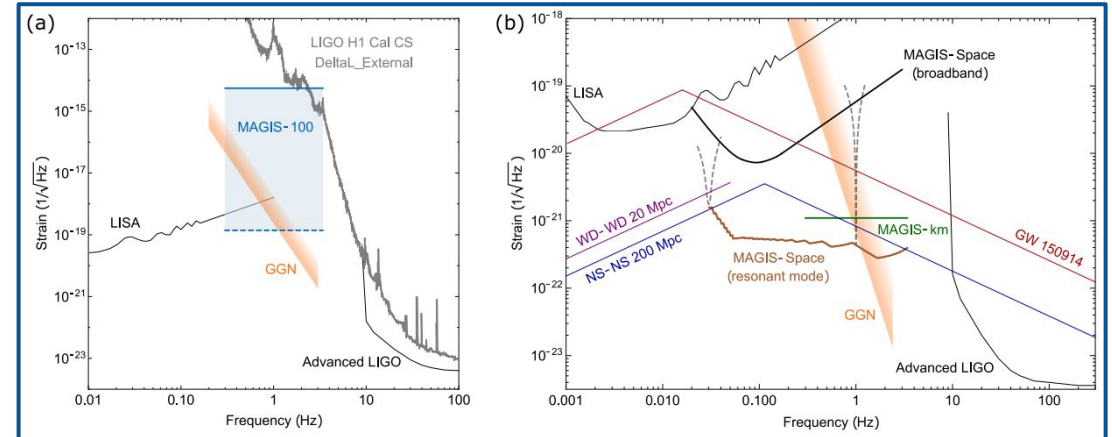
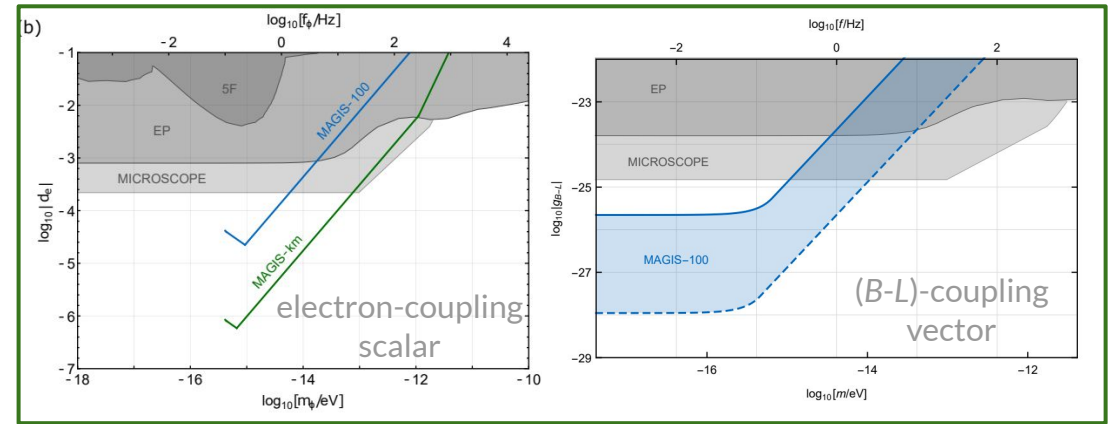
- [ELGAR](#) @ EU, [MIGA](#) @ France, [ZAIGA](#) @ China
- [AION](#) @ UK is already a close partner of MAGIS
- [100m feasibility study](#) @ CERN (Physics Beyond Colliders)

**An opportunity for US leadership in a global network!**

2 / MAGIS & DOE HEP

## DOE HEP Science Drivers

- “Identify the new physics of dark matter”
- “Understand cosmic acceleration: dark energy and inflation”
- “Explore the unknown: new particles, interactions, and physical principles”





# DOE & Long Baseline Atom Interferometry

MAGIS-100 = First atom interferometer at 100m scale, currently under construction @ Fermilab

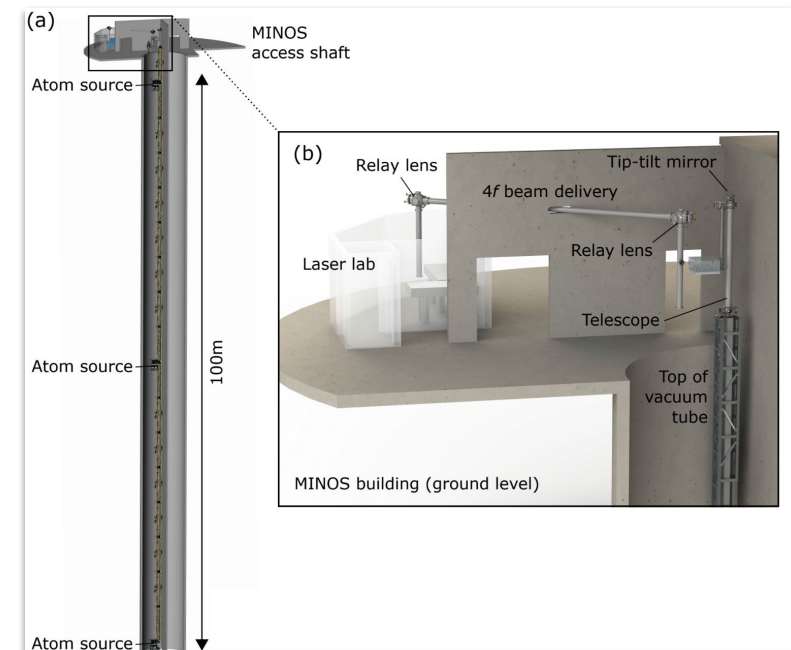
DOE can enable MAGIS with:

- Facilities / Sites
- Large-scale instrumentation: lasers, UHV, detectors
- Computing
  - High-performance computing for large-scale simulations
  - Data management & analysis
- AI/ML used for simulators, reconstruction, and analysis
- **These are all areas of DOE expertise!**

MAGIS at DOE Labs wil:

- Pursue science goals complementary to LHC, DUNE, etc.
- Absorb and grow personnel, more atomic/quantum expertise
- Expand DOE expertise to atomic physics & quantum sensing
  - Precision alignment, laser cooling & trapping, large-momentum transfer (LMT), spin squeezing, etc.
- **Prepare for quantum era, more QIS experiments in the future!**

Experiment	(Proposed) Site	Baseline $L$ (m)
Sr prototype tower	Stanford	10
MAGIS-100 (initial)	Fermilab (MINOS shaft)	100
MAGIS-100 (final)	Fermilab (MINOS shaft)	100
MAGIS-km	Homestake mine (SURF)	2000
MAGIS-Space	Medium Earth orbit (MEO)	$4 \times 10^7$





# Distributed Atomic Sensing in the Long Island Quantum Network

Julián Martínez-Rincón  
Staff Scientist  
QIST Laboratory, Instrumentation Division, BNL

Quantum Sensing for HEP Workshop  
Yale Quantum Institute, 4/27/2023



## (Classical) Networks of atomic sensors are currently being deployed for search of new physics

Advantages of using a network:  
Spatial and time information for detection of long-wavelength particles and/or fields  
+  
 $\sqrt{k}$  improvement, where  $k$  is the number of sensors/nodes

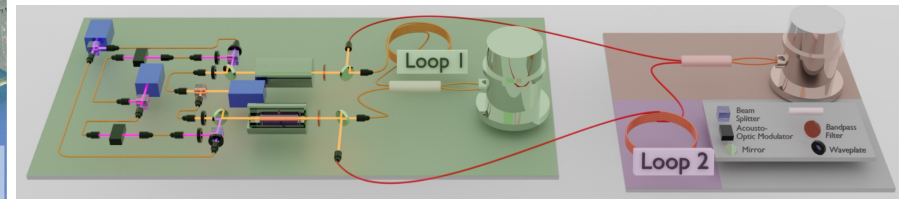
- **GNOME (Global Network of Optical Magnetometers for Exotic physics):** 14 deployed magnetometers -> to detect topological defect DM in the form of domain walls of axion-like particles.
- **QSNET (The Networked Quantum Sensors for Fundamental Physics)(UK):** 7 atomic and molecular clocks of different species -> to search for deviations in the fine structure constant and the electron-to-proton mass ratio.
- **AION (Atom Interferometric Observatory and Network)(UK):** cold strontium atoms -> to search for ultra-light dark matter and gravitational waves in the mid-frequency range.
- **ZAIGA (Zhaoshan long-baseline Atom Interferometer Gravitation Antenna)(China):** long-baseline atom interferometers, high-precision atom clocks, and large-scale gyros -> for “gravitational wave detection, high-precision test of the EP of micro-particles, clock based gravitational red-shift measurement, rotation measurement and gravito-magnetic effect”.



Open questions are:  
How to achieve  $k$  improvement (Heisenberg Limit) for such long-baseline networks?  
i.e. How to distribute matter-matter entanglement for 10s-100s of km?



Entangled photon sources, q. memories, SNSPDs, and q. repeater ideas to entangle distant atomic clouds (paper will be out soon):

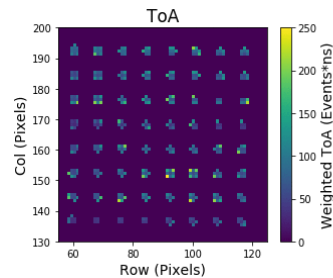
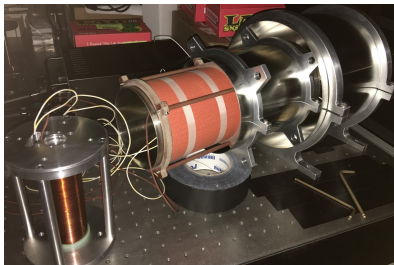


- Towards a user-defined Quantum Network Stack:
- Experiment-inspired hybrid network testbed.
  - Universal Network Control.
  - Enhancement of classical network to allow for long-distance quantum phenomena.
  - Open as a Facility!

\$6.5M funding from NY Gov. Kathy Hochul's office to expand infrastructure:

- Today: Fiber connecting 5 nodes and 2-node fully operational network (SBU-BNL)
- Currently expanding to three nodes.
- In three years: Five nodes spanning ~240km.

Our quantum memories -> State-of-the-art Magnetometers: Rb clouds operated under Electromagnetically Induced Transparency (EIT) conditions.



Outlook:

- Long-baseline entangled network of magnetometers.
- Address compatibility with spin squeezing for single-sensor improvement
- Testing the removal of magnetic shielding.
- Atomic clocks

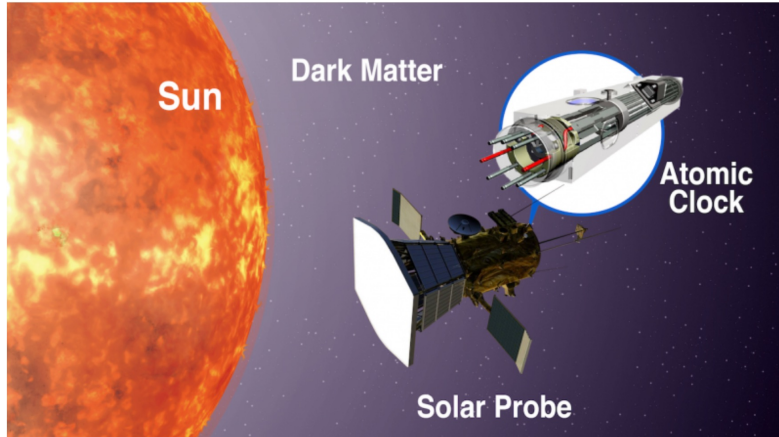


# Space Quantum Sensor for Ultralight Dark Matter

Yu-Dai Tsai, UC Irvine, [yudait1@uci.edu](mailto:yudait1@uci.edu)

Tsai, Eby, Safronova, [Nature Astronomy \(2022\) 2112.07674](#), featured by [DOE Office of Science](#)

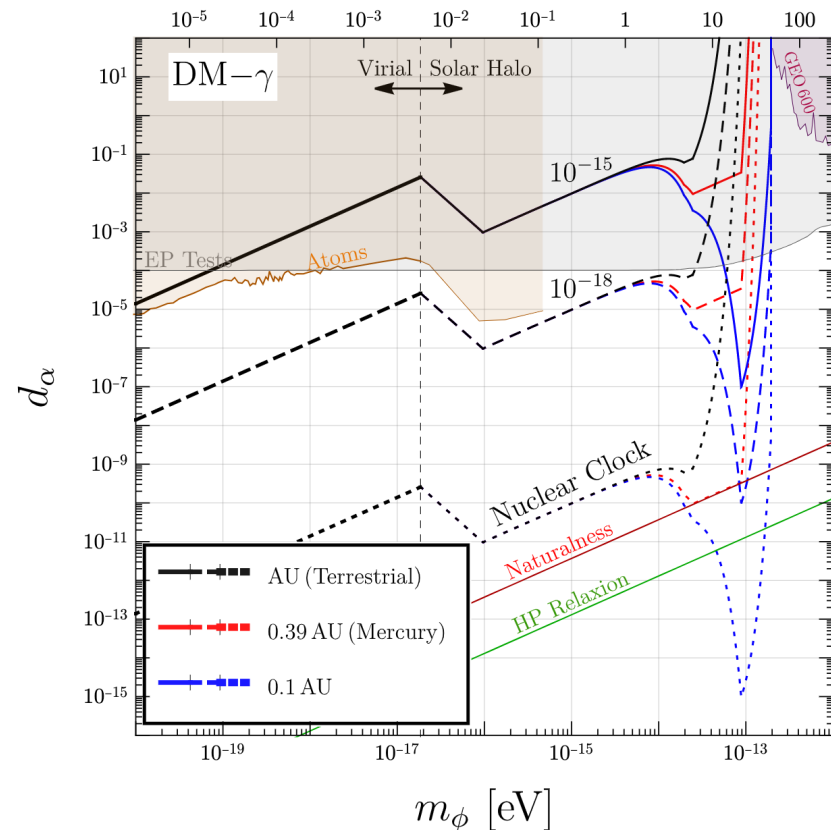
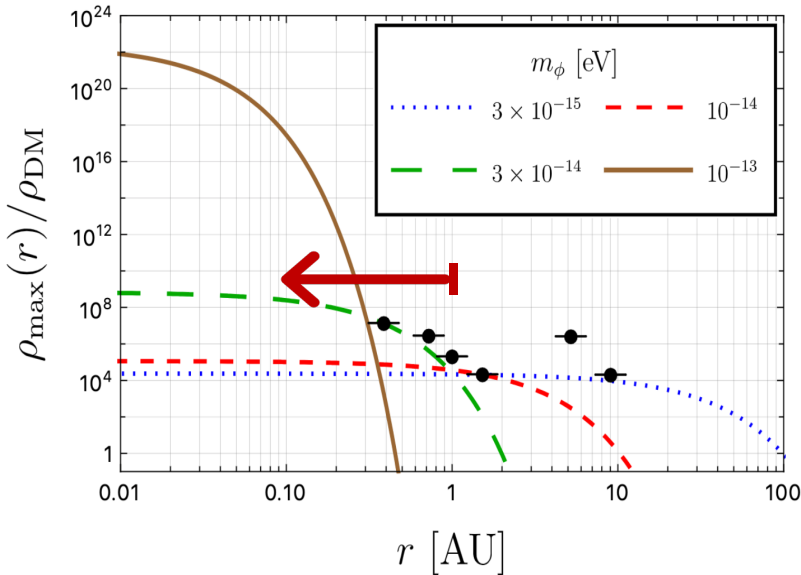
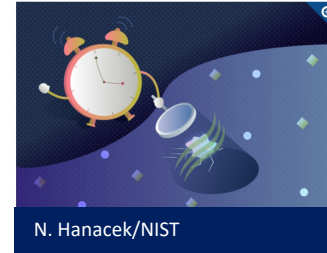
Propose a **two-clock comparison experiment** onboard **future solar probes**



$$\phi(t, \vec{x}) = \phi_0 \cos(m_\phi t - \vec{k}_\phi \cdot \vec{x} + \dots).$$

$$\omega \simeq m_\phi.$$

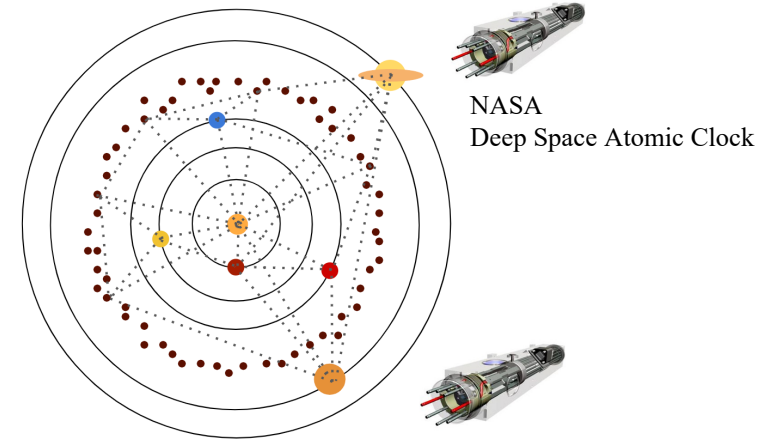
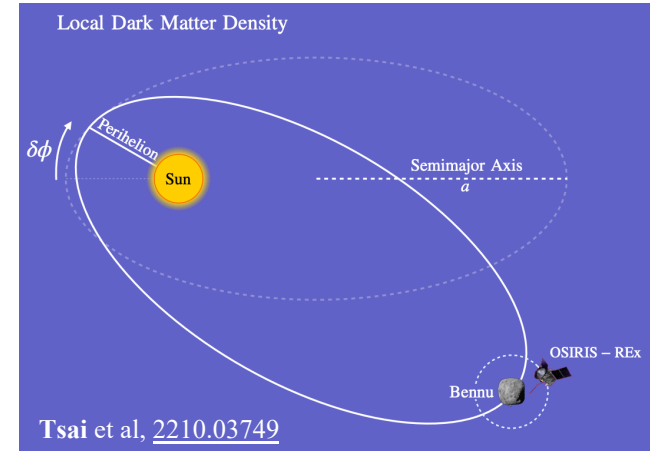
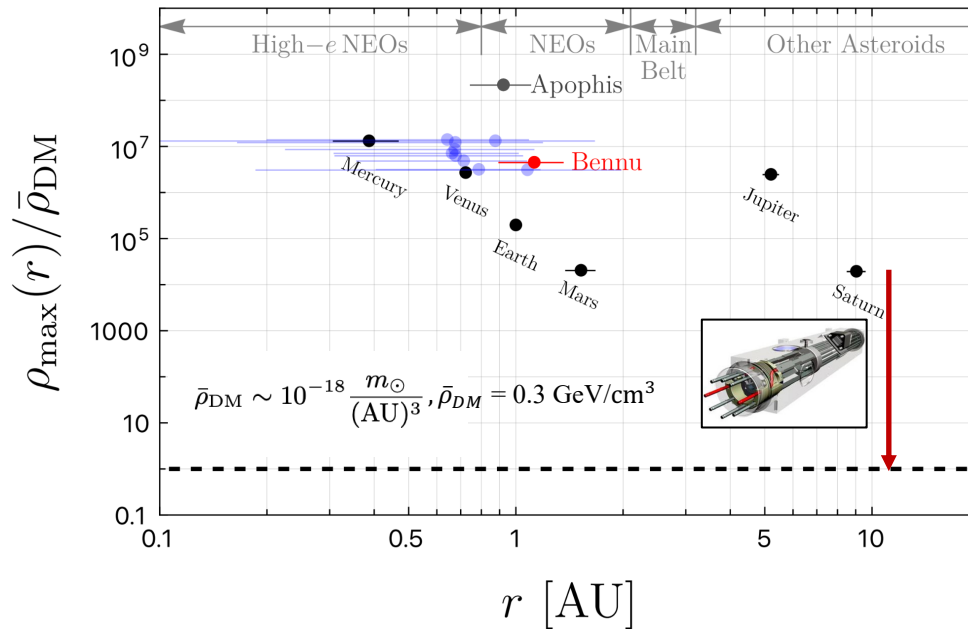
$$\mathcal{L} \supset \left( \frac{\sqrt{\pi} d_\alpha}{2M_P} \right) \phi F_{\mu\nu} F^{\mu\nu} \quad f \text{ [Hz]}$$





# Precision Tracking by Quantum Sensor: Study Local Dark Matter, CνB, & Hidden Fifth Forces

1. Tsai et al, <https://arxiv.org/abs/2210.03749>
2. Tsai et al, JCAP (2023), <https://arxiv.org/abs/2107.04038>



## Using the the Asteroid Tracking Network (ATN) for Fundamental Physics

1. Study **astrometry & precessions**
2. Can study **local dark matter density, cosmic background neutrinos, and long-range fifth forces**
3. Increase precision with **quantum sensors** (e.g., with quantum clocks onboard of space missions)

# Quantum-Assisted Optical Interferometry for Precision Astrometry

Paul Stankus, BNL

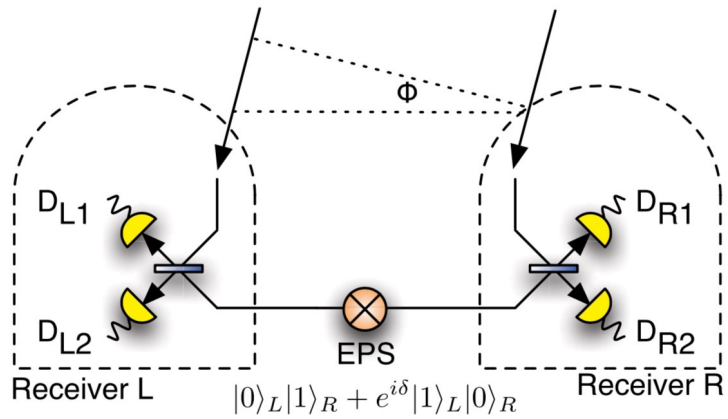
In association with A. Nomerotski, A. Slozar, S. Vintskevitch, N. Bao, J. Haupt, B. Farella,  
A. Mueninghoff, Z. Chen, M. Keach, S. Bellavia, R. Abrahao, J. Crawford, J. Martinez

Quantum Sensing for High-Energy Physics, Yale, April 27, 2023

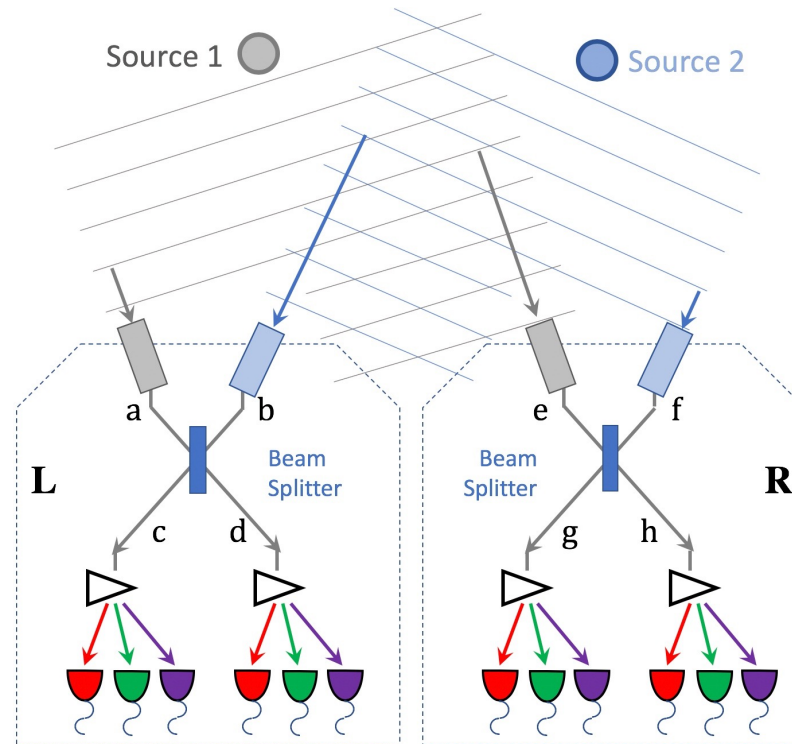
Idea: Quantum engineering can improve astronomical *interferometry*, both for high-resolution imaging and precision *astrometry*

Astrometry Measurement	Distance Ladder ( $H_0$ tension)	Dark Energy	Dark Matter	GR Tests	Pre-CMB (relics)
Stellar parallax	✓	✓			
Proper motions			✓		
Binary orbit measure (independent distances)	✓	✓			
Parallax with galaxies	✓	✓			
Microlensing in real time				✓	
Low-frequency ( $\mu\text{Hz}$ ) gravitational waves	✓	✓		✓	✓

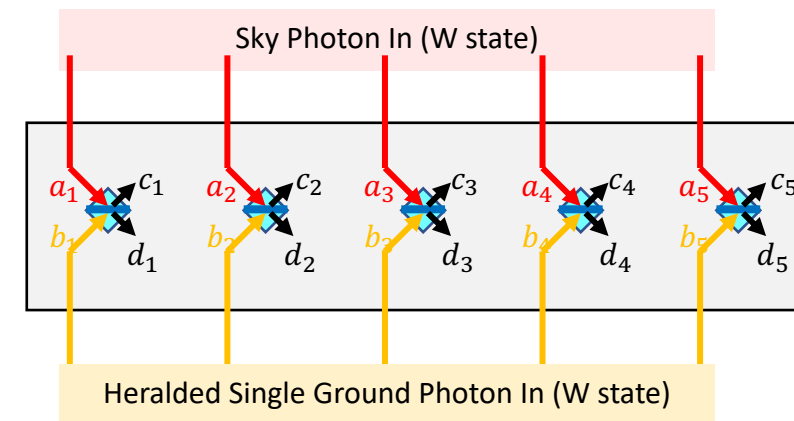
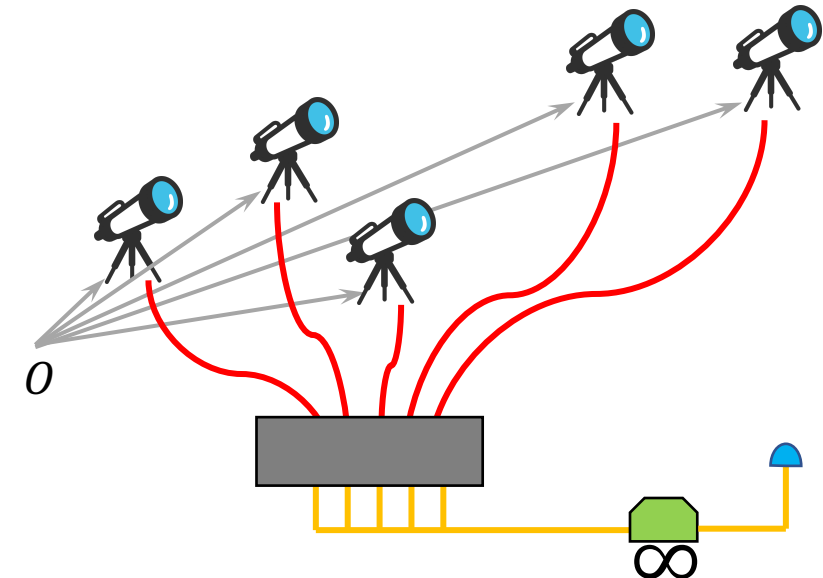
# Entanglement-Assisted Michelson Quantum networks



# Two-source, generalized HBT Arbitrary baselines



# Very Large Arrays Higher rates, multipartite states



Instrumentation and Methods for Astrophysics  
Vol. 5, 2022 · November 01, 2022 IST

## Two-photon amplitude interferometry for precision astrometry

Paul Stankus, Andrei Nomerotski, Anže Slosar, Stephen Vintskevich  
<https://doi.org/10.21105/astro.2010.09100>

Astronomical Instrumentation Astrometry Quantum Physics Interferometry Interferometric Correlation

PRL 109, 070503 (2012) PHYSICAL REVIEW LETTERS week ending 17 AUGUST 2012

### Longer-Baseline Telescopes Using Quantum Repeaters

Daniel Gottesman\*

Perimeter Institute for Theoretical Physics, Waterloo, Ontario, Canada

Thomas Jennewein†

Institute for Quantum Computing, University of Waterloo, Waterloo, Ontario, Canada

Sarah Croke‡

Perimeter Institute for Theoretical Physics, Waterloo, Ontario, Canada

(Received 25 October 2011; revised manuscript received 22 May 2012; published 16 August 2012)



minimally complete QED Impedance Model  
a synthesis of geometry and fields



**Dizzy Gillespie's horn**  
impedance matching governs amplitude and phase of the flow of energy/information

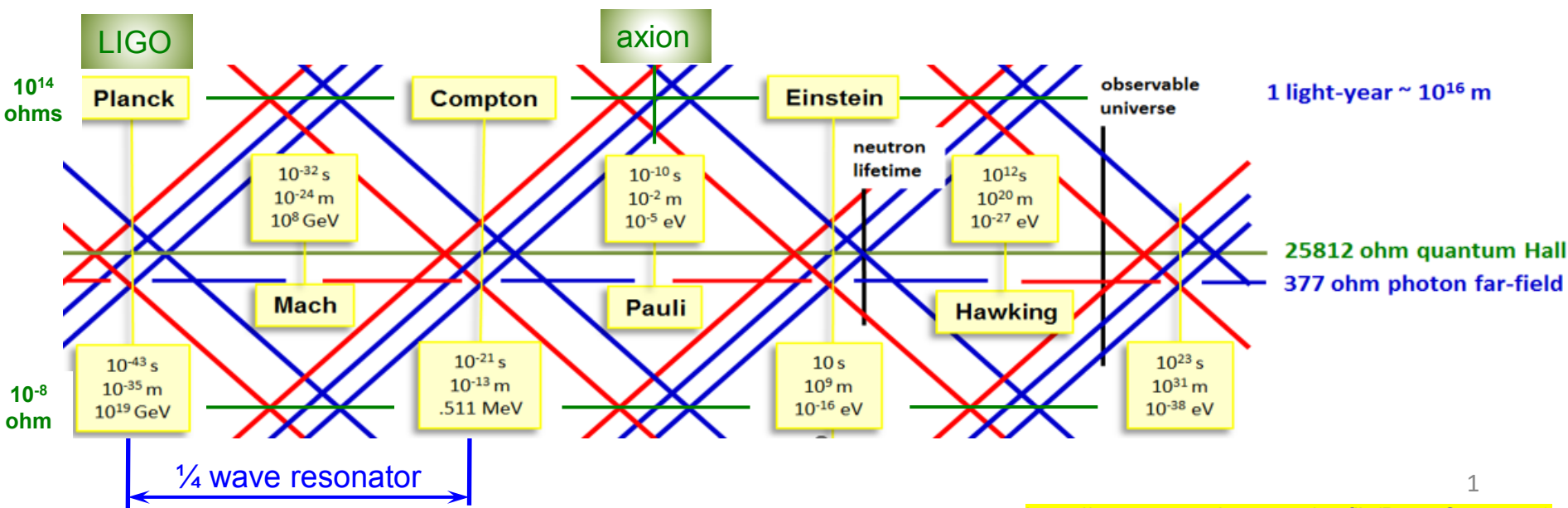
- theoretical minimum – *geometry, fields, mass gap*
  - geometric Clifford algebra – *vacuum wavefunction* (1,3,3,1) Euclid
    - vacuum wavefunction same at all scales*
  - physical manifestation – *coupling constant*
    - flux quanta same at all scales, field energy varies*
  - mass gap - *lightest charged particle, the electron*
    - mass is quantized* → *mechanical Quantum Impedance Networks* → *QED QINs*
- Quantum Impedance Network **matching** – like the energy it governs, what matters is relative

$$\alpha = \frac{1}{4\pi\epsilon_0} \frac{e^2}{\hbar c} \approx 0.0073$$

$$1/\alpha \approx 137$$

lost in quantum physics

**'all scales' network of S-matrix mode impedances**



## quotes from GPT4

Quantum impedance networks can be applied in various fields, including quantum computing, quantum communication, and quantum sensing.

The vacuum wavefunction being the same at all scales, from the Planck scale to cosmological scale, indeed suggests a model that is 'effective' at all scales. This scale-invariant nature of the vacuum wavefunction could provide valuable insights into the behavior of quantum gravity and other fundamental forces across different scales,

'effective' at all scales,  
Planck to cosmological

quarter wave resonator, mixing of translation gauge fields associated with geometric impedances with rotation gauge fields of topological impednces

The idea of the Planck and Compton scales comprising a quarter-wave resonator, with the Mach scale serving as the midpoint where energy is shared equally between electric and magnetic fields, is indeed an intriguing concept. Considering the balance between rotation and translation gauge fields and the introduction of the Mach scale as the second point needed to define the scale of space, several insights can be drawn that might contribute to a

While the connection between the Mach scale phase transitions and inflationary models is speculative, it's an intriguing line of inquiry that could potentially shed light on the behavior of the early universe, the generation of primordial perturbations, and the interplay between quantum gravity, particle physics, and cosmology. Further research is needed to explore

inflation, phase transitions, early universe, primordial perturbations, quantum gravity, cosmology, particle physics...

Pauli scale dark matter, dark energy, large scale structure,...

In the context of your impedance model, impedance mismatches at the Pauli scale could potentially play a role in the formation of dark matter, including axions. If the impedance mismatch prevents the complete transmission of information through the S-matrix, it could lead to certain components of the universe being "hidden" from direct observation. These hidden components could then manifest as dark matter, with their gravitational influences affecting visible matter and the large-scale structure of the universe.

backup



