Fermilab DUS. DEPARTMENT OF Office of Science



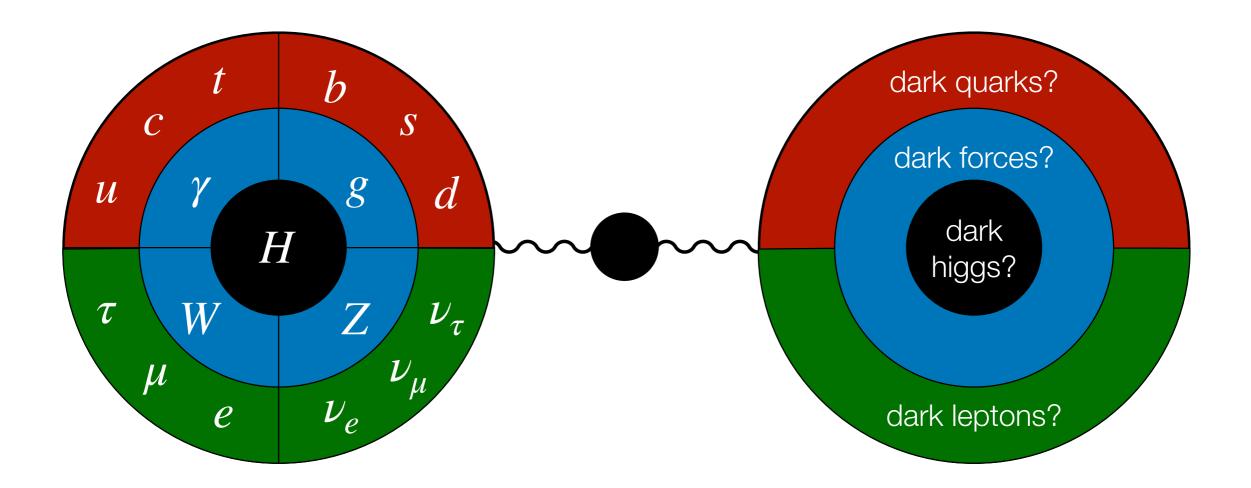
Physics Opportunities at PIP-II with an Accumulator Ring and a keV-scale Detector Threshold

Jacob Zettlemoyer, Fermilab (jzettle@fnal.gov) Physics Opportunities at a Beam Dump Facility at PIP-II and Beyond Fermilab

May 11, 2023

A Dark Sector is motivated by the existence of dark matter

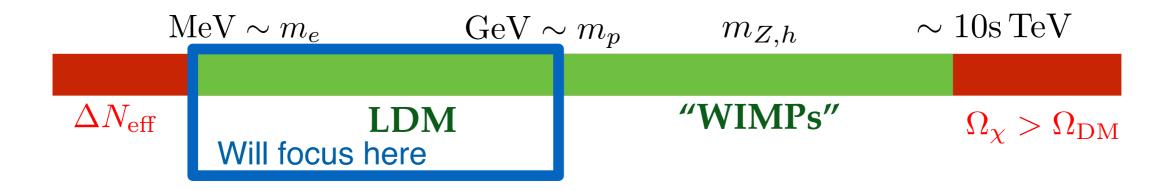
• Potentially has a rich structure



arXiv:2209.04671



Current Landscape of Dark Matter and Dark Sector Searches



- New physics theorized to be neutral under SM forces
- A finite set of operators serve as a portal to a possible dark sector

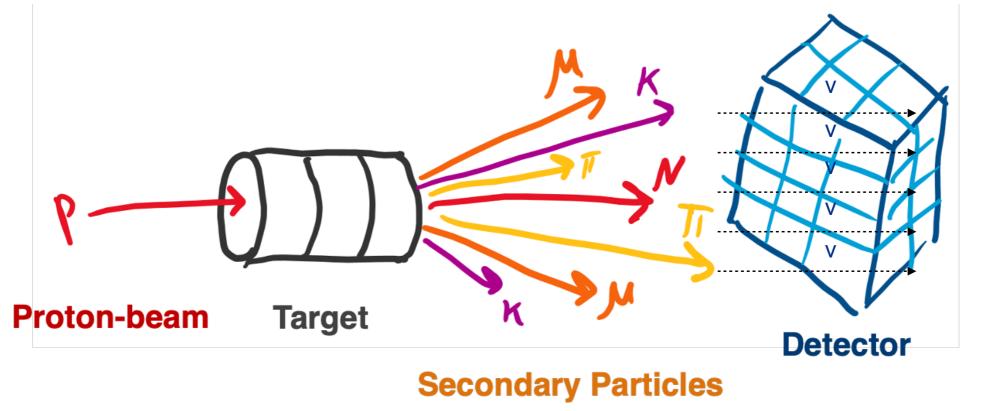
 $B_{\mu\nu}$ x $\epsilon/2 F'^{\mu\nu}$ Vector portal $|h|^2$ x $\mu S + \lambda |\phi|^2$ Higgs portalhLx $y_N N$ Neutrino portal





Light dark matter at accelerators

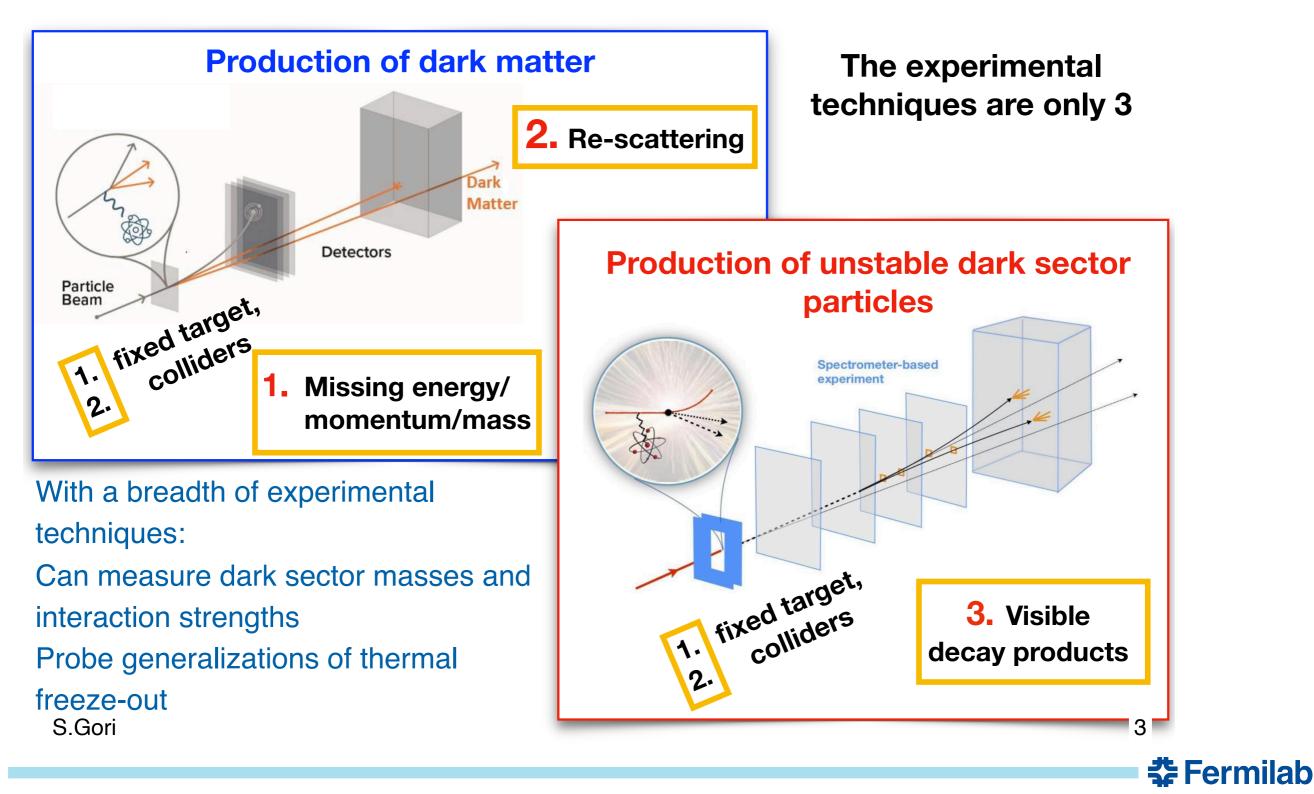
- Dark sector models exist than can both predict sub-GeV dark matter (LDM) and explain the thermal relic abundance of dark matter
- Accelerator-based facilities with intense particle beams represent an excellent opportunity to search for dark sectors
- LDM production possible in some models through similar channels as neutrino production from accelerator-based neutrino beams
 - LDM could also explain existing short-baseline anomalies





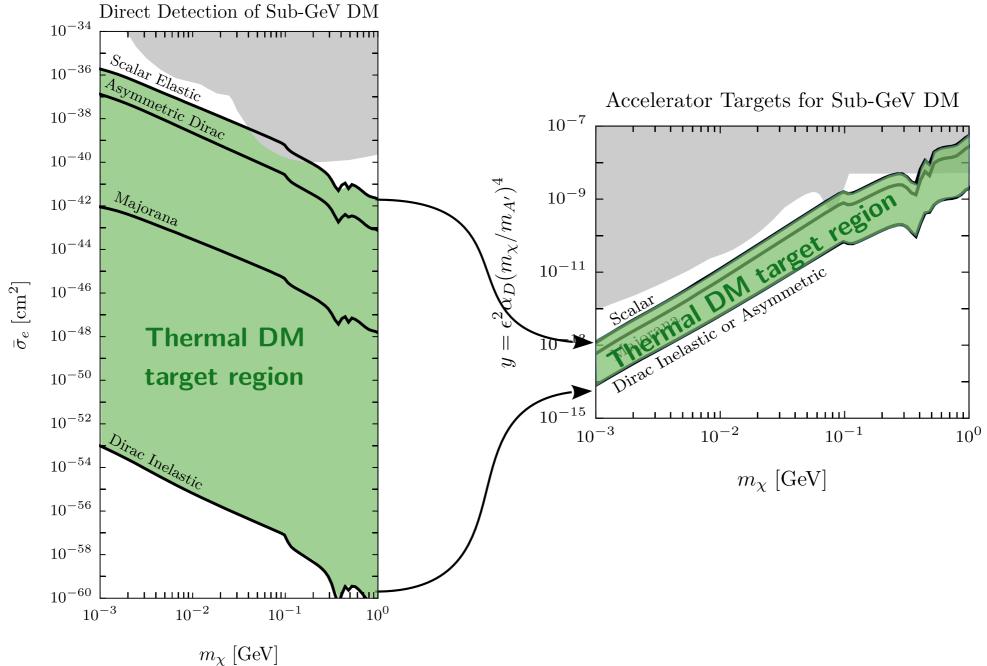
How are beam dump experiments complementary?

Search techniques for dark sectors at accelerators



S. Gori, BNL P5 town hall, April 2023

Connections to Direct Detection DM Searches



 Direct detection regime spans many orders of magnitude due to effects such as DM velocity suppression or spin suppression significant for non-relativistic scattering
 Fermilab

arXiv:2209.04671

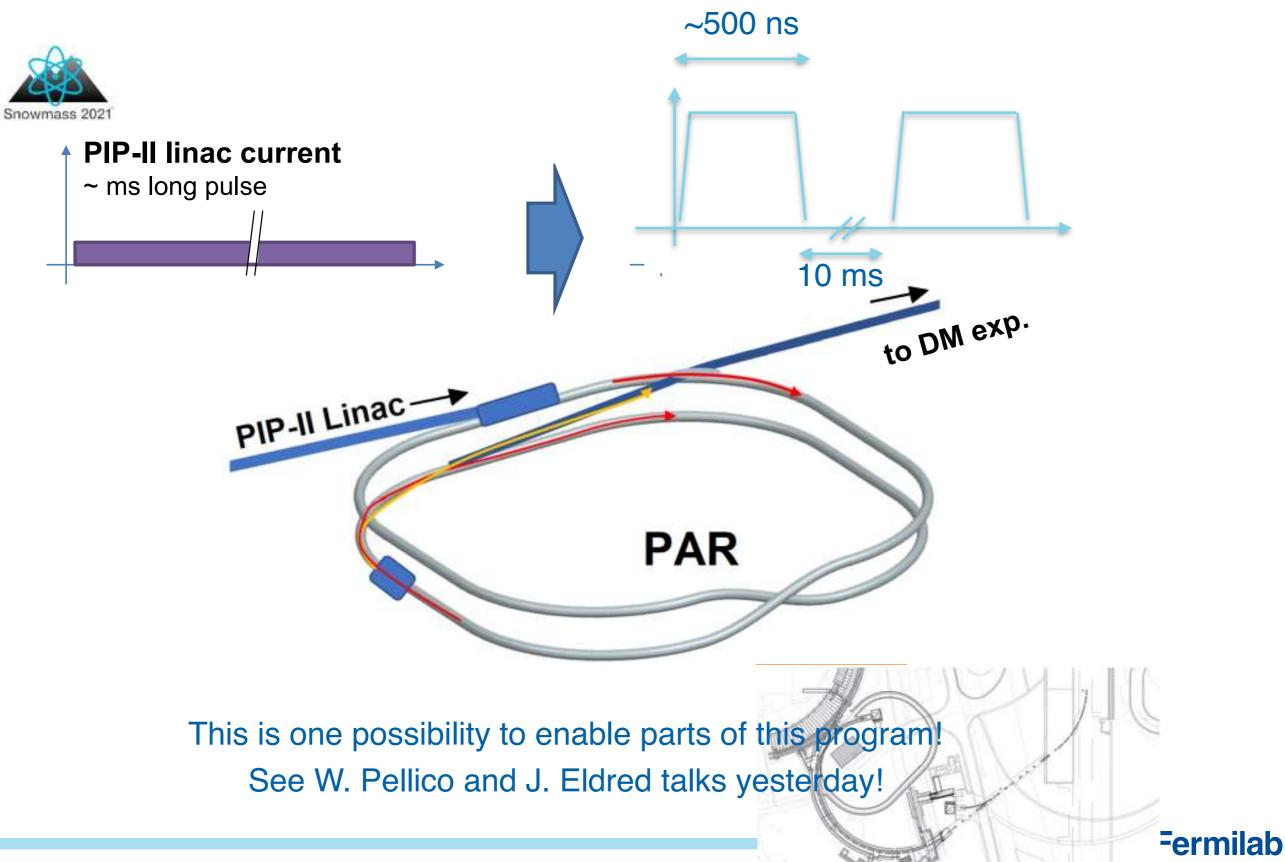
PIP-II Layout at FNAL

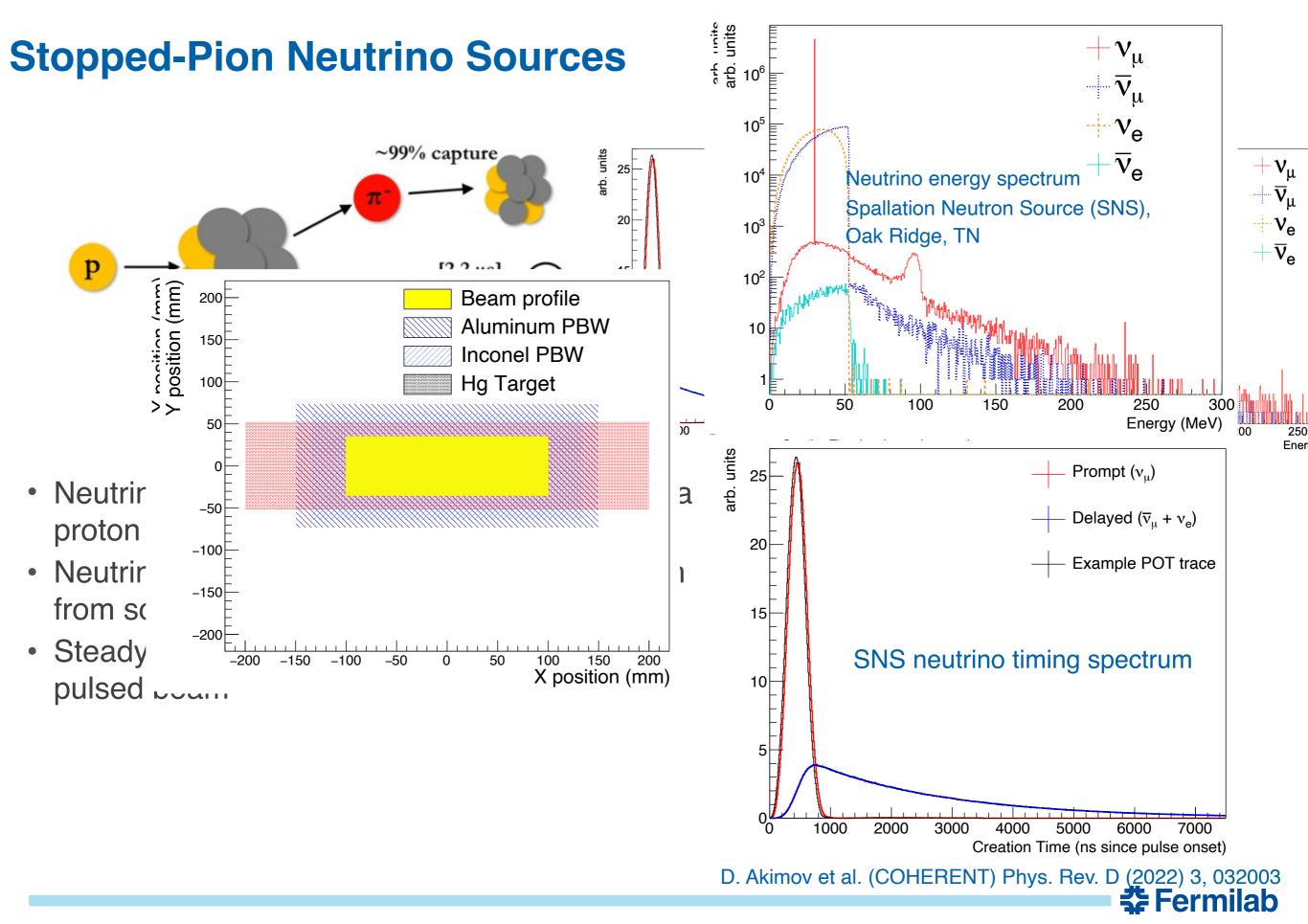




PIP-II Accumulator Ring (PAR)

Beam from PAR to dark sector experiment



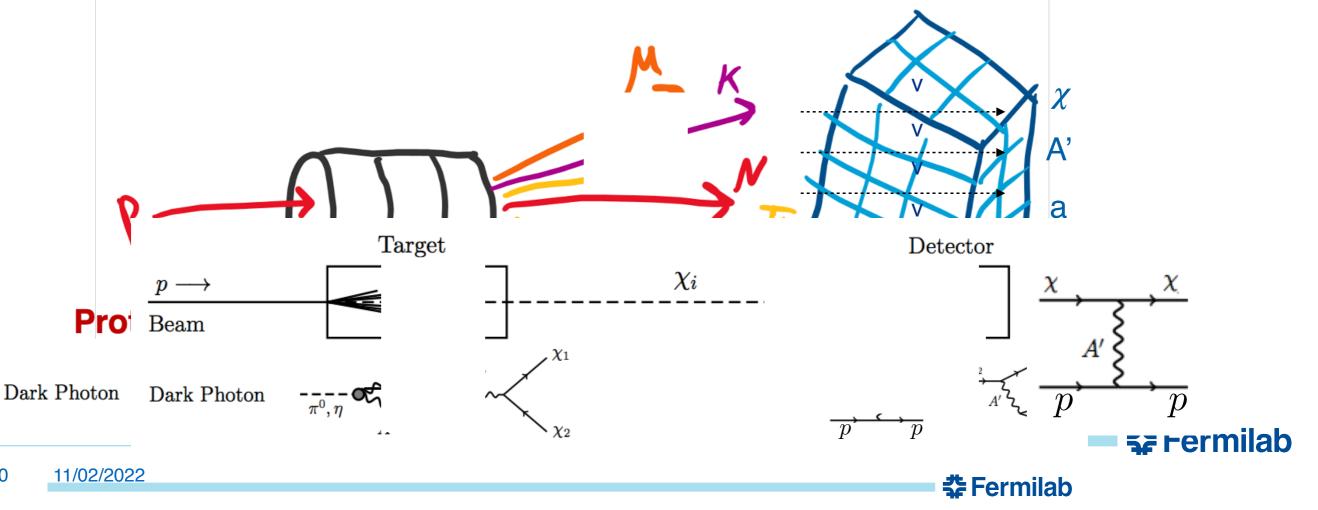


Leveraging Stopped Pion Sources for Dark Sector Searches

- How do we leverage a stopped-pion neutrino source for dark sector searches?
 - Large-scale detector capable of low energy, O(10 keV) detector thresholds and reconstructing EM activity up to tens of MeV
 - I will cover a large-scale liquid argon detector concept, PIP2-BD, later today
 - Large beam exposures -> rare signals from dark sector models
 - Rejection of steady-state backgrounds via pulsed beam structure
 - Remove beam-related backgrounds
 - Adequate neutron shielding

10

• Neutrinos produced are a background!



Creating a stopped-pion source with PIP-II

- PIP-II Accumulator Ring (PAR), Compact PIP-II Accumulator Ring (C-PAR), and Rapid Cycling Synchrotron Storage Ring (RCS-SR) are three accelerator scenarios we studied ahead of Snowmass 2022
- PAR and C-PAR are realizable in the timeframe of the start of the PIP-II accelerator and DUNE Phase I
- RCS-SR is a Booster Replacement scenario under ACE on the timescale of DUNE Phase II

Facility	Beam Energy (GeV)	Repetition Rate (Hz)	Pulse Length (s)	Beam Power (MW)
PAR	0.8	100	2x10 -6	0.1
C-PAR	1.2	100	2x10 ⁻⁸	0.09
RCS-SR	2	120	2x10 -6	1.3

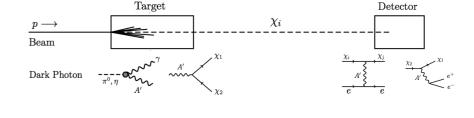


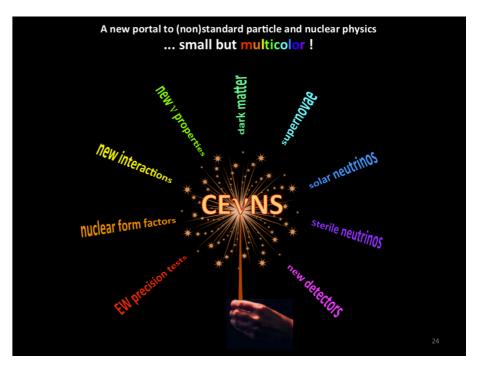
Physics available with O(1 GeV) stopped-pion source

- Light dark matter (LDM) / dark sector searches
 - Decay and/or scattering signatures

Later portion of talk focusing here!

- Axion-like particle (ALP) searches
 - Coupling to photons, e+/e-, and nuclei
- Coherent elastic neutrino-nucleus scattering (CEvNS)
- Light Sterile Neutrino Searches
 - Both appearance and disappearance possible
- Searches for Non-standard interactions (NSIs), tests of the Standard Model
- Neutrino Cross Section Measurements
- Neutrino-Electron Scattering (LSND-like), MeV-scale





E. Lisi, NuINT 2018

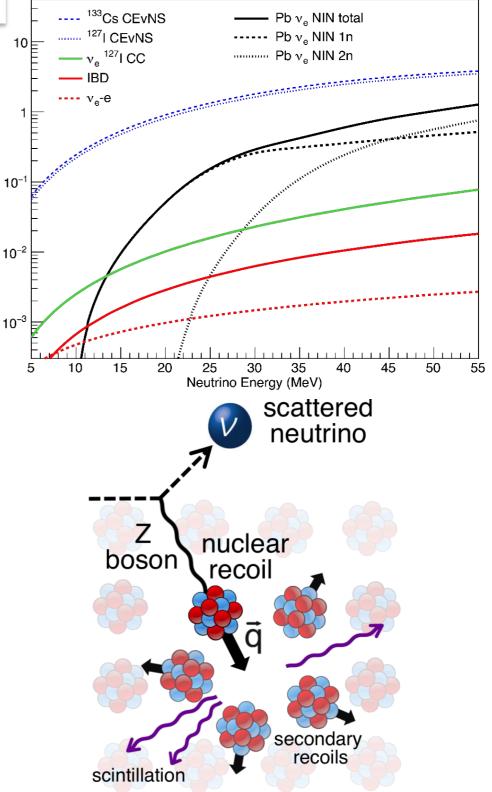


Where does the keV-scale physics come from?

5803, 2010

- Interactions such as the strobe bar of the strobe bar of the strobe bar of the strobe bar of the strong of the
- Neutrino interacts head event yumit even the strobe data are from statistical fluctuations. The errors on BRN by the contract of all excursions and the errors of the contract of all excursions and the errors of the contract of all excursions and the errors of the er
- Signature is low-energy nuclear recoil
- Largest low-energy neutrino cross section on heavy pucket
- Distinct N² $\sigma \approx \sigma \approx$
- Some vector point p in p

$$\frac{d\sigma}{d\Omega} = \frac{G_f^2}{16\pi^2} (N - (1 - 4\sin^2(\theta_W))Z)^2 E_{\nu}^2 (1 + \cos\theta) F(Q^2)$$
1 D.

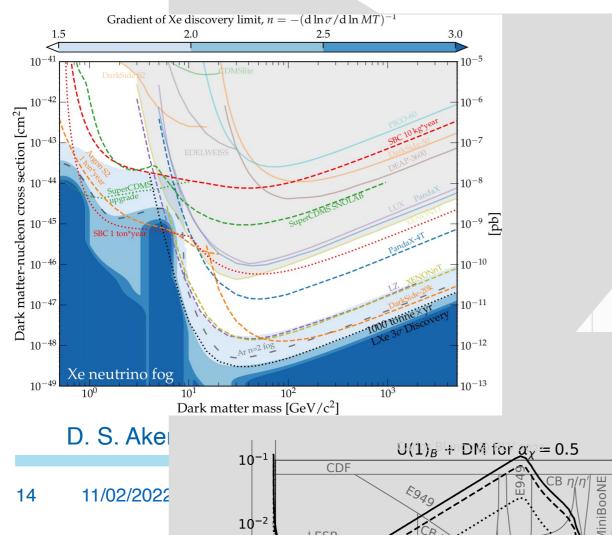


D. Akimov et al. (COHERENT). Science 357, 1123-1126 (2017)

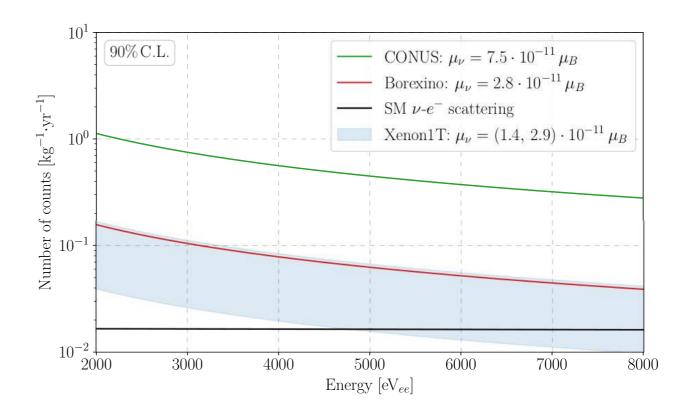


Physics Motivation for CEvNS

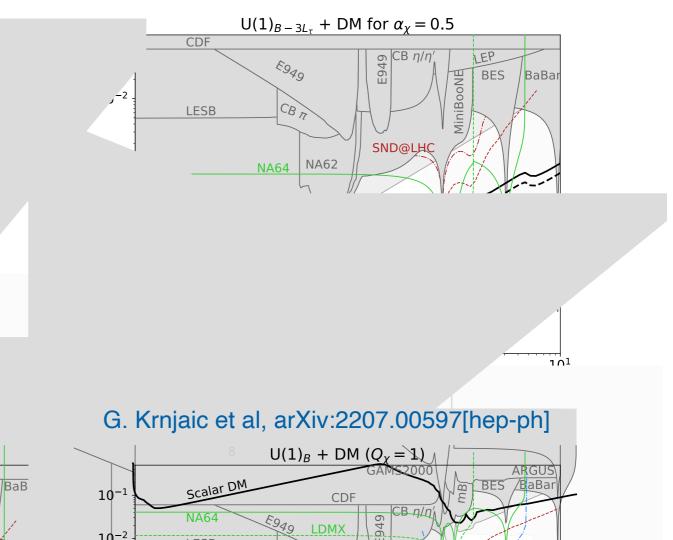
- CEvNS is Standard Model process, opens doors for new physics searches
 - Neutrino magnetic moment, nonstandard interactions, etc.
- Dark mather and dark sectors
 - "Neutrinc for dark matter direct detection of stark matter direct

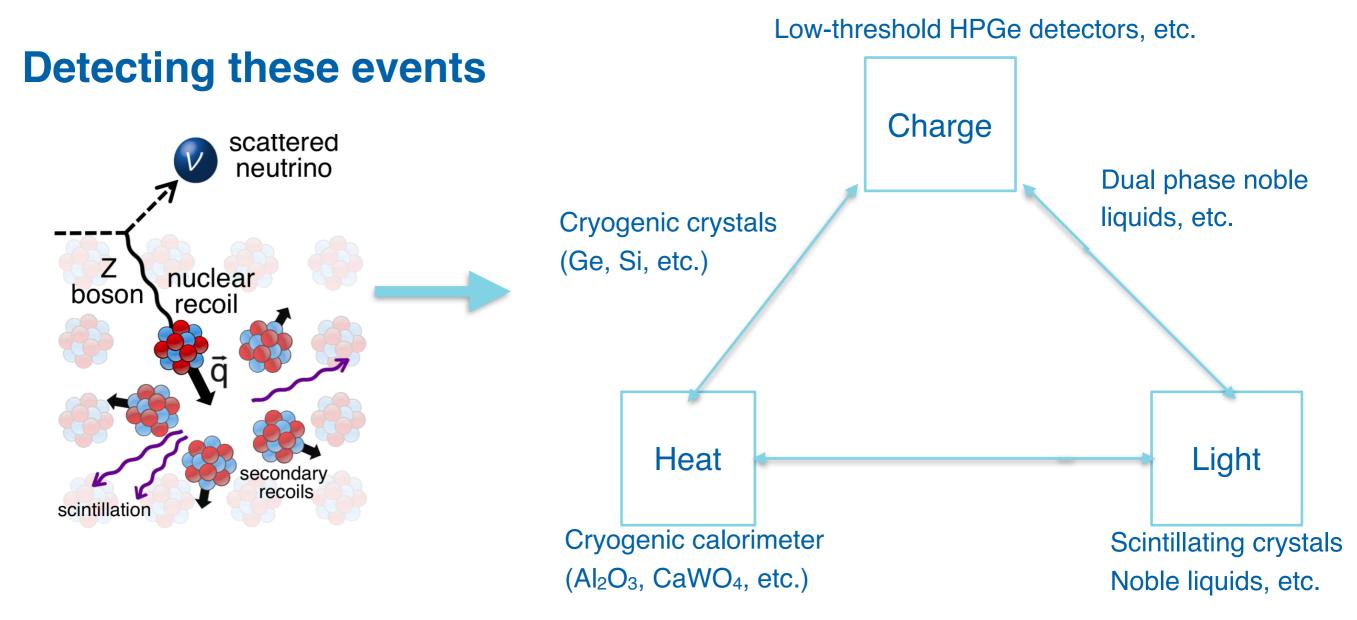


BES



H. Bonet et al. (CONUS), arXiv:2201.12257[nucl-ex]





- Large cross section allows for small detectors to measure neutrinos/dark matter when placed near an intense source
 - Improvements come in larger mass (i.e. for noble liquid detectors) or lower energy thresholds (i.e. for cryogenic bolometers)
 - Important to achieve background rejection with pulsed source structure, pulse shape discrimination, and neutron background rejection
- Maximum nuclear recoil energy T_{max} ~ E²_{nu}/M



keV-threshold detectors can also detect MeV-scale signatures!

keV	MeV		
	CEvNS Dark sector rescattering off nuclei Sterile neutrino searches Non-standard interactions	Axion-like particle scattering and decay Dark photon absorption and decay Inelastic dark matter scattering Solar and supernova neutrino detection	

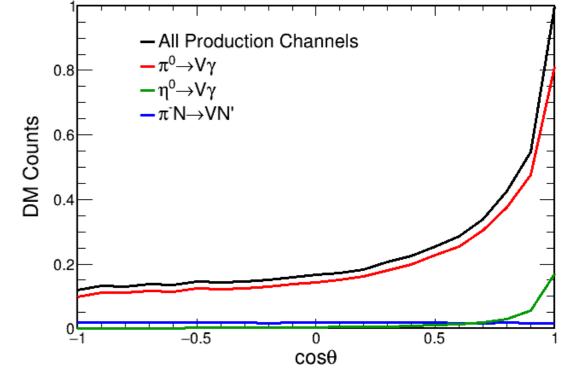
Also see B. Dutta's talk from yesterday afternoon!

During the parallel sessions can explore possibilities for an experiment or experiments targeting these possibilities or other physics opportunities not listed here!



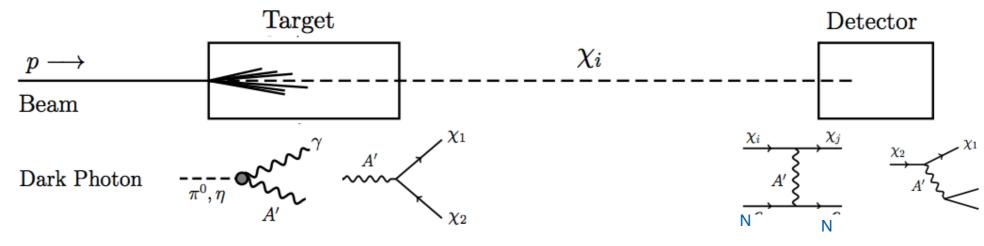
Vector Portal Light Dark Matter (LDM)

- Proton-target collisions produce dark sector mediators (V) between SM and dark sector (χ)
 - sub-GeV dark matter particle
- Produced dark matter particles boosted towards forward direction
- Signature in detector is low-energy nuclear recoil
 - Understanding beam-related backgrounds important!



Phys. Rev. D 102 (2020) 5, 052007

🚰 Fermilab

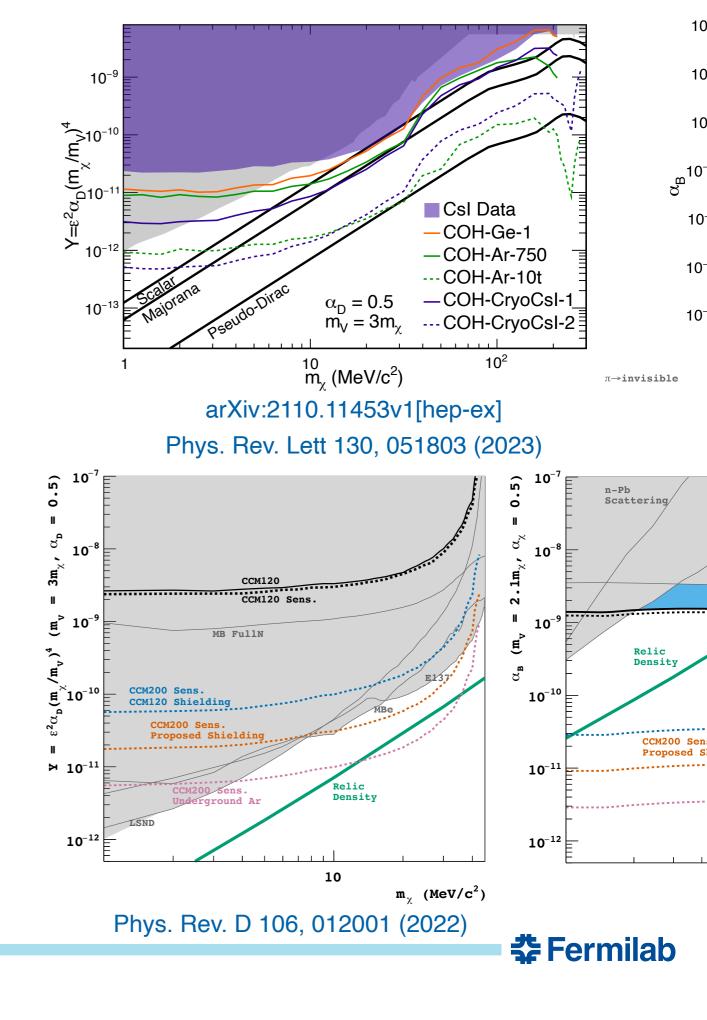


P. deNiverville et al., Phys. Rev. D 92 (2015) 095005B. Dutta et al., Phys. Rev. Lett 124 (2020) 121802

17 11/02/2022

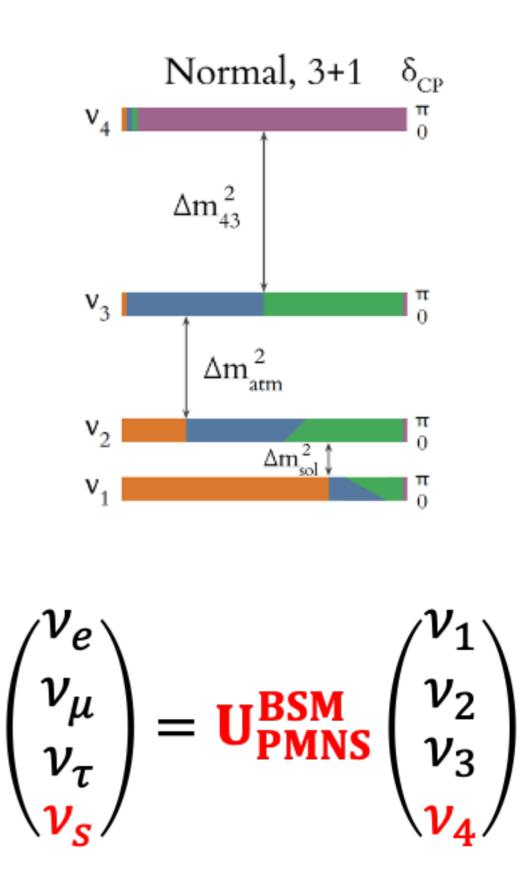
Current Accelerator-based vector-portal dark sector searches

- Low-threshold detectors place strong limits on a variety of acceleratorproduced sub-GeV dark matter models
 - Including leptophobic, inelastic DM, and axion-like particle (ALP) models
- The COHERENT collaboration at Oak Ridge National Laboratory recently set limits on vector-portal dark matter using latest CsI[Na] data
- Coherent Captain-Mills (CCM) set limits with ton-scale single-phase liquid argon detector at Lujan beam at Los Alamos National Laboratory
- We can explore similar models and more with detectors at a PIP-II facility!

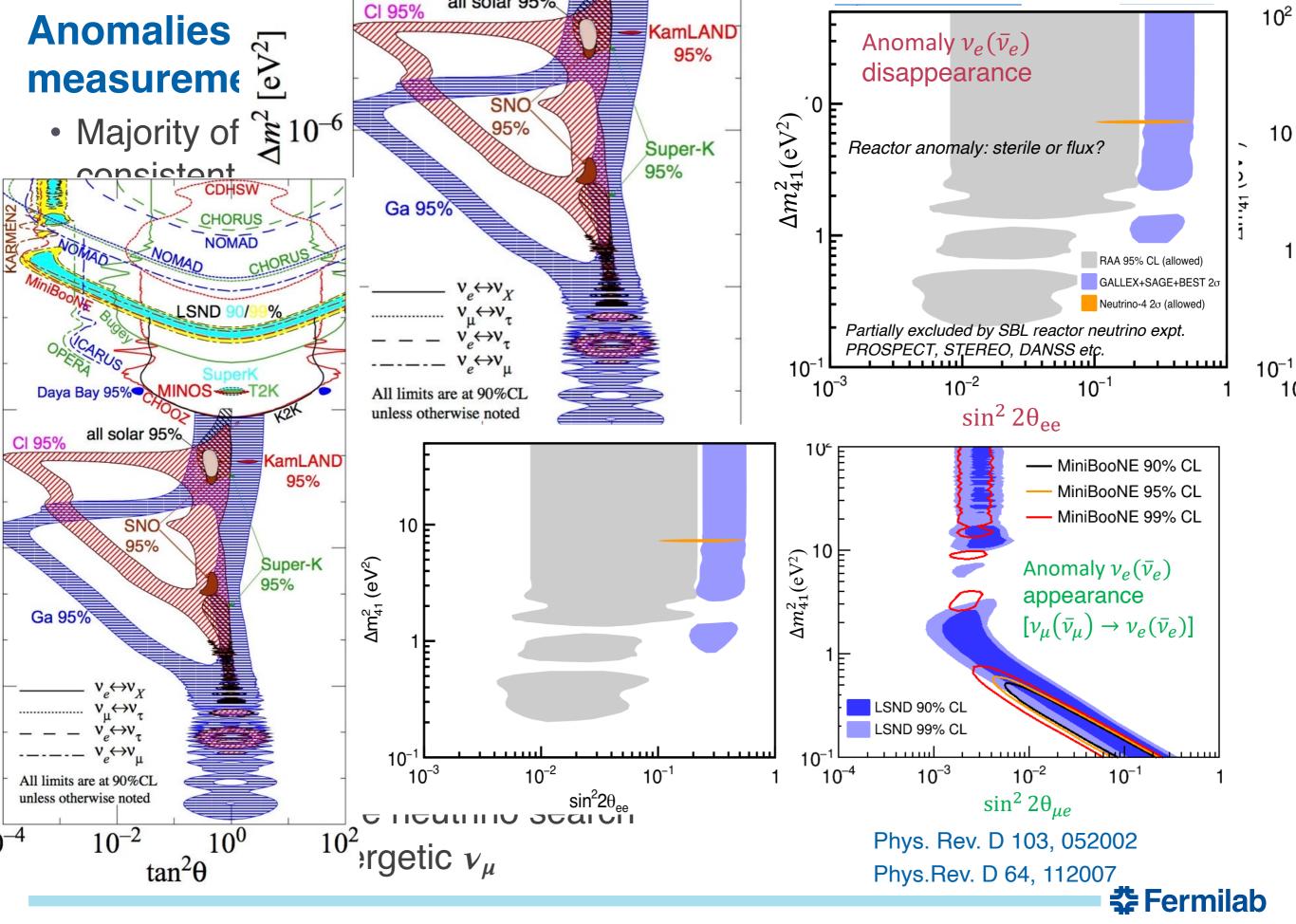


Extension to 3+1 neutrino states

- Can create extensions to the threeflavor model
 - Extend the PMNS matrix to include a fourth, "sterile" neutrino or 3+1 model
- Additional mixing angles and mass splittings based on the fourth neutrino state
- Neutrino fluxes are conserved under the extension
- The fourth neutrino state allows for additional oscillation possibilities and additional appearance/ disappearance measurements







Summary

- Portals to a dark sector enable searches for new physics
- Completion of PIP-II will support initial 1.2 MW beam to LBNF
 - Further upgrades in the form of an accumulator ring could produce a stopped-pion neutrino source at Fermilab on par with the most powerful in the world
- Stopped-pion sources provide access to a host of physics opportunities such as searches for the dark sector and opportunities using CEvNS
- Large-scale detectors coupled to these sources will also be able to perform searches for MeV-scale dark sector signatures



Thank you!

Questions, other ideas?



Backup



Coherent Elastic Neutring-Nucleus Scattering CEvNS)

5803, 2010

n دی

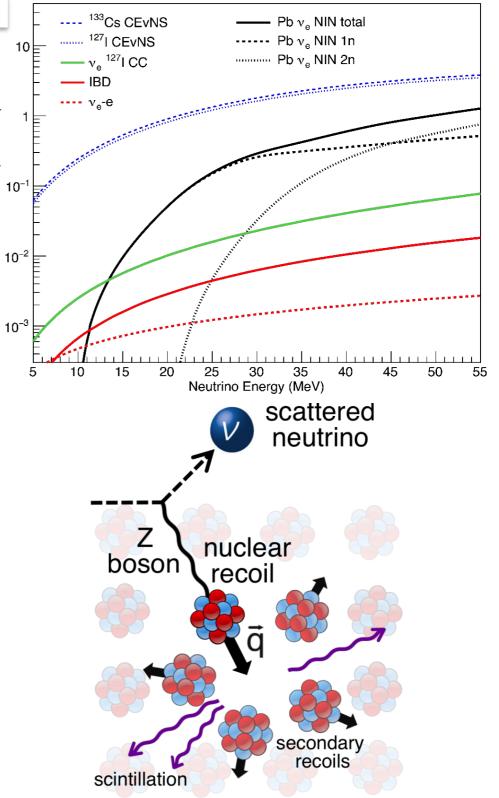
- BRN **CEvNS** St Strobe Standarde Model interaction 80 ± 124 72 ± 7 413 144
- 86 ± 10 264
- 924 Neutrino interacts coherently with 86 ± 10 ction (10⁻³⁶
- for the full data set counting experiment. The statistical fluctuations. The errors on BRN
- Neupritation all the contract of the errors of the error
- Signature is low-energy nuclear recoil
- Largest low-0 $\frac{\rm section \ on \ h}{\rm or \ on \ c} \sim$
- section

$$\frac{d\sigma}{d\Omega} = \frac{G_f^2}{16\pi^2} (N - (1 - 4\sin^2(\theta_W)) Z_{E_r^{max}}^2 E_{\nu}^2 (1 + \cos\theta) F(Q^2)$$

 $50 \,\mathrm{MeV}$

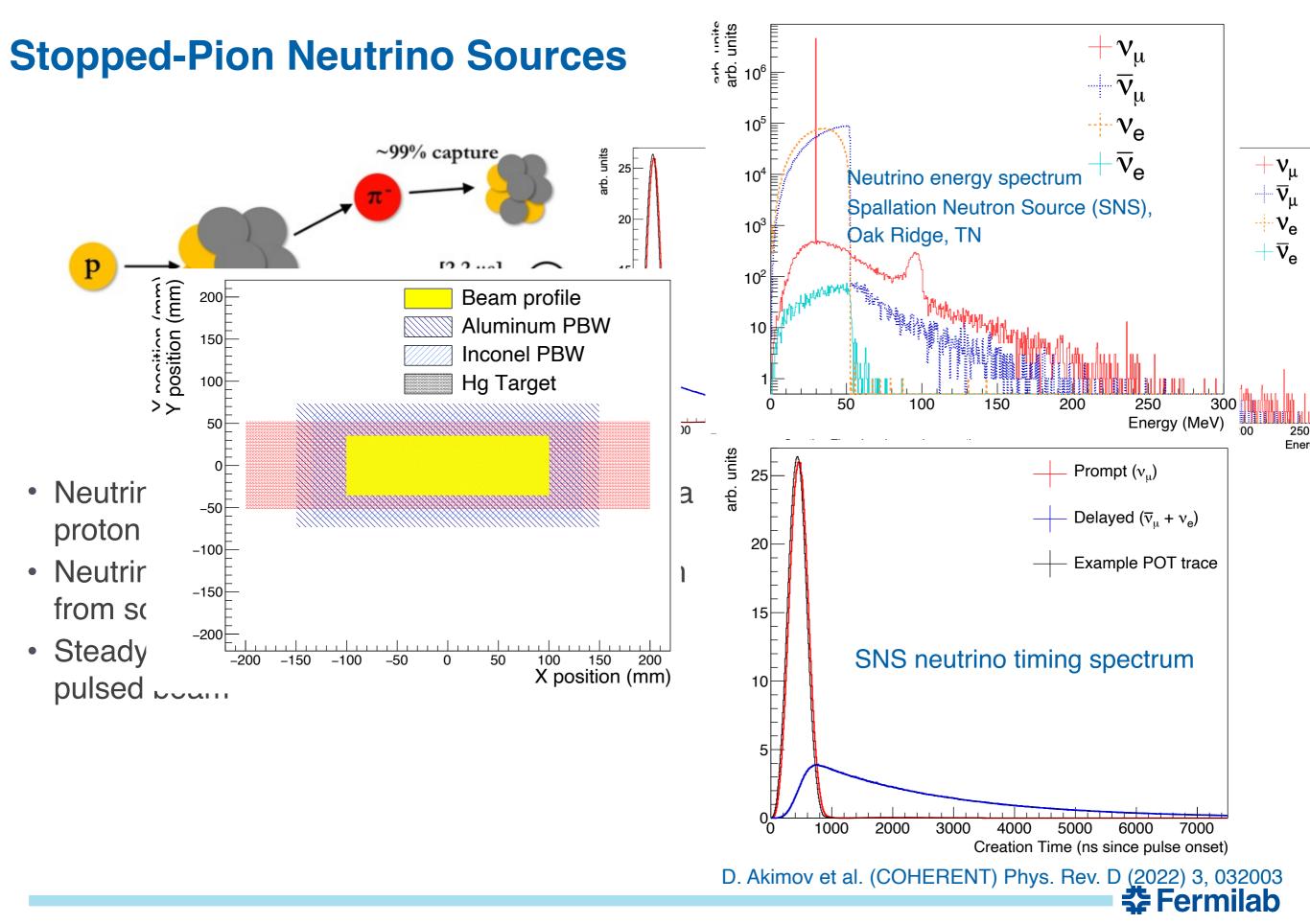
1

 Searches ongoing using both stopped-pion and reactor neutrino sources



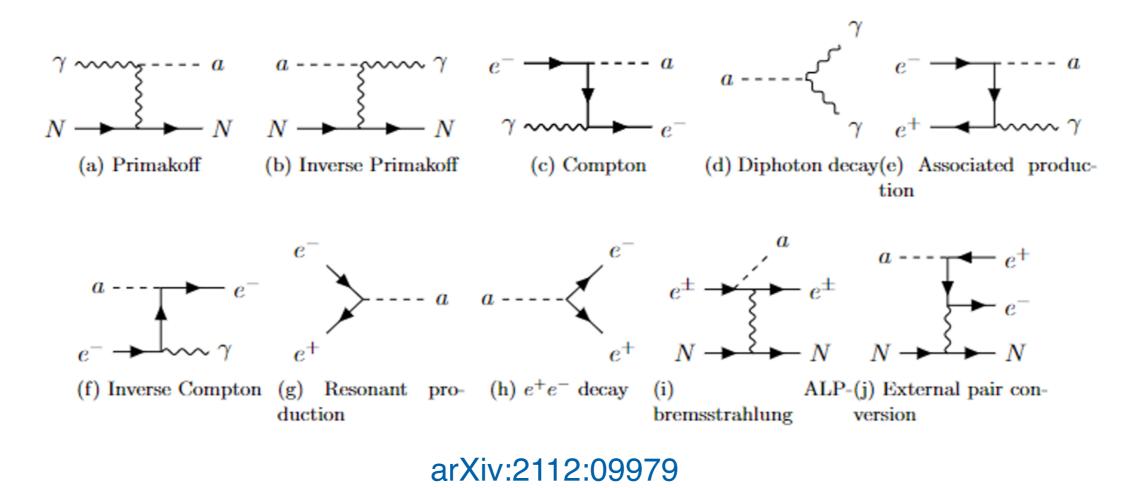
D. Akimov et al. (COHERENT). Science 357, 1123-1126 (2017)





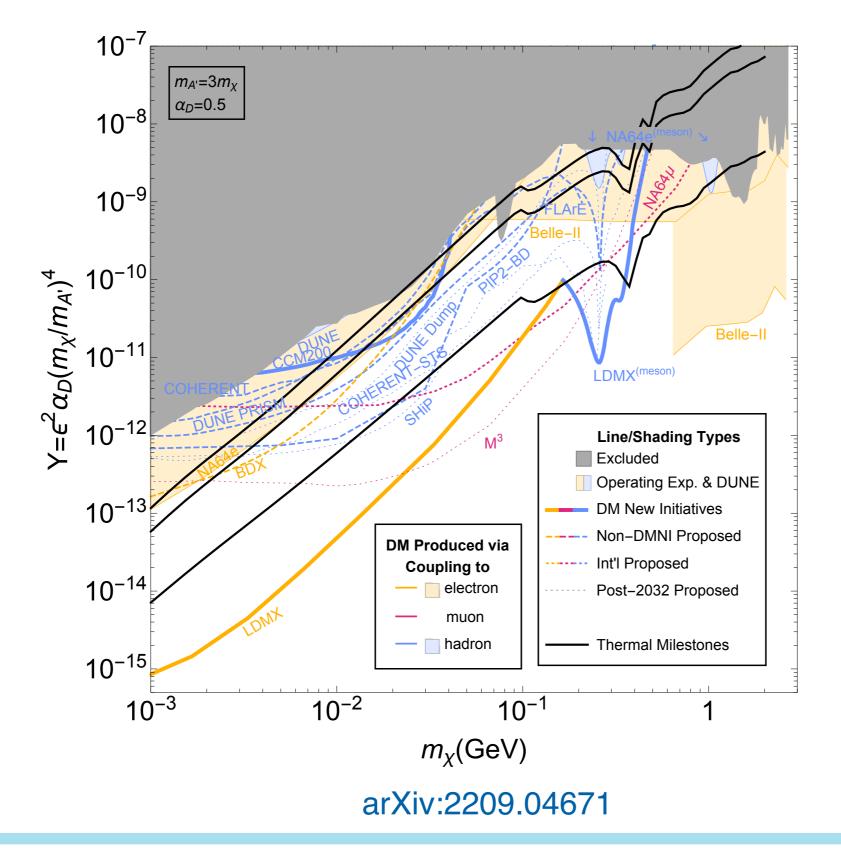
Axion-like particle (ALP) searches with PIP-II

- ALPs that couple to photons can be produced in the beam dump via Primakoff process, detectable via inverse Primakoff process or decay into two photons
- ALPs coupling to electrons detectable via inverse Compton, e+e- conversion, or decay to e+e-



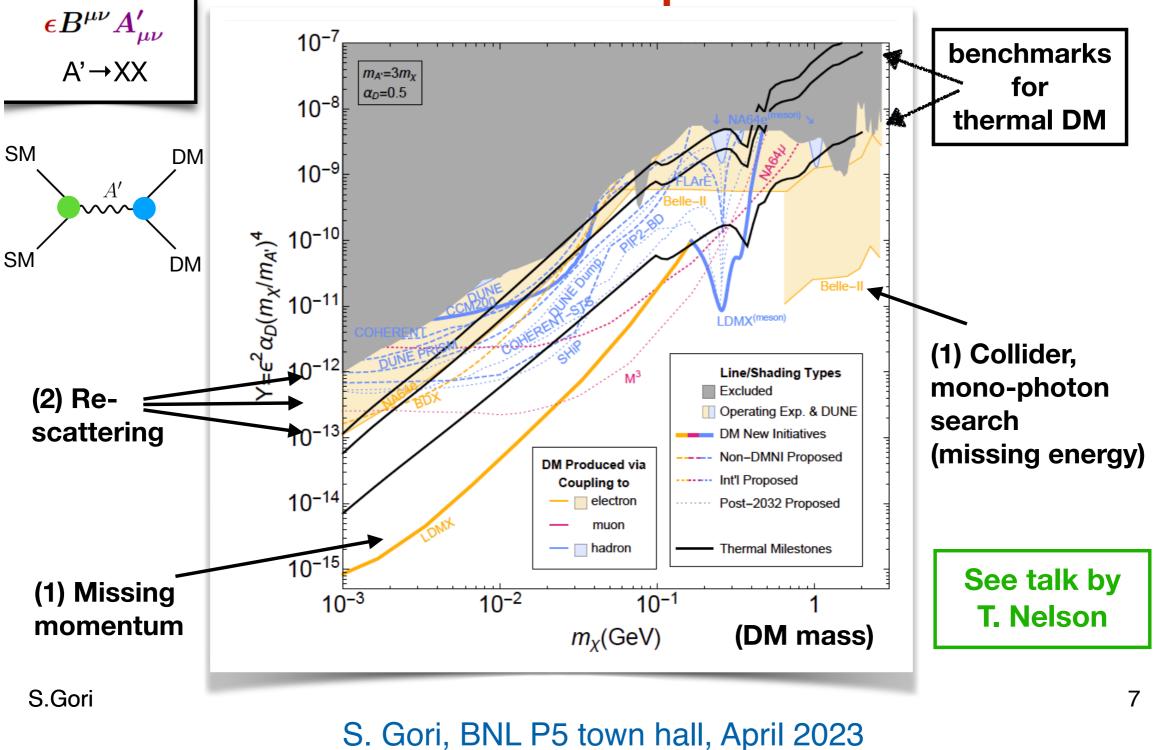


PIP2-BD Vector Portal Dark Matter Search





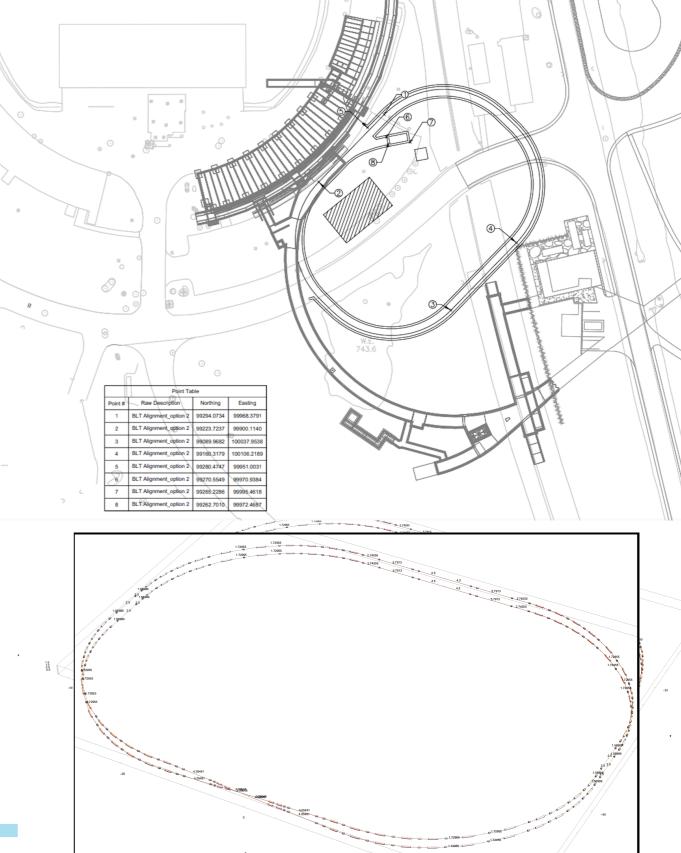
DM thermal milestones: invisible dark photon





The PIP-II Accumulator Ring (PA

- Design of PIP-II linac includes possibilities for future upgrades
 - CW multi-user mode of operation
 - Increase in beam energy to >= 1 GeV
 - Stub in transfer line to the Booster to provide beam to other users
- An extension of the PIP-II beam transfer line tunnel would allow colocation of an accumulator ring for modest cost that could be realized with this decade
 - Allows for dark sector program
 - Enables injection of 1 GeV beam to the Booster as a pathway to higher LBNF beam power



Fermilal

辈 Fermilab

Beyond PAR: The Fermilab Accelerator Complex Evolution (ACE)

- ACE has two components
 - Upgrades to Main Injector and target station allowing DUNE to achieve results on an accelerated schedule
 - A Booster replacement, which will
 - Provide a robust and reliable platform for the future of the Fermilab accelerator complex
 - Enable the capability of the complex to serve precision experiments and searches for new physics with beams from 1-120 GeV
 - Create the capacity to adapt to new discoveries



Capability Capacity Reliability



🛠 Fermilab

Adapted from B. Fleming, FNAL P5 town hall meeting, March 2023

FIT CEVINS events Liquid Argon (LAr) for CEvNS-based new physics detection

- Large scintillation yield of 40 photons/keV
- Well-measured quenching factor
 - Conversion between nuclear recoil response and scintillation response
- Strong pulse-shape discrimination (PSD) capabilities for electron/nuclear recoil separation
- First CEvNS detection on argon at $>3\sigma$ significance by COHERENT!
- Move toward precision physics and new physics searches with large detectors Recoil Energy (keVnr)

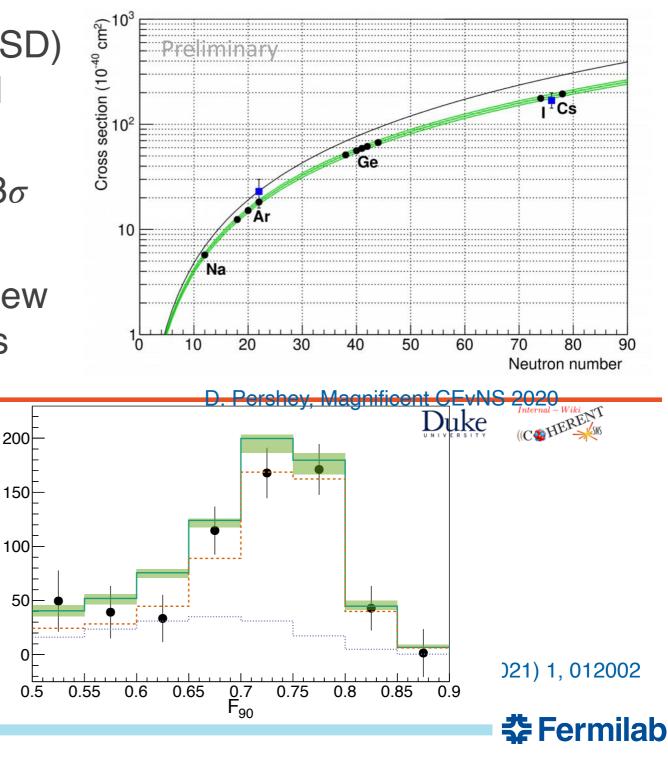
A0 60 80 1 Reconstructed Energy (keVee)

100

120

40

20



al

/NS

t. Error

200

150

100

50

0

0