

# The $\nu$ BDX-DRIFT Experiment to Study CE $\nu$ NS and New Physics at Fermilab

Dan Snowden-Ifft / Occidental College  
May 11, 2023

# The $\nu$ BDX-DRIFT Collaboration



**Occidental College**  
Dan Snowden-Ifft  
Jean-Luc Gauvreau



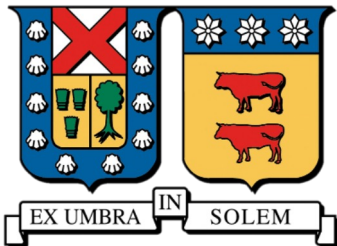
**University of New Mexico**  
Dinesh Loomba



**Texas A&M University**  
Louis Strigari  
Bhaskar Dutta  
Doojin Kim



**INFN-GE**  
**Genoa University**  
Marco Battaglieri  
R. DE Vita  
M. Bondi  
M. Spreafico



**Universidad Technica Federico Santa Maria**  
Diego Aristizabal Sierra



**Canisius College**  
Mike Wood



**INFN-CT**  
**University of Messina**  
M. de Napoli  
G. Mandaglio  
S. Sgrazzi  
A. Fulci  
A. Pilloni



**INFN-RM2**  
**University of Roma Tor Vergata**  
A. D'Angelo

**Zuckerman Post-Doctoral Scholar**

J. Barrow

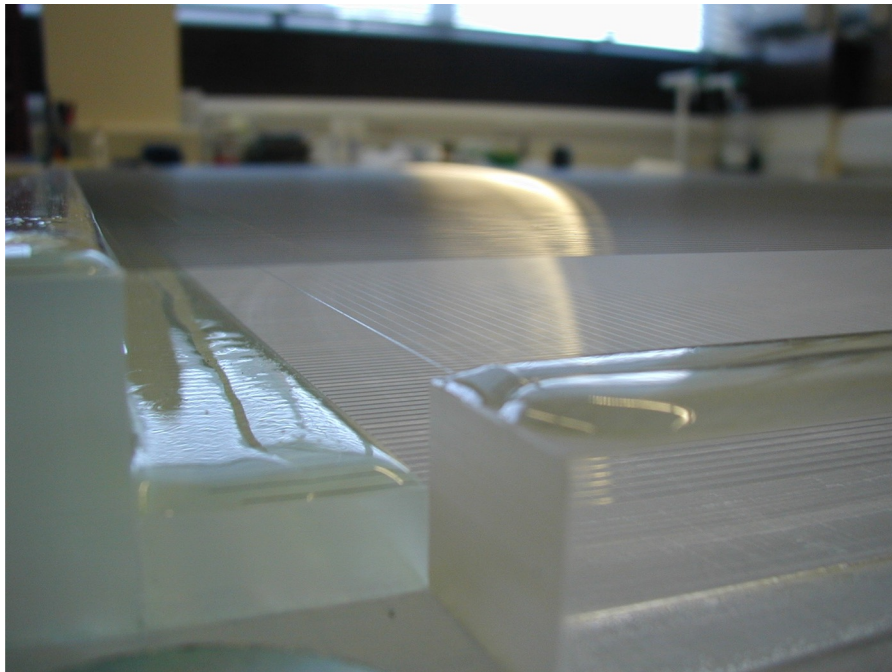


# DRIFT: Lightning Summary

Started = 1998, US/UK

Directional WIMP dark matter  
detector

40 Torr, 1 m<sup>3</sup> gaseous detector



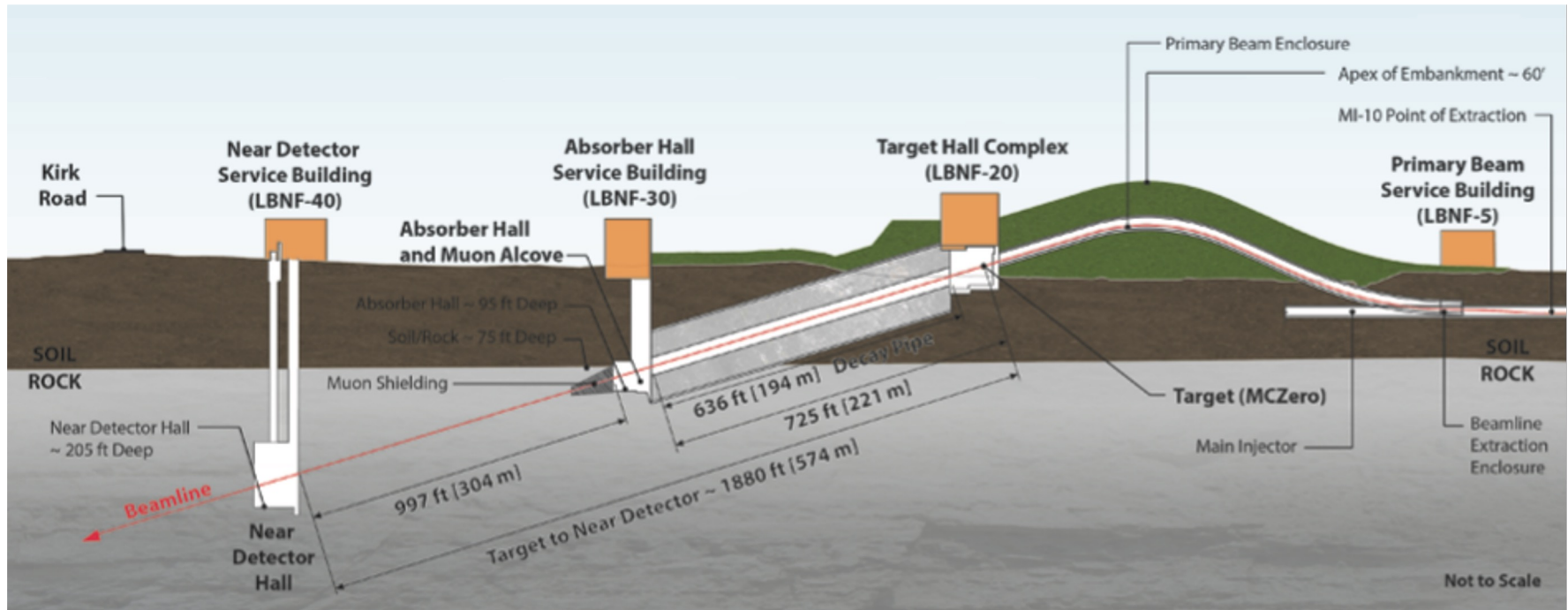
Unique and robust technology

Low energy (35 keV) threshold for  
nuclear recoils

Low background

AstroPle, 91, 2017

# $\nu$ BDX-DRIFT at DUNE



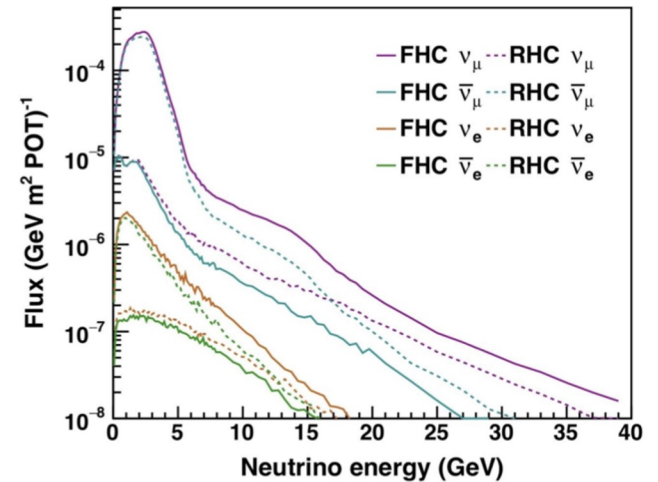
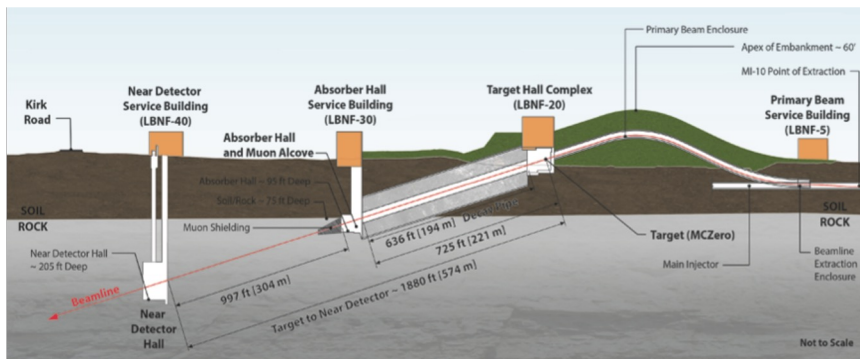
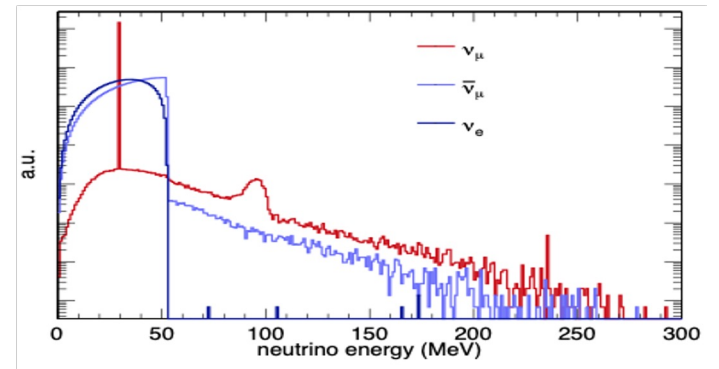
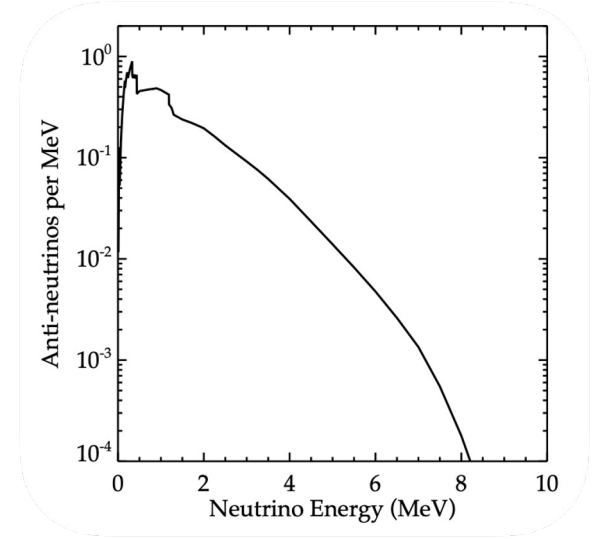
# Sources for CE $\nu$ NS



**COHERENT**

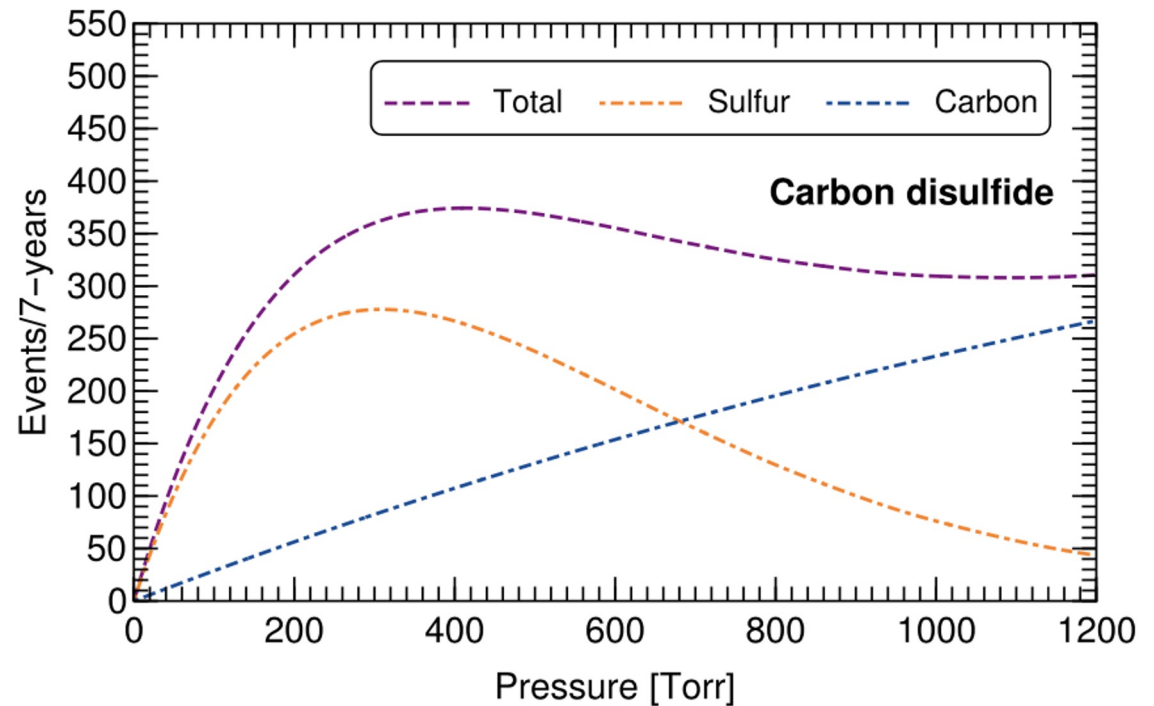


**Spallation Neutron Source**

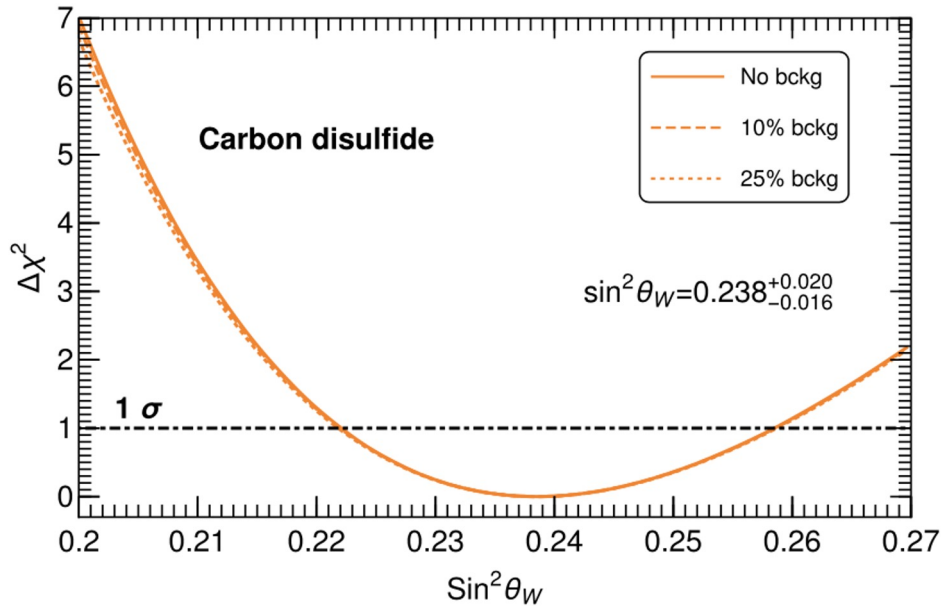


# CE $\nu$ NS Physics – Event Rates

- Recoil energy threshold  
proportional to pressure
- Max rate @ 400 Torr CS<sub>2</sub>
- Significant detection, ~400  
events for 7 yrs with 10 m<sup>3</sup>
- Different nuclei possible
- Directionality for signature

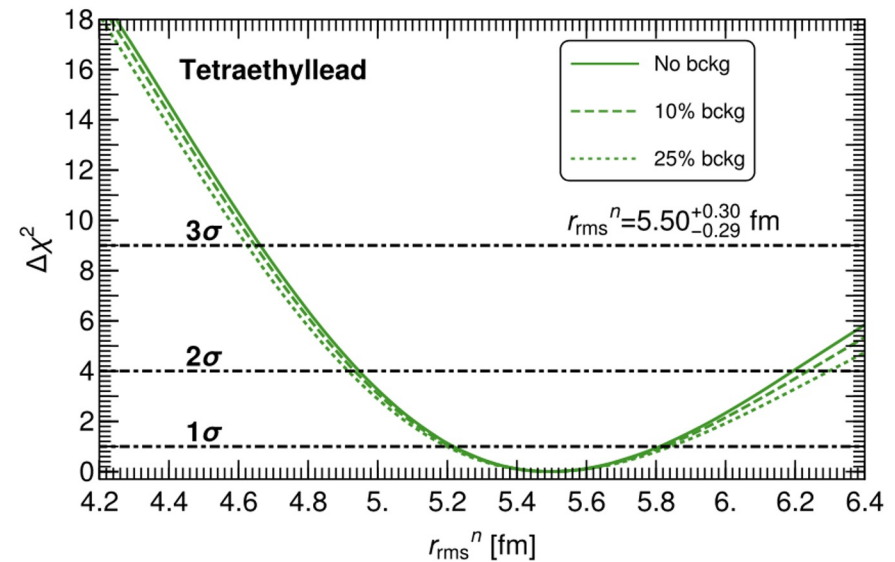


# CE $\nu$ NS Physics - Measurements



- Measurement of the weak mixing angle on S to 8%

- Measurement neutron distribution in Pb to 5% and skin
- Larger neutrino energy => Sufficient stats at high Q
- Systematics on form factor are small
- Comparison with PREX



## Other Neutrino – Nucleus Events

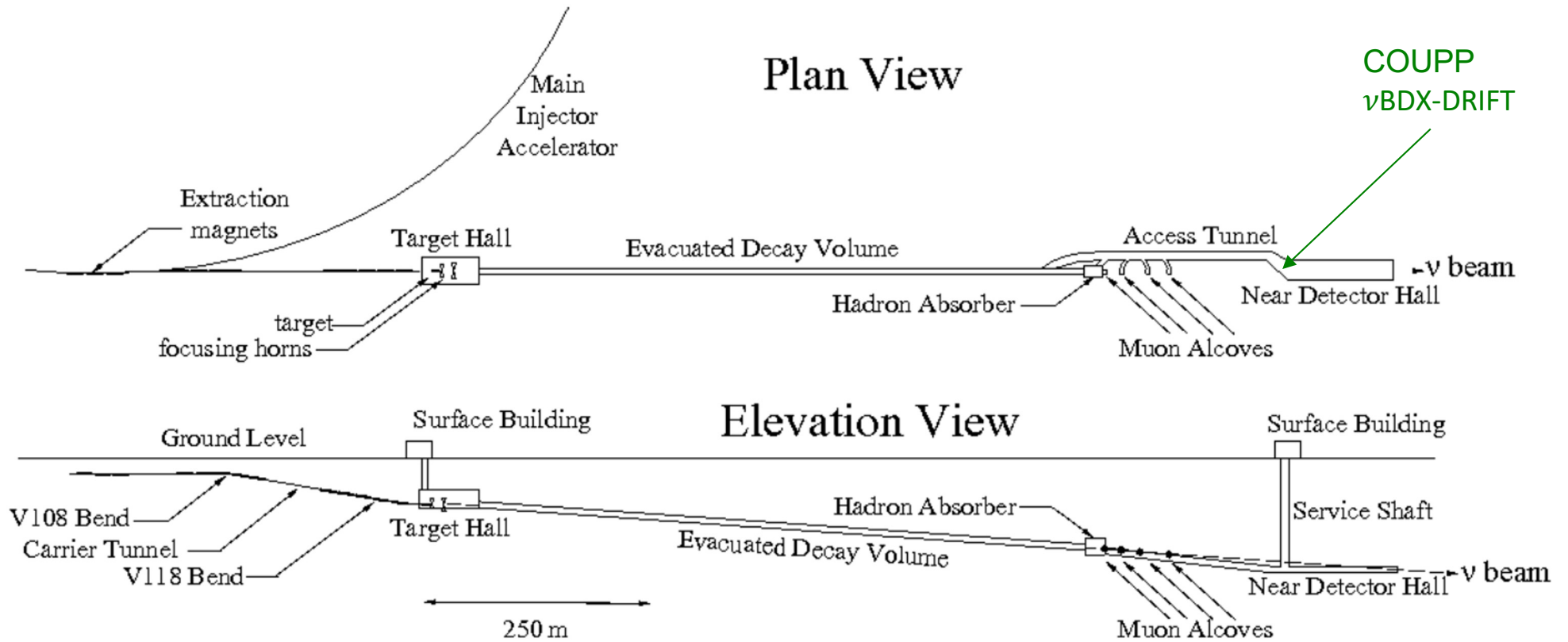
- $\nu$ BDX-DRIFT can image more than the nuclear recoils
- i.e. non-CE $\nu$ NS events are also targets for investigation
- i.e. Neutrino – Nucleus interactions at high momentum transfer
- i.e. “inelastic” events (quasielastic reactions, others possible)
- Study kinematics of nuclear remnants, multidifferential xsecs
- Can be compared to GENIE, other generators—constrain FSIs



# Beyond the Standard Model physics

- Low background and directional sensitivities can be utilized to investigate BSM physics
- Light dark-matter and axion-like particles
- Dark matter and neutrino inelastic up-scattering processes including directionality: primary recoil + secondary (displaced) decay
- Various types of dark matter and neutrino interactions
- Decays of various types of light mediators
- Complementarity with DUNE

# $\nu$ BDX-DRIFT at NuMI



# Backgrounds - Benchmarking

- Used COUPP 2009 nuclear recoil data to benchmark the simulation

2009 Run Event Rates (per Kg-day) vs. Time (NuMI Tag, No Veto Tag)

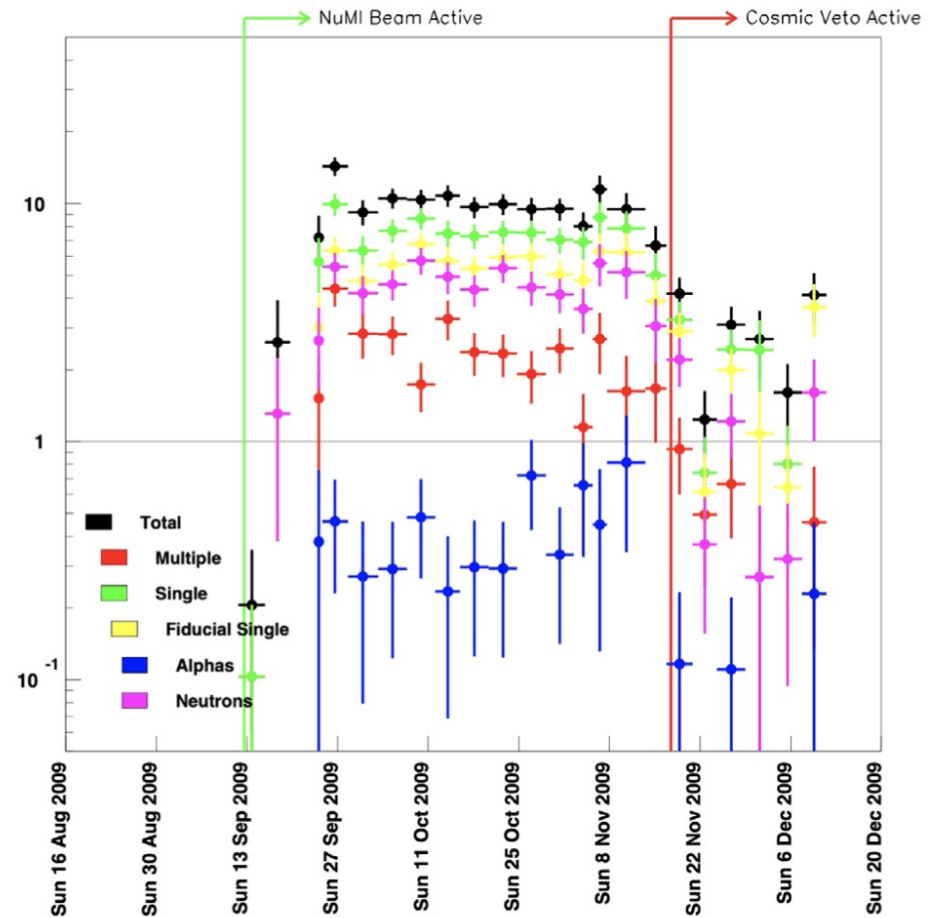


FIG. 1. The 2009 COUPP bubble formation data tagged to the beam pulse. Published here with the permission of the COUPP Collaboration.

# Backgrounds - Benchmarking

- Used COUPP 2009 nuclear recoil data to benchmark the simulation
- COUPP (unpublished) measured  $4.56 \pm 0.19$  events/kg/day

2009 Run Event Rates (per Kg-day) vs. Time (NuMI Tag, No Veto Tag)

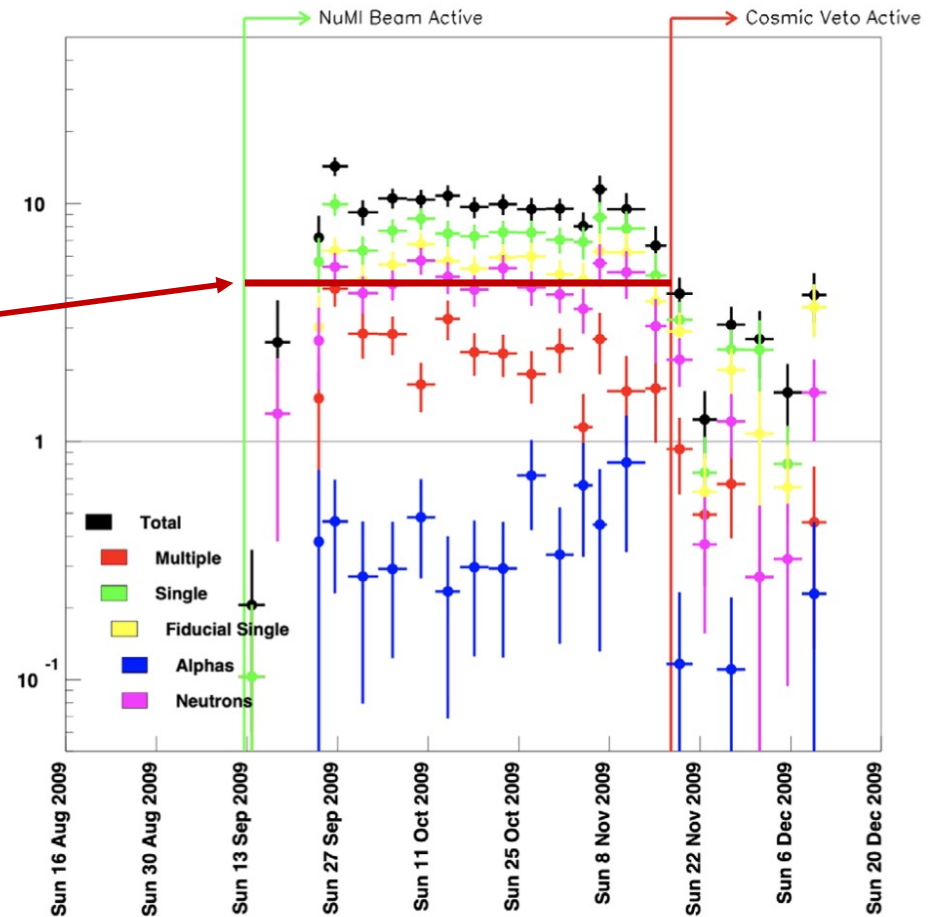


FIG. 1. The 2009 COUPP bubble formation data tagged to the beam pulse. Published here with the permission of the COUPP Collaboration.

# Backgrounds - Benchmarking

- Used COUPP 2009 nuclear recoil data to benchmark the simulation
- COUPP (unpublished) measured  $4.56 \pm 0.19$  events/kg/day
- Genie/GEANT predict  $(3.006, 3.356) \pm 0.023$  events/kg/day
- 30% low

2009 Run Event Rates (per Kg-day) vs. Time (NuMI Tag, No Veto Tag)

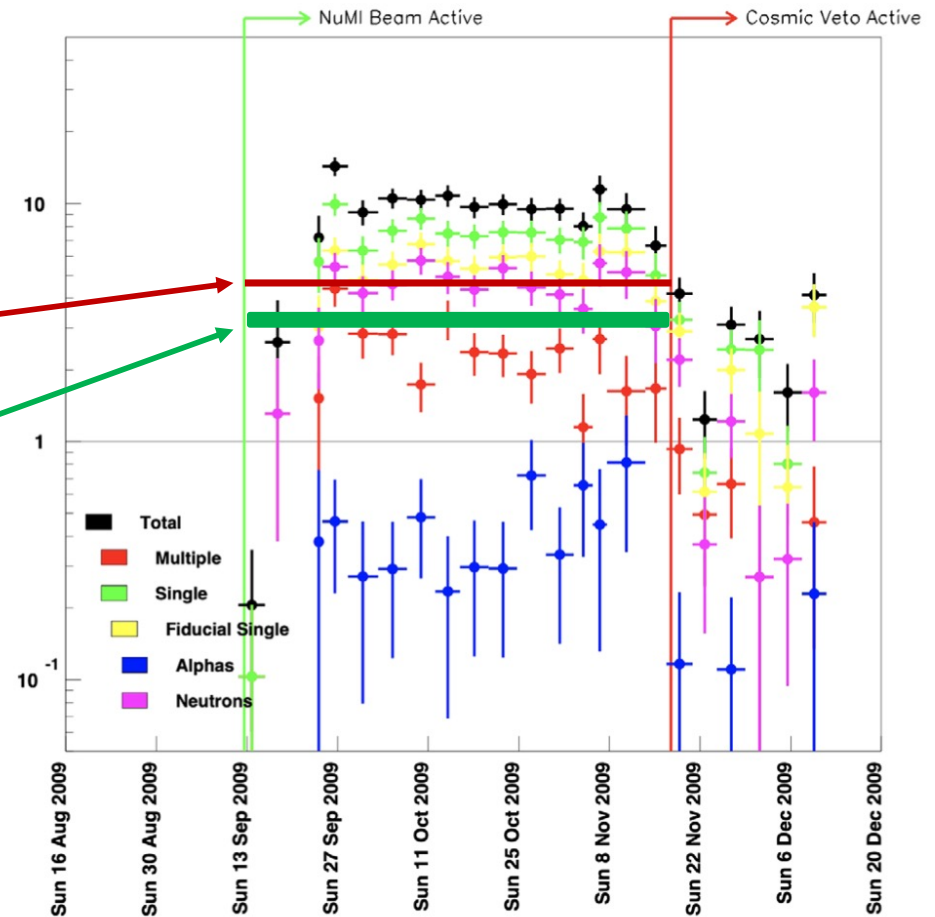


FIG. 1. The 2009 COUPP bubble formation data tagged to the beam pulse. Published here with the permission of the COUPP Collaboration.

# Backgrounds - Predicted

TABLE IV. Output table shows neutron flux from different walls and background in the signal region. For details see Sec. IV.

Simulations output for neutron flux from the walls				
Beamline and mode	Upstream [ $n^0/s/m^2$ ]	Sides [ $n^0/s/m^2$ ]	Downstream [ $n^0/s/m^2$ ]	Background [events/ $m^3/year$ ]
NuMI LE	0.0355	0.0204	0.0110	$8.61 \pm 0.62$
NuMI HE	0.209	0.131	0.0727	$54.9 \pm 3.8$
DUNE on-axis at 1.2 MW	0.101	0.0276	0.0524	$23.3 \pm 1.3$
DUNE 39 m off-axis at 1.2 MW	0.000381	0.0000831	0.000162	$0.0396 \pm 0.0031$

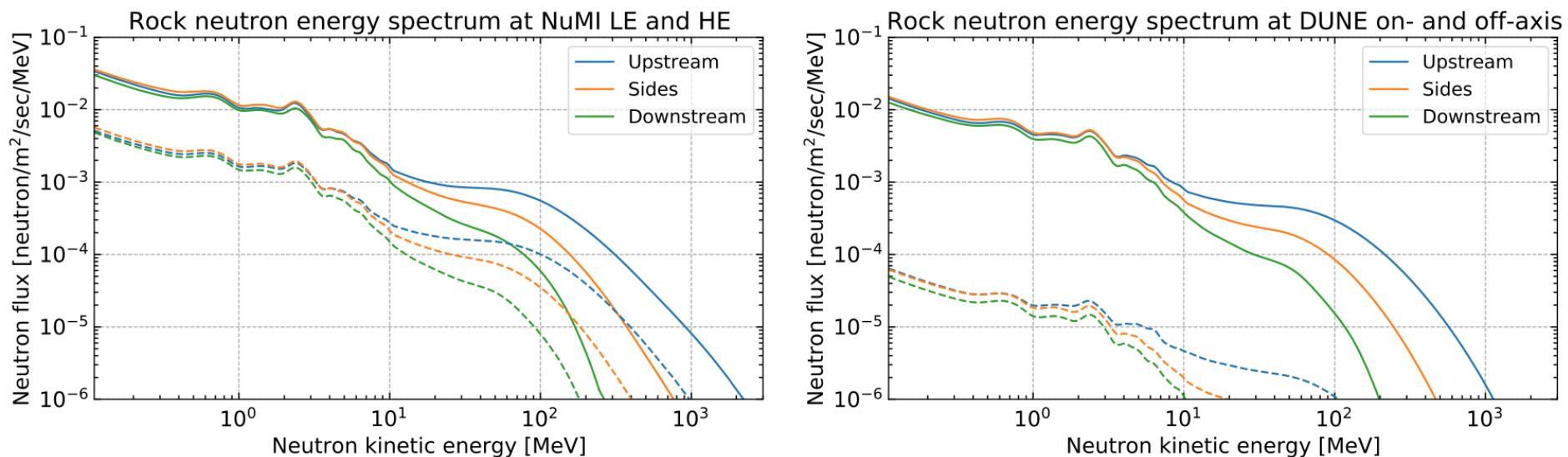


FIG. 6. The energy distribution of rock-neutrons generated in the four simulations of Table I. The blue lines show the spectra coming from the upstream wall. The orange lines show the spectra coming from the side walls. And the green lines show the spectra coming from the downstream wall. In the left graph, the solid (dashed) curves correspond to results obtained with the NuMI HE (LE) neutrino mode. In the right graph—instead—to results derived with the DUNE on-axis (off-axis) configuration.

# Backgrounds - Predicted

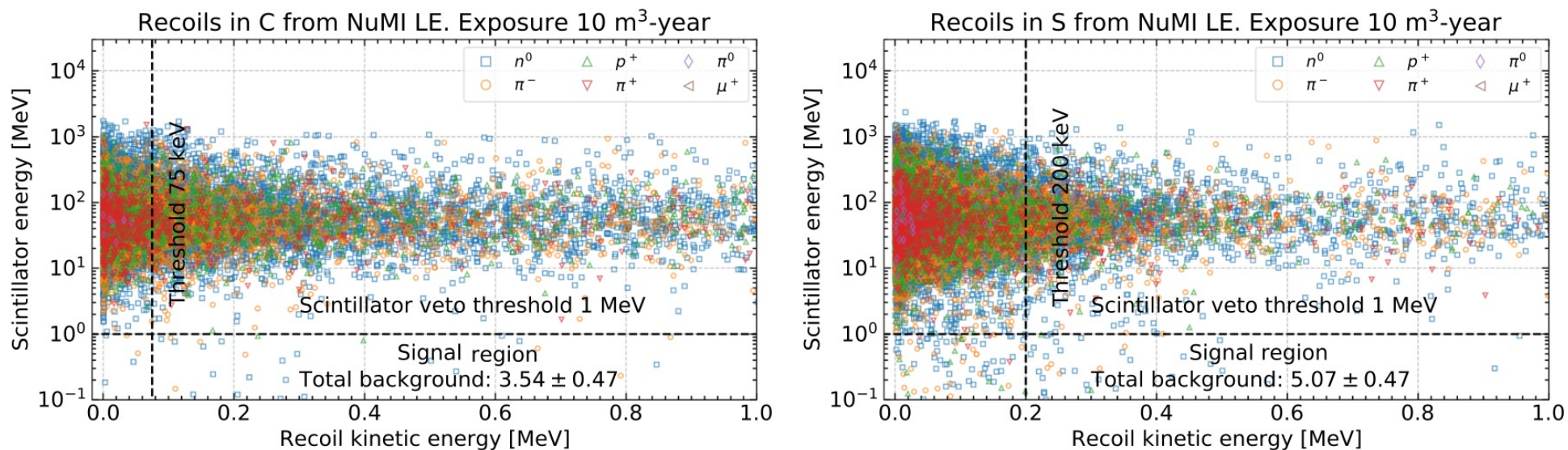
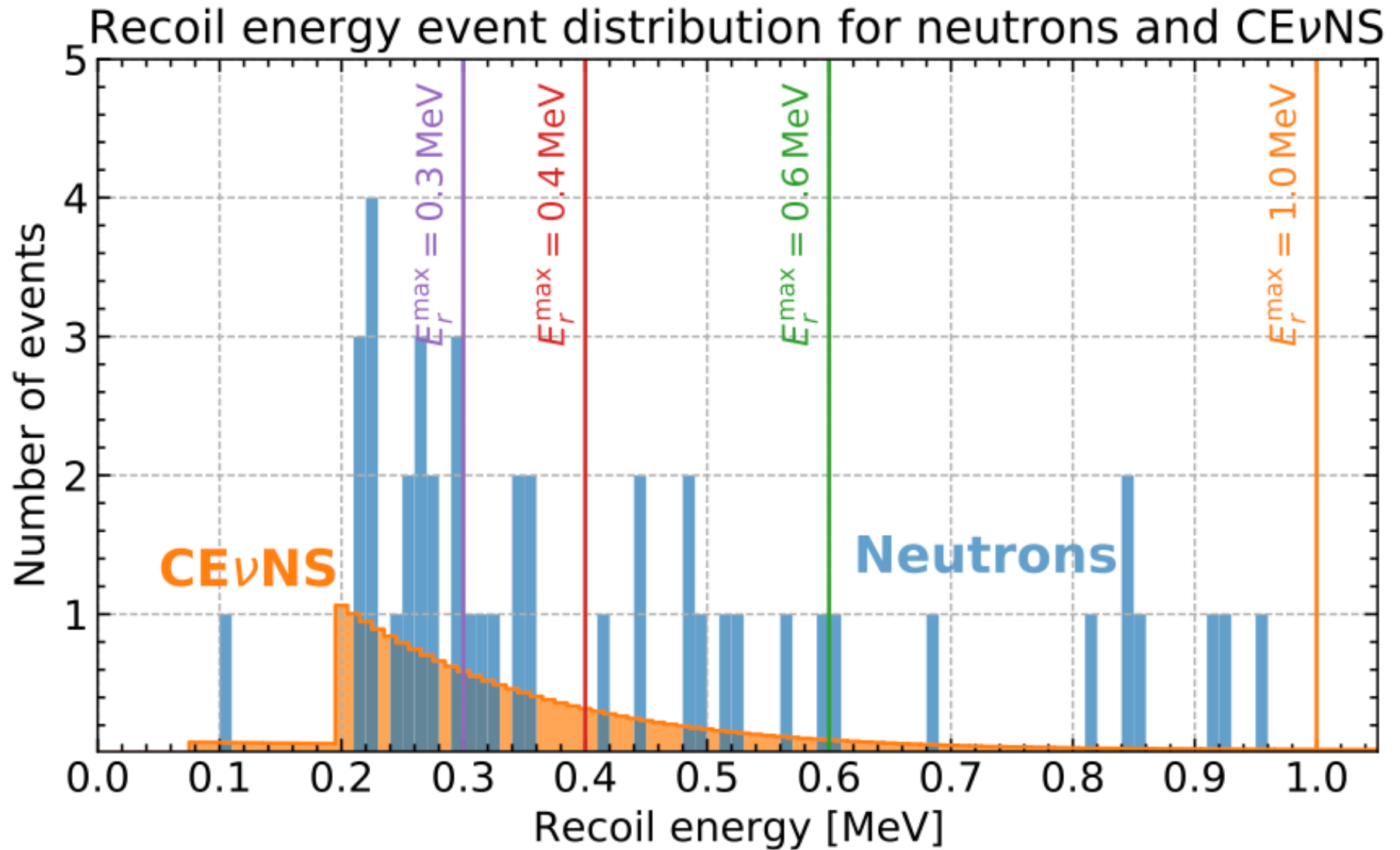


FIG. 8. Plots showing the distribution of recoil energies vs energy deposited in the scintillator with the neutrino-induced end-state particle responsible for the recoil shown in different colors. The left graph shows the results for C recoils while the right graph shows the S recoils. Both are heavily dominated by neutron end-states (about 63% for both target nuclei). The vertical dashed black lines indicate the recoil thresholds, 75 keV for C and 200 keV for S. The horizontal dashed black lines show the threshold for the scintillator veto, 1 MeV; events with larger energies are vetoed. The lower-right region therefore shows the signal region where either CE $\nu$ NS or BSM recoils events would occur. The background rate, in events per m<sup>3</sup> per year, are shown there. The sum is shown in the fifth column of Table IV.

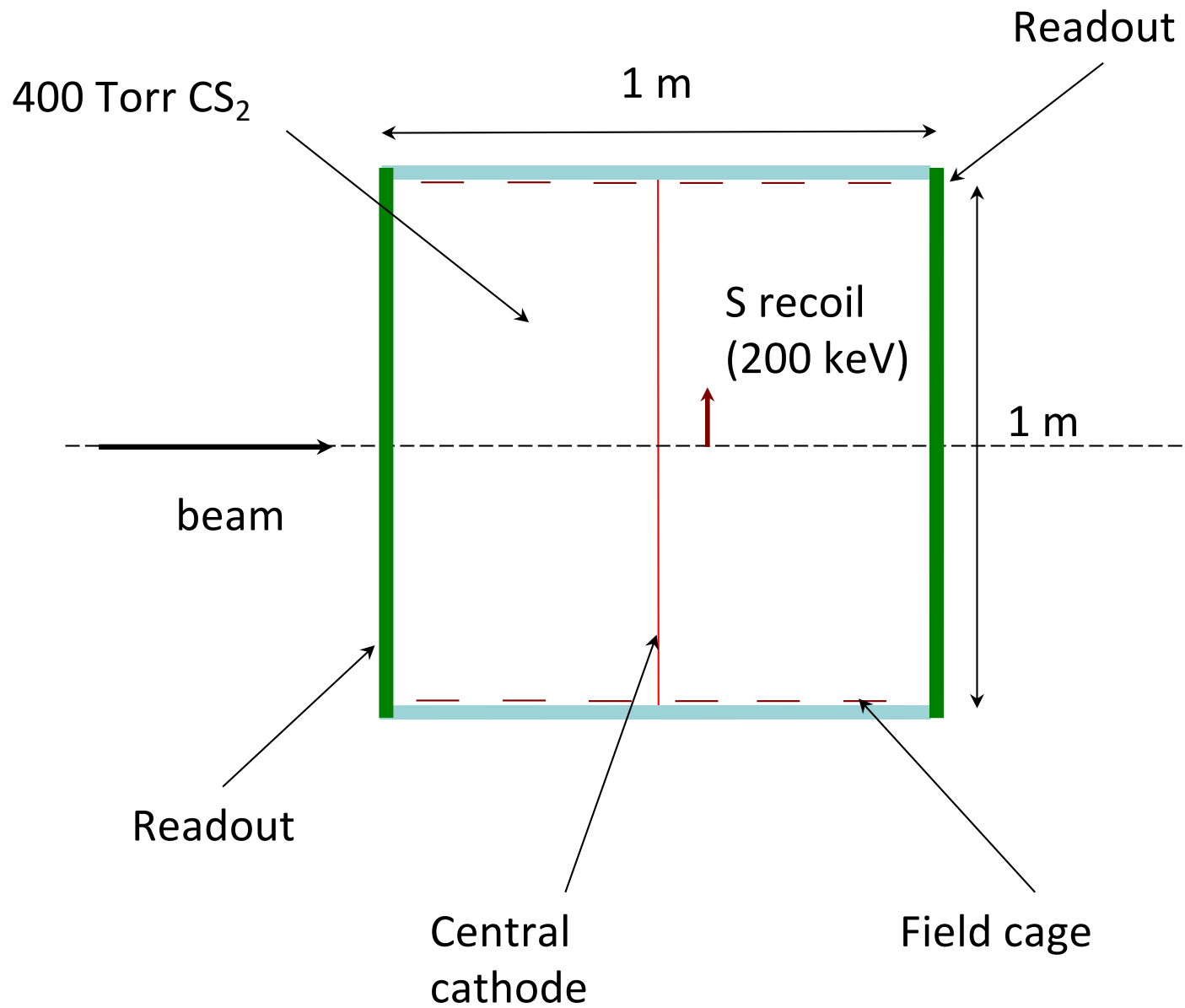
# Backgrounds – Predicted

10 m<sup>3</sup> year exposure to NuMI LE



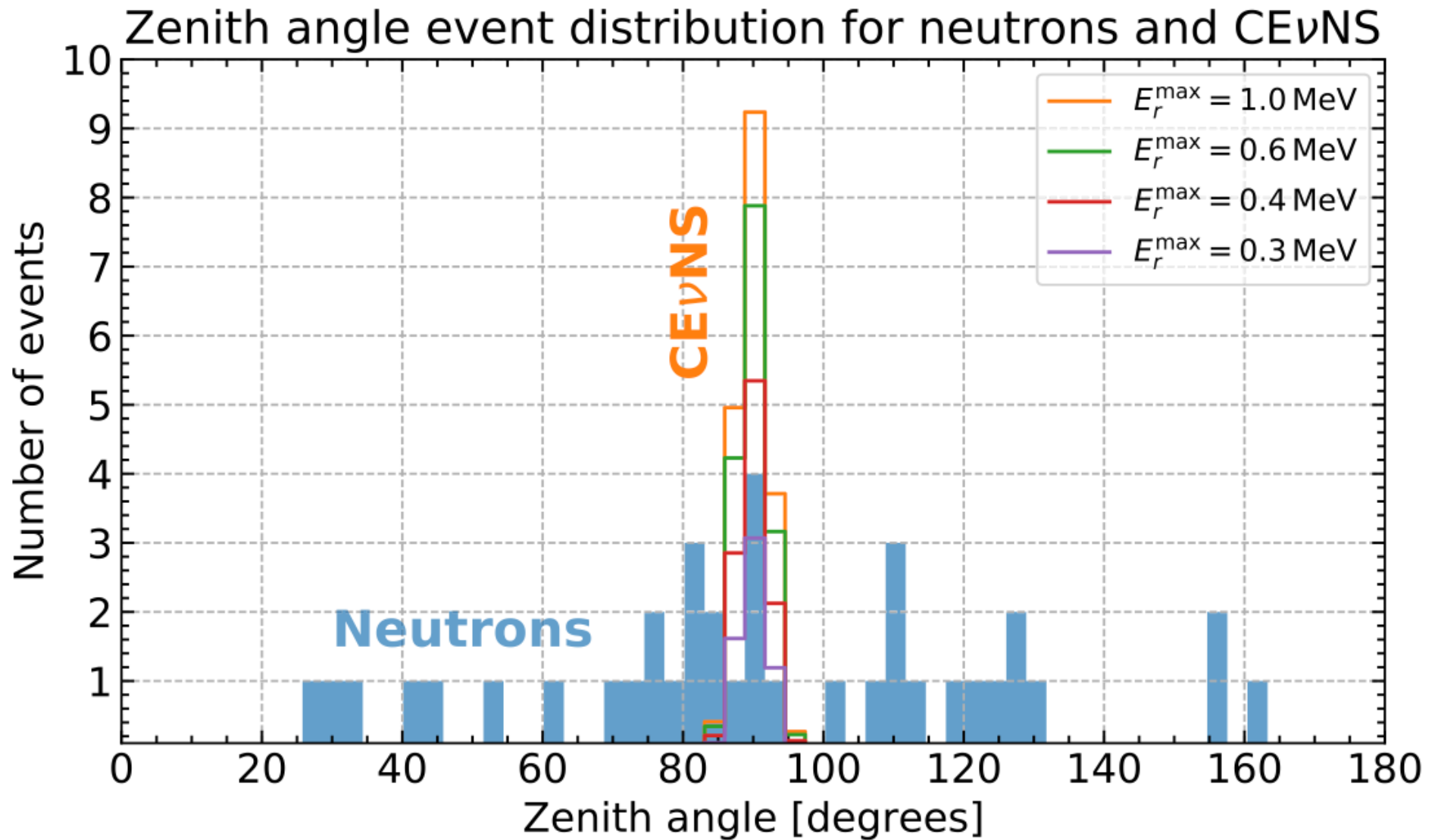


# Directionality



# Backgrounds - Predicted

10 m<sup>3</sup> year exposure to NuMI LE



# Conclusion

- $\nu$ BDX-DRIFT brings a unique, proven, halo-dark-matter detector to CE $\nu$ NS research
- A 10 m<sup>3</sup>  $\nu$ BDX-DRIFT detector in the Near Detector hall in DUNE could detect 400 CE $\nu$ NS events in a 7-year run
- $\nu$ BDX-DRIFT offers interesting new capabilities and complementarity for CE $\nu$ NS research
- Measurements of WMA and n distribution available
- NSI, BSM and new physics
- Backgrounds are expected to be under control
- In the near term we hope to deploy an existing 1 m<sup>3</sup>  $\nu$ BDX-DRIFT detector in the NuMI beam at Fermilab to test these ideas out