Sterile Neutrino Oscillation Searches using VALOR at SBN

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Overview



- Sterile Neutrinos
- The SBN Programme
- What is VALOR
- Current Sensitivities
- SBND-PRISM

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Sterile Neutrinos

SBN Program

- Limit of 3 active flavours from the Z-boson resonance width and cosmological data
 - 4th active flavour ruled out at a 98% confidence level by ALEPH in 1989^[1]
 - Cosmological limits give $N_{eff} = 3.32 \pm 0.27 (68\% \text{ CL})^{[2]}$
- Well motivated, could be a consequence of neutrino mass
- Experimental anomalies may hint at a fourth neutrino
 - Reactor: deficit of anti- ν_{e} flux at short baseline
 - Gallium: deficit of v_e flux from Ar-37 and Cr-51 electron capture decays.
 - Accelerator (LSND and MiniBooNE): excess of v_e flux from $v_{\mu} \rightarrow v_e$
- Tensions between these results and disappearance analyses ^[3]
- 3 active + 1 sterile is benchmark hypothesis (excluded to high significance)
- Test existence via mixing with active flavours

[1] ALEPH, D. Decamp et al., Determination of the Number of Light Neutrino Species, Phys. Lett. B 231, 519 (1989)

[2]On the behalf of the Planck Collaboration. Cosmological constraints on neutrinos with Planck data. In Boston, Massachusetts, USA; 2015 p. 140001.

[3]Dasgupta B, Kopp J. Sterile Neutrinos. Physics Reports. 2021 Sep;928:1–63.



Short Baseline Neutrino (SBN) Programme

- 3 liquid argon TPC detectors along the Booster neutrino beam
 - Similar technology and same beam allows for systematic constraints
- Physics aims:

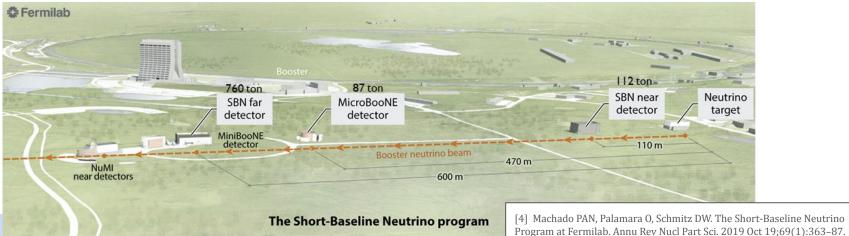
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 $\circ~$ Simultaneous measurements of $\nu_{\mu} disappearance$ and ν_{e} (dis)appearance to search for sterile neutrino oscillations: $\Delta m^{2}_{41} \sim 1 eV^{2}$

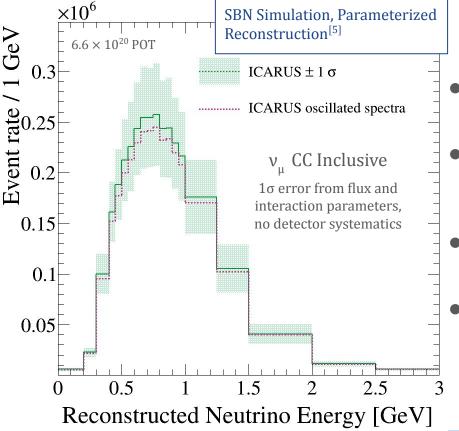
SBN

Program

- $\circ \quad {\rm Studying\ neutrino-argon\ interactions}$
- BSM searches ^[4]



Role of SBND in the SBN Programme



- SBN will definitively test the parameter space favoured by previous measurements
- Our predictions have uncertainties $\sim 30\%$
 - Too large to search for new physics
 - \circ Need to reduce to ~1%
- The role of SBND reduces uncertainty to enable new physics searches
- Will need powerful analysis framework to fully exploit the power of SBND samples

[5] Acciarri R, et al. A Proposal for a Three Detector Short-Baseline Neutrino Oscillation Program in the Fermilab Booster Neutrino Beam arXiv; 2015





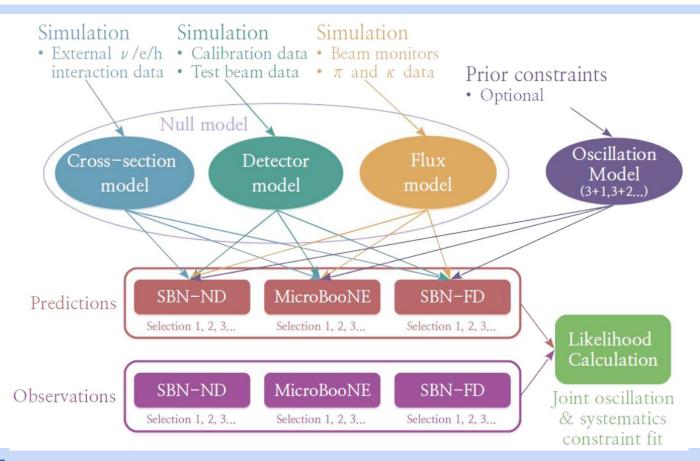


- Well established and tested neutrino fitting framework ^[6]
- Developed within T2K and used for many published results
- Can perform standalone (single oscillation channel) or joint multi-channel analyses
- VALOR can fit multiple different inclusive or exclusive samples for all detectors
 - Complementary information from different samples helps solve the degeneracies between systematic effects and/or new physics
- VALOR simultaneously fits for oscillation and systematic parameters
 - Provides explicit constraints on systematics

^[6] VALOR Neutrino Fit [Internet]. hep.ph.liv.ac.uk. Available from: https://hep.ph.liv.ac.uk/~costasa/valor/



VALOR: Analysis Strategy

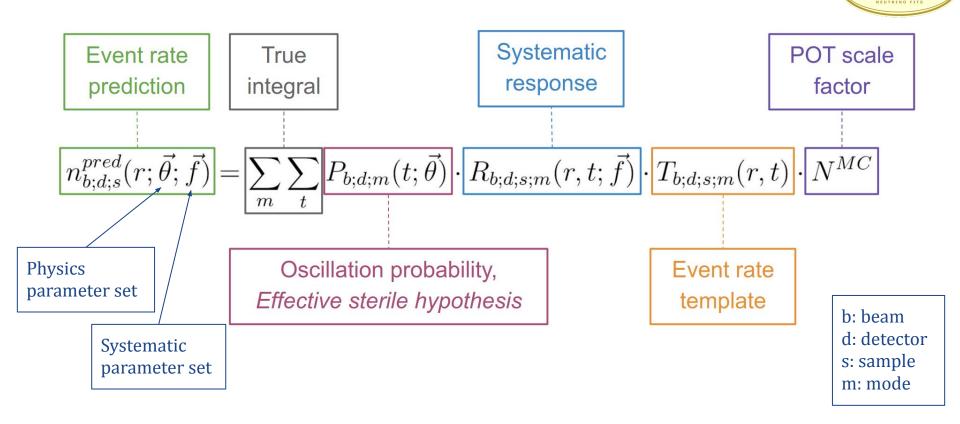




- The analysis strategy uses indirect extrapolation
- Event rate model is a convolution of flux, xsec, and detector models
- Explicit systematic constraints



Event Rate Prediction







• Main power of VALOR is to obtain explicit systematic constraints postfit

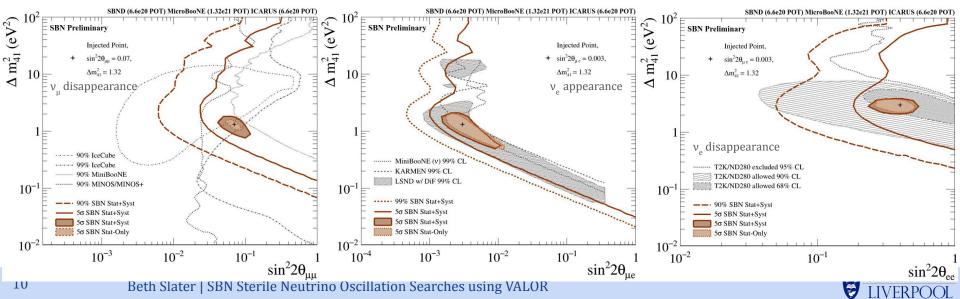
$$n_{b;d;s}^{pred}(r;\vec{\theta};\vec{f}) = \sum_{m} \sum_{t} P_{b;d;m}(t;\vec{\theta}) \cdot R_{b;d;s;m}(r,t;\vec{f}) \cdot T_{b;d;s;m}(r,t) \cdot N^{MC} \qquad \begin{array}{l} \text{b: beam } \\ \text{d: detector } \\ \text{s: sample } \\ \text{m: mode } \end{array}$$

- VALOR has high granularity in parameterising systematic effects
- Systematic parameters are currently eliminated via profiling
 - Option to use marginalisation (used in VALOR-T2K analysis)

Preliminary Sensitivities

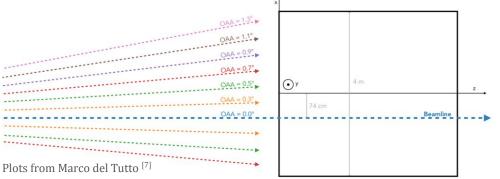


- VALOR has been used within SBN for several years
 - Implemented several oscillation sensitivity analyses
 - \circ Below are the standard sensitivities for the 3 standalone channels
 - Using inclusive samples and pseudo-reconstruction
- Effort is underway to build more sensitive analysis using exclusive samples and exploiting the PRISM effect

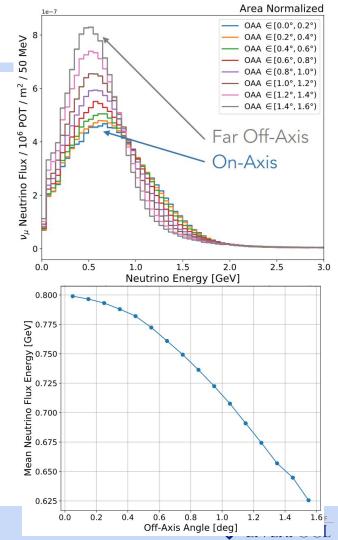


SBND-PRISM

- Takes measurements at different off-axis locations
 - Different energy spectra/composition
- Joint fit of all off-axis samples
 - Improved systematic constraints/degeneracy resolution
 - Enhanced oscillation sensitivity
- SBND split into 8 bins for illustration
 - The statistics in each bin are still large so the systematics dominate







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- SBN programme should improve understanding of sterile hypothesis
- SBND will have excellent statistics as the event rate is high
 - \circ ~ Used to constrain systematic uncertainties
- The use of SBND-PRISM should improve systematic constraints for the whole programme
 - This has been implemented for all 3 oscillation channels and is currently being validated
 - Results are encouraging
 - Work ongoing to find optimal number of off-axis bins and understand improvements to sensitivities
- Many other lines of work within VALOR to incorporate exclusive samples and evaluate uncertainties and biases within mock data





References:



- 1. ALEPH, D. Decamp et al., Determination of the Number of Light Neutrino Species, Phys. Lett. B 231, 519 (1989)
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- 6. VALOR Neutrino Fit [Internet]. hep.ph.liv.ac.uk. [cited 2023 Jun 22]. Available from: https://hep.ph.liv.ac.uk/~costasa/valor/
- Del Tutto M, Machado P, Kelly K, Harnik R. SBND-PRISM: Sampling Multiple Off-Axis Fluxes with the Same Detector. In: SBND-PRISM: Sampling Multiple Off-Axis Fluxes with the Same Detector [Internet]. US DOE; 2021 [cited 2023 Jun 22]. Available from: https://www.osti.gov/servlets/purl/1827399/



Likelihood Calculation

$$\chi_{0}^{2} = -2 \ln \mathscr{L}_{0}(\vec{\theta}; \vec{f}) = 2 \sum_{b,d,s,r} \left(n_{b;d;s}^{data}(r) \cdot \ln \frac{n_{b;d;s}^{data}(r)}{n_{b;d;s}^{pred}(r; \vec{\theta}; \vec{f})} + (n_{b;d;s}^{pred}(r; \vec{\theta}; \vec{f}) - n_{b;d;s}^{data}(r)) \right)$$

$$\chi^{2} = -2 \ln \mathscr{L}(\vec{\theta}; \vec{f}) = -2 \ln \mathscr{L}_{0}(\vec{\theta}; \vec{f}) - 2 \ln \mathscr{L}_{0}(\vec{\theta}; \vec{f}) - 2 \ln \mathscr{L}_{phys}(\vec{\theta}) - 2 \ln \mathscr{L}_{syst}(\vec{f})$$
Penalty term due to prior physics constraints
$$\chi_{phys}^{2} = -2 \ln \mathscr{L}_{phys}(\vec{\theta}) = 0$$
Penalty term due to prior systematic constraints
$$\chi_{syst}^{2} = -2 \ln \mathscr{L}_{syst}(\vec{f}) = -2 \ln \mathscr{L}_{syst}(\vec{f}) - 2 \ln \mathscr{L}_{syst}(\vec$$



SBND-PRISM: Flux



IVERPOO

- Muon neutrino flux decreases moving off axis
- Electron neutrino flux remains almost constant

