



# New ideas in neutrino physics

Pedro Machado

July 30th, 2023

Users Meeting 2023

# Why are we doing this?

Well, I can't answer that on your behalf.  
But I am here because I want to break the rules (that is, the standard model).  
I will try in as many ways as I can think of.  
Why neutrinos?



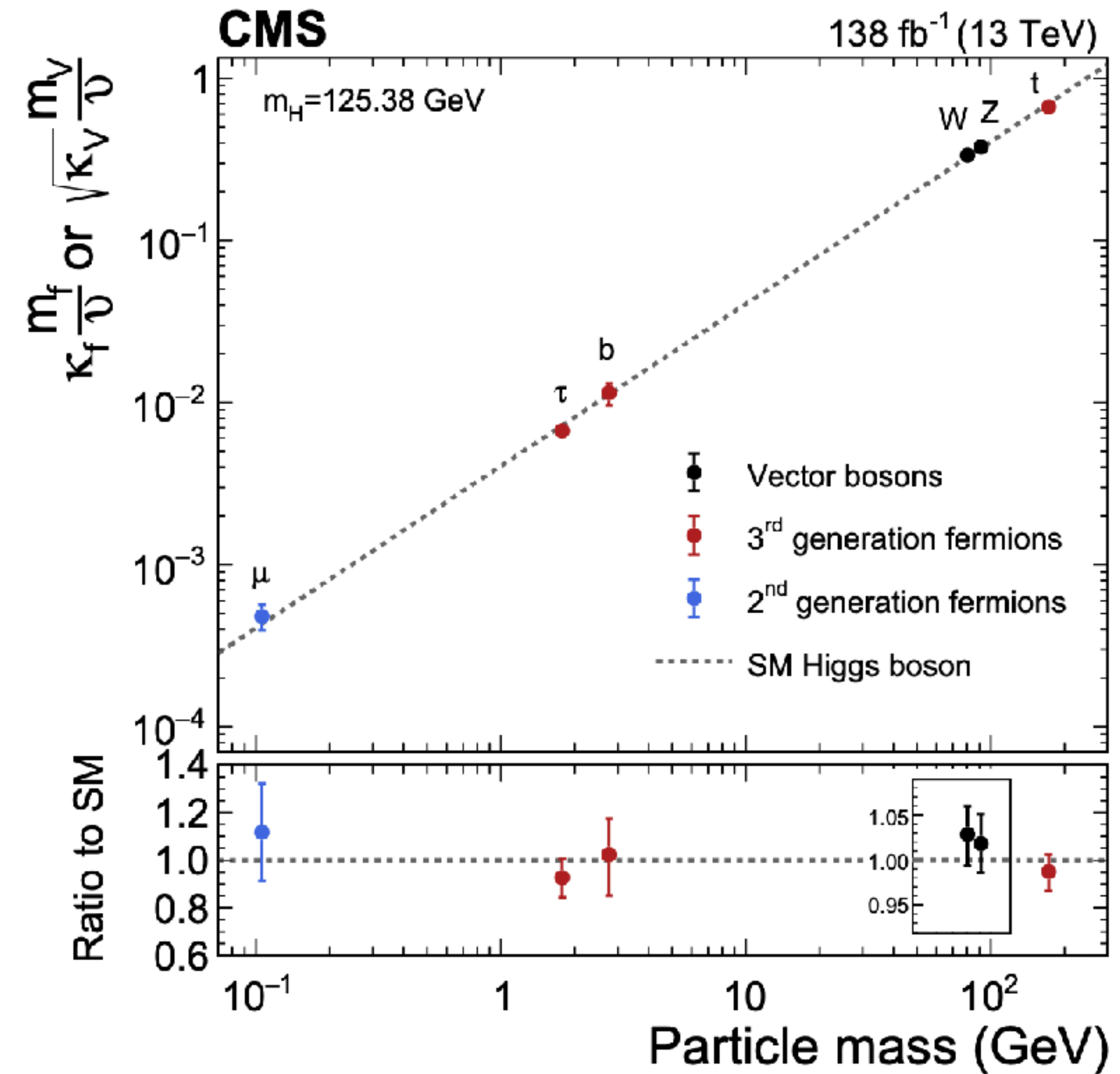
# Why neutrinos?

**The mechanism of neutrino masses is qualitatively different from charged fermions**

All particles within the framework of the standard model, except for neutrinos, get their masses directly and exclusively from the Higgs mechanism

Data points in that direction, at least for charged fermions of the 2<sup>nd</sup> and 3<sup>rd</sup> families and gauge bosons

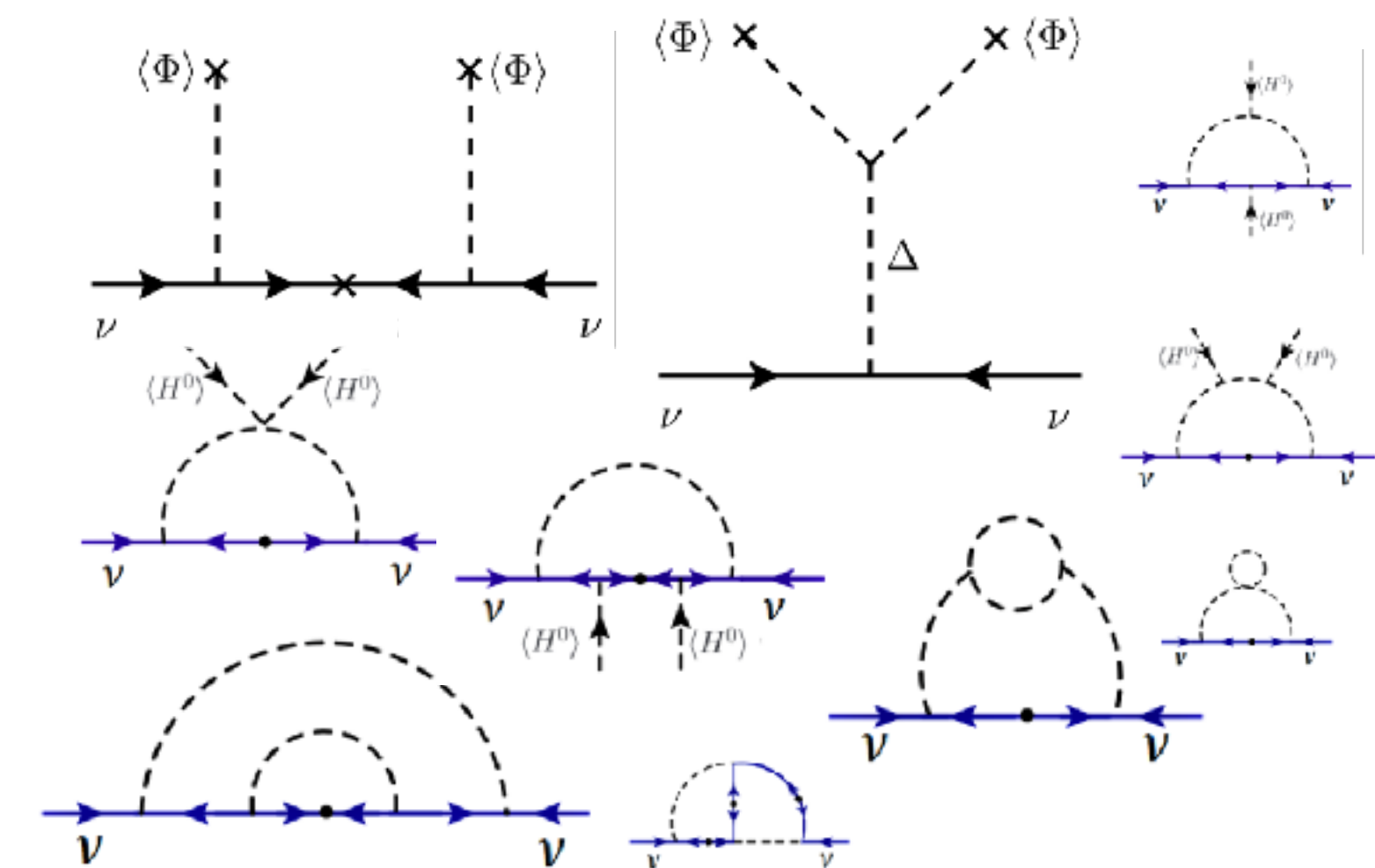
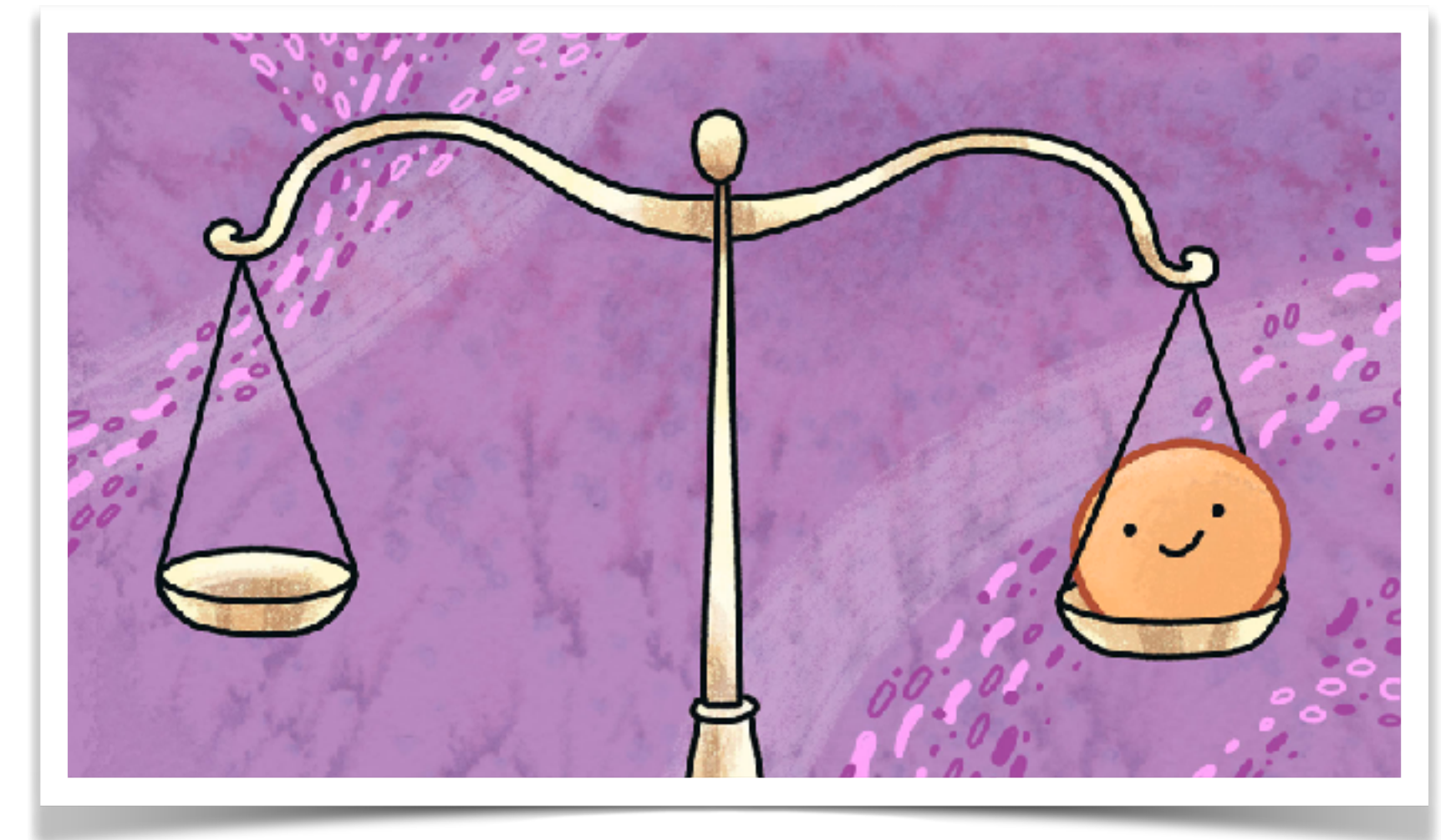
But neutrinos are very different



# Why neutrinos?

Just repeating the Higgs mechanism for neutrinos (invoking a right-handed neutrino) would **predict a particle that is completely different from all observed particles**: its mass has nothing to do with electroweak symmetry breaking

Possible realizations of the neutrino mass mechanism span at least 20 orders of magnitude in scale, from the sub-eV to grand unification, **and there is little to no experimental guidance on the right energy scale**



Miranda Valle 1602.00864  
Babu et al 1907.09498

# Why neutrinos?

One key point:

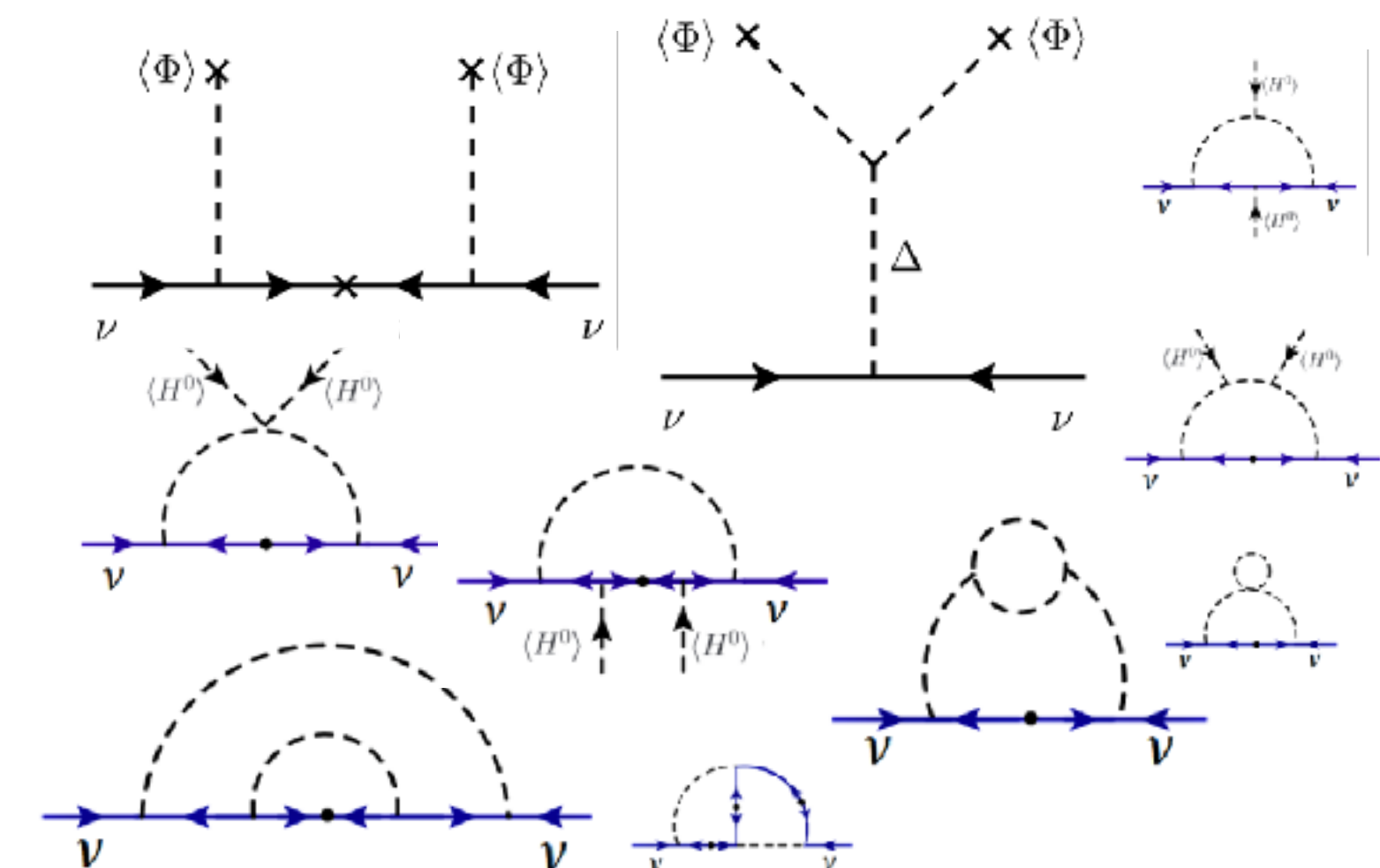
**The neutrino mass mechanism is much more than neutrino masses,** just as electroweak symmetry breaking is much more than the Fermi constant

Of course, we need to determine neutrino masses and the nature of neutrinos, but it is crucial that we go beyond these measurements

We need to approach the problem from many sides:

**Precision neutrino physics**

**BSM searches**



Miranda Valle 1602.00864  
Babu et al 1907.09498

# Precision neutrino physics program

*Redundancy, redundancy, redundancy*

# What is the deal about precision physics?

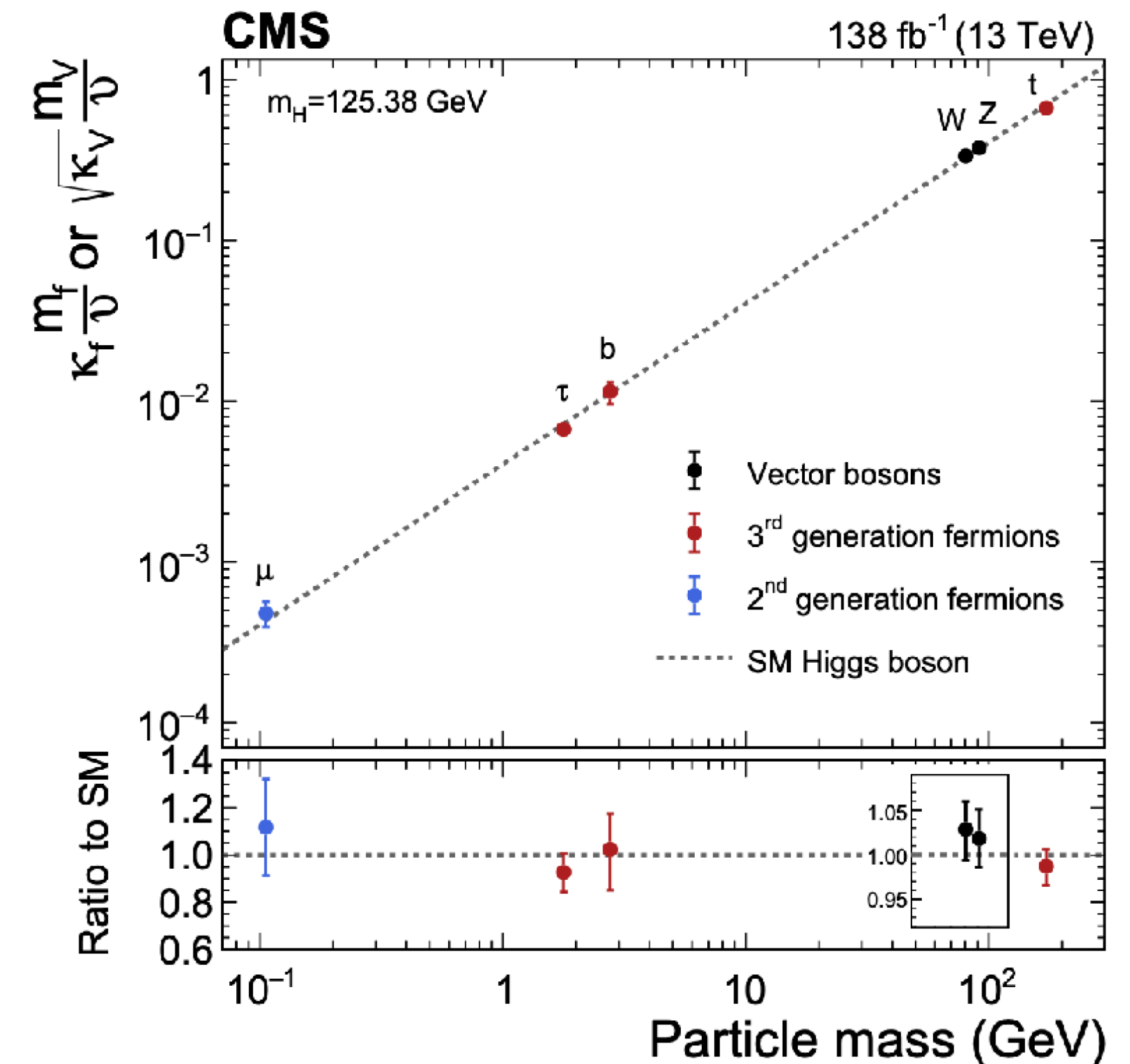
Let's go back to the Higgs example

Within the Higgs mechanism, the coupling of particles to the Higgs is given by the mass of the particle divided by the Higgs vev

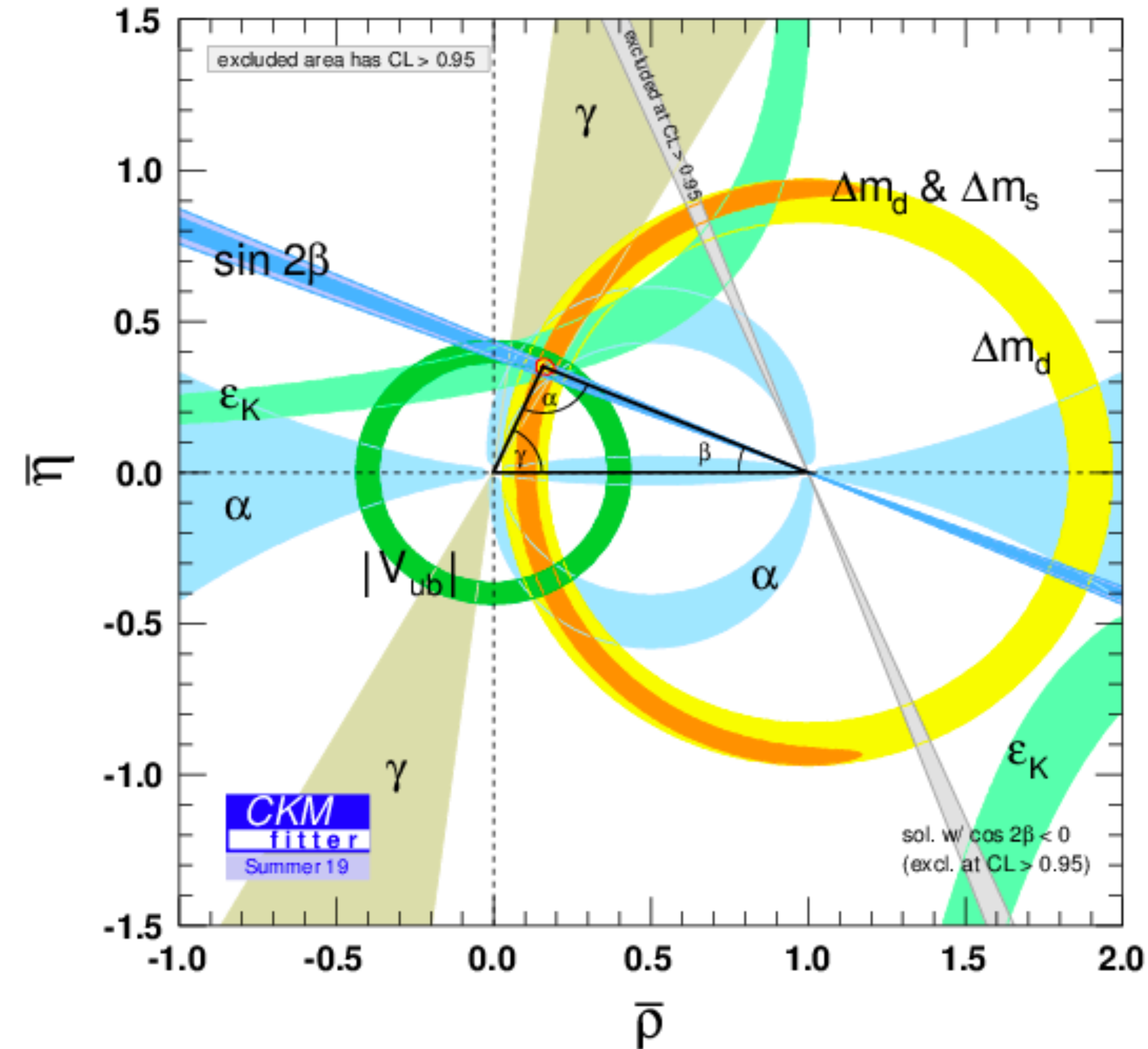
We measure the mass by looking at, e.g., the kinematic threshold in  $e^+e^-$  collisions

Now we have a clear prediction for the Higgs coupling

**Any deviation is a breaking of the standard model!**



# Precision neutrino physics

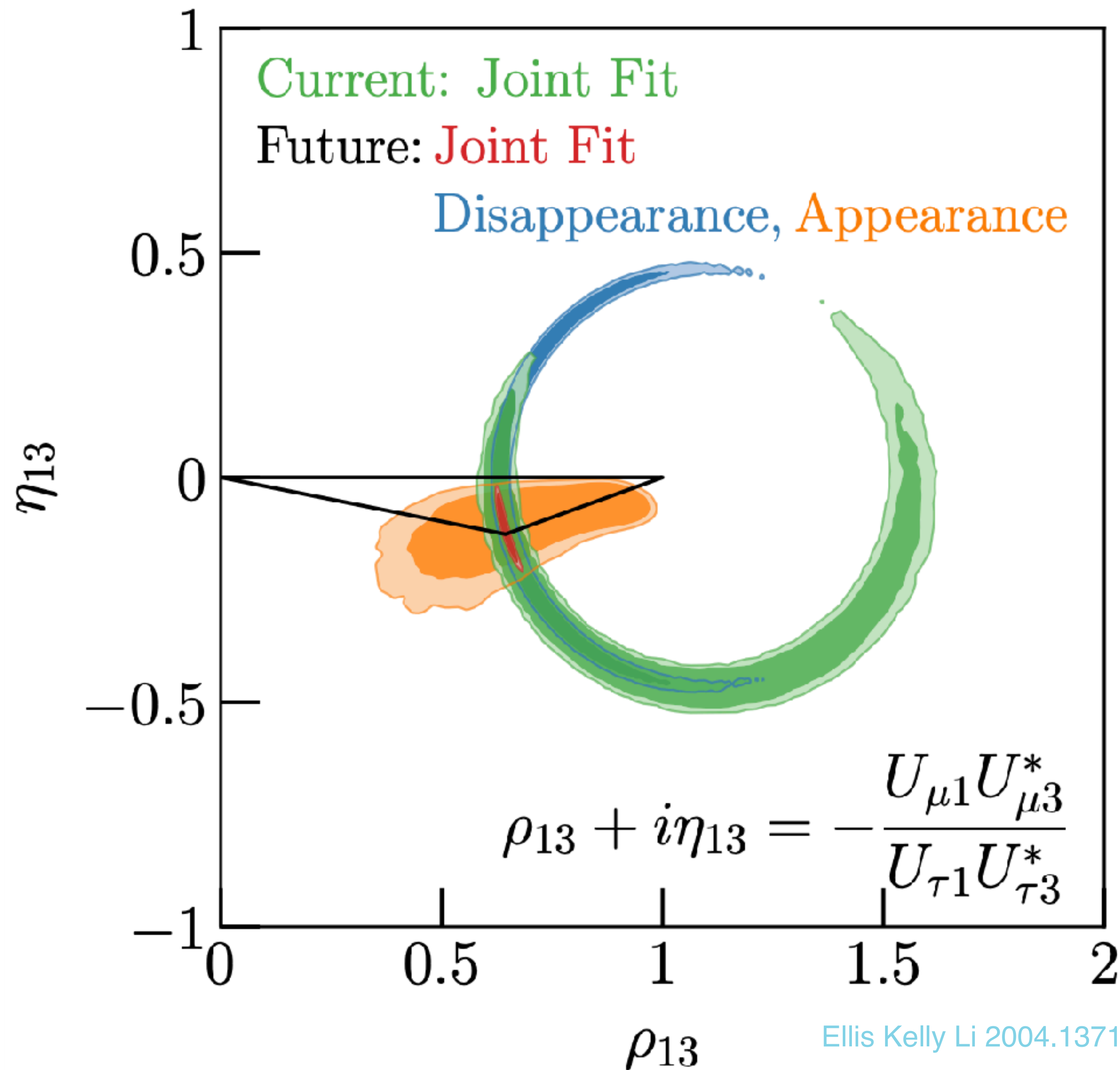


We have done it for quarks

We got to do it for neutrinos



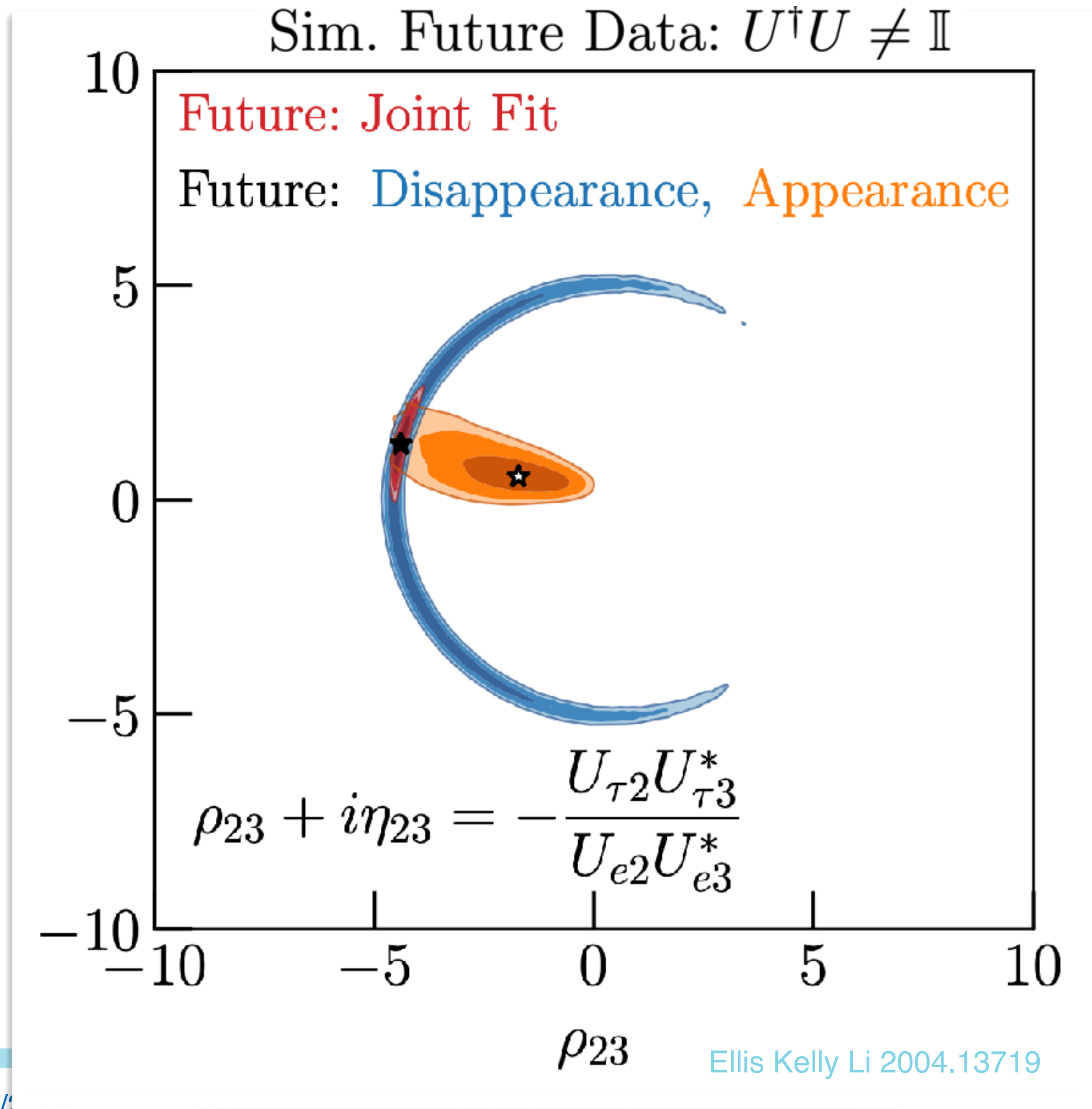
# Precision neutrino physics



Ellis Kelly Li 2004.13719

**DUNE, HK, JUNO and IceCube will enable a bona fide precision physics program in the neutrino sector**

# Precision neutrino physics



**DUNE, HK, JUNO and IceCube will enable a bona fide precision physics program in the neutrino sector**

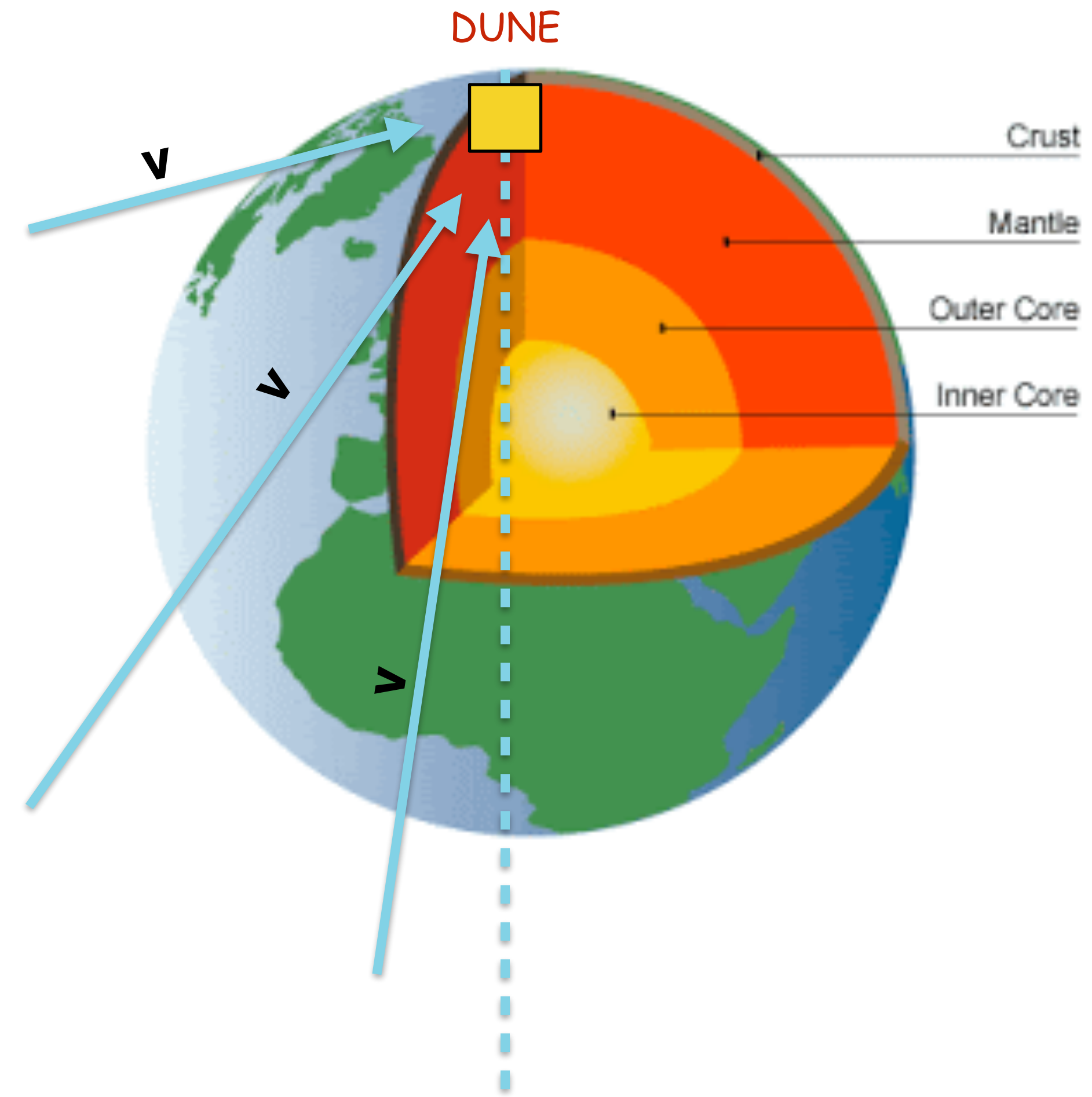
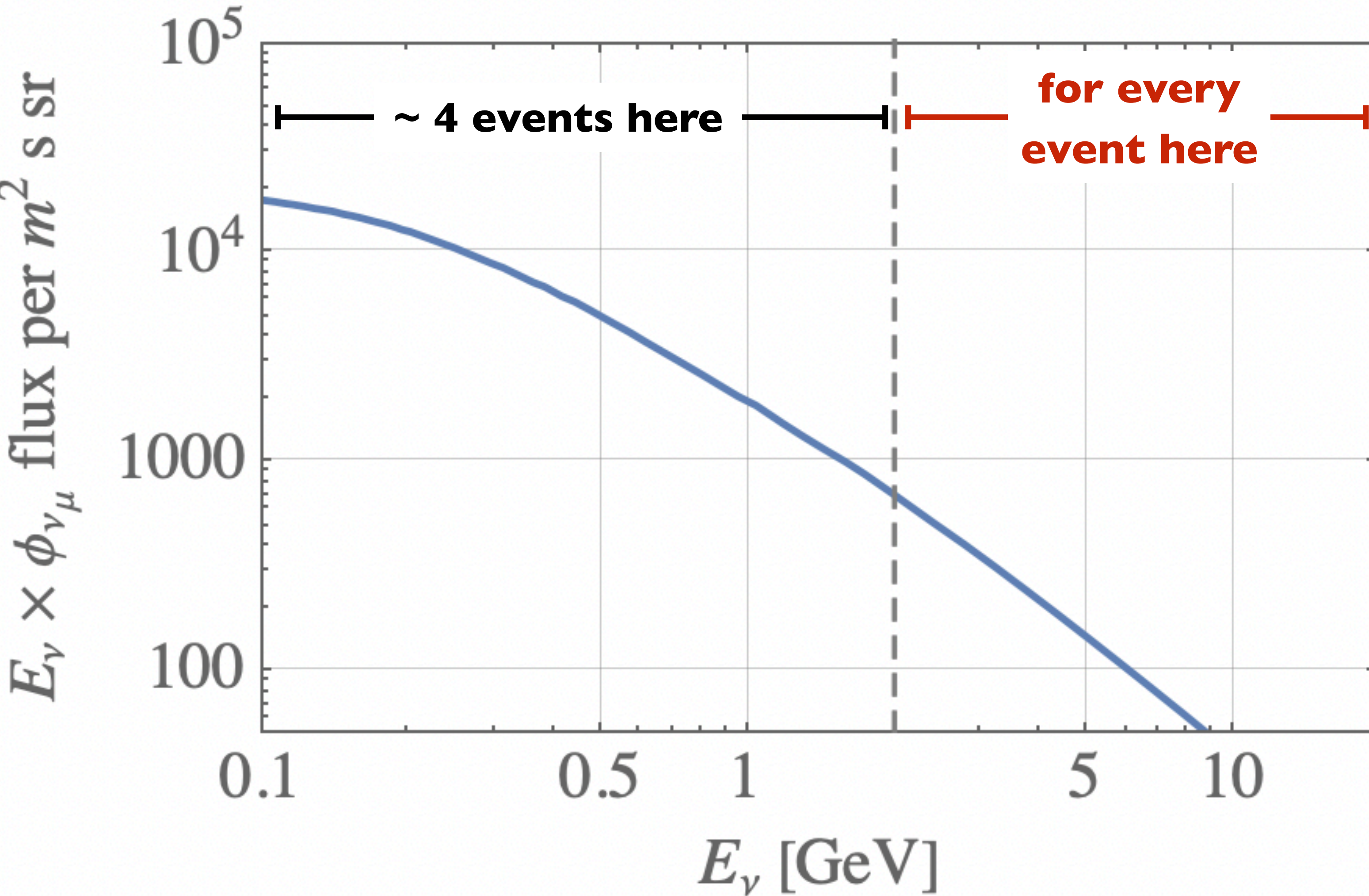
# Precision neutrino physics program

*New observables = new opportunities*

# Low energy atmospheric neutrinos

Based on Kelly PM et al 1904.02751

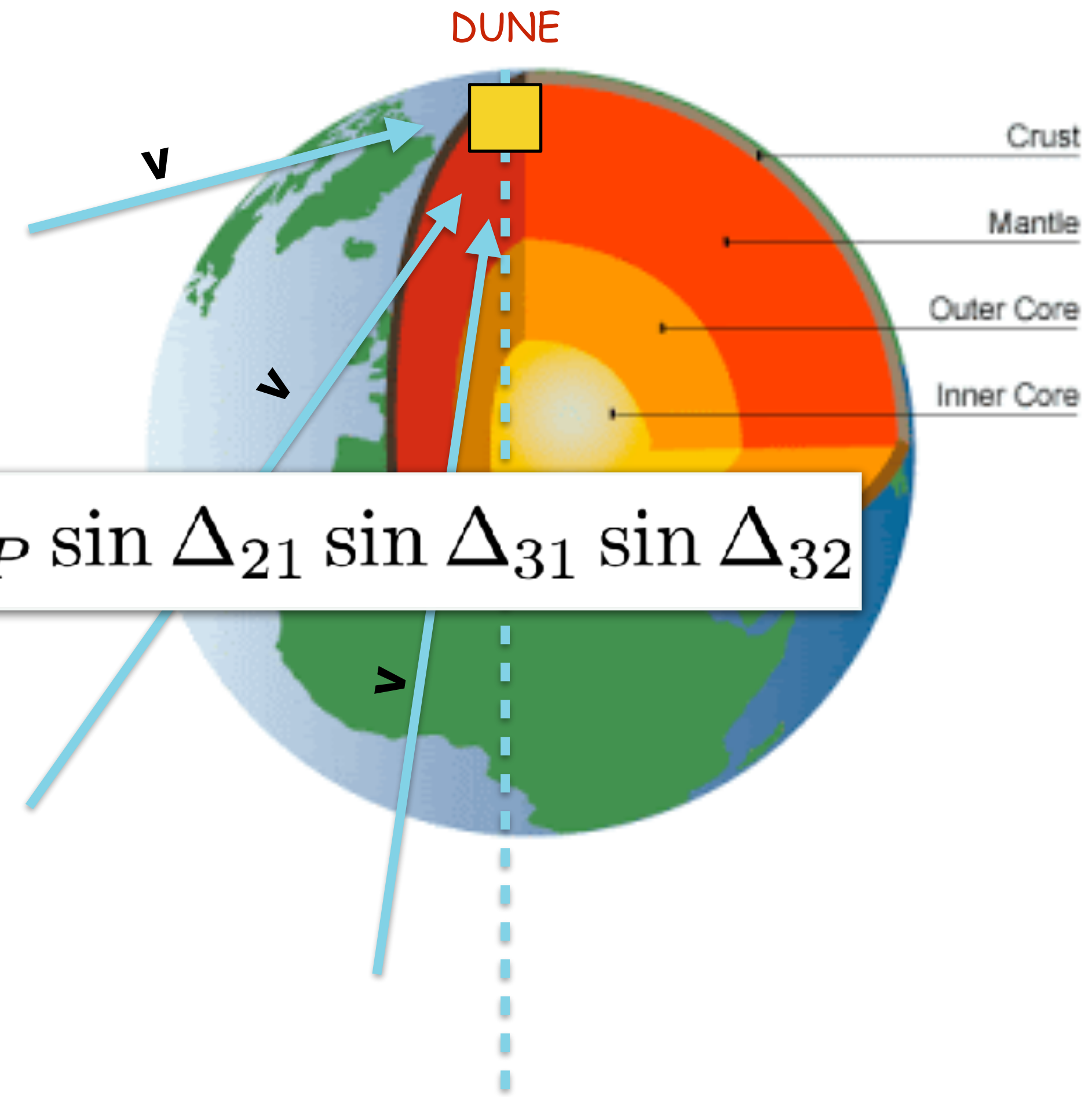
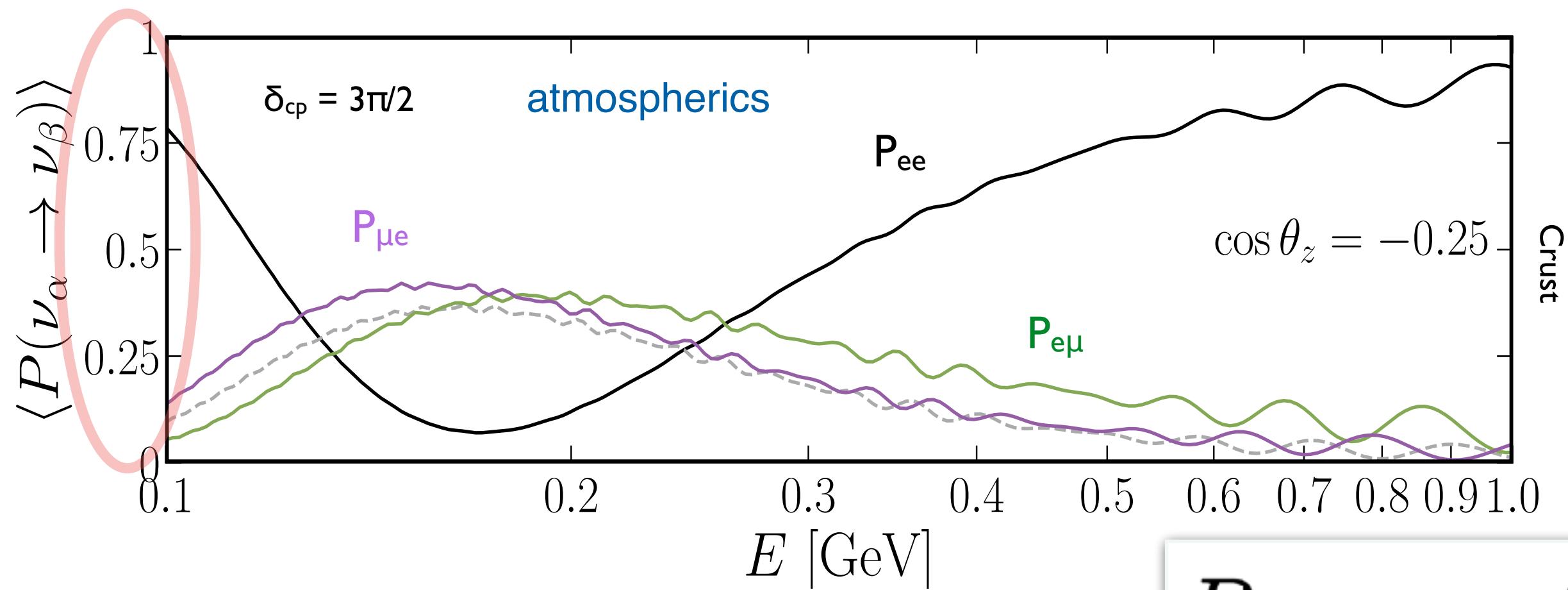
Lots of atmospheric neutrinos below the GeV!



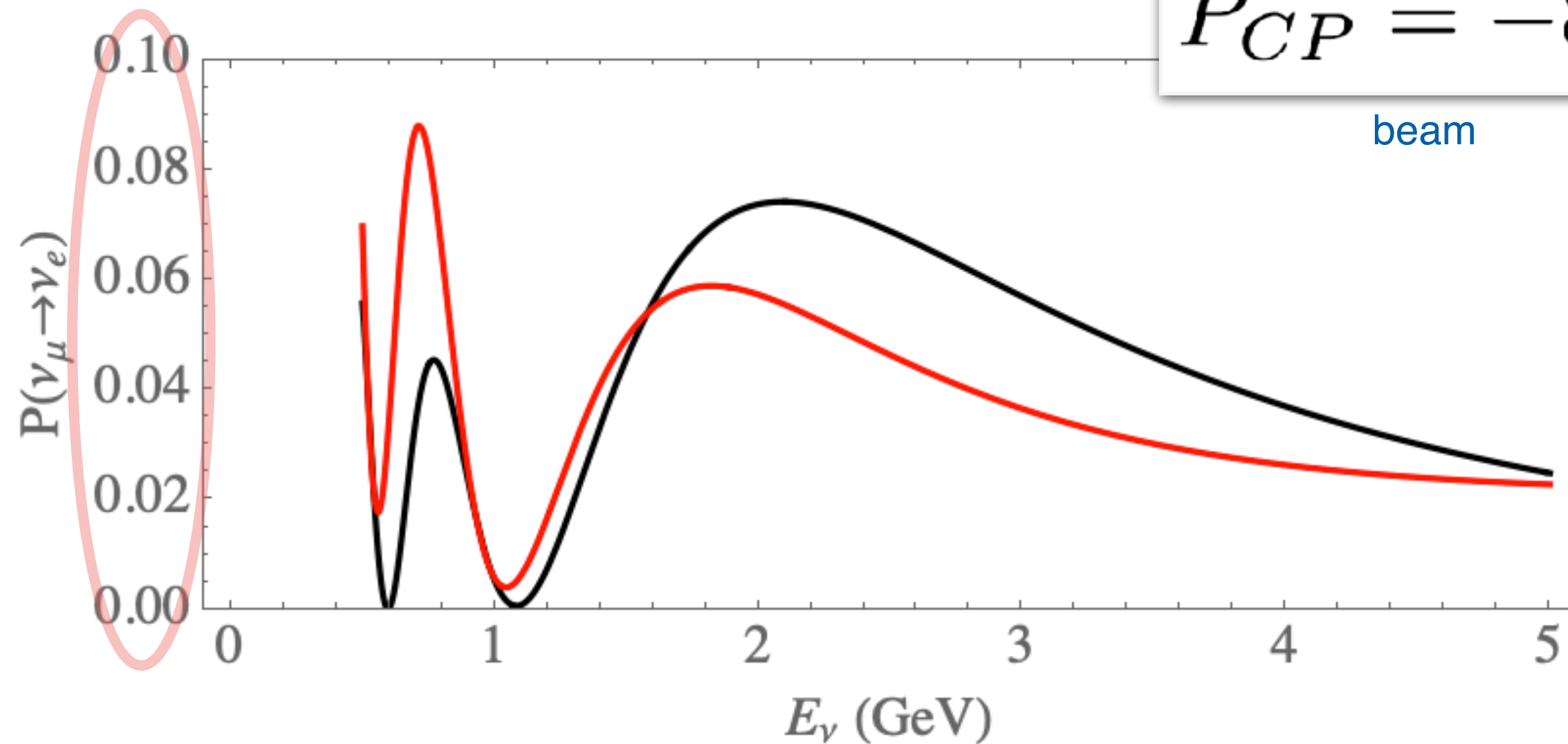
# Low energy atmospheric neutrinos

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**CP violation effect for atmospheric neutrinos is 10 times larger compared to beam neutrinos at DUNE!**



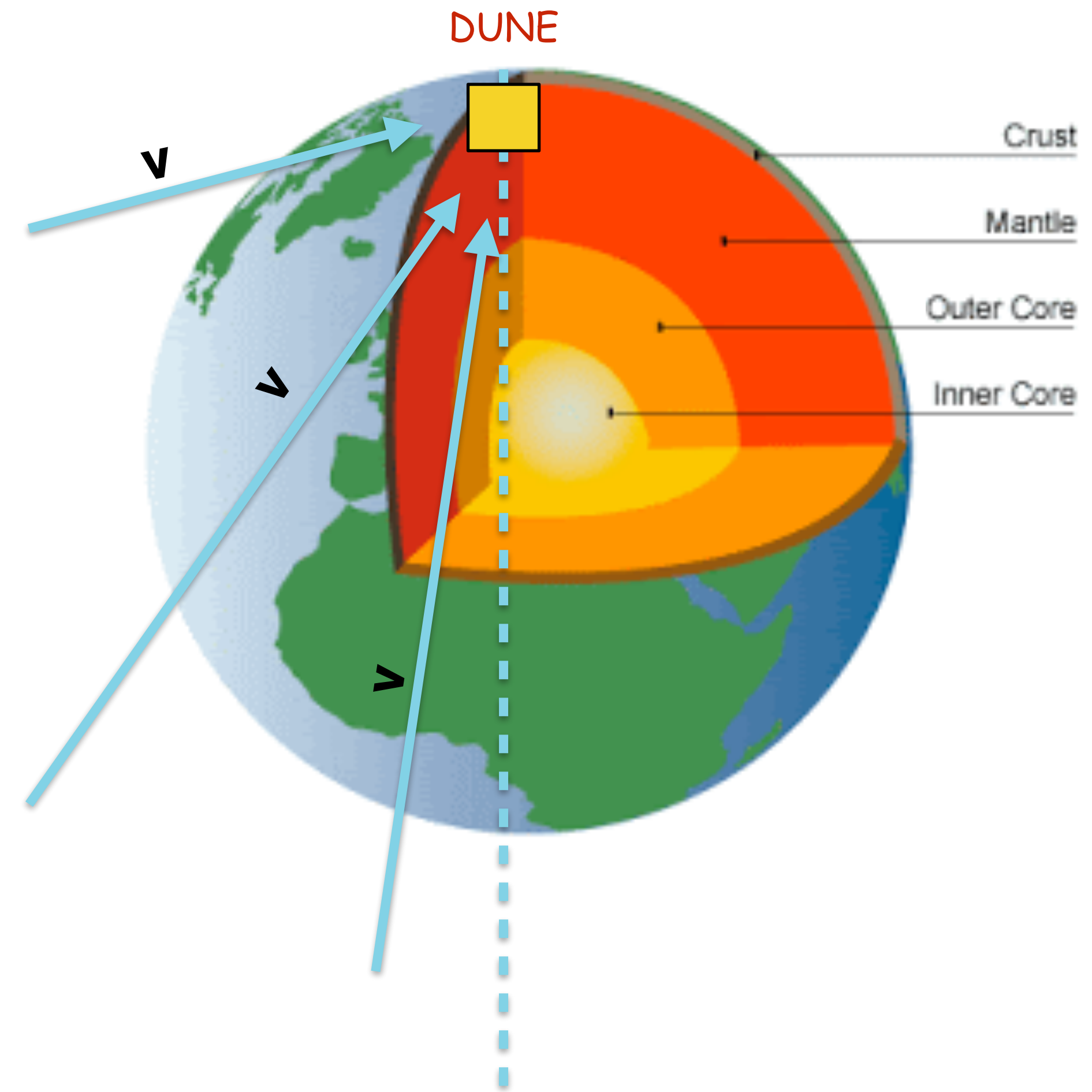
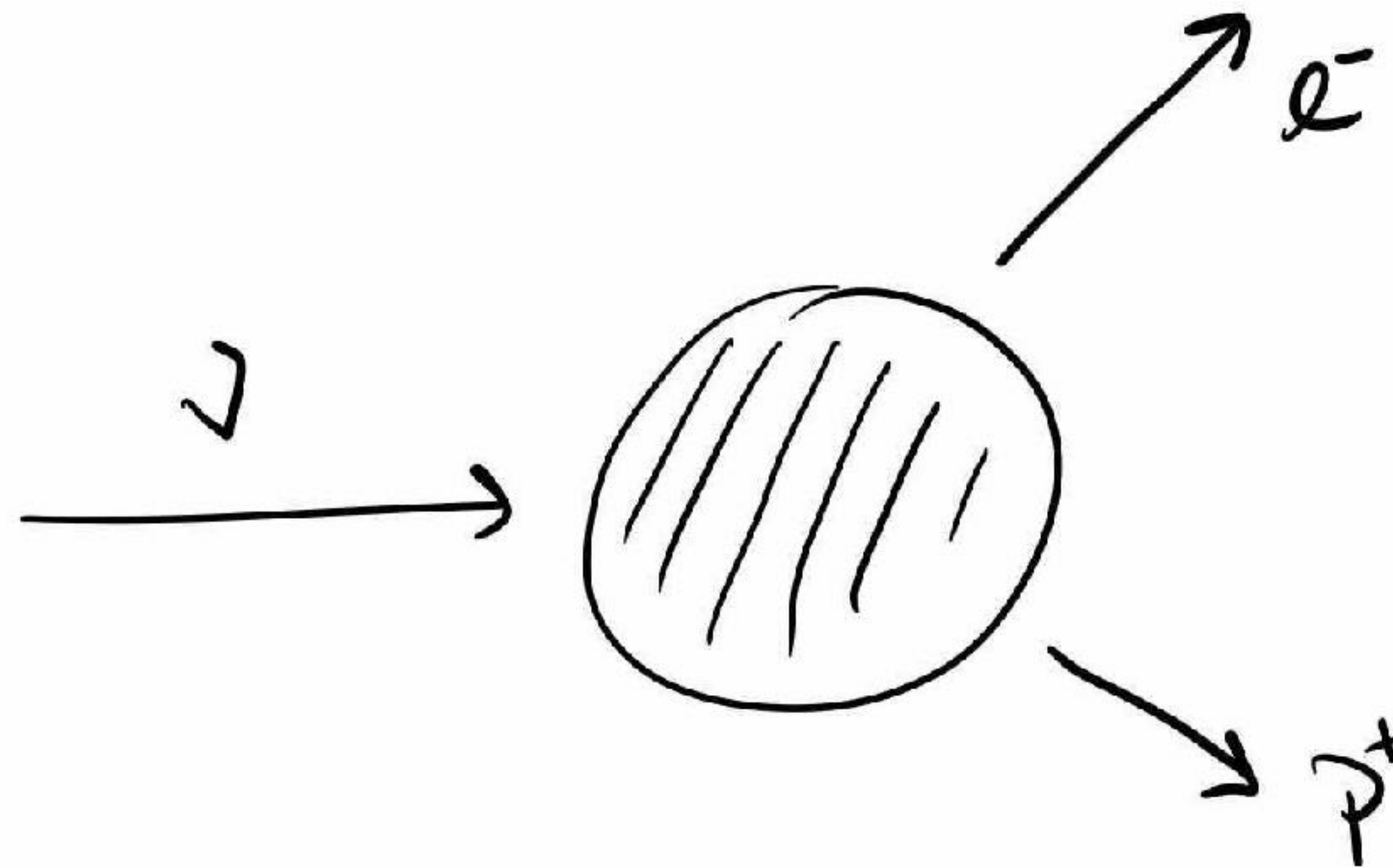
$$P_{CP} = -8J_r \sin \delta_{CP} \sin \Delta_{21} \sin \Delta_{31} \sin \Delta_{32}$$



# Low energy atmospheric neutrinos

Based on Kelly PM et al 1904.02751

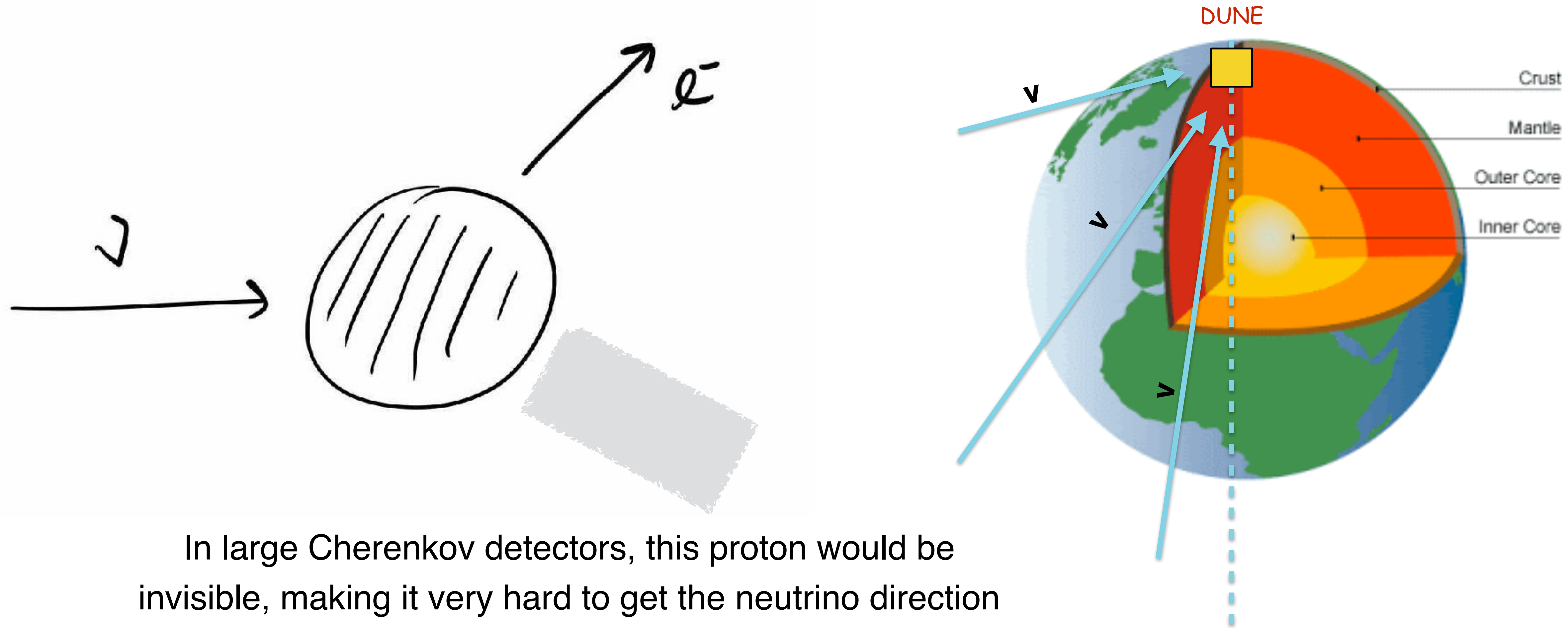
Reconstructing neutrino energy and direction for sub-GeV atmospheric neutrinos is also 10x harder...



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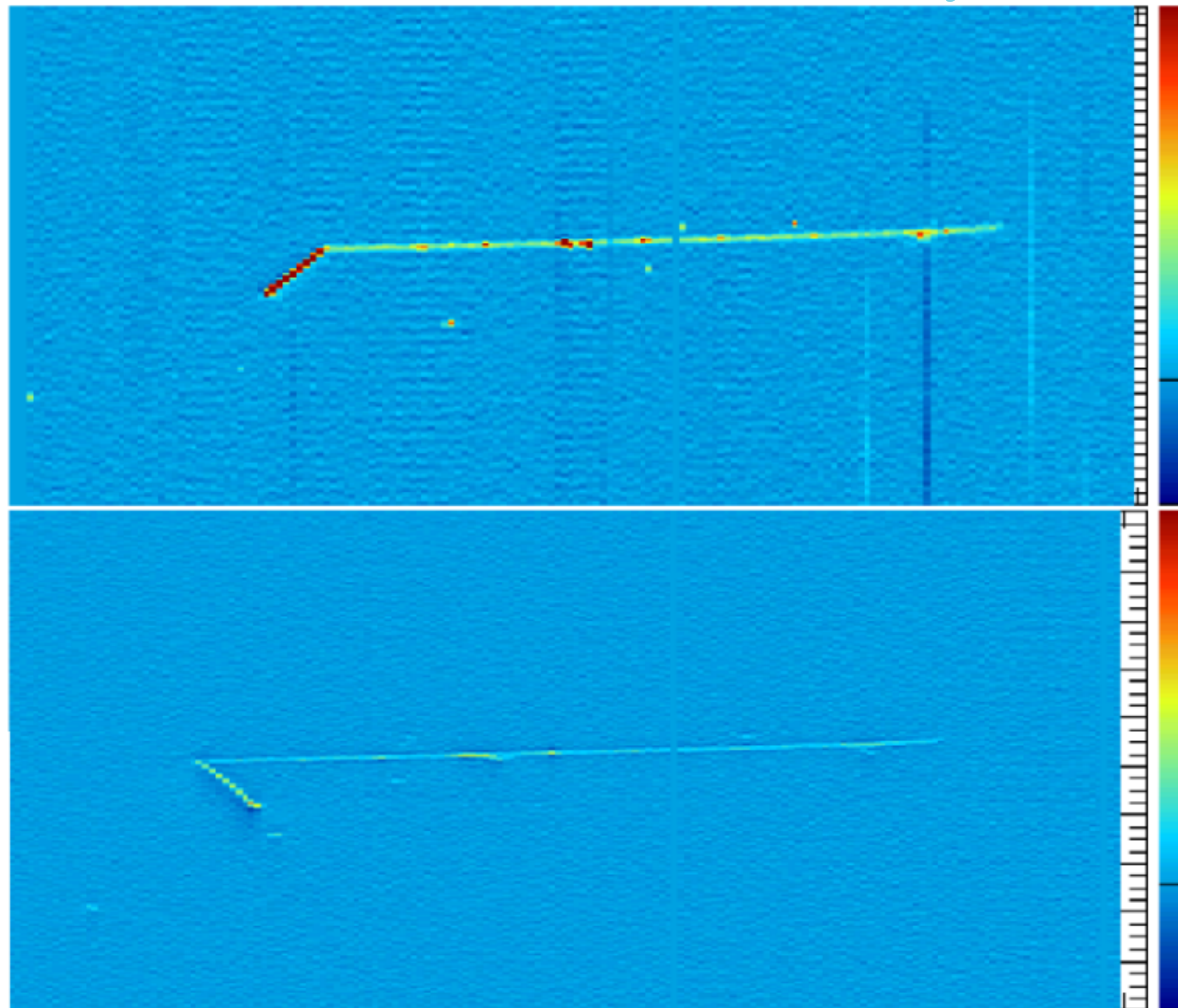


In large Cherenkov detectors, this proton would be invisible, making it very hard to get the neutrino direction

# Low energy atmospheric neutrinos

ArgoNeuT demonstrated the LAr capability to detect 21 MeV recoil protons.

ArgoNeuT 1810.06502

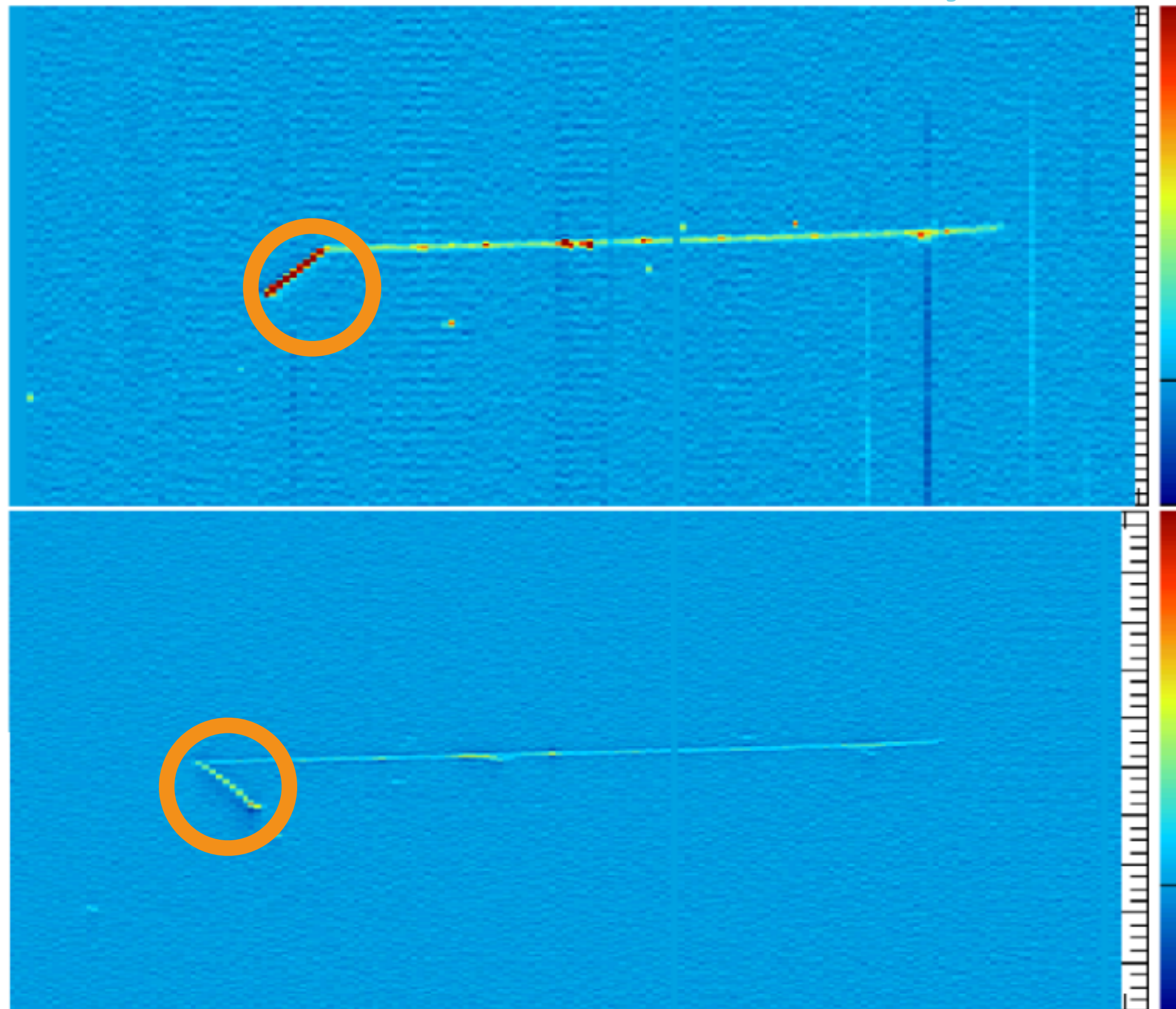




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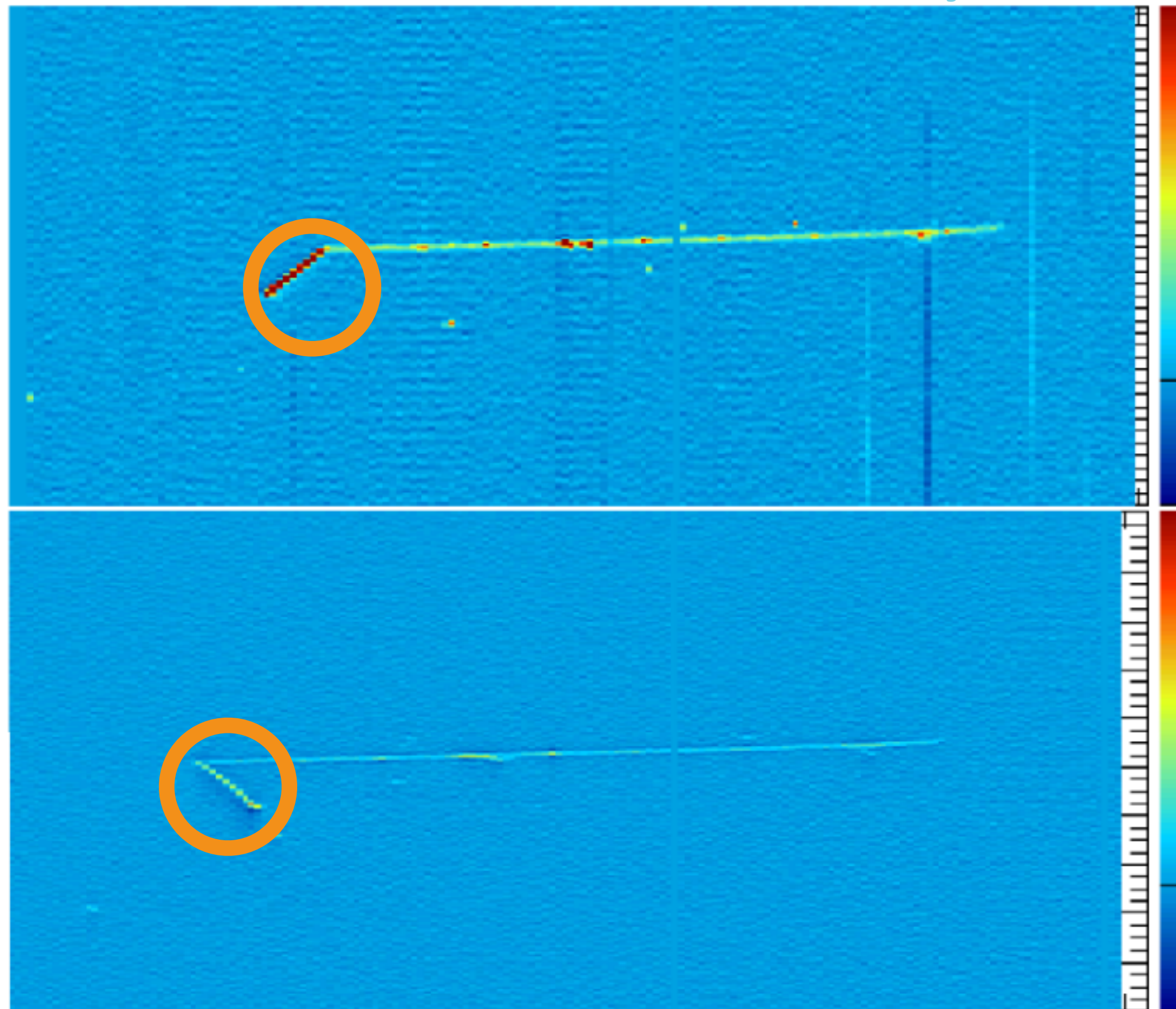
**Reconstruct, identify and point.**

For comparison, SK can only see protons that emit Cherenkov light, that is, protons with energy above  $\sim 1.4$  GeV

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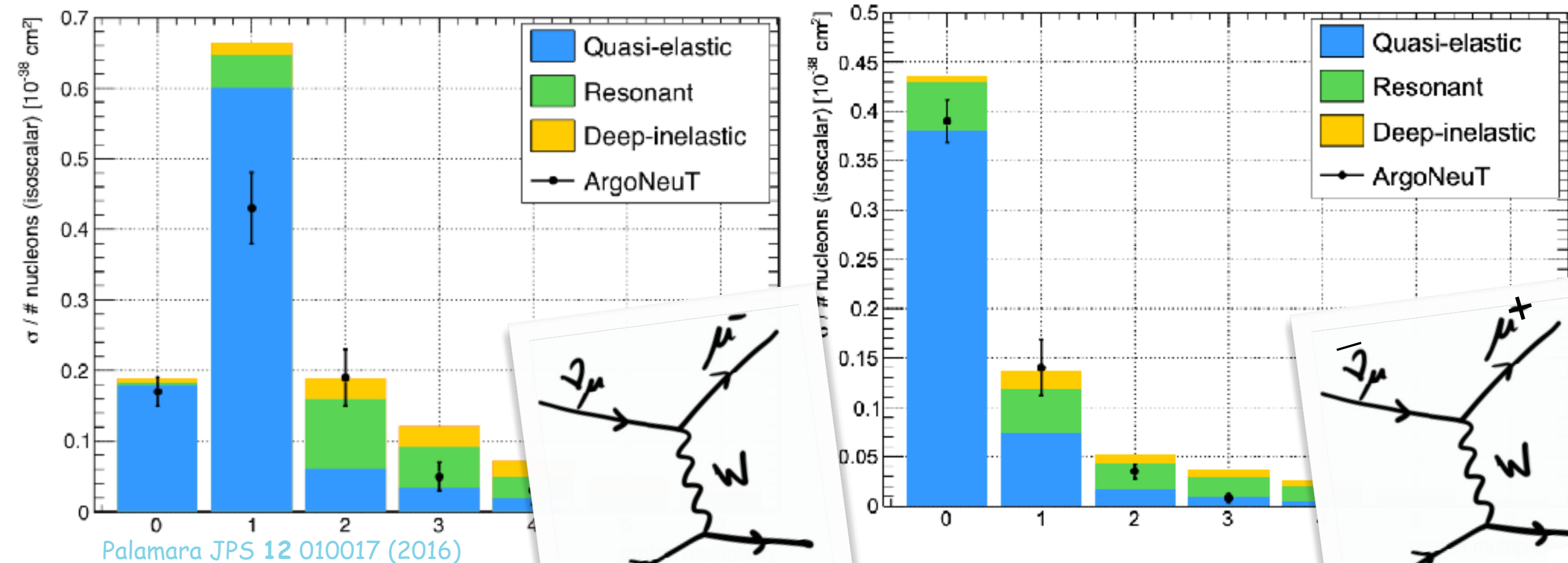
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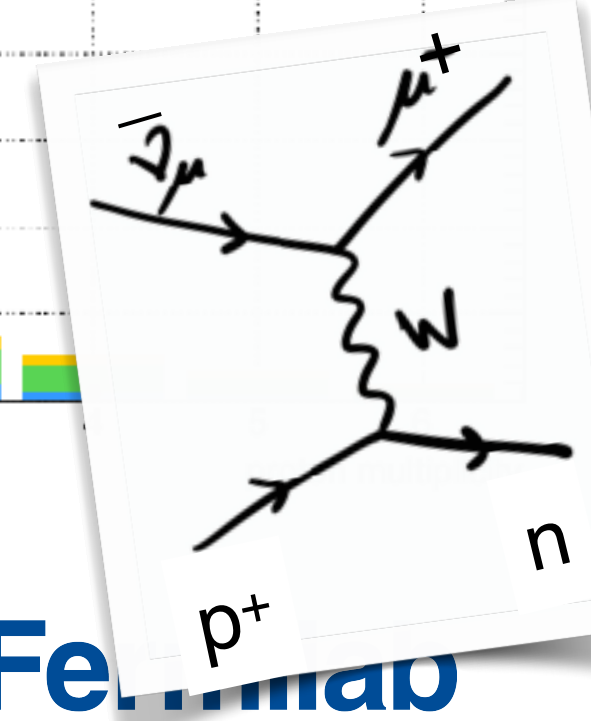
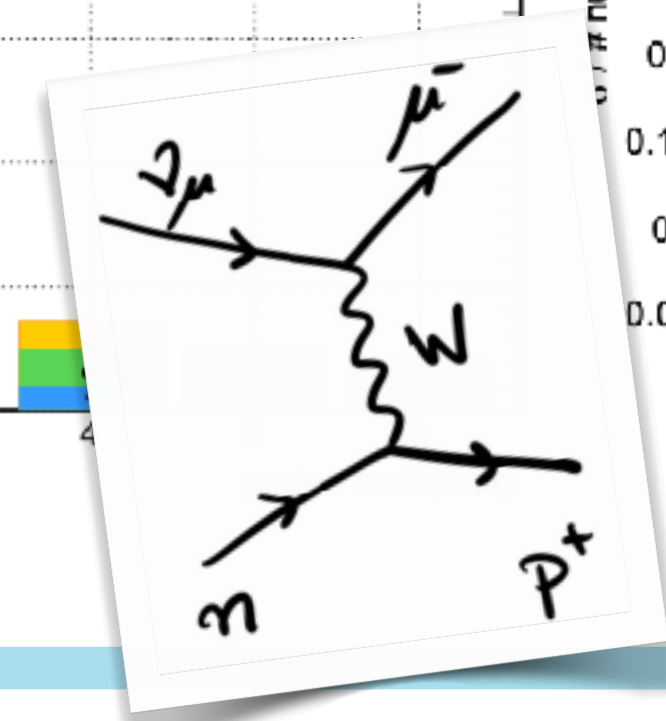
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**Event topology carries extra information**



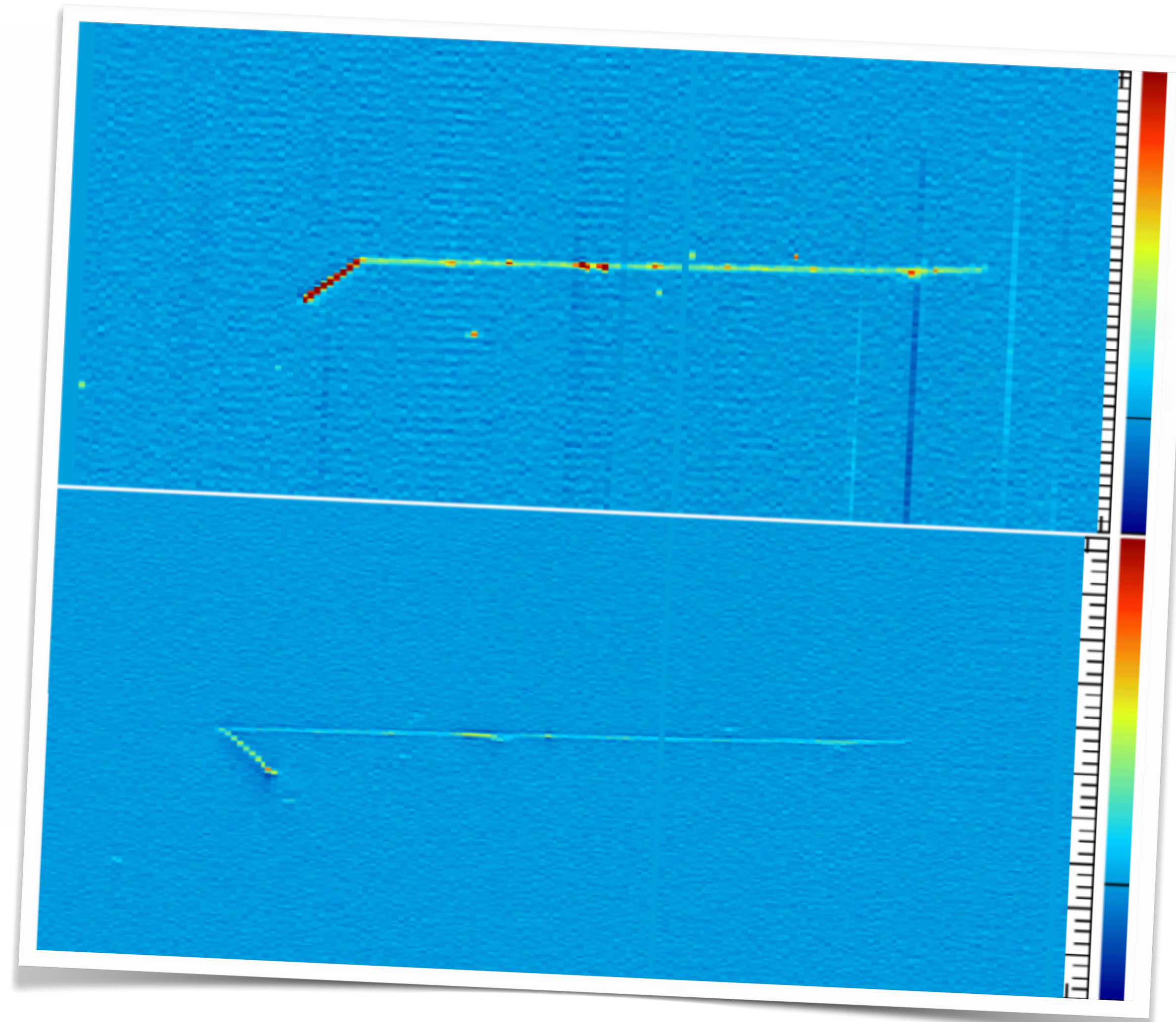
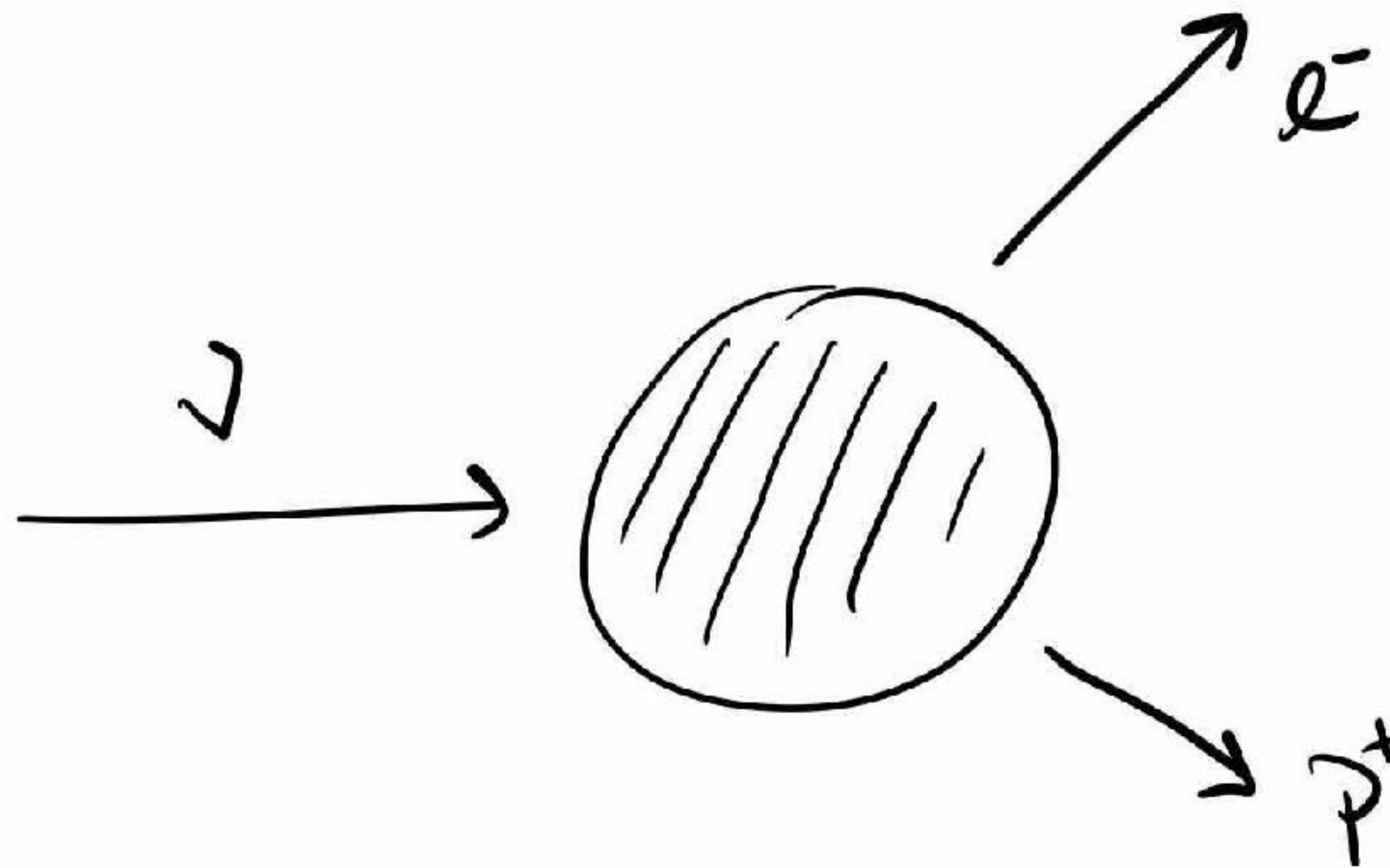
Palamara JPS 12 010017 (2016)



# Low energy atmospheric neutrinos

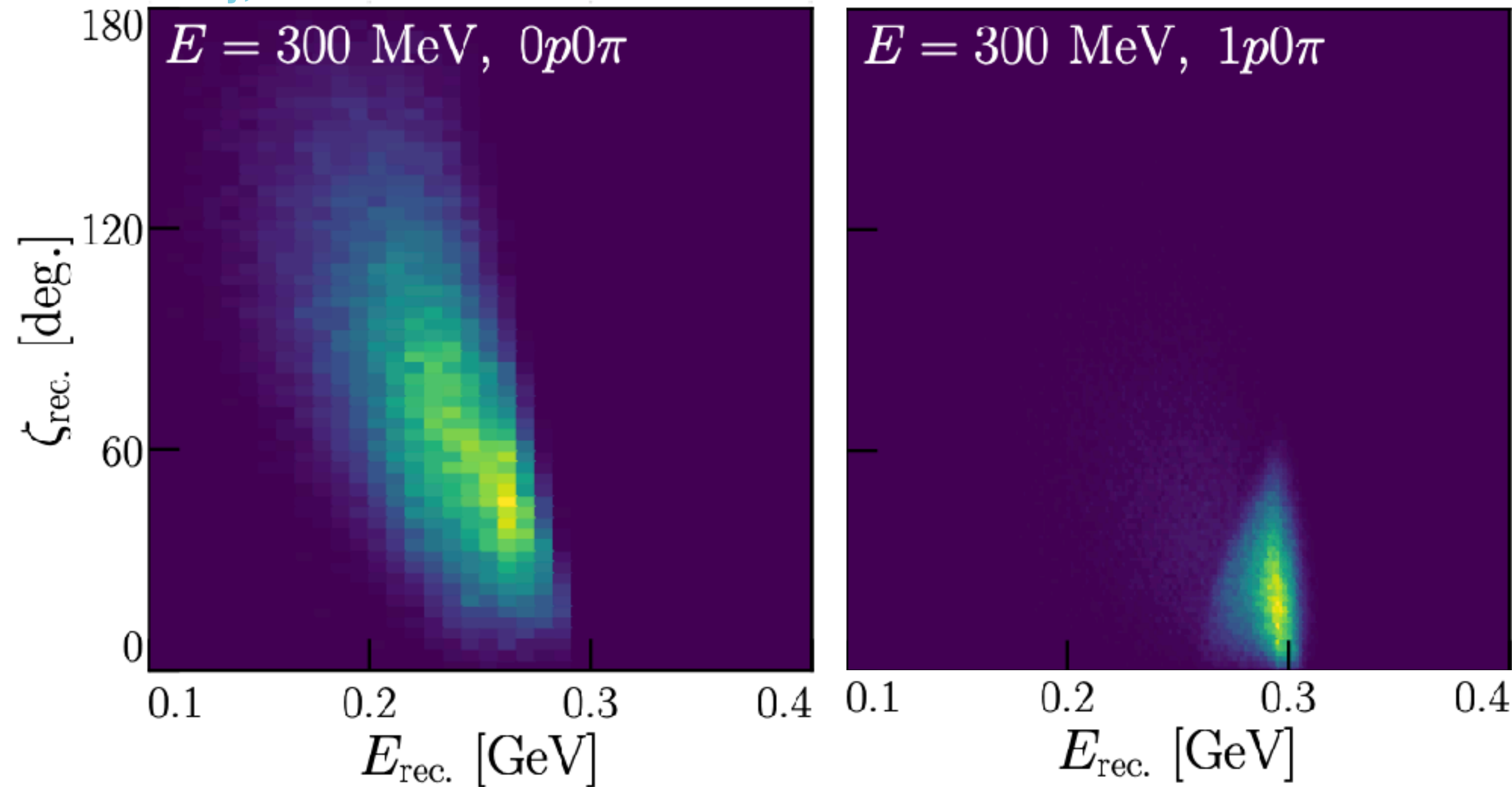
Based on Kelly PM et al 1904.02751

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# Low energy atmospheric neutrinos

Kelly, PM et al 2110.00003



**Details:**

Simulate neutrino-argon interactions with event generators

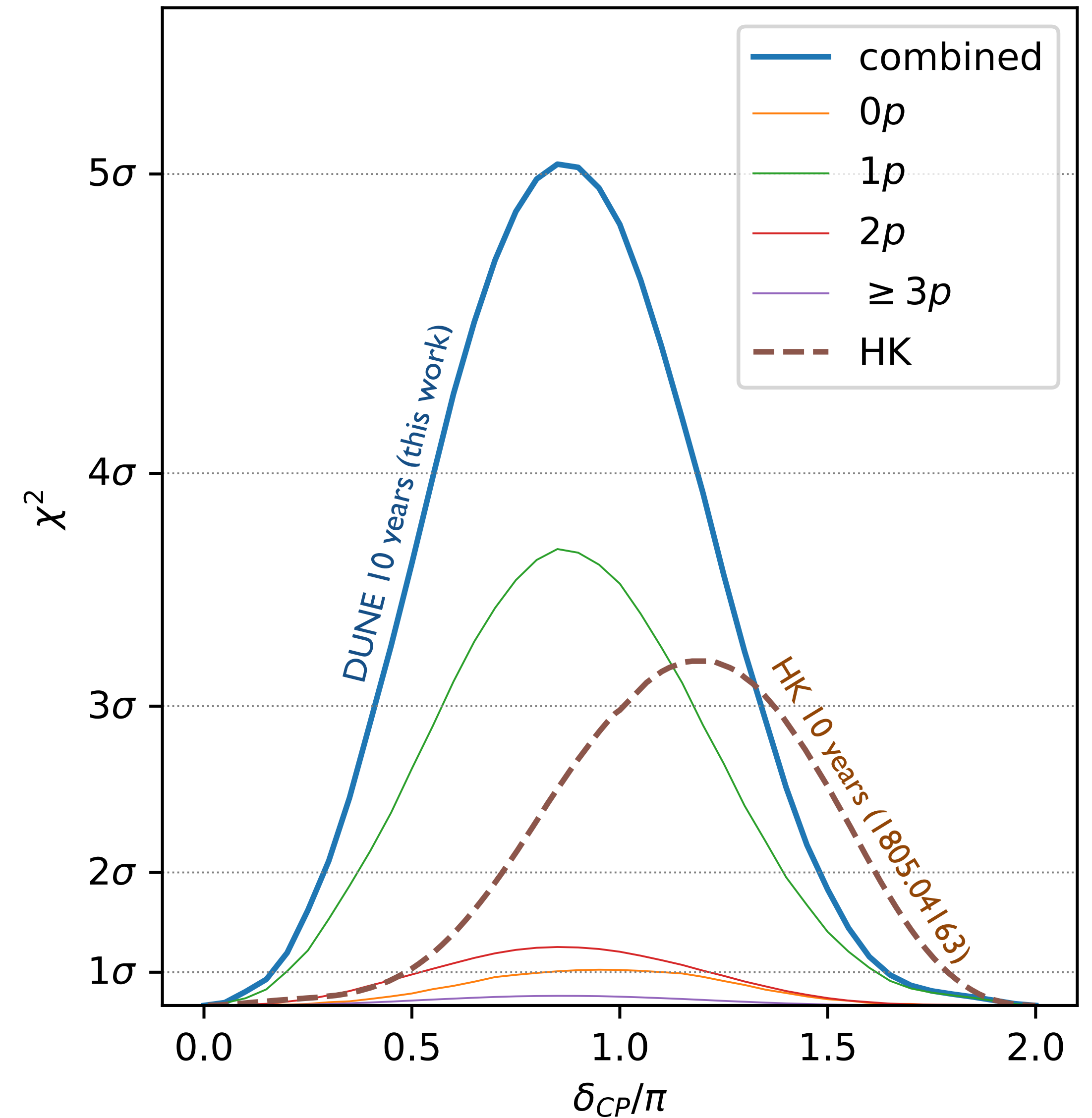
Use realistic atmospheric fluxes (Honda et al 1502.03916)

Account for uncertainties of atmospheric neutrino fluxes  $\Phi_{\alpha}(E) = \Phi_{\alpha,0} f_{\alpha}(E) \left(\frac{E}{E_0}\right)^{\gamma}$   
 40% normalization, 5% e/ $\mu$  ratio, 2% nu/nubar ratio,  $\pm 0.2$  spectral distortion coefficient

Realistic LArTPC capabilities

$\Delta p = 5\%, 5\%, 10\%$ ,  $\Delta\theta = 5^{\circ}, 5^{\circ}, 10^{\circ}$ , for e,  $\mu$ , p,  $K_p = 30 \text{ MeV}$

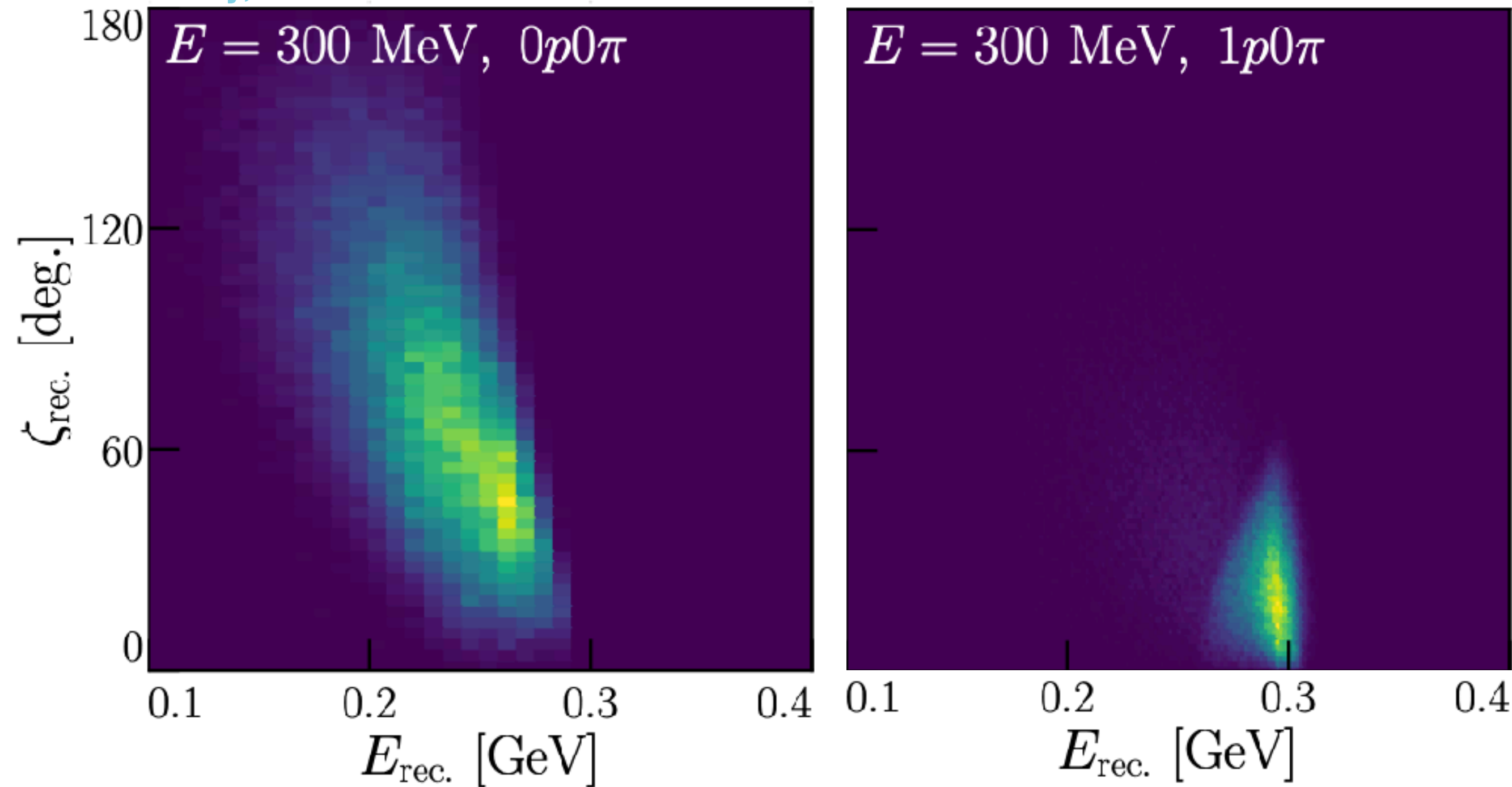
Classify events by final state topology (number of protons)



Kelly, PM et al Phys.Rev.Lett. 123 (2019) 8, 081801

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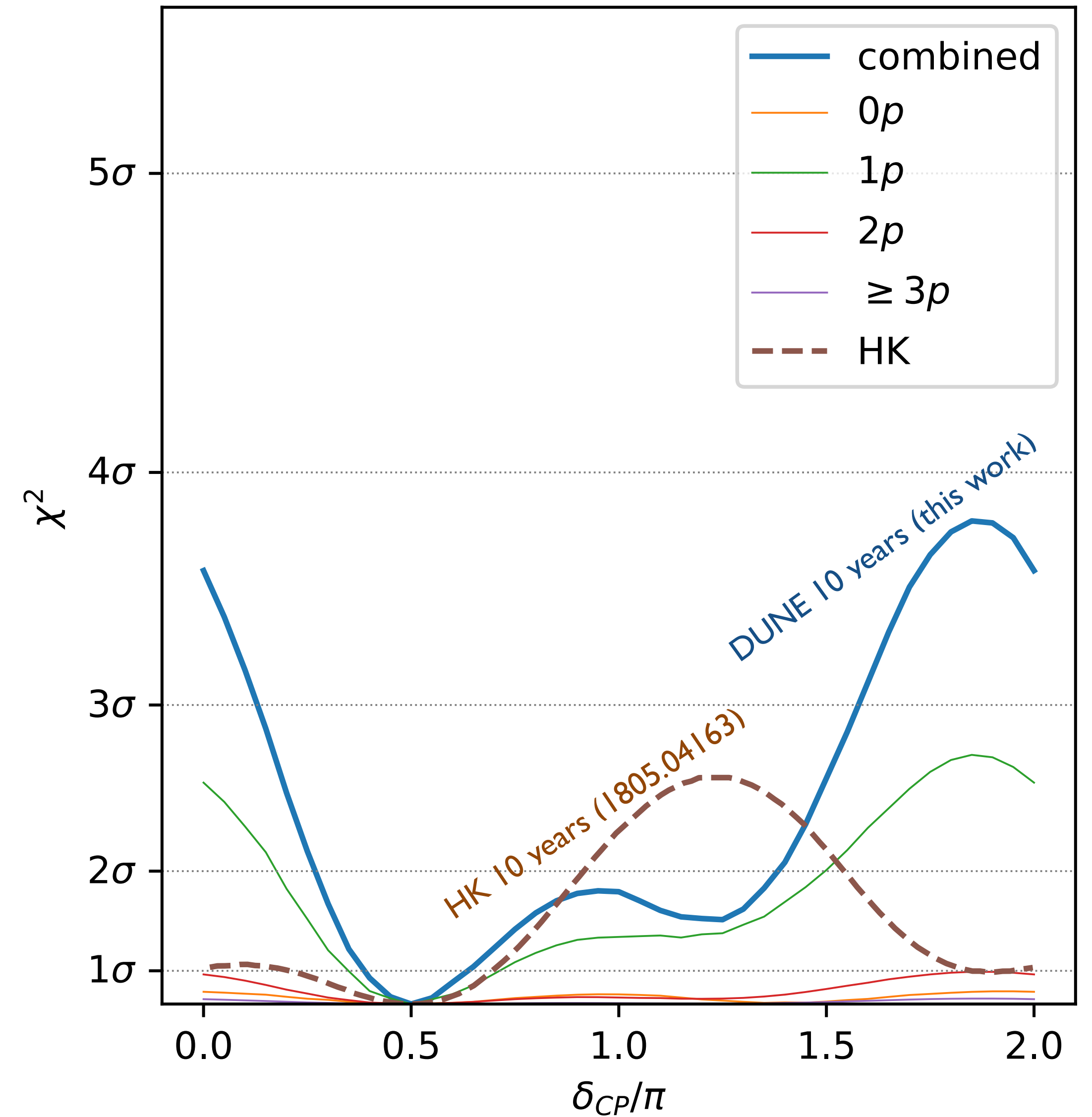
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# BSM searches in neutrino experiments

*What's new? Why now?*

# BSM in neutrino experiments

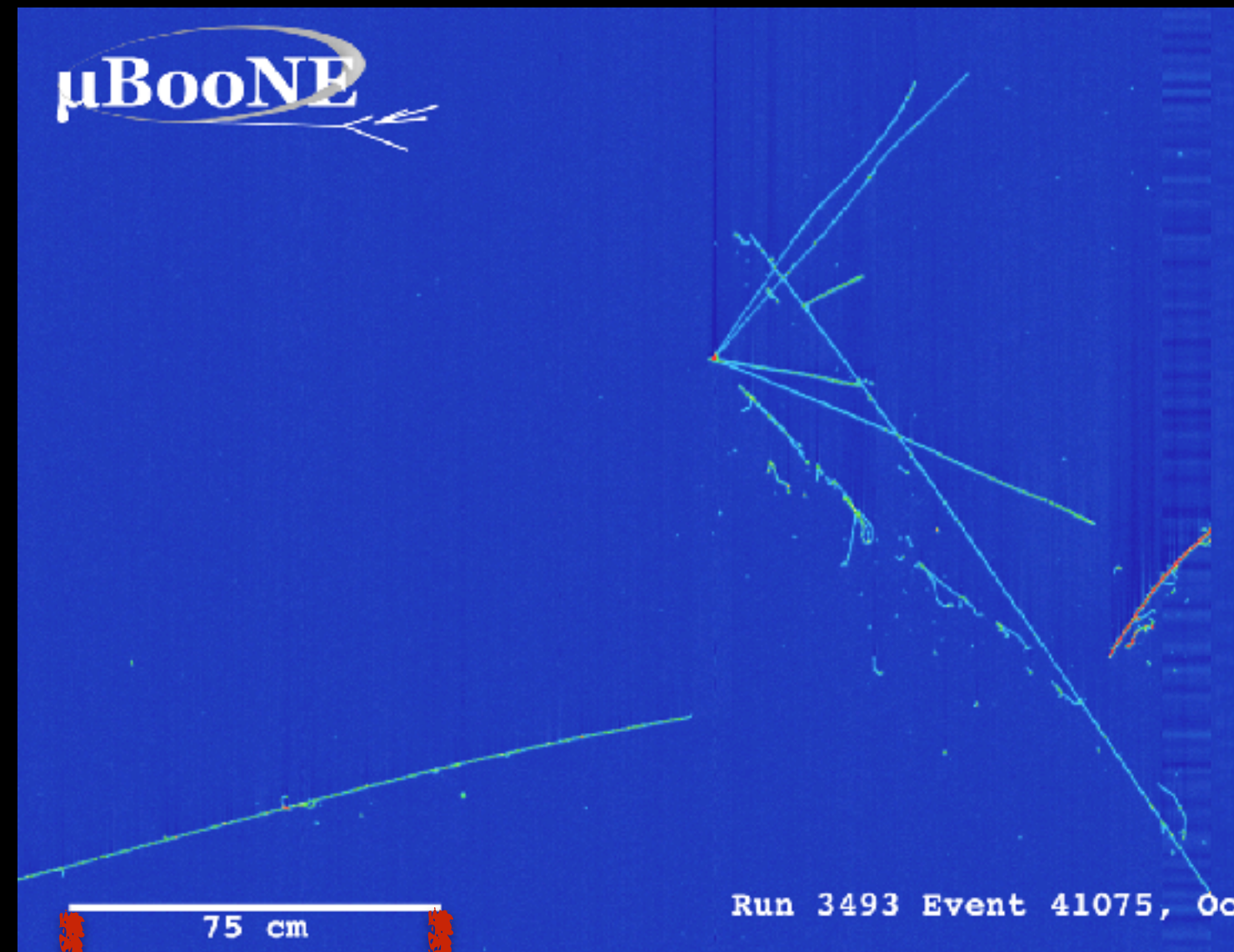
400

600

800

1000

1200



**LArTPCs are just amazing**  
**How can we leverage that**  
**to break the SM?**

From a theory perspective,  $(LH)$  is special: it is a gauge-singlet

Neutrinos are one of the renormalizable portals to new physics

## The three renormalizable portals to new physics:

Neutrinos  $(LH)$

Higgs  $(H^\dagger H)$

Photon  $(F_{\mu\nu})$

The overarching physics program should comprehend precise measurements of these three portals

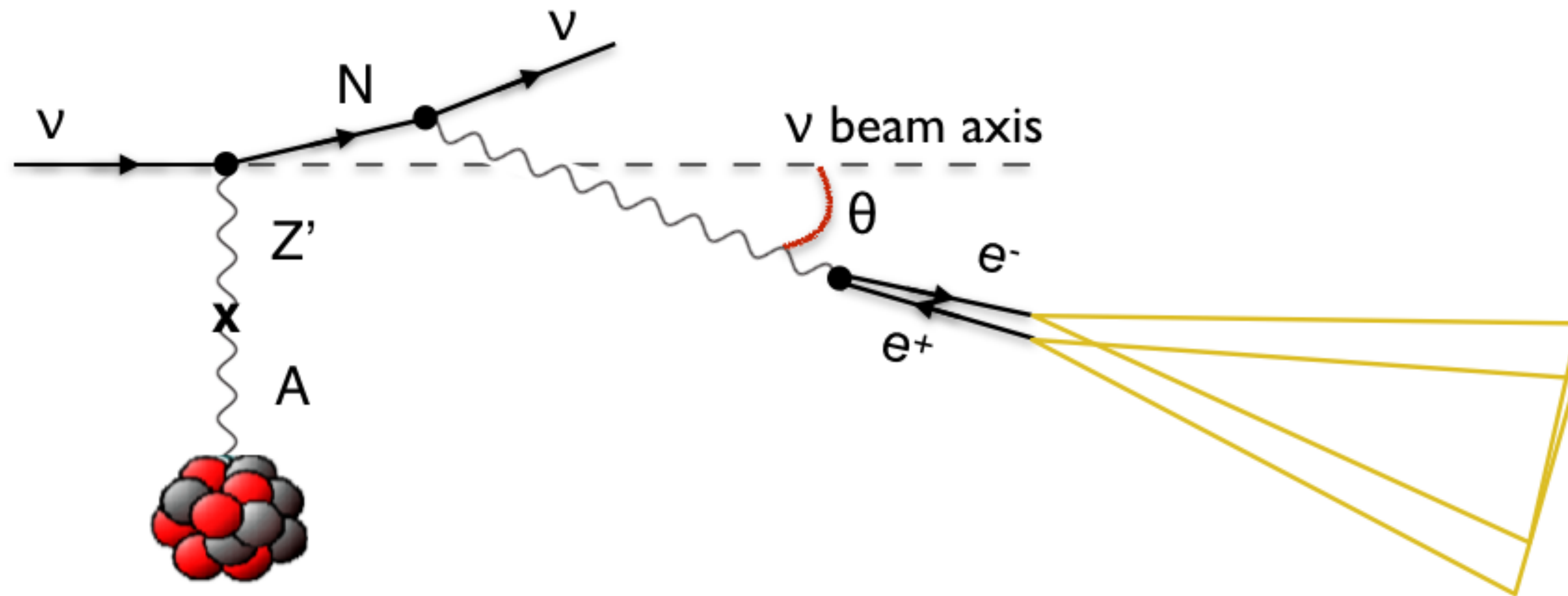
Dark sector

Weak sector

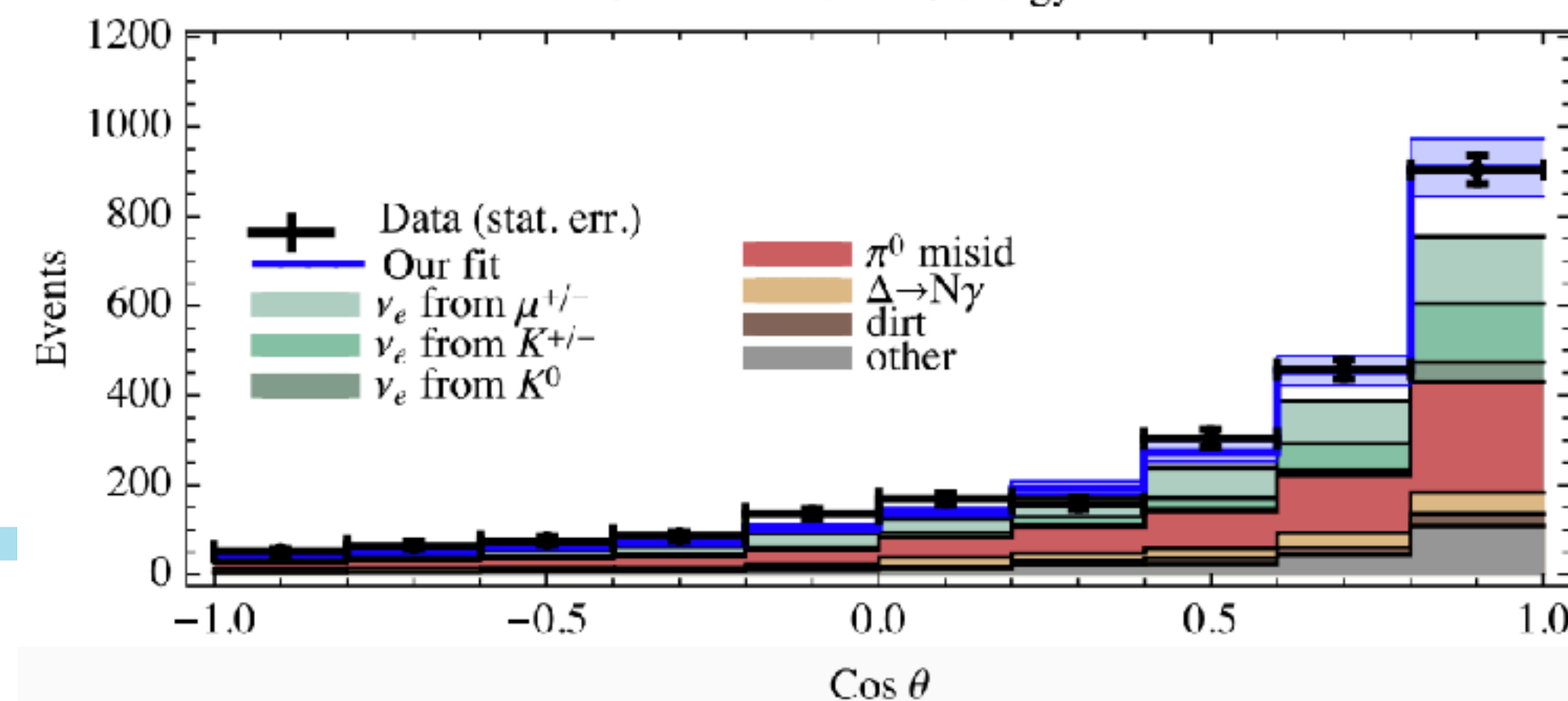
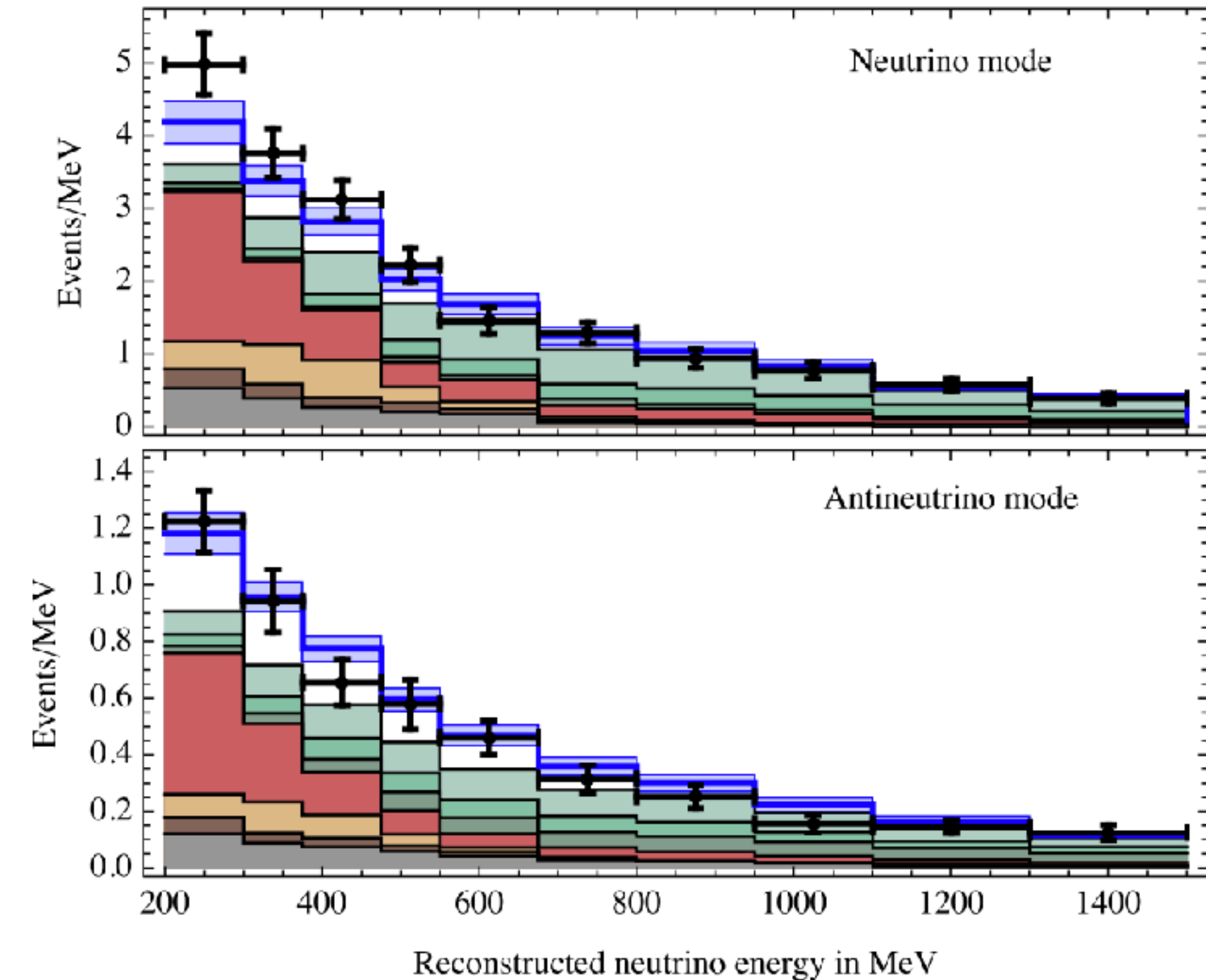


# Some novel developments in BSM searches in neutrino experiments

Bertuzzo, PM et al Phys.Rev.Lett. 121 (2018) 24, 241801



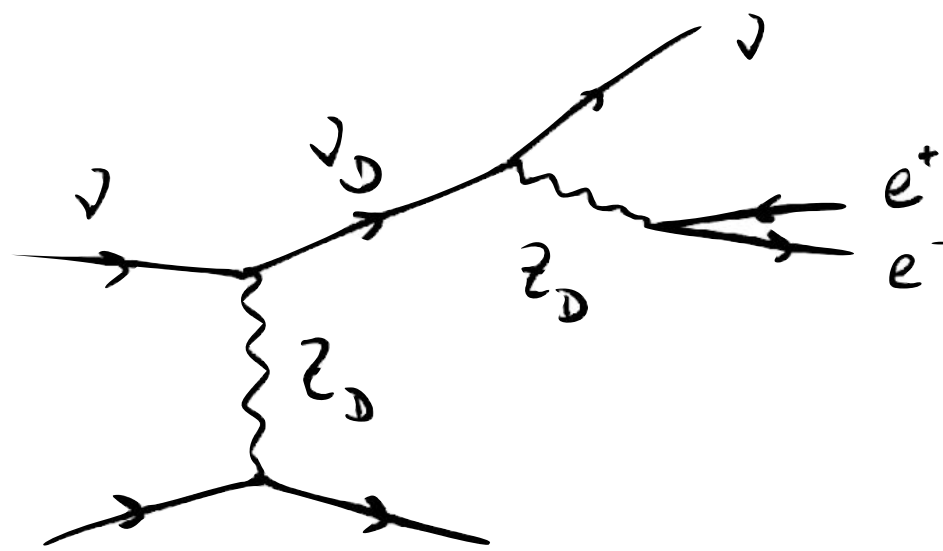
Novel connections between the MiniBooNE anomaly and the origin of neutrino masses



# Beyond the Standard Model

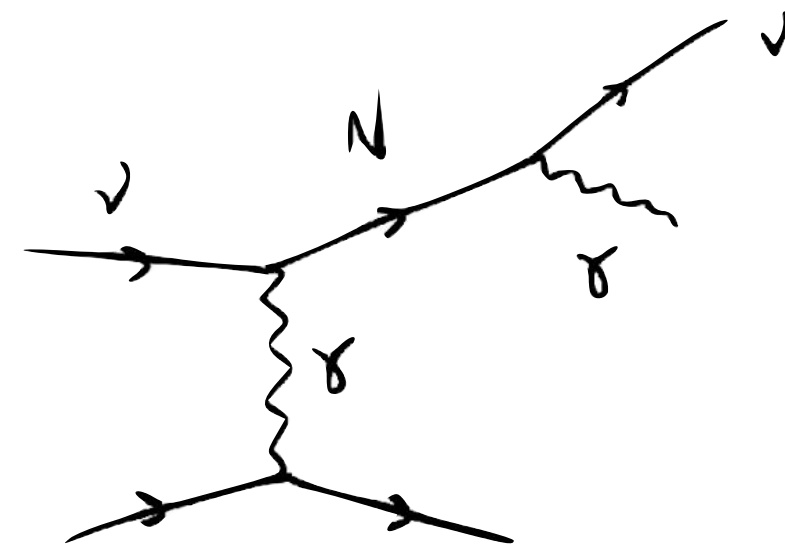


## Dark Neutrinos



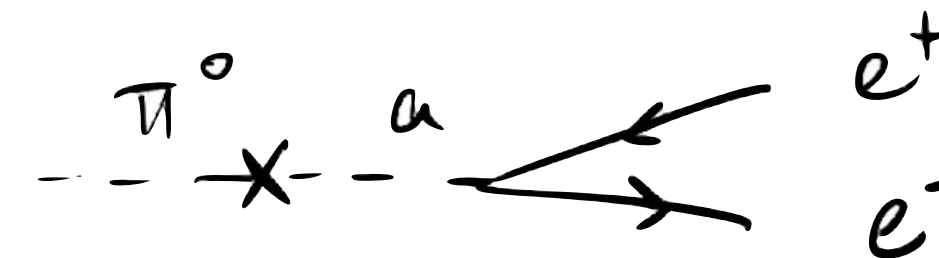
Bertuzzo Jana Machado Zukanovich PRL 2018, PLB 2019  
 Arguelles Hostert Tsai PRL 2019  
 Ballett Pascoli Ross-Lonergan PRD 2019  
 Ballett Hostert Pascoli PRD 2020

## Transition Magnetic Moment



Gninenko PRL 2009  
 Coloma Machado Soler Shoemaker PRL 2017  
 Atkinson et al 2021 Vergani et al 2021

## Axion-like Particles



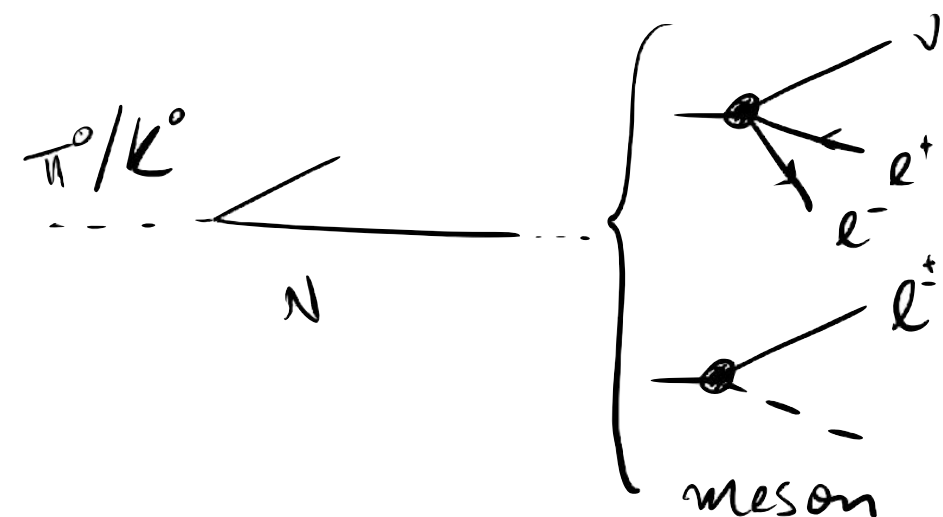
Kelly Kumar Liu PRD 2021  
 Brdar et al PRL 2021

Alternative explanations  
 to the MiniBooNE excess  
 and other BSM scenarios

Not an exhaustive list

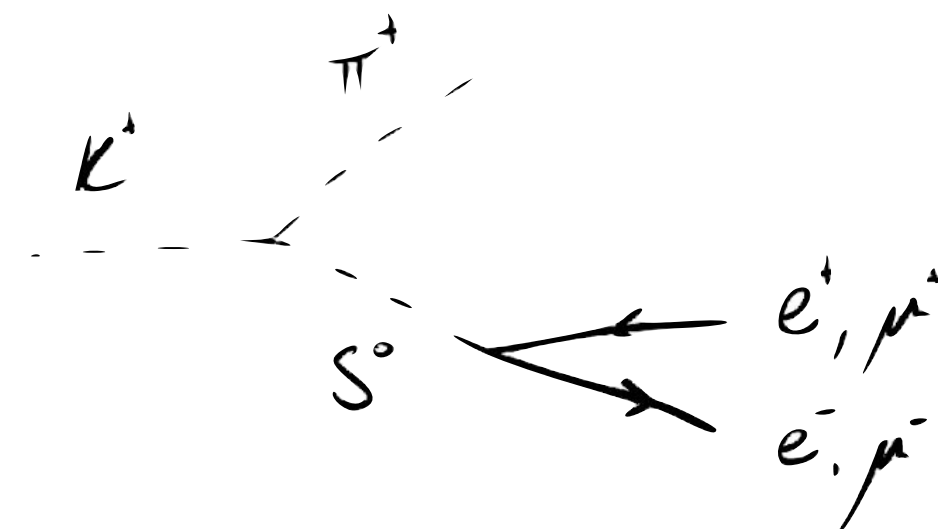
Some diagram credit: Pedro Machado  
 Slide credit: Marco Del Tutto  
 Slide credit 2: Rhiannon Jones

## Heavy Neutral Leptons



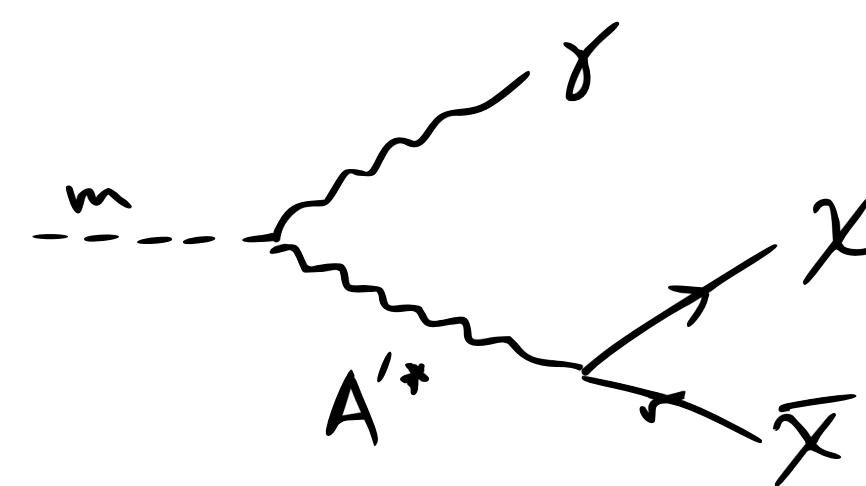
Ballett Pascoli Ross-Lonergan JHEP 2017  
 Kelly Machado PRD 2021

## Higgs Portal Scalar



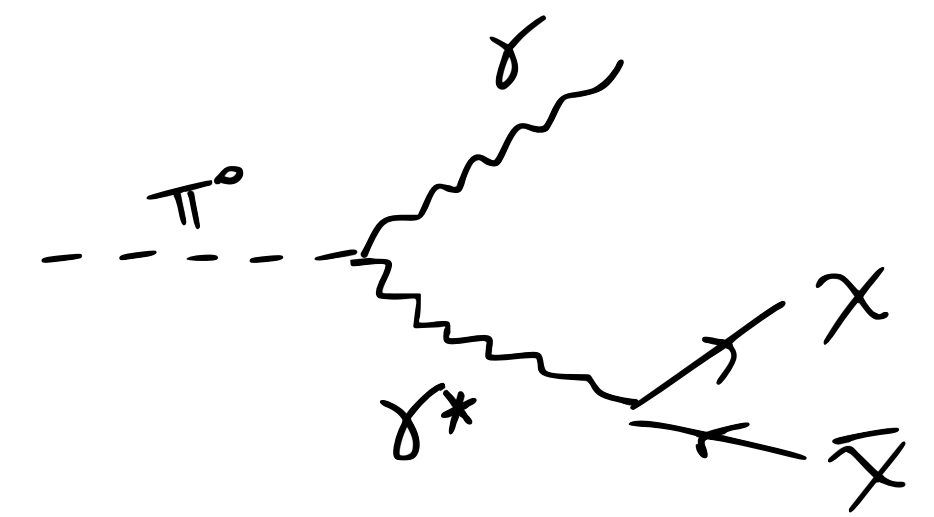
Pat Wilczek 2006  
 Batell Berger Ismail PRD 2019  
 MicroBooNE 2021

## Light Dark Matter



Romeri Kelley Machado PRD 2019

## Millicharged Particles

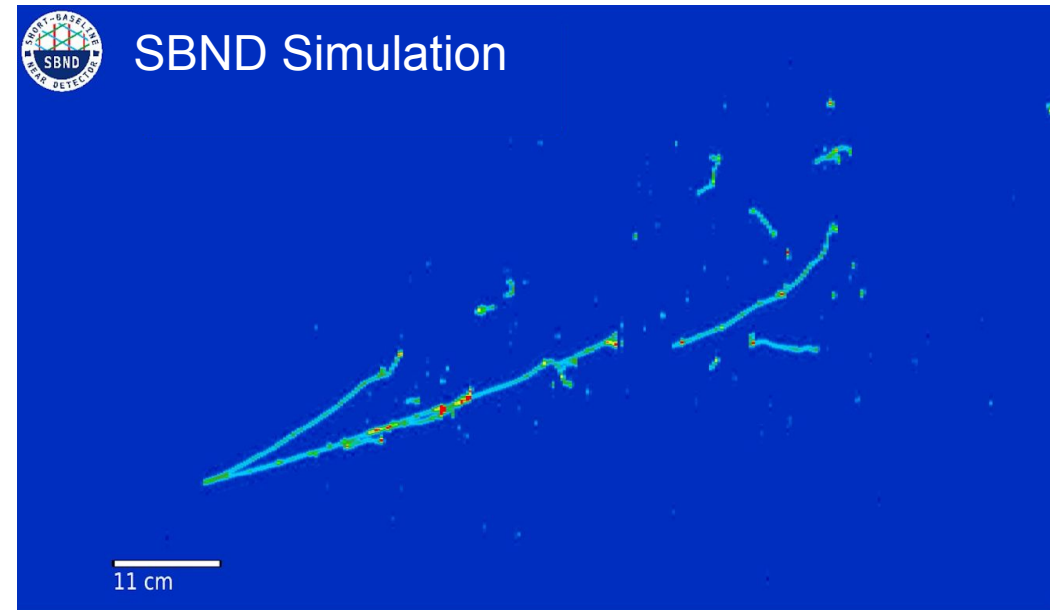


Magill, Plestid, Pospelov, Tsai, PRL 2019  
 Harnik Liu Palamara, JHEP 2019

# BSM signatures in SBND

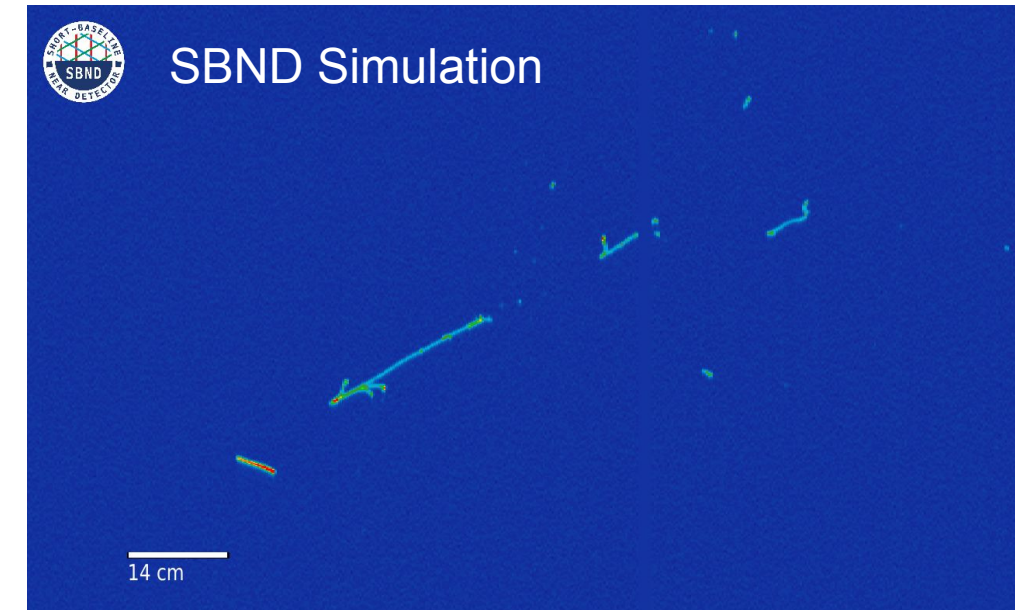


## Dark Neutrinos



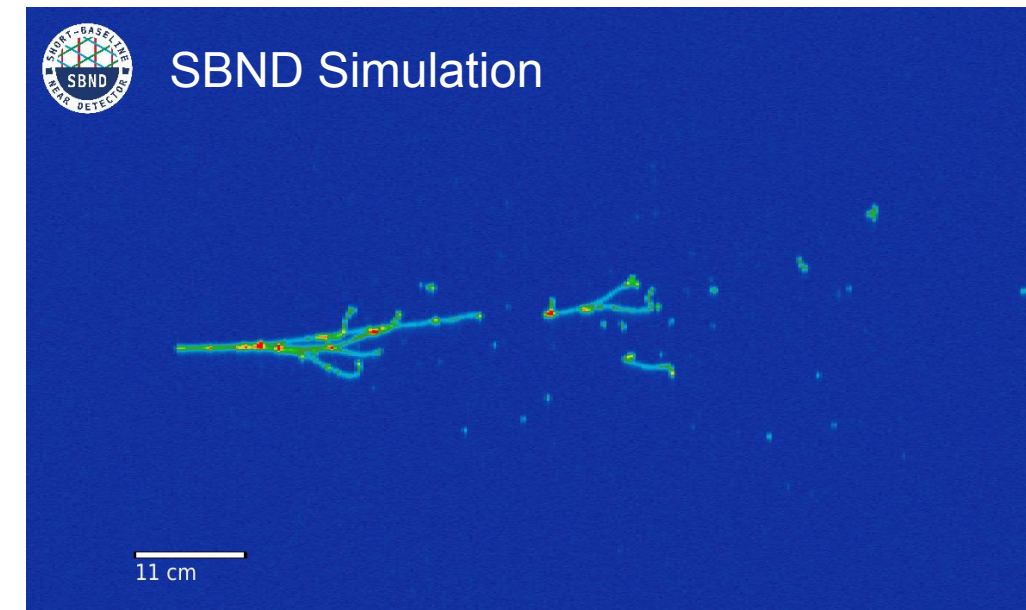
$e^+e^-$  pair with or w/o hadronic activity

## Transition Magnetic Moment



Photon shower and hadronic activity

## Axion-like Particles

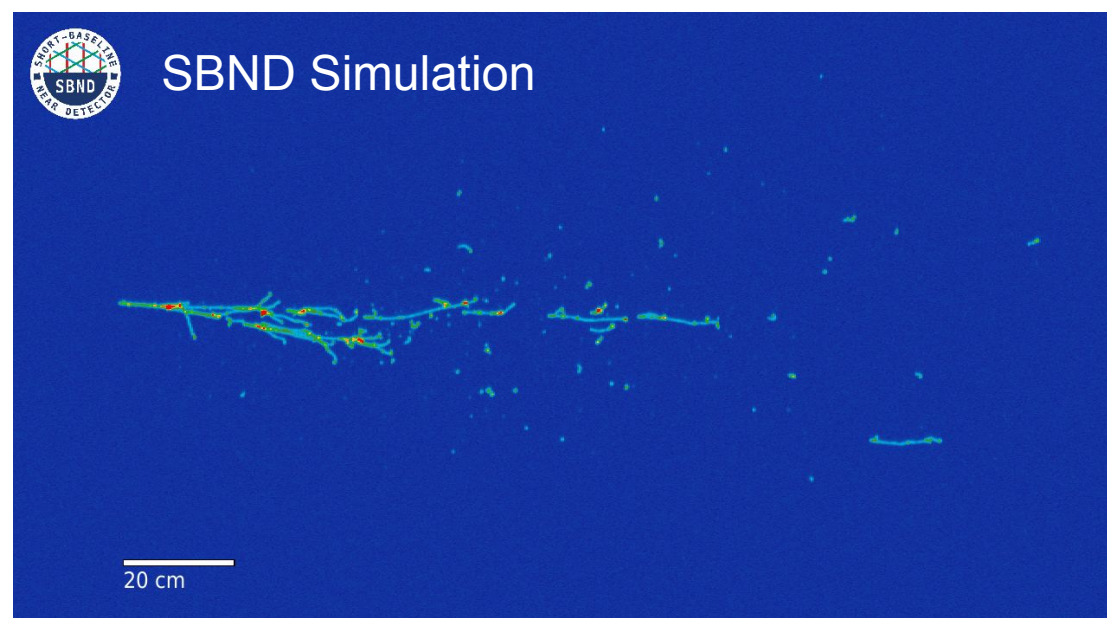


High-energy  $e^+e^-$  or  $\mu^+\mu^-$

Example signatures and event displays for various BSM scenarios

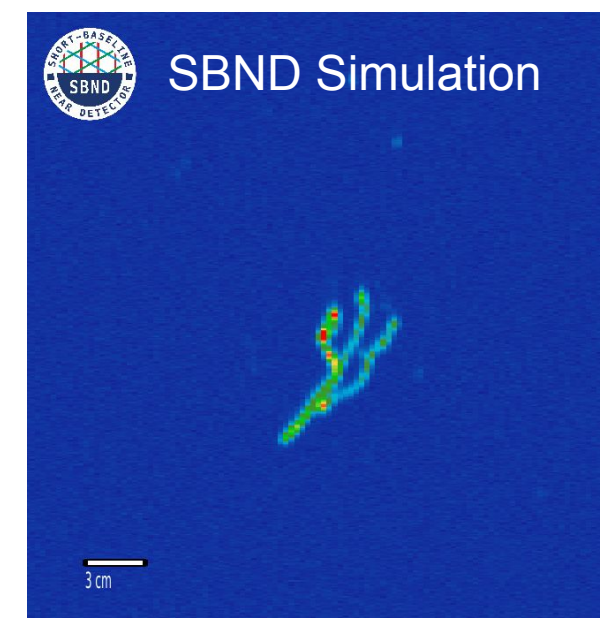
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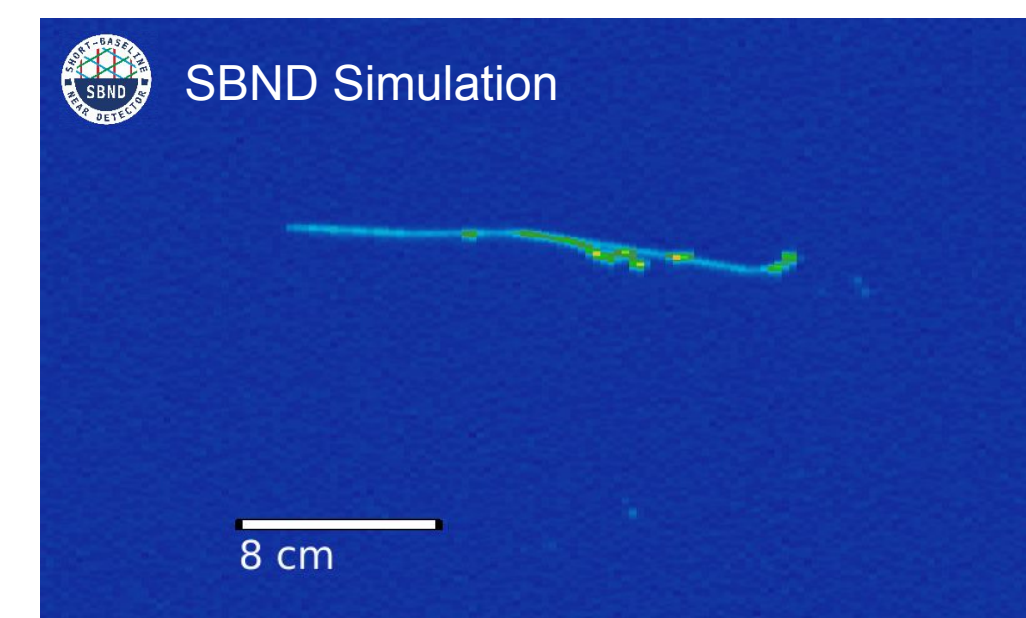
$e^+e^-$ ,  $\mu^+\mu^-$  or  $\mu\pi$

## Higgs Portal Scalar



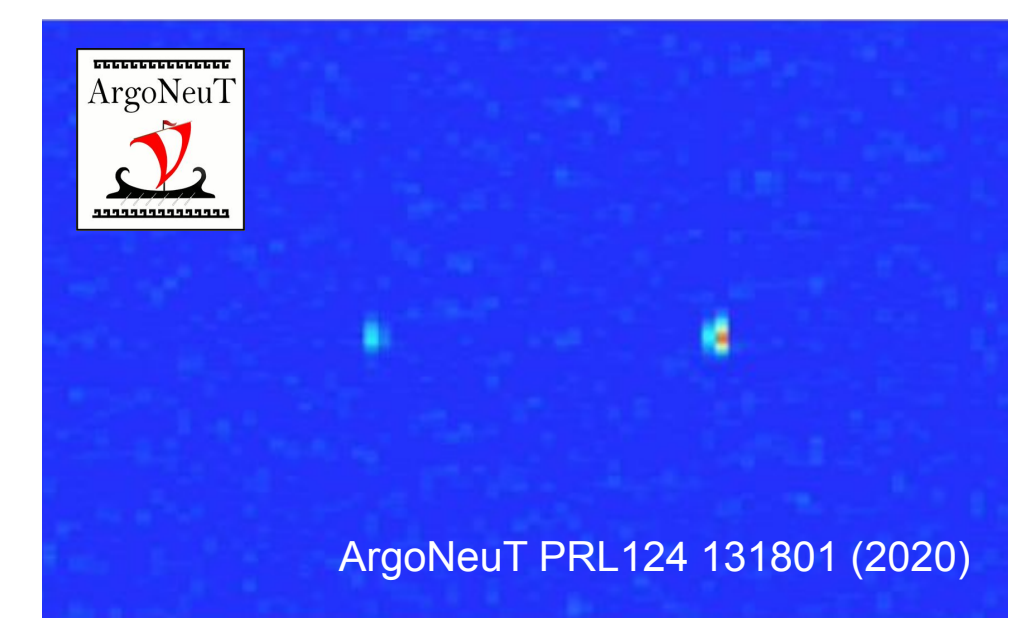
$e^+e^-$  or  $\mu^+\mu^-$ , no hadronic activity

## Light Dark Matter



Electron scattering

## Millicharged Particles



Blips/faint tracks

Take NSI as an example: what does that really mean?

$$\mathcal{L}_{NSI}^{NC} = -2\sqrt{2}G_F\epsilon(\bar{\nu}_L\gamma_\mu\nu_L)(\bar{f}\gamma^\mu f)$$

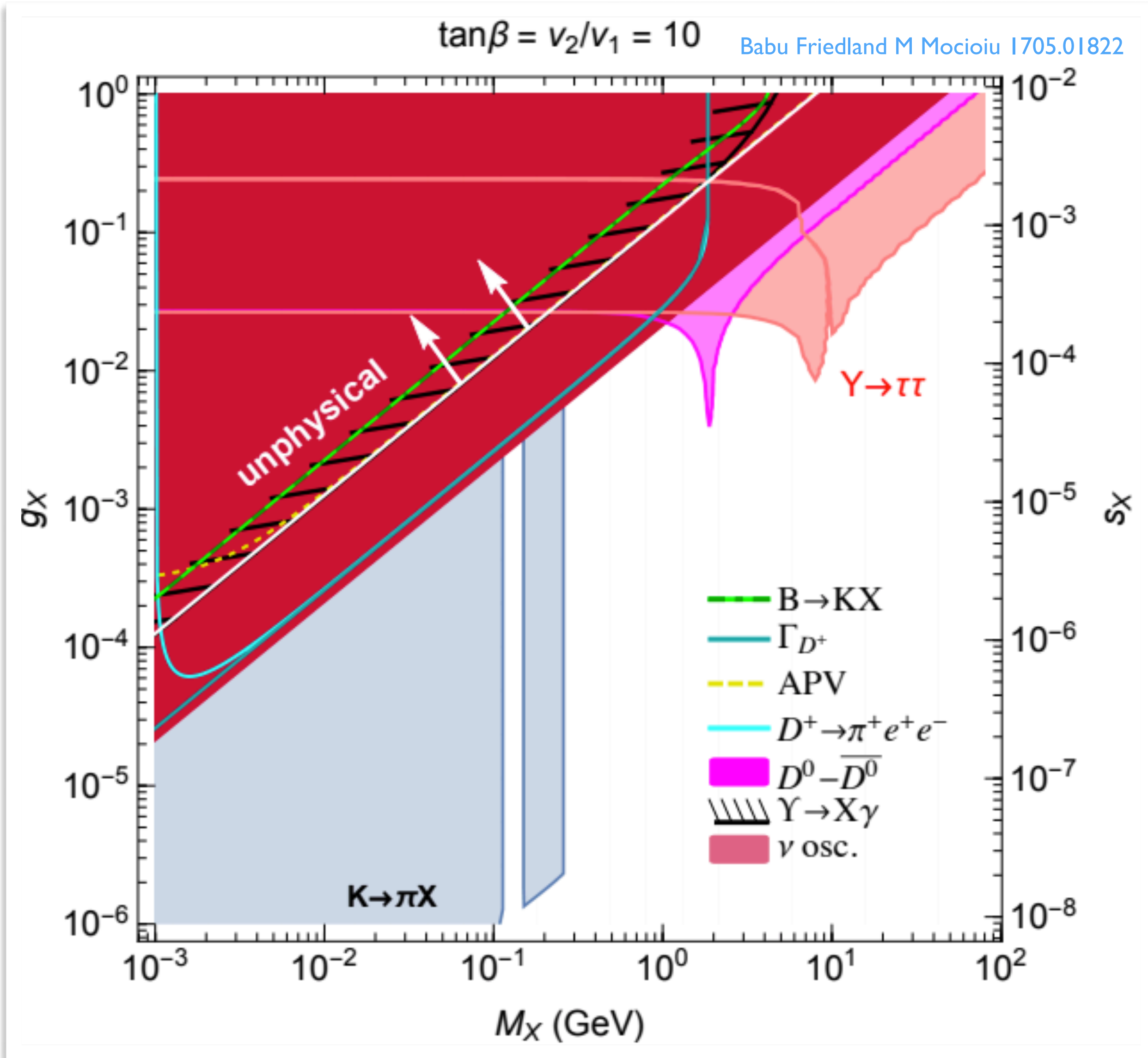
A way to answer that is to build up a full model that gives NSI and see what do you learn with it

Be bold:

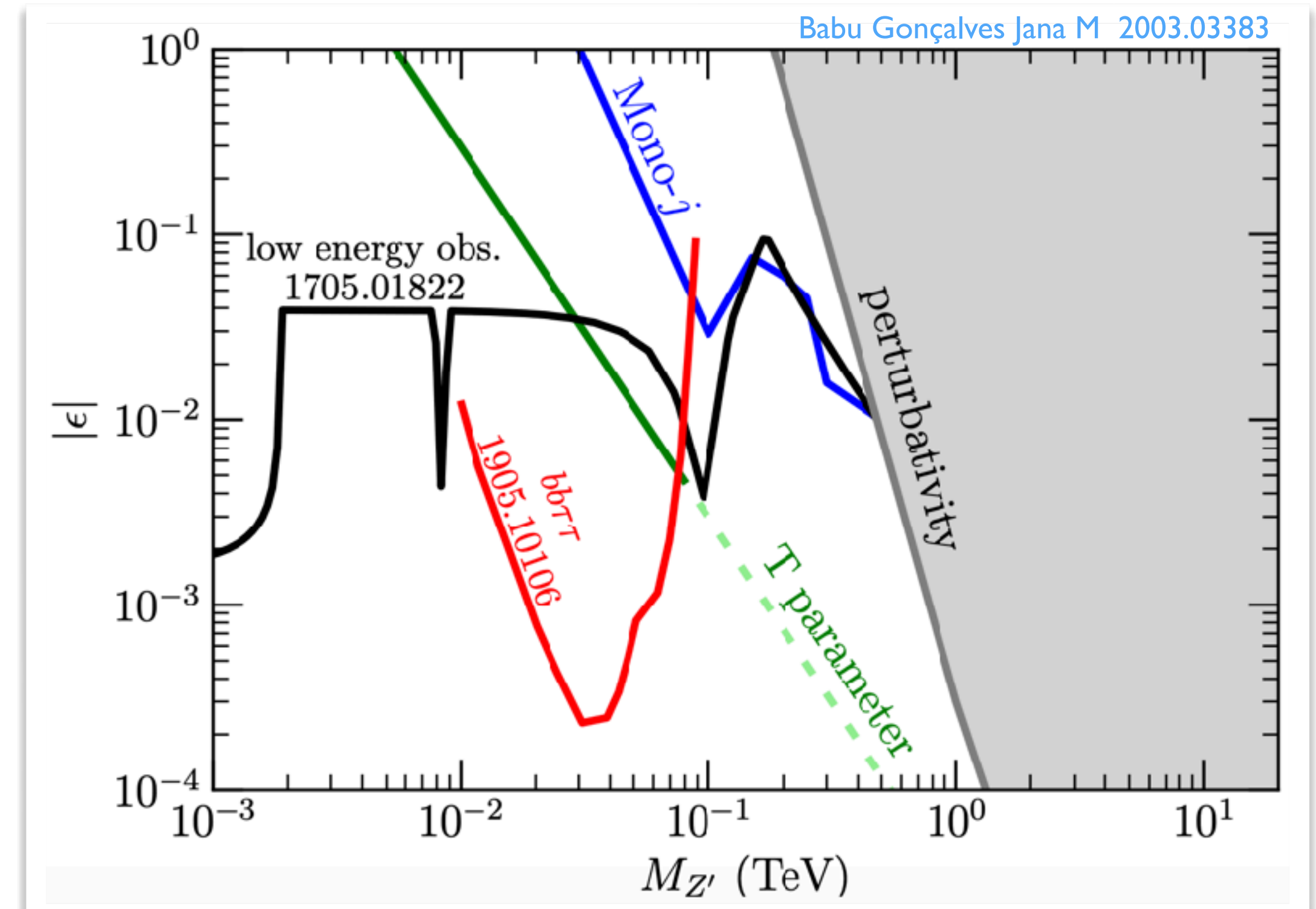
let's have new physics that violates flavor  
in both quark and lepton sector **at a low scale**

Gauge B – L of the third family only

# We should not discard the “old” BSM



Complementarity:  $\nu$  oscillations, meson decay and oscillation, parity violation, kaon physics, LHC, ...



# Beyond **DUNE**?

Given the **Accelerator Complex Evolution** and the large community interest in hosting a Muon Collider in the US in the future, it seems timely to re-address the physics case of a neutrino factory.  
(neutrino factory = a neutrino experiment in which the flux comes from muon decay)  
Or a muon storage ring to measure neutrino-nucleus interactions.

What can we learn with it?  
How much precision neutrino physics can we extract out of that?  
Which BSM searches could shine in a neutrino factory setup?  
What are the complementarity to other experiments?  
What is the physics case beyond neutrinos?

# Conclusions

**We do not know where new physics is**

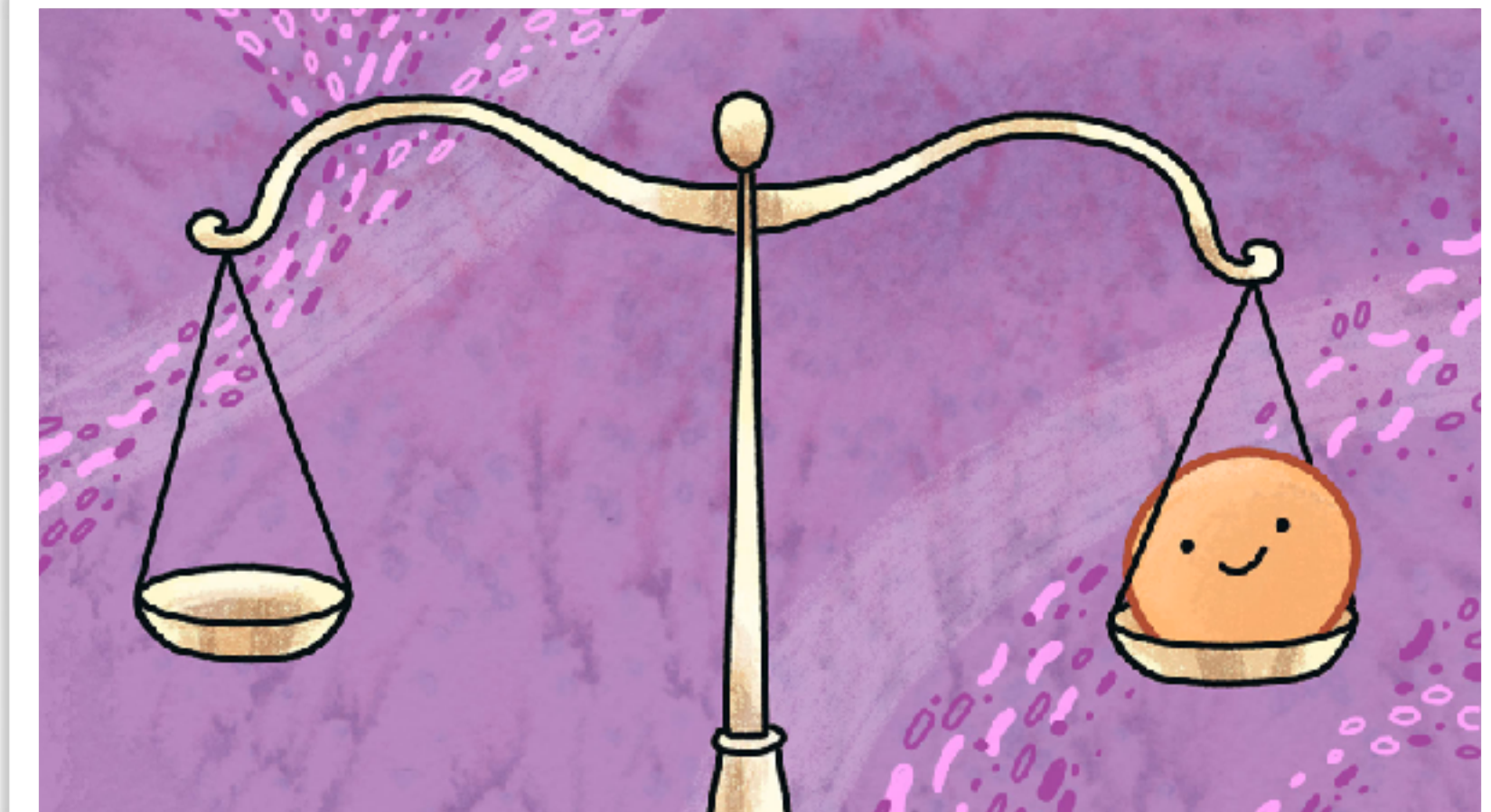
But we know that **there needs to be new physics** that address the outstanding questions of the standard model, in particular the **mechanism of neutrino mass generation**

LArTPCs offer novel opportunities

Improving neutrino-nucleus interaction modeling will further enable this program

Neutrino experiments are multipurpose experiments  
***neutrino experiments >> neutrinos***

*A precision neutrino physics program can stress-test the least known sector of the standard model*

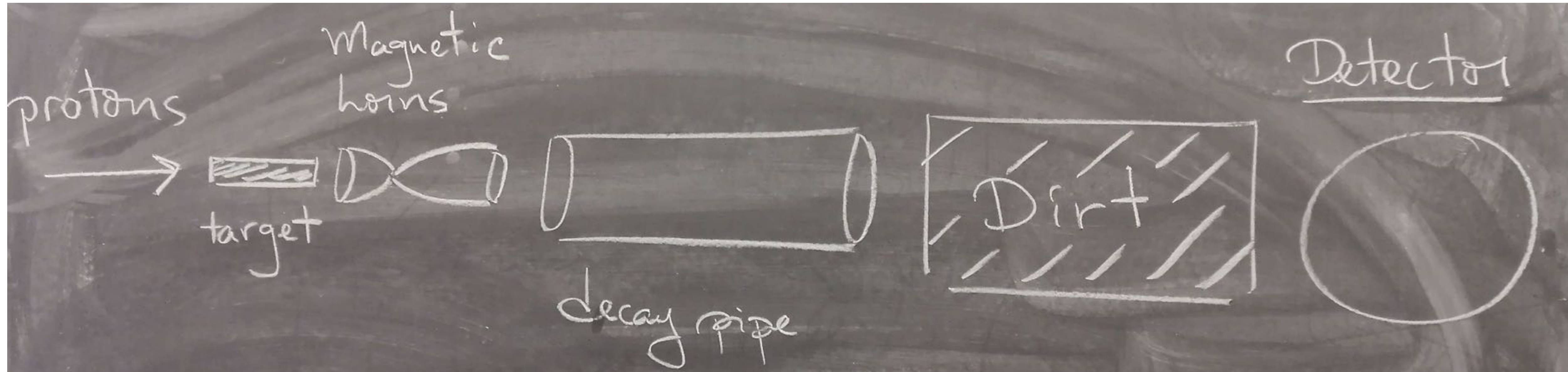


# Backup



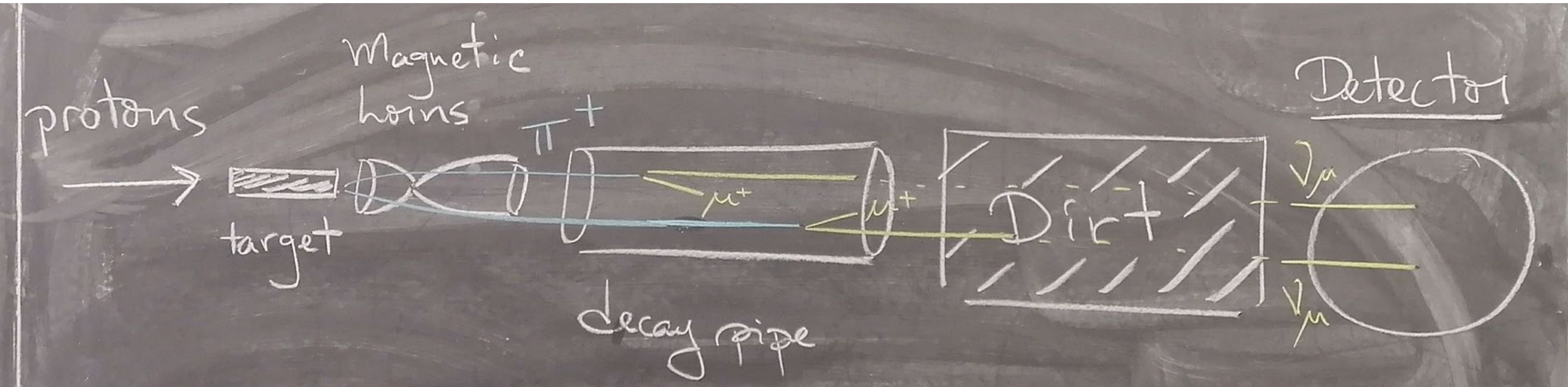
# Dark matter at DUNE

Based on de Romeri Kelly PM 1903.10505



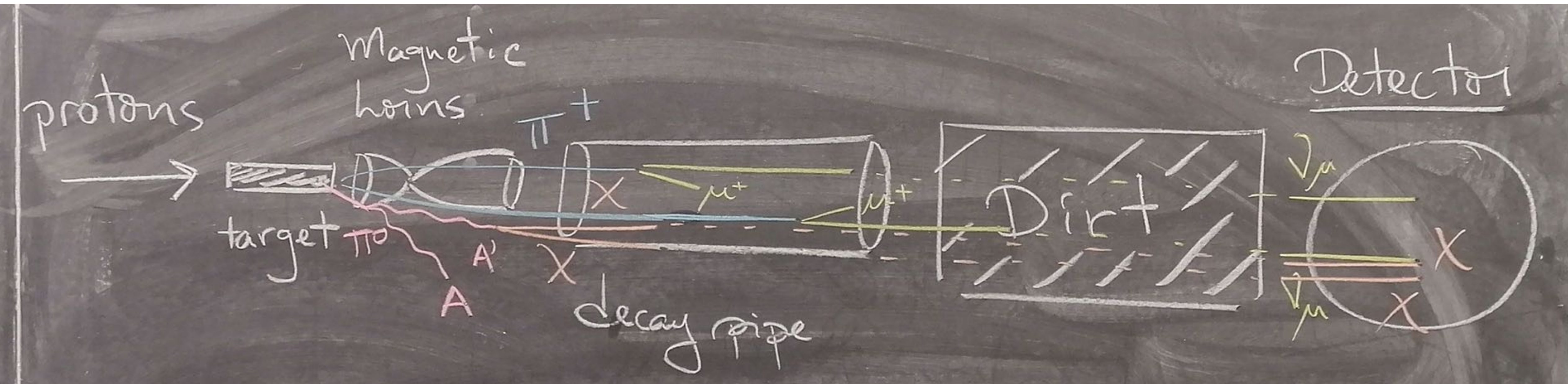
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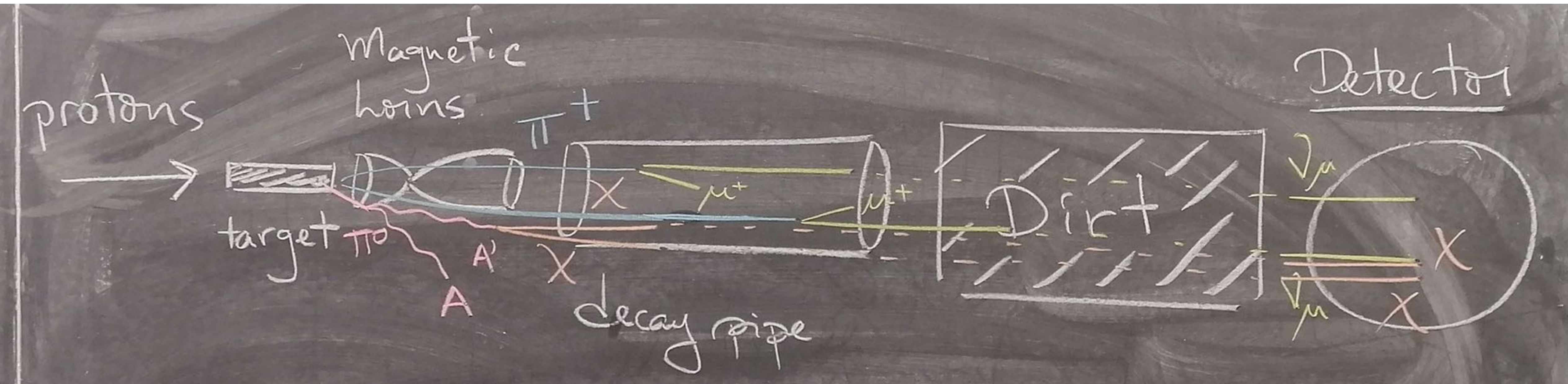
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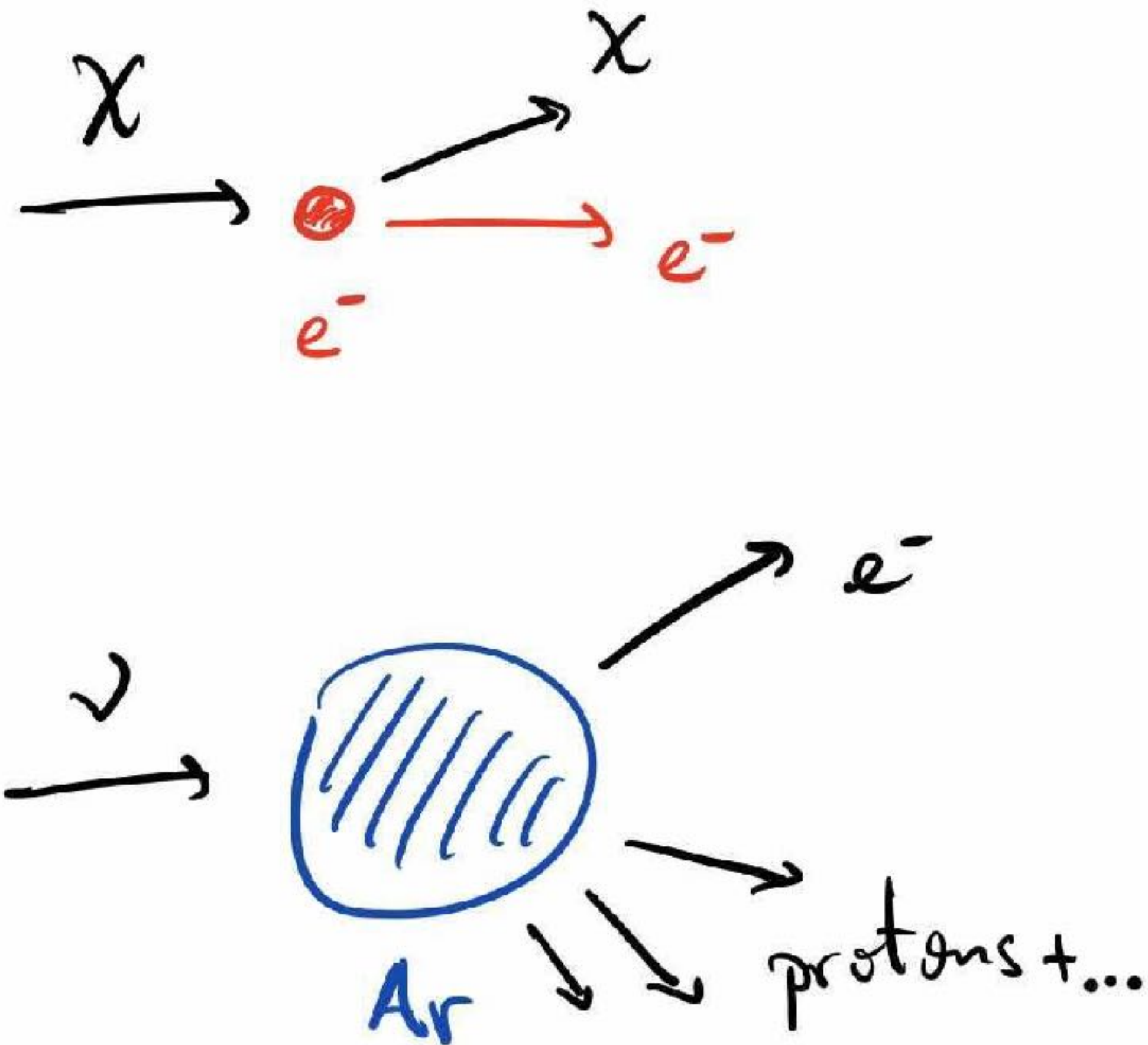


Maybe dark matter is been produced in neutrino beams as of right now

To search for it, we need very detailed detectors to disentangle the DM signal from the neutrino scattering background

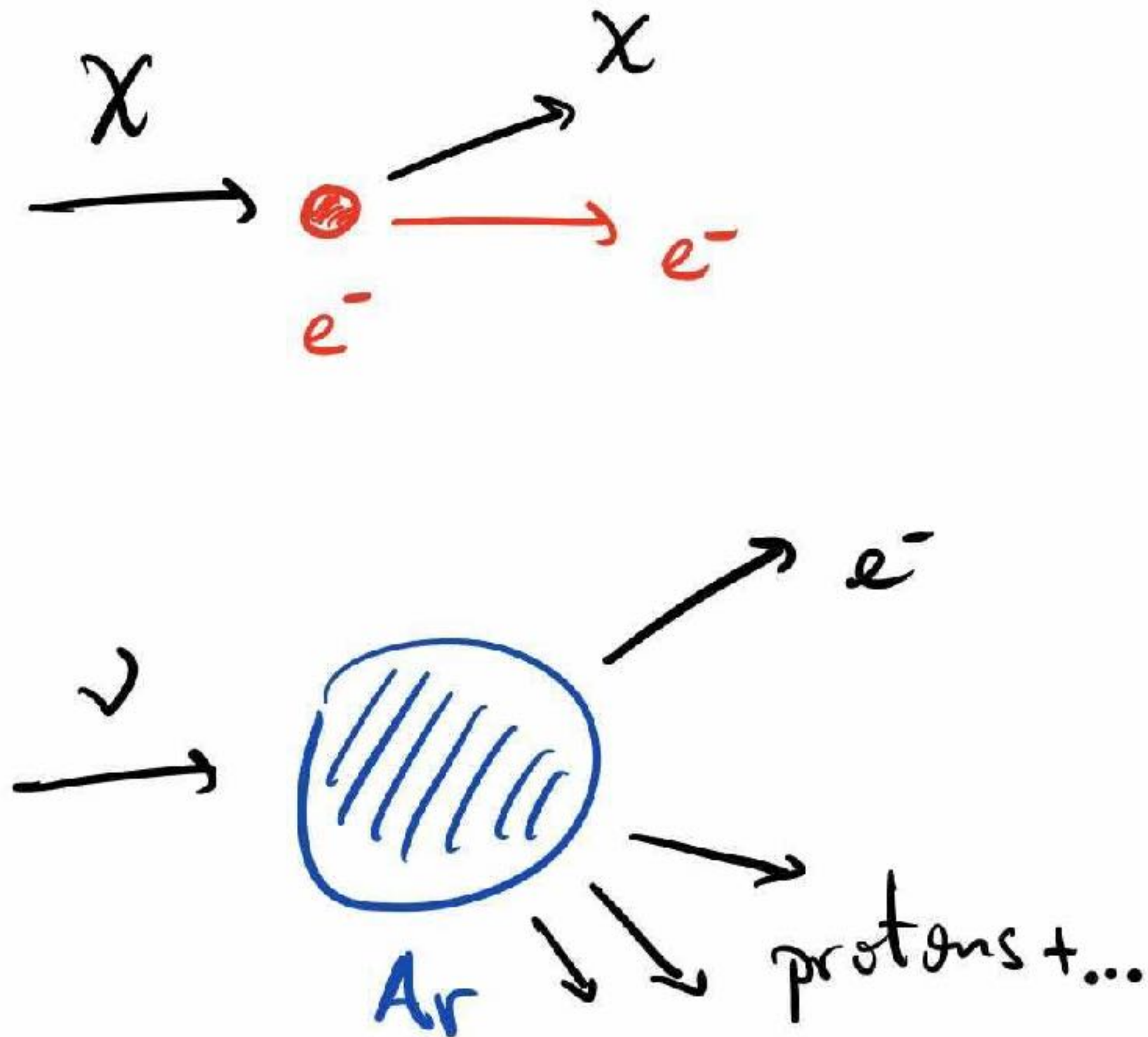
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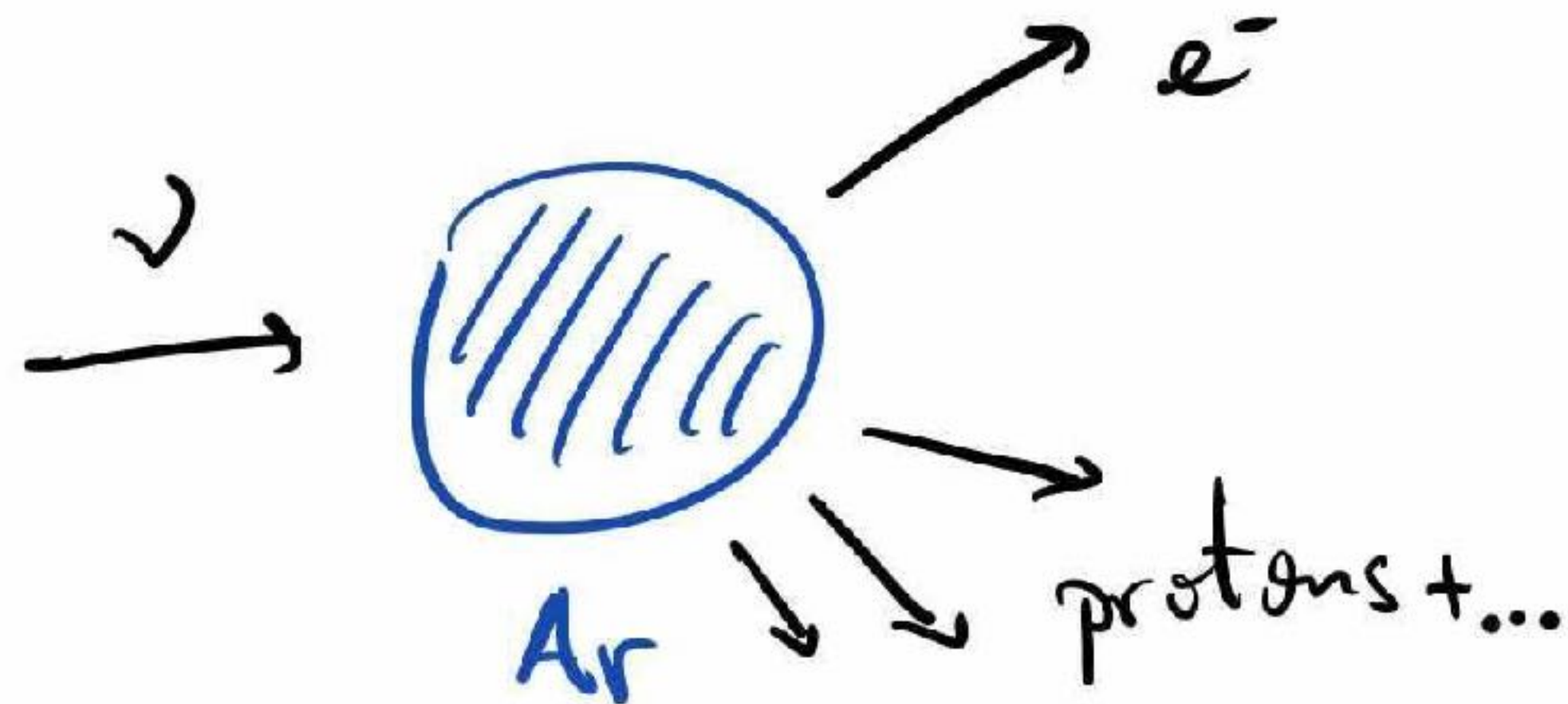
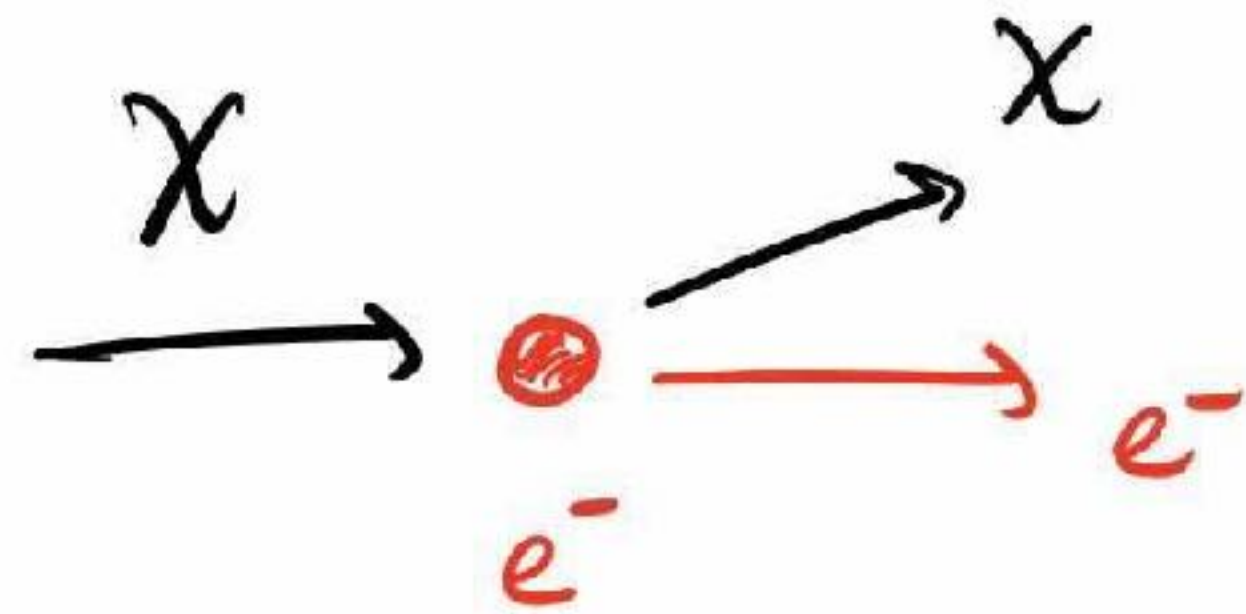
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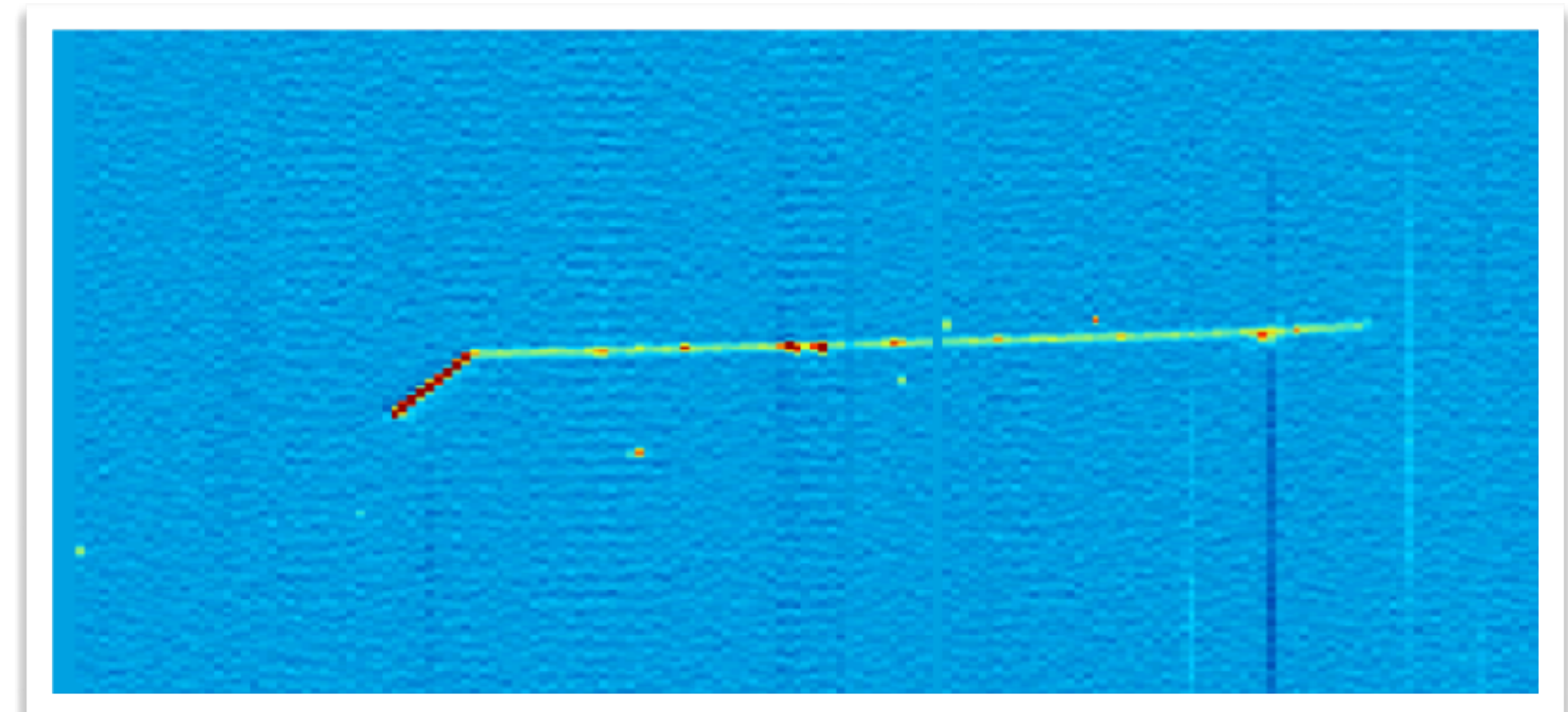




When DM scatters off electrons, the outgoing electrons are very forward

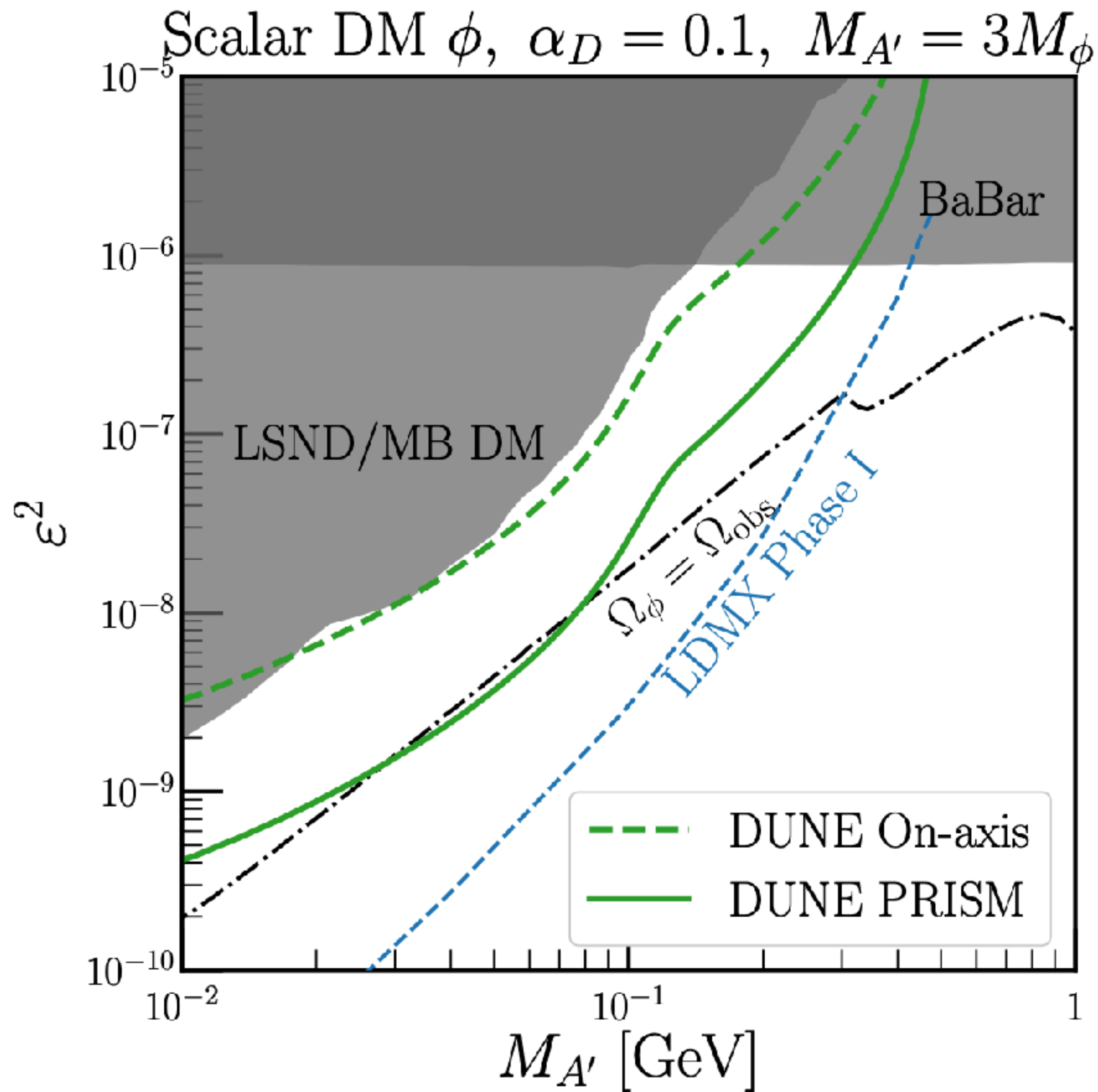
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LArTPC topological capabilities can play a crucial role in rejecting backgrounds



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