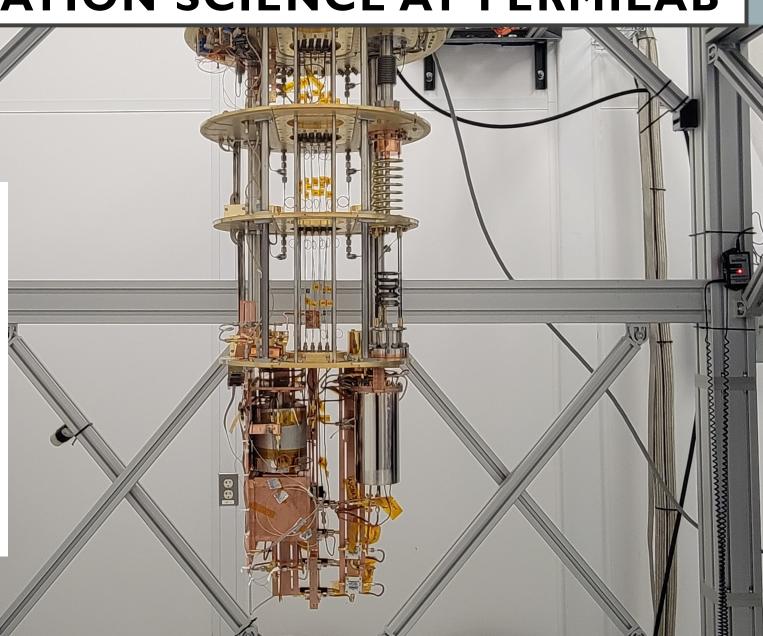
QUANTUM INFORMATION SCIENCE AT FERMILAB

Rakshya Khatiwada

Illinois Institute of technology

& Fermilab 06/29/2023



OUTLINE

Quantum Science Program at Fermilab

→ DOE Quantum Science Enabled Discovery (QuantiSED)

 \rightarrow Quantum Science Center (QSC)

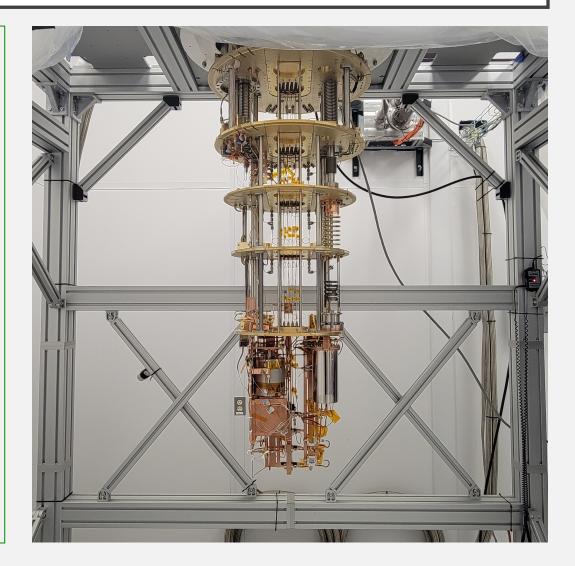
one of the five DOE Quantum Science Centers funded under National Quantum Initiative act passed by congress (2018). Led by Oak Ridge National Lab

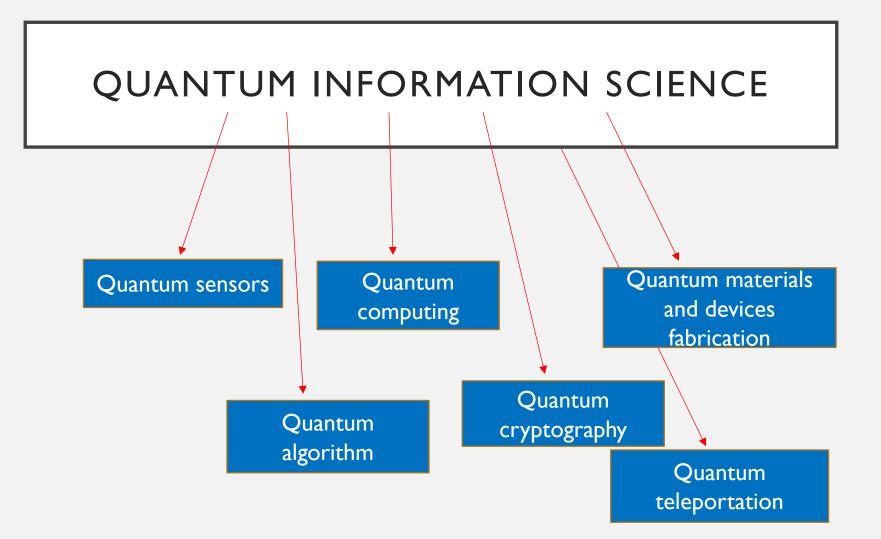
All of this is relatively new and very future focused!

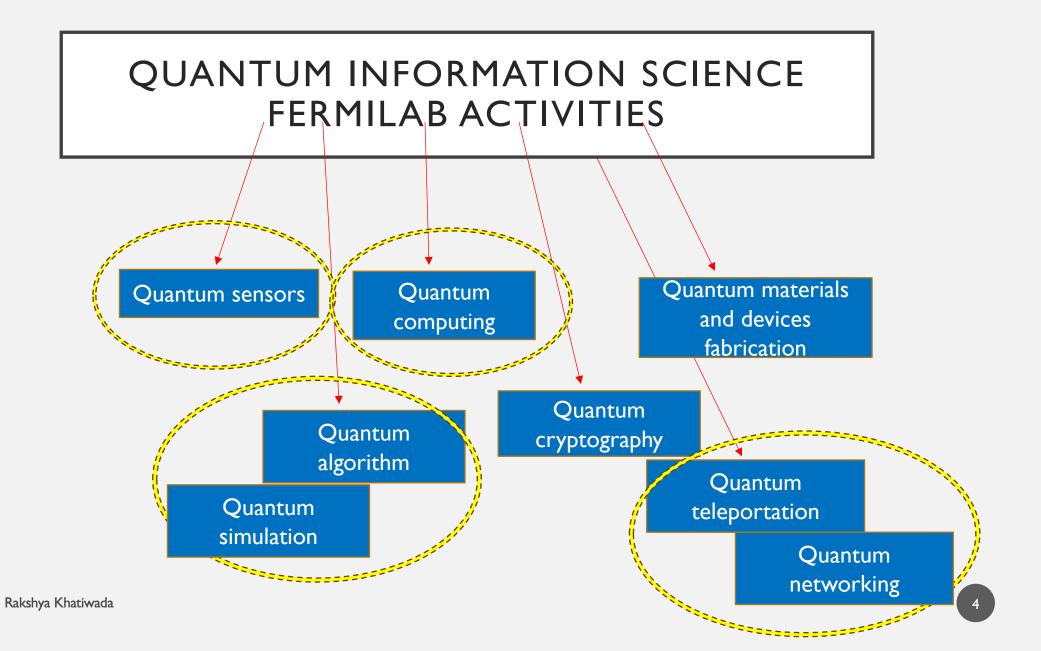
"New Ideas for future projects at Fermilab"

User's meeting 2023 theme

Rakshya Khatiwada

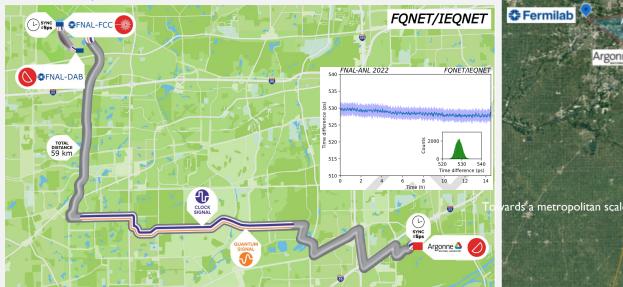




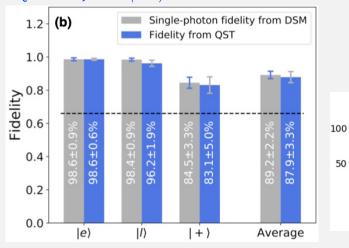


FERMILAB QUANTUM NETWORK (FQNET)

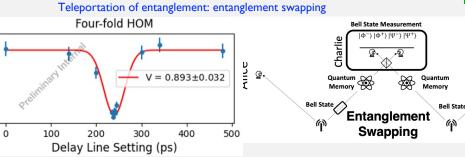
Picosecond level entanglement distribution and clock synchronization *EEE JQE 59*,1-7 (2023) doi: 10.1109/JQE.2023.3240756



High-fidelity quantum teleportation PRX QUANTUM 1, 020317 (2020) Quick-Quantum Network (QICK-QN): <u>https://arxiv.org/abs/2304.0</u>









- → Collaboration: FNAL, Argonne, Caltech, Northwestern University.
- → Illinois Express Quantum Network (IEQNET)
- → Hosts photonics and optical fiber based local quantum teleportation nodes, between Chicagoland institutions
- → High fidelity quantum teleportation was achieved between multiple nodes (50 km apart) at Argonne and Fermilab (2022) using entangled photons
- → Picosecond level entanglement distribution and clock synchronization between two nodes.
- → Step forward in building Quantum networking over metropolitan distances

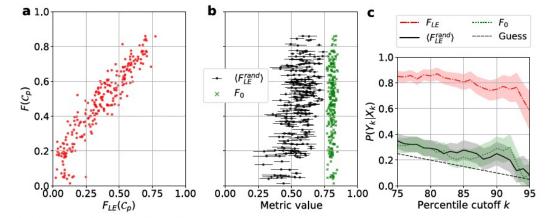
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QUANTUM COMPUTING AND ALGORITHMS

Qubit assignment using time reversal

Evan Peters,^{1, 2, 3, *} Prasanth Shyamsundar,¹ Andy C. Y. Li,¹ and Gabriel Perdue^{1, †}
 ¹Fermi National Accelerator Laboratory, Batavia, IL 60510
 ²Institute for Quantum Computing, University of Waterloo, Waterloo, Ontario, N2L 3G1, Canada
 ³Department of Physics, University of Waterloo, Waterloo, Ontario, N2L 3G1, Canada
 (Dated: October 8, 2022)

- In order to run quantum circuits on hardware, we need to map logical qubits to the physical hardware (qubit assignment problem).
- For NISQ-era devices, we need to perform this task in a noise-aware manner.
- Sub-tasks:
- · Efficiently estimate the performance of different qubit choices (Loschmidt Echoes)
- Efficiently search the space of possible qubit assignments (simulated annealing)



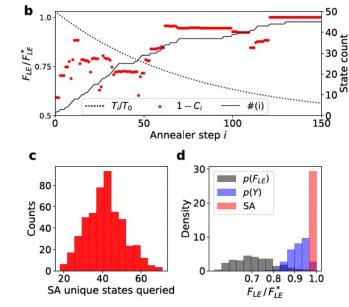
- 8 qubit GHZ circuit on Google Rainbow (23 qubits)
- Loschmidt Echo (red) fidelity estimator using the target circuit tracks more closely to true
 process fidelity for a given qubit assignment than a random circuit estimator (black) or
 historical calibration data (green).
- Far right plot shows the probability of an assignment truly being within the kth percentile of fidelities given a metic observation in the kth percentile.



DOE: Quantum Information Science Enabled Discovery

(QuantISED)

Key funding support driving pioneering investigations of quantum information in High Energy Physics



- Simulated annealing is a good optimization strategy.
- Top panel shows the performance progression one run of SA.
- Bottom left shows the number of unique states queried in a 10,000 circuit simulation.
- Bottom right shows the fidelities over all assignments (grey), random sampling (using SA query counts and keeping the best, blue) and simulated annealing (red)

QUANTUM SENSING

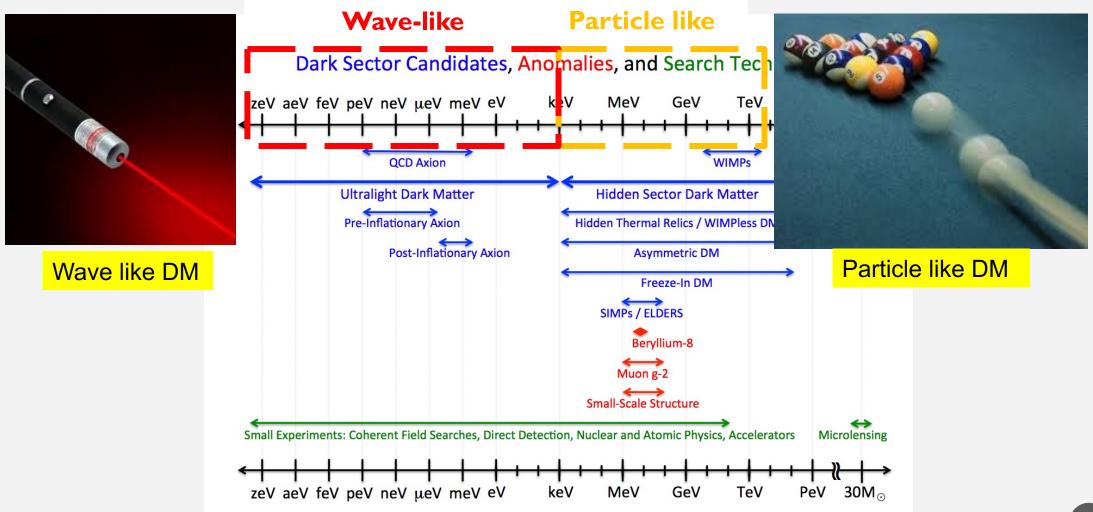
QUBIT BASED DETECTORS FOR ULTRALIGHT AND LIGHT DARK MATTER SEARCHES

ADVANTAGES OF QUBITS OVER CURRENT DARK MATTER SEARCH TECHNOLOGY

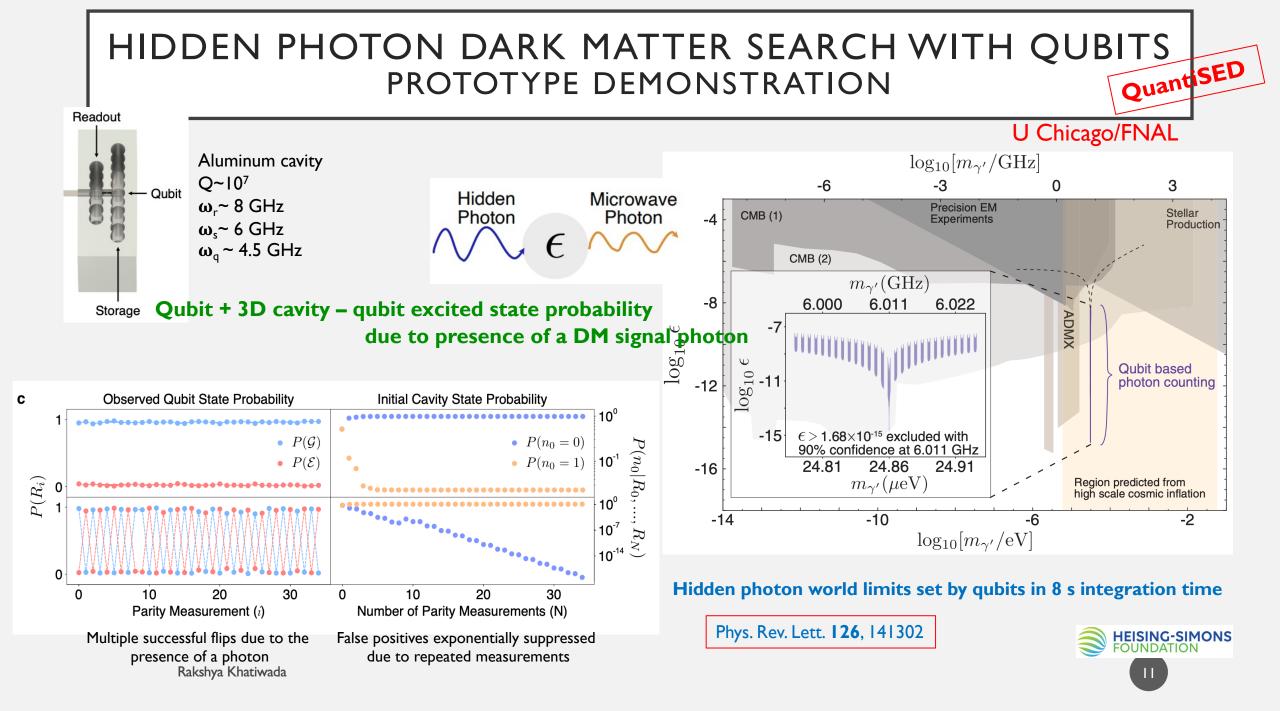
Sensitivity to sub-eV energy threshold

- Dark Matter can be coupled to *phonons* (lattice vibrations of the substrate) or *as photon absorption* in the superconducting part of qubit
- Broken cooper pairs generated in the qubit when $E > 2\Delta$ (superconducting bandgap energy) → can be used as Dark Matter signal
- \circ **Easy signal readout** with a qubit readout protocol (T₁, T₂, charge parity measurements)
- Qubit superconducting systems in *mK cryostat*, ideal for *thermal noise reduction* for Dark Matter searches.

DARK MATTER CANDIDATES



WHAT HAS BEEN DEMONSTRATED WITH QUBIT BASED DARK MATTER DETECTORS?



NOVEL TECHNOLOGY DEVELOPMENT UTILIZING QuantiSED **QUANTUM TOOLS**

400 200

-2500 -2000 -1500 -1000 -500

Ramsey of qubit

Cavity freq to flu

0.0

Flux_quantum

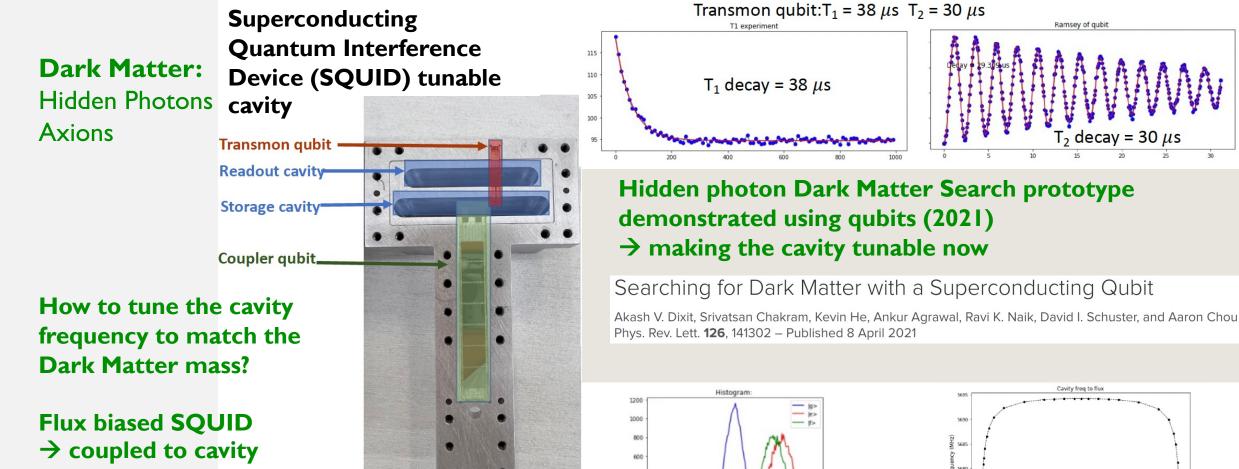
0.2

5670

-0.4

-0.2

 T_2 decay = 30 μ s



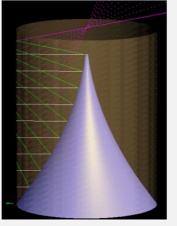
 \rightarrow tunes the frequency

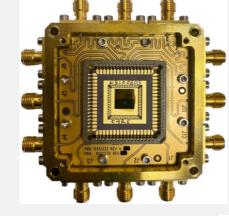
DARK MATTER DETECTOR R&D

Broadband Axion Antenna (BREAD) and Superconducting Nanowire Single Photon Detector for Dark Matter

Caltech, JPL, FNAL etc.

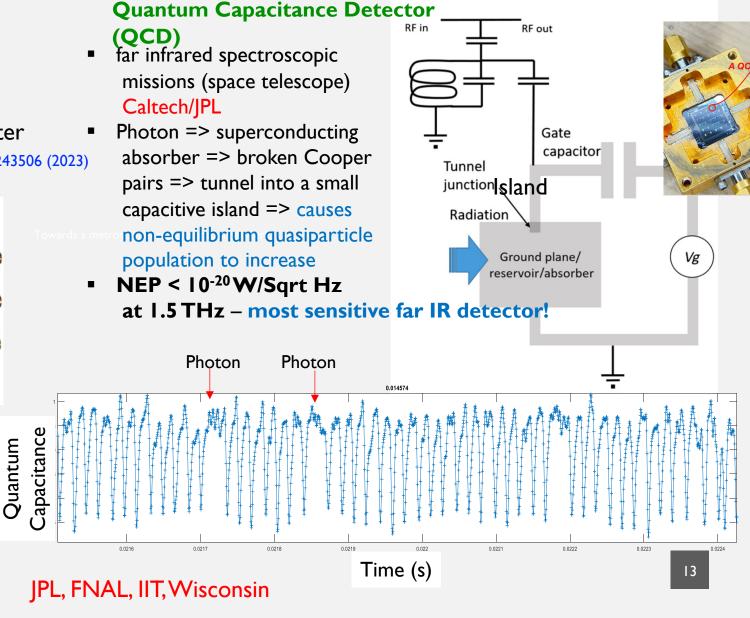
SNSPD R&D for axion detection experiments Appl. Phys. Lett. 122, 243506 (2023)





BREAD concept

SNSPD



QuantiSEL

QUANTUM SCIENCE CENTER (QSC)

 \rightarrow High throughput cryogenic facility for development and readout of

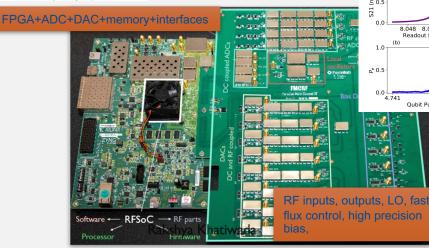
large arrays of quantum sensors (Commissioned Fall 2022)

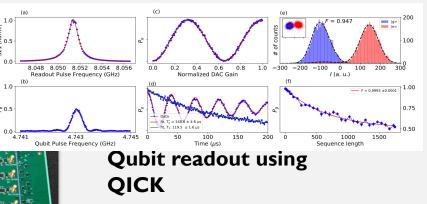
Dark Matter and radiation sensing with sub-eV single photon/phonon resolution qubit-based sensors

→multiplexed readout of array of quantum sensors using FNAL developed Quantum Instrumentation Control Kit (QICK) based on Radio Frequency on Chip (RFSoC) fpga technology.

- >30 labs in the US and abroad
 20 science talks in APS March

meeting!



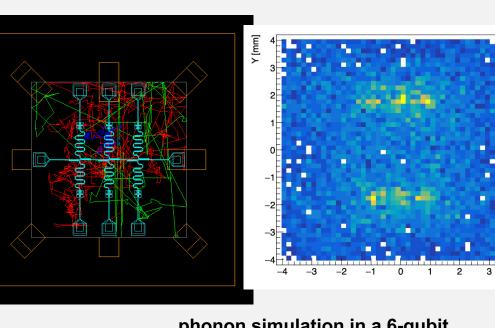


Open source: including hardware schematics/layout firmware, software. See <u>https://github.com/openq</u> uantumhardware





phonon simulation in a 6-qubit silicon chip using G4CMP



- Application of particle physics simulation tools like G4CMP • to understand qubits (various substrates and geometry)
- Cryogenic photon source development (**0.62 6.9 eV**)

UNDERSTANDING ENERGY DISSIPATION IN QUBITS

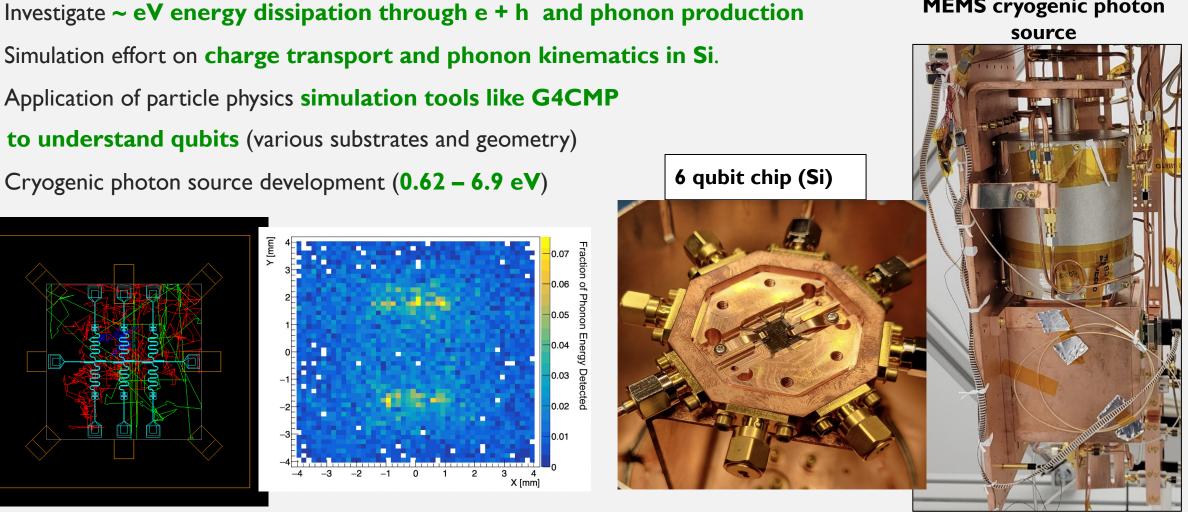
raction of Phonon 0.06

Energy).04

0.03 Detected

0.01

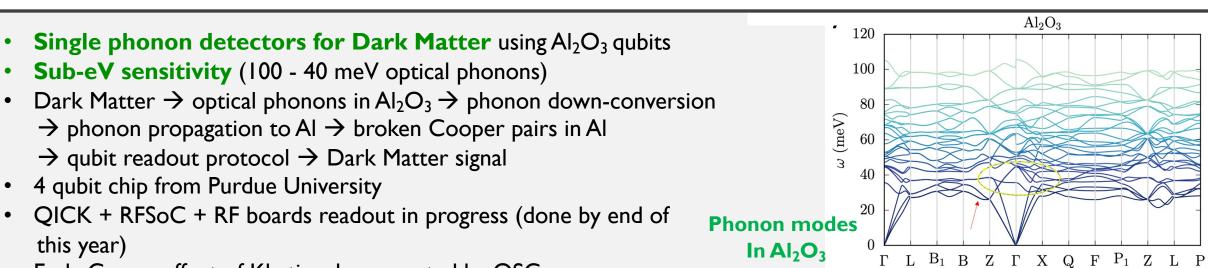
0.05



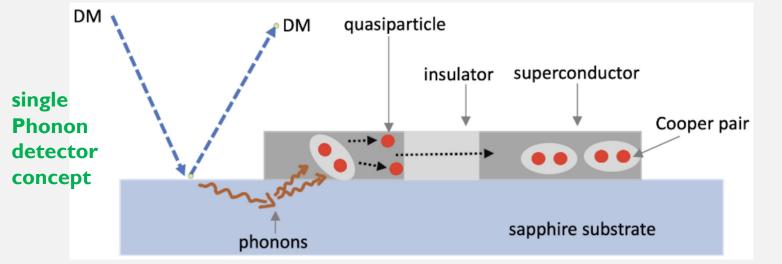
MEMS cryogenic photon

QSC

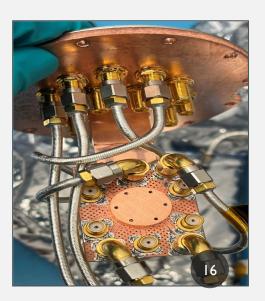
QUBITS FOR LOW ENERGY THRESHOLD DARK MATTER DETECTION



• Early Career effort of Khatiwada supported by QSC



4 qubit chip array In LOUD



Brillouin zones

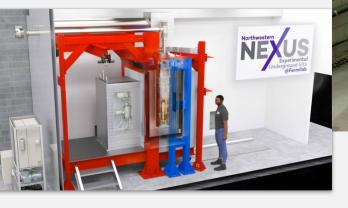
UNDERGROUND, LOW BACKGROUND FACILITY AT FERMILAB FOR FUNDAMENTAL SCIENCE

UNDERGROUND FACILITY AT FERMILAB

Neutrino tunnel at Fermilab: perfect place to study radiation effects on qubits

Rakshya Kha

NEXUS facility for SuperCDMS



QUIET low background facility by Quantum Science Center (QSC)

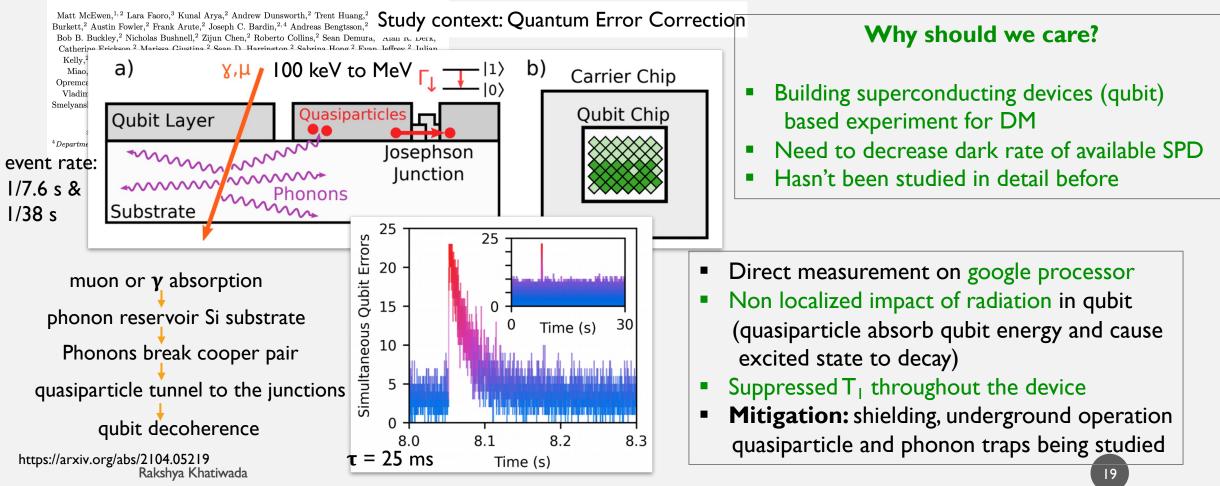
Dark Matter Impact of radiation in superconducting detectors



IMPACT OF COSMIC AND TERRESTRIAL RADIATION ON QUANTUM COMPUTERS

High energy radiation: source of quasiparticle in qubits

Resolving catastrophic error bursts from cosmic rays in large arrays of superconducting qubits

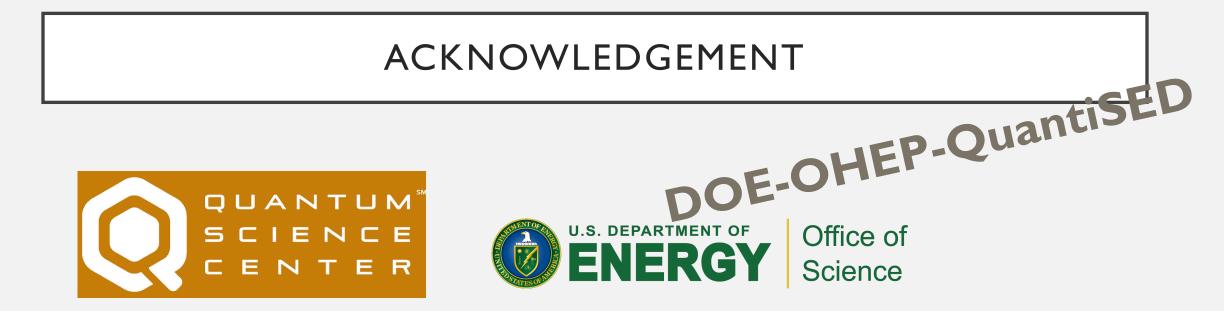


SUMMARY

• Various activities under Quantum Information Science at FNAL

 \rightarrow FNAL at the forefront of novel Quantum Science based research

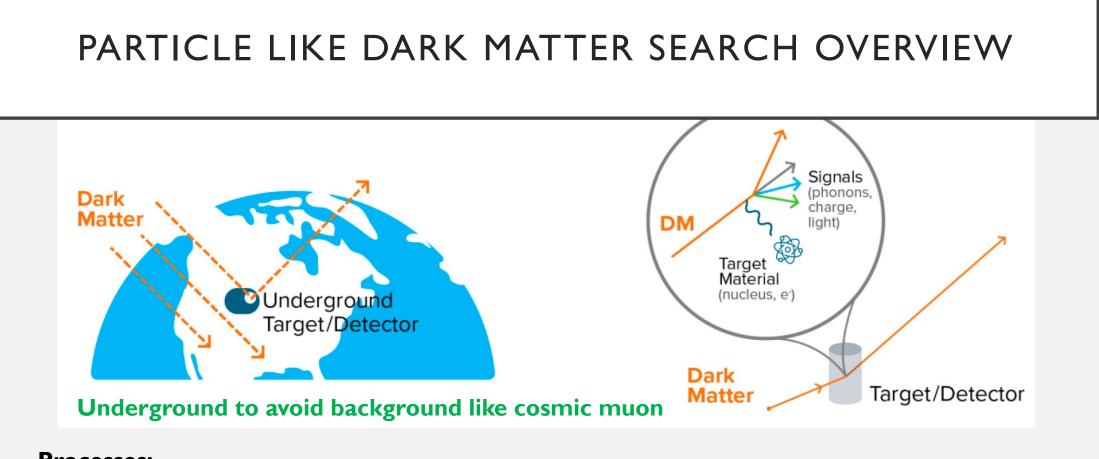
- Aboveground and Underground cryogenic facilities at Fermilab great platforms for studies of impact of radiation in superconducting devices and quantum sensing for Fundamental science
- Dark Matter community developed resources and expertise being utilized for Quantum Science research
- These activities expected to expand further in the coming years



This material is based upon work supported by the U.S. Department of Energy, Office of Science, National Quantum Information Science Research Centers, Quantum Science Center.

Fermilab

• This manuscript has been authored by Fermi Research Alliance, LLC under Contract No. DE-AC02-07CH11359 with the U.S. Department of Energy, Office of Science, Office of High Energy Physics.



Processes:

- --DM Scattering off of nuclei
- --DM Scattering off of electrons
- *Fraction of DM Energy transferred to the target material (nuclear, electron recoil)

--<u>Absorption of DM</u> DM Energy absorbed by the target material

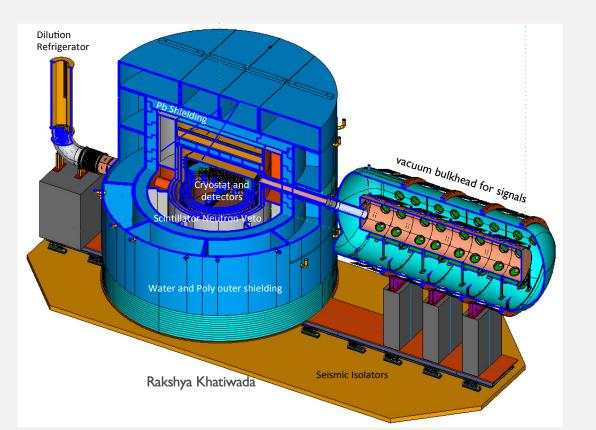


SOME EXPERIMENTS

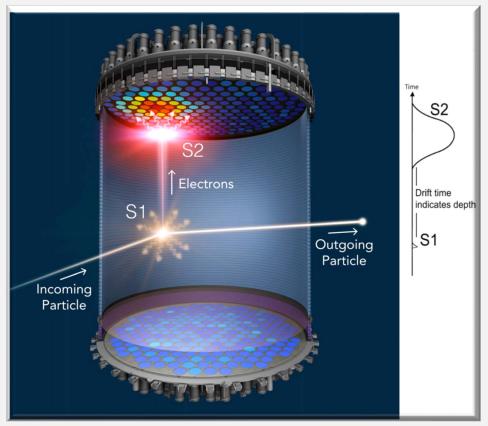
Super-CDMS-SNOLAB

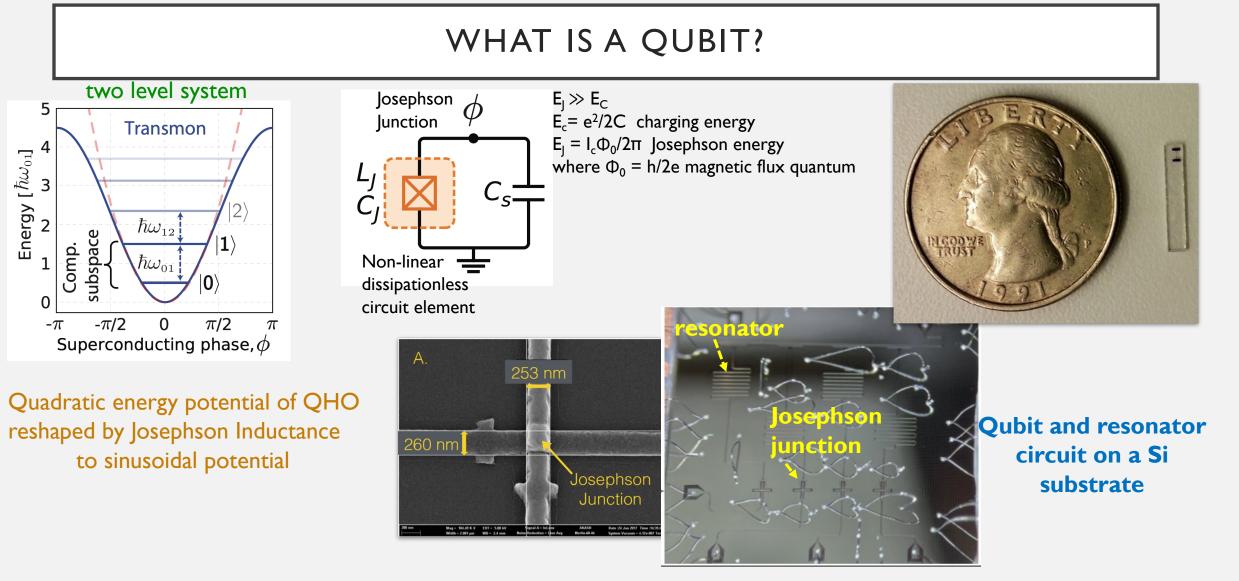
LZ (LUX-ZEPLIN)

-- DM-nuclei scattering (signal nuclear recoil) produces phonons (Ge/Si crystal lattice vibrations) and electrons through ionization (charge)



-- DM-Xe nuclei interaction produces electrons through ionization and photons that drift to the top causing flash of light (PMT)



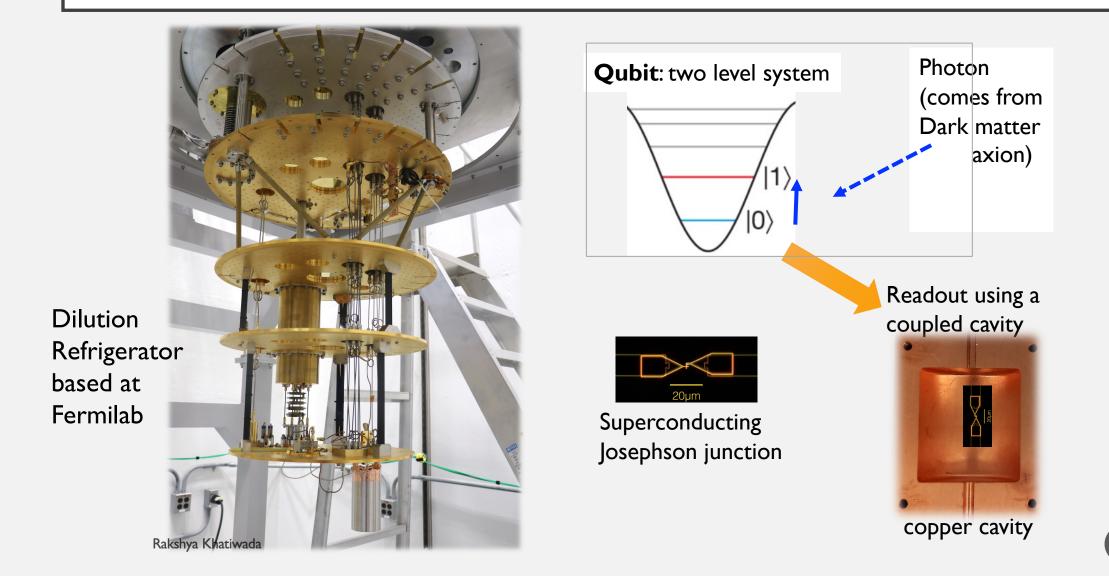


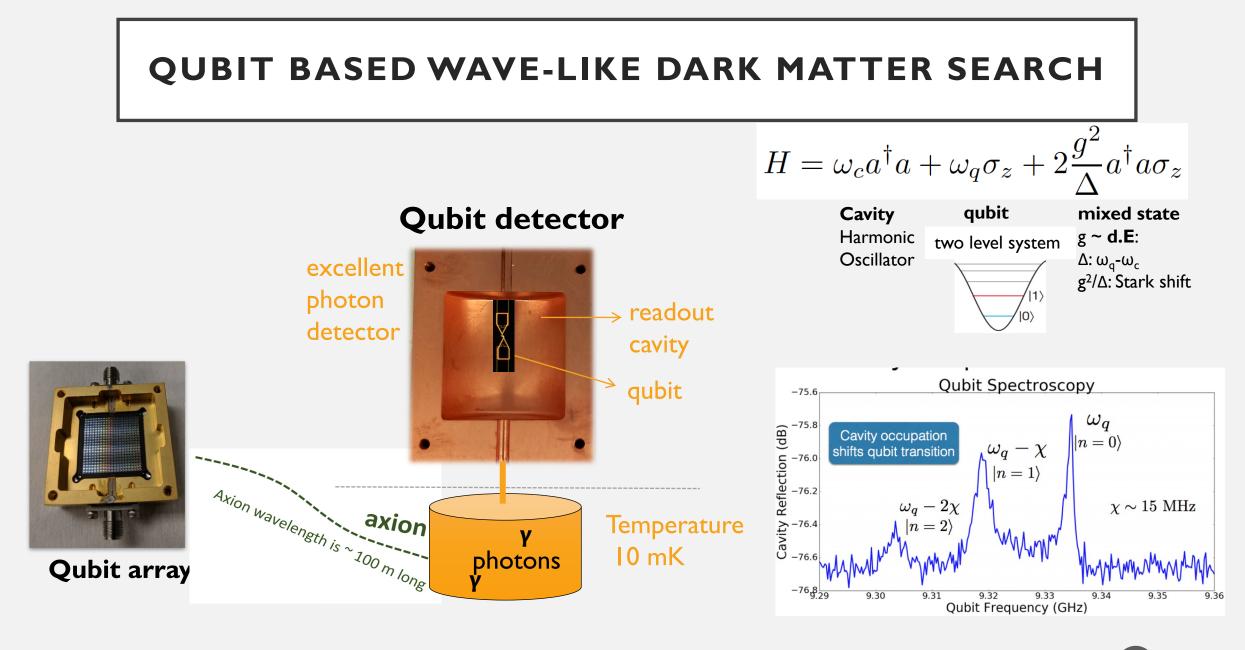
--Superconducting Transmon and its variants can be utilized for Dark Matter detection through several mechanisms of coupling

App. Phys. Rev., 6, 2, 10.1063 (2019).

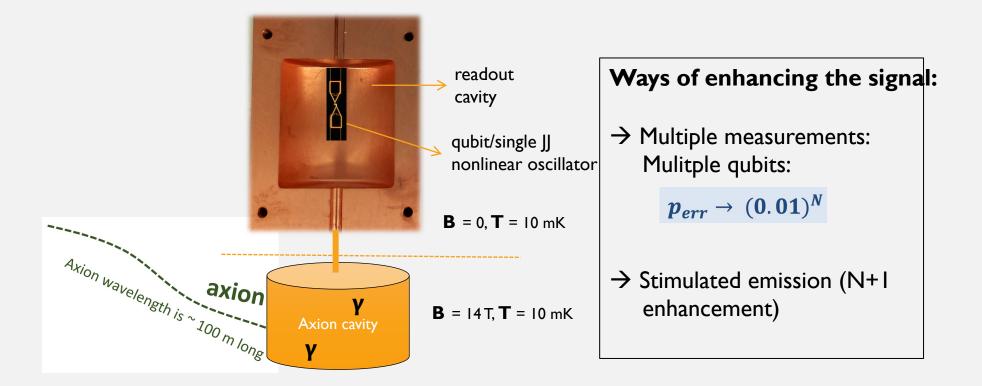
Rakshya Khatiwada

DETECTING DARK MATTER WITH QUBITS



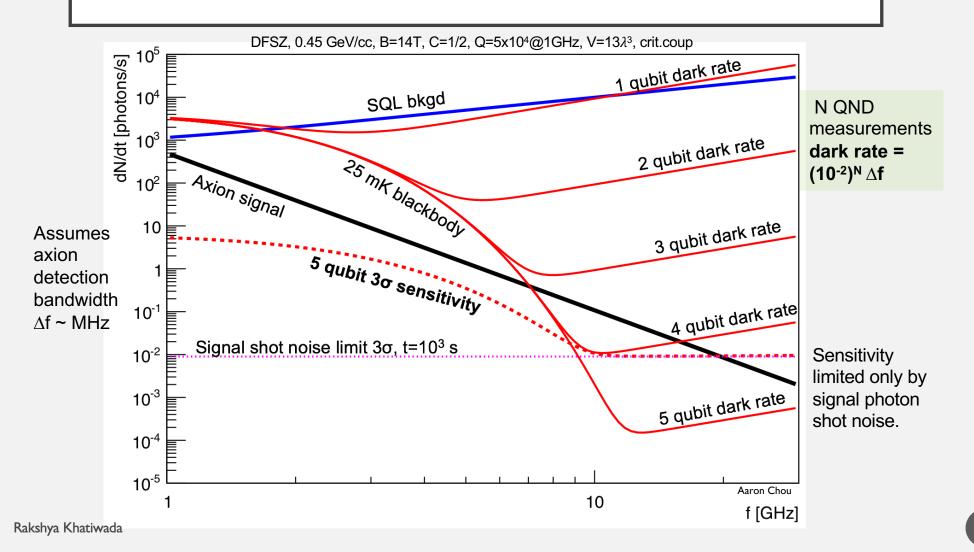


QUBIT BASED AXION DETECTOR



Photon # counting evades the quantum noise limit

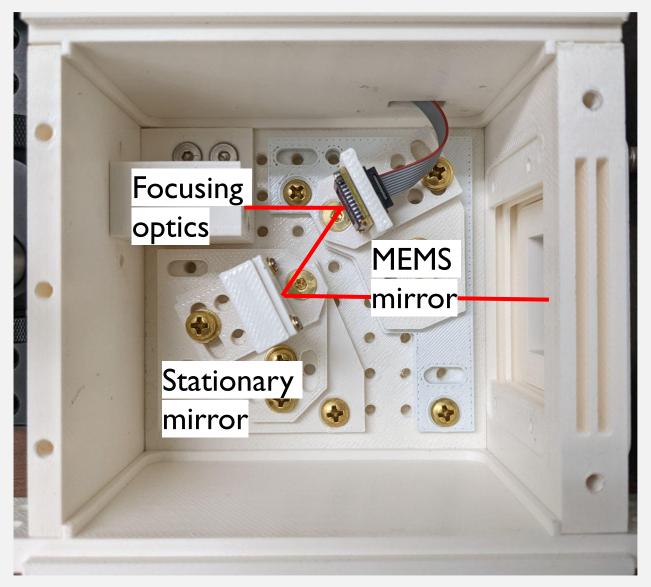
SIGNAL AND NOISE RATE



NQI Program Component Areas

- **Quantum Sensing and Metrology (QSENS)** refers to the use of quantum mechanics to enhance sensors and measurement science. This can include uses of superposition and entanglement, non-classical states of light, new metrology regimes or modalities, and advances in accuracy and precision enabled by quantum control, for example with atomic clocks.
- **Quantum Computing (QCOMP)** activities include the development of quantum bits (qubits) and entangling gates, quantum algorithms and software, digital and analog quantum simulators using programmable quantum devices, quantum computers and prototypes, and hybrid digital plus analog, as well as quantum plus classical computing systems.
- **Quantum Networking (QNET)** includes efforts to create and use entangled quantum states, distributed over distances and shared by multiple parties, for new information technology applications and fundamental science; for example, networking of intermediate scale quantum computers (modules) for enhanced beyond-classical computing capabilities.
- **QIS for Advancing Fundamental Science (QADV)** includes foundational efforts to invoke quantum devices and QIS theory to expand fundamental knowledge in other disciplines; for example, to improve understanding of biology, chemistry, computation, cosmology, energy science, engineering, materials, nuclear matter, and other aspects of fundamental science.
- Quantum Technology (QT) catalogues several topics: work with end-users to deploy quantum technologies in the field and develop use cases; basic R&D on supporting technology for quantum information science and engineering, e.g., infrastructure and manufacturing techniques for electronics, photonics, and cryogenics; and efforts to understand and mitigate risks raised by quantum technologies, e.g., post-quantum cryptography (see Box 4.1).

MEMS mirror allows for desired operating specifications:



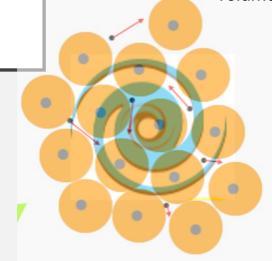
- ~1.5" x 1.5" scanning area
- <100µm spot size</p>
- ~10µm position resolution
- O(100)Hz scanning speed
- O(µs) pulse width
- >10mK operating temperature

NATURE OF DARK MATTER

For **mass < 70 eV**, Pauli exclusion principle causes dark matter clumps to swell up to be larger than the size of the smallest dwarf galaxies. (Randall, Scholtz, Unwin 2017)



Fermions: 1 DM particle per mode volume $(\lambda_{deBroglie})^3$





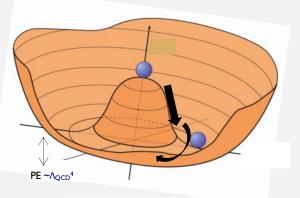
Particle like DM

Rakshya Khatiwada

AXION PRODUCTION

- Global symmetry broken at scale f_a
 - -- axion produced through misalignment mechanism
 - -- during QCD phase transition, trough tilted by $\Lambda_{\text{QCD}}{}^4$
- PE $\sim \Lambda_{QCD}^4$ released, makes up dark matter
- -- oscillation of the QCD θ angle about its minimum--vacuum energy to axions
- QCD axion mass $m_a \sim \Lambda_{QCD}^2/f_a$ ~ (200 MeV)²/f_a

--- f_a unknown \Rightarrow GHz frequencies at $f_a \sim 10^{13}$ GeV scale



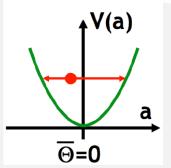
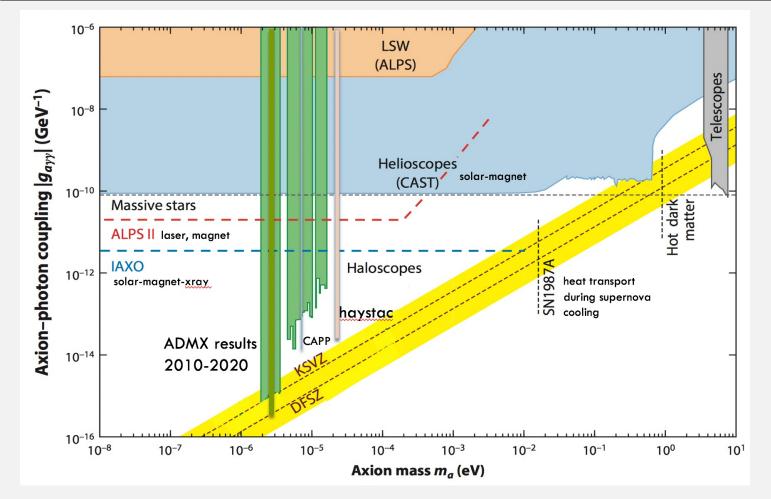
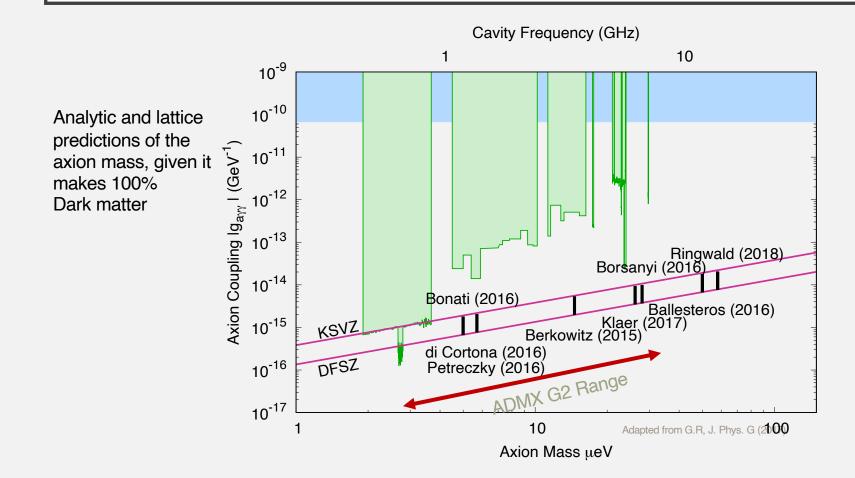


Fig 1:J. Ellis et al; arxiv:1201.6045v1

AXION SEARCHES OVERVIEW



AXION SEARCHES OVERVIEW CONTD.

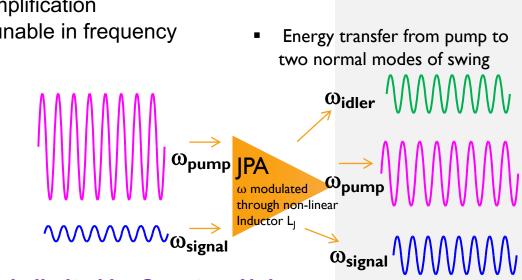


QUANTUM AMPLIFIERS

Why quantum amps.?

Intrinsically low noise (superconducting technology)

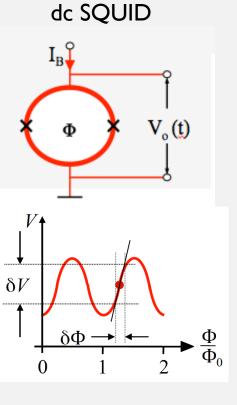
- \Rightarrow low resistance elements
- \Rightarrow low thermal dissipation
- \Rightarrow Add very low added noise during amplification
- => Tunable in frequency



Only limited by Quantum Noise Rakshya Khatiwada



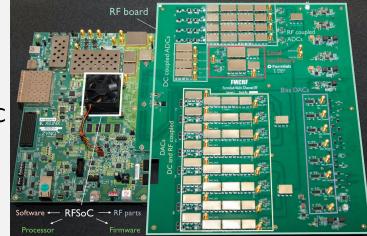
Josephson Parametric Amplifier JPA



QUANTUM INSTRUMENTATION CONTROL KIT (QICK)

- Primary support through QSC
- Broadly used: >30 labs in the US and abroad
- > 20 science talks in APS March meeting!
- 16 DAC outputs from DC to 10GHz all digital mixers.
 - This is critical need for quantum. => Better experiments.
 - Minimizes calibrations.
- 16 ADC inputs from DC to near 10GHz without analog mixers.
 - Input noise 120K (i.e. 1.2dB)
- 8 x 20-bit DAC outputs for bias (BW of 10 or 100Hz) ultra low noise, ±10V outputs.
- <200ns latency for error correction (fastest in the market).
- Optimal filtering, no dead-time 3ns pulses.
- 100 pico-second pulses for QN.
- Accurate timing readout 2ps resolution.
- RF multiplexing of 10K detectors per board.

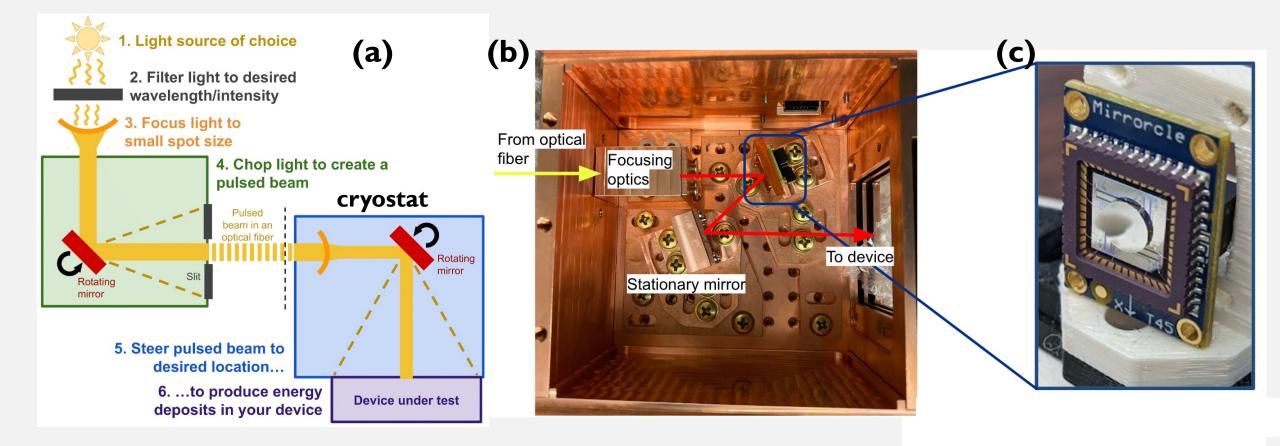
AMD-RFSoC ZCUIII +QICK custom RF/DC amplifiers, bias





AMD-RFSoC ZCU216 QICK custom RF/DC amplifiers, bias under design

Cryogenic photon source for Detector Characterization



WHAT'S CAUSING THESE DARK COUNTS?

A superconductor free of quasiparticles for seconds

E. T. Mannila,^{1,*} P. Samuelsson,² S. Simbierowicz,^{3,†} J. T. Peltonen,¹ V. Vesterinen,³ L. Grönberg,³ J. Hassel,³ V. F. Maisi,² and J. P. Pekola¹

2022

Aug

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¹QTF Centre of Excellence, Department of Applied Physics, Aalto University, FI-00076 Aalto, Finland ²Physics Department and NanoLund, Lund University, Box 118, 22100 Lund, Sweden ³VTT Technical Research Centre of Finland Ltd, QTF Centre of Excellence, P.O. Box 1000, FI-02044 VTT, Finland (Dated: February 2, 2021)

Eliminated cosmic muon and radioactive background from suspects since qp poisoning suppressed over longer ~ a week cooldown period

Used similar device to charge parity device like QCD

Microfractures due to GE Varnish and mounting glue on Si substrate causing phonon bursts breaking cooper pair -> qp poisioning

A Stress Induced Source of Phonon Bursts and Quasiparticle Poisoning

R. Anthony-Petersen,¹ A. Biekert,^{1,2} R. Bunker,³ C.L. Chang,^{4,5,6} Y.-Y. Chang,¹ L. Chaplinsky,⁷ E. Fascione,^{8,9} C.W. Fink,¹ M. Garcia-Sciveres,² R. Germond,^{8,9} W. Guo,^{10,11} S.A. Hertel,⁷ Z. Hong,¹² N.A. Kurinsky,¹³ X. Li,² J. Lin,^{1,2} M. Lisovenko,⁴ R. Mahapatra,¹⁴ A.J. Mayer,⁹ D.N. McKinsey,^{1,2} S. Mehrotra,¹ N. Mirabolfathi,¹⁴ B. Neblosky,¹⁵ W.A. Page,^{1,*} P.K. Patel,⁷ B. Penning,¹⁶ H.D. Pinckney,⁷ M. Platt,¹⁴ M. Pyle,¹ M. Reed,¹ R.K. Romani,^{1, *} H. Santana Queiroz,¹ B. Sadoulet,¹ B. Serfass,¹ R. Smith,^{1,2} P. Sorensen,² B. Suerfu,^{1,2} A. Suzuki,² R. Underwood,⁸ V. Velan^{1,2} G. Wang⁴ Y. Wang^{1,2} S.L. Watkins¹ M.R. Williams¹⁶ V. Yefremenko⁴ and J. Zhang⁴ ¹Department of Physics, University of California, Berkeley, CA 94720, USA ²Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA ³Pacific Northwest National Laboratory, Richland, WA 99352, USA ⁴High Energy Physics Division, Argonne National Laboratory, 9700 S. Cass Avenue, Argonne, Illinois 60439, USA ⁵Department of Astronomy and Astrophysics, University of Chicago, 5640 South Ellis Avenue, Chicago, Illinois 60637, USA ⁶Kavli Institute for Cosmological Physics, University of Chicago, 5640 South Ellis Avenue, Chicago, Illinois 60637, USA ⁷Department of Physics, University of Massachusetts, Amherst, Massachusetts 01003. USA

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