

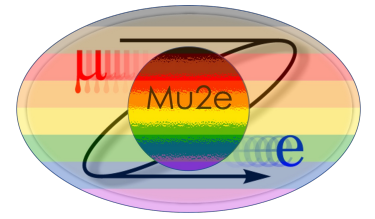
Mu2e: A Search for Charged Lepton Flavor Violation at FNAL

Design and Current Status

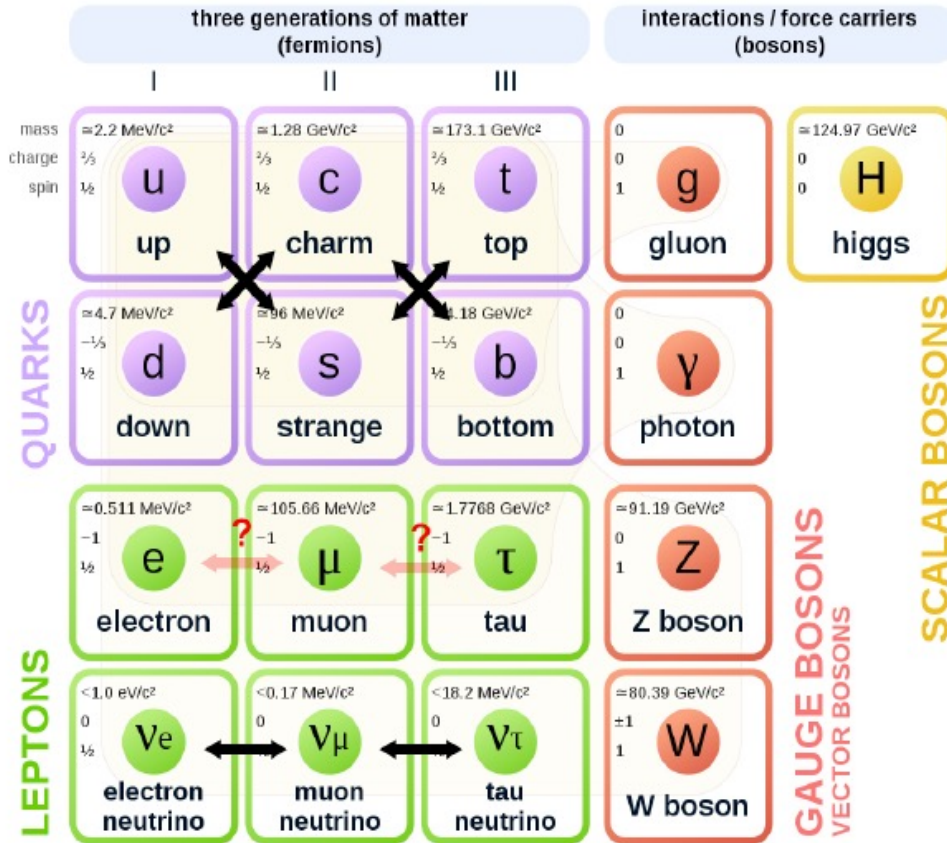
Sophie Charlotte Middleton

on behalf of the Mu2e experiment

Flavor Physics

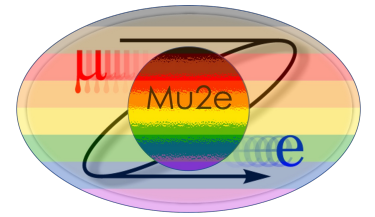


Standard Model of Elementary Particles

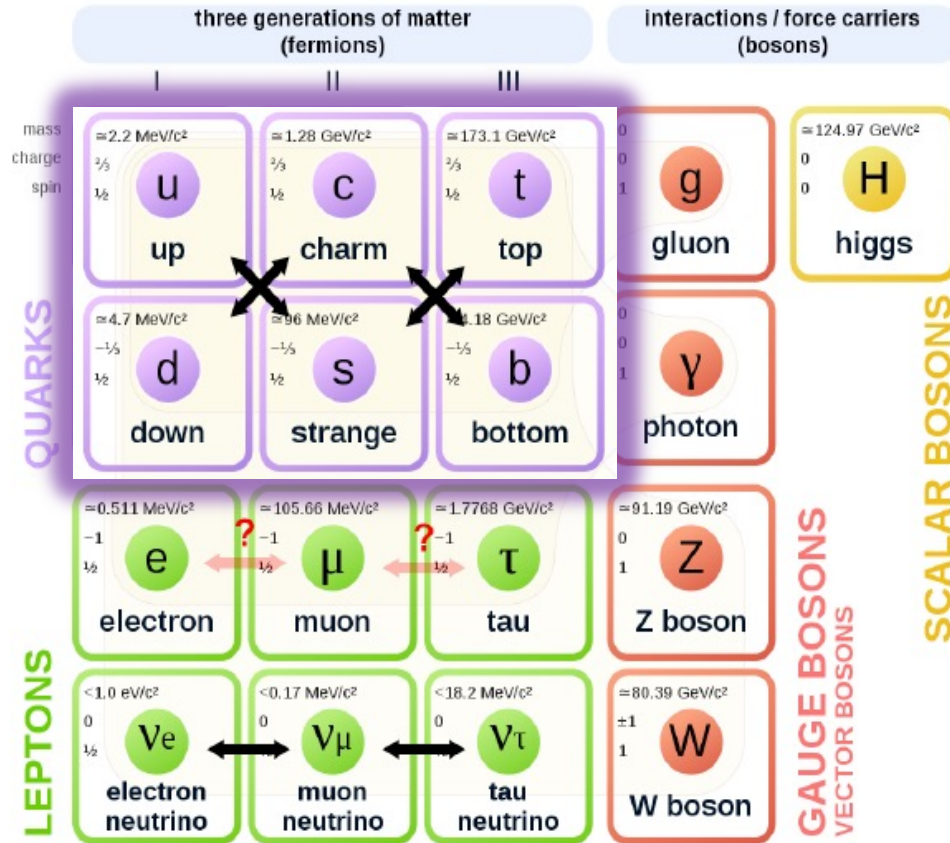


- Flavor is not conserved in:
 - quarks (via quark mixing);
 - neutrinos (via neutrino oscillations)

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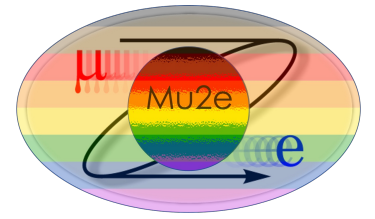


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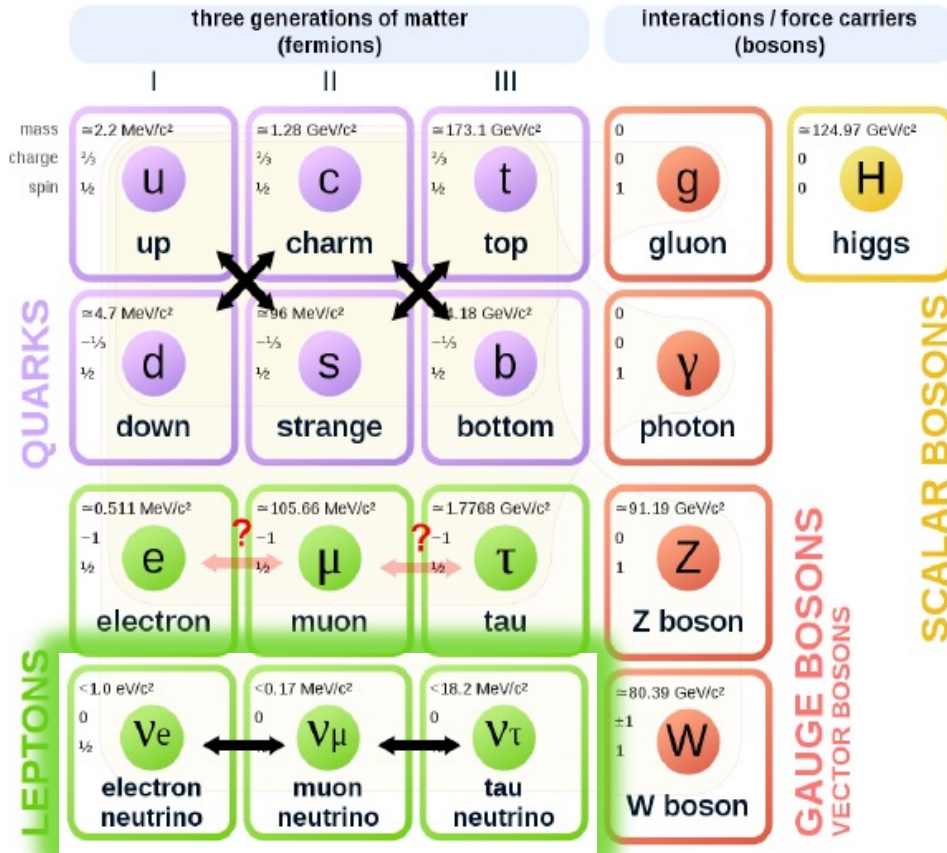


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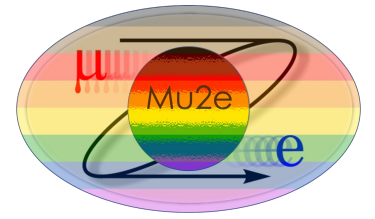


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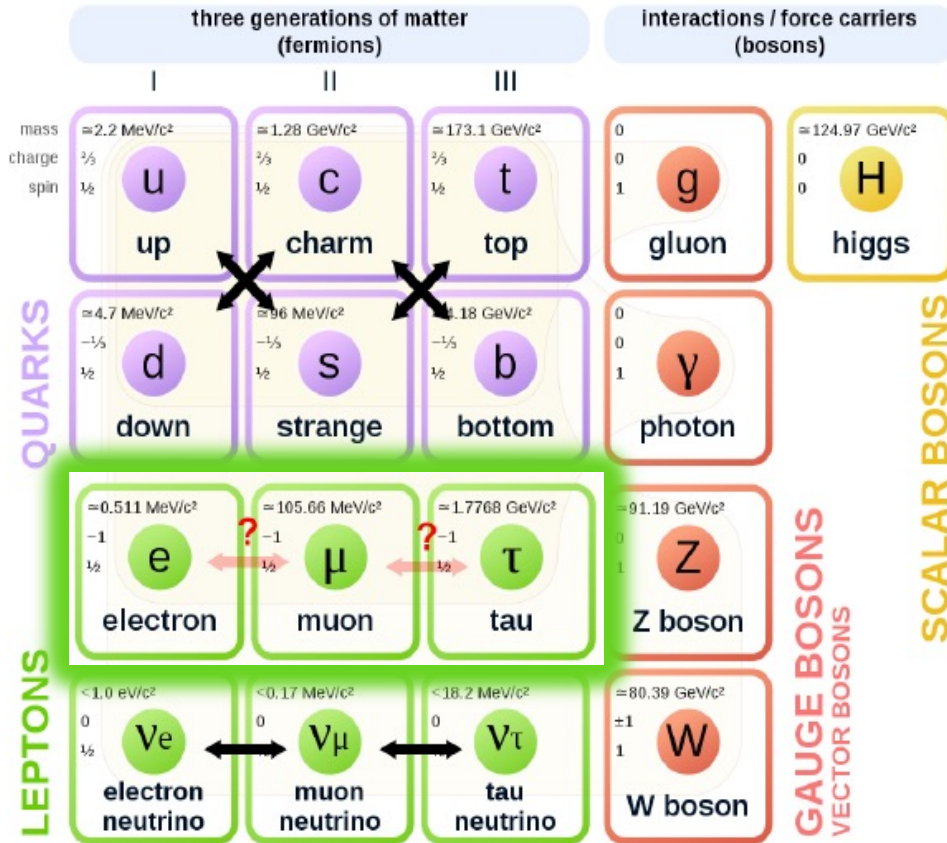


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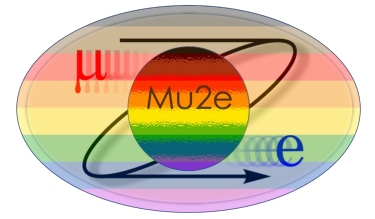


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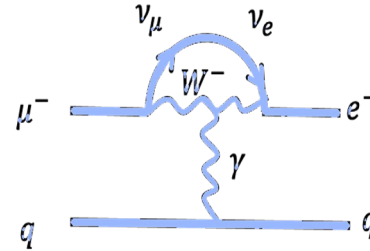
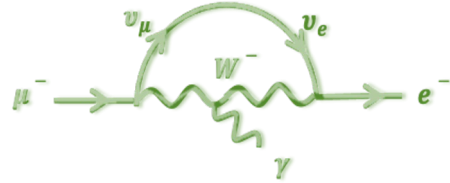


- Flavor is not conserved in:
 - quarks (via quark mixing);
 - neutrinos (via neutrino oscillations)
- What about Charged Lepton Flavor Violation?**

Charged Lepton Flavor Violation (CLFV)



- The minimal extension of the Standard Model, including masses of neutrinos, allows for CLFV at loop level, mediated by W bosons.



No outgoing neutrinos!

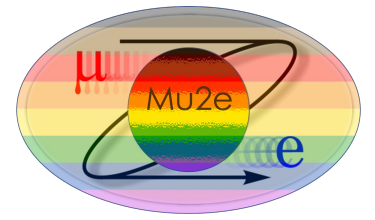
- Rates heavily suppressed by GIM suppression and are far below any conceivable experiment could measure:

$$B(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left(\frac{1}{4}\right) \sin^2 2\theta_{13} \sin^2 \theta_{23} \left| \frac{\Delta m_{13}^2}{M_W^2} \right|^2$$

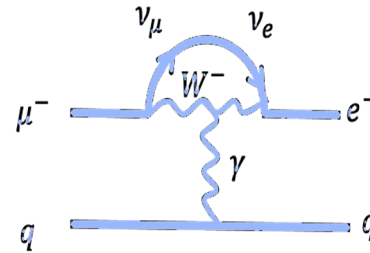
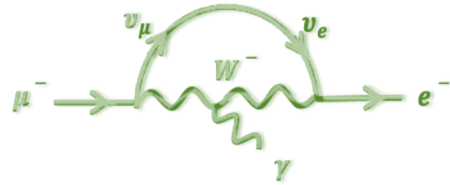
$$B(\mu \rightarrow e\gamma) \sim \mathcal{O}(10^{-54})$$

S.T. Petcov, Sov.J. Nucl. Phys. 25 (1977) 340

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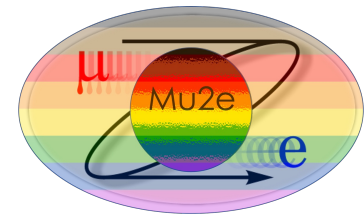
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- ...but many Beyond Standard Model (BSM) theories (e.g. SO(10) SUSY, scalar leptoquarks, seesaw models) predict enhanced rates of CLFV just below current limits $\mathcal{O}(10^{-13})$.

Mu2e is an indirect search for New Physics which offers a deep probe of many well-motivated BSM theories.

Current Experimental Searches for CLFV

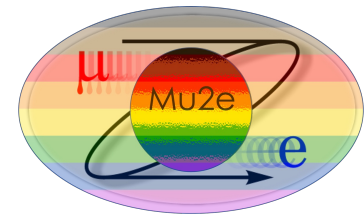


- **Mu2e is an important part of a world-wide search for CLFV.**
- Muons are a very powerful probe thanks to the availability of very intense beams and their relatively long lifetime.

Mode	Current Upper Limit (at 90% CL)	Projected Limit (at 90% CL)	Upcoming Experiment/s
$\mu^+ \rightarrow e^+ \gamma$	4.2×10^{-13}	4×10^{-14}	MEG II
$\mu^+ \rightarrow e^+ e^+ e^-$	$\sim 10^{-12}$	$10^{-15} \sim 10^{-16}$	Mu3e
$\mu^- N \rightarrow e^- N$	7×10^{-13} (SINDRUM-II, 2006)	10^{-15} 10^{-16} 10^{-17}	COMET Phase-I Mu2e Run-I Mu2e Run-II/ COMET Phase-II

- Synergies with tau CLFV searches. Comparing different lepton CLFV elucidates source of flavor violation.
- Need to explore the entire CLFV-sector like with quark and neutrino searches.

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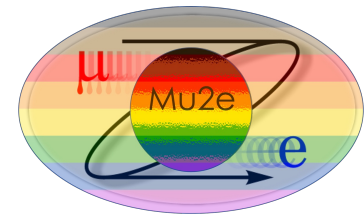


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Effective Physics Reach

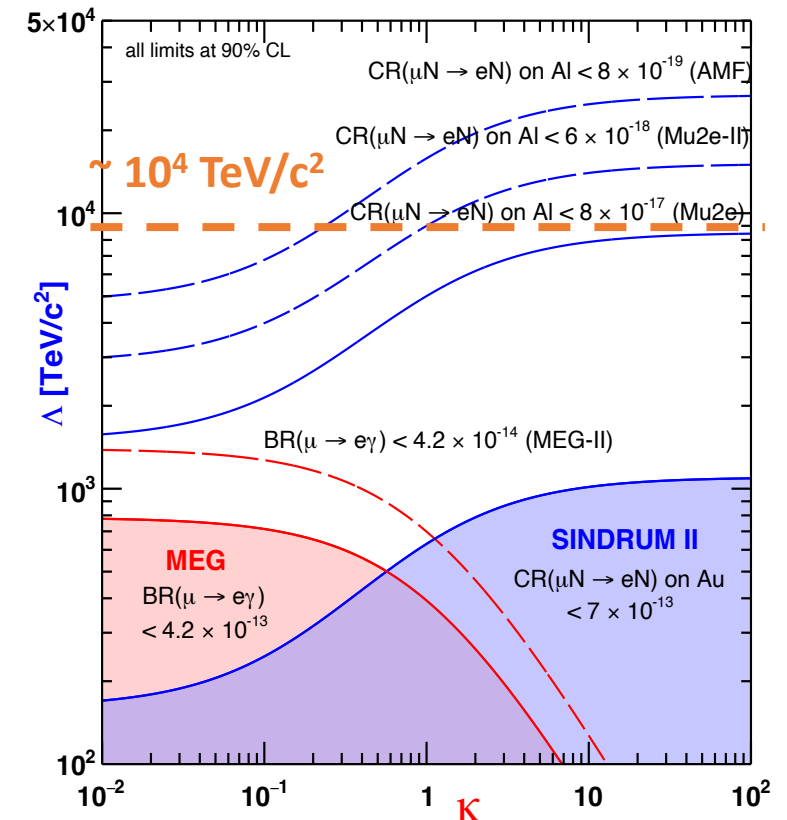


$$\mathcal{L}_{CLFV} = \frac{m_\mu}{(1+\kappa)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1+\kappa)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L \left(\sum_{q=u,d} \bar{q}_L \gamma_\mu q_L \right)$$

“Photonic”

“Contact”

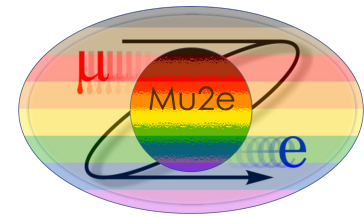
* Mu2e-II and AMF are proposed future experiments/facilities.



Λ : Effective mass scale of New Physics (NP),

κ : Determines relative sizes of contributions and to what extent NP is photonic ($\kappa \ll 1$) or 4-fermion ($\kappa \gg 1$)

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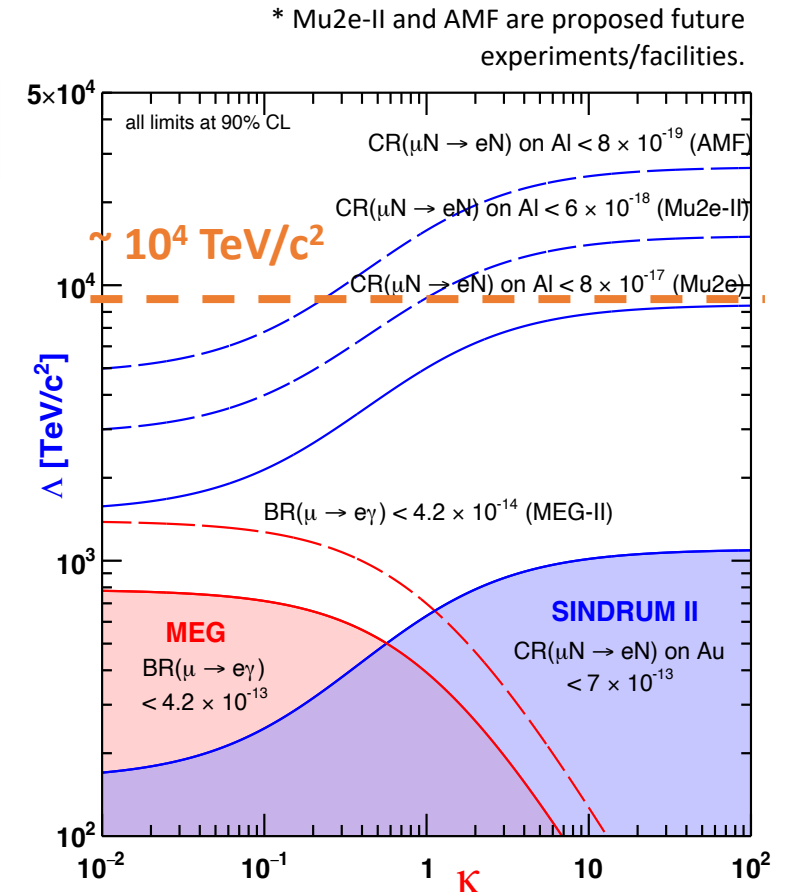
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$$\mu^+ \rightarrow e^+ \gamma, \mu^+ \rightarrow e^+ e^+ e^-.$$

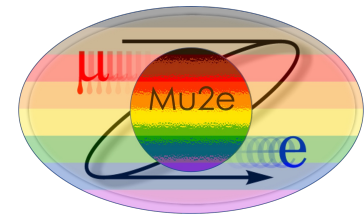
$\mu^- N \rightarrow e^- N$ at loop level

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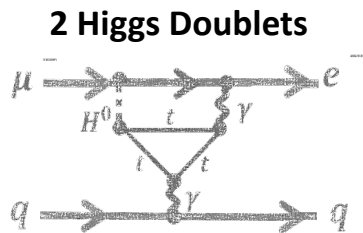
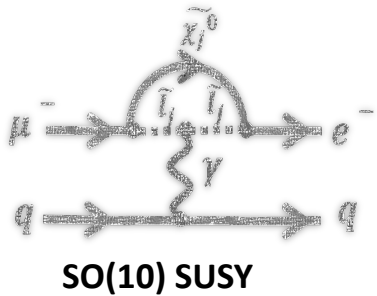
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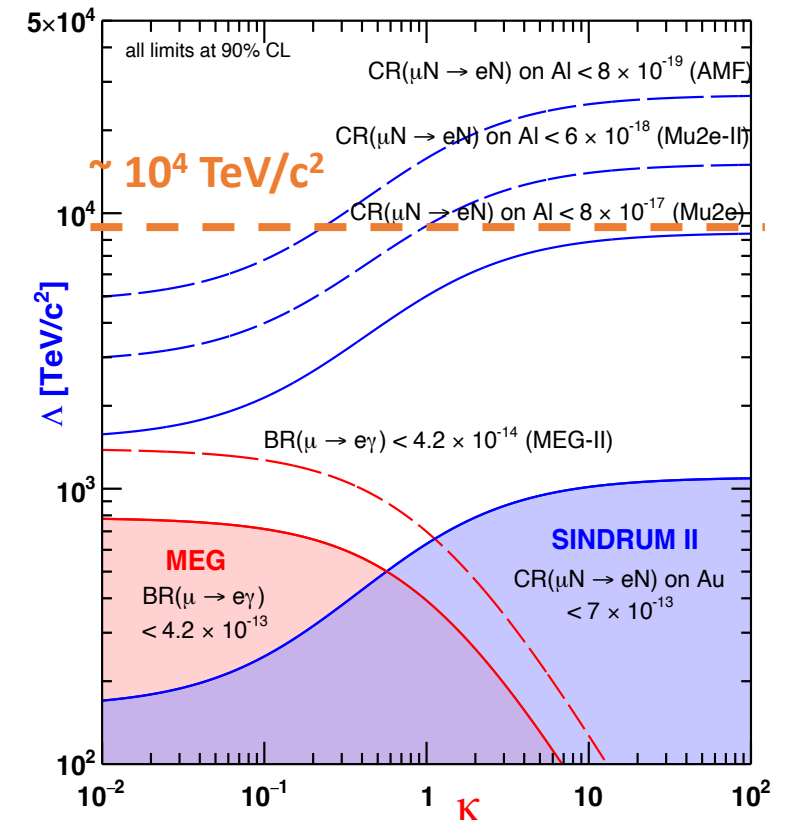
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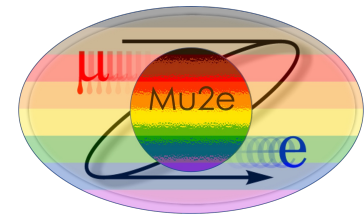
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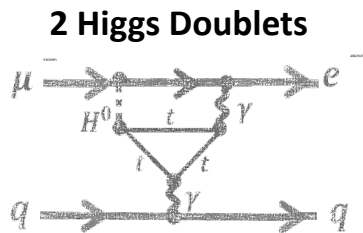
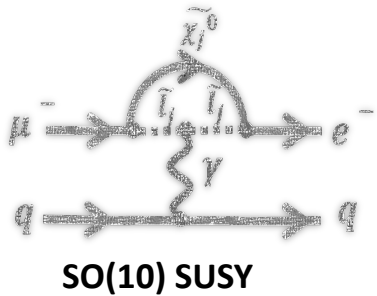
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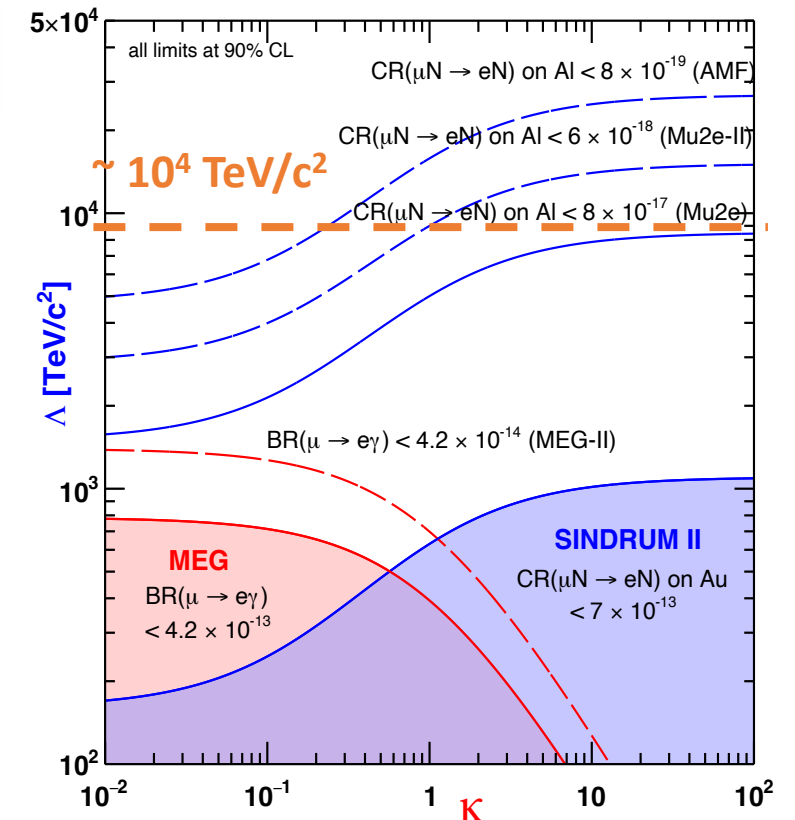
i.e. 4 Fermion Term

$\mu^- N \rightarrow e^- N$ at leading order.
Heavily suppressed in $\mu^+ \rightarrow e^+ \gamma$

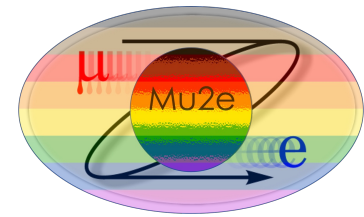
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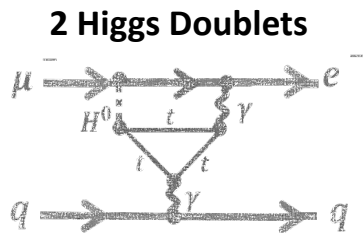
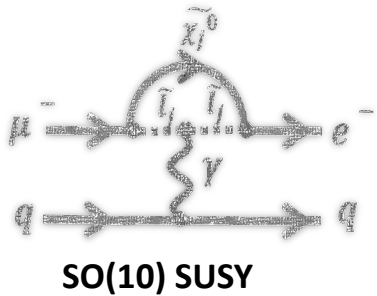
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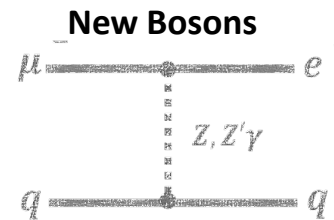
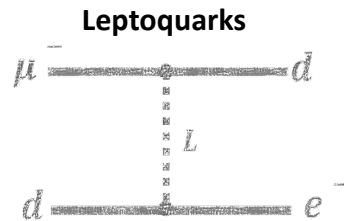


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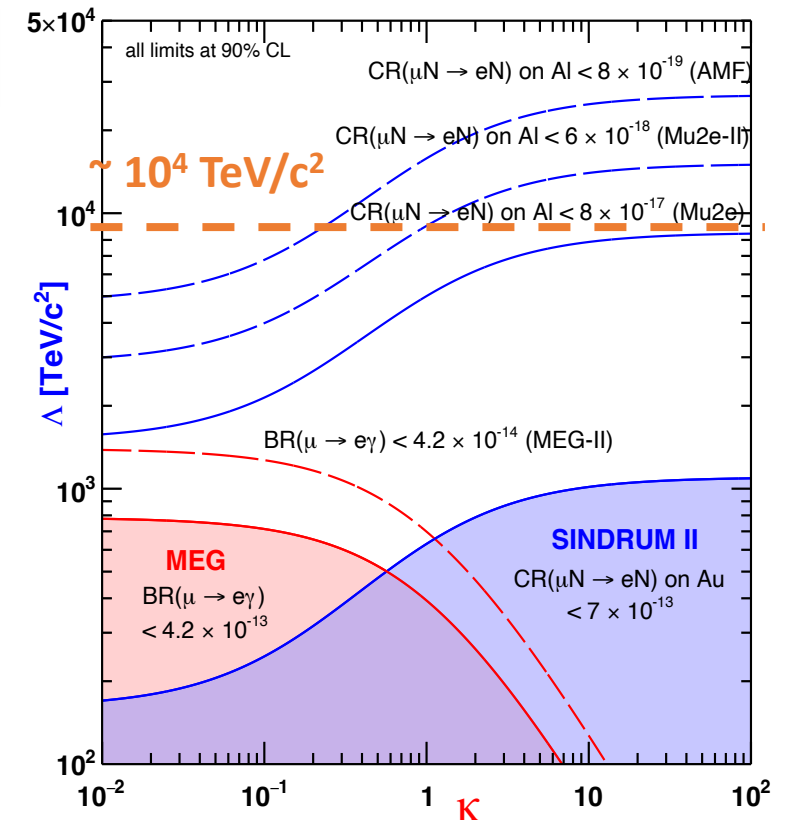
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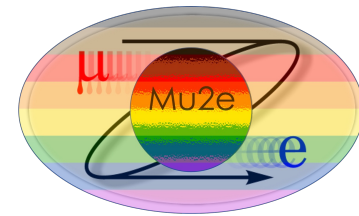
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$N\mu^- \rightarrow Ne^-$: The Mu2e Experiment



- The Mu2e experiment will search for this process in AI and improve on this limit by four orders of magnitude!
- Need to stop $O(10^{18})$ and have $\ll 1$ background event

2026 – 27 Run-I:

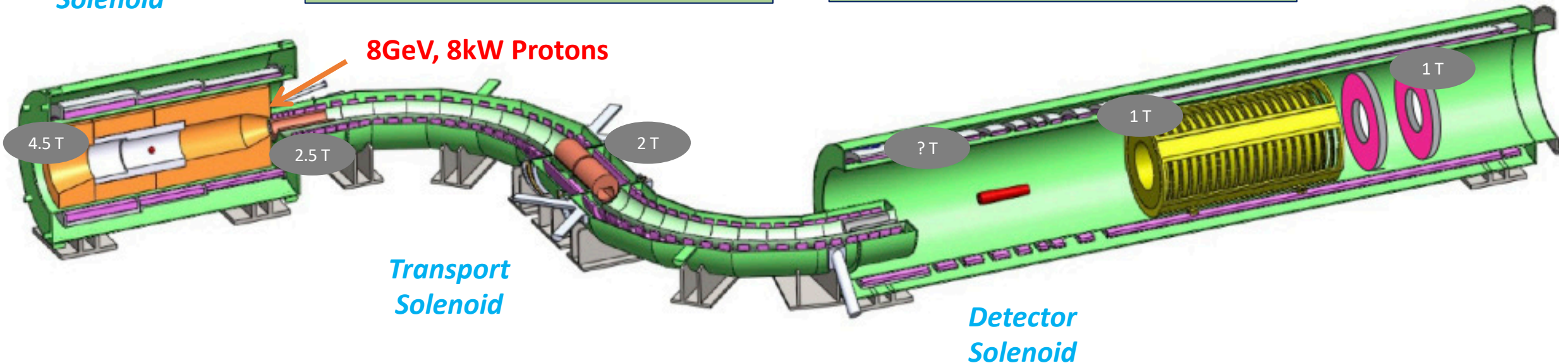
- 1×10^{-15} 5σ discovery,
- Single-Event-Sensitivity = 2×10^{-16}
- U.L : 6×10^{-16} (90% C.L.)
 - 1000 x current limit.
 - Universe 2023, 9, 54.

Total (Run-I + Run-II) end-goal (2033):

- 2×10^{-16} 5σ discovery,
- Single-Event-Sensitivity = 3×10^{-17}
- U.L : 8×10^{-17} (90% C.L.)
 - 10000 x current limit.

Production
Solenoid

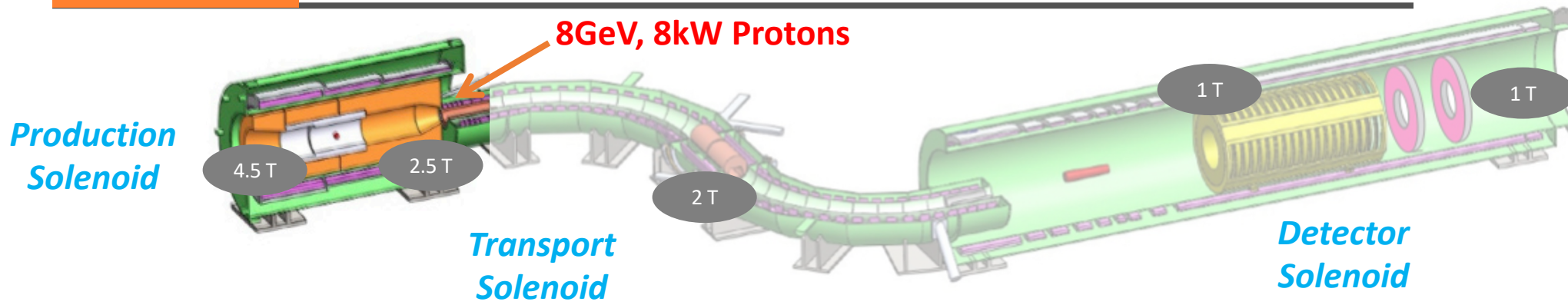
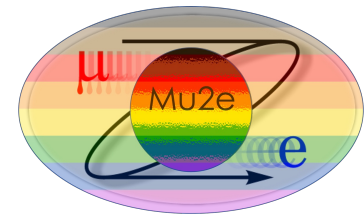
8GeV, 8kW Protons



Transport
Solenoid

Detector
Solenoid

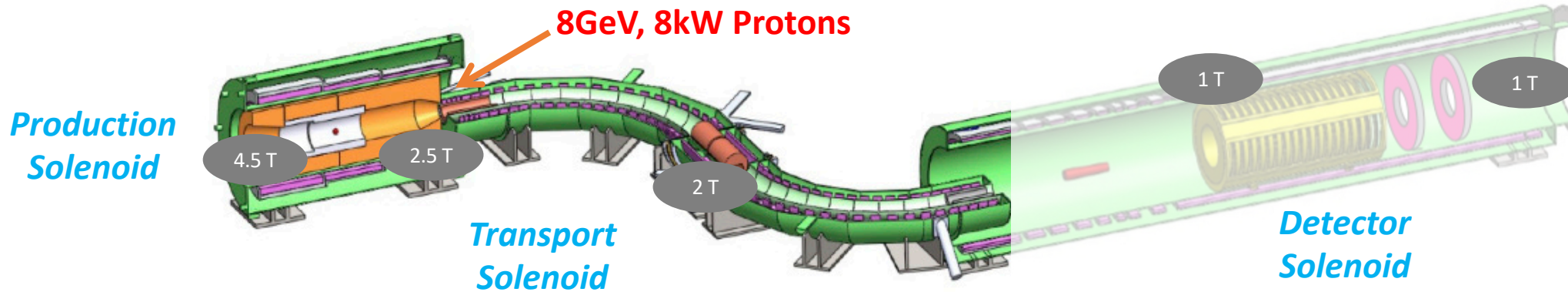
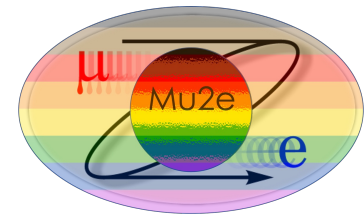
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Production Solenoid:

- Pulsed 8 GeV Protons enter, hit Production Target. π produced, decay to μ .
- Graded magnetic field reflects muons to transport solenoid.

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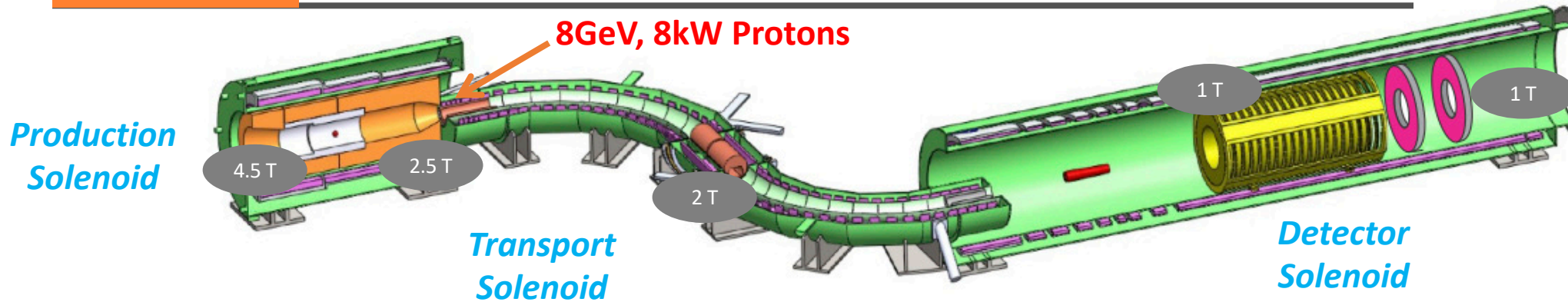
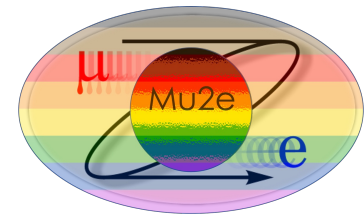
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- "S" shape removes line of sight backgrounds.
- Collimators select low momentum, negative muons.

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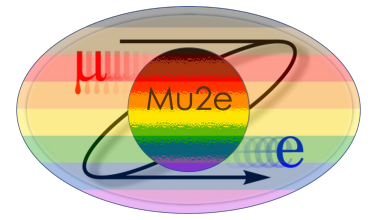
Transport Solenoid:

- “S” shape removes line of sight backgrounds.
- Collimators select low momentum, negative muons.

Detector Solenoid:

- Thin aluminum foil target captures the muons.
- Possible signal electrons are detected by a tracker and a calorimeter.
- Cosmic ray veto covers the whole detector solenoid and half the transport solenoid.

$N\mu^- \rightarrow Ne^-$: Signal



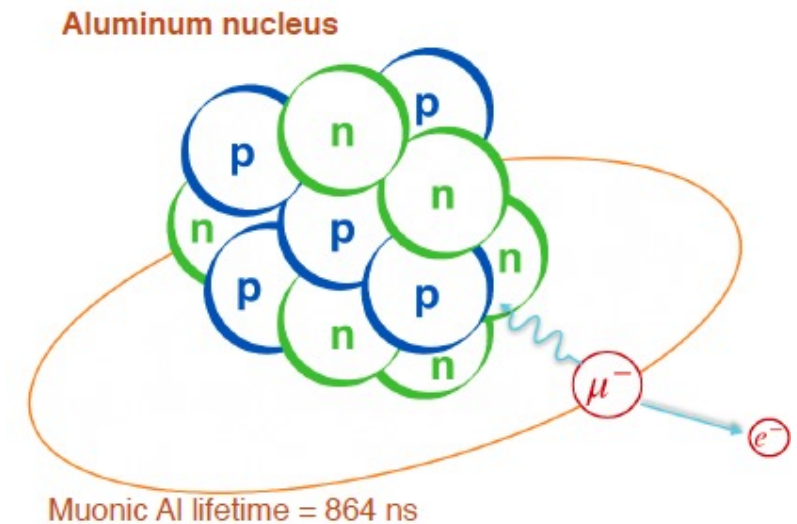
- The $\mu \rightarrow e$ conversion rate is measured as a ratio to the muon capture rate on the same nucleus:

$$R_{\mu e} = \frac{\Gamma(\mu^- + A(Z, N) \rightarrow e^- + A(Z, N))}{\Gamma(\text{all-captures})}$$

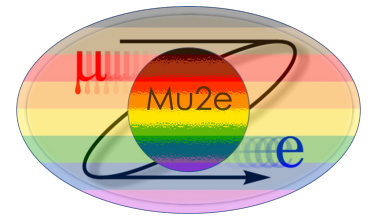
- Low momentum (-) muons are captured in the target atomic orbit and quickly (\sim fs) cascades to 1s state.
- In aluminum:
 - 39 % - Decay**: $\mu + N \rightarrow e + \bar{\nu}_e + \nu_\mu$ (**Background**)
 - 61 % - Capture**: $\mu + N \rightarrow \nu_\mu + N'$ (**Normalization**)
 - $< 7 \times 10^{-13}$ - Conversion**: $\mu + N \rightarrow e + N$ (**Signal**)
- Signal is monoenergetic electron consistent with:

$$E_e = m_\mu - E_{\text{recoil}} - E_{1S B.E.}, \text{ e.g For Al: } E_e = 104.97 \text{ MeV.}$$

- Coherent = nucleus stays intact.
- Will be smeared by scattering and energy losses



$N\mu^- \rightarrow Ne^-$: Removing Backgrounds



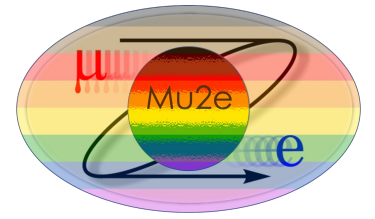
Beam delivery and detector systems optimized for high intensity, pure muon beam – must be “background free”:

Type	Source	Mitigation	Yield (for Run-I only)*
Intrinsic	Decay in Orbit (DIO)	Tracker Design/Resolution	0.038 ± 0.002 (stat) $^{+0.025}_{-0.015}$ (sys)
Beam Backgrounds	Pion Capture	Beam Structure/Extinction	(in time) 0.010 ± 0.002 (stat) $^{+0.001}_{-0.003}$ (sys) (out time) $(1.2 \pm 0.001$ (stat) $^{+0.1}_{-0.3}$ (sys)) $\times 10^{-3}$
Cosmic Induced	Cosmic Rays	Active Veto System	0.046 ± 0.010 (stat) ± 0.009 (sys)

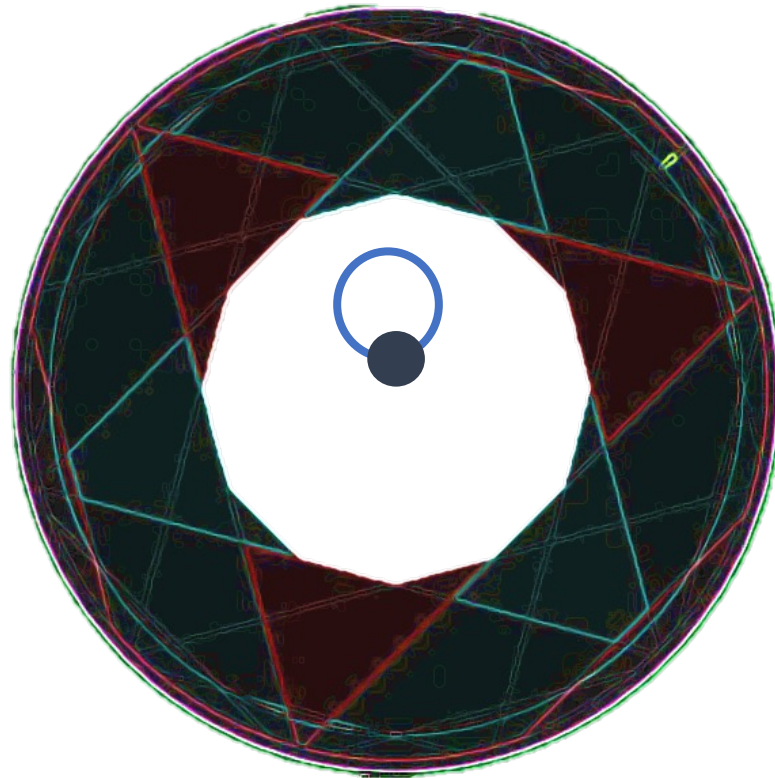
* assumes signal region of $103.6 < p < 104.9$ MeV/c and $640 < t < 1650$ ns

Run-I Sensitivity of Mu2e:
Universe 2023, 9, 54.

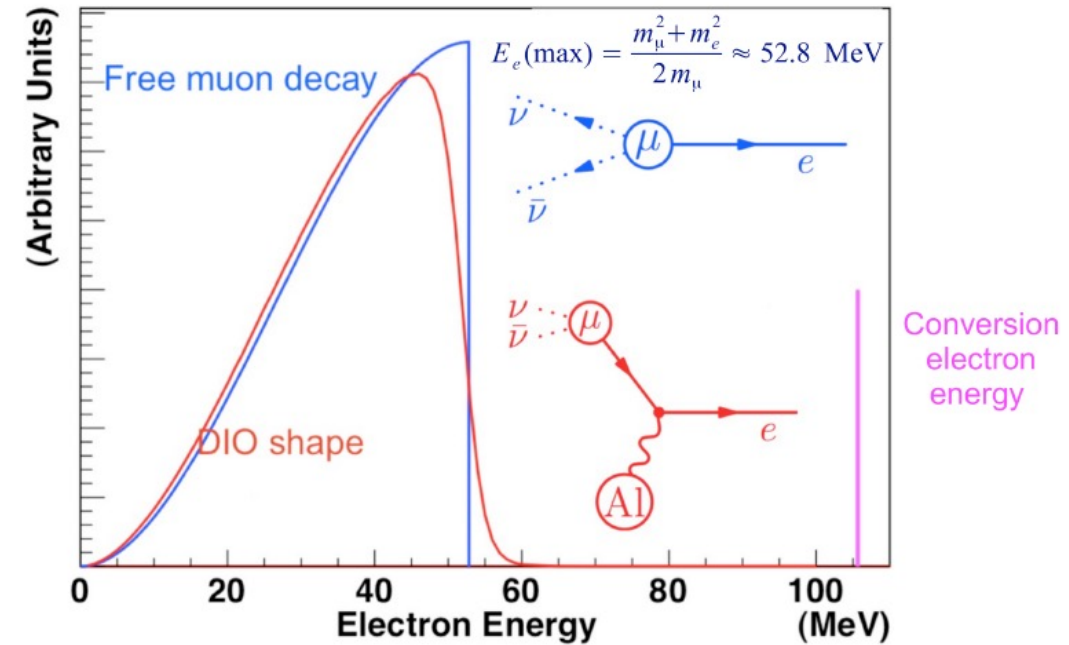
Decay in Orbit (DIO) Backgrounds



- Annular tracker: Removes > 97% of DIO (all Michel peak electrons).

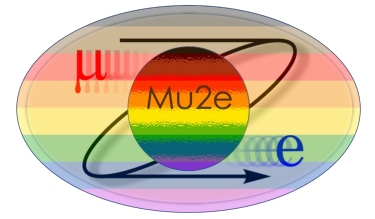


- Michel Electron (< 52MeV/c)



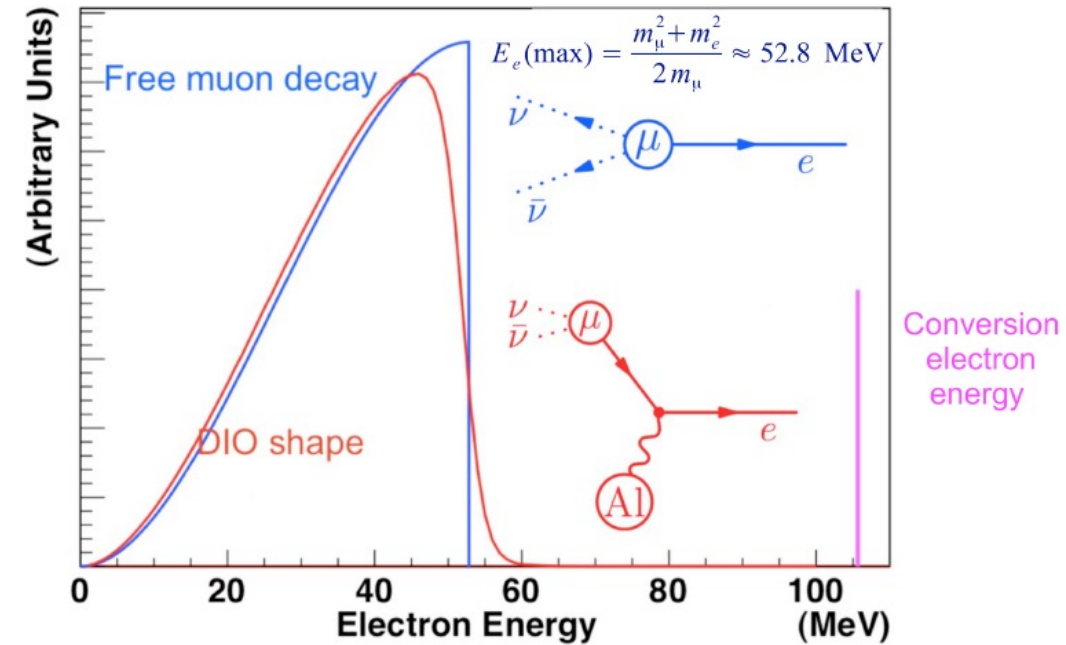
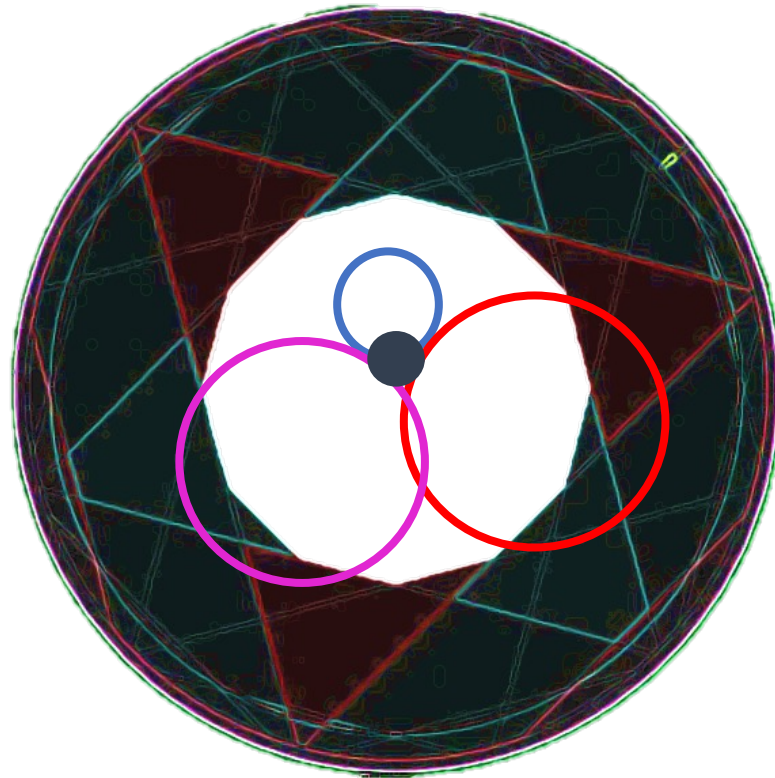
Cartoon of DIO Energy Spectrum
(see Szafron, Czarnecki PhysRevD.94.051301 + others)

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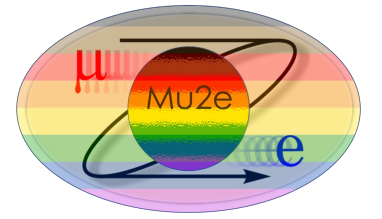
- Annular tracker: Removes > 97% of DIO (all Michel peak electrons).
- However, when decay happens in orbit, exchange of momentum produces recoil tail close to signal region (105 MeV/c).
- To remove remaining backgrounds necessitates < 200 keV/c momentum resolution.

- Michel Electron (< 52MeV/c)
- Problematic Tail (>100MeV/c)
- Signal (105MeV/c)



Cartoon of DIO Energy Spectrum
(see Szafron, Czarnecki PhysRevD.94.051301 + others)

Decay in Orbit Backgrounds



- Need a high-resolution momentum measurement to distinguish tail DIO from signal:
 - Minimize energy loss by operating in vacuum and using low mass straws of $15\ \mu\text{m}$ diameter filled with 80:20 Ar:CO₂ ;
 - Include extra hit position information with high-angle stereo overlaps and readout on both ends of straw.



96 Straws



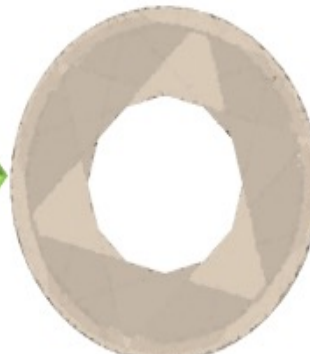
6 Panels



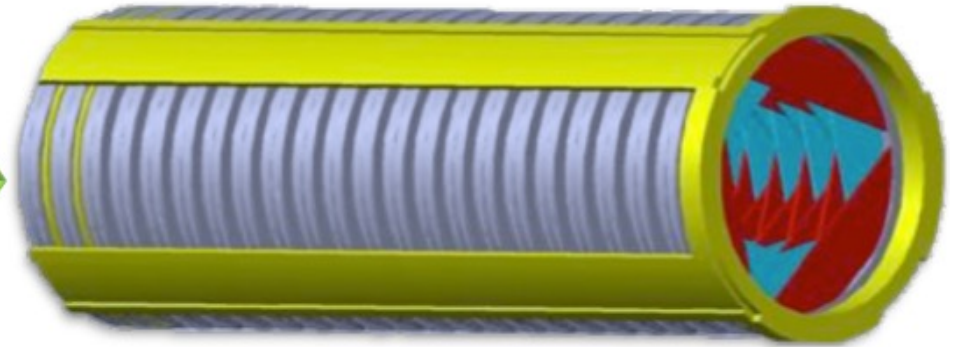
2 Plane



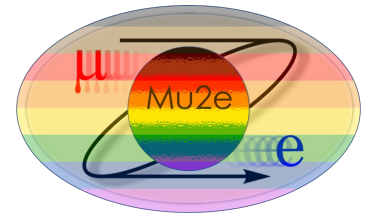
1 Station



18 Station = 1 Tracker

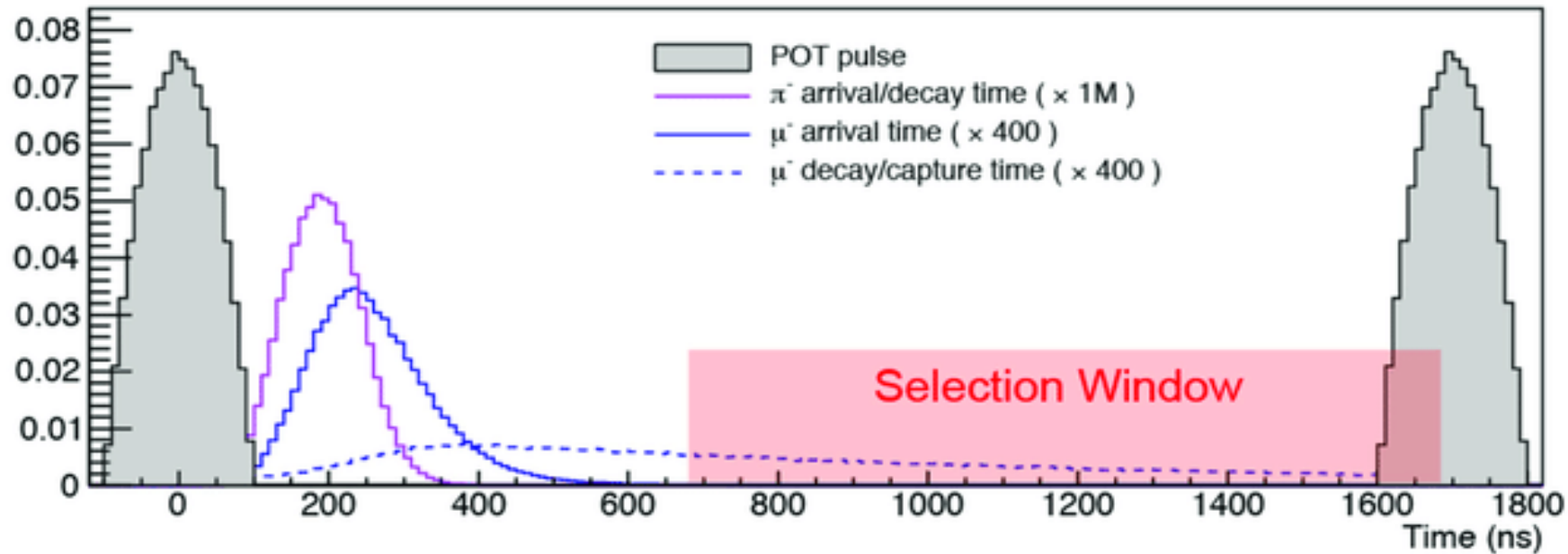


Radiative Pion Capture Backgrounds

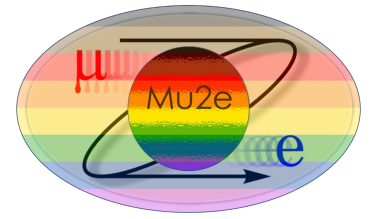


- Radiative pion capture backgrounds: $\pi^- + N (A, Z) \rightarrow \gamma^{(*)} + N (A, Z - 1)$ followed by $\gamma^{(*)} \rightarrow e^+ + e^-$.
- Pion lifetime 26 ns at rest. Pulsed proton beam (250 ns wide, pulses 1695 ns apart) \rightarrow wait out pion decay.
- In addition, upstream extinction removes out-of-time protons.

Delayed live-gate helps remove pion and beam backgrounds.

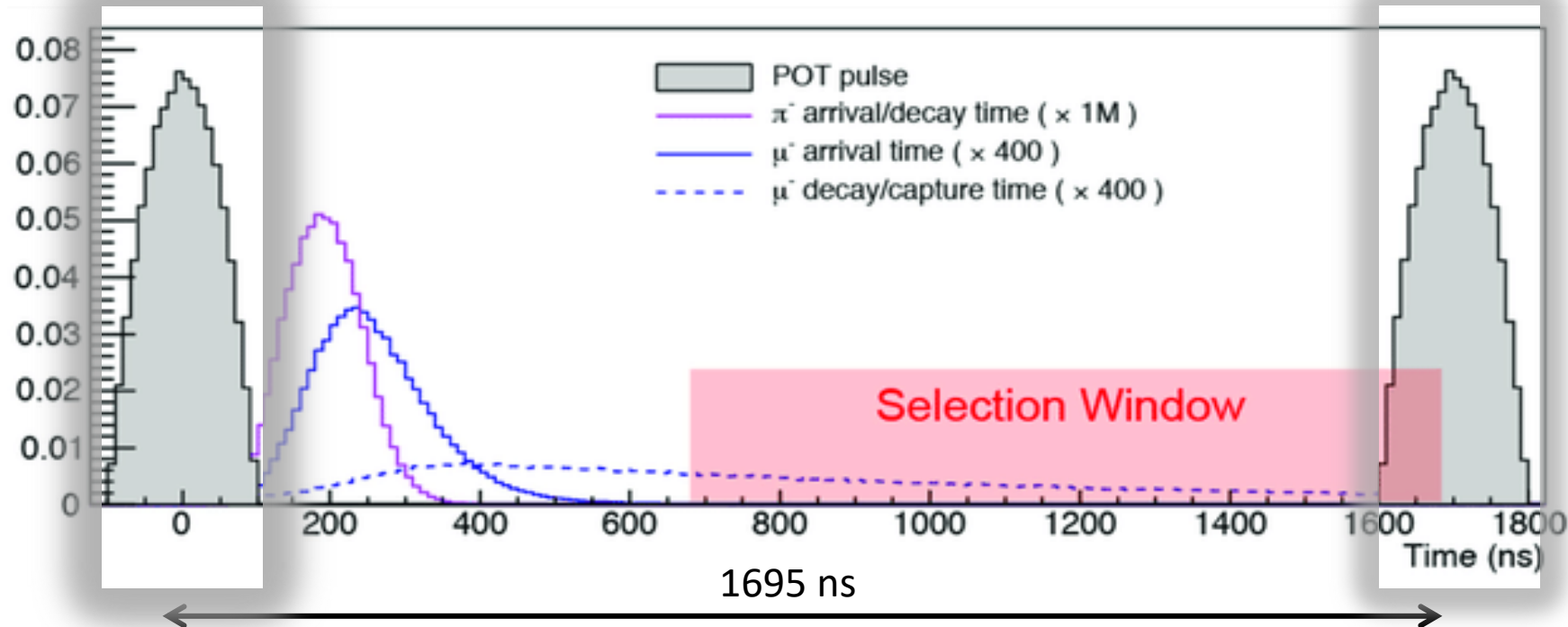


Radiative Pion Capture Backgrounds

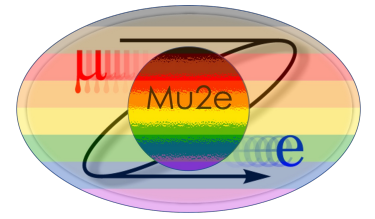


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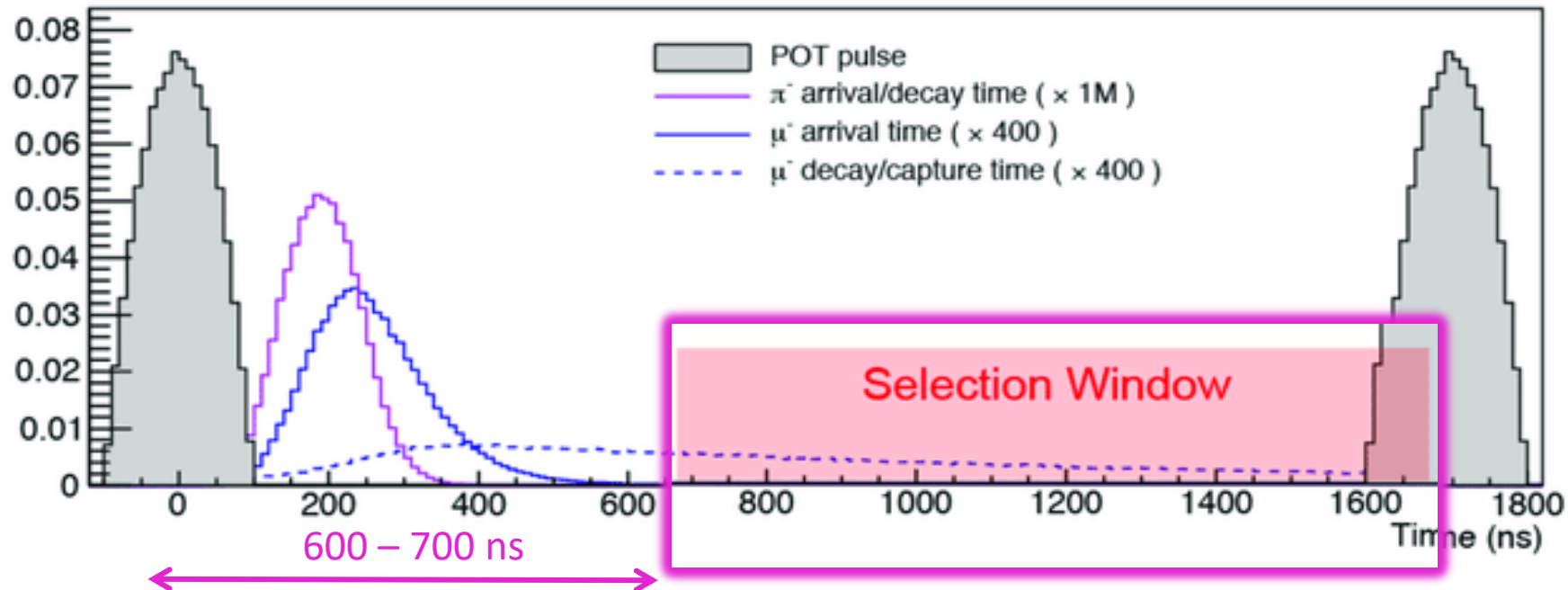


Radiative Pion Capture Backgrounds

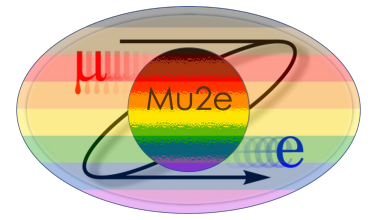


- Radiative pion capture backgrounds: $\pi^- + N(A, Z) \rightarrow \gamma^{(*)} + N(A, Z - 1)$ followed by $\gamma^{(*)} \rightarrow e^+ + e^-$.
- Pion lifetime 26 ns at rest. Pulsed proton beam (250 ns wide, pulses 1695 ns apart) \rightarrow wait out pion decay.
- In addition, upstream extinction removes out-of-time protons.

Delayed live-gate helps remove pion and beam backgrounds.

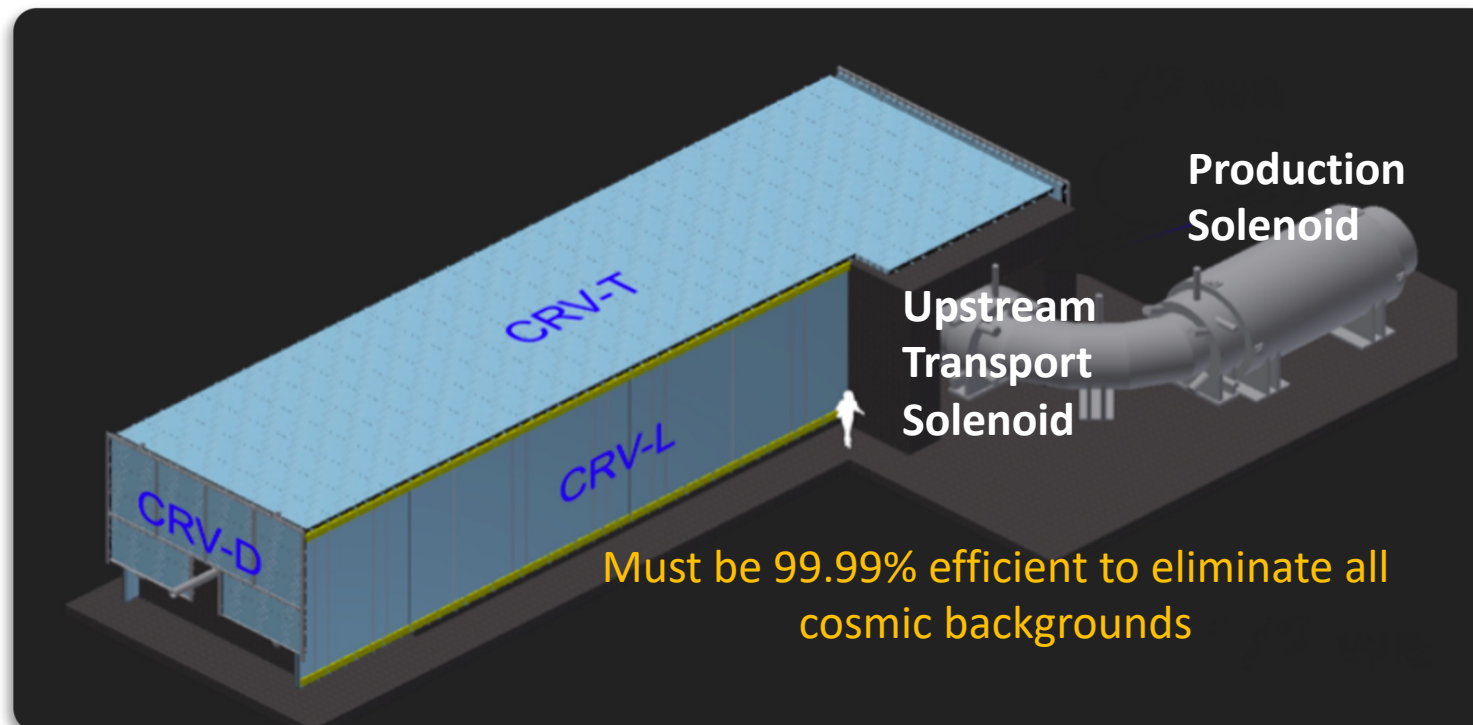


Cosmic Induced Backgrounds

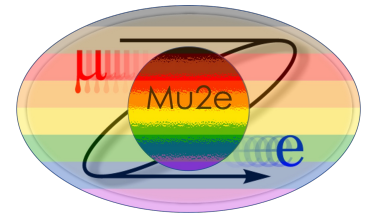


- Cosmic-ray muons can initiate 105 MeV particles that appear to emanate from the stopping target.
- Remove using active veto (CRV) + overburden and shielding concrete surrounding the Detector Solenoid.

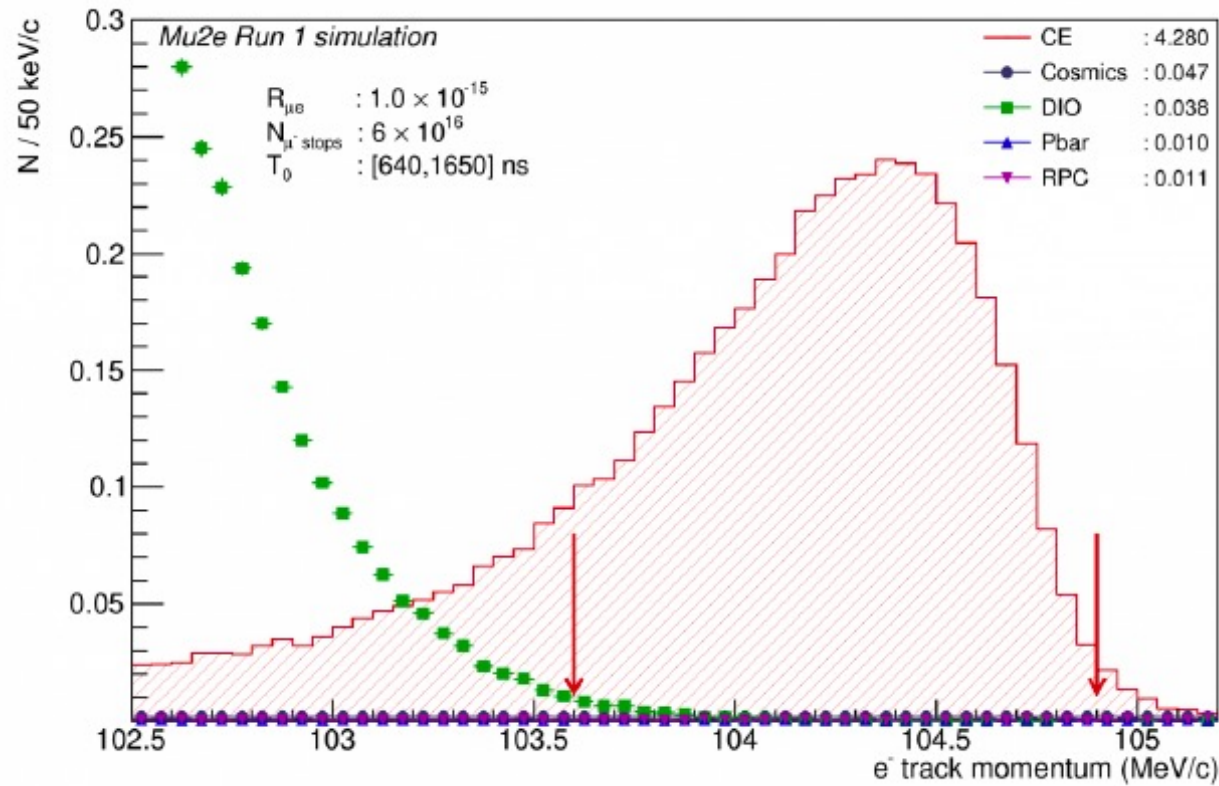
Active Cosmic Ray Veto system is key to eliminating cosmic induced backgrounds.



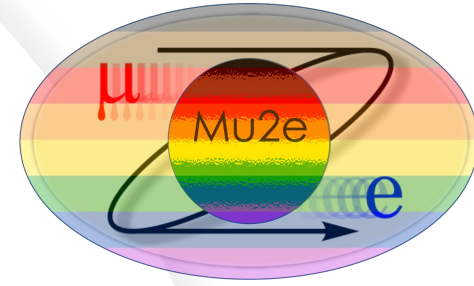
Signal & Backgrounds



Expected signal ($R_{\mu e} = 10^{-15}$) and DIO spectra from Run_I simulation (~10% of final dataset, includes resolution and energy loss effects):



Run-I Sensitivity of Mu2e:
Universe 2023, 9, 54.

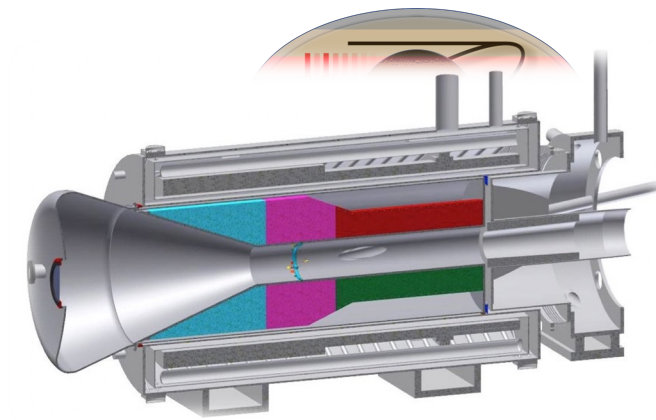


Status Update

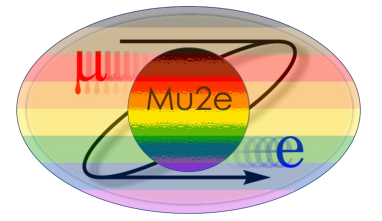
Update on current status of experimental apparatus

Solenoids

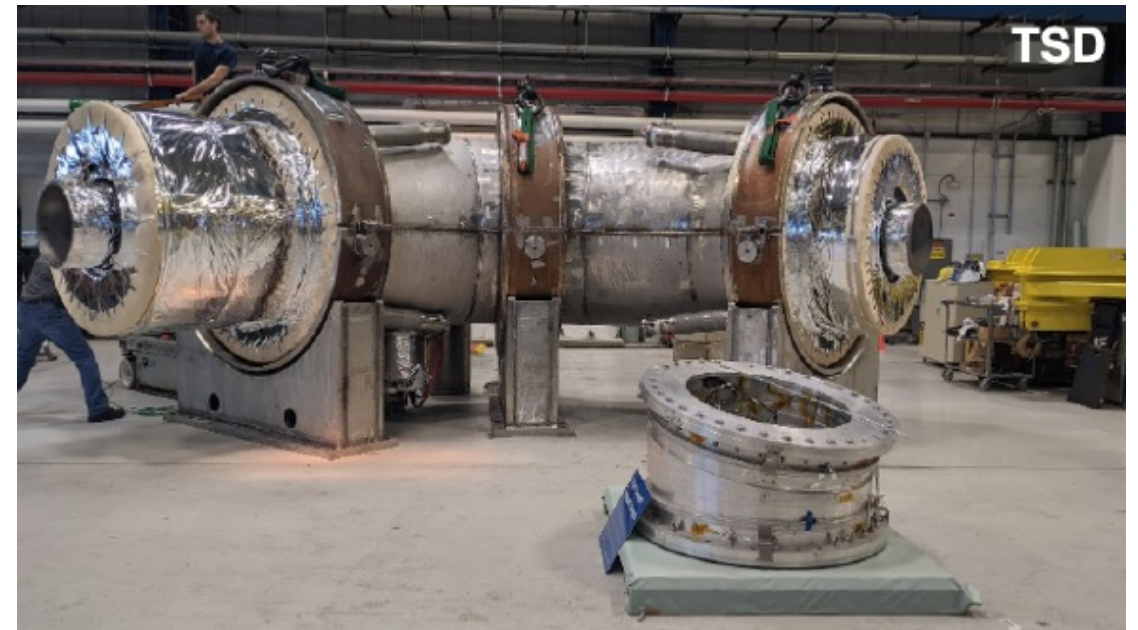
- Production Solenoid:
 - Consists of 3 coils, all wound at vendor.
 - Undergoing final assembly.
 - Arrives at Fermilab in Fall 2023.



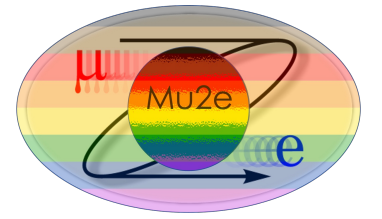
Solenoids



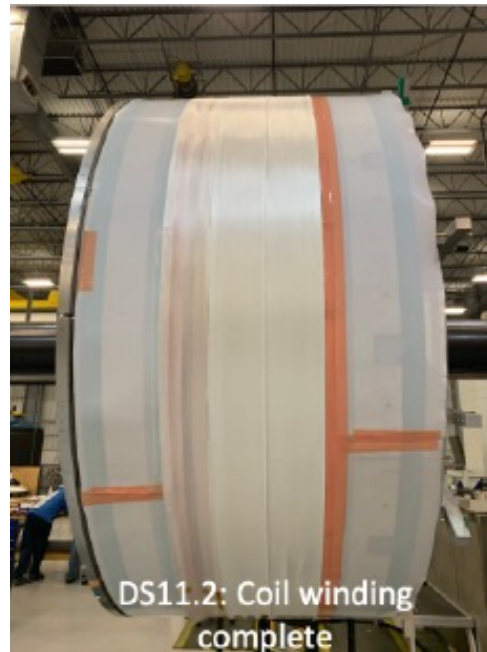
- **Transport Solenoid:**
 - Assembly being completed on-site (located in HAB).
 - Moved to Mu2e Hall in Fall 2023.



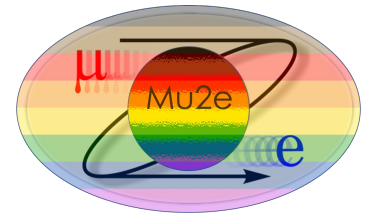
Solenoids



- **Detector Solenoid:**
 - All coils fabricated at vendor.
 - Cold mass cryo. supports prepared
 - Delivery to Fermilab expected early 2024.



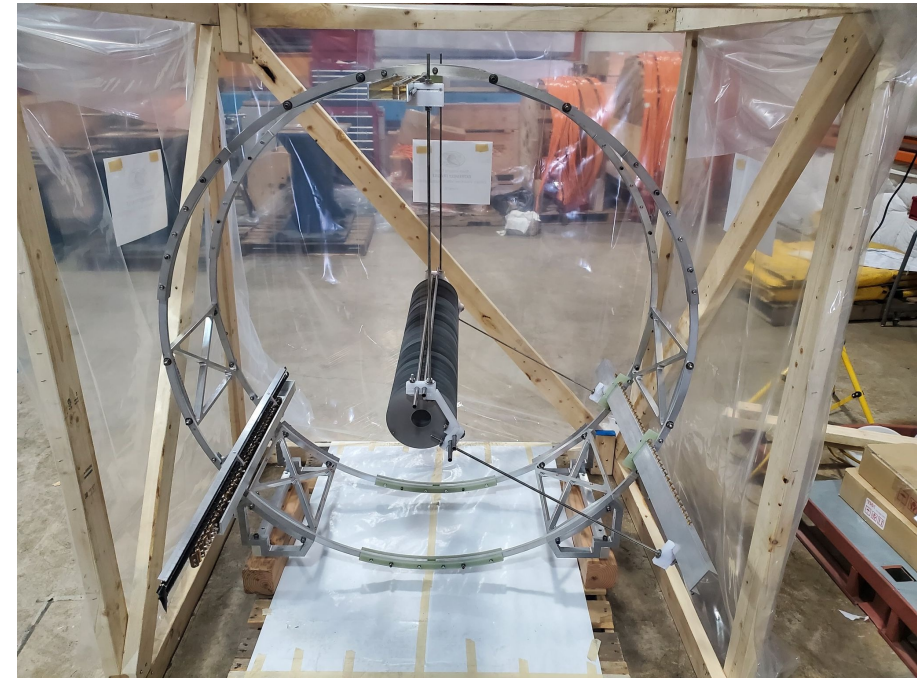
Targets



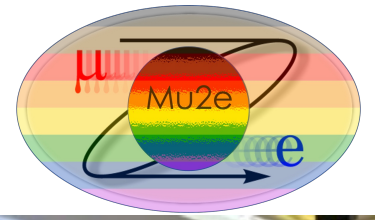
Production target: resides in Production Solenoid, stops 8 GeV protons, produces pions.



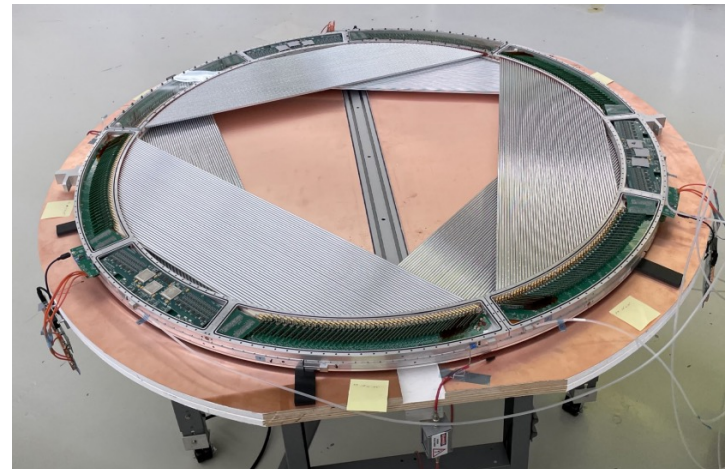
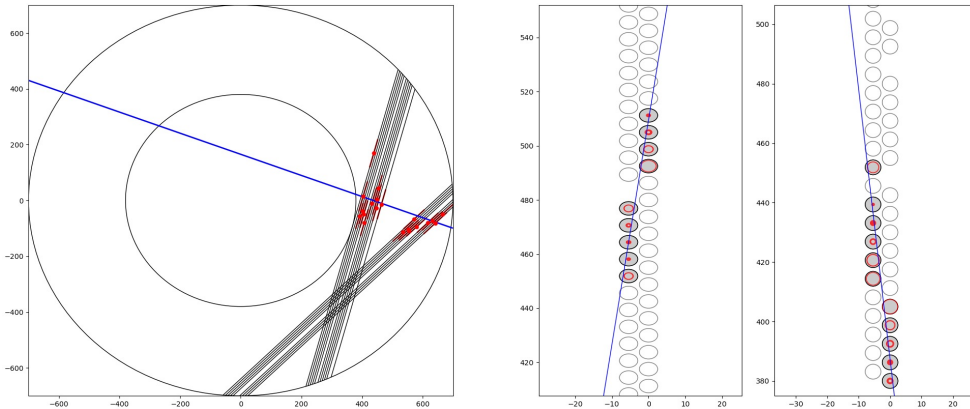
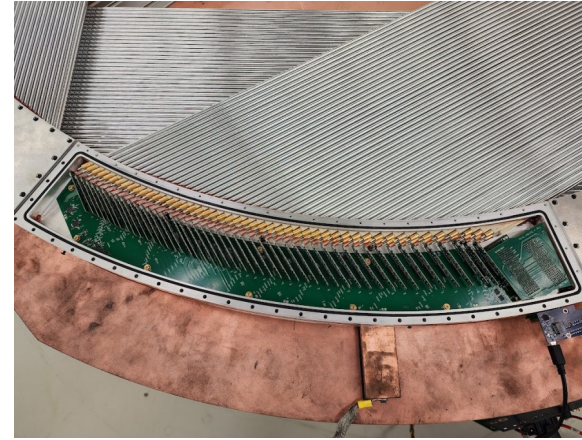
Muon Stopping target: resides in Detector Solenoid, stops muons, potentially produces signal conversion electrons.



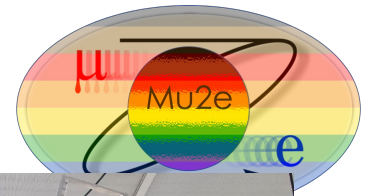
Tracker



- All 20736 straws produced.
- All 216 panels produced. Now working through QC.
- 25 / 36 planes are built.
- Cosmic ray tests carried out with a single plane and full readout system for 3 years.



Calorimeter



Calorimeter is vital for providing:

- Particle identification,
- Fast online trigger filter,
- Seed for track reconstruction.

Design:

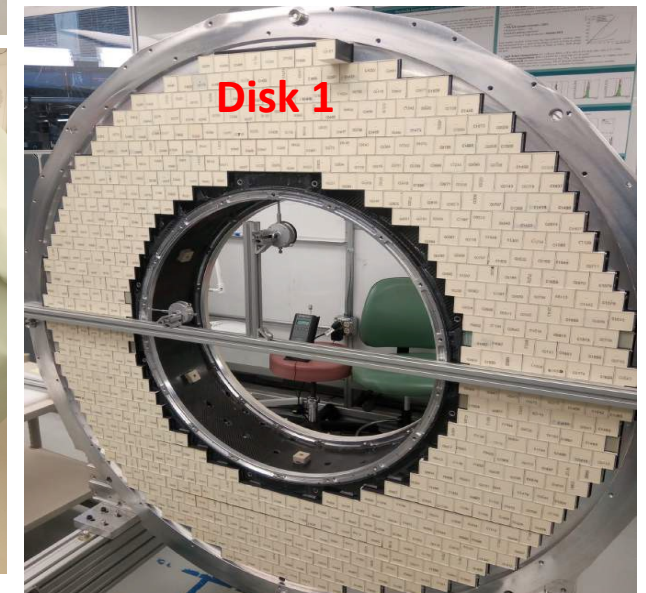
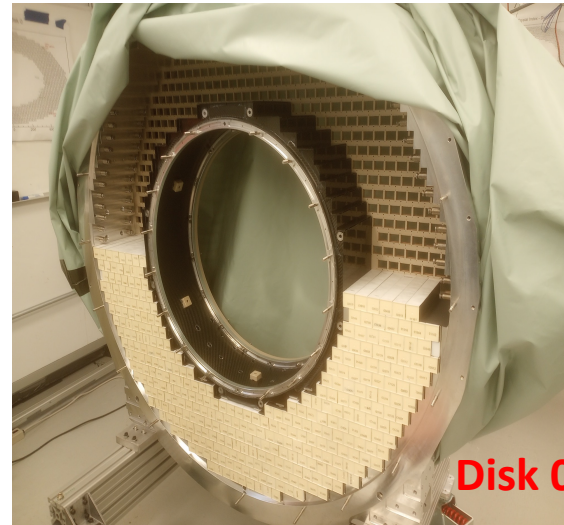
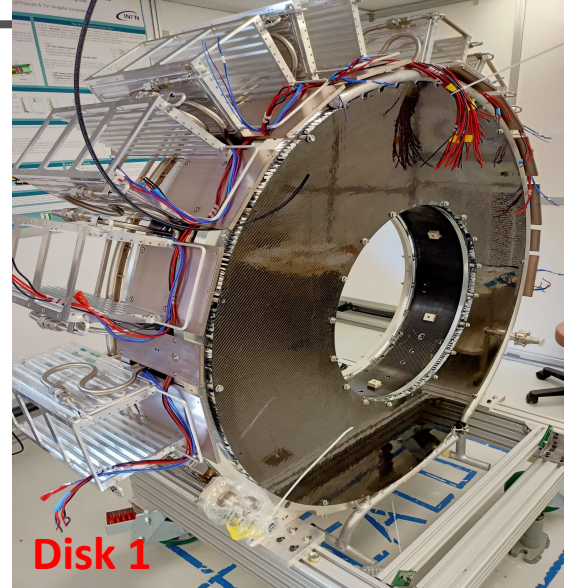
- 2 x 674 CsI crystals in 2 disks, each coupled to 2 SiPMs.

First disk (Disk – 1):

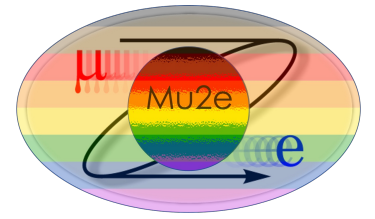
- Has crystals and SiPMs installed (2022).
- 450 ROU's installed.

Second disk (Disk – 0):

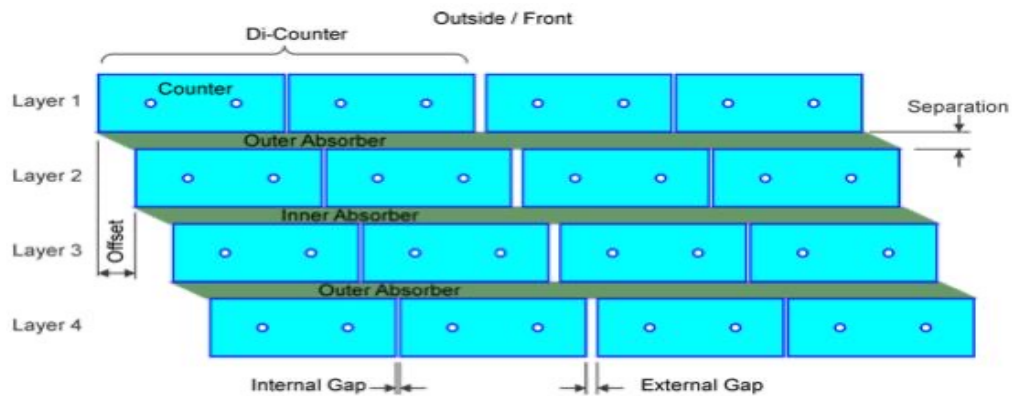
- Being assembled this summer (2023).



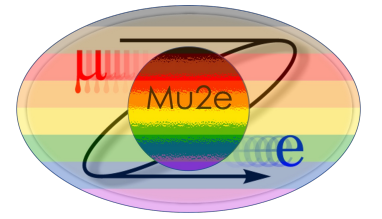
Cosmic Ray Veto System



- All 5344 di-counters produced.
- 81 / 83 modules produced.
- Cosmic ray tests underway at Wideband.



Summary



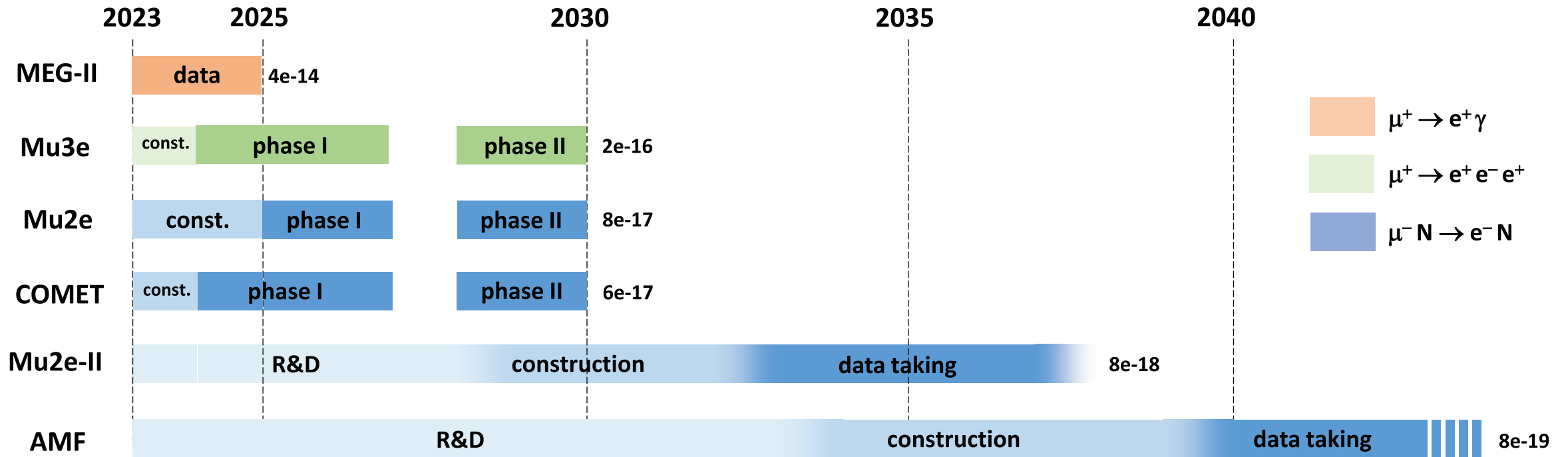
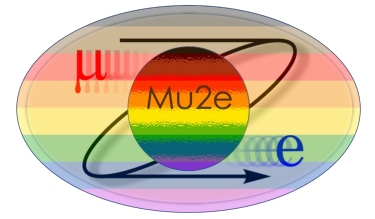
- Mu2e will search for the CLFV in muon to electron conversion with a 90% CL upper limit of $< 8 \times 10^{-17}$.
- Muon CLFV channels offer deep indirect probes into BSM. Discovery potential over a wide range of BSM models.
- Mu2e commissioning with cosmics begins in 2024, commissioning with beam in 2025 and physics data taking begins in 2026.
- Looking further ahead the proposed Mu2e-II and AMF experiments will help elucidate any signal and push to higher mass scales.

Plenty of opportunities for postdocs and new collaborators to commission a new experiment! <https://mu2e.fnal.gov/>

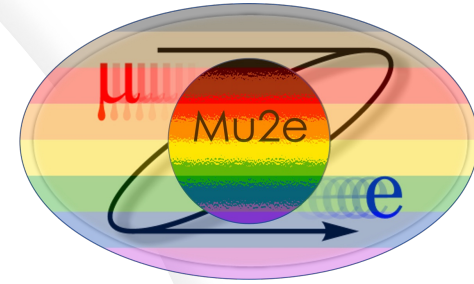
Thank you for listening!

Any Questions?

Timelines



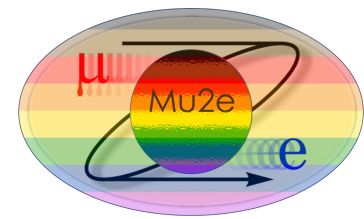
- $\mu^+ \rightarrow e^+ \gamma$
- $\mu^+ \rightarrow e^+ e^- e^+$
- $\mu^- N \rightarrow e^- N$



Additional Material

Things I couldn't fit into this talk...

Complementarity amongst channels



- All three channels are sensitive to many new physics models → discovery sensitivity across the board.
- Relative Rates however will be model dependent and can be used to elucidate the underlying physics.

Model	$\mu \rightarrow eee$	$\mu N \rightarrow eN$	$\frac{\text{BR}(\mu \rightarrow eee)}{\text{BR}(\mu \rightarrow e\gamma)}$	$\frac{\text{CR}(\mu N \rightarrow eN)}{\text{BR}(\mu \rightarrow e\gamma)}$
MSSM	Loop	Loop	$\approx 6 \times 10^{-3}$	$10^{-3} - 10^{-2}$
Type-I seesaw	Loop*	Loop*	$3 \times 10^{-3} - 0.3$	0.1–10
Type-II seesaw	Tree	Loop	$(0.1 - 3) \times 10^3$	$\mathcal{O}(10^{-2})$
Type-III seesaw	Tree	Tree	$\approx 10^3$	$\mathcal{O}(10^3)$
LFV Higgs	Loop [†]	Loop* [†]	$\approx 10^{-2}$	$\mathcal{O}(0.1)$
Composite Higgs	Loop*	Loop*	0.05 – 0.5	2 – 20

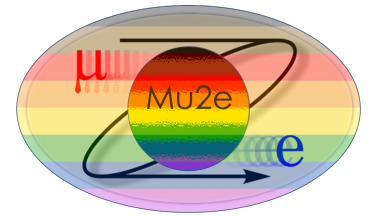
arXiv:1709.00294v2[hep-ph]

from L. Calibbi and G. Signorelli, Riv. Nuovo Cimento, 41 (2018) 71

- In seesaw models CLFV rates aren't suppressed by smallness of neutrino mass.
- Different seesaw models given very different predicted rates of CLFV.
- Measuring CLFV can help us understand neutrino mass origin!**

} Synergies with neutrino physics:
 Nuclear Physics B (Proc. Suppl.) 248–250
 (2014) 13–19

Possibilities



Outcomes of current era of CLFV searches



Signal in Mu2e and/or MEG-II and/or Mu3e



No signals



Nice papers, exciting times!



Make further measurements.



Elucidate nature of physics and flavor structure by comparing rates in different channels.



Push to higher effective mass scale, opening other BSM scenarios.



Some models excluded or heavily constrained.

$N\mu^- \rightarrow Ne^-$: Complementarity in Target Materials

Kitano et al 2002: arXiv:hep-ph/0203110v4

$$BR(\mu \rightarrow e) \propto |DC_{DL} + S^p C_{S,L}^p + V^p C_{V,R}^p + S^n C_{S,L}^n + V^n C_{V,R}^n|^2 + (L \leftrightarrow R)$$

Overlap with nucleus probes form factors and reveals the nature of the interaction.

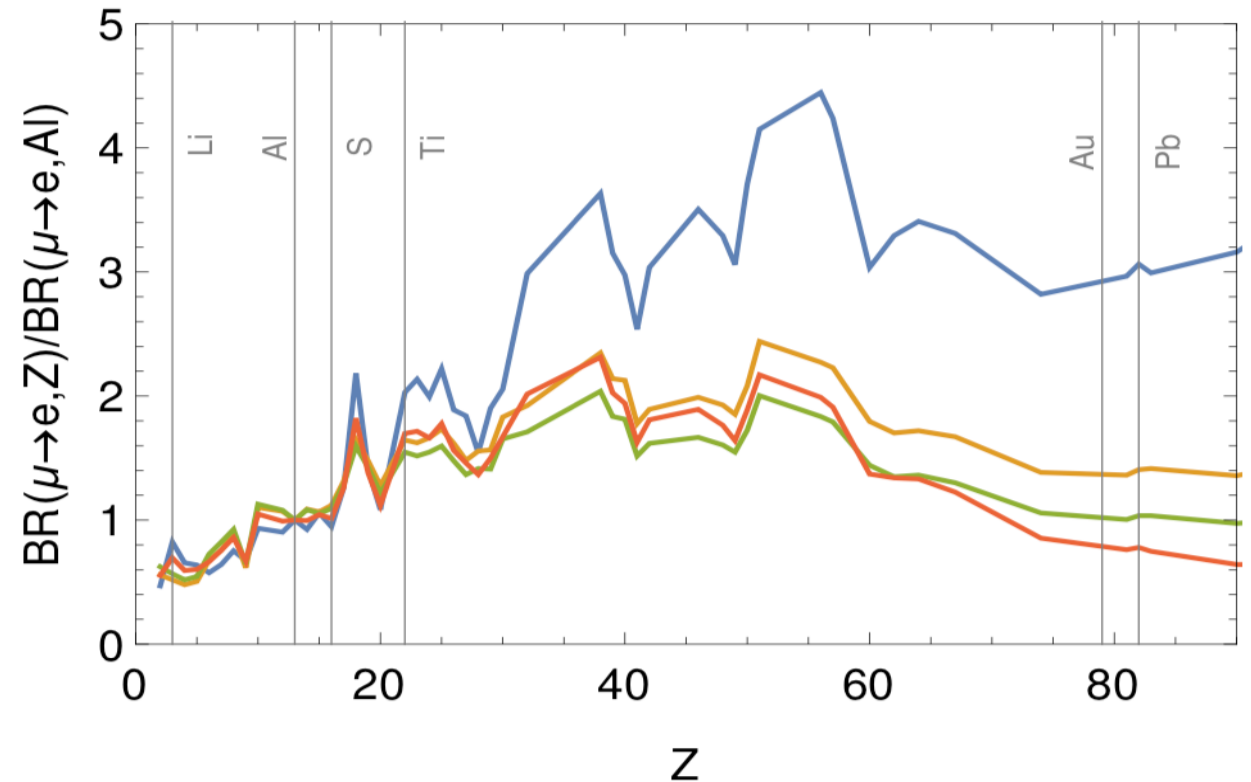
→ can elucidate type of physics through looking at relative conversion rate.

	S	D	V ¹	V ²
$\frac{B(\mu \rightarrow e, Ti)}{B(\mu \rightarrow e, Al)}$	$1.70 \pm 0.005_y$	1.55	1.65	2.0
$\frac{B(\mu \rightarrow e, Pb)}{B(\mu \rightarrow e, Al)}$	$0.69 \pm 0.02_{\rho_n}$	1.04	1.41	$2.67 \pm 0.06_{\rho_n}$

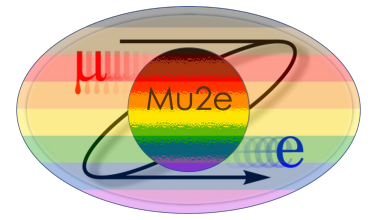
y = nuclear scalar form factor, ρ_n = nuclear neutron density

Higher Z target provides most splitting!

— Z Penguin — Charge Radius — Dipole — Scalar



Next Generation Searches



Proposed multi-decade muon CLFV at Fermilab which would utilize PIP-II and ACE 2GeV ring:

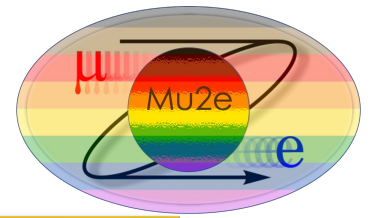
Mu2e-II [see: [arXiv: 2203.07569](https://arxiv.org/abs/2203.07569) [hep-ex]] (mid-2030s):

- ***Similar design to Mu2e, reuses much of the hardware but requires new production target and detector systems.***
- Uses pulsed beam as necessary to remove pion backgrounds.
- Lots of R&D on-going including 2 LDRD proposals: tracker and production target.

AMF [see: [arXiv: 2203.08278](https://arxiv.org/abs/2203.08278) [hep-ex]] (mid 2040s):

- ***A multi purpose muon facility which would search for all three muon CLFV channels at Fermilab.***
- Would utilize a fixed field alternating (FFA) gradient synchrotron which would provide:
 - **Monoenergetic beam of central momentum 20-40 MeV/c:** thin target, minimizing material effects, retaining momentum resolution.
 - **Pure muon beam:** don't need the pulsed beam and delayed signal window.
 - Can utilize a high Z material to elucidate physics if signal at Mu2e/COMET or Mu2e-II.
 - Has smaller decay branching fraction.
- R&D required and lots of opportunities to get involved.

Stopping Target Monitor (STM)



- Need an accurate measure of total number of stopped muons in the target (within 10%) .
- Placed far downstream of target (~34 m).
- STM uses HPGe and LaBr₃ detectors to measure X/γ-rays produced by stopped muons in Al target:
 1. Prompt X-ray emitted from muonic atoms at 347keV;
 2. Delayed gamma ray at 844keV;
 3. Semi-prompt gamma ray at 1.809MeV

$$R_{\mu e} = \frac{\Gamma(\mu^- + A(Z, N) \rightarrow e^- + A(Z, N))}{\Gamma(\text{all} - \text{captures})}$$



Normalization = from X-rays emitted when muon stops in Al.

