

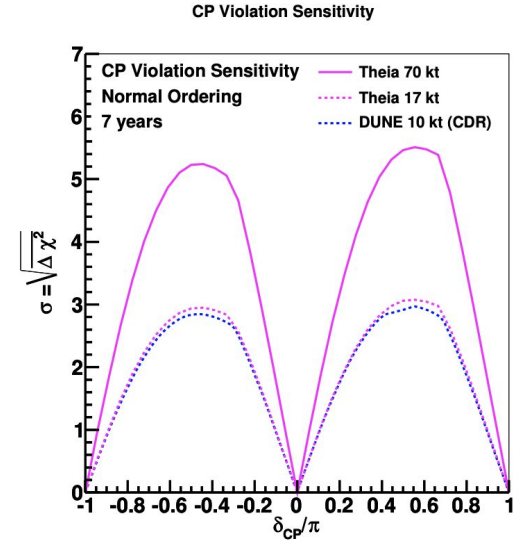
WbLS in ND-GAr:

An **Even-More** Multipurpose ND for
DUNE Phase 2?

Mike Wilking
ND-GAr Meeting
May 9th, 2023

FD Module of Opportunity

- DUNE Phase 1 will create 4 far detector caverns, but only 2 far detectors
- DUNE Phase 2 goal is to produce 2 new FDs, upgrade the beam, and upgrade the ND
 - Ongoing US P5 process may end up weighing in on these elements with regard to US funding
- Additional funding sources / new collaborators can substantially improve the odds of completing DUNE Phase 2
 - e.g. a WbLS detector (Theia) would grow the collaboration and substantially broaden the physics program (DSNB, SN burst, solar CNO, $0\nu\beta\beta$, ...) and provide a different target nucleus for the LBL program
 - Other options are also being explored (e.g. LiquidO)
- New FD detector technologies require new ND capabilities
- Expanding the capabilities of ND-GAr (e.g. to include WbLS targets) can grow the number of ND-GAr stakeholders and funding opportunities

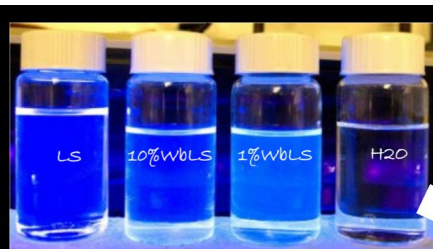


Brief Introduction to Theia

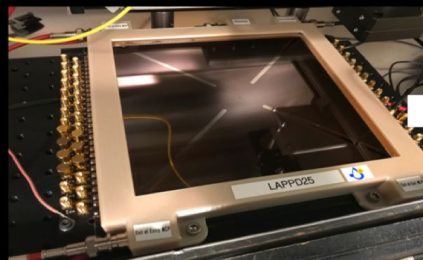
[The European Physical Journal
C 80 no. 5, \(May, 2020\)](#)

(slide from M. Wurm)

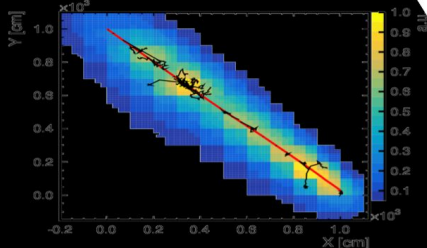
- Observe both Cherenkov and scintillation light for each event
- At High-E, operates much like Super-K
 - with additional information about below-Cherenkov-threshold hadrons
- Low-E threshold is lowered
 - More photons near and below Cherenkov threshold



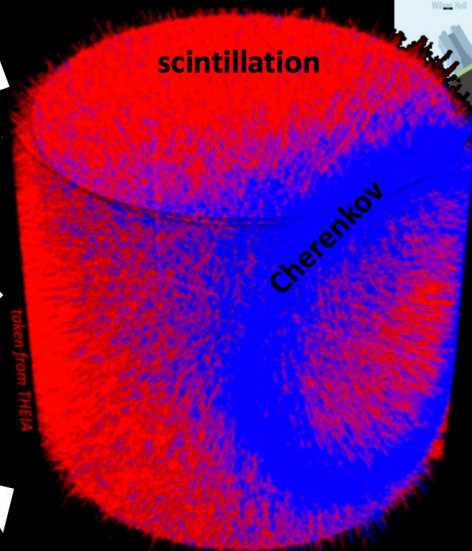
Novel target medium:
Water-based Liquid Scintillator



Novel light sensors:
LAPPDs, dichroicons

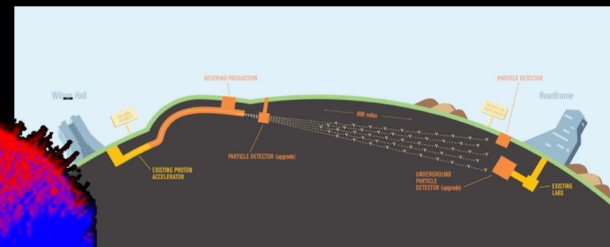


Novel reconstruction techniques



Large volume detector
able to exploit both
Cherenkov+Scintillation
signals

→ **Enhanced sensitivity to broad physics program**



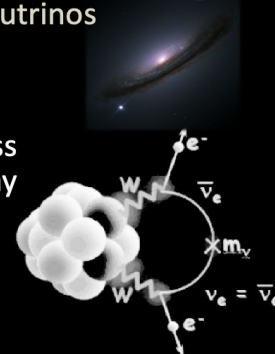
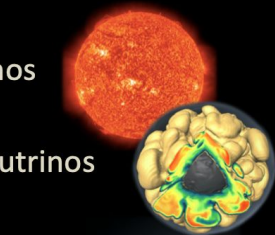
→ **Long-Baseline Oscillations**

→ Solar neutrinos

→ Supernova neutrinos

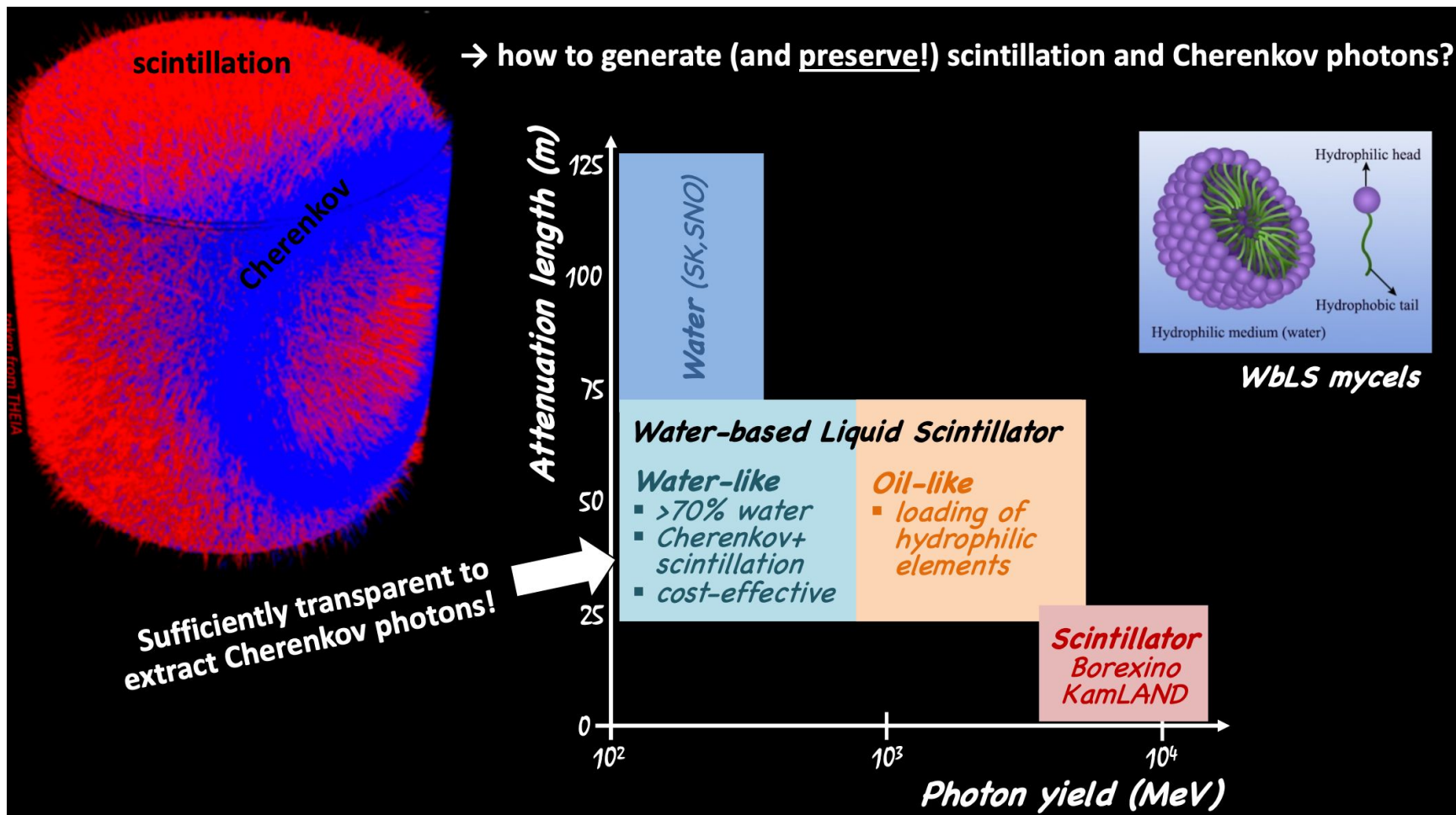
→ Diffuse SN neutrinos

→ Neutrinoless
Double-Beta Decay



WbLS Concentrations

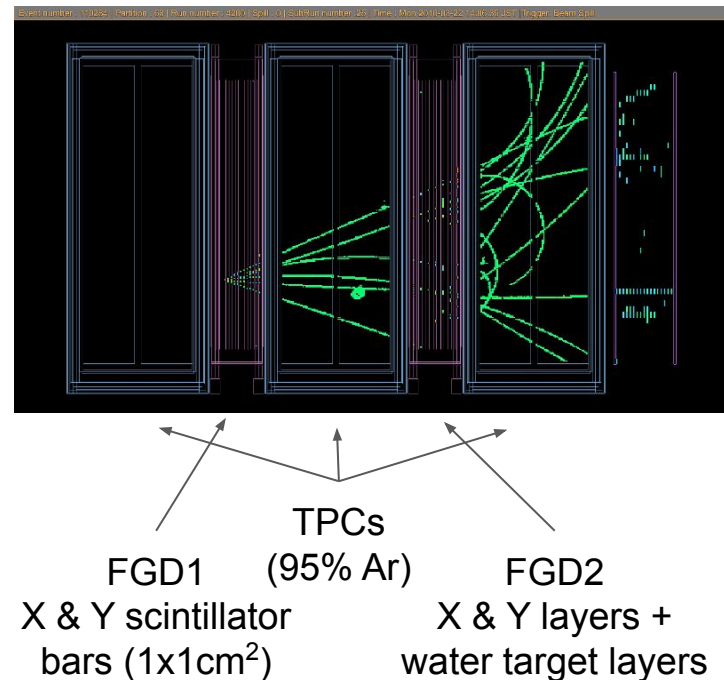
(slide from M. Wurm)



WbLS Near Detector Considerations

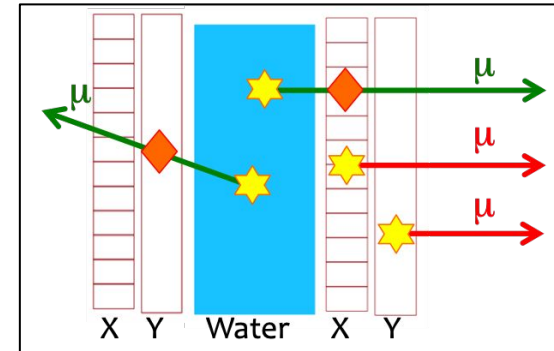
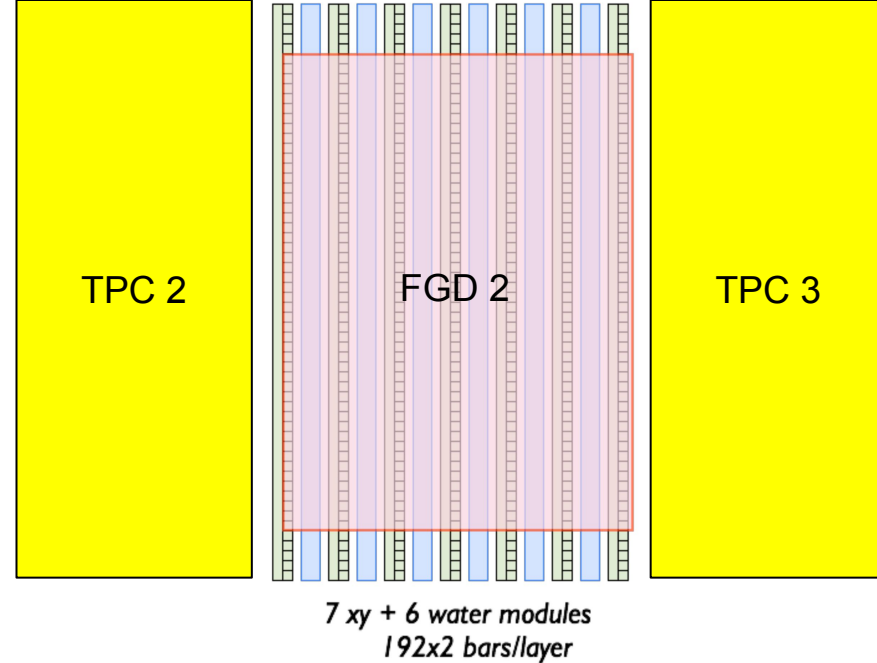
- A key component of LAr detectors is hadron calorimetry
 - Neutrino energy is the sum of the reconstructed lepton energy and the (corrected) deposited hadronic energy
- For water Cherenkov detectors, E_ν reconstruction is performed with above-Cherenkov particles
 - The Theia LBL sensitivity studies were performed without utilizing scintillation light
- The primary requirement for a Theia near detector is to measure above-Cherenkov-threshold particles
 - This is the approach used for the primary T2K / Hyper-K near detector
 - Additional external measurements of Cherenkov/scintillation ratio may be needed
 - Large R&D program with several WbLS detectors is currently underway

T2K Near Detector (ND280)



T2K Fine-Grained Detector (FGD2)

- T2K already employs water targets embedded within X & Y layers of scintillator bars
 - This reduced T2K's neutrino interaction uncertainties on water by $\sim 30\%$
- One of the most important detector uncertainties is disentangling events occurring within water to events occurring in adjacent scintillator layers
- The key difference using WbLS is the water layers themselves can be instrumented
 - Surrounding scintillator layers are not longer a strict requirement
 - Must ensure a sufficient light yield to record MIPs

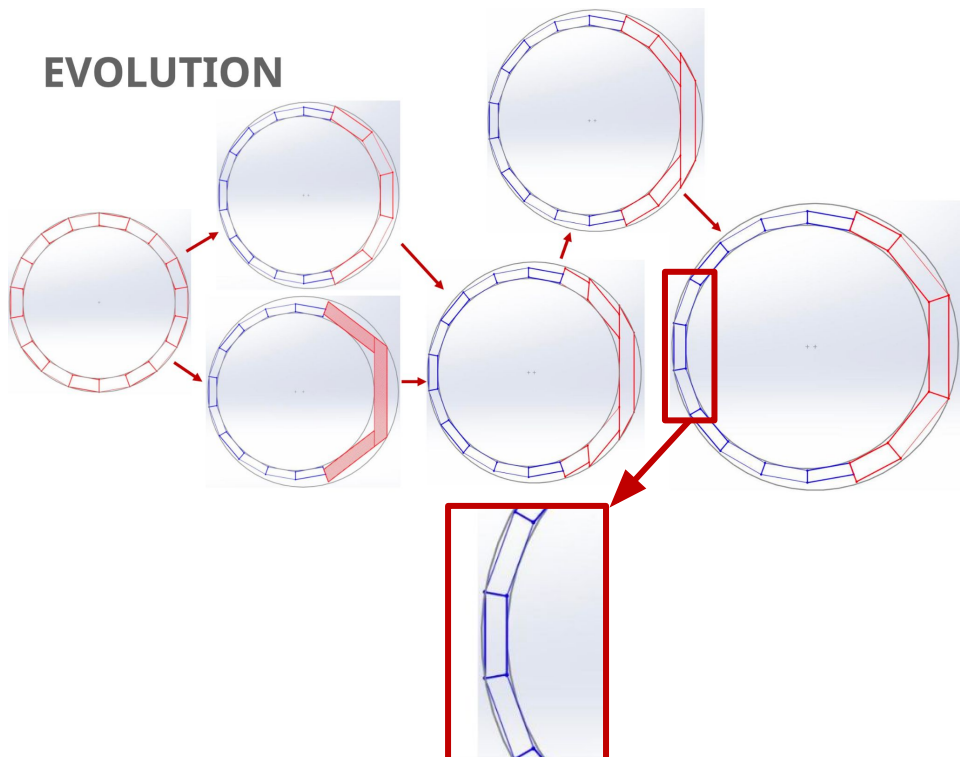


ND-GAR “Thin” Upstream ECAL

- The ND-GAR ECAL is most needed in the downstream direction
- The upstream portion has been redesigned to be “thin”
- Questions for ND-GAR ECAL experts:
 - How does the thickness requirement vary as a function of angular position? Can the most upstream portions be thinner?
 - Could the downstream-most layers of the upstream ECAL be composed of WbLS? What would be the light yield requirements in this case?
 - Some additional effort may be available to perform relevant studies

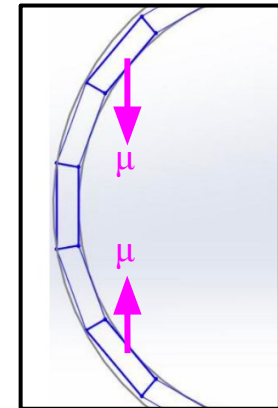
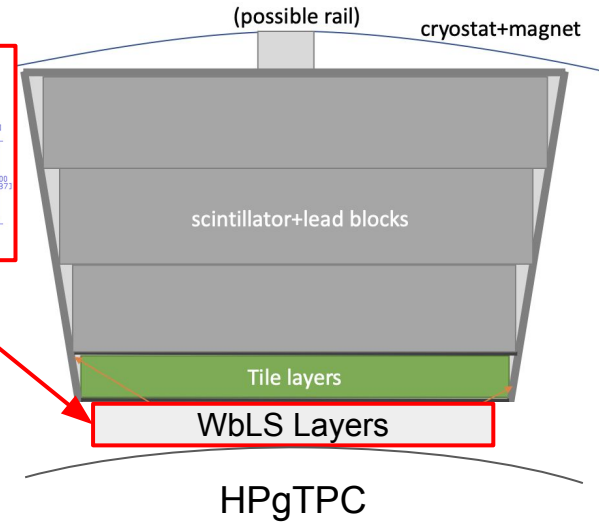
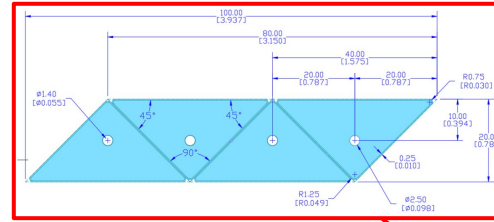
ND-GAR ECAL Design Evolution

https://indico.fnal.gov/event/50217/contributions/241513/attachments/155287/202160/220517_DUNE_CM_Talk_ECAL_Concepts.pdf



WbLS Inside ND-GAr ECAL

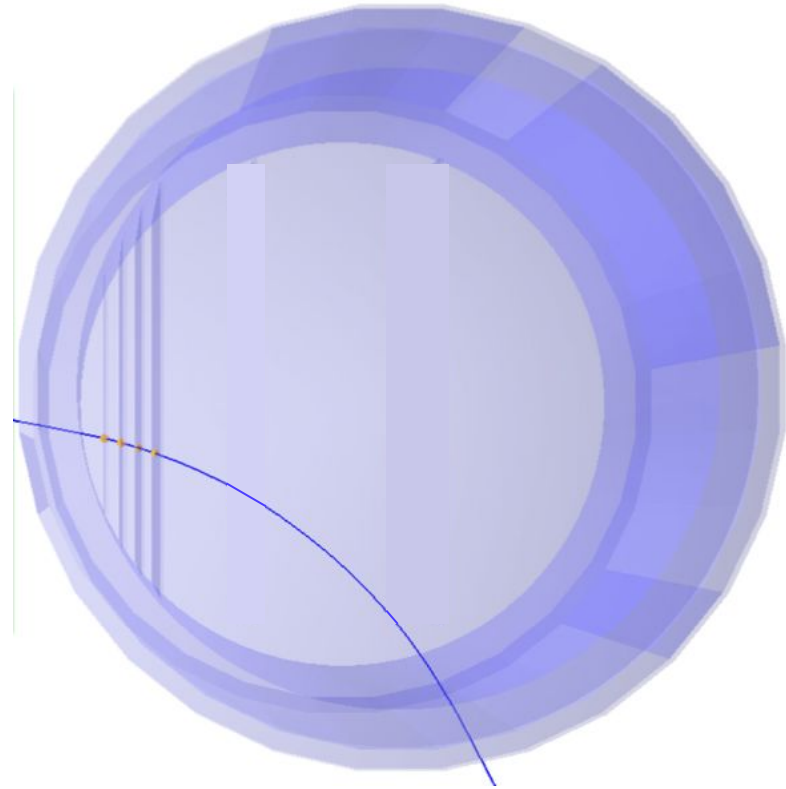
- WbLS layers would need to track X & Y positions
 - Optically segmented X & Y bars or 3D cubes
 - Or perhaps a non-segmented LiquidO detector with X & Y fibers
- Order of magnitude size estimate:
5 m long TPC * ~4 m total layer width * 5 cm thickness
≈ 1 ton WbLS detector mass
 - Similar to the target mass for the GAr TPC
- Additional benefit: variation in detector configurations allows for sampling all of the muon angle phase space
 - The lack of muon acceptance near 90° was an important limitation of the T2K FGD+TPC configuration



WbLS + GAr TPC?

Figure adapted from
ND-GAr-Lite proposal

- A more invasive approach: integrate WbLS layers into the GAr TPC
 - Lower threshold particle detection
 - Precision tracking upstream + downstream of each layer
- Important to not harm ND-GAr physics
 - Minimize reduction of GAr-TPC fiducial volume
 - Minimize loss of performance in upstream ECAL (these layers would effectively form the first ECAL layers, but would consist of <1 radiation length)
- Many technical issues to consider
 - Additional penetrations and support structure in the pressure vessel
 - Field shaping on the surfaces of the WbLS layers
 - Potential liquid leaks
 - The T2K FGD2 water layers were held at lower pressure so that small leaks would produce gas ingress rather than water outflow



Summary and Next Steps

- Purpose of today's talk is to begin a discussion with ND-GAr experts on potential synergies between the ND-GAr and WbLS communities for DUNE Phase 2
 - Many potential benefits to such a collaboration (both politically and to enhance DUNE physics)
 - Additional potential person-power exists to study the options discussed today
- Multiple options for integrating WbLS into ND-GAr could be considered
 - Minimally invasive: place targets in downstream most layers of upstream ECAL
 - More invasive: integrate WbLS layers into GAr-TPC
- Based on the discussion here and any subsequent feedback from ND-GAr experts, I'll present a slightly more detailed concept at the DUNE Phase 2 ND workshop at Imperial College June
 - More discussion is also welcome at the upcoming DUNE collaboration meeting