Coherent Elastic Neutrino-Nucleus Scattering Status and Prospects

Enectali Figueroa-Feliciano \ NuInt São Paulo \ April 2024

Enectalí Figueroa-Feliciano Northwestern



Neutrino Energy





10 MeV

Reactor neutrinos



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See Lucas' talk





Accelerator based neutrinos

See Zarko's tal



Spallation neutrinos



See Yuri's talk





Neutrino Energy





10 Mey

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E(v)

Accelerator based neutrinos

See Zarko's tal

Fermilab Accelerator Complex 2012

Spallation neutrinos



See Yuri's talk

Fermilab





Neutrino Energy







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Accelerator based neutrinos

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Fermilab





Neutrino Energy



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See Lucas' talk

Hadron Beam Facility

r based neutrinos

E(v)

Fermilab Accelerator Complex 2012



Yuri's talk

Credit: Netflix Stranger Things





Neutrino Energy



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Hadron Beam Facility

r based neutrinos

E(v)

Fermilab Accelerator Complex 2012



Yuri's talk

Credit: Netflix Stranger Things





Life is simple at low energies



 $CE\nu NS$

Dessel et al., Universe 2023, 9(5), 207; https://doi.org/10.3390/universe9050207

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Quasi-Elastic

CCQE, NCQE





Coherent Elastic v-Nucleus Scattering

- σ: Cross Section
- T: Recoil Energy
- E_v : Neutrino Energy
- G_F: Fermi Constant
- Q_W: Weak Charge
- M_A: Atomic Mass
- Fw: Form Factor



No flavor-specific terms!!! Same rate for v_e , v_μ , and v_τ



Coherent Elastic v-Nucleus Scattering

 $CE\nu NS$

PHYSICAL REVIEW D

VOLUME 9, NUMBER 5

1 MARCH 1974

Coherent effects of a weak neutral current

Daniel Z. Freedman[†]

National Accelerator Laboratory, Batavia, Illinois 60510 and Institute for Theoretical Physics, State University of New York, Stony Brook, New York 11790 (Received 15 October 1973; revised manuscript received 19 November 1973)

Our suggestion may be an act of hubris, because the inevitable constraints of interaction rate, resolution, and background pose grave experimental difficulties for elastic neutrino-nucleus scattering. We will discuss these problems at the end of this note, but first we wish to present the theoretical ideas relevant to the experiments.

> Also: D. Z. Freedman et al., "The Weak Neutral Current and Its Effect in Stellar Collapse", Ann. Rev. Nucl. Sci. 1977. 27:167-207



 $\frac{d\sigma}{dT} = \frac{G_F^2}{4\pi} Q_W^2 M_A \left(1 - \frac{T}{E_\nu} - \frac{M_A T}{2E_\nu^2} \right) F_W(q^2)^2$ Principles and applications of a neutral-current detector VOLUME 30, NUMBER 11 A. Drukter and L. Stodolsky Max-Planck-Institut für Physik und Astrophysik, Werner-Heisenberg-Institut für Physik, Munich, Federal Republic of Germany PHYSICAL REVIEW D PHYSICAL REVIEW LETTERS Bolometric Detection of Neutrinos Blas Cabrera, Lawrence M. Krauss, and Frank Wilczek Blas Cabrera, Lawrence IVI, Krauss, and Frank Wilczek Department of Physics, Stanford University, Stanford, California 94305 Laboratory of Physics, Hanvard Halvereity Combridge Massachusette (). Department of Physics, Stanford University, Stanford, California 94305 Lyman Laboratory of Physics, Harvard University, Cambridge, Massachusetts 01238 stitute for Theoretical Physics. University of California. Santa Barbara. California 93106 Lyman Laboratory of Physics, Harvard University, Cambridge, Massachusetts 01238 Institute for Theoretical Physics, University of California, Santa Barbara, California 93106 VOLUME 55, NUMBER 1





Coherent Elastic v-Nucleus Scattering

 $CE\nu NS$

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Form factor means you want to look at low energies



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Dessel et al., Universe 2023, 9(5), 207; https://doi.org/10.3390/universe9050207







Neutrino Sources: the Sun & other cosmic sources

- Need lots of exposure (mass * time)
- Probably need to do this underground...
- We'll leave these to the dark matter experiments







Neutrino Sources for $CE\nu NS$





Energy Thresholds Needed for mono-energetic Neutrinos







10 N

Detector Response: Need Low-thresholds!



Large cross section allows for "tabletop" neutrino experiments with kg-scale mass!

11 N





- 12 N





- 12 N





- 12 N





- 12 N

$CE\nu NS$ Science

- **EW Precision tests:**
 - Weak Mixing Angle
- New neutrino interactions
 - Non-standard interactions
 - Generalised interactions
 - New mediators
- Neutrino Properties
 - Neutrino Charge Radius
 - Magnetic Moments
- Nuclear Physics
 - Nuclear form factors
 - Neutron radius and "skin"

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Brdar and Rodejohann, arXiv:1810.03626; Chang and Liao, arXiv:2002.10275; Li et al, arXiv:2005.01543; CONUS, arXiv:2110.02174; Cadeddu et al, arXiv:1710.02730, arXiv:2005.01645, arXiv:1908.06045; Aristizabal Sierra et al, arXiv:1902.07398; Huang and Chen, arXiv:1902.07625; Papoulias et al, arXiv:1903.03722, arXiv:1907.11644; Miranda et al, arXiv:2003.12050; Papoulias et al, arXiv:1711.09773, arXiv:1907.11644; Cadeddu et al, arXiv:1808.10202, arXiv:2005.01645, arXiv:1908.06045, arXiv:2205.09484; Huang and Chen, arXiv:1902.07625; Miranda et al, arXiv:1902.09036, arXiv:2003.12050; Khan and Rodejohann, arXiv:1907.12444; COHERENT, arXiv:2110.07730; Papoulias and Kosmas, arXiv:1711.09773; Blanco et al, arXiv:1901.08094; Miranda et al, arXiv:1902.09036, Cerdeño et al, arXiv:1604.01025; Farzan et al, arXiv:1802.05171; Aristizabal Sierra et al, arXiv:1806.07424; Khan and Rodejohann, arXiv:1907.12444; Aristizabal Sierra et al, arXiv:1910.12437; Miranda et al, arXiv:2003.12050; Aristizabal Sierra et al, JHEP 09 (2019) 069; Suliga and Tamborra, arXiv:2010.14545; CONUS, arXiv:2110.02174; Li and Xia, arXiv:2201.05015; Atzori Corona et al, arXiv:2202.11002; Liao et al, arXiv:2202.10622; Coloma et al, arXiv:2202.10829; Lindner et al, arXiv:1612.04150; Aristizabal Sierra et al, arXiv:1806.07424; Aristizabal Sierra et al, JCAP 01 (2022) 01, 055













COHERENT Limits on Weak Mixing Angle and Skin Radius





De Romeri, Miranda, DKP, Sanchez, Tortola, Valle: arXiv: 2211.11905





$CE\nu NS$ Future Science Example: Ricochet



Spectral information improves searches for new physics!



Many CEvNS Efforts Worldwide [incomplete]

Experiment	Technology	Location	Source
COHERENT	CsI, Ar, Ge, Nal	USA	πDAR
ССМ	Ar	USA	πDAR
ESS	CsI, Si, Ge, Xe	Sweden	πDAR
BULLKID	Si/Ge	Italy	Reactor
CONNIE	Si CCDs	Brazil	Reactor
CONUS	HPGe	Germany	Reactor
NEWS-G	Ar+2%CH4	Canada	Reactor
MINER	Ge/Si cryogenic	USA	Reactor
NEON	NaI(TI)	Korea	Reactor
NUCLEUS	CaWO ₄ , Al ₂ O ₃ cryogenic	Europe	Reactor
νGEN	Ge PPC	Russia	Reactor
RED-100	LXe dual phase	Russia	Reactor
Ricochet	Ge, Zn, Al, Sn cryogenic	France	Reactor
TEXONO	p-PCGe	Taiwan	Reactor
Dresden II	PCGe	USA Reactor	
SBC	Scintillating Bubble Chamber	Fermilab Reactor (R&D)	

+DM detectors, +directional detectors +Solar/SN detectors... many novel low-background, low-threshold technologies!!





DAR Experiments



17 N



D. Akimov et al. (COHERENT). Science 357, 1123–1126 (2017)



COHERENT Measurements on Ar and Csl





COHERENT Germanium Detectors

- 8 2.2-kg Ge PPC detectors from Mirion
- First deployment 2022 ullet
- Physics run 2023 ullet
- 110-150 eV FWHM ulletpulser resolution/noise





- High-density polyethylene
- Copper box
- Copper cryostat
- Preamplifier
- Liquid nitrogen
- Structural frame







First Ge Results (preprint soon!)





COHERENT Cross Sections: Filling Out the Curve!





22 N

COHERENT CEvNS Detector Status

Nuclear Target	Technology	Mass (kg)	Distance from source (m)	Status
Csl[Na]	Scintillating crystal	14.6	19.3	First result 2017; final result 2021; decommissioned
COH-Ar-10 (CENNS-10)	Single-phase LAr	24	27.5	First result 2020; x2 data analysis underway; decommissioned
COH-Ge-18 (GeMini)	HPGe PPC	18	22	First result 2023/24; continued running
COH-Nal-3000 (NalvETe)	Scintillating NaI[TI] crystal	3400	25	Staged deployment up to 3400 kg; first modules taking commissioning data
COH-Ar-750	Single-phase LAr	750	27.5	First data expected 2025
COH-CryoCsl-10	Cryogenic scintillating CsI crystal	10-15	~20	Proposed









+D₂O for flux normalization

Reactor Experiments







 $E_{\rm M}[\rm keV]$







- 2 Skipper-CCDs (0.4 g mass), exposure 18.4 g-days.



- ullet



NEW [arXiv: 2403.15976]

See Poster by Pedro Ventura

cross-section, obtained by a surface-level experiment.









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<u>10.1140/epjc/s10052-022-11150-x</u> <u>10.1140/epjc/s10052-024-12433-1</u>



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Institut Laue-Langevin (ILL) Nuclear Research Reactor



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- 58 MW nominal thermal power
- Located ~9 m from reactor core
- ~11 evts/day/kg (50 eV threshold)
- 3 to 4 cycles per year: excellent ability to subtract uncorrelated backgrounds
- Significant overburden (~15 m.w.e) to reduce cosmics
- Ricochet commissioning at ILL started February 2024!



28 N

Ricochet Detectors Installed and Taking Data!



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11 CEvNS events/day and 10³ Electron-recoil rejection factor ~4 (20) sigma CEvNS detection after one reactor cycle including low-energy excess background (or not)

- Provides gamma discrimination

RICOCHET 1st commissioning run at ILL

Other DAR Experiments

CCM **Coherent CAPTAIN-Mills** At LANSCE proton beam

ESS & NCC-1701

- summer of 2025.
- lacksquare
- •

Cryogenic undoped Csl

ERC-Advanced grant

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Three fully-funded technologies for CEvNS@ESS. Start of onsite activities

Reinstallation of upgraded NCC-1701 in Vandellós' tendon gallery in progress. Expected signal/bckg of >40 (~1/4 @ Dresden)

Additional QF measurement at OSURR (1E5 Ge NRs below 700 eVnr collected) Update: https://indico.ess.eu/event/3495/

p-type point contact Ge

high pressure gas TPC

31 N

CONUS+ **21 m from KKL Power Plant**

- CONUS shield modified to include additional second muon veto.
- CONUS Ge detectors upgraded. Energy resolution and threshold improved.

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vGEN **11.5 m from KNPP Power Plant**

New results with more statistics and optimized measurement modes are expected soon. Search for evidence of other effects in data are ongoing.

32 N

RED-100 Moving from Xe to Argon

RED-100 Moving from Xe to Argon

RED-100 Moving from Xe to Argon

Conclusions

- COHERENT is leading the way with detections in CsI, Ar, and Ge! Physics impact already very strong.
- New upgrades in the COHERENT program plus other DAR sources (ESS) will bring lots of new data in the coming years.
- Reactor experiments have made some contested claims of detection, but we will have several definitive measurements with higher statistics within a year or two(ish).
- Quenching factor measurements are crucial for the interpretation of reactor data, and many new measurements are being done to provide clarity on QF values (especially for Ge).
- New technologies are being brought to bear, and we can look forward to future percent-precision measurements of CEvNS spectra.

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Additional Slides on Specific Experiments

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NUCLEUS

NUCLEUS: Study $CE\nu NS$ from reactor antineutrinos

VNS

*Comics courtesy Chloe Goupy, CEA, 2022

Status and prospects of NUCLEUS

Milestone result in 2022: Calibration at 100 eV by observing the capture of thermal neutrons on a NUCLEUS target crystal

First observation of a 100eV scale nuclear recoil peak \rightarrow

Spring 2023 – Full setup started commissioning in the TUM-UGL

- Mechanical integration tests lacksquare
- Calibrations at keV energies and below
- Detector performance ${ \bullet }$

Experimental room at the Power Plant ready

The NUCLEUS collaboration

Please check out the NUCLEUS homepage for further info: https://nucleus-experiment.org/

\approx 50 collaborators

CONUS

CEVNS search at Brokdorf nuclear power plant with CONUS

ents

e

SN

Ш О

- full background decomposition
- reactor thermal power correlated neutrons negligible inside CONUS shield
- precise quenching measurement in Ge at 88 K:
 - validity of Lindhard confirmed in (0.4-6.3) keV_{nr}
 - quenching factor k=0.162+-0.04 (stat. + syst.)

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5 years of successful operation of 4 x 1kg Ge detectors @ 17 m distance from a 3.9 GW_{th} reactor core center

Key parameters: intense neutrino flux: $\Phi = 2.3 \times 10^{13} \text{ v/s/cm}^2$ ultra low energy threshold: ~200 eV_{ee} ultra low bkg in ROI: O(10) cts/d/kg

41 N

CONUS and **Dresden-II**

Lindhard, k=0.162 CONUS meas.

CONUS+

CONUS+ will be installed KKL power plant during summer 2023.

4 upgraded x 1kg Ge detectors @ 21 m distance from a 3.6 GW_{th} reactor core center

Key parameters: intense neutrino flux: $\Phi = 1.5 \times 10^{13} \text{ v/s/cm}^2$ ultra low energy threshold: $<200 \text{ eV}_{ee}$

- Full background characterization of detector location:
- Gamma background with HPGe detector. Contribution over 2.7 MeV \rightarrow 25 times smaller than in KBR.
- Neutron background measured with Bonner Sphere array. Neutron flux 30 times larger than in KBR. Correlated with thermal power.
- Cosmic muons measured with liquid scintillator detector. 4 times larger than in KBR \rightarrow overburden on 6 m.w.e.
- Radon level 200 Bq/cm3 \rightarrow radon filtering system for detector chamber flushing.

- CONUS shield modified to include additional second muon veto.
- CONUS Ge detectors upgraded. Energy resolution and threshold improved.

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VGeN

44 N

experiment at Kalinin nuclear power plant

- applications, including reactor monitoring. to detect desired signals.
- \checkmark The passive and active shielding covers detectors from all sides suppressing backgrounds. \checkmark Active antivibration platform suppresses microphonic noises. \checkmark Nitrogen flushing decrease background from radon. \checkmark Movable platform allows to change the distance to the reactor, thus changing the v flux. Special acquisition system was developed to suppress noise and achieve energy resolution of 101.6(5) eV (FWHM).

Alekseev et al., Phys. Rev. D 106, L051101 (2022)

✓ The vGeN project studies neutrino scattering at Kalinin Nuclear Power Plant (KNPP, Russia). Main interests: coherent elastic neutrino-nucleus scattering (CEvNS), the search for the magnetic moment of neutrino (MMN), search for New Physics beyond the SM, and many other

The experimental setup is constructed under reactor unit #3 of KNPP at a distance of 11.1-12.5 m from the center of the 3.1 GW_{th} core under enormous antineutrino flux (4.4-3.6)×10¹³ v/cm²/s at ~50 m w.e. overburden. The low-threshold HP Ge detector with a mass of 1.4 kg is used

obtained with 154 days ON and 39 days OFF statistics. kgd raw data.

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NEWS-G3

47 N

• $\Phi = 60 \text{ cm}$

2

- Aiming for 20 bar pressure
- Candidate detector materials: copper (C10100), stainless steel
- Candidate gas targets: Ar, Ne, Xe

上

First successful detector operation at 10 bar

3

Fe55: 5.9 keV ~ 9000 ADU

Queen's

Expected number of CEvNS events in at various detector-reactor distances for three candidate gas targets

• In all cases, 60cm diameter detector is considered

Complementary cumulative distribution

CCM (Coherent CAPTAIN-Mills)

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51 N

Impactful Neutrino Physics and Dark Sector Searches at LANSCE with the **Coherent CAPTAIN-Mills (CCM) Experiment** (LANL, FNAL, Columbia, MIT, TAMU, UNM, UFlorida, ERAU, UNAM-Mexico, Edinburgh-UK)

- Successfully built and commissioned CCM detector at Lujan 800 MeV stopped pion source (FY2019-FY2022 LDRD, and HEP-DMNI funding). Have begun HEP funded three-year beam run (2023-25) to collect 20E21 Protons on Target (POT).
- Large energy dynamic range (100 keV to 10 GeV) with fast beam and ~nsec detector response can probe unexplored parameter space for dark matter, axions, and meson portal model test of MiniBooNE. Expect to complete beam run end of 2025 with ~20E21 POT.

CCM: 10-ton Liquid Argon (LAr) scintillation detector instrumented with 200 8" Photomultiplier tubes, veto region, shielding, fast electronics.

Recent CCM@LANL Highlights Feb 2024

- CCM has three papers published, and one accepted by PRD (based on 2019 CCM120 data 1.8E21 POT)
- Physical Review Letters Vol. 129, No. 2 (2022), "First Leptophobic Dark Matter Search from Coherent CAPTAIN-Mills" (51 citations).
- *Physical Review D* 106, 1, (2022), "First dark matter search results from Coherent CAPTAIN-Mills" (23 citations).
- *Physical Review D* 107, 9, (2023), "Prospects for Detecting Axionlike Particles at the Coherent CAPTAIN-Mills Experiment" (21 citations).
- arXiv:2309.20599 (accepted by PRD), "Testing Meson Portal Dark Sector Solutions to the MiniBooNE Anomaly at CCM" (1 citation).

- results on ALP, DM, and meson portal model test of MiniBooNE by Fall 2024.
- sector search beam experiments at LANL and/or FNAL with order magnitude improved reach.

CCM200 detector is running well and has collected 9E21 POT beam data from 2021-2023. Expect new preliminary

Expect six to eight impactful dark sector publications and >3 Phd thesis based on final data set (20E21 POT). Continue collaborating with theorists developing new dark sector models to test. Contemplating follow-on dark

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BULLKID

BULLKID phonon detector array

A. Cruciani, et al, Appl. Phys. Lett. 121, 213504 (2022)

Surface background of a single die

Array of silicon phonon-mediated nuclear recoil detectors sensed by KIDs

Multiplexed Kinetic Inductance Detectors (KIDs) convert the phonons from the nuclear recoil into an electronic signal

Acquiring the neighbours of a given die allows to veto events interacting elsewhere

Mass and threshold improvement

SBC (Scintillating Bubble Chamber)

Liquid Noble Bubble Chambers

Objective:

Quasi-background-free detection of subkeV Nuclear Recoils

Signal:

Single bubble with little or no coincident scintillation

Backgrounds: ER's (beta, gamma): No bubbles

NR's (fast neutron): Multiple bubbles Strong coincident scintillation

Depends on NR threshold and target fluid:

- Freon-based chambers ER-blind @ ~3 keV
- Liquid-noble chambers ER-blind @ < 500 eV,
 - (target 100 eV)

Liquid Noble Bubble Chambers

Objectives for SBC-LAr10 in MINOS:

- **Demonstrate operation** of physics-scale liquid-noble bubble chamber
- **Determine maximum** superheat for ER-blind operation
- **Calibrate Threshold** for NR detection, @ 100 eV, with 10 eV resolution

Top-view from image acquisition system LED reflections-(less pronounced in liquid-filled chamber)

Silica Vessel for LAr target (gas Ar in this image) Hammamtsu VUV4 SiPMs

Calibration Strategies (for the ER-blind)

Photo-neutron Sources (> 500 eV recoils)

Thermal neutron Capture (200 – 500 eV recoils)

Photon-nucleus Scattering (< 300 eV recoils)

