

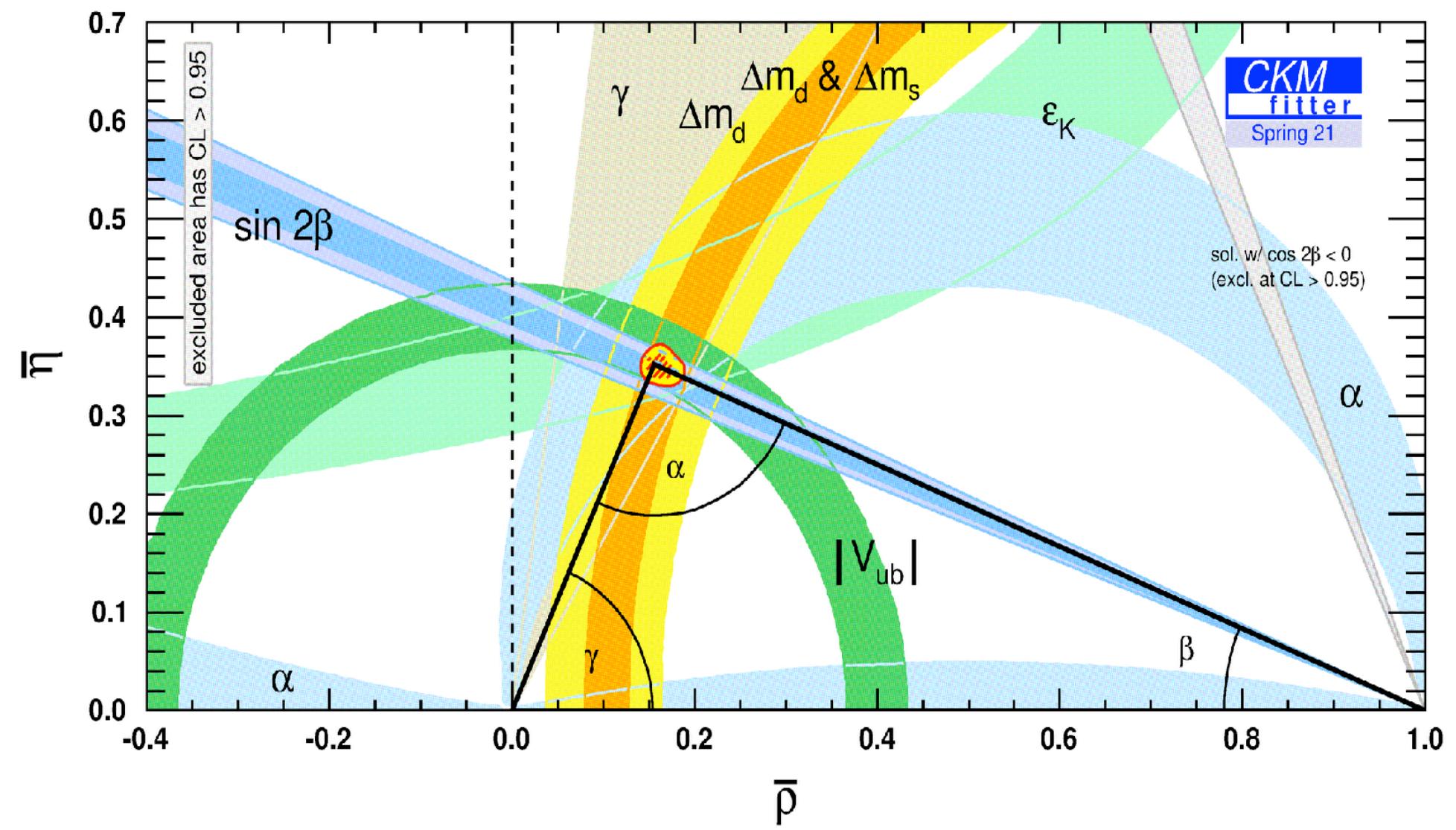
# Impact of Neutrino Interaction Uncertainties on BSM Searches

Joachim Kopp (CERN & JGU Mainz)  
Nulnt 2024 • São Paulo • April 2024

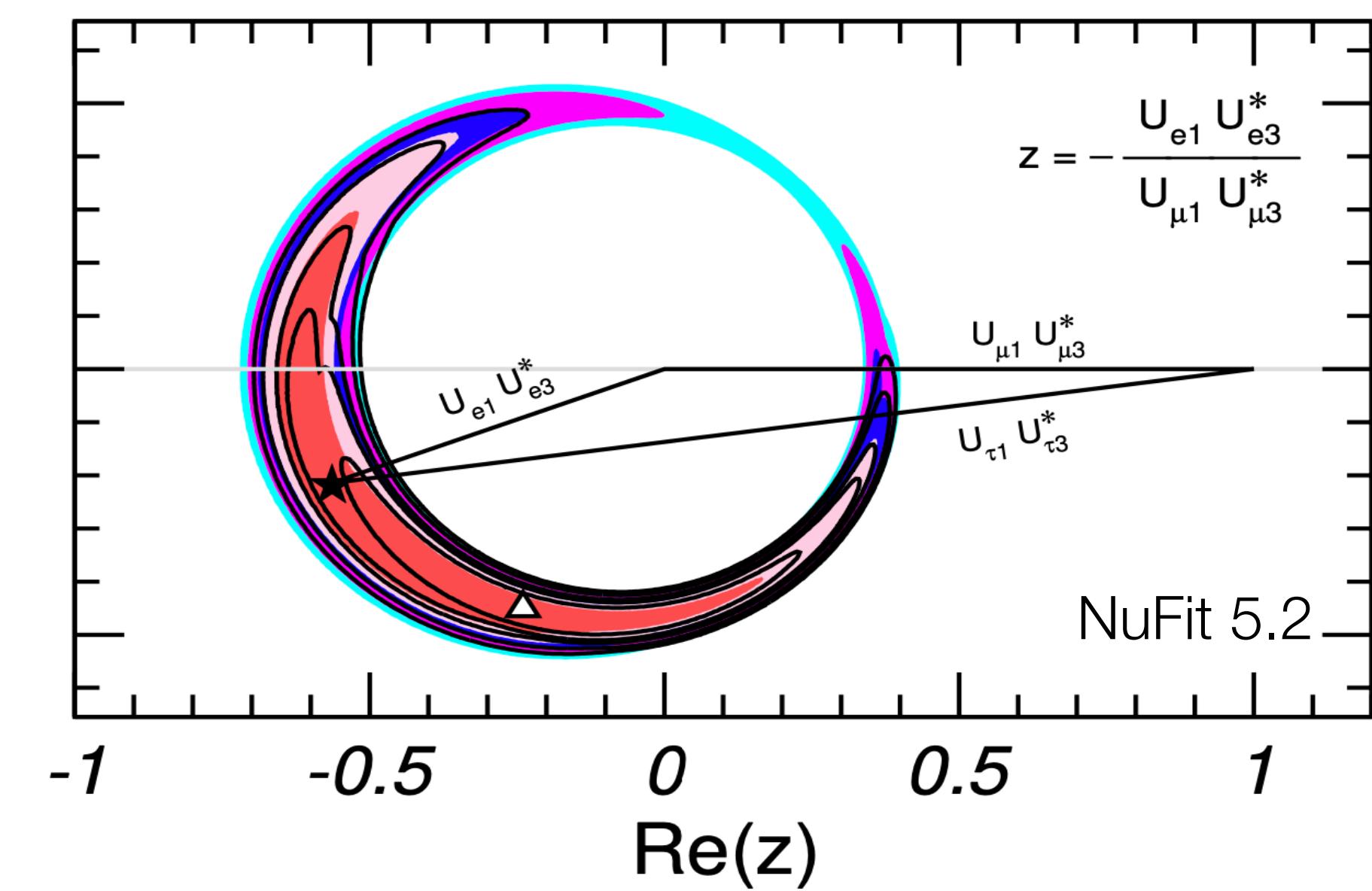


# Precision Neutrino Physics

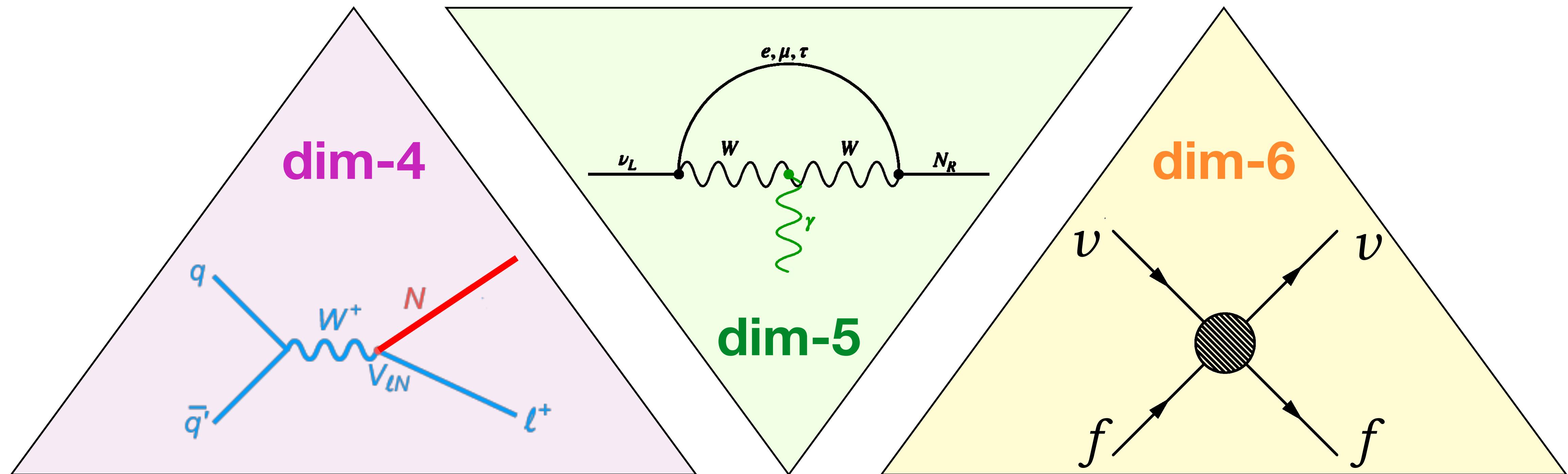
## Quarks



## Leptons

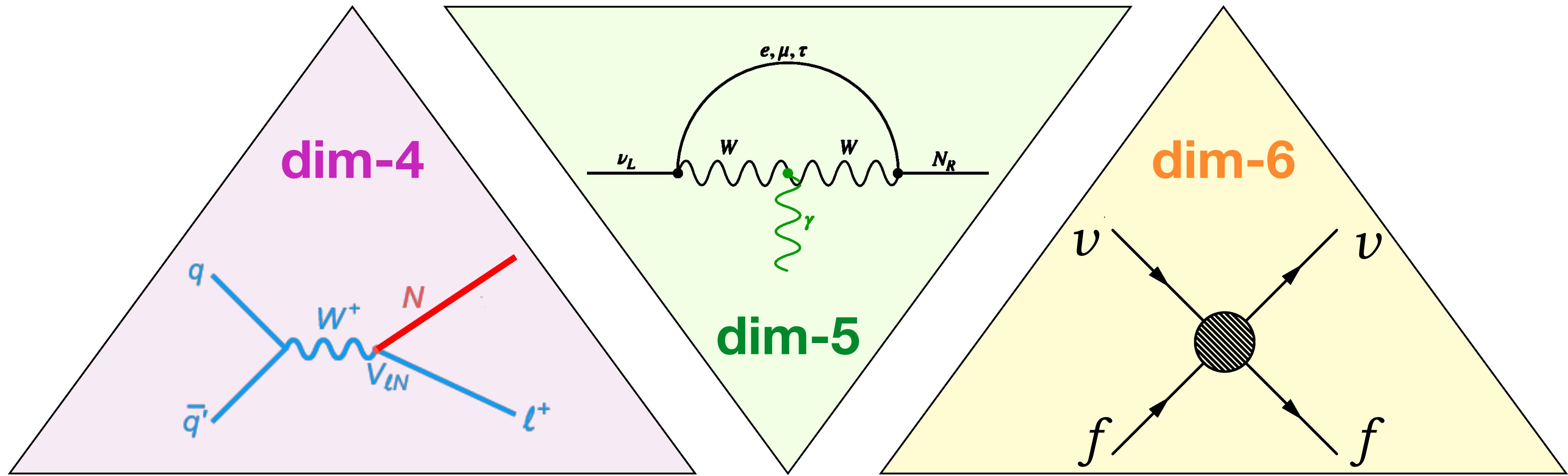


# Neutrino Physics Beyond the Standard Model



# Neutrino Physics Beyond the Standard Model

e.g. neutrino magnetic moments



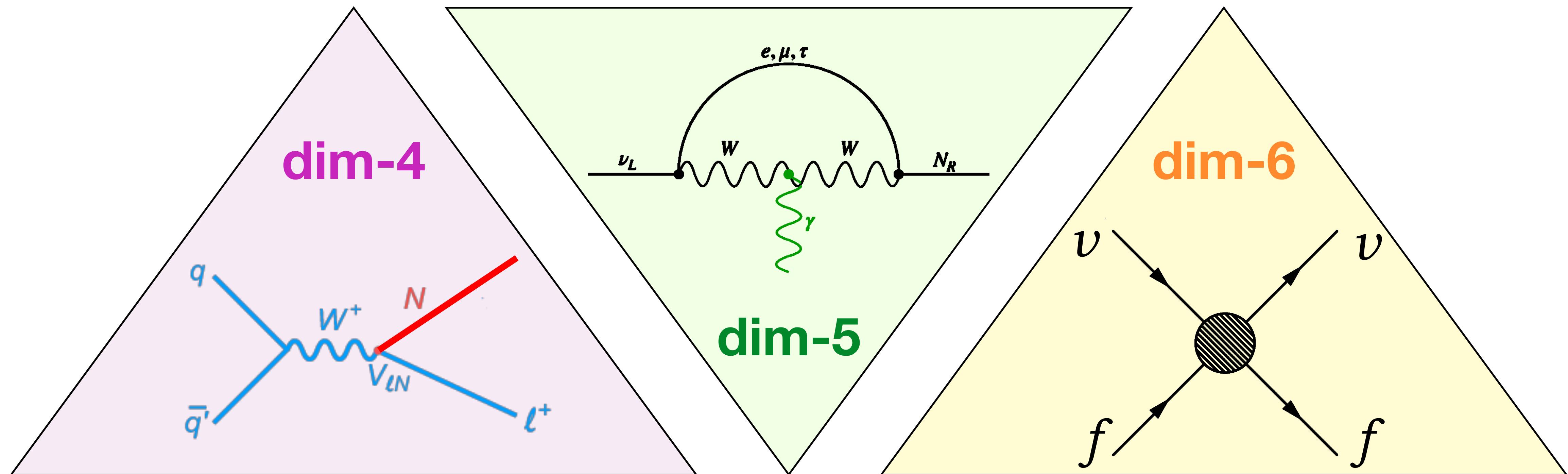
e.g. sterile neutrinos

e.g. non-standard interactions

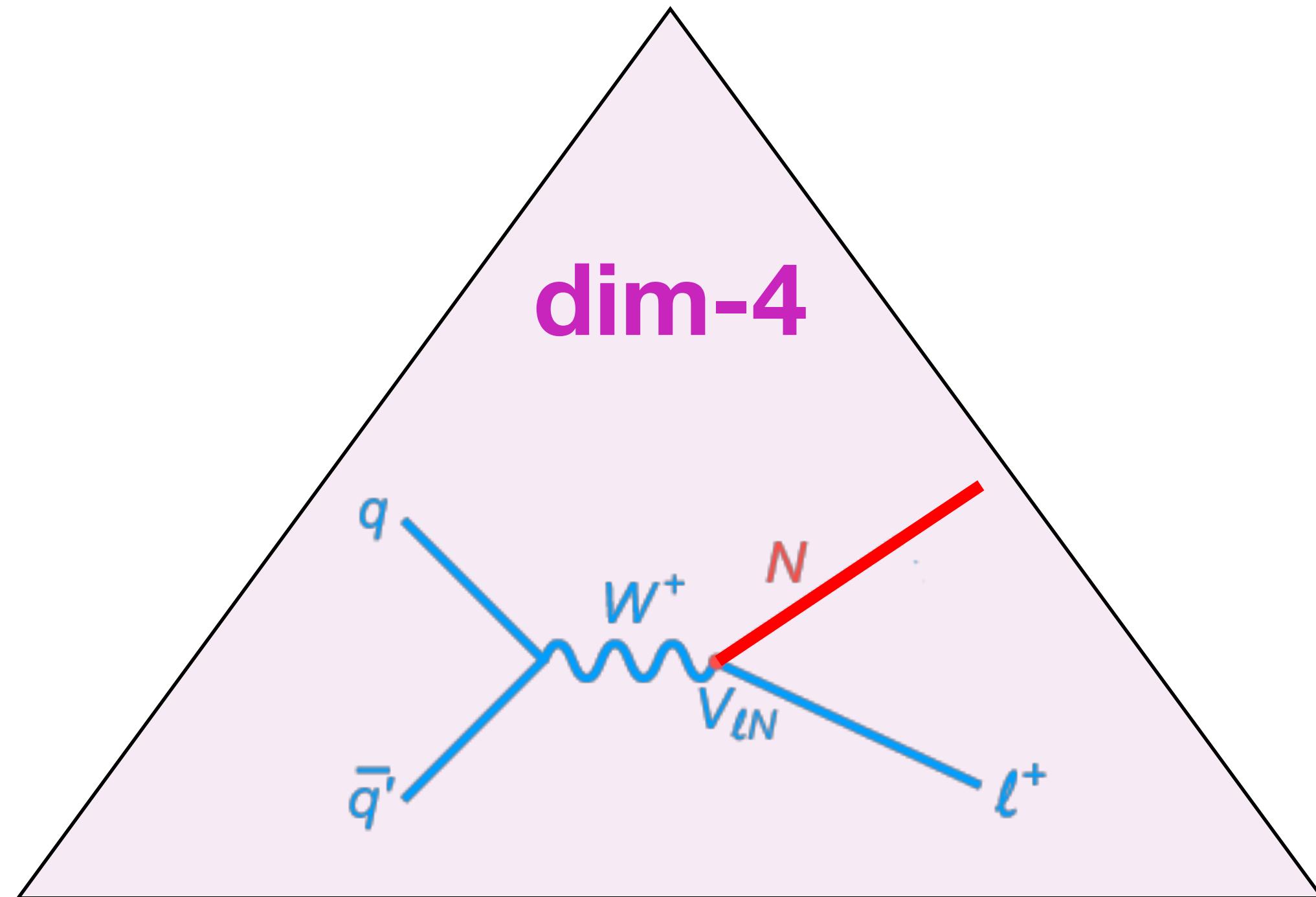
“With great precision comes great responsibility.”

*Tim Linden, WIN 2021*

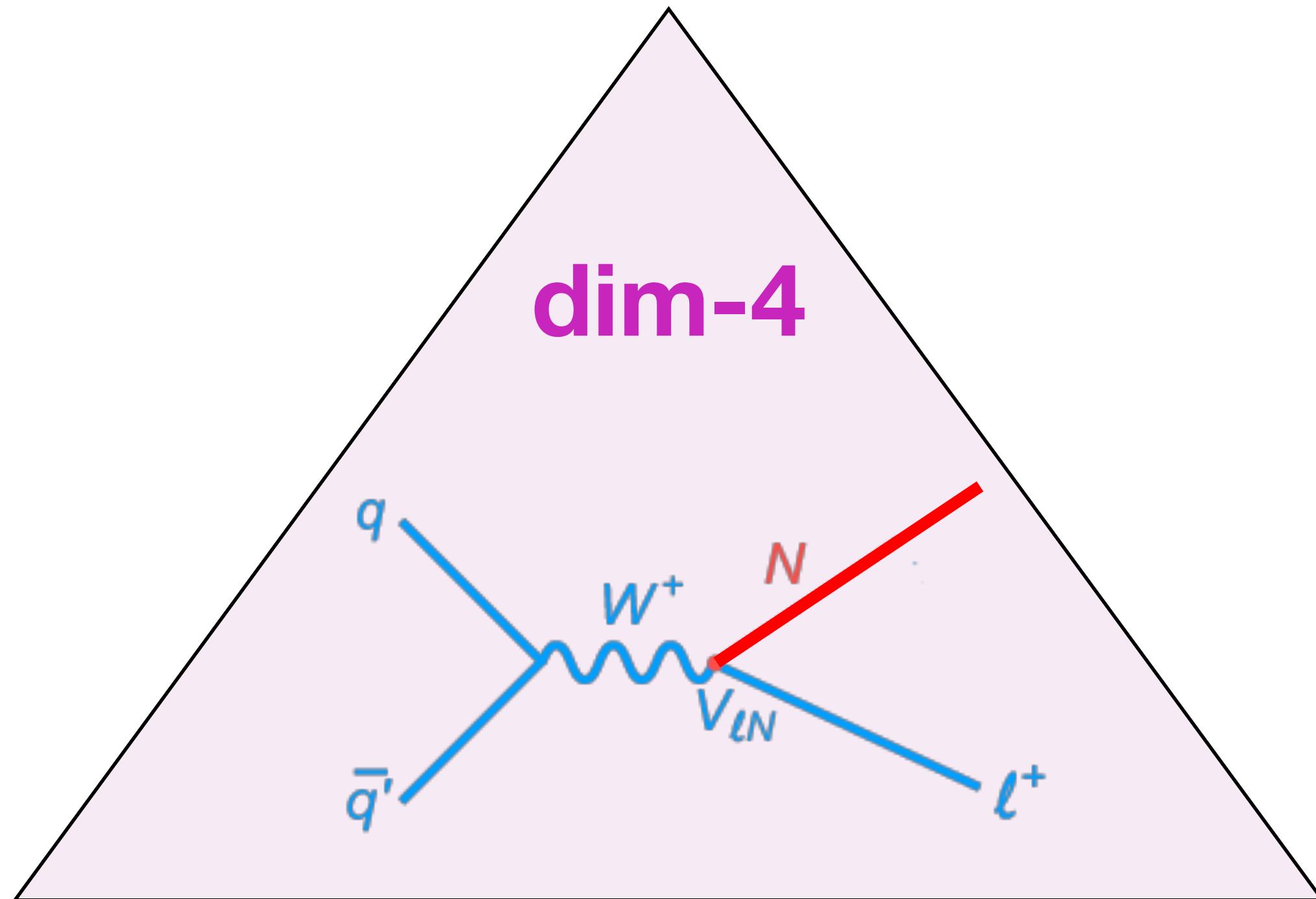
# Neutrino Physics Beyond the Standard Model



# Sterile Neutrinos = new, uncharged fermions

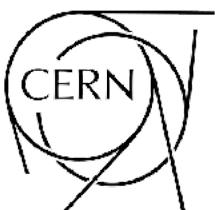


# Sterile Neutrinos = new, uncharged fermions

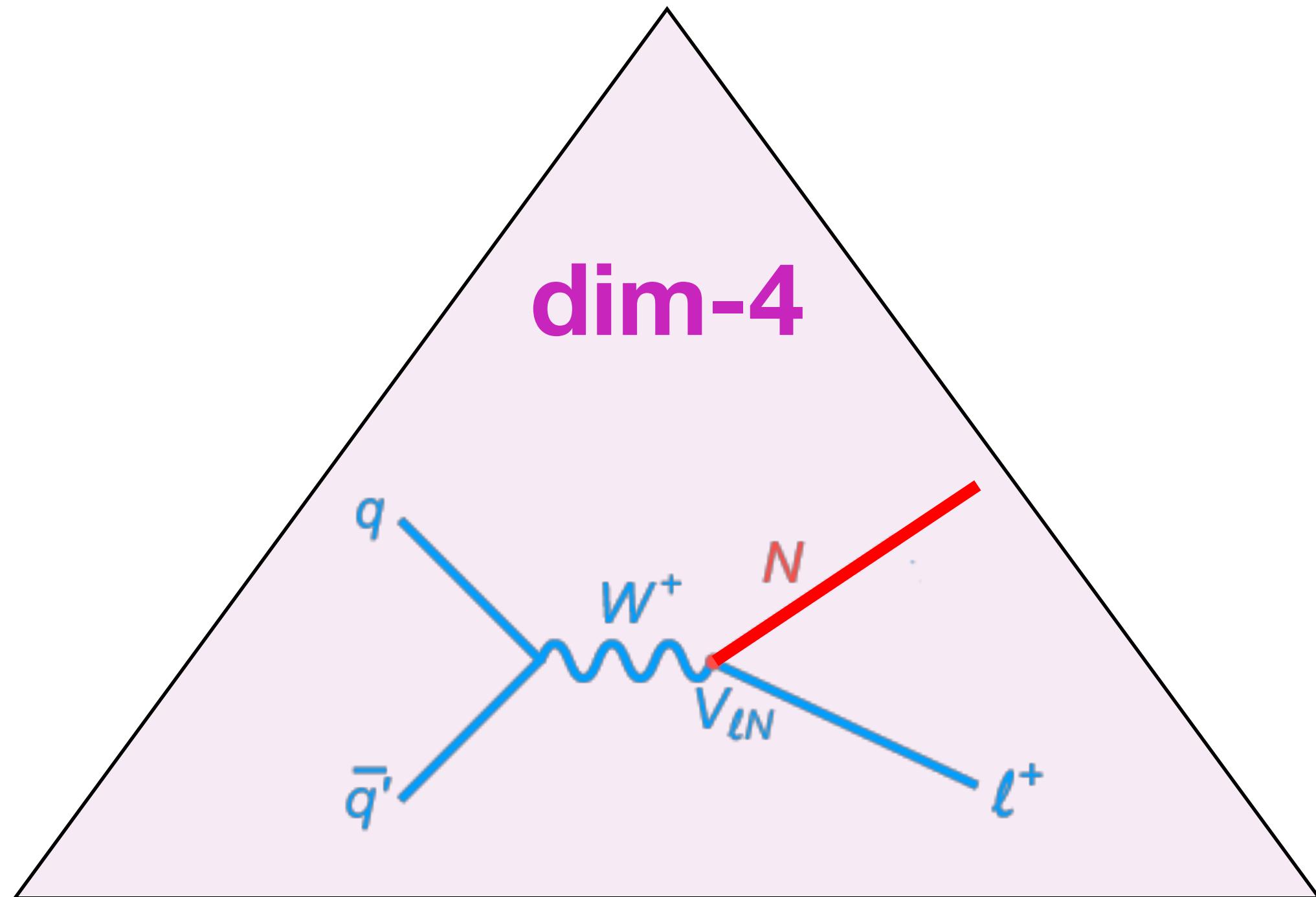


## Standard Model of Elementary Particles

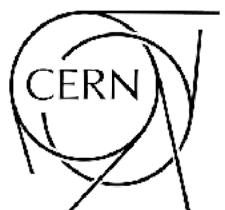
three generations of matter (fermions)				
	I	II	III	
mass	$\approx 2.4 \text{ MeV}/c^2$	$\approx 1.275 \text{ GeV}/c^2$	$\approx 172.44 \text{ GeV}/c^2$	
charge	2/3	2/3	2/3	0
spin	1/2	1/2	1/2	1
	<b>u</b> up	<b>c</b> charm	<b>t</b> top	<b>g</b> gluon
				<b>H</b> Higgs
QUARKS				
mass	$\approx 4.8 \text{ MeV}/c^2$	$\approx 95 \text{ MeV}/c^2$	$\approx 4.18 \text{ GeV}/c^2$	
charge	-1/3	-1/3	-1/3	0
spin	1/2	1/2	1/2	1
	<b>d</b> down	<b>s</b> strange	<b>b</b> bottom	<b><math>\gamma</math></b> photon
LEPTONS				
mass	$\approx 0.511 \text{ MeV}/c^2$	$\approx 105.67 \text{ MeV}/c^2$	$\approx 1.7768 \text{ GeV}/c^2$	
charge	-1	-1	-1	0
spin	1/2	1/2	1/2	1
	<b>e</b> electron	<b><math>\mu</math></b> muon	<b><math>\tau</math></b> tau	<b>Z</b> Z boson
				<b>W</b> W boson
SCALAR BOSONS				
mass	$< 2.2 \text{ eV}/c^2$	$< 1.7 \text{ MeV}/c^2$	$< 15.5 \text{ MeV}/c^2$	
charge	0	0	0	±1
spin	1/2	1/2	1/2	1
	<b><math>\nu_e</math></b> electron neutrino	<b><math>\nu_\mu</math></b> muon neutrino	<b><math>\nu_\tau</math></b> tau neutrino	



# Sterile Neutrinos = new, uncharged fermions

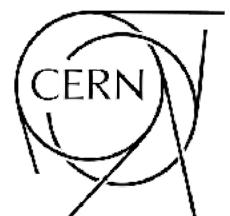
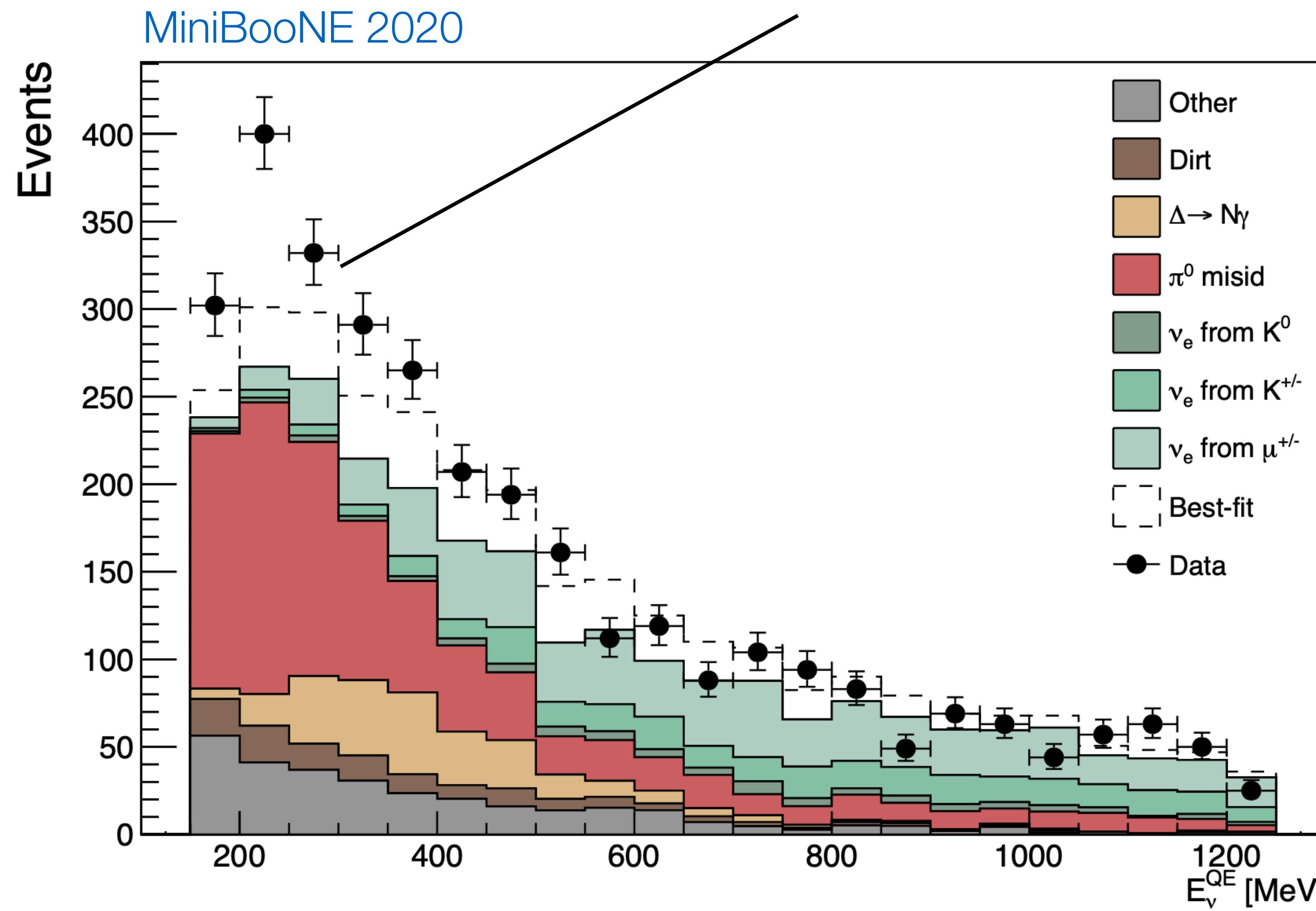


Standard Model of Elementary Particles				
three generations of matter (fermions)				
	I	II	III	
mass	$\approx 2.4 \text{ MeV}/c^2$	$\approx 1.275 \text{ GeV}/c^2$	$\approx 172.44 \text{ GeV}/c^2$	
charge	$2/3$	$2/3$	$2/3$	
spin	$1/2$	$1/2$	$1/2$	
	<b>u</b> up	<b>c</b> charm	<b>t</b> top	<b>g</b> gluon
	<b>d</b> down	<b>s</b> strange	<b>b</b> bottom	<b>H</b> Higgs
	<b>e</b> electron	<b><math>\mu</math></b> muon	<b><math>\tau</math></b> tau	<b><math>\gamma</math></b> photon
	<b><math>\nu_S</math></b> sterile neutrino	<b><math>\nu_e</math></b> electron neutrino	<b><math>\nu_\mu</math></b> muon neutrino	<b><math>\nu_\tau</math></b> tau neutrino
	$<2.2 \text{ eV}/c^2$	$<1.7 \text{ MeV}/c^2$	$<15.5 \text{ MeV}/c^2$	$\approx 80.39 \text{ GeV}/c^2$
	0	0	0	$\approx 91.19 \text{ GeV}/c^2$
	$1/2$	$1/2$	$1/2$	$1$
LEPTONS				
SCALAR BOSONS				
GAUGE BOSONS				



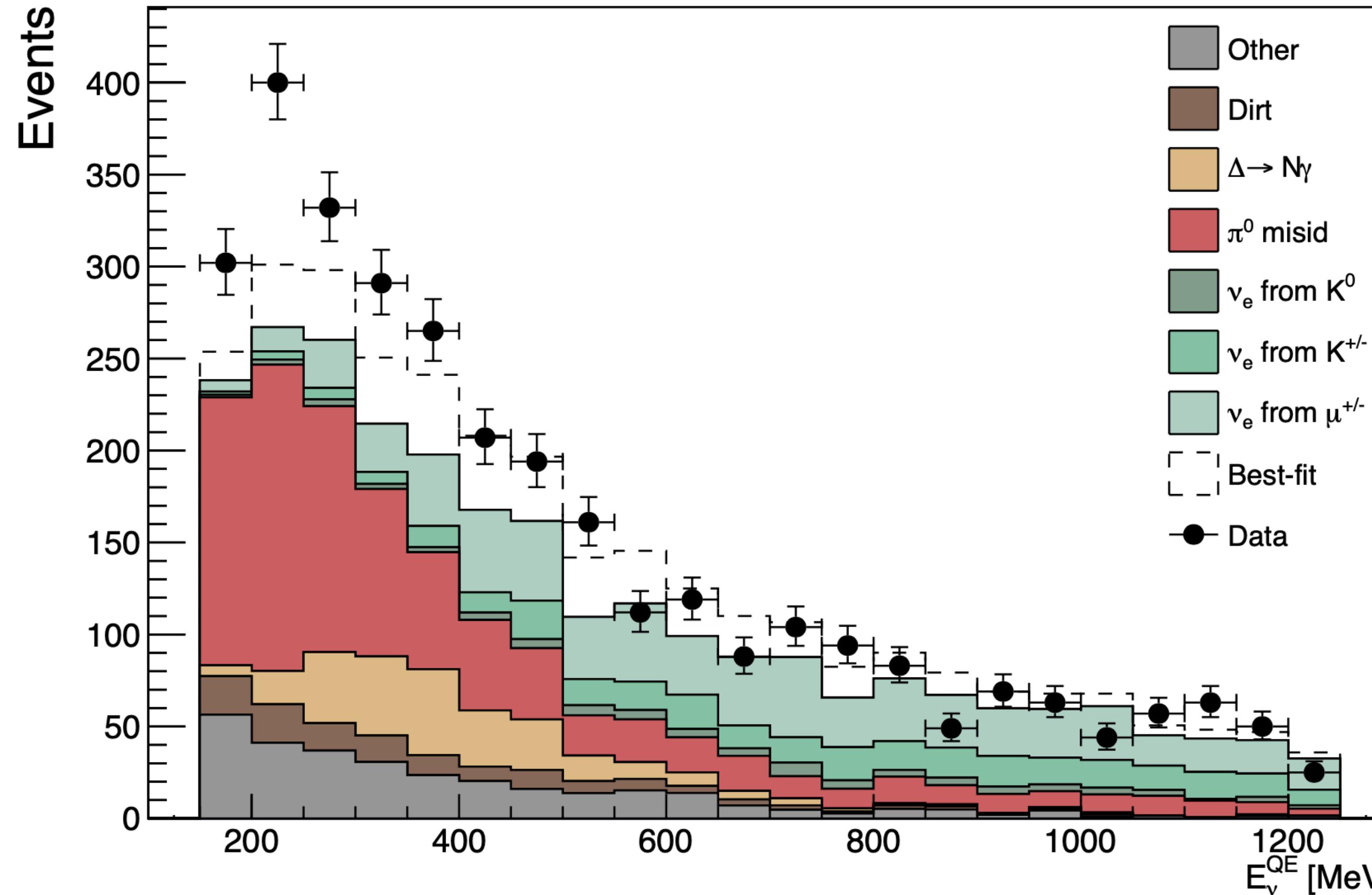
# MiniBooNE

4.8  $\sigma$  excess of  $\nu_e$  in a  $\nu_\mu$  beam

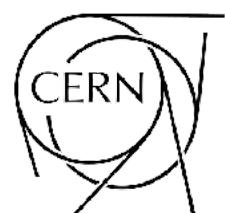


# MiniBooNE

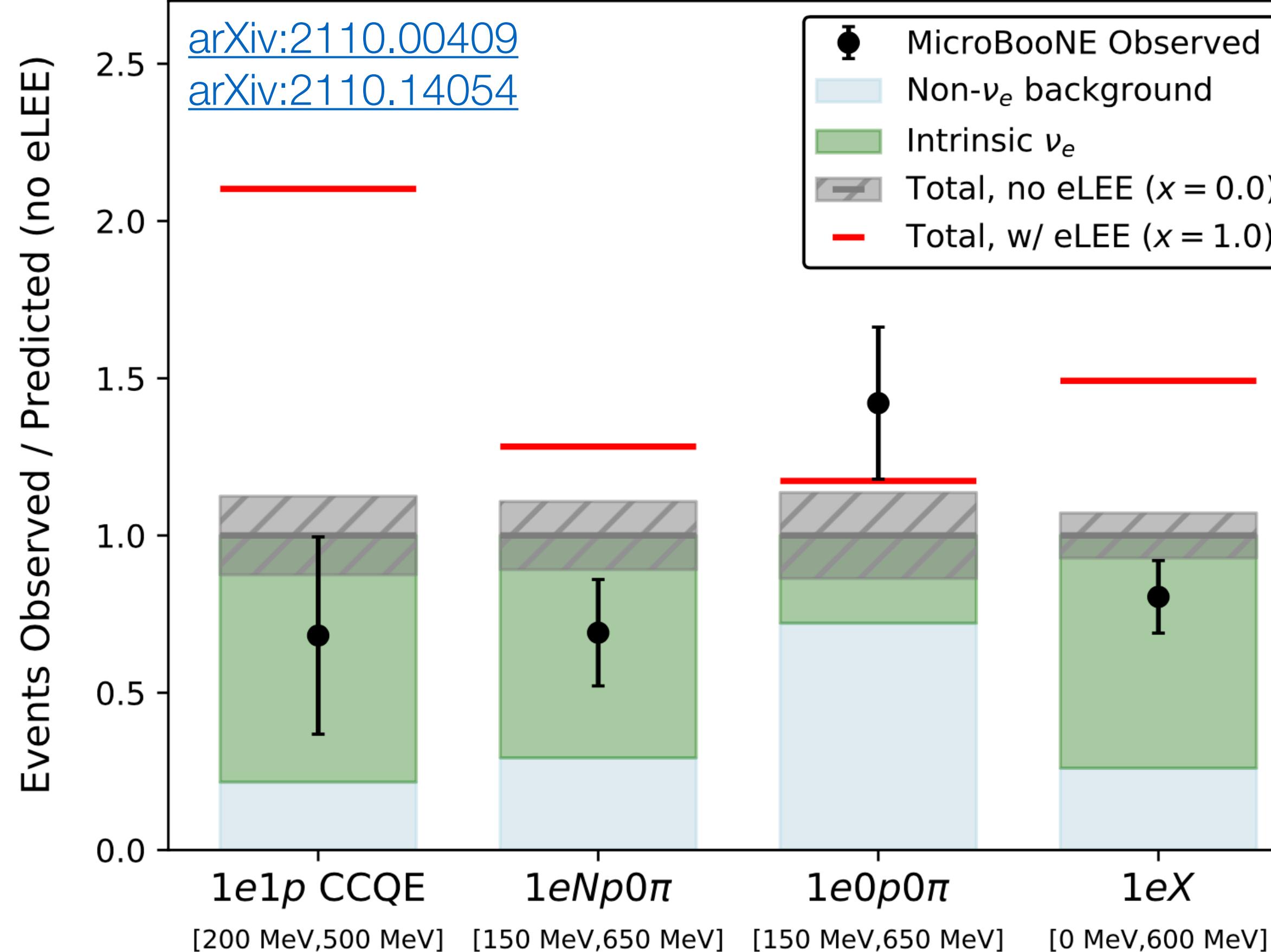
MiniBooNE 2020



- baseline too short for std. oscillations
- but could be explained by **eV-scale sterile neutrino** (“ $\nu_\mu \rightarrow \nu_s \rightarrow \nu_e$ ”)
- ~ consistent with other anomalies
- but inconsistent with null searches

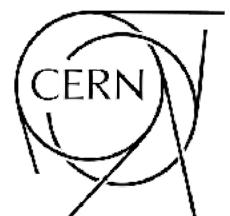
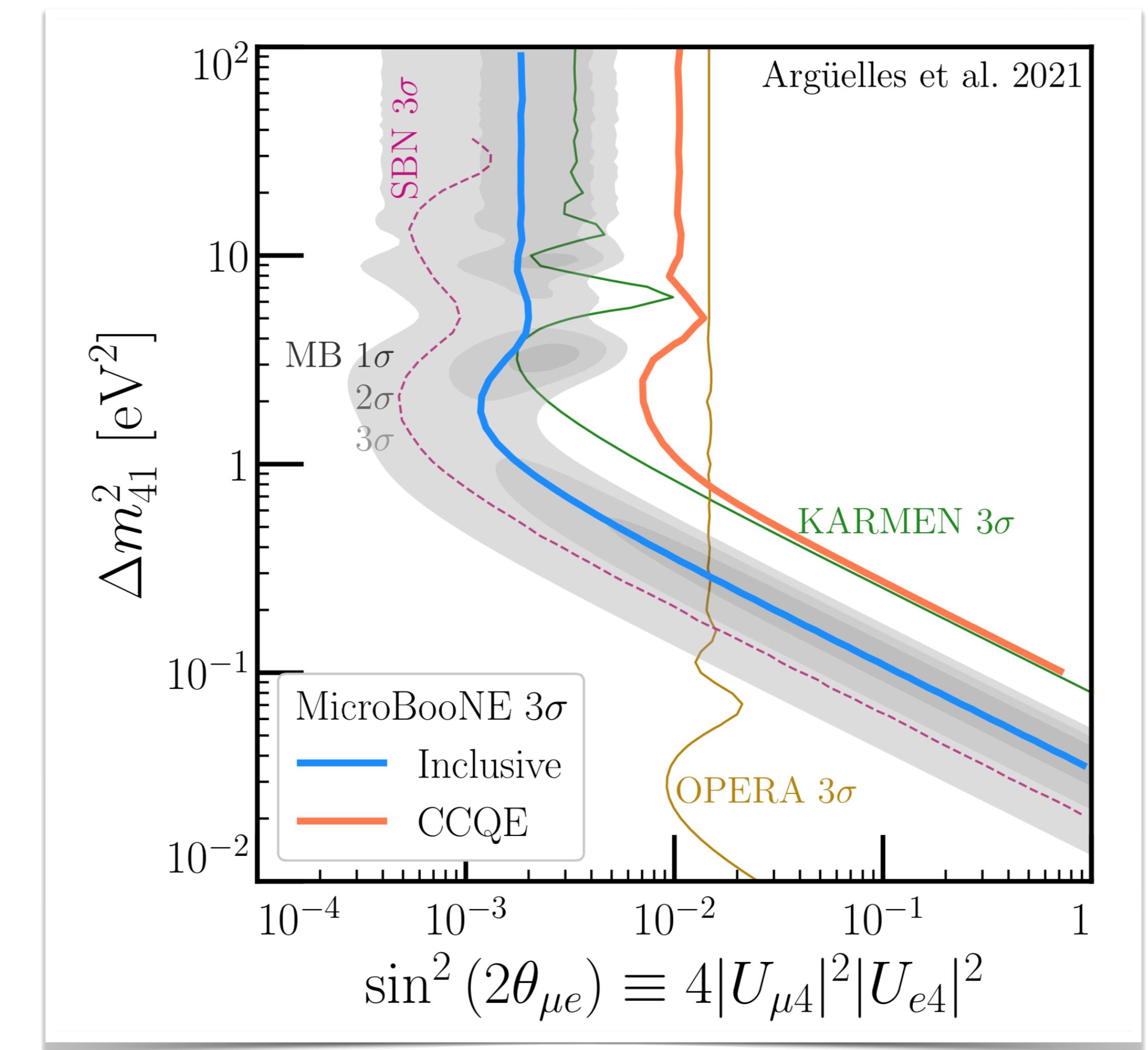
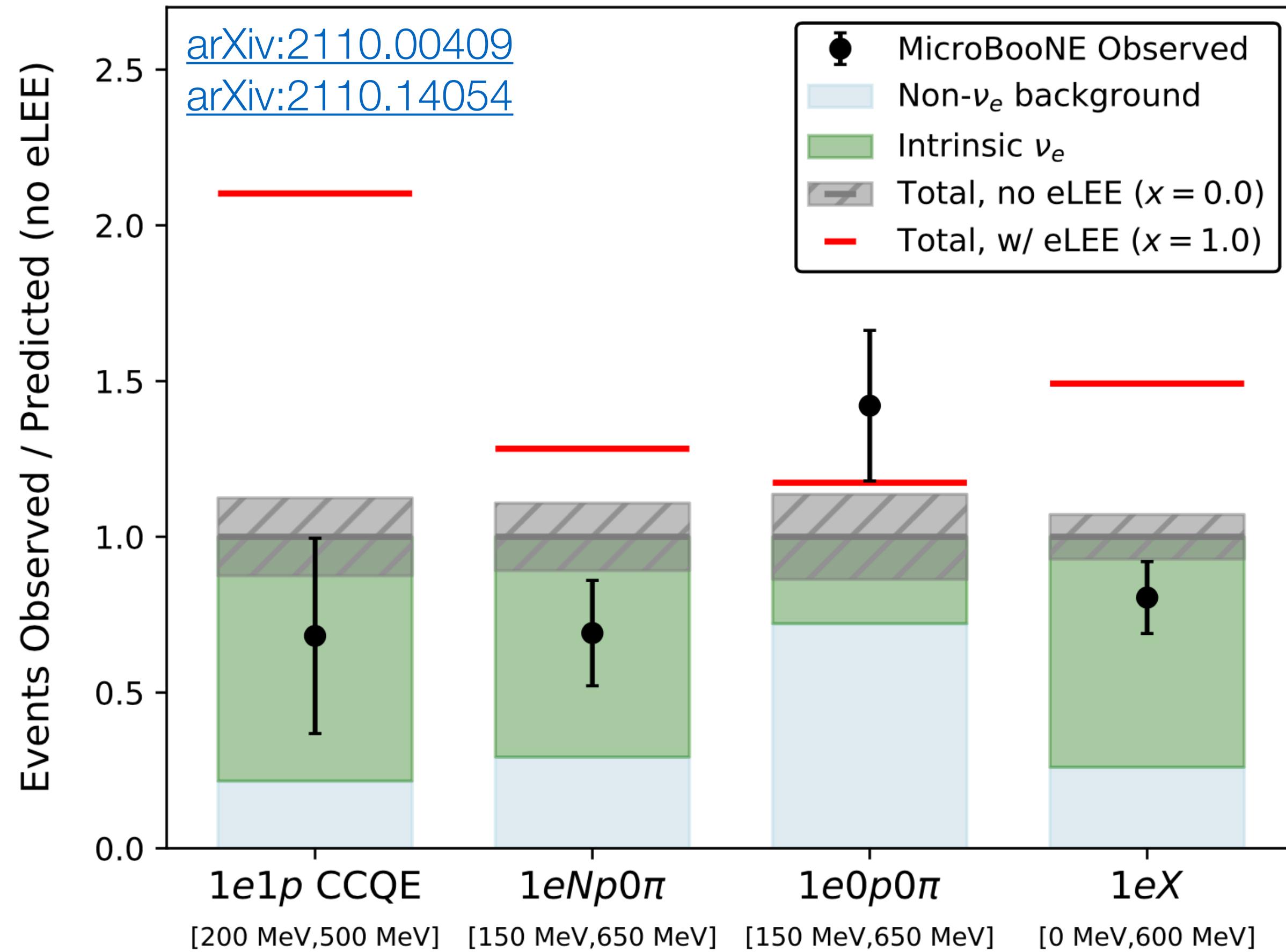


# MicroBooNE



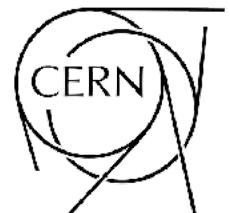
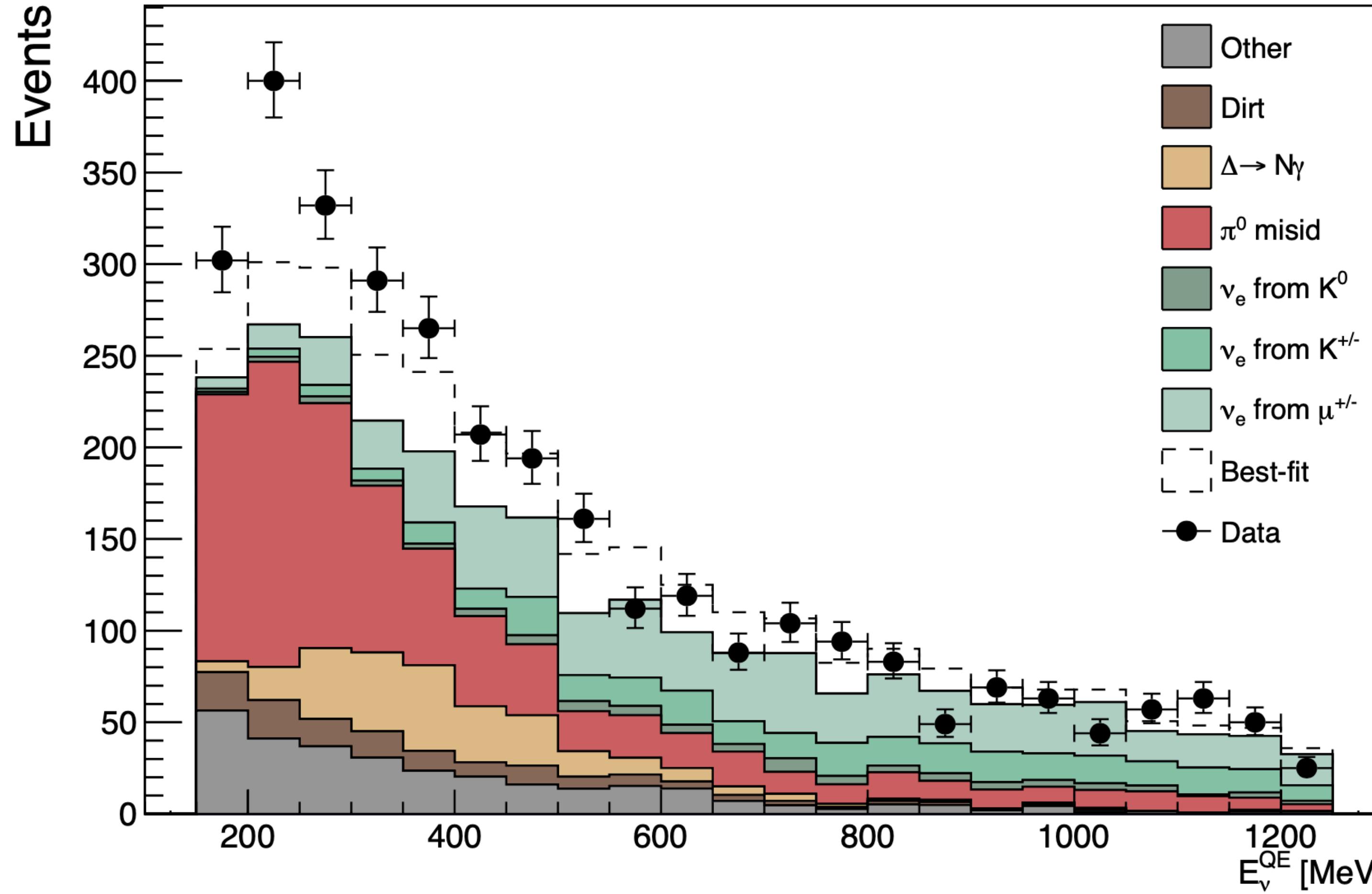
- LAr TPC → superior event reconstruction
- no excess seen so far  
(but still consistent with MiniBooNE)

# MicroBooNE



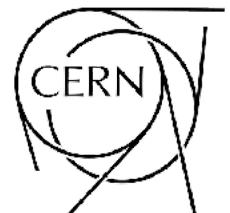
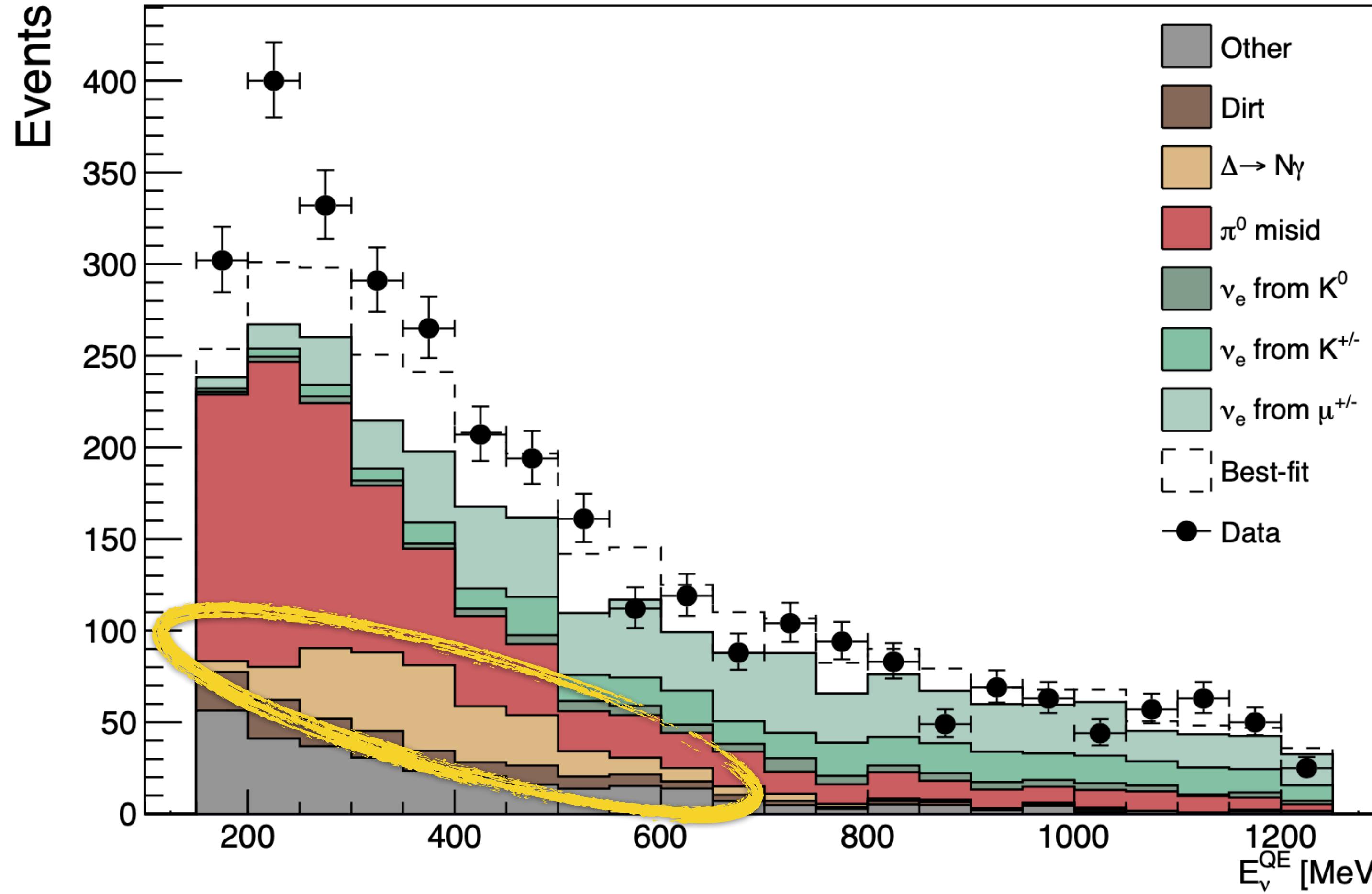
# MiniBooNE

MiniBooNE 2020

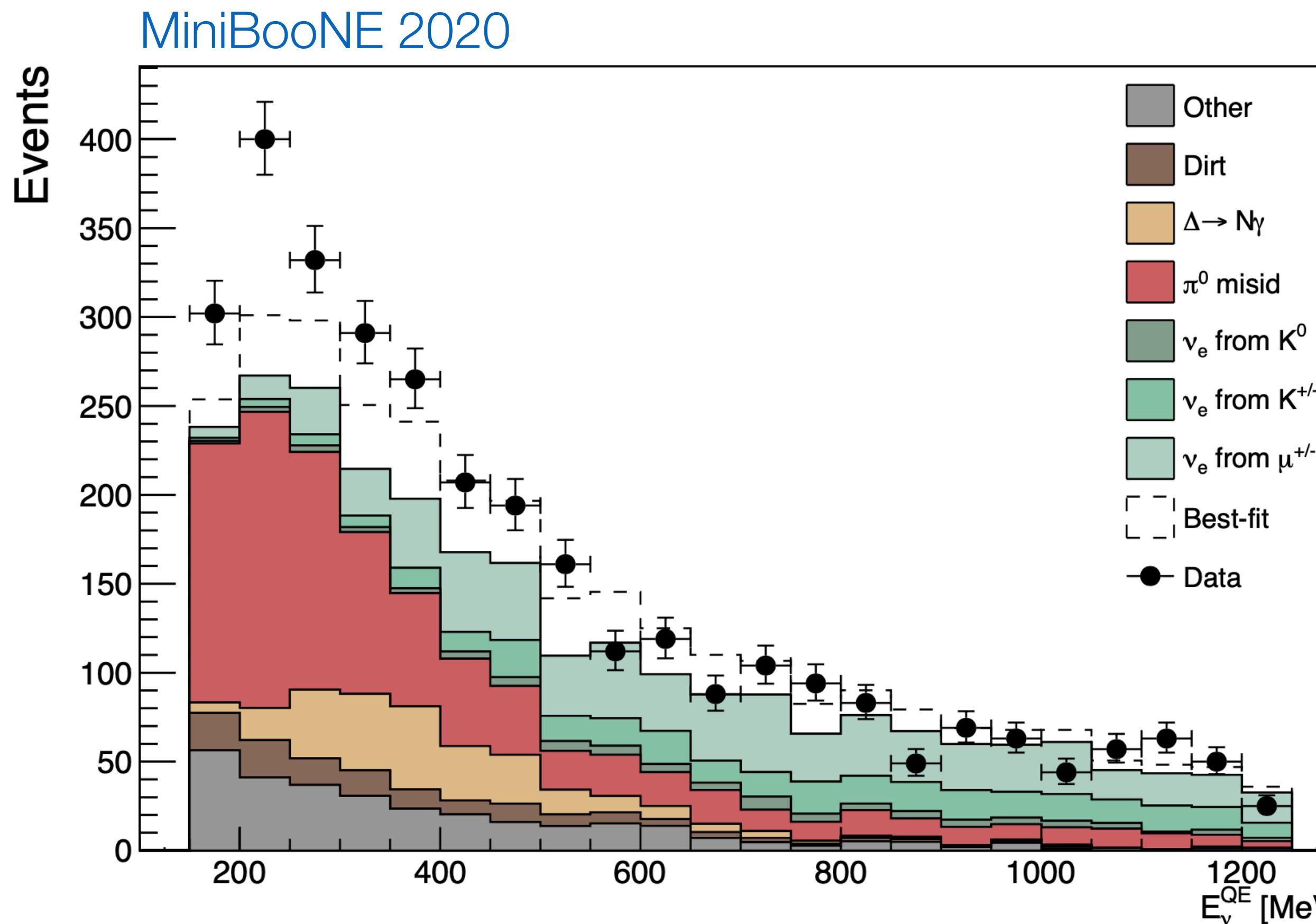


# MiniBooNE

MiniBooNE 2020

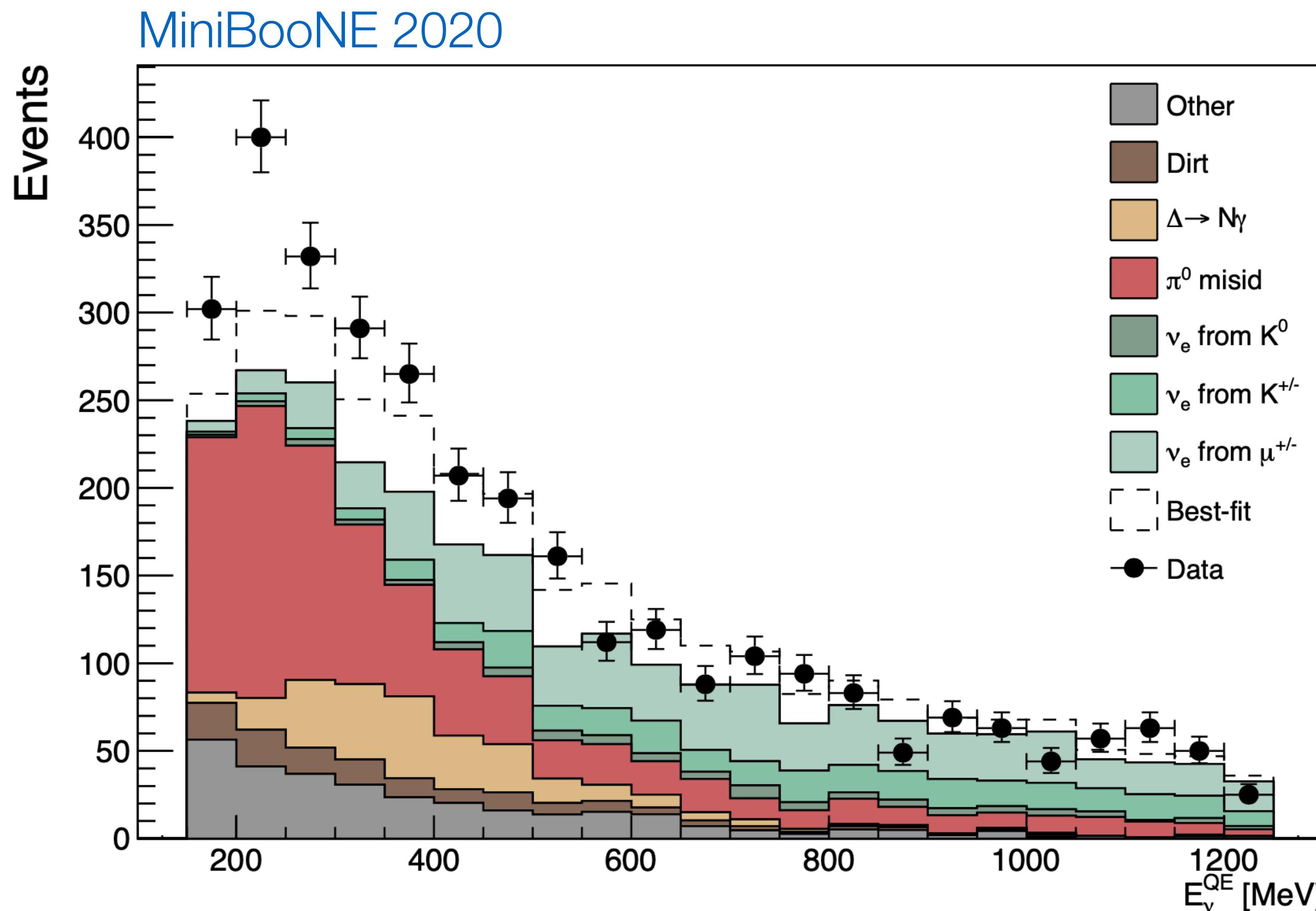


# $\Delta \rightarrow N\gamma$ Background



- NC interaction:  
 $\nu + N \rightarrow \nu + \Delta(1232)$
- Most  $\Delta(1232)$  decay to  $\pi + N$
- But rare decay exists to  $\gamma + N$
- MiniBooNE cannot distinguish single- $\gamma$  background from CC  $\nu_e$  signal

# $\Delta \rightarrow N\gamma$ Background



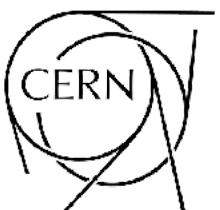
- $\Delta$  production rate can be estimated from  $\Delta \rightarrow \pi N$
- Pions may be **absorbed** on their way out of the nucleus
  - may **excite another  $\Delta(1232)$** 
    - ⇒  $\Delta \rightarrow \gamma N$  enhanced
  - or may be **absorbed**
    - ⇒ control region suppressed

Ioannisian [1909.08571](#)

Giunti Ioannisian Ranucci [1912.01524](#)

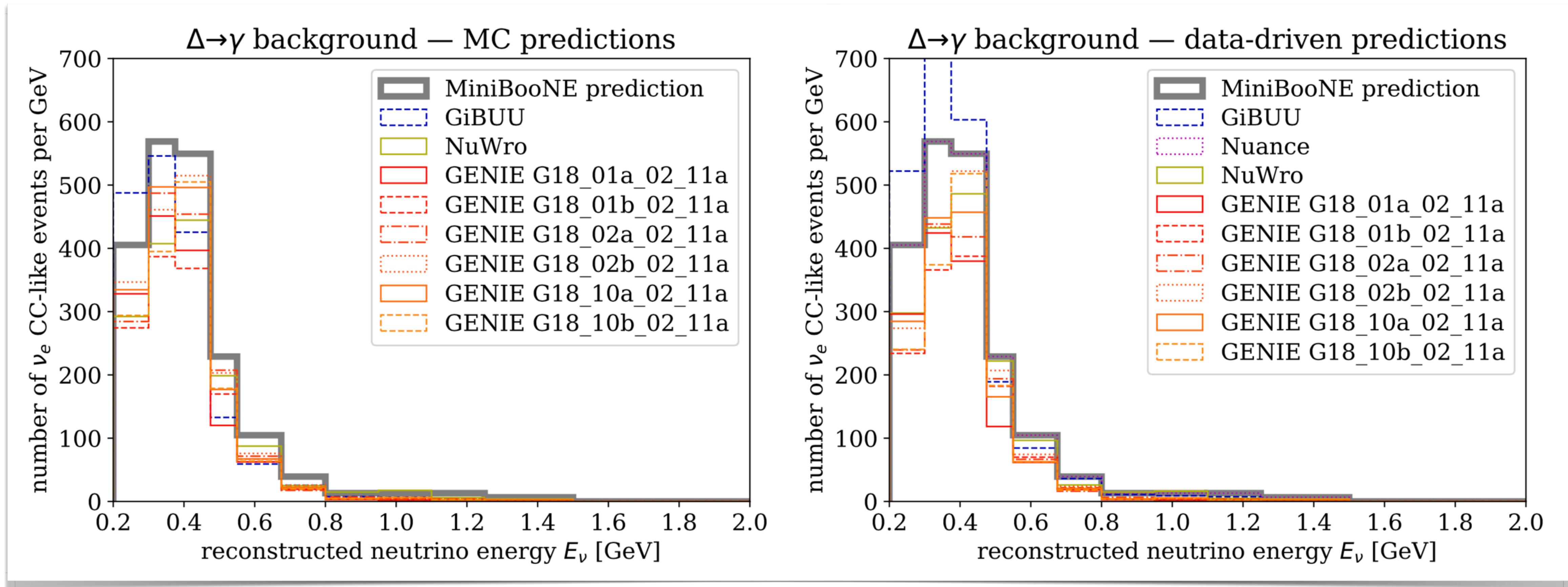
(These effects **have been**  
**taken into account** by MiniBooNE)

MiniBooNe, [arXiv:2006.16883](#)

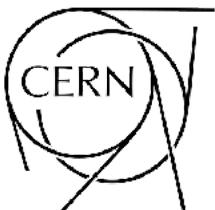


# $\Delta \rightarrow N\gamma$ – Comparison of Generators

Brdar JK, arXiv:2109.08157



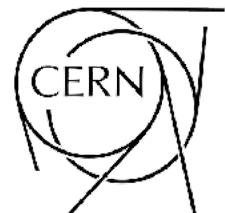
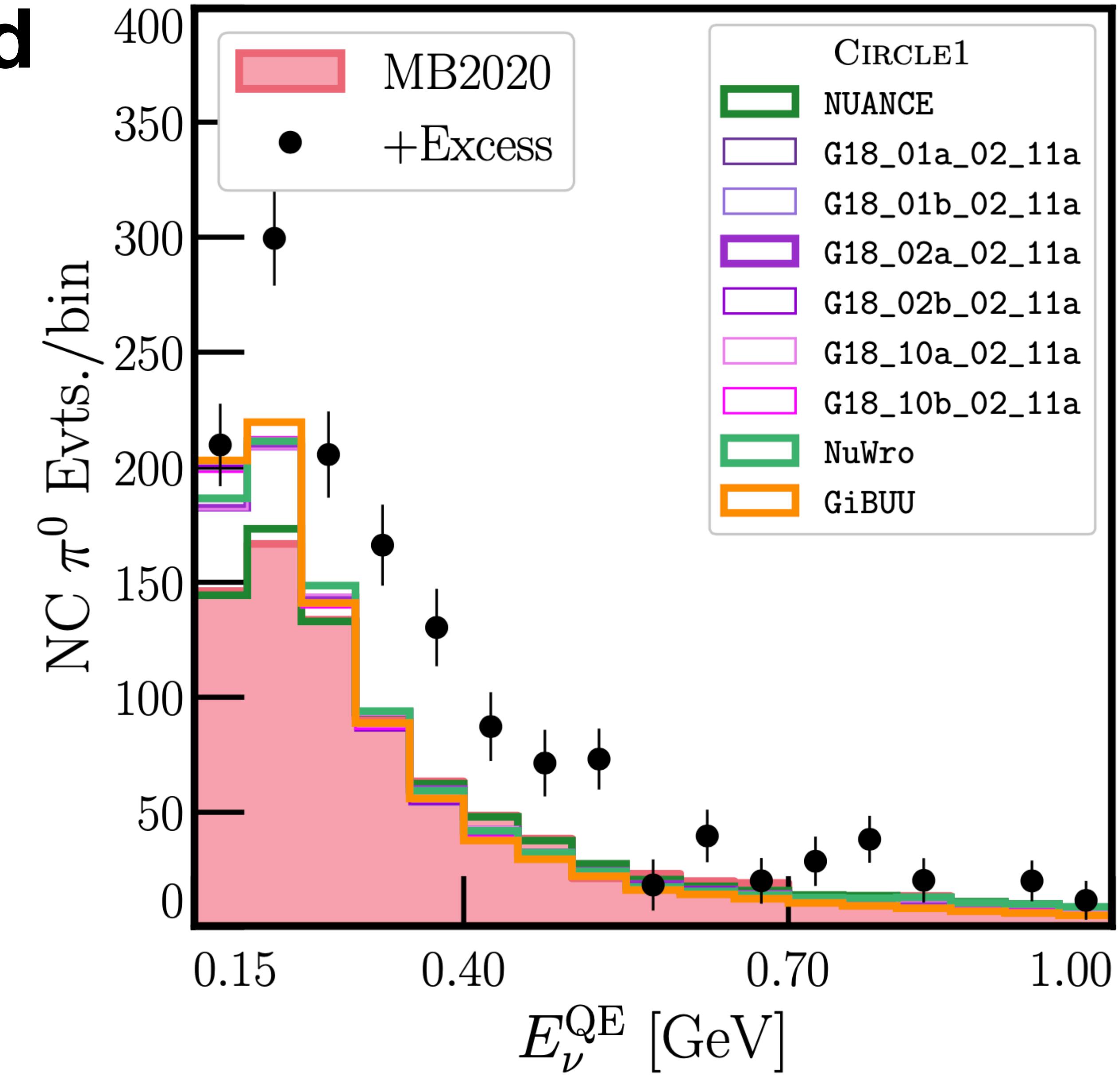
- NUANCE prediction matched to MB prediction, others rescaled in the same way
- using our **own implementation of radiative resonance decays** in GiBUU, NuWro, NUANCE



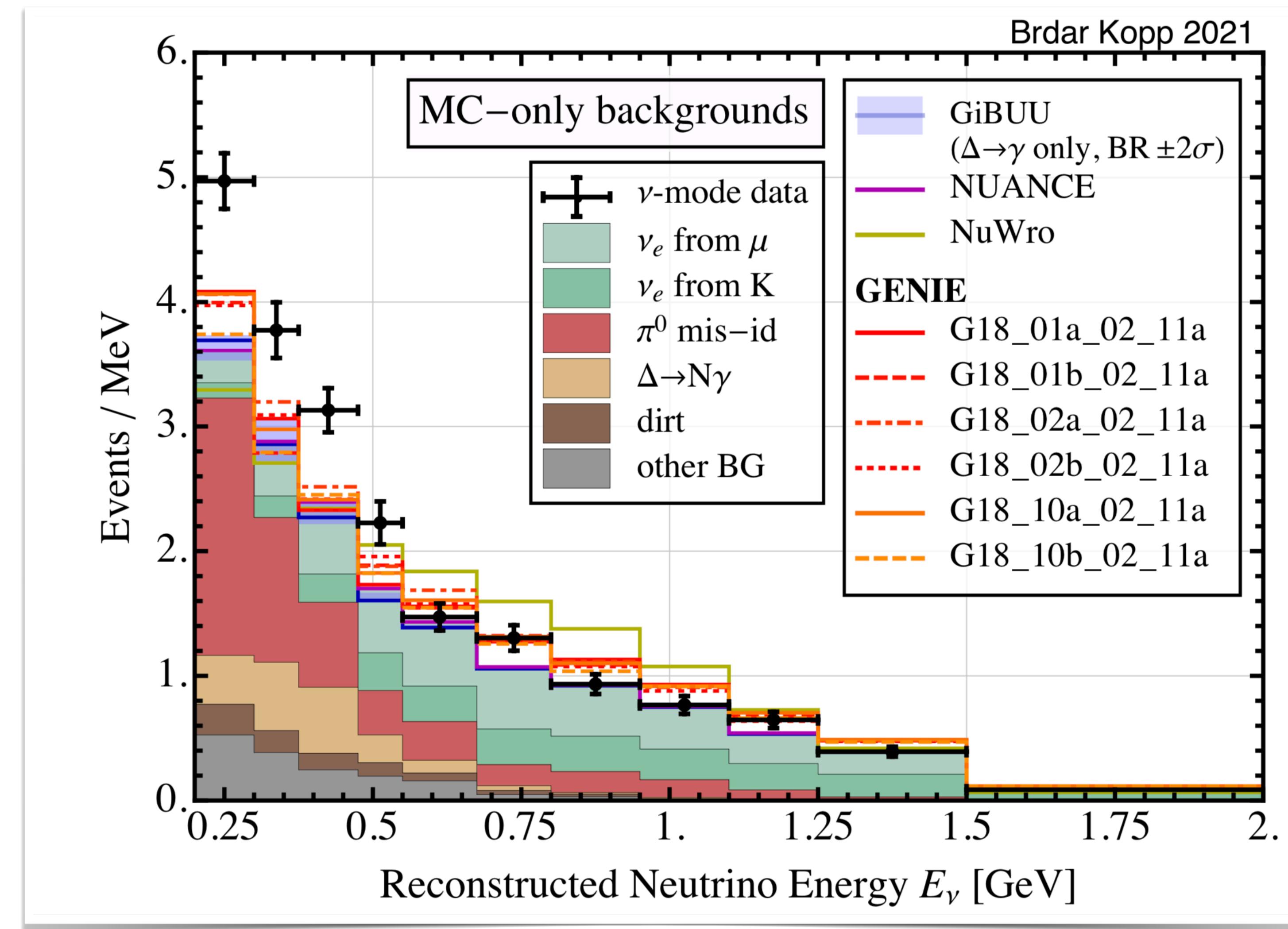
# MiniBooNE $\pi^0$ Background

- Attempt to reproduce MiniBooNE's  $\pi^0$  mis-ID probability
- using our own implementation of radiative resonance decays in GiBUU, NuWro, NUANCE

Kelly JK, arXiv:2210.08021

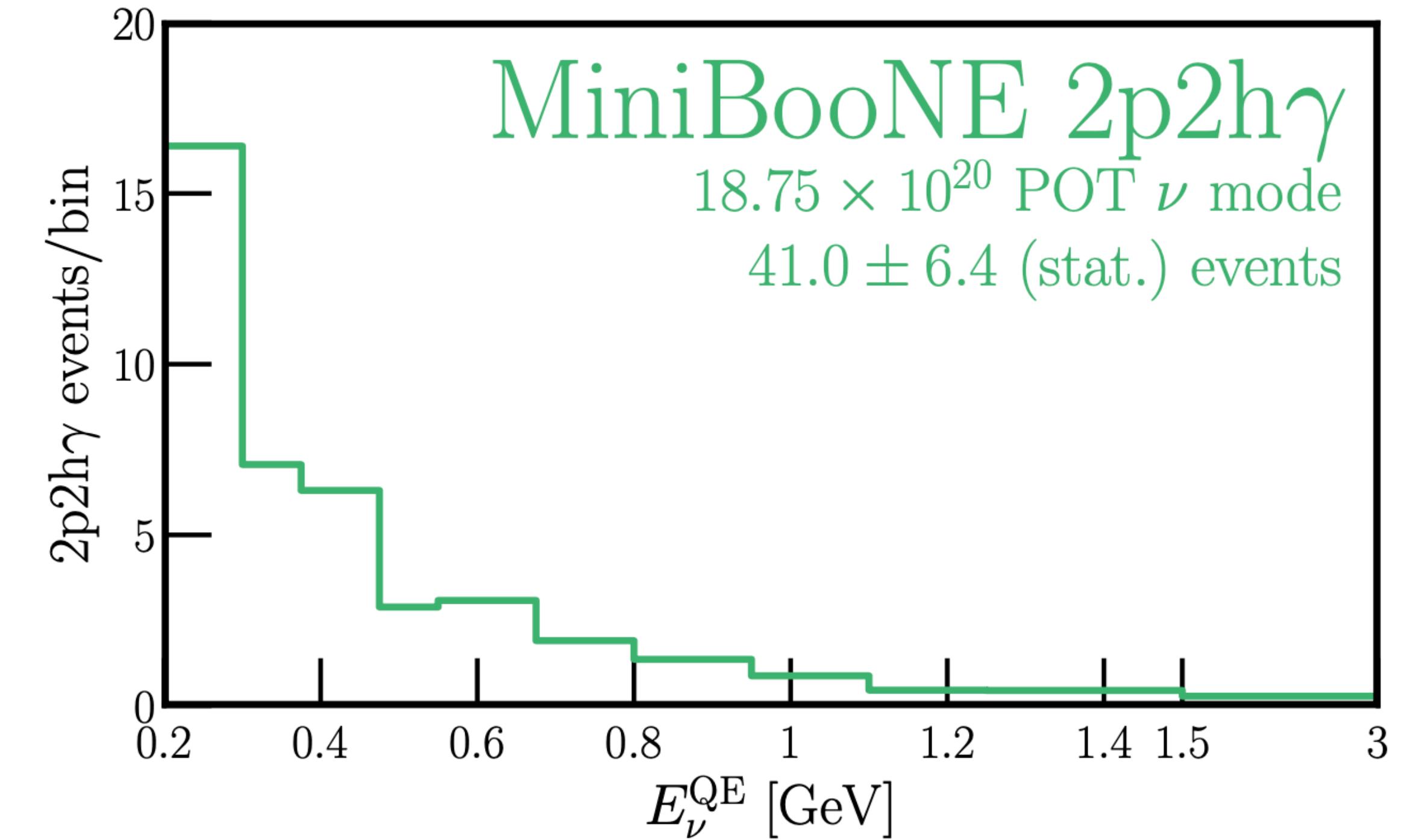


# All Backgrounds – Comparison of Generators



# New Processes? – 2p2h+ $\gamma$

- intuition: neutrino scatters on 2-nucleon system
  - ➡ radiated  $\gamma$  sees both nucleons coherently
  - ➡ enhanced by factor 2
- very naïve implementation  
(elastic scattering on 2-nucleon system)
- event rate falls far short of explaining the anomaly



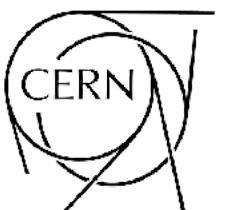
Kelly JK, arXiv:2210.08021

# My Bet for MiniBooNE

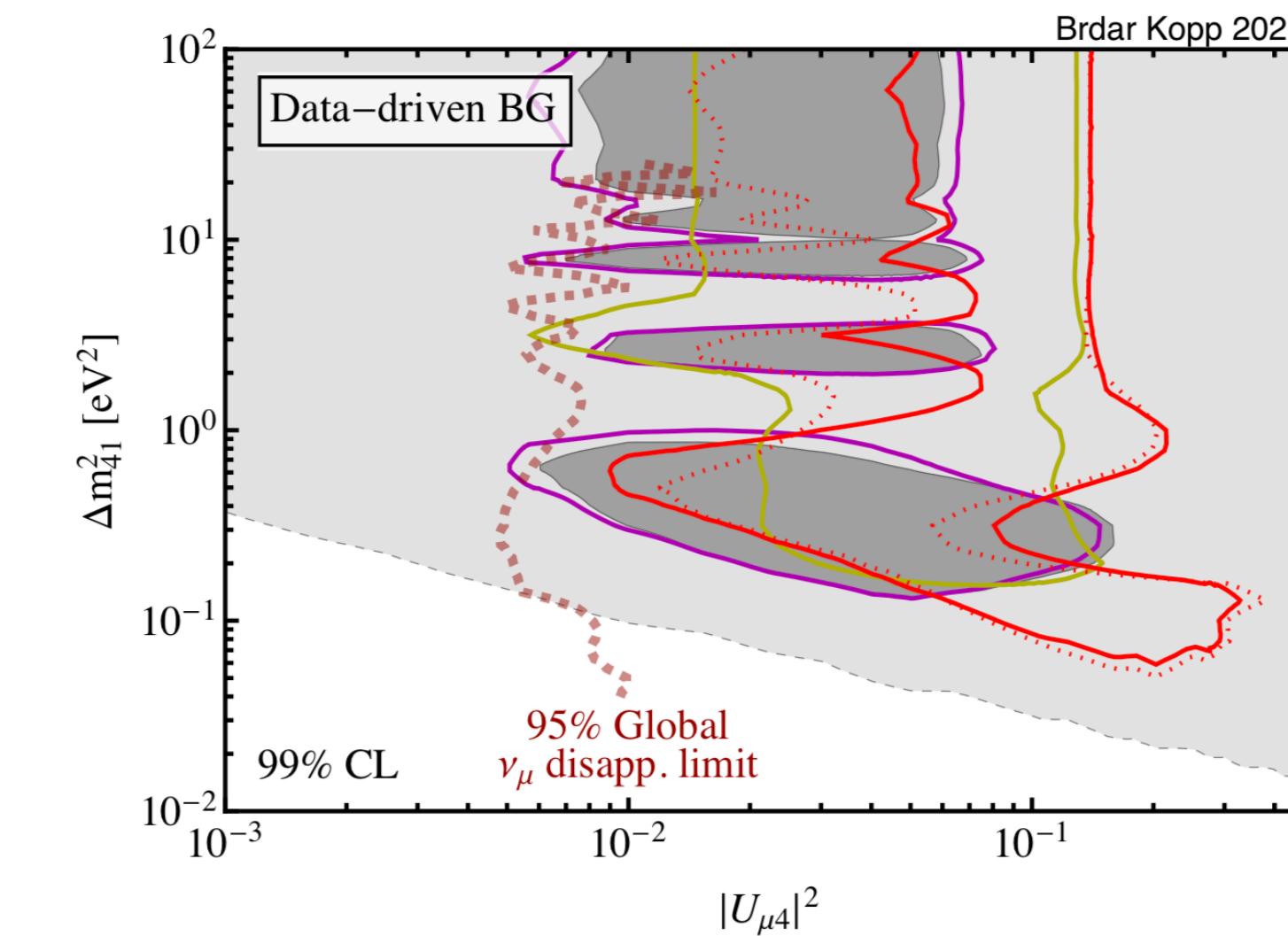
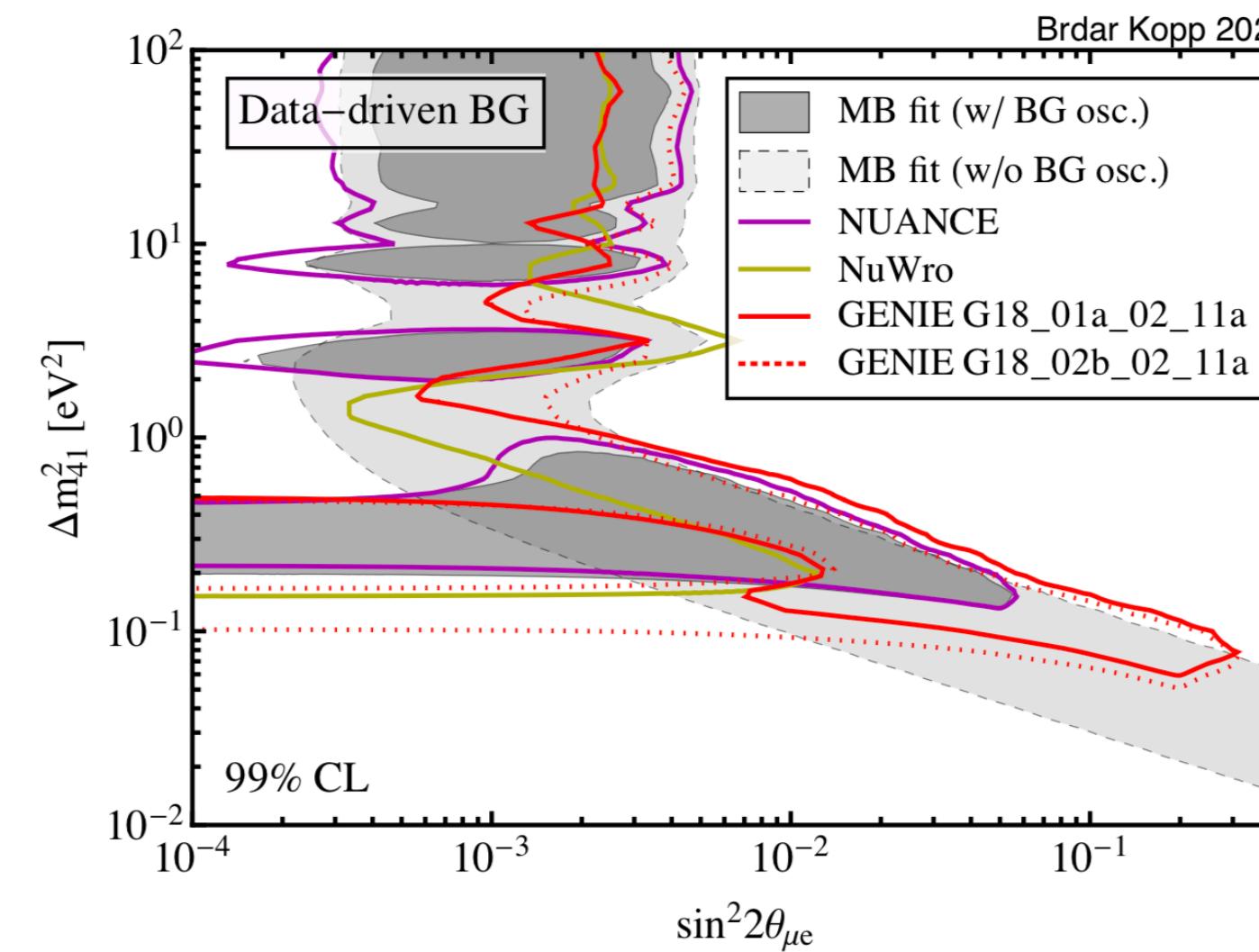
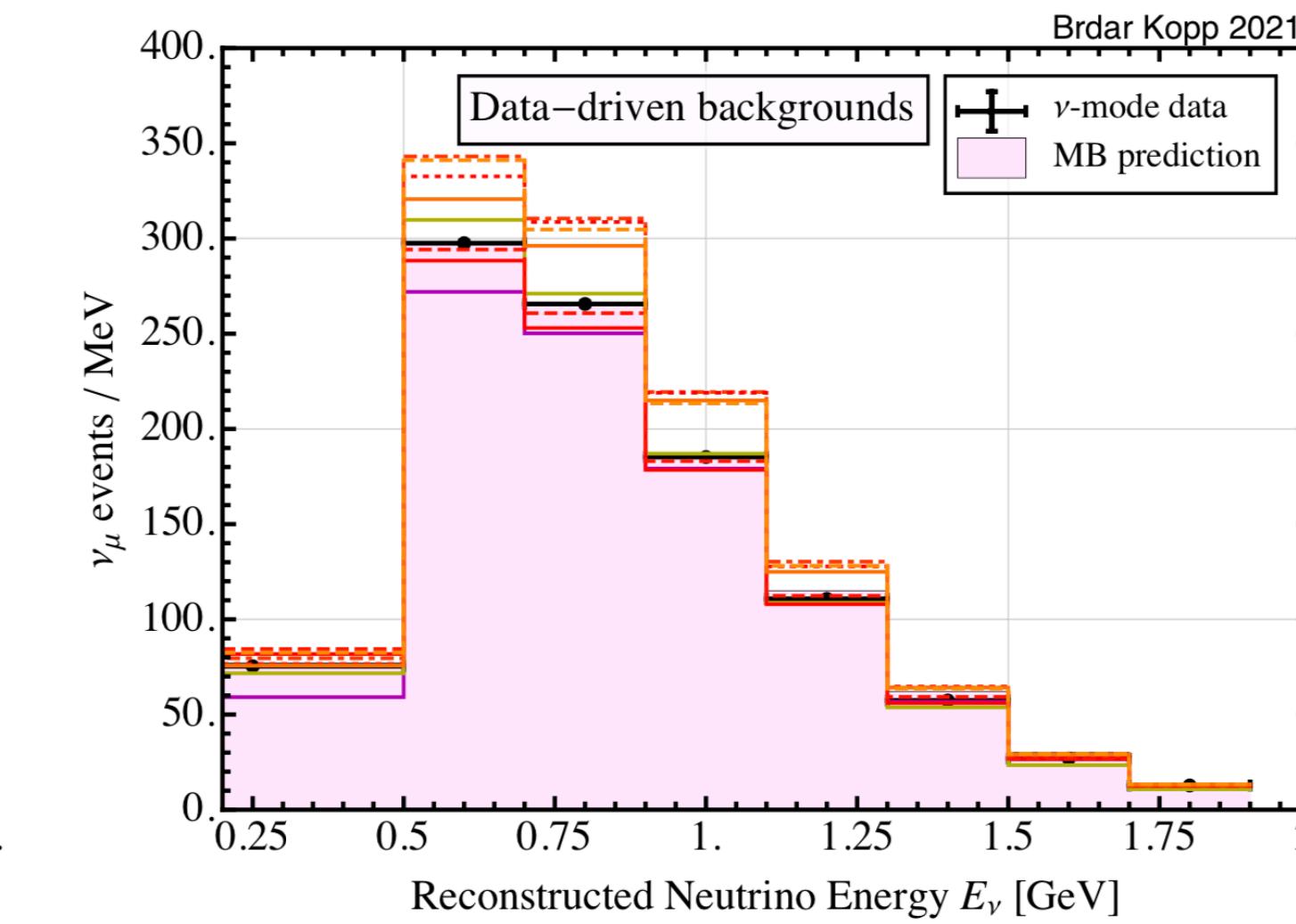
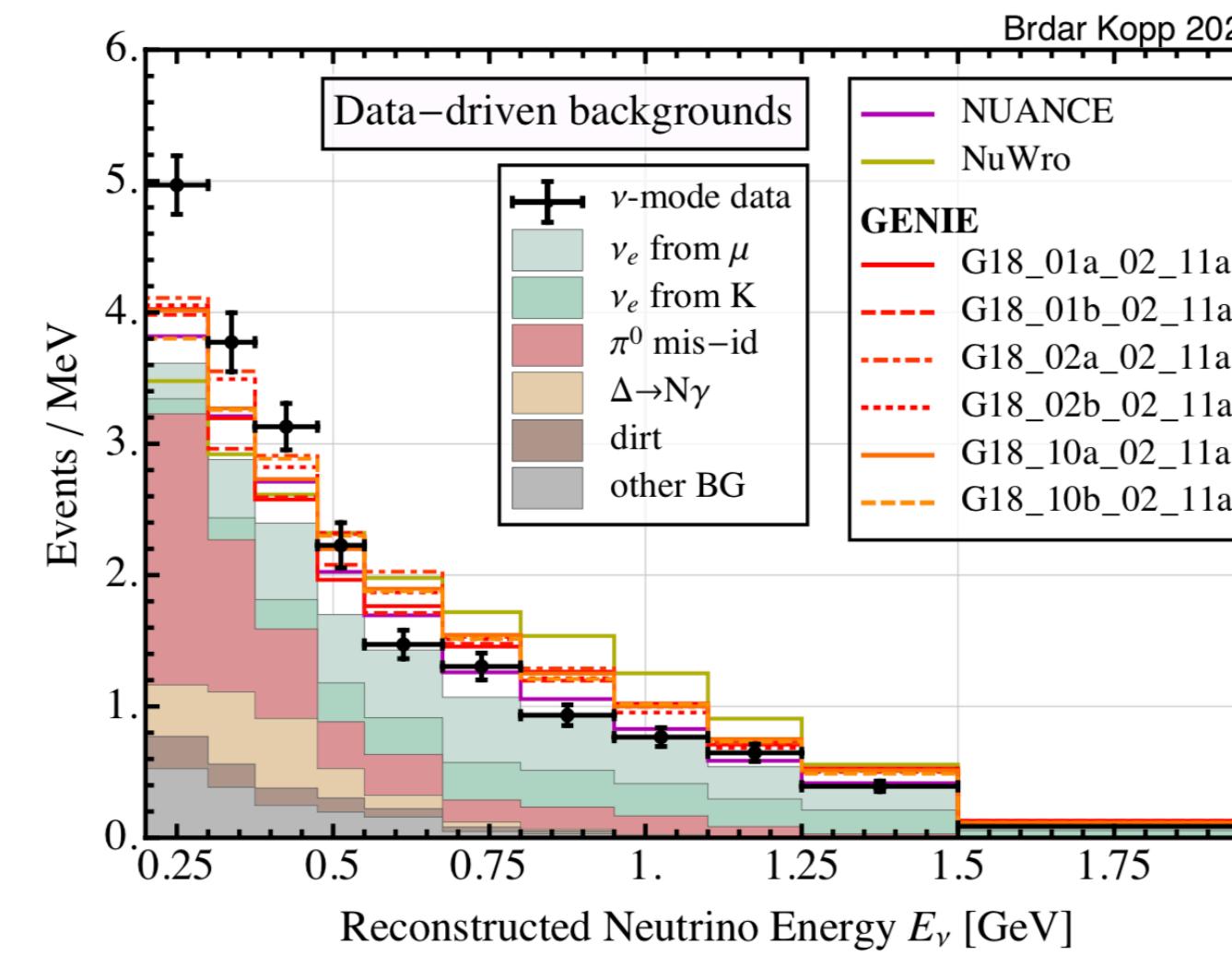
an  
**“Altarelli Cocktail”**



Image: ChatGPT  
(who refused to draw an Italian physicist w/o a beard)



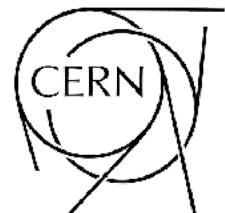
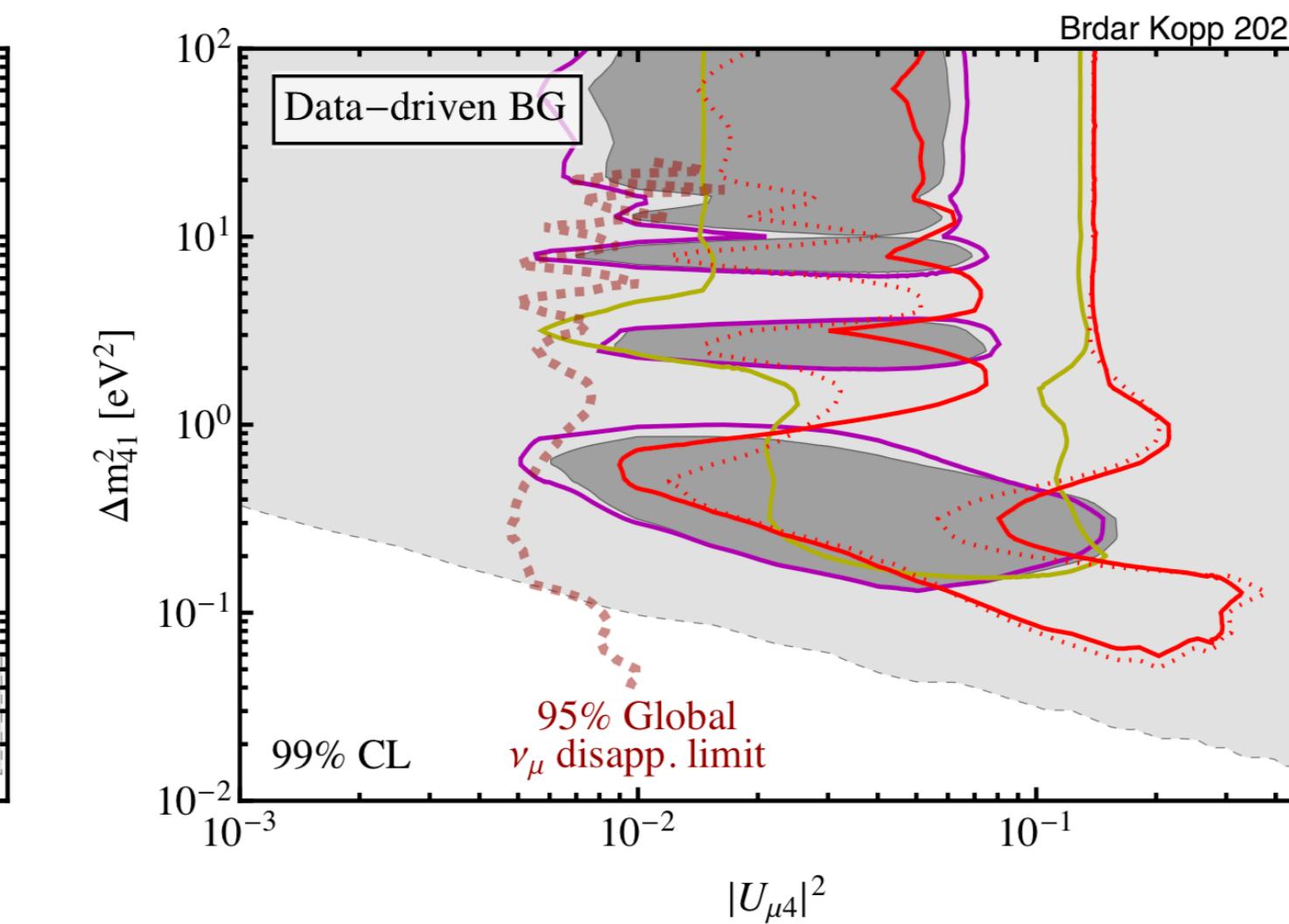
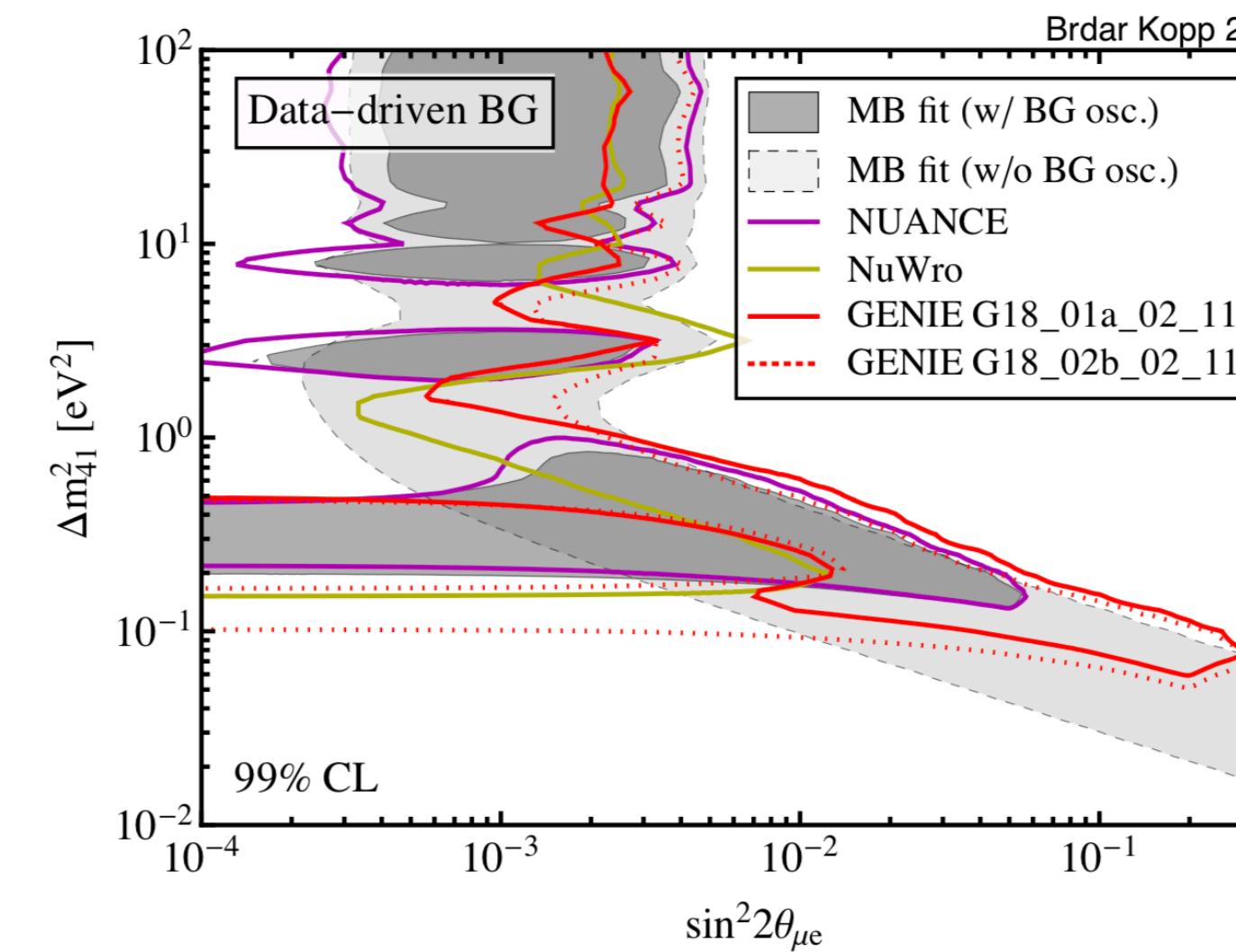
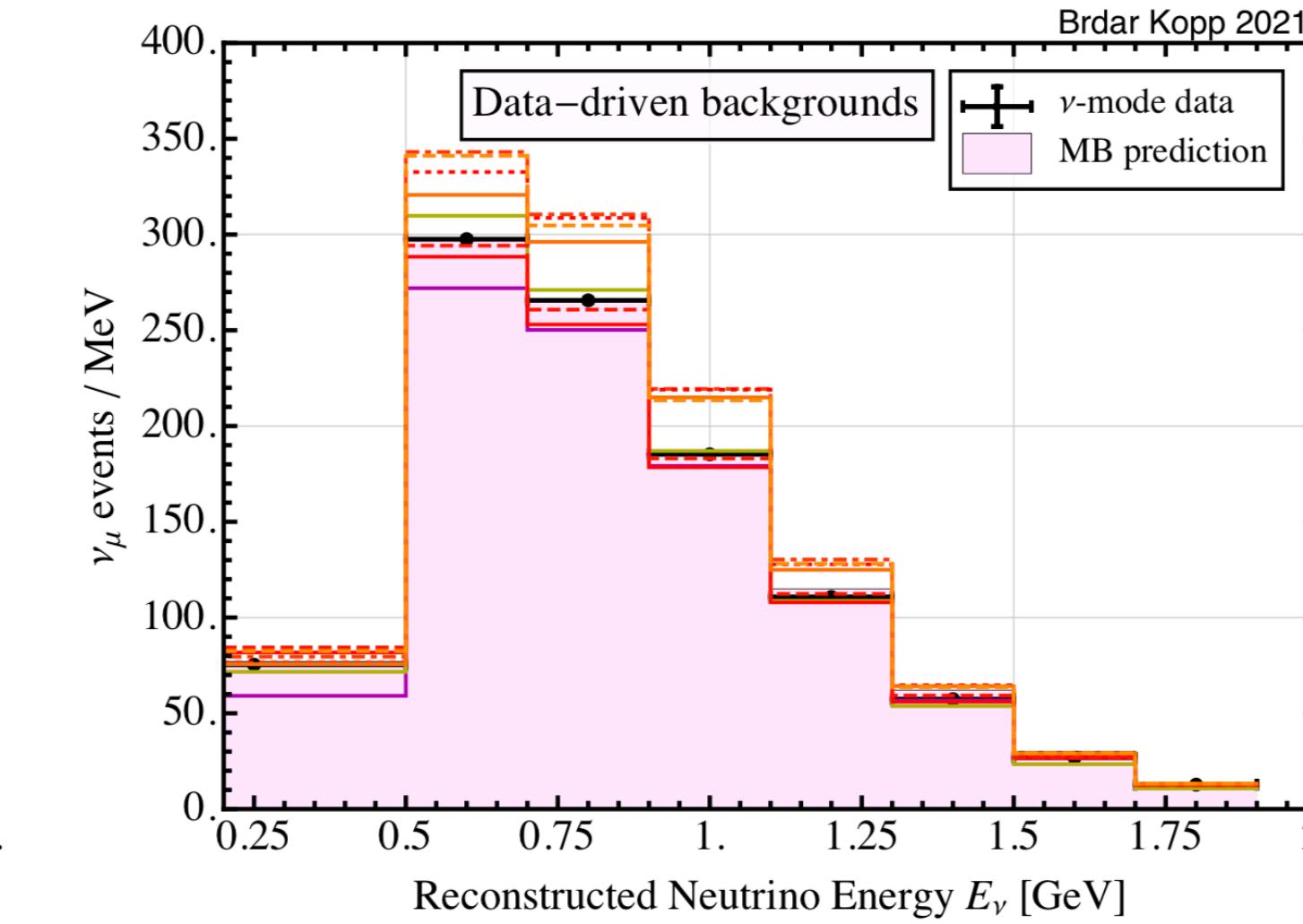
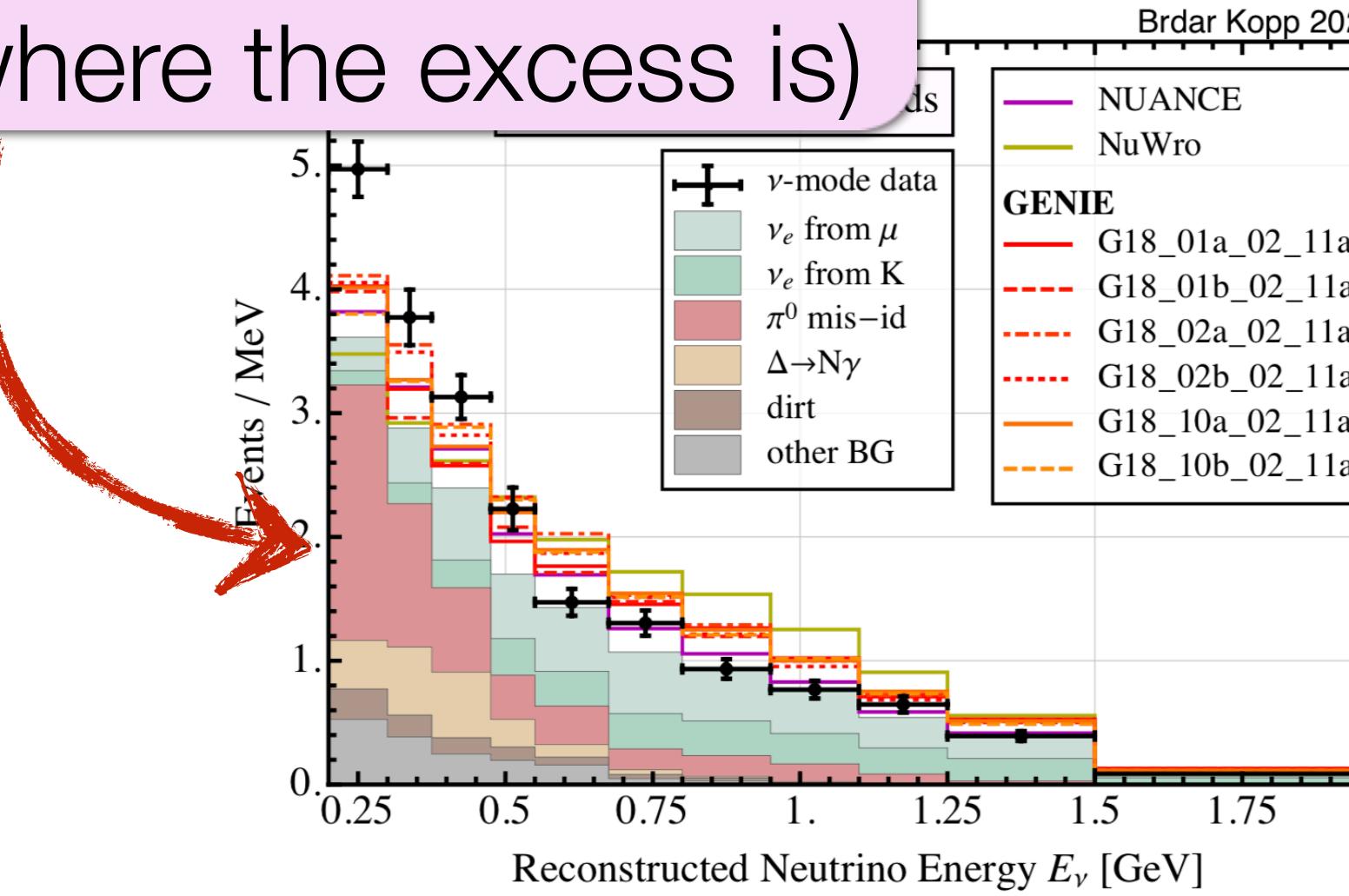
# 3+1 Models in MiniBooNE – Comparison of Generators



# 3+1 Models in MiniBooNE – Comparison of Generators

$\nu_e$  spectrum

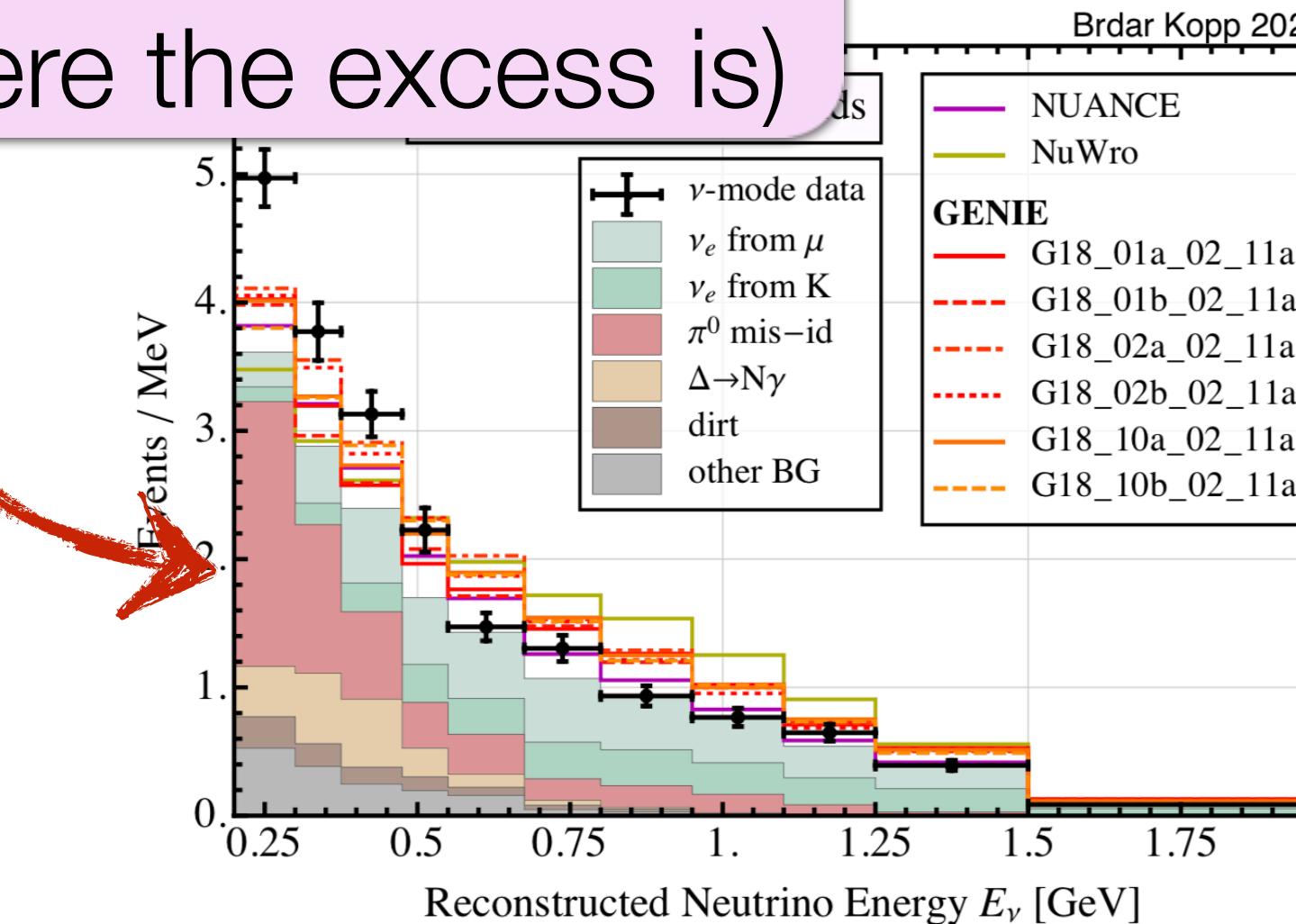
(where the excess is)



# 3+1 Models in MiniBooNE – Comparison of Generators

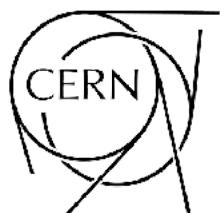
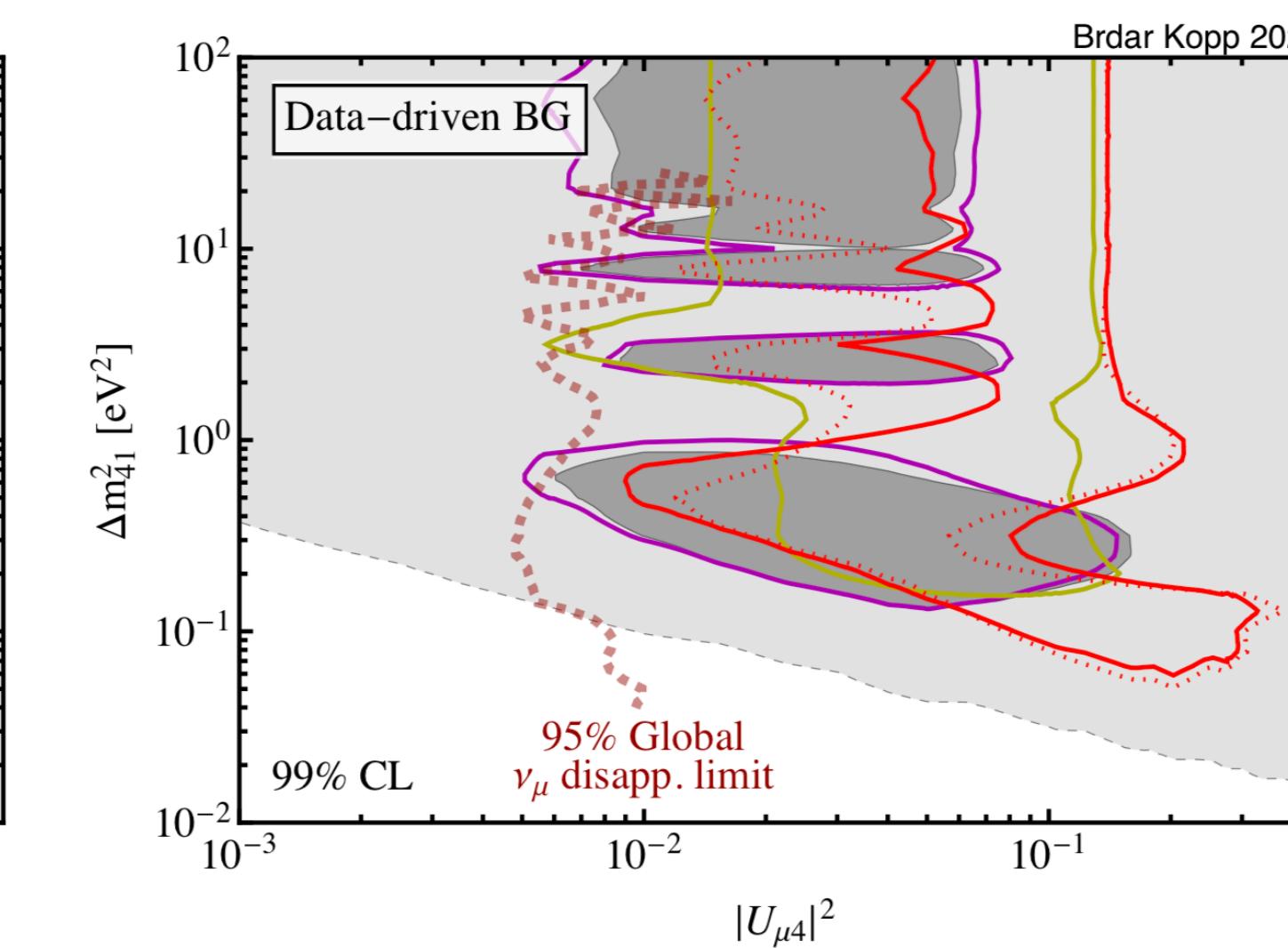
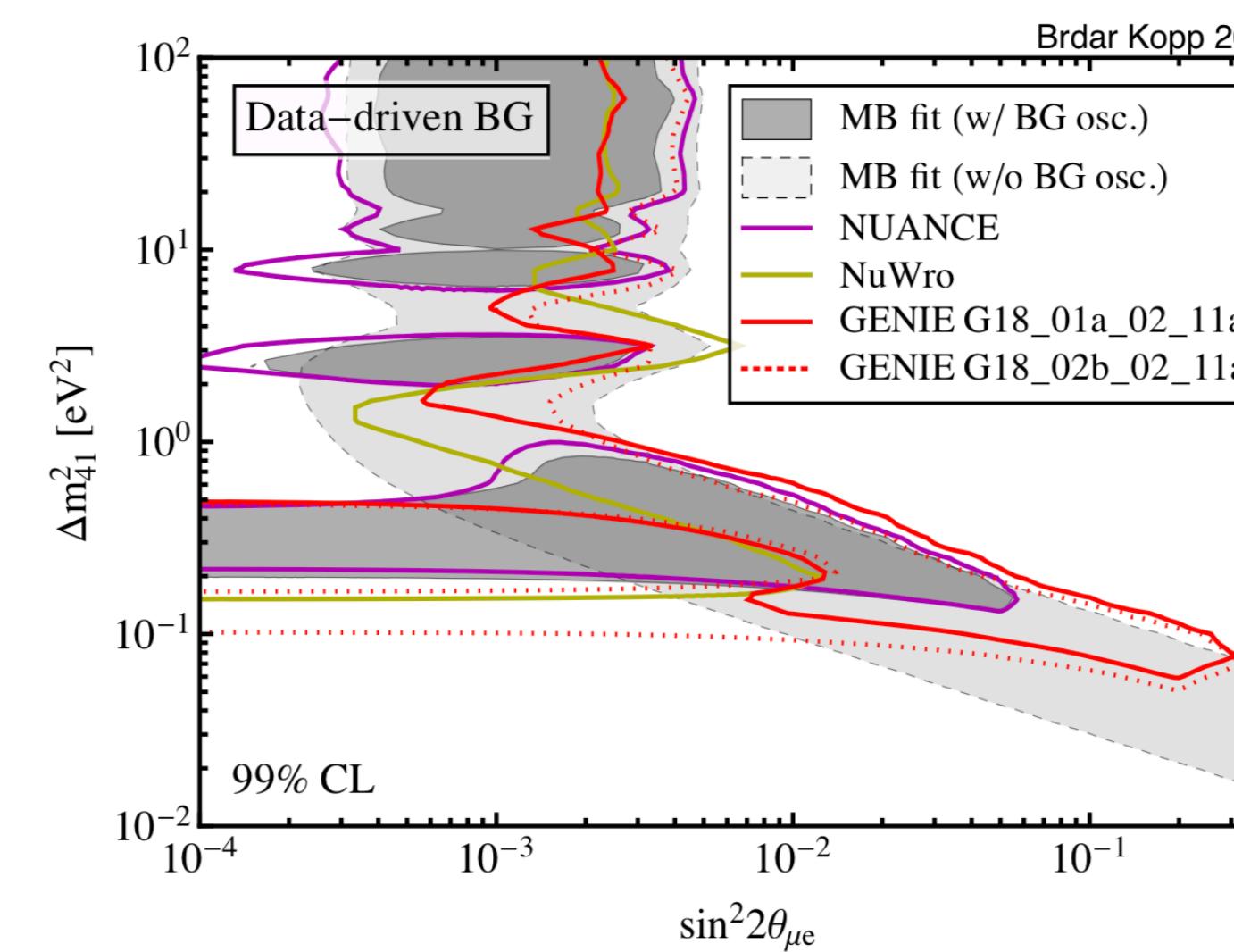
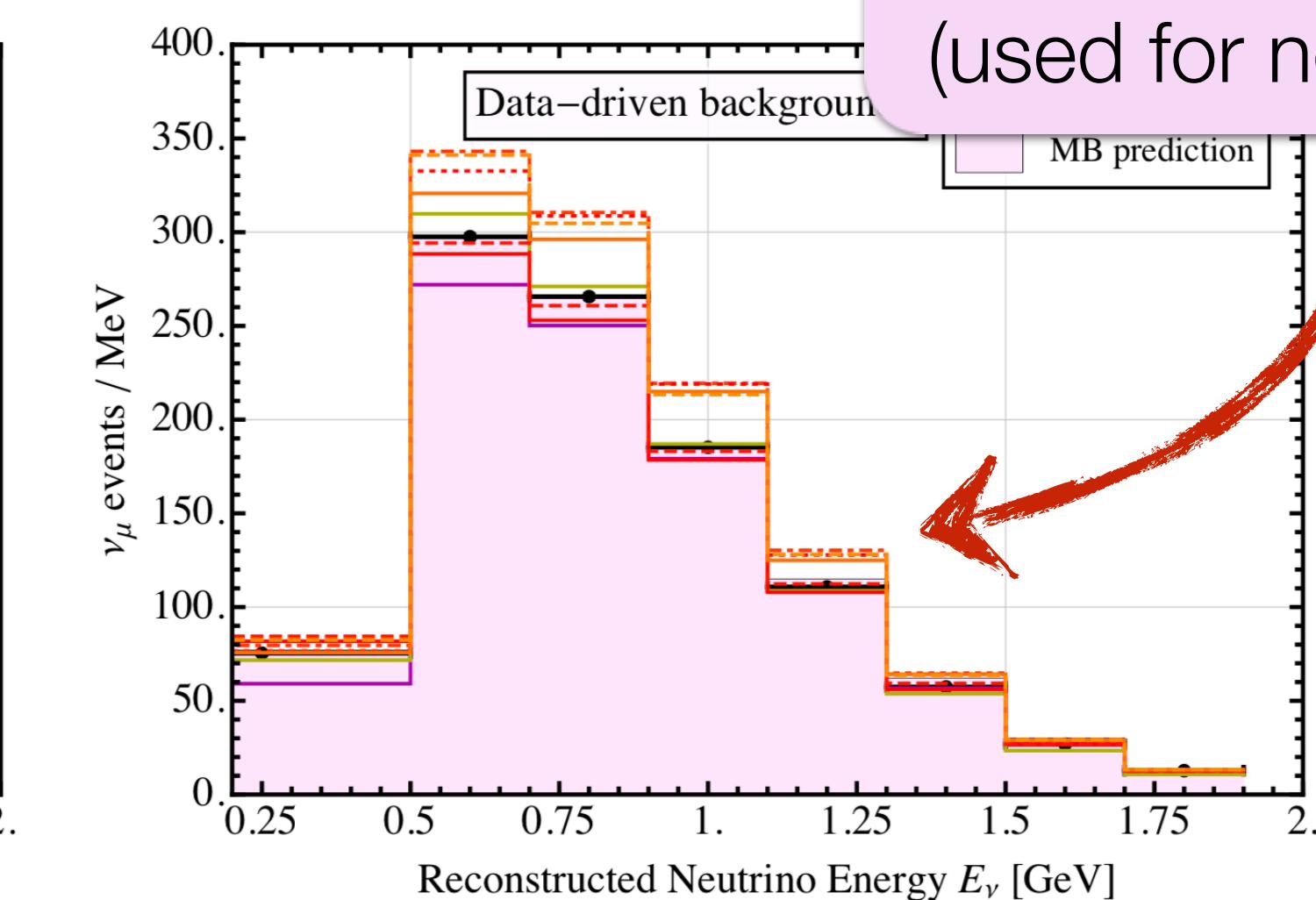
## $\nu_e$ spectrum

(where the excess is)



## $\nu_\mu$ spectrum

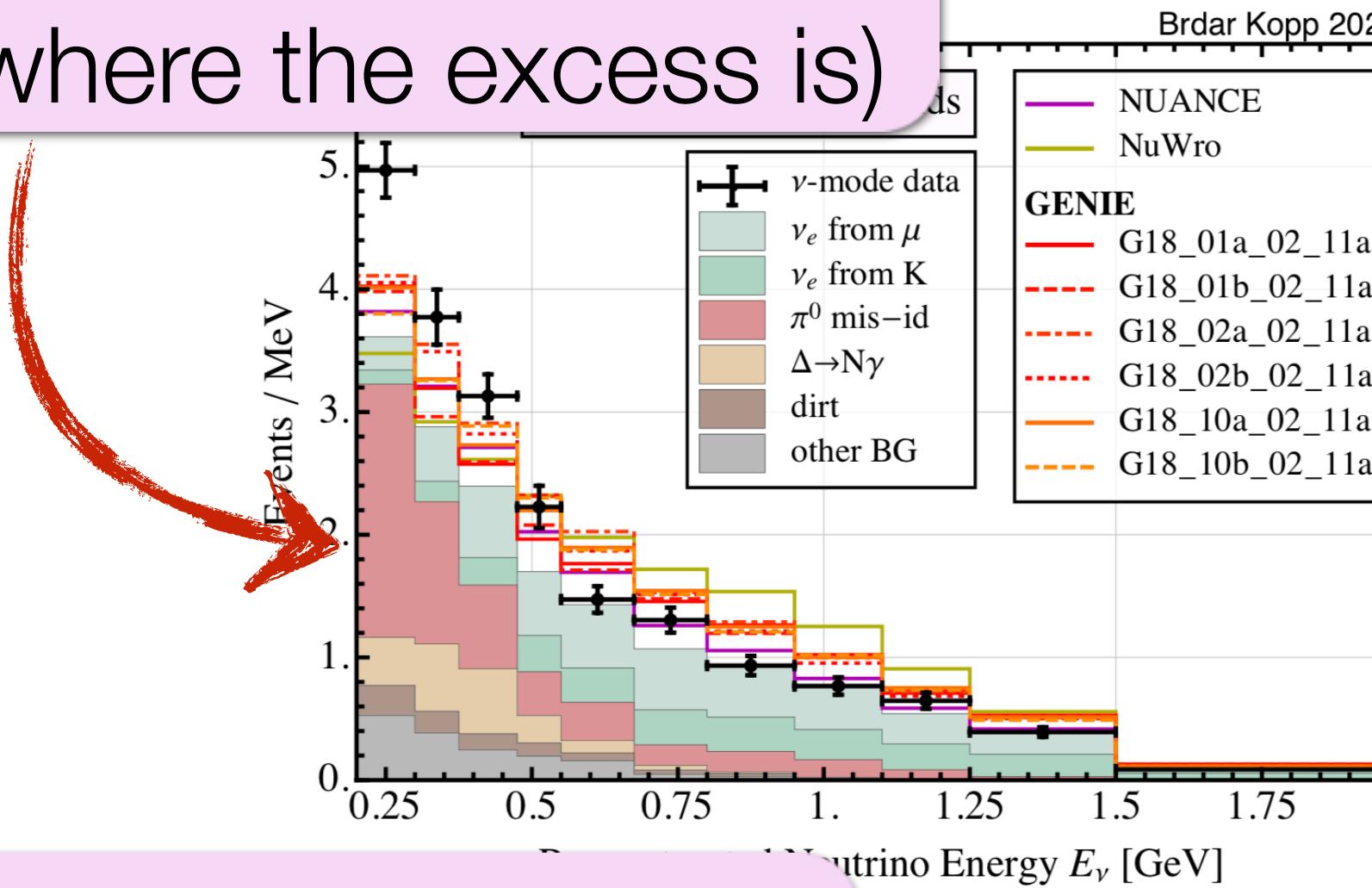
(used for normalization)



# 3+1 Models in MiniBooNE – Comparison of Generators

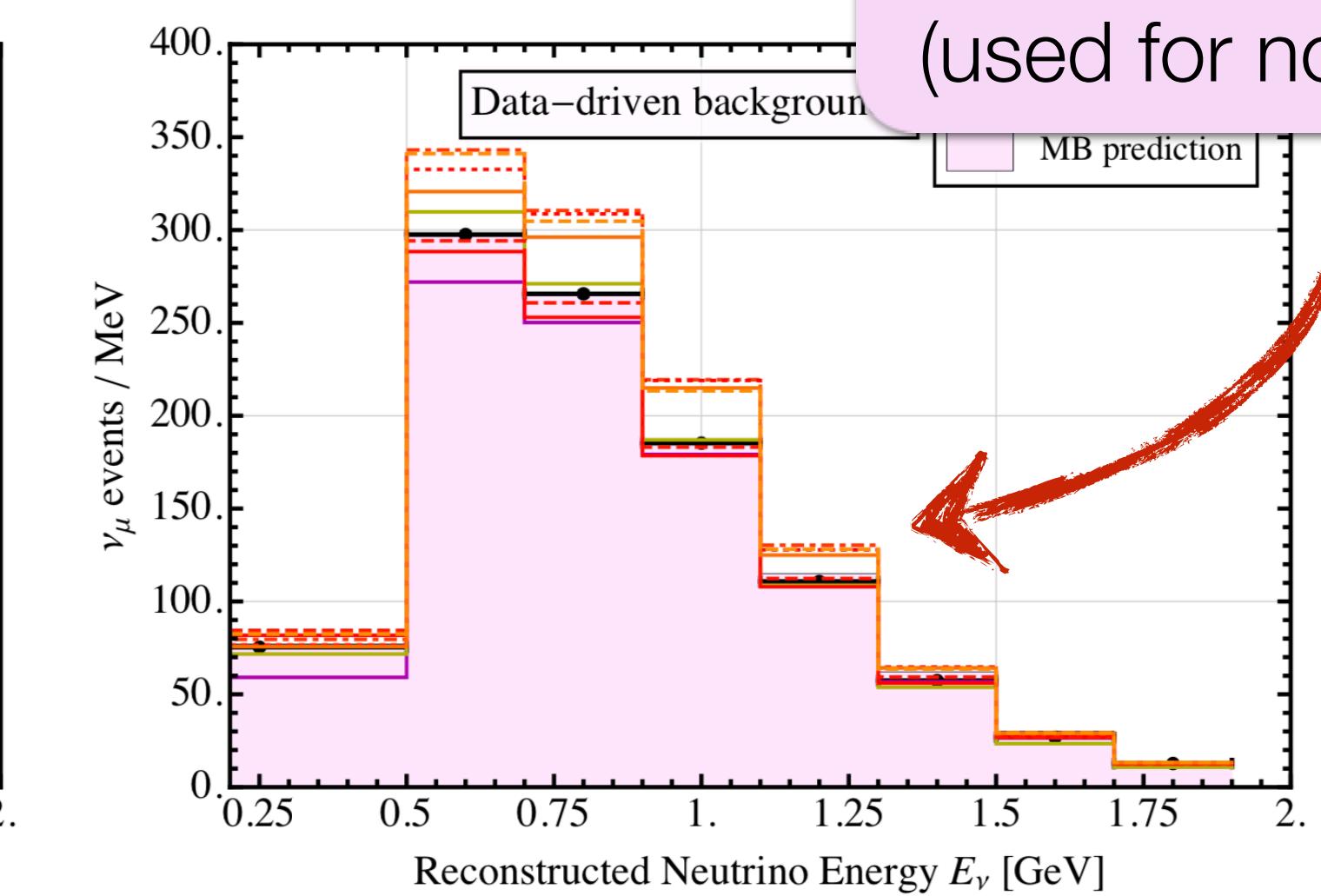
$\nu_e$  spectrum

(where the excess is)

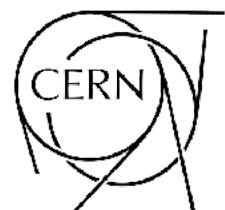
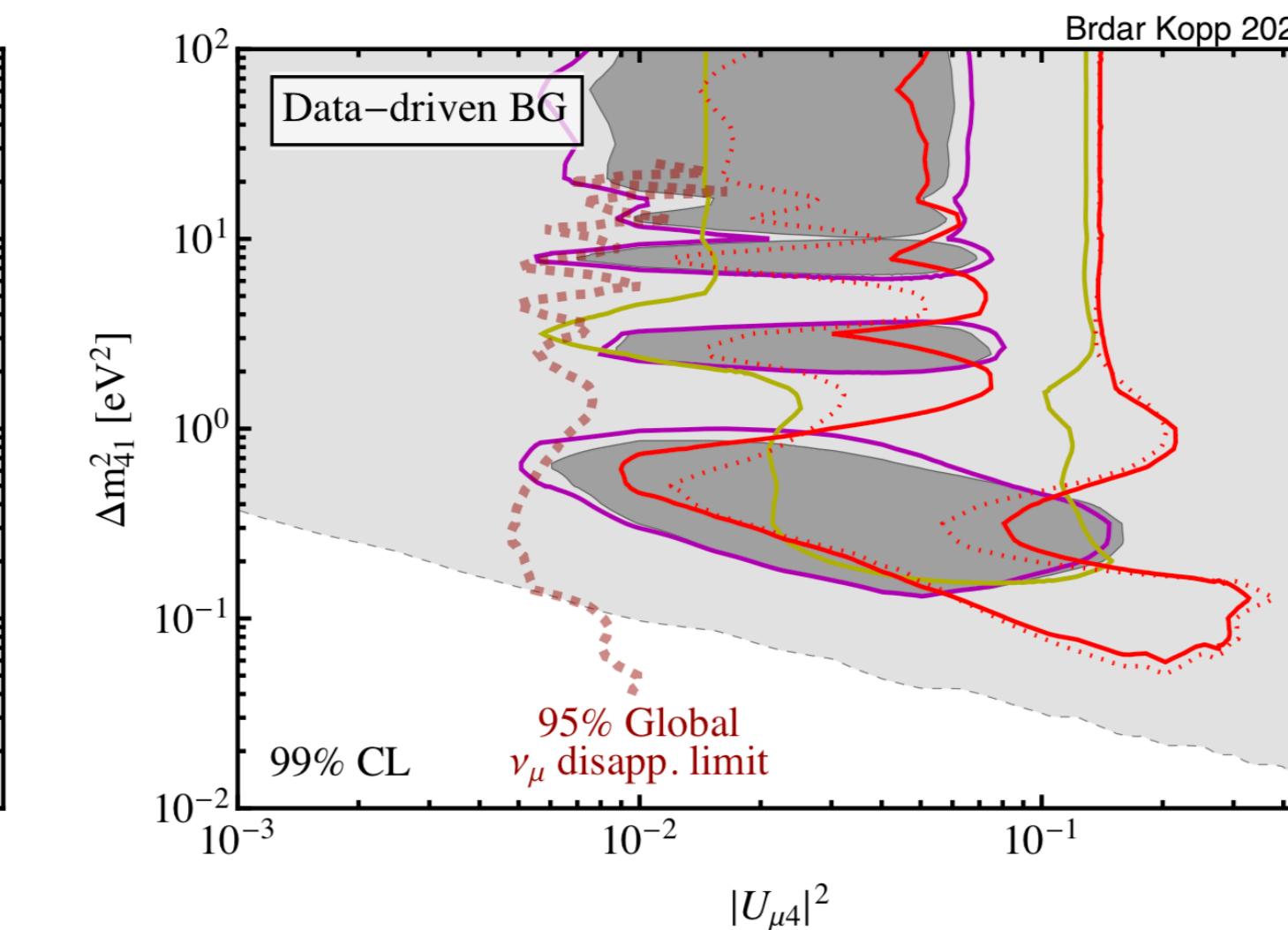
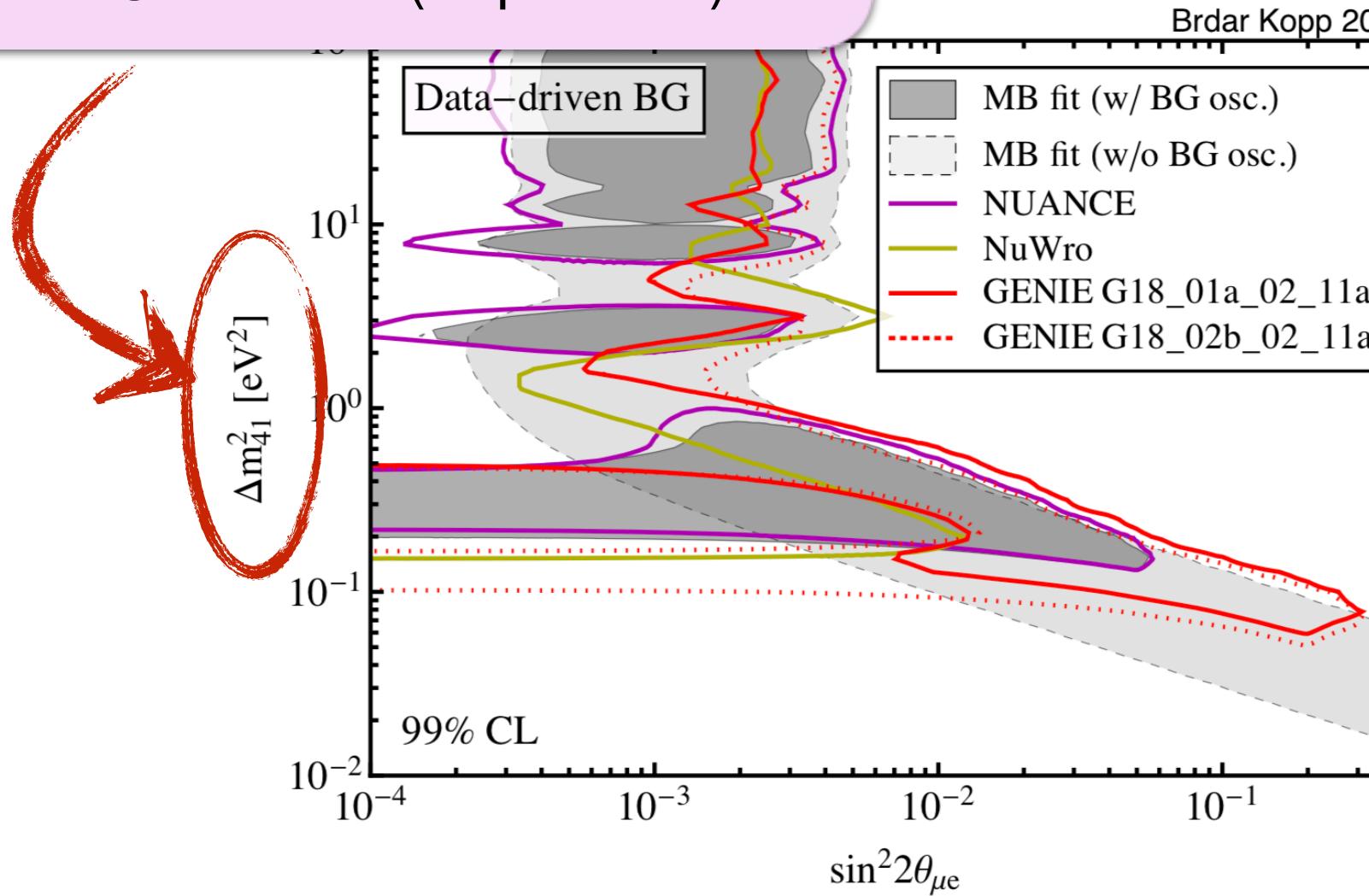


$\nu_\mu$  spectrum

(used for normalization)



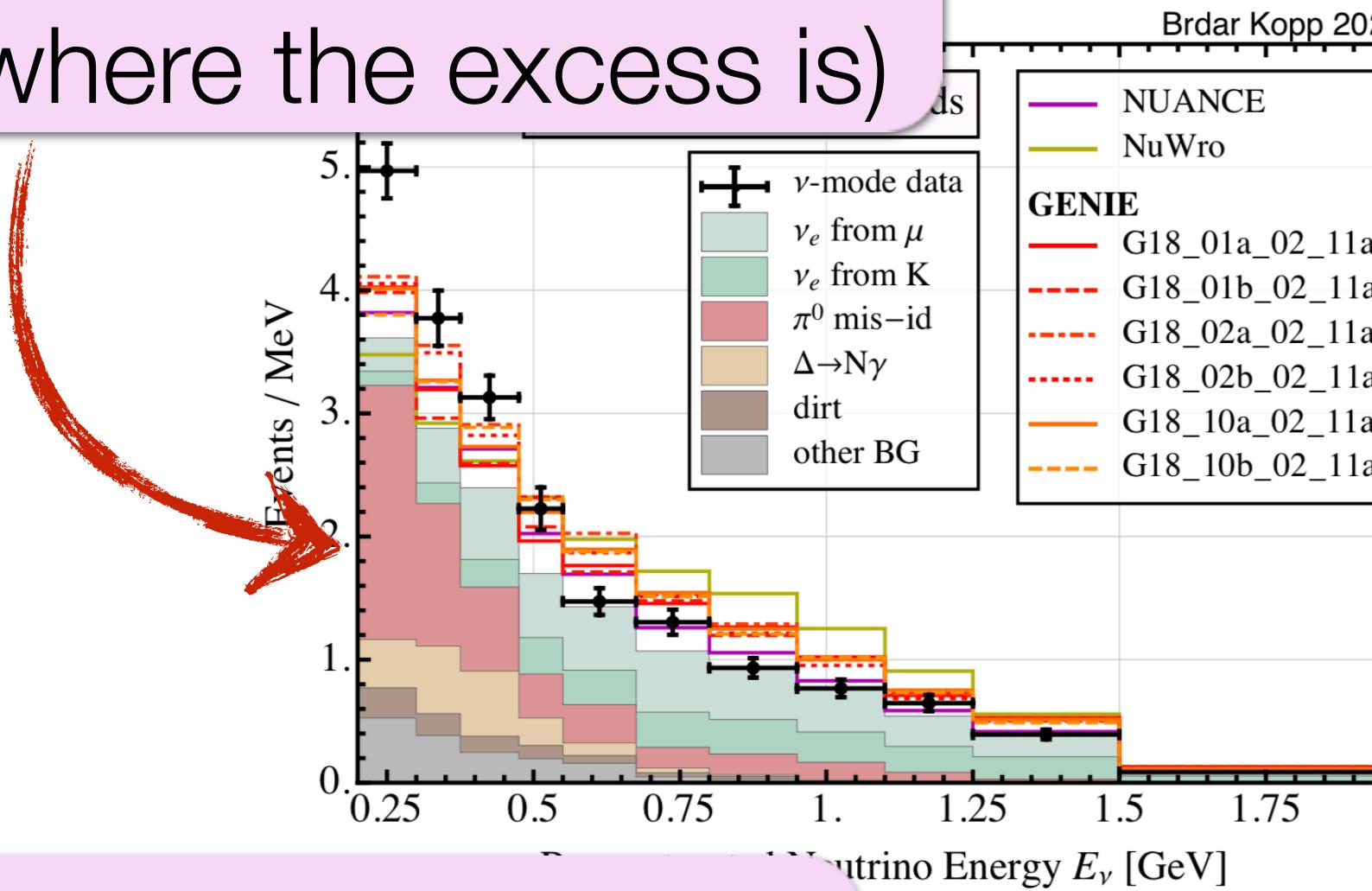
$\nu_s$  mass (squared)



# 3+1 Models in MiniBooNE – Comparison of Generators

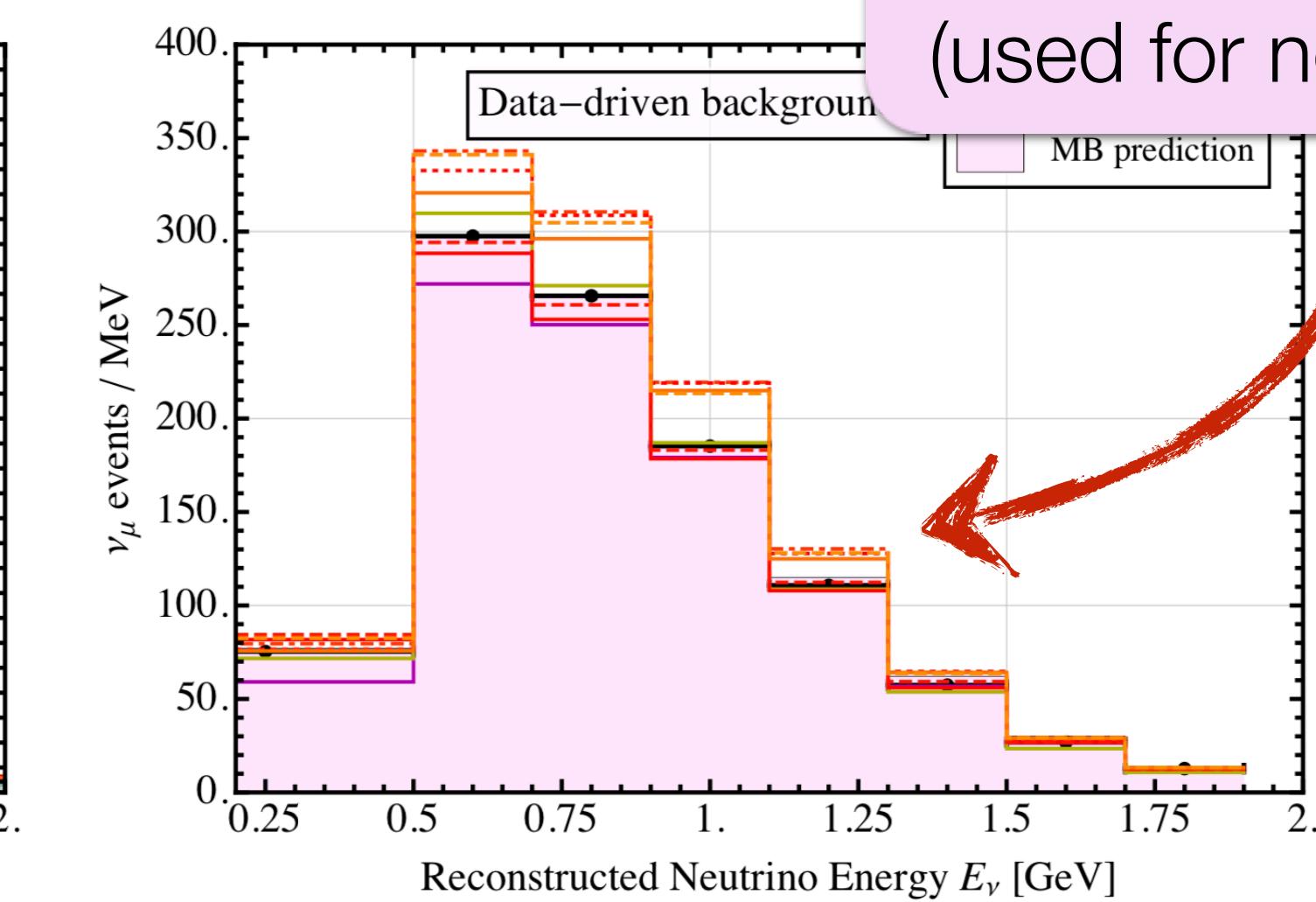
$\nu_e$  spectrum

(where the excess is)

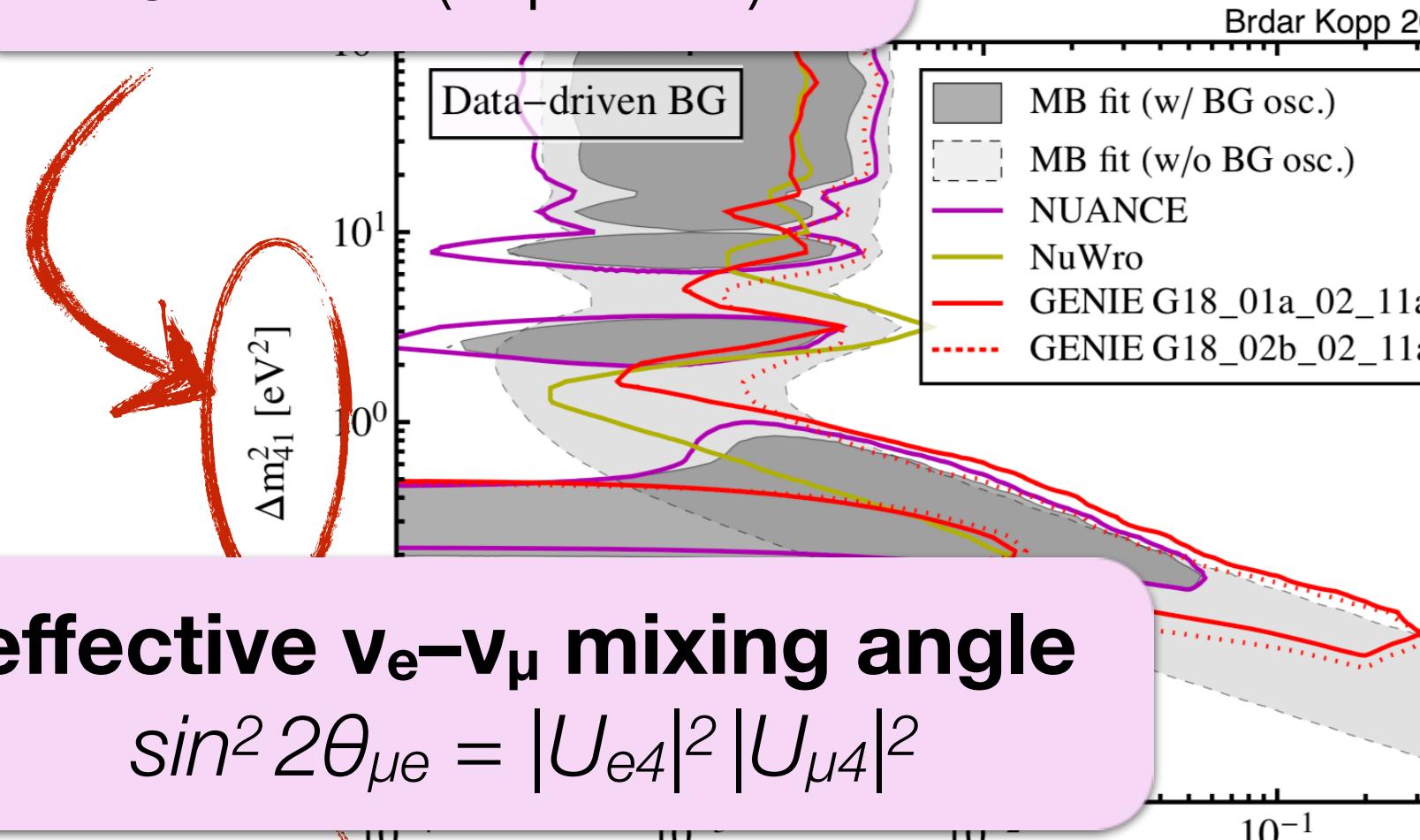


$\nu_\mu$  spectrum

(used for normalization)

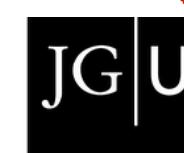
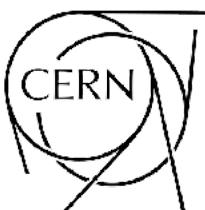
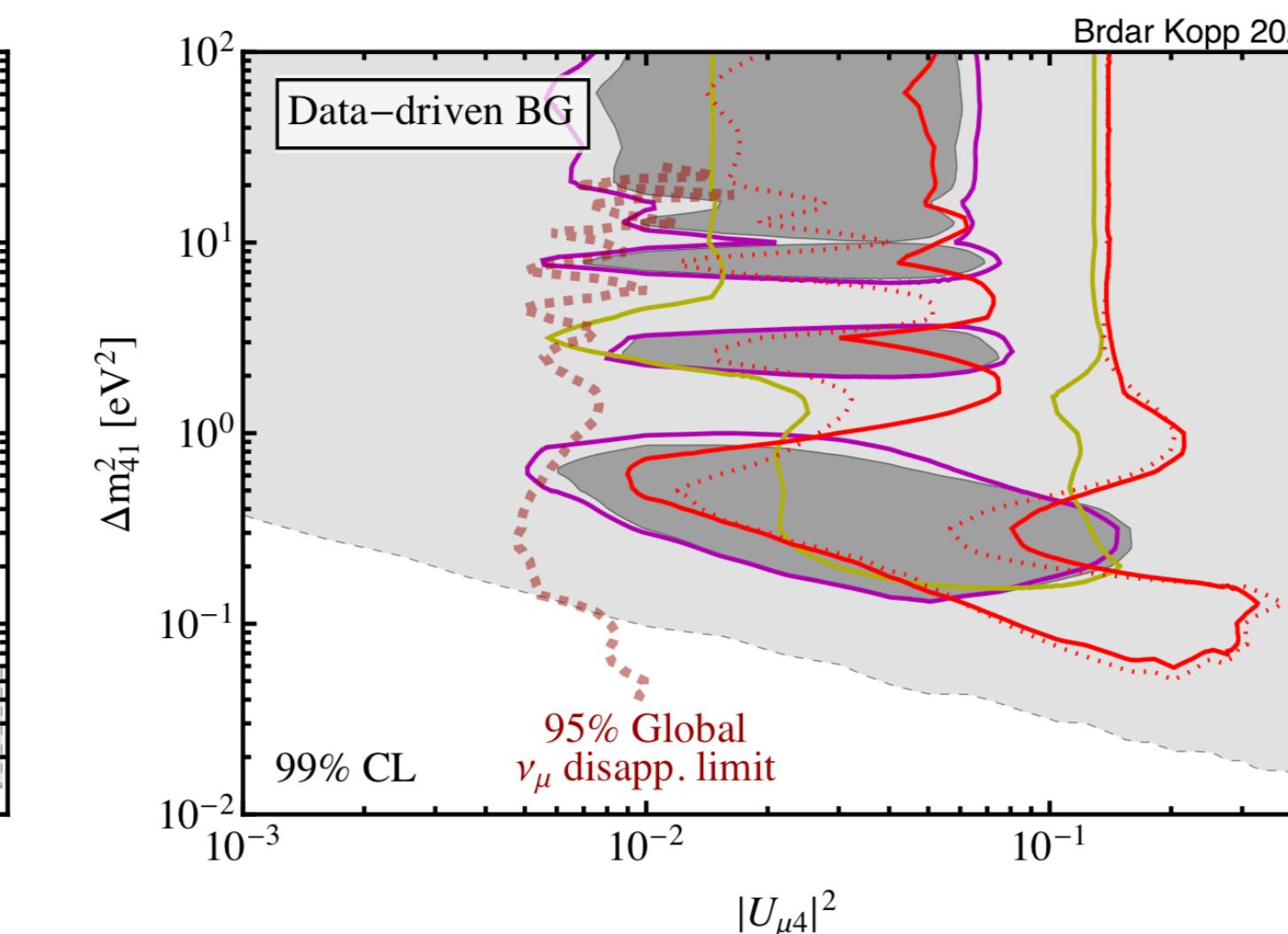


$\nu_s$  mass (squared)



effective  $\nu_e$ - $\nu_\mu$  mixing angle

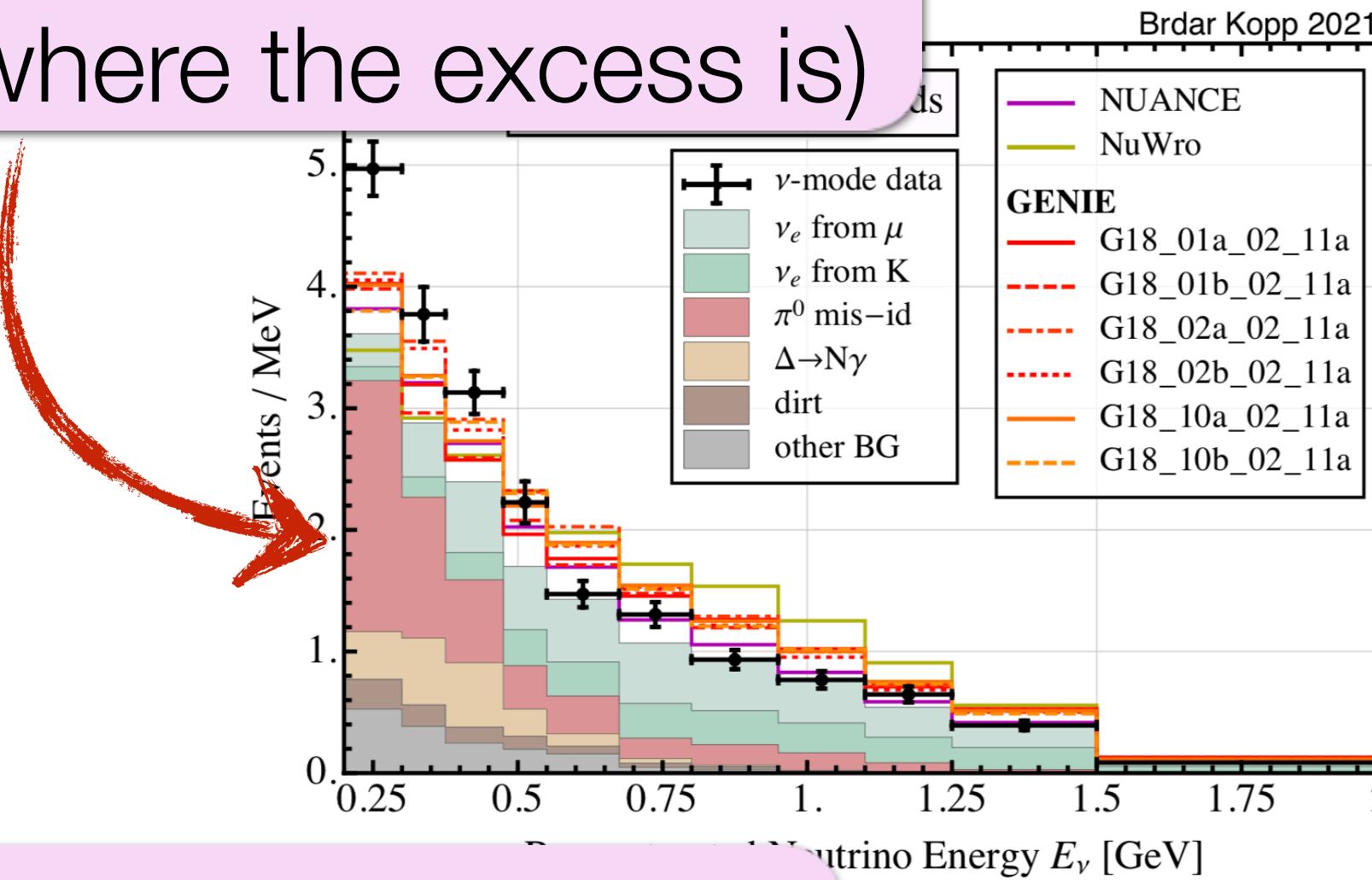
$$\sin^2 2\theta_{\mu e} = |U_{e4}|^2 |U_{\mu 4}|^2$$



# 3+1 Models in MiniBooNE – Comparison of Generators

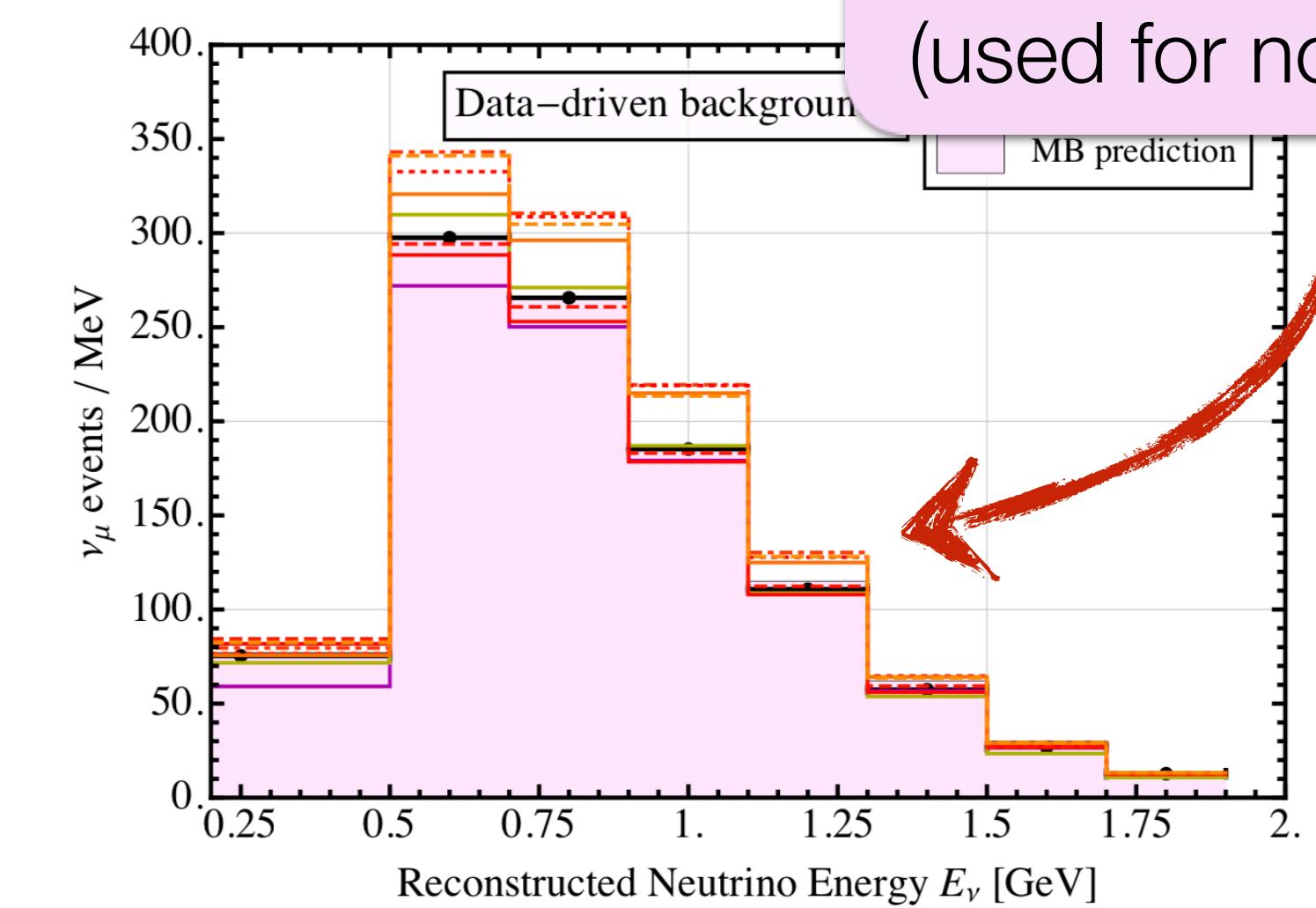
$\nu_e$  spectrum

(where the excess is)

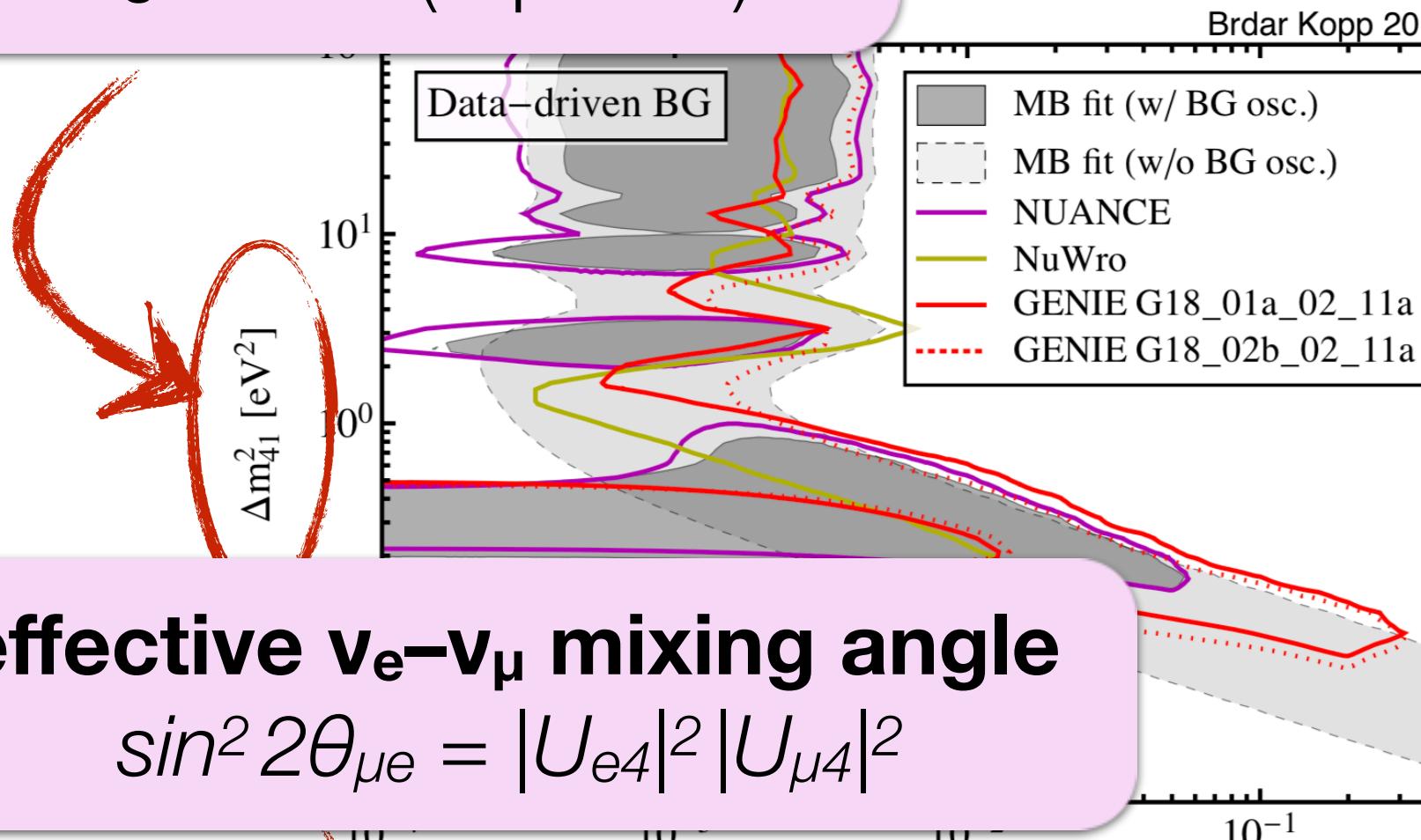


$\nu_\mu$  spectrum

(used for normalization)

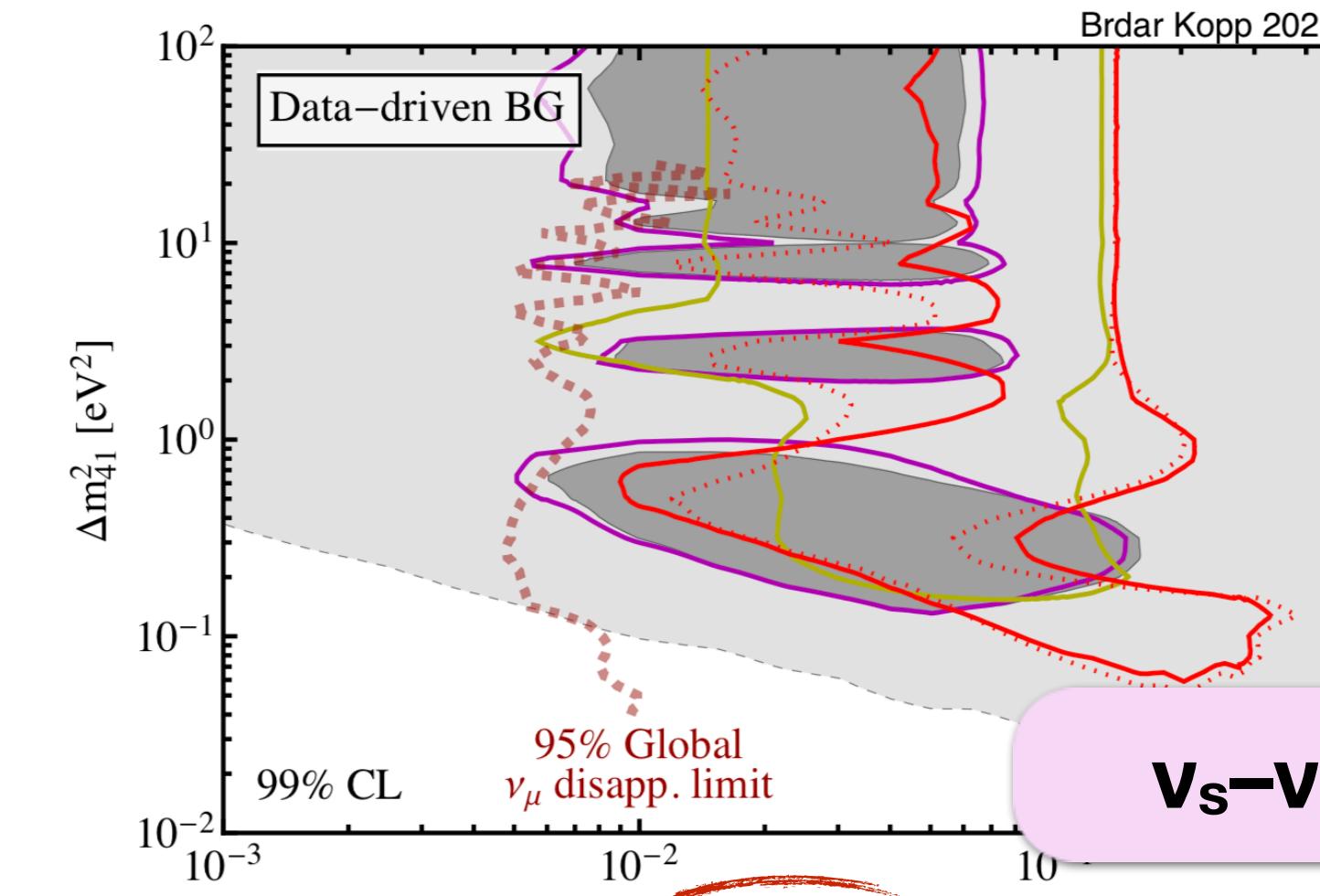


$\nu_s$  mass (squared)

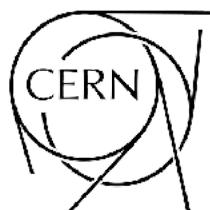


effective  $\nu_e$ - $\nu_\mu$  mixing angle

$$\sin^2 2\theta_{\mu e} = |U_{e4}|^2 |U_{\mu 4}|^2$$



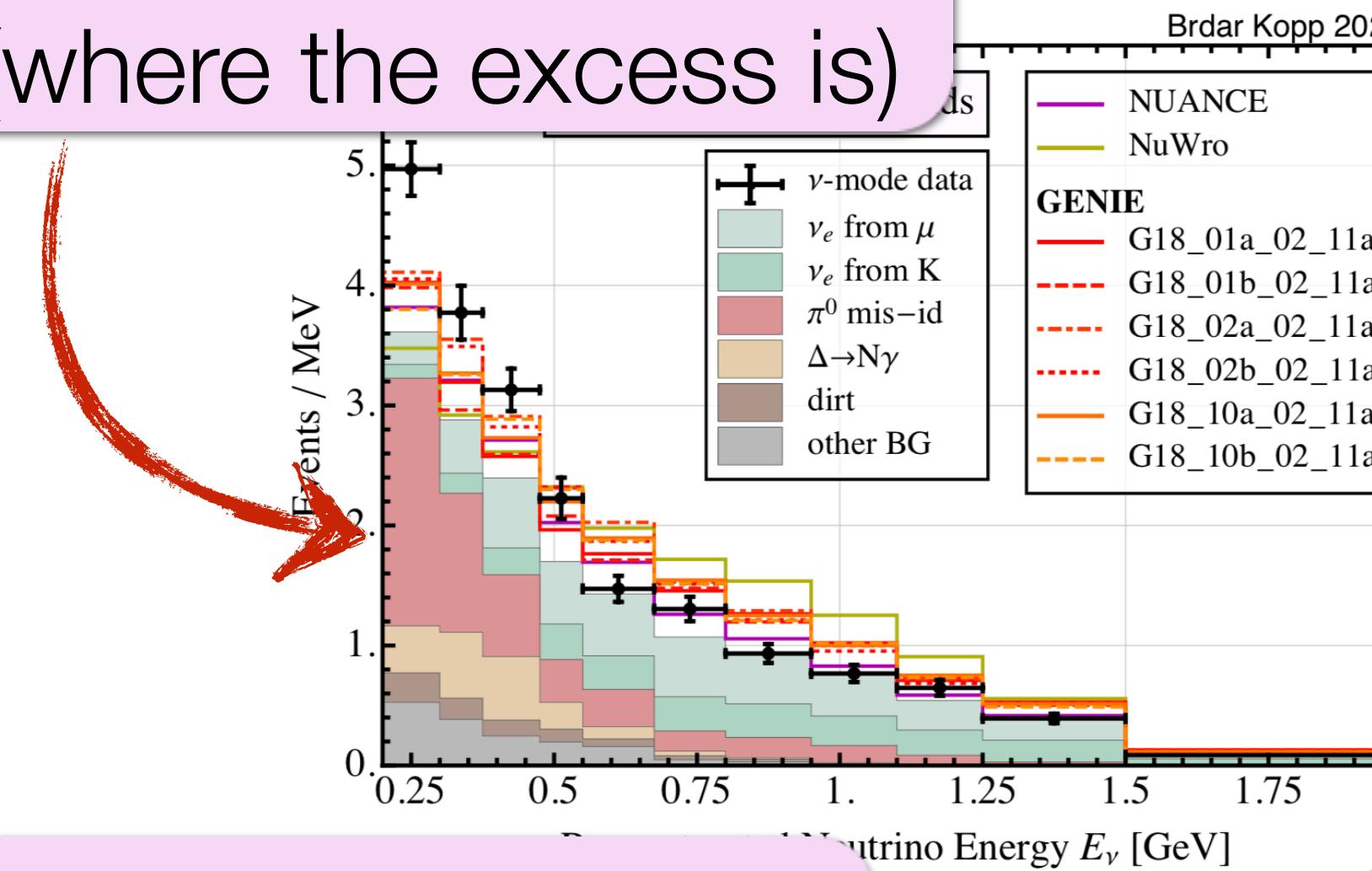
$\nu_s$ - $\nu_\mu$  mixing



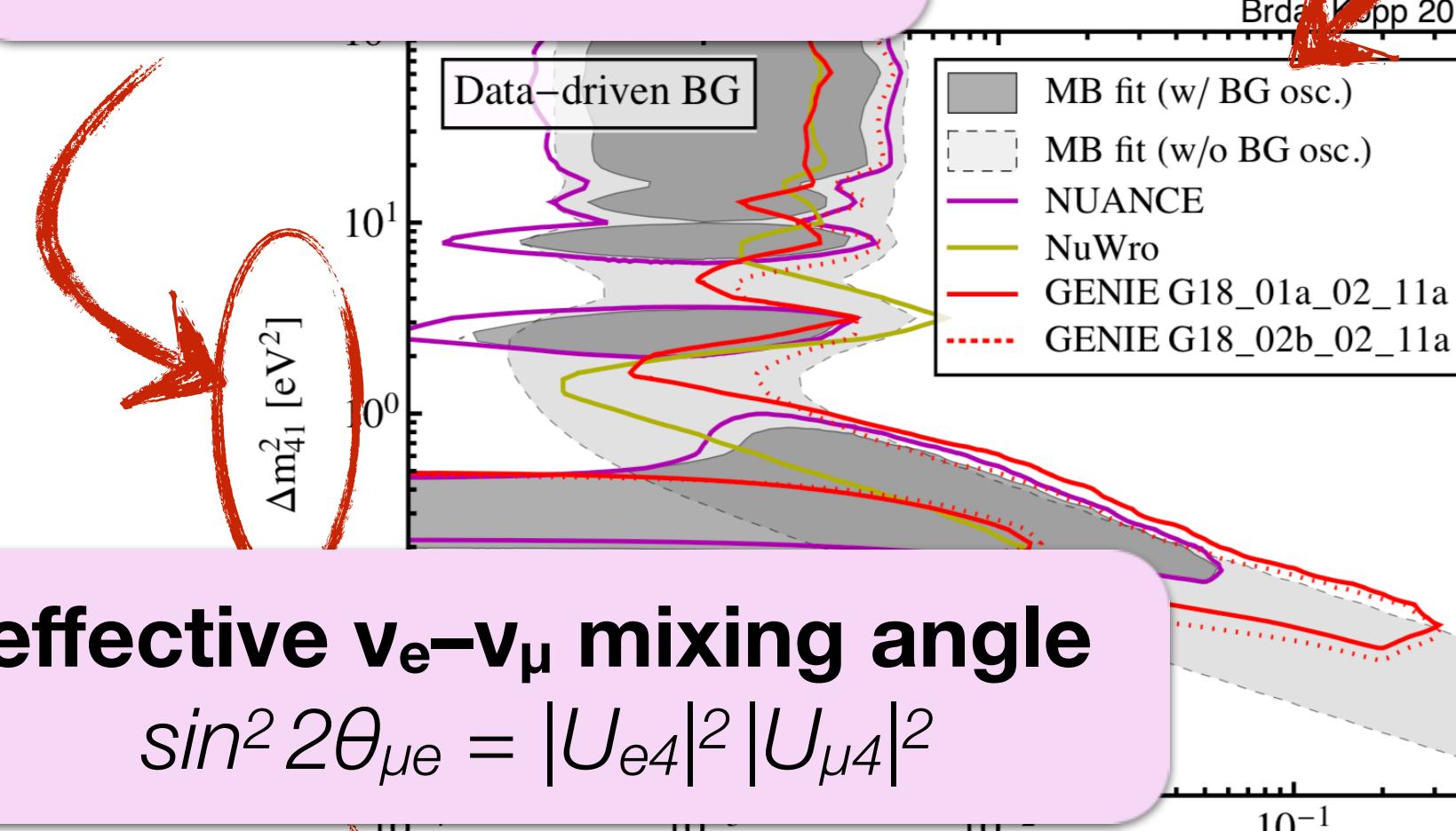
# 3+1 Models in MiniBooNE – Comparison of Generators

**$\nu_e$  spectrum**

(where the excess is)



**$\nu_s$  mass (squared)**



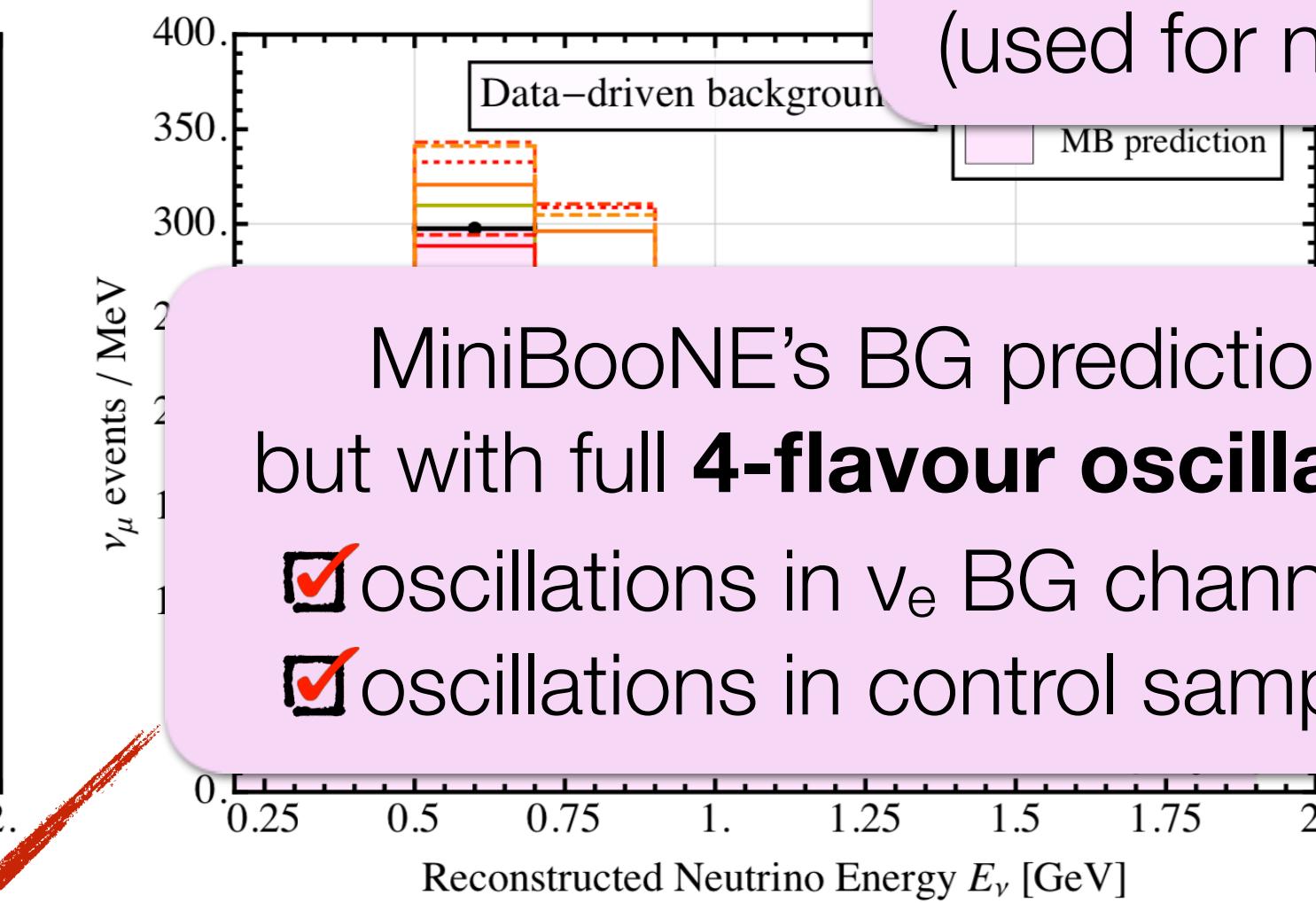
**effective  $\nu_e$ - $\nu_\mu$  mixing angle**

$$\sin^2 2\theta_{\mu e} = |U_{e4}|^2 |U_{\mu 4}|^2$$



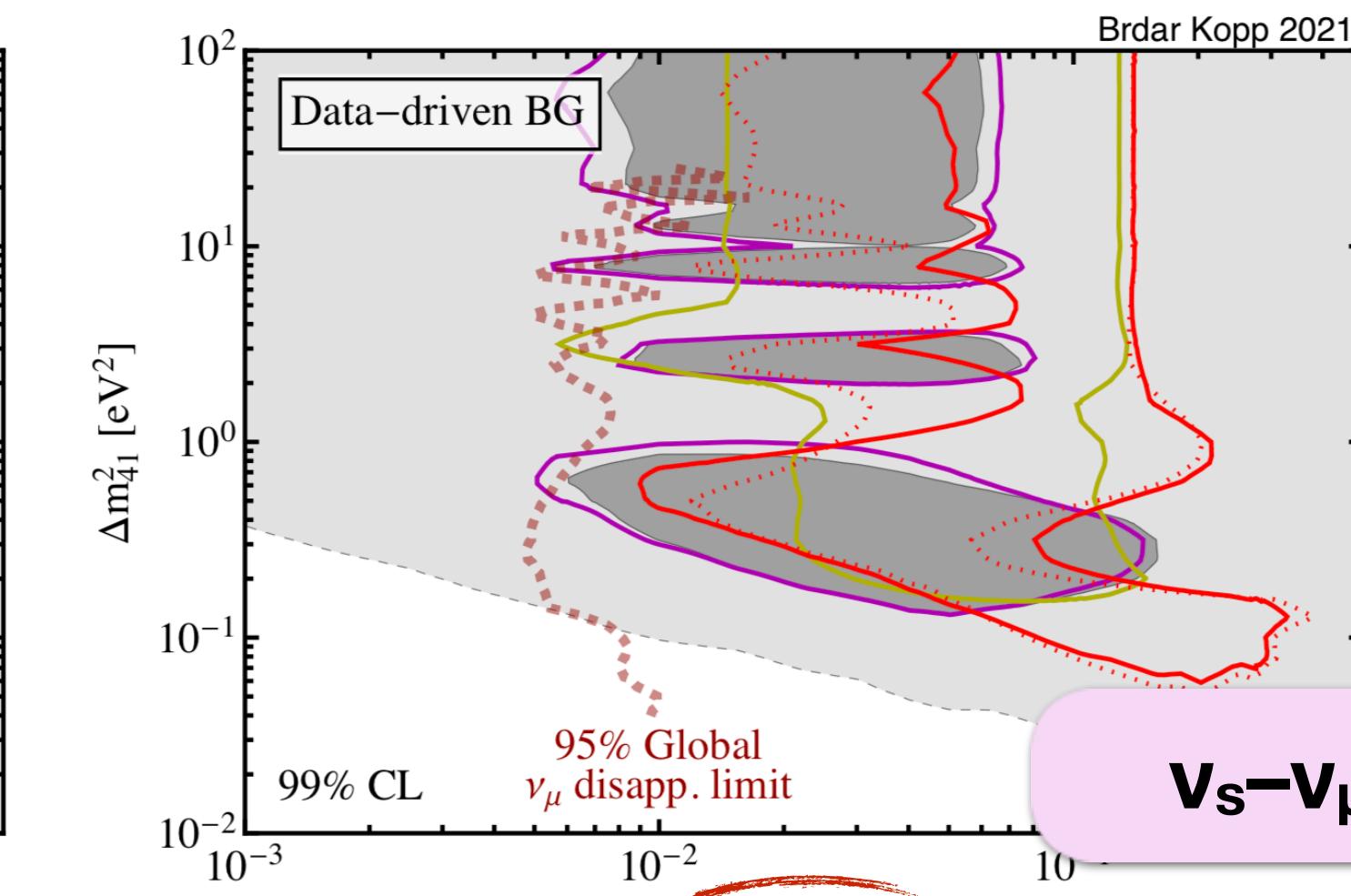
**$\nu_\mu$  spectrum**

(used for normalization)

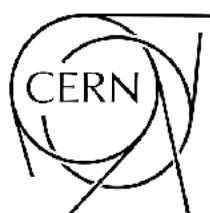


MiniBooNE's BG predictions,  
but with full **4-flavour oscillations**

- oscillations in  $\nu_e$  BG channels
- oscillations in control sample



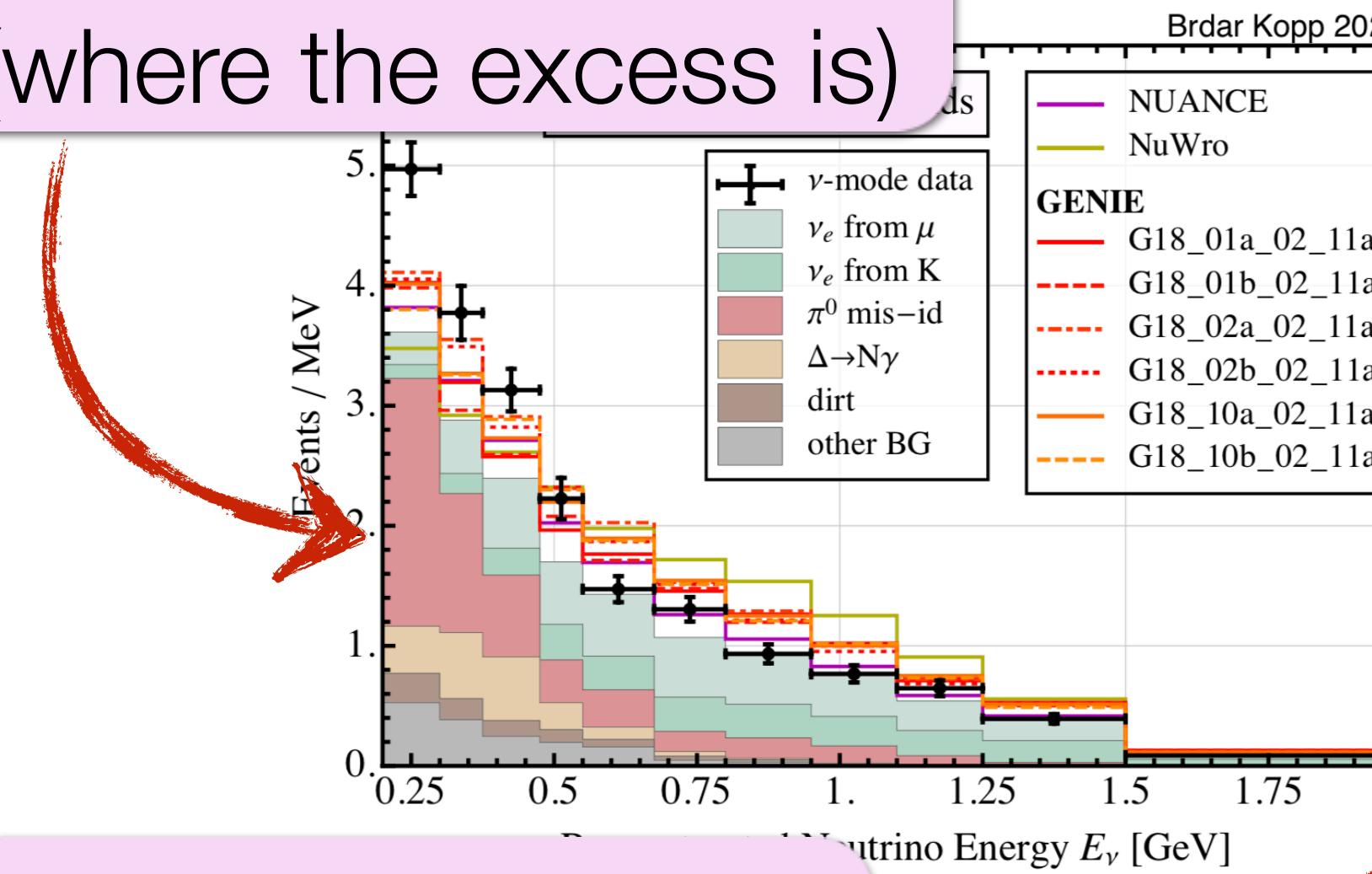
**$\nu_s$ - $\nu_\mu$  mixing**



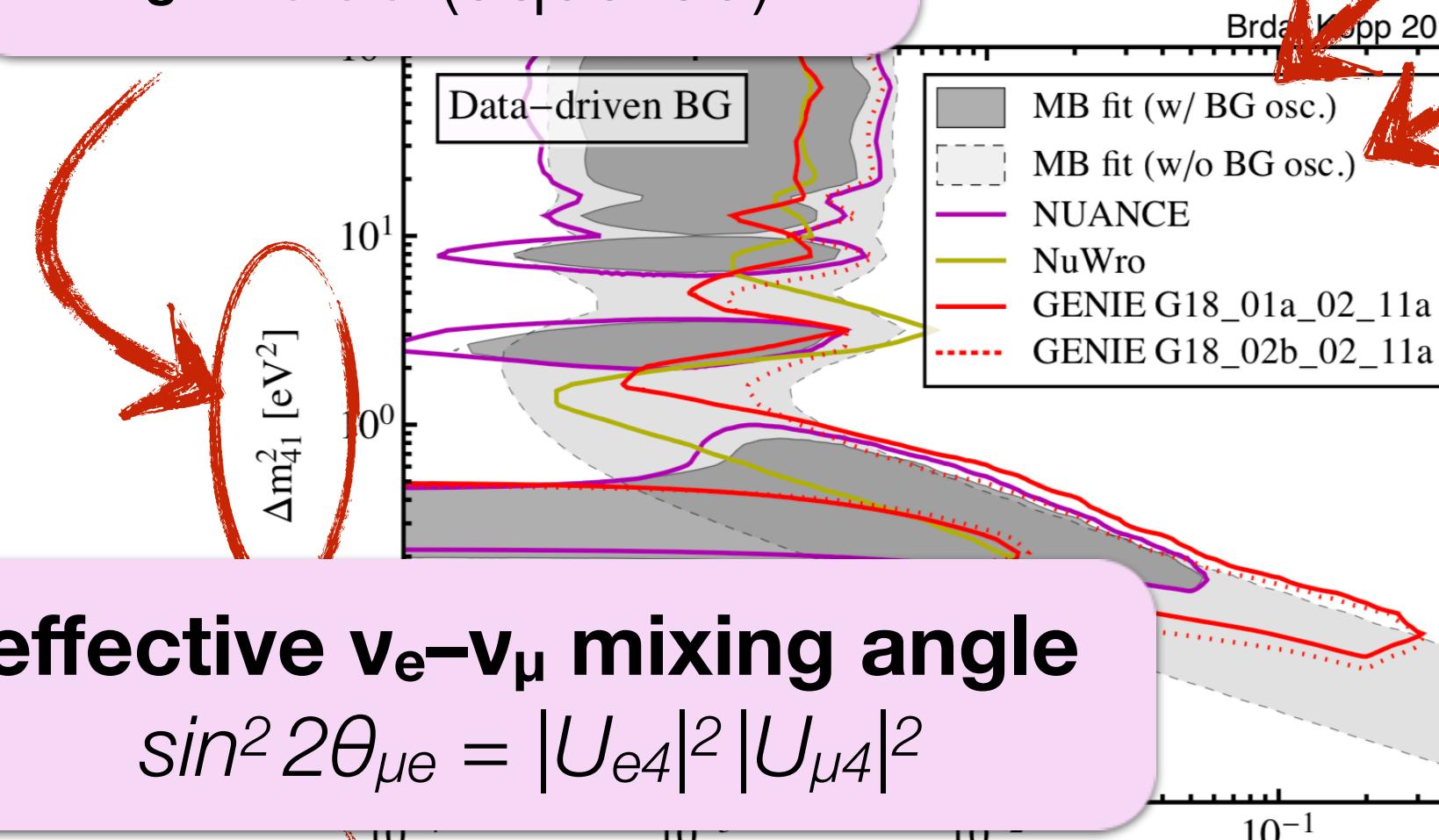
# 3+1 Models in MiniBooNE – Comparison of Generators

$\nu_e$  spectrum

(where the excess is)



$\nu_s$  mass (squared)

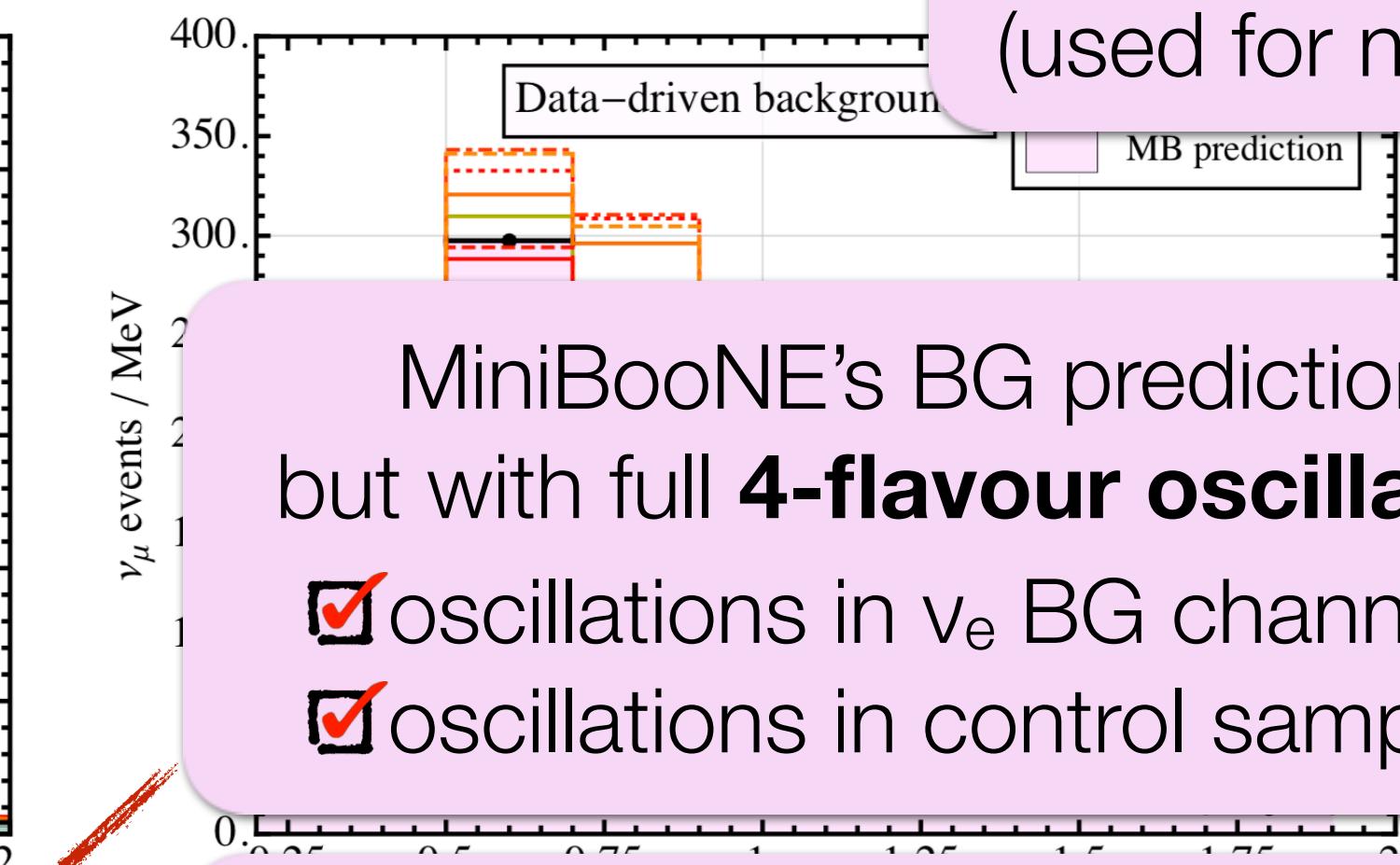


$$\sin^2 2\theta_{\mu e} = |U_{e4}|^2 |U_{\mu 4}|^2$$



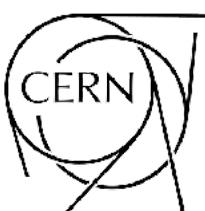
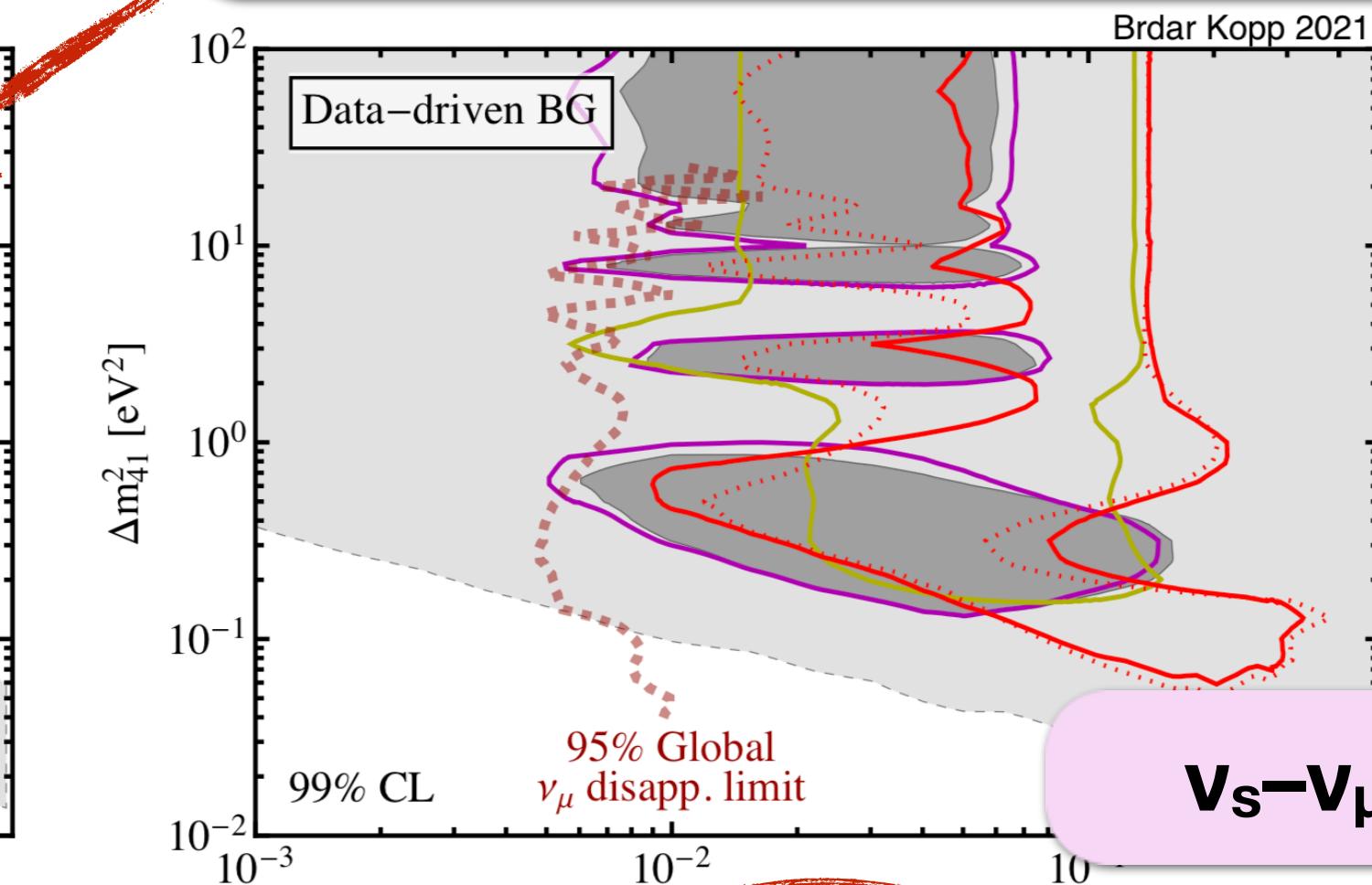
$\nu_\mu$  spectrum

(used for normalization)



- ✓ oscillations in  $\nu_e$  BG channels
- ✓ oscillations in control sample

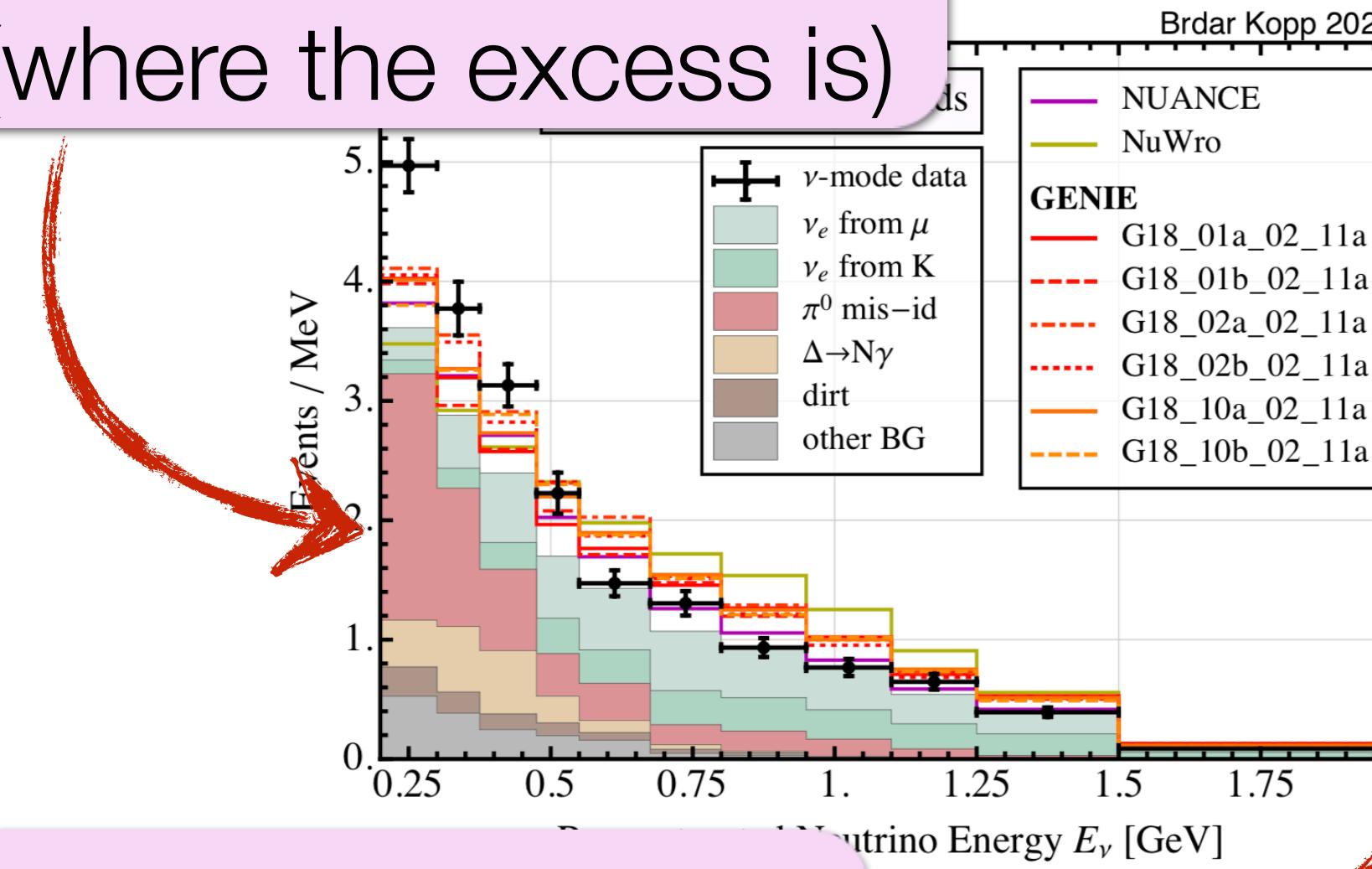
MiniBooNE's fit (2-flavour oscillations)



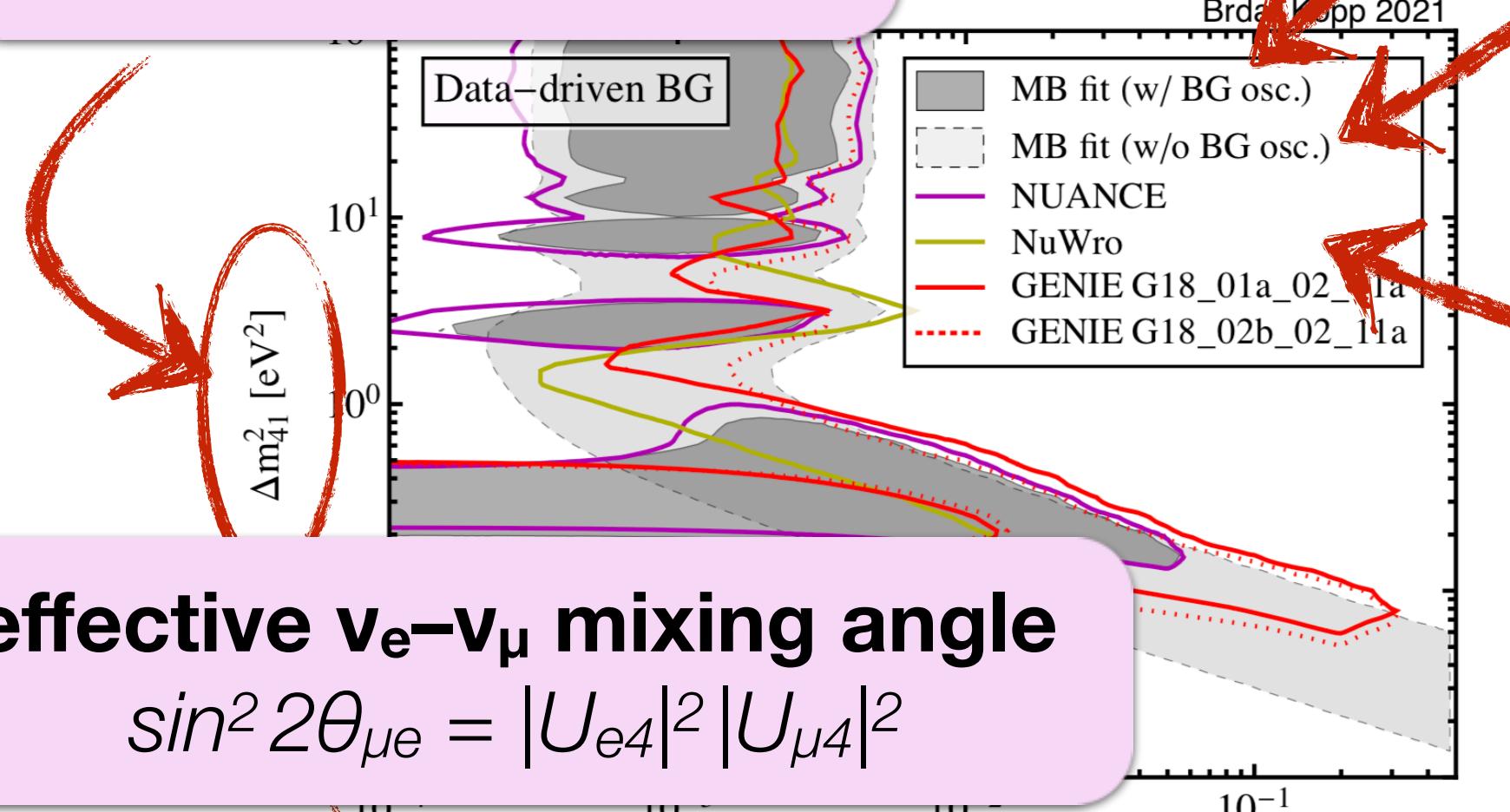
# 3+1 Models in MiniBooNE – Comparison of Generators

$\nu_e$  spectrum

(where the excess is)



$\nu_s$  mass (squared)



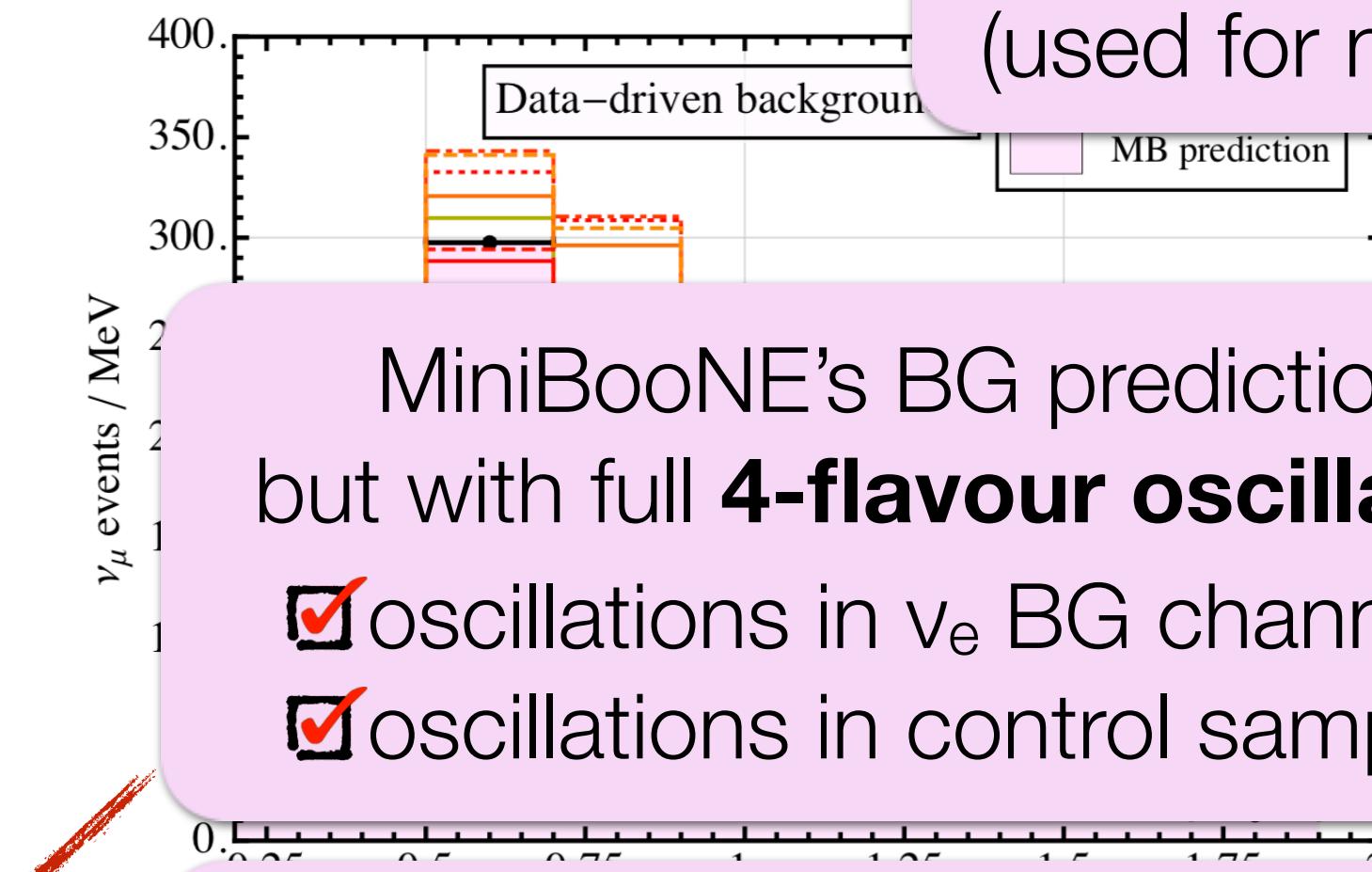
effective  $\nu_e$ - $\nu_\mu$  mixing angle

$$\sin^2 2\theta_{\mu e} = |U_{e4}|^2 |U_{\mu 4}|^2$$



$\nu_\mu$  spectrum

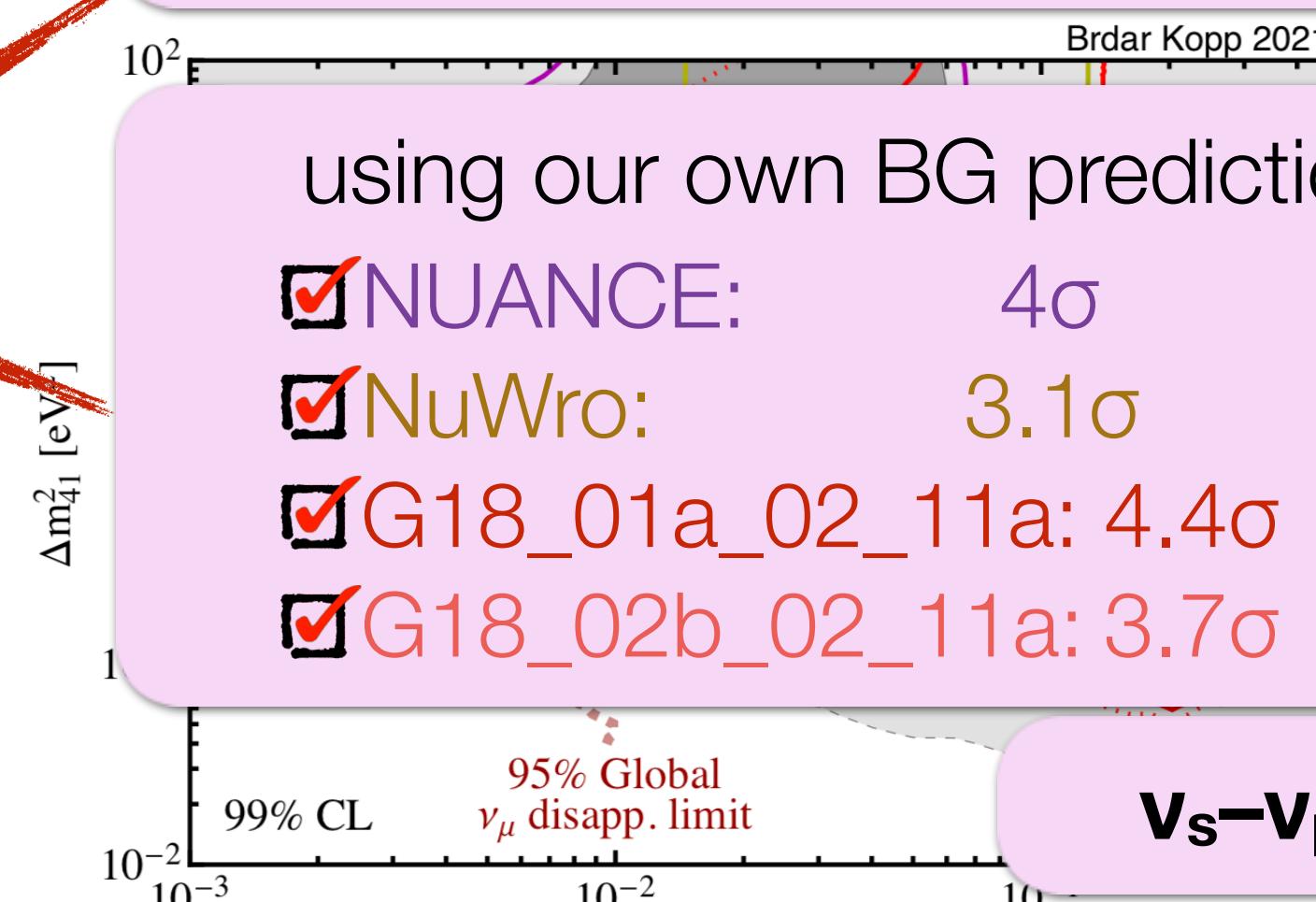
(used for normalization)



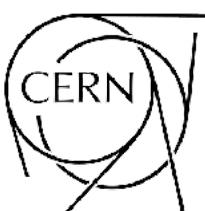
MiniBooNE's BG predictions,  
but with full **4-flavour oscillations**

- oscillations in  $\nu_e$  BG channels
- oscillations in control sample

**MiniBooNE's fit** (2-flavour oscillations)



$\nu_s$ - $\nu_\mu$  mixing

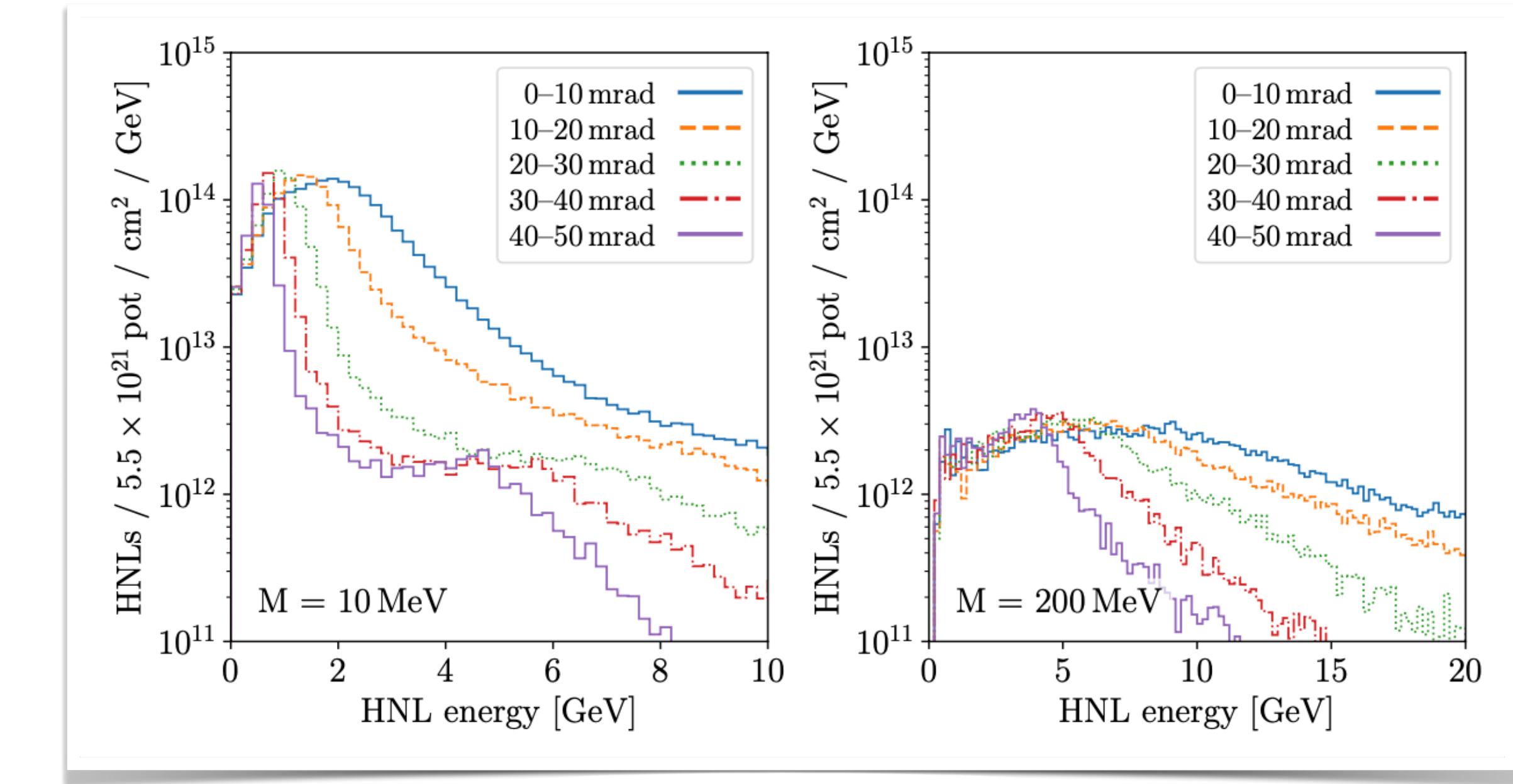


# Heavy Neutral Leptons

- Neutrino portal leads to mixing between  $\nu$  and heavy sterile neutrino  $N$

$$\mathcal{L} \supset y \bar{L} (i\sigma^2 H^*) N$$

- any process that makes  $\nu$  in the SM can also make  $N$  (suppressed by a mixing angle)
- production in meson decays

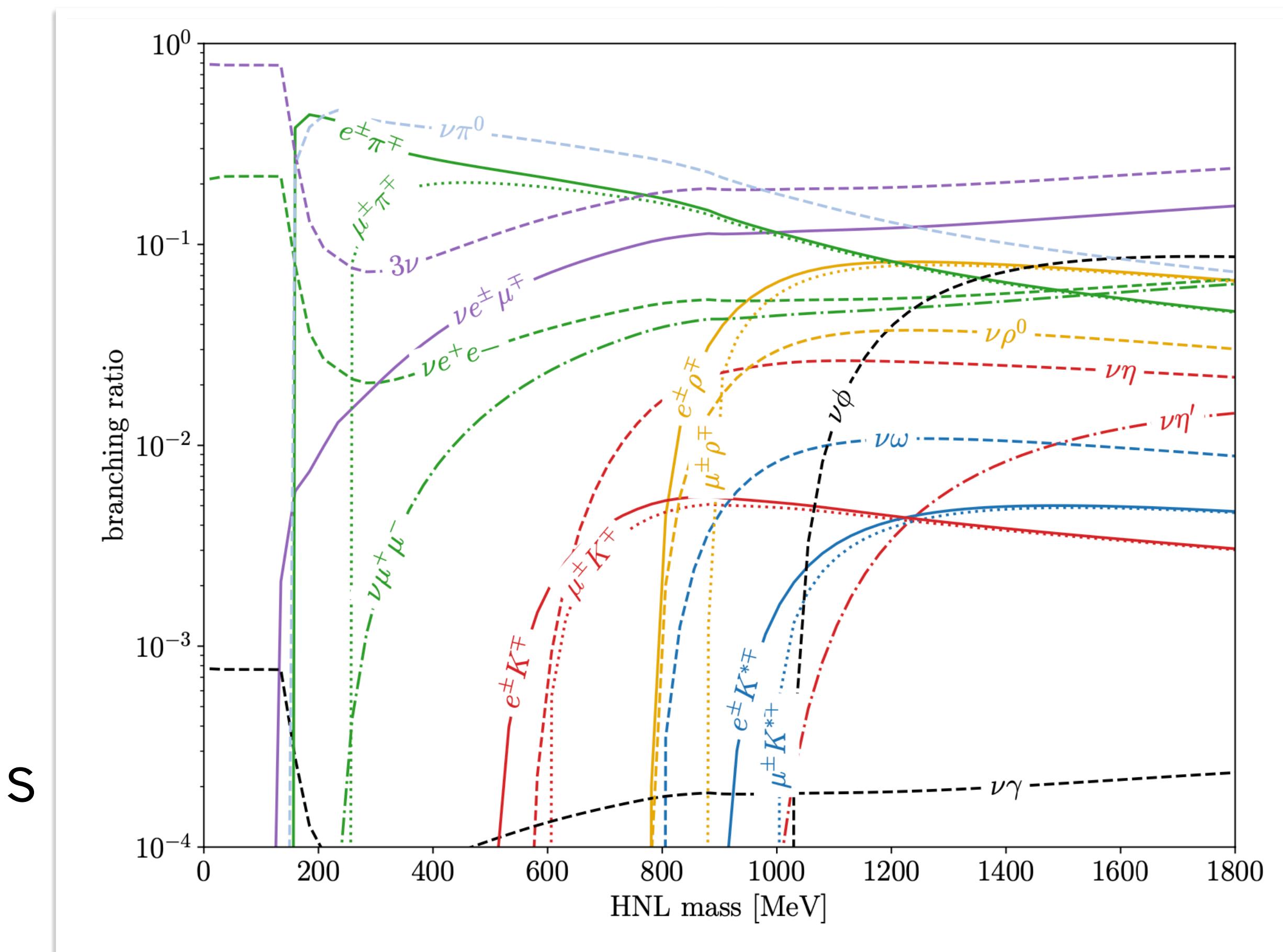


# Heavy Neutral Leptons

- Neutrino portal leads to mixing between  $\nu$  and heavy sterile neutrino  $N$

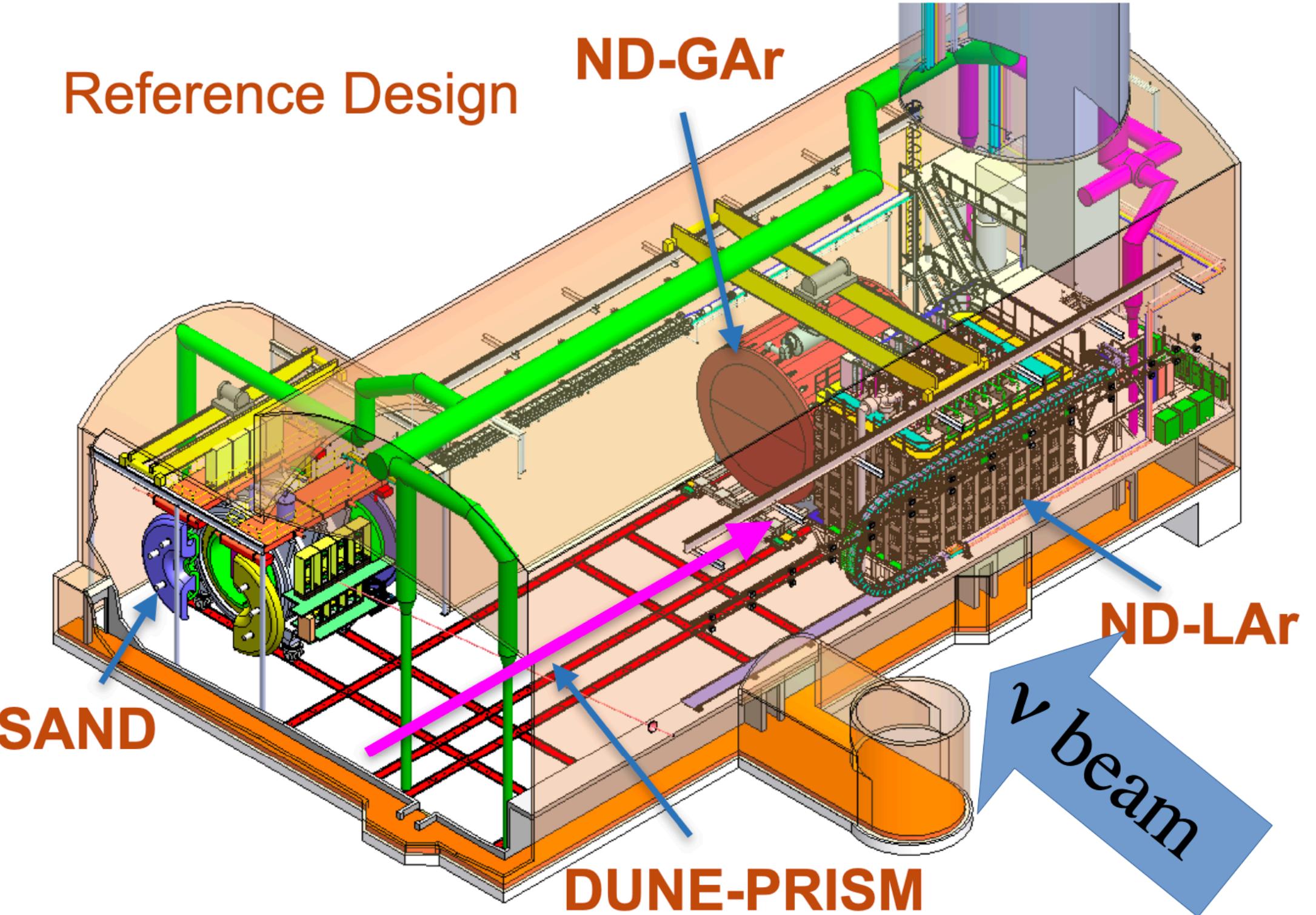
$$\mathcal{L} \supset y \bar{L} (i\sigma^2 H^*) N$$

- any process that makes  $\nu$  in the SM can also make  $N$  (suppressed by a mixing angle)
- production in meson decays
- many 2-body and 3-body decay channels
- background estimates require detailed understanding of **exclusive final states**



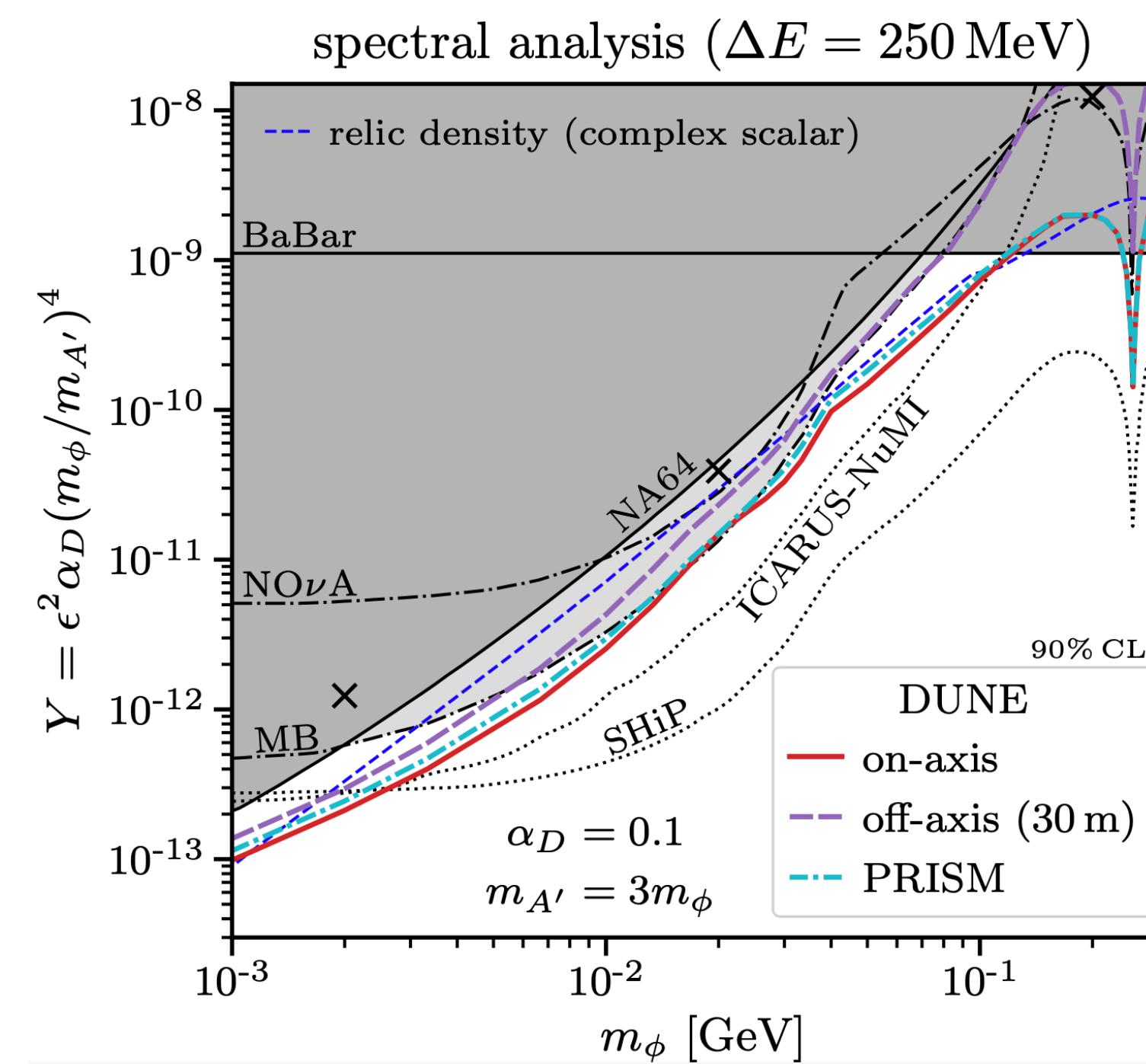
# DUNE-PRISM

- Movable near detectors measure event rates at different off-axis angles  
→ disentangle flux and x-sec uncertainties
- Interesting opportunities for BSM searches
  - heavy particle less boosted
  - backgrounds lower off-axis

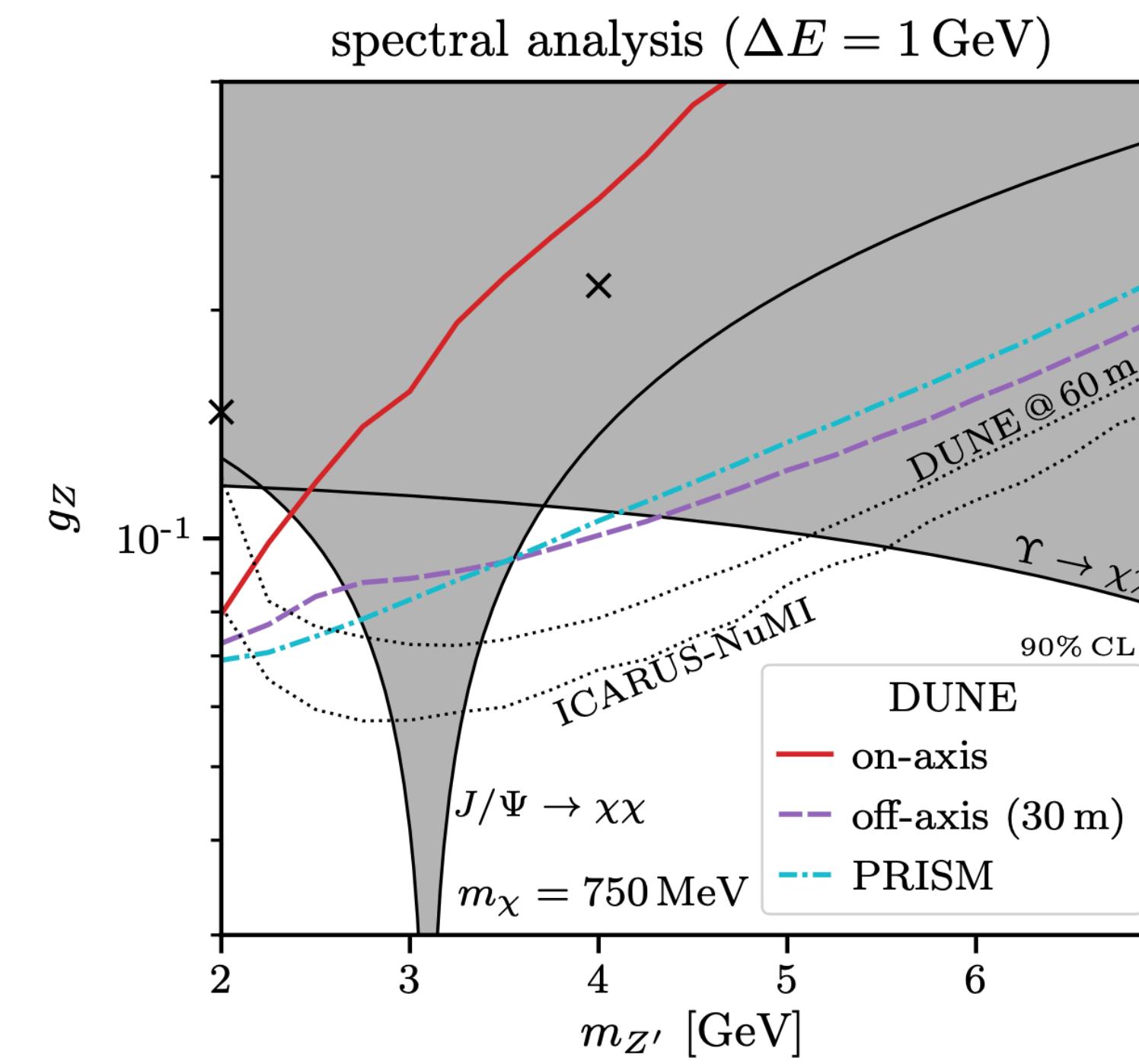


# DUNE-PRISM

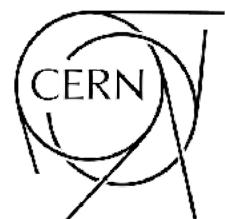
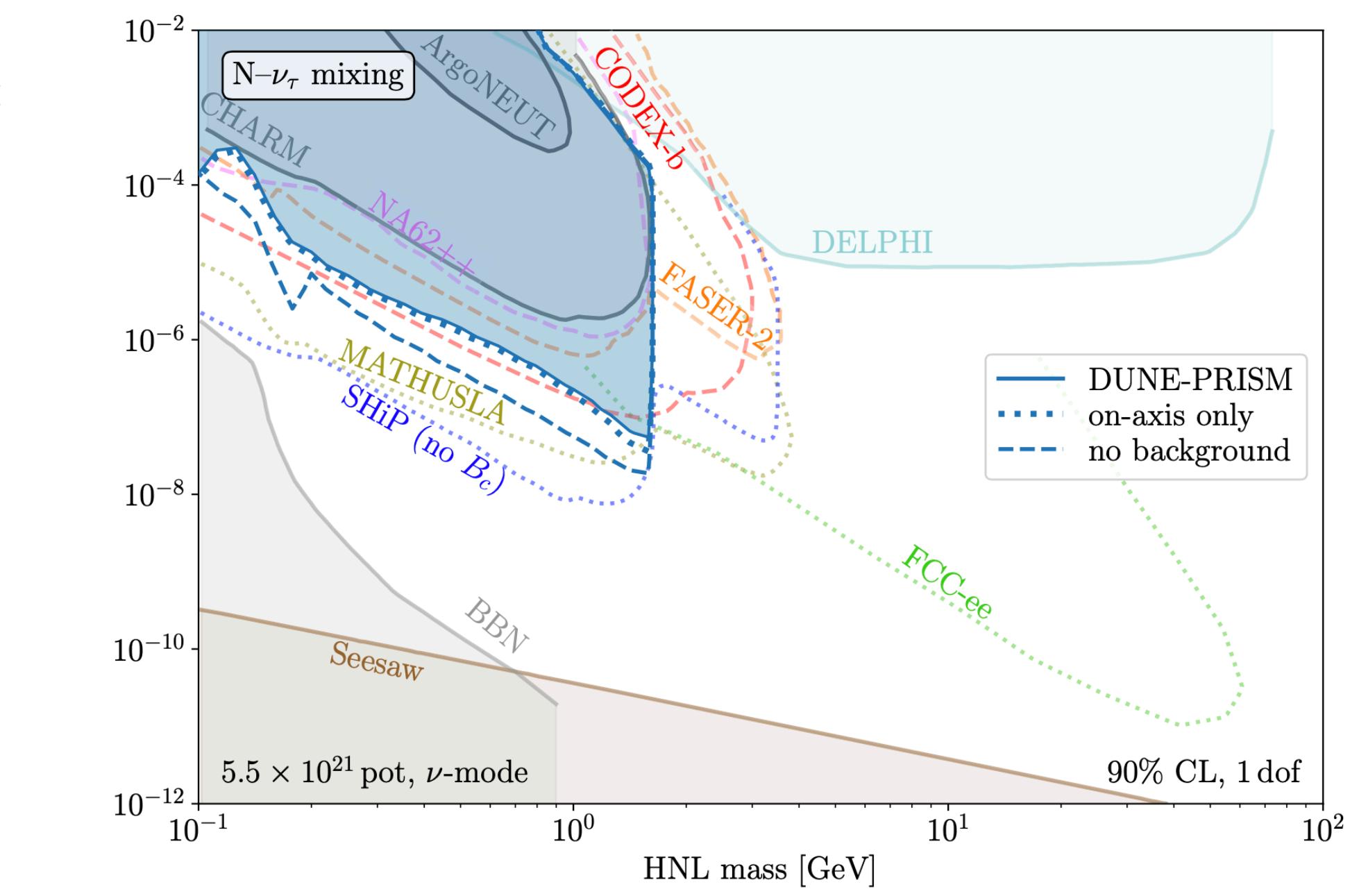
## A'-mediated DM



## leptophobic DM

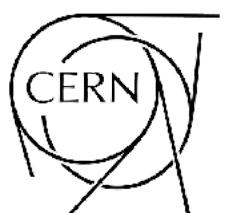
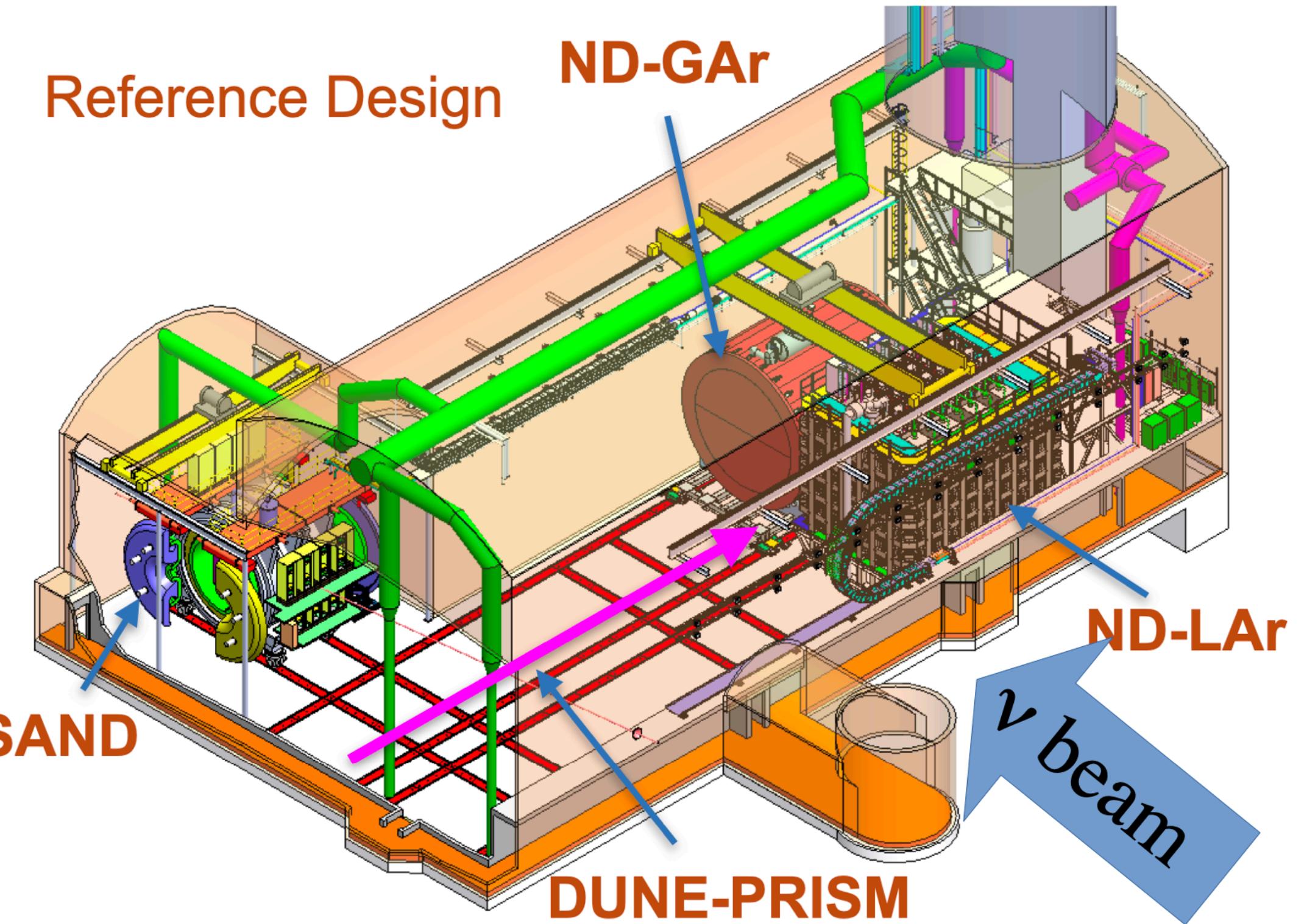


## heavy neutral leptons



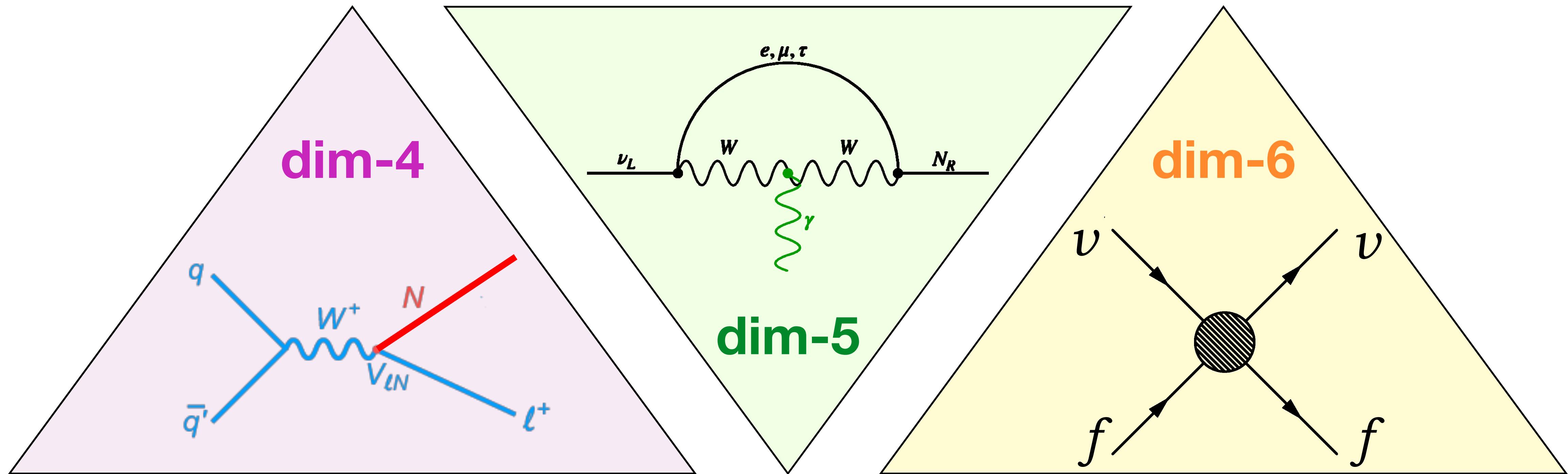
# DUNE-PRISM

- Movable near detectors measure event rates at different off-axis angles  
→ disentangle flux and x-sec uncertainties
- Interesting opportunities for BSM searches
  - heavy particle less boosted
  - backgrounds lower off-axis
- BSM searches are typically robust w.r.t. running strategy



# Neutrino Physics Beyond the Standard Model

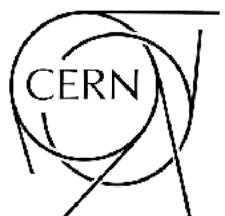
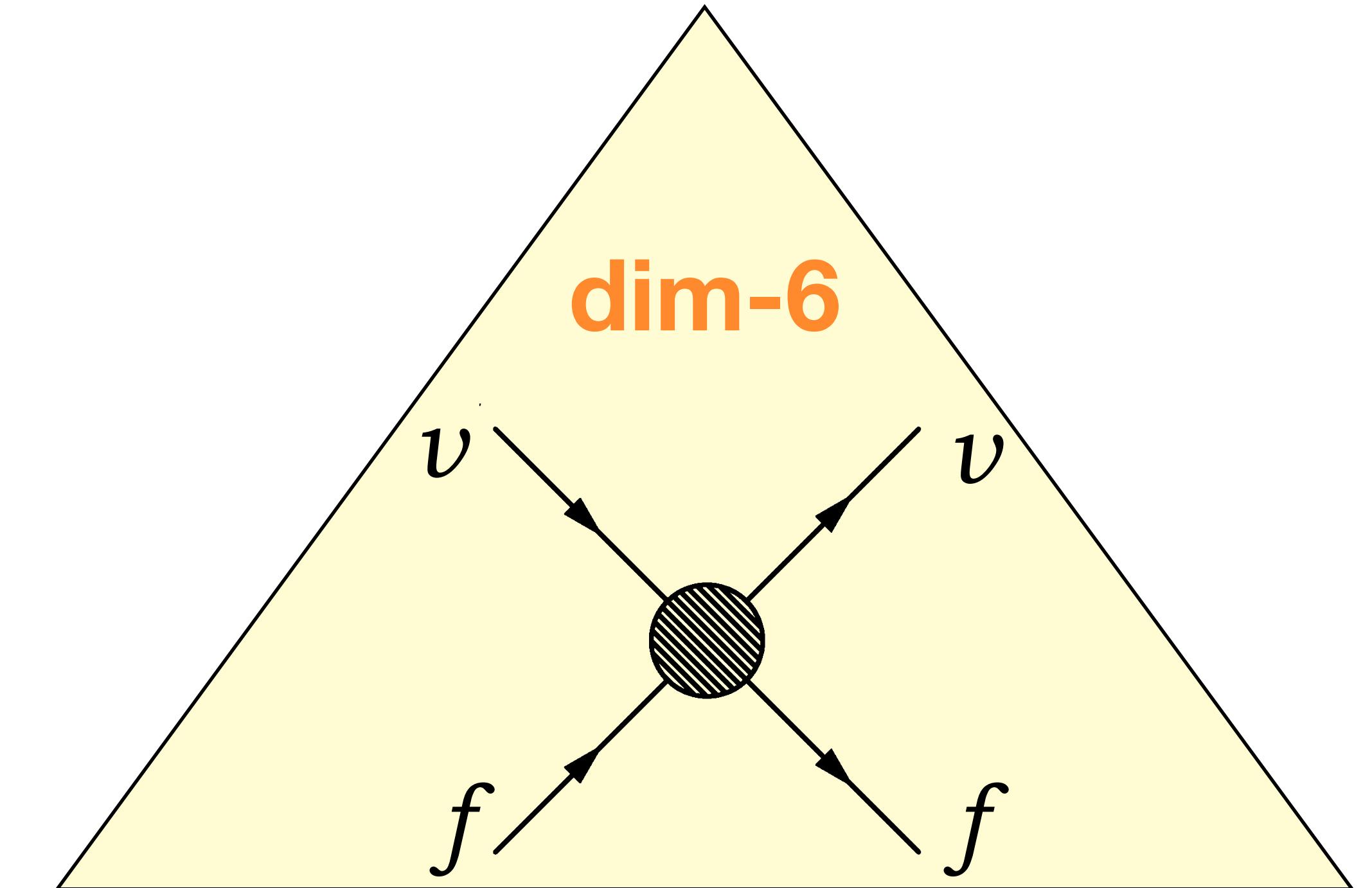
e.g. neutrino magnetic moments



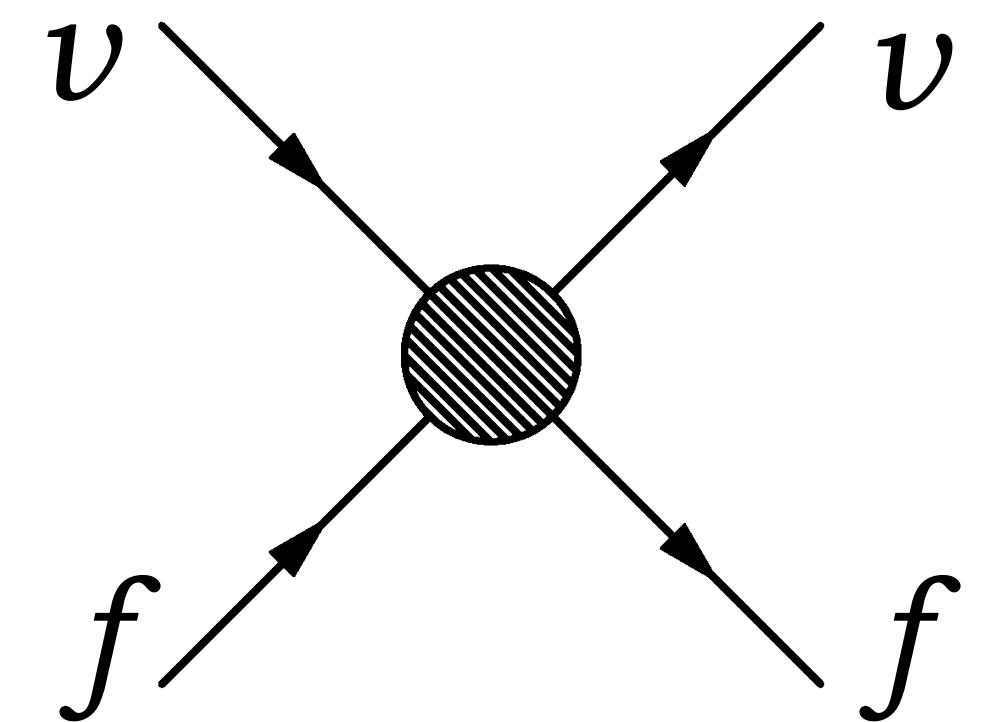
e.g. sterile neutrinos

e.g. non-standard interactions

# Effective Field Theories for Neutrinos



# Effective Field Theories for Neutrinos



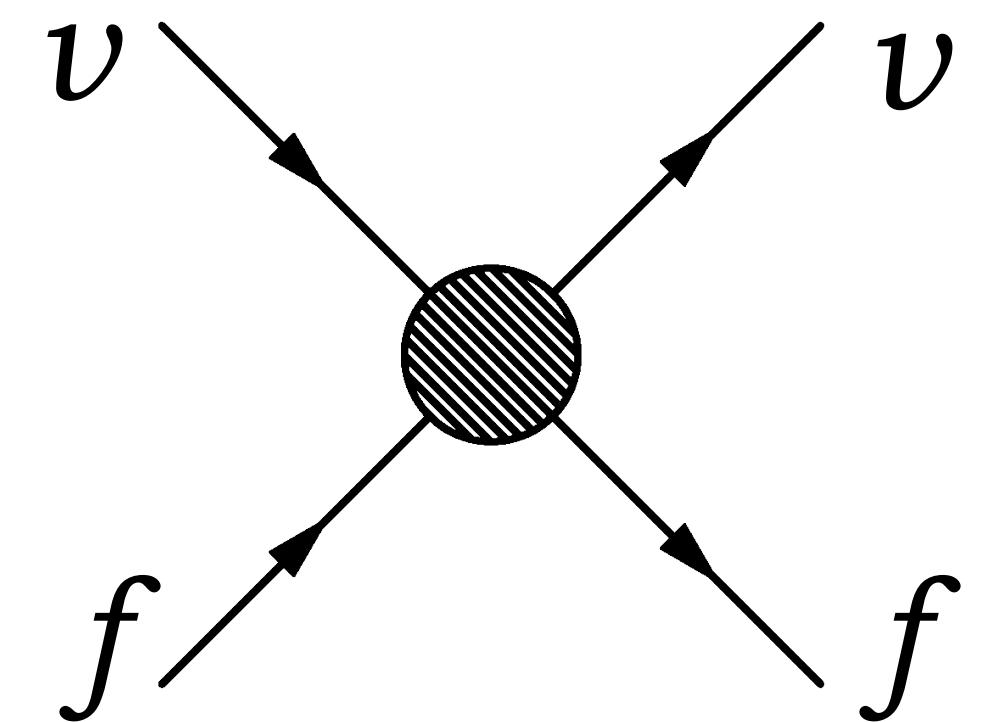
$$\begin{aligned} \mathcal{L}_{\text{WEFT}} \supset & -\frac{2V_{ud}}{v^2} \left\{ [1 + \epsilon_L]_{\alpha\beta} (\bar{q}_u \gamma^\mu P_L q_d) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) + [\epsilon_R]_{\alpha\beta} (\bar{q}_u \gamma^\mu P_R q_d) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) \right. \\ & + \frac{1}{2} [\epsilon_S]_{\alpha\beta} (\bar{q}_u q_d) (\bar{\ell}_\alpha P_L \nu_\beta) - \frac{1}{2} [\epsilon_P]_{\alpha\beta} (\bar{q}_u \gamma_5 q_d) (\bar{\ell}_\alpha P_L \nu_\beta) \\ & \left. + \frac{1}{4} [\epsilon_T]_{\alpha\beta} (\bar{q}_u \sigma^{\mu\nu} P_L q_d) (\bar{\ell}_\alpha \sigma_{\mu\nu} P_L \nu_\beta) + \text{h.c.} \right\}. \end{aligned}$$

# Effective Field Theories for Neutrinos

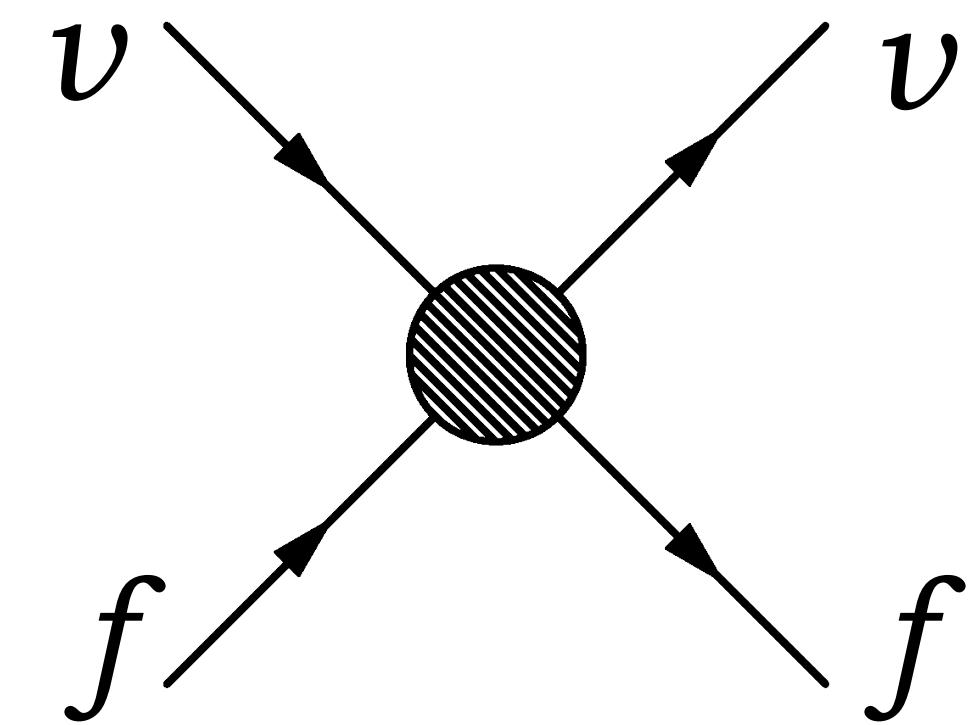
**dimensionless coefficients**

(interaction strength  
relative to SM weak interactions)

$$\begin{aligned} \mathcal{L}_{\text{WEFT}} \supset & -\frac{2V_{ud}}{v^2} \left\{ [1 - \epsilon_L]_{\alpha\beta} (\bar{q}_u \gamma^\mu P_L q_d) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) + [\epsilon_R]_{\alpha\beta} (\bar{q}_u \gamma^\mu P_R q_d) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) \right. \\ & + \frac{1}{2} [\epsilon_S]_{\alpha\beta} (\bar{q}_u q_d) (\bar{\ell}_\alpha P_L \nu_\beta) - \frac{1}{2} [\epsilon_P]_{\alpha\beta} (\bar{q}_u \gamma_5 q_d) (\bar{\ell}_\alpha P_L \nu_\beta) \\ & \left. + \frac{1}{4} [\epsilon_T]_{\alpha\beta} (\bar{q}_u \sigma^{\mu\nu} P_L q_d) (\bar{\ell}_\alpha \sigma_{\mu\nu} P_L \nu_\beta) + \text{h.c.} \right\}. \end{aligned}$$



# Effective Field Theories for Neutrinos



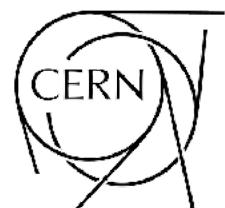
**dimensionless coefficients**

(interaction strength  
relative to SM weak interactions)

$$\begin{aligned} \mathcal{L}_{\text{WEFT}} \supset & -\frac{2V_{ud}}{v^2} \left\{ [1 - \epsilon_L]_{\alpha\beta} (\bar{q}_u \gamma^\mu P_L q_d) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) + [\epsilon_R]_{\alpha\beta} (\bar{q}_u \gamma^\mu P_R q_d) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) \right. \\ & + \frac{1}{2} [\epsilon_S]_{\alpha\beta} (\bar{q}_u q_d) (\bar{\ell}_\alpha P_L \nu_\beta) - \frac{1}{2} [\epsilon_P]_{\alpha\beta} (\bar{q}_u \gamma_5 q_d) (\bar{\ell}_\alpha P_L \nu_\beta) \\ & \left. + \frac{1}{4} [\epsilon_T]_{\alpha\beta} (\bar{q}_u \sigma^{\mu\nu} P_L q_d) (\bar{\ell}_\alpha \sigma_{\mu\nu} P_L \nu_\beta) + \text{h.c.} \right\}. \end{aligned}$$

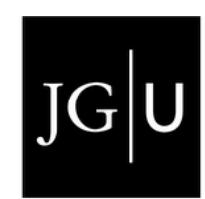
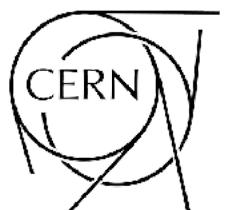
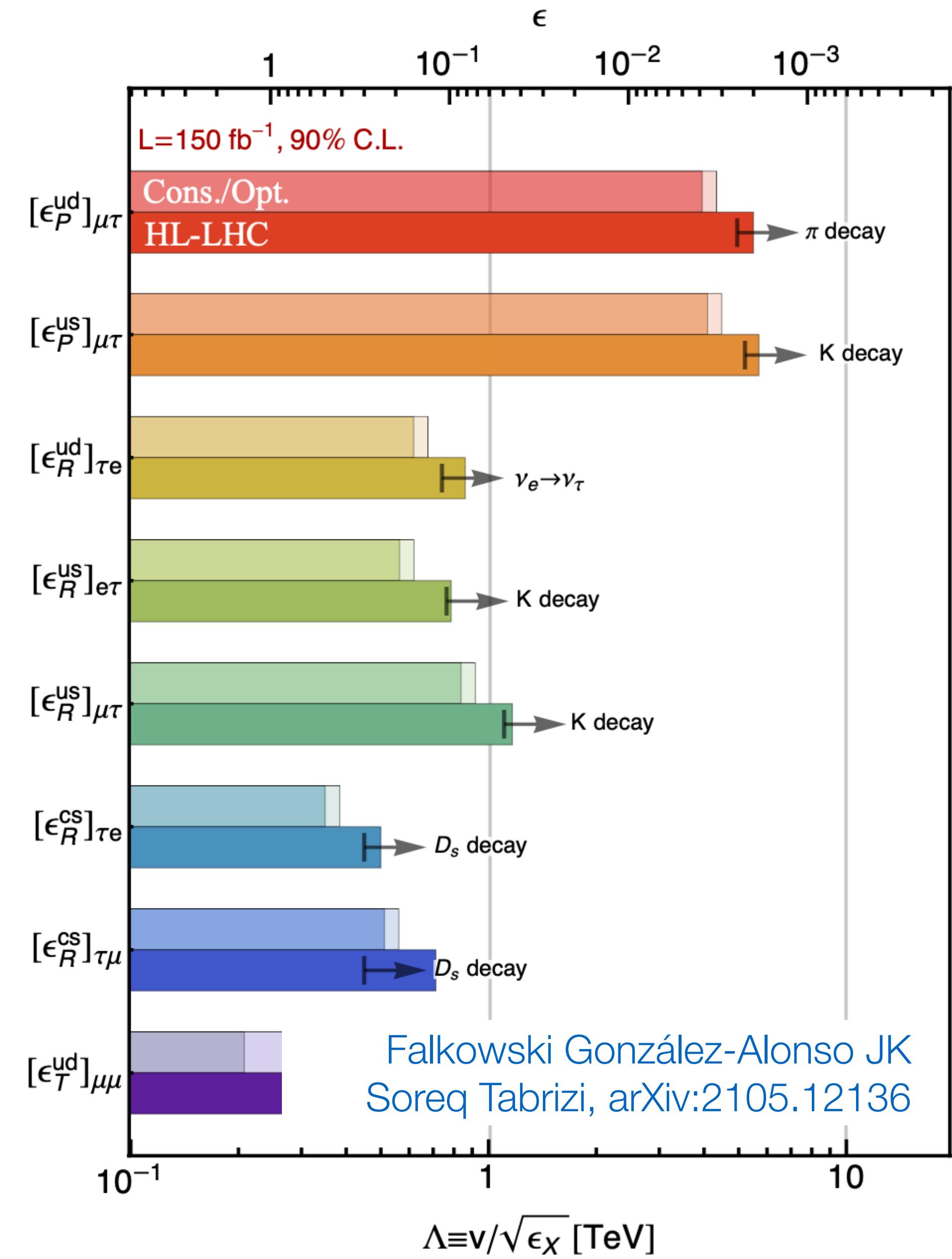
**dim-6 operators**

with different Lorentz structures



# New Interactions in DIS

- straightforward – ν scattering described in terms of PDFs
- though flux uncertainties remain a problem

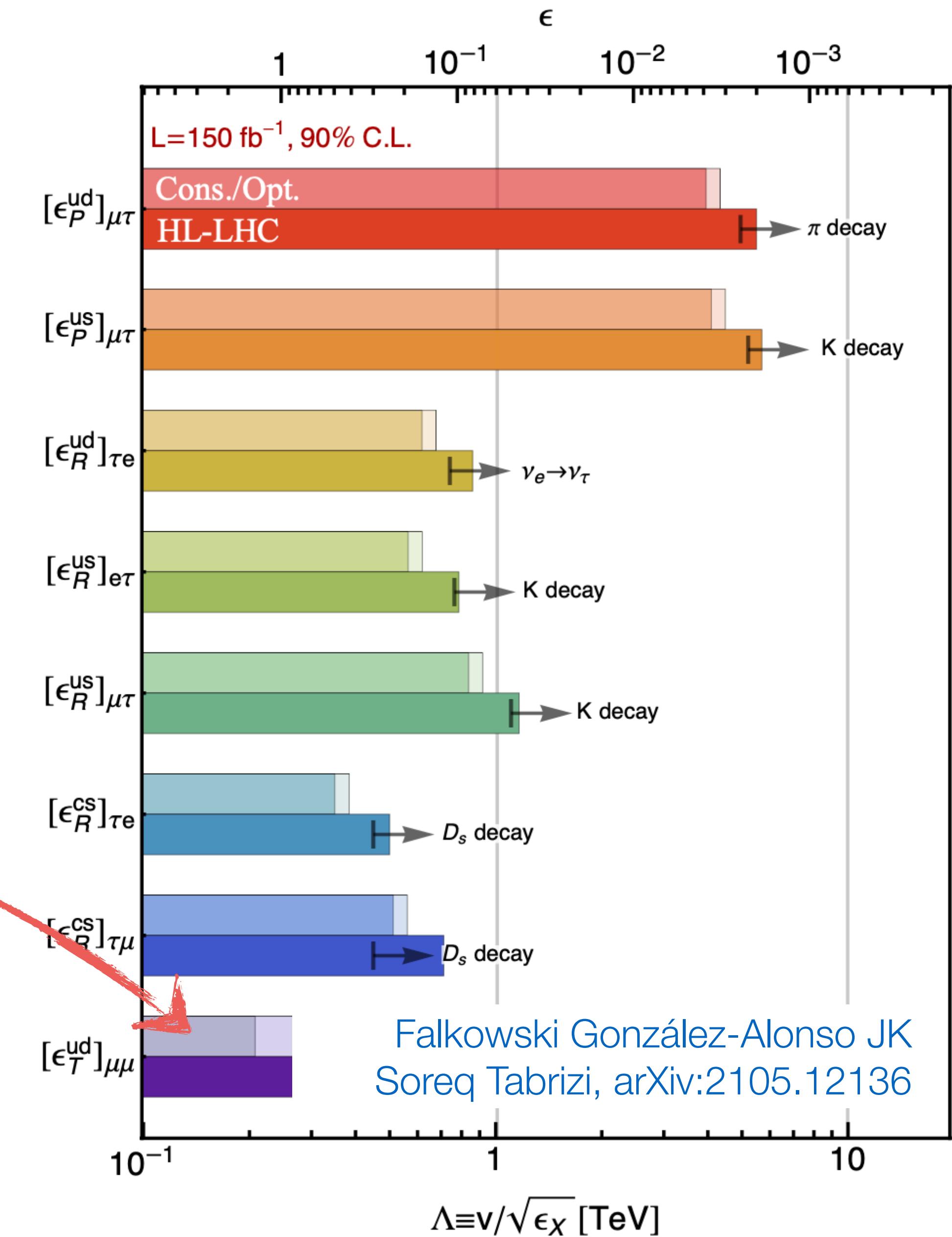


# New Interactions in DIS

- straightforward –  $\nu$  scattering described in terms of PDFs
- though flux uncertainties remain a problem

## conservative scenario

flux uncertainties 30% / 40% / 50%  
for  $\nu_e$  /  $\nu_\mu$  /  $\nu_\tau$



# New Interactions in DIS

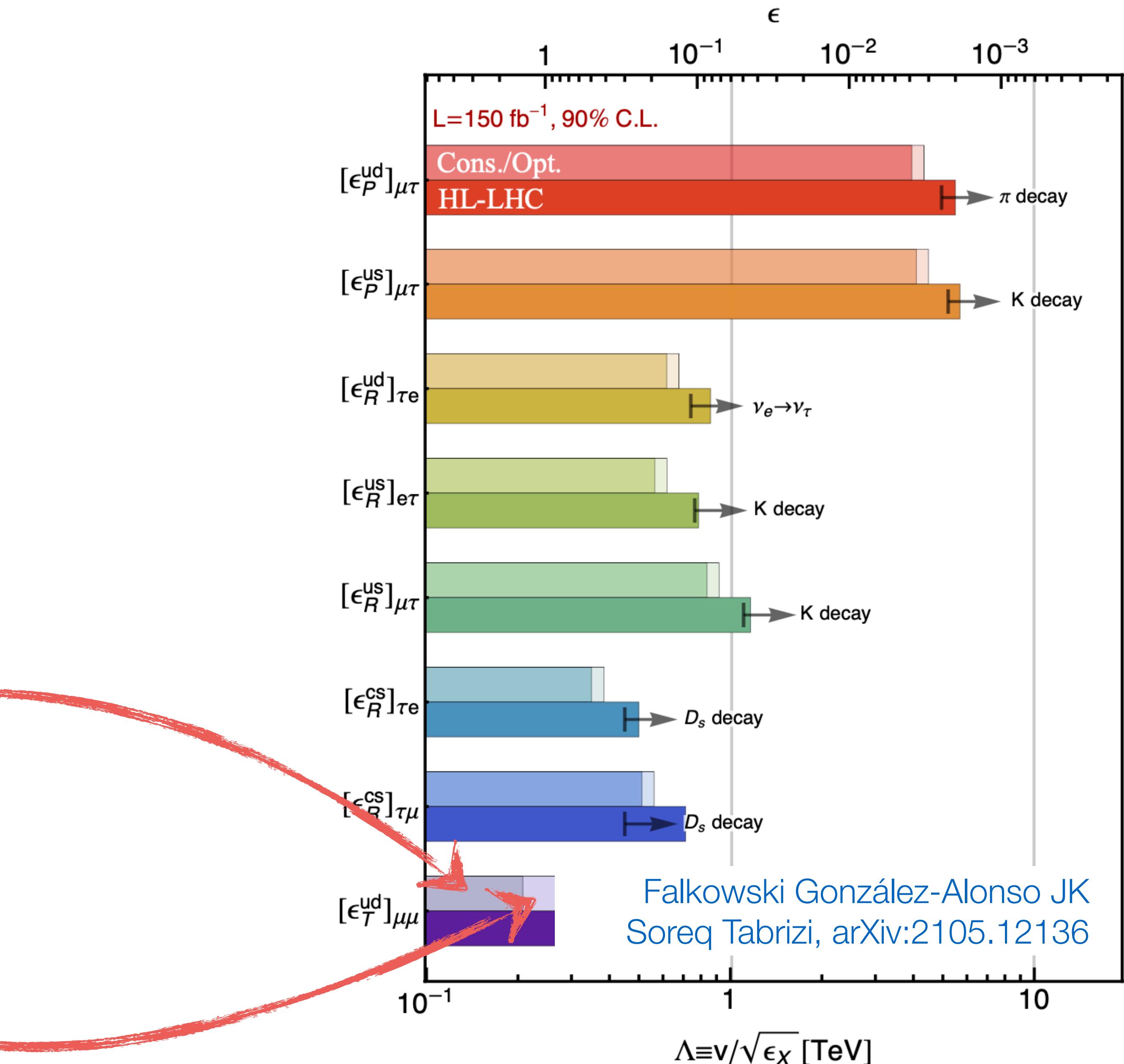
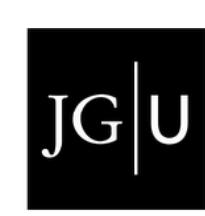
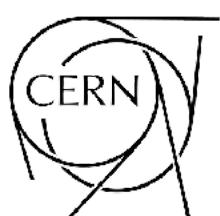
- straightforward – ν scattering described in terms of PDFs
- though flux uncertainties remain a problem

## conservative scenario

flux uncertainties 30% / 40% / 50%  
for  $\nu_e$  /  $\nu_\mu$  /  $\nu_\tau$

## optimistic scenario

flux uncertainties 5% / 10% / 15%  
for  $\nu_e$  /  $\nu_\mu$  /  $\nu_\tau$



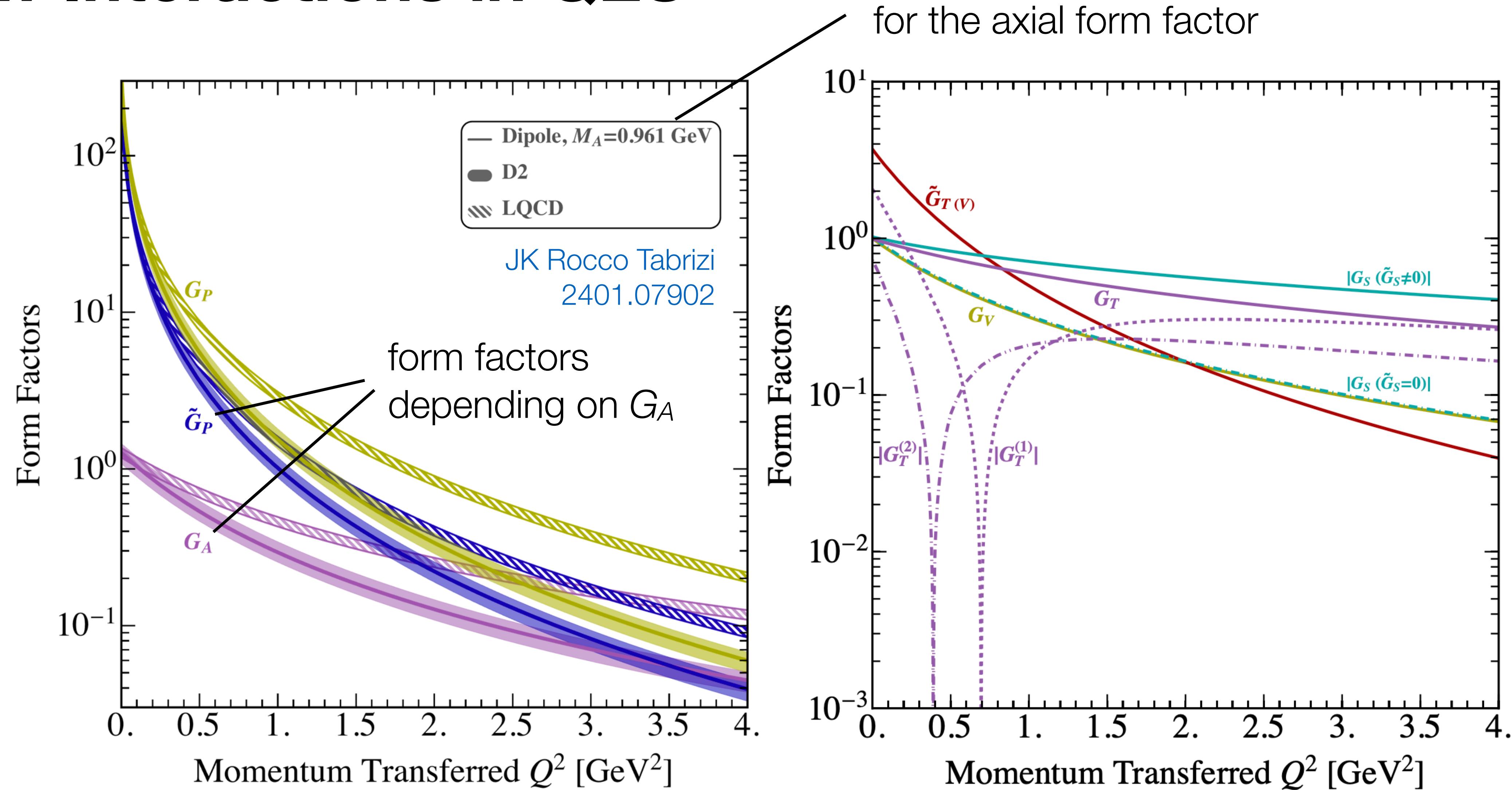
# New Interactions in QES

- For V–A: same uncertainties as in the SM
- For other Lorentz structures: significant nucleon form factor uncertainties

$$\langle p(p_p) | \bar{q}_u \gamma_\mu q_d | n(p_n) \rangle = \bar{u}_p(p_p) \left[ G_V(Q^2) \gamma_\mu + i \frac{\tilde{G}_{T(V)}(Q^2)}{2M_N} \sigma_{\mu\nu} q^\nu - \frac{\tilde{G}_S(Q^2)}{2M_N} q_\mu \right] u_n(p_n),$$
$$\langle p(p_p) | \bar{q}_u \gamma_\mu \gamma_5 q_d | n(p_n) \rangle = \bar{u}_p(p_p) \left[ G_A(Q^2) \gamma_\mu \gamma_5 + i \frac{\tilde{G}_{T(A)}(Q^2)}{2M_N} \sigma_{\mu\nu} q^\nu \gamma_5 - \frac{\tilde{G}_P(Q^2)}{2M_N} q_\mu \gamma_5 \right] u_n(p_n),$$
$$\langle p(p_p) | \bar{q}_u q_d | n(p_n) \rangle = G_S(Q^2) \bar{u}_p(p_p) u_n(p_n),$$
$$\langle p(p_p) | \bar{q}_u \gamma_5 q_d | n(p_n) \rangle = G_P(Q^2) \bar{u}_p(p_p) \gamma_5 u_n(p_n),$$
$$\langle p(p_p) | \bar{q}_u \sigma_{\mu\nu} q_d | n(p_n) \rangle = \bar{u}_p(p_p) \left[ G_T(Q^2) \sigma_{\mu\nu} - \frac{i}{M_N} G_T^{(1)}(Q^2) (q_\mu \gamma_\nu - q_\nu \gamma_\mu) \right. \\ \left. - \frac{i}{M_N^2} G_T^{(2)}(Q^2) (q_\mu P_\nu - q_\nu P_\mu) - \frac{i}{M_N} G_T^{(3)}(Q^2) (\gamma_\mu \not{q} \gamma_\nu - \gamma_\nu \not{q} \gamma_\mu) \right] u_n(p_n),$$

# New Interactions in QES

three different estimates  
for the axial form factor



# New Interactions in QES

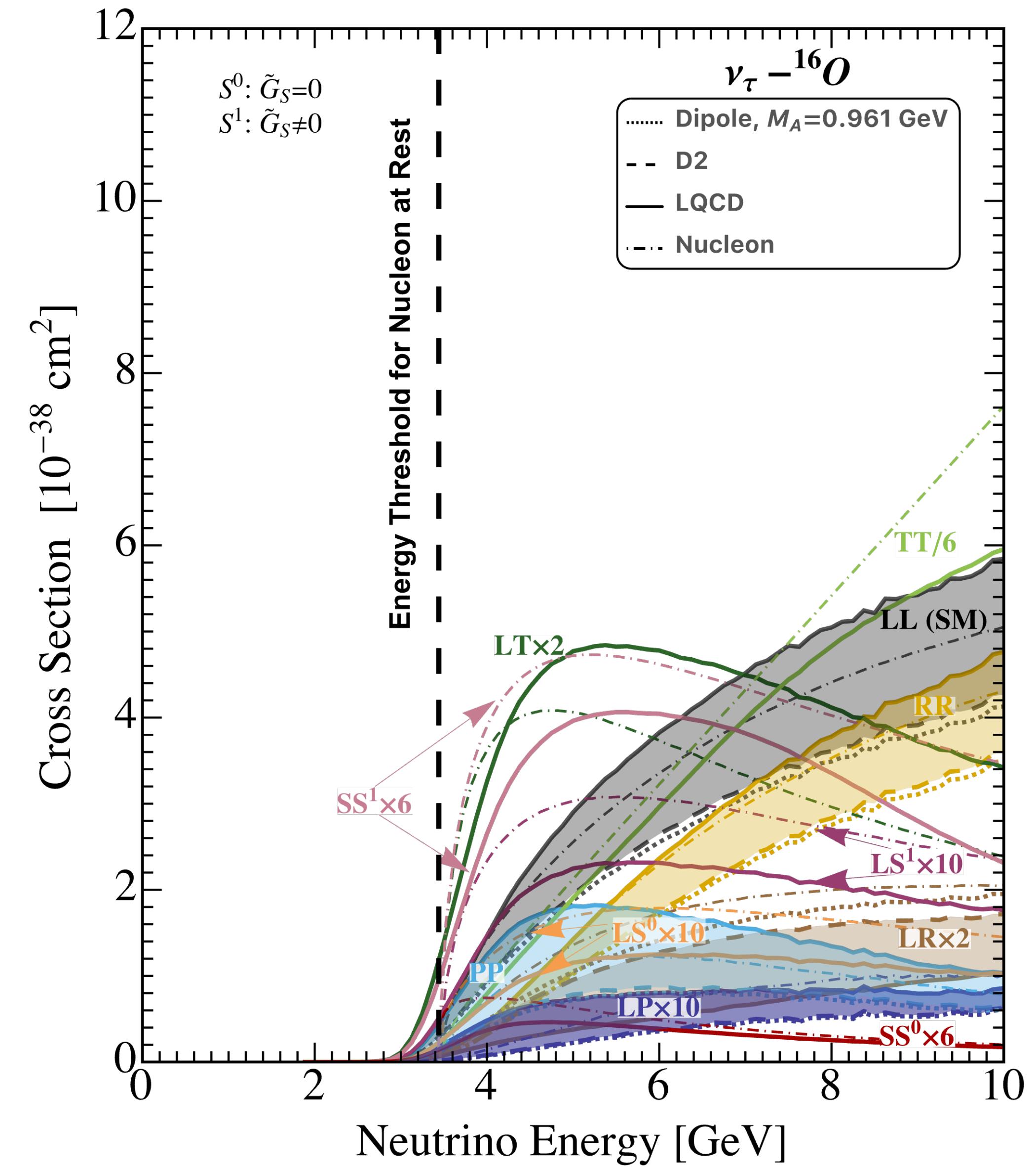
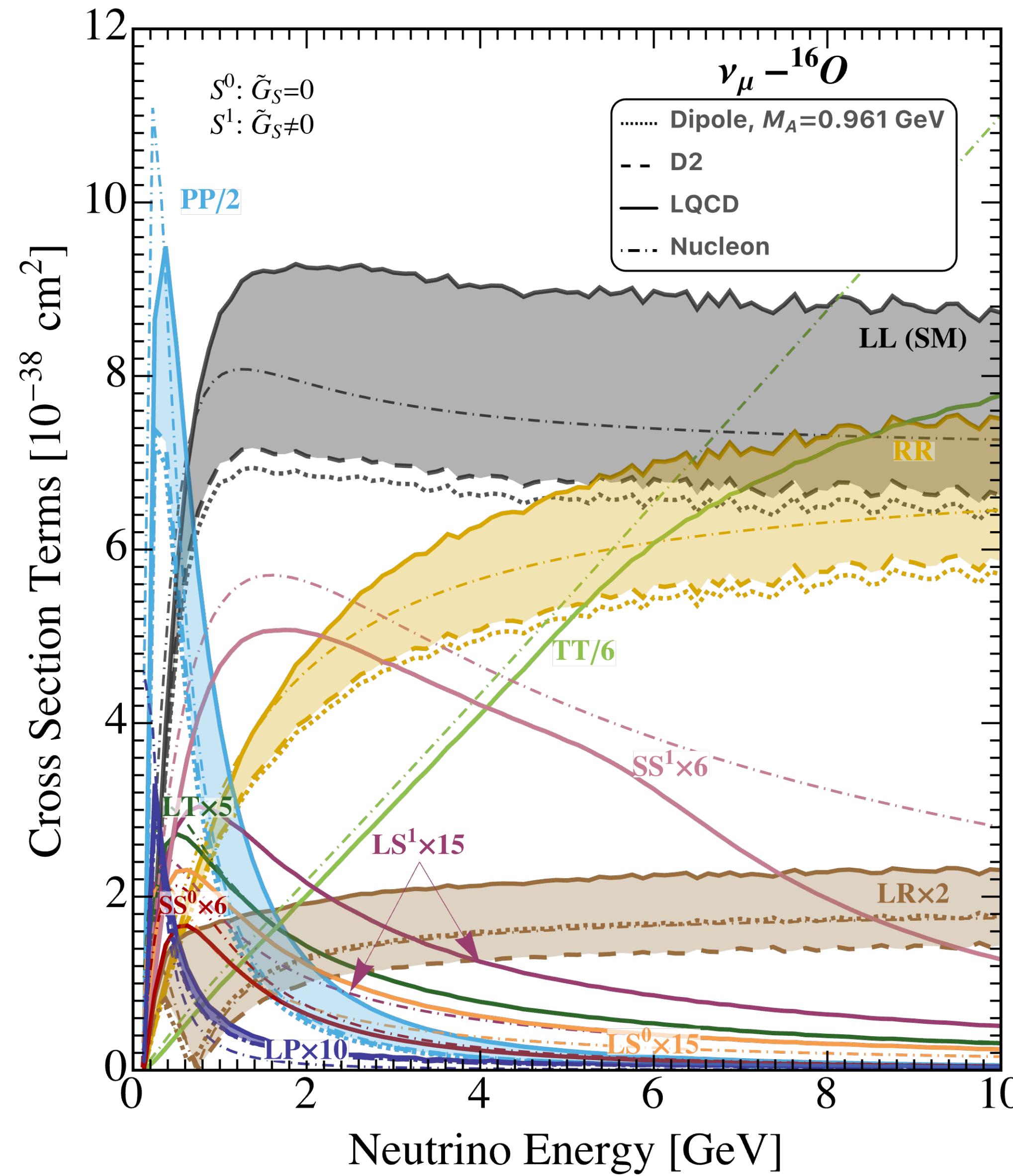
- in addition: nuclear effects, parameterized via a spectral function

$$\begin{aligned} \frac{1}{2} \sum_{\text{spin}} \mathcal{A}_{X,\alpha} \mathcal{A}_{Y,\alpha}^* &= \frac{1}{2} \int \frac{d^3 p_N}{(2\pi)^3} P_h(\mathbf{p}_N, E^*) \frac{m_N}{e(\mathbf{p}_N)} \frac{m_N}{e(\mathbf{q} + \mathbf{p}_N)} \\ &\times \sum_{\text{spin}} \sum_N A_{X,\alpha} A_{Y,\alpha}^* \delta(\tilde{\omega} + e(\mathbf{p}_N) - e(\mathbf{q} + \mathbf{p}_N)), \end{aligned}$$

- cross-section decomposed into contributions with different Lorentz structures

$$\frac{d\sigma_{\alpha\beta}}{dQ^2} = \frac{d\hat{\sigma}_{LL,\alpha}}{dQ^2} \delta_{\alpha\beta} + \sum_X \left( [\epsilon_X]_{\alpha\beta} \frac{d\hat{\sigma}_{LX,\alpha}}{dQ^2} \delta_{\alpha\beta} + h.c. \right) + \sum_{X,Y,\beta} [\epsilon_X]_{\alpha\beta} [\epsilon_Y]_{\alpha\beta}^* \frac{d\hat{\sigma}_{XY,\alpha}}{dQ^2}$$

# New Interactions in QES





# Thank You!



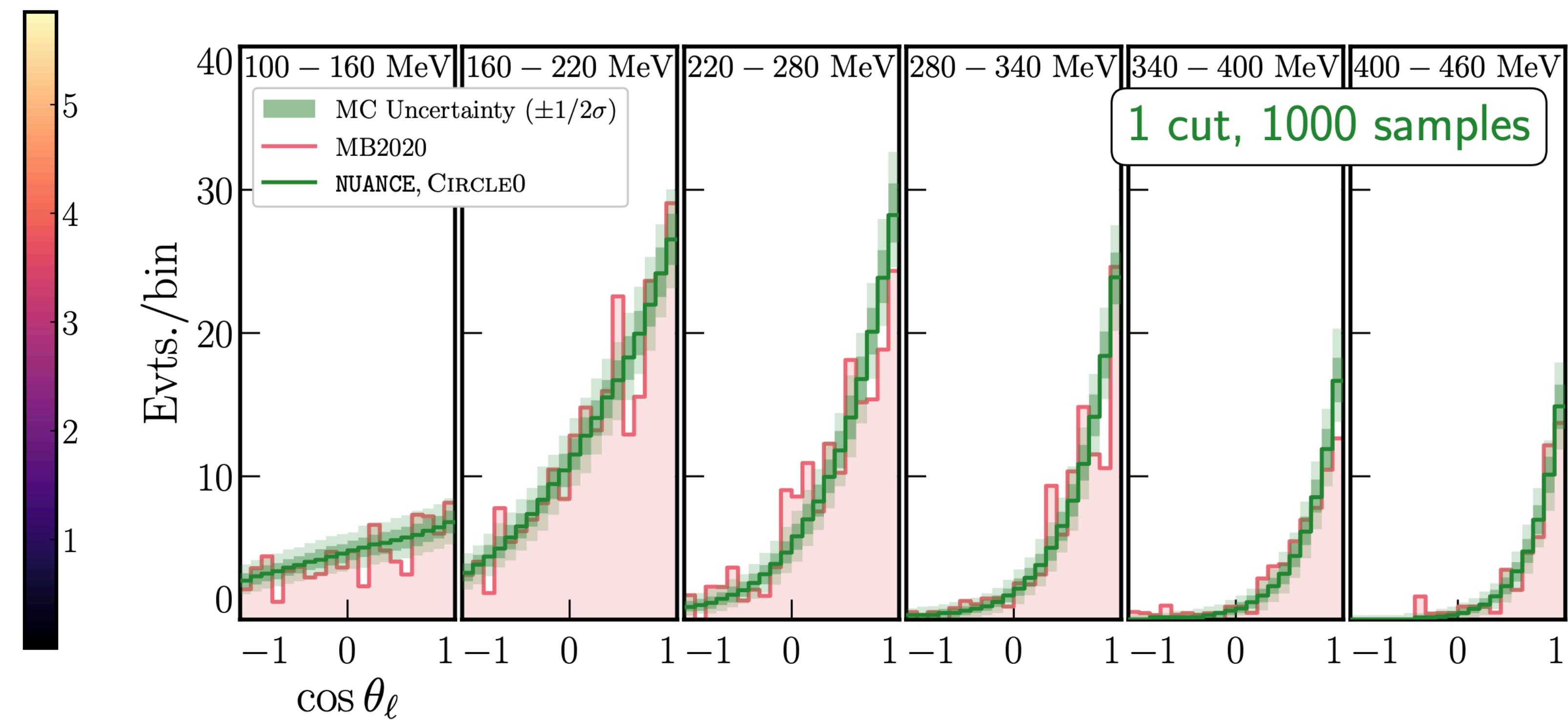
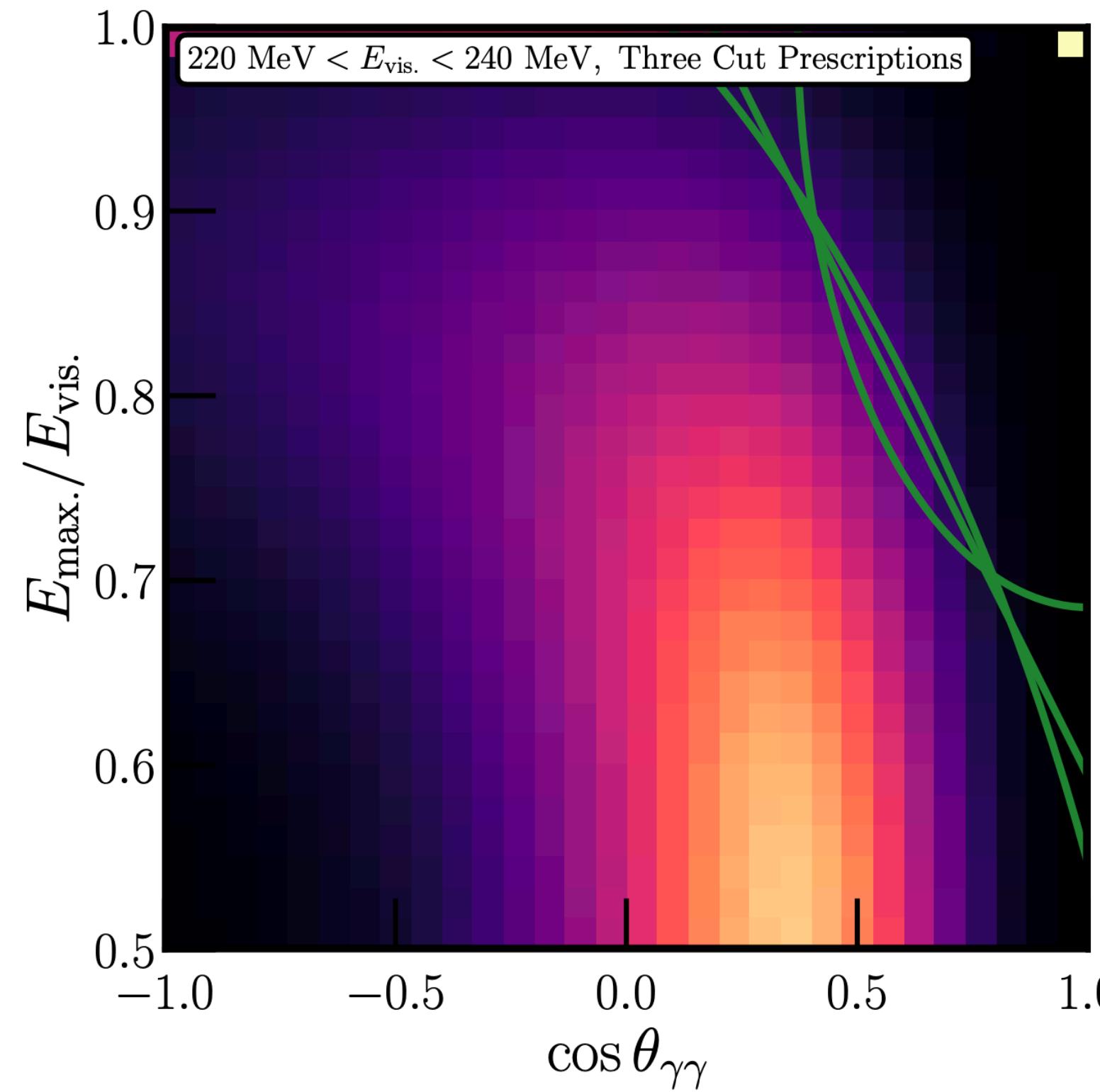




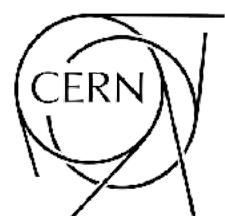
# Bonus Slides



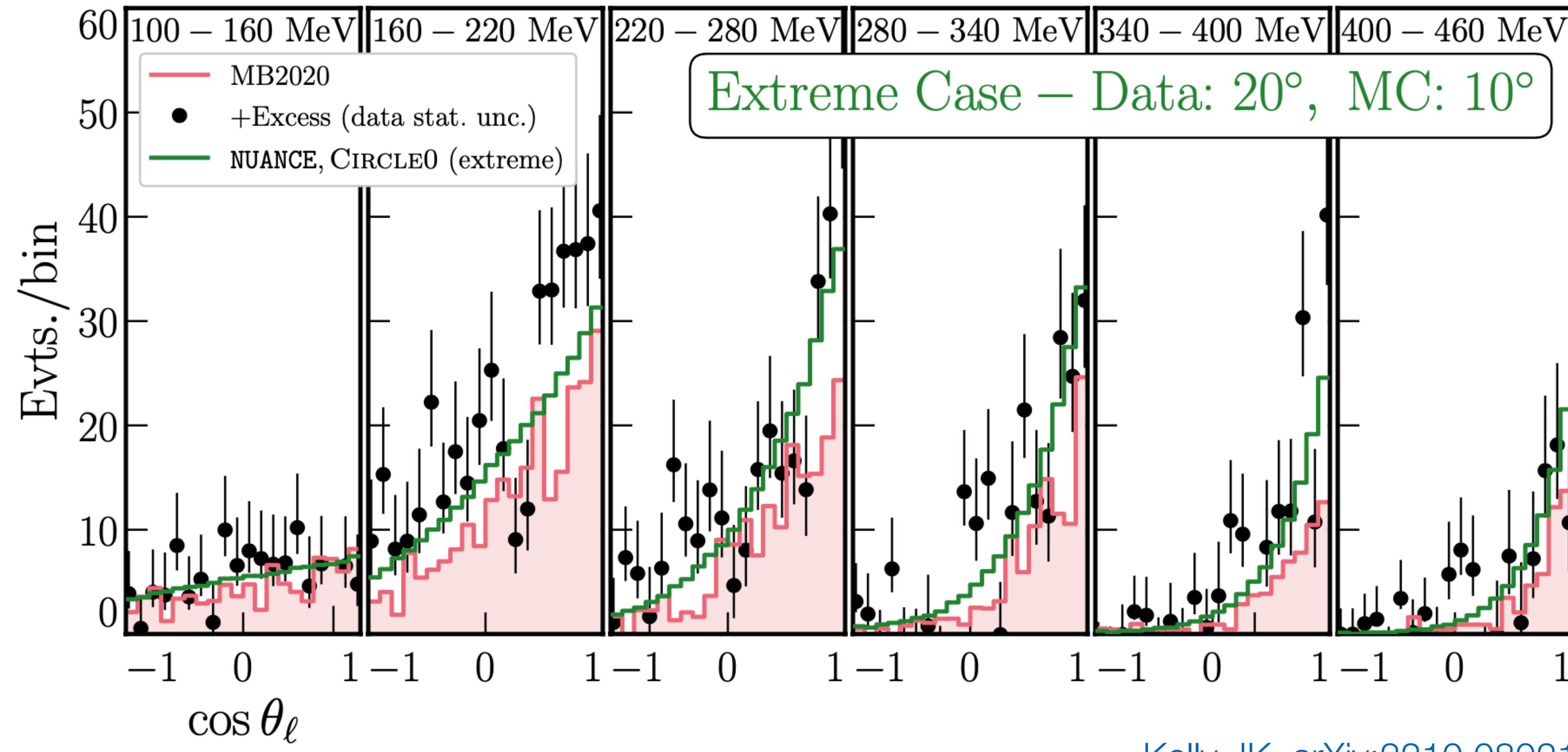
# Modeling $\pi^0$ Mis-ID in MiniBooNE



Kelly JK, arXiv:2210.08021



# Can mis-modeling of $\pi^0$ mis-ID explain the anomaly?



Kelly JK, arXiv:2210.08021

