

# Physics Topics about the Neutrons in DUNE

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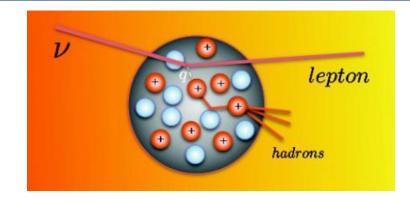




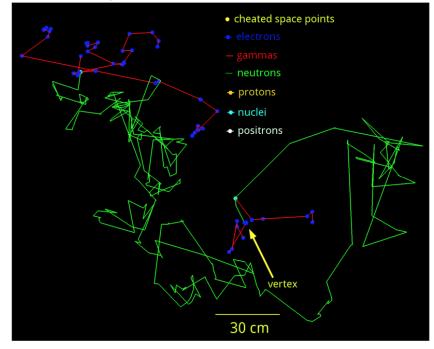
June 7, 2023
Pulsed Neutron Source
Working Group Meeting

# Neutrons are Important

- To turn neutrino physics into precision science, we must understand the neutrino-nucleus interactions
  - Neutrons carry away a large fraction of energy
  - Neutron yield is model dependent
  - Very hard to detect in liquid argon
- Understanding the neutrons are also important for low-energy physics
  - Part of the signals for supernova/solar neutrino event
  - Dominant background



#### Particle trajectories from a simulated SN event in DUNE

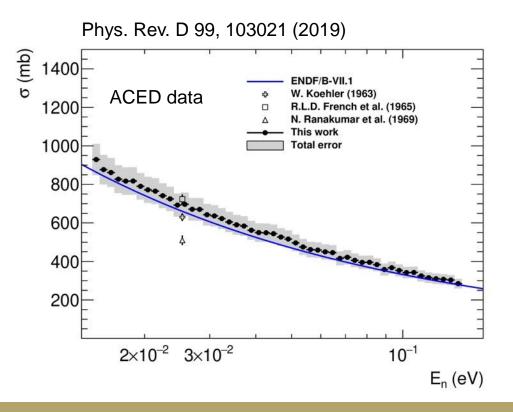


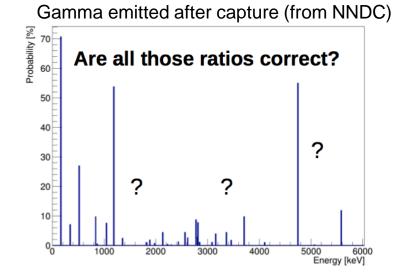
#### **Neutron Cross Section Measurements**

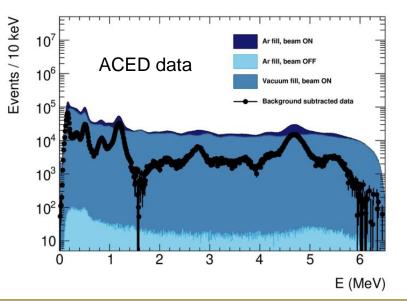
- ACED (LANL): Measured neutron capture cross sections at thermal energies
- ARTIE-I (LANL): Measured neutron total cross section for 20-70 keV
- ARTIE-II (LANL): Measure neutron total cross section from 0 to 200 keV; Proposal approved by LANL.
- MArEX@n-TOF (CERN): Multiple experiments planned to measure neutron total cross section, neutron capture cross section from 0 to 500 MeV. LOI approved by CERN.
- Near detector 2x2 Demonstrator: inelastic cross sections at several hundred MeV
- ProtoDUNE-I and II: inelastic cross sections at several hundred MeV
- Question: How are the neutron cross sections going to affect the physics simulation and data analysis?

### **ACED Neutron Capture Result**

- ACED experiment measured capture cross sections at thermal energies
- Gamma cascade measured but data not analyzed

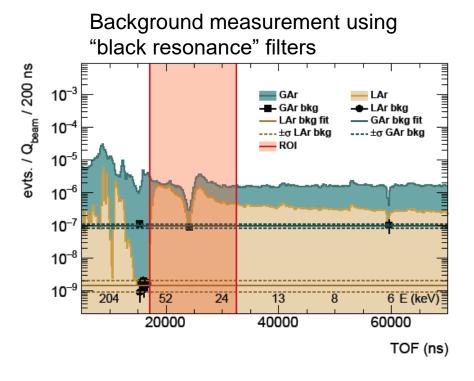


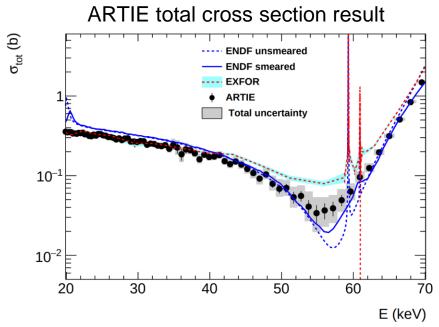




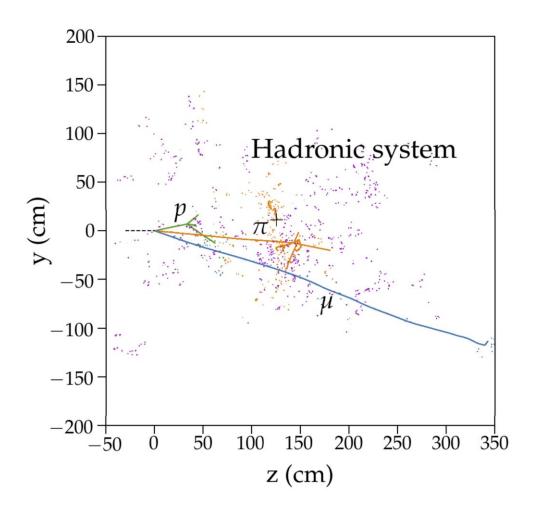
#### **ARTIE Neutron Transmission Result**

- ARTIE experiment confirmed the existence of the cross section dip around 57 keV
- Paper draft available on arXiv (<a href="https://arxiv.org/pdf/2212.05448.pdf">https://arxiv.org/pdf/2212.05448.pdf</a>)
- Submitted to PRC





#### **Neutrino Energy in DUNE**



Hadronic composition of CC  $\nu_{\mu}$ +<sup>40</sup>Ar scattering event (E $_{\nu}$  = 4 GeV). For 10,000  $\nu_{\mu}$ +<sup>40</sup>Ar CC scattering events with E $_{\nu}$ =4 GeV, there was 19% energy loss to neutrons Simulated with GENIE 2.12.8 default tune.

 It is necessary to be able to account for neutron "missing energy" in order to make precision oscillation measurements

Alexander Friedland, et. al., arXiv:1811.06159 [hep-ph]

# Missing Energy Problem

Neutron multiplicity measurements will be very useful to verify models

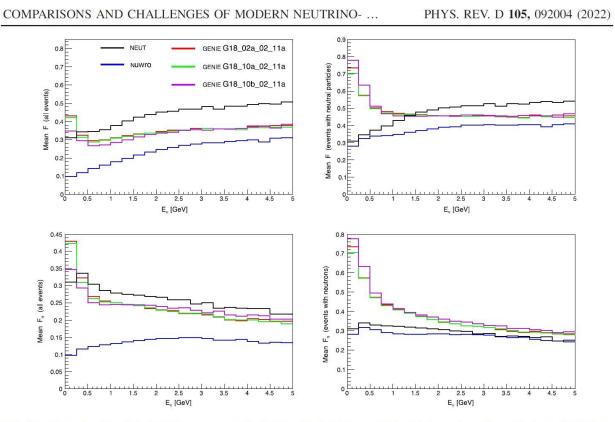


FIG. 37. Mean fraction of the leptonic energy transfer imparted to final-state neutral particle species. All panels show distributions calculated for charged-current  $\nu_{\mu}$  interactions on  $^{40}$ Ar. Top left: Predictions including all neutral particles and all events in the sample. Top right: Predictions including all neutral particles for events containing at least one final-state neutral particle. Bottom left: Predictions including final-state neutrons only and all events in the sample. Bottom right: Predictions including final-state neutrons only for events containing at least one final-state neutron.

# **Neutron Multiplicity Simulations**

Eur. Phys. J. Spec. Top. (2021) 230:4449-4467



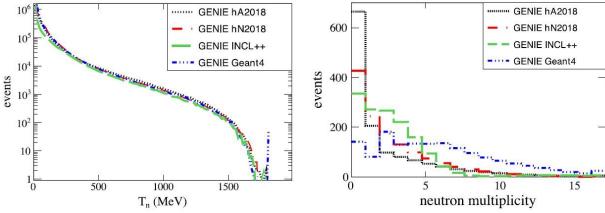
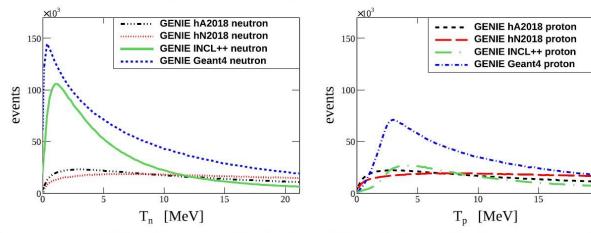


Fig. 8 Neutron distributions from a simulation of 2 GeV  $\nu_{\mu}$  <sup>40</sup>Ar, kinetic energy (left) and multiplicity (right). In each case, results from all 4 models described in the text are shown

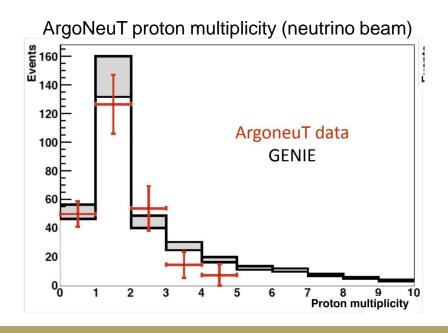


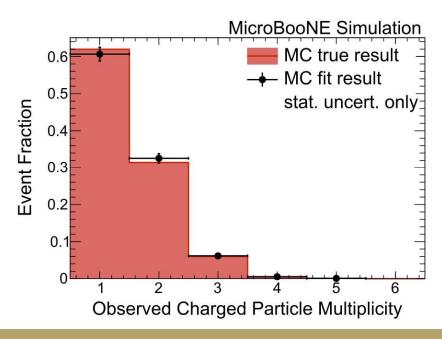
**Fig. 9** Left: Neutron (left) and proton (right) kinetic energy distributions from a simulation of 2 GeV  $\nu_{\mu}$  <sup>40</sup>Ar focusing on low energy responses, in each case, results from

all 4 models described in the text are shown. Results from INCL++ and Geant4 models described in the text are shown

### **Existing Measurements**

- Existing Measurements of neutrons
  - MINERvA measured neutron multiplicity on hydrocarbon target
  - SNO measured neutron multiplicity on heavy water and CI-mixed water
  - Super-K measured neutron multiplicity on water
- Existing liquid argon experiments measured protons (or charged particles), but neutrons are hard to detect (most will escape DUNE far detector module)





#### **SNO Measurement**

arXiv:1904.01148v3 [hep-ex] 19 Jun 2019

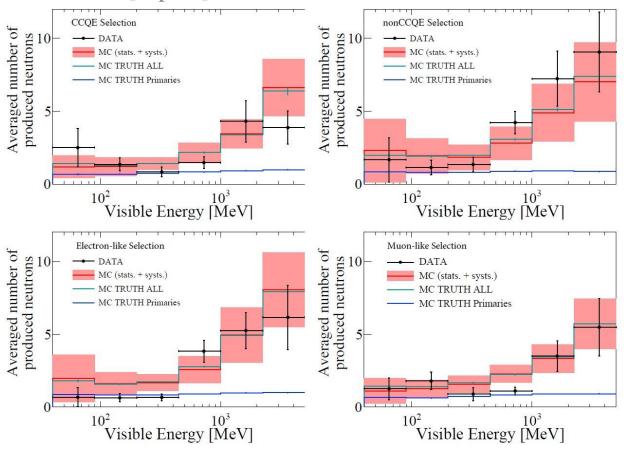
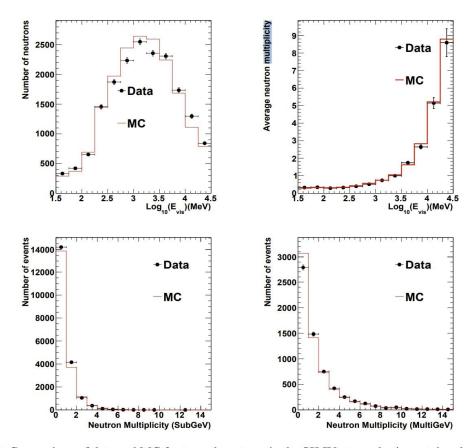


FIG. 15. Averaged number of produced neutrons vs visible energy for both phases together. We show the different selections: CCQE (top left), nonCCQE (top right), electronlike (bottom left) and muonlike (bottom right). The points represent data with statistical uncertainties. The reconstructed MC is shown with red boxes with the size corresponding to the systematic uncertainties. The green line represents the average total number of neutrons given by the MC truth, and the blue line corresponds to the average number of primary neutrons given by the MC truth.

# Super-K Measurement



**Figure 13.** Comparison of data and MC for tagged neutrons in the SK-IV atmospheric neutrino data. The top left (right) plot shows the total number of neutrons (average neutron multiplicity) as a function of visible energy ( $E_{\rm vis}$ ). The bottom left plot shows the neutron multiplicity for sub-GeV events ( $E_{\rm vis} < 1.33\,{\rm GeV}$ ) and the bottom right plot shows that for multi-GeV events ( $E_{\rm vis} \ge 1.33\,{\rm GeV}$ ). These plots are normalized to the number of neutrino events observed in the data. Only statistical errors are shown.

#### Minerva Measurement

arXiv:1901.04892v2 [hep-ex] 20 Sep 2019

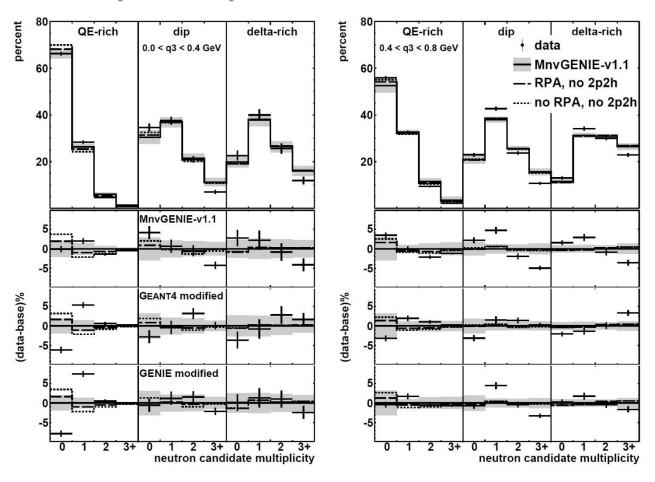
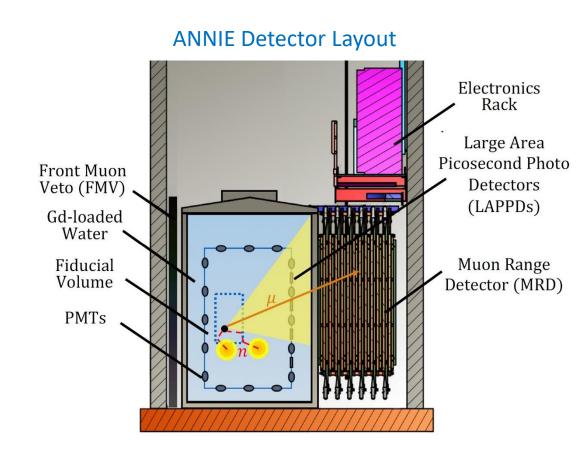


FIG. 10: Candidate multiplicity distribution for all six subsamples,  $0 < q_3 < 0.4$  (left) and  $0.4 < q_3 < 0.8$  GeV/c (right), with subpanels for the QE-rich, dip, and  $\Delta$ -rich regions. The top plot shows the reference MnvGENIE-v1.1 simulation with a solid line and error band, and two variations that turn off completely the 2p2h component and then also turn off the RPA component. The next row shows the difference from the reference simulation. The middle (lower) row of difference plots uses the modified GEANT4 benchmark (modified GENIE benchmark) for all distributions.

#### Possible to measure in ANNIE?

- Accelerator Neutrino Neutron Interaction Experiment (ANNIE) is a 26 ton Gdloaded water detector at the BNB beam
- Possible to deploy an argon target
  - Vacuum-insulated liquid argon target, or
  - High-pressure gaseous argon target at the fiducial volume location
- Need small-size photodetectors inside the target to trigger.
- Neutrons can go through the target to get captured by Gdwater
- Neutrons are counted by existing ANNIE detector



#### Interesting Topics to Discuss

- Measure total/inelastic/capture cross sections over all energy ranges through many experiments
- Recover missing energy by measuring the neutron activities
- Identify neutrino beam wrong sign contamination by counting neutrons
- Test/model the light response from neutron interactions in argon
- Test supernova trigger efficiency with the PNS runs.
- Test neutron capture identification on different materials with PNS data at ProtoDUNE
- Test supernova/solar neutrino event selection algorithms with PNS data at ProtoDUNE
- Perform muon capture analysis in existing liquid argon TPCs or propose new experiments.

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