

# n-Ar Cross Section @ ProtoDUNE-ND

Nicholas Carrara on behalf of the analysis team

2x2 First Analysis Meeting

July 21, 2023

# Two Major Changes for MiniRun5

Neutron Capture Gammas in edep-sim:

- Geant4 has an incorrect gamma cascade for neutron captures on Ar40.
- David Rivera is working on implementing the fix (from Jingo Wang in LArSoft) into edep-sim.

Microphysics Recombination:

- Sam Fogarty and Nick Carrara have been working on integrating LArNEST into larnd-sim/LArSoft.

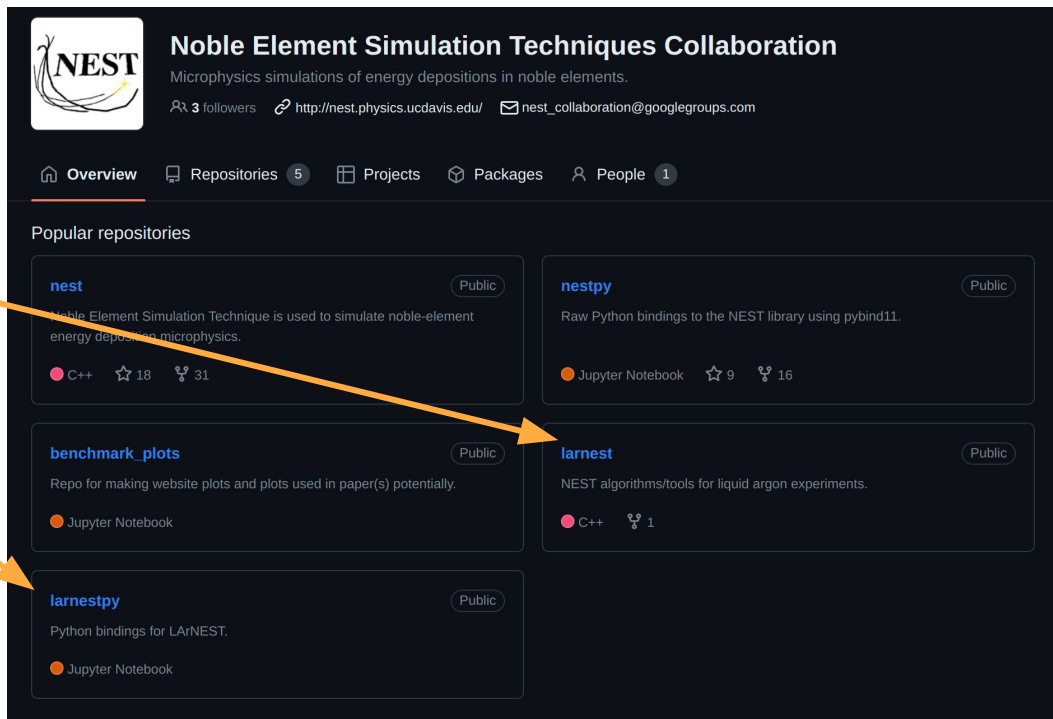
Data fit from various experiments for low energy NR, ER and Alphas

- ARIS
- SCENE
- Joshi
- WARP
- CREUS
- Regenfus
- MicroCLEAN
- Scalettar
- DarkSide
- Kimura
- Bondar
- Doke
- Lippincott
- Sangiorgio

# LArNEST

Two repos for LArNEST,

- c++ version for LArSoft  
(<https://github.com/NESTCollaboration/larnest>),
- python bindings for larnd-sim  
(<https://github.com/NESTCollaboration/larnestpy>)



The screenshot shows the GitHub profile for "Noble Element Simulation Techniques Collaboration". The profile includes a logo, a description of the organization's focus on microphysics simulations, and statistics such as 3 followers and a website link. Below the profile information, there are navigation tabs for Overview, Repositories (5), Projects, Packages, and People (1). The "Popular repositories" section lists five repositories:

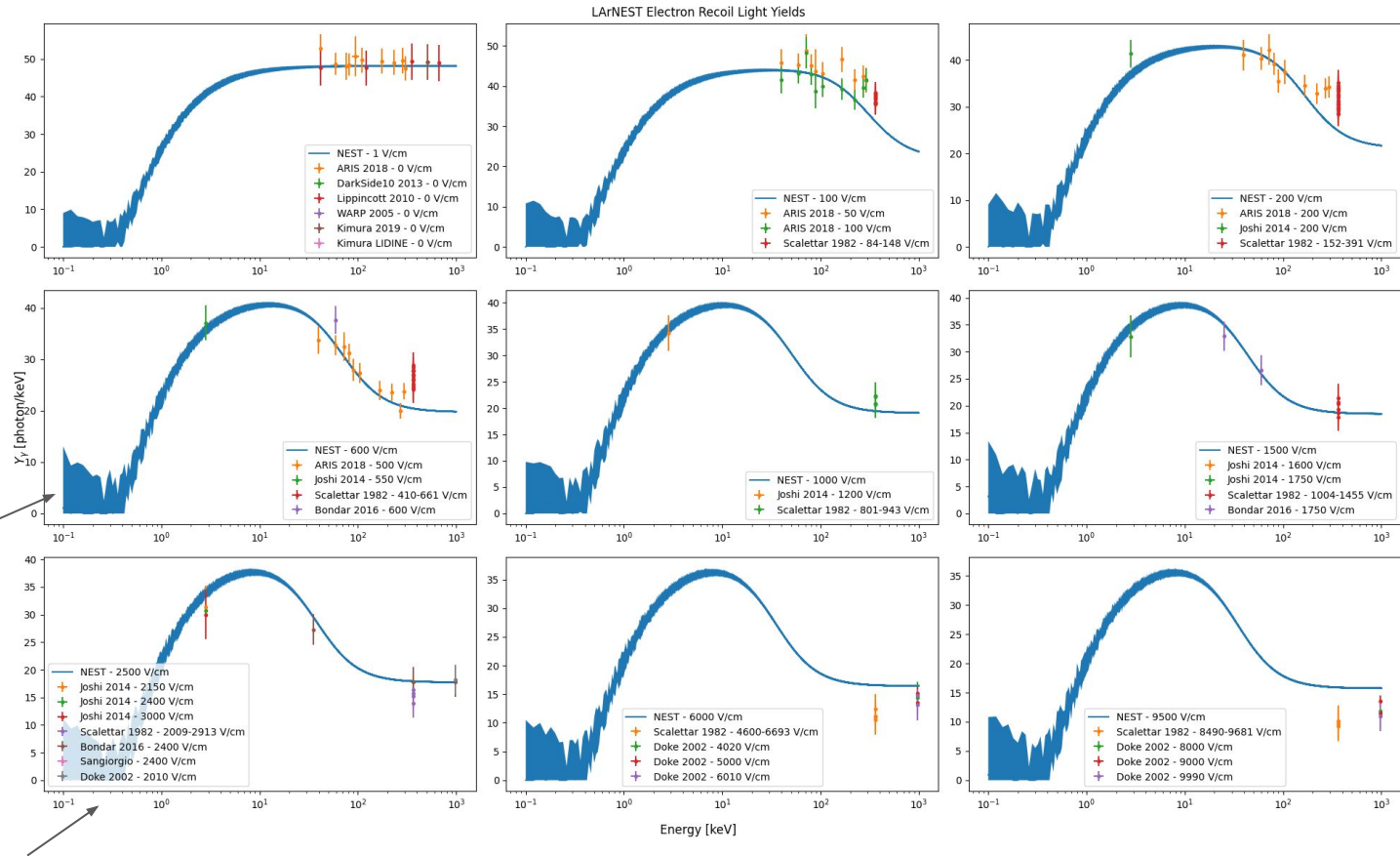
- nest** (Public): Noble Element Simulation Technique is used to simulate noble-element energy deposition microphysics. 18 stars, 31 forks.
- nestpy** (Public): Raw Python bindings to the NEST library using pybind11. 9 stars, 16 forks.
- benchmark\_plots** (Public): Repo for making website plots and plots used in paper(s) potentially. 1 fork.
- larnest** (Public): NEST algorithms/tools for liquid argon experiments. 1 fork.
- larnestpy** (Public): Python bindings for LArNEST.

Two orange arrows originate from the text on the left. One arrow points from the <https://github.com/NESTCollaboration/larnest> link to the "larnest" repository card. The other arrow points from the <https://github.com/NESTCollaboration/larnestpy> link to the "larnestpy" repository card.

# LArNEST

Electron Recoil model is fit to the available data using functions from first principles (**work done by Justin Mueller and Ekaterina Kozlova**).

Fluctuations are Poisson/Fano.



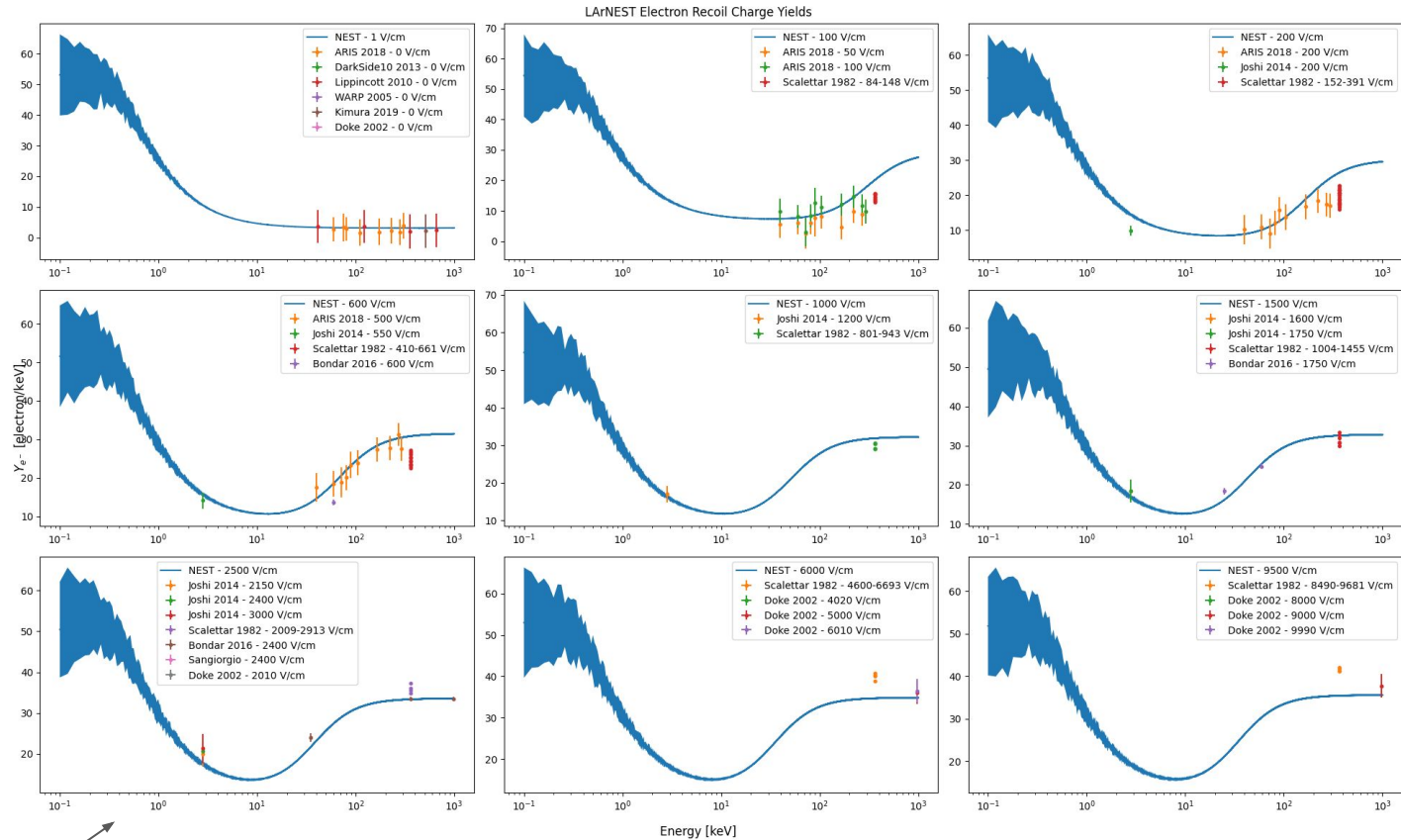
Electron Recoil Light Yields [photons/keV]  
for different electric field values [V/cm]

# LArNEST

The function for electron yields takes (energy, electric field, density) as inputs, and is fit with **31** parameters.

It is adapted from the beta model for liquid Xenon in NEST.

Light yields are assumed anti-correlated for electronic recoils.



Electron Recoil Quanta Yields [electrons/keV]  
for different electric field values [V/cm]

# LArNEST

Parameter	LXe	LAr
$A_\alpha$	32.988	32.988
$B_\alpha$	-32.988	-552.988
$C_\alpha$	1	17.2346
$D_\alpha$	0	-4.7
$E_\alpha$	0.026715	0.025115
$F_\alpha$	-0.33926	0.26536
$G_\alpha$	0.6705	0.242671
$A_\beta$	1	0.778482
$B_\beta$	0.4607	25.9
$C_\beta$	1	1.105
$D_\beta$	621.74	0.4
$E_\beta$	-2.2717	4.55
$F_\beta$	53.502	-7.502
$A_\gamma$	1	0.659509
$B_\gamma$	1000	1000
$C_\gamma$	6.5	6.5
$D_\gamma$	1	5
$E_\gamma$	-11	-0.5
$F_\gamma$	47.408	1047.408
$G_\gamma$	1.9851	0.01851
$A_\epsilon$	1652.264	1052.264
$B_\epsilon$	1.415935e10 - 1652.264	1.415935e10 - 1652.264
$C_\epsilon$	1	-5
$D_\epsilon$	0.02673144	0.157933
$E_\epsilon$	1.564691	1.83894
$p_1$	1	1
$p_2$	1.304	10.304
$p_3$	2.1393	13.0654
$p_4$	0.35535	0.10535
$p_5$	1	0.7
$\delta$	28	15.7489
$\ell$	-2	-2.07763

$$Y_e(E|\alpha, \beta, \gamma, \delta, \epsilon, p_k, \ell) = \alpha\beta + \frac{\gamma - \alpha\beta}{(p_1 + p_2(E + 0.5)^{p_3})^{p_4}} + \frac{\delta}{p_5 + \epsilon E^\ell}, \quad (61)$$

where the  $p_k$ ,  $\delta$  and  $\ell$  are constants<sup>3</sup>,  $\alpha$  is a function of the electric field and the density,

$$\alpha(\vec{E}, \rho_Z) = A_\alpha + B_\alpha \left( C_\alpha + \left( \frac{|\vec{E}|}{D_\alpha + E_\alpha \exp\left[\frac{\rho_Z}{F_\alpha}\right]} \right)^{G_\alpha} \right)^{-1}, \quad (62)$$

while  $\beta$  is a function of only the electric field:

$$\beta(\vec{E}) = A_\beta + B_\beta \left( C_\beta + \left( \frac{|\vec{E}|}{D_\beta} \right)^{E_\beta} \right)^{F_\beta}. \quad (63)$$

The  $\gamma$  function depends on the electric field and the work function of LAr:

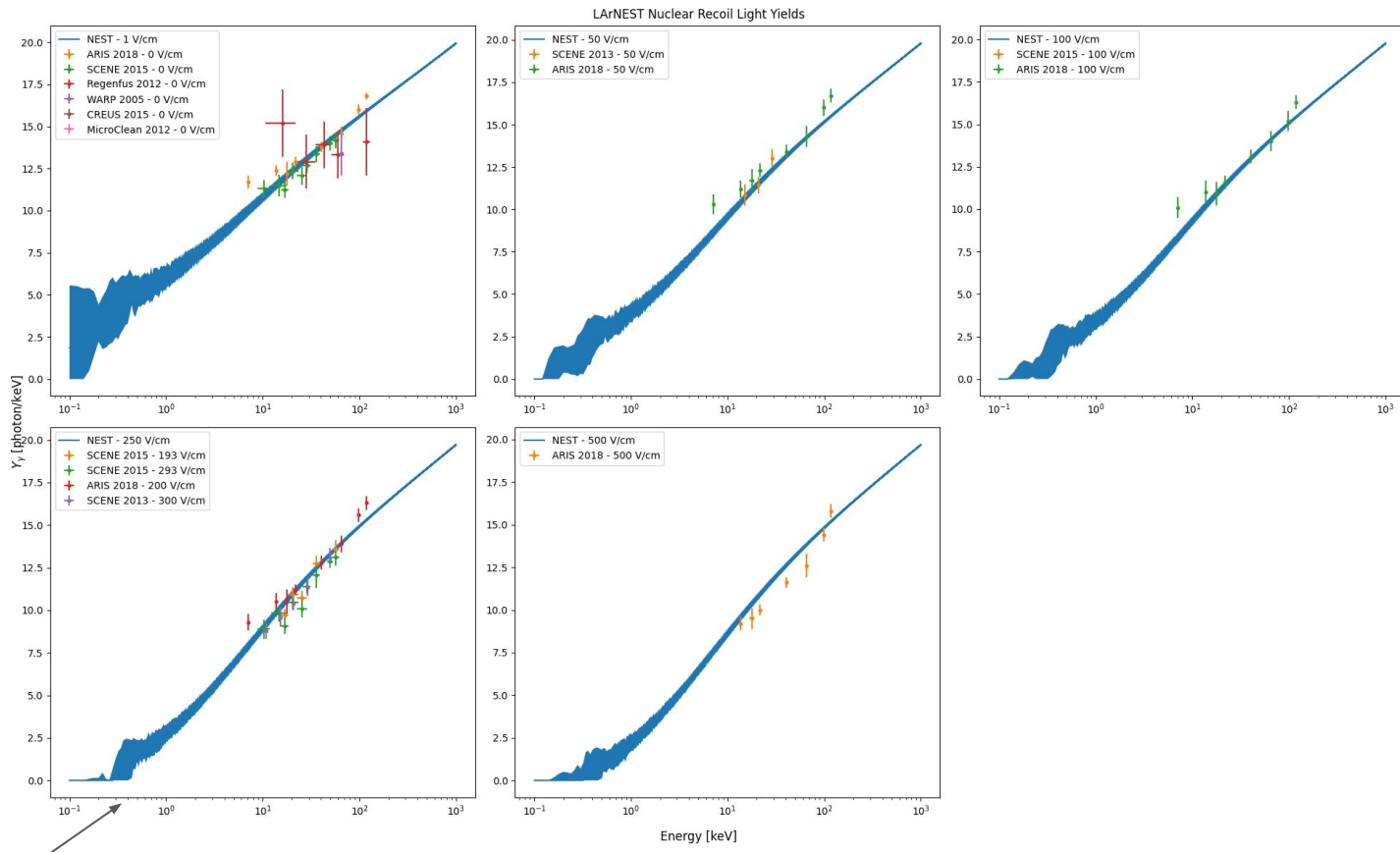
$$\gamma(\vec{E}, W) = A_\gamma \left( \frac{B_\gamma}{W} + C_\gamma \left( D_\gamma + \frac{E_\gamma}{\left( \frac{|\vec{E}|}{F_\gamma} \right)^{G_\gamma}} \right) \right). \quad (64)$$

The  $\epsilon$  function is the Doke-Birks function for LAr which is given by:

$$\epsilon(\vec{E}) = A_\epsilon + \frac{B_\epsilon}{\left( C_\epsilon + \left( \frac{|\vec{E}|}{D_\epsilon} \right)^{E_\epsilon} \right)}. \quad (65)$$

# LArNEST

Similarly for nuclear recoils, a model for light yields and electron yields are fit according to available data, **but with a slight breaking of the anti-correlation.**



**Nuclear Recoil Light Yields [photons/keV] for different electric field values [V/cm]**

# LArNEST

## Nuclear Recoil Total Yields [quanta/keV] for different electric field values [V/cm]

The equation describing the total yield used in this study is given by a simple power law,

$$Y_q(E, \vec{E}|\alpha, \beta) = Y_q(E|\alpha, \beta) = \alpha E^\beta, \quad (57)$$

where  $\alpha$  is a positive scalar with units  $[\alpha] = (\text{keV}^\beta)^{-1}$  and  $\beta$  is a dimensionless real number.

$$\begin{aligned} \tilde{Y}_\gamma(E, \vec{E}|\alpha, \beta, \gamma, \delta, \epsilon) &= \alpha E^{\beta-1} - \left( \frac{1}{\gamma |\vec{E}|^\delta} \right) \left( \frac{1}{\sqrt{E + \epsilon}} \right) \\ &= Y_q - Y_e \left( 1 - \frac{1}{1 + \left( \frac{E}{\zeta} \right)^\eta} \right)^{-1}. \end{aligned} \quad (59)$$

Slight breaking of a strict anti-correlation

## Nuclear Recoil Light Yields [photons/keV] for different electric field values [V/cm]

Parameter	LXe	LAr
$\alpha$	$11^{+2.0}_{-0.5}$	$11.10 \pm 1.4$
$\beta$	$1.1 \pm 0.05$	$1.087 \pm 0.01$

Parameter	LXe	LAr
$\gamma$	$0.0480 \pm 0.0021$	$0.1 \pm 0.005$
$\delta$	$-0.0533 \pm 0.0068$	$-0.0932 \pm 0.0095$
$\epsilon$	$12.6^{+3.4}_{-2.9}$	$2.998 \pm 1.026$
$\zeta$	$0.3 \pm 0.1$	0.3 (Fixed)
$\eta$	$2 \pm 1$	$2.94 \pm 0.12$

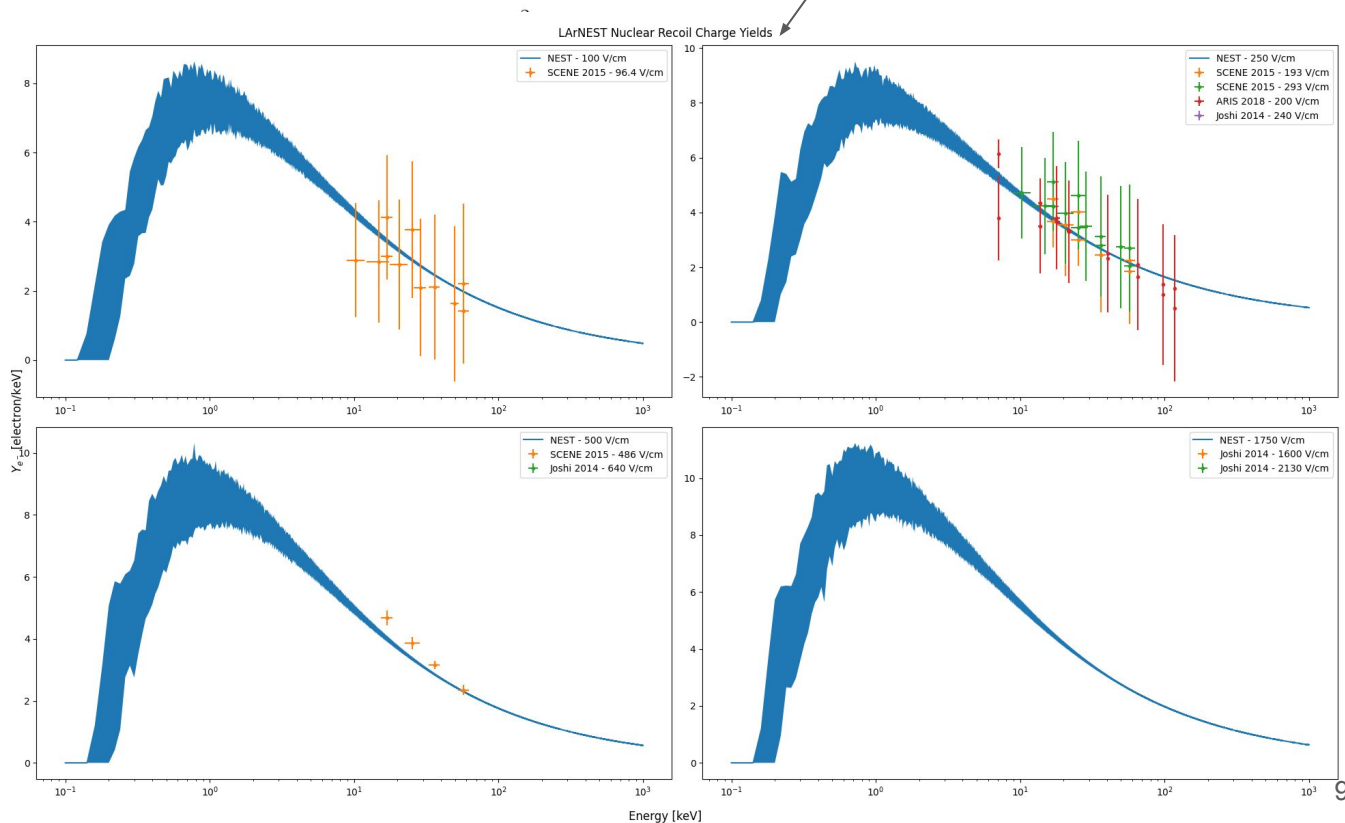


# LArNEST

The equation describing the exciton (or charge) yield is given below:

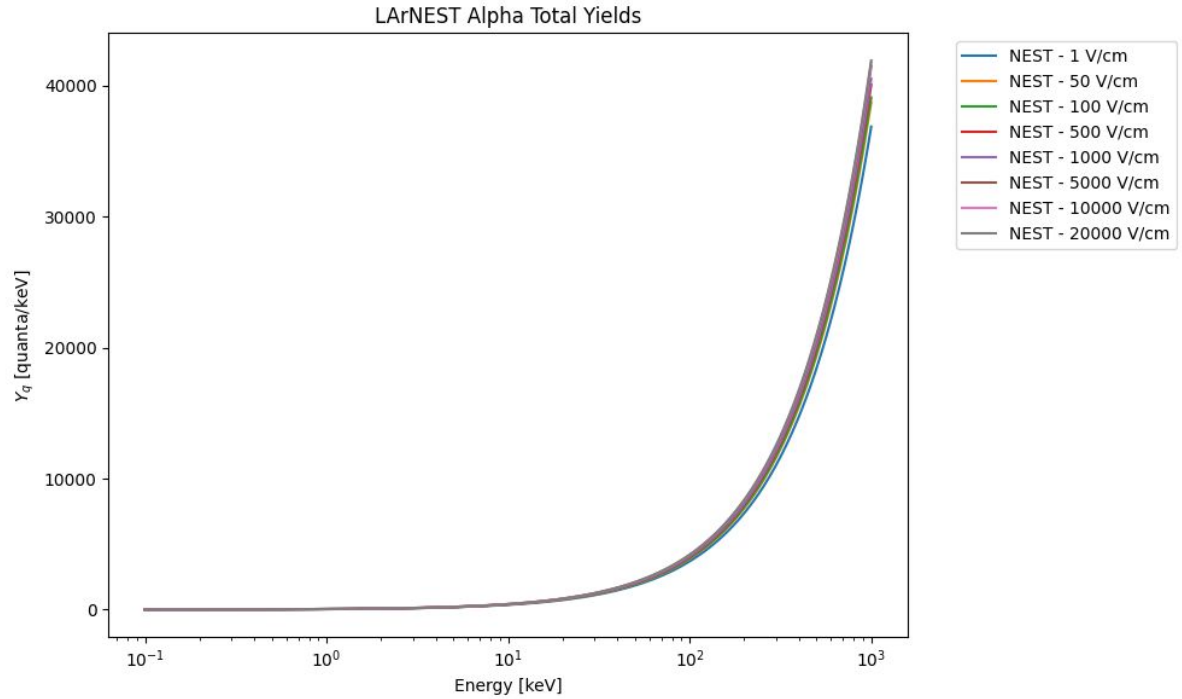
$$Y_e(E, \vec{E} | \gamma, \epsilon, \delta, \zeta, \eta) = \left( \frac{1}{\gamma |\vec{E}|^\delta} \right) \left( \frac{1}{\sqrt{E + \epsilon}} \right) \left( 1 - \frac{1}{1 + \left( \frac{E}{\zeta} \right)^\eta} \right), \quad (58)$$

**Nuclear Recoil Electron Yields [electrons/keV] for different electric field values [V/cm]**



# LArNEST

We use a slightly different model for alphas (of which there is very little data!).



# LArNEST

## Alpha Light/Electron Yields [(photons/electrons)/keV]

The equation describing the electron yield is,

$$Y_e(\vec{E}) = A \left( B - \left( BC + \frac{B}{D} \left[ 1 - \frac{E \log \left( 1 + \frac{B}{D} F \frac{(G + |\vec{E}|^H)^I}{J} \right)}{\frac{B}{D} F (G + |\vec{E}|^H)^I} \right] \right) \right), \quad (66)$$

where  $A, B, C, D, E, F, G, H, I$  and  $J$  are scalar parameters. The fits for these values are given below:

Parameter	LAr
$A$	1.0/6200.0
$B$	64478398.7663
$C$	0.173553719
$D$	1.21
$E$	0.028752
$F$	0.01
$G$	4.71598
$H$	7.72848
$I$	-0.109802
$J$	3.0

$$Y_\gamma(\vec{E}) = \left( \frac{1}{A|\vec{E}|^B} \right) C \left( DE + \frac{D}{F} \left[ 1 - \frac{G \log \left( 1 + \frac{D}{F} H \frac{(I + (\frac{|\vec{E}|}{J})^K)^L}{M} \right)}{\frac{D}{F} H (I + (\frac{|\vec{E}|}{J})^K)^L} \right] \right), \quad (67)$$

where  $A, B, C, D, E, F, G, H, I, J, K, L$  and  $M$  are scalar parameters. The fits for these values are given below:

Parameter	LAr
$A$	1.5
$B$	-0.012
$C$	1.0/6500.0
$D$	278037.250283
$E$	0.173553719
$F$	1.21
$G$	2.0
$H$	0.653503
$I$	4.98483
$J$	10.0822
$K$	1.2076
$L$	-0.97977
$M$	3.0

# LArNEST

LArNEST currently has eight options for calculating yields/fluctuations. The first three (**NR**, **ER**, **Alpha**) are described in the previous slides.

The **LeptonLET**, **LET** and **Legacy** versions were used in the early days of LBNE (M. Szydagis).

The **BOX** and **BIRKS** models are taken from the current larnd-sim quenching.

```
enum class LArInteraction
{
    NR = 0,
    ER = 1,
    Alpha = 2,
    dEdx = 3,
    LeptonLET = 4,
    LET = 5,
    BOX = 6,
    BIRKS = 7,
    Legacy = 8
};
```

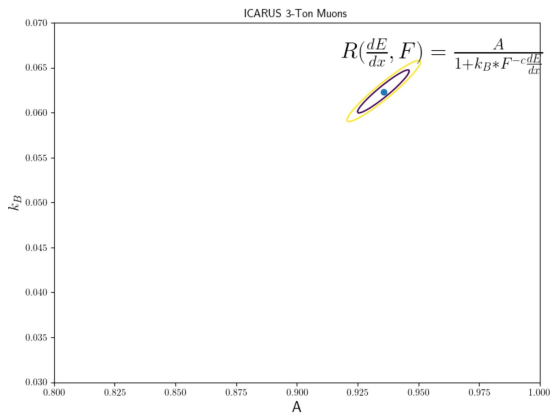
```
You, 3 weeks ago | 1 author (You)
struct LArNRYieldsParameters
{
    double alpha = {11.10};
    double beta = {0.087};
    double gamma = {0.1};
    double delta = {-0.0932};
    double epsilon = {2.998};
    double zeta = {0.3};
    double eta = {2.94};
};
```

```
You, 3 weeks ago | 1 author (You)
struct LArERElectronYieldsAlphaParameters
{
    double A = {32.988};
    double B = {-552.988};
    double C = {17.2346};
    double D = {-4.7};
    double E = {0.025115};
    double F = {0.265360653};
    double G = {0.242671};
};
```

The several dozen parameters are all tunable at run time.

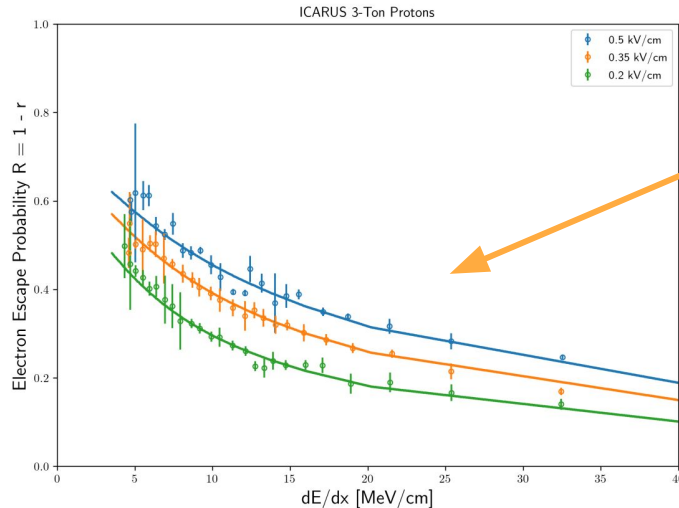
# LArNEST

LArEdxParameters were fit by **Justin Mueller** for a Birks type model on ICARUS/DM data.



Prefers  $c = 0.90$

BOX and BIRKS correspond to larnd-sim models.



**Best Result**

**Prefers  $c = 0.85$**

```
You, 33 minutes ago | 1 author (You)
struct LArEdxParameters
{
    double A = {0.87};
    double kb = {0.045};
    double c = {0.9};
};

You, 3 weeks ago | 1 author (You)
struct ThomasImelParameters
{
    double A = {0.1};
    double B = {-0.0932};
};

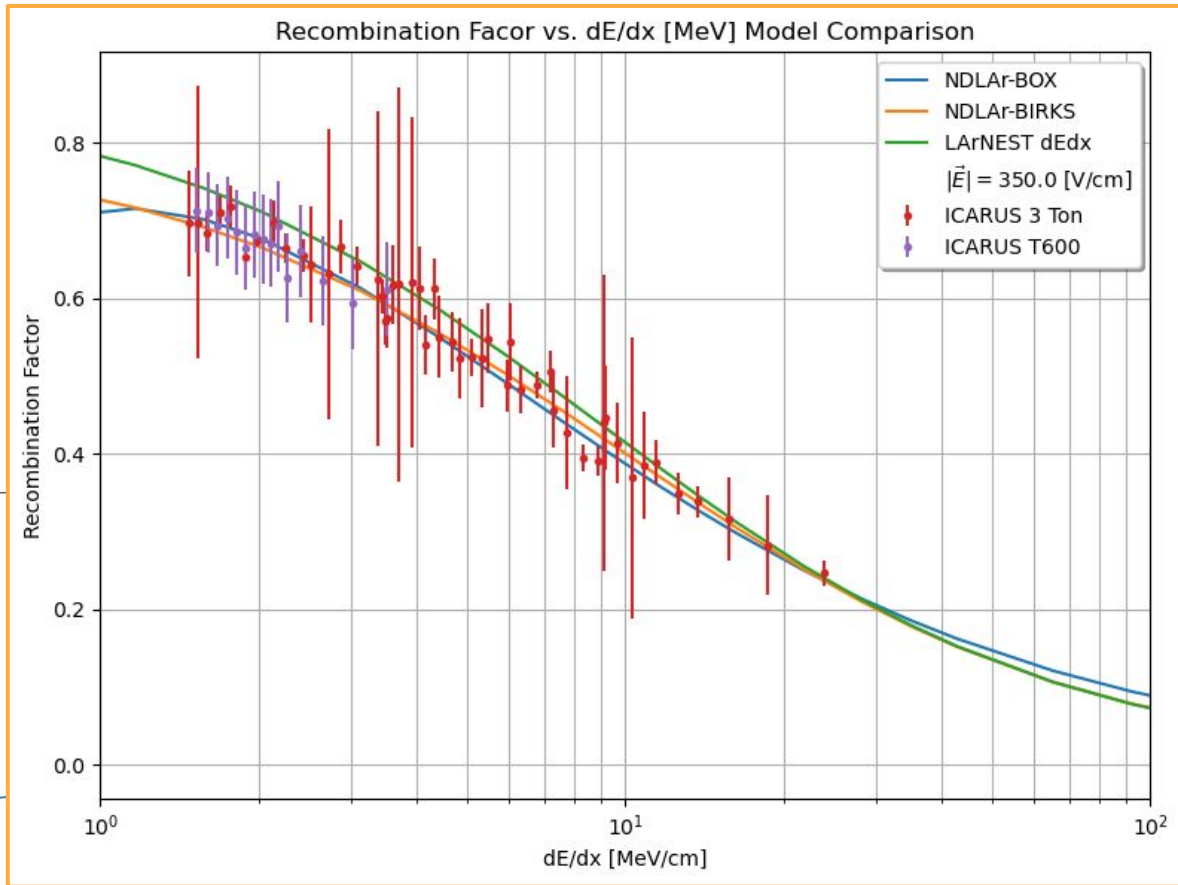
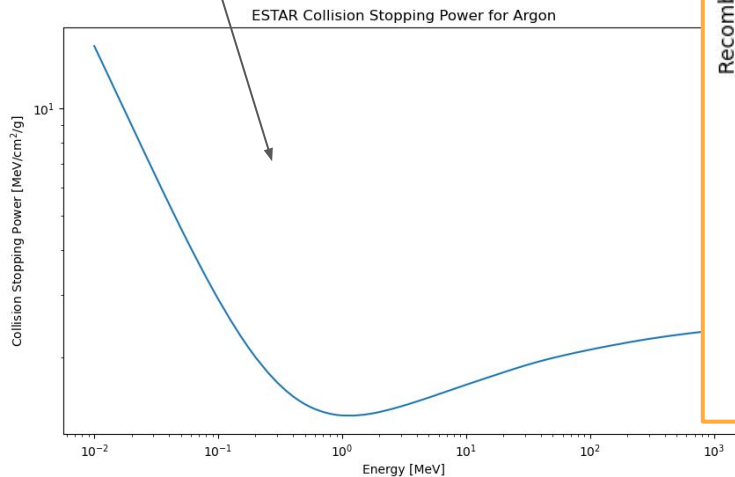
You, 3 days ago | 1 author (You)
struct BOXParameters
{
    double alpha = {0.93};
    double beta = {0.207};
};

You, 3 days ago | 1 author (You)
struct BIRKSPParameters
{
    double Ab = {0.800};
    double kb = {0.0486};
};
```

# LArNEST Model Comparison

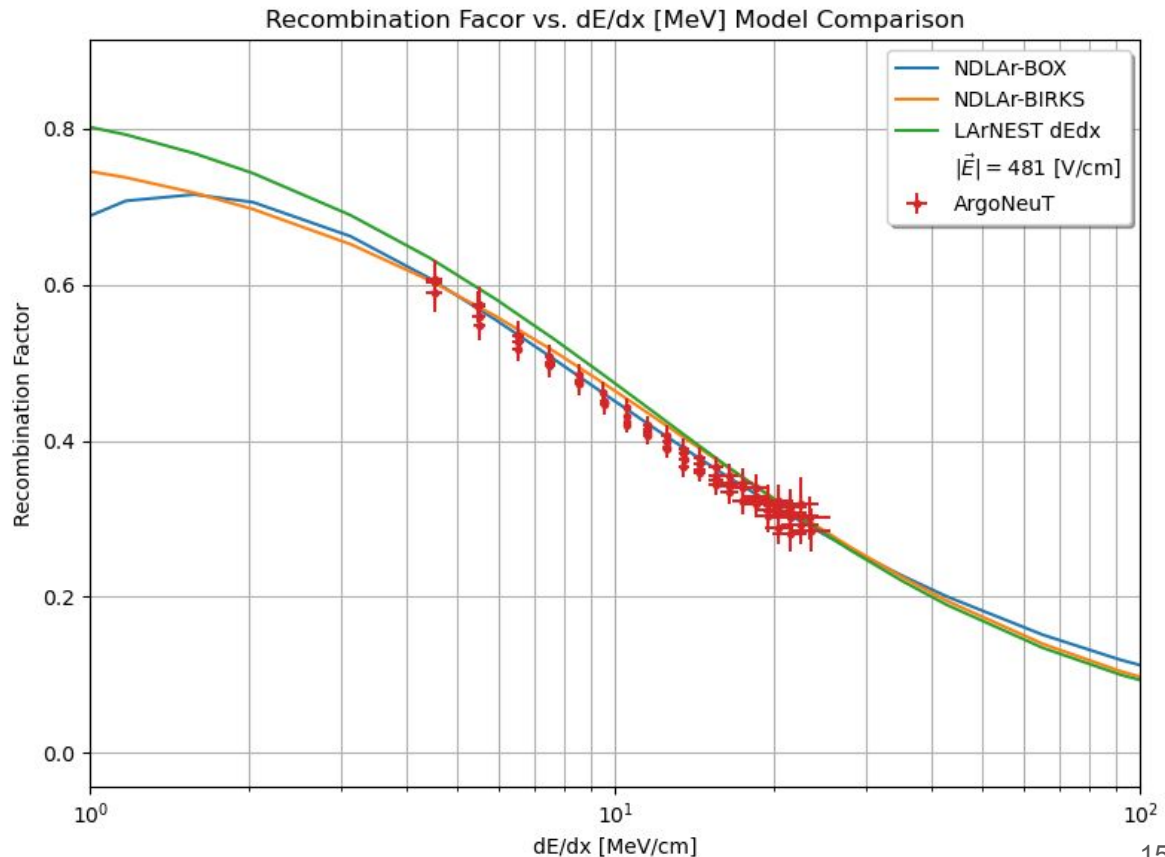
Comparison of larnd-sim  
BOX/BIRKS models with Justin's  
dEdx model against ICARUS  
data.

Using the associated scattering  
length from ESTAR stopping  
power of electrons in Argon.



# LArNEST Model Comparison

ArgoNeuT parameters need to be tuned slightly for the dEdx model.



# LArNEST implementation - progress & validations

## Current Recombination Models in larnd-sim

- Currently [larnd-sim](#) only supports using a single recombination model at a time, and that is either the **Box model** (Baller, 2013 JINST 8 P08005) or the **Birks model** (Amoruso, et al NIM A 523 (2004) 275)
- It is of interest to allow larnd-sim to use models from LArNEST, namely the **ER model** for low-energy electrons, the **alpha model** for alphas, and the **NR model** for nuclear recoils

Currently in quenching.py:

```
if mode == physics.BOX:
    # Baller, 2013 JINST 8 P08005
    csi = physics.BOX_BETA * dEdx / (detector.E_FIELD * detector.LAR_DENSITY)
    recomb = max(0, log(physics.BOX_ALPHA + csi)/csi)
elif mode == physics.BIRKS:
    # Amoruso, et al NIM A 523 (2004) 275
    recomb = physics.BIRKS_Ab / (1 + physics.BIRKS_kb * dEdx / (detector.E_FIELD * detector.LAR_DENSITY))
```

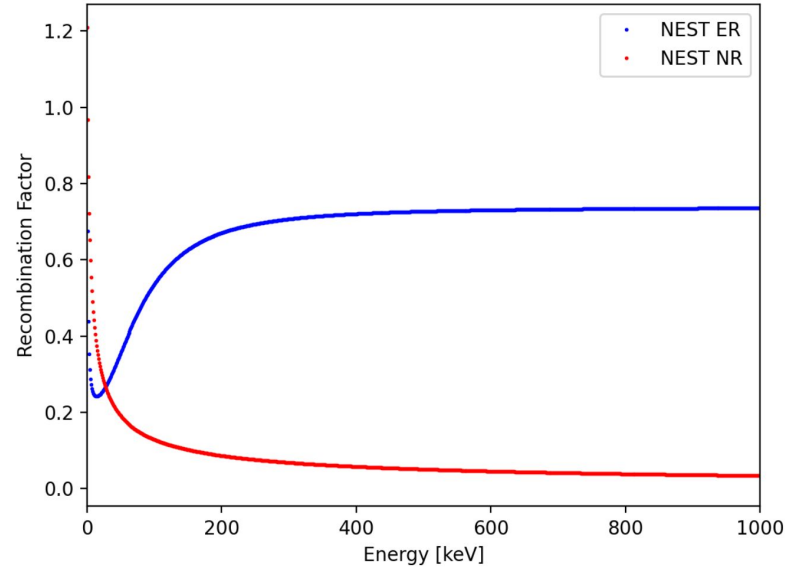


# Implementing LArNEST into larnd-sim

- Implementing LArNEST into larnd-sim is complicated by the fact that larnd-sim requires many of its functions to work with the CUDA library (allowing the program to use GPUs to speed up computations).
  - The quench function, which is where recombination factors are calculated, uses CUDA. So LArNEST would need to play nice with CUDA.
  - Many python functions do not work with CUDA (such as ones in larnestpy)
  - LArNEST is written in C++, so it is not clear how to make it compatible with CUDA
- Potential solution: Rewrite LArNEST in pure python so that it works with CUDA
  - Would obviously get around the CUDA compatibility issues
  - Would be a challenge, as many python functions do not work with CUDA
  - This is a path that we can go down if we want to, but it will obviously take some time to get working

# Implementing LArNEST into larnd-sim (alternative option)

- An alternative option to using LArNEST directly in larnd-sim is to `cache` the LArNEST data for various energies:
  - Run LArNEST for each model (ER, alpha, NR) for the range of potential energies just once prior to running larnd-sim
  - Save the results to a file (energies and recombination factors for each model), then load the data into larnd-sim
  - Interpolate the data when calculating recombination factors
- Benefits to this alternative option:
  - We do not need to add larnesty as a dependency to larnd-sim, and we wouldn't need to run larnesty for every simulation
  - We get around CUDA compatibility issues
- This implementation has been made and is being tested:  
[https://github.com/sam-fogarty/larnd-sim/tree/feature\\_cached\\_LArNEST](https://github.com/sam-fogarty/larnd-sim/tree/feature_cached_LArNEST)



↓  
Store in  
npz file

# 'Cached-LArNEST' Implementation

Added to simulation properties file:

```
pdg_to_recombination_model: # Box 1; Birks 2; NEST_ER 3; NEST_ALPHA 4; NEST_NR 5
  11: 3 # electron
  13: 2 # muon
  2212: 2 # proton
  321: 2 # kaon
  211: 2 # pion
  1000020040: 4 # alpha
  1000180400: 5 # 40Ar
  1000010020: 2 # deuteron
```

Dictionary to specify which recombination model to use for which particle

Energy threshold for NEST ER model

er\_energy\_threshold: 1.0 # MeV

default\_recombination\_model: 2 # 1 or 2

Default model to use if a simulated particle isn't in this dictionary

# 'Cached-LArNEST' Implementation

- Added logic to pick recombination model for each segment
- Includes option to bypass all logic and just use default model
- If no particular model is picked, the default model is used (Box or Birks)

```
@njit
def pick_model(model, E, dEdx, er_energy_threshold, default_model, use_default_model, \
              E_ER, E_NR, R_ER, R_NR):
    """
    Function to pick a recombination model for a particular segment.

    Args:
    model (int): recombination model number, as defined in consts.physics.
    E (float): energy in MeV, either corresponding to particle starting energy or segment dE
    dEdx (float): segment dE/dx in MeV/cm
    er_energy_threshold (float): threshold energy in MeV for using NEST ER model
        (NEST ER is used if particle energy is less than this threshold). Only relevant for electrons. but must still be specified.
    default_model (int): number corresponding to the model that should be used
        if the segment pdgID is not in the pdg->model dictionary in the simulation properties file.
    use_default_model (bool): if True, bypasses the if statements and always uses the default model.

    """
    recomb = 0
    if use_default_model:
        recomb = DEFAULT_MODEL(default_model, dEdx)
    elif model == physics.BOX:
        recomb = BOX(dEdx)
    elif model == physics.BIRKS:
        recomb = BIRKS(dEdx)
    elif model == physics.NEST_ER and E < er_energy_threshold:
        recomb = NEST_ER(E, E_ER, R_ER)
    elif model == physics.NEST_ALPHA:
        recomb = NEST_ALPHA(E)
    elif model == physics.NEST_NR:
        recomb = NEST_NR(E, E_NR, R_NR)
    else:
        recomb = DEFAULT_MODEL(default_model, dEdx)

    return recomb
```

# 'Cached-LArNEST' Implementation

Box and Birks models are unchanged, just put into their own functions

```
@njit
def BOX(dEdx):
    """
    Box recombination model. Baller, 2013 JINST 8 P08005

    Args:
        dEdx (float): Segment dE/dx in MeV/cm
    """
    csi = physics.BOX_BETA * dEdx / (detector.E_FIELD * detector.LAR_DENSITY)
    return max(0, log(physics.BOX_ALPHA + csi)/csi)

@njit
def BIRKS(dEdx):
    """
    Birks recombination model. Amoruso, et al NIM A 523 (2004) 275

    Args:
        dEdx (float): Segment dE/dx in MeV/cm
    """
    return physics.BIRKS_Ab / (1 + physics.BIRKS_kb * dEdx / (detector.E_FIELD * detector.LAR_DENSITY))
```

# 'Cached-LArNEST' Implementation

LArNEST ER and NR functions linearly interpolate the cached data for a particular energy

Alpha recombination factor is a constant w.r.t. energy, so it is implemented as just a constant for now

```
@njit
def NEST_ER(E, er_energies, er_recomb_factors):
    """
    LArNEST electron recoil (ER) recombination model used for low-energy electrons.
    https://github.com/NESTCollaboration/larnestpy

    Args:
        E (float): Starting energy in MeV of the trajectory corresponding to the current segment.
        er_energies (:obj:`numpy.ndarray`): ER energies from LArNEST.
        er_recomb_factors (:obj:`numpy.ndarray`): ER recombination factors from LArNEST.
    """
    recomb = linear_interpolation(E, er_energies, er_recomb_factors, physics.ER_ASYMPTOTE_AVG)
    return recomb

@njit
def NEST_NR(E, nr_energies, nr_recomb_factors):
    """
    LArNEST nuclear recoil (NR) recombination model. https://github.com/NESTCollaboration/larnestpy

    Args:
        E (float): Segment dE in MeV.
    """
    recomb = linear_interpolation(E, nr_energies, nr_recomb_factors, physics.NR_ASYMPTOTE_AVG)
    return recomb

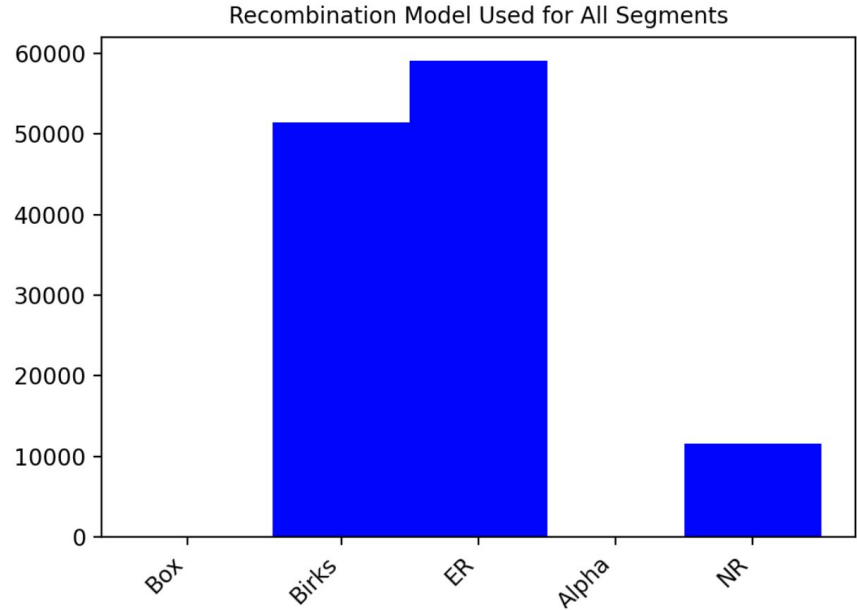
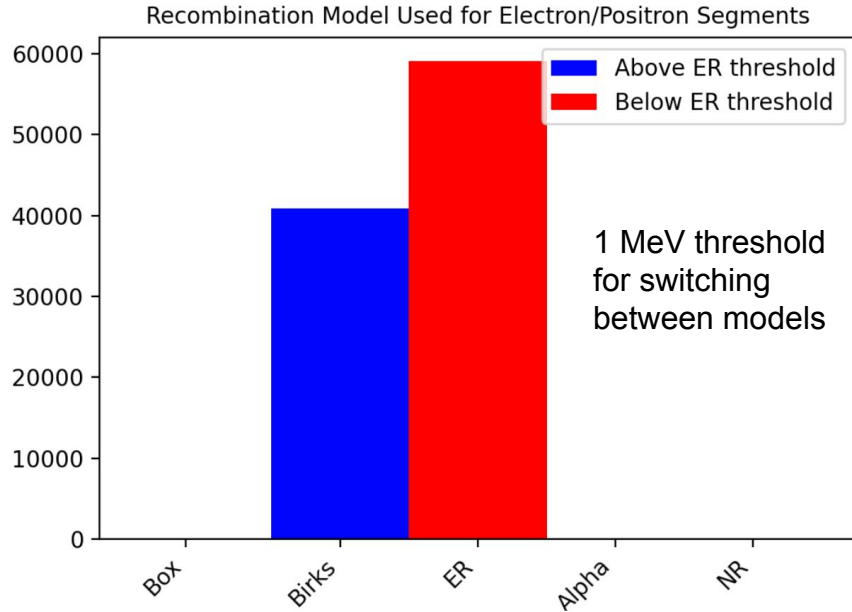
@njit
def NEST_ALPHA(E):
    """
    LArNEST alpha recombination model. https://github.com/NESTCollaboration/larnestpy

    Args:
        E (float): Segment dE in MeV.
    """
    return physics.ALPHA_R_FACTOR
```

# 'Cached-LArNEST' Implementation

- Benchmarking tests
  - On a MiniRun3 1e19 RHC file:
    - Cached-LArNEST: **0.536** seconds to run quenching
    - larnd-sim develop branch: **0.170** seconds to run quenching
    - 3.15x slower than original larnd-sim
  - On a 10k events 39Ar beta decay file (electrons < 565 keV)
    - Cached-LArNEST: **0.507** seconds to run quenching
    - larnd-sim develop branch: **0.165** seconds to run quenching
    - 3.07x slower than original larnd-sim
  - Note: The quenching speed depends on the number of data points in the cached LArNEST data due to the interpolation function.
  - Is this slow down acceptable?

# 'Cached-LArNEST' Implementation: MiniRun3 Tests



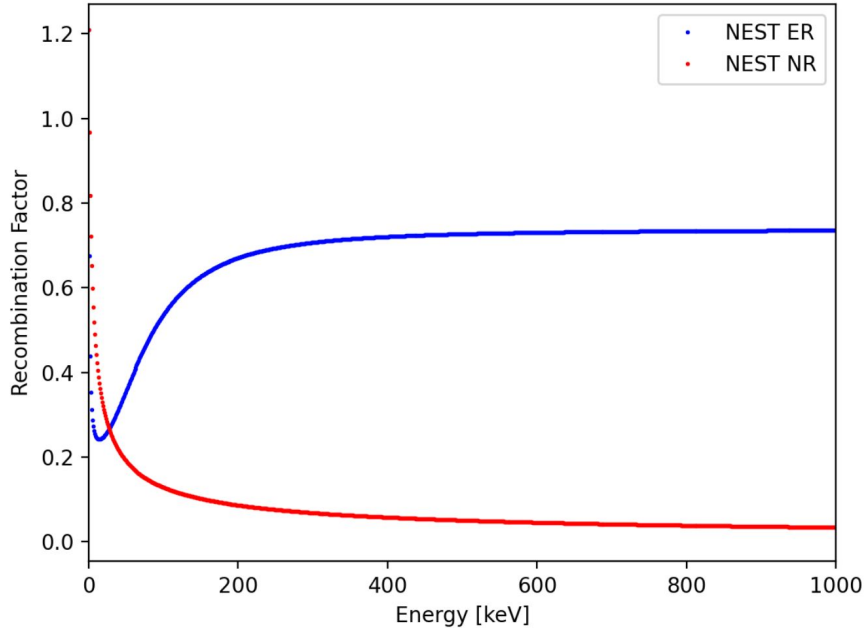
## Questions:

- At what energy should we switch from ER to Box/Birks?
- Should we smooth the transition between models, rather than having a hard cut off?



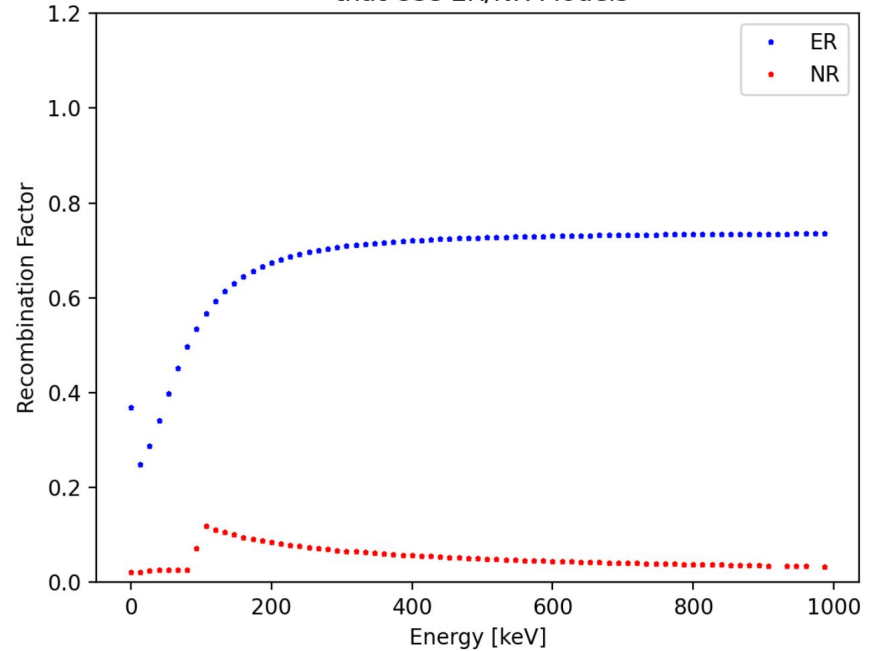
# 'Cached-LArNEST' Implementation: MiniRun3 Tests

Recombination Factors from LArNEST



Input to larnd-sim

Recombination Factors of Segments that Use ER/NR Models



From larnd-sim output

# Final Thoughts

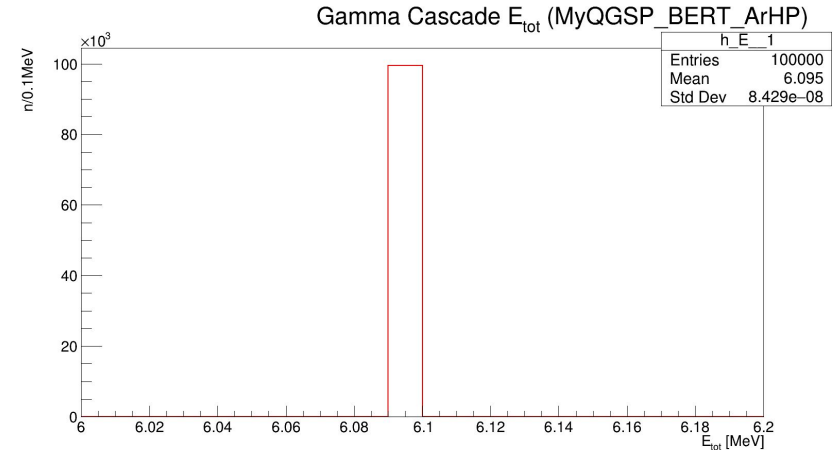
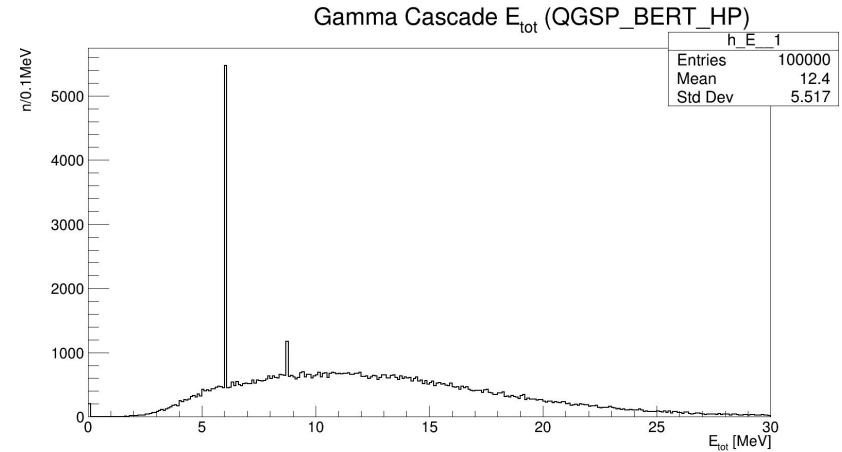
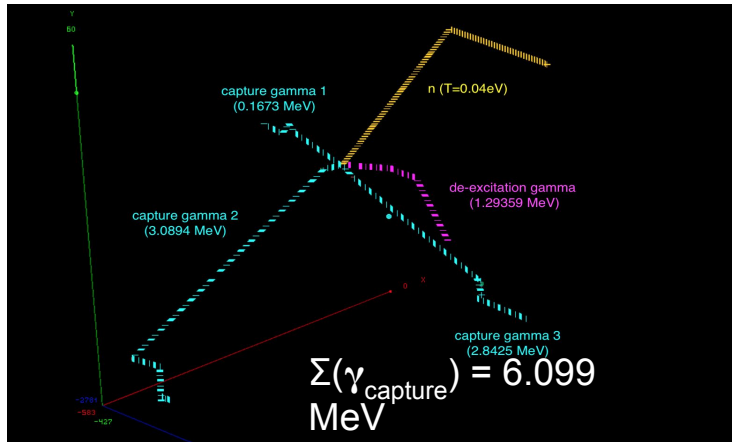
- We now have a version of larnd-sim that:
  - Can utilize the LArNEST electron recoil, alpha, and nuclear recoil recombination models
  - Allows for specifying different recombination models for different particles
- What kinds of tests would others like to see of this implementation?

# Modified *edepsim* physics list - profiling & validations

- Three physics lists considered:
  - **QGSP\_BERT**: Default physics list used in edepsim
  - **QGSP\_BERT\_HP**: High precision physics for neutrons up to 20 MeV
  - **MyQGSP\_BER\_ArHP**: QGSP\_BERT\_HP with corrected version of gamma cascade production for neutron captures on Argon

# Validation of gamma cascade energies

- Expect 6.1 MeV total energy for the gammas
- Custom physics list matches expectation



# Output file (n, $\gamma$ ) thresholds

- Neutrons and gammas have momentum thresholds for trajectory info storage in output files
- Hit info is kept, but trajectory info for particles including the physics process that led to its creation may be lost

```
--- a/src/edepsim-defaults-1.0.mac
+++ b/src/edepsim-defaults-1.0.mac
@@ -24,8 +24,8 @@
 /edep/db/set/requireEventsWithHits false
 /edep/db/set/lengthThreshold 1 mm
 /edep/db/set/trajectoryAccuracy 1 mm
-/edep/db/set/neutronThreshold 50 MeV
-/edep/db/set/gammaThreshold 10 MeV
+/edep/db/set/neutronThreshold 0 MeV
+/edep/db/set/gammaThreshold 0 MeV

#####
```

# Resource usage

- Head-to-head comparison of *same event* using example.gdml in edep-sim
- Single evt, with 10GeV proton burst (100 p) above the LArTracker
- Set the random seed for G4 simulation using: /edep/random/randomSeed

	QGSP_BERT (default cuts)	QGSP_BERT (zero cuts)	QGSP_BERT _HP (default cuts)	QGSP_BERT _HP (zero cuts)	MyQGSP_BE RT_ArHP (default cuts)	MyQGSP_BE RT_ArHP (zero cuts)
# Trajectories	45312	46122	53028	53901	91078	92450
Real Time (s)	260.98	257.25	1662.51	1748.50	1946.36	2007.01
Peak Memory (GB), Heap	1.40	1.402	2.634	2.668	3.532	3.528
File Size (MB)	17	17	21	22	24	28

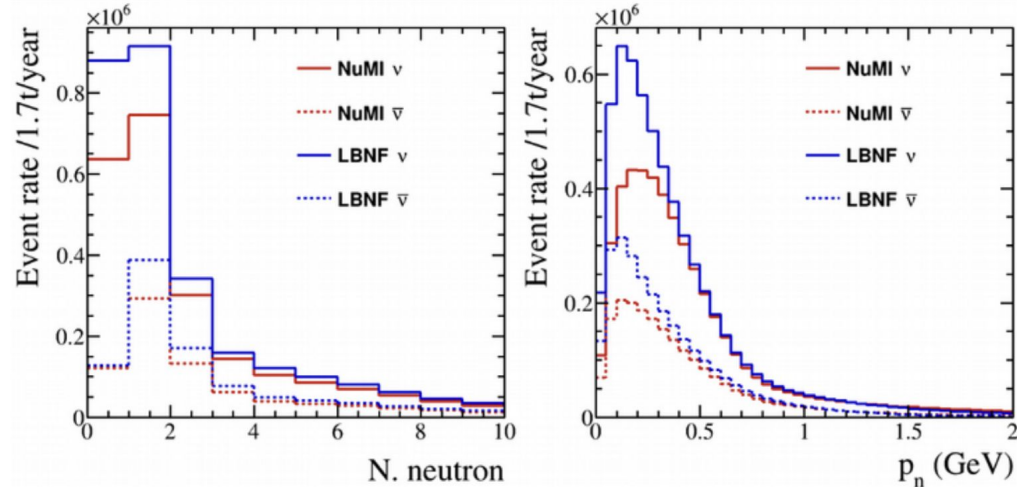
# n-Ar measurement

DUNE ND CDR  
arXiv:2103.13910

Current idea is to use the PDS system to tag neutron TOF to reconstruct KE.

Slides from Mike Mooney

([https://indico.fnal.gov/event/48610/contributions/212284/attachments/141944/179049/DUNE\\_NDLAr\\_AnalysisMeeting\\_21\\_04\\_08.pdf](https://indico.fnal.gov/event/48610/contributions/212284/attachments/141944/179049/DUNE_NDLAr_AnalysisMeeting_21_04_08.pdf))



- ◆ Many neutrons from LBNF beam at DUNE ND-LAr (67 t)
  - Above plots scaled to ArgonCube 2x2 demonstrator (1.7 t)
- ◆ Sizable fraction of neutron scatters will produce proton of high enough energy to tag neutron in TPC

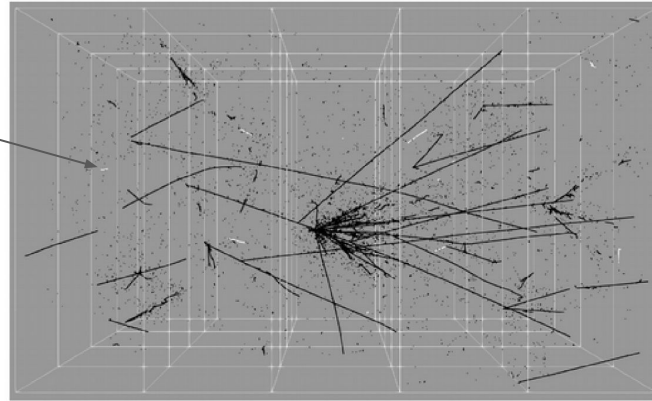
# n-Ar measurement

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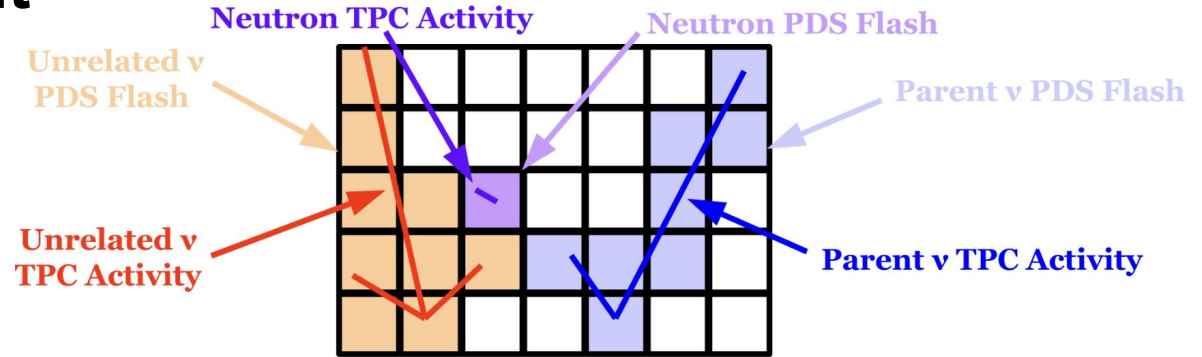


- ◆ Tag neutrons via scattering proton TPC tracks (white tracks in above), associate with parent neutrino interaction via light
  - Current PDS timing resolution specification ( $< 10$  ns) allows unambiguous matching to associated neutrino via scintillation light
  - *Can we do more with even better PDS timing precision?*



# n-Ar measurement

Current idea is to use the PDS system to tag neutron TOF to reconstruct KE.



Slides from Mike Mooney

([https://indico.fnal.gov/event/48610/contributions/212284/attachments/141944/179049/DUNE\\_NDLAr\\_AnalysisMeeting\\_21\\_04\\_08.pdf](https://indico.fnal.gov/event/48610/contributions/212284/attachments/141944/179049/DUNE_NDLAr_AnalysisMeeting_21_04_08.pdf))

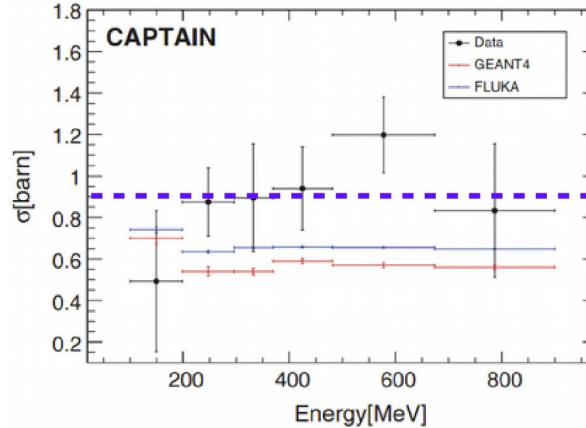
- ◆ Yes! With PDS timing resolution  $< 3$  ns, can measure neutron time of flight (TOF)  $\rightarrow$  neutron energy measurement
  - Energy from proton track in TPC less reliable – detector effects, nuclear effects such as Fermi motion wash things out
  - Reconstructed proton track coupled with TOF neutron energy measurement will provide data sample to tune above effects in MC
- ◆ Measure distance between  $\nu$  vertex and neutron scatter (proton) track, divide by time of flight (difference of PDS times)  $\rightarrow v \rightarrow E$ 
  - Modular ND-LAr design helps: keeps out late light from  $\nu$  interaction

# n-Ar measurement

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PhysRevLett.123.042502

$$dN_B/dx = -T\sigma_T N_B \rightarrow N_B(x) = N_0 e^{-T\sigma_T x}$$

$$T = \rho_{LAr} \times N_{Avogadro} / m_{Ar}$$

$$T \sim 2.11 \times 10^{22} \text{ cm}^{-3}, \sigma \sim 0.9 \text{ barns}$$
$$\rightarrow P_{\text{survival}} \sim e^{-x/D}, D \sim 50 \text{ cm}$$

- ◆ Use Mini-CAPTAIN measurements to determine neutron survival probability as function of distance from  $\nu$  vertex
  - Use  $\sigma \sim 0.9$  barns for all energies (approximation/simplification)
  - Corresponds to attenuation length  $D \sim 50$  cm
  - Weight events in average  $\Delta E/E$  plots according to  $P_{\text{survival}} \sim e^{-x/D}$
- ◆ Use minimum distance cut,  $x > 50$  cm, to improve  $\Delta E/E$

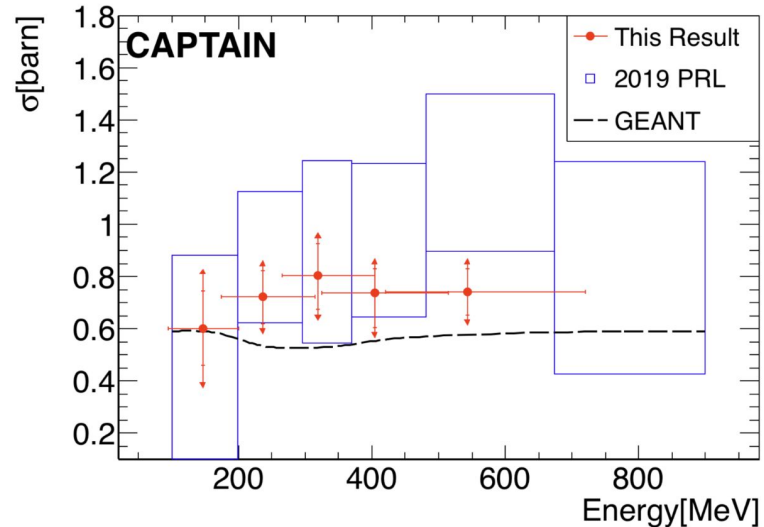
# n-Ar measurement

Current idea is to use the PDS system to tag neutron TOF to reconstruct KE.

- The described measurement is consistent with the hypothesis of small cross section change across considered energy range;
- Current PRD result [<https://doi.org/10.1103/PhysRevD.107.072009>]
- Previous PRL result [<https://doi.org/10.1103/PhysRevLett.123.042502>]

Slides from Sergey Martynenko

(<https://drive.google.com/drive/u/0/folders/1KzXxFFPBjOUBwub7Ta8pSR9jxfXXFc6M>)



# High purity proton sample - Bern cosmic ray data

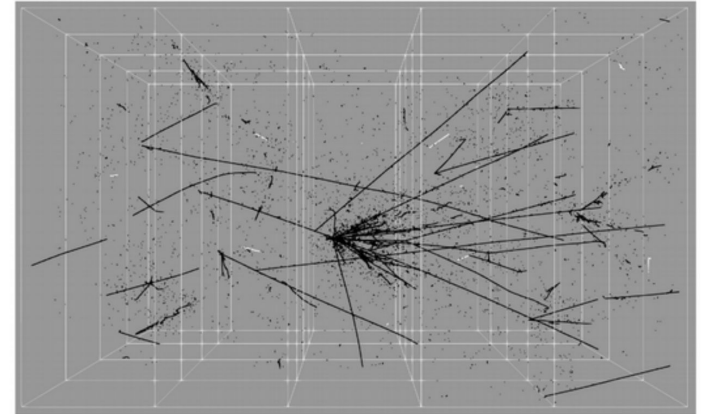
Initial investigation into proton-related detector systematics using Bern module data

Preliminary proton selection using <4 hours of Module-1 data on hand

- Charge + light matching
- Track fitting with Hough transform
- Field response unfolding
- PIDA to discriminate HIPs from MIPs
- Void (inactive channel) analysis
- ML approach using Blip

Work in progress in coordination with  $\nu_{\mu} C\bar{C}0\pi$  analysis

- Recast analysis from “flow” files
- Analysis validation



**BACKUP**

# Custom neutron physics list

- **Name:** **MyQGSP\_BERT\_ArHP**
- Used in ProtoDUNE-SP for simulations
- Integrated w/ edep-sim ([feature/neutron\\_physics\\_DR](#))
- Based on QGSP\_BERT\_HP reference physics list in GEANT4
  - HP refers to high precision neutron physics
  - Applies to low energy neutrons (< 20 MeV)
  - Custom neutron capture physics (from J. Wang) corrects the gamma cascade production to match NNDC tables for Ar40
- Other changes:
  - Set proton range cut to 0
    - This keeps low energy nuclear recoils

# National Nuclear Data Center tools

- <https://www.nndc.bnl.gov/capgam/>
- CapGam by Target provides data on thermal neutron captures on Ar-40 and Ar-41

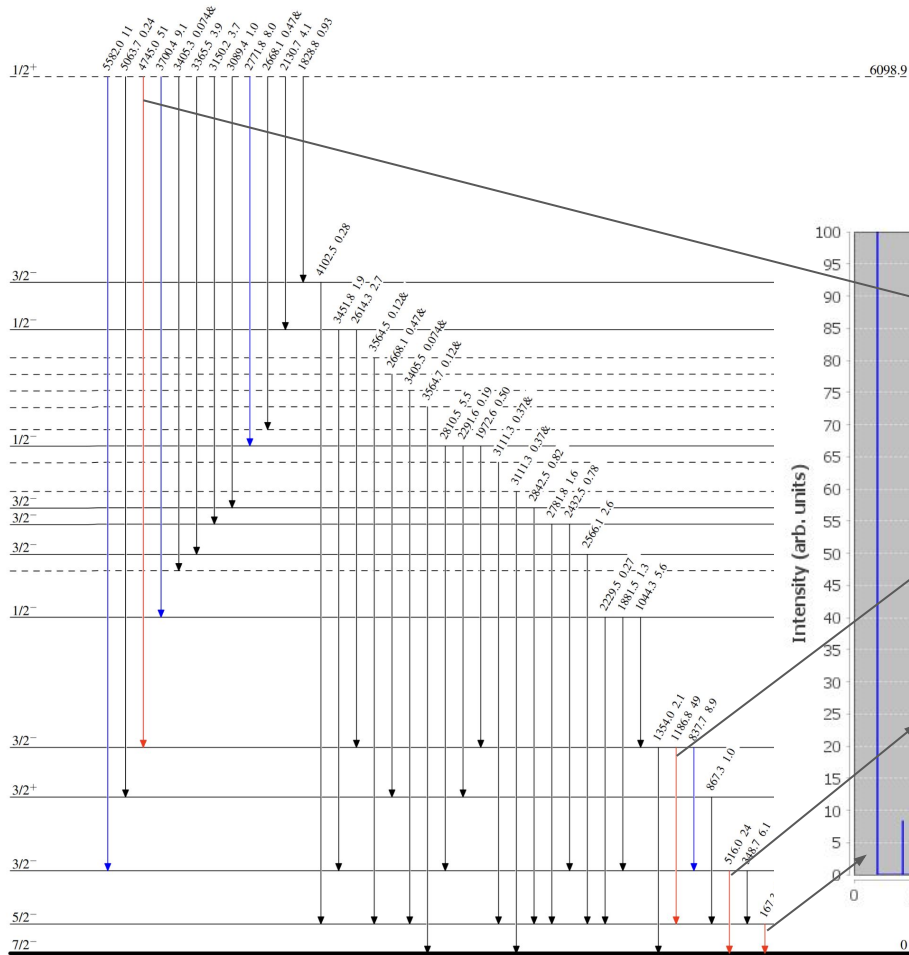
- [NNDC Site Index](#)
- **CapGam**
  - [About CapGam](#)
  - [CapGam by Energy](#)
  - [CapGam by Target](#)
- **Resources**
  - [ENSDF](#)
  - [Nuclear Data Sheets](#)
- **Networks**
  - [USNDP](#)
  - [NSDD](#)

## Thermal Neutron Capture $\gamma$ 's (CapGam)

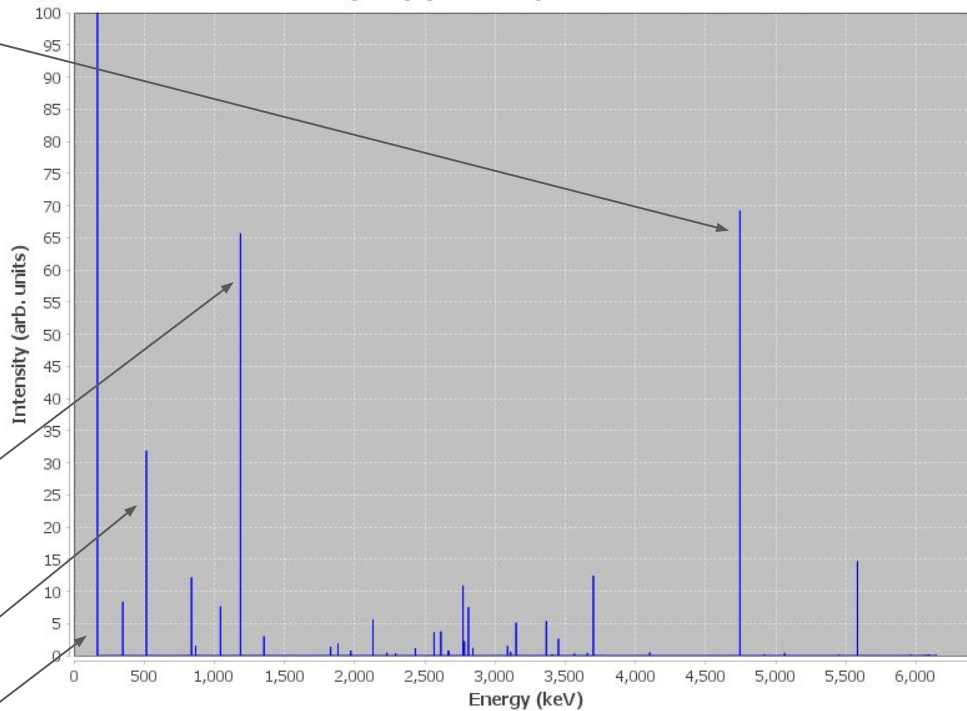
The energy and photon intensity with uncertainties of gamma rays as seen in thermal-neutron capture are presented in two tables, one in ascending order of [gamma energy](#) and a second organized by [Z, A of the target](#). In the energy-ordered table the three strongest transitions are indicated in each case. The nuclide given is the target nucleus in the capture reaction. The gamma energies given are in keV. The gamma intensities given are relative to 100 for the strongest transition. %I<sub>γ</sub> (per 100 n-captures) for the strongest transition is given, where known.

All data are taken from [Evaluated Nuclear Structure Data File \(ENSDF\)](#), a computer file of evaluated nuclear structure data and [eXperimental Unevaluated Nuclear Data List \(XUNDL\)](#), both maintained by the National Nuclear Data Center, Brookhaven National Laboratory, on behalf of the [U.S. Nuclear Data Program](#) and [Nuclear Structure and Decay Data network](#). The data for  $A > 20$  are published in the [Nuclear Data Sheets](#), Elsevier. The data for  $A \leq 20$  are published in [Nuclear Physics A](#).

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### 40AR(N,G),(POL N,G) E=THERMAL





# Simulated event for profiling

