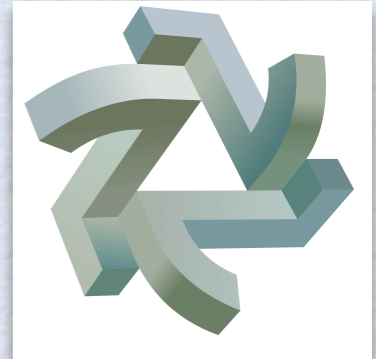


# Study of Neutrino-Antineutrino Transitions in MINOS



Richa Sharma,  
for the MINOS Collaboration

Panjab University, Chandigarh  
India-Fermilab Neutrino Collaboration

Advisor: Dr. Vipin Bhatnagar, Panjab University, Chandigarh  
Co-advisor: Prof. Brajesh C Choudhary, University of Delhi, Delhi

Working closely with: Dr. Robert K Plunkett, Fermilab



# Outline



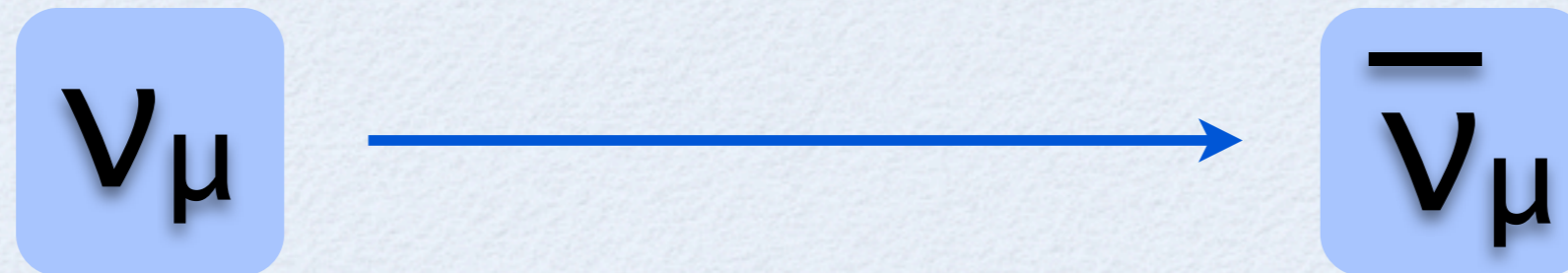
- Introduction
- Discussion of the Model
- Overview of the MINOS Experiment
- Analysis Procedure
- Preliminary Results and Future Plans



# Introduction



- Lorentz and CPT symmetry are inherent in Standard Model and General Theory of Relativity.
- Tiny violations of Lorentz and CPT symmetry might be observable in experiments and can lead to interesting phenomena.
- Lorentz and CPT violating couplings in the Standard Model Extension (SME) can describe neutrino-antineutrino oscillations.\*



- The aim of the present work is to study the probability of oscillation of muon neutrinos to muon antineutrinos and to find the sensitivity of MINOS to such oscillations.

\*V. Alan Kostelecky and Matthew Mewes, PhysRevD.69, 016005 (2004)

\*Sebastian Hollenberg, Octavian Micu, and Heinrich Päs, PhysRevD.80.053010 (2009)



# The Model

- Considering two neutrino generations, the Schrodinger equation reads:

$$i \frac{d}{dt} \begin{pmatrix} \nu_\mu \\ \nu_\tau \\ \bar{\nu}_\mu \\ \bar{\nu}_\tau \end{pmatrix} = h_{eff} \begin{pmatrix} \nu_\mu \\ \nu_\tau \\ \bar{\nu}_\mu \\ \bar{\nu}_\tau \end{pmatrix}$$

- If we assume Lorentz and CPT violation, and consider a set of four coefficients  $b_\mu$ ,  $b_\tau$ ,  $c_{\mu\mu}$ ,  $c_{\tau\tau}$  which govern it, we can get **neutrino-antineutrino oscillations**, and as well as **altered dispersion relations** in the standard  $\nu_\mu \rightarrow \nu_\tau$  and  $\bar{\nu}_\mu \rightarrow \bar{\nu}_\tau$  sectors.

$$h_{eff} = \text{diag} \left( E + \frac{\sum m^2}{4E} \right) + \begin{pmatrix} -\frac{\Delta m^2}{4E} \cos 2\theta - \frac{C_{\mu\mu}}{2E} & \frac{\Delta m^2}{4E} \sin 2\theta & \frac{B_\mu}{2E} & 0 \\ \frac{\Delta m^2}{4E} \sin 2\theta & \frac{\Delta m^2}{4E} \cos 2\theta - \frac{C_{\tau\tau}}{2E} & 0 & \frac{B_\tau}{2E} \\ \frac{B_\mu}{2E} & 0 & -\frac{\Delta m^2}{4E} \cos 2\theta - \frac{C_{\mu\mu}}{2E} & \frac{\Delta m^2}{4E} \sin 2\theta \\ 0 & \frac{B_\tau}{2E} & \frac{\Delta m^2}{4E} \sin 2\theta & \frac{\Delta m^2}{4E} \cos 2\theta - \frac{C_{\tau\tau}}{2E} \end{pmatrix}$$

$$B_\mu(E) = b_\mu E^2$$

$$C_{\mu\mu}(E) = c_{\mu\mu} E^2$$

$$B_\tau(E) = b_\tau E^2$$

$$C_{\tau\tau}(E) = c_{\tau\tau} E^2$$

# Resonant Mixing

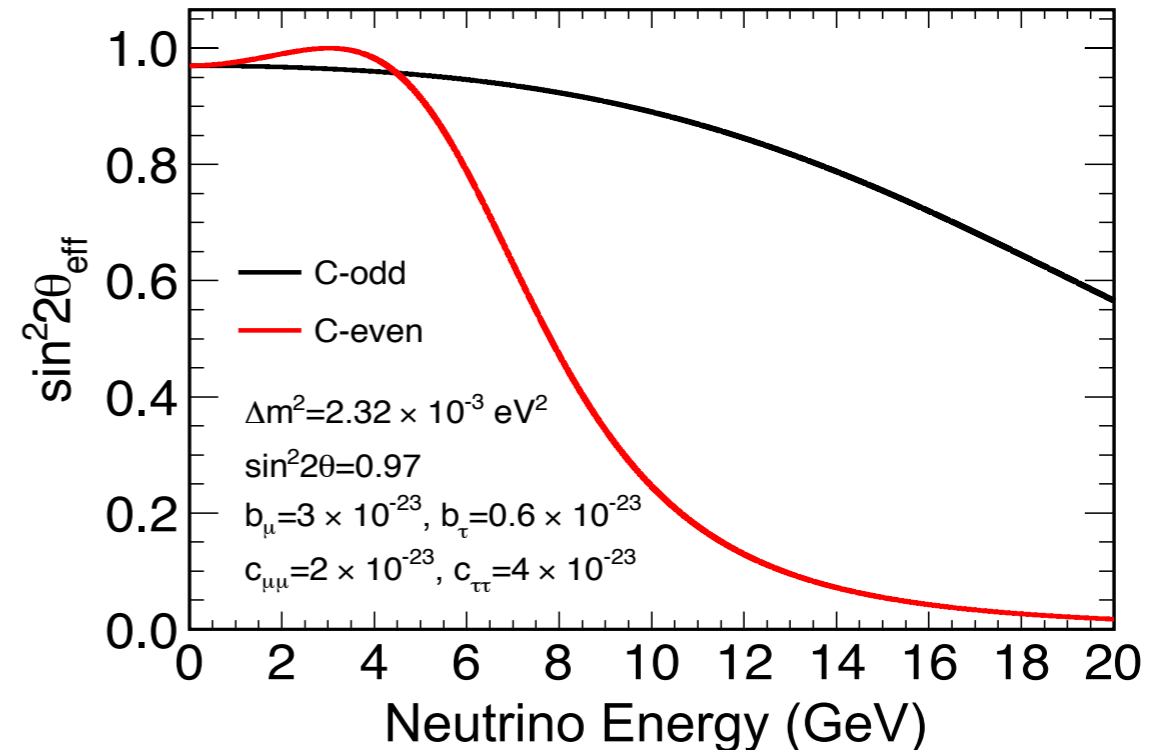


- The effective hamiltonian can be diagonalised:

$$i \frac{d}{dt} \begin{pmatrix} \nu^- \\ \nu^+ \end{pmatrix} = \begin{pmatrix} h_{\text{eff}}^{C\text{-odd}} & 0 \\ 0 & h_{\text{eff}}^{C\text{-even}} \end{pmatrix} \begin{pmatrix} \nu^- \\ \nu^+ \end{pmatrix}.$$

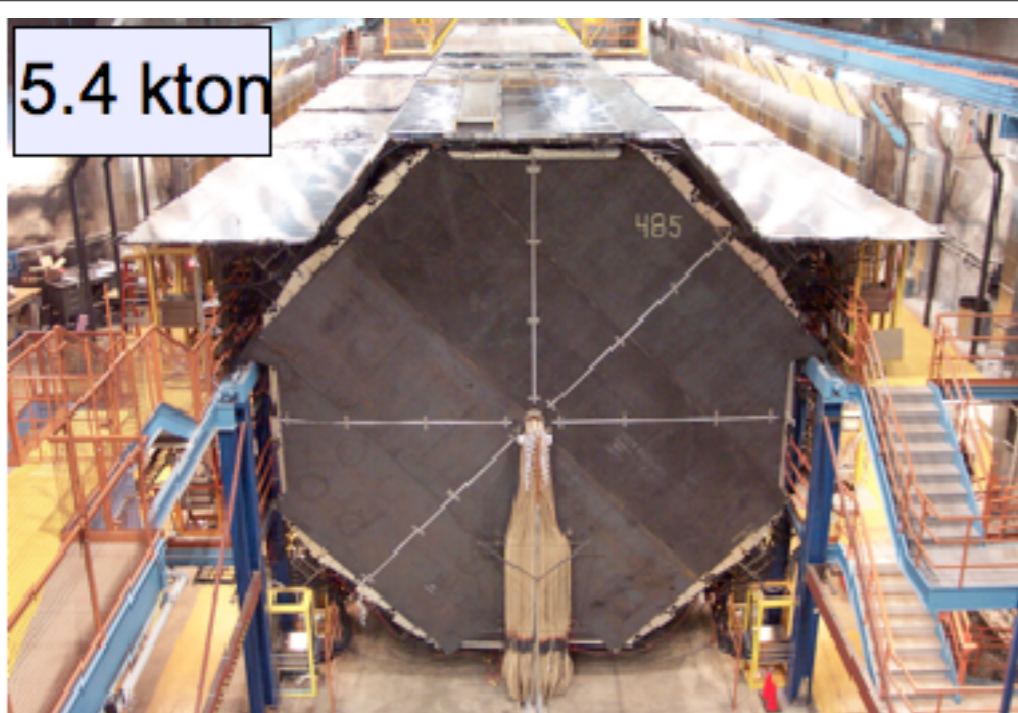
$C\nu^- = -\nu^-$  ; C-odd states

$C\nu^+ = \nu^+$  ; C-even states



- For certain values of parameters  $b_\mu$ ,  $b_\tau$ ,  $c_{\mu\mu}$ ,  $c_{\tau\tau}$ , resonant mixing for C-odd and C-even states occurs.
- The neutrino-antineutrino mixing occurs only between C-even states and C-odd states
- The condition for resonance needs to be satisfied.

5.4 kton

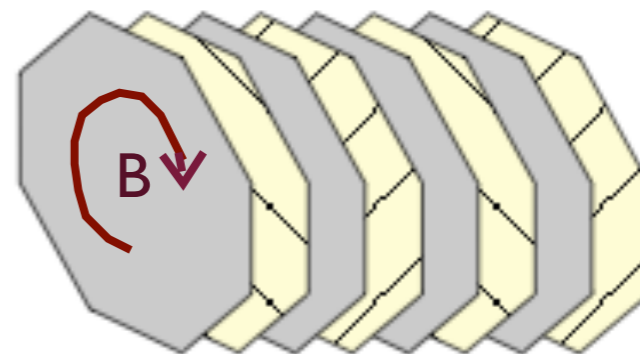


# The MINOS Experiment

- NuMI high intensity neutrino beam at Fermilab (Average power ~340 kW)
- **Near Detector:**
  - 100m deep, 1 km from source.
  - Measure beam composition and energy spectrum.
- **Far Detector:**
  - 700m deep, 735 km from source.
  - Search for evidence of oscillations.



- Alternating layers of 2.54cm steel and 1cm plastic scintillator with WLS and clear fibre.
- ~1.3T magnetic field.



## Data collected:

- $10.71 \times 10^{20}$  POT (neutrino-optimised mode)
- $3.36 \times 10^{20}$  POT (antineutrino-optimised mode)

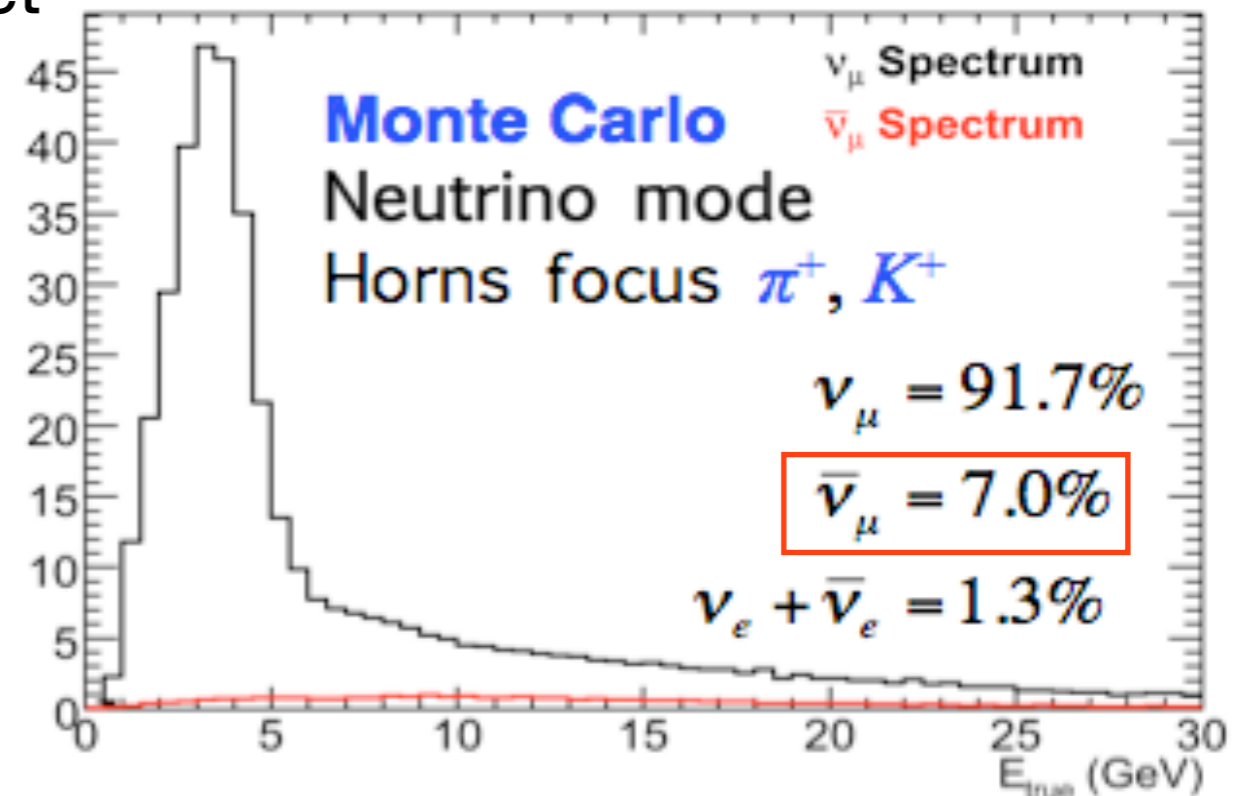
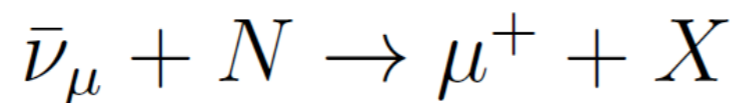


# Selecting Antineutrinos

Used the neutrino-optimised MC dataset

- ▶ 7% antineutrinos at the Far Detector.
- ▶ Higher energy than neutrinos.

Antineutrinos interact as:



Positively charged tracks with interaction vertex inside the detector.

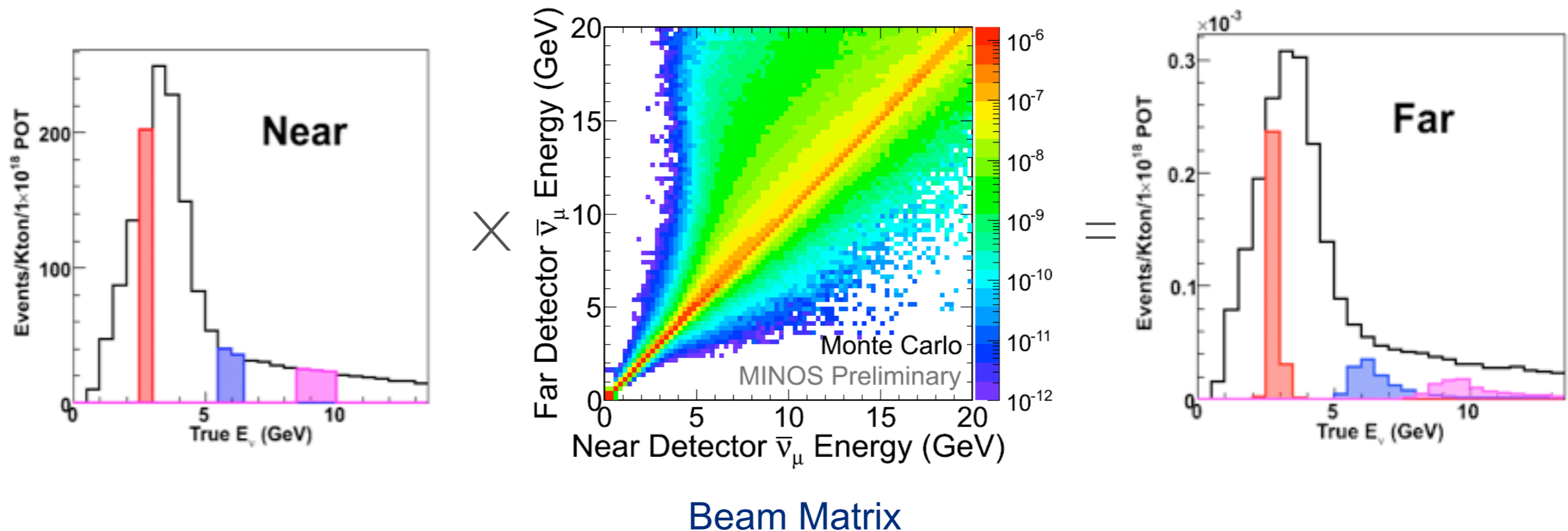
To reduce the misidentified NC and  $\nu_{\mu}$  CC background:

- ▶ A discriminant variable - formed from 3 variables describing track properties.
- ▶ Confidence of charge-sign determination from track fit.
- ▶ Compare the track direction at the vertex to that at the end of the track.



# Near To Far Extrapolation

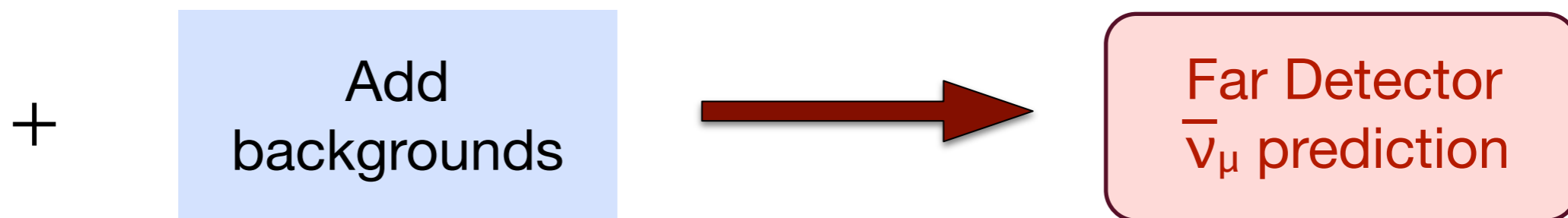
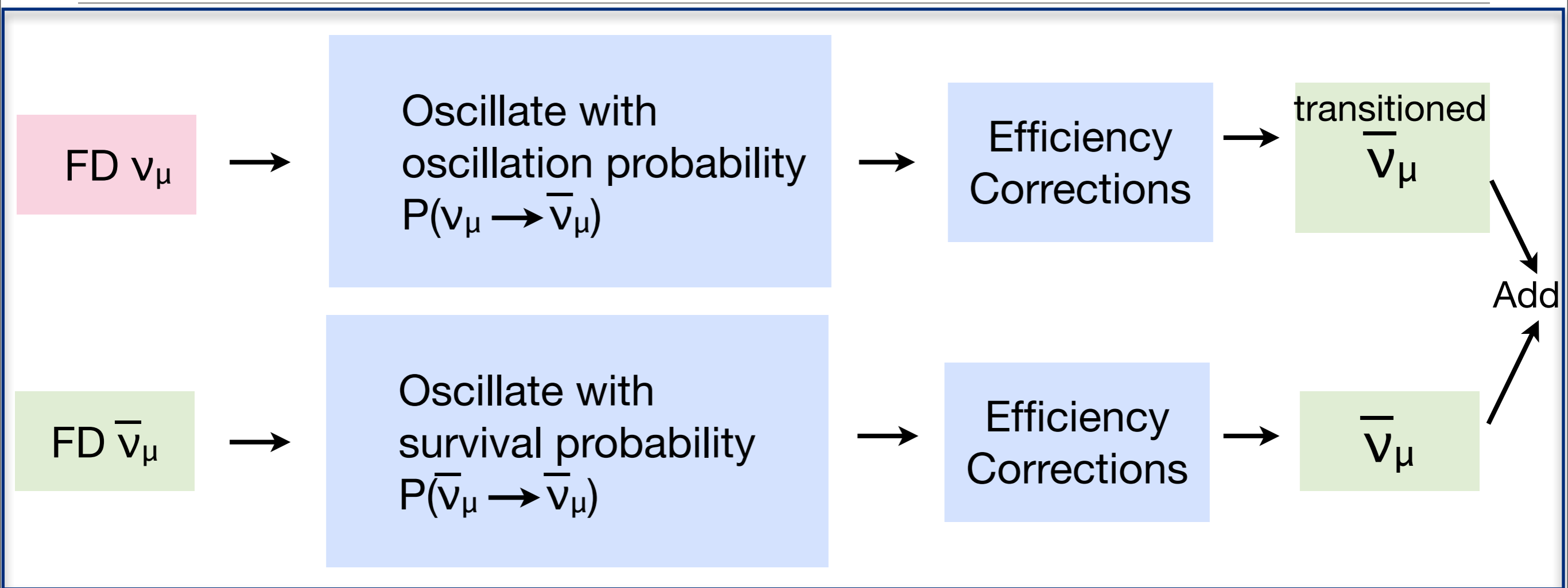
- The ND spectrum is used to predict the FD spectrum.
- Flux and cross-section uncertainties cancel.
- Using the Matrix Method extrapolation framework adapted for antineutrino oscillation analysis.







# Generating FD $\bar{\nu}_\mu$ Prediction

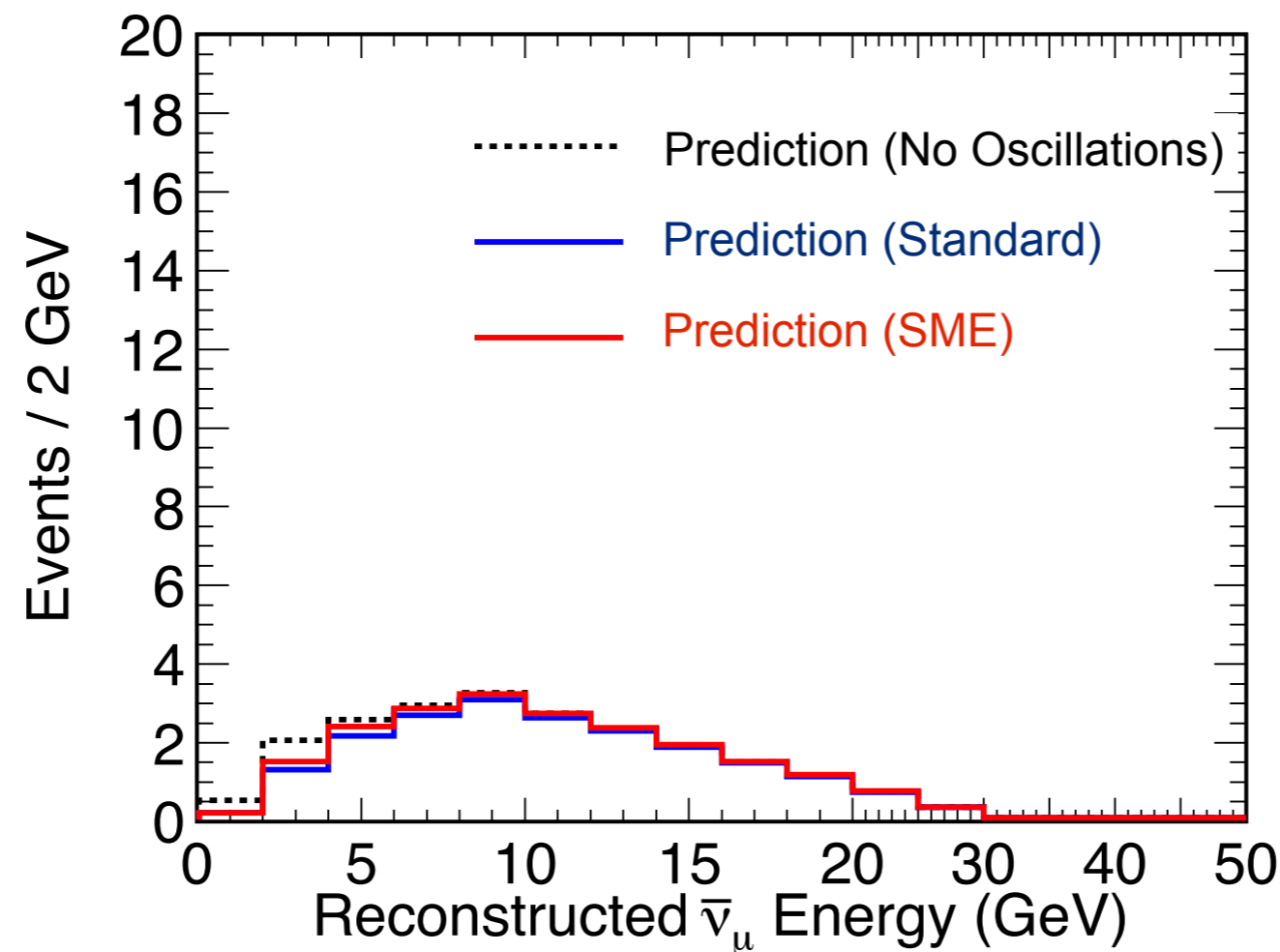




# Far Detector $\bar{\nu}_\mu$ Prediction

The Far Detector prediction is made using the following parameters:

- ▶  $\Delta m^2 = 2.32 \times 10^{-3} \text{ eV}^2$ ,  $\sin^2 2\theta = 0.97$
- ▶  $b_\mu = 3 \times 10^{-23}$ ,  $b_\tau = 0.6 \times 10^{-23}$ ,  $c_{\mu\mu} = 2 \times 10^{-23}$ ,  $c_{\tau\tau} = 4 \times 10^{-23}$



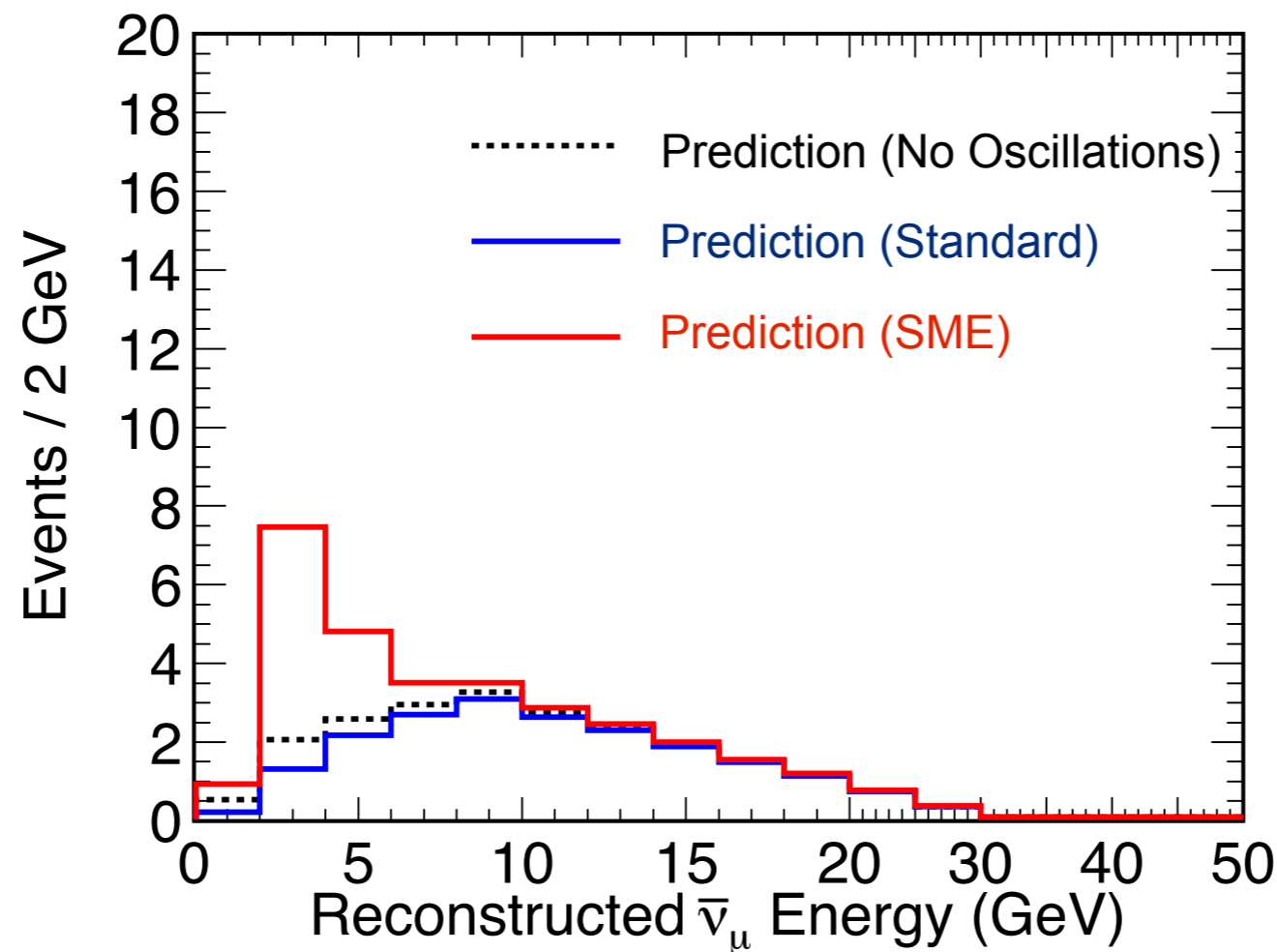
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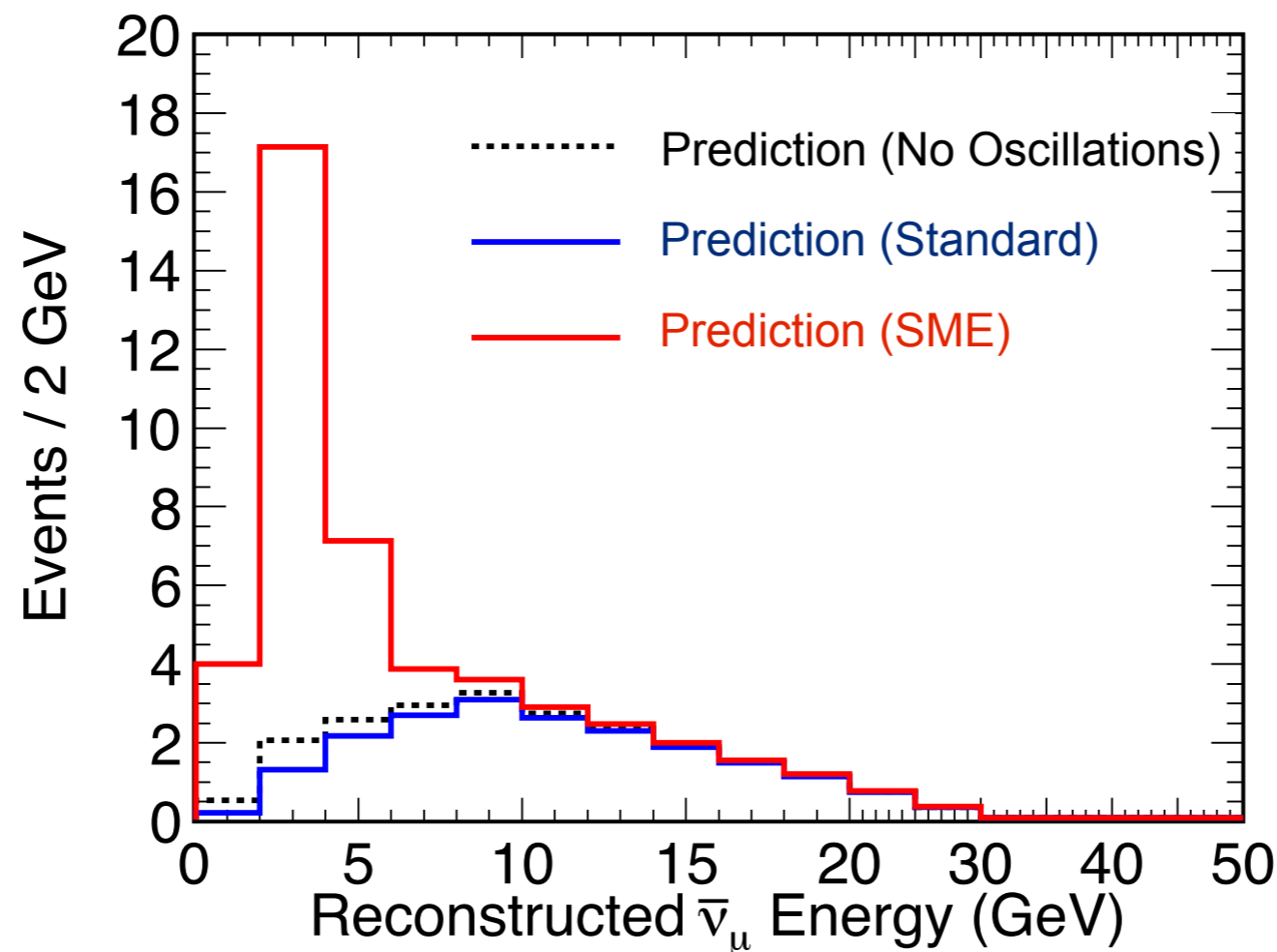
The number of  $\nu_\mu$  transitioning to  $\bar{\nu}_\mu$  increases as the values of the parameters is increased.



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# Summary and Outlook



- Demonstrated that the  $\nu_\mu$  to  $\bar{\nu}_\mu$  transitions are governed by the size of Lorentz and CPT violating parameters.
- It is possible to see the signal for such transitions in MINOS.
- Working towards obtaining the sensitivity of MINOS to transitions and to obtain the parameters  $b_\mu$ ,  $b_\tau$ ,  $c_{\mu\mu}$ ,  $c_{\tau\tau}$  that govern them.
- This would be the major part of my thesis. First analysis with this model.
- I have also worked on the analysis of the 7%  $\bar{\nu}_\mu$  component in the  $7.1 \times 10^{20}$  POT neutrino-optimised data.
- This analysis excluded the  $(3.37 < |\Delta\bar{m}^2| < 1000) \times 10^{-3} \text{ eV}^2$  at 90% C.L., assuming  $\sin^2 2\bar{\theta} = 1$ .
- The results have been published : P. Adamson et al. (MINOS Collaboration), **“Search for the disappearance of muon antineutrinos in the NuMI neutrino beam”**, Phys. Rev. D 84, 071103(R) (2011)

# THANK YOU



# BACKUP



# Relation of $b_\mu$ , $b_\tau$ , $c_{\mu\mu}$ and $c_{\tau\tau}$ coefficients to more fundamental coefficients

The LV equations of motion can be written in analogy to the Dirac Equation:

$$(i\Gamma_{AB}^\nu \partial_\nu - M_{AB})\nu_B = 0,$$

where

$$\Gamma_{AB}^\nu = \gamma^\nu \delta_{AB} + c_{AB}^{\mu\nu} \gamma_\mu + d_{AB}^{\mu\nu} \gamma_5 \gamma_\mu + e_{AB}^\nu + if_{AB}^\nu + \frac{1}{2} g_{AB}^{\lambda\mu\nu} \sigma_{\lambda\mu},$$

$$M_{AB} = m_{AB} + im_{5AB} \gamma_5 + a_{AB}^\mu \gamma_\mu + b_{AB}^\mu \gamma_5 \gamma_\mu + \frac{1}{2} H_{AB}^{\mu\nu} \sigma_{\mu\nu}.$$

$$c_{\mu\mu} = 2(c_L)^{\tau\tau}_{\mu\mu} [1 + \cos^2\Theta]$$

$$c_{\tau\tau} = 2(c_L)^{\tau\tau}_{\tau\tau} [1 + \cos^2\Theta]$$

$$b_\mu = -i \sin 2\Theta \tilde{g}^{Z\tau}_{\mu\bar{\mu}}$$

$$b_\tau = -i \sin 2\Theta \tilde{g}^{Z\tau}_{\tau\bar{\tau}}$$

$$c_L = c + d$$

$$\tilde{g}^{Z\tau} \equiv g^{0Z\tau} + (i/2) \epsilon^{0Z}_{\lambda\rho} g^{\lambda\rho\tau},$$

$\Theta$  is the celestial colatitude



- Maximal experimental sensitivities attained for coefficients in the neutrino sector of minimal SME
- Experimental limits for  $(c_L)^{\tau\tau}_{\mu\mu}$ ,  $(c_L)^{\tau\tau}_{\tau\tau}$ ,  $\tilde{g}^{Z\tau}_{\mu\bar{\mu}}$  and  $\tilde{g}^{Z\tau}_{\tau\bar{\tau}}$  have not been obtained yet.

$d = 4$	Coefficient	$e\mu$	$e\tau$	$\mu\tau$	Coefficient	$e\mu$	$e\tau$	$\mu\tau$
	$\text{Re}(c_L)^{XY}$	$10^{-21}$	–	$10^{-23}$	$\text{Im}(c_L)^{XY}$	$10^{-21}$	–	$10^{-21}$
	$\text{Re}(c_L)^{XZ}$	$10^{-21}$	–	$10^{-23}$	$\text{Im}(c_L)^{XZ}$	$10^{-21}$	–	$10^{-21}$
	$\text{Re}(c_L)^{YZ}$	$10^{-21}$	–	$10^{-23}$	$\text{Im}(c_L)^{YZ}$	$10^{-21}$	–	$10^{-21}$
	$\text{Re}(c_L)^{XX}$	$10^{-21}$	–	$10^{-23}$	$\text{Im}(c_L)^{XX}$	$10^{-21}$	–	$10^{-21}$
	$\text{Re}(c_L)^{YY}$	$10^{-21}$	–	$10^{-23}$	$\text{Im}(c_L)^{YY}$	$10^{-21}$	–	$10^{-21}$
	$\text{Re}(c_L)^{ZZ}$	$10^{-19}$	–	–	$\text{Im}(c_L)^{ZZ}$	–	–	–
	$\text{Re}(c_L)^{TT}$	$10^{-19}$	–	–	$\text{Im}(c_L)^{TT}$	–	–	–
	$\text{Re}(c_L)^{TX}$	$10^{-22}$	–	$10^{-27}$	$\text{Im}(c_L)^{TX}$	$10^{-22}$	–	$10^{-22}$
	$\text{Re}(c_L)^{TY}$	$10^{-22}$	–	$10^{-27}$	$\text{Im}(c_L)^{TY}$	$10^{-22}$	–	$10^{-22}$
	$\text{Re}(c_L)^{TZ}$	$10^{-20}$	–	–	$\text{Im}(c_L)^{TZ}$	–	–	–



$$P(\nu_\mu \rightarrow \bar{\nu}_\mu)$$

$$\begin{aligned} P(\nu_\mu \rightarrow \bar{\nu}_\mu) &= \frac{1}{4} * \left[ \left( 1 - \sin^2 2\theta_{c-odd} \sin^2 \left( \Delta m^2 \frac{L}{4E} \right) \right) + \left( 1 - \sin^2 2\theta_{c-even} \sin^2 \left( \Delta m^2 \frac{L}{4E} \right) \right) \right] \\ &- \frac{1}{2} * \left[ \cos^2 \theta_{c-odd} \cos^2 \theta_{c-even} + \sin^2 \theta_{c-odd} \sin^2 \theta_{c-even} \right. \\ &\quad + \cos^2 \theta_{c-odd} \sin^2 \theta_{c-even} \cos \left( \Delta m^2 \frac{L}{4E} \right) \\ &\quad \left. + \sin^2 \theta_{c-odd} \cos^2 \theta_{c-even} \cos \left( \Delta m^2 \frac{L}{4E} \right) \right] = P(\bar{\nu}_\mu \rightarrow \nu_\mu) \end{aligned}$$

The relation between the effective and the standard mixing angles for C-odd and C-even states is given by:

$$\tan 2\theta_{c-odd} = \frac{\Delta m^2 \sin 2\theta}{((b_\mu - b_\tau + c_{\mu\mu} - c_{\tau\tau}) E^2 + \Delta m^2 \cos 2\theta)}$$

$$\tan 2\theta_{c-even} = \frac{\Delta m^2 \sin 2\theta}{((-b_\mu + b_\tau + c_{\mu\mu} - c_{\tau\tau}) E^2 + \Delta m^2 \cos 2\theta)}$$

# Modified probabilities

Standard:

- $\Delta m^2 = 2.32 \times 10^{-3} \text{ eV}^2, \sin^2 2\theta = 0.97$
- $b_\mu = 0, b_\tau = 0, c_{\mu\mu} = 0, c_{\tau\tau} = 0$

New:

- $\Delta m^2 = 2.32 \times 10^{-3} \text{ eV}^2, \sin^2 2\theta = 0.97$
- $b_\mu = 3 \times 10^{-21}, b_\tau = 0.6 \times 10^{-21}, c_{\mu\mu} = 2 \times 10^{-21}, c_{\tau\tau} = 4 \times 10^{-21}$

