



Broadband Parametric Amplifiers for Dark Matter Searches and QIS

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4 October 2023 -- BREAD Collaboration Meeting





Particle physicists developing qubits and quantum sensors for low-threshold particle detection

QSC@FNAL:

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Lauren Hsu
Daniel Baxter
Rakshya Khatiwada
Daniel Bowring
Gustavo Cancelo
Sho Uemura
Sami Lewis
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Sara Sussman
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Sunil Golwala
Karthik Ramanathan
Osmond Wen
Taylor Aralis
Brandon Sandoval

SLAC:

Noah Kurinsky
Kelly Stifter
Hannah Magoon
Sukie Kevane

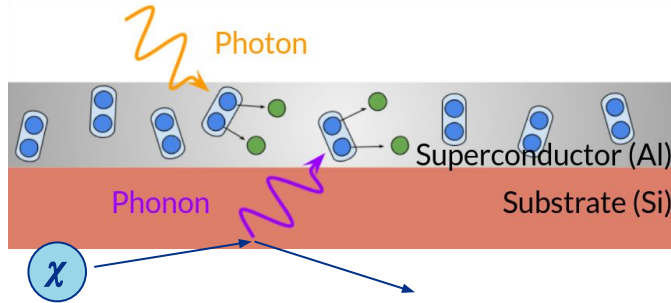
QSC@Purdue

Alex Ma
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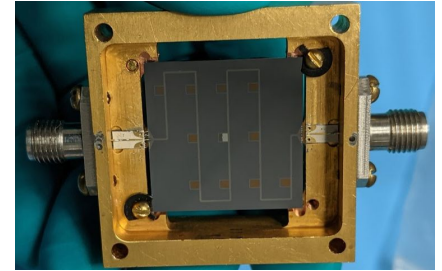
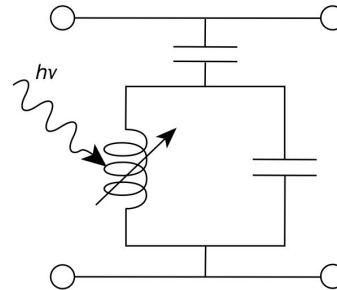
UW Madison

Robert McDermott
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Superconducting RF Sensors for Particle DM Detection



Kinetic Inductance Phonon-Mediated Detectors



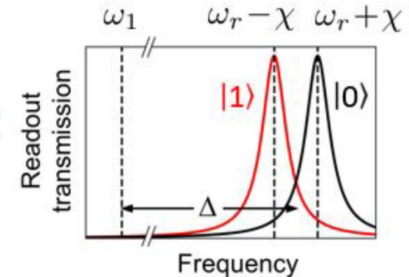
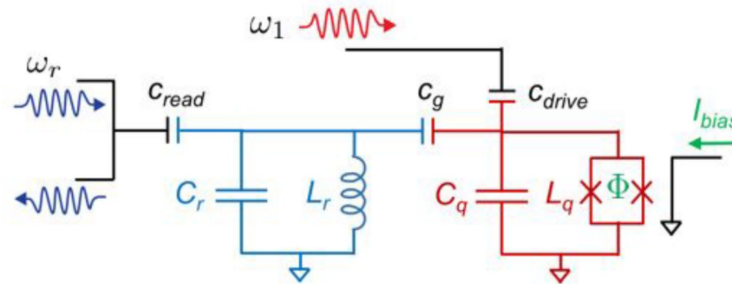
RF readout, frequency-multiplexed devices

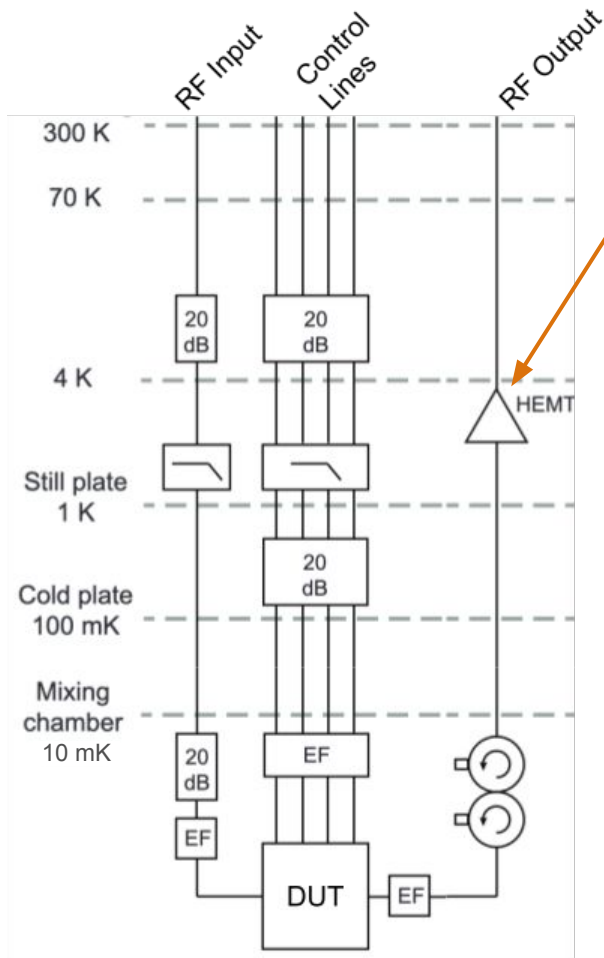
- Common readout electronics/scheme
- **Common challenge: noise**

Superconducting Qubits

KIPMDs → energy resolution

Qubits → single shot fidelity





Typical Cryogenic Setup

First stage amplification: HEMT

- Widely commercially available
- Low noise ($T_n \sim 4\text{-}10\text{ K}$)

Reducing amplifier noise temperature \rightarrow improved SNR

Quantum-limited amplifier as first stage amplification

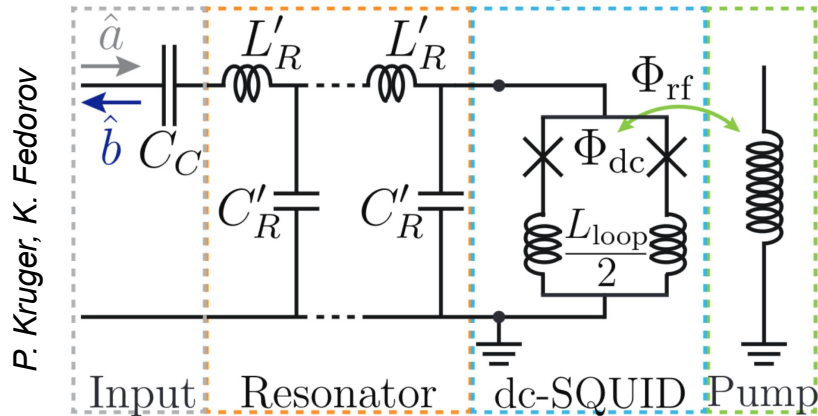
- SQL @ 6 GHz: 150 mK
- Noise scales linearly with T_n

\rightarrow Direct improvement in KIPMD energy resolution & qubit single shot fidelity (SSF)

Option 1: Josephson Parametric Amplifier

Single resonance feature provides amplification by extracting energy of pump tone

- Pro: commercially available
- Con: narrow gain band: few -- 700 MHz
 - Flux tuning allows user to select frequency of maximal gain
- Con: low saturation power



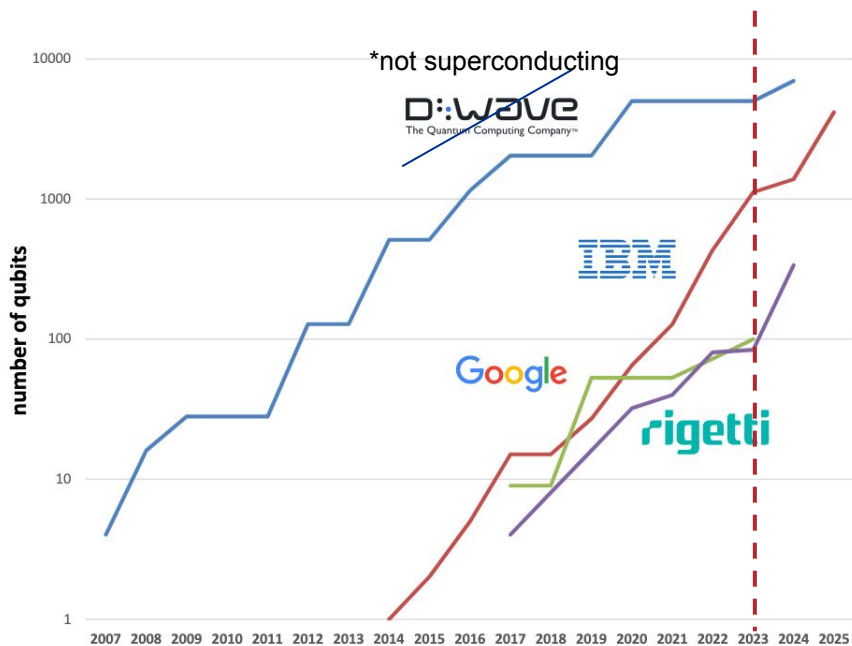
Raytheon BBN JPA

Typical Performance Characteristics

Parameter	Typical Value	Units
Frequency Range	5.0-7.0	GHz
Bandwidth	300	MHz
Gain	20	dB
Noise Temperature	295	mK
Input Power 1dB Compression (P1dB)	-107.5	dBm
Flux Bias Current Periodicity ($1\Phi_0$)	4	mA
3 Wave Operation Pump Power	-45	dBm
4 Wave Operation Pump Power	-75	dBm

Untenable solution for large arrays required for computation, particle detection.

Scaling Up: Qubits & KIPMDs

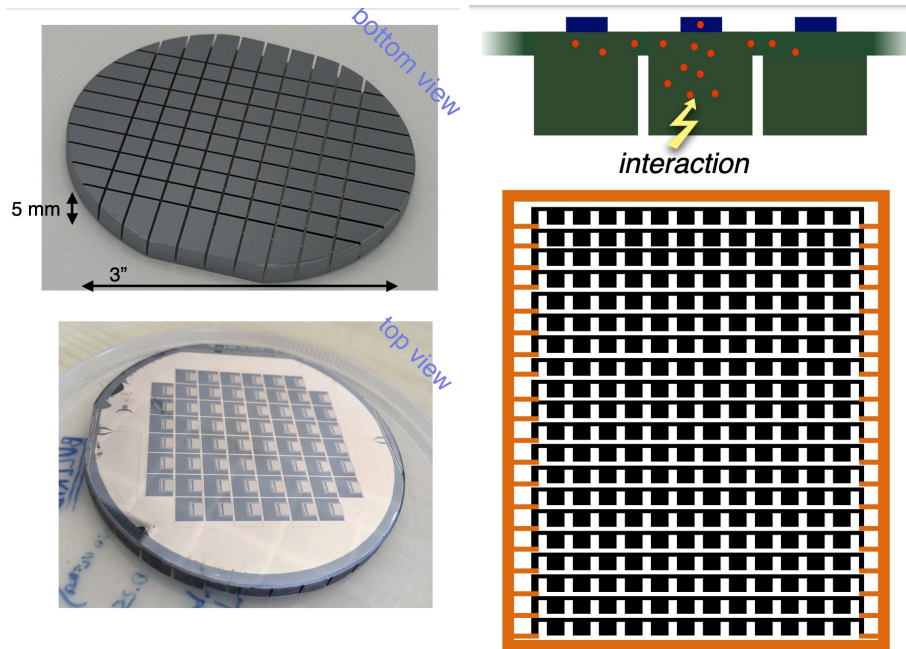


(cc) Olivier Ezratty, 2023

20 x 10⁶ qubits to break RSA-2048 encryption

arXiv:1905.09749

BULLKID Concept (M. Vignati)



60 resonators/wafer x 20 wafers/stack

→ **1200 resonators!**

Option 2: Traveling Wave Parametric Amplifiers

Signal amplification obtained by injecting a pump tone into a nonlinear transmission medium which modulates the line's inductance per unit length.

- Josephson effect (J-TWPA): Long array of Josephson junctions/SQUIDs (~ 100 s)
- Kinetic inductance effect (KI-TWPA): high KI material transmission line (NbTiN)

Offer wide (GHz) bandwidth, high saturation power, and near quantum-limited noise performance.

J-TWPA developed by MIT-LL, primary source \rightarrow supply more restricted now

- Group at Washington University St. Louis developing devices based on MIT-LL design

Need: An avenue for expanding the supply of TWPA to the experimental community

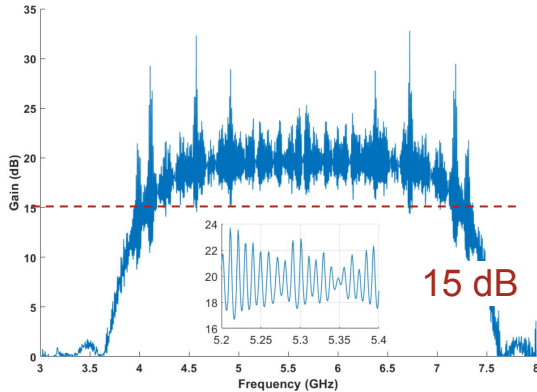
J-TWPA: C. Macklin et al, *Science*, 2015.

Kinetic Inductance TWPA

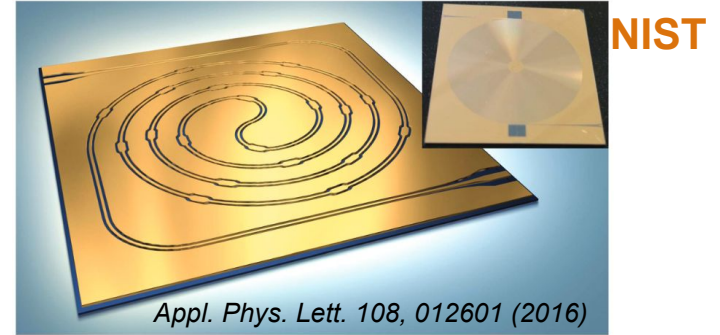
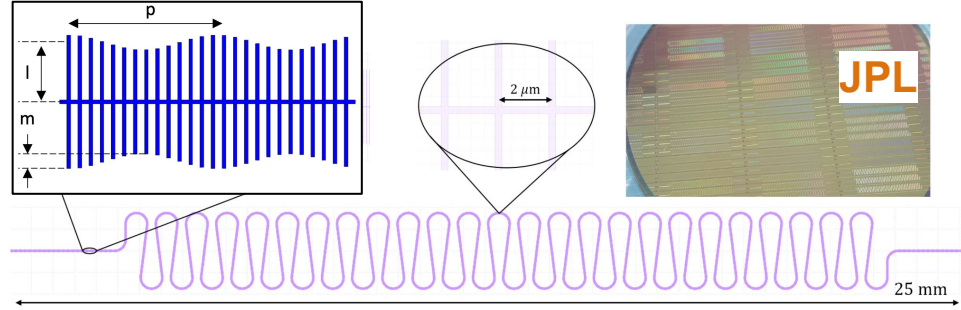
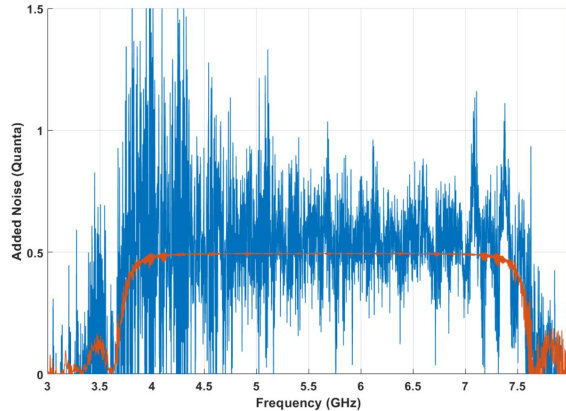
Developed by groups at JPL and NIST

Figures of merit:

- Gain: > 15 dB
- Bandwidth: > 4 GHz
- Saturation power: > -60 dBm
- Operates at SQL



N. Kilmovich et al. arXiv:2306.11028 (2023)



Other KI-TWPA references:

G. Che et al. arXiv:1710.11335 (2017)

S. Shu et al. Phys. Rev. Research 3, 023184 (2021)

M. Malnou et al. PRX Quantum 2, 010302 (2021)

M. Malnou et al. arXiv:2110.08142 (2021)

Fermilab LDRD Proposal

Partnering with Peter Day (JPL) to produce and characterize KI-TWPAs

Our proposal would establish:

1. The KI-TWPA as a viable *community* option for broadband, quantum-limited RF amplification
2. The pathway from JPL to Fermilab for the supply of these amplifiers
3. The infrastructure and expertise at Fermilab to make use of these amplifiers and [disseminate them to multiple Fermilab efforts](#)

Anticipated throughput: 6 devices/year during LDRD

- More going forward, depending on demand, funds

Fermilab Low-Background Cryogenic Facilities



Located in the MINOS experimental hall at Fermilab, 100 meters underground

Fermilab LDRD Proposal: Technical & Scientific Goals

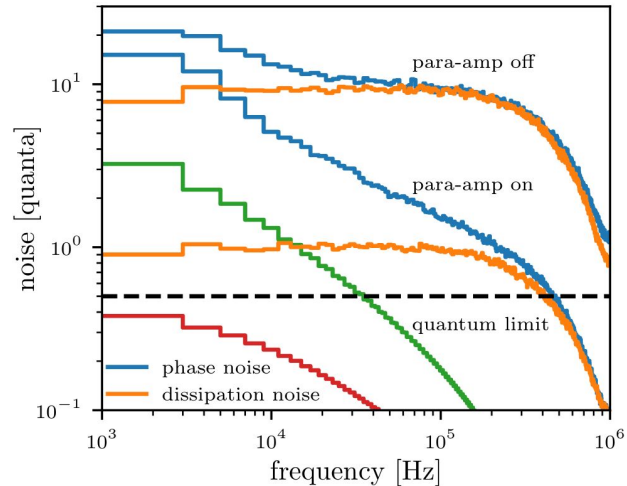
Instrument the NEXUS & QUIET facilities for KI-TWPA characterization and application

Apply KI-TWPA readout to achieve the following:

1. Demonstration of a KIPMD with an eV-scale energy resolution
2. A DM search using a low-threshold KIPMD
3. Demonstration of near quantum-limited amplification of an array of frequency-multiplexed superconducting qubits

KI-TWPA Application: MKIDs & KIPMDs

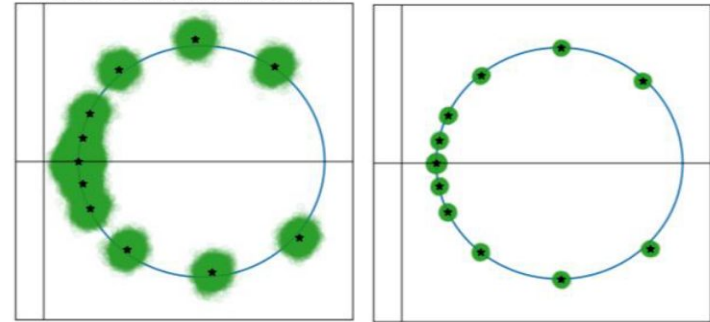
Optical MKID device w/ KI-TWPA



Zobrist *et al.* Appl. Phys. Lett. **115**, 042601 (2019)

Caltech device w/ KI-TWPA

Idealized transmission



KIPA Off → KIPA On

x4.2 improvement in RMS

“Ramanathan, Wen, in prep. (Golwala group, Caltech)”

KI-TWPA Application: KIPMDs at FNAL

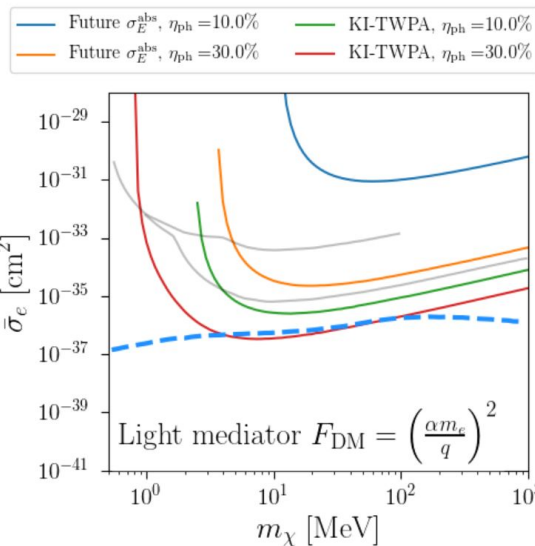
Current device: $\sigma_E^{\text{abs}} = 2.6$ eV with $\eta_{\text{ph}} \sim 1\%$

- Previous device: $\eta_{\text{ph}} = 8\%$
- Improved RF transmission: 4x improvement in σ_E^{abs}

Adding a KI-TWPA: 4-7x improvement in σ_E^{abs}

KI-TWPA Readout for KIPMDs enables:

- the first demonstration of eV-scale resolution
- the first DM search with these devices
- reaching unexplored parameter space for DM.



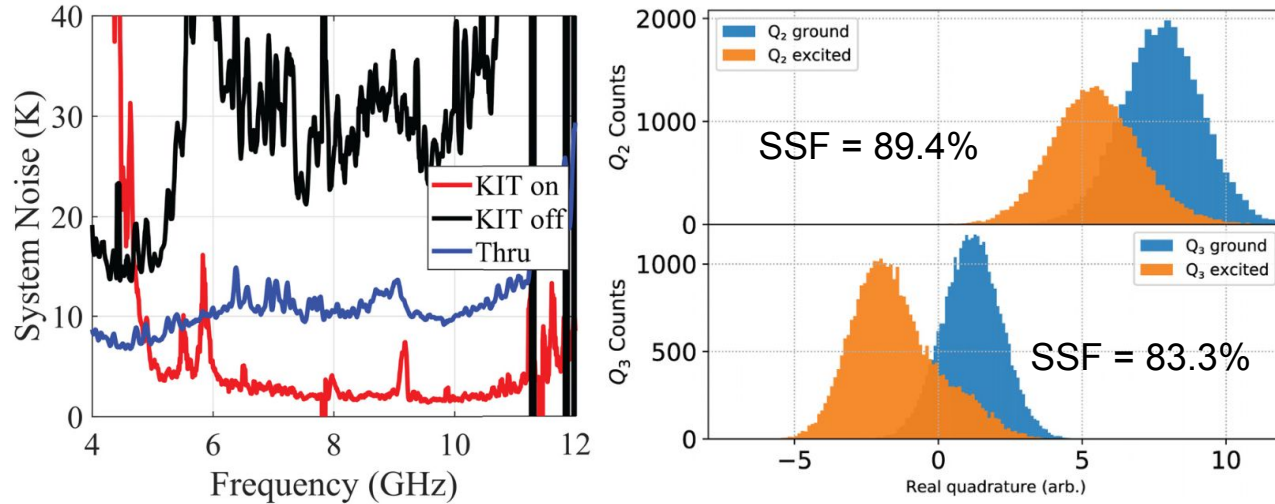
Gram-Month Exposure: $E_{\text{thresh}} = 5 \times \sigma_E$

	Current	$\eta_{\text{ph}} = 10\%$	$\eta_{\text{ph}} = 30\%$	KI-TWPA + $\eta_{\text{ph}} = 10\%$	KI-TWPA + $\eta_{\text{ph}} = 30\%$
σ_E^{abs} [eV]	2.6	0.65	0.65	0.14	0.14
σ_E [eV]	320	6.5	2.2	1.4	0.47
E_{thresh} [eV]	1600	32.5	11	7	2.35
ER M_χ^{min} [MeV]	--	15	4	3	0.9

$\eta_{\text{ph}} \equiv$ energy absorbed in MKID/energy deposited in substrate

KI-TWPA Application: Qubits

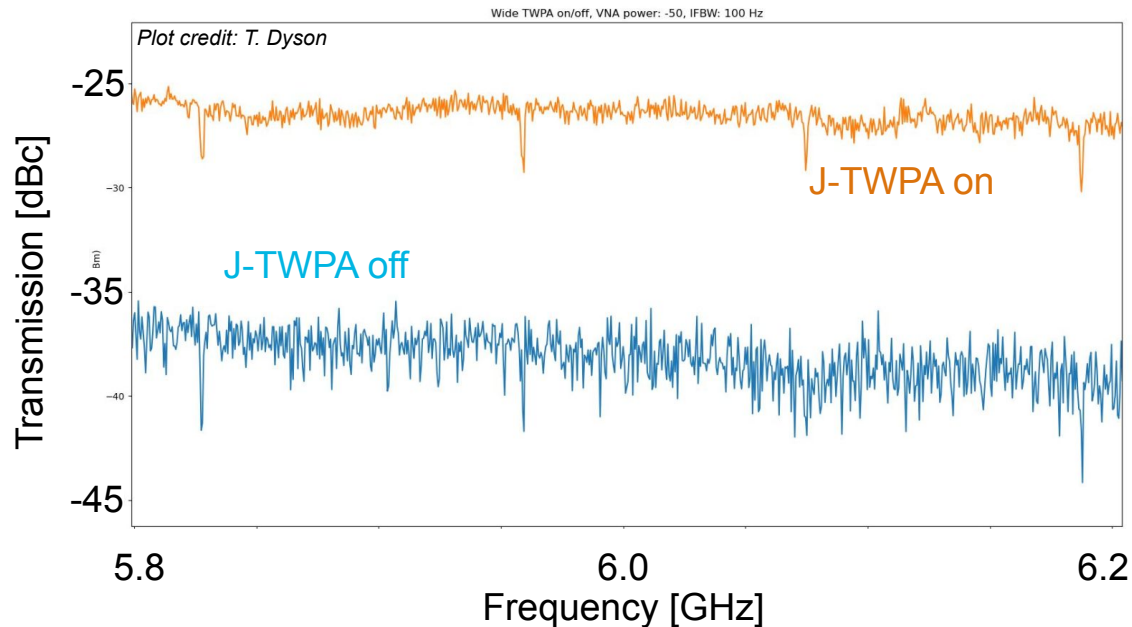
NIST & Raytheon BBN demonstrated KI-TWPA use with superconducting qubits



- ~10x noise reduction across >4 GHz bandwidth
 - 10(6) GHz: 1.4(4.1) photons of added noise at 10 GHz
- Simultaneous readout of 2 qubits with fidelity approaching 90%

L. Ranzani et al. Appl. Phys. Lett. 113, 242602 (2018)

KI-TWPA Application: Qubits at FNAL



Daniel Bowring's ECA

- **MIT-LL J-TWPA**
- 4 transmon chip

Demonstrated ~10 dB of gain

Clear reduction of noise

Objective: demonstrate simultaneous readout of 4 qubits with SSF $\geq 90\%$

Setup allows direct comparison of J-TWPA and KI-TWPA with identical chip, transmission line, and readout electronics

Conclusion

- Kinetic Inductance Traveling Wave Parametric Amplifiers offer a promising alternative to JPAs and Josephson TWPAs.
 - Shown to operate near the SQL of noise across 4 GHz of bandwidth.
- Future arrays of qubits and KIPMDs will require these amplifiers, while other RF experiments will benefit from their availability.
- Seeking LDRD support to outfit Fermilab cryogenic facilities for KI-TWPA characterization and application.
 - Will enable first of its kind low-threshold dark matter search with KIPMDs
- Will establish a pool of KI-TWPAs at FNAL for experimenters, resolving the current challenges due to limited J-TWPA availability.



Thank You!



D. Temples
FNAL



D. Bowring
FNAL



S. Golwala
Caltech

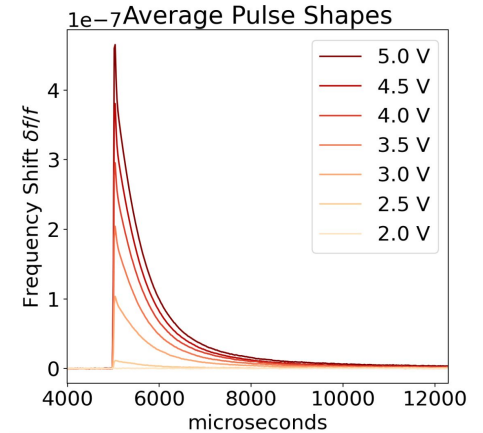
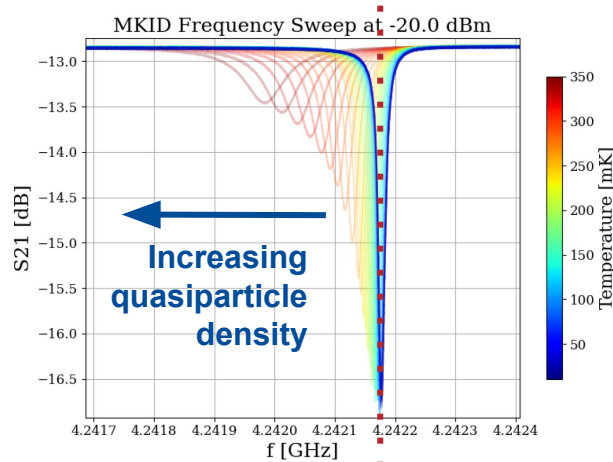
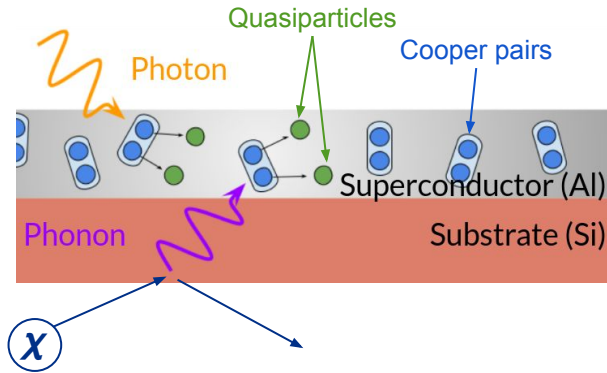
Advertisement: Have a need for cryogenic broadband low-noise RF amplifiers?

Acknowledgement: KI-TWPAs sourced from Peter Day, JPL.

Kinetic Inductance Phonon-Mediated Detectors

Phonon-mediated MKIDs: superconducting microwave LC resonators deposited on a crystalline insulating substrate (e.g., silicon).

We aim to demonstrate these devices as eV-scale microcalorimeters given their sensitivity to athermal phonons produced by particle interactions in the substrate propagating to, and breaking Cooper pairs in, the superconductor.



Kinetic Inductance Phonon-Mediated Detectors for Dark Matter

Dark matter detection down to fermionic thermal relic mass limit of a few keV requires sensitivity to eV-scale depositions (sub-eV quanta).

MKID-based microcalorimeters:

- **Fast.** 10 μs rise time and O(ms) fall times demonstrated at NEXUS.
 - Directly sense athermal phonons rather than waiting for a thermal distribution to be reached.
- **Scalable.** Natively frequency multiplexable: no cryogenic multiplexers, identical warm electronics for a 1 resonator device up to $O(10^3)$ resonators.

Currently performance limited by energy resolution

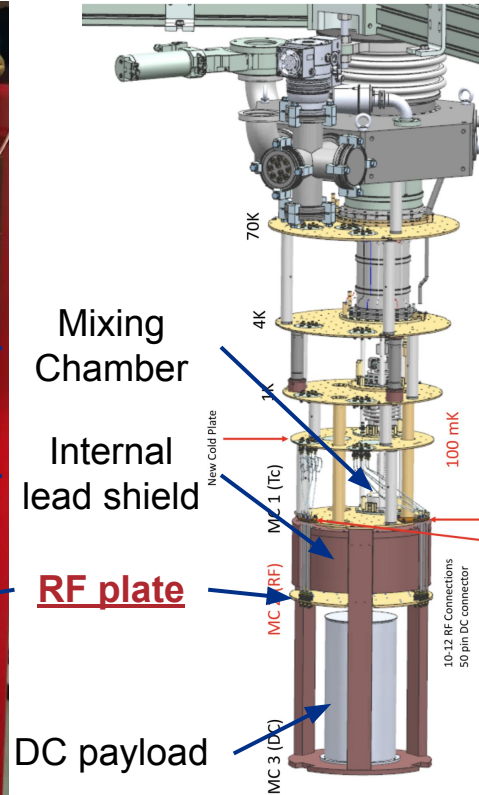
- TES sensors: $\sigma_E = 2.7 \text{ eV}$ (SuperCDMS)
- Phonon-mediated MKID sensors: $\sigma_E = 26 \text{ eV}$ (BULLKID)

To date, no DM searches yet performed with phonon-mediated MKID devices.

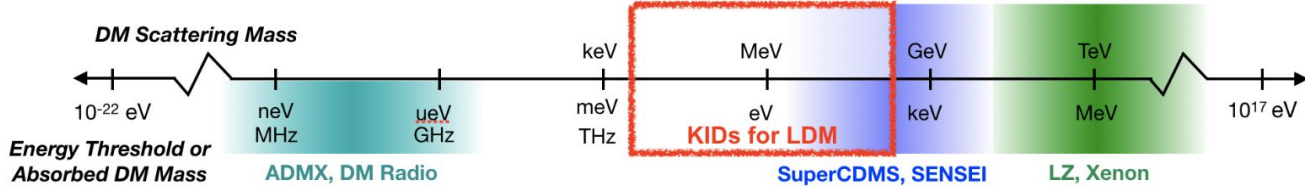
NEXUS: Northwestern EXperimental Underground Site

Key facility features:

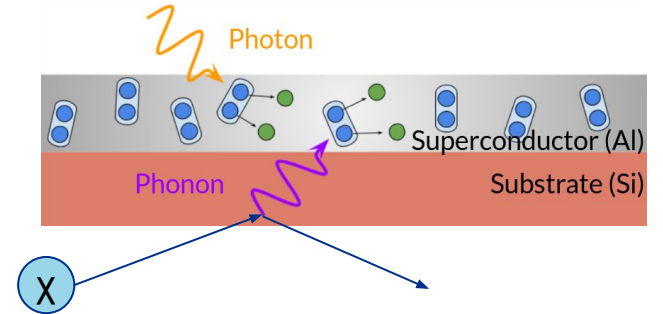
- CryoConcept HEXADRY dilution refrigerator
- 107 m rock overburden (300 mwe)
 - Muon flux reduced by $\sim 500\times$
- Moveable lead shield & **B** shielding
- $\mathcal{O}(100)$ dru background rate (NaI)
- Optical photon & DD neutron calib. systems



Phonon-mediated Kinetic Inductance Detectors for Dark Matter



- **Alternative to TES.** KIDs are faster, highly multiplexable with no extra electronics, and non-dissipative (lower heat load & amenable to QIS techniques)
- State of the art energy resolution (threshold):
 - KID (phonon-mediated): $\sigma_E = 26$ eV (~ 125 eV) [CALDER, Supercond. Sci. Technol. 31, 075002 (2018)]
 - TES: $\sigma_E = 2.65$ eV (12.5 eV) [SuperCDMS, Phys. Rev. D 104, 032010 (2021)]
- No direct DM search with KID devices performed to date.
- Ultimate KID sensitivity lies between qubits and TES. Provides the ability to study effects of interest for qubits with a better understood detector (stress microfractures, Cherenkov radiation).



Interactions in substrate produce meV-scale athermal phonons that propagate freely until they are absorbed by the superconductor, breaking Cooper pairs ($E_{ph} > 2\Delta \sim \mathcal{O}(100) \mu\text{eV}$)

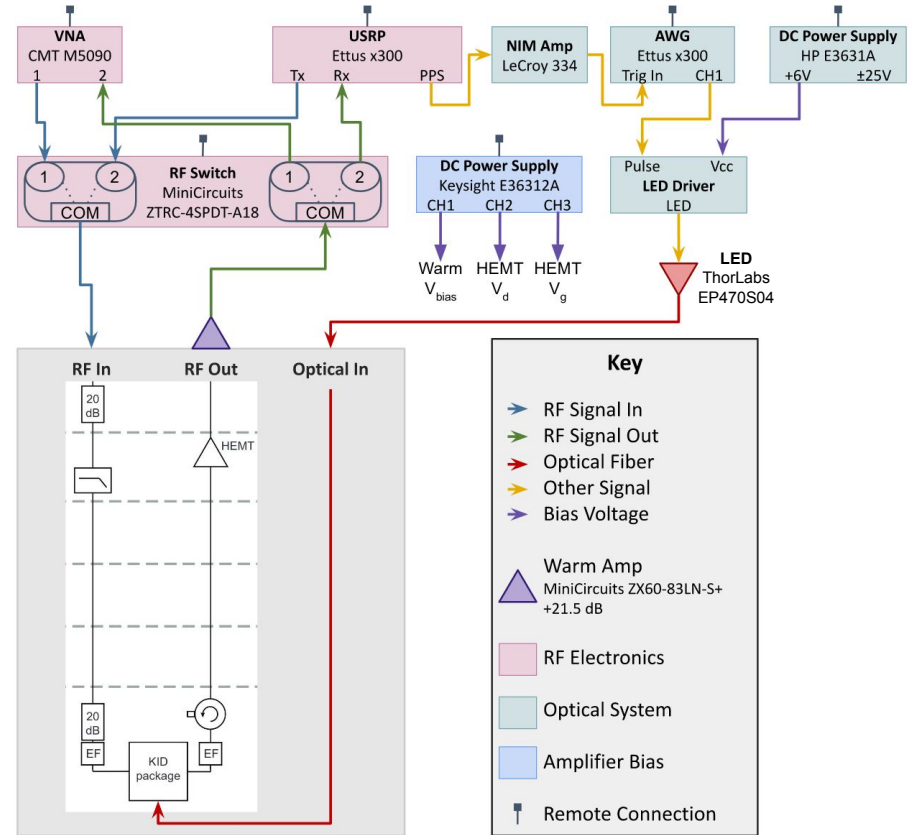
MKID Warm Electronics and LED Calibration System

Fully remote-controllable system

- Ettus x300 USRP & GPU SDR
 - DAQ stability issues largely resolved
 - Server, client running on same PC
- DAQ PC GPU GeForce RTX 2080 Ti
 - Planned upgrade to RTX 6000 Ada
 - >4x more memory, cores

Optical Calibration System

- 470 nm LED (ThorLabs EP470S04)
- Up to ~5 mW optical power
- Fiber pointing at rear of device
- Pulse burst synced to USRP clock
- Pulse profiles:
 - $1 \mu\text{s}$ @ 10 Hz / $2 \mu\text{s}$ @ 5 Hz
 - Limited to 0.001% duty cycle



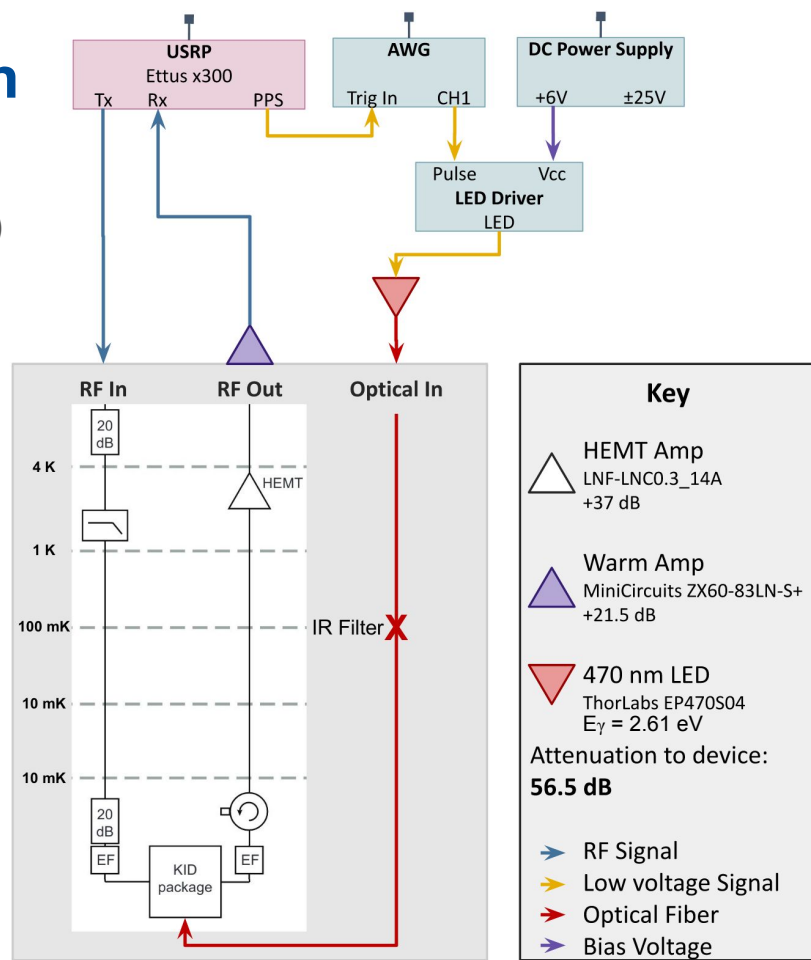
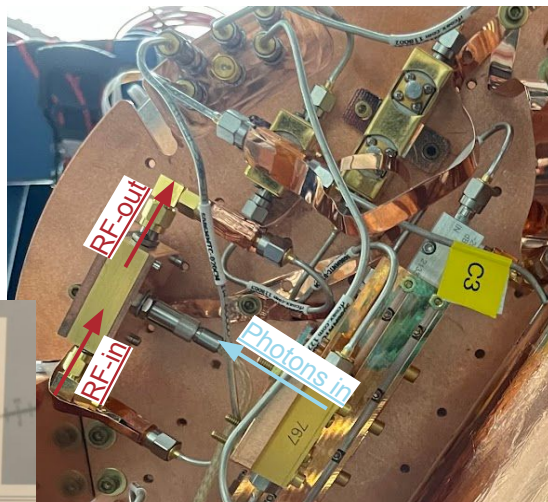
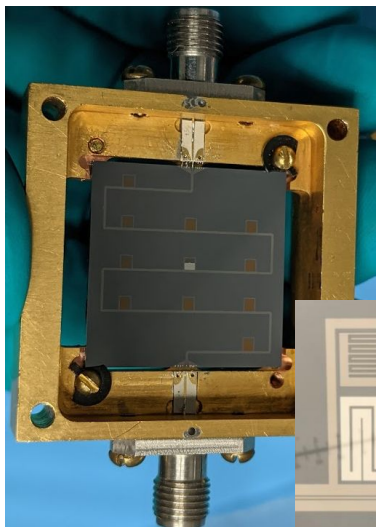
Readout & Optical Calibration System

FNAL-I Device (O. Wen) 3.9--4.4 GHz

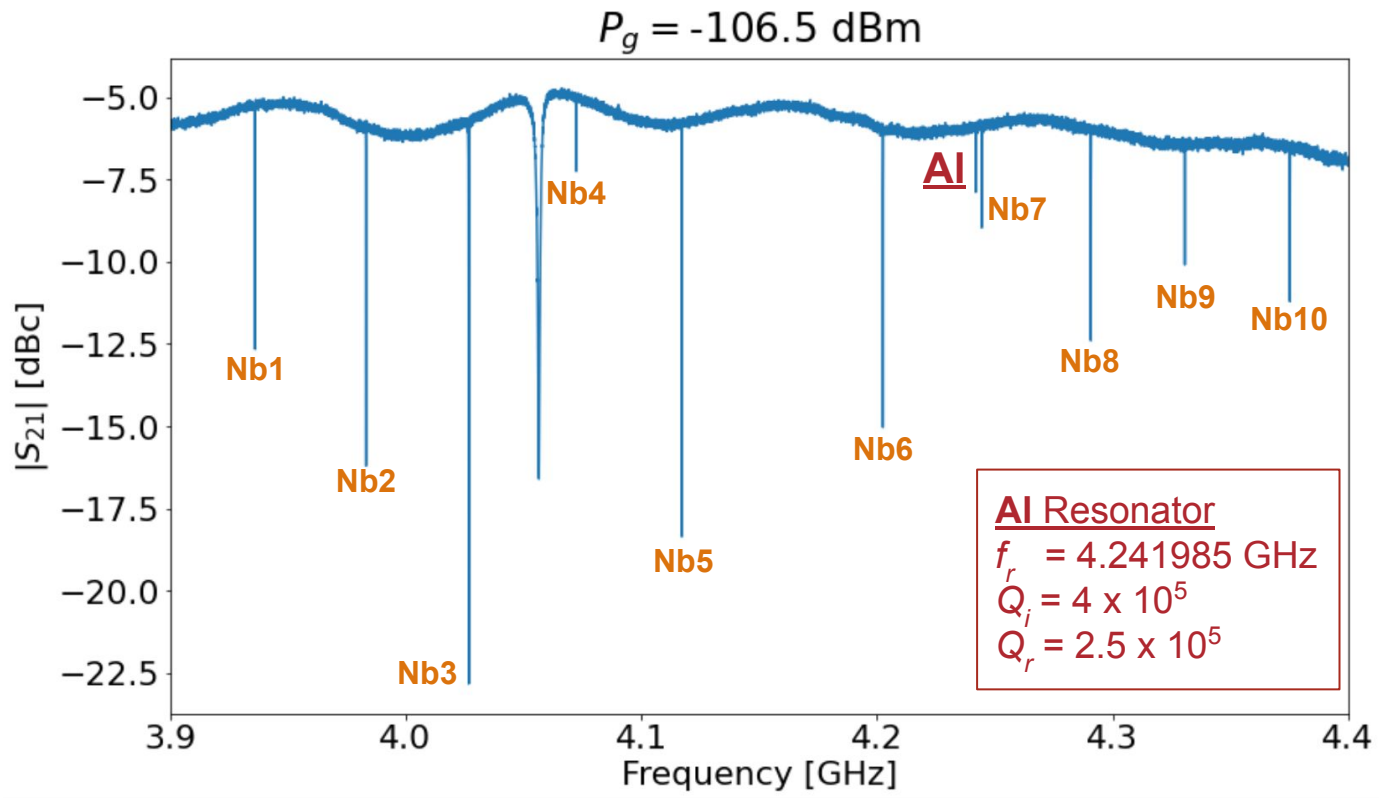
- Phonon sensitive resonator 30 nm Al inductor (4.24 GHz)
- 10x Nb resonators
- ~1g high-resistivity Si substrate

DAQ: Ettus Research x300 USRP

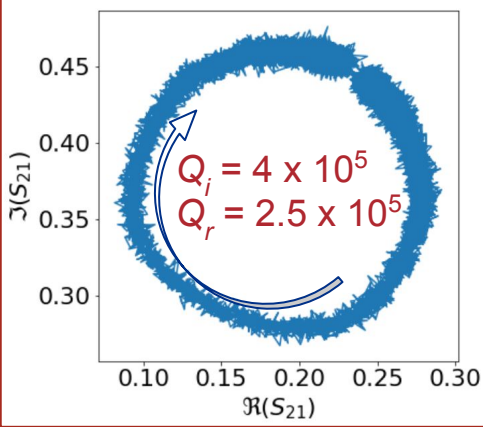
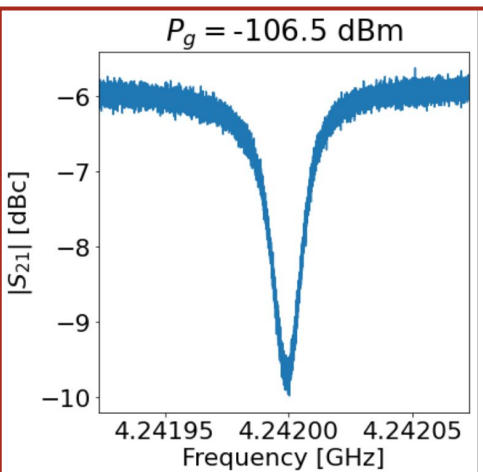
Correlated noise removal with off-resonant tones



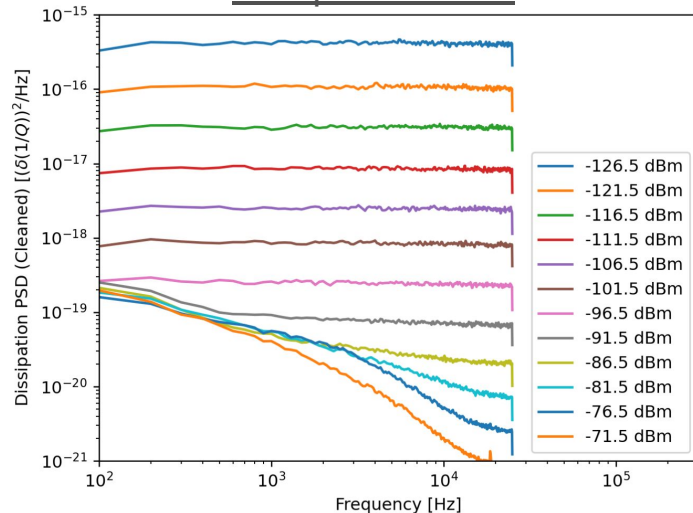
“Fermilab I” Resonance Spectrum



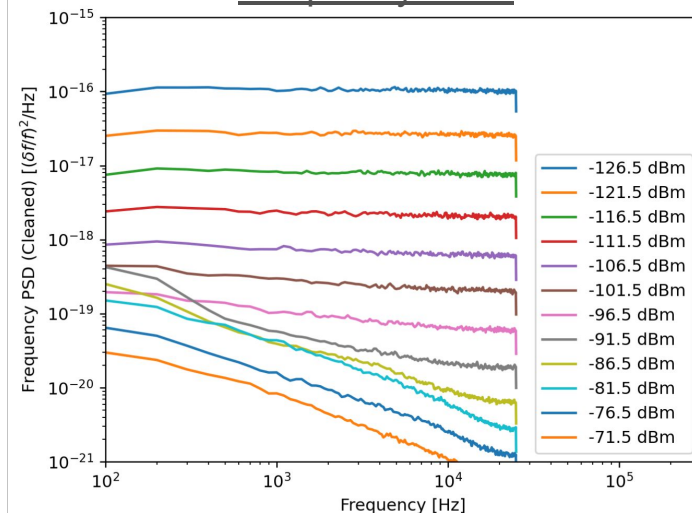
FNAL-I Device Performance



Dissipation noise



Frequency noise



- White amplifier noise dominated at low readout powers
- Frequency quadrature $1/f$ TLS noise (O. Wen's talk M-2-2)
- Correlated κ_1 - κ_2 noise of unknown origin at highest readout powers

$$\sigma_{dff} = 3.6 \times 10^{-9}$$

$$\sigma_{qp}^{\kappa_2} = 0.49 \mu\text{m}^{-3}$$

$$\sigma_{1/Q} = 7.4 \times 10^{-9}$$

$$\sigma_{qp}^{\kappa_1} = 0.85 \mu\text{m}^{-3}$$

Determined via optimal filter on noise data (LED pulse signal template)

idealized transmission

