

Updates on AMS-02



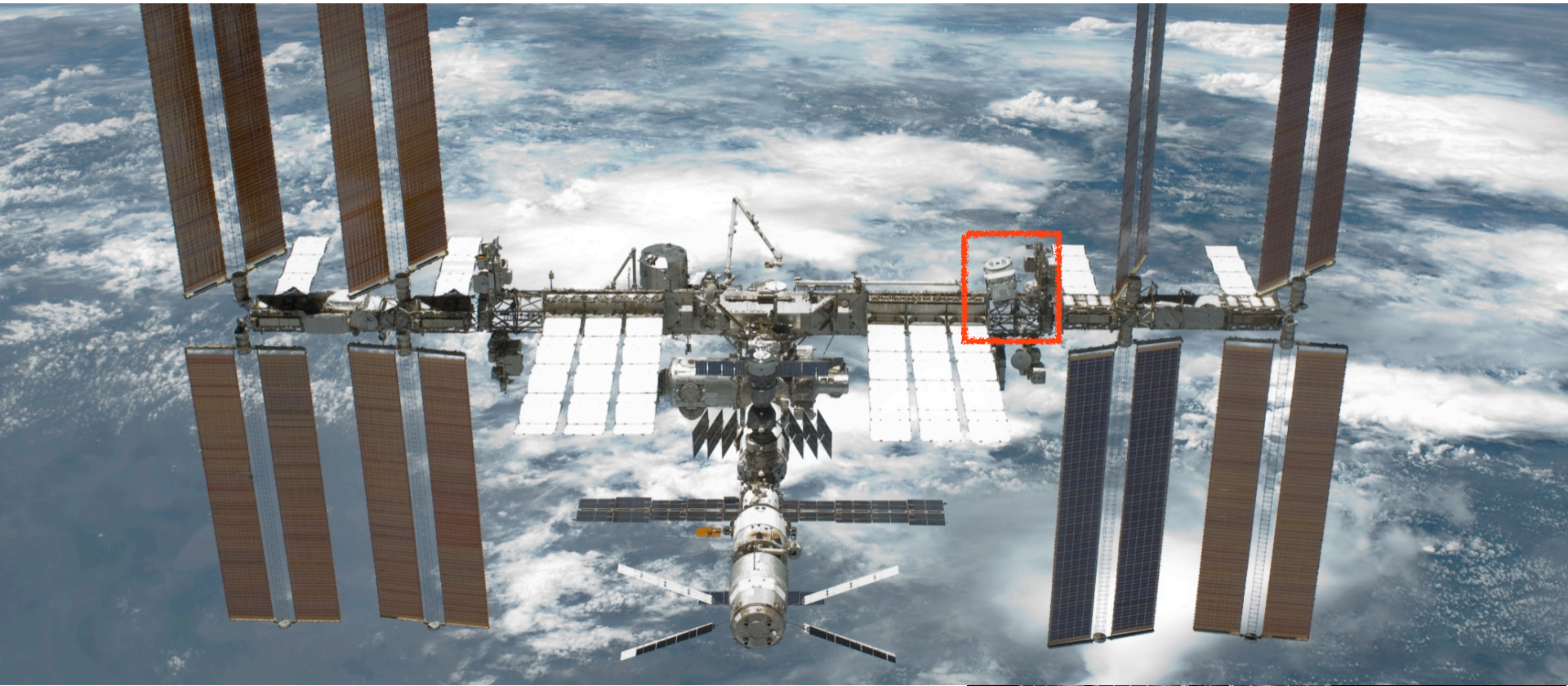
Veronica Bindi

Physics and Astronomy Department University of Hawaii at Manoa



AMS is a US DOE led International Collaboration

Spokesperson: Nobel laureate Prof. Dr. S. Ting from MIT



**AMS-02 has been installed on
the International Space Station
on May 19th 2011**



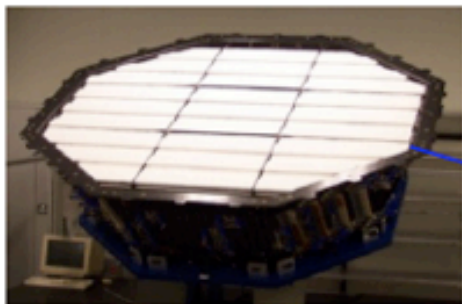


5 m x 4 m x 3 m

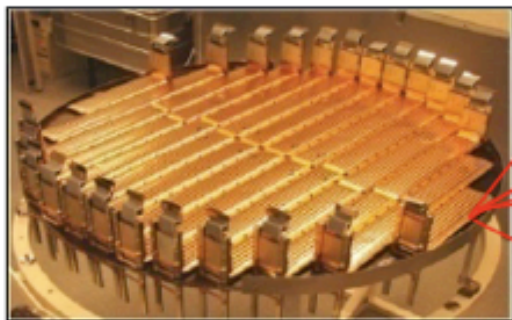
7.5 tons

AMS consists of 5 sub-detectors which provide redundant information for particle identification

Identify e^+ , e^-



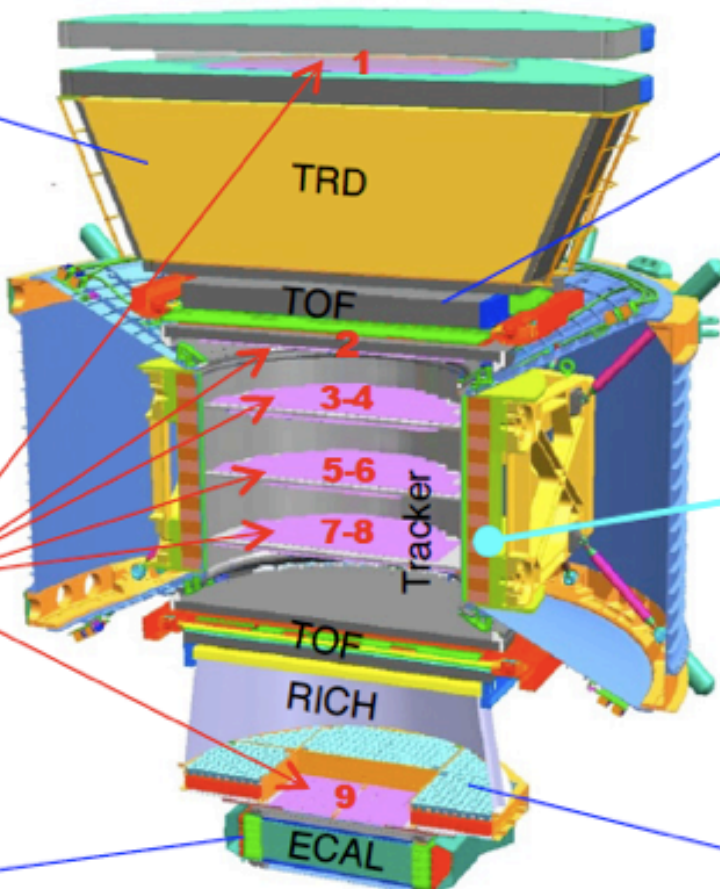
Silicon Tracker
 Z, P



ECAL
 E of e^+ , e^- , γ



Particles and nuclei are defined by their charge (Z) and energy ($E \sim P$)

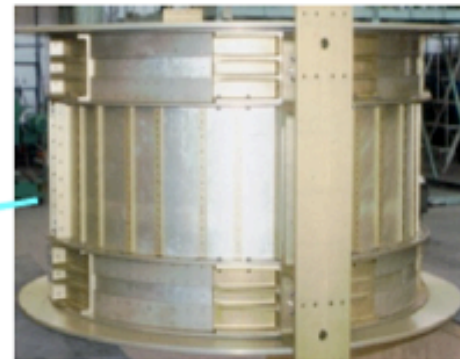


Z, P are measured independently by the Tracker, RICH, TOF and ECAL

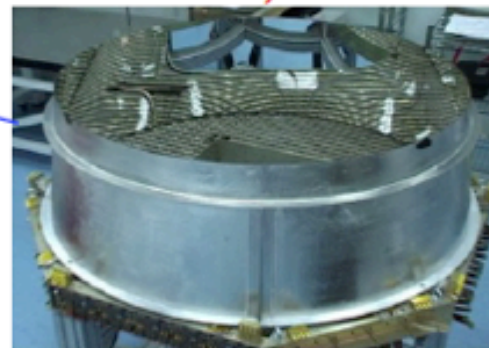
TOF
 Z, E



Magnet
 $\pm Z$

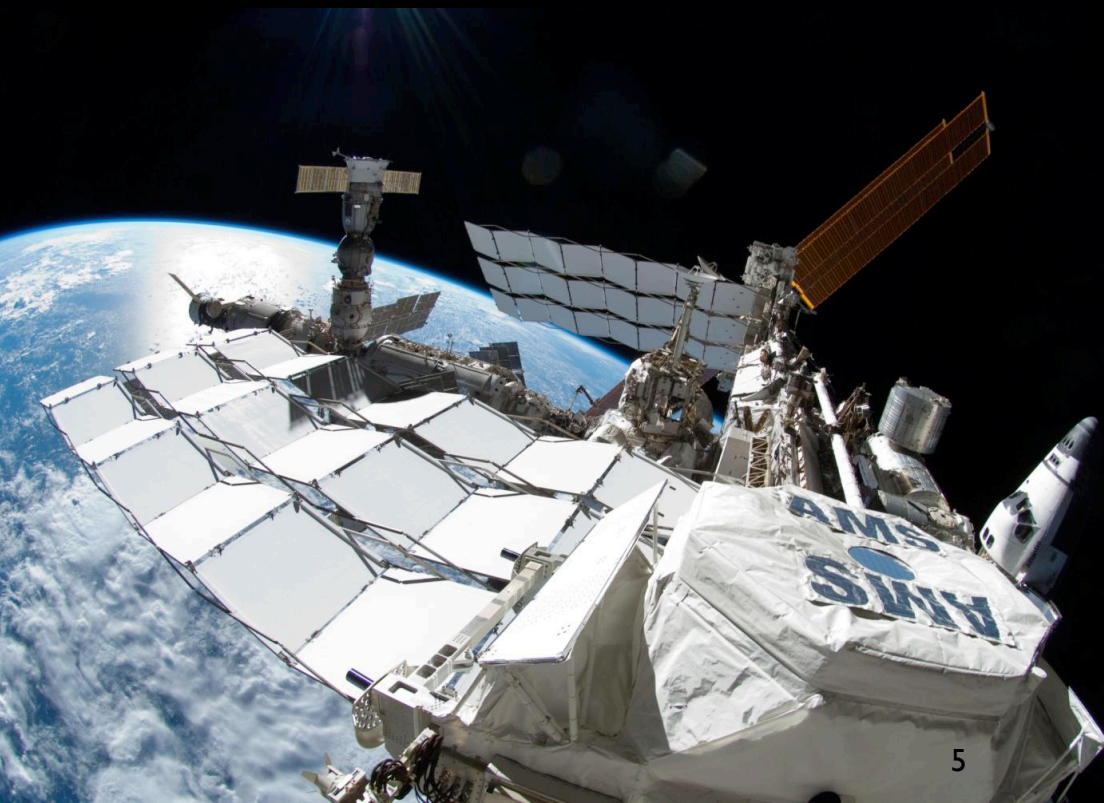


RICH
 Z, E



Scientific goals of AMS on the International Space Station

- Indirect search of Dark Matter: simultaneous observation in several signal channels... e^+ , antiprotons, γ , antideuteron
- Measuring CR spectra up to the iron – refining propagation models;
- Solar modulation on CR spectra over 11 year solar cycle
- Solar activity
- Direct search of primordial antimatter: Anti He, Anti C ...
- New forms of matter: strangelets
- Identification of local sources of high energy photons: SNR, Pulsars, ...

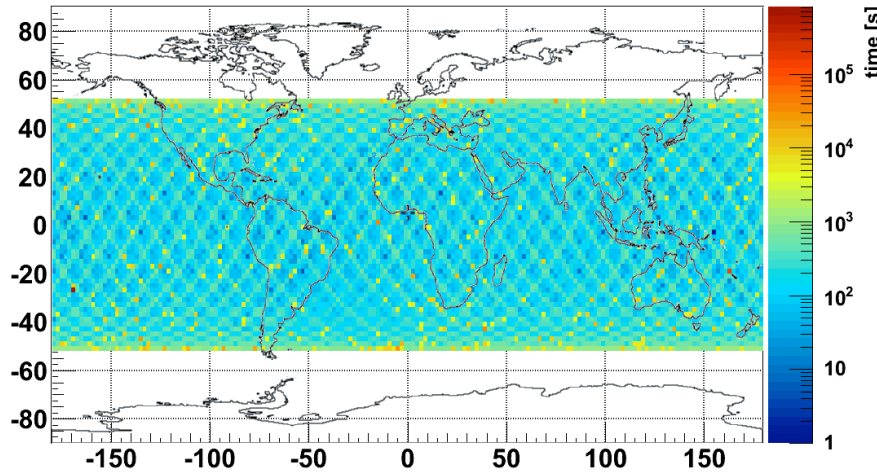


Main analysis currently on going:

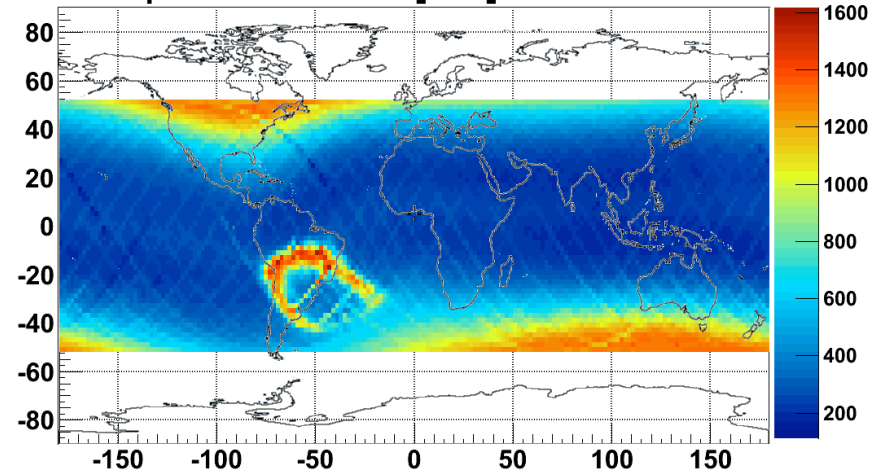
- Positron fraction
- B/C
- P, He, electron ... fluxes
- Monitor of the solar activity

AMS-02 Orbital parameters

Time at location [s]

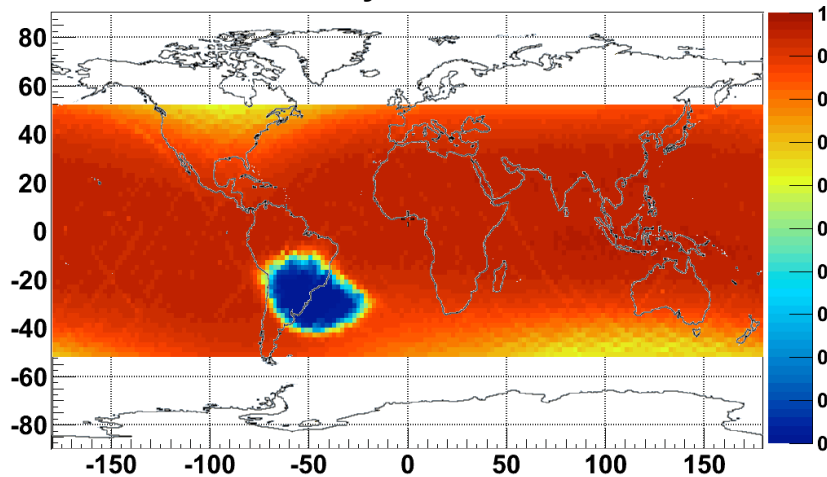


Acquisition rate [Hz]



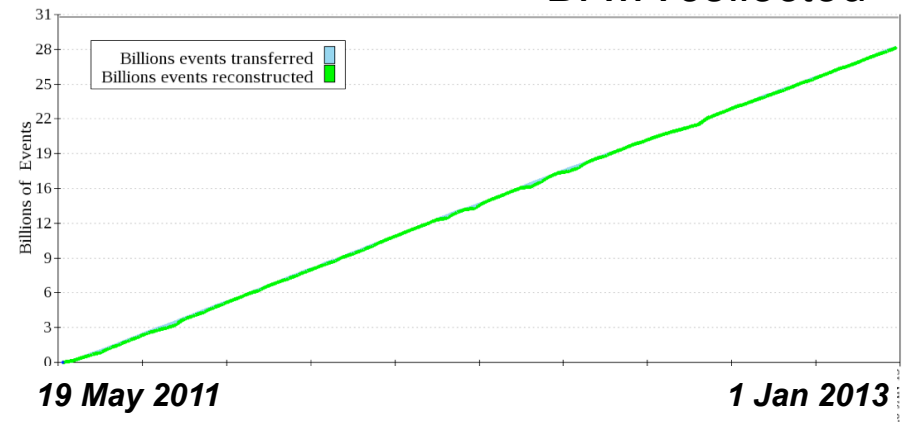
Particle rates vary from 200 to 2000 Hz per orbit

DAQ efficiency



Average DAQ efficiency 85%
Average DAQ rate ~800Hz

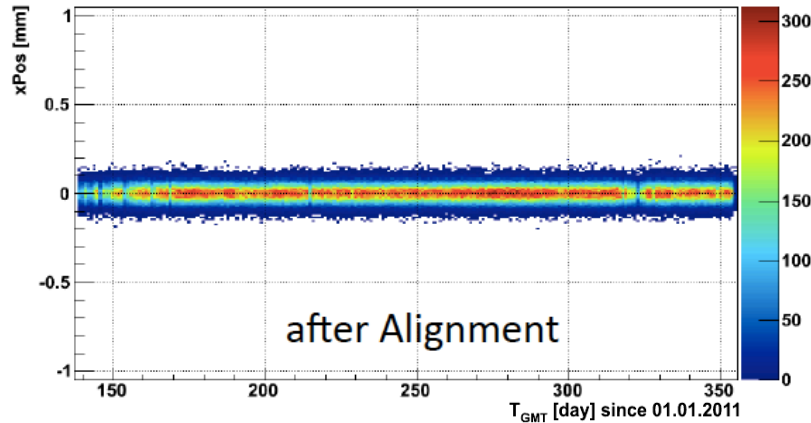
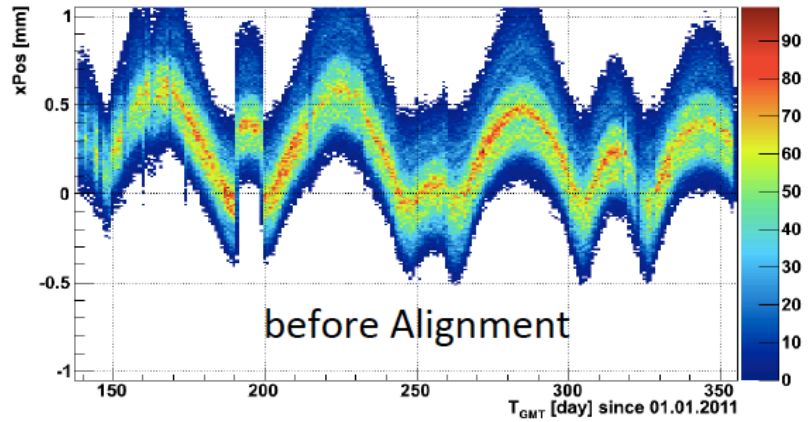
DATA collected



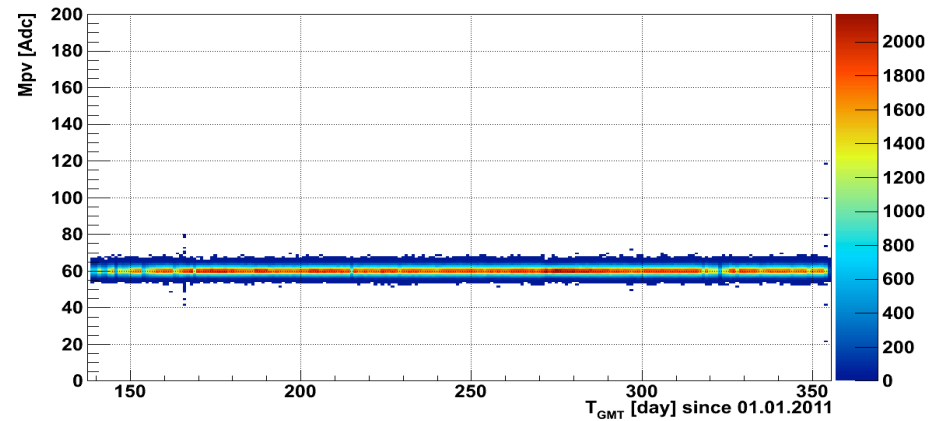
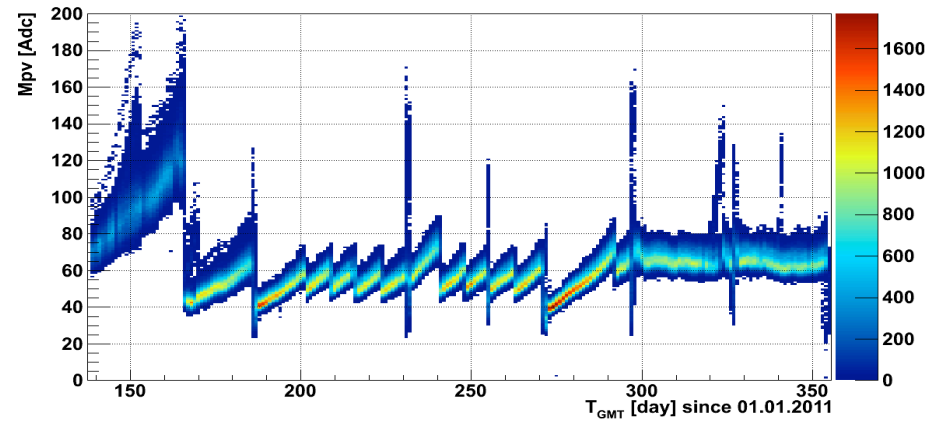
28 billion events collected in 18 months
60 TB raw events (Downlink 10 Mbit/s)

TRD offline calibration

TRD alignment

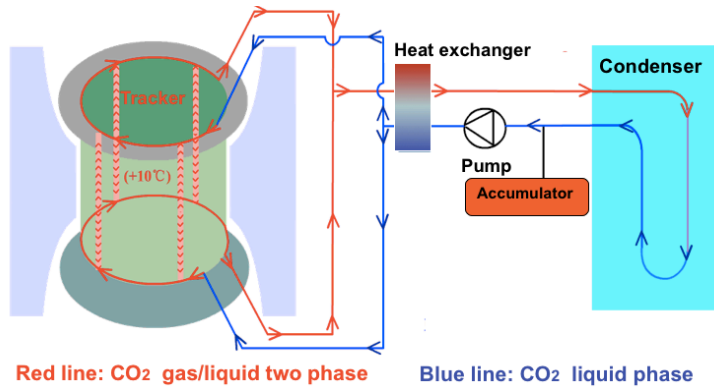


TRD gain calibration



Cosmic protons are used for alignment to an accuracy of 0.04 mm for each straw module and used to calibrate the detector response to 3% accuracy.

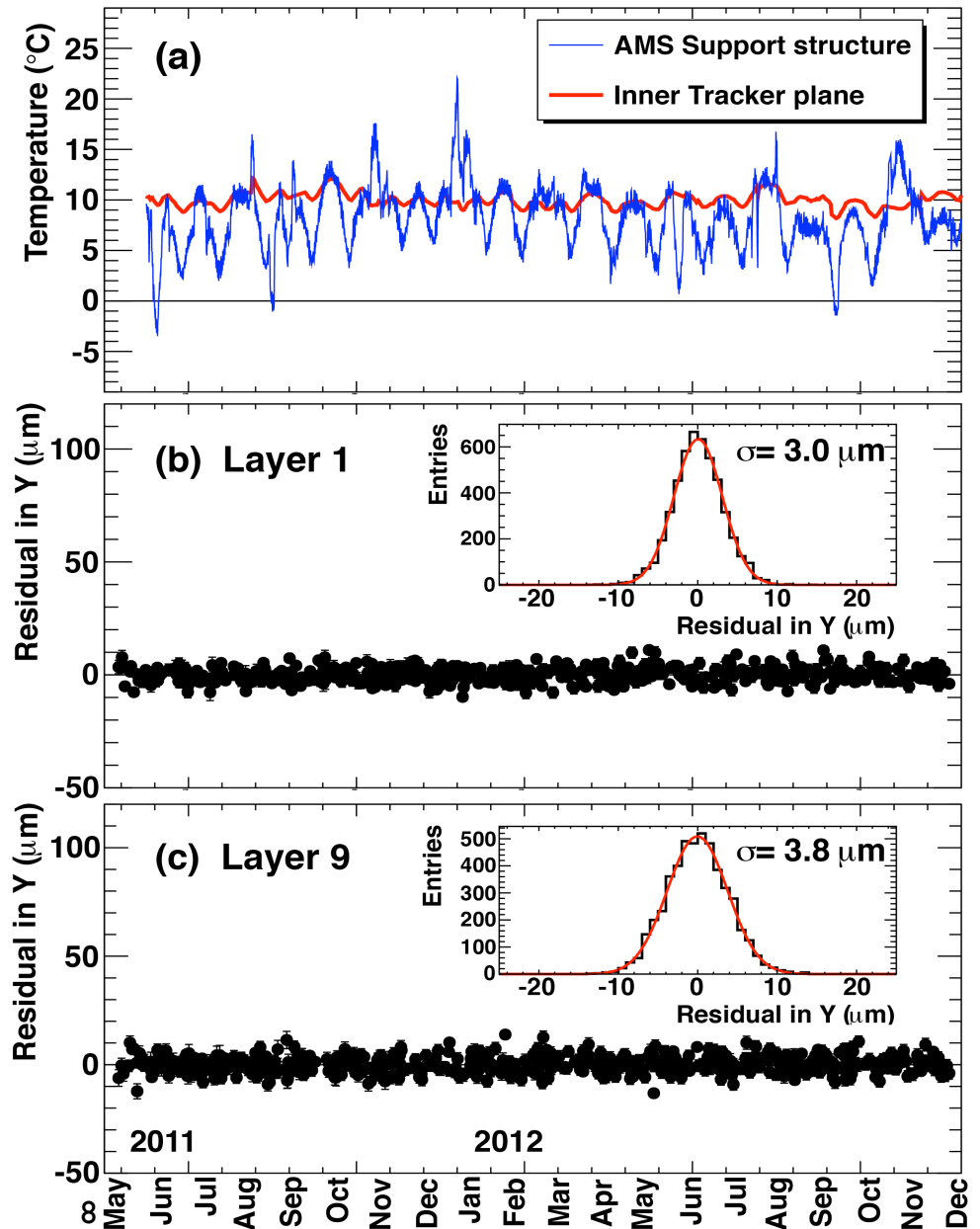
Tracker layers thermal stability



Tracker Thermal Control System

Coordinate resolution on each plane is measured with $10\text{ }\mu\text{m}$ in the bending direction.

Position of ladders in the external layers are dynamically aligned to an accuracy of $3\text{ }\mu\text{m}$.



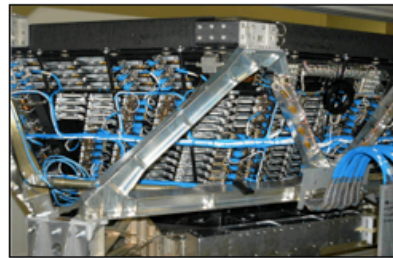
Positron identification and Proton rejection

e^+ low signal and high P background: $P \sim (10^3 \div 10^4) e^+$

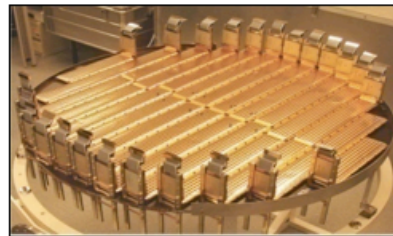
P rejection factor: $10^5 \div 10^6$ to identify e^+ with an error at % level

TRD

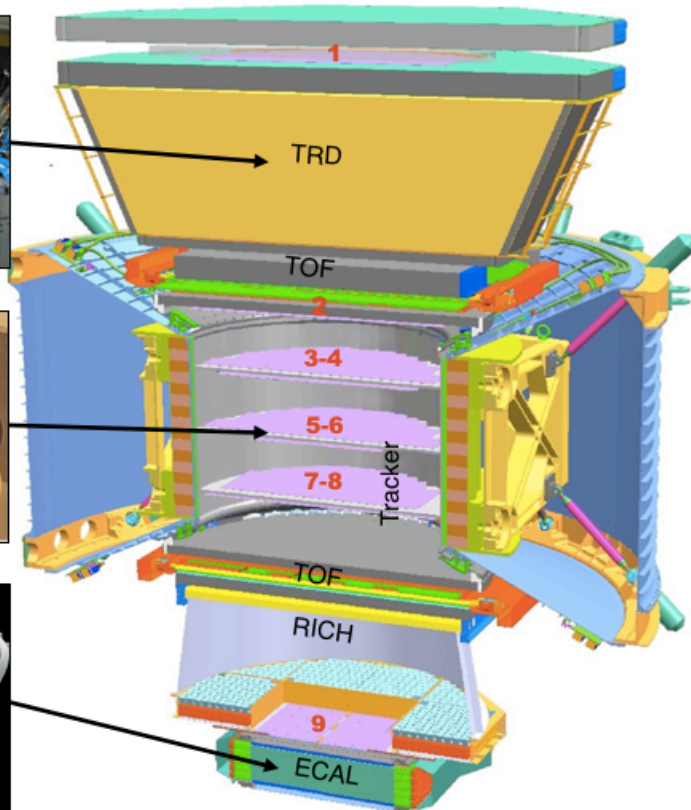
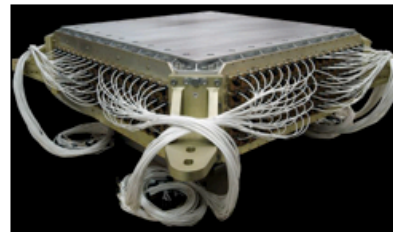
Distinguish between
electrons and protons



SILICON TRACKER and MAGNET
measure the sign and the rigidity



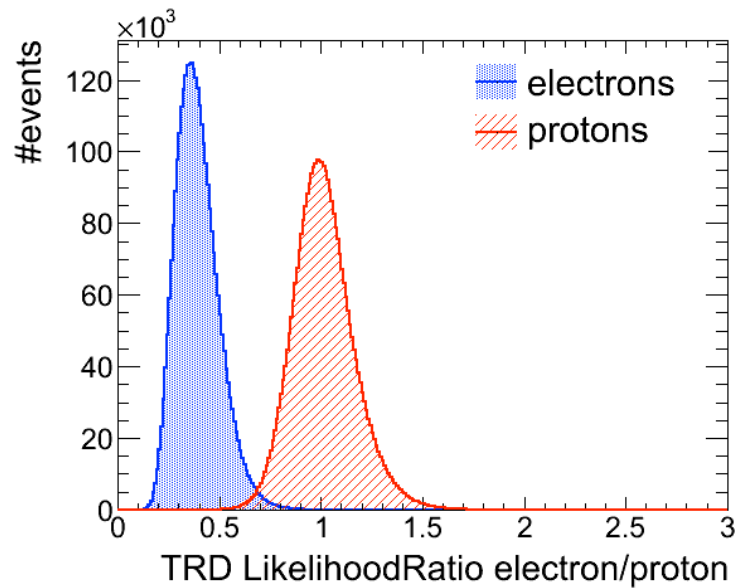
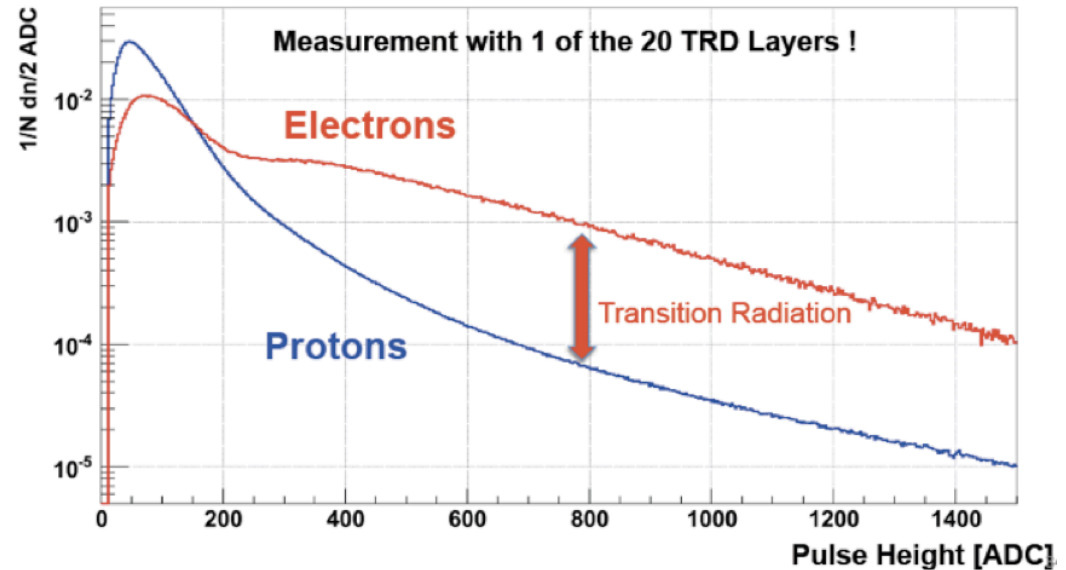
ECAL
measures the energy,
Identifies 3D characteristic positron
shower and rejects hadronic
showers



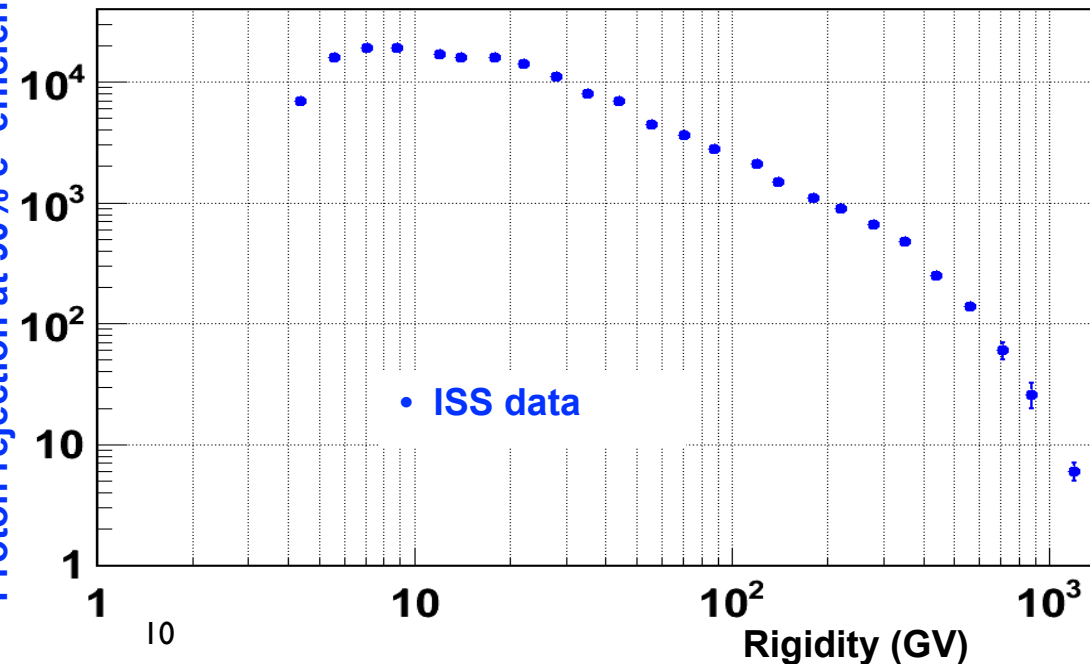
Total rejection of proton 1,000,000
Verified at test beam at CERN

TRD Proton rejection

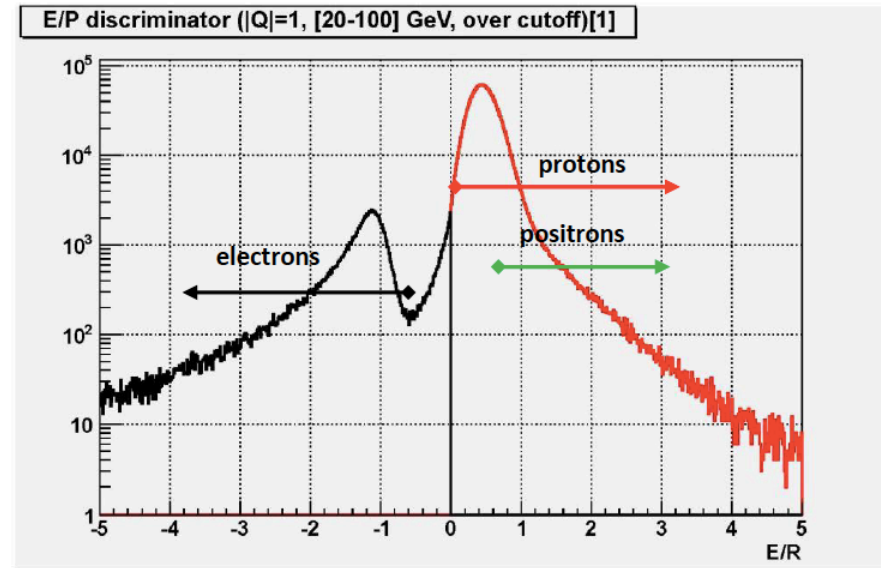
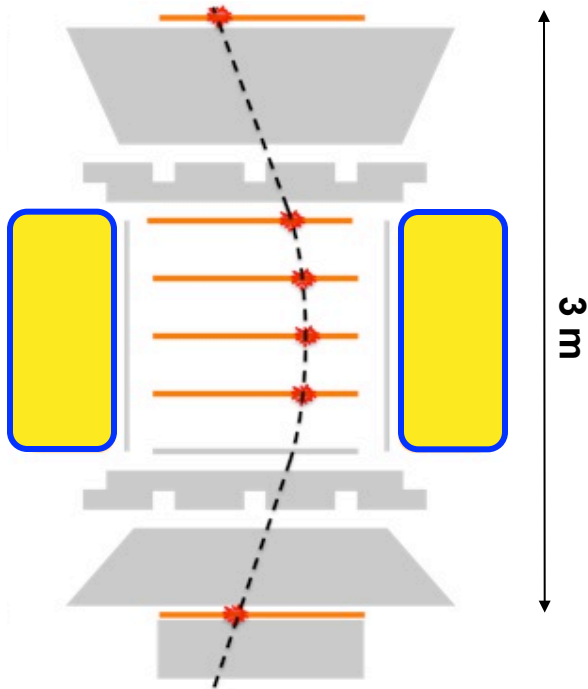
Signals from 20 layers are combined in a likelihood estimator which allows an efficient discrimination of proton background



Proton rejection at 90% e^+ efficiency



Tracker e⁺ and e⁻ identification and P rejection



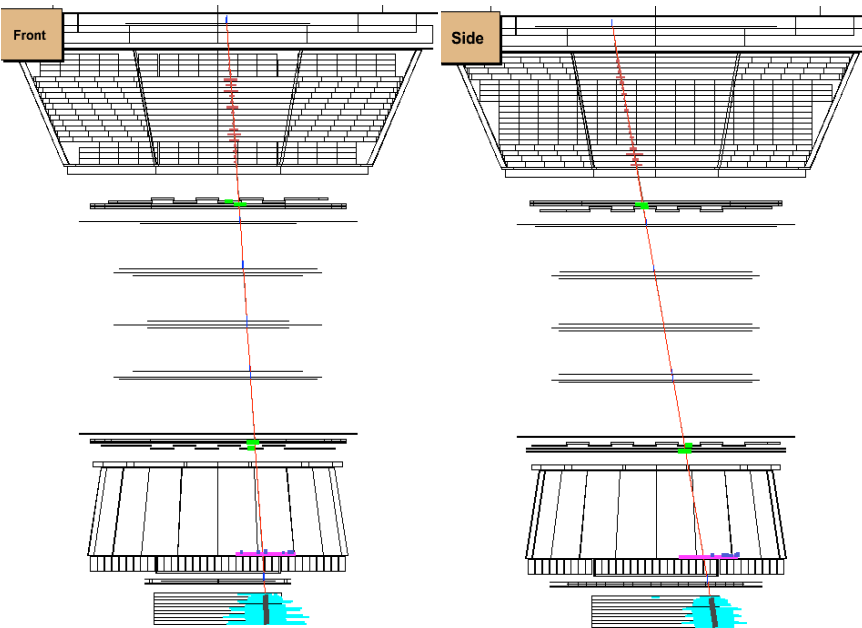
A fit of particle trajectory is used to measure the sign of the particle and its rigidity:

- 1) Used to suppress e⁻
- 2) compared to the energy E measured by ECAL to suppress P.

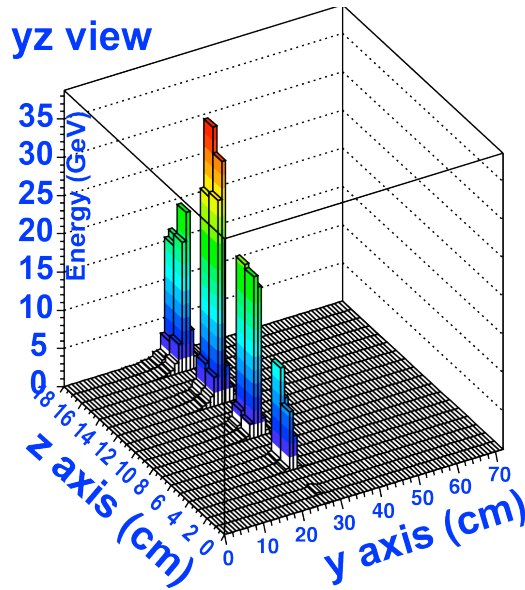
The Electromagnetic Calorimeter

Positron $E=636$ GeV

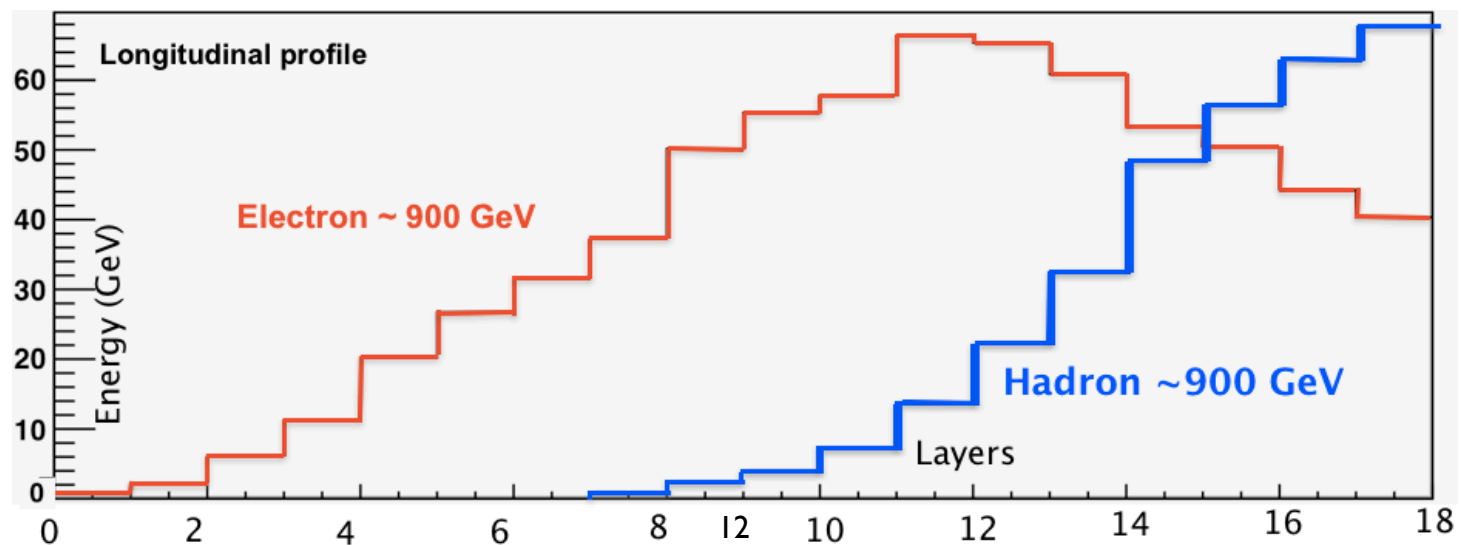
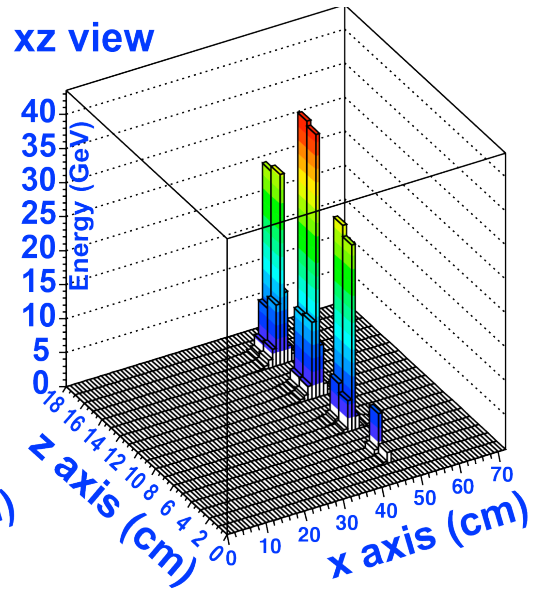
17 radiation length



yz view



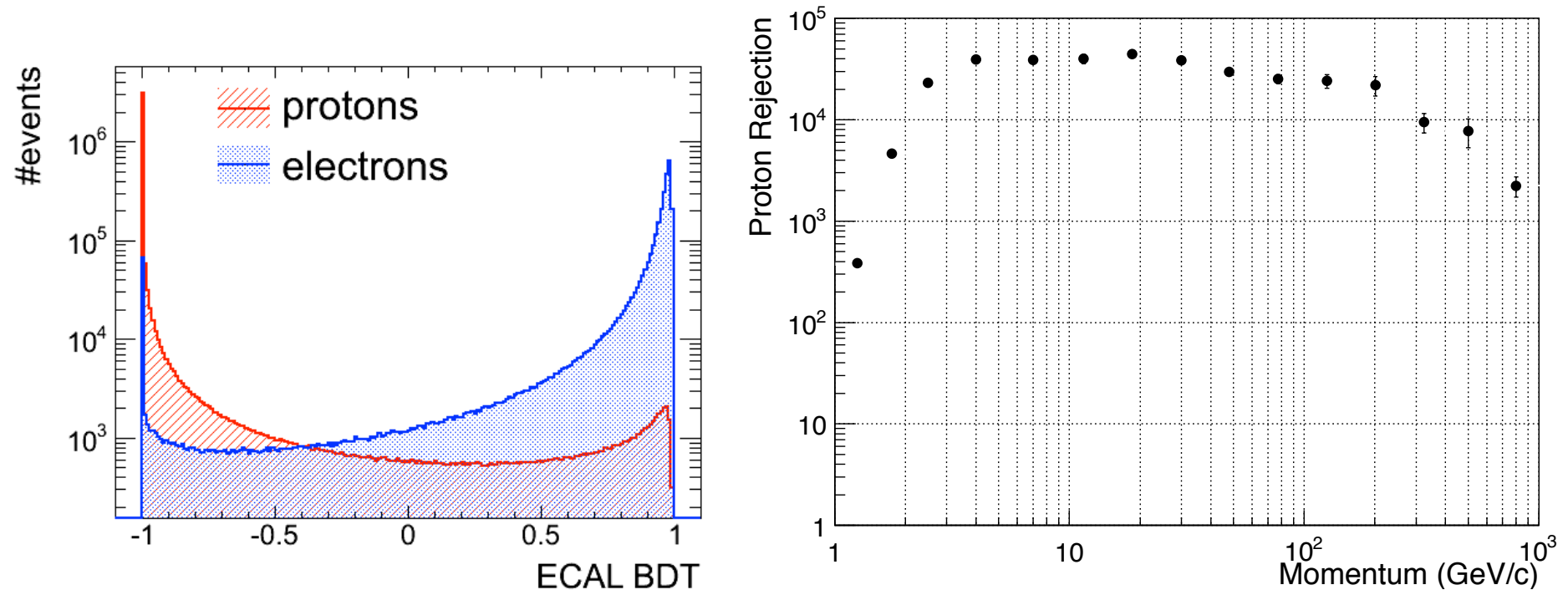
xz view



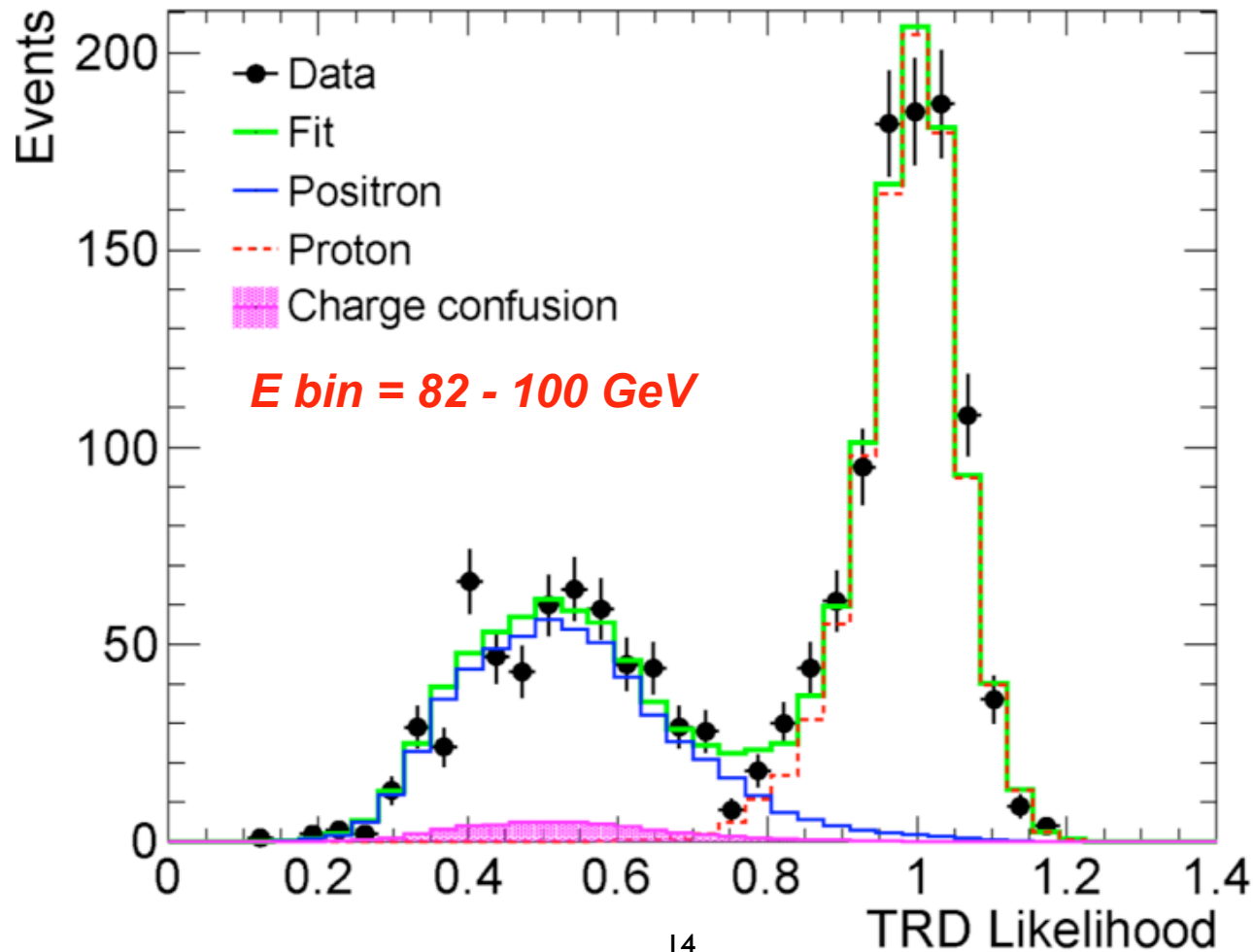
ECAL performance on ISS

A Boosted Decision Tree (BDT) is constructed on the basis of the shower shape in the ECAL to distinguish protons and electrons.

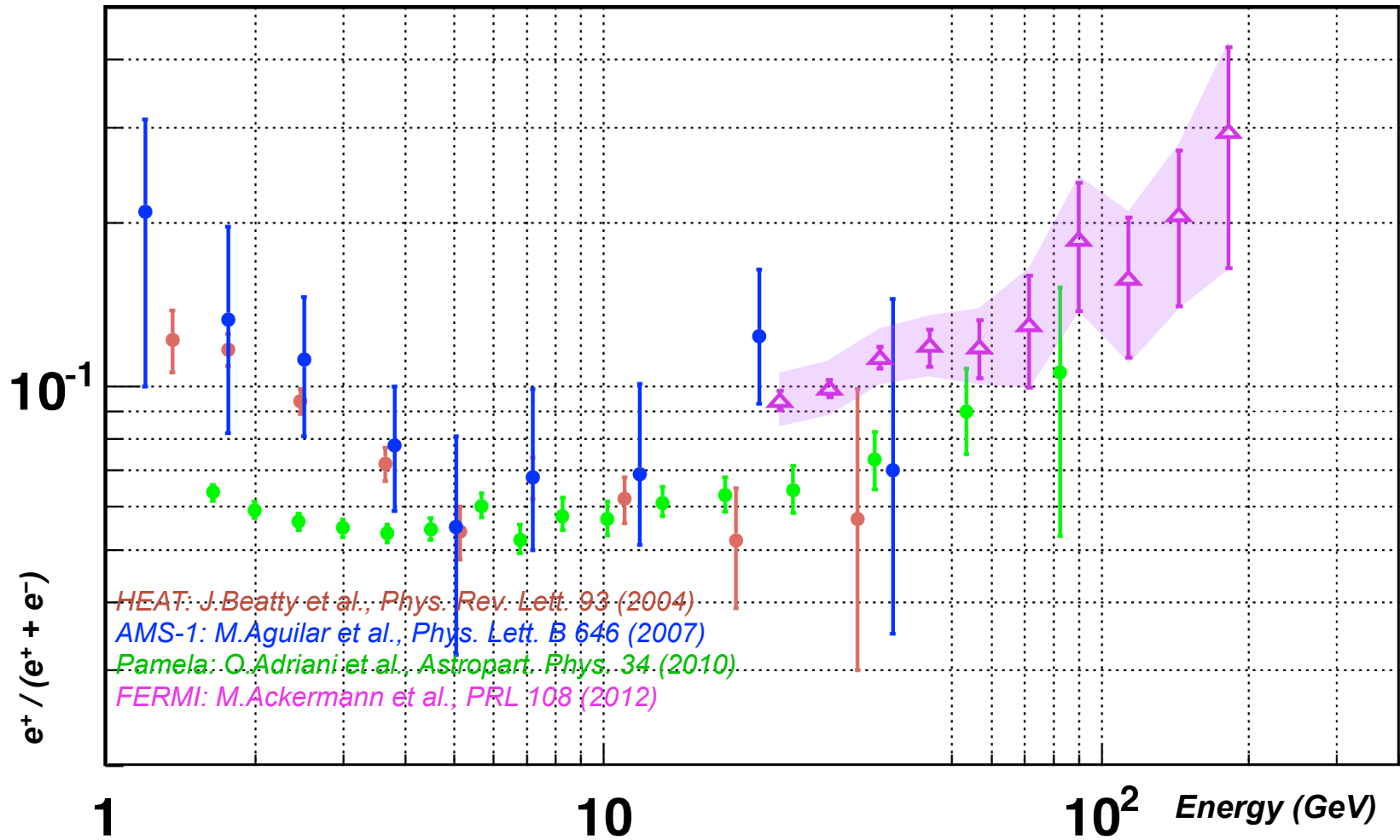
BDT and E/p matching combined give a Proton Rejection of 10^4



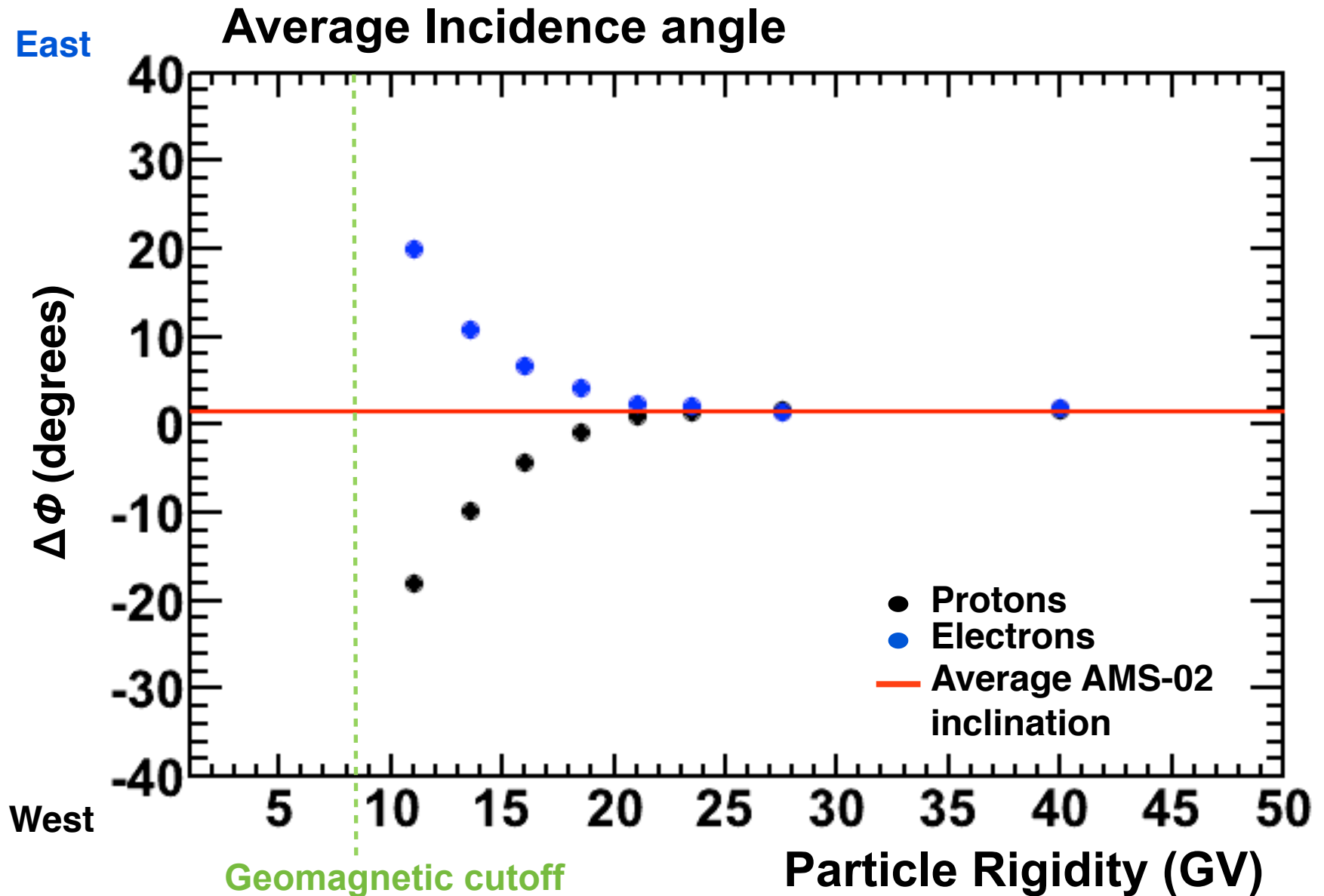
The TRD Likelihood shows clear separation between protons and positrons with a small charge confusion background



AMS-02 Positron Fraction will be published soon



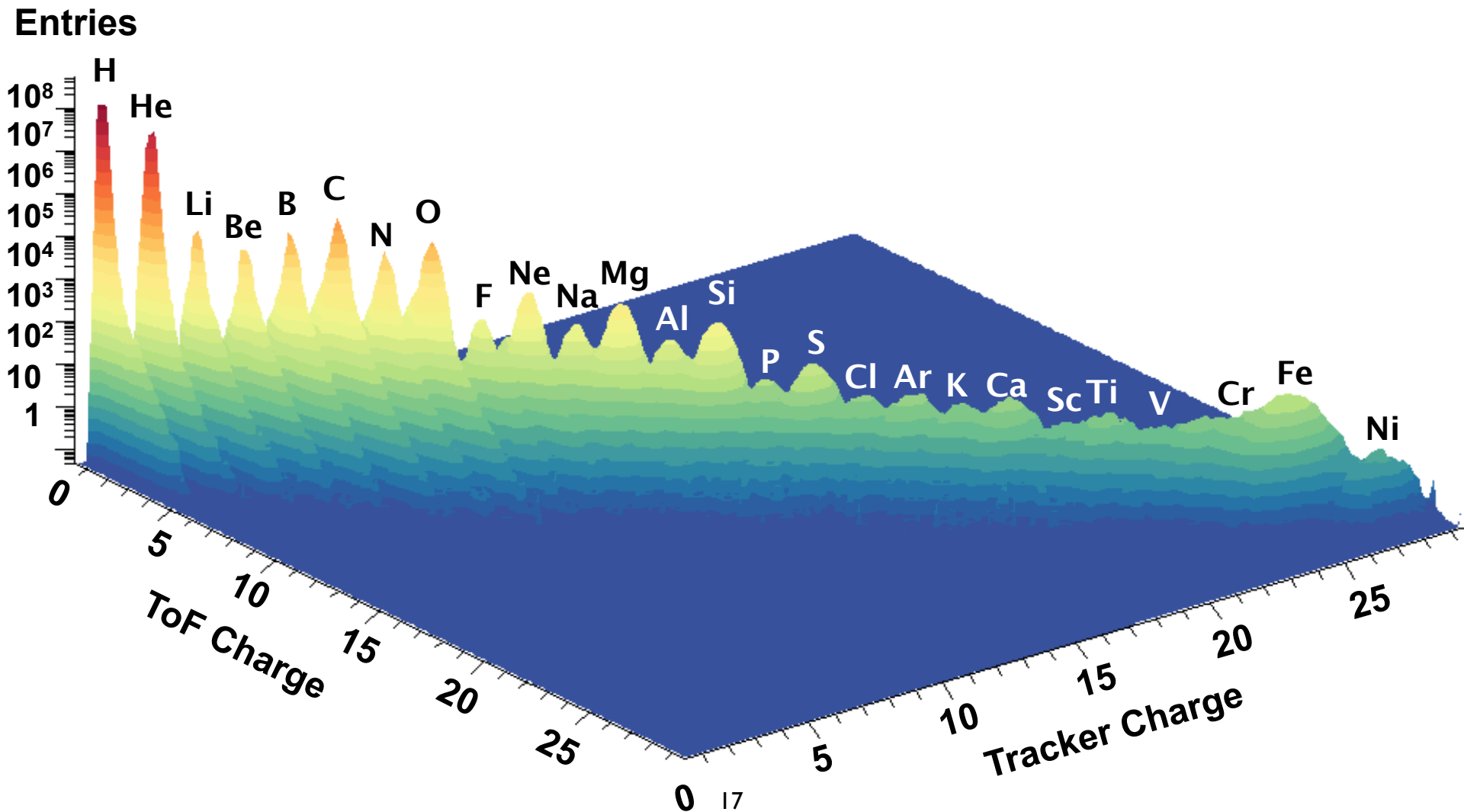
East - West effect



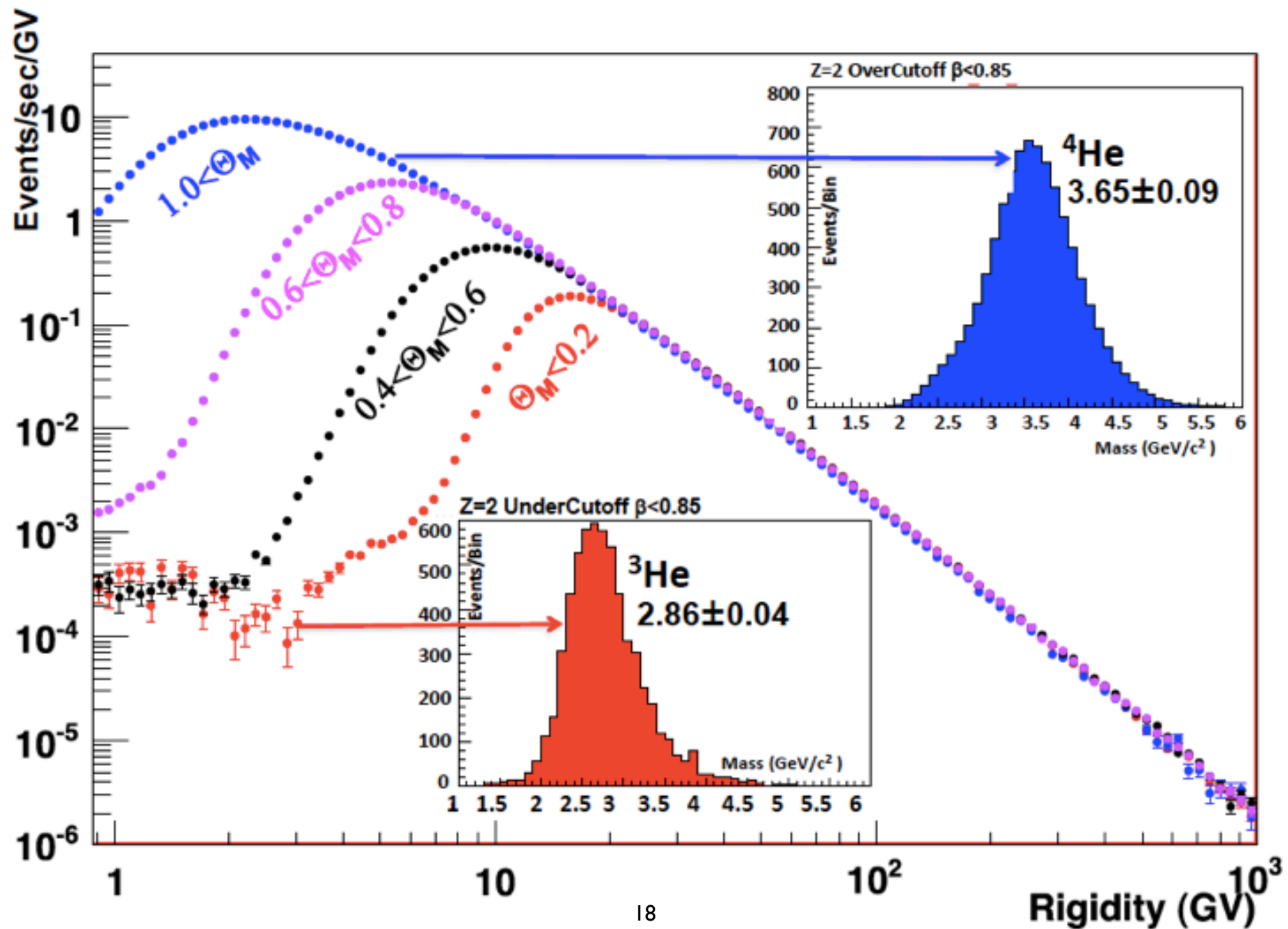
AMS Nuclei Measurement on ISS

Accurate Study of the composition of the cosmic rays

Multiple Independent Measurements of the Charge ($|Z|$)

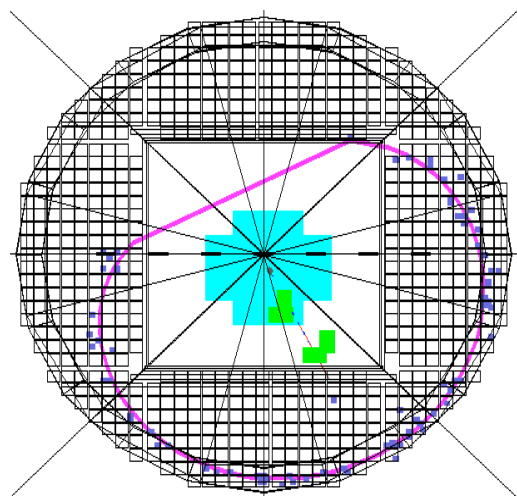
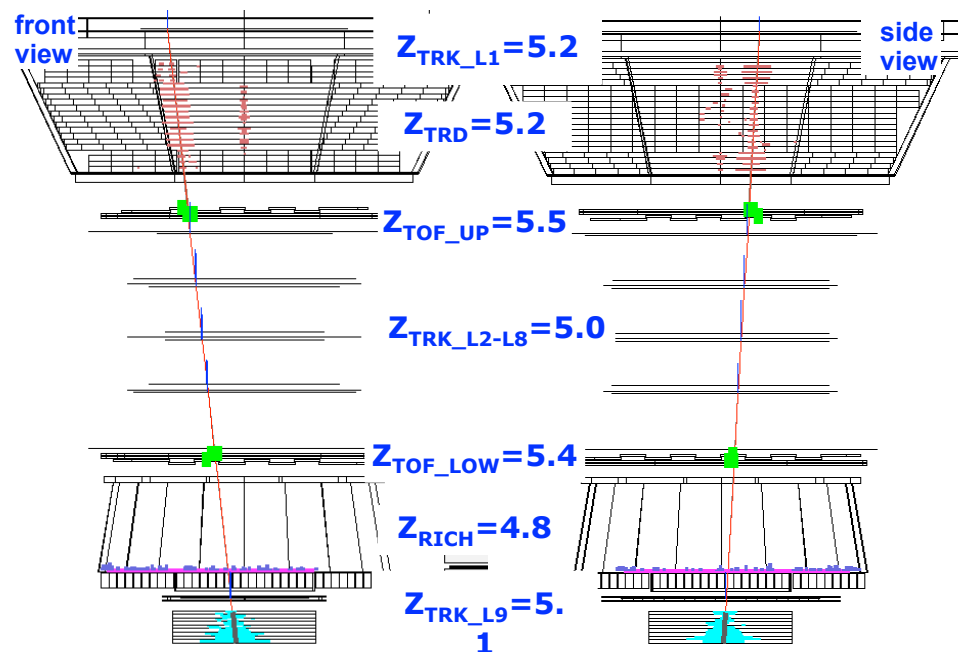


Helium rate

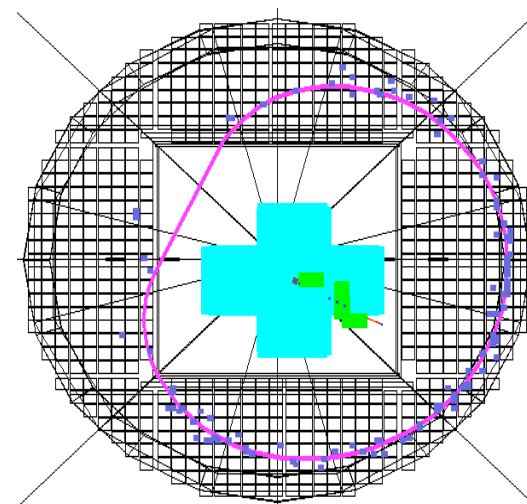
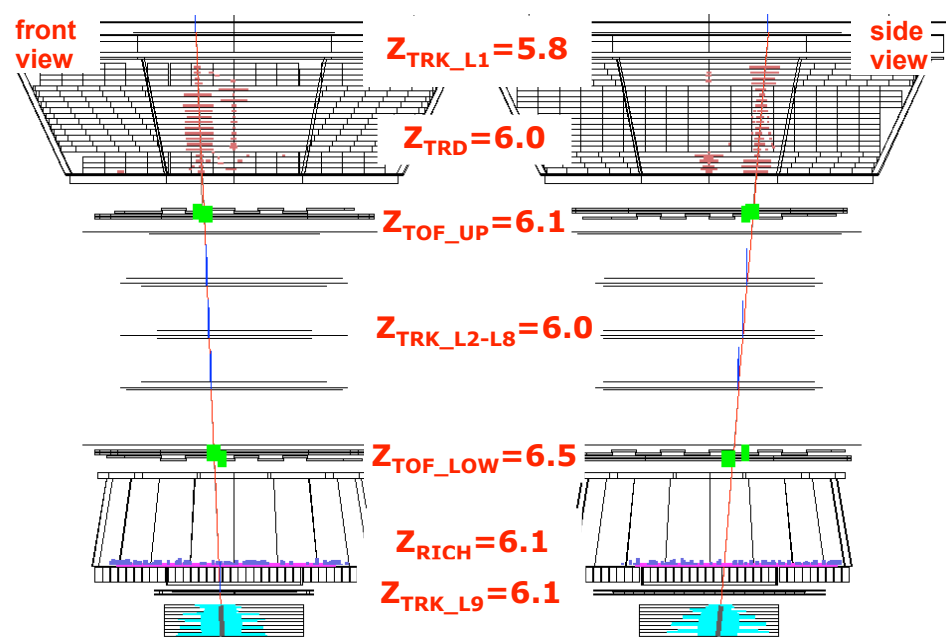


Boron and Carbon

Rigidity=680 GV

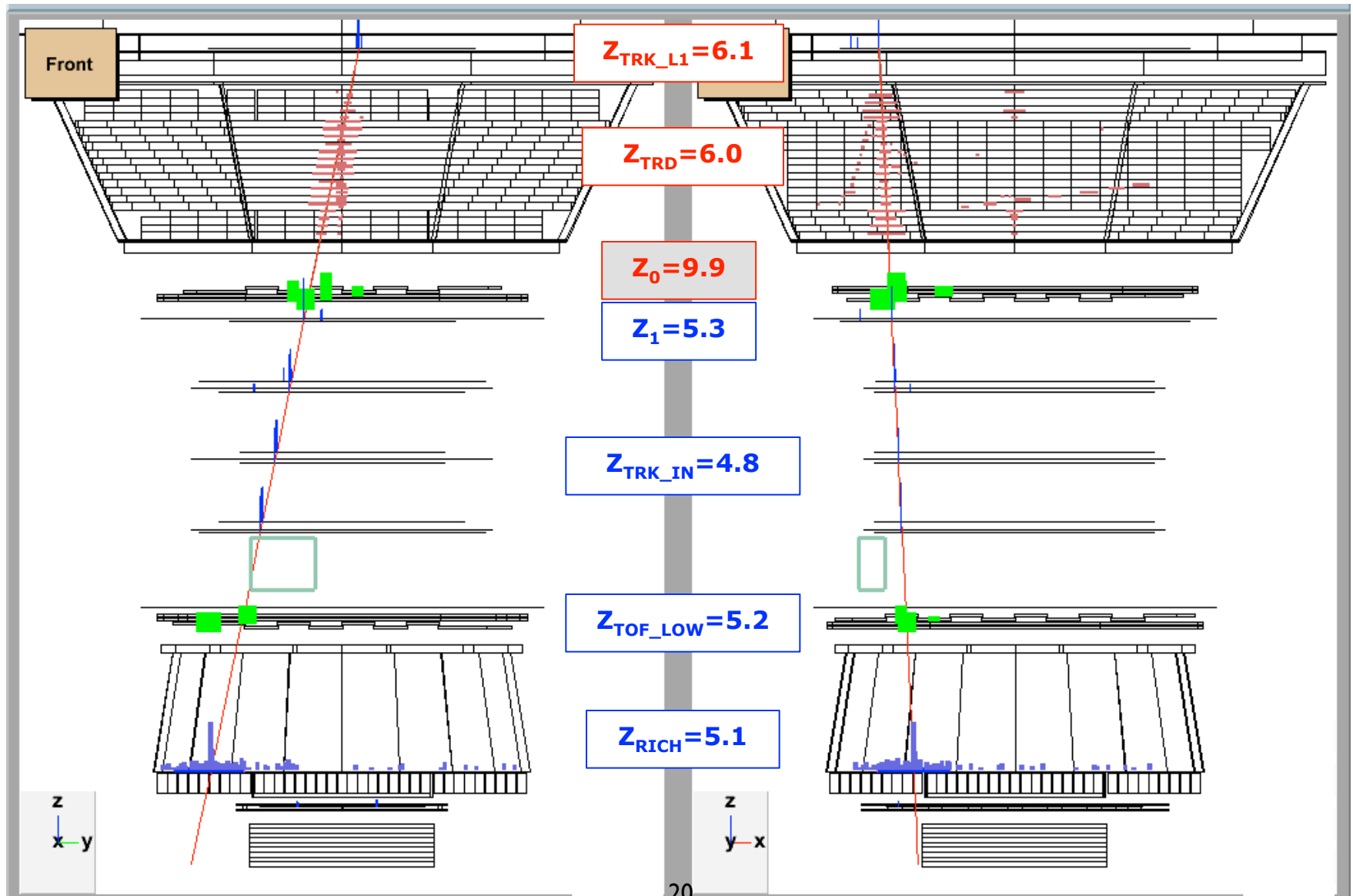


Rigidity=666 GV

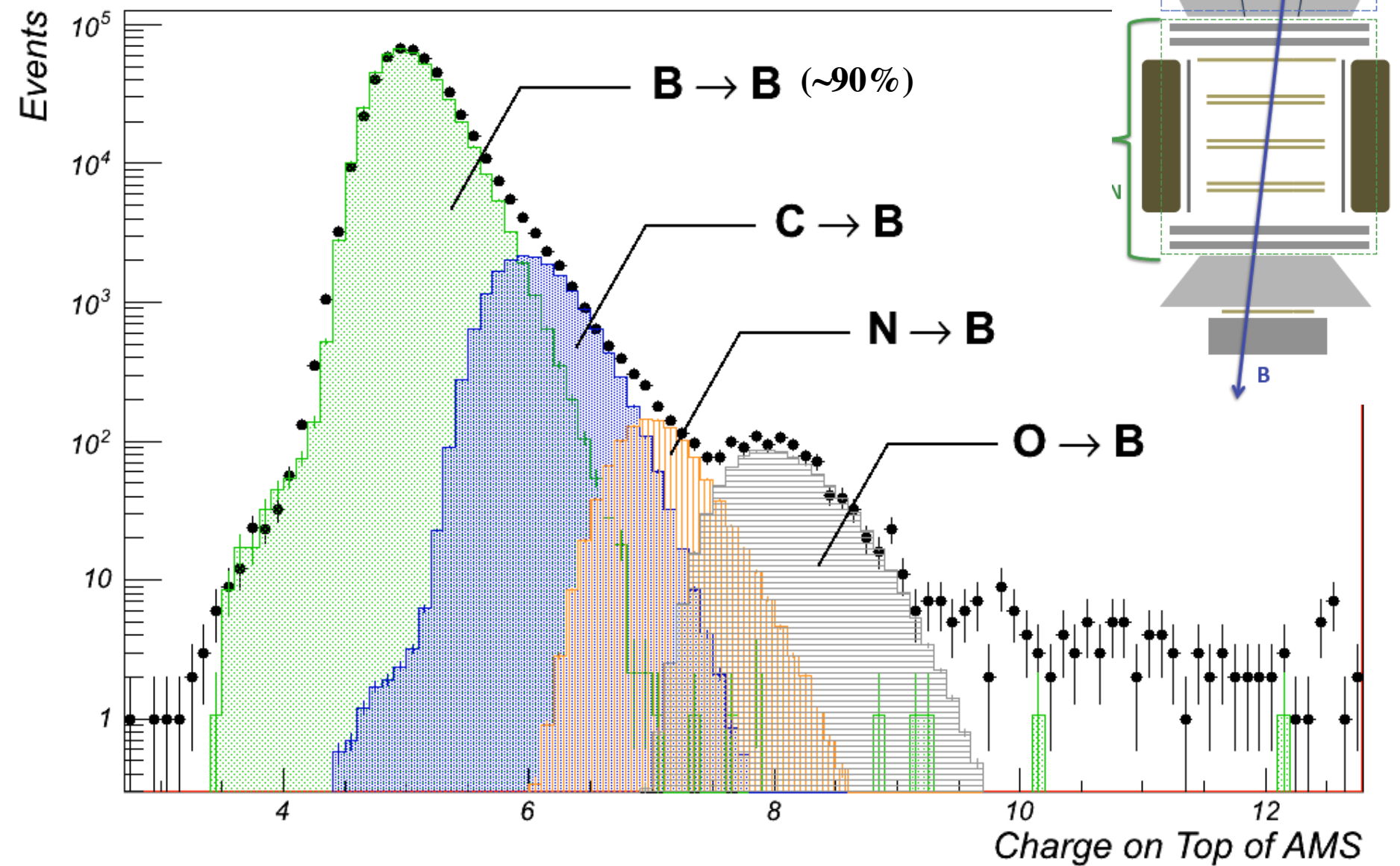


Carbon Fragmentation to Boron in Upper TOF

Rigidity 10.6 GV



Boron measured by AMS



Conclusions

- AMS02 is in on the ISS since May 19th 2011 and all the detectors are properly functioning
- Detector calibration (alignment, e/p rejection, charge id, etc.) are well advanced
- Data analysis is in progress (positron fraction, P and He fluxes, B/C ratio, gamma)
- 10+ years on board the ISS at 10^9 events/year will provide enormous sensitivity and statistic: great physics potential

**We want to thank NASA and DOE
for making AMS possible!**

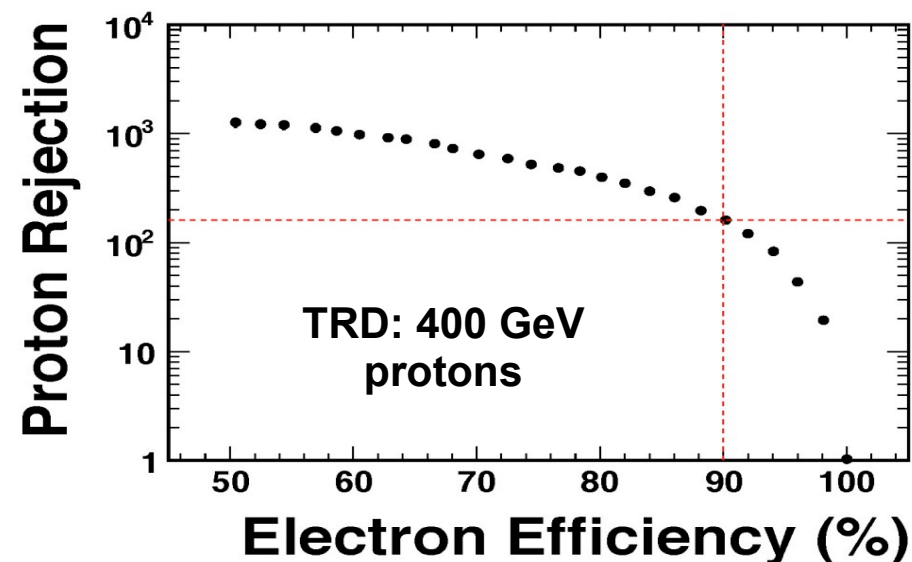
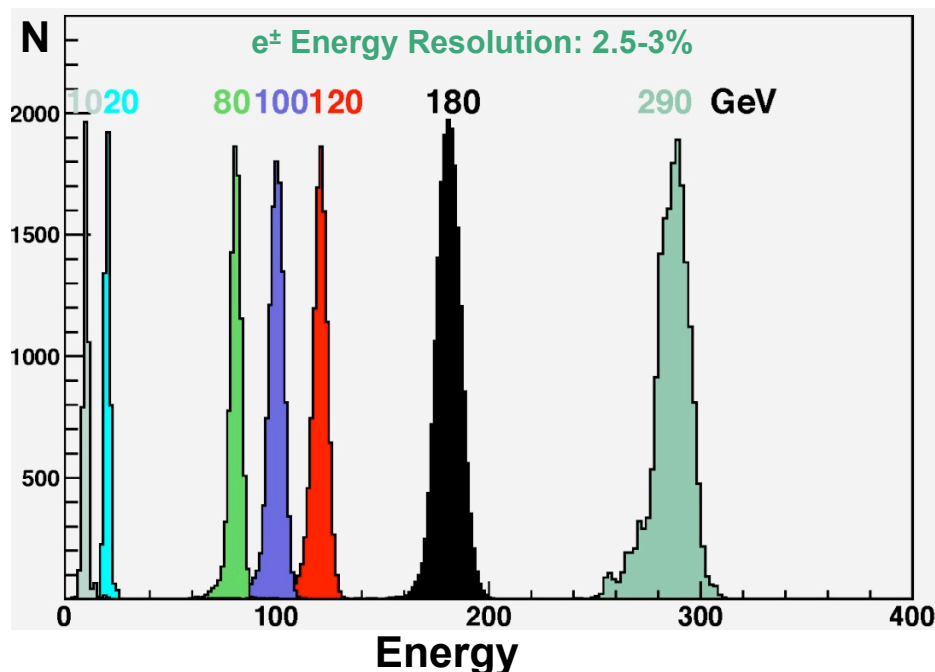
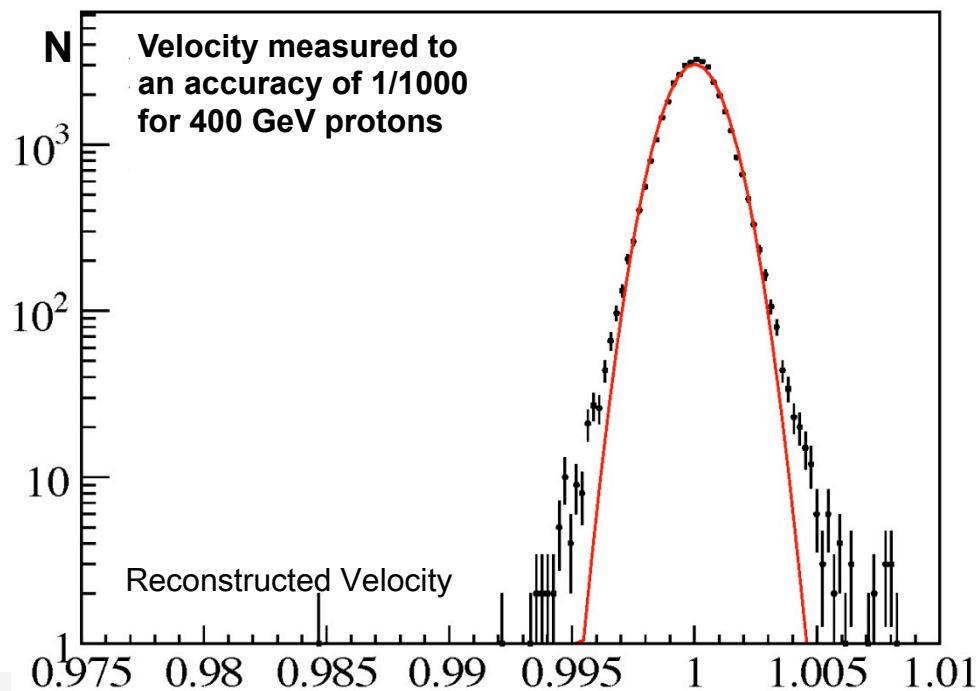
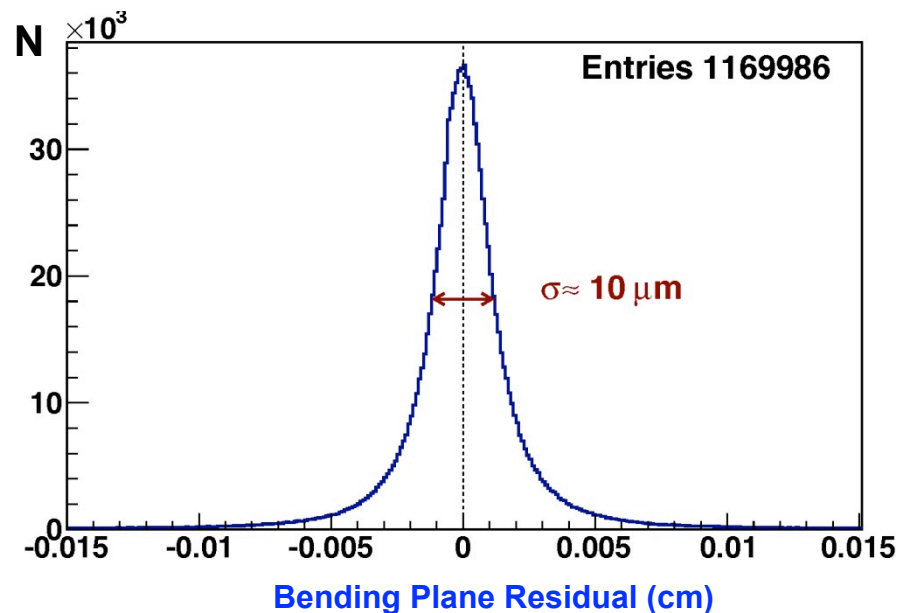




Science coming soon!!! Stay tuned!!!

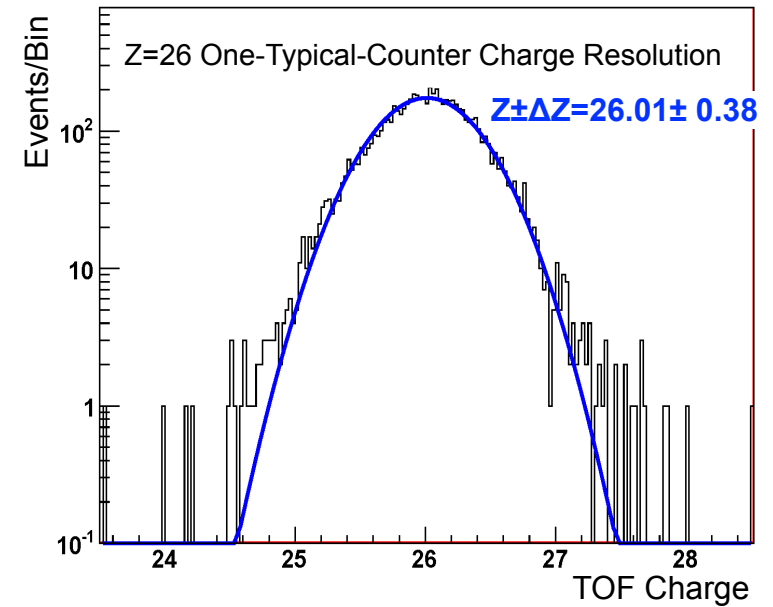
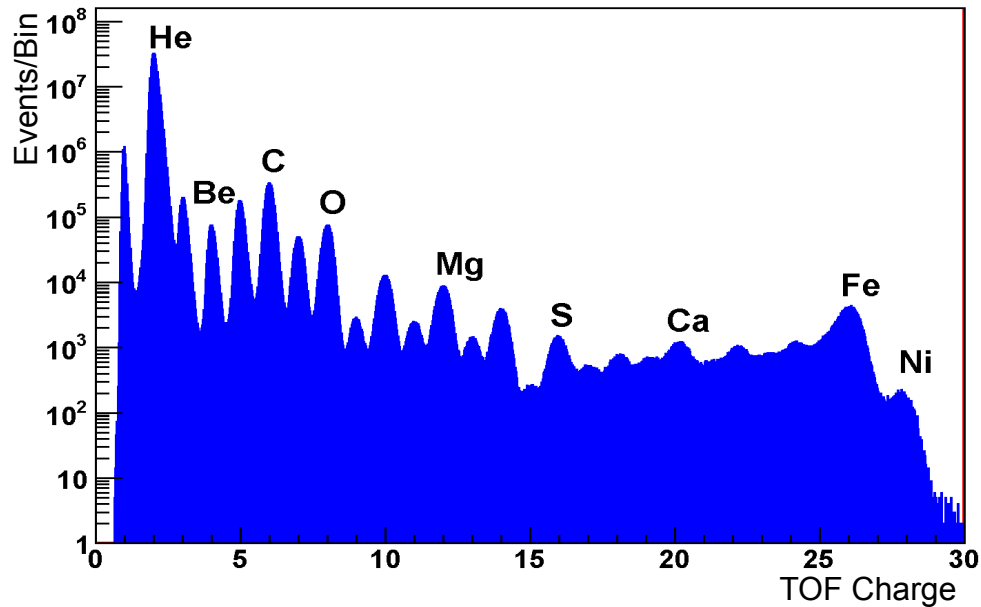
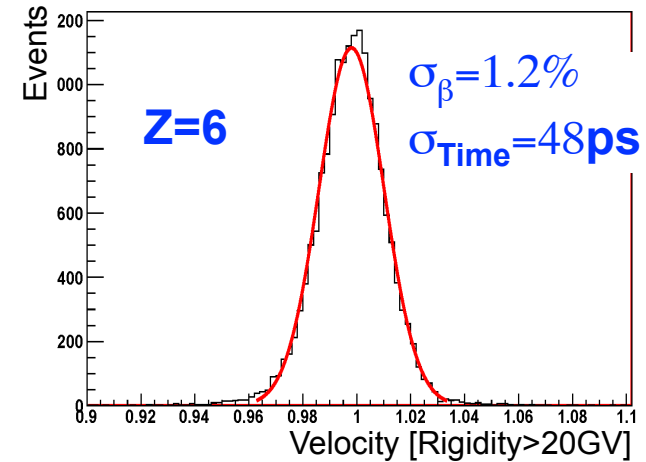
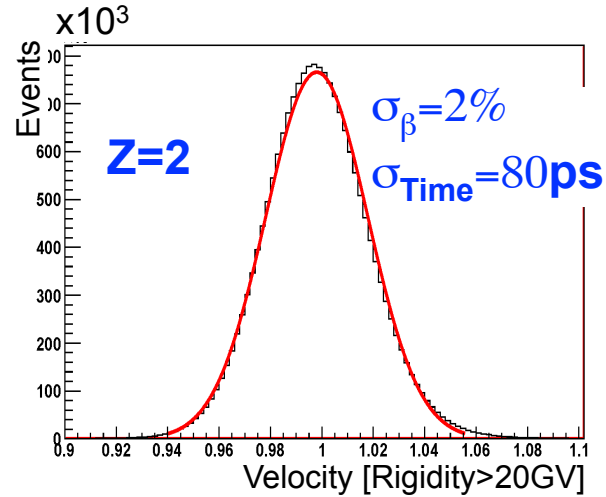
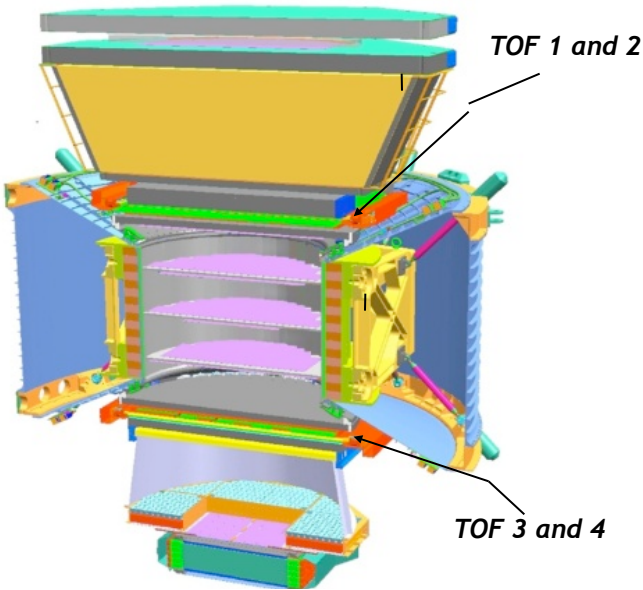
Back-ups

Test Beam Results used as reference points



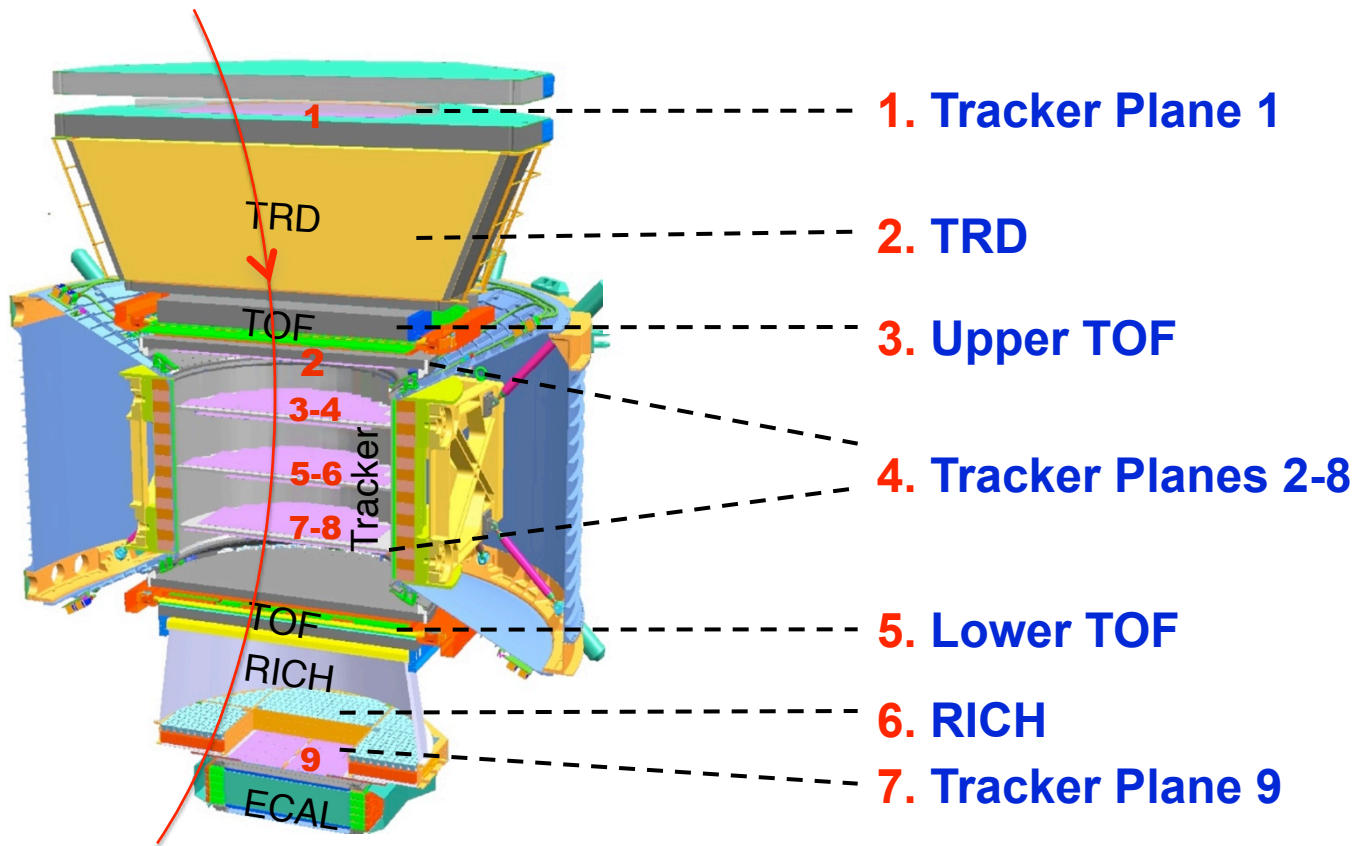
The Time of Flight System - Data from ISS

Measures Velocity and Charge of particles



Accurate Study of the composition of the cosmic rays

Multiple Independent Measurements of the Charge ($|Z|$)



Search for New Matter in the Universe

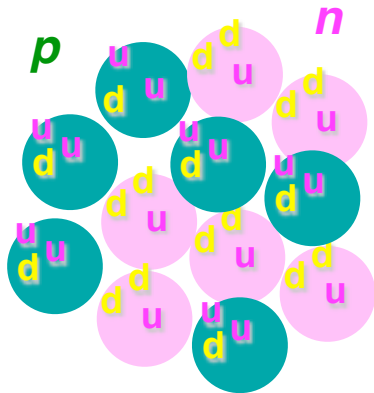
After many years, the question of the existence of strange quark matter still remains without a definitive answer.

There are six types of Quarks found in accelerators (u, d, s, c, b, t).

All matter on Earth is made out of only two types (u, d) of quarks. “Strangelets” are new types of matter composed of three types of quarks (u, d, s) which should exist in the cosmos.

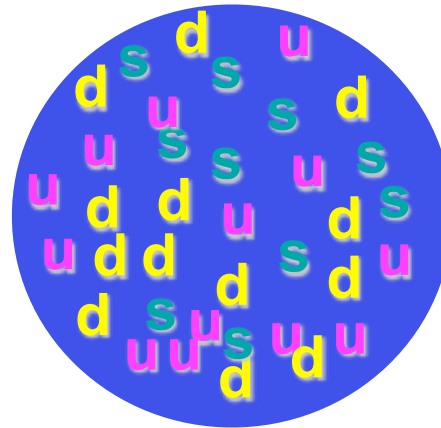
Carbon Nucleus

$Z/A \sim 0.5$



Strangelet

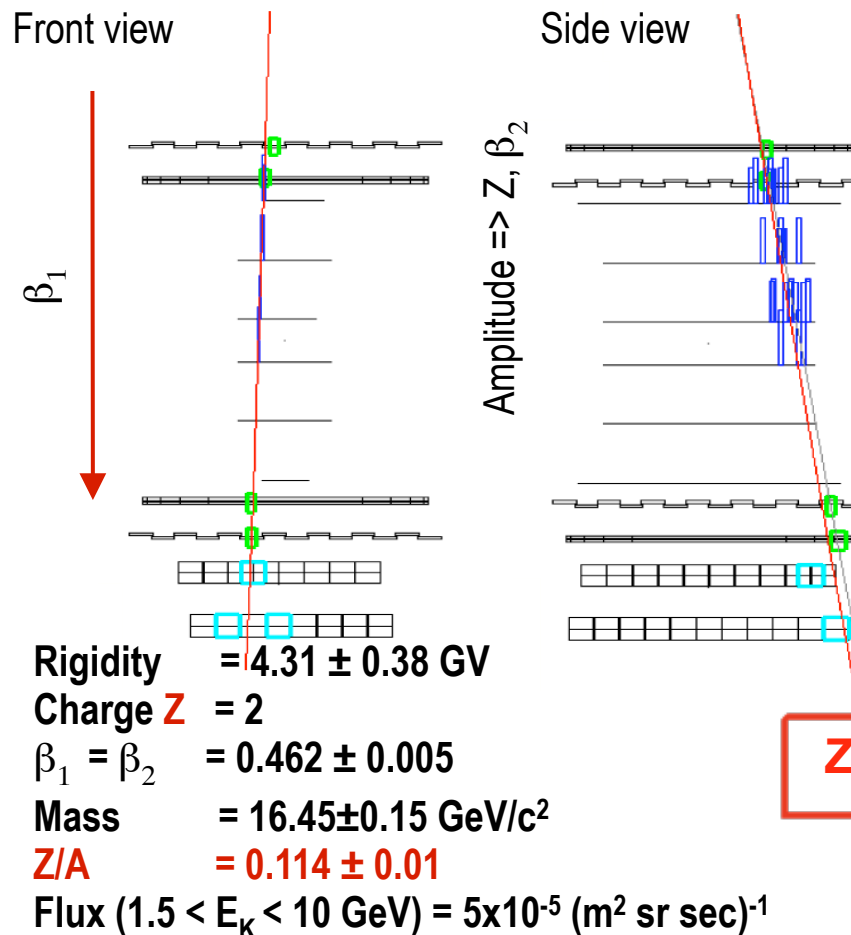
$Z/A \sim 0.1$



Strangelets

Jack Sandweiss (Yale) is leading the AMS search.

Candidate observed with AMS-01 5 June
1998 11:13:16 UTC

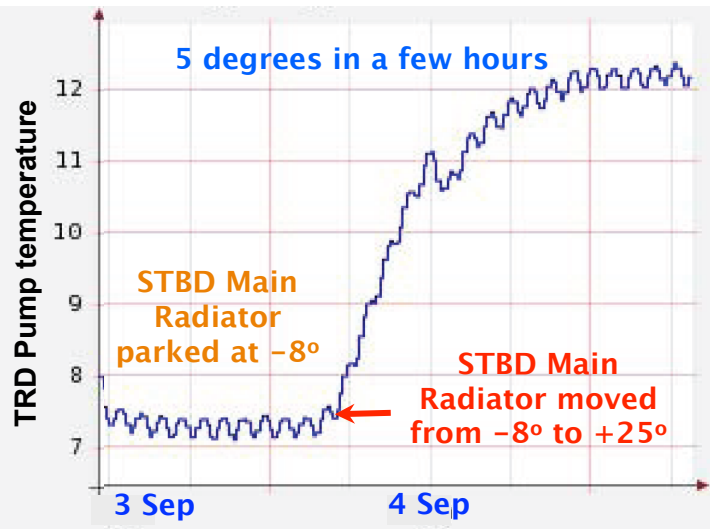
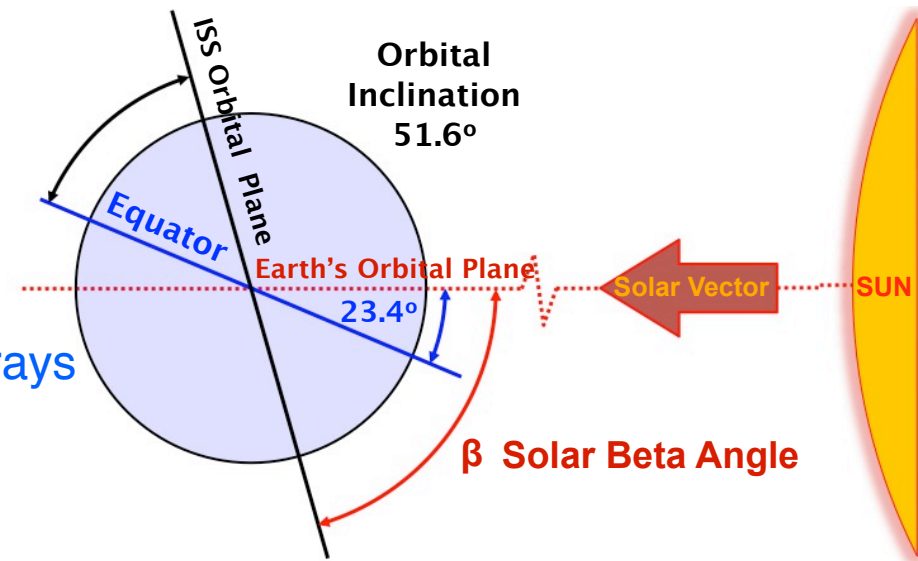


On-orbit thermal control

The thermal environment on ISS is constantly changing due to:

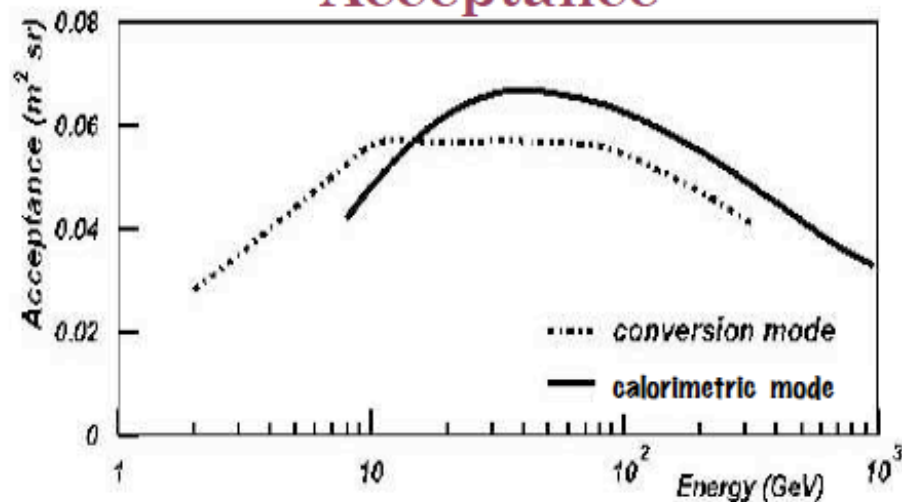
- Solar Beta Angle (beta)
- Position of the ISS Radiators and Solar Arrays
- ISS Attitude

Over 1,100 temperature sensors and 298 heaters are monitored to assure components stay within thermal limits and avoid permanent damage.

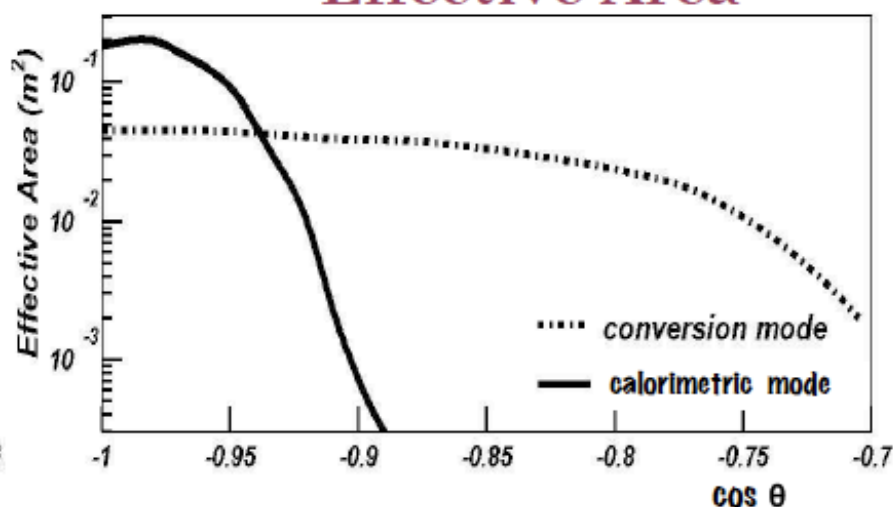


Tracker and Ecal Performances

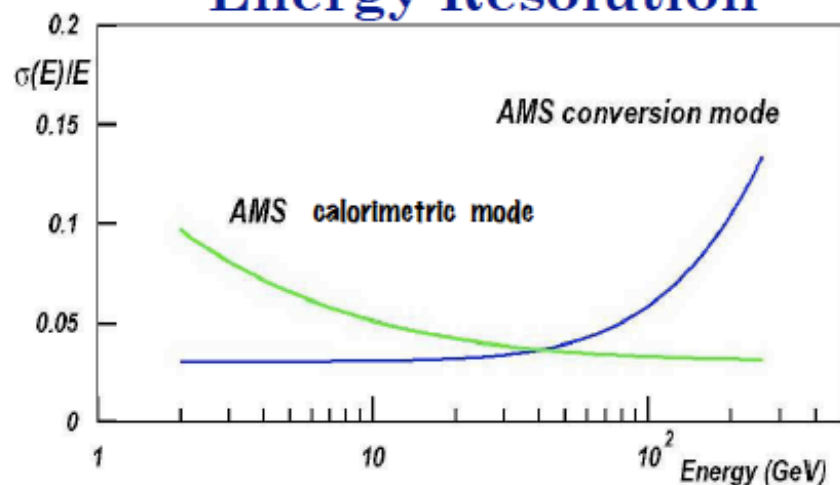
Acceptance



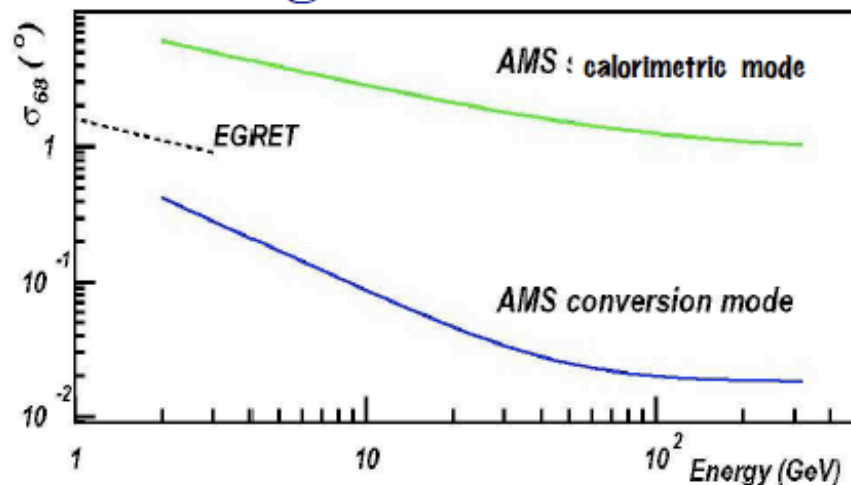
Effective Area



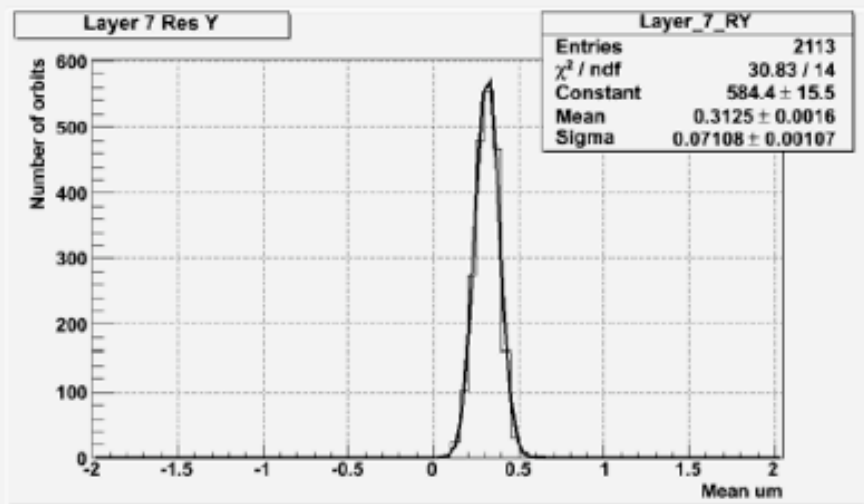
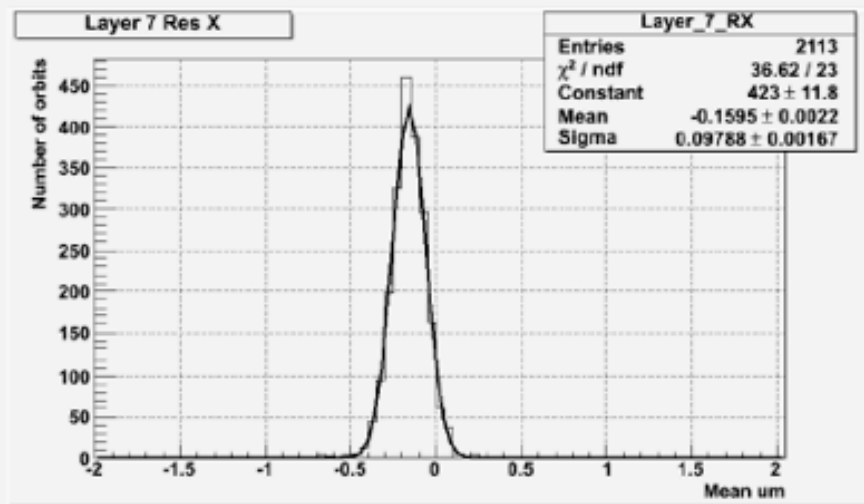
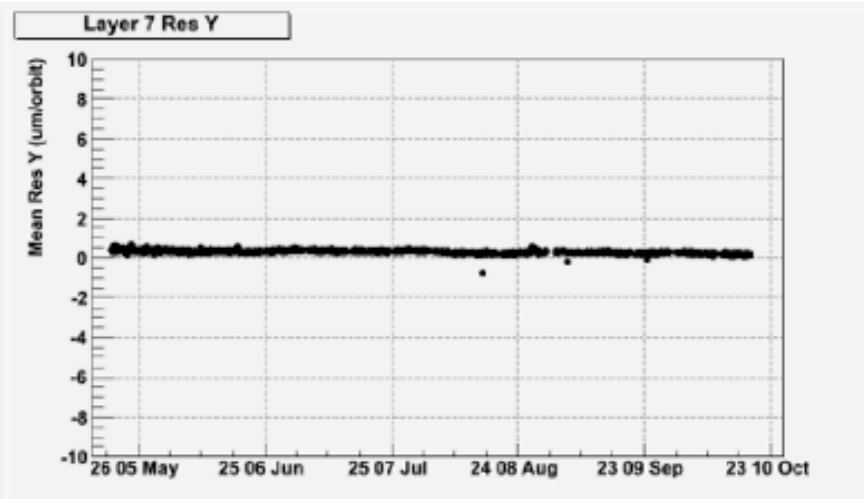
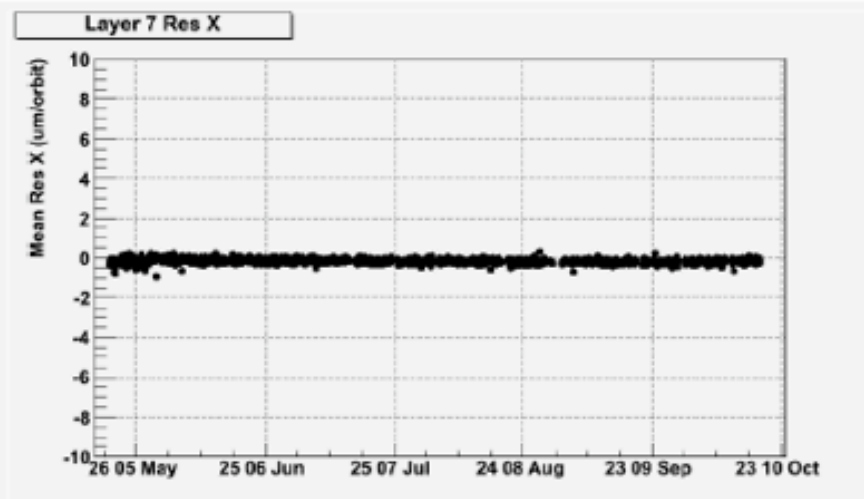
Energy Resolution



Angular resolution



Typical stability of an inner tracker plane is $0.1\ \mu\text{m}$ as measured with Protons on the ISS



Material on AMS-02

AMS Hadronic Tomography

with the cosmic-ray p/He ratio

Exposure Time: May 20 2011 - May 20 2012

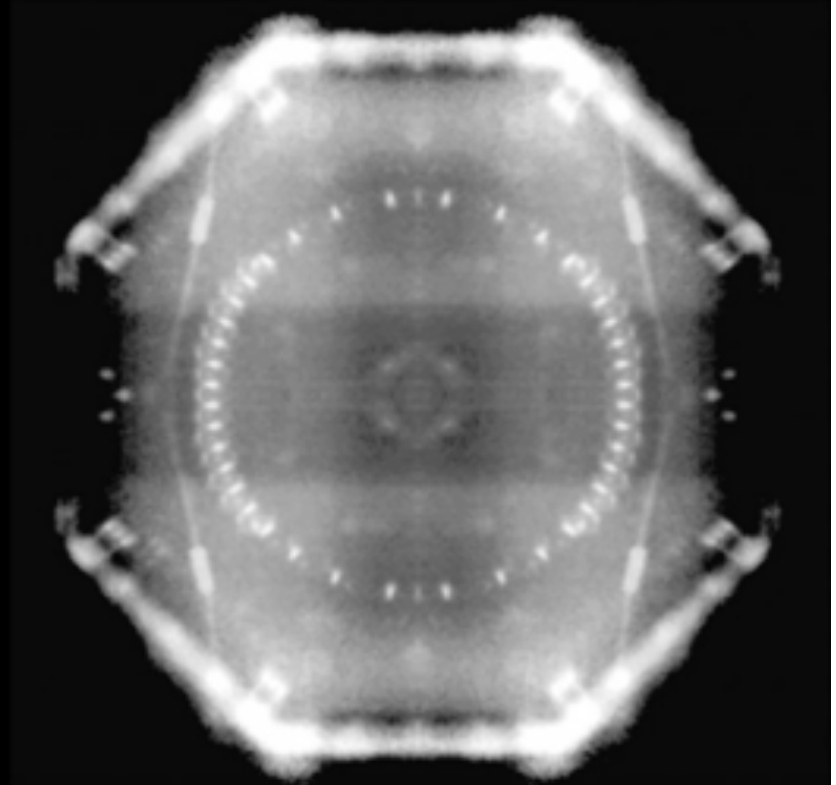
Number of Protons: 3,676,863,217

Number of Helium nuclei: 620,303,906

Rigidity range: 2 GV - 2000 GV

Tomographic plane: $Z = +165$ cm

XY pixel area: 1 cm^2



As estimations:

3 g/cm^2 in the TRD

2 g/cm^2 in the TOF

AMS Data/MC Volumes Projected

DATA

Per Year Of Operation:

- 1.6×10^{10} Events
- 35 TB Raw Events
- 130 TB Reconstructed Ev.

MC

Per Year Of Operation:

- $\sim 2 \times 10^{10}$ Simulated Events
- ~ 200 TB Simulated Data Volume