

# TARA: Radar Detection of UHECRs



John Belz, University of Utah  
for the *Telescope Array Radar* Project  
Cosmic Frontier Workshop, SLAC  
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# The TARA Project

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*and the Telescope Array Collaboration*

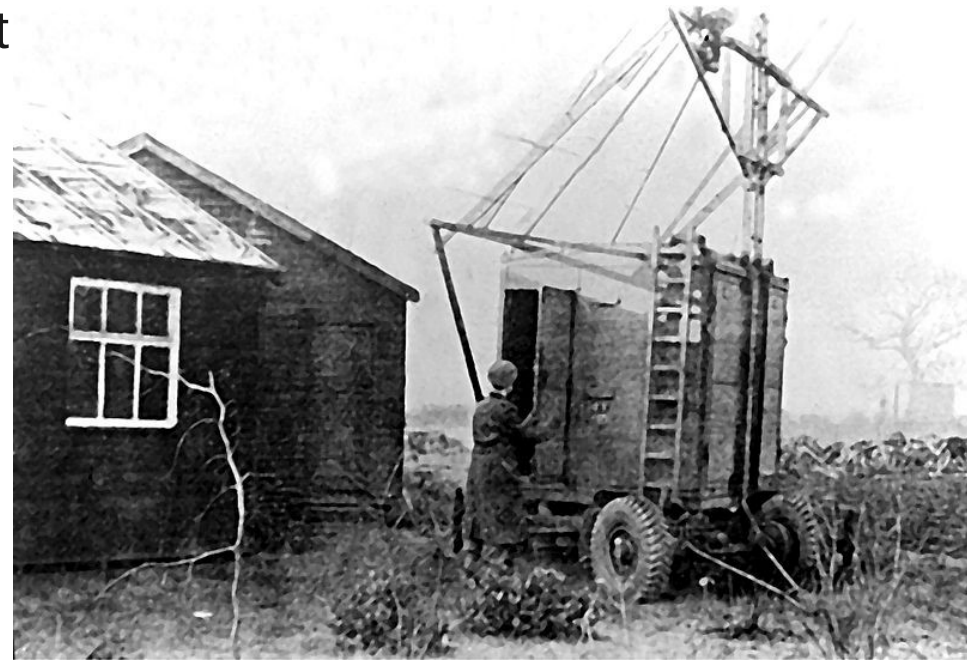
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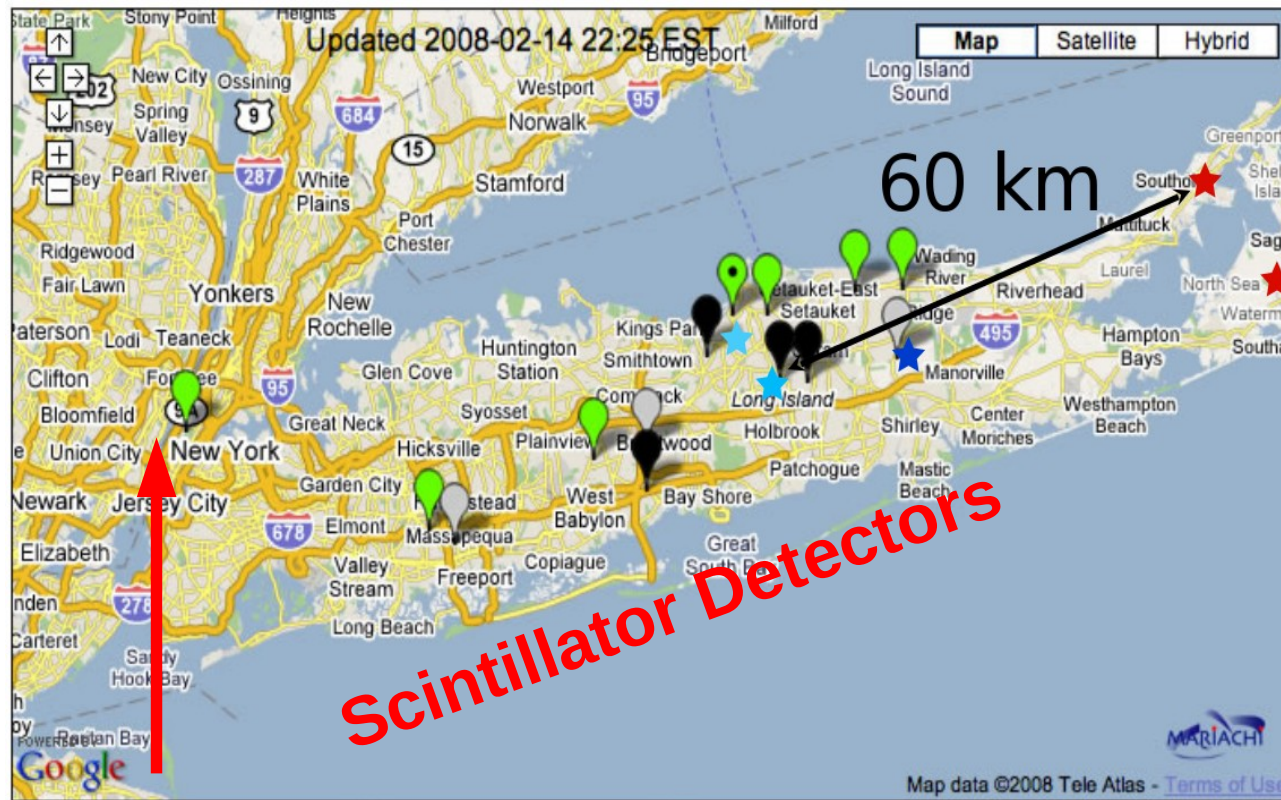
# Cosmic Rays and Radar Detection

- 1940 – Blackett and Lovell; Used to explain anomalies in atmospheric radar data. Built “facility” (right), no signals reported.
- 1968 – Suga et al; Propose experiment, but no results.
- 2000 – Gorham; revisits, updates calculations
- 2003 – Iyono et al; propose measurements with LAAS array. No results reported.
- 2007 – Wahl et al (Jicamarca), anomalies seen; consistent with CR?
- 2009 – Terasawa et al (Shigaraki), anomalies seen; consistent with CR?
- 2009 – MARIACHI → *Parasitic Bistatic radar*

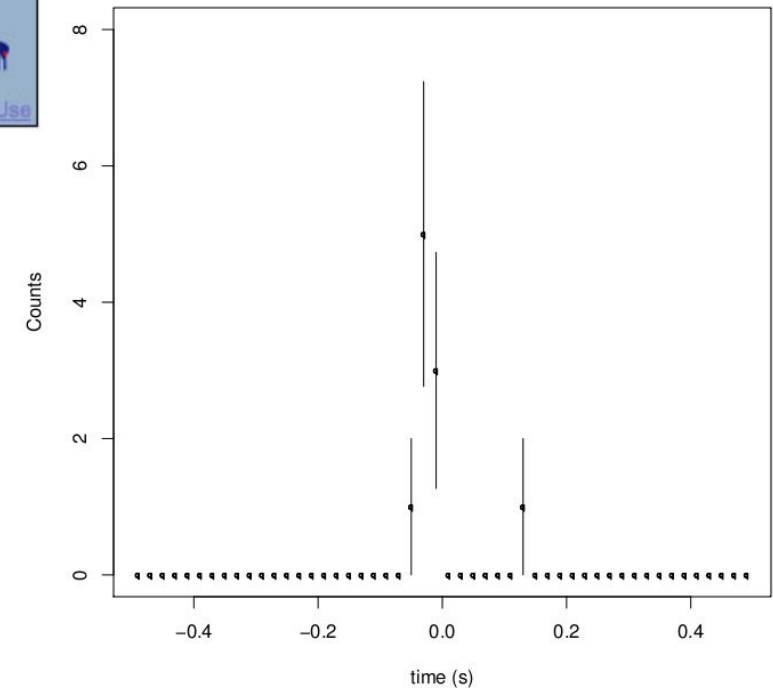


Jodrell Bank radar cosmic ray observatory, 1945

# MARIACHI



Time distribution of radar echos relative to nearest MARIACHI scintillator firing

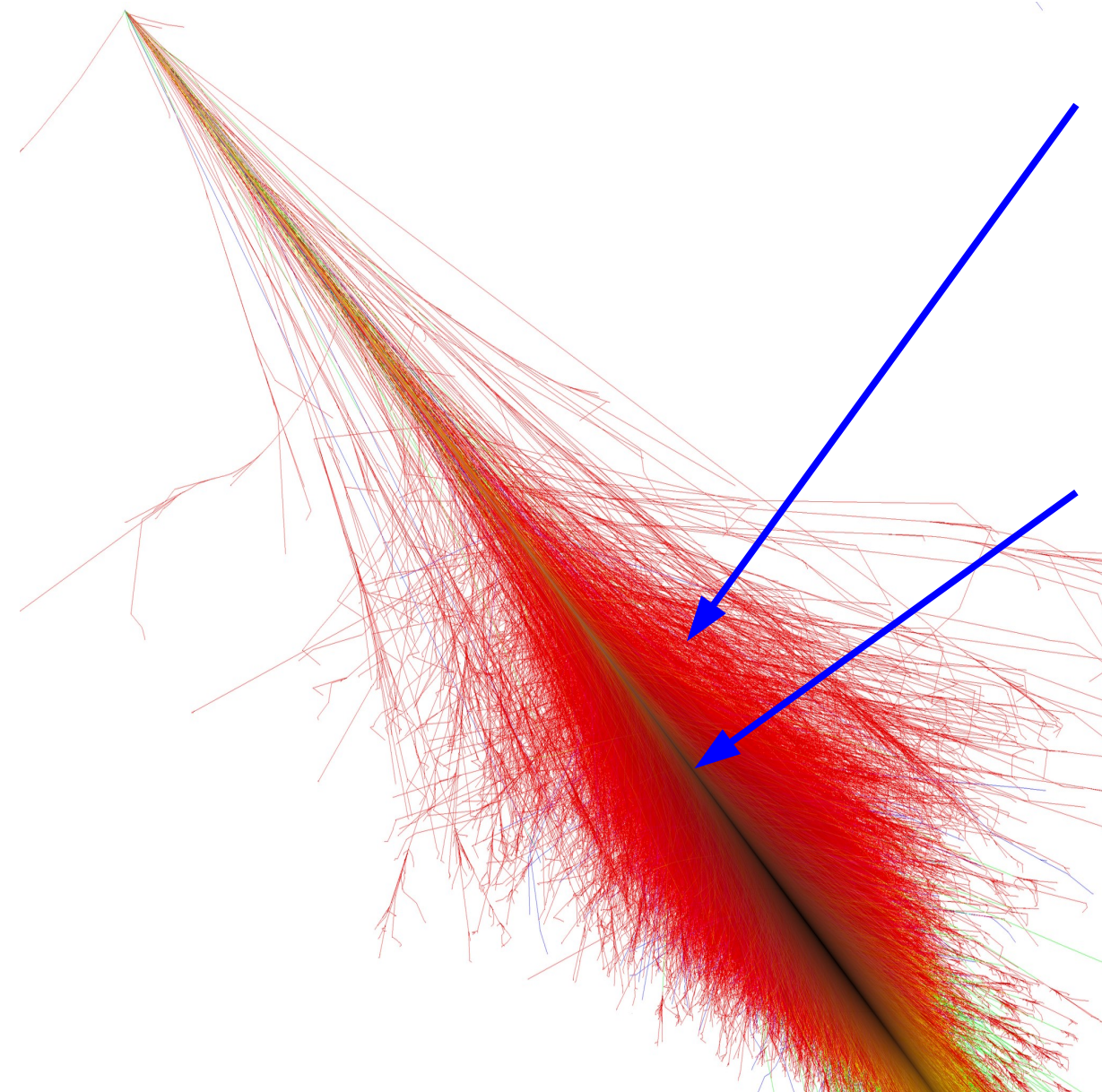




# MARIACHI → TARA

- **MARIACHI** coincidences intriguing but inconclusive
  - Low sampling rate receivers, crude surface detectors
  - Order-of-magnitude uncertainties on sounding radiation
- **TARA** is the ideal laboratory for testing radar technique
  - Single transmitter in radio quiet environment
  - Access to world class cosmic ray detector facility
  - High sampling rate receivers, smart triggering
- *TARA is the most ambitious effort to date to test the 70+ year-old idea that extensive air showers should reflect RF.*

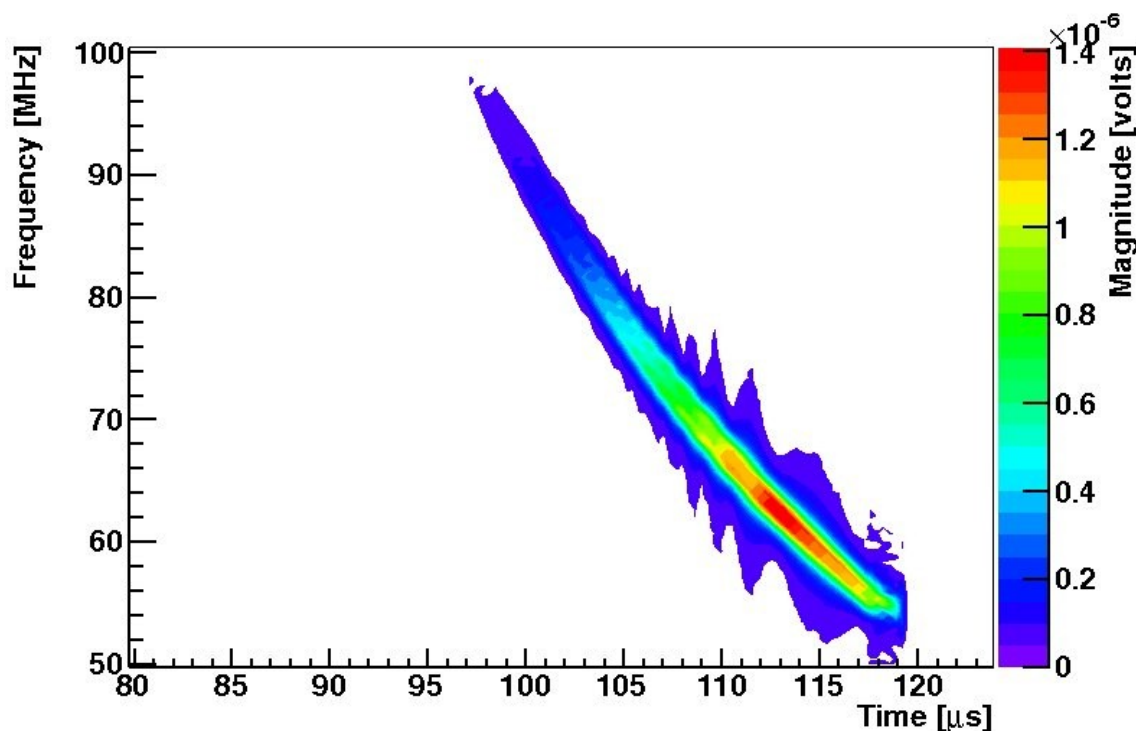
# What is the Radar Cross Section?



- Underdense Region
  - Most of the particles
  - $\sigma = (\text{Thomson} \times N_e)$
  - “Collisional Damping”
- Overdense Region
  - Macroscopic target
  - How big? Skin depth?

$$\nu_e = \sqrt{\frac{n_e e^2}{m \epsilon_0}} \frac{1}{2\pi}$$

# Signal Characteristics: The “Chirp”



Calculation: Isaac Myers

- Predicted signal for  $10^{19}$  eV shower,  $30^\circ$  from zenith; frequency vs time.
- Rapid movement of “target” produces Doppler-like frequency shift.
- Shape from **geometry**, not electron lifetime.
- Requires very high bandwidth/sampling rate receiver.
- Unique signature for air shower echoes.



# TARA1.5

## Telescope Array Locations General Reference Map



0 3,000 6,000  
Meters

WF2XHR

~50 km

U.S. Highway 6 & 50

Route 257

Middle Drum

Hinckley

Delta

Long Ridge

Black Rock Mesa

- TA Locations
- Communication Towers
- Fluorescence Locations
- ▲ Central Laser Facility
- Streams
- Lakes
- Town Boundaries
- State Land
- Private Land
- BLM Land
- Military Airspace: Sevier B

Department of Geography  
University of Utah  
April 2004



# TARA1.5: Low-Power Transmitter

- \* April 2011 to July 2012,  
54.1 MHz sine wave at 1.5 kW
- \* Low-gain (8.5 dBi) antenna





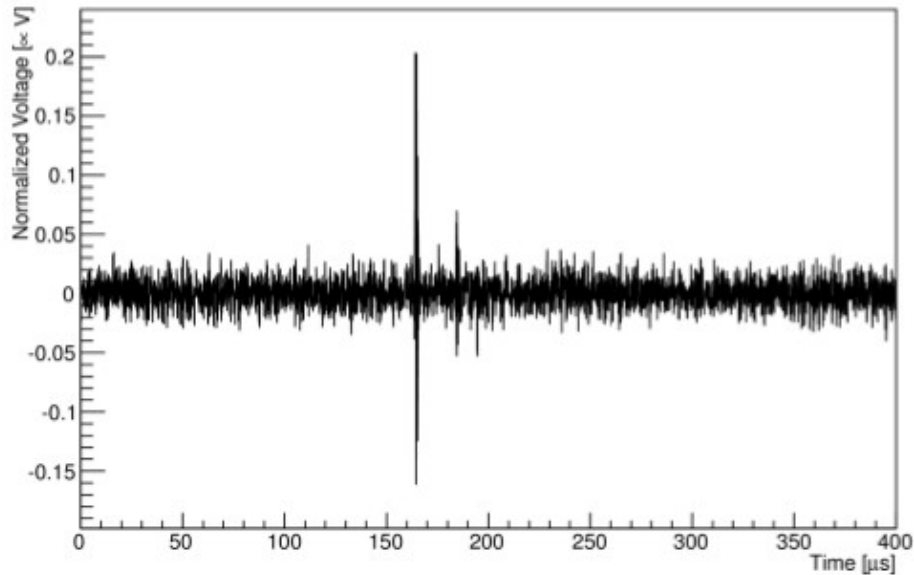
# TARA1.5 Receiver Station at Long Ridge

- Array of low-VHF log-periodic antennas (KU)
- USRP2 Software-defined radio RX
- FD Trigger, read at 12.5 Ms/s.
- $5\sigma$  threshold trigger, read at 6.25 Ms/s.

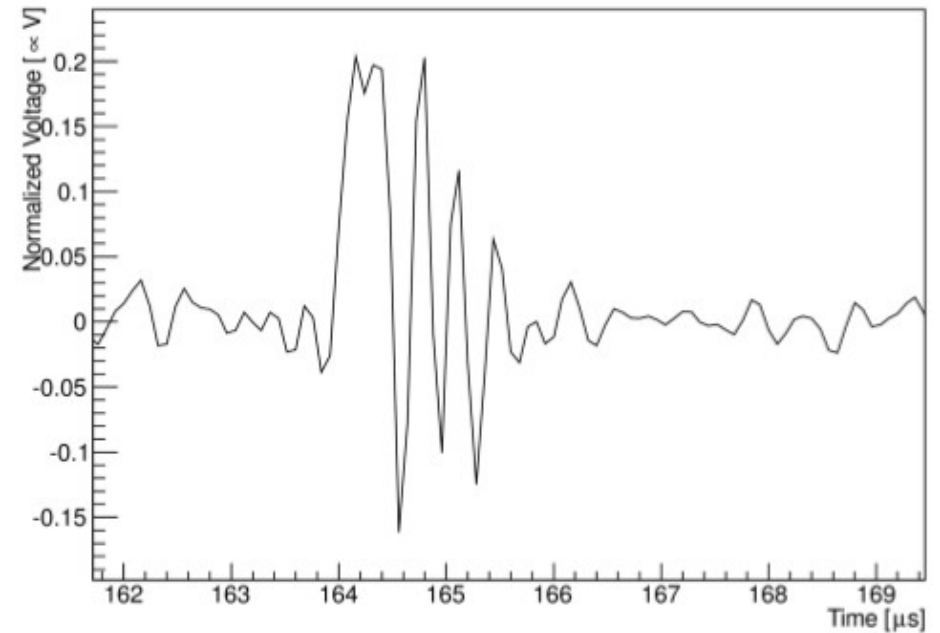


# Interesting Event, TARA1.5 FD-Trigger Stream

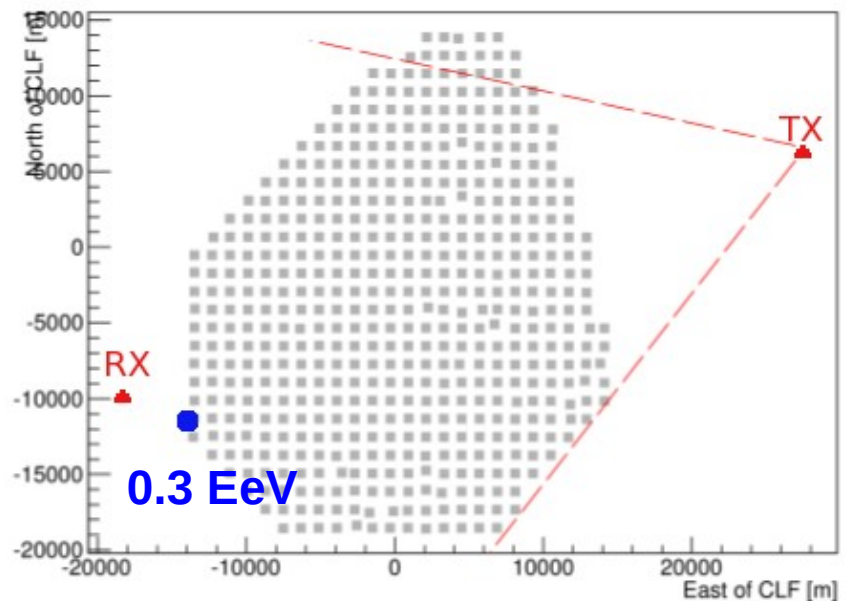
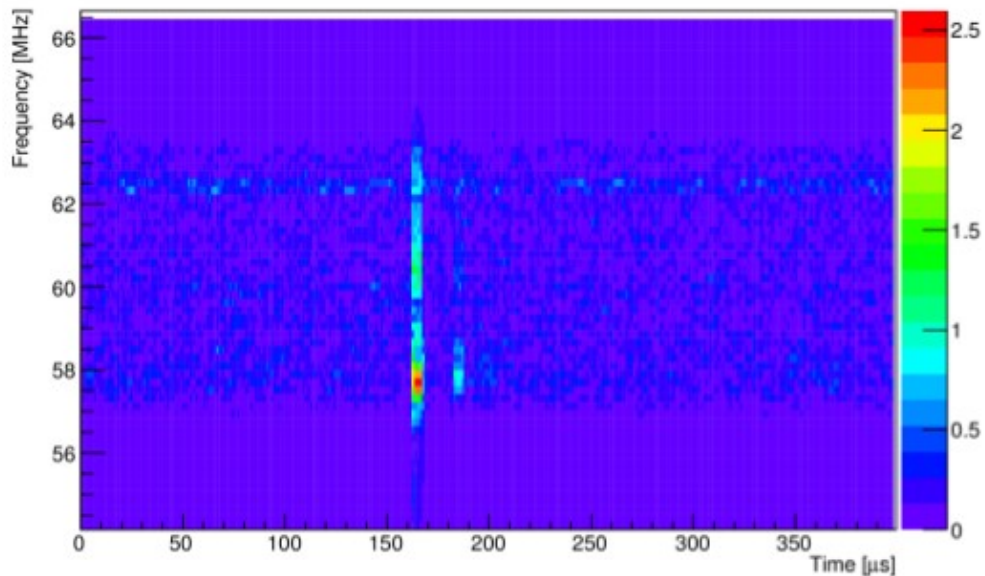
2012-01-17 06:13:36.620959245Z



2012-01-17 06:13:36.620959245Z



Spectrogram





# Spring 2013: TARA40

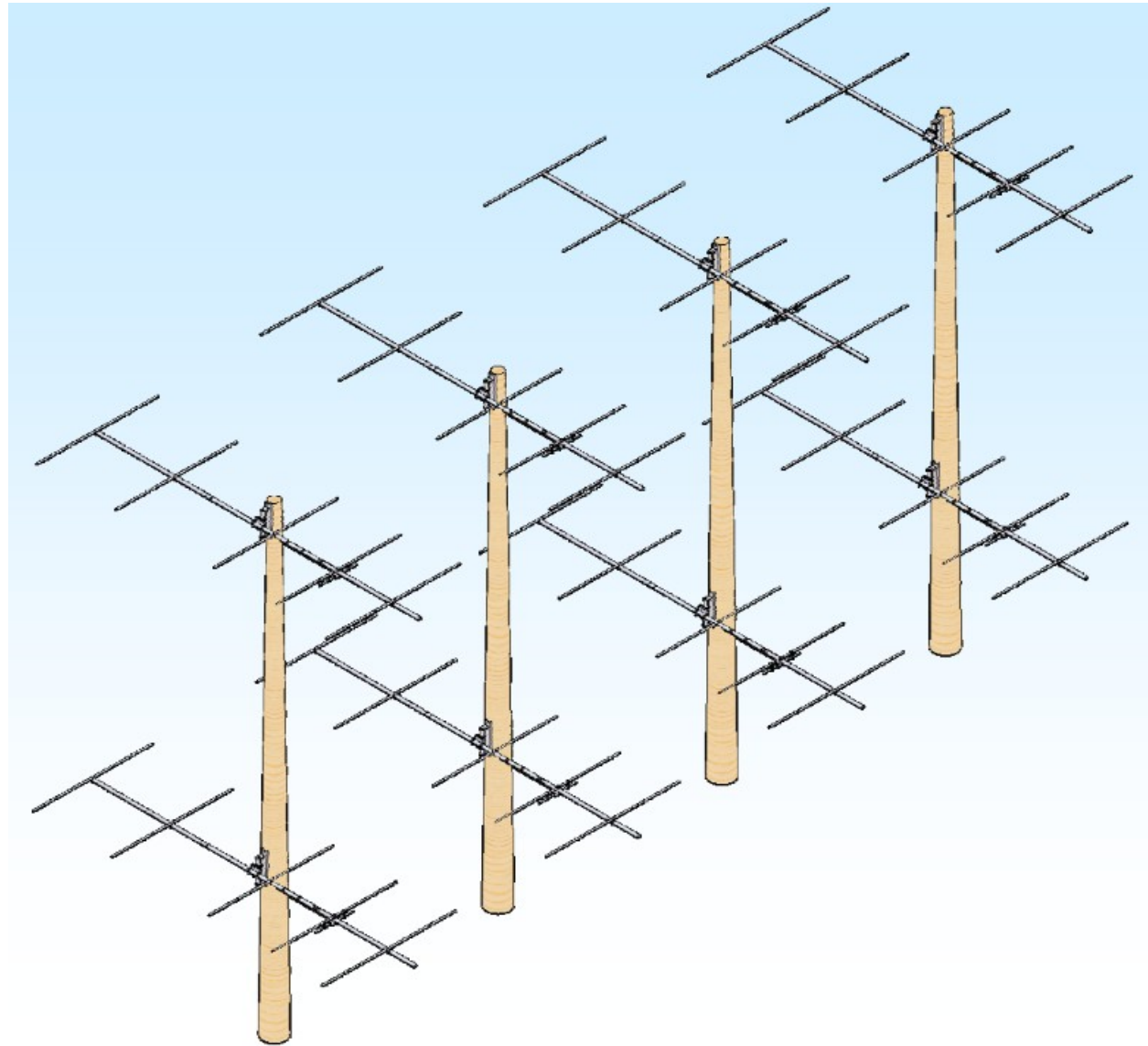
- 40 kW transmitter (x27)
- High-gain antenna (x5)
- Closer to ground array (x2.5)
- Improve RX DAQ



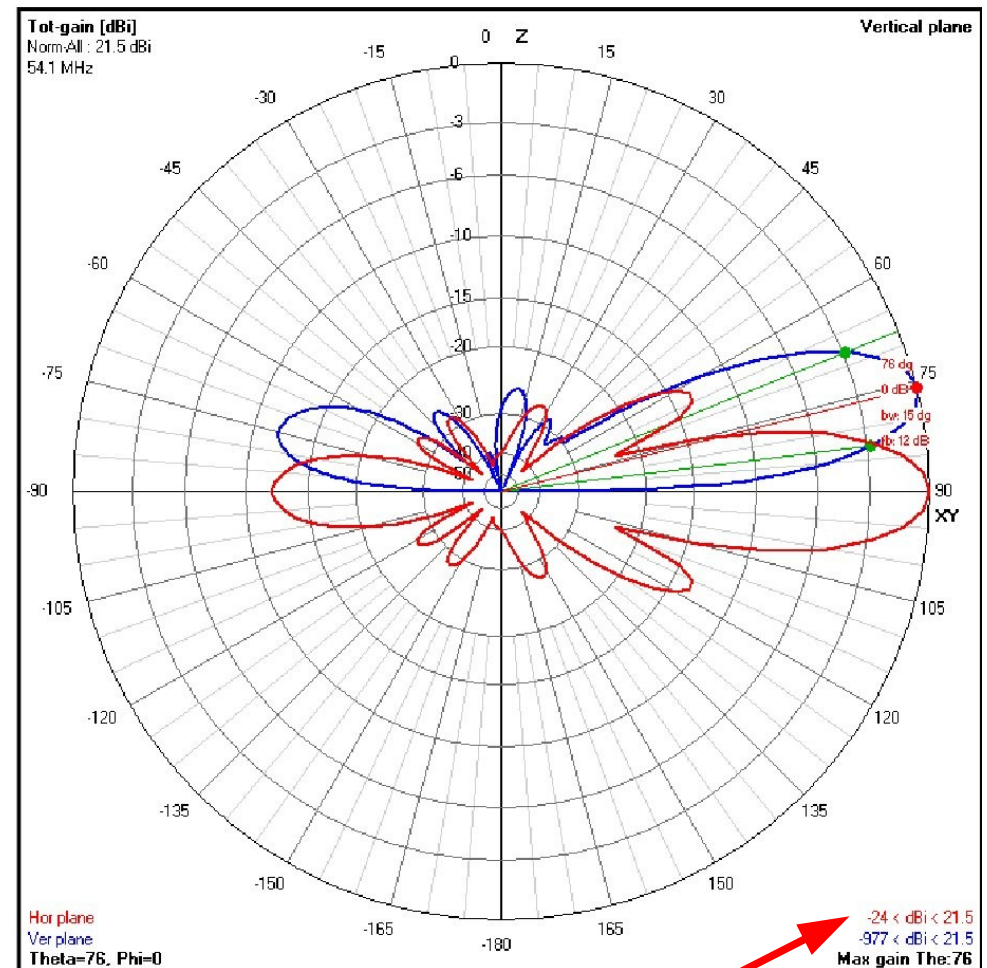
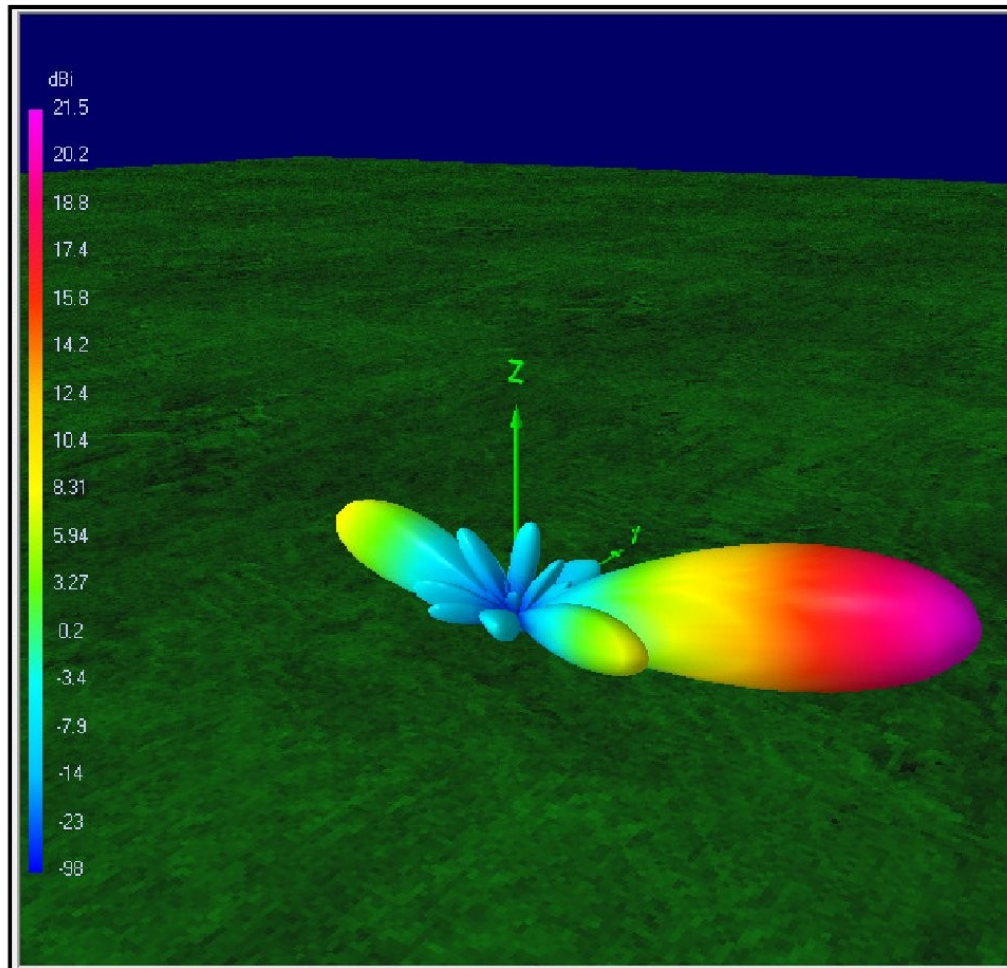


# TARA40: 8-Yagi Array

- Replace single 3-element Yagi with eight 5-element Yagis in phased array.
- Gain of  $\sim 21$  dB above isotropic.
- 8 MW ERP, continuous!



# TARA40: 8-Yagi Array



**21.5 dBi = Gain of 140!**

# Spring 2013: TARA40

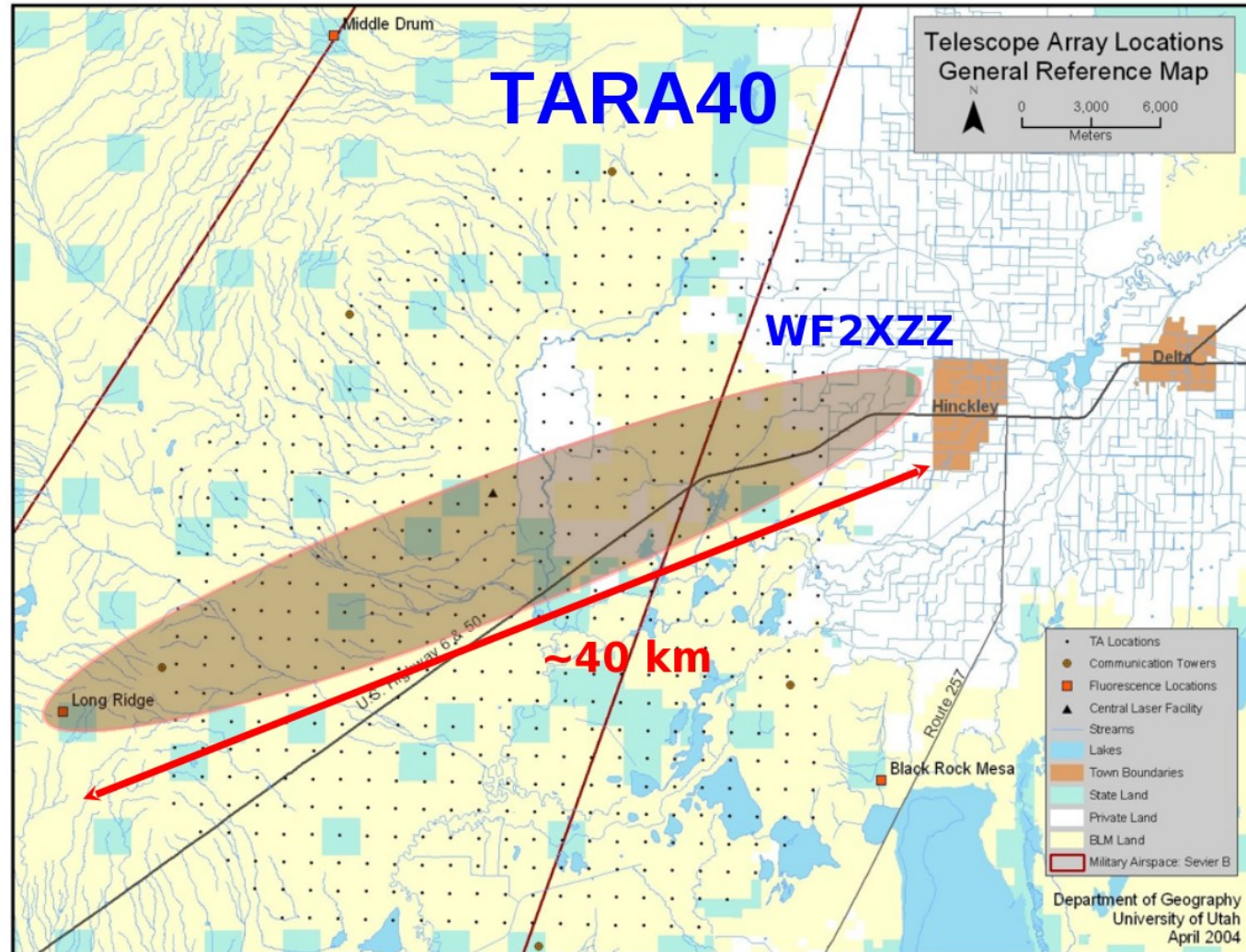
- 40 kW transmitter (x27)
- High-gain antenna (x5)
- Closer to ground array (x2.5)
- Improve RX DAQ





# Spring 2013: TARA40

- 40 kW transmitter (x27)
- High-gain antenna (x5)
- Closer to ground array (x2.5)
- Improve RX DAQ



# Spring 2013: TARA40

- 40 kW transmitter (x27)
- High-gain antenna (x5)
- Closer to ground array (x2.5)
- **Improve RX DAQ**

- Real-time “Matched Filter” trigger algorithm, search for  $\sim 1 \text{ MHz}/\mu\text{sec}$  “chirps”
- Analysis (for now) focused on
  - Verifying trigger
  - Understanding noise triggers
- Future: Hough transforms...



## NI PXIe / FlexRIO

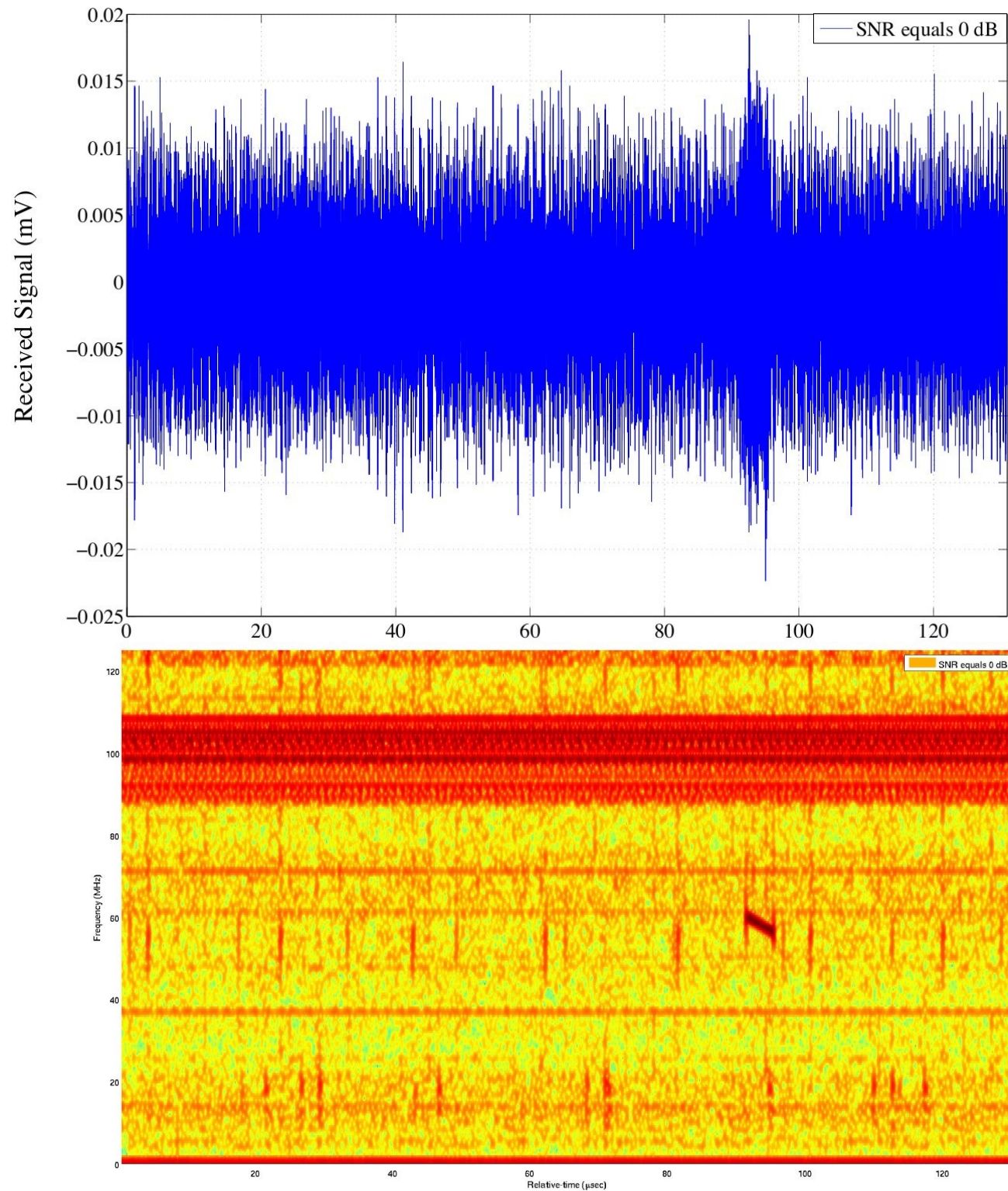
- **250 MHz ADC**
- complete control of FPGA
- **large FPGA**
- **256 MB DRAM (data buffering)**
- computer and FPGA in same chassis
- 400 MHz clock



# RX DAQ Upgrades

- Top: FADC trace of artificially generated chirp, which triggered FlexRIO
- Bottom: Spectrogram

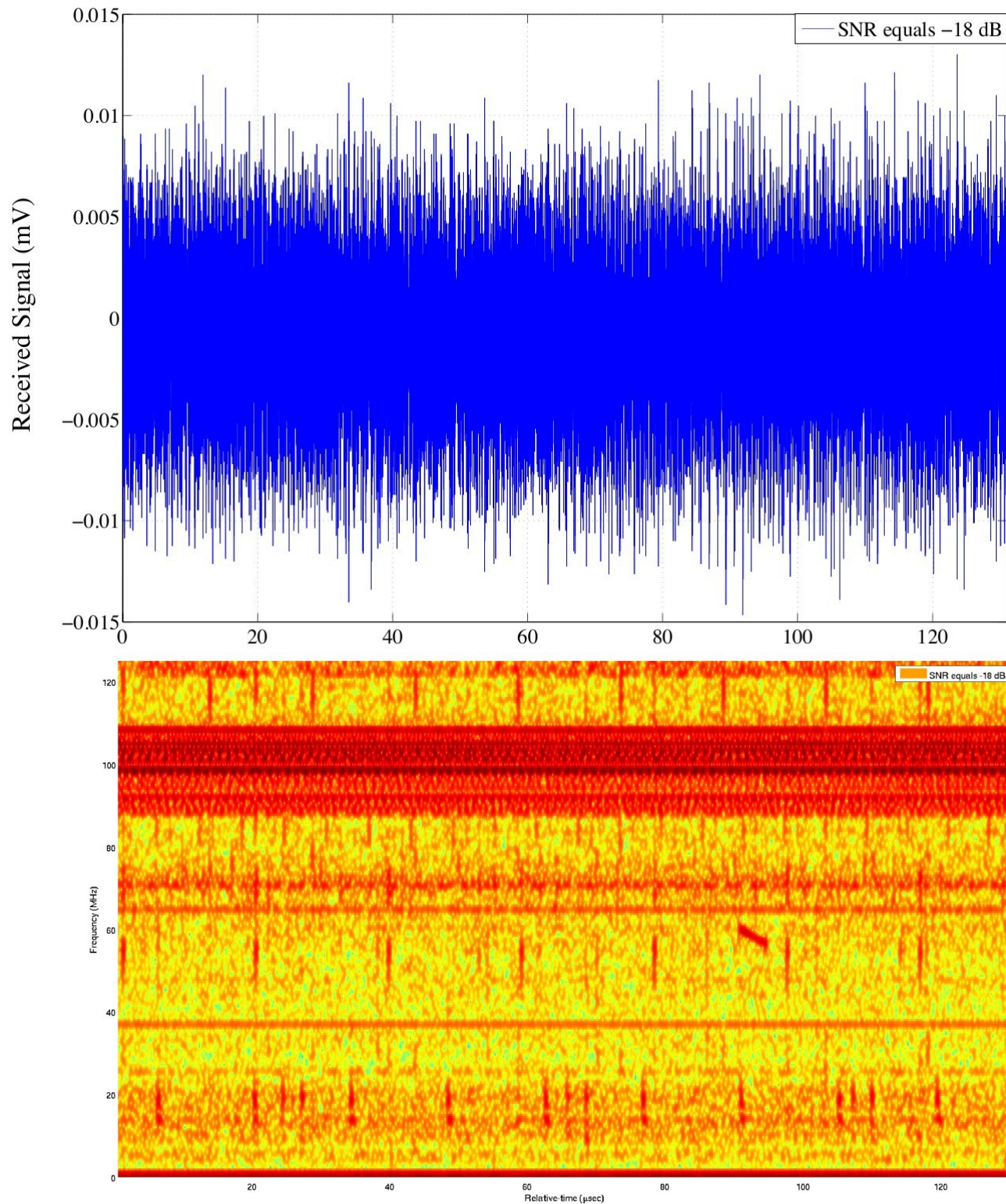
**0 dB SNR**



# RX DAQ Upgrades

- Top: FADC trace of artificially generated chirp, which triggered FlexRIO
- Bottom: Spectrogram

**-18 dB SNR**



# TARA1.5 → TARA40 Enhancements

<b>TX Power</b>	<b>27</b>
<b>TX Antenna Gain</b>	<b>5</b>
<b>TX Location</b>	<b>2.5</b>
<b>TX Total</b>	<b>337</b>
<b>RX Trigger SNR</b>	<b>~100</b>



# Summary

- TARA is the ideal testbed for the bistatic radar idea:
  - Single transmitter under our control
  - Radio-quiet environment
  - Co-located with state-of-the-art “conventional” cosmic ray observatory
- TARA1.5 has been operating for ~1 year.
- TARA40 upgrades underway spring 2013 will *substantially* increase sensitivity.