### ALP-Assisted Electroweak Phase Transition

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# A Quick View

Probing ALP in various experiments

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> Introduction: challenges of electroweak baryogenesis
> Axion-like particle assisting EWPT
> Experimental probes: EDM, beam-dump, collider

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### Baryogenesis problem

Solution Matter > Antimatter  $n_B \sim 9 * 10^{-11}$ 

> Why?



Planck 2018b

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#### Necessary condition for baryogenesis

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> *B* violation: sphaleron process
> *C* and *CP* violation: model-dependent
> Out of equilibrium: ?

**Electroweak** phase transition?

[Sakharov, 1967]

### Sphaleron



B and L number are anomalous!

$$\partial_{\mu}J^{B}_{\mu} = \partial_{\mu}J^{L}_{\mu} = \frac{3g^{2}}{32\pi^{2}}W\tilde{W}$$
  
 
$$B - L \text{ is conserved.}$$

[G. 't Hooft, Phys. Rev. Lett. 37, 8 (1976)]

#### Electroweak phase transition

- > Higgs potential
- > T = 0:  $V = -\frac{1}{2}\mu^2 h^2 + \frac{1}{4}\lambda h^4$ , symmetry breaking
- T > 0: finite temperature correction.
- Very high T: "symmetry restoration"
- T<sub>c</sub>: h = 0 and h = v degenerate, "critical temperature"



Figure from: J Cline: hep-ph/0609145

#### 1st order vs 2nd order

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> 1st: well-defined barrier, bubble nucleation

> 2nd: smooth crossover

#### Electroweak baryogenesis



[V. A. Kuzmin, V. A. Rubakov and M. E.
Shaposhnikov, Phys. Lett. B 155, 36 (1985)]
[M. E. Shaposhnikov, JETP Lett. 44, 465 (1986), Nucl. Phys. B 287, 757 (1987)]

#### Phase transition strength

- Finite-T: thermal correction to the effective potential V.
- Boson contribution:  $\frac{T^4}{2\pi^2} n_B (\frac{\pi^2}{12} (\frac{m}{T})^2 - \frac{\pi}{6} (\frac{m}{T})^3 + \dots), n_B: \text{d.o.f.}$
- Fermion contribution:  $\frac{T^4}{2\pi^2}n_F(\frac{\pi^2}{24}(\frac{m}{T})^2 + ...),$  $n_F: \text{ d.o.f.}$

> Total: 
$$V = DT^2h^2 - \frac{1}{2}\mu^2h^2 - ETh^3 + \frac{1}{4}\lambda h^4$$
,

cubic from bosons

 $\frac{v(T_c)}{T_c} = \frac{2E}{\lambda} \text{ (strongly 1st order)}$ 



Only for intuition. Needs full form to compute!

### Challenges on EWBG

> Is EWPT strongly 1st order? How to assist?
> CP-violating effect: EDM constraint!

### Challenge 1: PT strength

Lattice result: EWPT crossover for  $m_H \gtrsim 70$  GeV. How to enhance?

General way: introduce singlet S(1,1,0). Self interaction:  $\mathscr{L} \supset \frac{1}{2}\mu_S^2 S^2 + \frac{1}{4}\lambda_S S^4$   $Z_2$  symmetric:  $S^2 H^2$ , enhance E from thermal effect.  $Z_2$  breaking: additional  $S^3$ , SHH, etc. Enhance from tree-level. [K. Jansen, hep-lat/9509018], [K. Kajantie et al, hep-lat/9510020] [K. Rummukainen, hep-lat/9608079], [K. Kajantie et al, hep-ph/9605288.

[M. Gurtler et al, hep-lat/9704013], [F. Csikor et al, hep-ph/9809291]
[M. Laine and K. Rummukainen, hep-ph/9804255, hep-lat/9804019]
[K. Rummukainen et al, hep-lat/9805013], [Z.Fodor, hep-lat/9909162]

#### Intensitve scalar search



### Extra hierarchy problem





Higgs mass correction

$$\delta m_H^2 = -\frac{y_f^2}{8\pi^2}\Lambda_{\rm UV}^2$$

Huge quantum corrections!

Scalar mass correction:

Leading: 
$$\frac{1}{4}\lambda h^2 S^2$$

$$\delta m_S^2 = \frac{\lambda_{hs}}{16\pi^2} \Lambda_{\rm UV}^2$$

Extra hierarchy problems

## Challenge 2: EDM

CP-violating required!

In general

$$d_f \simeq \left(\frac{m_f}{M^2}\right) \frac{\alpha_f}{4\pi} \sin \delta_{CP} \ e \ \mathrm{cm}$$

 $\alpha_f$ : coupling that produces EDM for fermion f

[D. Morrissey, M. Ramsey-Musolf, 1206.2942]

### Example: local EWBG

$$O = \frac{\alpha_2}{M^2} H^2 W \tilde{W} = \frac{H^2}{M^2} \dot{n}_B \equiv \theta \dot{n}_B, \text{ CP violating.}$$

Integratal by part:  $\partial_{\mu}J^{B}_{\mu} = \partial_{\mu}J^{L}_{\mu} = \frac{3g^{2}}{32\pi^{2}}W\tilde{W}$   $O = \dot{\theta}n_{B}$ This induces a minimum value:  $n_{B0} = \dot{\theta}T^{2}$ 

$$\dot{n}_B = -3\Gamma_{\rm sph}(n_B - \dot{\theta}T^2).$$

Start from  $n_B = 0$ , difficult to achieve  $n_{B0}$ , ignore it.  $n_B = 3\Gamma_{\rm sph}T^2\Delta\theta$ .

> [Dine, et al, 1991] [Dine, hep-ph/9206220]

#### Example: local EWBG



 $\Delta \theta$ : field value shift  $\delta h^2/M^2$  during PT before sphaleron rate starts to drop  $\delta h \simeq 0.3T_c$ 

Small *bh* Small  $n_B$ Require small M EDM! M. D'Onofrio et al, 1404.3565

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### Facing the challenge

What do we need?

> (Naturally) light (avoid extra hierarchy problem)
> Weakly interacting
> Large UV scale (for EDM), large field value shift

The answer is...

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### The axion-like particle

 $P = (f + \Phi)^2 \exp(iS/f)$  S: angular mode, axion or ALP

Couple with Higgs:

$$V = -\frac{1}{2}(\mu_H^2 - Af\cos\delta)h^2 + \frac{1}{4}\lambda h^4$$
$$+\mu_S^2 f^2 \left(1 - \cos(\frac{S}{f})\right) - \frac{1}{2}Af\left(h^2 - 2\nu^2\right)\cos(\frac{S}{f} - \delta).$$
Convention:  $\langle h \rangle = \sqrt{2}\nu \ \delta \in [0, \pi/2]$ 

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[Jeong et al, 2018] [K.Harigaya and IRW, 2309.00587

#### Vacuum structure



### Mass spectrum

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At 
$$\langle h \rangle = \sqrt{2}v, \langle S \rangle = 0$$
  
 $m_{hh}^2 = 6\lambda v^2 - \mu_H^2$   
 $m_{SS}^2 = \mu_S^2$   
 $m_{hS}^2 = -Ah \sin \delta$   
mixing angle!

$$A = \frac{1}{2\sqrt{2}\nu} (m_h^2 - m_S^2) \sin(2\theta) \csc(\delta),$$
  

$$\lambda = \frac{1}{8\nu^2} (m_h^2 + m_S^2 + (m_h^2 - m_S^2) \cos(2\theta)),$$
  

$$\mu_H^2 = \frac{1}{4} (m_h^2 + m_S^2 + (m_h^2 - m_S^2) \cos(2\theta)),$$
  

$$\mu_S^2 = \frac{1}{2} (m_h^2 + m_S^2 - (m_h^2 - m_S^2) \cos(2\theta)).$$

[K.Harigaya and IRW, 2309.00587]

### Enhancing the PT Leading order effect

 $\mu_S^2 f^2 \cos(\frac{S}{f}) \to \frac{1}{2} \mu_S^2 S^2$  Mass term

 $-\frac{1}{2}Afh^2\cos(\frac{S}{f}-\delta) \rightarrow -\frac{1}{2}Ah^2S\sin\delta + \frac{1}{4}\frac{A}{f}h^2S^2\cos\delta$ 

Z<sub>2</sub> breaking

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[K.Harigaya and IRW, 2309.00587]

### Enhancing the PT Leading order effect

$$S = \frac{Af(h^2 - 2v^2)\sin\delta}{2f\mu_s^2 + A(h^2 - 2v^2)\cos\delta} \quad \text{along "valley"}$$

$$\approx \frac{Afv^2\sin\delta}{Av^2\cos\delta - f\mu_s^2} + \frac{Af^2\mu_s^2\sin\delta}{2(Av^2\cos\delta - f\mu_s^2)^2}h^2 + \frac{A^2f^2\mu_s^2\sin\delta\cos\delta}{4(Av^2\cos\delta - f\mu_s^2)^3}h^4 + O(h^6).$$
Quartic:  $\frac{1}{4}\lambda h^4 \rightarrow \frac{1}{4}(\lambda - \frac{A^2f^3\mu_s^4\sin^2\delta}{2(\mu_s^2f - Av^2\cos\delta)^3})h^4$ 
Negative for parameter
Additional:  $O(h^6)$ 
Positive, guarantee vev stability

[K.Harigaya and IRW, 2309.00587]

### Thermal potential

 $V_{\rm eff} = V_0 + V_{\rm CW} + V_{\rm FT}$ 

$$V_{\rm CW} = \frac{1}{64\pi^2} \left( \sum_B n_B \left( \log \left( \frac{m_B^2(h, S)}{Q^2} \right) - c_B \right) - \sum_F n_F \left( \log \left( \frac{m_F^2(h, S)}{Q^2} \right) - c_F \right) \right)$$

$$V_{\rm FT} = \frac{T^4}{2\pi^2} \left( \sum_B n_B J_B \left( \frac{m_B^2(h, S)}{T^2} \right) + \sum_F n_F J_F \left( \frac{m_F^2(h, S)}{T^2} \right) \right)$$
$$J_{B,F}(x^2) = \pm \int_0^\infty dy \ y^2 \log \left( 1 \mp \exp\left( -\sqrt{y^2 + x^2} \right) \right).$$

Resummation: all boson masses replaced by thermal mass:  $m^2 \rightarrow m^2 + \Pi, \Pi \sim g^2 T^2$ 

### Thermal phase transition

High-*T* expansion:

Again, only for intuition. Needs full form to compute!

$$\begin{split} V_{\text{highT}} &= -\frac{1}{2}(\mu_{H}^{2} - Af\cos\delta)h^{2} + D_{\text{SM}}T^{2}h^{2} - E_{\text{SM}}Th^{3} + \frac{1}{4}\lambda h^{4} \\ &- \frac{1}{2}Af\cos\left(\frac{S}{f} - \delta\right)\left(h^{2} - 2v^{2} + \frac{1}{3}T^{2}\right) & \text{One-step PT.} \\ &- f^{2}\mu_{S}^{2}\cos\left(\frac{S}{f}\right). & (0,\langle S\rangle(0,T)) \\ &- f^{2}\mu_{S}^{2}\cos\left(\frac{S}{f}\right). & \gamma(v(T),\langle S\rangle(v(T),T)) \\ &\langle S\rangle(h,T) &= f\arctan\left(\frac{A(3(h^{2} - 2v^{2}) + T^{2})\sin(\delta)}{6f\mu_{S}^{2} + A(3(h^{2} - 2v^{2}) + T^{2})\cos(\delta)}\right) & \text{Parametrize } f \text{ in } \\ &\text{terms of } S \text{ shift } \\ &\text{when } h: 0 \rightarrow \sqrt{2}v \text{:} \\ &\text{Large } f: \langle S\rangle(h,T) \simeq \frac{1}{2}\frac{A}{\mu_{S}^{2}}(h^{2} - 2v^{2} + \frac{1}{3}T^{2})\sin\delta. & f_{c} \equiv \frac{A}{\mu_{S}^{2}}v^{2}\sin\delta, f = cf_{c} \\ &\text{Light } S \text{ has large field value shift.} \end{aligned}$$

#### Bubble nucleation

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 $rac{Prom}{T} = T_c$ : transition allowed.

> Transition rate:  $\Gamma \simeq T^4 \exp(-S_3/T)$  $S_3 \equiv \int r^2 (\frac{1}{2} \left(\frac{d\phi}{dr}\right)^2 + V(\phi)) dr,$ 

decrease as universe cools down. *r*: radius of the 3d space.

> Nucleation temperature  $T_n: \Gamma/H^3 \simeq H$ 

 $\sim$  Condition:  $\frac{S_3}{T} \simeq 140$ .

> Inverse time duration:  $\frac{\beta}{H} \equiv T_n \frac{d}{dT} \left(\frac{S_3}{T}\right)|_{T_n}$ 



#### Bubble nucleation

2d: correct. 1d: ignore S.



$$S_3 = \int dr \frac{1}{2} \left(\frac{dh}{dr}\right)^2 + \frac{1}{2} \left(\frac{dS}{dr}\right)^2 + V$$

Huge kinetic energy

Huge  $\beta/H$ . GW signal hopeless.  $T_n$  not so far from  $T_c$ [K.Harigaya and IRW, 2309.00587





#### Results: $\delta = \pi/5$

Constraints to be explained later



#### Results: $\delta = \pi/20$

Constraints to be explained later

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### Possible signals

- UV model independent: scalar production, Higgs exotic decay, rare meson decay, CMB...
- > UV model dependent: EDM (electron+atomic), heavy fermions, local EWBG...

### Collider Signals

#### Vertices: hSS, hhS, $h^3$ ...

with mixing angle insertion. Signals:

- > Direct production:  $SM \rightarrow S + SM$
- > Higgs exotic decay:  $h \rightarrow SS$



The production channel of the current best bound.

[L3 collaboration, Phys. Lett. B 385 (1996) 454–470.] [K. Harigaya and IRW, 2309.00587]

### Rare meson decay and $\Delta N_{\rm eff}$

> Extra decay channel for *B* and *K*:

>  $B^0 \rightarrow K^0S, B^+ \rightarrow K^+S$ , searched by LHCb at 200 MeV <  $m_S$  < 4 GeV

>  $K^+$  →  $\pi^+S, K^0$  →  $\pi^0S$ , searched by NA62, KLEVER.... for MeV scale.

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MeV scale m<sub>s</sub>: large energy density when neutrino decouples.

> S decays into  $\gamma$ : negative  $\Delta N_{\rm eff}$ 

 ative ΔN<sub>eff</sub>
 [E. Goudzovski et al, 2201.07805]

 [NA62 Collaboration, 2010.07644, 2103.15389]

 [KLEVER Project Collaboration, 1901.03099]

 [M. Ibe et al, 2112.11096], [Planck Collaboration, 1807.06209]

 [LHCb Collaboration, 1508.04094, 1612.07818, 1703.08501]

 [CMB-S4 Collaboration, 1610.02743], [K. Harigaya and IRW, 2309.00587]

#### Local EWBG

Consider  $O \propto \frac{S}{M}W\tilde{W}$ 

In general, *M* is not necessarily *f* Completely depends on UV model.  $n_B \propto \frac{\Delta S}{M}$ 

Enhance:  $\Delta S \gg \delta h$ , and M suppression is linear.  $\langle S \rangle(h,T) \simeq \frac{1}{2} \frac{A}{\mu_S^2} (h^2 - 2v^2 + \frac{1}{3}T^2) \sin \delta$ .

### Electric dipole moment

> Electron:  $d_e/e \simeq \frac{y_e}{(16\pi^2)^2} \frac{A\sin\delta}{Mv}$ .

>  $d_e < = 10^{-29} e \text{ cm}$ 

Atomic: 
$$g_N = \frac{\sin \theta}{3\sqrt{2}} \frac{m_N}{v}, g_f = \frac{m_f}{M}.$$
  
 $g_N g_e \le 10^{-31} \left(\frac{m_S}{eV}\right)^2$ 

Depend on the existence of CPviolating g<sub>f</sub> in UV model.

Bound computed with *M* that produces the

correct  $n_B$ 

[V.A.Dzuba et al, 1805.01234 [ACME Collaboration, 2018]

### Heavy fermions?

In general, UV model of *S* contains heavy fermions. > Model dependent.

Senerally, expect f > 1 TeV to avoid collider bounds.

> Not a very solid bound. Depends on model.



Constraints to be explained later



#### Results: $\delta = \pi/5$

Constraints to be explained later



#### Results: $\delta = \pi/20$

Constraints to be explained later







### Summary

> ALP can assist the EWBG by:
> Enhancing the EWPT strength.
> Enhancing n<sub>B</sub> produced by local EWBG mechanism
> Can be probed by: Electron and atom EDM, collider search, rare meson decay, CMB bound on ΔN<sub>eff</sub>