

Thoughts on Reconstruction with PDS

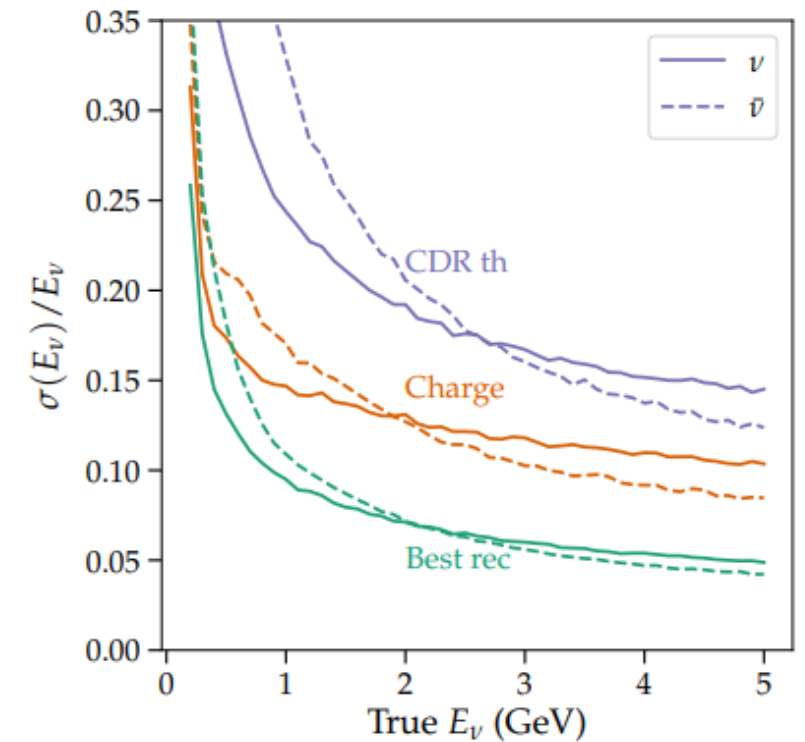
11/30/2023

Chao Zhang

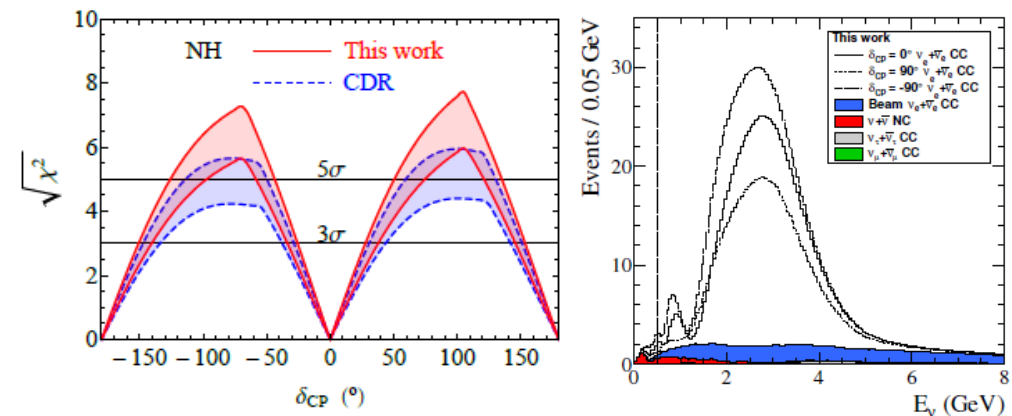
Motivation

Study how much a better PDS can improve **oscillation physics**

- ❑ Understand the limiting factors that affect neutrino energy resolution
 - [previous study by Shirley Li and Friedland](#)
- ❑ Study if **photon detection** can help neutrino energy reconstruction and improve energy resolution
 - Calorimetry from light
 - **Timing from PDS** to improve particle ID
- ❑ Study the CP sensitivity gain if neutrino energy resolution can be improved
 - Two recent talks at the DUNE collaboration meeting ([Marta Torti](#), [Luis Gustavo](#)), but more study is needed
- ❑ Study the requirement for an optimized APEX design



Phys. Rev. D 99, 036009 (2019)



JHEP 09 (2016) 030: $\delta E \sim 8\%$

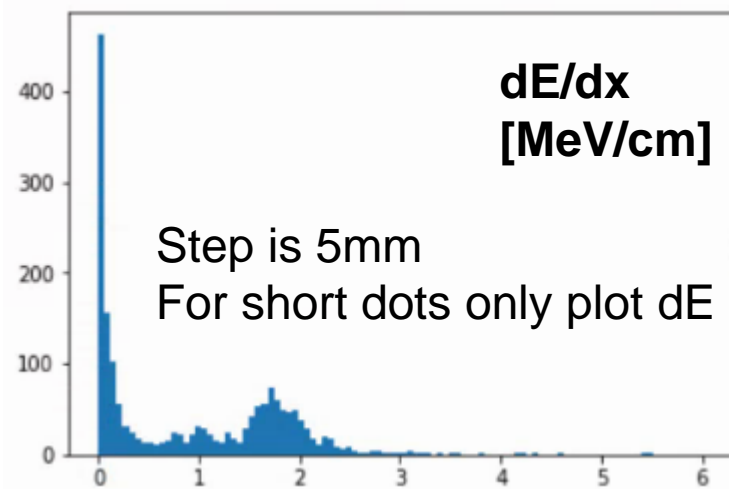
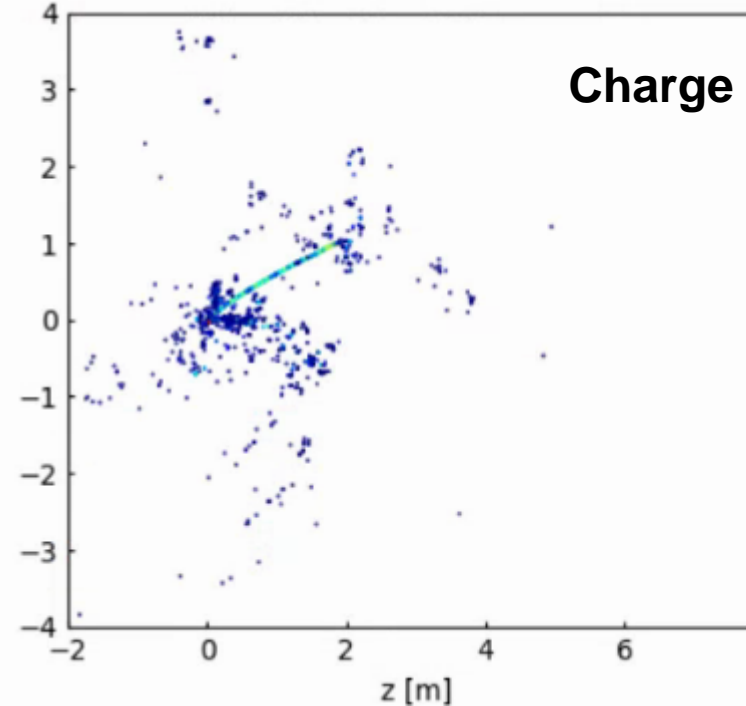
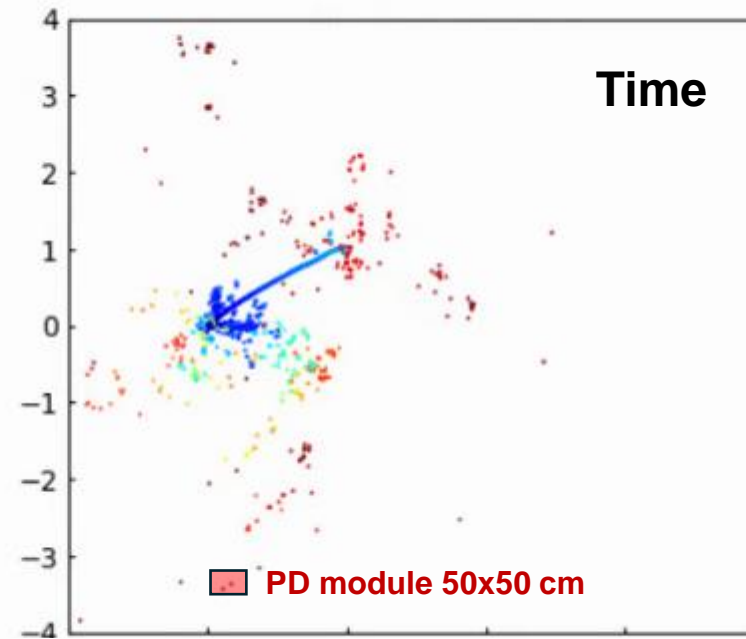
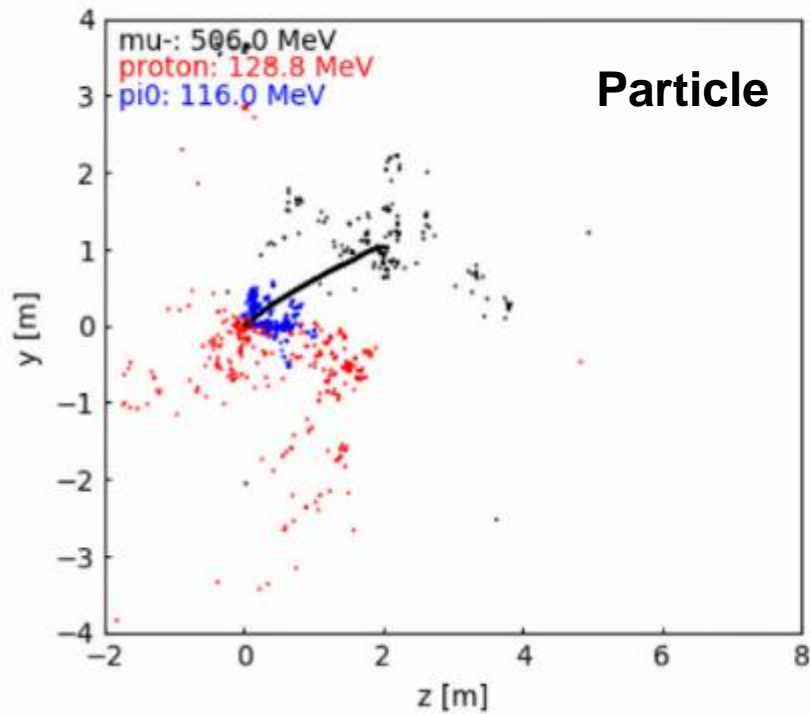
Simulation Setup

- Use Genie and [edep-sim](#) to simulate mono-energetic neutrino CC events from 0.5 – 5 GeV.
 - 1000 events per 0.5 GeV
 - Record all truth-level information including dE/dx along the tracks
 - Geometry: LArBath (a huge LAr Volume so that all events are contained)
- [Simulation setup instruction](#) on FNAL clusters (from Wei Shi)
- Further analysis code: <https://github.com/czczc/PyEdep>
- Next slides will show a few event displays

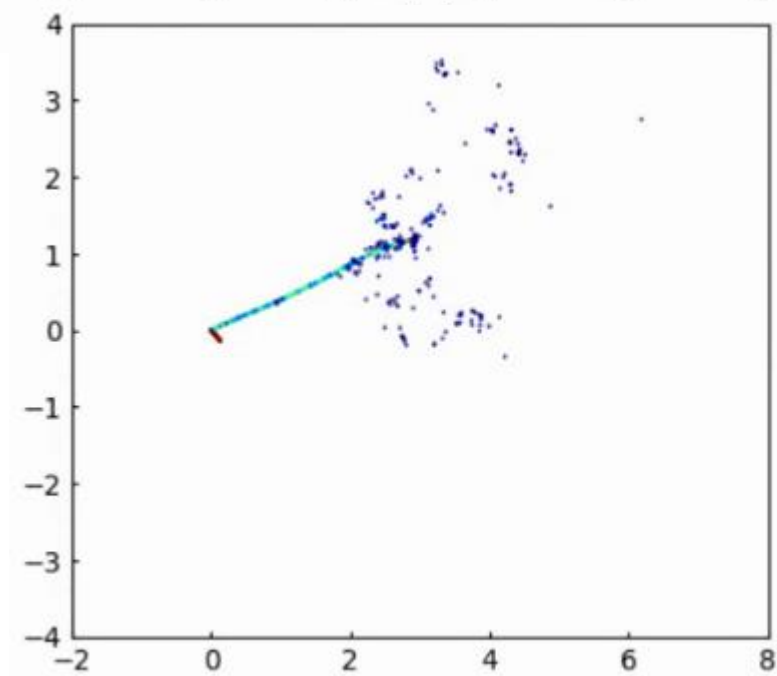
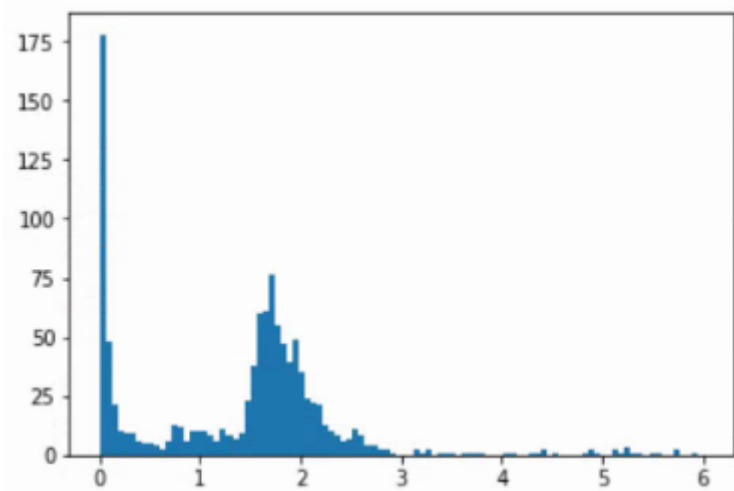
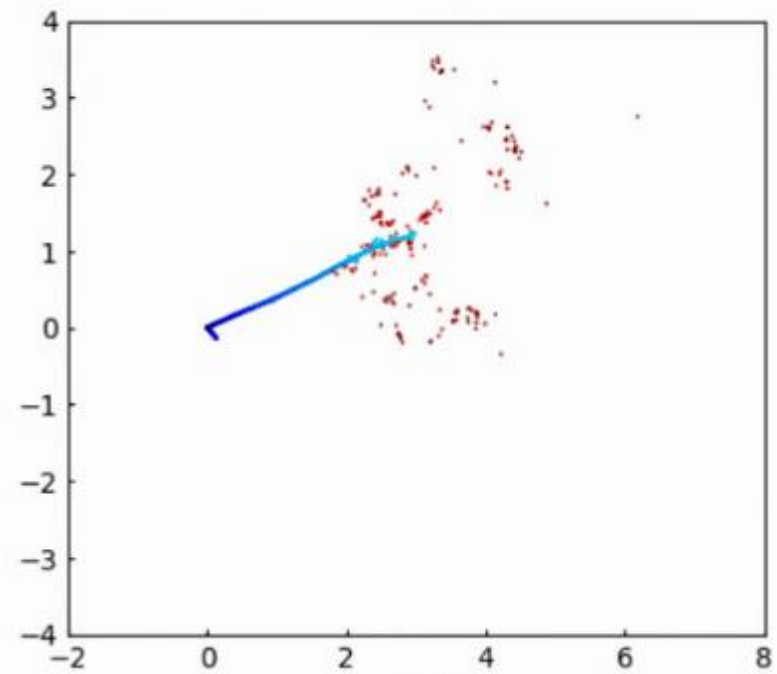
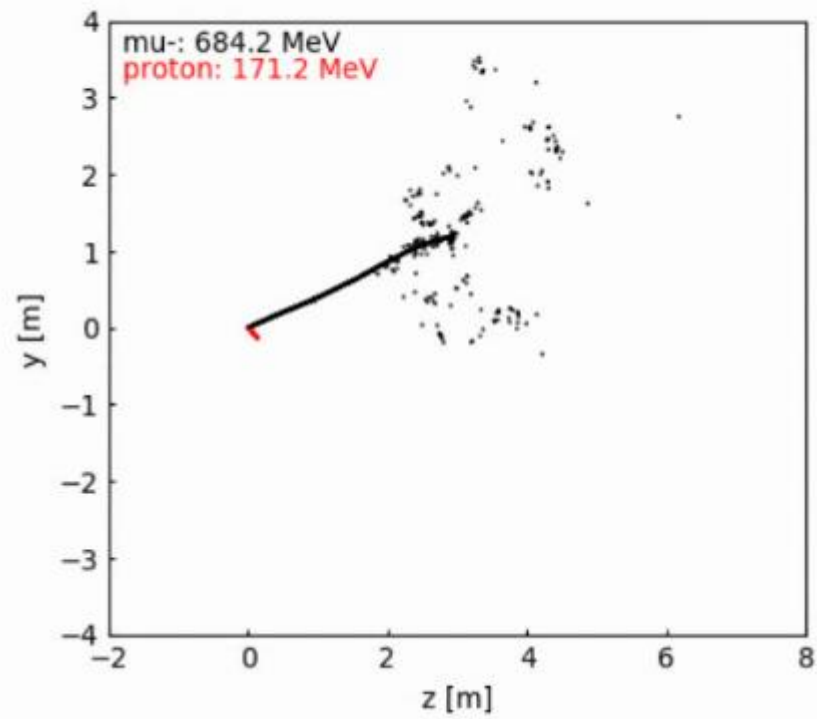
| Branch | unit | description |
|-------------|-----------------|--|
| nu_pdg | int | neutrino pdg code |
| nu_xs | cm ² | interaction cross section |
| nu_proc | int | 10+X: CC; 20+X: NC; (X: 1: QES; 2: RES; 3: DIS; 4: COH, 5: MEC) |
| nu_nucl | int | interact with proton or neutron |
| E_nu | MeV | true neutrino energy |
| E_avail | MeV | available energy from initial state particles, including mass if meson or lepton |
| E_availList | array, MeV | 0: mu/e; 1: proton; 2: neutron; 3: pi+/-; 4: pi0; 5: others |
| E_depoTotal | MeV | total deposited energy |
| E_depoList | array, MeV | similar to E_availList but for deposited energy including all children |

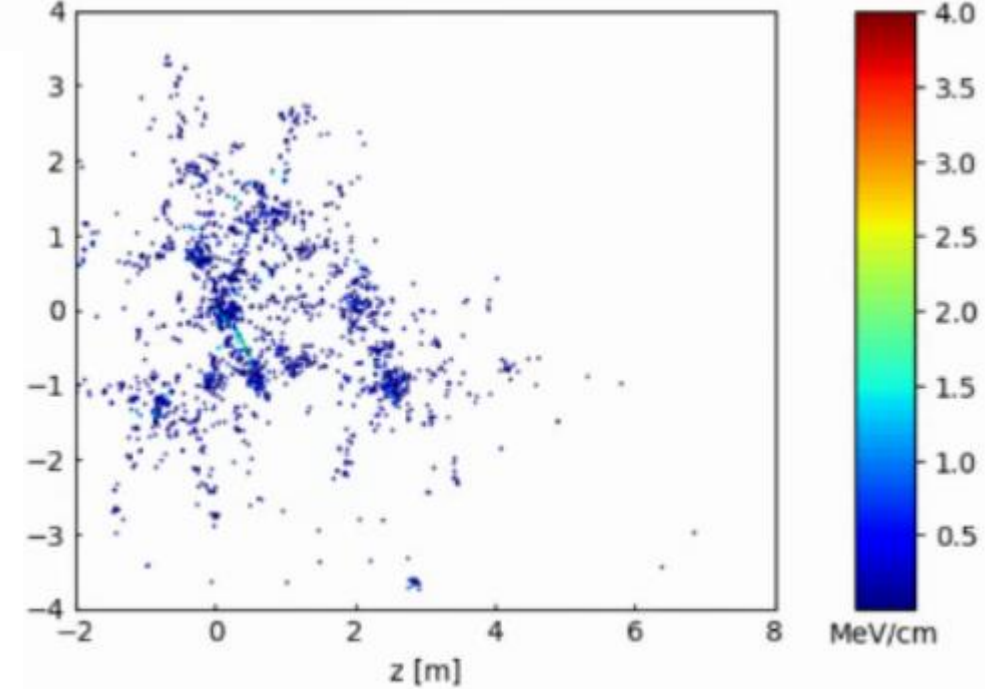
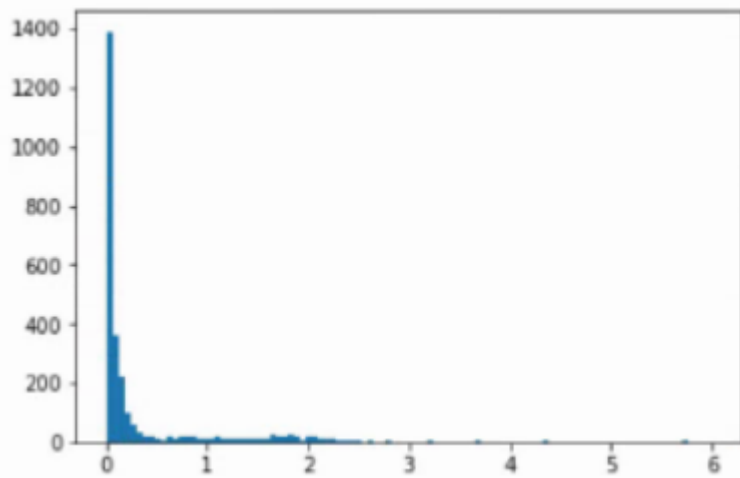
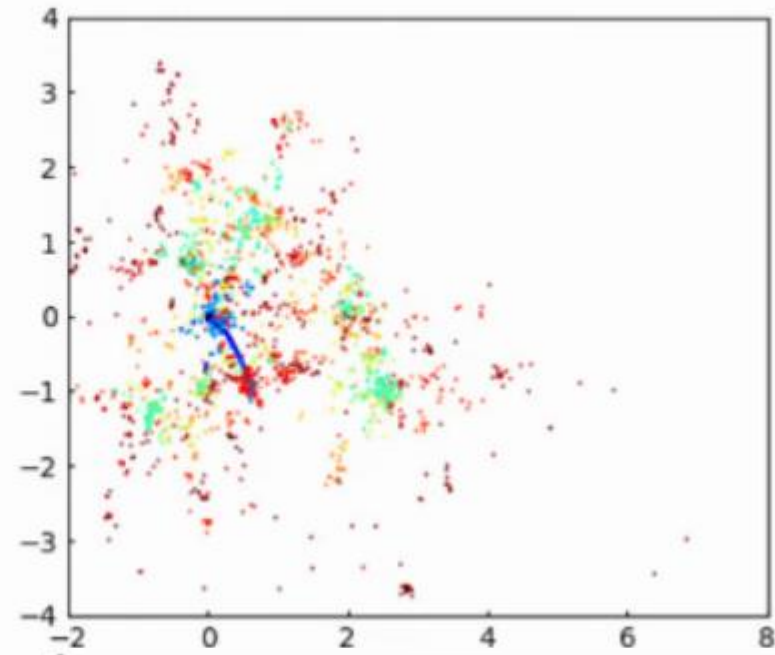
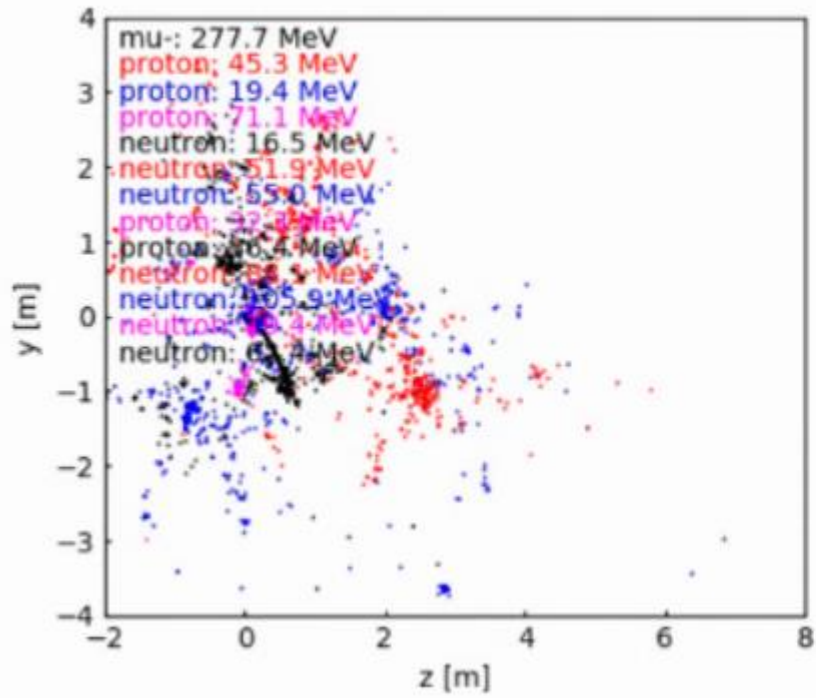
Possible contributors to ν energy resolution

- Generator models
- “Missing energy”
 - Nuclear scatter
 - Muon/pion decay
 - muon capture
 - Detection threshold
- dE/dx -> dQ/dx
- Reconstruction and PID
- Others.

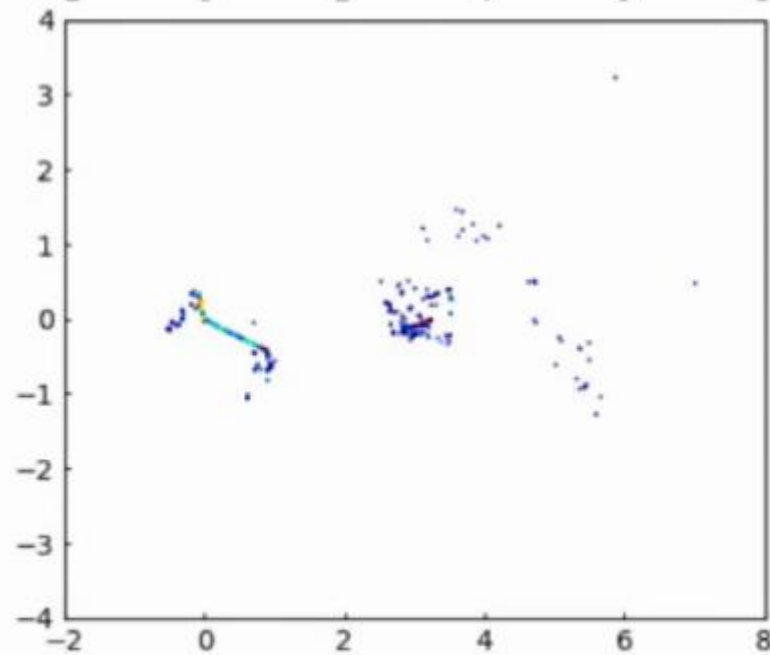
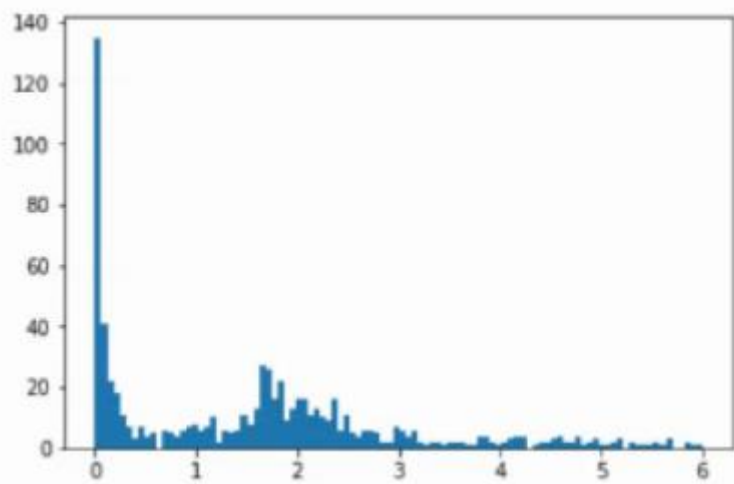
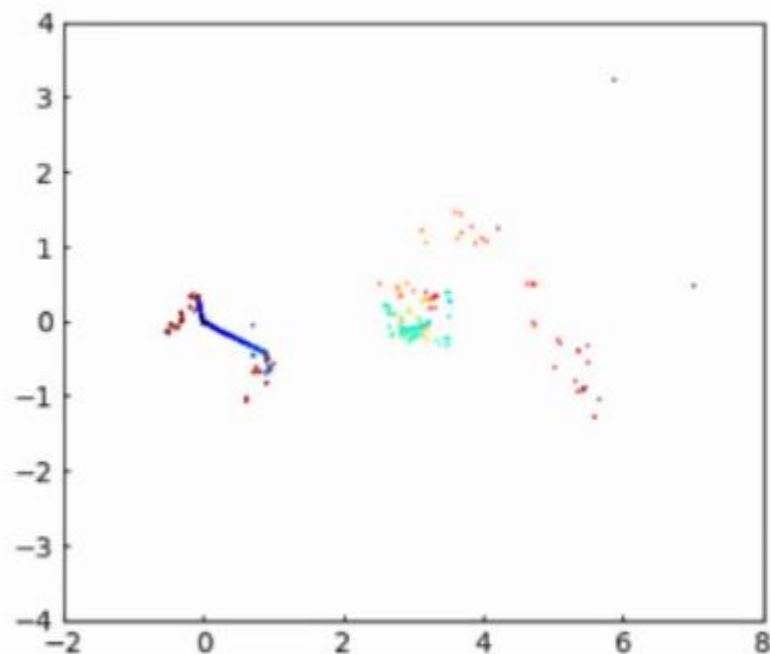
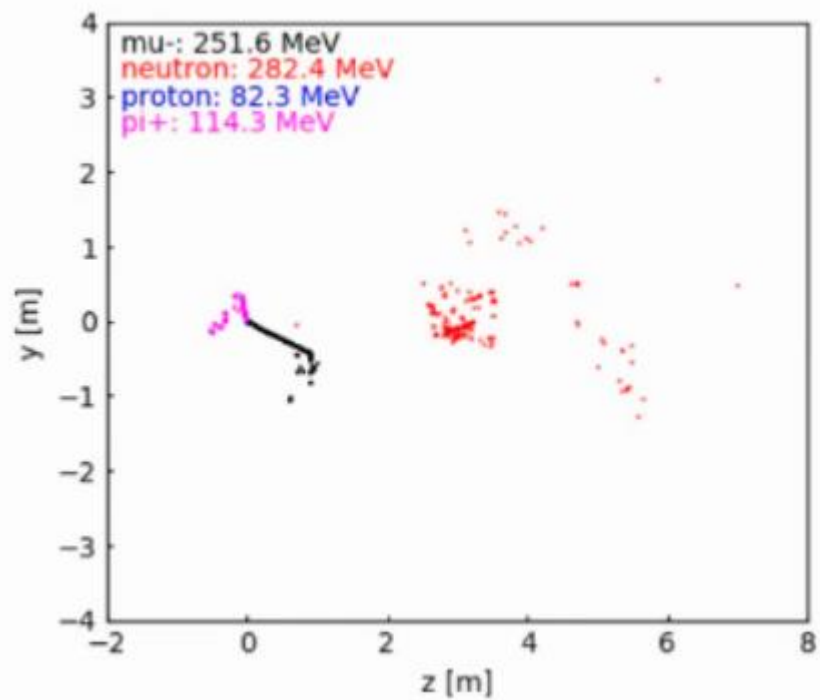


nu:14;tgt:1000180400;N:2112;proc:Weak[CC],QES;





nu:14;tgt:1000180400;N:2112;proc:Weak[CC],RES;res:1;

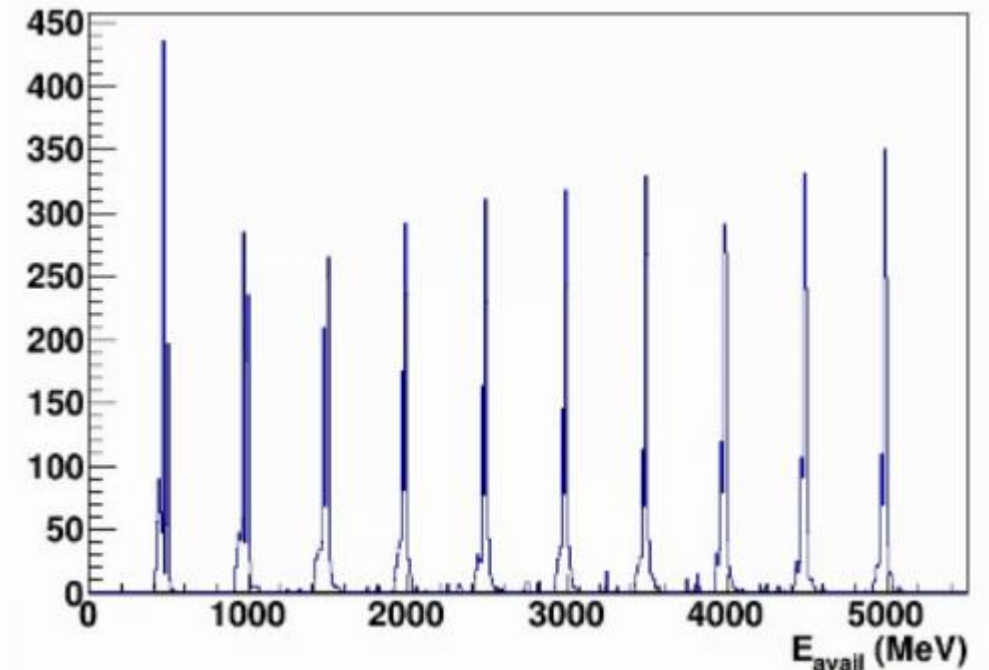
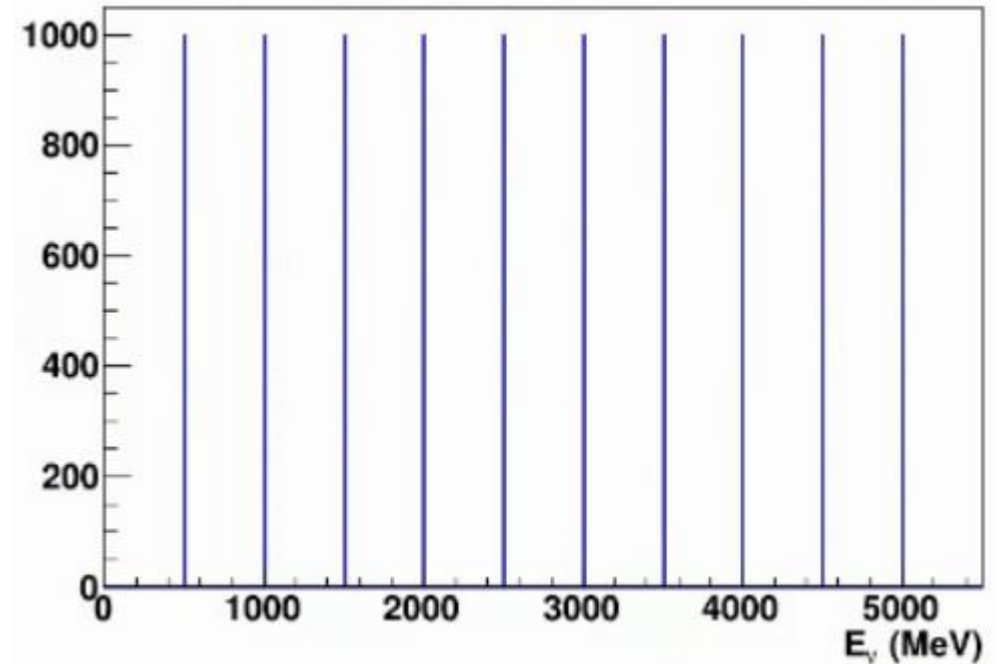


Observations

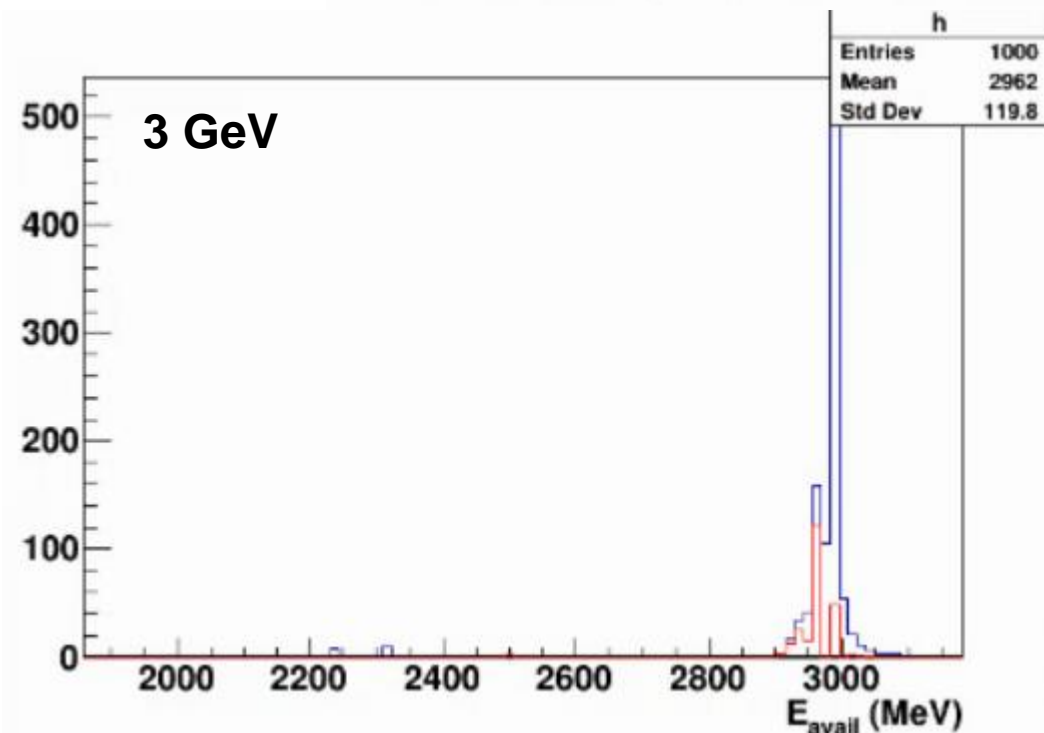
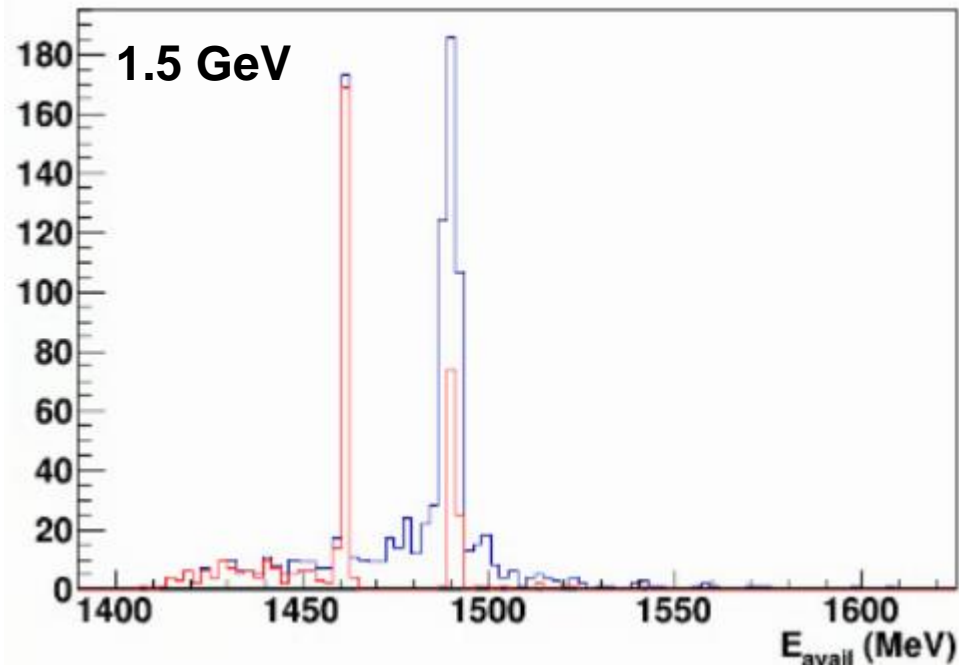
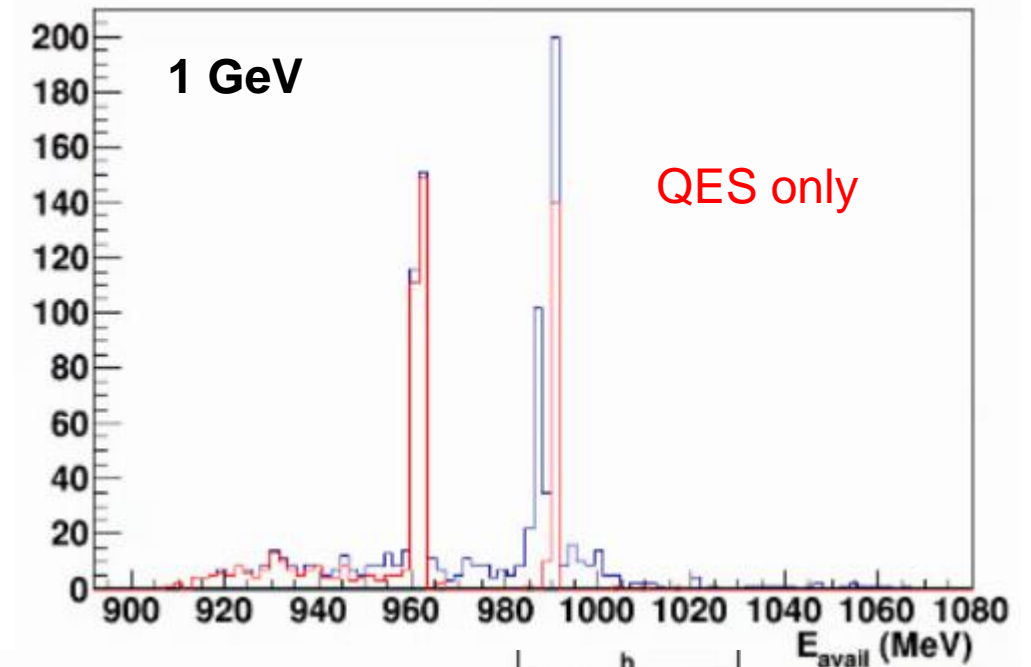
- ❑ Time distribution has more features than charge distribution
 - Tracks near vertex, EM showers (π^0): prompt (<few ns):
 - Fast track propagation: ~4 m per 10 ns
 - Neutron slow down and propagation: tens of ns
 - Particle decay: muon: 2 us; pion 26 ns
 - Muon capture: could emits neutrons that result in delayed components
- ❑ Using time distribution can help neutrino energy reconstruction in the following ways:
 - **Distinguish dots between EM and Neutron interactions**, so that we can apply different calibration constants
 - **Identify decay/capture products** so that we can exclude them from calorimetry (adding back the parent mass)
 - Reconstruct the **direction** of particle tracks for background rejection

Nuclear smearing at Generator level

- ❑ Neutrino energy
 - 1000 events per 0.5 GeV
- ❑ Available energy:
 - KE for: p, n, nuclei
 - Total energy (KE + mass) for leptons and mesons: mu, e-, pi, pi0, etc
 - Smearing can be caused by
 - binding energy
 - intra-nuclear transport
 - Observation: smearing is less than tens of MeV, and is highest at 1.5 GeV where RES process is the highest

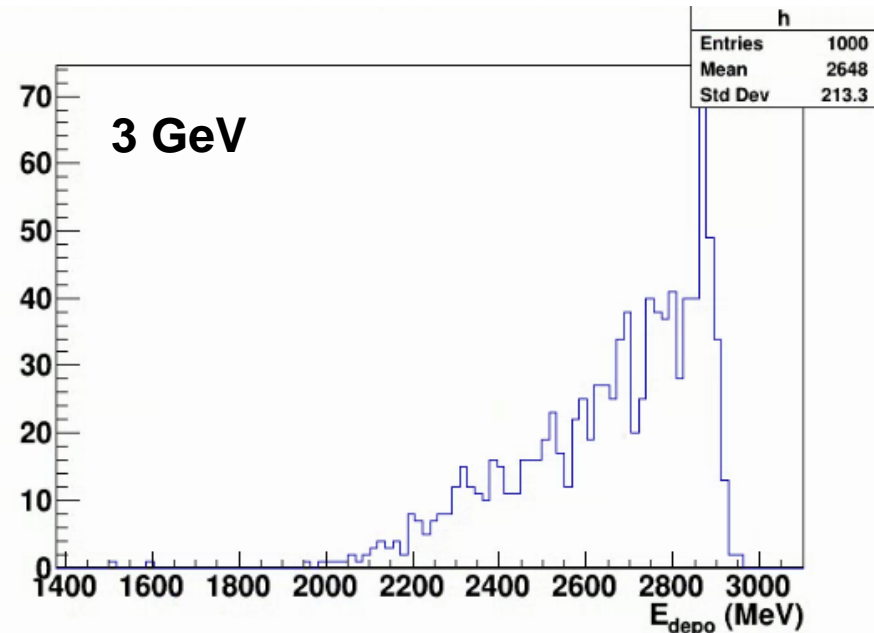
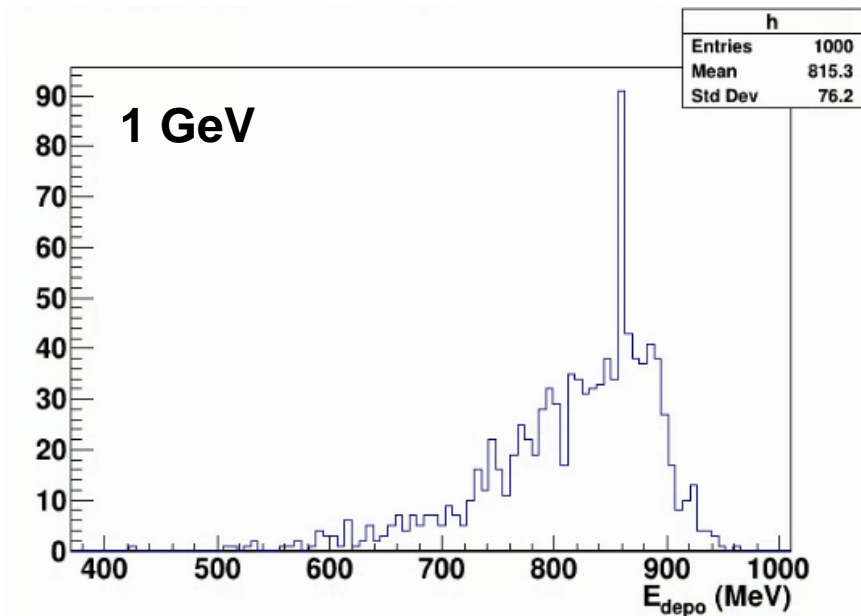
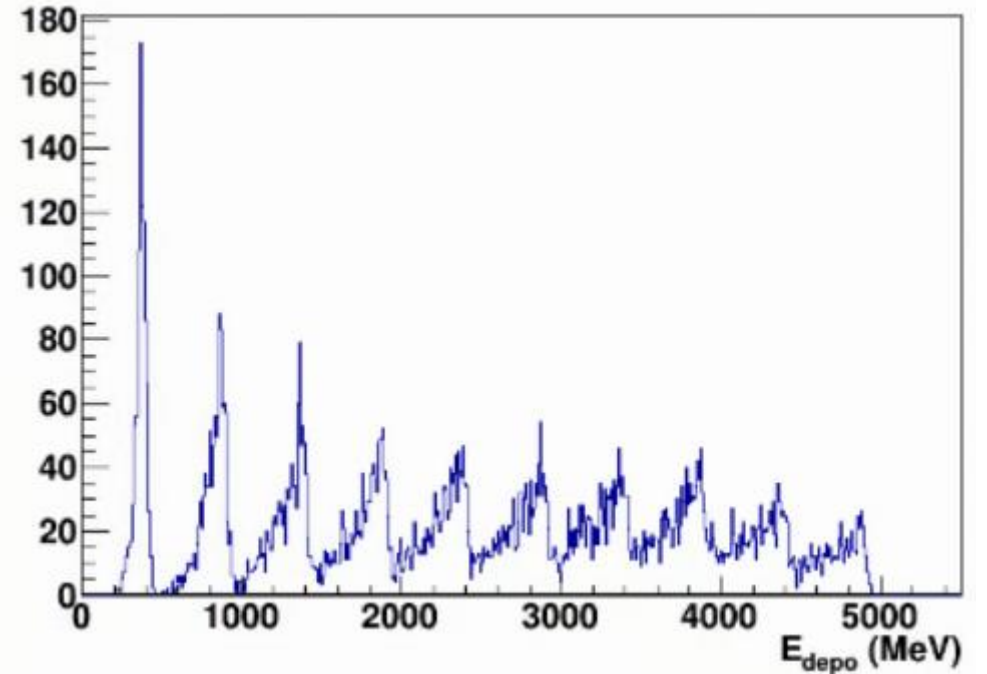


- Zoom in: available energy appears to have 2 peaks, which come from the Quasi Elastic Scattering process
 - reasons unclear yet, may be related to the Genie generator



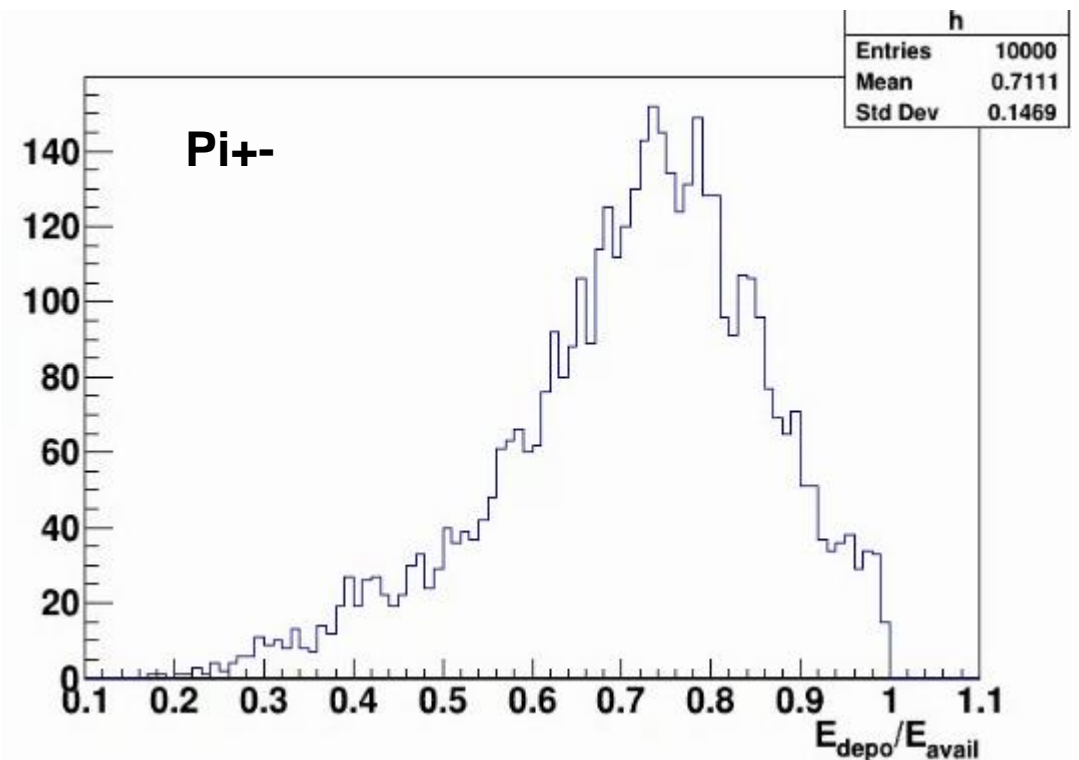
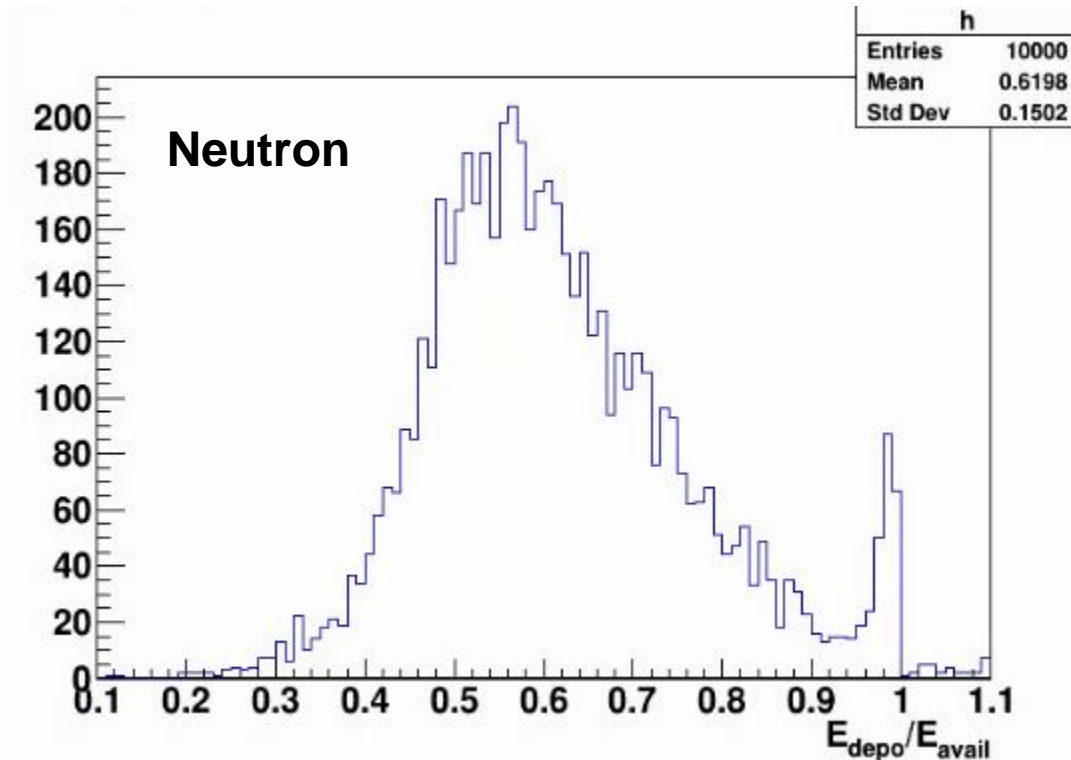
Smearing at energy deposition level

- Deposited energy:
 - All energy deposition including all daughter particles (scatter, decay, de-excitation, etc.)
 - Observation: large smearing ($\sim 10\%$) with a long low-energy tail.

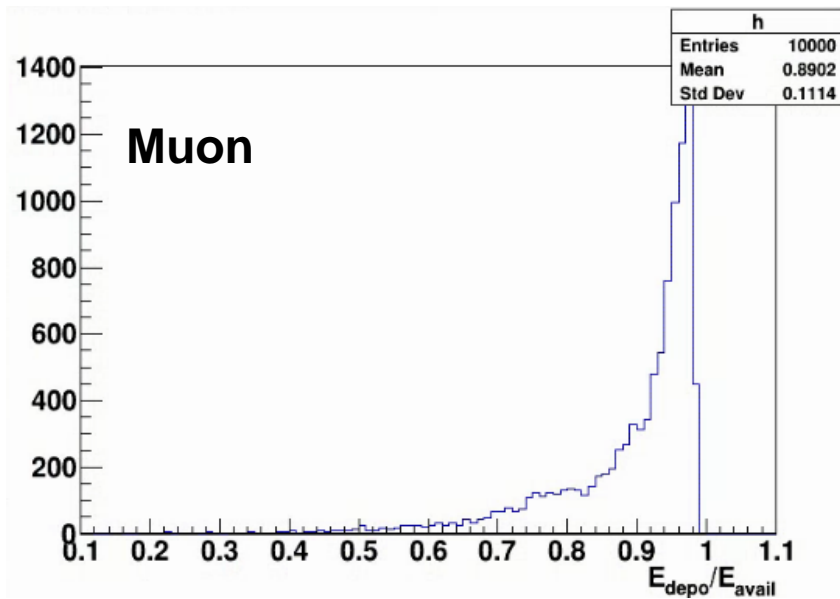
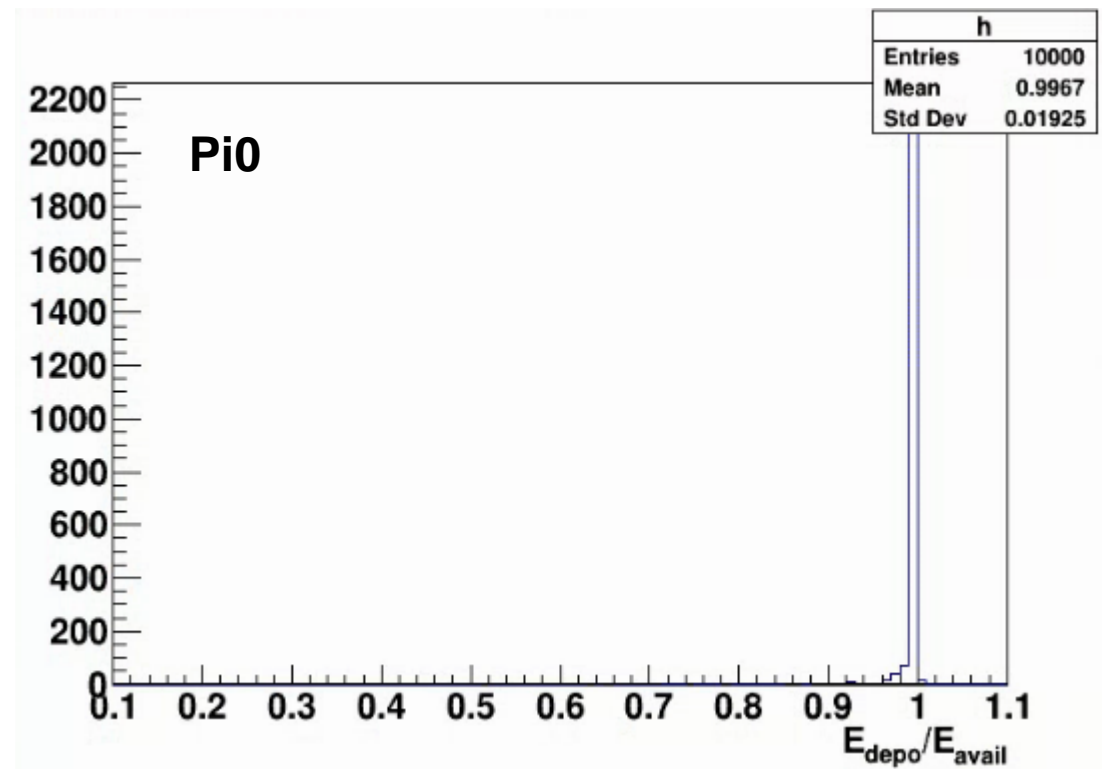
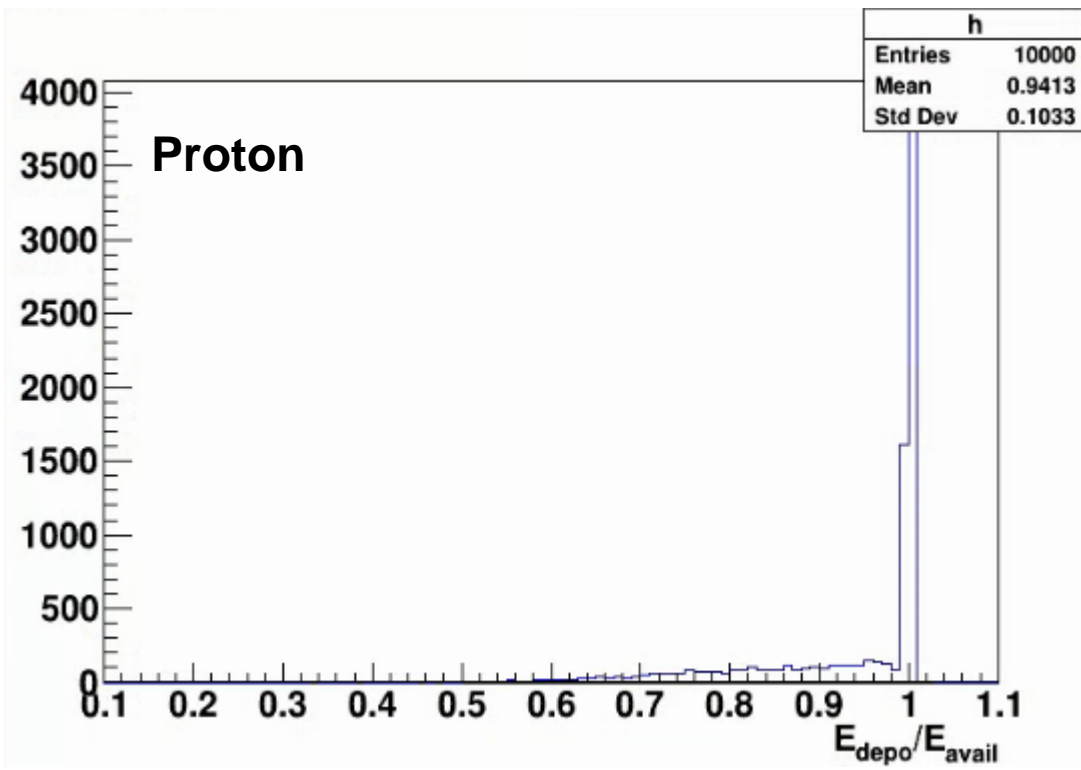


□ Cause of the low energy tail

- Energy lost to nucleus during nuclear scattering: neutrons, pions
- Ratio of $E_{\text{depo}} / E_{\text{avail}}$ appears to be largely independent of initial energy: can be used as a calibration constant if there is PID



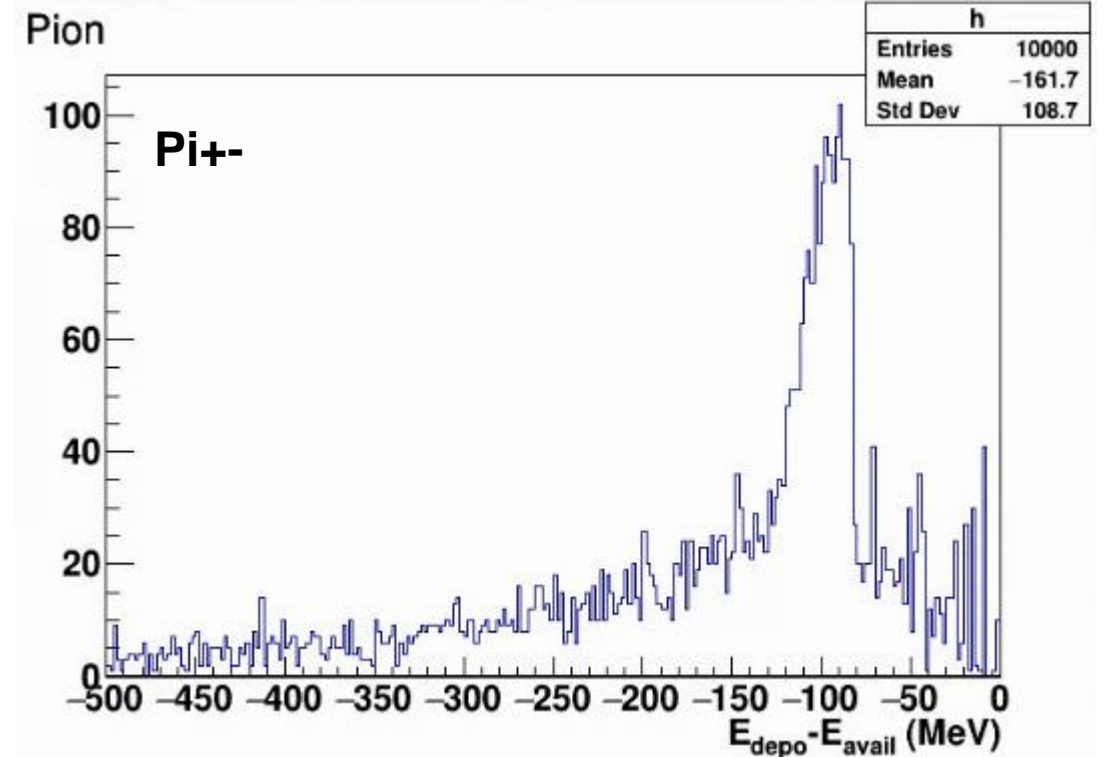
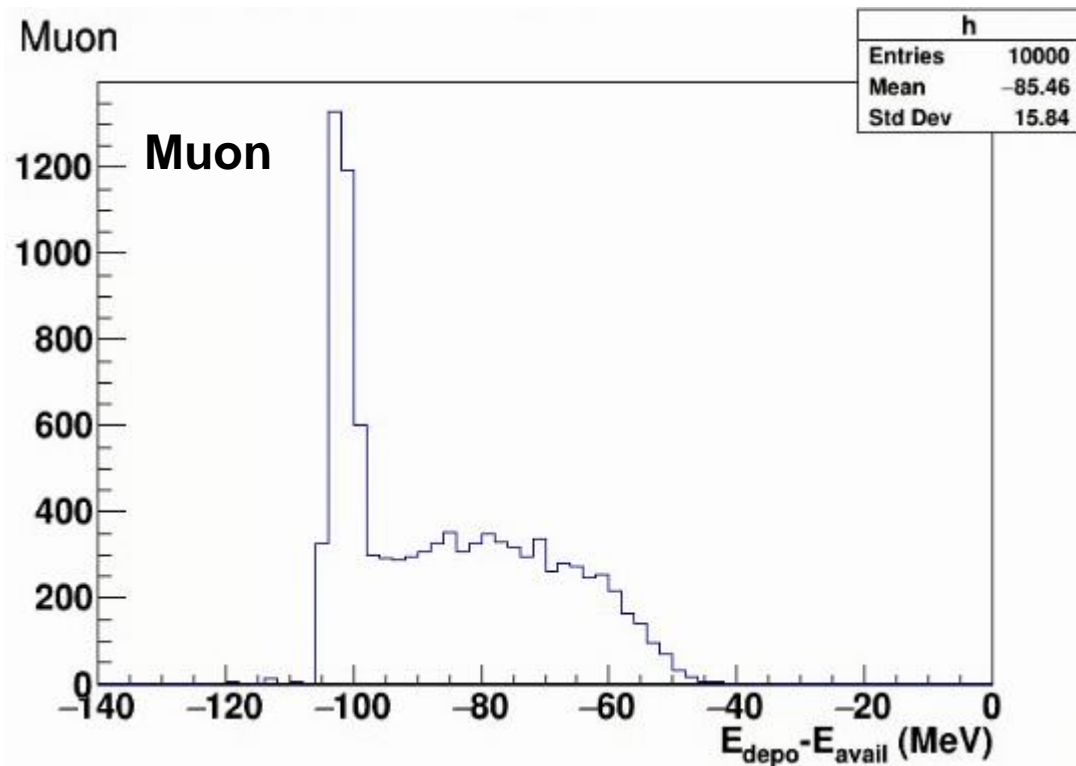
Some of Pion's tail is also from decay/capture



Muon's tail is from decay/capture,
see next slide

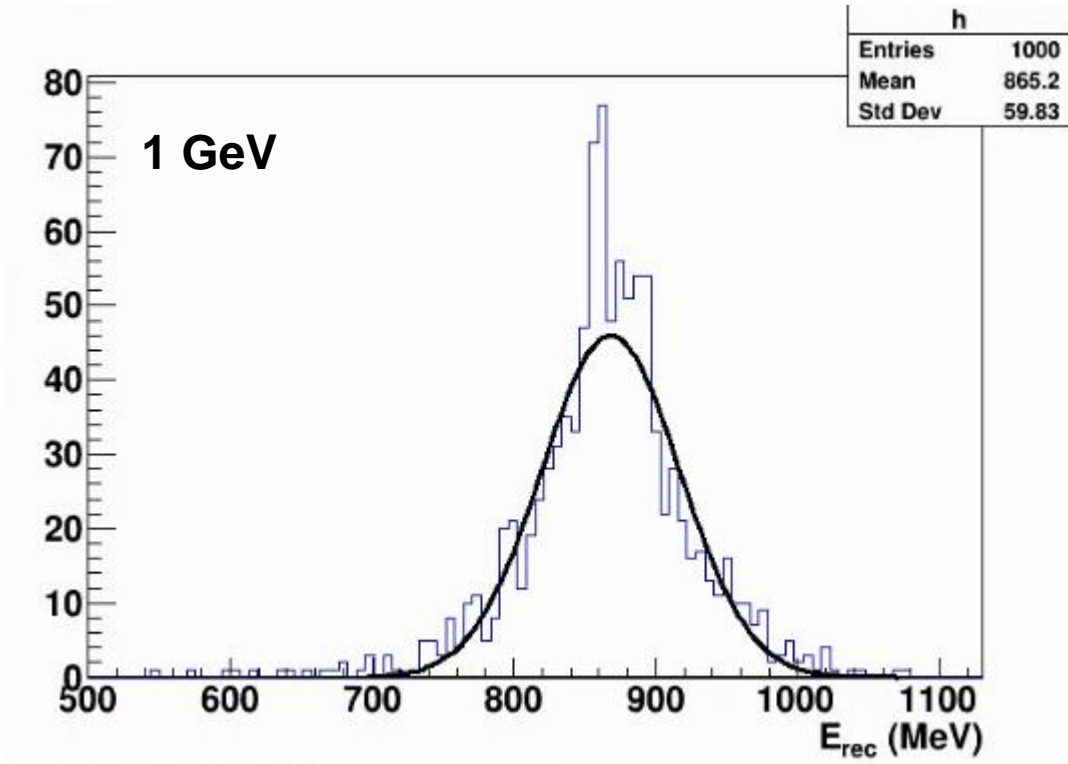
□ Cause of the low energy tail

- Energy lost to neutrinos during decay: muons, pions
- Energy lost from muon- capture

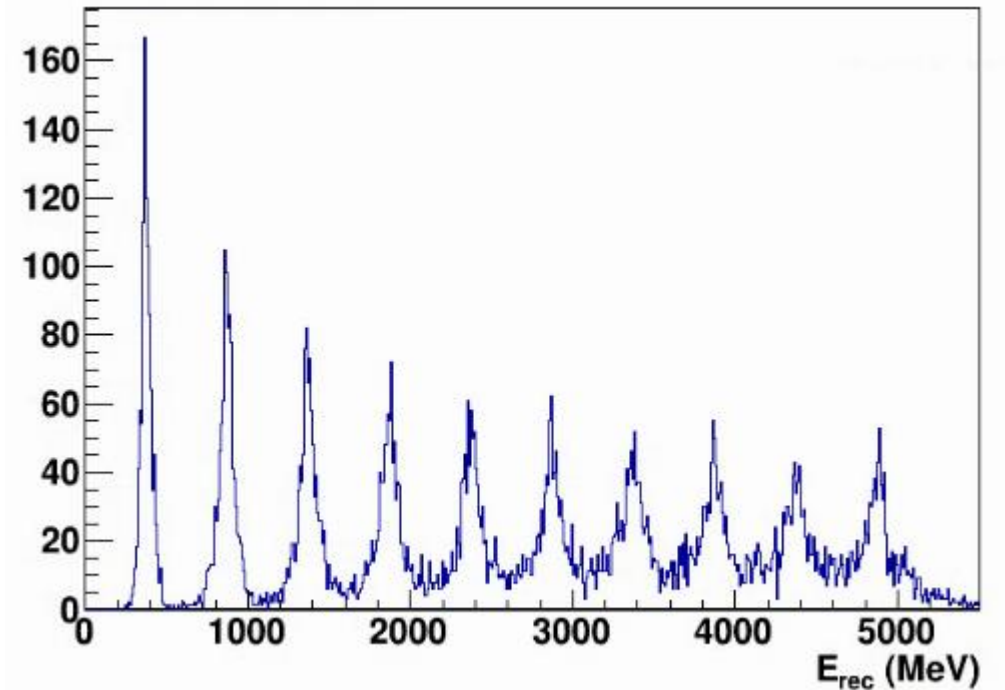
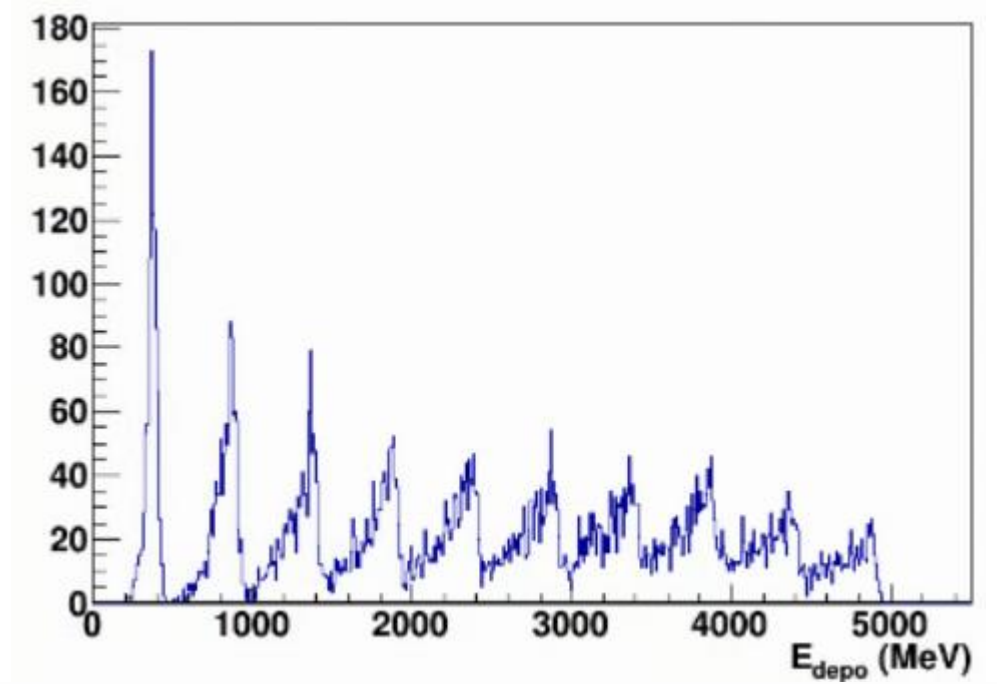
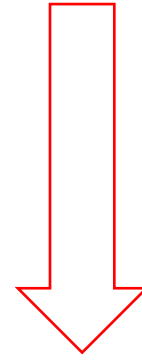


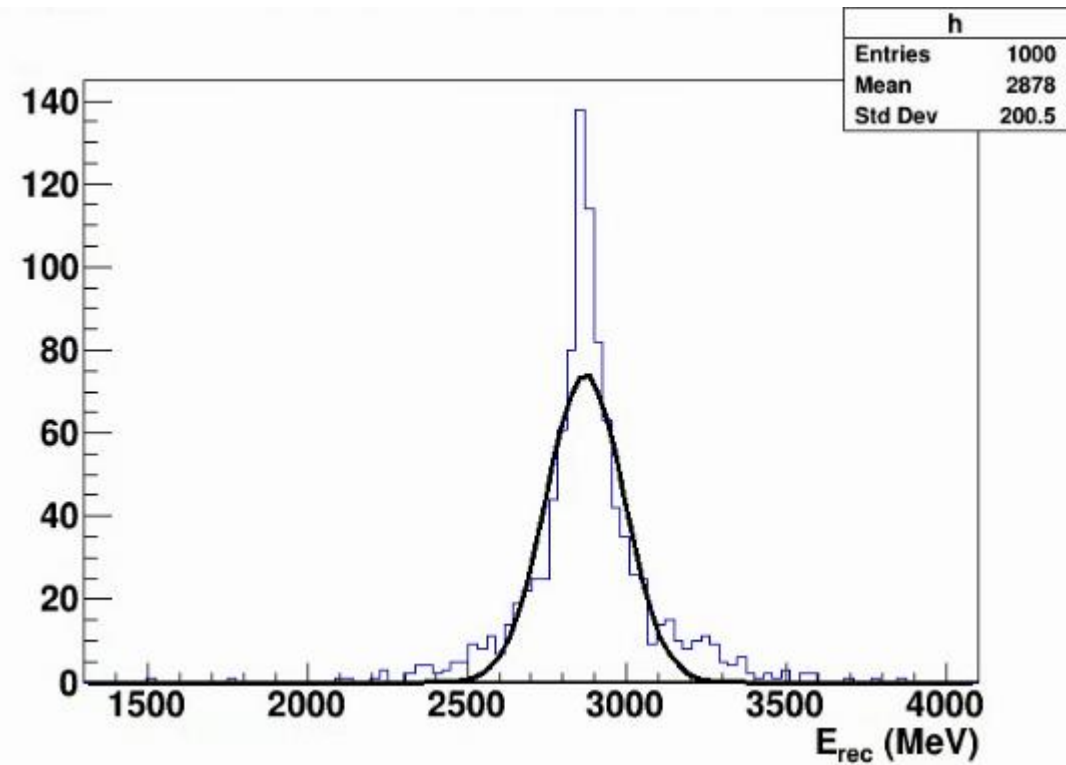
- Simple reconstruction

- Neutron: scale E_{depo} by $1/0.62$
- Pi^{\pm} : scale E_{depo} by $1/0.71$



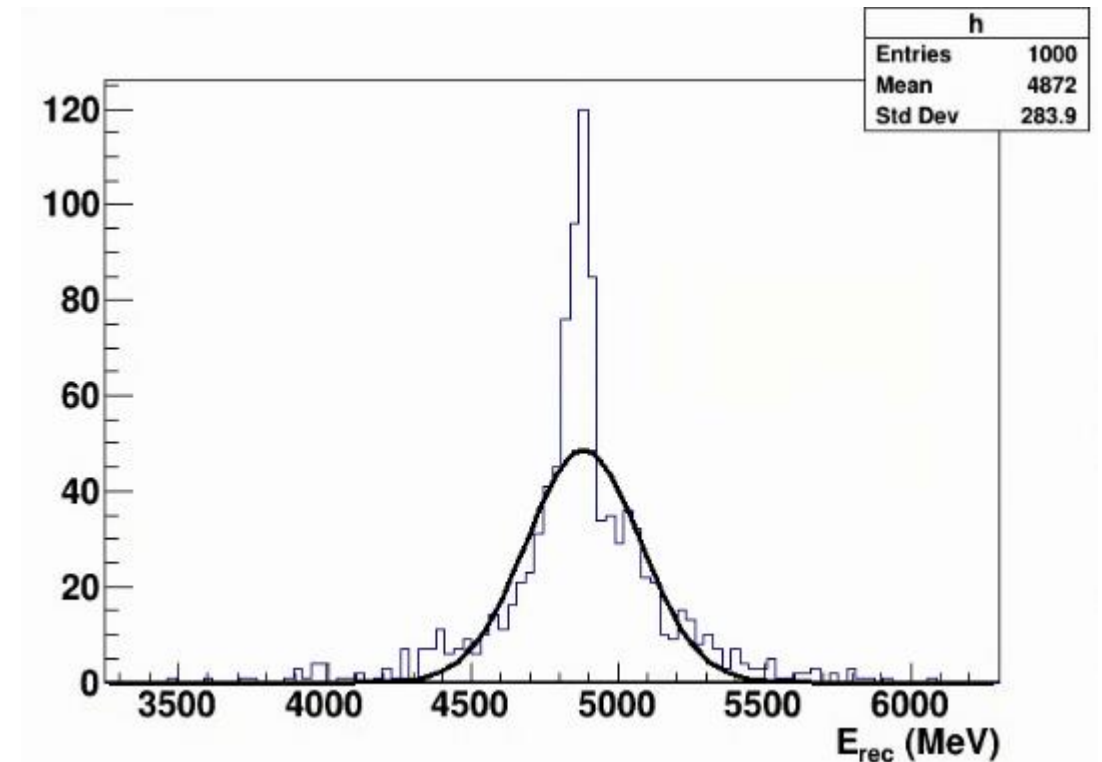
Fitted sigma = 50 MeV, ~5%





3 GeV
Fitted sigma = 120 MeV, ~7%

5 GeV
Fitted sigma = 200 MeV, ~9%



Next Steps

- ❑ Correct decay products: improve resolution
- ❑ Apply thresholds: worsen resolution
- ❑ Apply $dE/dx \rightarrow dQ/dx$: worsen resolution
- ❑ Apply $dE/dx \rightarrow dL/dx$: improve resolution

- ❑ Finally
 - Estimate “realistically achievable energy resolution” with photon information
 - If significantly better than TDR (~13%), study the improvement on CP sensitivity.

Thoughts on PDS timing requirement and optical simulation

- ❑ LAr (doped with LXe) property
 - Scintillation time, fast and slow components
 - Rayleigh scattering
 - Absorption
 - Reflection from detector material
- ❑ PDS timing
 - Single photon time response -> t_0 resolution
 - SiPM response, photon propagation inside the PDS module
 - Multiple photons: multiple-pulse separation
 - Sampling frequency
 - Module granularity

| Parameter | Value |
|----------------------------------|--|
| LAr Photon Yield (mip, 500 V/cm) | 25,000 ph/MeV |
| Xe doping in Ar | 10 ppm |
| Rayleigh Scattering Length | $\lambda_R(128 \text{ nm}) = 1 \text{ m}$ |
| | $\lambda_R(176 \text{ nm}) = 8.5 \text{ m}$ |
| Absorption Length | $\lambda_{Abs}(N_2@128 \text{ nm}) = 20 \text{ m}$ |
| | $\lambda_{Abs}(N_2@176 \text{ nm}) = 80 \text{ m}$ |
| X-Arapuca Tile det. Efficiency | $\epsilon_D = 2\%$ |
| Field Cage Reflectivity | R=70% |
| Cryostat Reflectivity | R=30% @128 nm, R=40% @176 nm |
| Anode | R=0% @128 nm, R=20% @176 nm |

53% of total light emitted @176nm
 35% of light loss @128nm

