



# Nucleon Decay at LBNE

pre-Snowmass IF meeting at Argonne



Eric Church, Yale, 25-27-Apr-2013

## LBNE is...

- A new neutrino beam at Fermilab
  - 700 kW proton beam, 2.3 MW capable
- A near neutrino detector
- An optimal 1300 km baseline: Fermilab-SURF
- A 34 kt Liquid Argon TPC with 4850' overburden
  
- An optimized cost/time effective path to the science
  
- This conceptual design...
  - Completed a successful CD-1 Director's Review (March 2012)
  - Updated cost estimate (July 2012):  
    ~\$1.5B (incl. contingency + escalation)

# LBNE update, continued

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December 2012: CD-1 approval for \$867M first phase DOE funding

- We have completed an extensive cost/schedule for 10 kt LAr far detector (LBNE10) on the surface but the design is **not** fixed
- ***CD-1 approval explicitly allows for scope change enabled by new partners***

**First phase goal:** greater than 10 kt far detector underground and a full capability near detector

In the past 3 months there has been considerable progress towards international partnerships (encouraged by European Strategy statement)

# Underground!

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- ▶ A nucleon decay search is vital for the physics case for the flagship US program.
- ▶ Hence, LBNE is as motivated by nucleon decay studies as it is in filling out the PMNS matrix and finding the Mass Hierarchy
- ▶ As B-mode Cosmic Microwave Background polarization would be evidence of Inflation at the Planck scale of  $10^{19}$  GeV, so is nucleon decay a window into Grand Unification Theories (GUTs) on scales of  $10^{16}$  GeV.

# ideas of the GUT modelers in this room!

OSU-HEP-13-02  
MAN/HEP/2012/017  
UMD-PP-013-003

OSU-HEP-12-08

## Post-Sphaleron Baryogenesis and an Upper Limit on the Neutron-Antineutron Oscillation Time

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Such modes imply  $n\bar{n}$  oscillations ...

## B - L Violating Proton Decay Modes and New Baryogenesis Scenario in SO(10)

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Abstract

breaking scenario with  $O(1000 \text{ TeV})$  color sextet particles would allow  $d=7$   $B/L$  modes.  $B/L$  means low temperature baryogenesis is possible

## Type II Seesaw Dominance in Non-supersymmetric and Split Susy SO(10) and Proton Life Time

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(Date: October 31, 2011)

Recently type II seesaw in a supersymmetric SO(10) framework has been found useful in explaining large solar and atmospheric mixing angles as well as a larger value of  $\theta_{13}$ , while unifying quark and lepton masses. An important question in these models is whether there exists consistency between coupling unification and type II seesaw dominance. Scenarios where this consistency can be demonstrated have been given in a SUSY framework. In this paper we give examples where type II dominance occurs in SO(10) models without supersymmetry but with additional  $U(1)$  gauge particles and also in models with split-supersymmetry. Grand unification is realized in a two step process via breaking of SO(10) to SU(5) and then to a TeV scale standard model supplemented by extra fields and an SU(3) Higgs multiplet  $15_\Delta$  at a scale of about  $30^4$  GeV to give type II seesaw. The predictions for proton lifetime in these models are in the range  $\tau_p^p = 2 \times 10^{29}$  yrs. to  $6 \times 10^{30}$  yrs. A number of recent searches for  $d=7$  nucleon number violation can be accommodated within

Type II seesaw, no SUSY, is/was possible. Some TeV scale particles needed, in any case.

## Proton decay and $\mu \rightarrow e + \gamma$ Connection in a Renormalizable SO(10) GUT for Neutrinos

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Abstract

Type I or II seesaw is potentially distinguishable

[hep-ph] 27 Mar 2013

[hep-ph] 24 Jul 2012

MIFFA-13-0  
UMD-PP-013-00  
February, 2013

28 Oct 2011

[hep-ph] 11 Feb 2013

# GUT phenomenology

require/get	discovered/ accommodated	needed	bonus
<p>coupling unification</p> <p>(depending on type I/II seesaw) Cabibo angle, <math>m_b = m_\tau, \dots</math> at <math>M_G</math></p>	<p><math>m_\nu</math></p> <p>large <math>\theta_{13}</math></p>	<p><b>nucleon decay</b></p> <p>normal hierarchy</p> <p><math>O(1-1000 \text{ TeV})</math> particles</p>	<p>nnbar oscillations</p> <p>B,L (<math>\delta_{CP}</math>) asymmetry</p> <p><math>O(1-1000 \text{ TeV})</math> particles!</p>

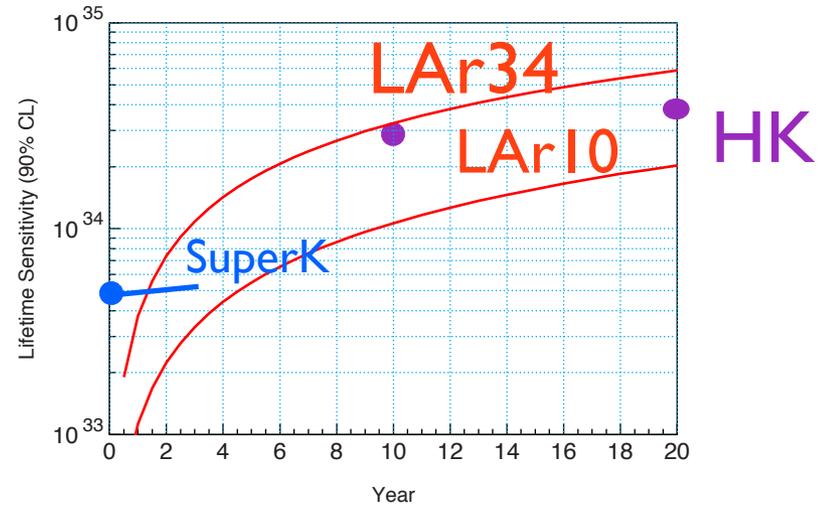
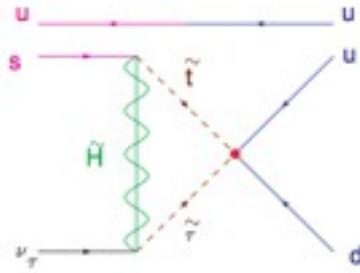
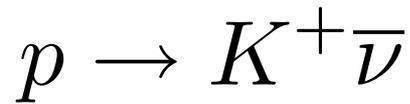
# example pdk Modes

	Mode	SK Eff (%)	SK Bkgd (1/MT-yr)	LAr Eff (%)	LAr Bkgd (1/MT-yr)
B-L	$p \rightarrow e^+ \pi^0$	45	2	45	1
	$p \rightarrow K^+ \bar{\nu}$	6	7	97	1
	$p \rightarrow \mu^+ K^0$	10	5-10	47	<2
B+L	$p \rightarrow \mu^- \pi^+ K^+$	?	?	97	1
	$p \rightarrow e^- K^+$	10	3	96	<2

unofficial SK!

Bueno, et al.

# golden mode:



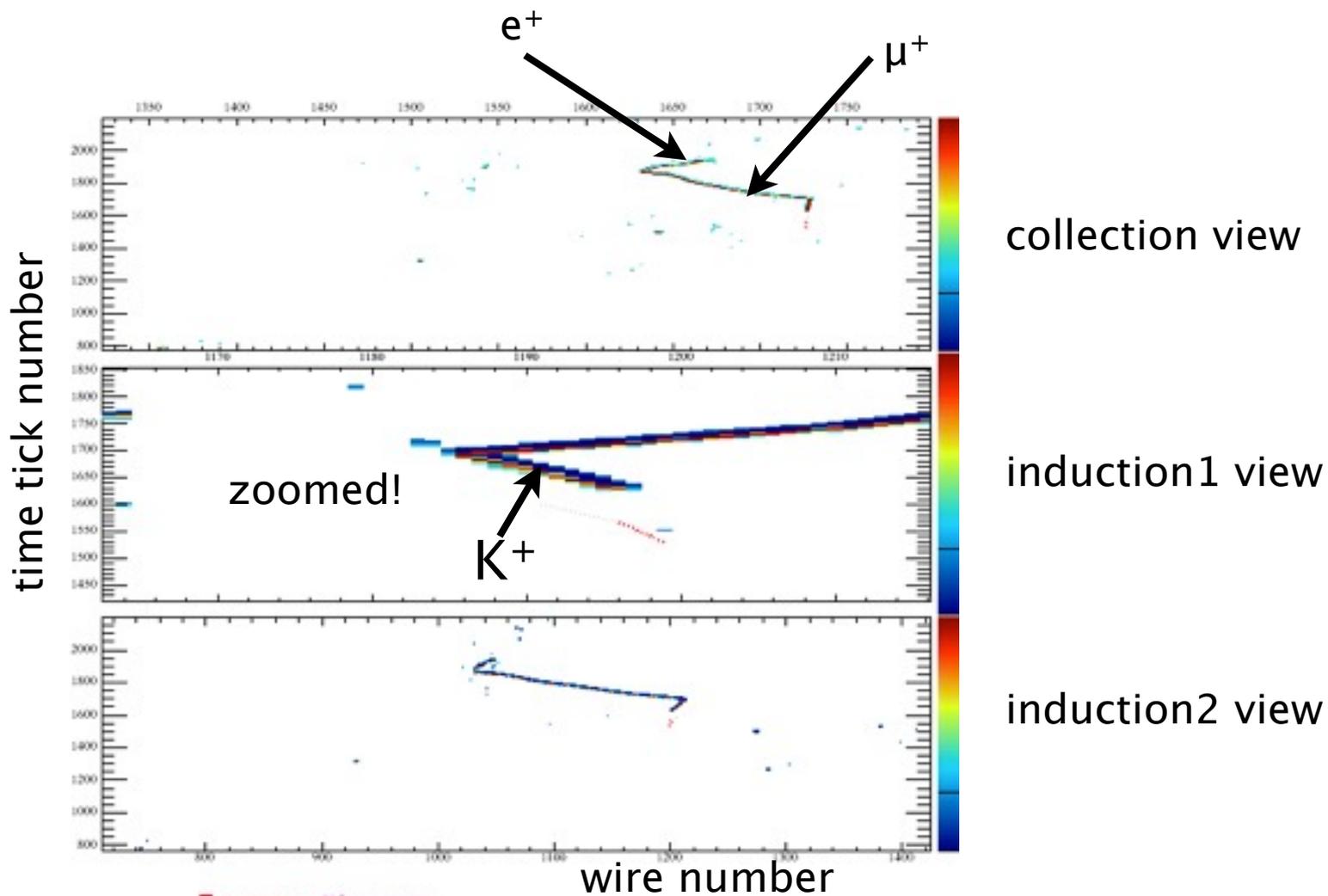
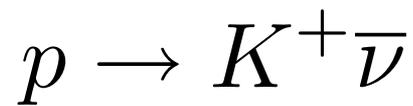
Ed Kearns

Assuming  $<0.1$   
 $K^+$  bgd/100kt-yr,  
 we get  $\sim 6x(1.5x)$  still-  
 running SK(HK)  
 sensitivity.

$$\frac{\tau}{B} = \frac{N_0 \Delta t \epsilon}{n_{obs} - n_{bg}}$$

S →

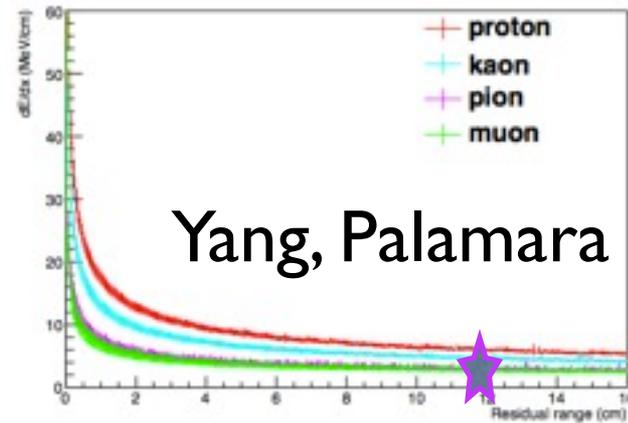
$$\frac{\sum_{n=0}^{n_0} P(n, b+S)}{\sum_{n=0}^{n_0} P(n, b)} = \alpha$$



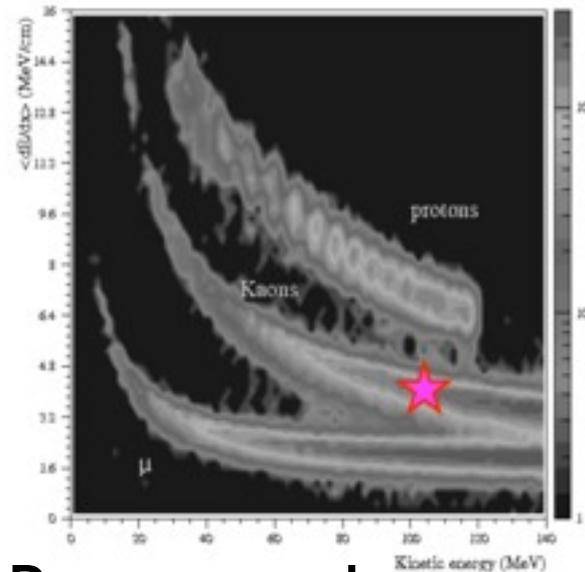
$\bar{\nu}_\mu$  [0.3 GeV/c] +  $K^+$  [0.3 GeV/c]  
MicroBooNE Event Display

# K<sup>+</sup> from pdk

Kaons: particle ID  
well-separated at this  
momentum, even with  
Fermi smearing.



ArgoNeuT  
MC  
truth



Bueno, et al. truth?

•**Bill Willis impressed the following upon me. (In fact, this is his slide, almost unedited.)**

Bill made important contributions and provided insightful leadership to novel detector techniques and liquid argon calorimetry and LAr TPCs, particularly. There is a symposium in his honor at BNL today.

**... due to isospin – in all a considerable increase in rate, about 12 times the pi-zero rate in Hyper K. (2–3x Bill Marciano’s claim.)**

- Keeping the expected number of background events for a given reaction  $\ll 1$  is important, to allow discovery with just a few events.
- some reactions allow near zero background events, others (like  $K^+$  neutrino) allow an almost perfect fake.
- Capturing all modes is important: a good example is the set of decays on following page (the channels with negative pions are expected to have double the rate)

# Bill Marciano and Bill Willis suggest inclusive searches

e.g., B-L conserving modes

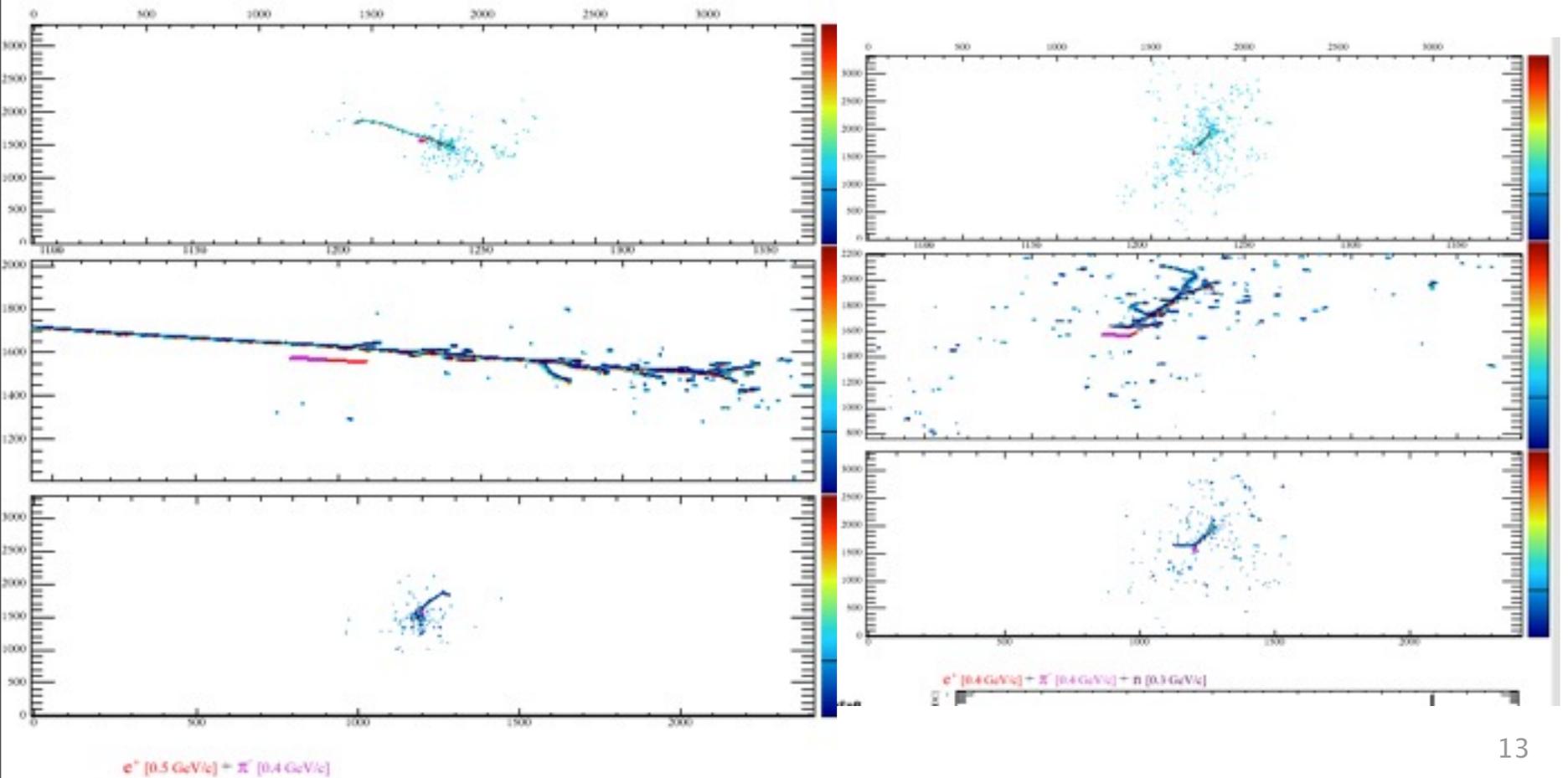
- Meaning, instead of  $p \rightarrow e^+ \pi^0$
- limited by inherent 45% absorption, which only gets worse for Argon, we can hope to see  $n \rightarrow e^+ \pi^-$  too and  $p, n \rightarrow \bar{\nu} \pi^{+,0}$  and  $p \rightarrow e^+ \omega \rightarrow \pi \pi \pi$

(and swap  $\mu$  in for e everywhere too.)



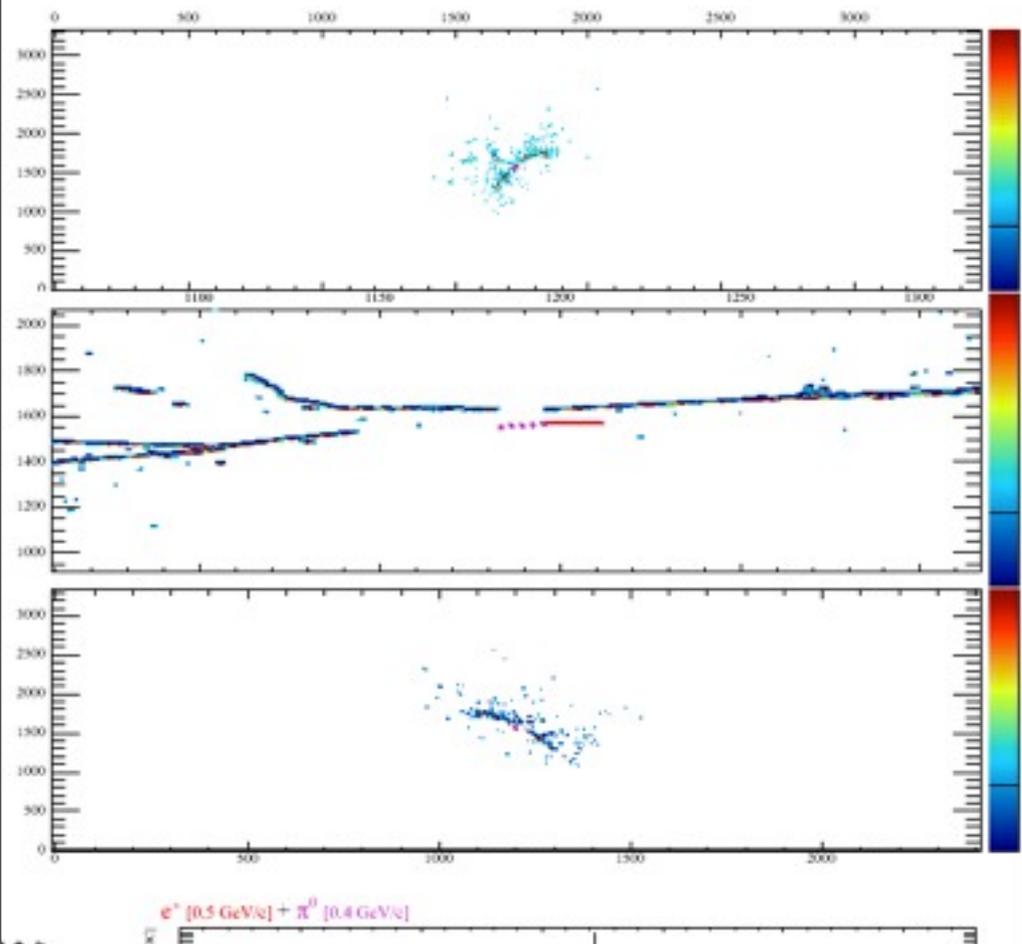
Easy

Hard

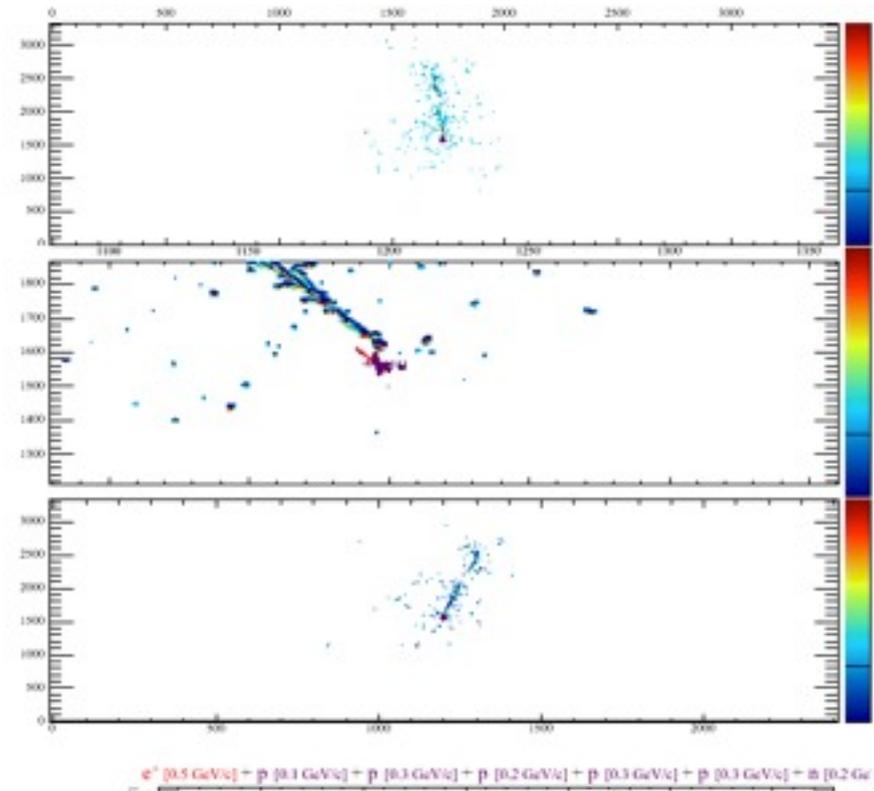


$$p \rightarrow e^+ \pi^0$$

Easy



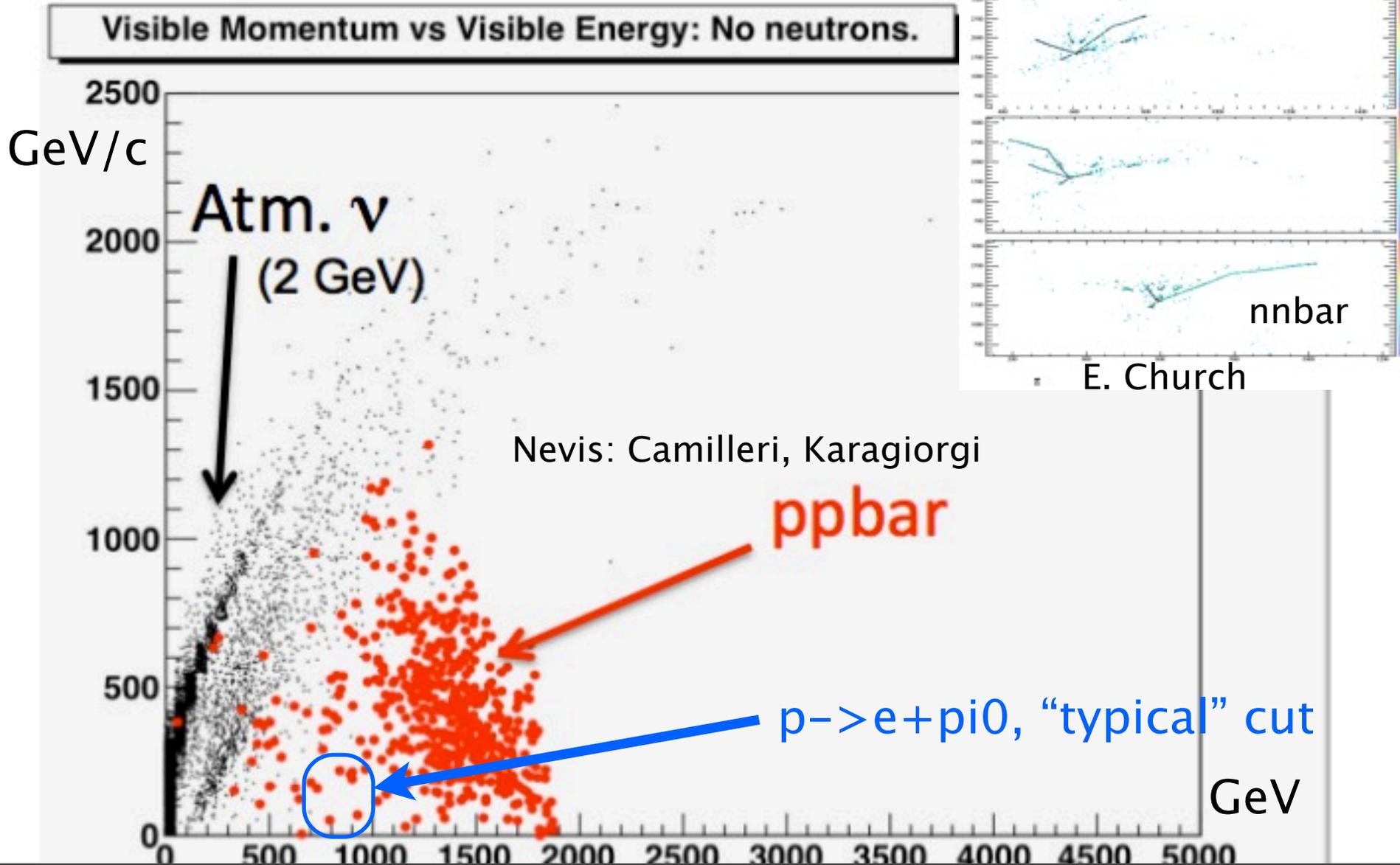
Hard

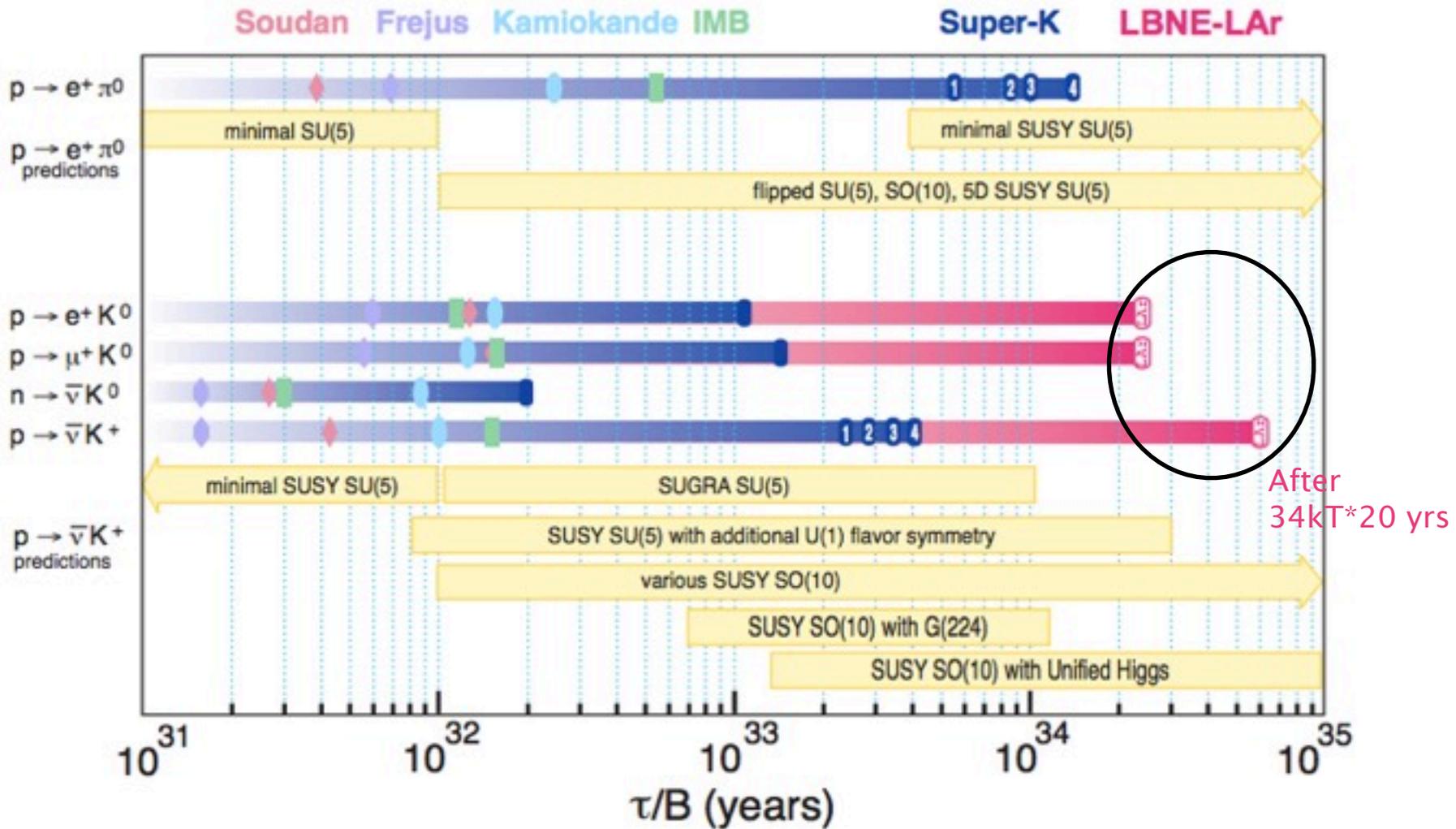


# At 4850'

- Atmospheric generator works in LArSoft (the Argon simulation and reconstruction software package).
- Expect atmospheric neutrino interactions to dominate
- We have generated muons at 4850'.  $K_0$  and cosmic muon-Ar spallation products need to be understood too.
- Then we need to walk through the modes
- We have a Ndk generator (GENIE) now for most interesting modes.
  - Not clear that FSI for Kaons in GENIE make sense.
  - No  $n \rightarrow K_0, \nu$

# momentum-energy balance in $n\bar{n}$ (qua $p\bar{p}$ ) annihilation





The theory space is vast and the sense of satisfactorily vanquishing any one class of models is elusive, but of course our job is to do the experiment(s).

# Ndk summary, per Bill M.

- Generically, partial lifetime for  $p \rightarrow e^+ \pi^0$  goes like
- $\sim \text{sqrt}(1 \text{ TeV} / M_{\text{susy}}) \cdot 10^{35 \pm 1} \text{ yr}$
- As LHC push on  $M_{\text{susy}}$  from below, we push down on lifetime from above.

# Bkup

# 800 feet (not longer considered), $K^0_L$ bkgd

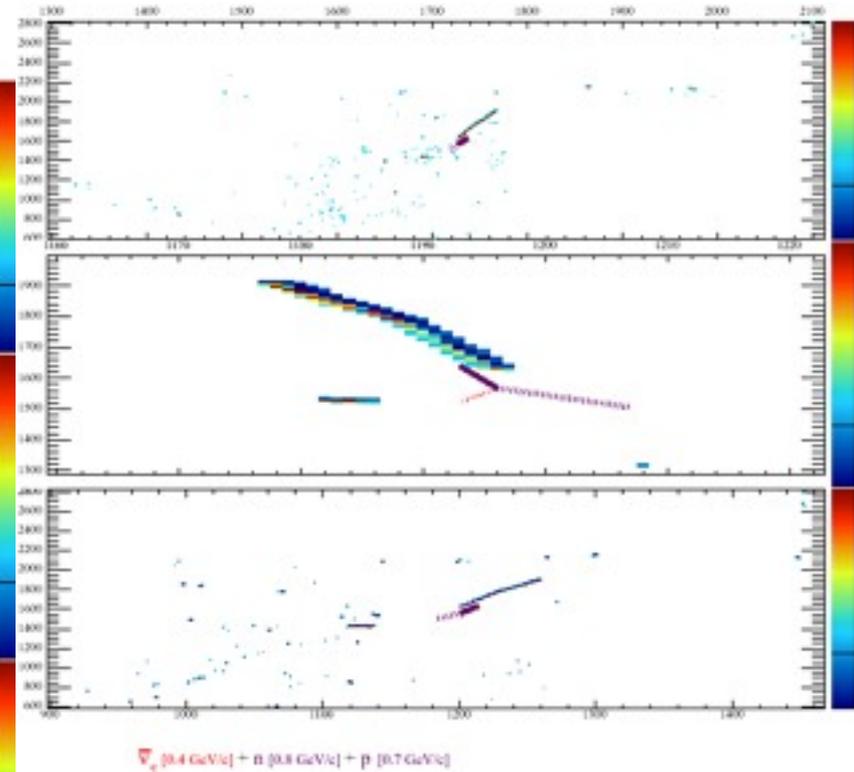
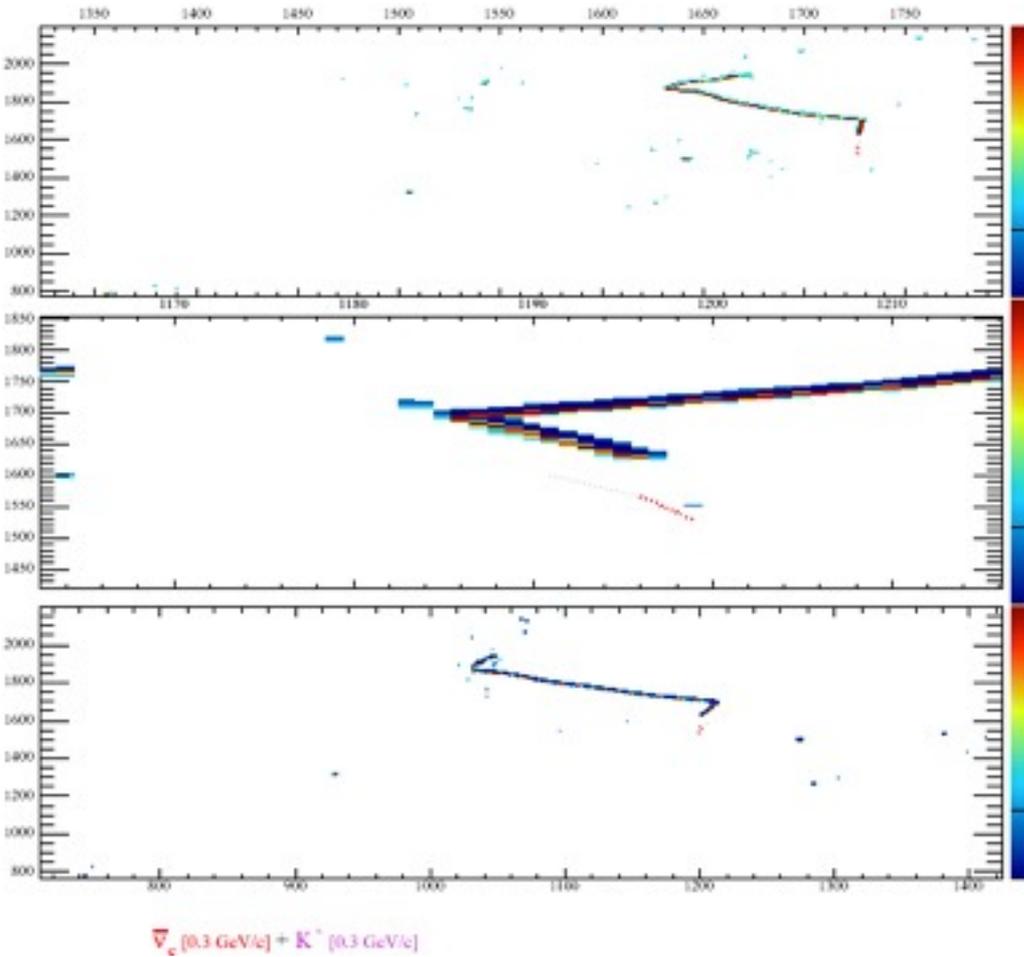
study/Authors	Muon Simulation Package	Exposure	Bgd/100 kt-yr
Bueno, et al	Geant4 and FLUKA	10+ years	<0.1
LBNE.Sheffield	MUSIC/Geant4	X	<0.1
LBNE.JdJ+EC	Geant4	200 kt-years	<<0.1

Slide 8 again

# $p \rightarrow K^+ \text{ nubar}$

less nice event.  
K+ lost in FSI,  
GENIE says.

Nice “prototypical” event.



# Does GENIE FSI make sense?

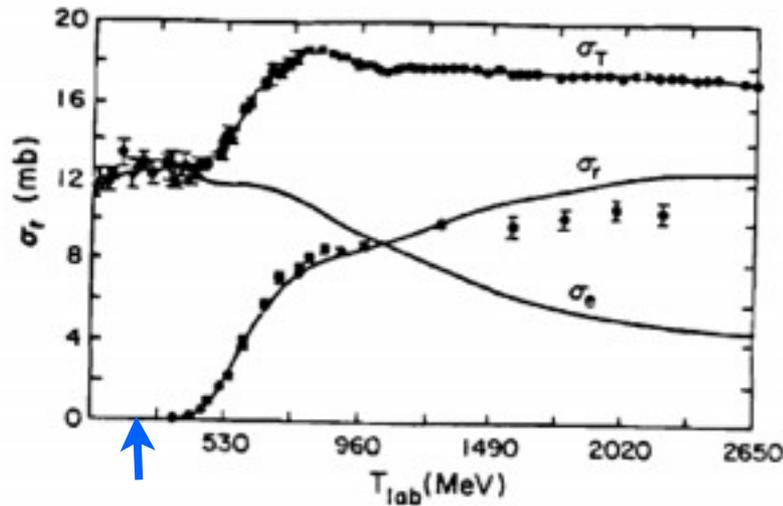


FIG. 4.4. Total ( $\sigma_T$ ), inelastic ( $\sigma_r$ ), and elastic ( $\sigma_e$ )  $K^+ p$  cross sections. [From Arndt and Roper (85).]

arrow at  $T(p)_{K^+} = 106 \text{ MeV}$   $337 \text{ (MeV/c)}$

1 mb => 1.5 Argon radius interaction length

charge exchange

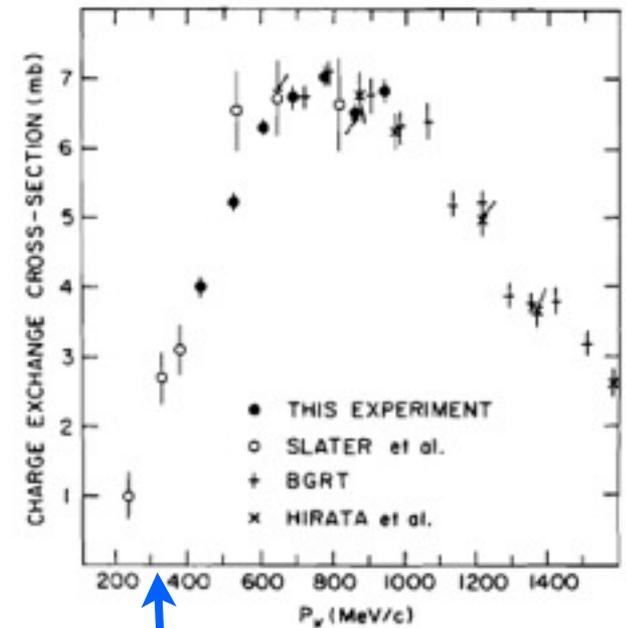


FIG. 4.7. Cross section for  $K^+ n \rightarrow K^0 p$  charge exchange. [From Darnerell, et al. (85).]