

# What can we learn from Charged Leptons vs. Neutrinos?

David McKeen  
University of Victoria

work with Batell, Pospelov, Ritz  
& many others' work...

# Outline

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- High-scale SUSY

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- **Muon anomalies**

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- Stueckelberg Portal

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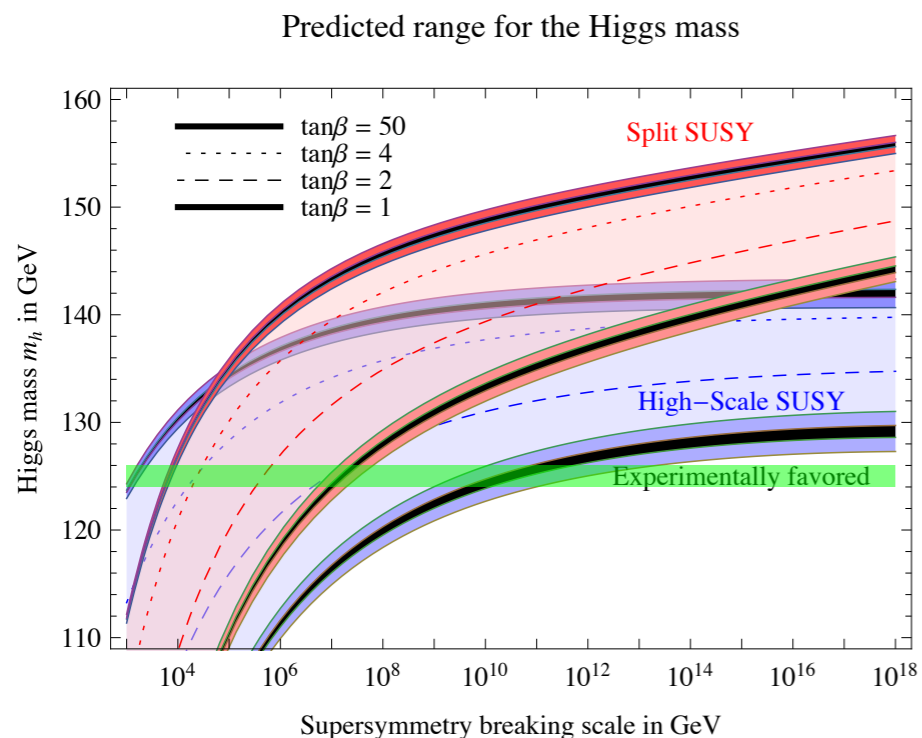
- High-scale SUSY
- Muon anomalies
- Stueckelberg Portal
- Leptonic Higgs Portal

Arkani-Hamed, Kaplan, Weiner,  
Giudice, Strumia, Kane, Hall,  
Nomura, etc. etc....

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- Higgs at 126 GeV, absence of flavor violation & squarks at LHC suggest SUSY broken at a high scale



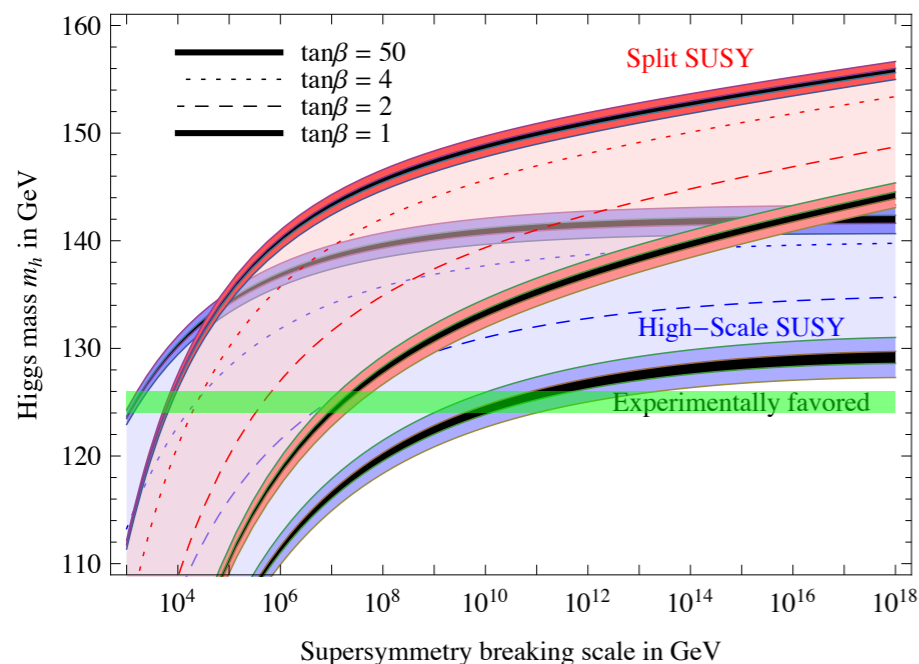
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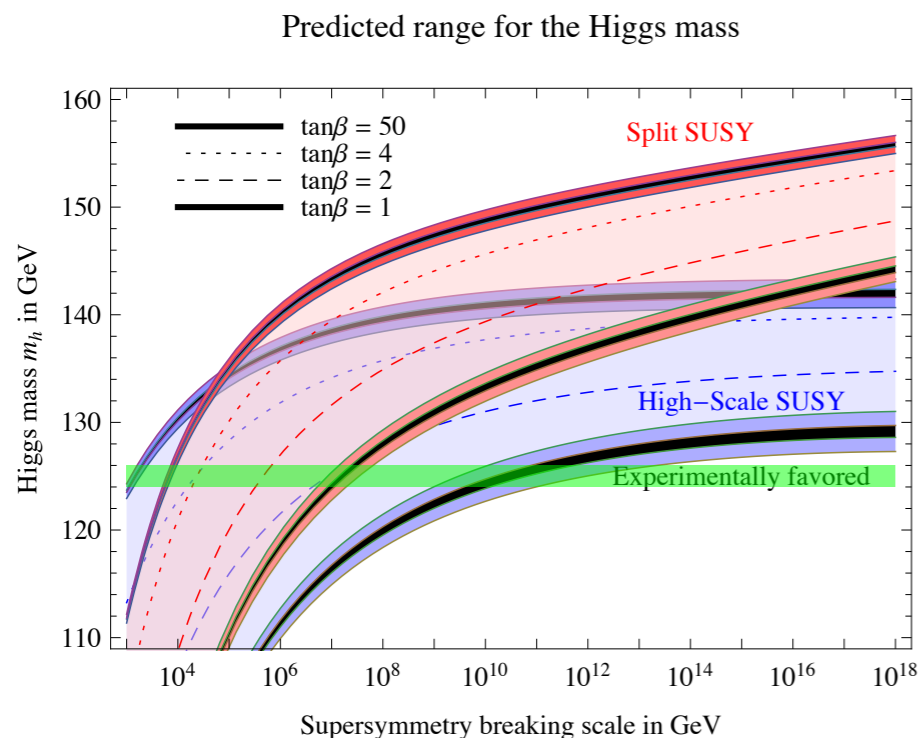
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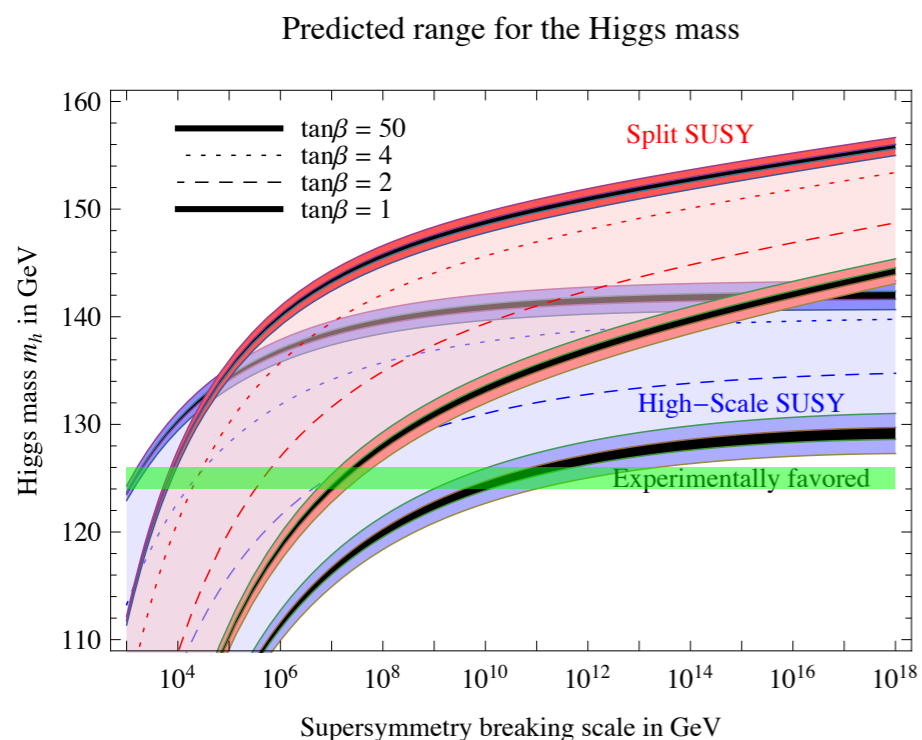
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- Flavor constraints satisfied for soft masses with “anarchic” flavor structure

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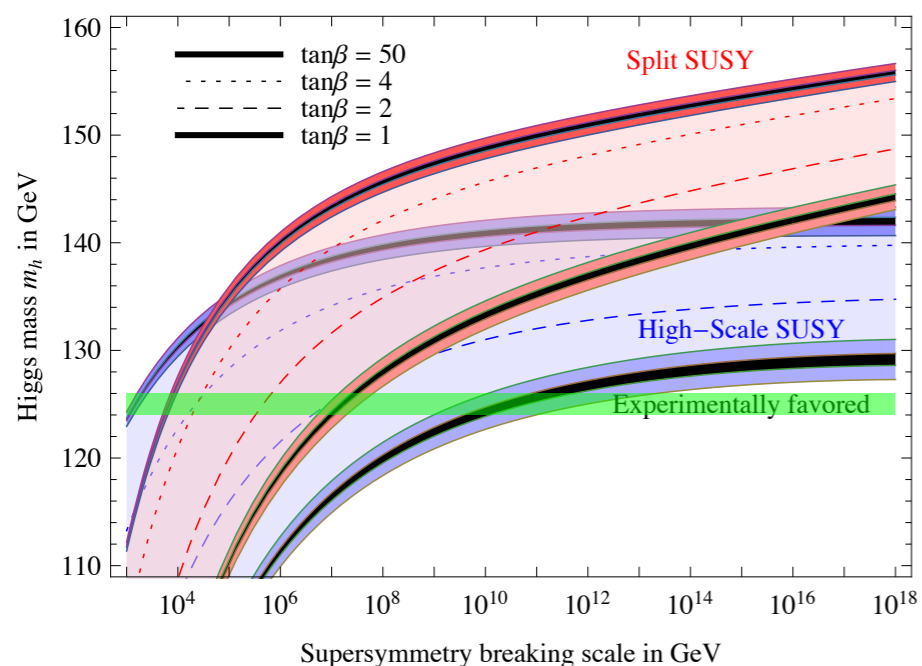


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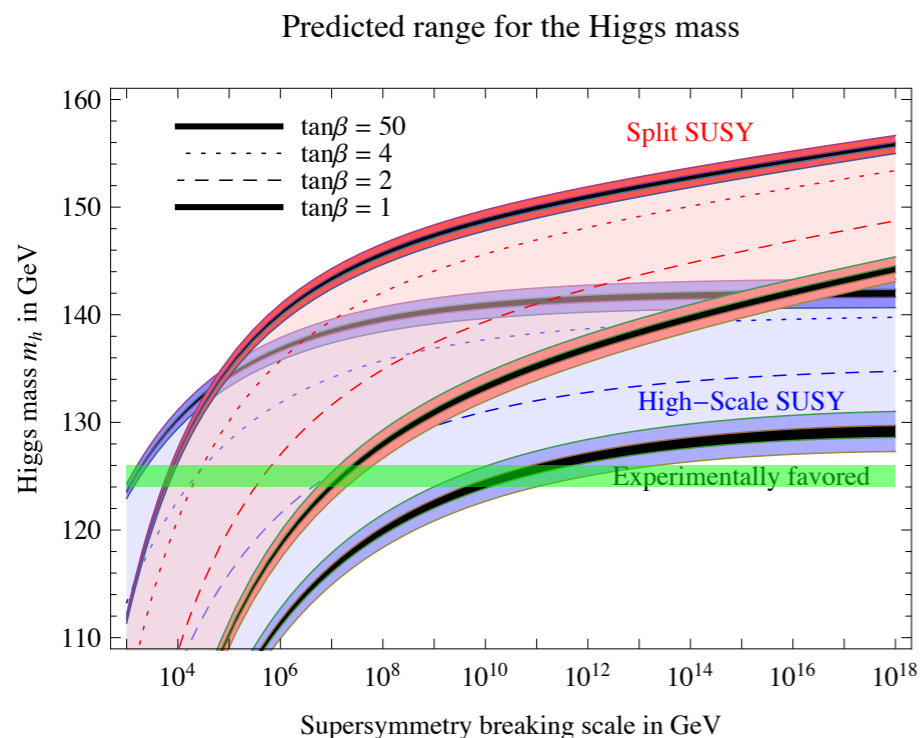
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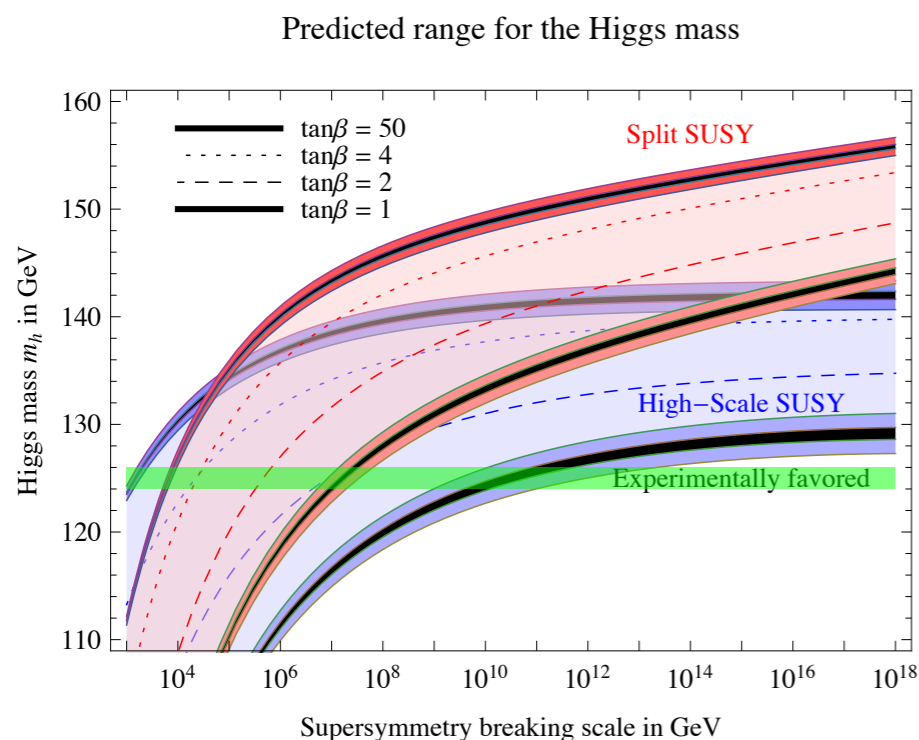
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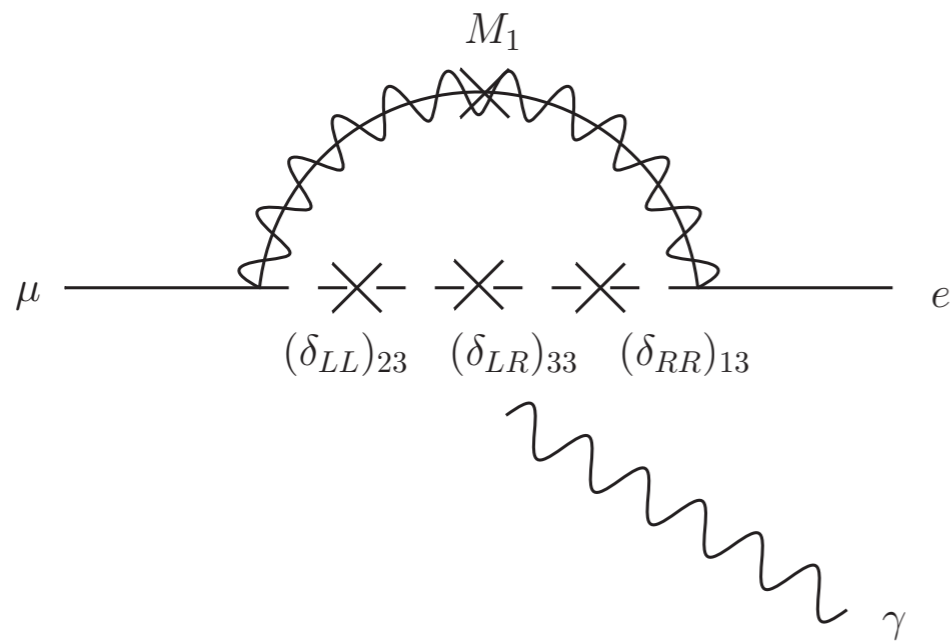
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- Increased importance of LR observables

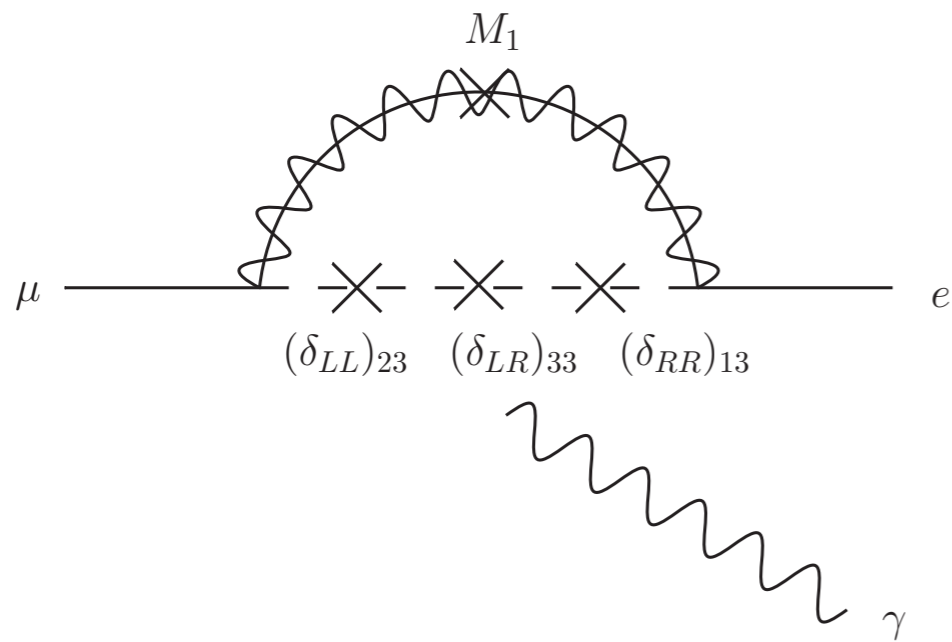
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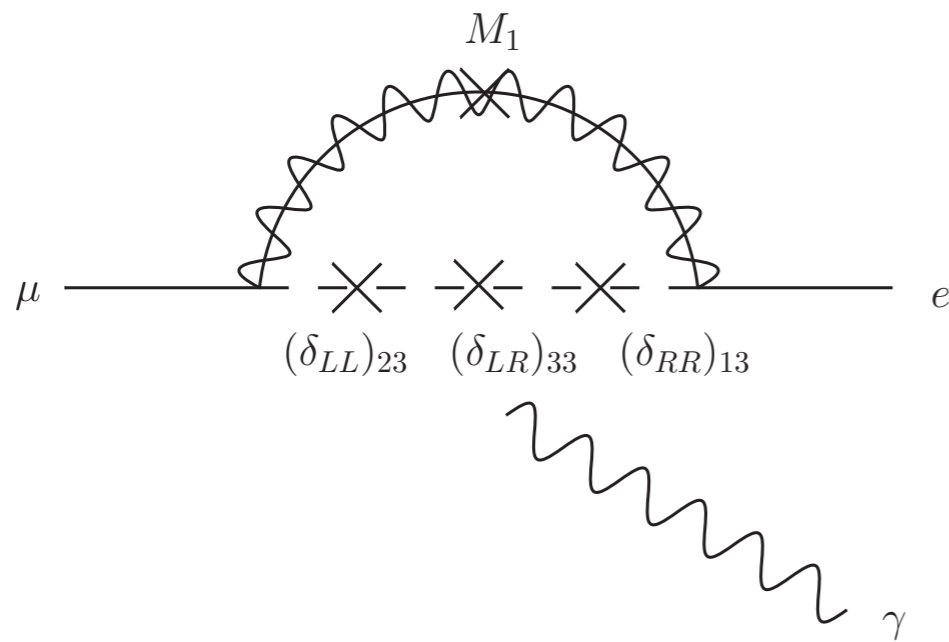


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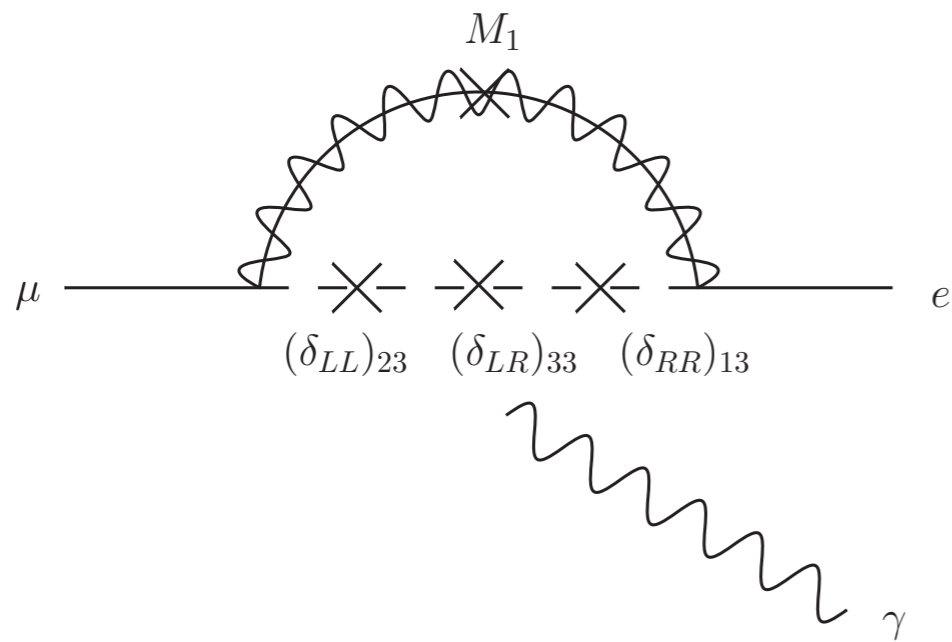


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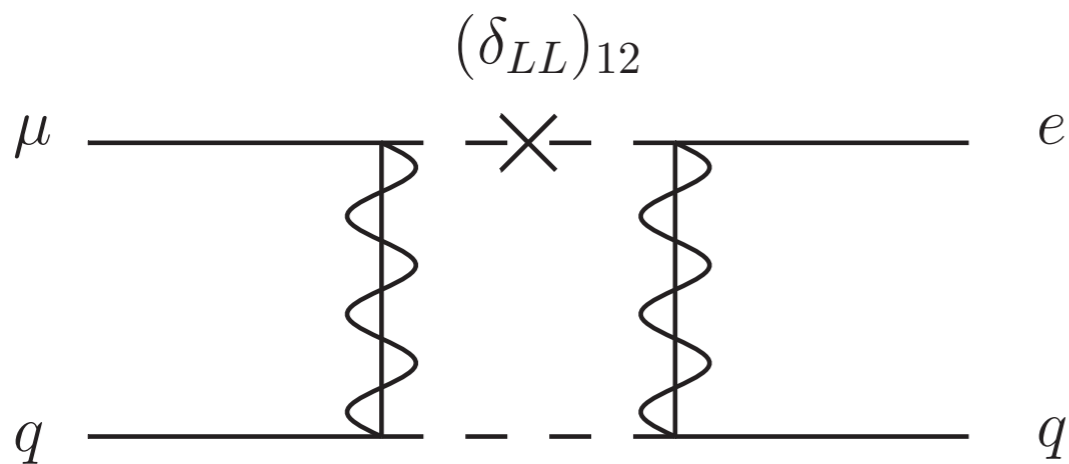
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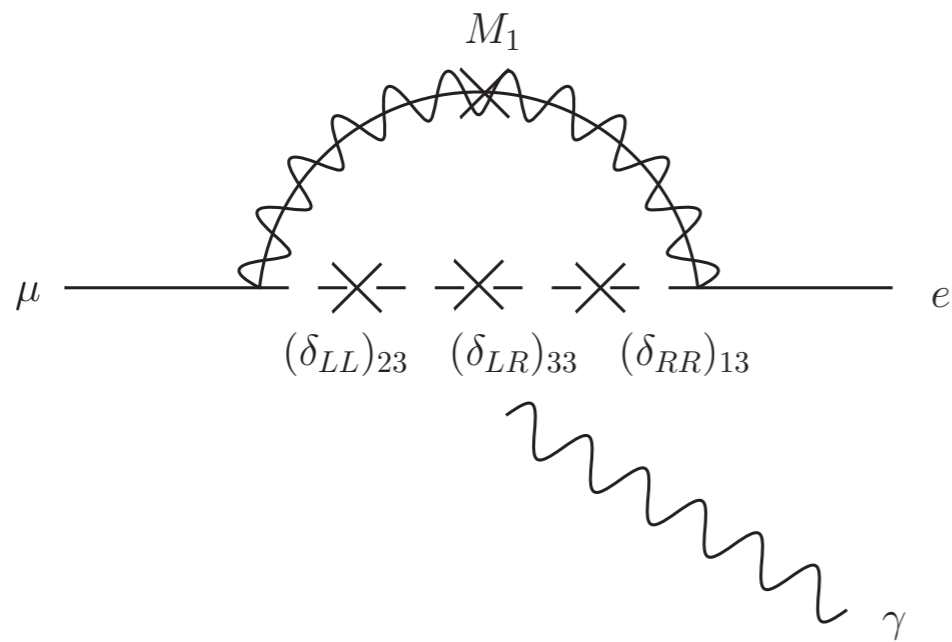
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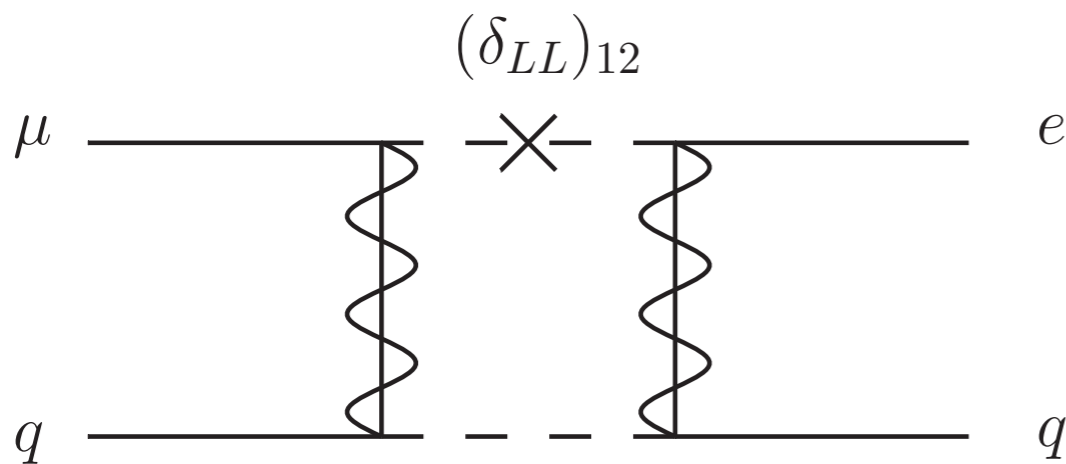
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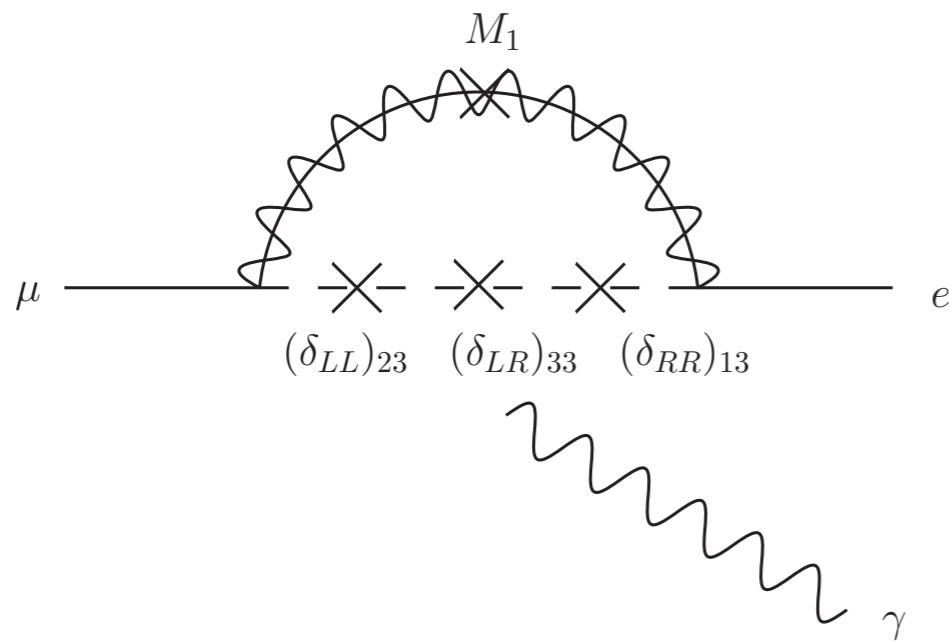
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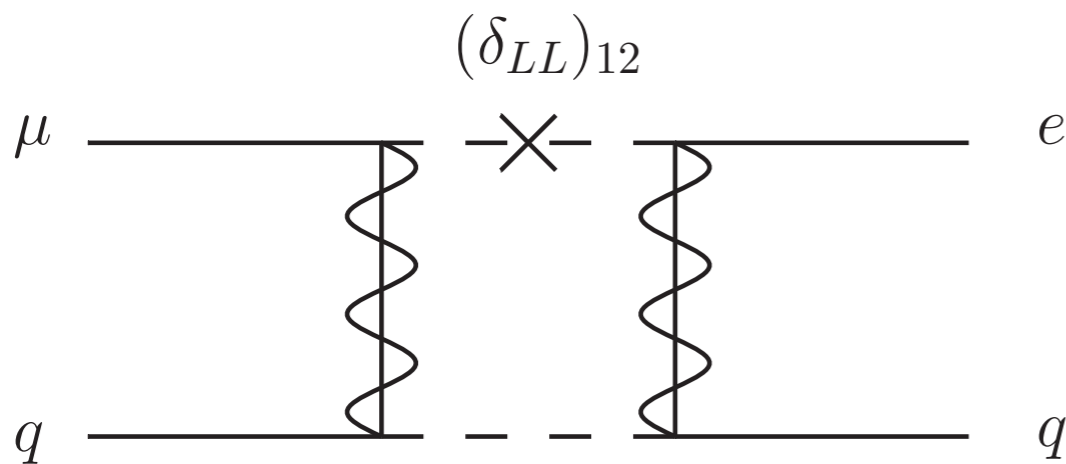
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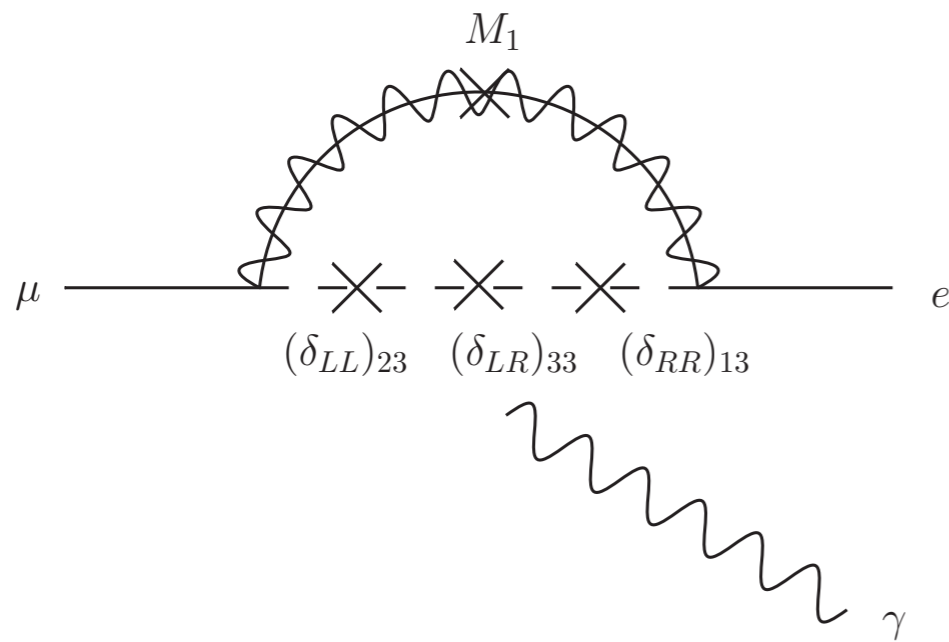
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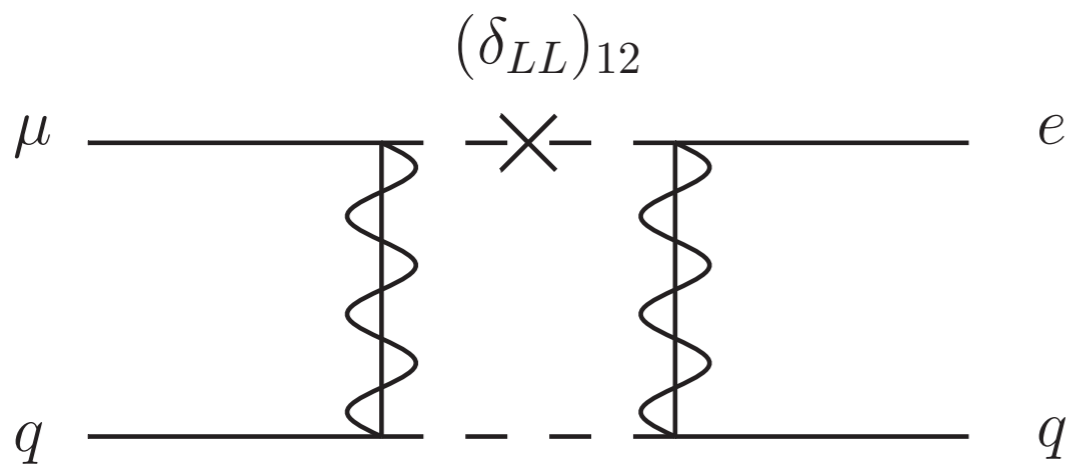
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We'll talk a bit about this

**Proton charge radius in  $\mu\text{u-H}$**

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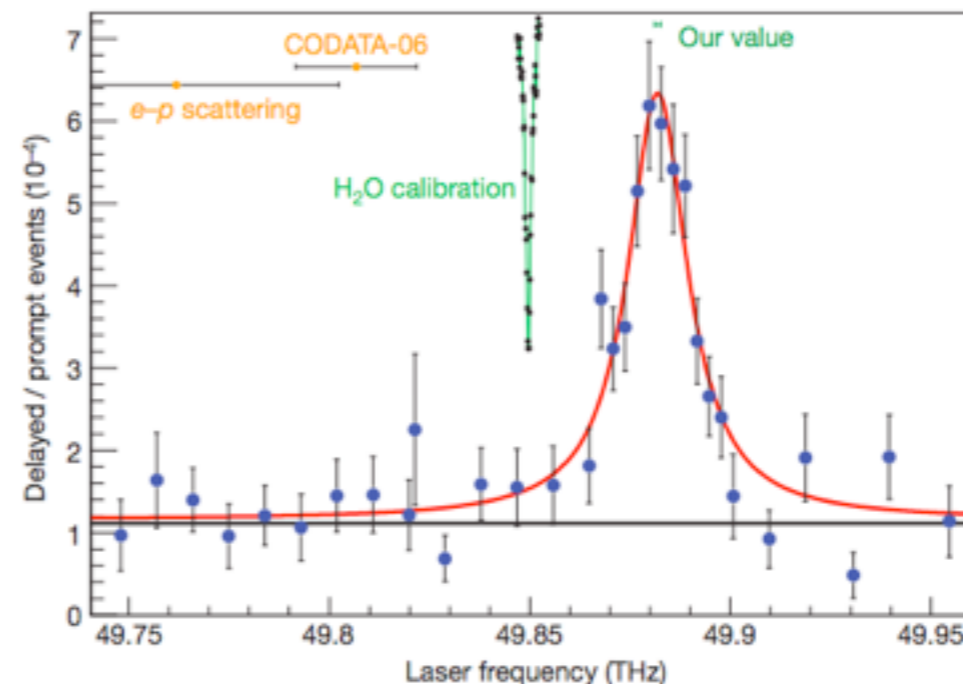
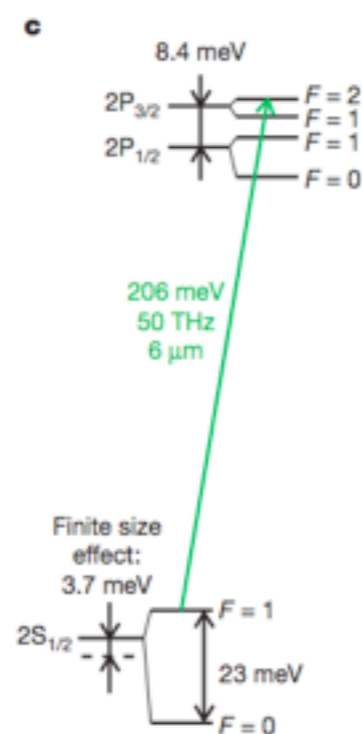
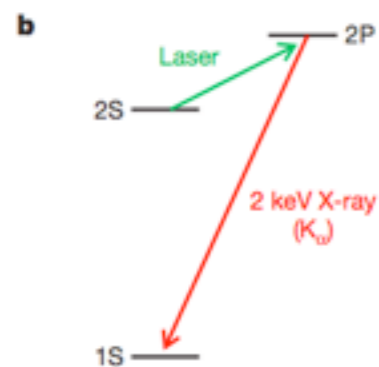
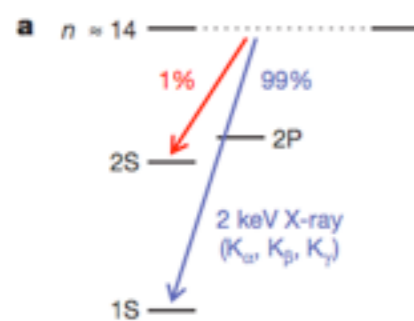
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● Muonic H:  $\Delta E = E(2P_{3/2}^{F=2}) - E(2S_{1/2}^{F=1}) = 206.2949(32)$  meV

$$\Delta E_{\text{theory}} = (209.9779(49) - 5.2262 r_p^2 + 0.0347 r_p^3) \text{ meV}$$

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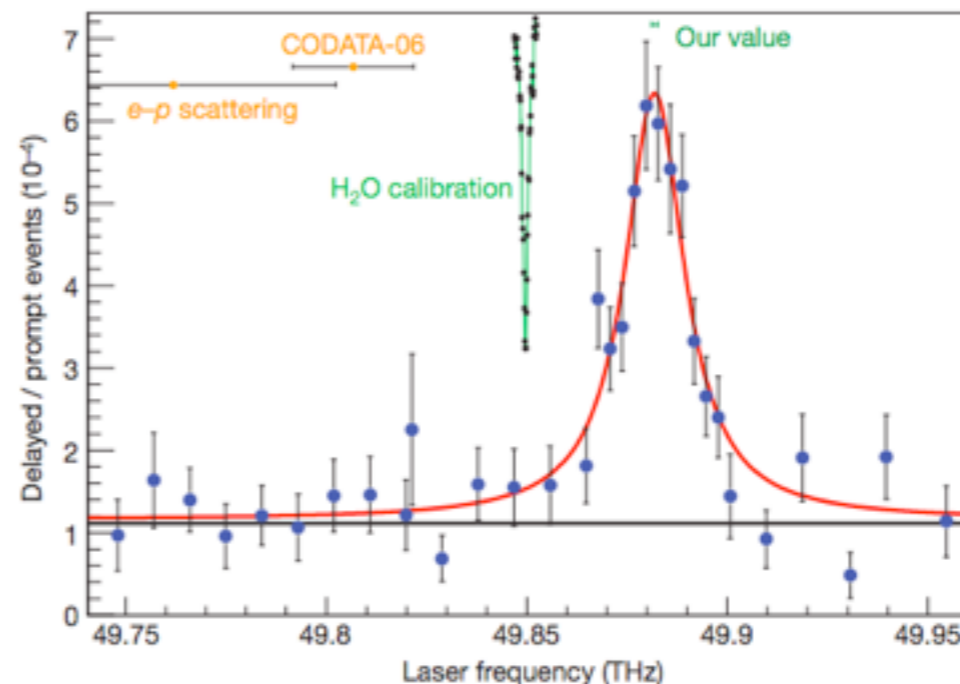
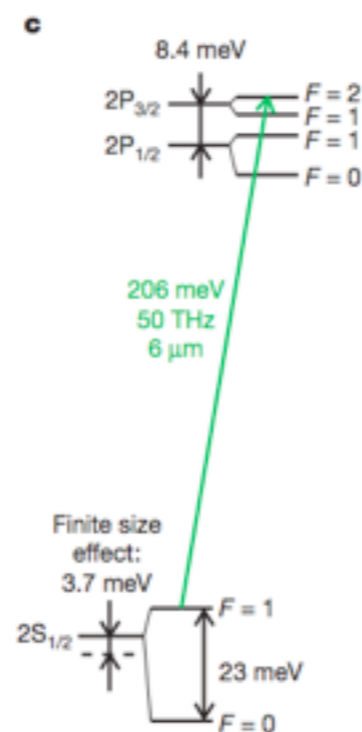
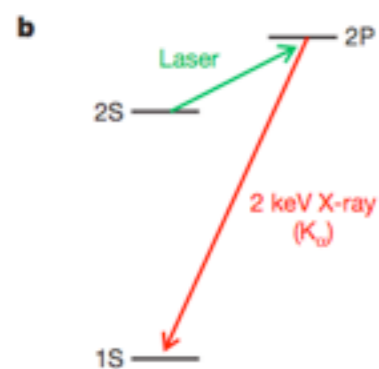
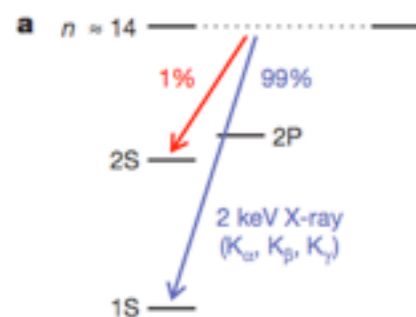
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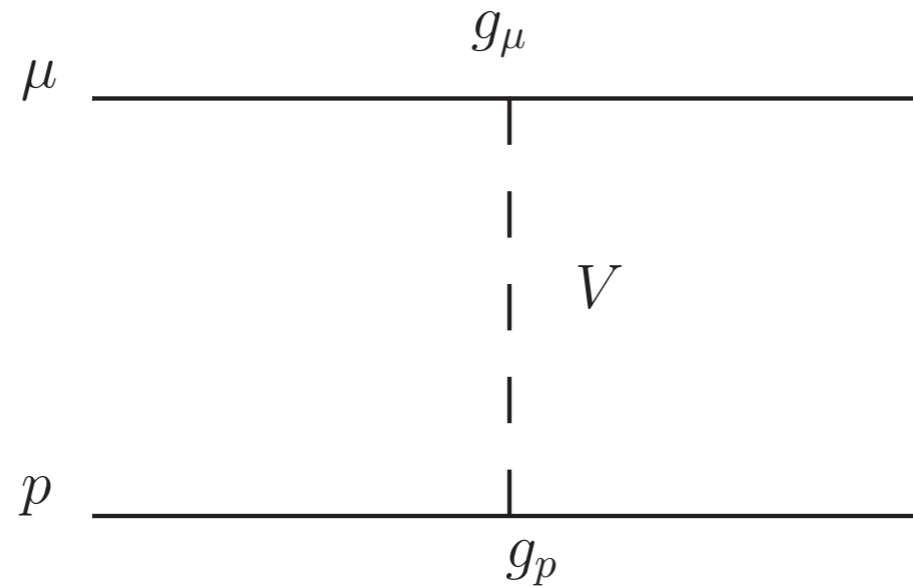
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**Recent update:**  $r_p = 0.84087(39)$  fm

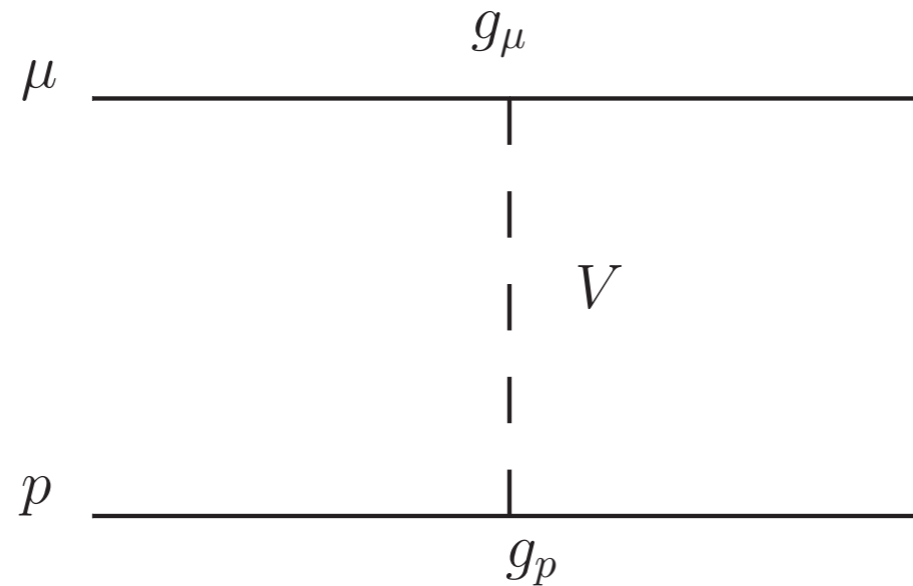


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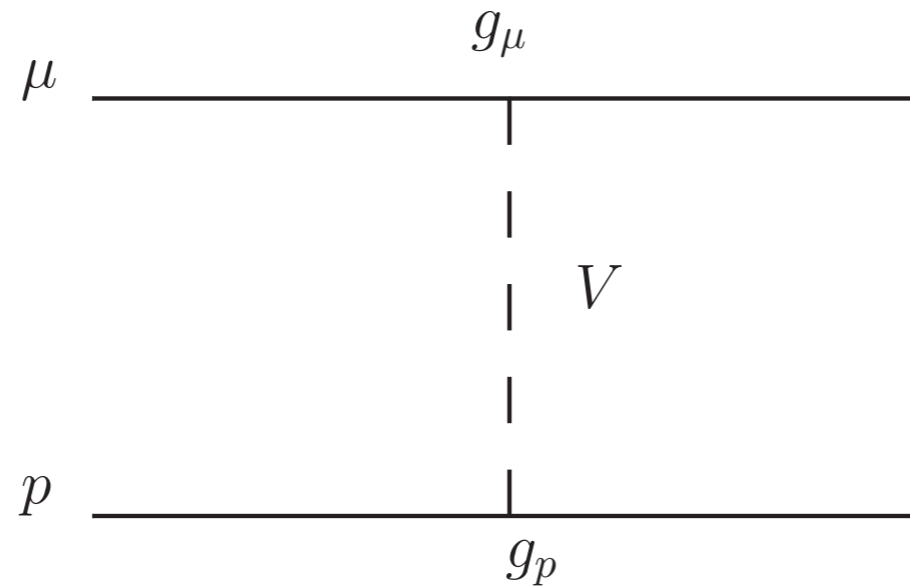
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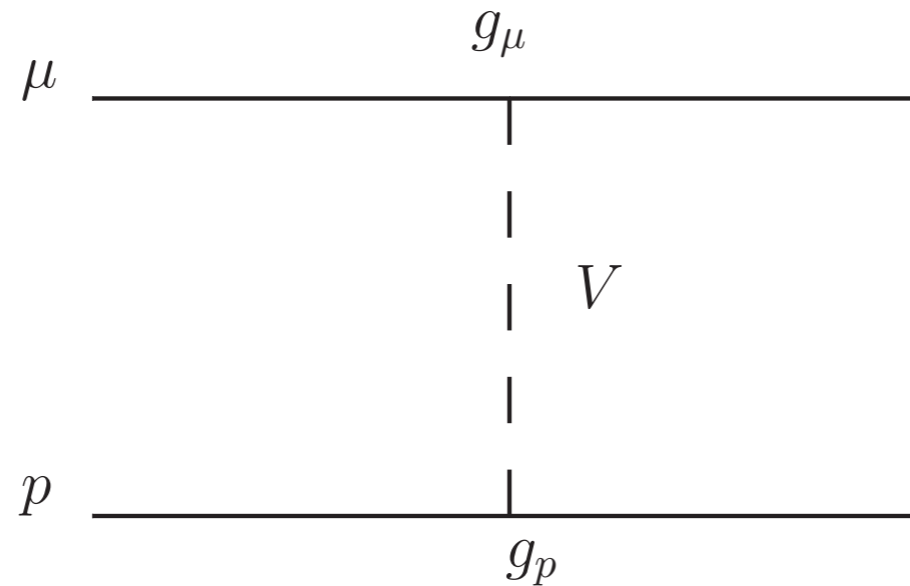


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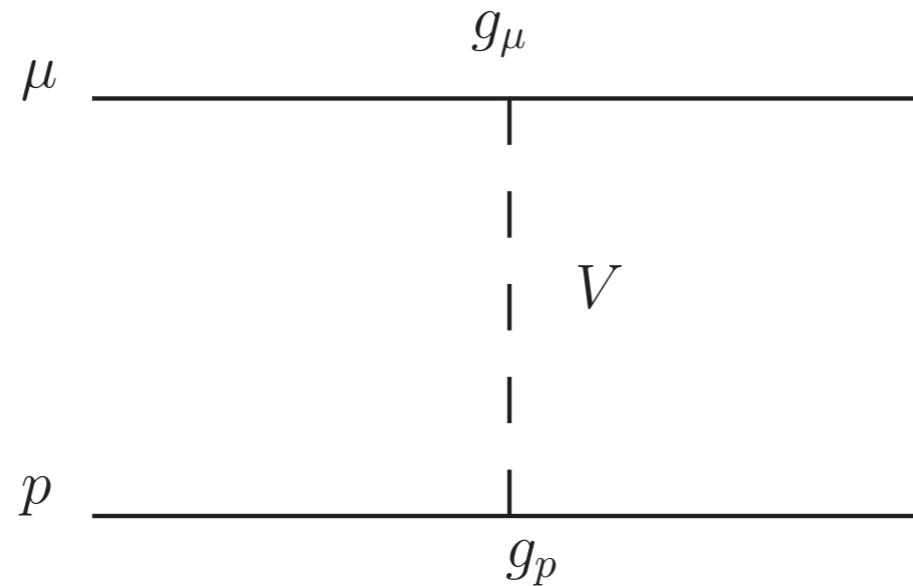
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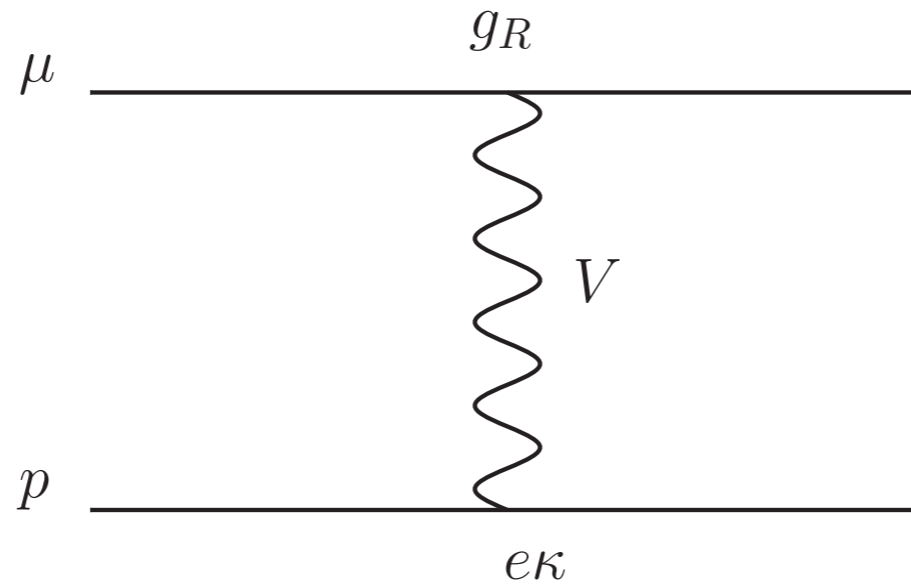
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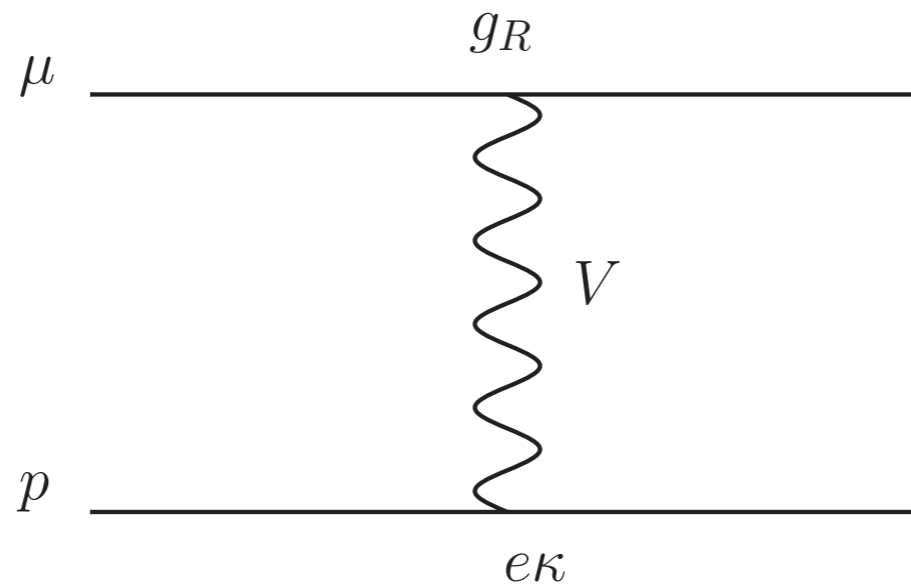
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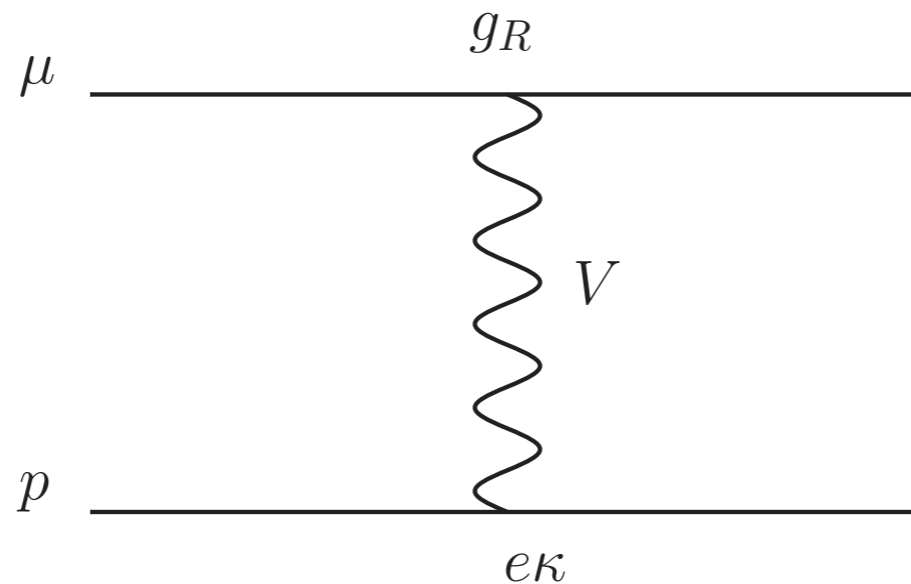


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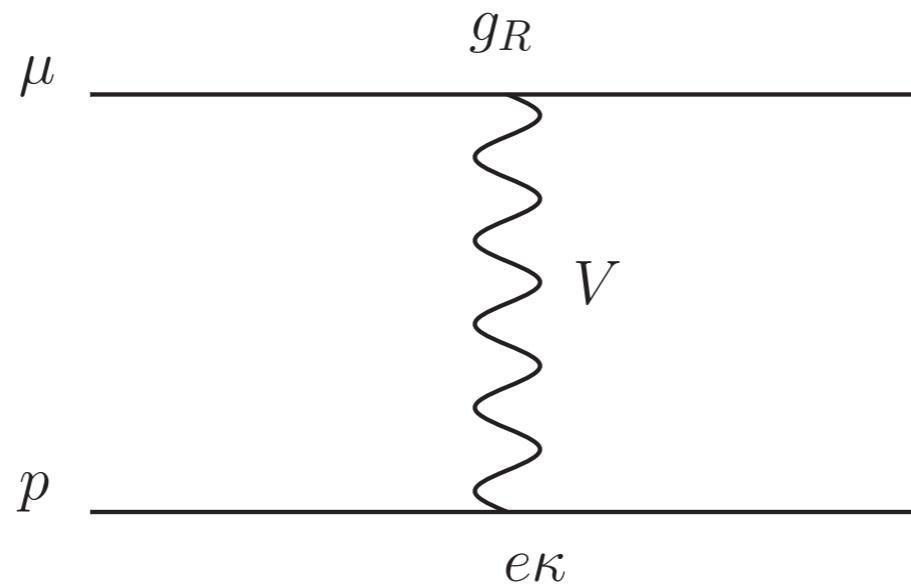
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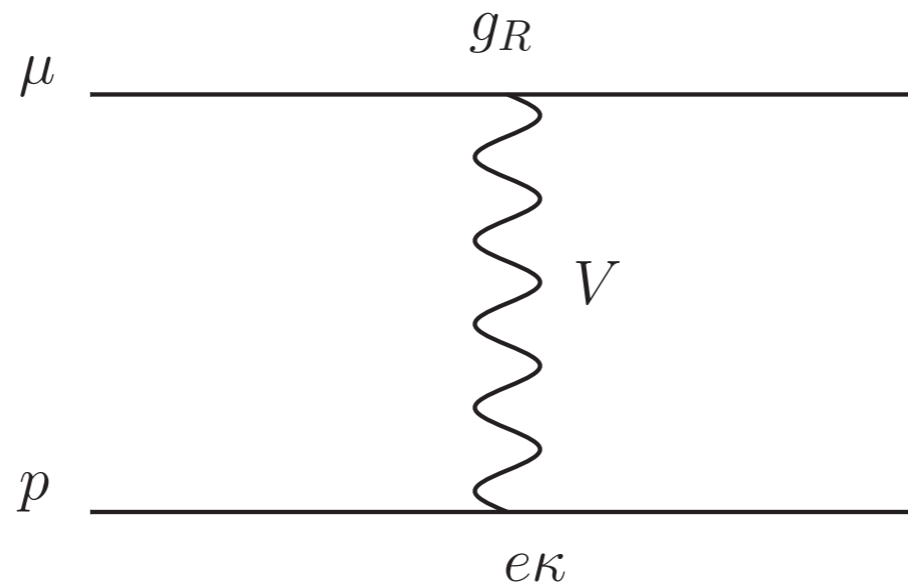
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$$K \rightarrow V \mu \nu$$

Enhanced due to axial coupling

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Higher-dim. operators for lepton masses: see Roni's talk tomorrow for effects on Higgs' properties

- Neutrino & (Quark) Flavor constraints help motivate "Stueckelberg Portal"

$$\mathcal{L} \supset -g_R V_\alpha \bar{l}_R \gamma^\alpha l_R - \frac{\kappa}{2} V_{\alpha\beta} F^{\alpha\beta}$$

- Can couple in a MFV way: look at taus in particular

- Rare decays
- 2-loop APV

$$\left( \frac{g_R \kappa}{m_V^2} \text{ held fixed} \right)$$

$$\frac{A_V}{A_Z} \simeq 10^3 \frac{m_V^2}{m_V^2 + Q^2}$$

Enhanced PV-asymmetries

- What to expect?

$$K \rightarrow V \mu \nu$$

Enhanced due to axial coupling

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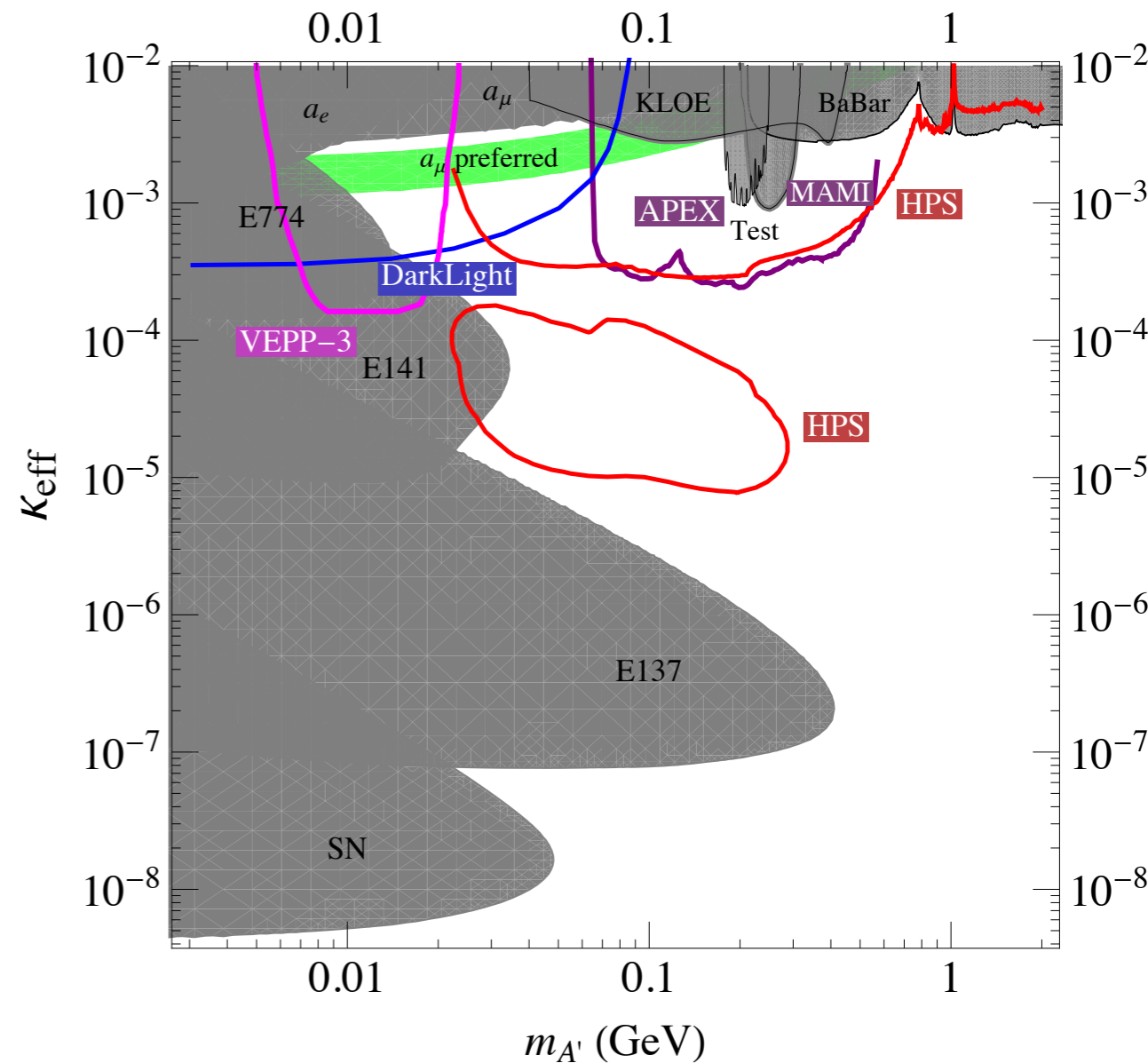
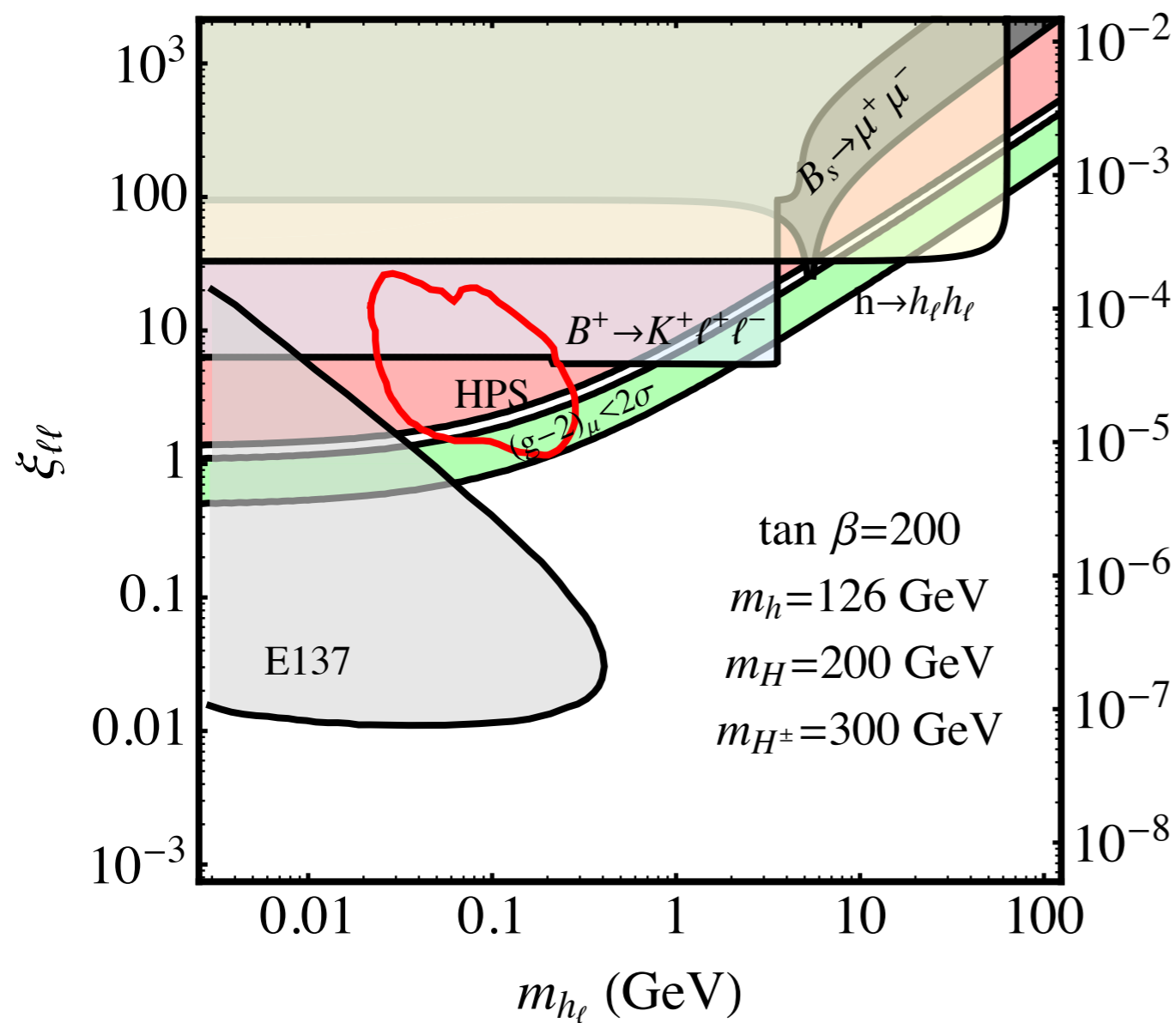
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The light scalar's interaction is:

$$-\mathcal{L}_{Y,h_\ell} \rightarrow \xi_{\ell\ell}\frac{m_\ell}{v}h_\ell\bar{\ell}\ell + \xi_{qq}\frac{m_q}{v}h_\ell\bar{q}q, \quad \xi_{qq} \sim \xi_{\ell\ell}\cot^2\beta \quad \Rightarrow \tan\beta \text{ is large}$$

# Comparison with Vector



$$eK_{\text{eff}} \equiv \xi_{ll} \frac{m_e}{v}$$

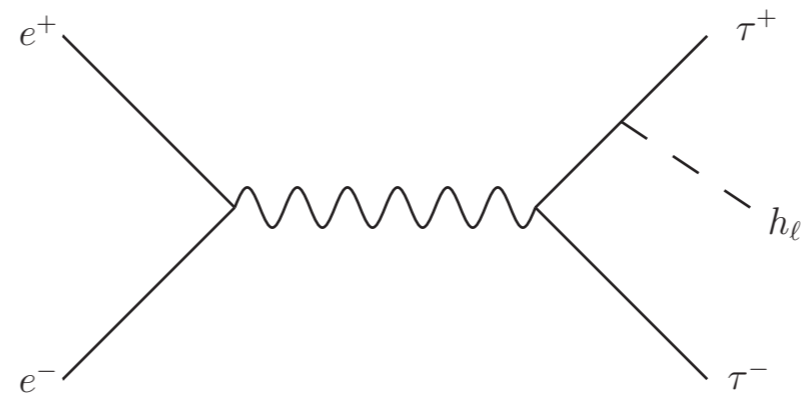
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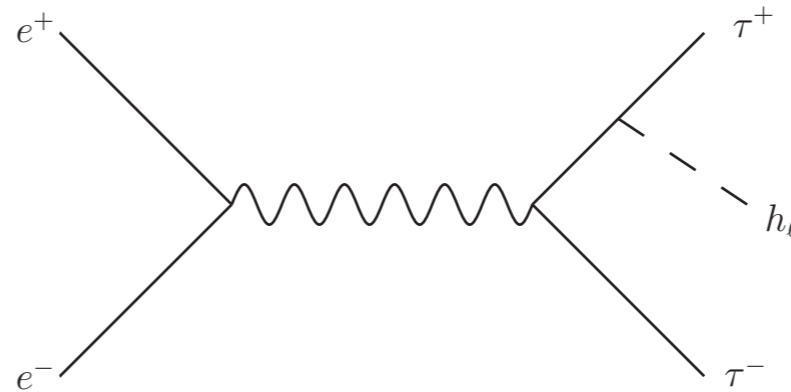
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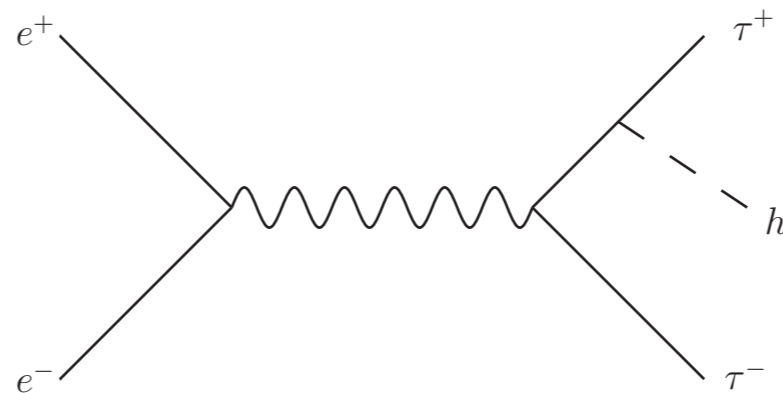
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$$N(e^+e^- \rightarrow \tau^+\tau^-h_\ell) \sim 2 \times 10^4 \left( \frac{\xi_{\ell\ell}}{2} \right)^2 \left( \frac{\int \mathcal{L} dt}{10 \text{ ab}^{-1}} \right)$$



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What else??

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- New ideas, thoughts, etc. welcome!