

Et Fermial

neutrino event generator

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40 A

40 K*



Model of Argon Reaction Low Energy Yields (MARLEY)

- Event generator for **MeV-scale** neutrino interactions
 - DUNE sensitivity to **supernova + solar neutrinos**
 - COHERENT backgrounds to CEvNS and efficiencies for inelastic cross sections
 - DUNE+SBN de-excitation y-rays in new GENIE model set ("AR23_20i_00_000")
- Current release is **version 1.2.1** (<u>marleygen.org</u>)
 - Interaction model:
 <u>Phys. Rev. C 103, 044604 (2021)</u>
 - C++14 implementation:
 <u>Comput. Phys. Commun. 269, 108123 (2021)</u>





Typical generator approach for the GeV scale

- Much attention given the needs of accelerator-based oscillation experiments



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MARLEY's treatment is designed for a different regime

- Complements without competing
- Standard two-step approach $\leq 100 \text{ MeV}$

 - Emission of de-excitation products



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Physics limitations of MARLEY v1.2.1

- High-lying nuclear energy levels form a **continuum**
 - Existing cross-section model pretends they are discrete
- "Allowed approximation" evaluates matrix elements at zero momentum transfer
 - Forbidden contributions neglected
 - Allowed cross section overestimated due to missing q dependence
- No tools for uncertainty quantification

Capability needed for future analyses





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A real continuum for MARLEY: The HF-CRPA model

- Low-energy cross sections above the nucleon knockout threshold (~8 MeV)
 - See WG2 talk by Vishvas on Thursday afternoon
 - Developed by University of Ghent group:
 <u>Phys. Rev. C 100, 055503 (2019)</u>
- Forbidden transitions become important at several tens of MeV
 - Curves other than "1+" shown in the plot
- Implemented in MARLEY using tabulated nuclear responses
 - Many thanks to Natalie Jachowicz, Alexis Nikolakopoulos, and Vishvas Pandey!





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Discrete transitions in MARLEY v1.2.1: Allowed approximation

Differential cross section can be written in the form

AA evaluates the nucleon current under two limits:

Long-wavelength limit: $q \rightarrow 0$

Slow nucleon limit:
$$\frac{|\mathbf{p}_{N_i}|}{m_N} \to 0$$

Tensor contraction evaluated using Fermi and Gamow-Teller strengths, from data where available

$$\mathbf{v}_{\mu\nu} \,\mathsf{R}^{\mu\nu} = \left(1 + \beta_{\ell} \cos \theta_{\ell}\right) B(\mathbf{F}) \\ + \left(1 - \frac{1}{3} \,\beta_{\ell} \cos \theta_{\ell}\right) B(\mathbf{GT})$$



Improved discrete treatment

- Dependence on momentum transfer $\kappa = |\mathbf{q}|$ needed for compatibility with HF-CRPA
 - \circ Current operators now evaluated to order 1/m_N
- Data-driven approach retained
 - "Old MARLEY" preserved in AA limits
 - \circ Input B(F) and B(GT) scaled by relevant form factors dependent on Q²
- jo dependence is roughly approximated in nuclear matrix element:

$$\left\langle J_f \left| \left| \sum_{k=1}^{A} j_0(\kappa r_k) t_{-}(k) \right| \right| J_i \right\rangle \simeq \left(\frac{1}{N} \sum_{k \in \text{neutrons}} j_0(\kappa r_k) \right) \left\langle J_f \right| \left| \sum_{k=1}^{A} t_{-}(k) \right\rangle \right\rangle$$

$$\mathbf{v}_{\mu\nu} \,\mathsf{R}^{\mu\nu} = F_1^2 \,\mathcal{B}_F \left[v_{CC} + \frac{\kappa}{m_N} \,v_{CL} + \frac{\kappa^2}{4 \,m_N^2} \,v_{LL} \right] \\ + F_A^2 \,\mathcal{B}_{GT} \left[\frac{\kappa^2}{12 \,m_N^2} \left(1 - 2 \,\omega \,F_{PA} + \omega^2 \,F_{PA}^2 \right) \,\phi \right]$$

$$+ \frac{\kappa}{3m_N} \left(1 - \left[\omega + \frac{\kappa^2}{2m_N} \right] F_{PA} + \frac{\omega \kappa^2}{2m_N} F_{PA}^2 \right. \\ \left. + \left(\frac{1}{3} - \frac{\kappa^2}{3m_N} F_{PA} + \frac{\kappa^4}{12m_N^2} F_{PA}^2 \right) v_{LL} \right. \\ \left. + \left(\frac{2}{3} + \frac{\kappa^2}{6m_N^2} F_{12A}^2 \right) v_T - h \frac{2\kappa}{3m_N} F_{12A} v_{T'} \right) \right]$$

v factors are dimensionless functions of lepton kinematics

FPA and F12A are expressions involving multiple nucleon form factors



Inclusive cross-section results

 $^{40}\operatorname{Ar}(\nu_e, e^-)X$



Reduction in $\sigma(E_v)$ seen at a few tens of MeV



Inclusive cross-section results



Exclusive predictions



Neutron emission reduced in new calculation = less missing energy for DUNE

SN₇ = Time-integrated supernova v_e spectrum from <u>Phys. Rev. D 97, 023019 (2018)</u>

Inclusive cross section folded with high-stats MC branching ratios to produce right-hand plot



Exclusive predictions



Still almost entirely **1n** and **y-only** final states in new calculation

••• ${}^{40}\text{Ar}(\nu_e, e^- \gamma){}^{40}\text{K v}1.2.1$ ••• 40 Ar $(\nu_e, e^- n)^{39}$ K v1.2.1 ••• ${}^{40}\text{Ar}(\nu_e, e^- p){}^{39}\text{Ar v1.2.1}$ ••• ${}^{40}\text{Ar}(\nu_e, e^- \alpha)^{36}\text{Cl v1.2.1}$ - ⁴⁰Ar($\nu_e, e^- \gamma$)⁴⁰K develop - ⁴⁰Ar($\nu_e, e^- n$)³⁹K develop - ⁴⁰Ar($\nu_e, e^- p$)³⁹Ar develop ${}^{40}\mathrm{Ar}(\nu_e, e^- \alpha){}^{36}\mathrm{Cl}$ develop

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Reweighting for uncertainty quantification

- Standard "tool of the trade" at GeV energies
 - Weights assigned according to likelihood ratio between alternate and nominal physics models
- **Systematics** in future MARLEY-based analyses
 - Efficiency / purity / energy resolution for supernova ve detection in DUNE
 - Neutrino backgrounds for MeV-scale exotic physics searches
- Some exploration of uncertainties already
 - No technical infrastructure in MARLEY v1.2.1
 - $\circ \sigma(E_v)$ uncertainty not rigorously quantified yet





MARLEY nuclear de-excitation model

Hauser-Feshbach formalism for computing differential decay widths:



Phenomenological ingredients based on global fits to nuclear reaction data

Level density model: Back-shifted Fermi gas Nucl. Data Sheets 110, 3107–3214 (2009)

Nuclear optical model potential: Koning & Delaroc Nucl. Phys. A 713, 231–310 (2003)

Gamma-ray strength function model: Standard Lore Nucl. Data Sheets 110, 3107–3214 (2009)



che,	Supplemented with tabulated discrete leve γ -rays for bound states (taken from TALYS
ntzian	Transitions from continuum to all accessibl levels are explicitly treated.







Optical model potential uncertainties

- Original KD fit did not report these
- Revisited by Pruitt et al. (LLNL) in 2023
 - Motivated by applications to FRIB, etc.
 - KDUQ optical potential includes covariances between all 46 empirical parameters
- How do these impact calculations of exclusive O(10 MeV) neutrino cross sections?
 - Expected to be sub-leading at present
 - Well-justified by data, related uncertainties (level densities, etc.) under development
 - Starting point for MARLEY uncertainty toolkit



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<u>Phys. Rev. C 107, 014602 (2023)</u>







Optical model potential uncertainties





y-only cross section dominated by discrete levels. OMP uncertainty nonzero but tiny!

1α has modestly higher fractional uncertainty → sensitive to both p and n OMP parameters







Towards MARLEY v2.0.0

- Many technical enhancements in addition to physics topics addressed in this talk
- NuHepMC: Universal neutrino generator event format
 Adopted internally and in MARLEY output
 - Directly enables reweighting, other new features
 - $\circ\,$ Automated, generator-agnostic analysis of results
 - Could enable interoperability
 (e.g., delegate de-excitations to MARLEY)
- mardecay: Standalone simulation of de-excitation step only
- marreweight: Apply new weights to existing MC sample
- Speedup of optical model calculations (~30x!)
- New major release (v2) anticipated in the coming months \circ Development team remains small (\ll 1 FTE)
 - New collaborators are always welcome! (gardiner@fnal.gov)







Backup

