

New results and status of the JSNS² / JSNS²-II

NuFact 2024

The 25th International Workshop on Neutrinos from Accelerators

Sep 16-21, 2024 Argonne National Laboratory

Dongha Lee (KEK)



JSNS² / JSNS²-II Collaboration

(J-PARC Sterile Neutrino Search at J-PARC Spallation Neutron Source)



KEK
JAEA
J-PARC
Tsukuba University
Osaka University
Tohoku University
Kitasato University
Kyoto Sangyo University



Soongsil University
Dongshin University
Seoyeong University
Kyung Hee University
Gwangju Institute of Science and Technology
Seoul National University of Science and Technology
Sungkyunkwan University
Chonnam National University
Jeonbuk National University
Kyungpook National University



Brookhaven National Laboratory
University of Michigan
University of Utah



Sun Yat-sen University



University of Sussex

5 countries
23 institutions
~60 collaborators

JSNS² / JSNS²-II Papers in recent

- Y.Hino et al, “Characterization of the correlated background for a sterile neutrino search using the first dataset of the JSNS² experiment”, Eur. Phys. J. C (2022) 82: 331 (arXiv:2111.07482 [hep-ex]) (Commissioning runs)

Correlated Bkg

- C.D.Shin et al, “The acrylic vessel for JSNS²-II neutrino target”, 2023 JINST 18 T12001 (arXiv:2309.01887 [hep-ex])

JSNS²-II

- D.H.Lee et al, “Study on the accidental background of the JSNS² experiment”, Eur. Phys. J. C 84, 409 (2024) (arXiv:2308.02722 [hep-ex])

Accidental Bkg

- T.Dodo et al, "Pulse Shape Discrimination in JSNS²", arXiv:2404.03679 [physics.ins-det] (Under review @ PTEP)

PSD

- D.H.Lee et al, "Evaluation of the performance of the event reconstruction algorithms in the JSNS² experiment using a 252Cf calibration source", arXiv:2404.04153 [hep-ex] (under review @ NIMA)

**Calibration/
Reconstruction**

- E.Marzec et al, "First Measurement of Missing Energy Due to Nuclear Effects in Monoenergetic Neutrino Charged Current Interactions" (arXiv:2409.01383 [hep-ex], submitted to PRL)

KDAR

Indication of a sterile neutrino ($\Delta m^2 \sim 1 \text{eV}^2$)

- Anomalies, which cannot be explained by standard neutrino oscillations for a few tens years are shown.

Experiments	Neutrino source	Signal	Significance	E (MeV)	L (m)
LSND	μ Decay-At-Rest	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	3.8σ	40	30
MiniBooNE	π Decay-In-Flight	$\nu_\mu \rightarrow \nu_e$	4.8σ	800	600
		$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$			
BEST	e capture	$\nu_\mu \rightarrow \nu_x$	4.2σ	<3	10
Reactors	Beta decay	$\bar{\nu}_\mu \rightarrow \bar{\nu}_x$	3.0σ	3	10-100

- JSNS² uses the same neutron source(μ), target(H) and the detection principle (IBD) as the **LSND**.
=> Even if the excess is not due to the oscillation, we can catch this directly.
=> Two advantages : short-pulsed beam / Gd-LS give excellent S/N ratio.

JSNS²: J-PARC E56

JSNS²-II: E82

Sterile ν search

@MLF

<http://research.kek.jp/group/mlfnu/eng>

J-PARC Facility (KEK/JAEA)

South to North

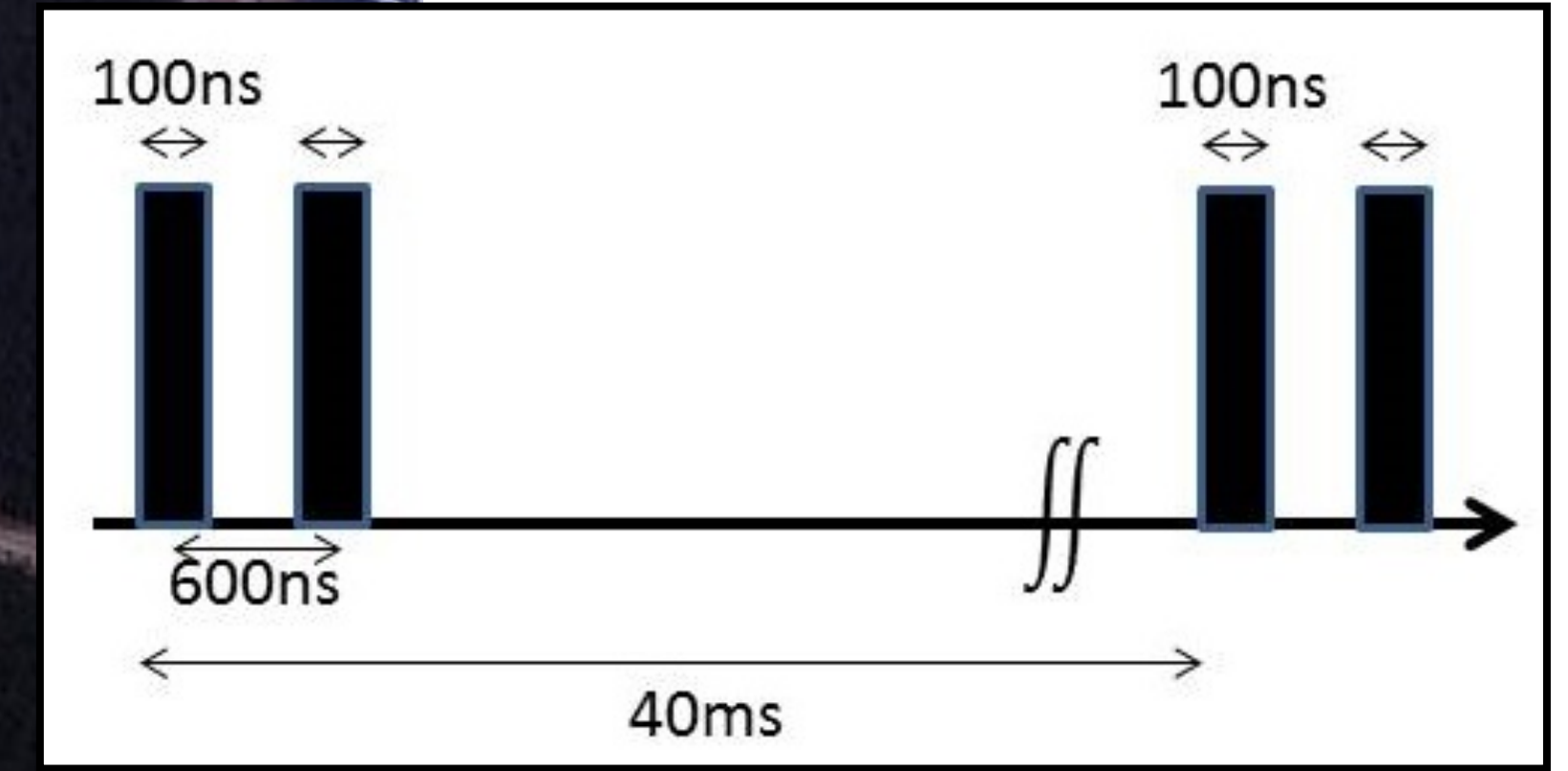
3GeV RCS

Neutrino Beams
(to Kamioka)

Materials and Life
Science Experimental
Facility (MLF)

30GeV MR

Hadron hall



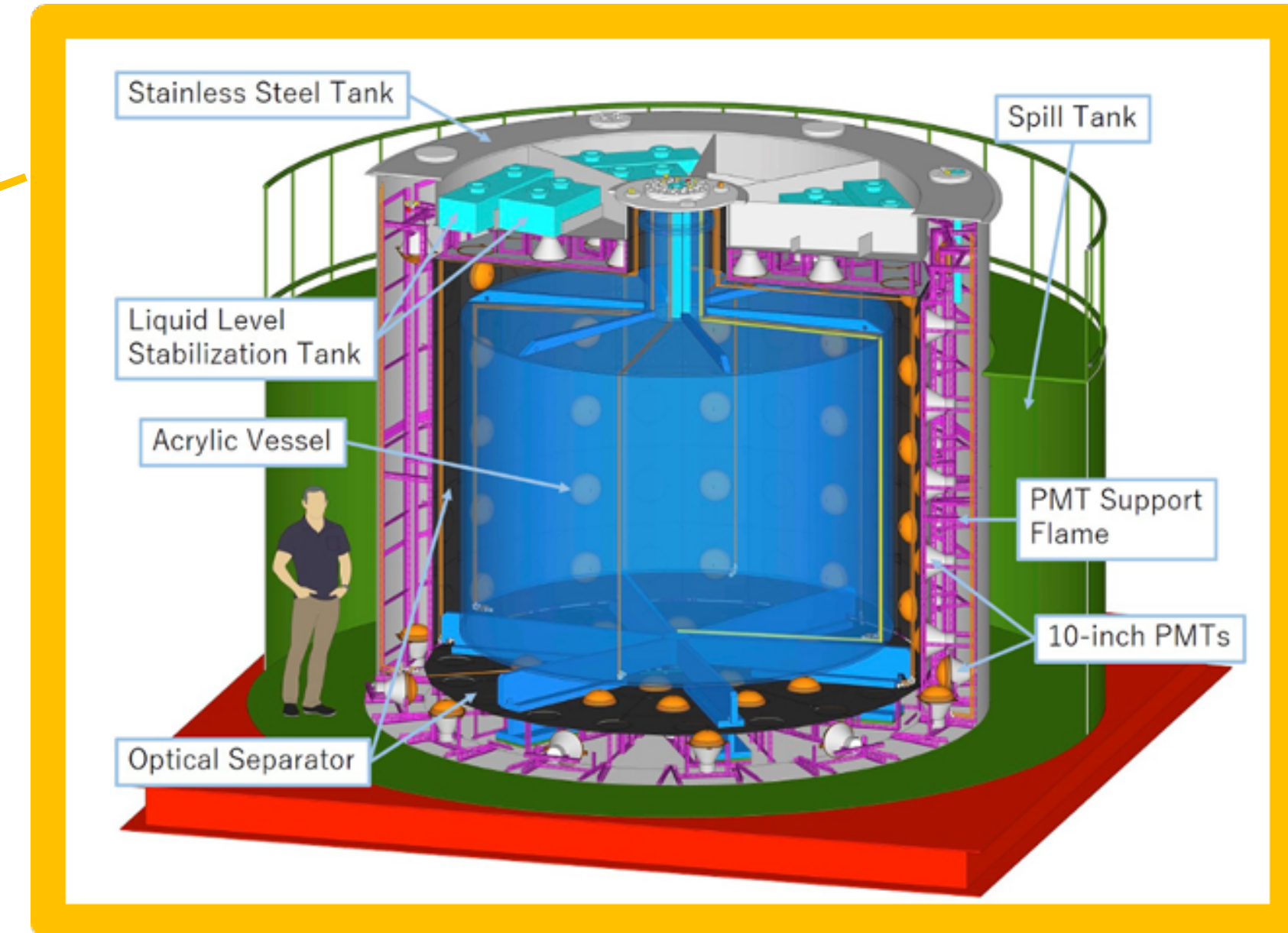
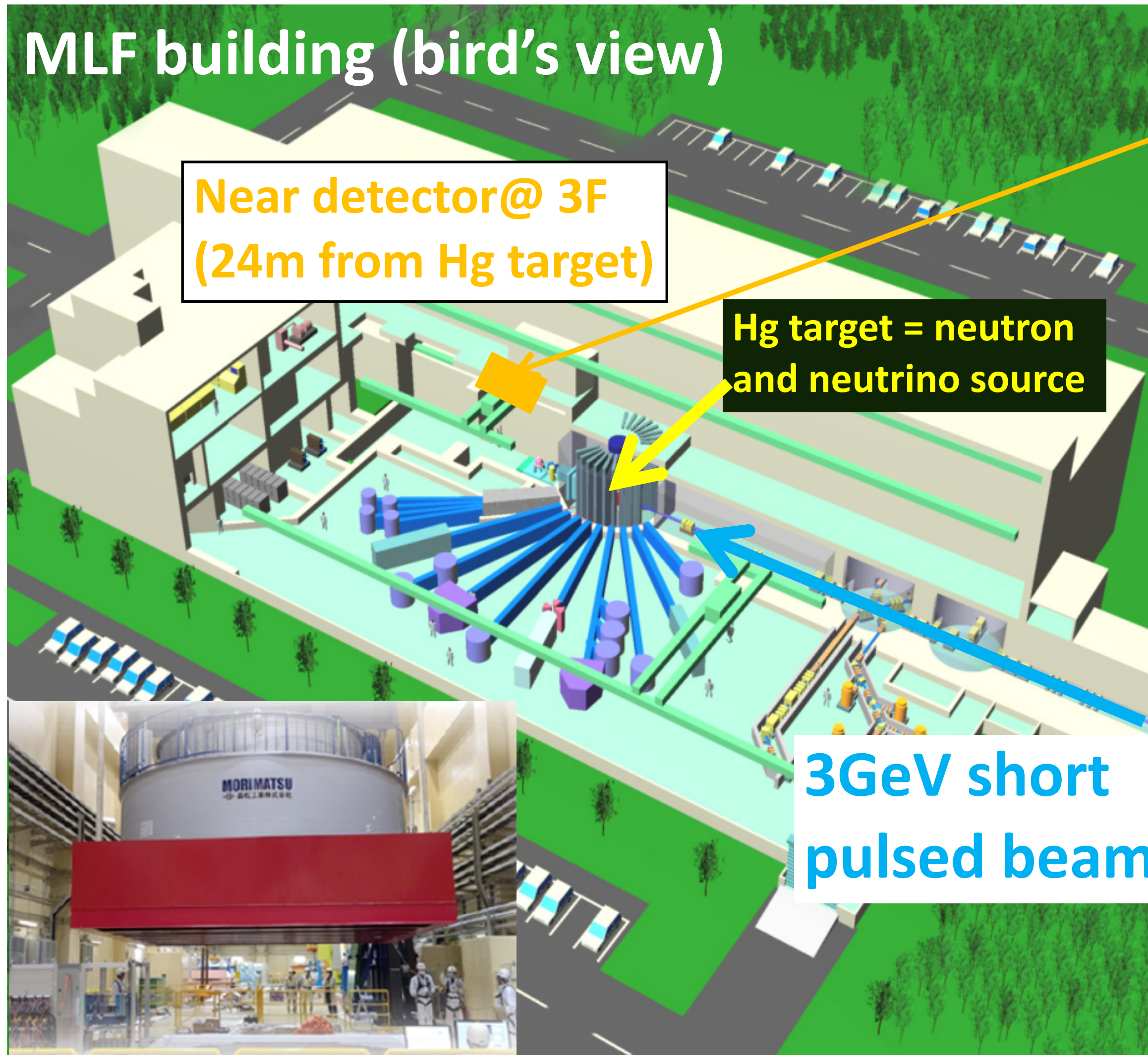
Low duty factor beam
(short pulse +
small repetition rate)
gives excellent S/N ratio.

25Hz, 1MW (design)

- 0.6-0.7MW (2021)
- 0.7-0.8MW (2022)
- 0.84MW (2023)
- 0.73-**0.95MW** (2024)

@MLF

JSNS² detector



Nucl. Instrum. Methods A 1014 165742 (2021)

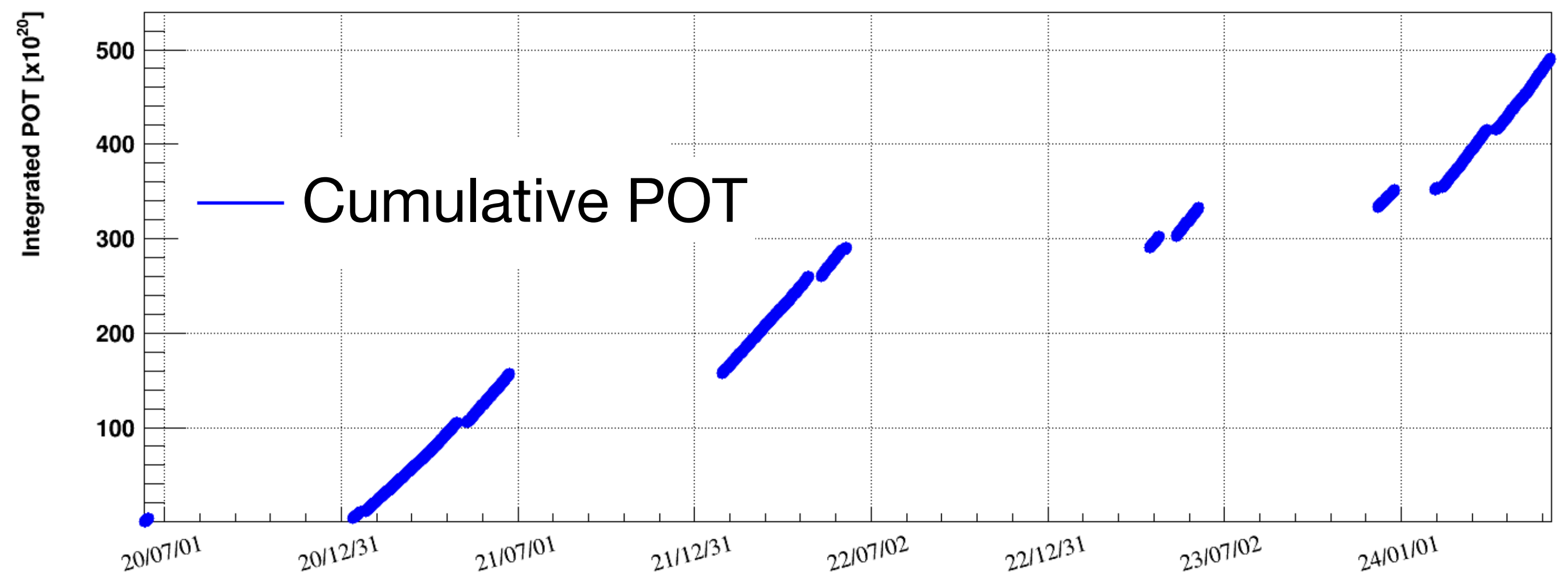
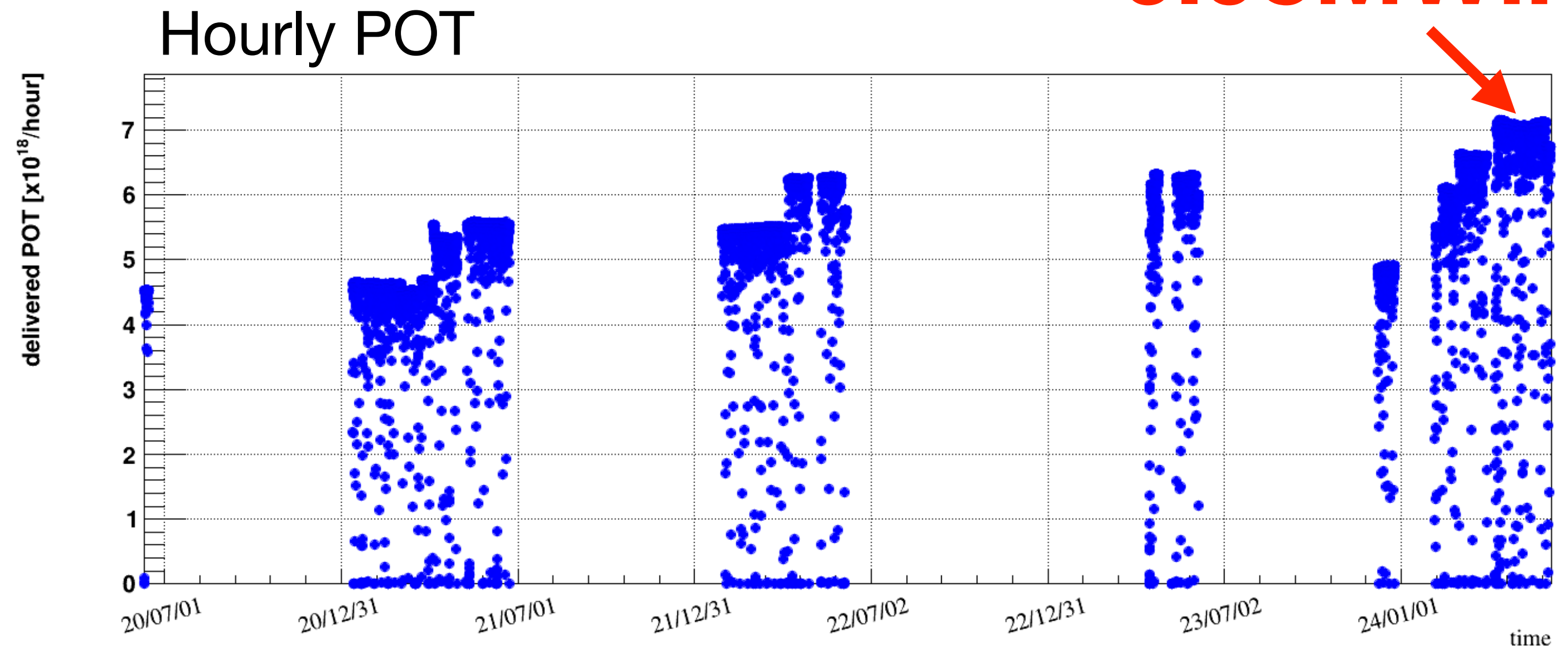
	Liquid	Volume
Target	Gd-LS + DIN(10%)	17 tons
Gamma-catcher & Veto	pure LS	31 tons

- 96 10" PMTs for the Inner Detector
- 24 10" PMTs for Veto

Operation of Near Detector

0.95MW!!

- Data taking
 - Commissioning (2020)
 - 4 long term physics run (2021-2024)
- 1MW beam power (Design)
 - Achieve 1MW @ RCS extraction point
 - 0.95MW @MLF
- 4.85×10^{22} POT so far
 - 43% of approved POT

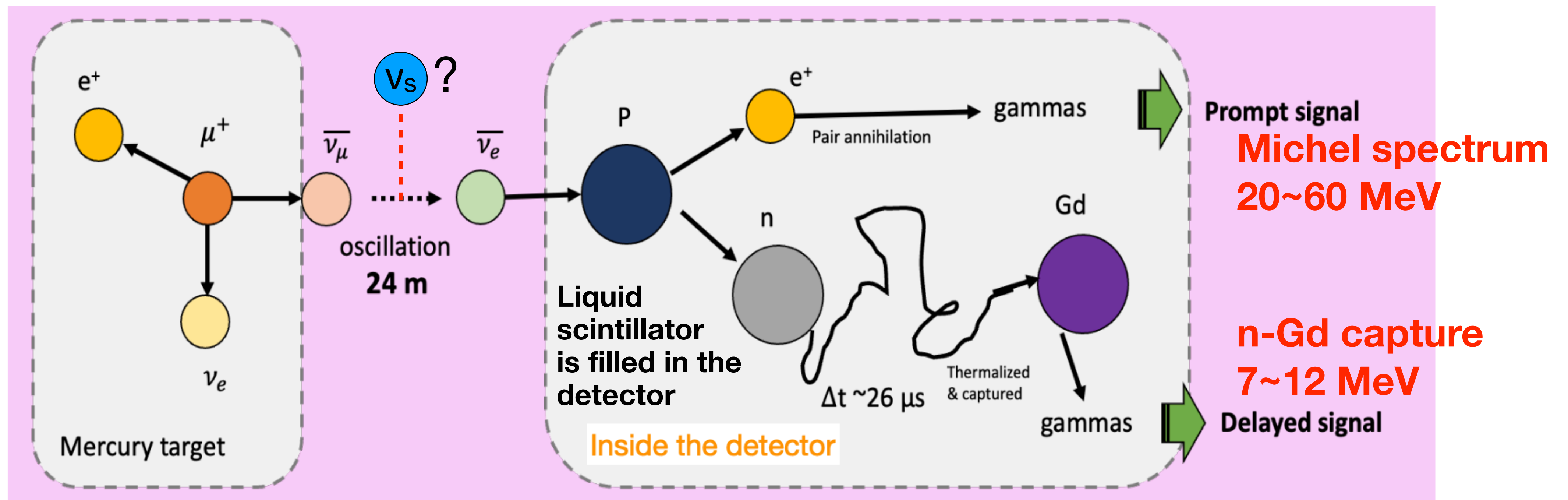


How to detect sterile neutrino signal

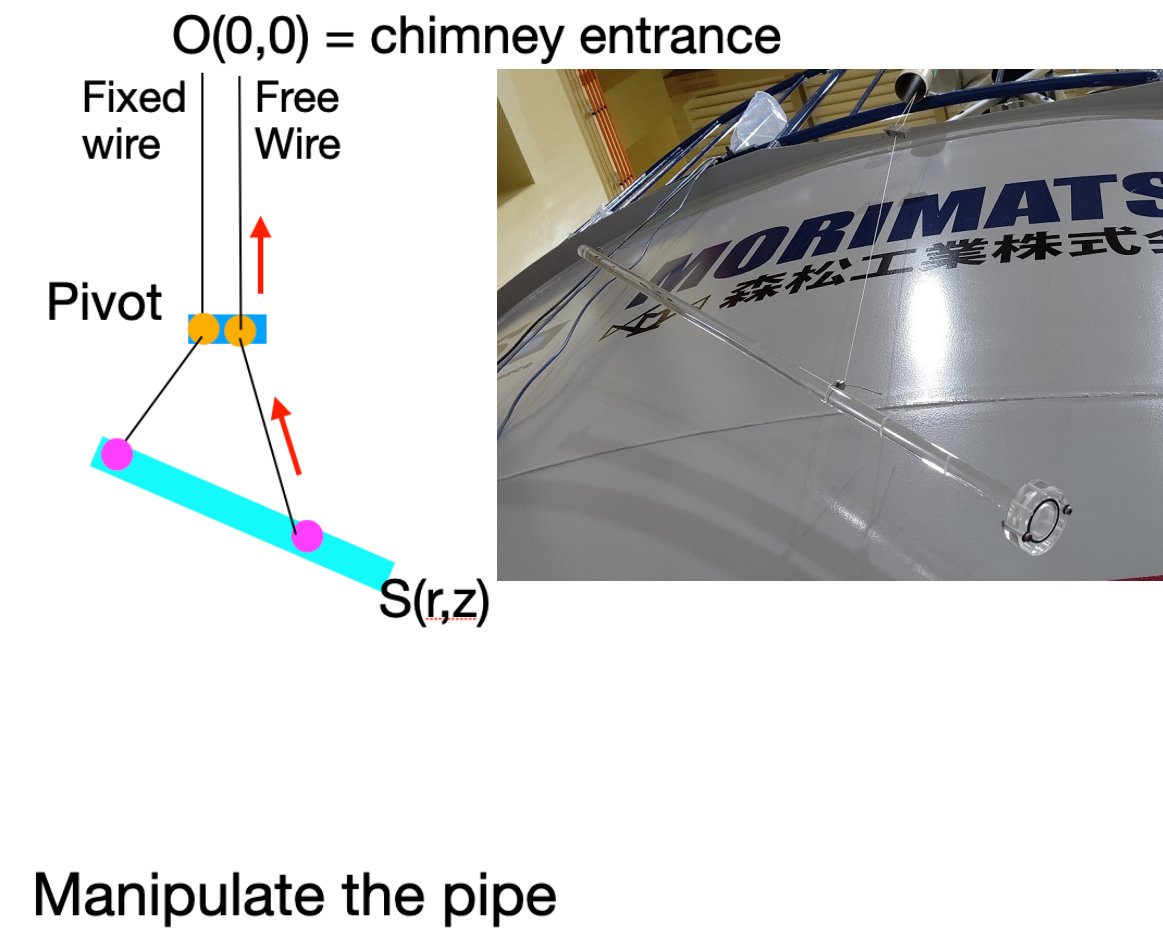
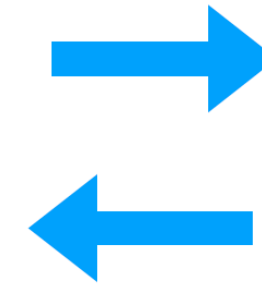
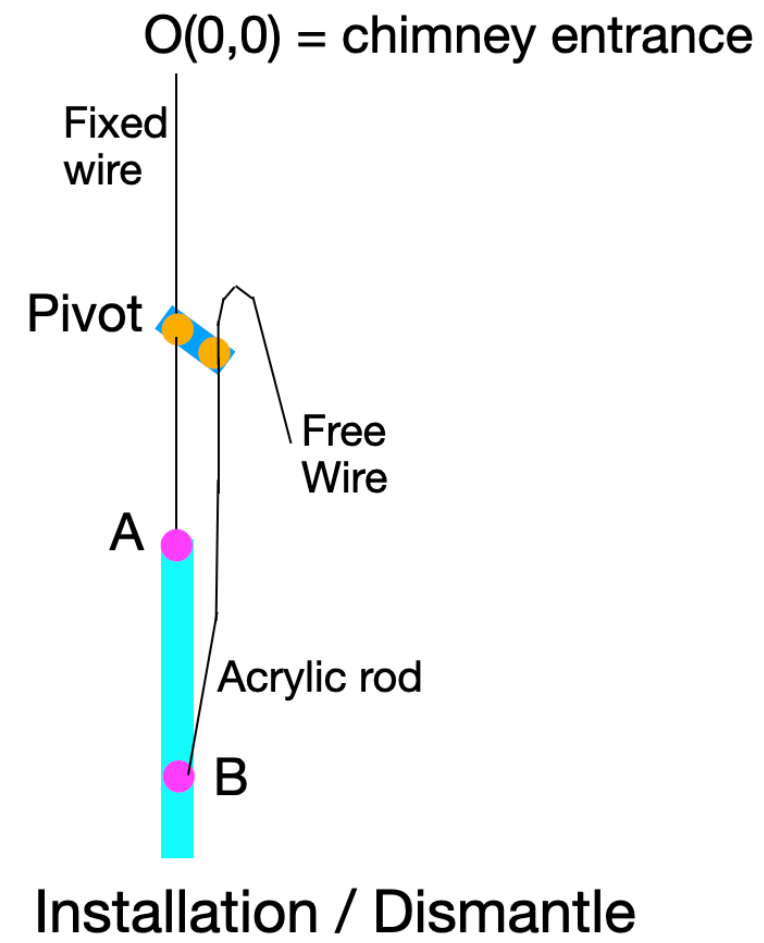
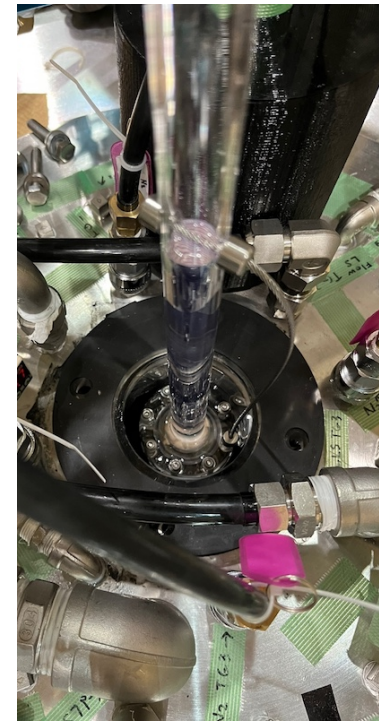
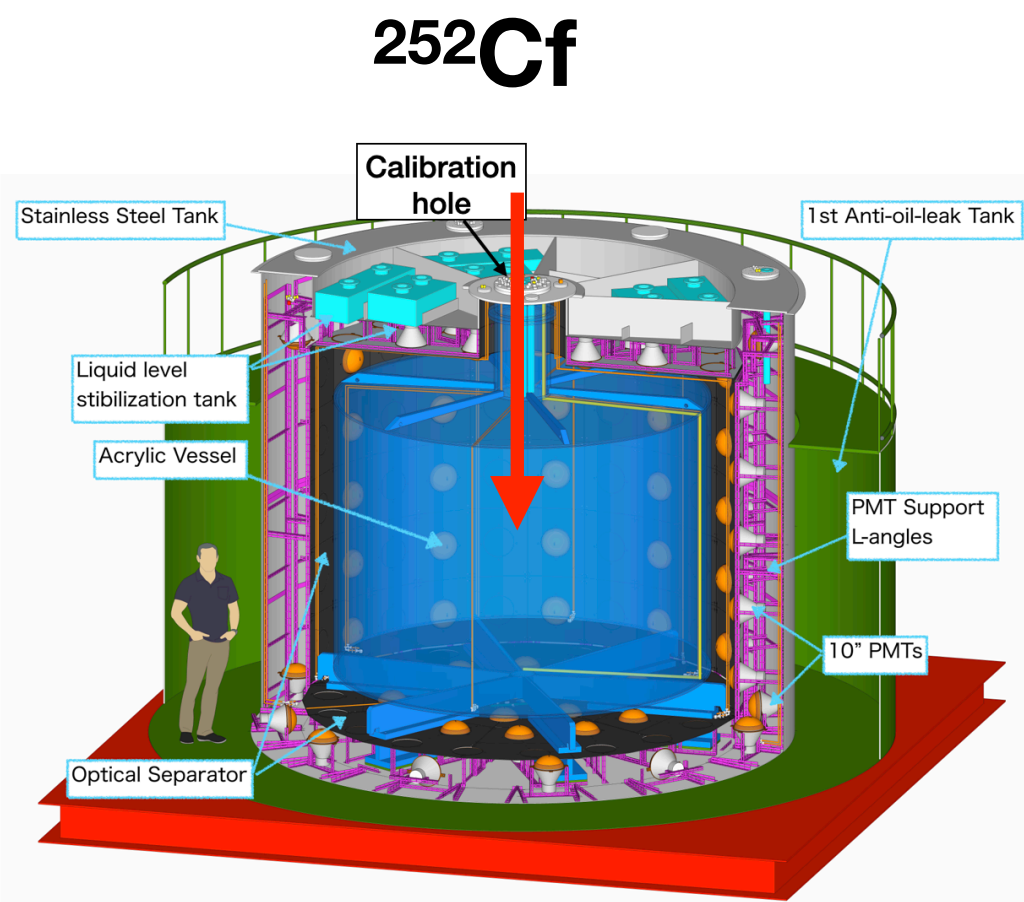
- **Sterile neutrino oscillation** during 24m baseline.
- Detection by coincidence of IBD (Inverse Beta Decay)
=> Prompt : e+ signal
=> Delayed : gammas from neutron capture on gadolinium (Gd)

	Timing	Energy
Prompt	$2 < T_{\text{beam}} < 10 \mu\text{s}$	20~60 MeV
Delayed	$\Delta T_{\text{p-d}} < 100 \mu\text{s}$	7~12 MeV

+ Spatial correlation cut
 $\Delta V_{\text{TX}_{\text{p-d}}} < 60\text{cm}$



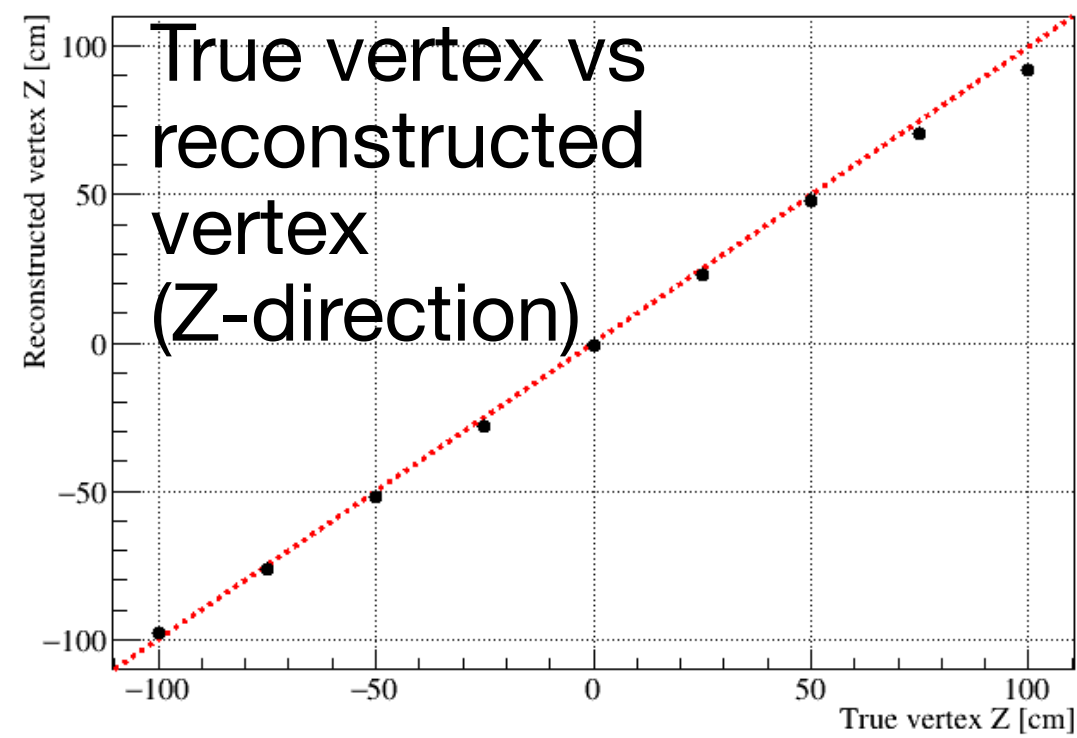
^{252}Cf 3D calibration : Energy/Vertex reconstruction



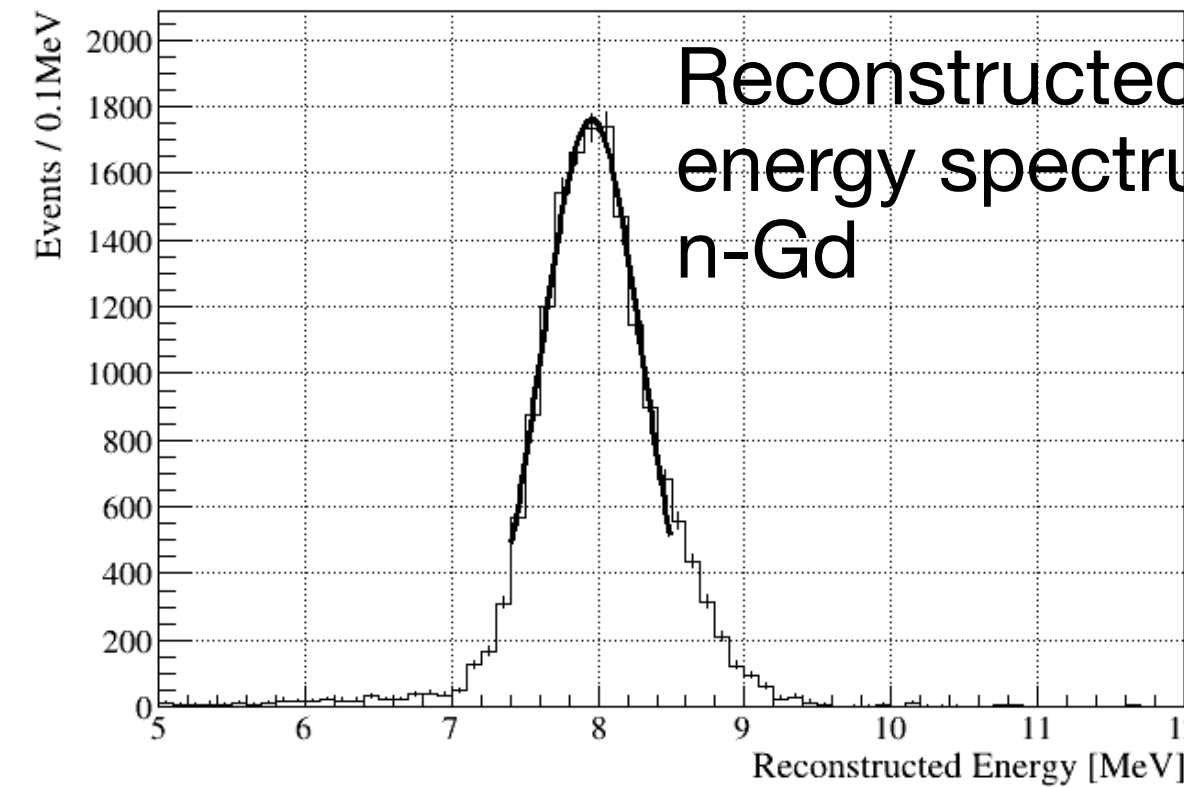
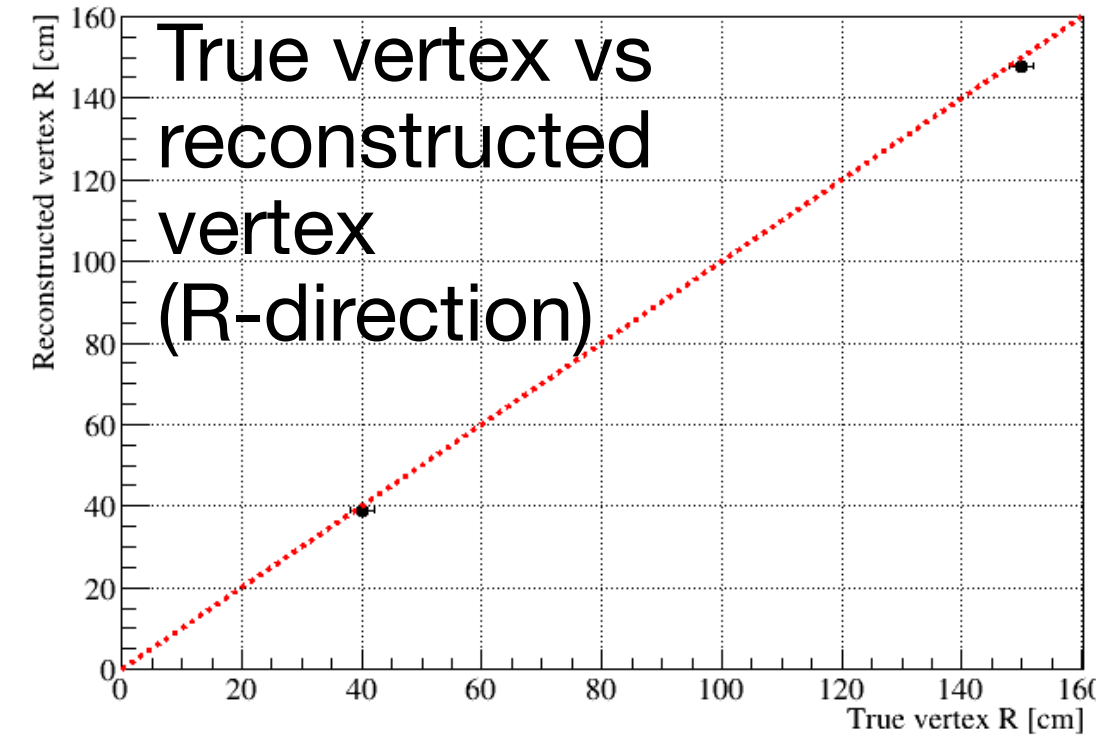
(NEW)
 ^{252}Cf 3D calibration system
 -> **Any of position (R&Z)** can be covered!!!

arXiv:2404.04153
 [hep-ex],
 under review @ NIMA

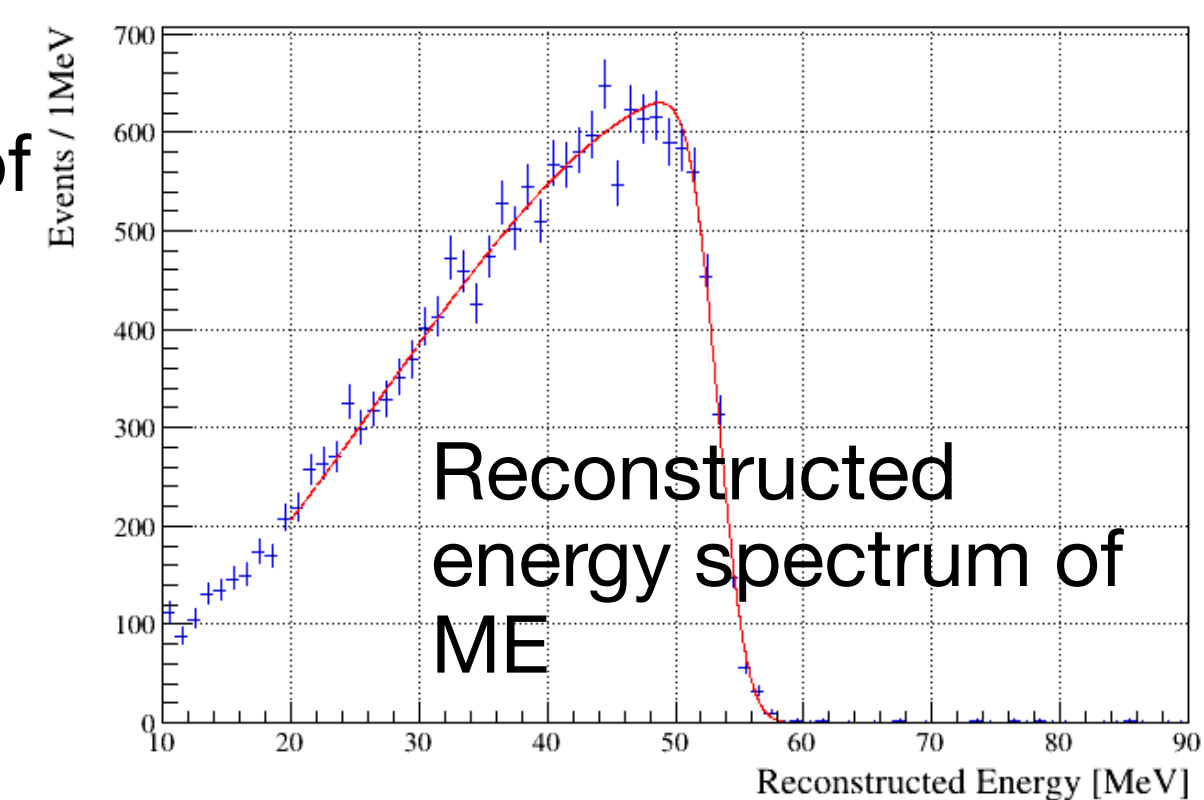
- Reconstruction Parameters were studied based on this calibration
 => Reconstruction performances were checked using ^{252}Cf and Michel electrons data sample



~10% fiducial volume uncertainty



~4% E resolution @ 8MeV of n-Gd



~3% E resolution @ ~53MeV of ME

Toward the sterile neutrino search (Blind analysis)

- PSD : arXiv:2404.03679 [physics.ins-det], under review @PTEP
- Accidental background : Eur. Phys. J. C 84, 409 (2024)

Toward the sterile neutrino search (Blind analysis)

- We are doing the blind analysis to search a sterile neutrino, by using the energy side-bands.
- The rates in the side-band regions will be predicted by the control samples driven by data
- **Now, side-band 4 (prompt 60-100MeV) data are opened!!!**
- Side-band 1&3 are being studied now.

Observation

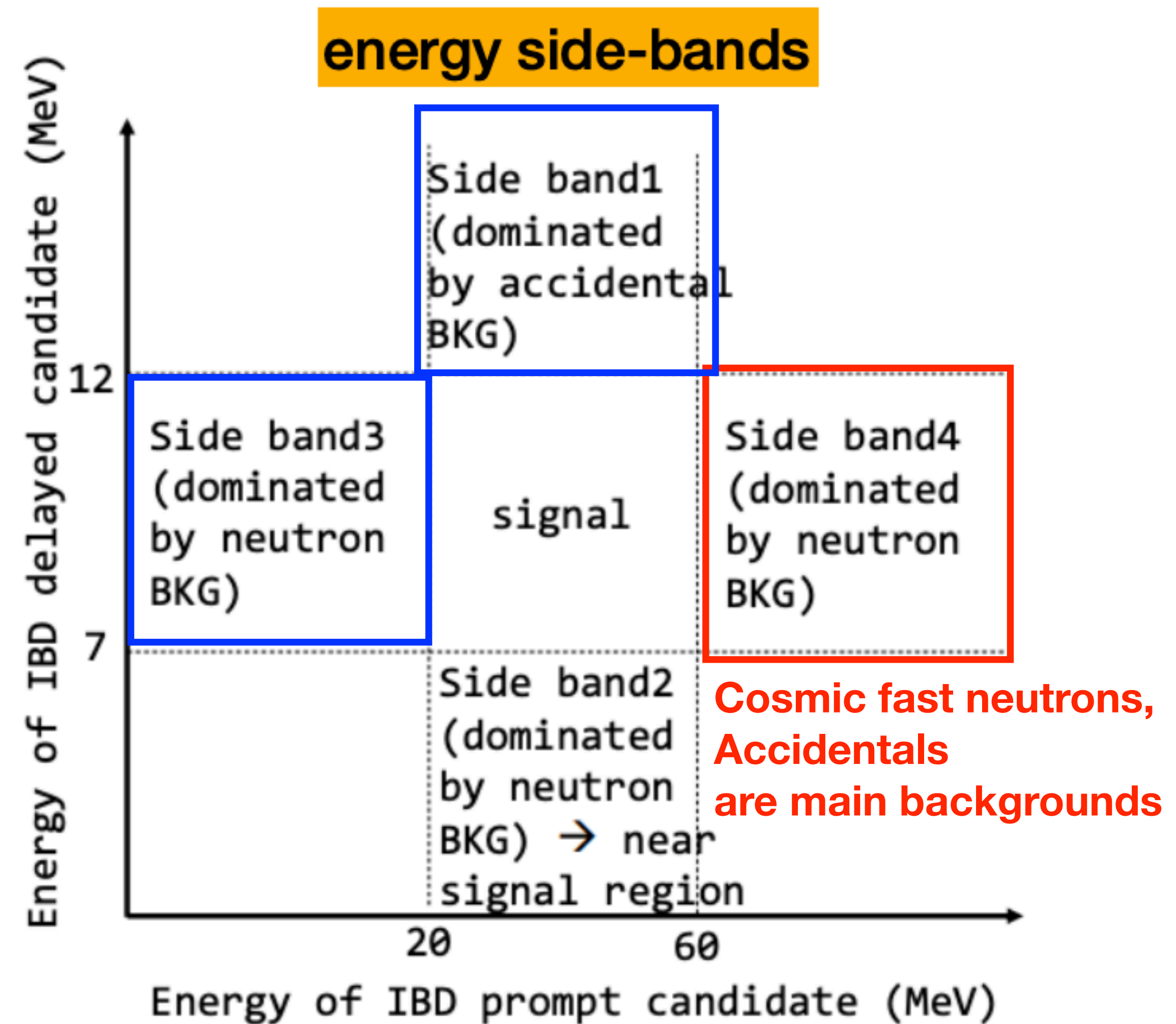
VS

Expectation

Physics data with criteria :
 $xx < E_p < yy$ MeV
 $zz < E_d < ww$ MeV
 $\Delta T_{p-d} < 100 \mu s$
 $\Delta VTX_{p-d} < 60 \text{cm}$

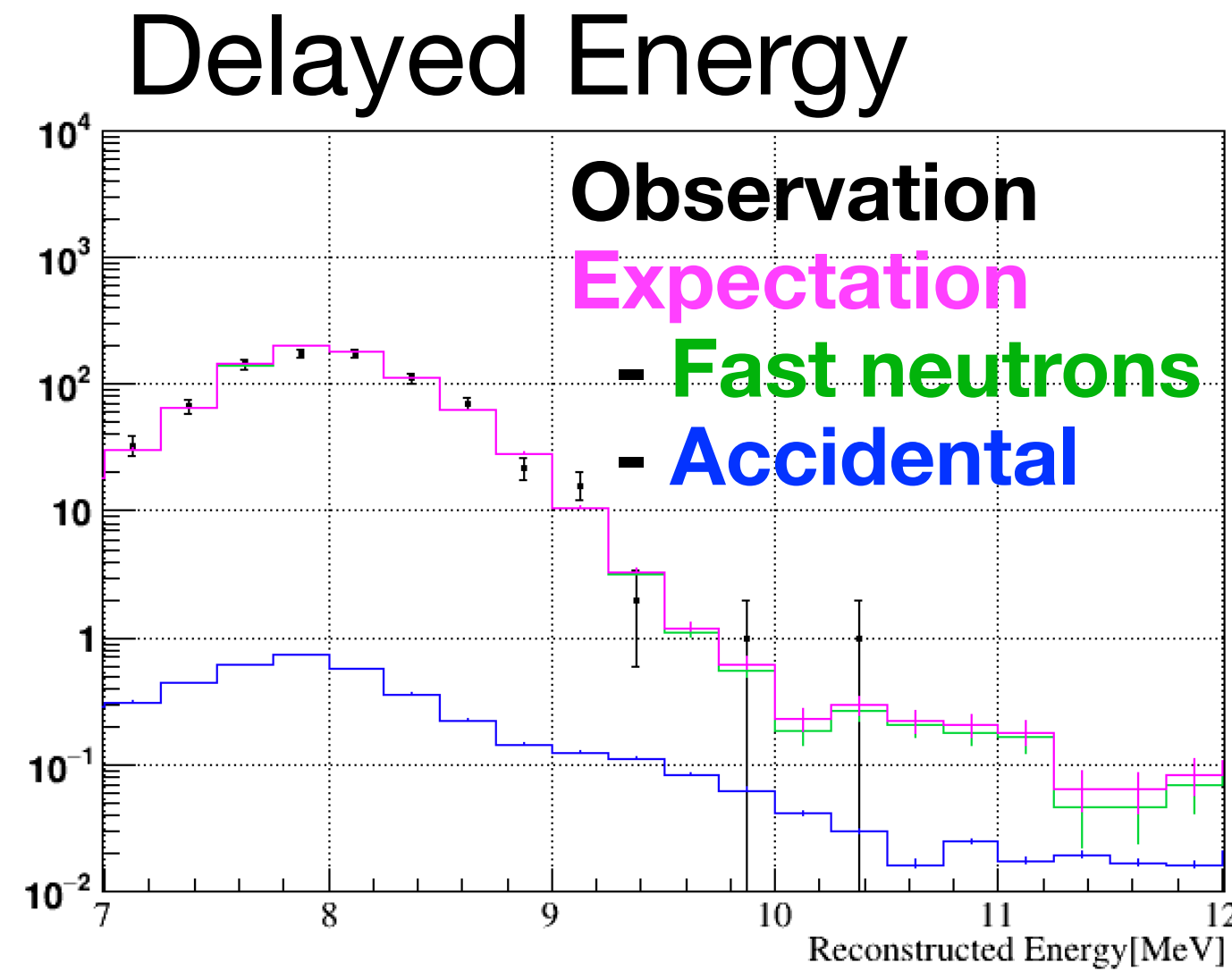
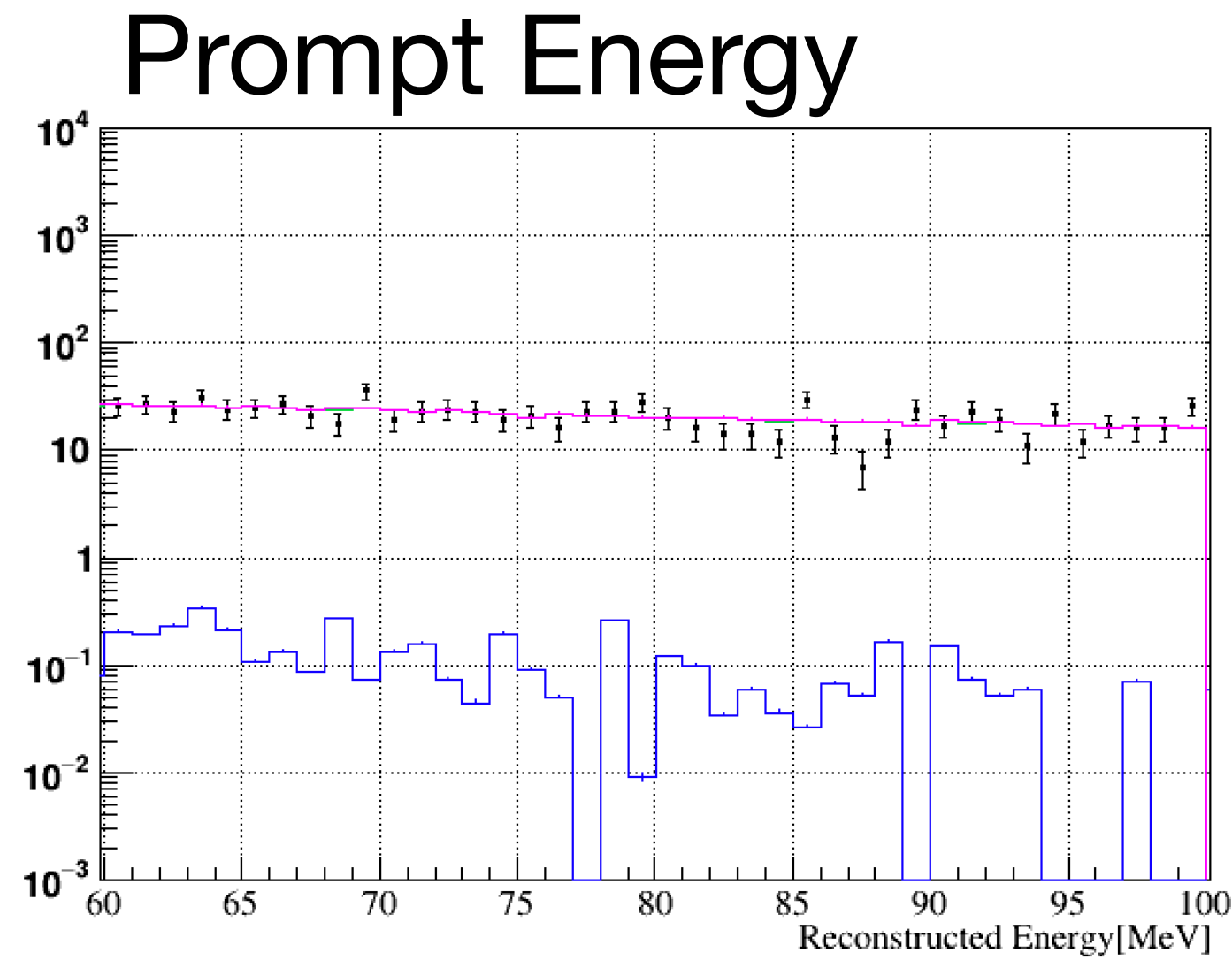
Purely data-driven control sample

- **Fast neutrons**
-> obtained at $T_{\text{beam}} > 1 \text{ms}$
- **Accidental background**
-> obtained with specific calibration runs

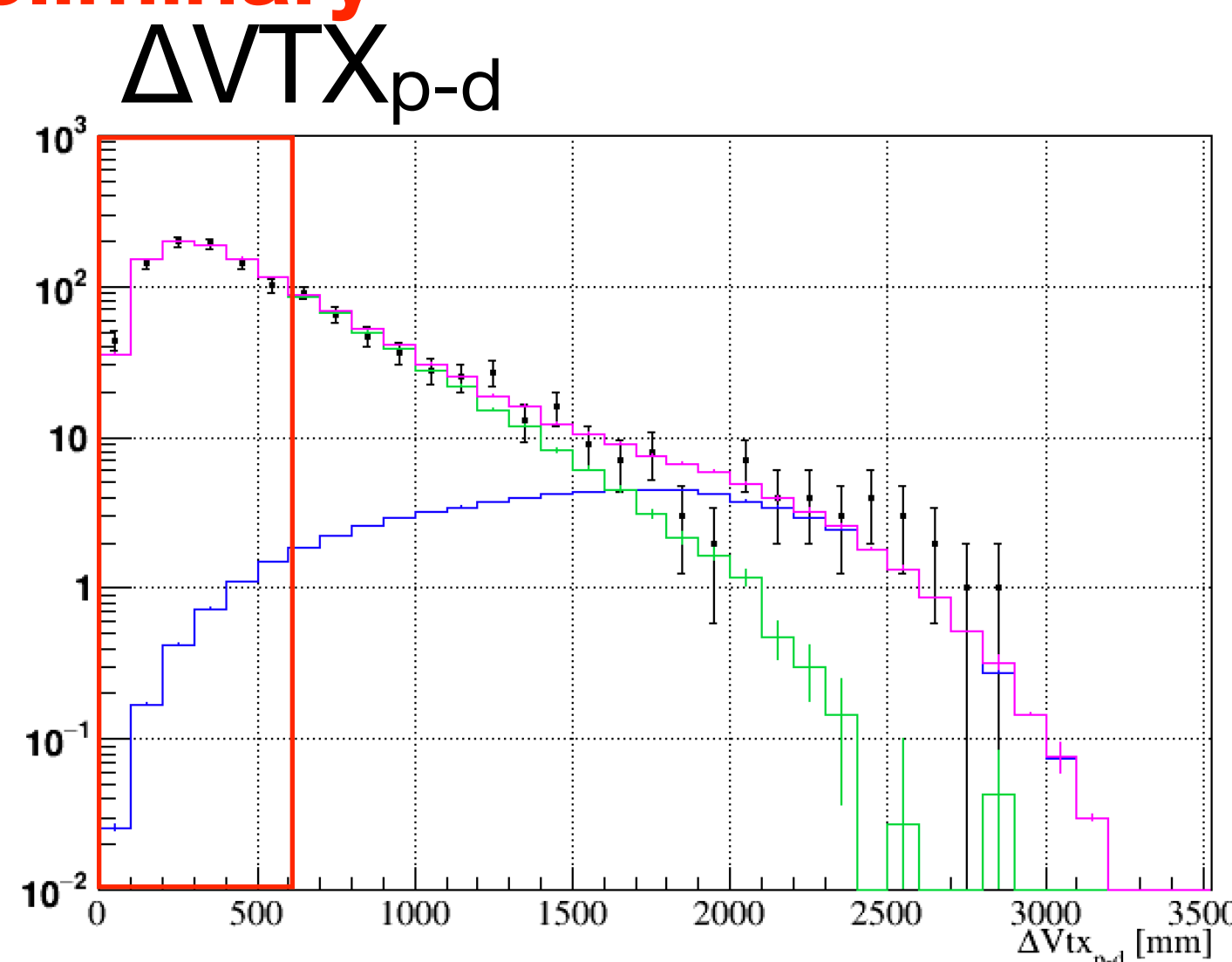
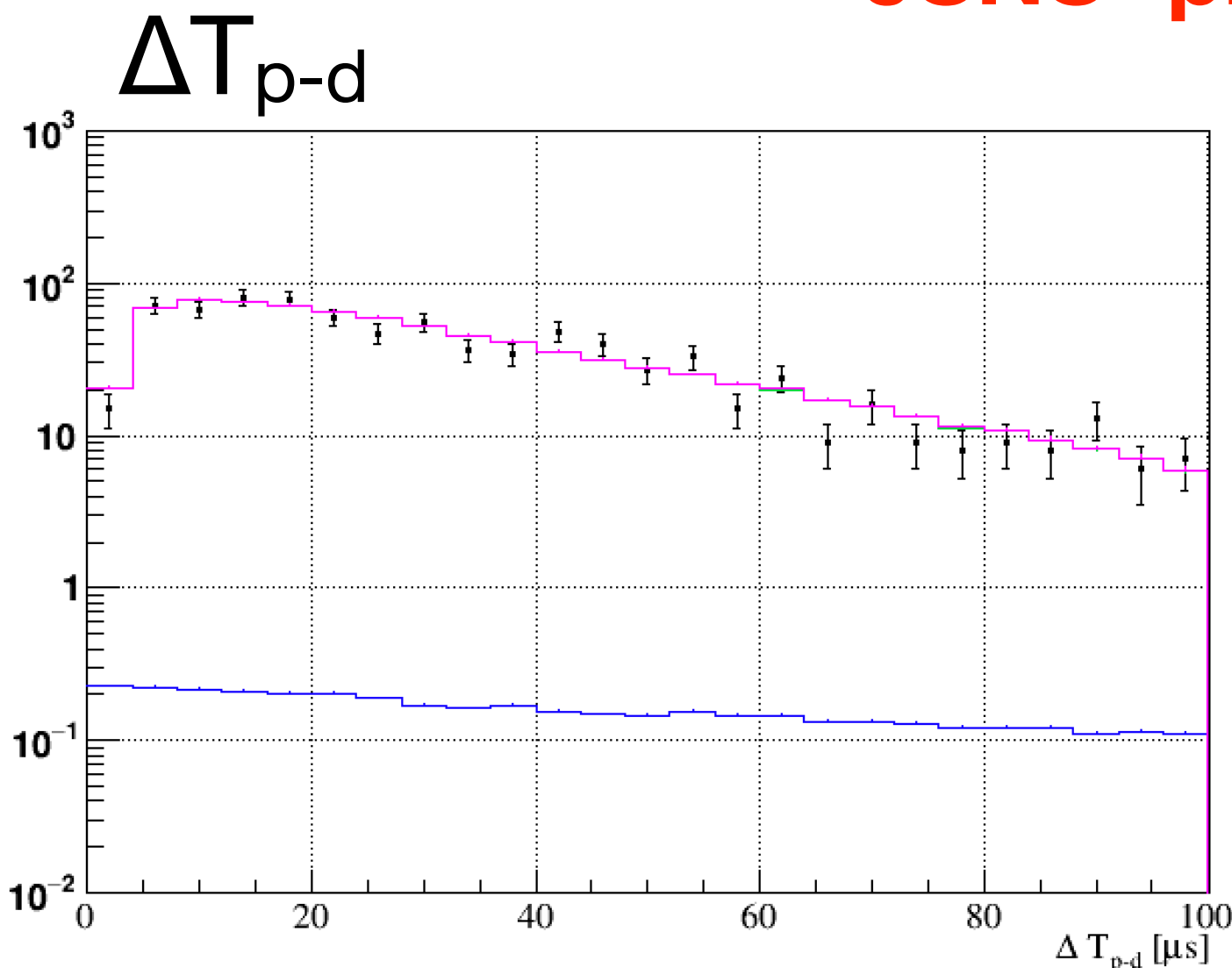


The comparison between the observation vs expectation for Side-band4 (prompt 60~100MeV)

*2022 physics data (0.8×10^{22} POT)



JSNS² preliminary



	# of Observation events	# of Expectation events
Total	818 \pm 28.6 (stat) \pm 6.0 (sys)	839.3 \pm 3.0 (stat) \pm 2.6 (sys)
Fast neutrons		835.3 \pm 3.0 (stat) \pm 2.5 (sys)
Accidental		4.0 \pm 0.1 (stat) \pm 0.07 (sys)

• **Good consistency between Observation vs Expectation !!!**

• Cosmogenic fast neutrons are dominated
 \Rightarrow Aim to $\sim 99\%$ rejection
 \Rightarrow **Pulse Shape Discrimination (PSD)** would reject most of them
(arXiv:2404.03679 [physics.ins-det])

• After applying PSD, **accidental bkg will be also crucial.** (Eur. Phys. J. C 84, 409 (2024))

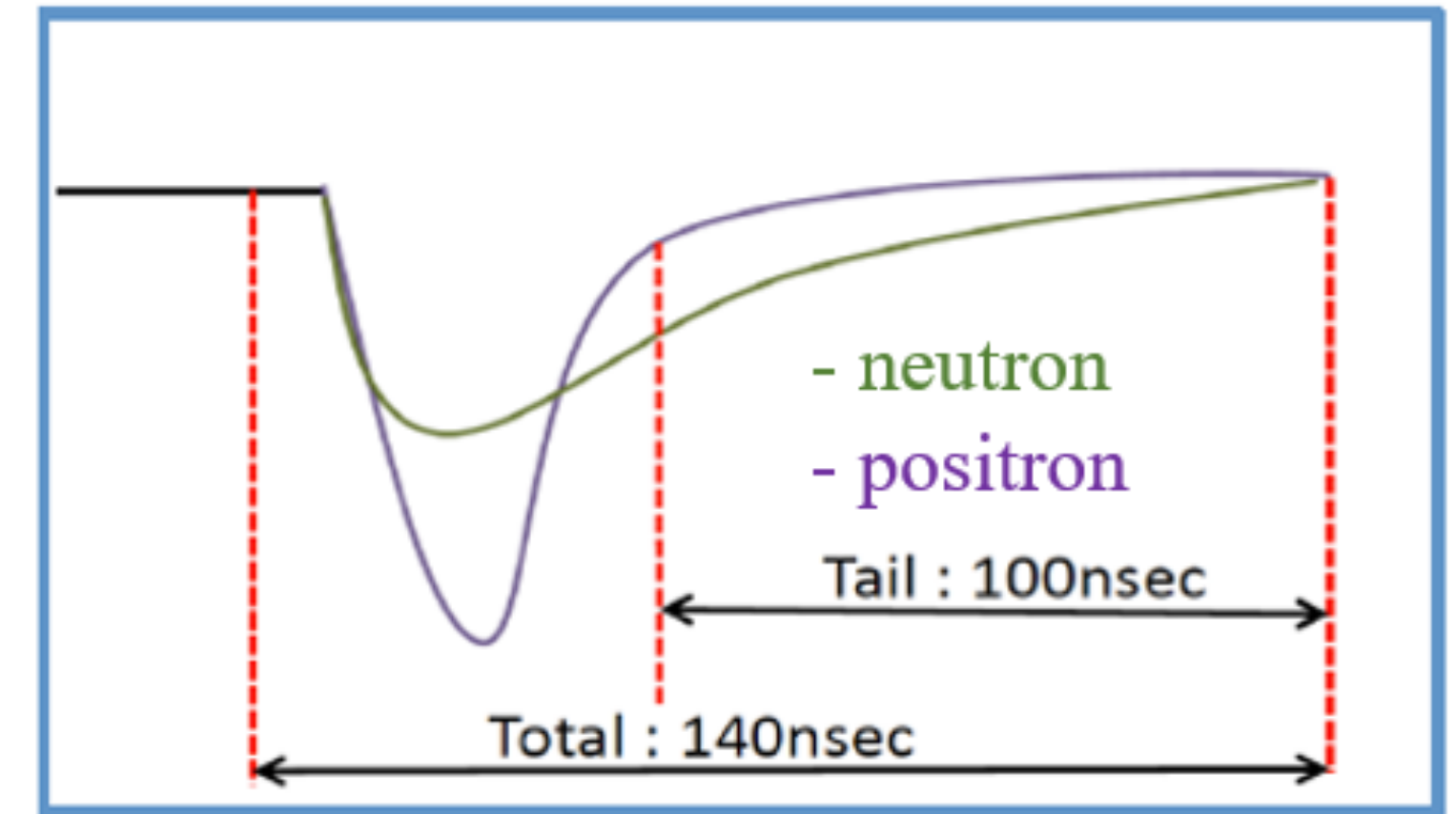
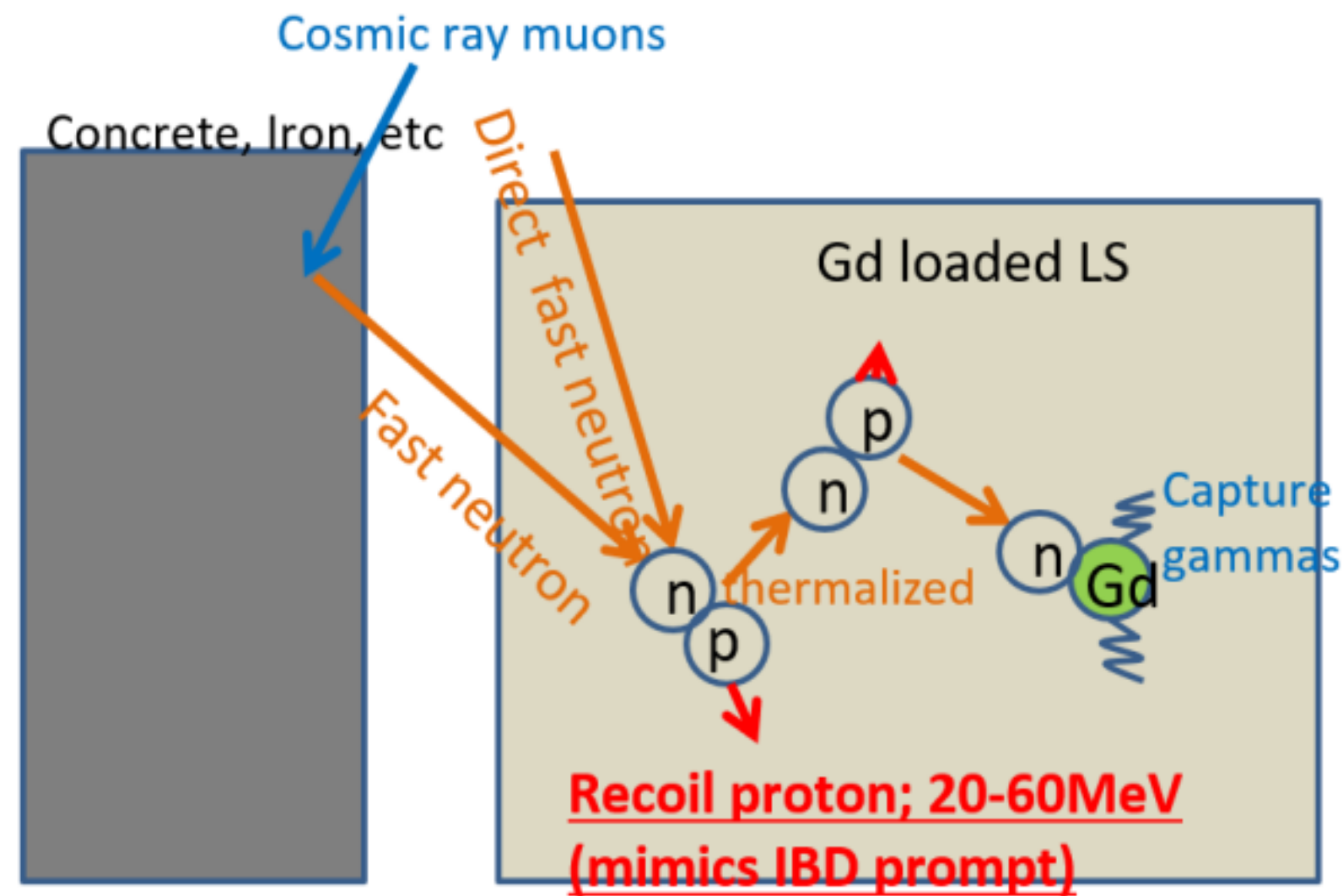
• (Preliminary) energy scale systematic uncertainties are much less than stat. error

Pulse Shape Discrimination (PSD)

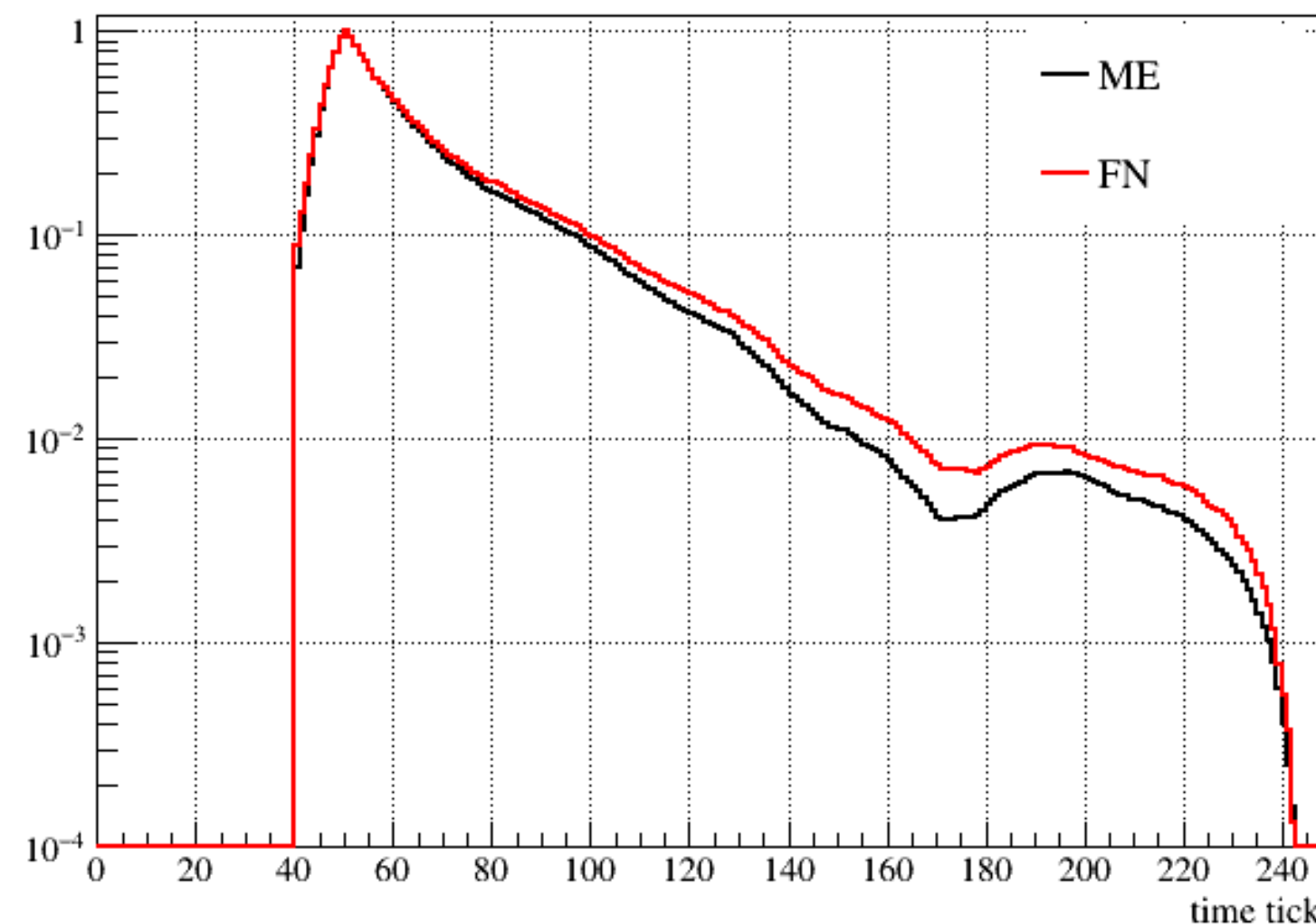
- **Fast neutrons** can mimic the IBD signals from electron anti-neutrino.

- **Pulse Shape Discrimination (PSD)** by using the difference of waveforms between e-like and n-like.

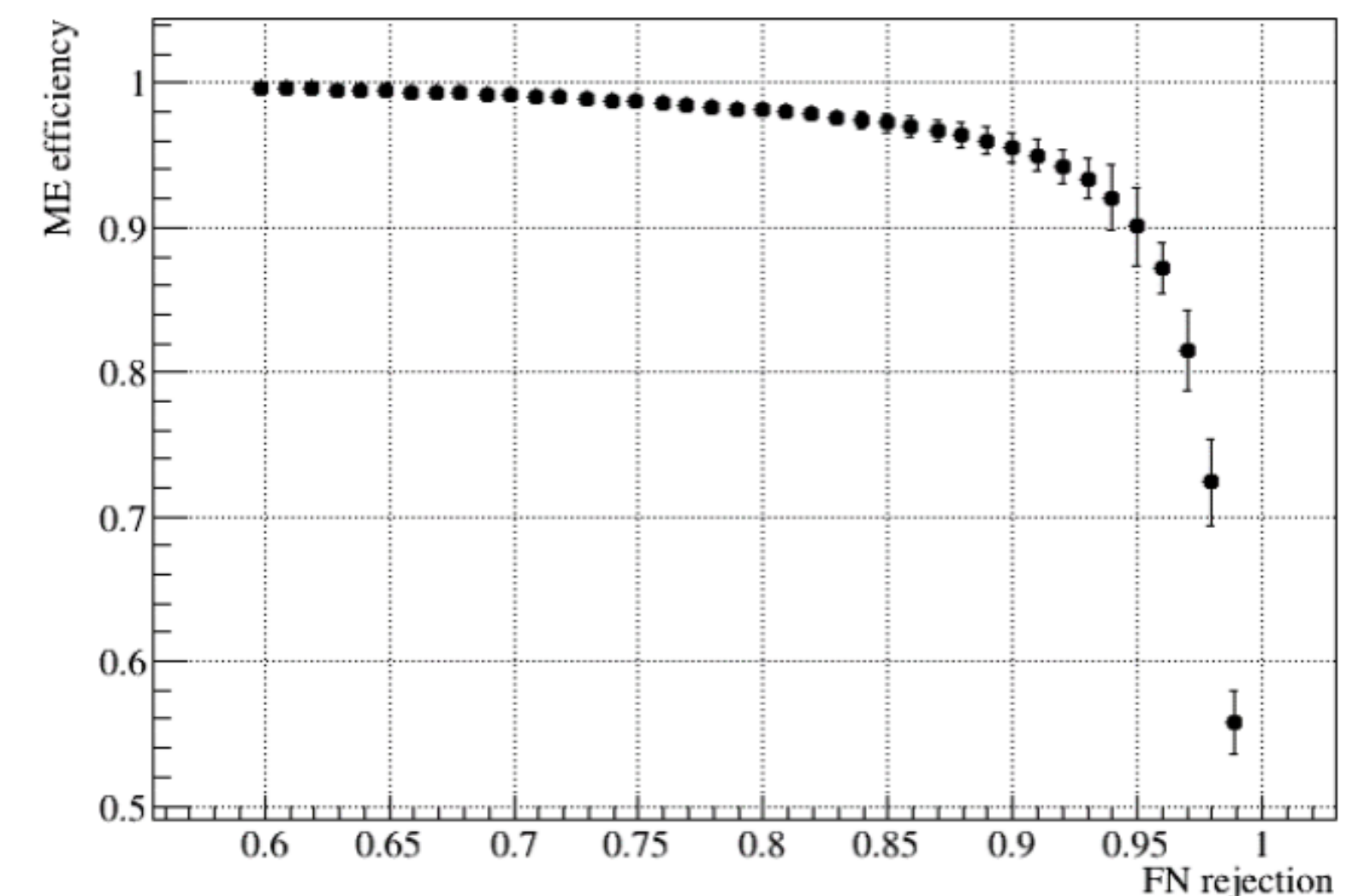
- **Data-driven likelihood method** (Control sample : Michel electron / Fast neutron)
=> Full information of waveform height are used.
=> Each PMT has its own PDF and separation power.



Average waveform (1 PMT)

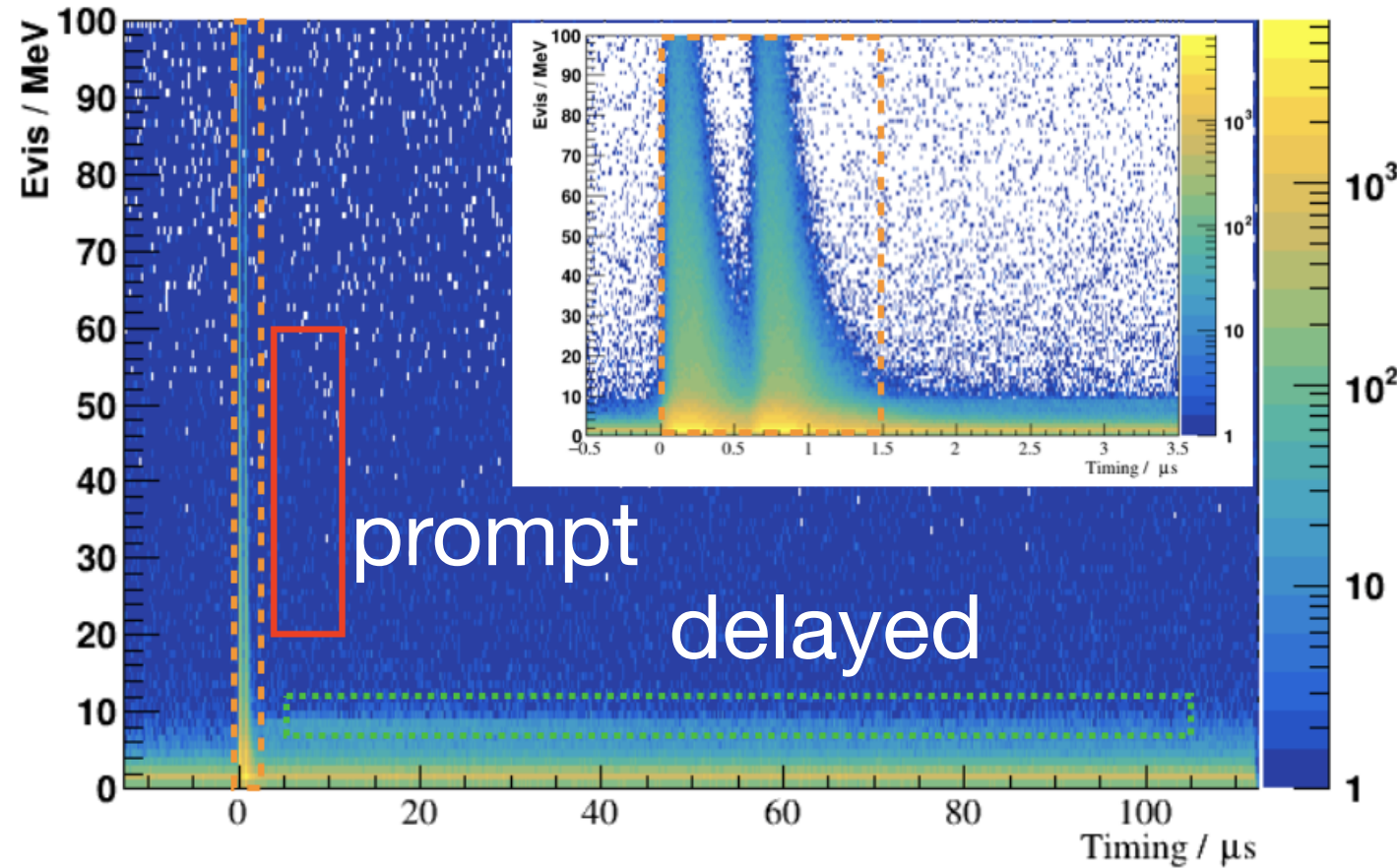


PSD capability



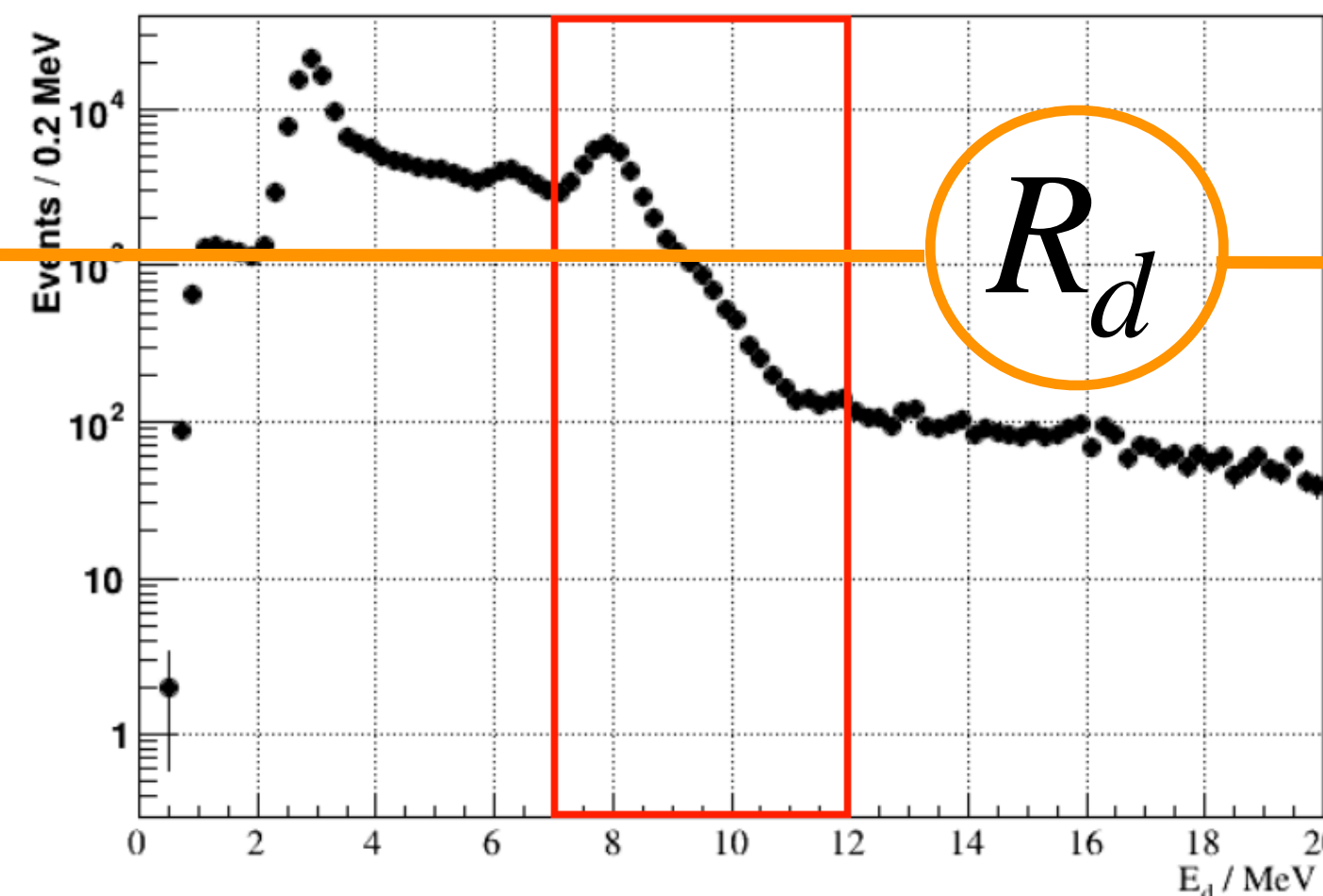
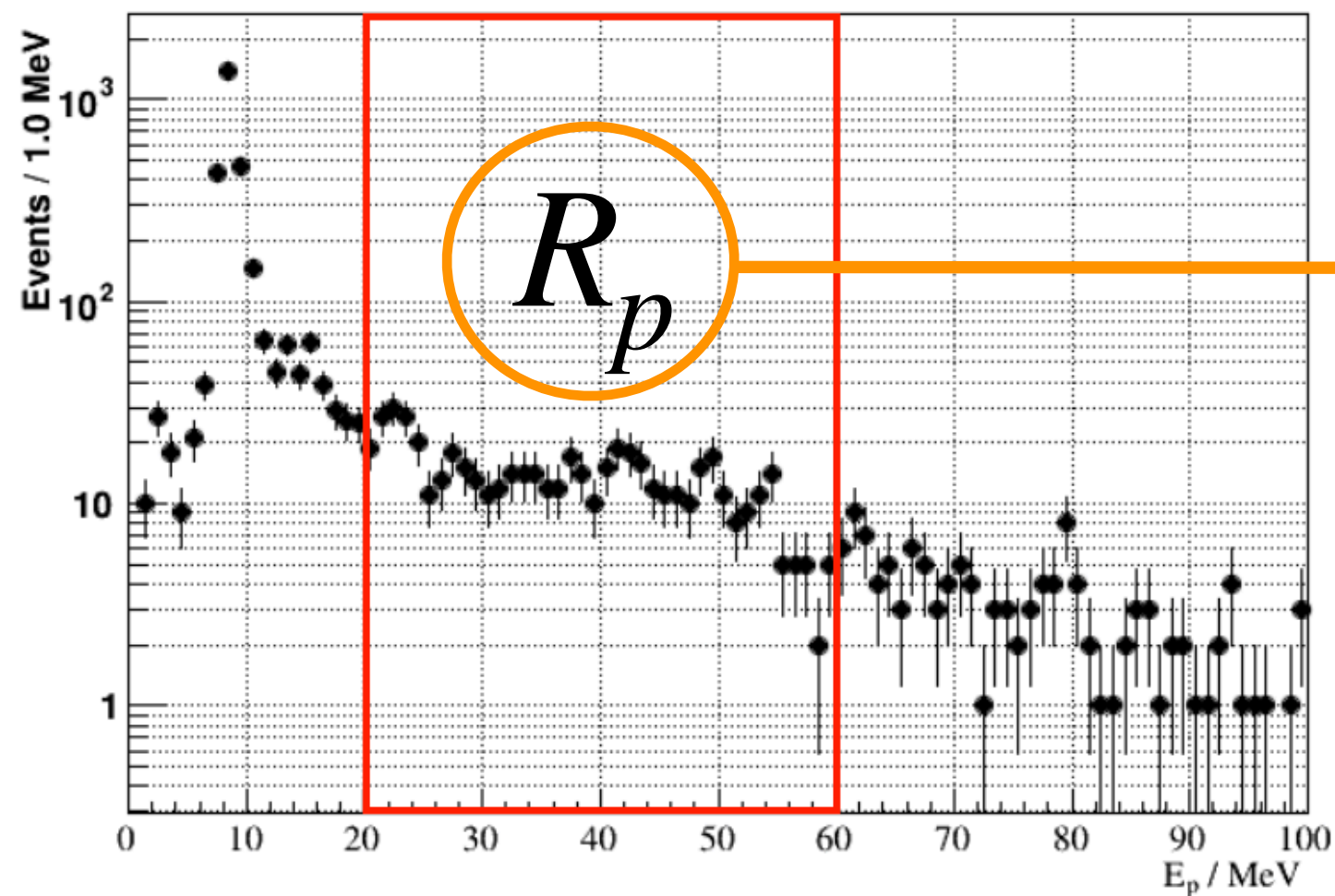
Study of the accidental background

125 μ s time window triggered data
(special calibration run)



- Accidental background means just randomly paired events of the IBD signal coincidence selections.
- Accidental background can be estimated by the multiplication of the **Single Rates of Prompt/Delayed**.

Single Rates(/spill) of Prompt/Delayed events



$$R_{acci} = R_p \times R_d \times \epsilon_{cut}$$

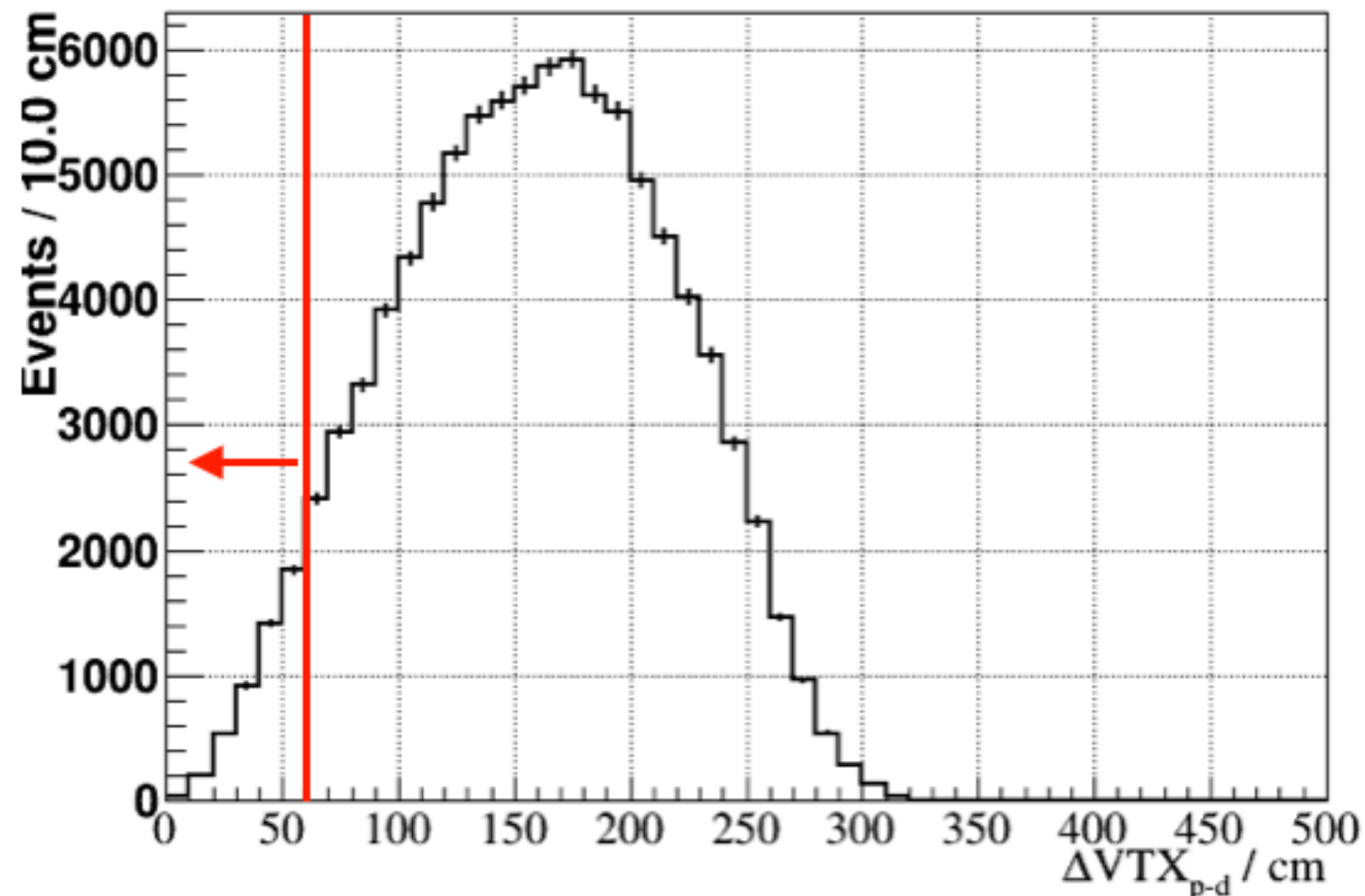
	Values
Prompt rate/spill	$(2.20 \pm 0.09) \times 10^{-4}$
Delayed rate/spill	$(1.80 \pm 0.01) \times 10^{-2}$

Eur. Phys. J. C 84, 409 (2024)

Study of the accidental background

- Not only the single rates, also any variable such as ΔVTX_{p-d} can be obtained from the data-driven control sample of random pairs.

ΔVTX_{p-d} Distribution



- For reduction of accidental background, the events $\Delta VTX_{p-d} < 60\text{cm}$ are selected.

$$R_{acci} = R_p \times R_d \times \varepsilon_{cut}$$

	Values
Prompt rate/spill	$(2.20 \pm 0.09) \times 10^{-4}$
Delayed rate/spill	$(1.80 \pm 0.01) \times 10^{-2}$
ε ($\Delta VTX_{p-d} < 60\text{cm}$)	$5.1 \pm 0.1\%$
ε (Timing Likelihood)	46%
Acci rate/spill/0.75MW	$(9.29 \pm 0.42) \times 10^{-8}$

Data
Data
Data
MC

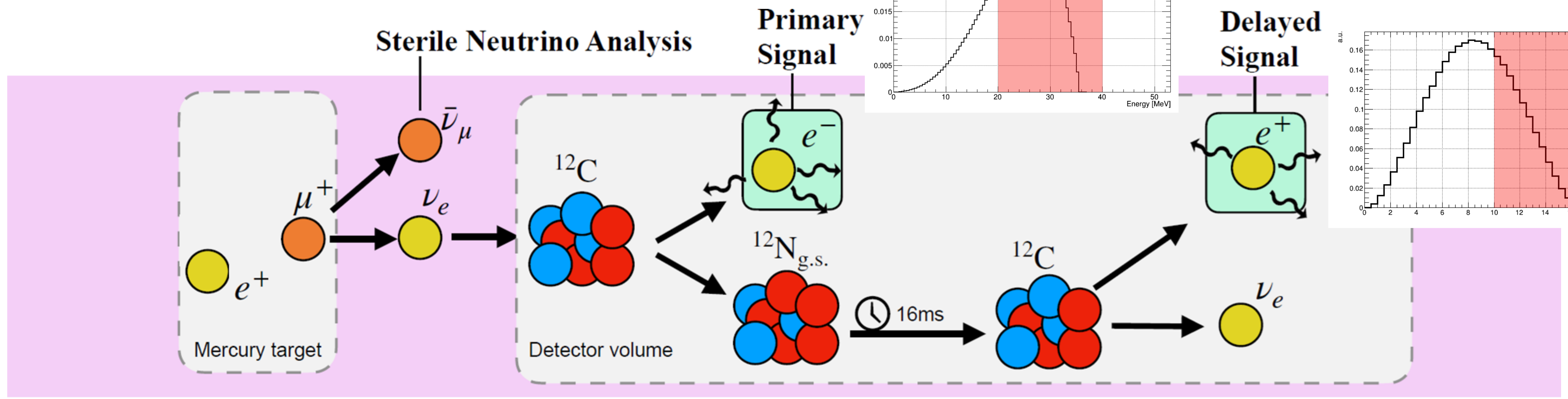
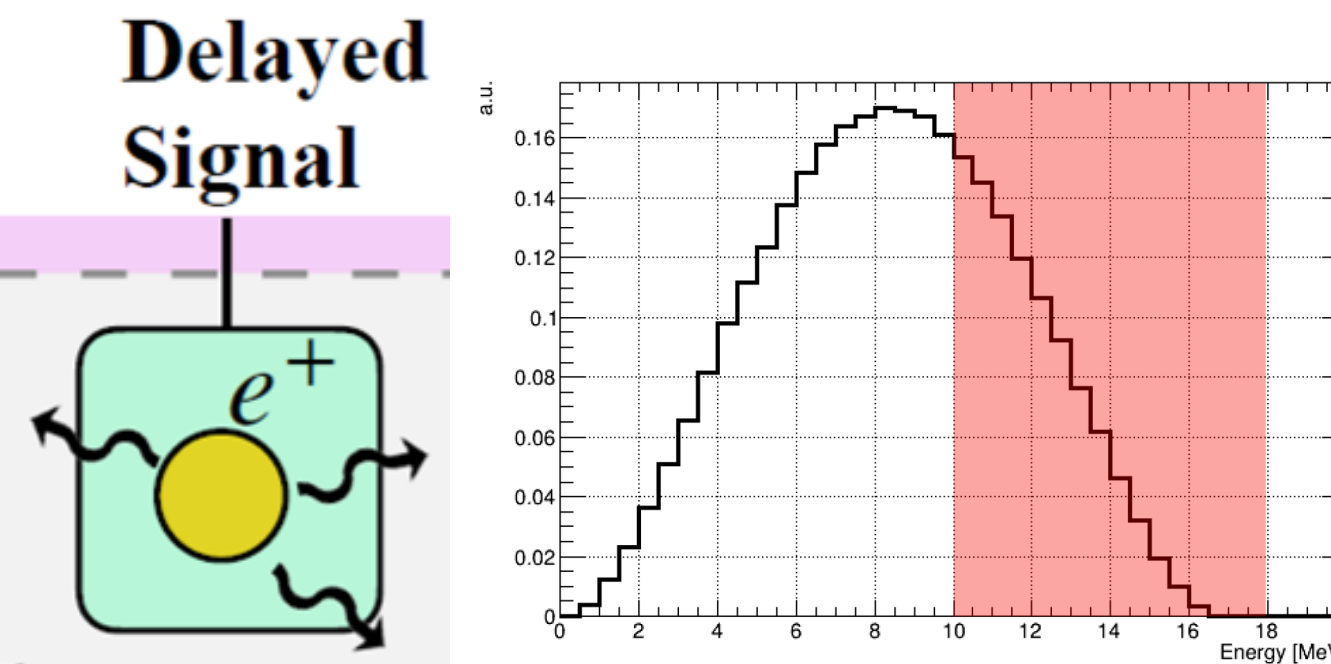
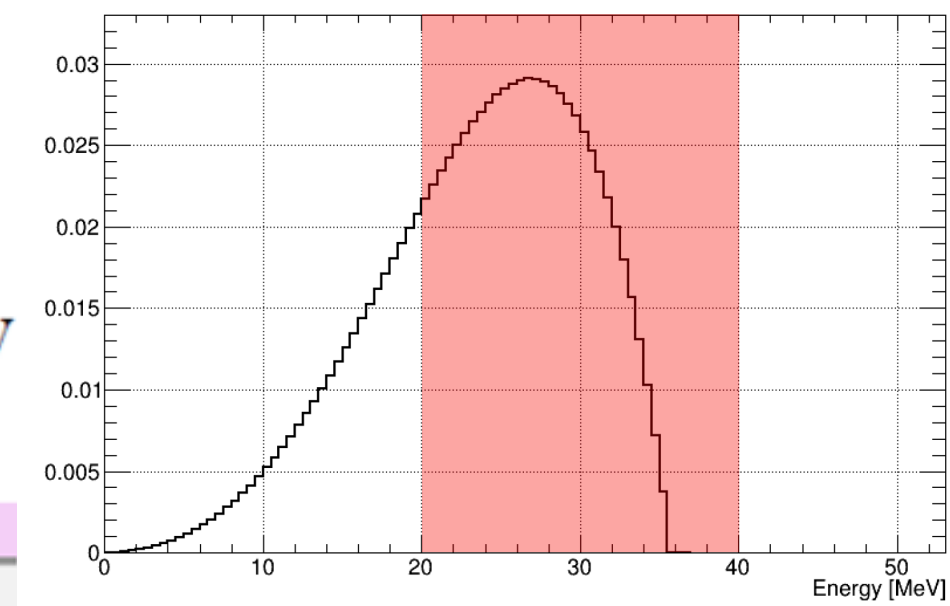
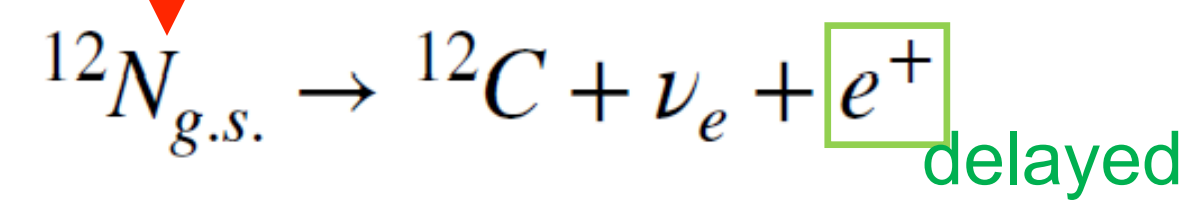
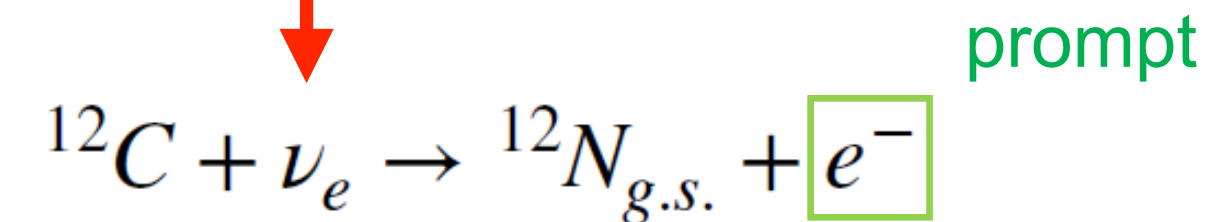
- JSNS² has a comparable accidental rate as the expected IBD signal rate **Still will be improved further !!!**
 $4.59 \times 10^{-8} / \text{spill} / 1\text{MW}$.
 (based on the LSND best fit oscillation parameters)

$^{12}\text{C}(\nu_e, e^-)^{12}\text{N}_{\text{g.s.}}$ measurement

- Preparing a paper now

$^{12}\text{C}(\nu_e, e^-)^{12}\text{N}_{g.s.}$ measurement

- No π or K production rate measurements with Hg - p (3GeV) reaction, so far.
- $^{12}\text{C}(\nu_e, e^-)^{12}\text{N}_{g.s.}$ measurements can provide a normalization of IBD signals.

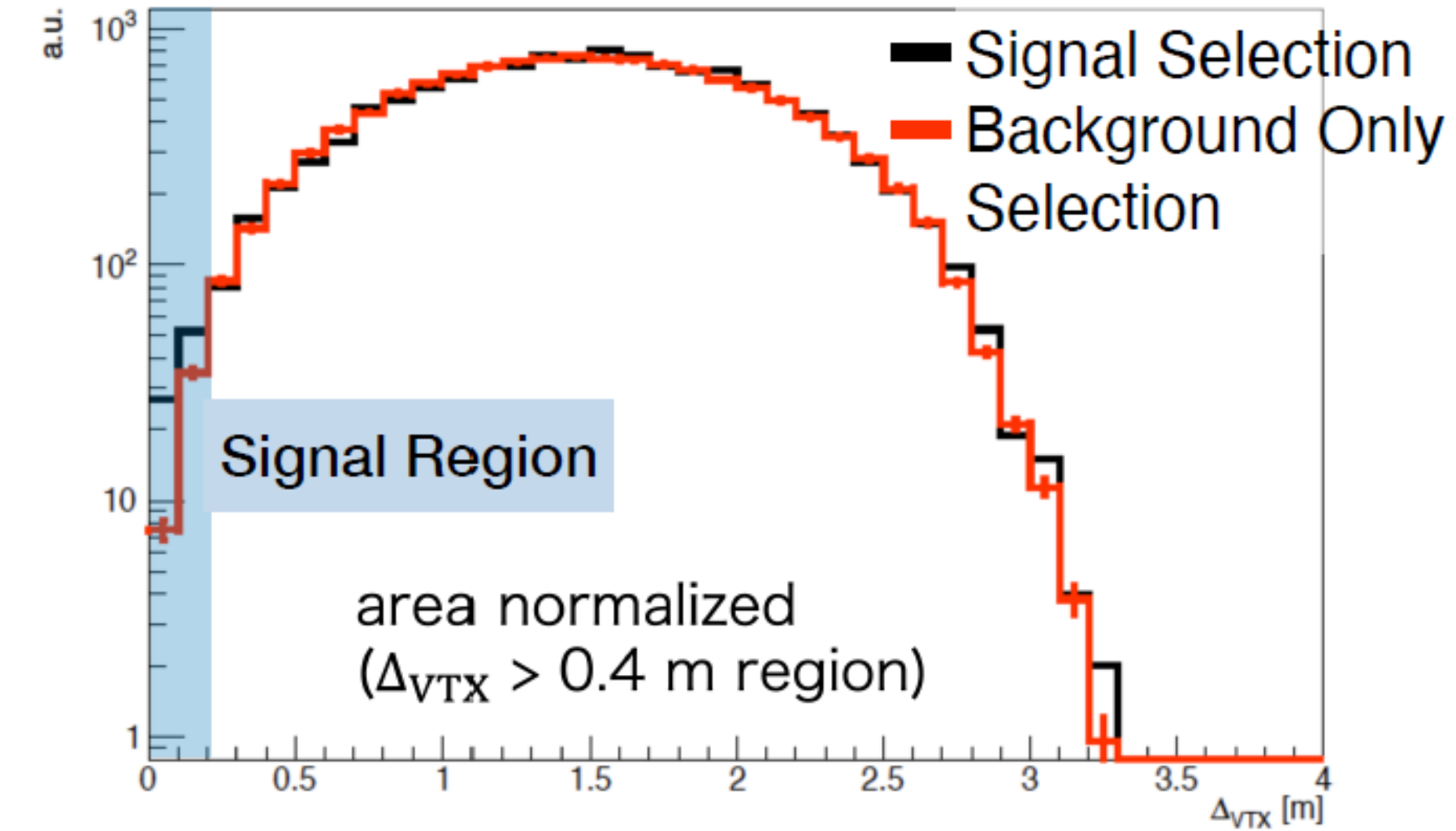


Observed events of $^{12}\text{C}(\nu_e, e^-)^{12}\text{N}_{g.s.}$

JSNS² preliminary

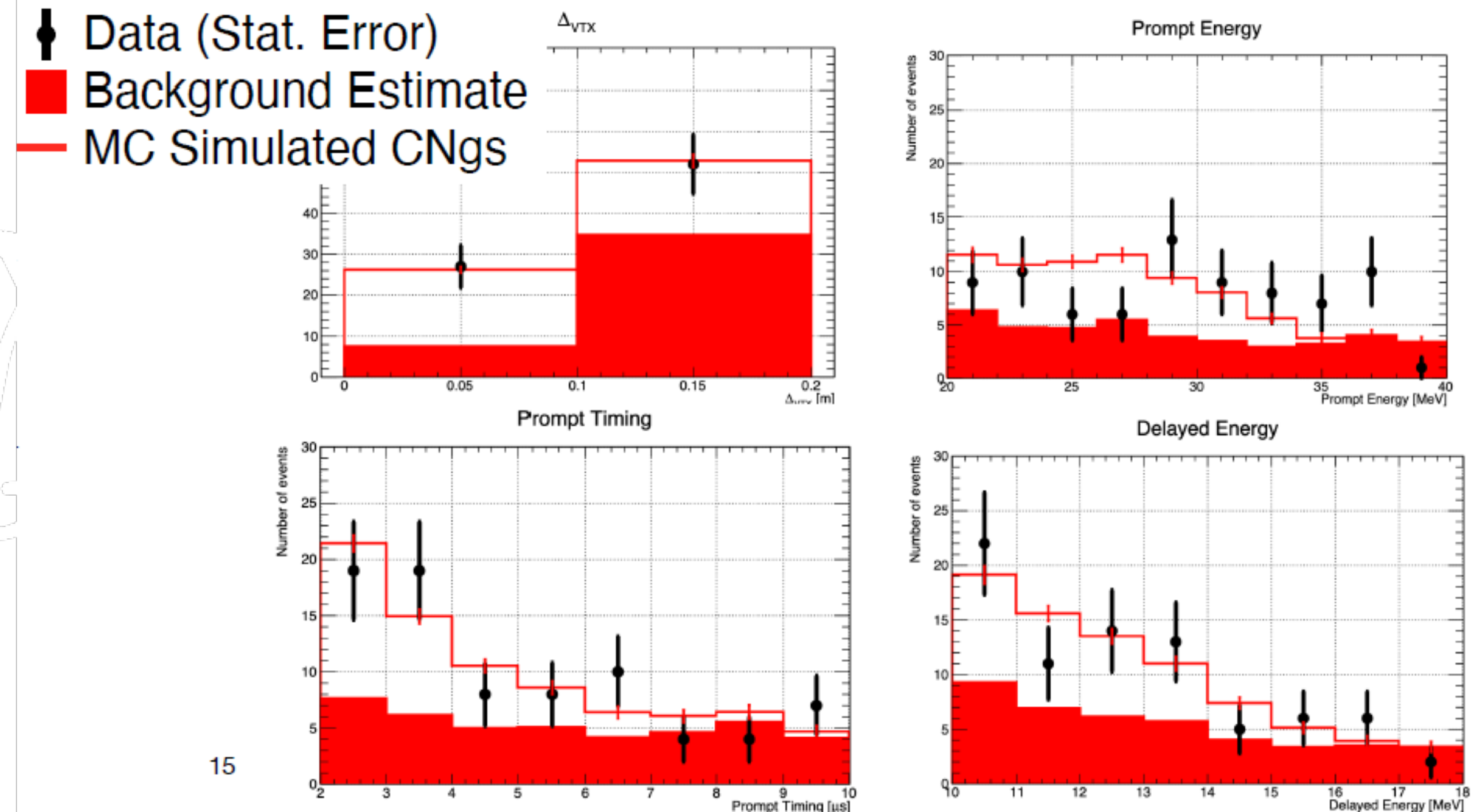
- 2021 & 2022 physics data (2.2×10^{22} POT)
- Observed CNgs candidates : 79 events
- Background events : 42.2 ± 6.5 (stat.) ± 1.7 (syst.)

<Event Selection>
 $20 < E_p < 40 \text{ MeV}$
 $10 < E_d < 18 \text{ MeV}$
 $0.2 < \Delta T_{p-d} < 12 \text{ ms}$ (2021)
 $< 25 \text{ ms}$ (2022)
 $\Delta V_{TX_{p-d}} < 20 \text{ cm}$



- The accidental background is dominant.
 \Rightarrow estimated by normalization from large ΔV_{TX} region.

- All distributions for selected variables seem to be reasonable.



Neutrino flux measurement

$$N_{CNGS} = \frac{\Phi_{\nu_e} \text{ POT}}{4\pi r^2} \epsilon \sigma N_C$$

Observed Event Rate $\rightarrow N_{CNGS}$
 Neutrino Flux $\rightarrow \Phi_{\nu_e}$
 Selection Efficiency $\rightarrow \epsilon$
 Cross Section $\rightarrow \sigma$
 Number of Carbons $\rightarrow N_C$
 POT (Protons on Target)

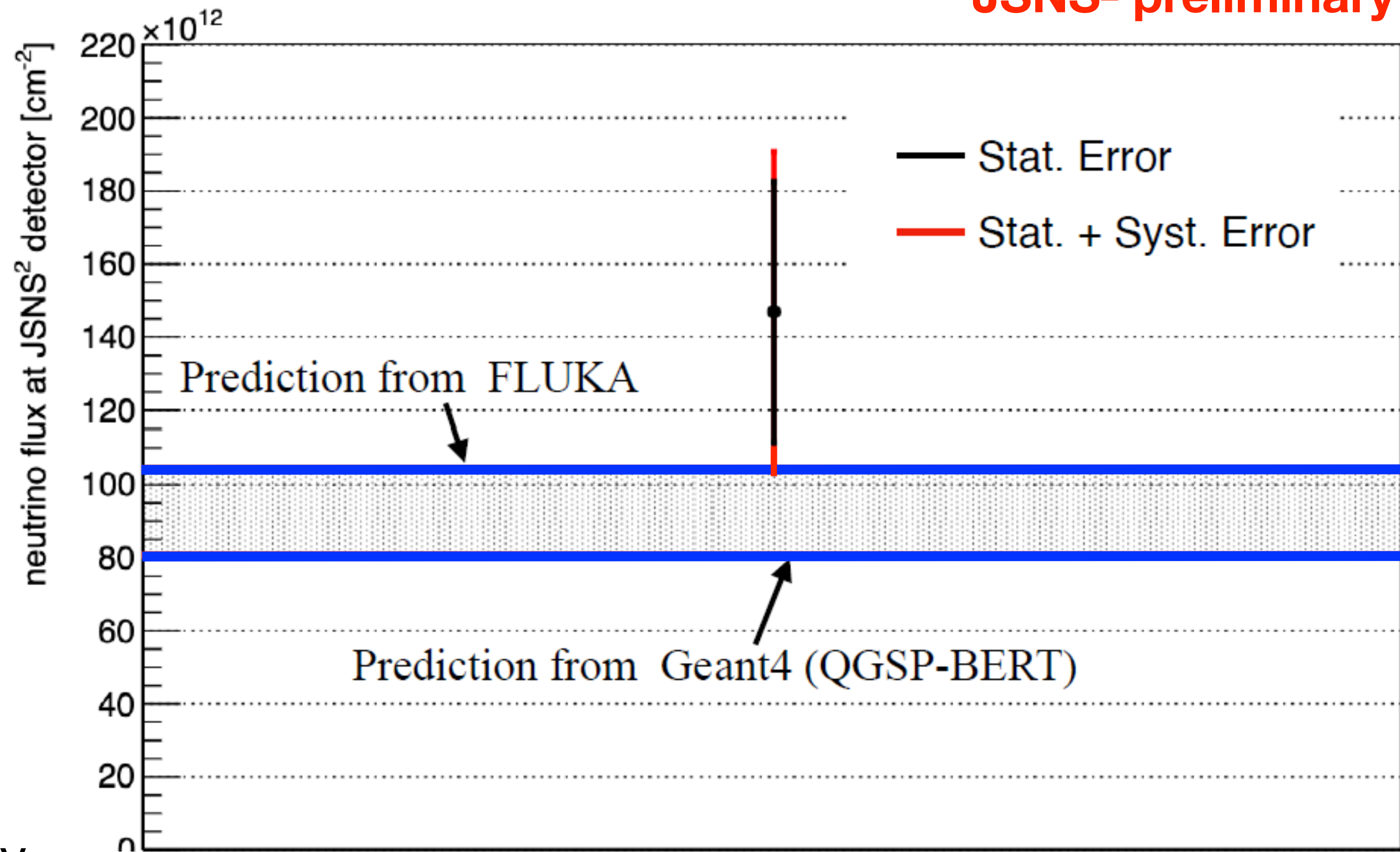
- Neutrino flux can be obtained by above equation.
 $\Rightarrow \epsilon_{\text{detection}} = 0.0588 \pm 0.0021$
 \Rightarrow Cross section (LSND+KARMEN)
 $= (9.1 \pm 0.7) \times 10^{-42} \text{ cm}^2$

- Good consistency with the predictions.

- POT will be twice with full data of 2021-2024.
 \Rightarrow can achieve ~20% total uncertainty.
 \Rightarrow to normalization of IBD signals.

Dongha Lee

JSNS² preliminary



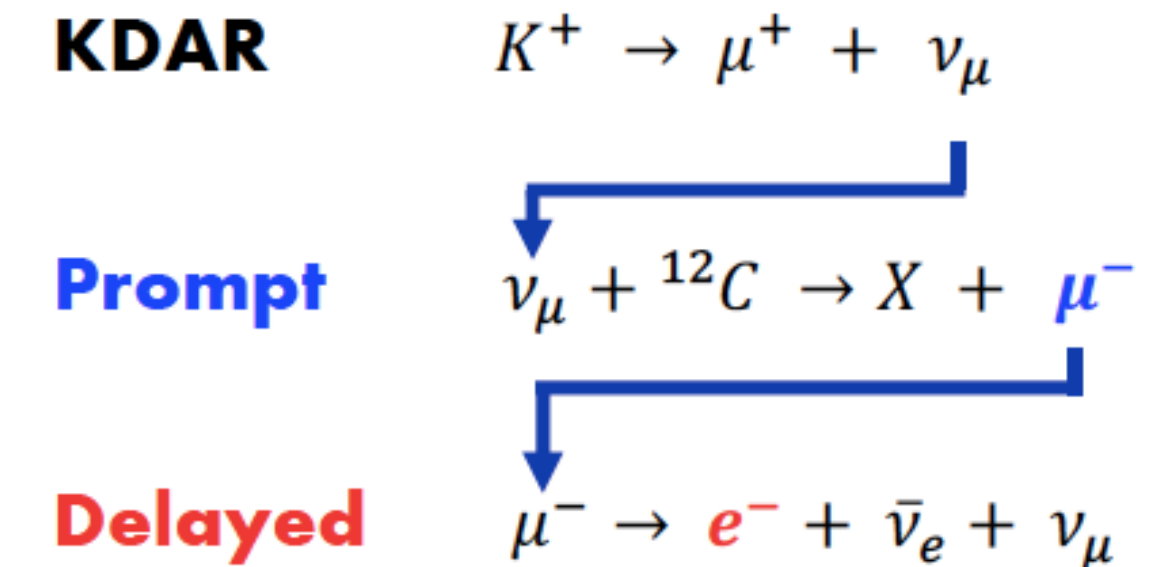
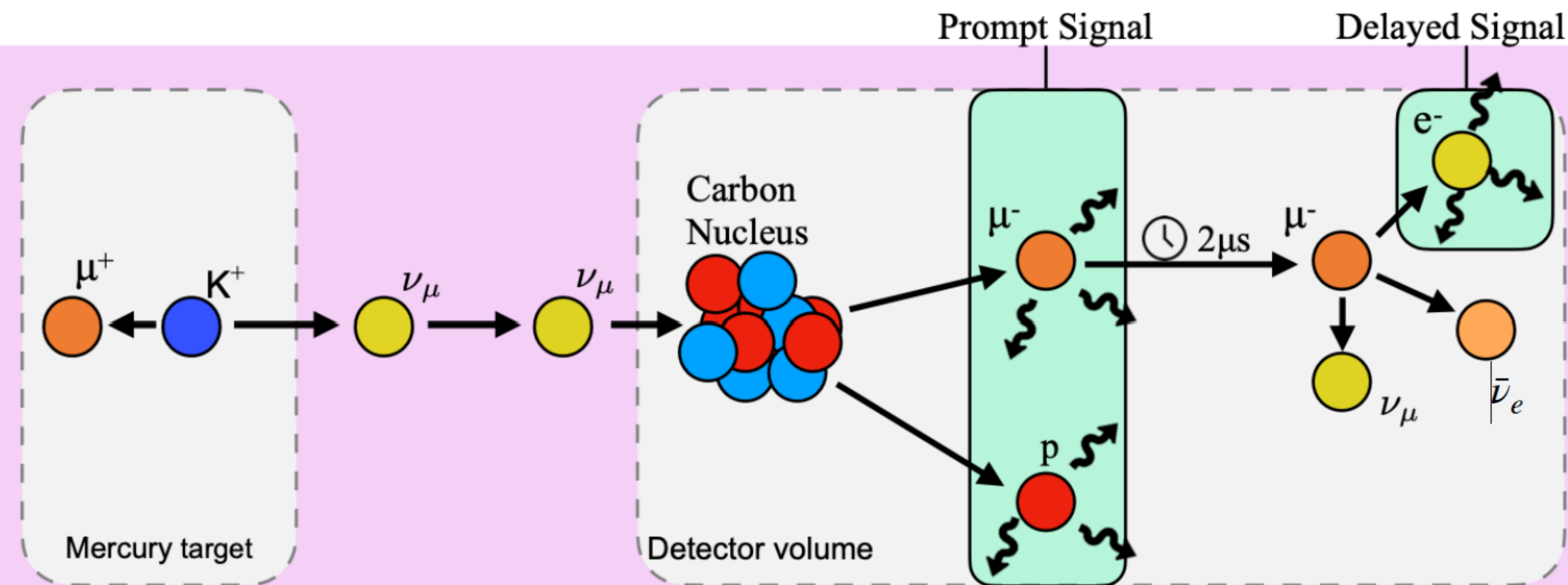
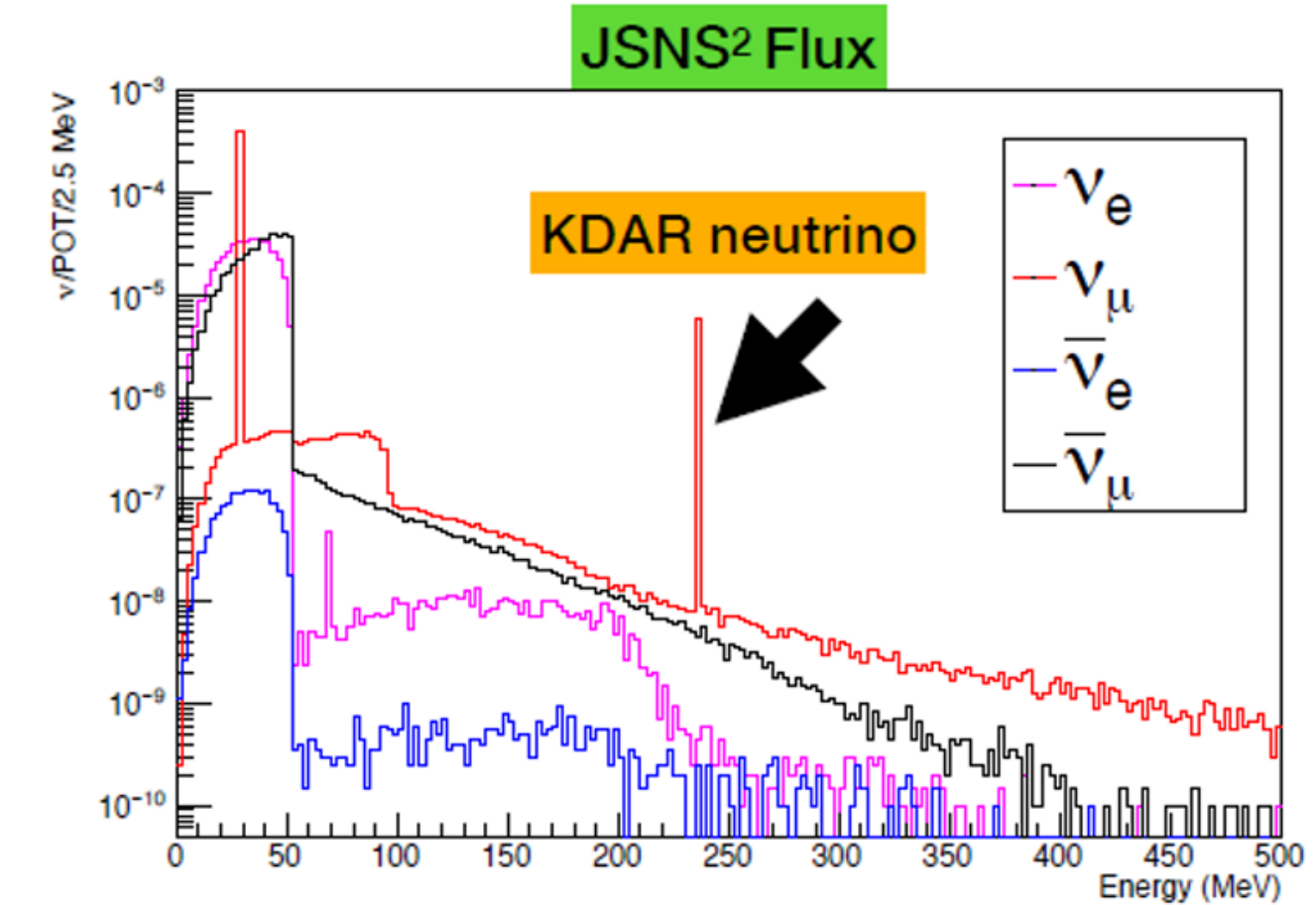
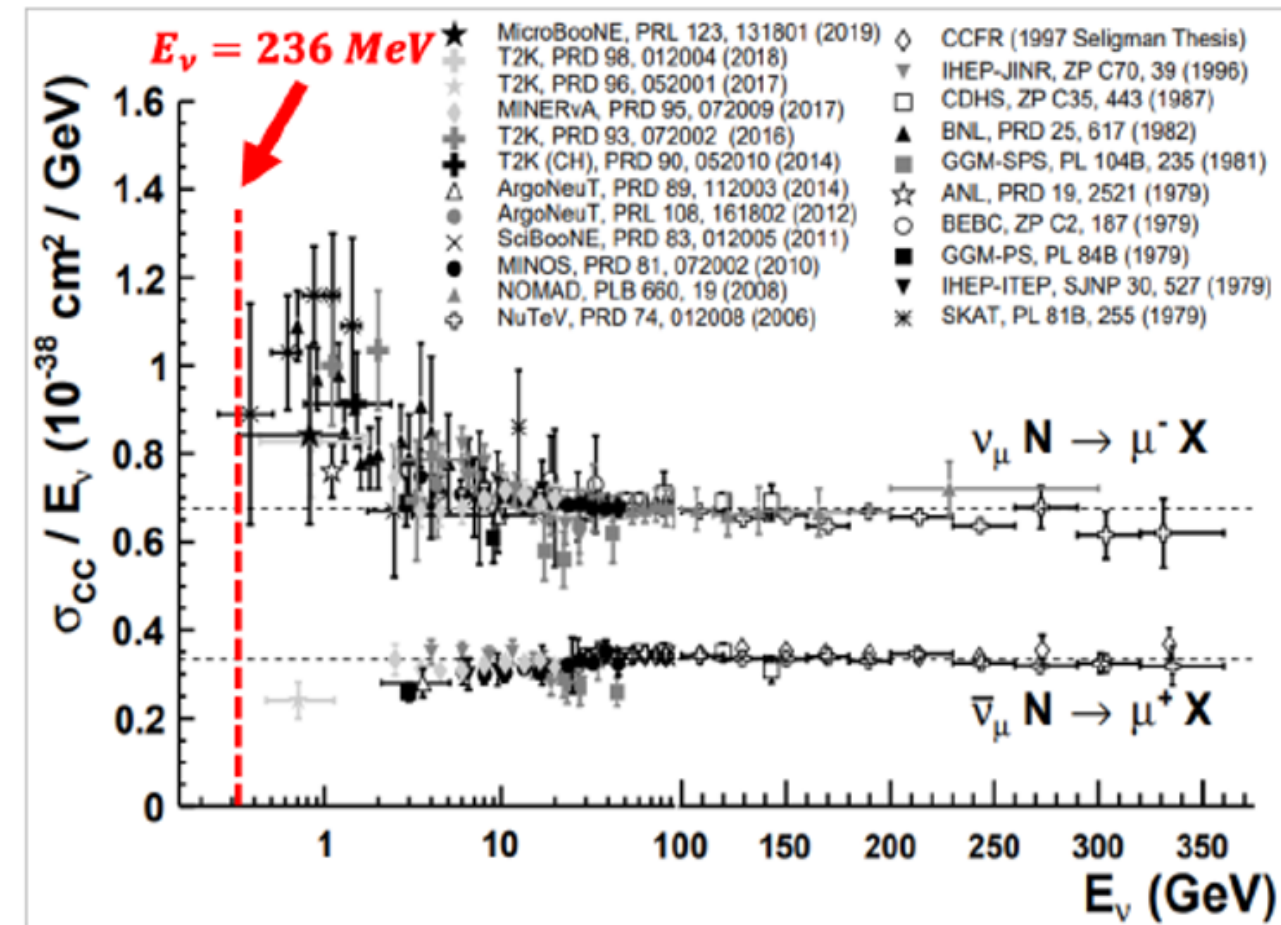
* For cross section, the combined value of LSND (Phys. Rev. C 64:065501 (2001)) +KARMEN (Prog. Part. Nucl. Phys. 40, 183 (1998)) is used.

KDAR neutrino measurement

- arXiv:2409.01383 [hep-ex], submitted to PRL

Kaon Decay-At-Rest neutrino measurement

- KDAR neutrinos have a **KNOWN mono-energy (235.5MeV)**
- Few cross-section measurements below 1GeV so far.
- Nuclear effects measurements
-> some progress in this study

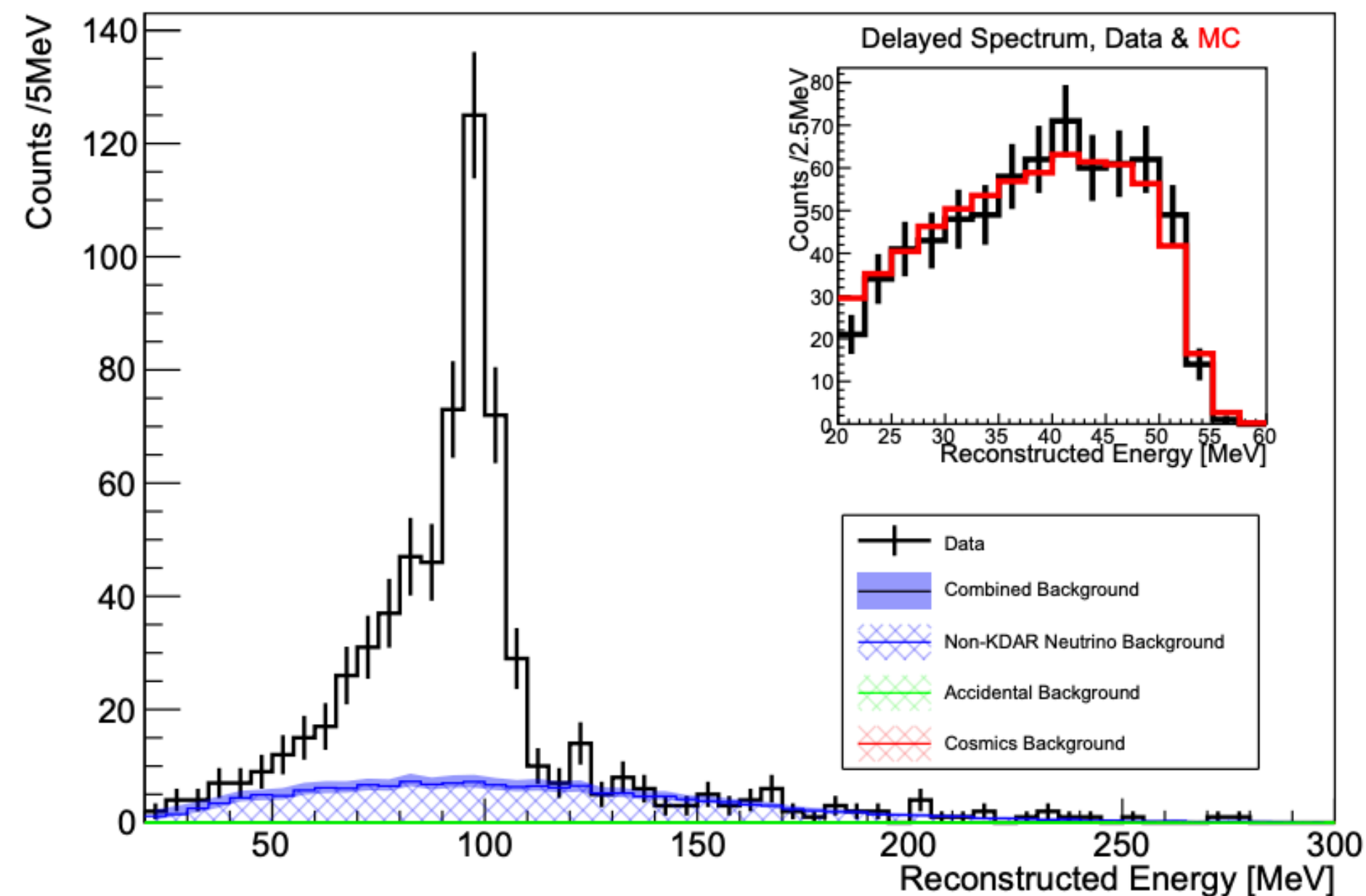


- Detection by double coincidence
=> Prompt : muon and proton
=> Delayed :
electron from muon decay

KDAR Results

- 2021 physics data is used. (1.4×10^{22} POT)
- KDAR candidates : 621 events
- Best Fit Bkg : $144.4^{+21.3}_{-21.1}$ events
(π DIF ν dominant)

	Prompt	Delayed
Energy	20~150MeV	20~60MeV
Timing	2 x 150ns (Beam centered windows)	$\Delta T < 10\mu\text{s}$
Position	R < 140cm -100cm < z < 50cm	$\Delta VTX < 30\text{cm}$

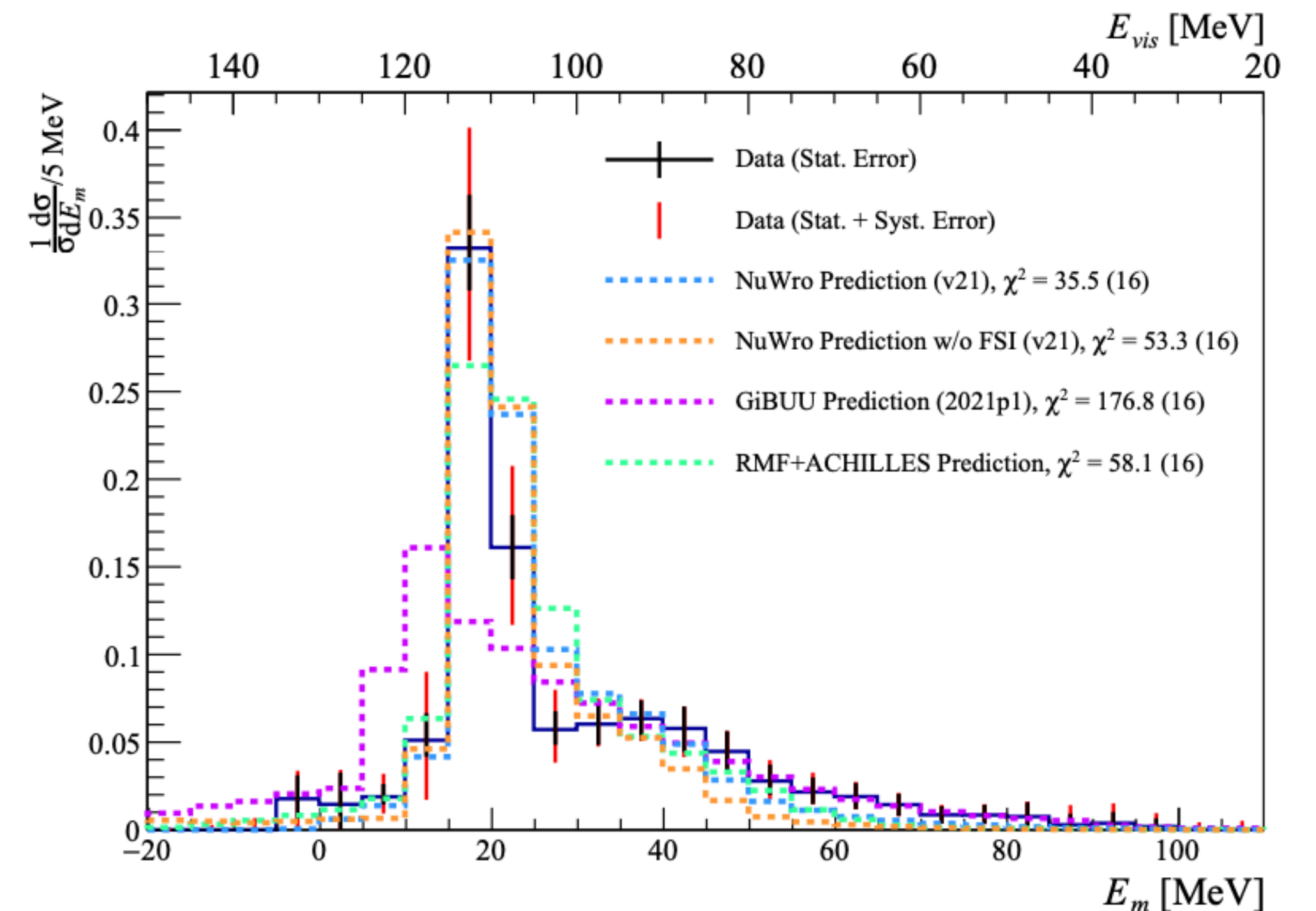


- The First measurement of Missing Energy due to nuclear effects (shape-only analysis)
=> By using the KNOWN neutrino energy.

$$(E_m \text{ (missing energy)}) = E_\nu - m_\mu - E_{vis}$$

=> Unfolded E_{vis} is used.

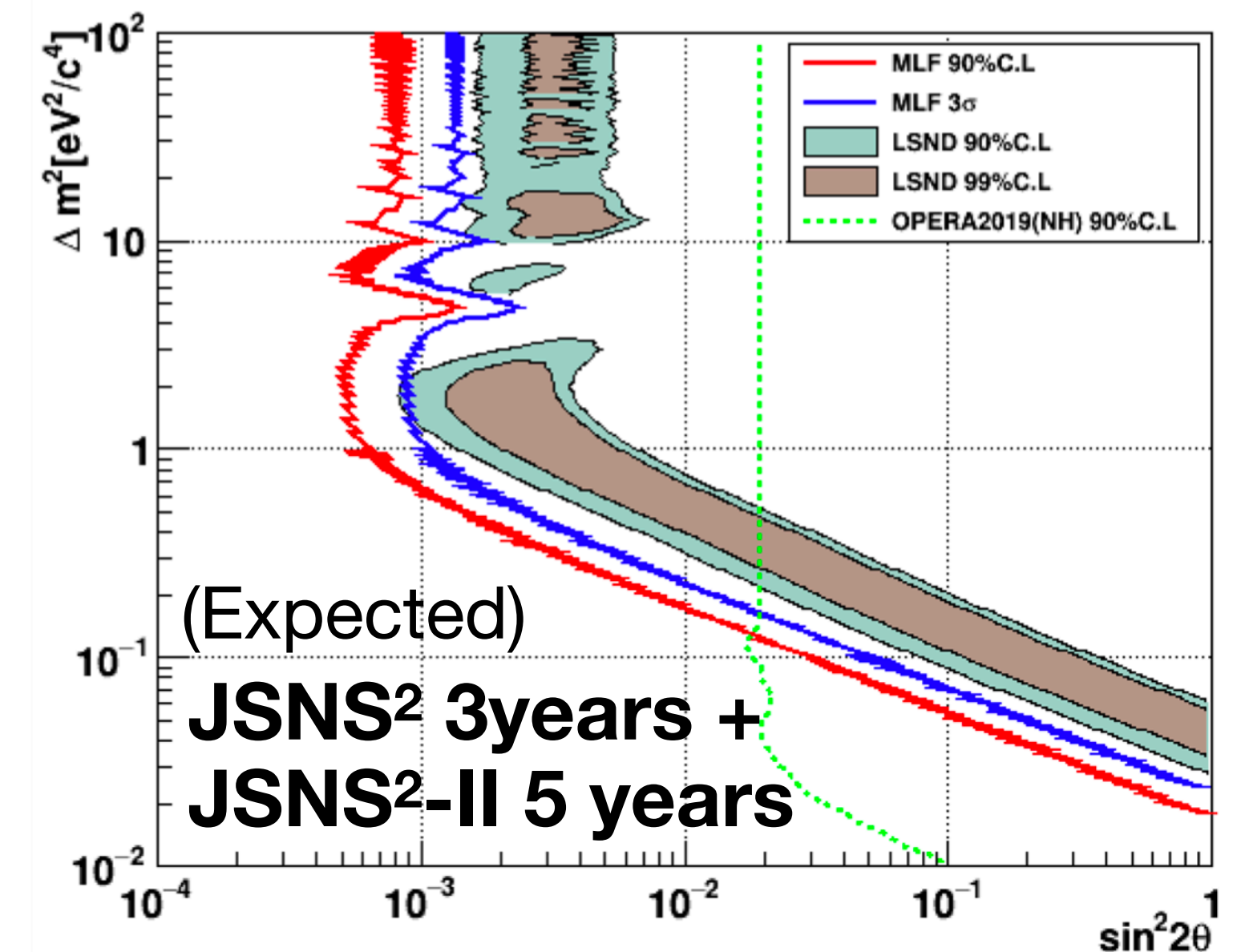
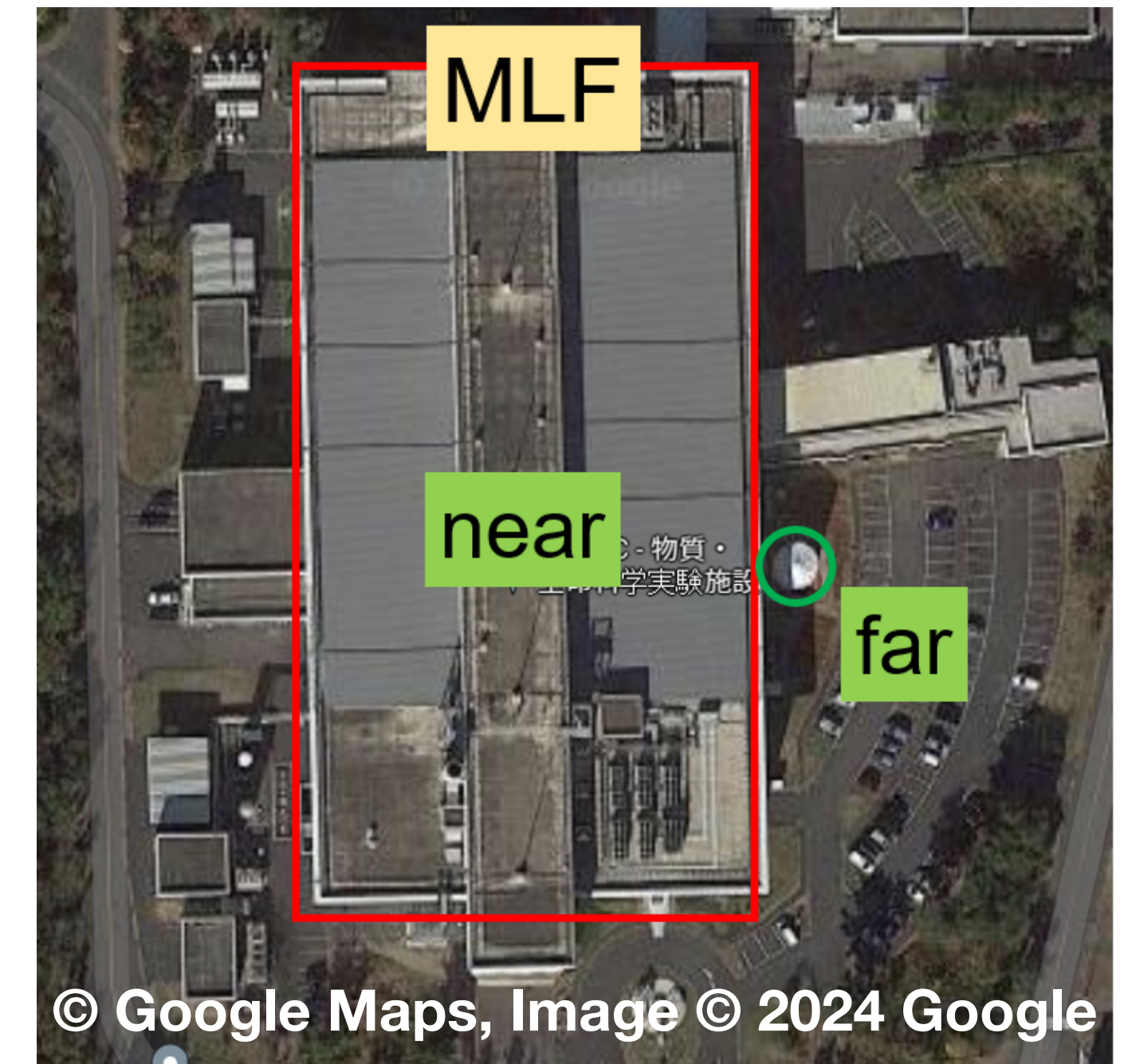
- Standard candle for improving the understanding of the low energy neutrino-nucleus interaction.
- arXiv:2409.01383 [hep-ex], submitted to PRL



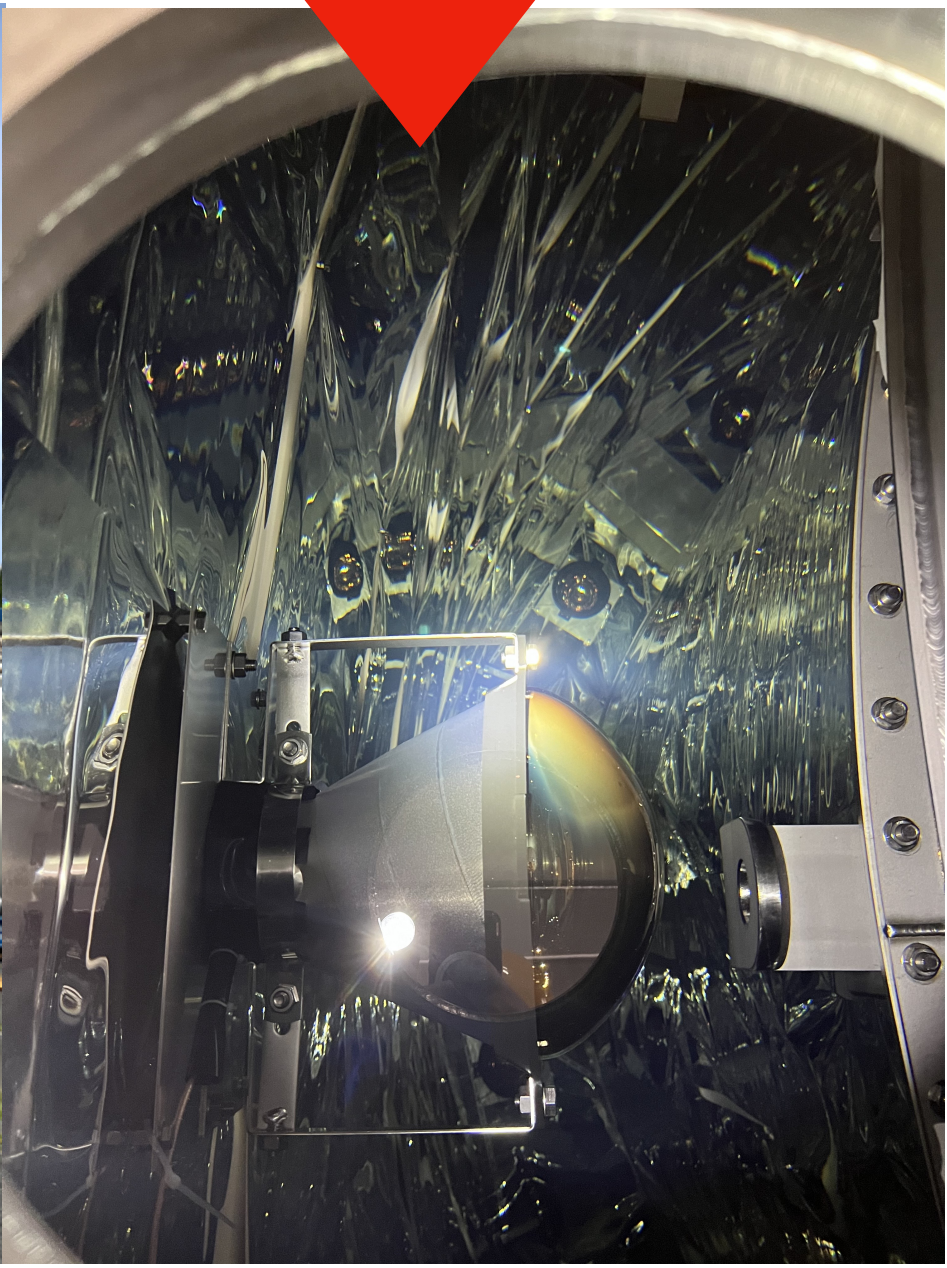
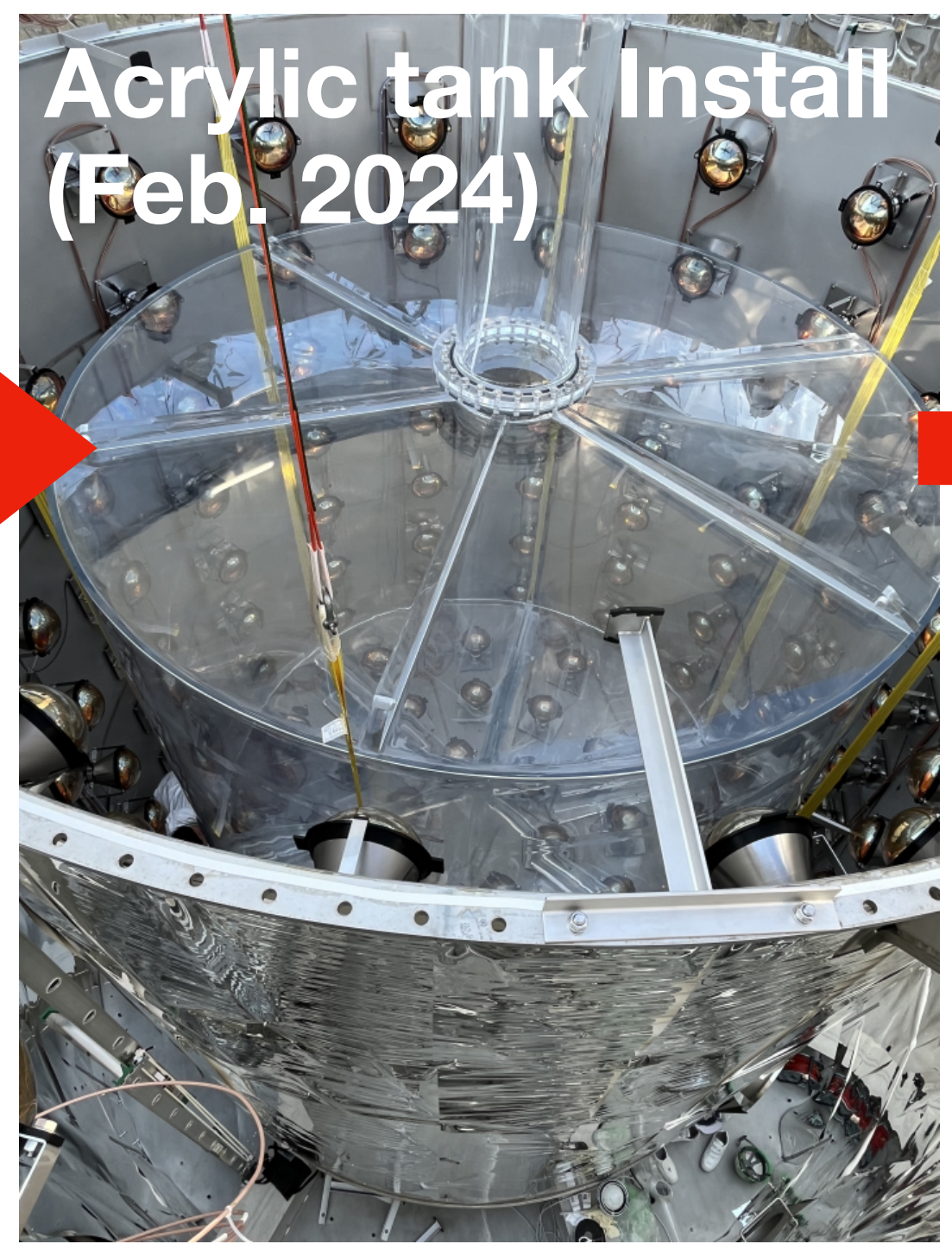
Status of JSNS²-II

JSNS²-II : the second phase of JSNS²

- New Far Detector
 - => Almost identical structure with ND.
 - => Target : 32tons of Gd-LS.
 - => 48m baseline, outside of MLF building.
- Two detectors (ND+FD) with two baselines (24m, 48m)
 - => The sensitivity at low Δm^2 region will be much improved.
 - => A solid conclusion on LSND anomaly.
- FD construction since 2021.
 - => Expected to begin data taking in early 2025.



JSNS2-II Far detector construction



- Smooth progress so far.
- LED test soon.

Summary

- 4 long term physics run (2021-2024)
=> 950kW @MLF in 2024 (Design 1MW is achieved at RCS extraction point)
- Blind analysis (energy side-bands) for the sterile neutrino search
=> Good consistency between Observation vs Expectation for a side-band so far.
- CNgs measurement
=> Neutrino flux measurement has a good consistency with prediction.
- KDAR measurement
=> With known neutrino energy (235.5MeV), the first measurement of the missing energy.
- JSNS²-II : the second phase
=> Now climax of New Far Detector construction.

Backup

Expected Sterile neutrino sensitivity (JSNS²)

Signal	$\sin^2 2\theta = 3.0 \times 10^{-3}$ $\Delta m^2 = 2.5 eV^2$ (Best fit values of MLF)	87
	$\sin^2 2\theta = 3.0 \times 10^{-3}$ $\Delta m^2 = 1.2 eV^2$ (Best fit values of LSND)	62
background	$\bar{\nu}_e$ from μ^-	43
	$^{12}C(\nu_e, e^-)^{12}N_{g.s.}$	3
	beam-associated fast n	≤ 2
	Cosmic-induced fast n	negligible
	Total accidental events	20

**1MW x 3years
(100% approved POT)**

