# Gibuu@faser

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- GiBUU : Quantum-Kinetic Theory and Event Generator based on a BM solution of Kadanoff-Baym equations, allows for off-shell propagation
- GiBUU propagates phase-space distributions, not particles
- Physics content and details of implementation in: Buss et al, Phys. Rept. 512 (2012) 1-124
- Code from <u>gibuu.hepforge.org</u>, new version GiBUU 2023 add. details in Gallmeister et al, Phys.Rev. C94 (2016) 3, 035502, Phys.Rev.D 109 (2024) 3, 033008, v 2024 later this year.





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### **FASER Flux**

v Faserflux



Flux from: Max Fieg, UC Irvine

### Peaks around 100 – 200 MeV, Tails out to 2 TeV

Target is heavy nucleus: W



# Physics@FASER

- DIS is dominant, QE, RES, etc negligible
  DIS in GiBUU is handled by PYTHIA string fragmentation
- At these high energies: are the targets relevant?? Long formation times will cause to create ,real' hadrons only outside the nucleus?

- Are final state interactions are essential? :
  - Production and formation times of these hadrons?
    Physical hadrons produced inside or outside the target?





### Hadronization

Central Question: how do the final state hadronic cross sections emerge after a hard elementary process?

Does the target (W) matter for exclusive events?







## String-breaking in PYTHIA



Leading particles connected with hard interaction 3 times for each hadron

1. Production V1 : String Breaking

2. Production V2 : String-Breaking

3. Formation: Quark lines meet





### **PYTHIA's formation times**





K. Gallmeister, T. Falter *Phys.Lett.B* 630 (2005) 40-48 Times extracted from PYTHIA

 $\mathsf{Z}=\mathsf{E}_{\mathsf{h}}\,/\,\nu$ 

The higher the hadron's energy the larger the formation time
 → high z hadrons form outside the nuclear target -> no FSI
 → Low z hadrons (high multiplicity events) form inside the nuclear target -> Final State Interactions may be relevant depends on ,prehadronic' interactions during formation





### Hadronization in DIS events

### GiBUU can (has) been used to study hadronization at

 HERMES (28 GeV) and EMC (280 GeV), uses nuclear radius to determine formation times and prehadronic cross sections.







### **Prehadron Cross Sections**





K. Gallmeister, U. Mosel, Nucl. Phys. A 801 (2008) 68-79, MDPI Physics 4 (2022) 2, 440-450

Only linear dependence on time fits data over wide kinematical range!

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 $z = E_h / v$ 

### **Cross Section during Hadronization**

Dominated by DIS, have to worry about Color Transparency in FSI:

$$\frac{\sigma^*(t)}{\sigma} = X_0 + (1 - X_0) \cdot \left(\frac{t - t_P}{t_F - t_P}\right), \quad X_0 = r_{\text{lead}} \frac{\text{const}}{Q^2}, \quad \text{Farrar},$$

Farrar, Strikman et al

describes JLAB and HERMES hadronization data, also EMC







# HERMES@27 GeV

### Airapetian et al.



### **FASER** inclusive kinematics













### **Multiplicities at FASER: FSI**





### FSI effects on multiplicities at I TeV

Dramatic ,avalanche effect' on baryon multiplicity pions are there from the first (string) decay on, less affected



Target: W, 1 TeV



### Hadron Spectra at FASER flux



#### Qualitatively similar to lower neutrino energies (MINERvA)

Protons show ,avalanche effect' (multiple collisions through final state interactions), Pions are there from the start, not much affected by FSI

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# $\Sigma$ + Spectrum







### vA Questions for FPF

#### JLAB 2021



Where isColor Transparency ???

### EMC effect for neutrinos, tension with electrons





K. Gallmeister, U.M., Nucl. Phys. A 801 (2008) 68 Time Dependent Hadronization via HERMES and EMC Data Consistency

### Final state interactions; Problems

Frozen density approximation, used e.g. in mean free path approaches ???

At very high energies target nucleus will break up
 Target is ,time-dependent', target nucleons also have to be propagated

GiBUU allows for such a ,time-dependent target' , and density change

 BUT: approximate this by changing so far only target A (saves computer time!)







# Summary

- FASER flux baryon spectra are sensitive to formation times and prehadronic cross section developments
  - $\rightarrow$  FSI are essential even at very high energies of FASER
- FASER may add to solutions of problems of
  - Color Transparency and
  - EMC effect for neutrinos
  - Hadronization times at low z
- Same physics at EIC: also there FSI are essential, but so far have not been considered in detail.



