



Design and status of the JUNO detector

Marco Beretta

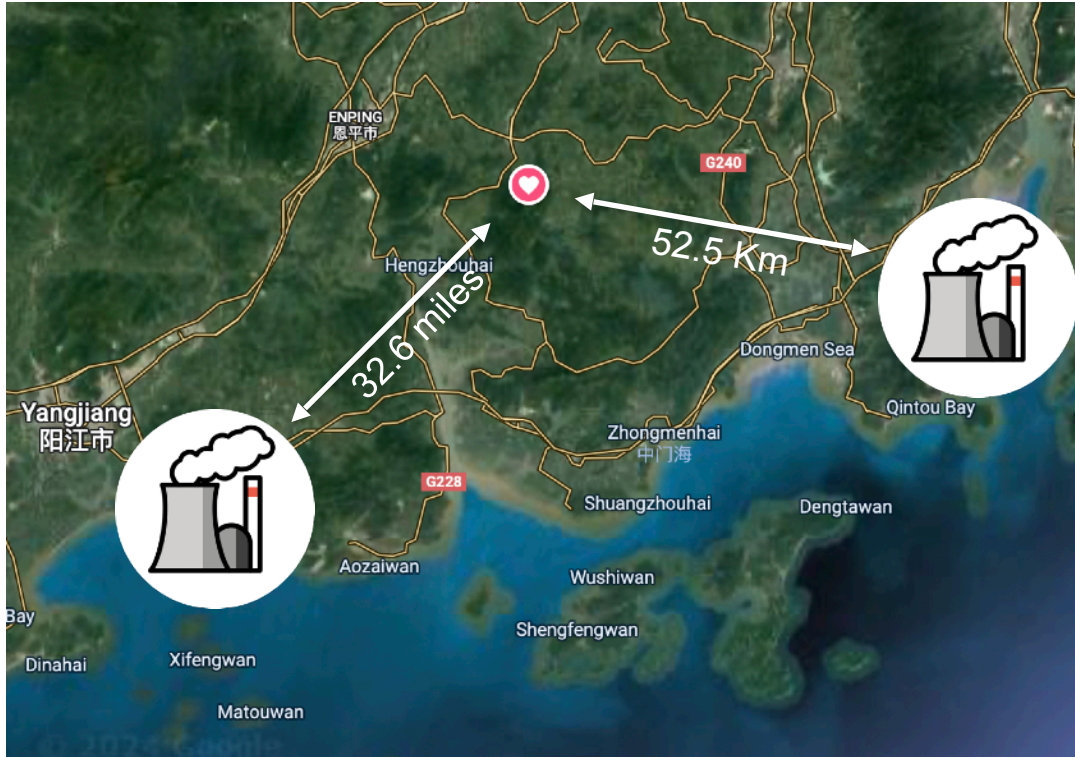
On behalf of the JUNO collaboration

marco.beretta@mi.infn.it



NuFact 2024

The Jiangmen Underground Neutrino Observatory



JUNO is a **20 kton** multi-purpose underground **liquid scintillator** detector.

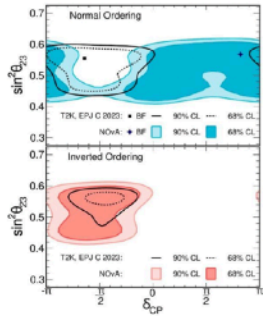
Baseline of about **32.6 miles** from **two nuclear plants** in the Guangdong Province of South China.

Why a reactor experiment at an accelerator devoted conference?

Neutrino physics

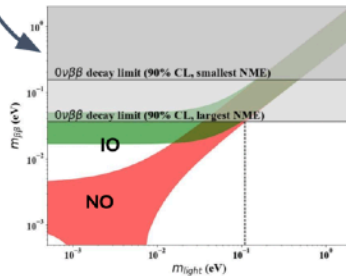
Why NMO is important?

- 1) Missing tile for the fundamental comprehension of neutrinos
- 2) Strictly connected to $< 1\%$ determination of $\theta_{12}, \theta_{31}, \Delta m^2$ splittings
- 3) Remove the degeneracy between leptonic CP violation (δ_{CP}) and MO \rightarrow T2K, NOvA, DUNE
- 4) Driving the strategy for the next-gen $0\nu\beta\beta$ experiments: can they determine their Dirac/Majorana nature?



Eur. Phys. J.
C (2023) 53:
782.

Phys. Rev. D
10, 02,005



Gómez-Cadenas et al. Riv. Nuovo Cim. 46 (2023) 10, 619-692

8

Why a reactor experiment at an accelerator devoted conference?

Neutrino physics

Why NMO is important?

- 1) Missing tile for the fundamental comprehension of neutrinos



Neutrinos from Natural Sources at JUNO

Iwan Morton-Blake

On behalf of the JUNO collaboration



NuFact 20/09/2024

Argonne National Laboratory, Chicago



Slide from Iwan's parallel talk

Why a reactor experiment at an accelerator devoted conference?

Neutrino physics

Detector technology

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1. Gigantic detector \longrightarrow engineering challenges
2. Strict requirements on energy resolution for NMO determination
3. Strict requirements on internal radiopurity especially for solar neutrino analysis

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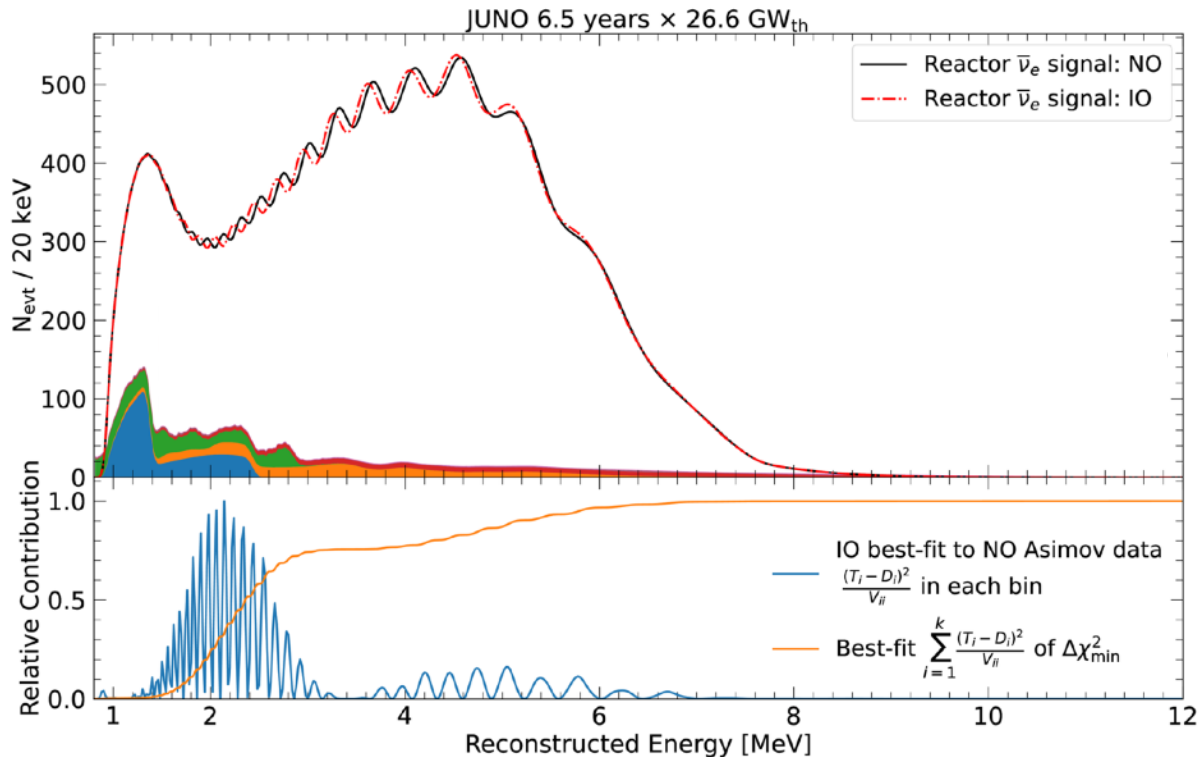
The aim of this talk is
describing how we will meet
the physics requirements

Slide from Iwan's parallel talk

What do we need to measure the NMO?

From: [2405.18008] Potential to Identify the Neutrino Mass Ordering with Reactor Antineutrinos in JUNO (arxiv.org)

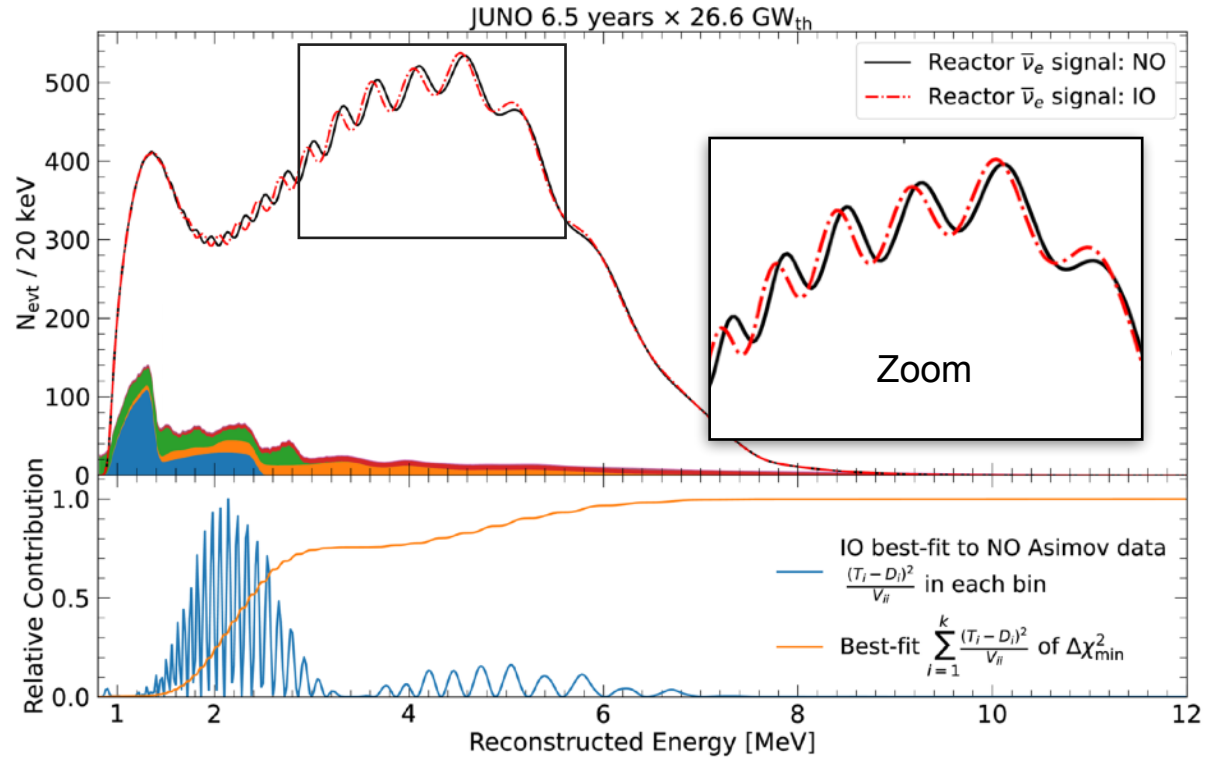
1) Energy resolution



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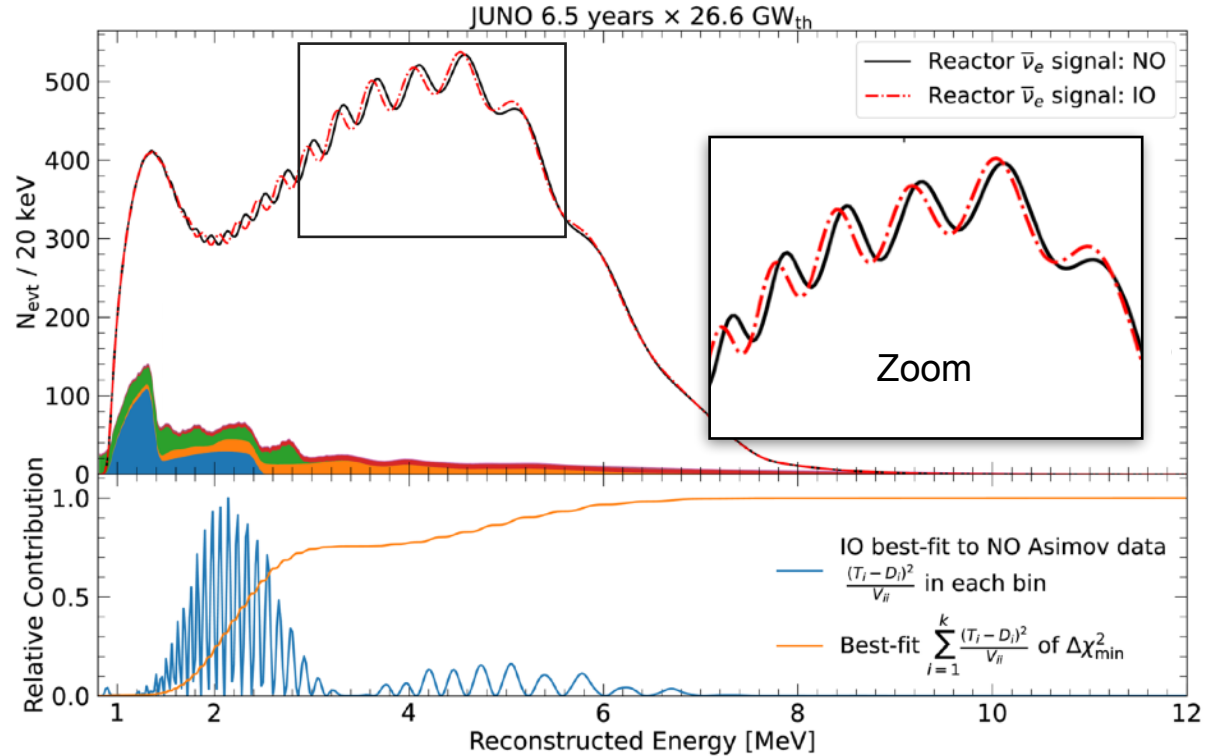


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→ High number of detected photons



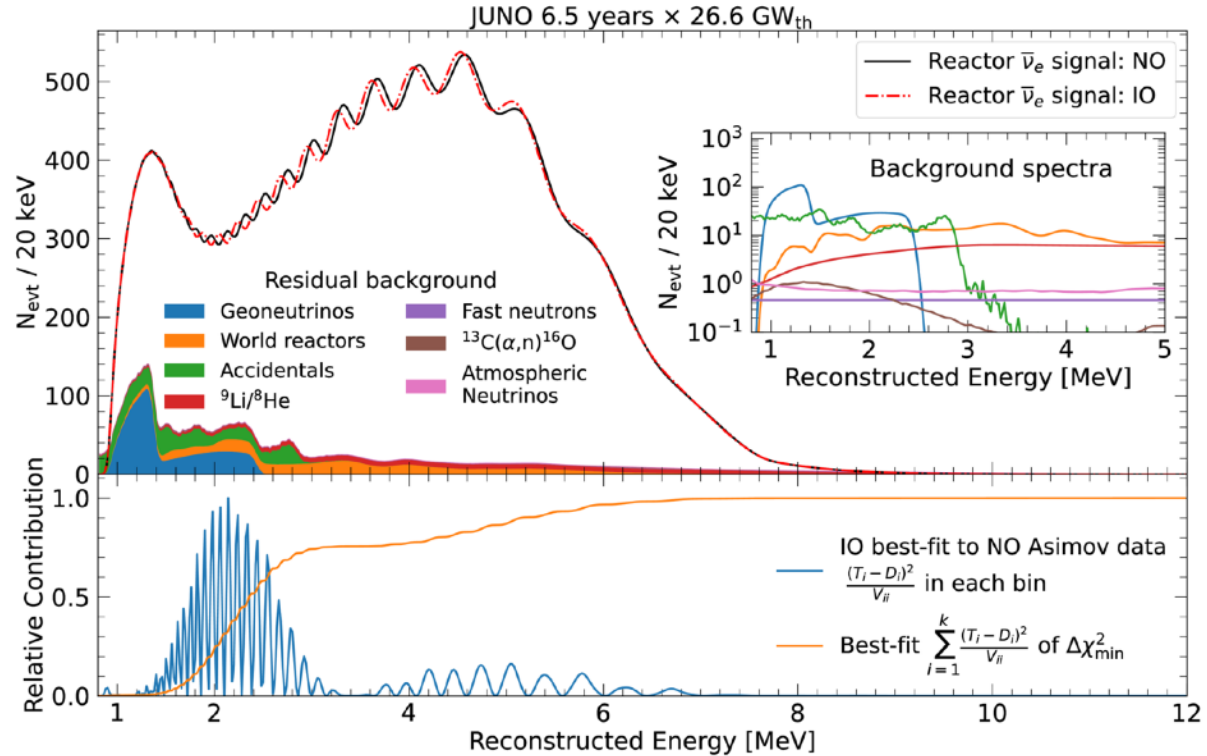
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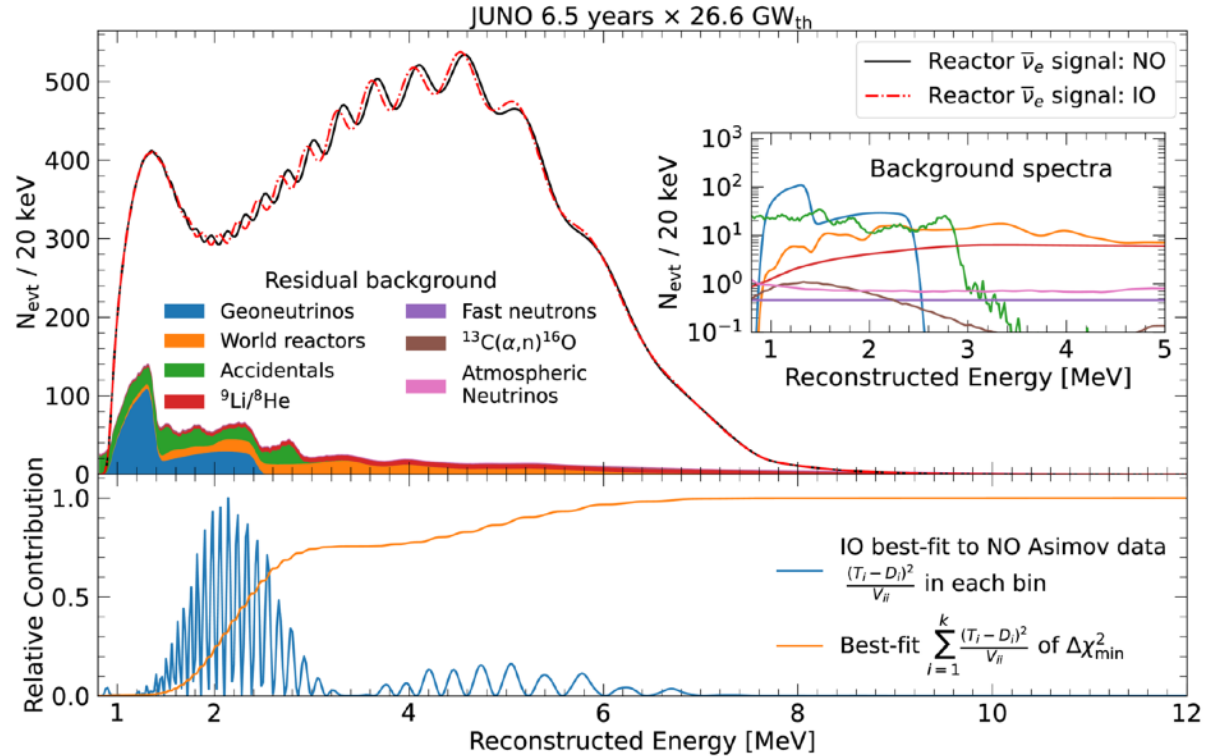
1) Energy resolution

→ High number of detected photons

2) High statistics

→ Low backgrounds

→ Huge mass



What do we need to measure the NMO?

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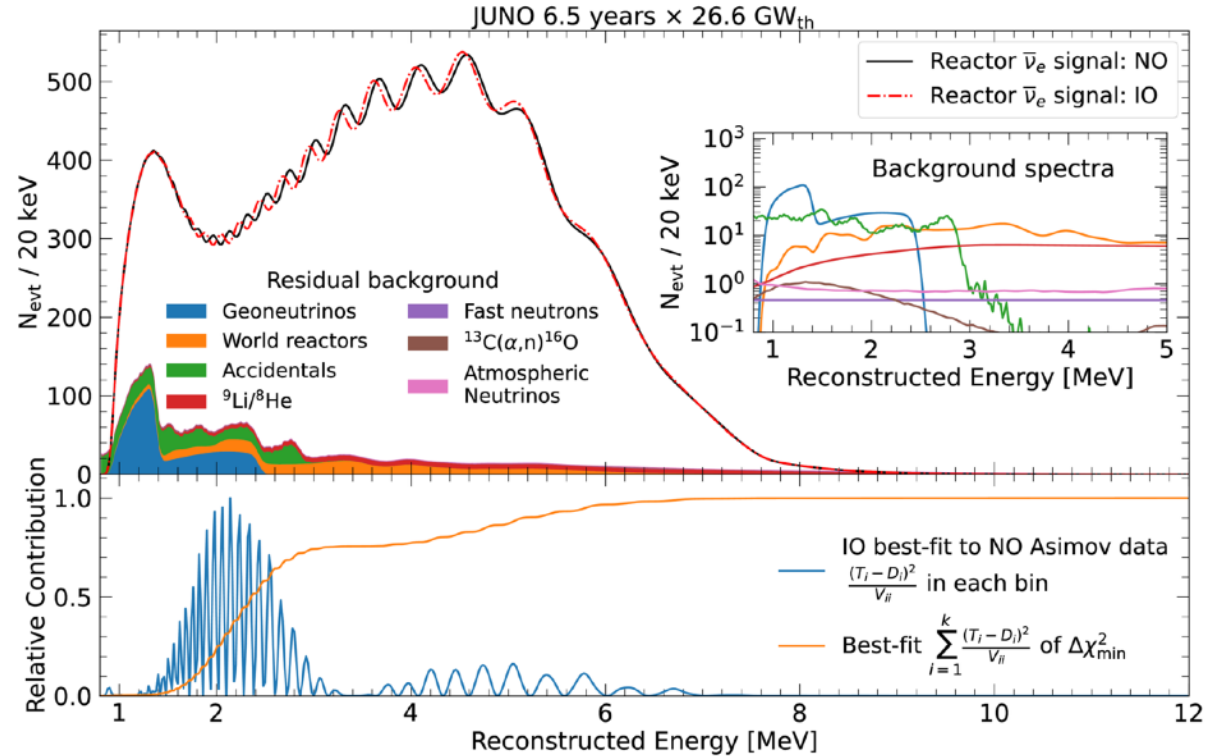
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3) Precise and accurate knowledge of the detector



What do we need to measure the NMO?

1) Energy resolution

- High number of detected photons

2) High statistics

- Low backgrounds
- Huge mass

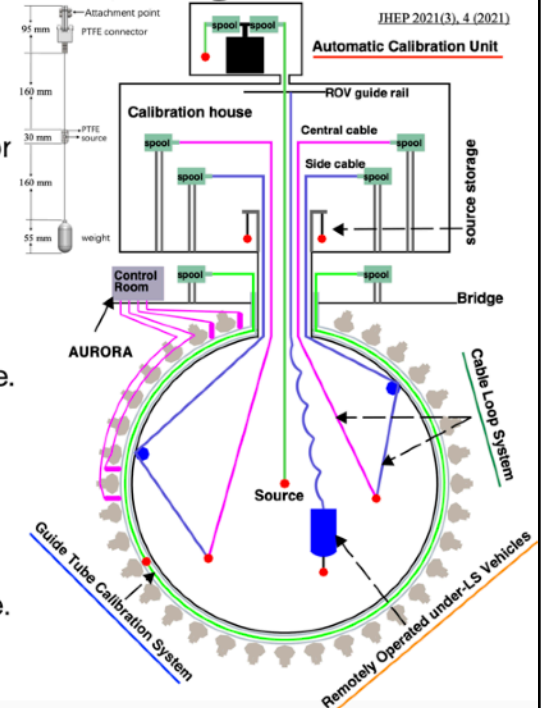
3) Precise and accurate knowledge of the detector

- Multi calibration campaign

JUNO Calibration System

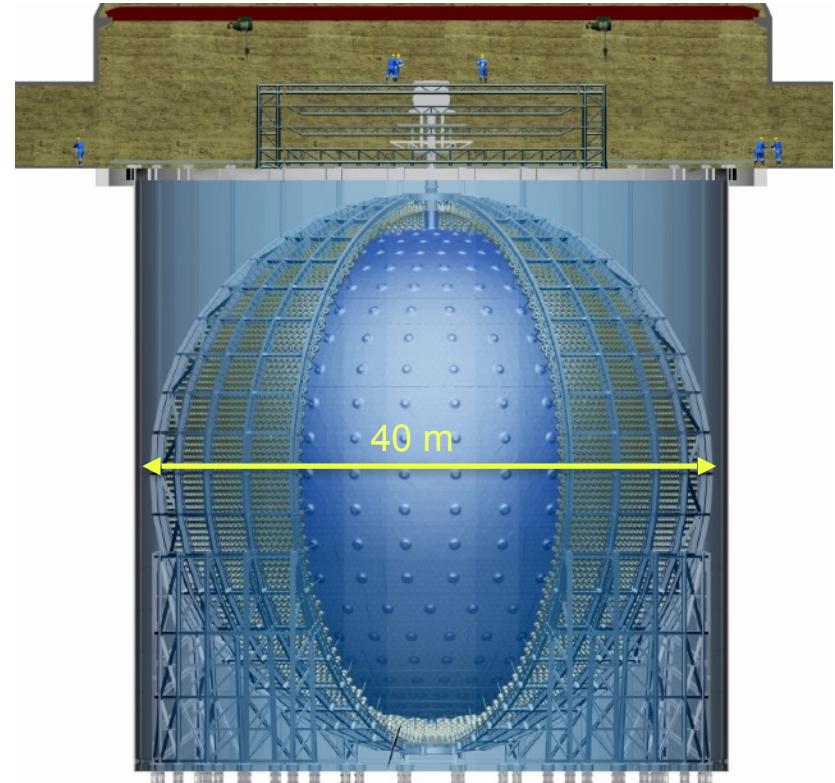
- Multiple calibration source deployment devices will be installed, placing a calibrator source at different positions:

- **Automatic Calibration Unit (ACU)** will cover the central axis.
- **Cable Loop System (CLS)** can cover the off-axis region in a two-dimensional plane.
- **Guide Tube Calibration System (GTCS)** will deploy the source on the outer surface of the acrylic sphere.
- **Remotely Operated Vehicle (ROV)** can access any position inside the LS volume.



Huge active mass

The Central Detector of the JUNO experiment is a gigantic sphere of 40 m of diameter which support all the parts of the detector:



Huge active mass

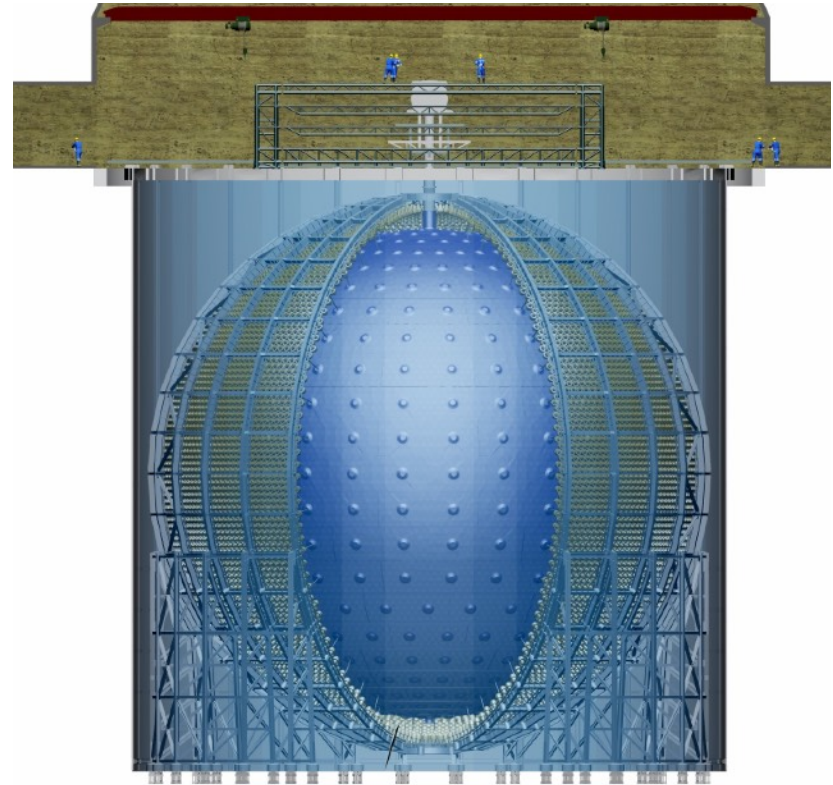
The Central Detector of the JUNO experiment is a gigantic sphere of 40 m of diameter which support all the parts of the detector:

_ More then **42000 Photo-Multiplier Tubes** with all the electronic boxes

_ An **acrylic sphere of 35.5 m of diameter** needed to contain the liquid scintillator

_ **20 000 tons** of an organic liquid scintillator:
LAB + 2.5 g/l PPO + 3 mg/l bis-MSB

All submerged in ultra-pure water



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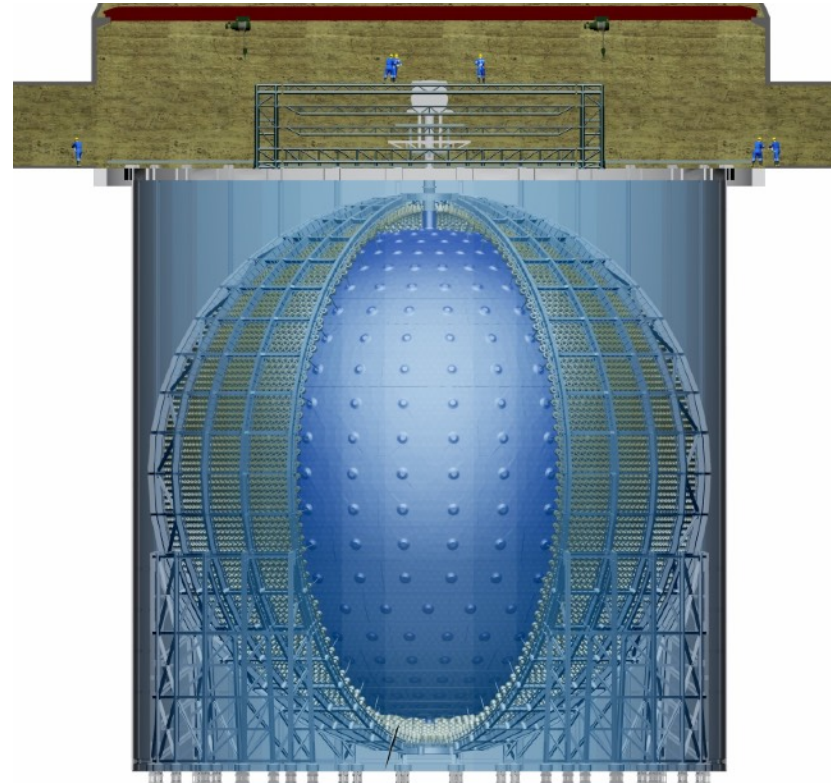
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→ **Detecting ~60 IBDs per days**



High Energy Resolution

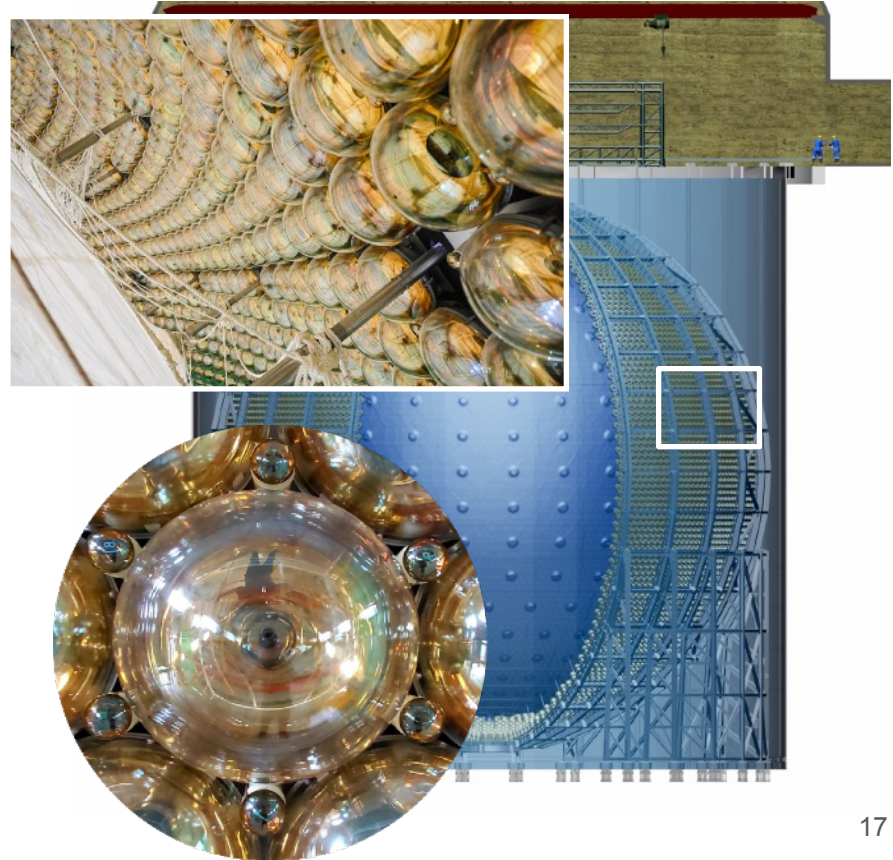
The energy resolution is related to the total number of photon detected. For these reason several strategies were adopted to increase this number as much as possible:

1. Exceptional optical coverage: 78 %

This is possible thanks to an enormous number of PMTs (42 000) divided in two system, small and large to fill the gaps between different PMTs

20" PMTs called Large PMTs

3" PMTs called Small PMTs



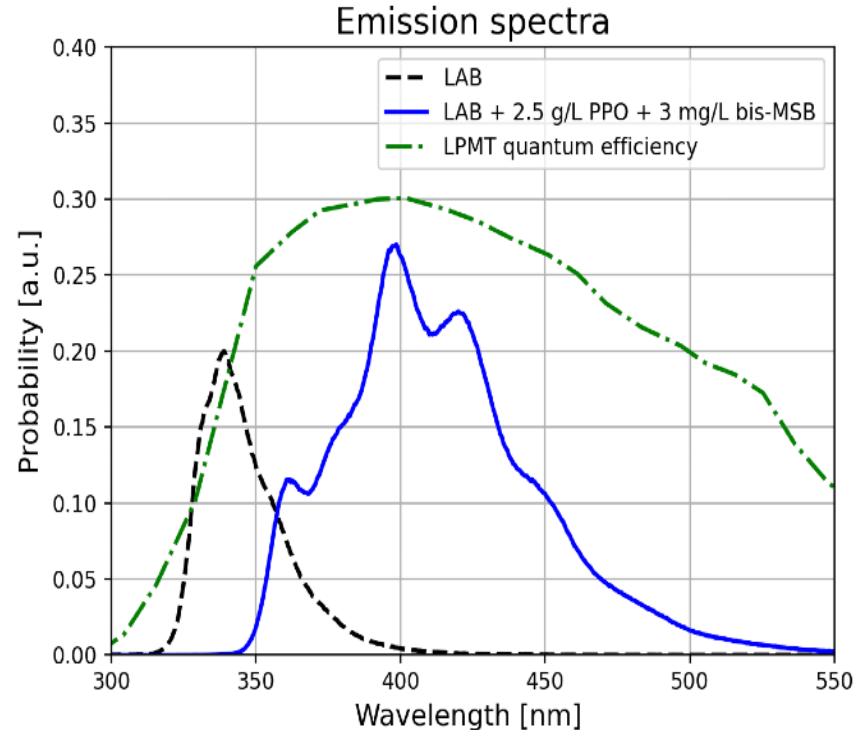
High Energy Resolution

The energy resolution is related to the total number of photon detected. For these reason several strategies were implied to increase this number as much as possible:

2. Good matching of the photon spectral emission with the PMTs detection efficiency

The JUNO liquid scintillator receipt: the addition of two elements (PPO and bis-MSB) move the emission of LAB in the optimal spectral region for Large PMTs

LAB + 2.5 g/l PPO + 3 mg/l bis-MSB



High Energy Resolution

The energy resolution is related to the total number of photon detected. For these reason several strategies were implied to increase this number as much as possible:

3. High transparency in the 400-420 nm region

Given the large JUNO dimensions, the scintillator absorption length must be larger than 20 m

A plant dedicated to the optical purification of the liquid scintillator is present, the Alumina Filtration Plant (AFP)



Low backgrounds

Reducing the backgrounds inside the JUNO detector is crucial not only for the purpose of measure the NMO but also for detecting geoneutrinos or solar neutrinos

Backgrounds could be divided in external or internal

Low backgrounds: external

Reducing the backgrounds inside the JUNO detector is crucial not only for the purpose of measure the NMO but also for detecting geoneutrinos or solar neutrinos

Backgrounds could be divided in external or internal

The main source of external backgrounds are muons and the material around the acrylic vessel

To reduce muons JUNO is build in an underground laboratory with **650 m of rock above**



Low backgrounds: external

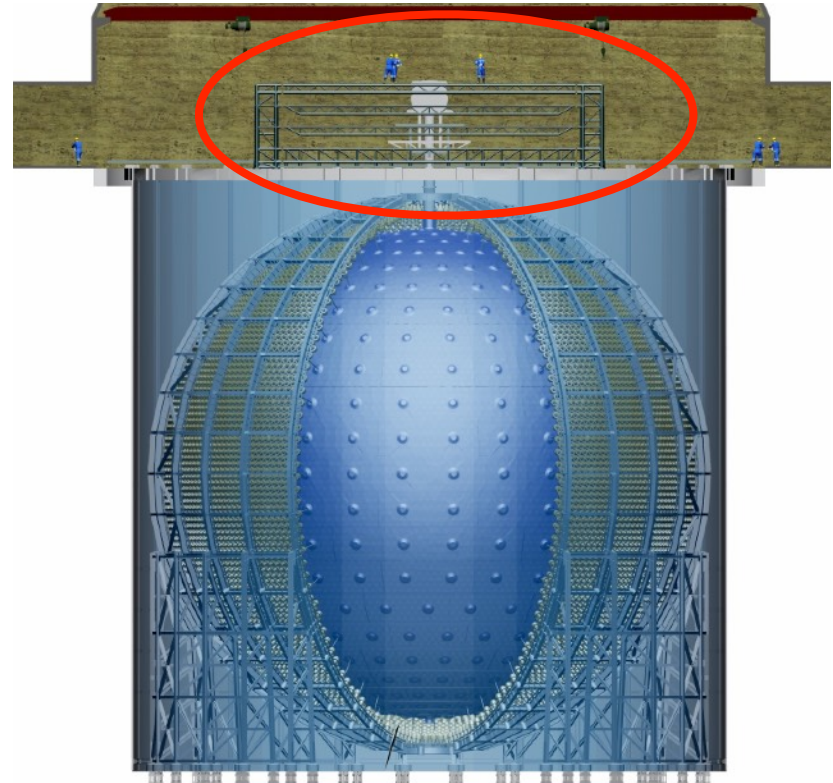
The main source of external backgrounds are muons and the material around the acrylic vessel

To reduce muons JUNO is build in an underground laboratory with **650 m of rock above**

A **top tracker** will be placed over the acrylic vessel to tag about 30 % of the muons

The water pool will be instrumented with more than **2400 LPMTs** to tag the **Cherenkov light**

→ Reaching the identification of **99.5 %** of muons

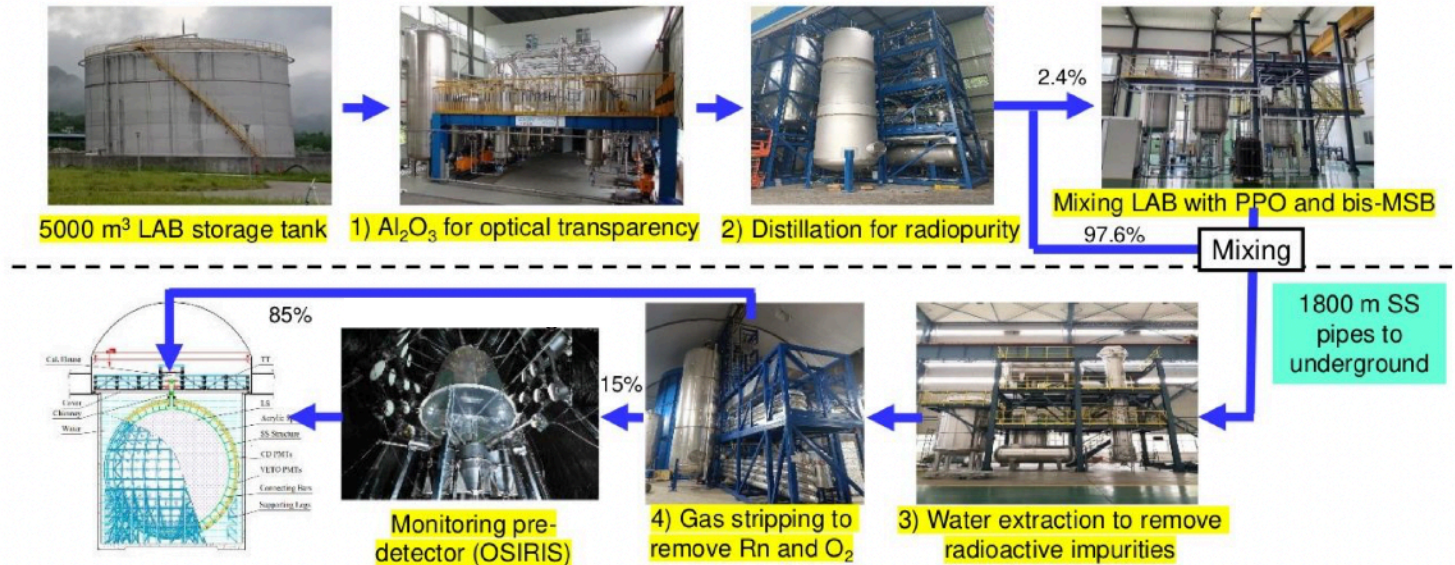


Low backgrounds: internal

To measure the **NMO** a limit of 10^{-15} g/g on the concentration of ^{238}U and ^{232}Th was set in the design phase

In addition a level of 10^{-16} - 10^{-17} g/g is aimed for the Solar neutrino campaign

→ An online purification chain is mandatory during the filling to reach those levels.



Status of the JUNO detector

From what we started ...

Experimental hall



... digging hole ...

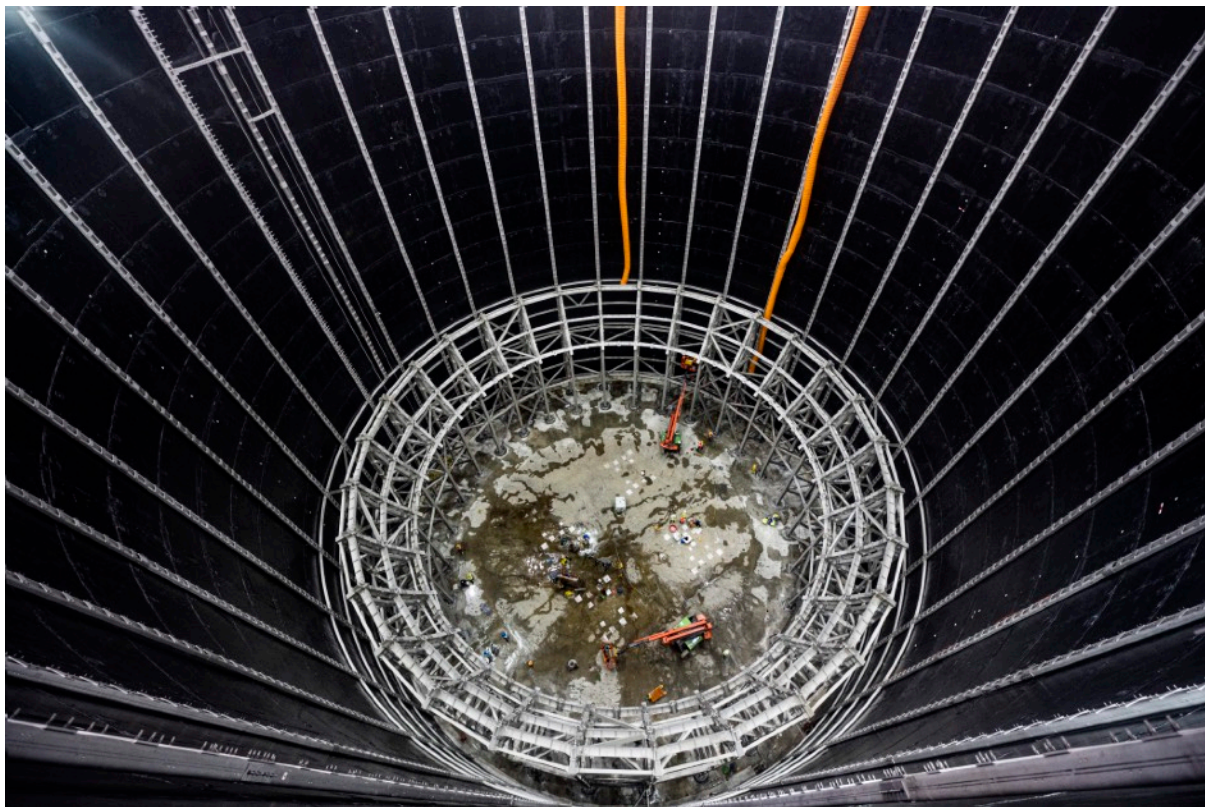
Experimental hall



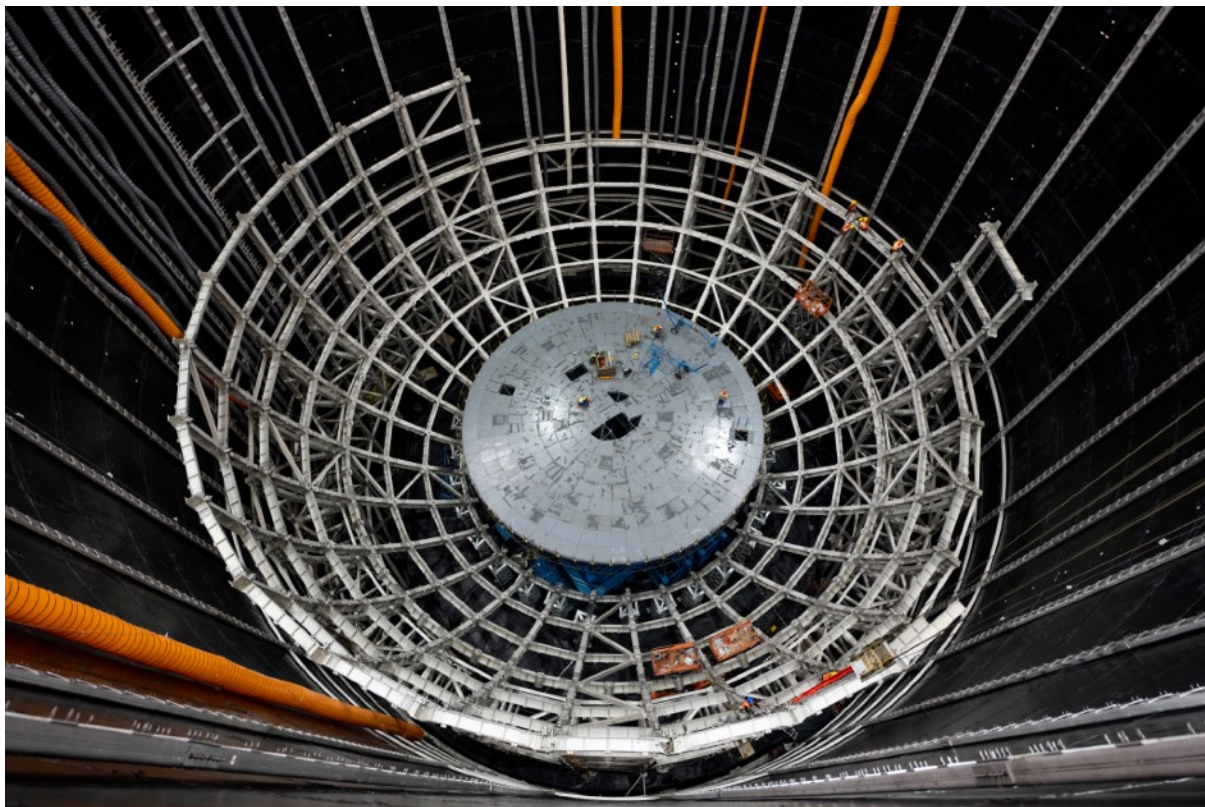
Excavation of the pool



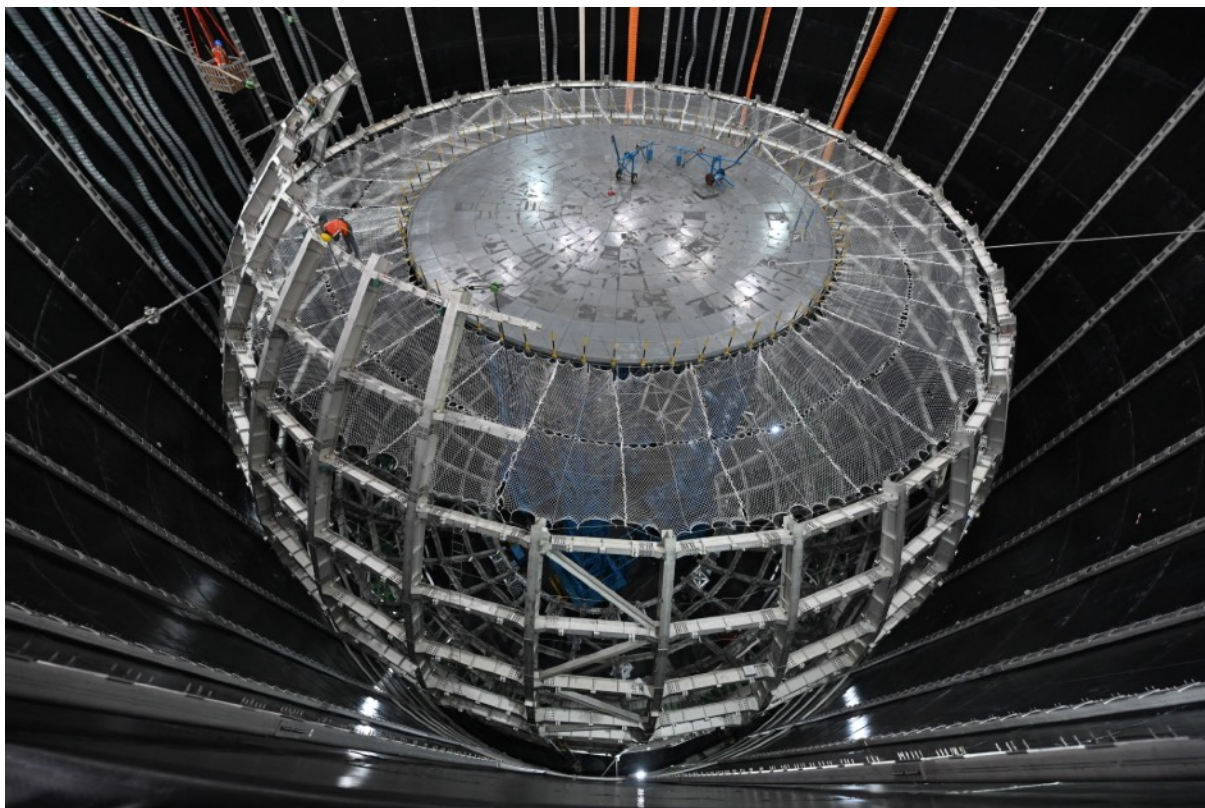
... we started to build the detector ...



... we started to build the detector ...



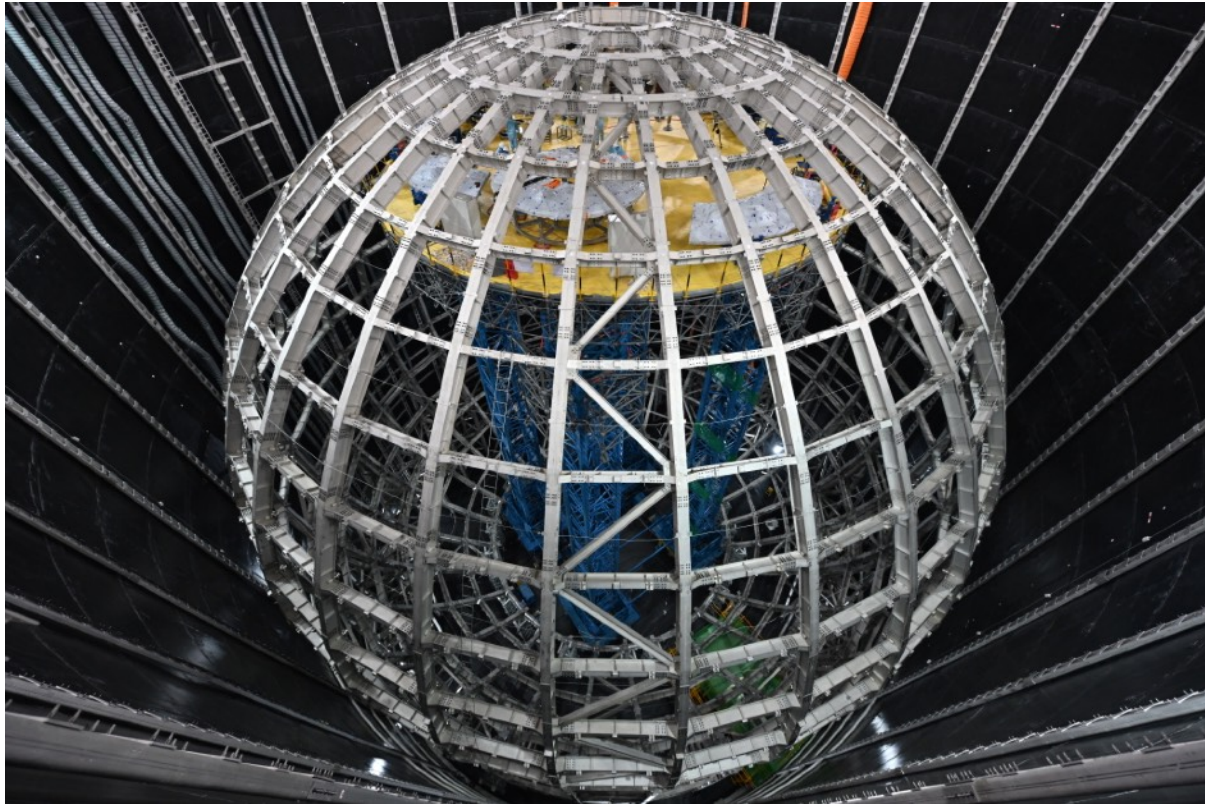
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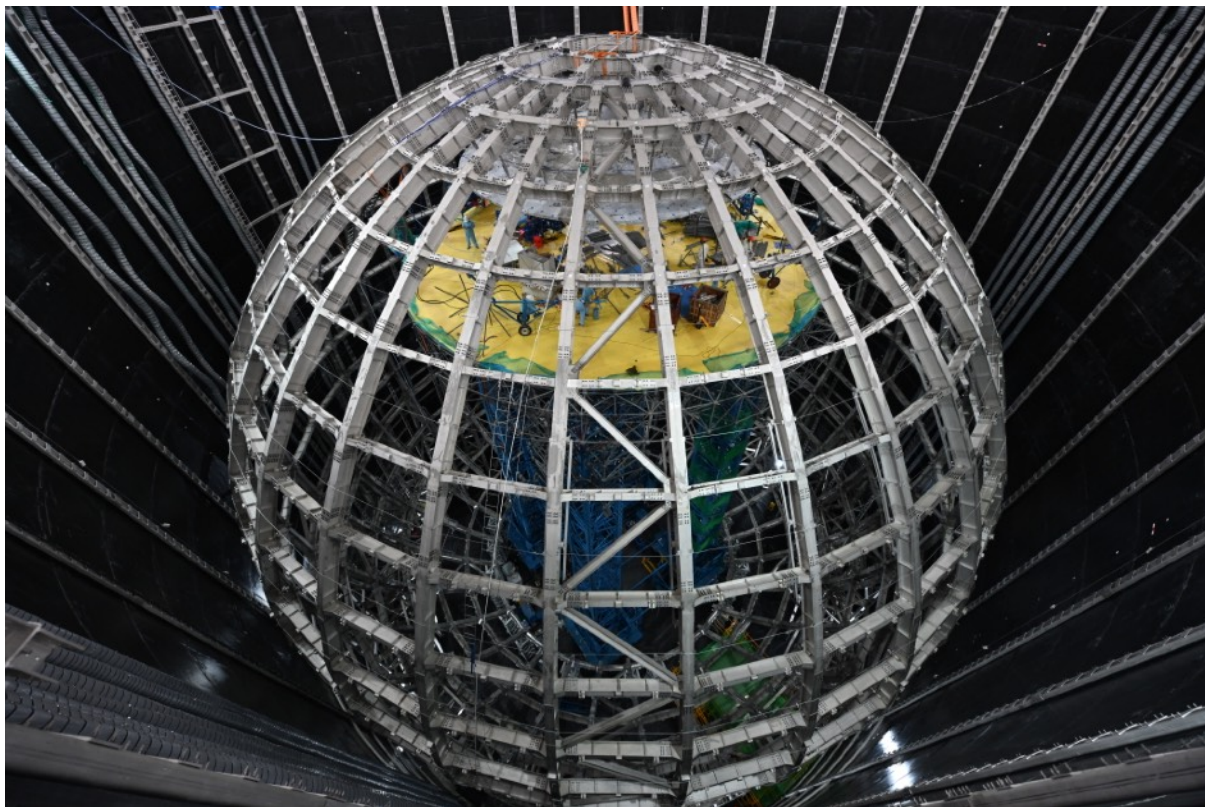
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... we started to build the detector ...



... to where we are now



Stainless **steel structure completed**
(Except 4 layers waiting for the acrylic)

17/23 acrylic layers completed:
_ production completed (<1 ppt U/Th/K contamination)
_ high transparency reached (>96%)

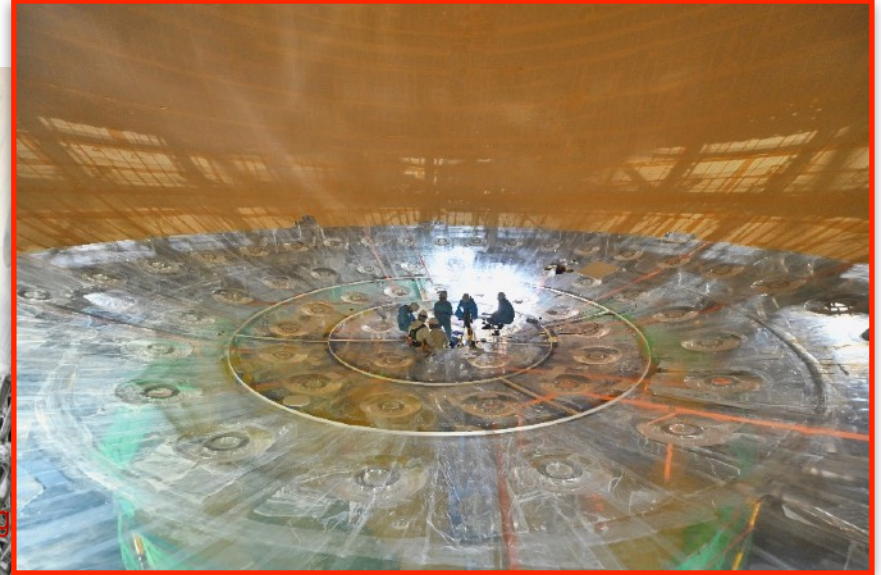
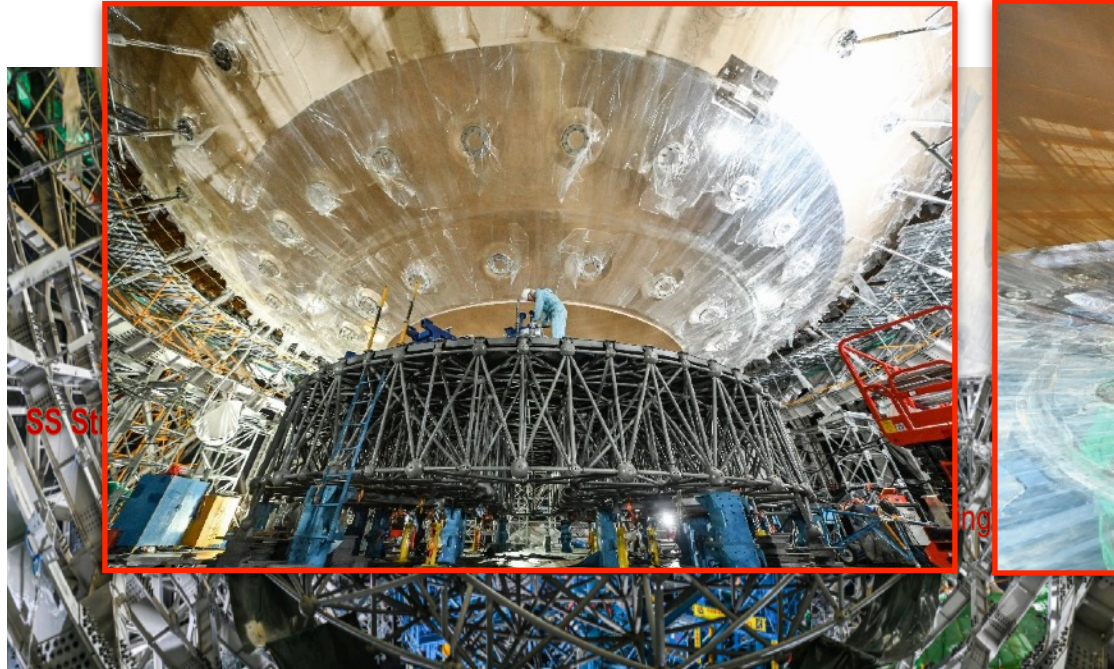
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Last week update

... to where we are now



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All PMTs tested and characterized

Dimension	Type	Number	Phot. Det. Eff.	Dark Noise	Transit time spread (1σ)
20" L-PMT	MCP-PMT (NNVT)	15,012	30.1%	31.2 kHz	7.0 ns
	Dynode PMT(Hamamatsu)	5,000	28.5%	17.0 kHz	1.3 ns
3" S-PMT	Dynode PMT (HZC XP72B22)	25,600	24.9%	0.5 kHz	1.6 ns

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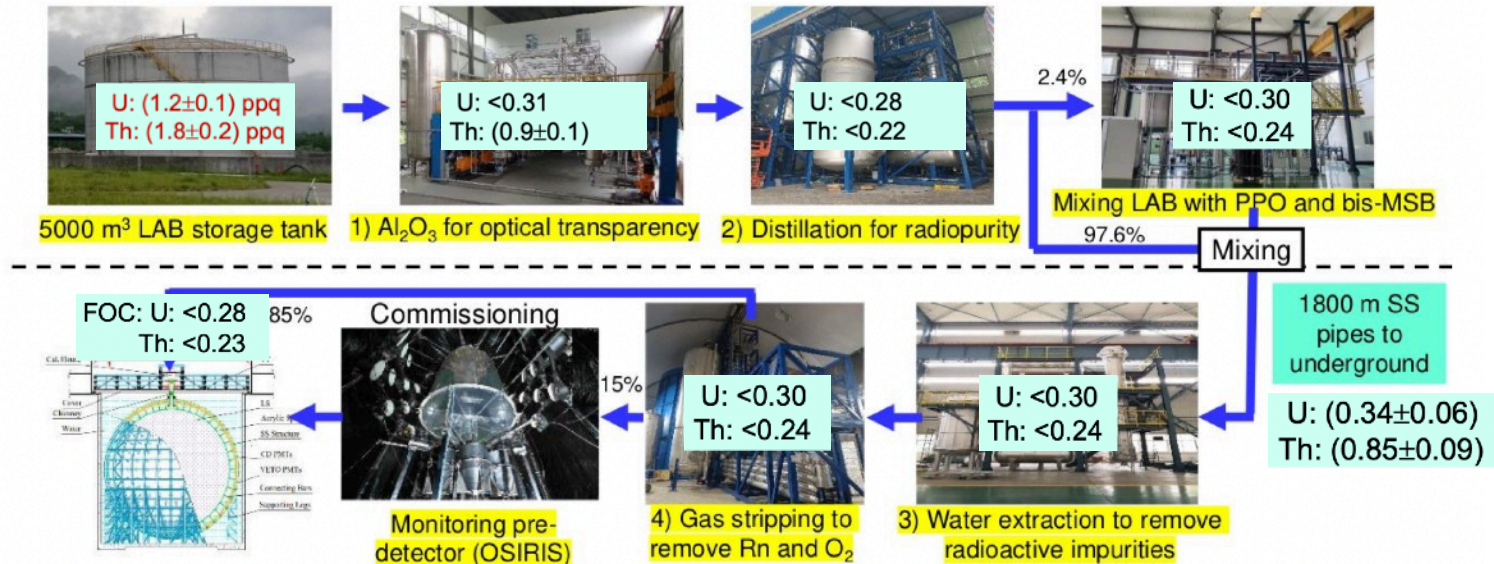
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Purification of the liquid scintillator

All the plants have been fully installed and commissioned.

4 campaigns of joint commissioning have been done testing the purification efficiency and the filling speed of $7 \text{ m}^3/\text{h}$ (six moth to fill JUNO).

Radiopurity of samples have been tested with NAA and ICMPs, with a sensibility of $\sim 10^{-15} \text{ g/g}$ level



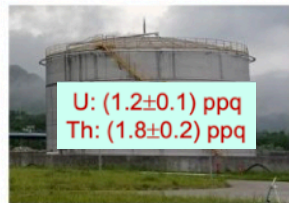
Purification of the liquid scintillator

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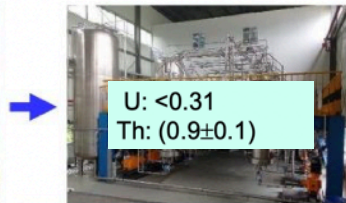
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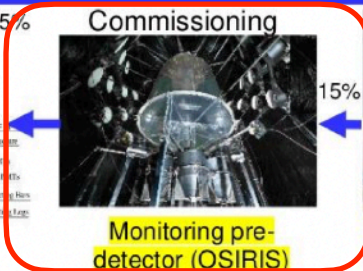
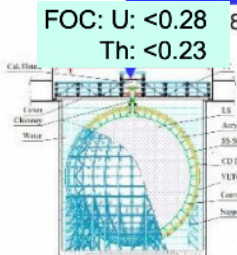
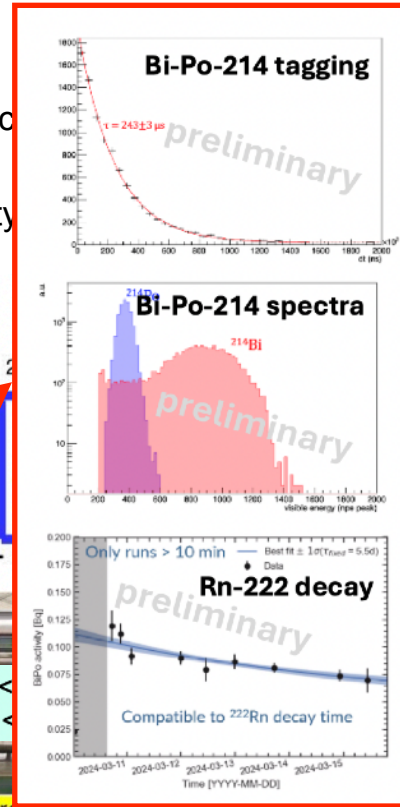
5000 m³ LAB storage tank



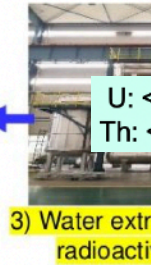
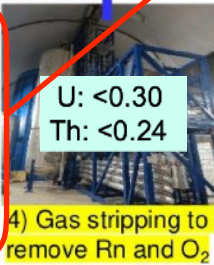
1) Al₂O₃ for optical transparency



2) Distillation for radiopurity



Monitoring pre-detector (OSIRIS)



Conclusion

JUNO is a colossal work of engineering

The construction is going to be completed in Fall 2024

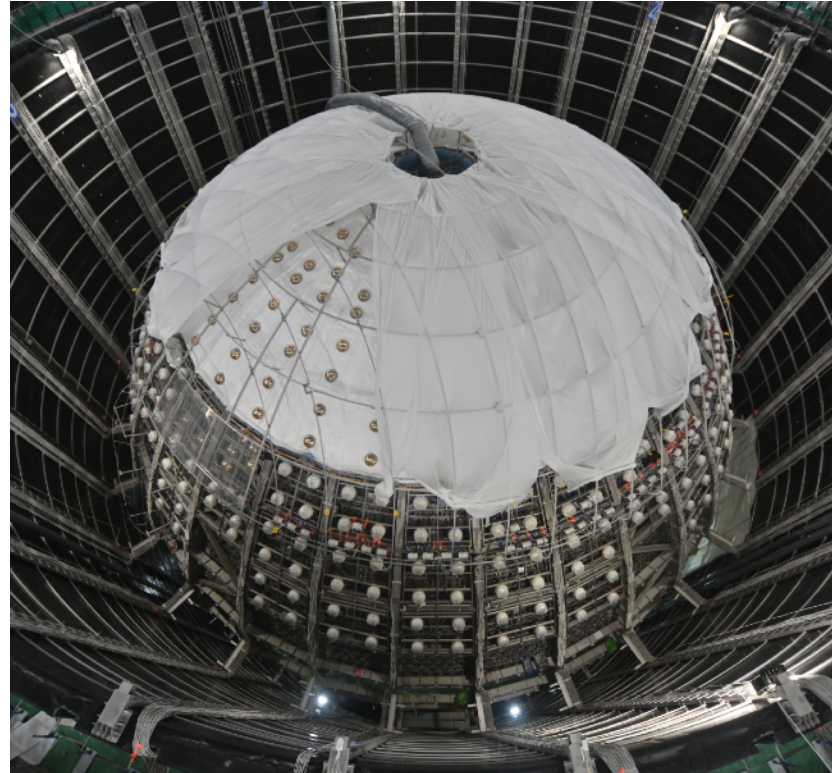
Filling with water will start before the end of 2024

**Filling with liquid scintillator will start in late Winter/
start of Spring 2025**

The data-taking will start in 2025

JUNO will be the biggest liquid scintillator

Stay tuned!



Thank you

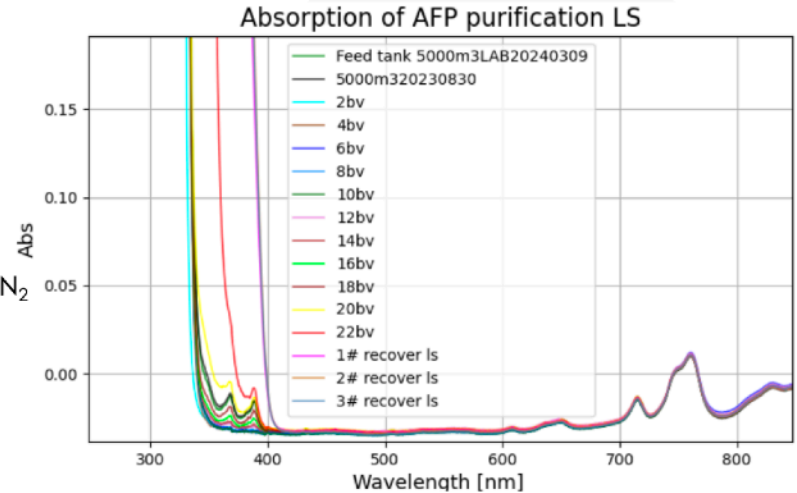


Since 2014, >700 collaborators from 74 institutions in 17 countries/regions

Backup: Alumina

► Main design parameters:

- Flow rate: 7 m³/h (2 BV/h/column, 8 total column with one always in maintenance)
- Al₂O₃ powder first batch purchased from Hunan company (first 20 tons of 500 tons required)
 - Particle size: between 30 and 300 μm
 - Unit surface area: >150 m²/g
 - Powder purity: = 99.5%
 - Average pore width: 5 nm ~ 6 nm
 - Active powder volume: 0.5 m³/column
 - Low powder radioactivity: < 0.3 Bq/kg in ²³⁸U and ²³²Th
 - Transportation and storage in double aluminum bags under N₂
- Wasted alumina recycled by vendor
- LAB pumping pressure: 15 bar
- Column Diameter: 600 mm
- Column H:D ratio = 3:1
- Two stage Filters: 220 nm + 50 nm (Pall / Cobetter)
- Radon contribution from Al₂O₃ purification system: < 10 mBq/m³.



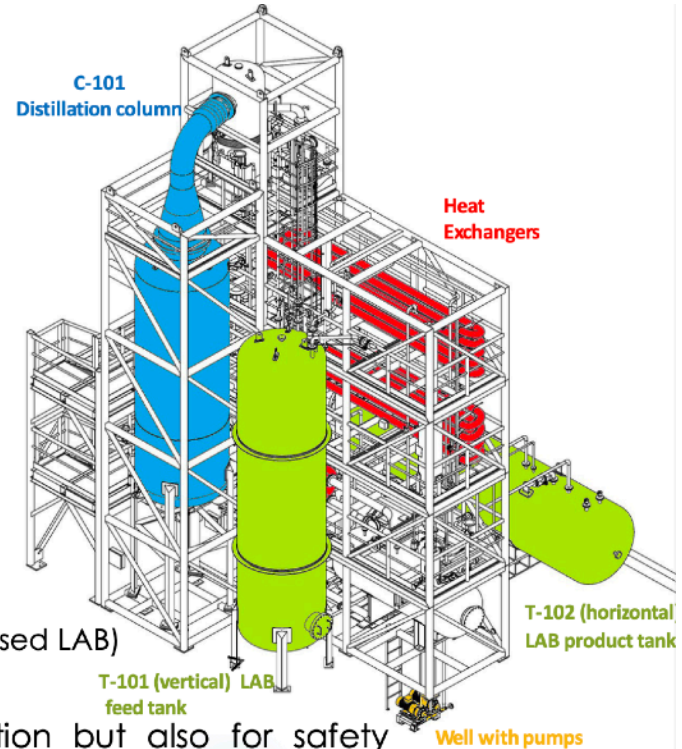
Backup: Distillation

► Main design parameters:

- Flow rate: 7 m³/h
- Column height: 7 m
- Column diameter: 2 m
- Number of trays: 6 Sieve trays
- Pressure on column top: 5 mbar
- Temperature at the reboiler = 220 °C
- Internal reflux: ~30%
- Bottom discharge: 1 – 2 %
- Number of theoretical stages: 4 - 5
- Heating Thermal Power(Hot Oil): 100 kW_{th}
- Water cooling tower: 1000 kW_{th}
- Heat exchanger energy recovery: 400 kW_{th} (feed / condensed LAB)
- Filters: 50 nm (Pall / Cobetter)
- Nitrogen blanket either to avoid oxidation/contamination but also for safety reason (LAB temperature > flash point only inside the distillation column)



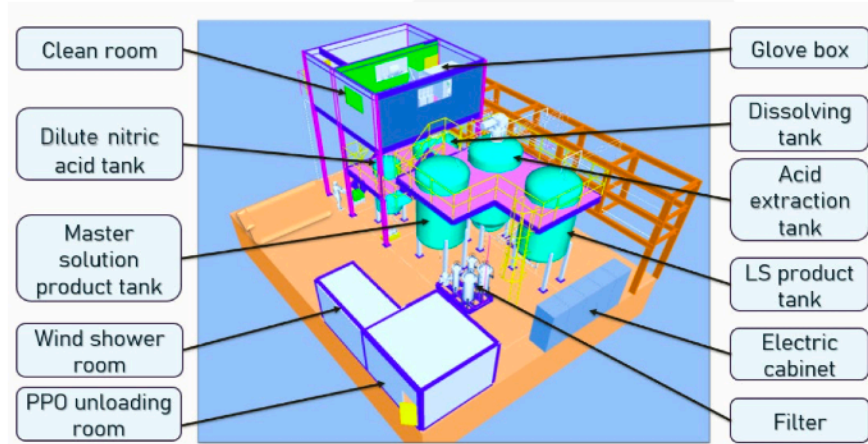
3500 holes (diameter 12 mm)



Backup: Mixing

► Main design parameters:

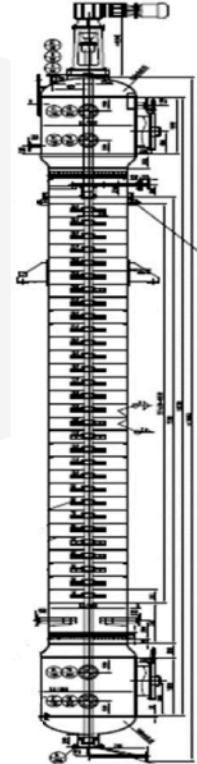
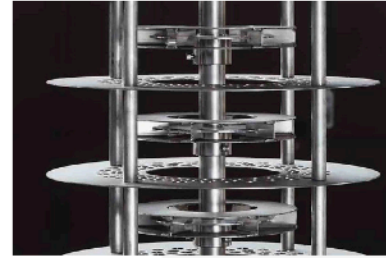
- Master Solution concentration: 105 g/L PPO
126 mg/L bis-MSB
- Dissolving temperature: 40 °C
- Method: batch mode with internal stirrer
- Acid washing:
 - 1 time with 1:2 (2 m³ acid solution)
 - 40 °C with 5% HNO₃
- Numbers of water washing= 2 times, 1:1
- bis-MSB transported and store in double aluminum bags under vacuum
- PPO transported and store in drums with plastic bags under vacuum



Backup: Water extraction

► Main design parameters:

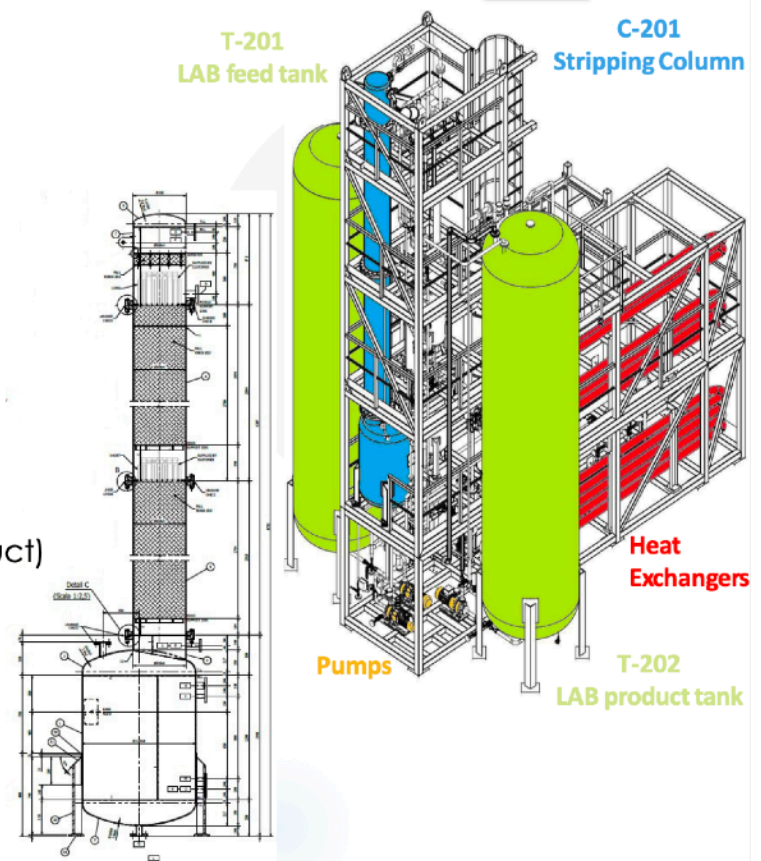
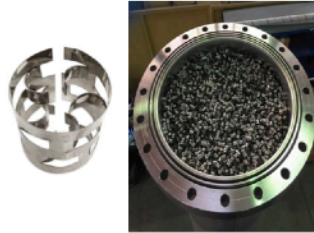
- Flow rate: 7 m³/h
- Column height: 13 m
- Column diameter: 1 m
- Column design: optimized Kühni turbine extraction tower with 30 stages of turbines connected in series on the shaft, separated into 30 chambers by 31 stages of porous trays.
- Ultra purity Water flow rate: 2.3 m³/h
- LS-water ratio: 3:1
- Extraction efficiency ≥ 5 theoretical equilibrium stages
- Rotation speed: 40 – 60 rpm
- LS temperature at the column = 40 °C
- HPN blanket either to avoid oxidation and Rn pollution
- Filters: 200 nm + 50 nm (Pall / Cobetter)



Backup: Stripping

► Main design parameters:

- Flow rate: 7 m³/h
- Column active height: 5.6 m
- Column diameter: 500 mm
- Unstructured packing: 13 mm stainless steel Pall rings
- Specific Interface Area: 430 m²/m³
- Number of theoretical stages: 3 - 4
- Pressure on column: 250 mbar
- LS Temperature at the column = 70 °C
- Heat exchanger energy recovery: 160 kW_{th} (feed / product)
- Heating Thermal Power (Hot Oil): 100 kW_{th}
- Chiller for cooling water: 200 kW_{th}
- Filters: 50 nm (Pall / Cobetter)
- HPN blanket either to avoid oxidation and Rn pollution



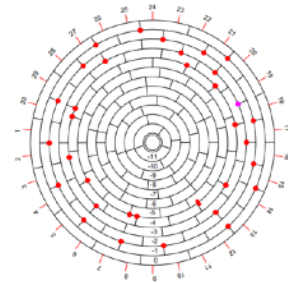
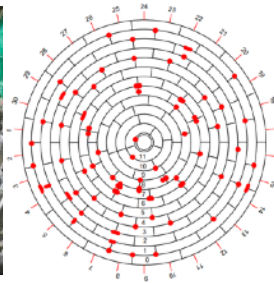
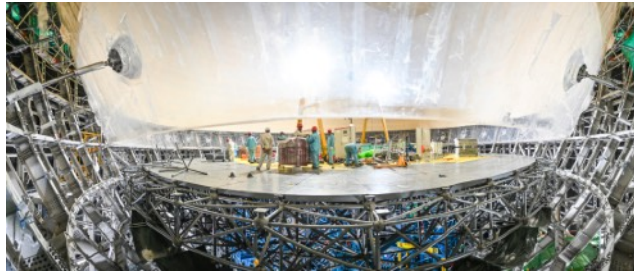
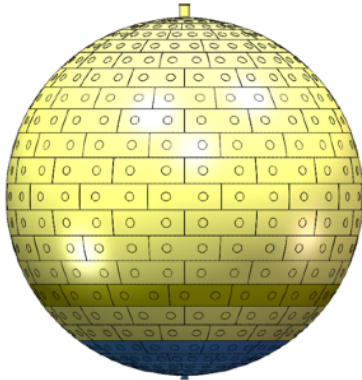
Backup: Acrylic

□ Production:

- Complete production of all panels in the factory: 263 panels + 2 chimneys;
- 256/263 panels transported onsite: layer -9# panels are ready onsite

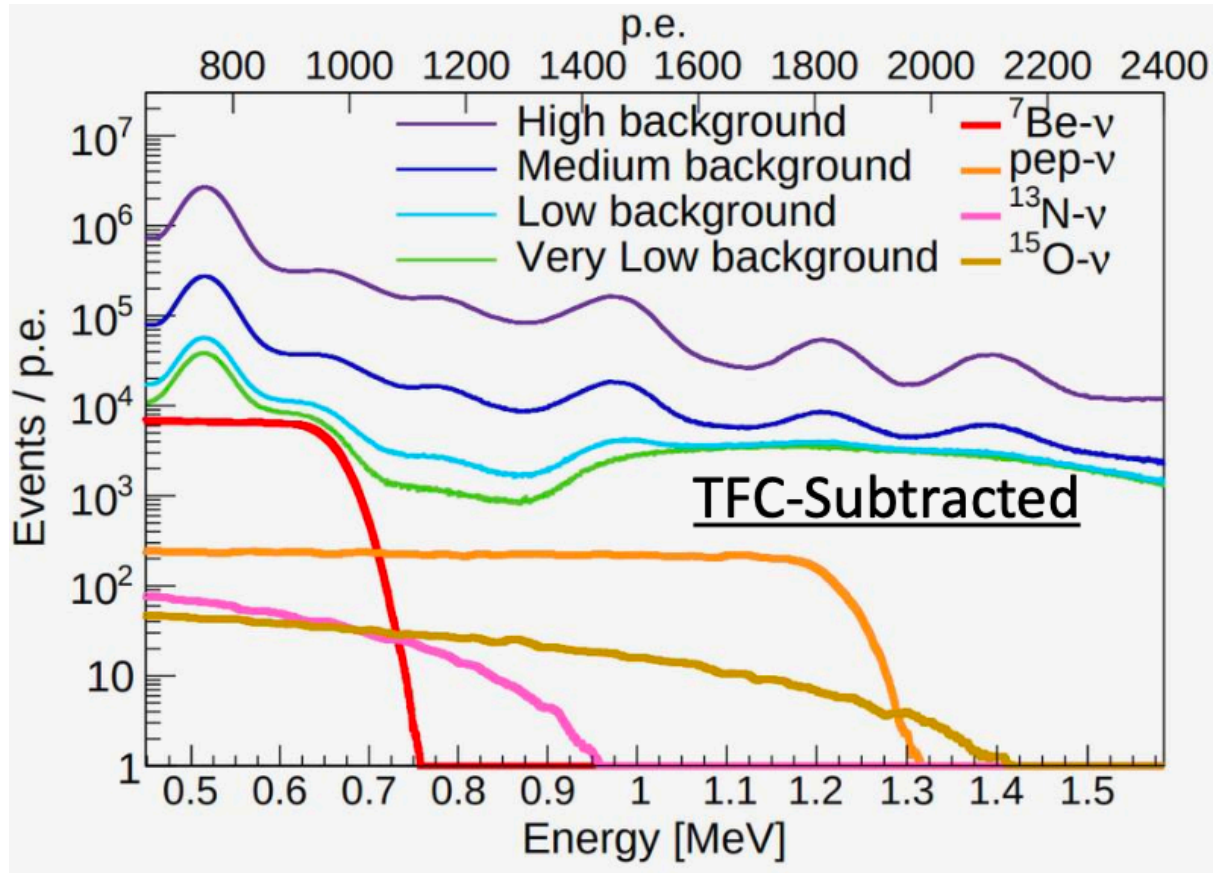
□ Construction:

- Completed the construction of layer -5# acrylic ring.
- Layer -6# is under bonding
- **Repairing works:** ~74 defects at the upper hemisphere, ~37 defects at the lower hemisphere. Now, only 1 defect is under repaired at the layer -5#
- The top layers of bonding have been thoroughly inspected, and regular inspections are planned.



Distribution of defects

Backup: Solar neutrino backgrounds



Backup: Energy resolution

arXiv:2405.17860

