Three-flavor oscillations and beyond the Standard Model physics with IceCube DeepCore

Funding from



Finn Mayhew on behalf of the IceCube Collaboration

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IceCube

[1]

- The IceCube Neutrino Observatory is a cubic-kilometer ice Cherenkov detector located at the south pole
- Consists of 5,160 photomultiplier tubes (PMTs) housed in digital optical modules (DOMs) buried 1.5-2.5 km deep under the glacial ice
- Operating continuously since 2011



IceCube DeepCore

[2]

- Cosmic rays strike the atmosphere and produce neutrinos (mostly v_µ's) that travel through the Earth and arrive at the detector from all upgoing directions ⇔ many baselines
- DeepCore consists of eight densely instrumented strings at the center of the array
- Gives the detector sensitivity to GeV-scale neutrinos => oscillation physics



Events

- Amount of light deposited in the detector gives us a handle on neutrino energy
- Relative hit timing from DOMs allows reconstruction of neutrino direction
 - \circ Direction parametrized as $\theta_{\rm zenith},$ direction's angle from vertical
 - Determines neutrino's baseline
- Topology/PID is the characteristic pattern of light in the detector
 - "Tracks" are characteristic of v_{μ} CC interactions, all other interactions are "cascades"



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Fitting

- Binning: neutrino energy, $\cos(\theta_{\text{zenith}})$, topology
- For non-analytic systematics, generate Monte Carlo (MC) at different hypotheses and interpolate between predicted event counts
- Minimize a likelihood function (forward

1.00 1.0 survival probability 0.75 0.8 0.50 $\begin{array}{c} 0.25 \\ 0.00 \\ 0.00 \\ -0.25 \end{array}$ 0.6 و 0.4 -0.500.2 -0.75 -1.00 0.0 10² 10¹ **10**⁰ E_v[GeV]



Atmospheric v



Systematic signature example

Effect of pulling the flux shape parameter $\Delta \gamma_{v} 0.1 (1\sigma)$



mid-pid

 NON-NECLIPATION PROPORTION PROPORTI PROPORTI PROPORTI PROPORTION PROPORTION PROPORTION PROPORTION PR

high-pid



 $\frac{N_{\text{pulled}} - N_{\text{nom}}}{\sqrt{N_{\text{nom}}}}$

Systematics

- Flux, cross section, detector response, ice properties
- Detector response and ice properties are the most impactful
- Orange systematics are left free in the CNN-based v_{μ} disappearance fit and purple are held fixed

Systematics ranked, for the CNN-based v_{μ} disappearance analysis, in order of the maximum difference in test statistic between fits where the physics parameters are free and fits where the physics parameters are fixed to the injected truth in the case that the systematic is held fixed at nominal in the fits while being off-nominal in the injected truth.



v_{μ} disappearance with the golden event sample

- Oscillation parameters $\sin^2(\theta_{23})$ and Δm^2_{32} constrained
- Using an 8-year smaller sample (18k v_{μ} CC events) optimized to contain mostly direct photons and high v_{μ} CC purity
- Validates modern calibration and achievability of good data/MC agreement with modern selection and analysis tools



New convolutional neural network (CNN) reconstruction

- Allows for higher statistics (88k v_{μ} CC events)
- Used in the most recent v_{μ} disappearance analysis

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	$N_{\rm events}(9.3 {\rm yrs})$	% of MC sample
$\nu_{\mu} { m CC}$	88306	58.8
$\nu_e \mathrm{CC}$	35296	23.5
$\nu_{\tau} \ CC$	8772	5.8
ν NC	16981	11.3
atm. μ	917	0.6
Total MC	150272	-
data	150257	-

Resolution for neutrino energy and $\cos(\theta_{\text{zenith}})$



v_{μ} disappearance with the CNN-based sample

• World-competitive constraints on oscillation parameters

[4]



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Data vs MC by L/E for the CNN-based v_{μ} disappearance analysis

• Good data/MC agreement

Left/top: cascade-like topology bin Middle: mixed topology bin Right/bottom: track-like topology bin

[4]





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Light sterile neutrinos

- Experimental anomalies can be explained by a noninteracting "sterile" neutrino with eV-scale mass splitting with respect to the other flavors
- Extension of the PMNS matrix:

[5]

$$U_{\rm PMNS}^{3+1} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

• Signal is less disappearance of v_{μ} 's that pass through Earth's core

[5] Signal in track-like topology bin $|U_{\mu4}|^2 = 0.3$ $|U_{\mu4}|^2 = 0.3$ $\cos(\theta_z)$ $|U_{\tau A}|^2 = 0$ $\delta_{24} = 180$ -0.5-1.00.0 $|U_{14}|^2 = 0$ $|U_{\mu 4}|^2 = 0$ $\cos(\theta_z)$ $\delta_{24} = 180$ -0.5-1.0 $|U_{\mu4}|^2 = 0.25$ 0.0 - $|U_{\mu4}|^2 = 0.25$ $\cos(\theta_z)$ $|U_{\tau A}|^2 = 0.25$ $\delta_{24} = 180^{\circ}$ -0.5-1.0 10^{2} 10^{1} 10^{1} 10² Ereco (GeV) Ereco (GeV)

-0.1

0.0

fractional change in event rate

0.1

0.2

-0.3

-0.2

0.3

Light sterile neutrinos

- New result with world-competitive constraints
- Does not measure the value of the mass splitting due to sensitivity to average oscillation effects only

[5]



Non-standard interactions (NSI)

- Given that neutrino masses are so small, it's possible that they get their mass from a mechanism outside of the standard model
- This mechanism may imply the existence of a new boson that interacts with neutrinos
- Neutrinos traveling through the Earth would then interact with the Earth through this boson, affecting their oscillations
 [6]



Effect of variations of three NSI parameters on neutrino oscillation

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[7]

Difference in atmospheric $v_{\mu} \rightarrow v_{\tau}$ probabilities between the standard interactions case and an NSI case

NSI, standard parametrization

• Parametrize possible effects generally as perturbations on the matter Hamiltonian (in the flavor basis):

$$\mathcal{H}_{\mathrm{mat}} = V_{\mathrm{CC}}(x) \begin{pmatrix} 1 + \epsilon_{ee}^{\oplus} - \epsilon_{\mu\mu}^{\oplus} & \epsilon_{e\mu}^{\oplus} & \epsilon_{e\tau}^{\oplus} \\ \epsilon_{e\mu}^{\oplus *} & 0 & \epsilon_{\mu\tau}^{\oplus} \\ \epsilon_{e\tau}^{\oplus *} & \epsilon_{\mu\tau}^{\oplus *} & \epsilon_{\tau\tau}^{\oplus} - \epsilon_{\mu\mu}^{\oplus} \end{pmatrix}$$

- ϵ 's are fit one by one, assuming all others are zero
- An analysis with 9.28 years of data is ongoing

[8]



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[8]



NSI, generalized matter potential parametrization

• Alternative parametrization:

$$\begin{split} H_{\rm mat}(x) &= Q_{\rm rel} U_{\rm mat} D_{\rm mat}(x) U_{\rm mat}^{\dagger} Q_{\rm rel}^{\dagger} ,\\ D_{\rm mat}(x) &= V_{\rm CC}(x) {\rm diag}(\epsilon_{\oplus}, \epsilon_{\oplus}', 0) ,\\ U_{\rm mat} &= R_{12}(\varphi_{12}) R_{13}(\varphi_{13}) \tilde{R}_{23}(\varphi_{23}, \delta_{\rm NS}) ,\\ Q_{\rm rel} &= {\rm diag} \left(e^{i\alpha_1}, e^{i\alpha_2}, e^{-i(\alpha_1 + \alpha_2)} \right) . \end{split}$$

- Assuming no CP violation and that the experiment has most sensitivity to v_{μ} disappearance, reduce to involving only ε , ϕ_{12} , and ϕ_{13} , which are fit simultaneously
- An analysis with 9.28 years of data is ongoing [8]



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Preview: probing the interior of the Earth

Three analyses described more fully in the backup:

- 1. Quantify DeepCore's ability to see the Earth through oscillations
- 2. Verify that the Earth has a layered structure
 - Result: Earth's uniformity excluded at 1.4σ
- 3. Use neutrino data to improve the constraints on the densities of Earth's layers placed by Earth's mass and moment of inertia

Conclusion

- IceCube DeepCore probes a variety of physics questions
- Many world-competitive results have come or are coming soon, stay tuned!
- NMO results coming soon
- Results in backup: decoherence, heavy neutral leptons, layered structure of the Earth
- Sensitivities in backup: v_{τ} appearance, long-range interactions, invisible decay, Earth's matter effect, mass of the Earth, density of layers of the Earth
- See Josh Peterson's talk tomorrow for IceCube's projected sensitivities with the coming Upgrade!

Backup

v_{τ} appearance

- Ongoing analysis to measure the appearance of v_{τ} 's following atmospheric v_{μ} oscillation
 - Standard oscillations case is $N_{\nu\tau} = 1$
- Signal is an excess of cascade-like events over the no- v_{τ} hypothesis
- Uses a high-statistics maximum likelihood reconstructed sample
 - An analysis using the CNN-based sample is also ongoing
- Projected to have world-leading sensitivity



Decoherence

- If the metric of spacetime has a quantum description, its fluctuations would introduce quantum decoherence in neutrinos that pass through them
- This would impact oscillations at long baselines and high energies
- Signal is less flavor transition and non-unitary flavor evolution
- Previous limits improved by a factor of 30
 [9]

[9]



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Heavy neutral leptons (HNLs)

- Search for steriles produced via v_{τ} upscattering
- Theoretical signature is a low-energy double cascade
 - This analysis searched for an excess of single cascades because identifying double cascades is not feasible with current tools
- Constraints placed for three HNL masses (m_4)
- Proof of concept for

HNL searches in IceCube

[10]

HNL mass	$ U_{\tau 4} ^2$	68 % CL	90 % CL	NH <i>p</i> -value
0.3 GeV	0.003	0.09	0.19	0.97
0.6 GeV	0.080	0.21	0.36	0.79
1.0 GeV	0.106	0.24	0.40	0.63







Layered structure of the Earth

- Gravitational and seismic studies give us the Preliminary Reference Earth Model (PREM) for Earth's radial density [11]
- Earth's matter effect is modified by its layered density structure
- Comparing the PREM hypothesis and the hypothesis that the Earth is uniformly dense, DeepCore rejects the uniform hypothesis at 1.4σ
- See [12] for more info

[13]



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Long-range interactions (LRI)

- If there is an undiscovered light boson that interacts differently with the different neutrino flavors and has a large interaction range, large quantities of matter (like the sun) could affect the oscillations of neutrinos
- Ongoing analysis to constrain the strengths of these interactions

[14]

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Invisible neutrino decay

- Ongoing analysis to constrain the possibility of neutrinos decaying into sterile neutrinos that do not mix with the known neutrinos
- Considering the case that v_3 decays into a sterile
- Modify the Hamiltonian to include $\alpha_3 = m_3/\tau_3$:

$$\begin{split} H_{\text{Total}} &= \frac{1}{2E} \left[U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} U^{\dagger} + U \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 - i\alpha_3 \end{pmatrix} U^{\dagger} \right] + \begin{pmatrix} V & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \\ V &= \pm \sqrt{2} N_e G_F \qquad \qquad \alpha_3 = \mathbf{m}_3 / \tau_3 \end{split}$$

[15]

Earth's matter effect

- Ongoing analysis to validate Earth's matter effect
- Compares the PREM hypothesis against the vacuum hypothesis
- Will quantify DeepCore's ability to see Earth's matter effect in oscillations data

[13]

Mass of the Earth

- Ongoing analysis to measure the mass of the Earth
- Compares the PREM hypothesis against a PREM hypothesis with all densities scaled
- M^v is Earth's mass measured from neutrino oscillations, M_e is Earth's mass from gravitational measurements

Density of layers of the Earth

- Ongoing analysis to study how neutrino data can improve upon constraints on the densities of layers of the Earth placed by Earth's mass and moment of inertia
- Compares the 5-layer PREM hypothesis against a 5-layer PREM hypothesis with the layer densities scaled while holding mass and moment of inertia constant

[13]

CNN-based sample v_{μ} rates

MC rates by particle and interaction type as a function of event selection stages from the CNN-based v_{μ} disappearance analysis

Parameter pulls in CNN-based v_{μ} disappearance analysis

Pearson correlation map

Correlation between systematics in CNN-based v_{μ} disappearance analysis

Nuisance parameters left free in the CNN-based v_{μ} disappearance analysis

The systematic uncertainty parameters included as nuisance parameters in the data analysis, along with their associated priors. The priors on parameters can either be Gaussian (in which case the value corresponding to $\pm 1\sigma$ is listed) or Uniform (in which case the allowed range is listed). [4]

[4]

Parameter	Nominal Prior width		Fit value	Pull (σ
Detector:				
DOM efficiency	+0%	$\pm 10\%$	+1.8%	0.18
Ice absorption	+0%	$\pm 5\%$	-3.5%	-0.71
Ice scattering	+5%	$\pm 10\%$	+1.8%	-0.32
Rel. eff. p_0	0.10	[-0.6, 0.5]	-0.14	-
Rel. eff. p_1	-0.05	[-0.2, 0.2]	-0.07	-
BFR efficiency	0.0	[0, 1]	0.48	-
Atm. flux:				
$\Delta \gamma_{\nu}$	0.0	± 0.1	-0.011	-0.11
$\Delta \pi^{\pm}$ yields I	0.0	$\pm 61\%$	+42%	0.68
$\Delta \pi^{\pm}$ yields G	0.0	$\pm 30\%$	-4.2%	-0.14
$\Delta \pi^{\pm}$ yields H	0.0	$\pm 15\%$	-12%	-0.81
ΔK^+ yields W	0.0	$\pm 40\%$	+4.2%	0.11
ΔK^+ yields Y	0.0	$\pm 30\%$	-6.9%	-0.23
Cross-section:				
$M_{\rm A}^{\rm CCQE}$	$0.99~{ m GeV}$	+25% -15%	-4.5%	-0.30
$M_{\rm A}^{\rm CCRES}$	$1.12 {\rm GeV}$	$\pm 20\%$	-3.9%	-0.20
DIS CSMS	0.0	± 1.0	0.12	0.12
Normalization:				
$A_{\rm eff}$ scale	+0%	[-90%, +100%]	-10%	-
Atm. muons:				
Atm. μ scale	+0%	$\pm 40\%$	-3.8%	-0.10

Brazil bands for the CNN-based v_{μ} disappearance analysis

Generated by injecting parameters at the best-fit point, applying statistical fluctuations to get pseudotrials, and running a scan of the test statistics over each physics parameter for each pseudotrial. The band shows the 68% and 90% regions of the test-statistic distribution at each scan point.

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