



Stockholm  
University



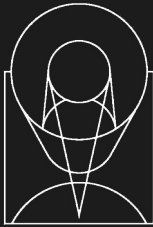
*Osaka Klein*  
centre



PRINCETON  
UNIVERSITY

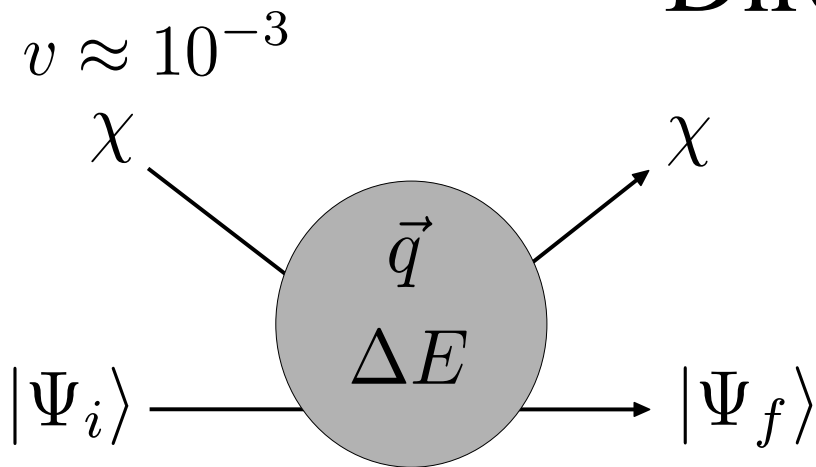
# X-Disciplinary Hunt for Dark Matter: ML and Material Science Meet Astroparticle Physics

CARLOS BLANCO



SPACE  
TELESCOPE  
SCIENCE  
INSTITUTE

# Direct Detection

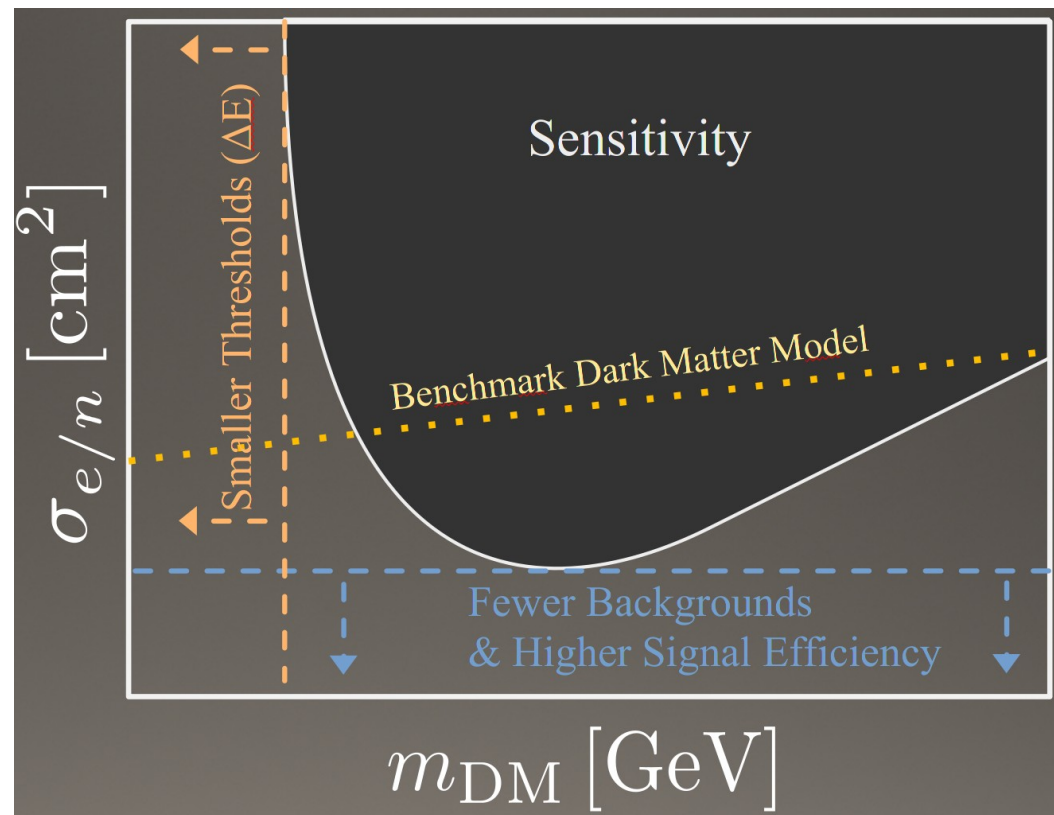


Energy Budget

$$\Delta E \sim v^2 m_\chi \approx \mathcal{O}(\text{eV}) \left( \frac{m_\chi}{1 \text{ MeV}} \right)$$

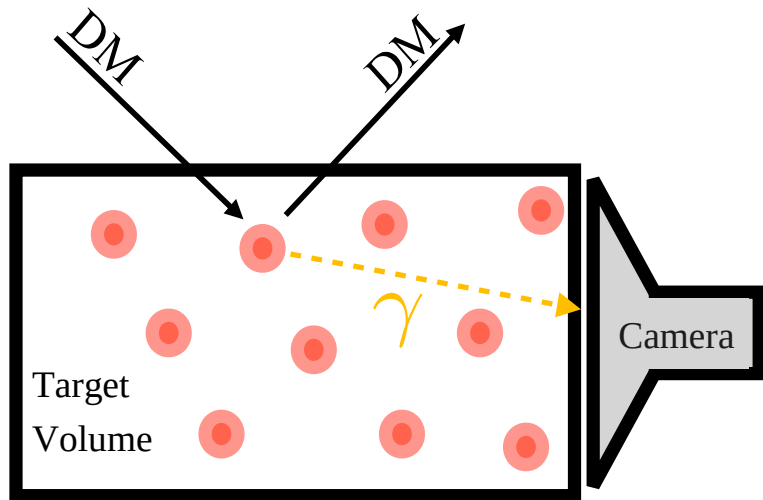
Momentum Budget

$$q \sim v m_\chi \approx \mathcal{O}(\text{keV}) \left( \frac{m_\chi}{1 \text{ MeV}} \right)$$



# Direct Detection: New Materials & Methods

Recoil-induced *fluorescence* (radiative deexcitation)



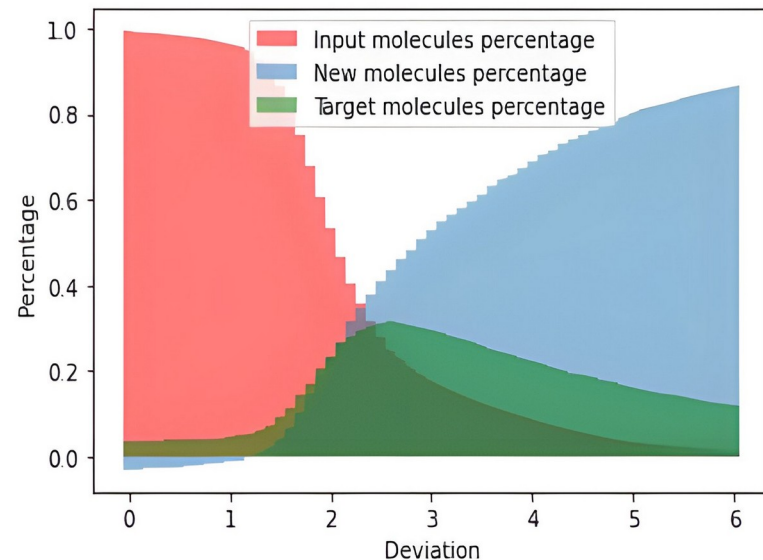
● unit that de-excites by emitting a photon

Could be:

- Nanostructures (quantum dots)
- Molecules in ordered crystal
- Hybrid material (QDs in Molecular matrix)



ML to explore vast data of material space



Generative ML models will identify novel materials that maximize signals

# Direct Detection: Interaction Rate

Rate *spectrum* (events in detector)

$$\Gamma \sim \int \frac{d^3\vec{q}}{q} \eta(v) |F_{\text{DM}}(q)|^2 |f_{i \rightarrow f}(q)|^2$$

Mean inverse velocity (Astrophysics)

$$\eta(v)$$

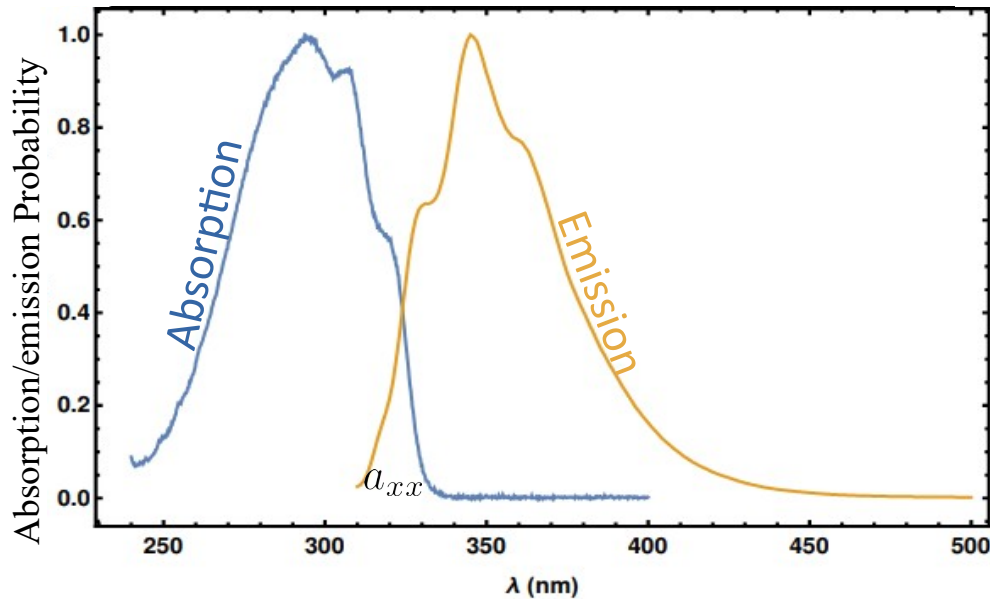
Transition form factor (Condensed matter / Chemistry)

$$f_{i \rightarrow f}(q) = \langle \tilde{\Psi}_f(k + q) | \tilde{\Psi}_i(k) \rangle$$

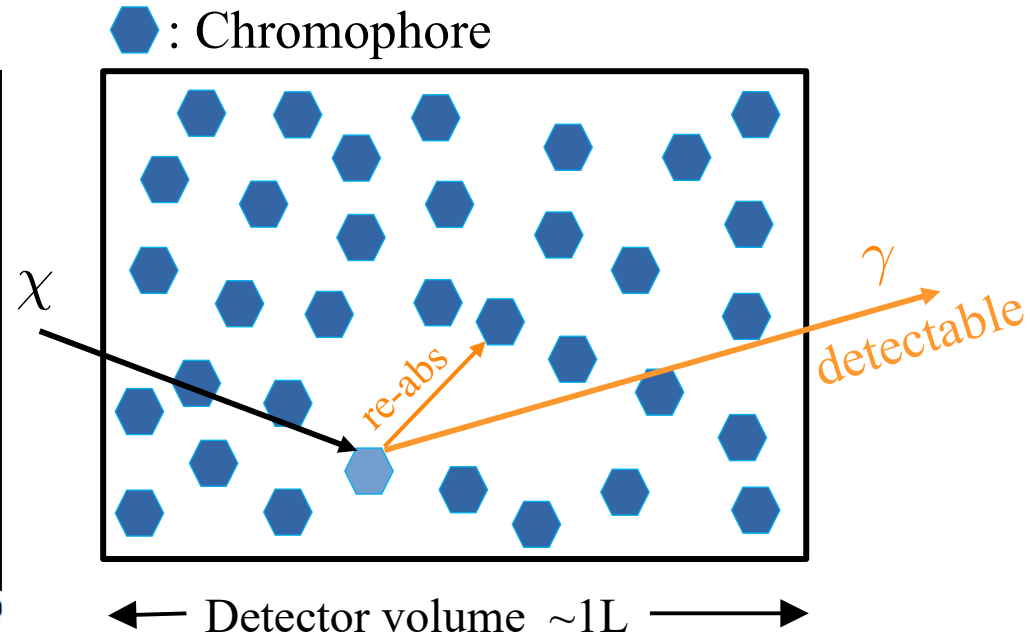
Dark matter form factor (Particle physics)

$$F_{\text{DM}}(q) \propto \begin{cases} 1 & , \text{Contact interaction} \\ \left(\frac{1}{q}\right)^2 & , \text{Long-range interaction} \end{cases}$$

# Fluorescence with DM



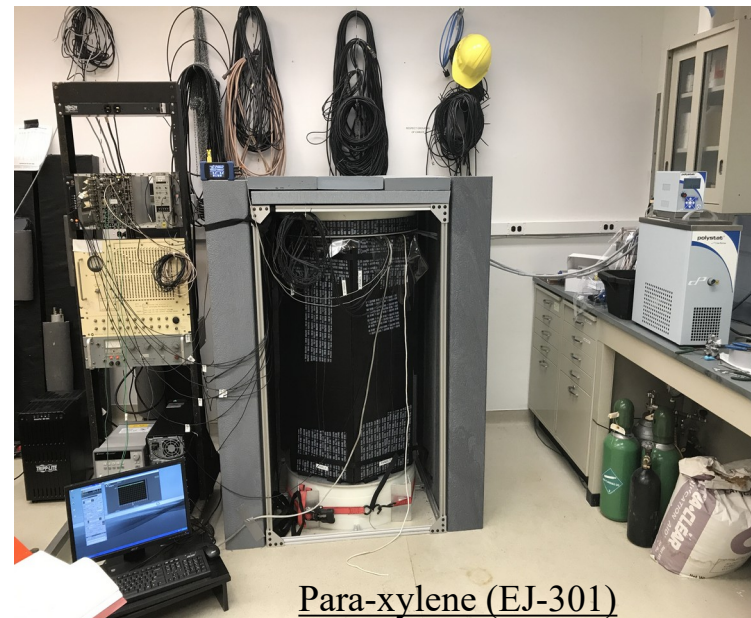
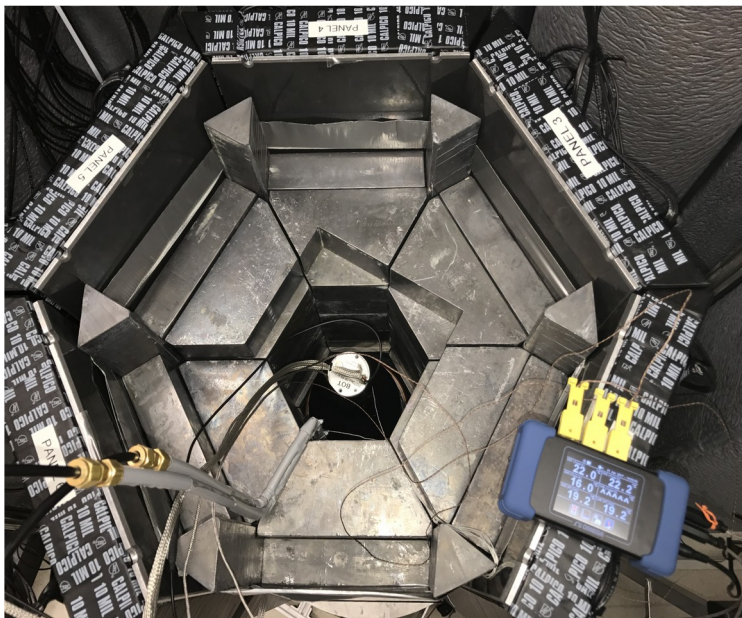
Decreasing energy (E) →



Probability for the photon to free stream

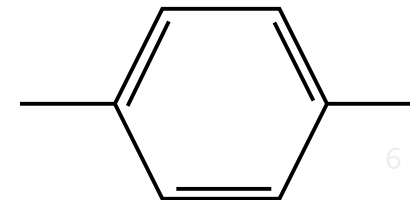
$$\Phi_{FB} \sim (1 - a_{xx}) \quad \text{e.g. molecular crystals: } \Phi_{FB} \approx 65\%$$

# First Experimental Setup

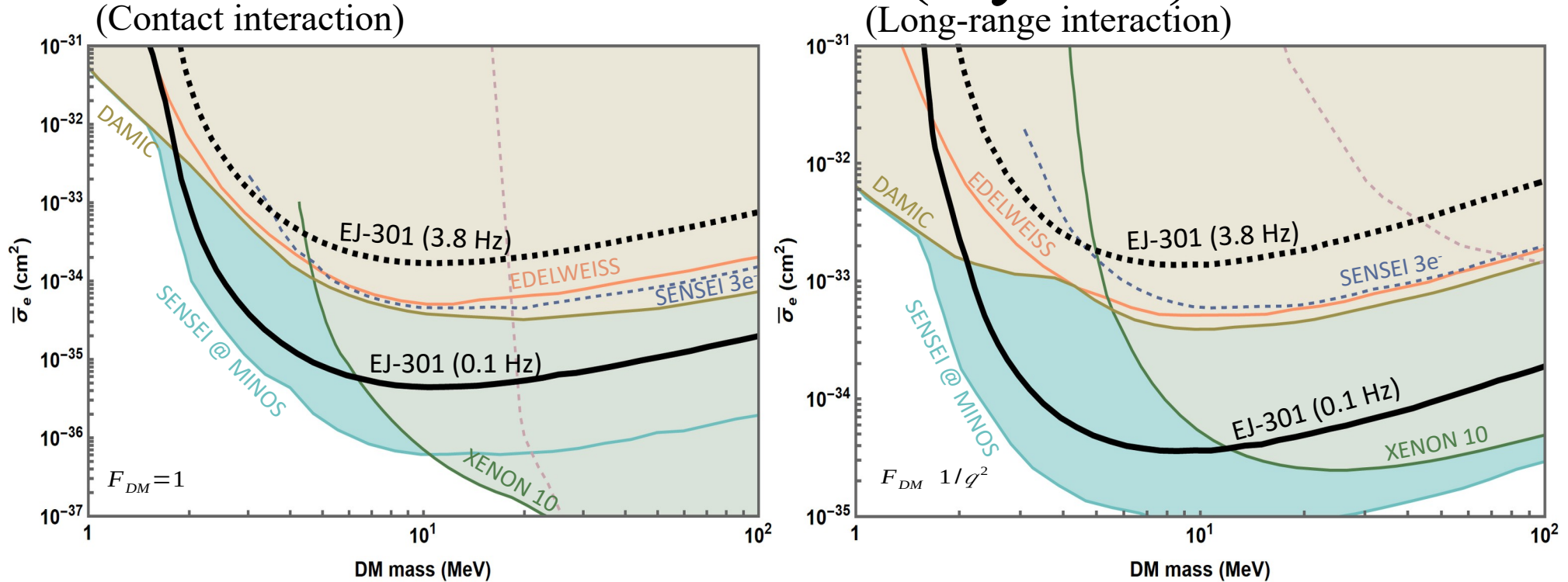


Para-xylene (EJ-301)

[CB, Collar, Kahn, Lillard: 1912.02822]



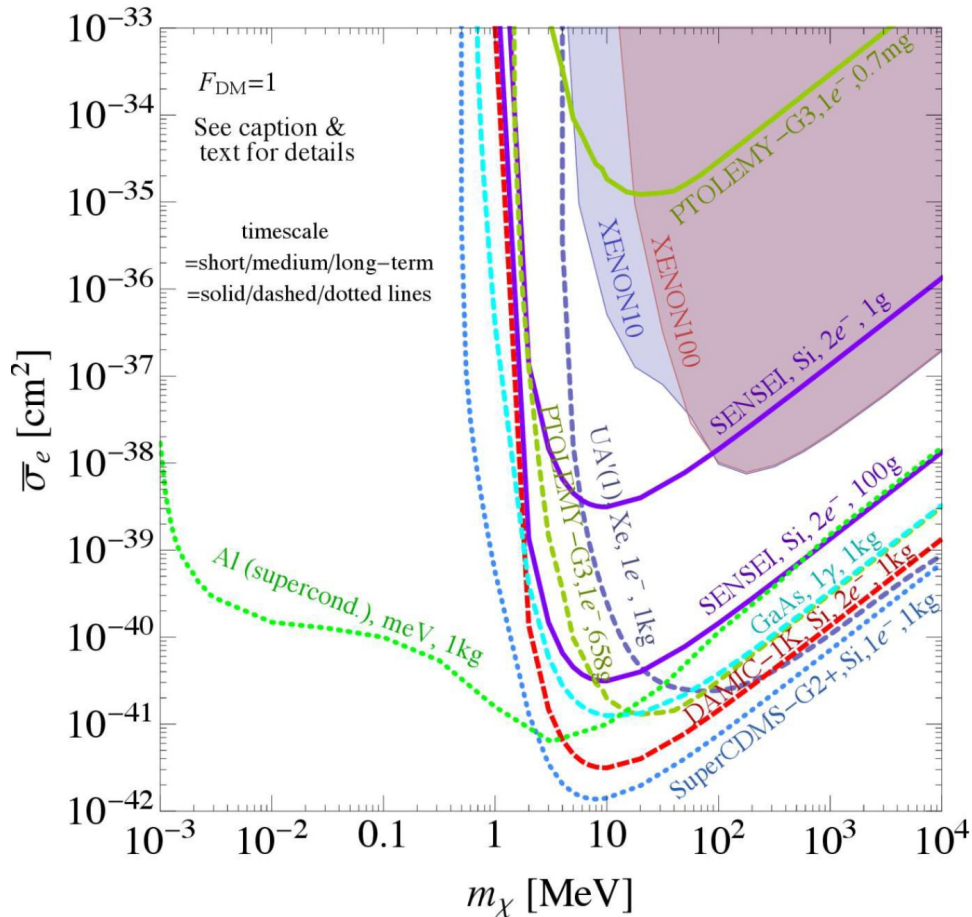
# Results: EJ-301 (Xylene)



[CB, Collar, Kahn, Lillard: 1912.02822]

About 6 months from theory development to results.

# The Field in Context



Many materials are proposed to probe the sub-GeV Space

In 2017:

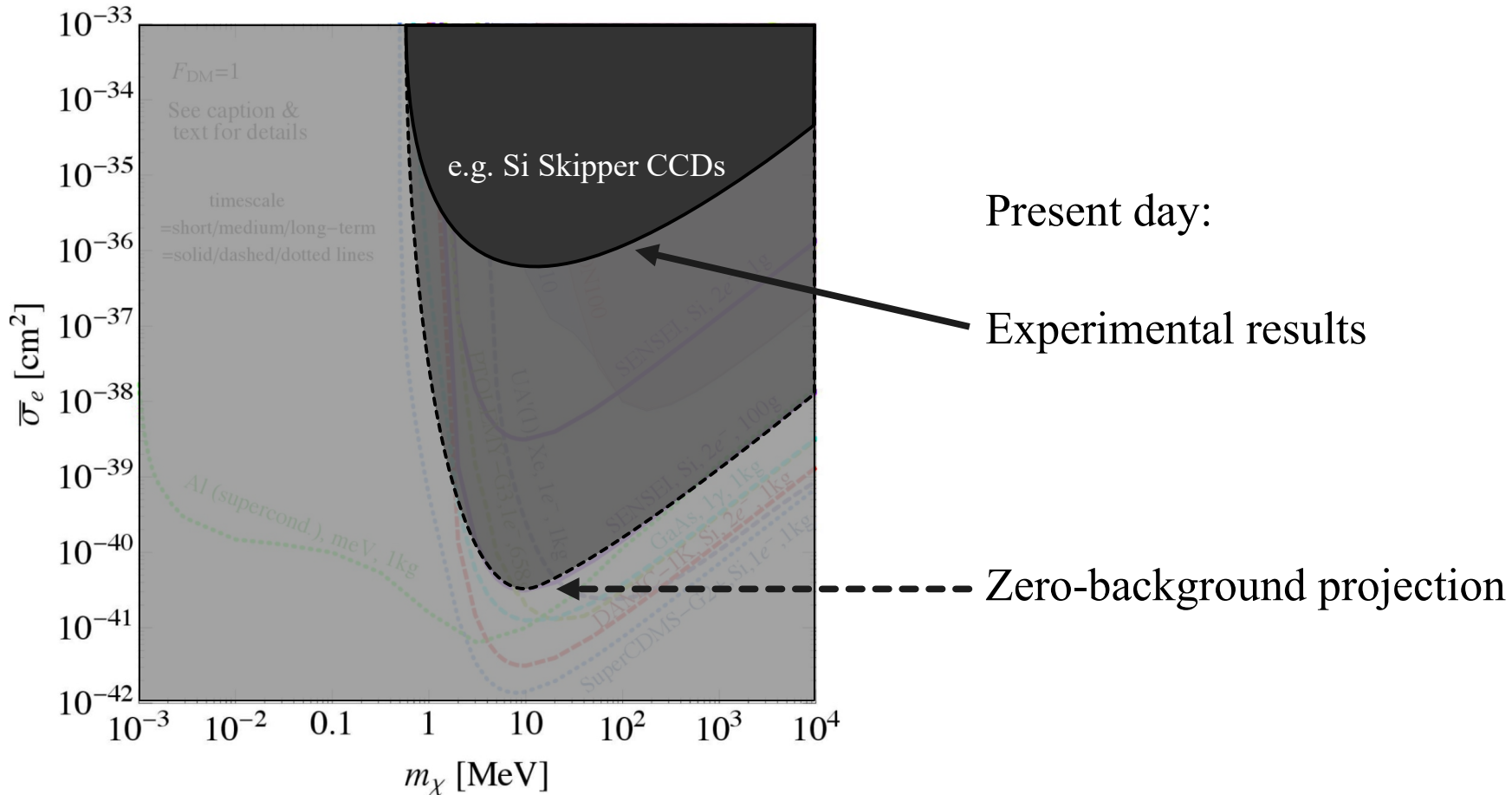
← *Short term (2 years)*

← - - - *Medium term (2-5 years)*

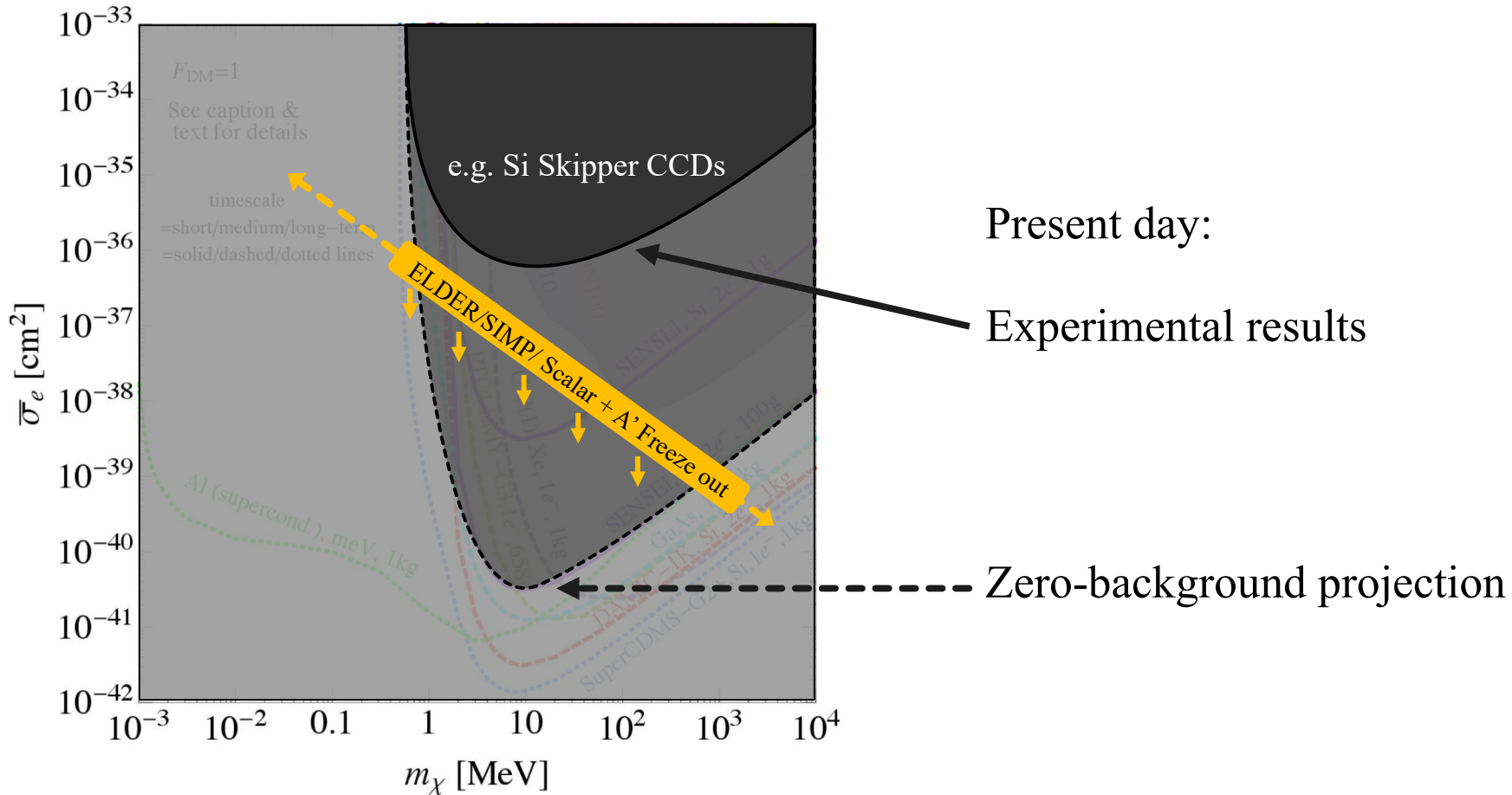
← . . . . *Long term*



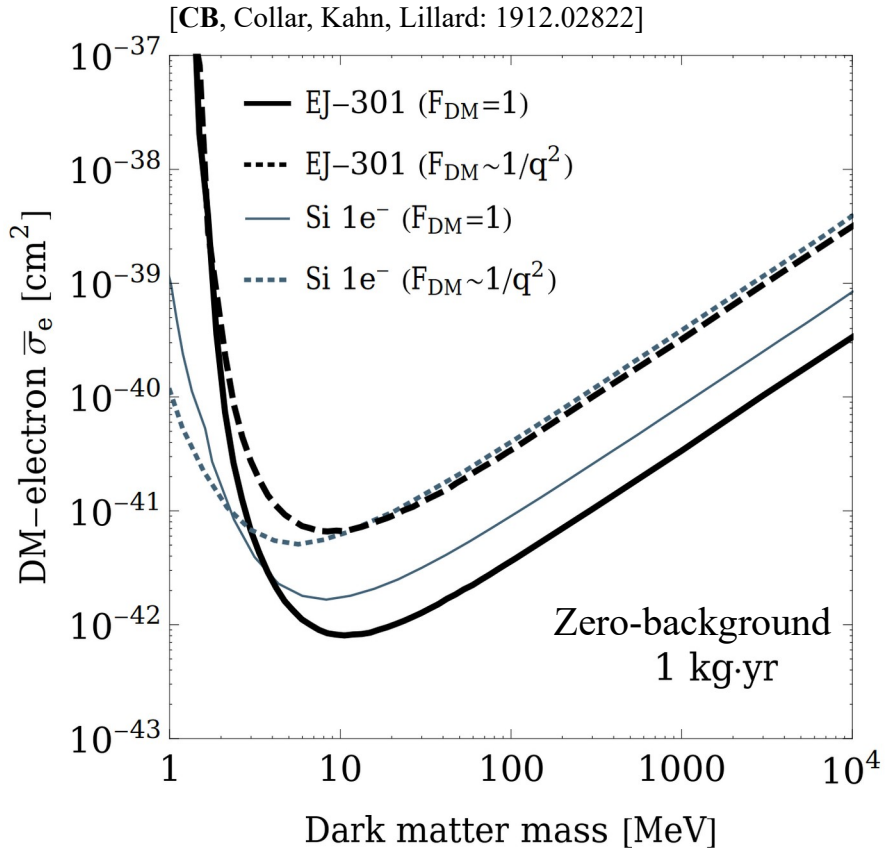
# Outlook and Potential Reach



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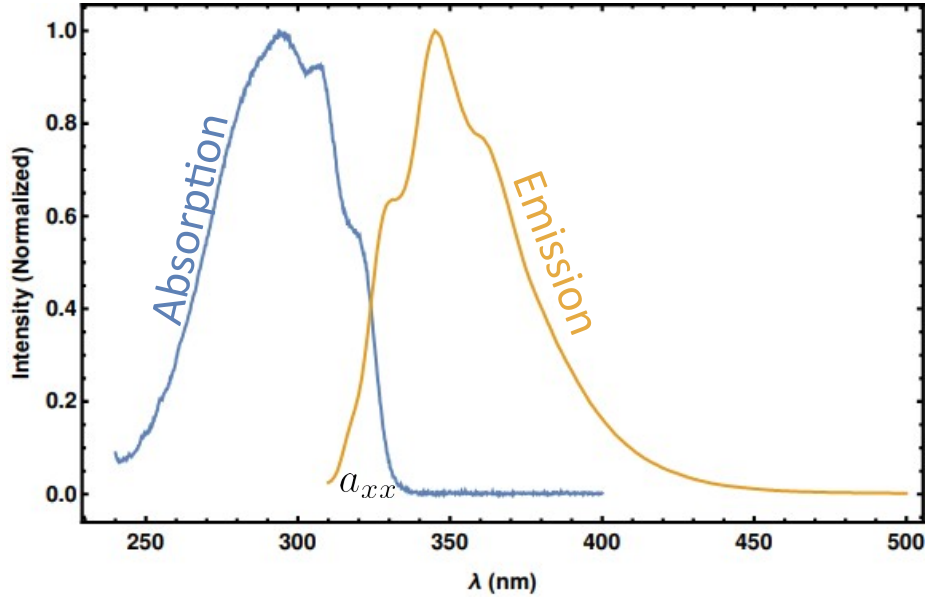
The obstacle

# Backgrounds

6 orders of magnitude of potential reach awaiting

Pound (kg) for Pound (kg) molecules produce about as much signal as e.g. Si.

# Fluorescence with DM



Next Step

## Option 1

Reduce background in the excitation.

*Molecular crystals*

Anisotropic excitation → Time-varying DM signal

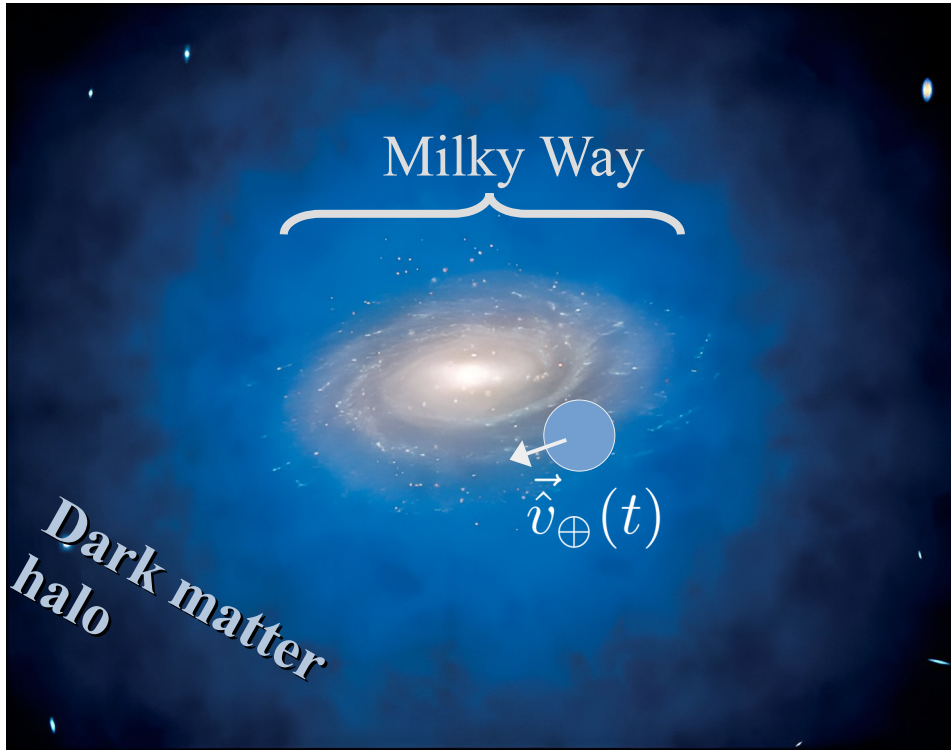
## Option 2

Reduce background in the emission.

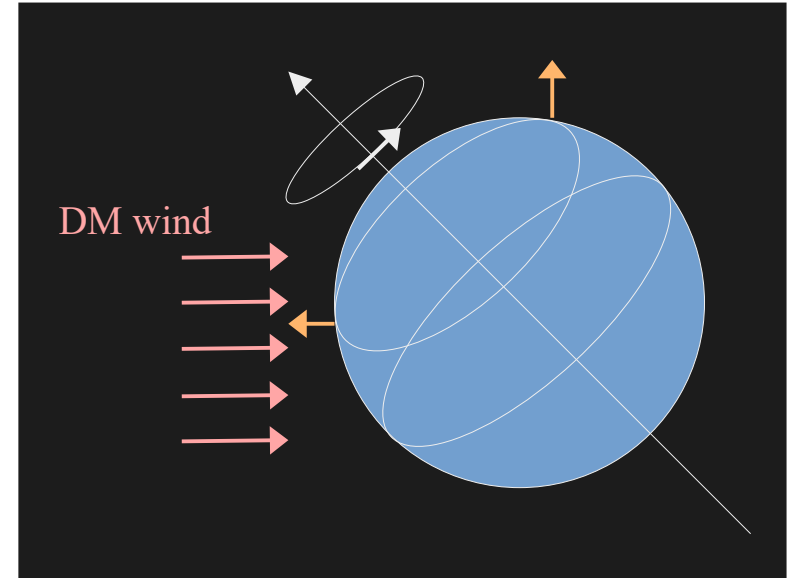
*Quantum dots*

Multiple excitons → Time-coincident DM signal

# Directional Detection

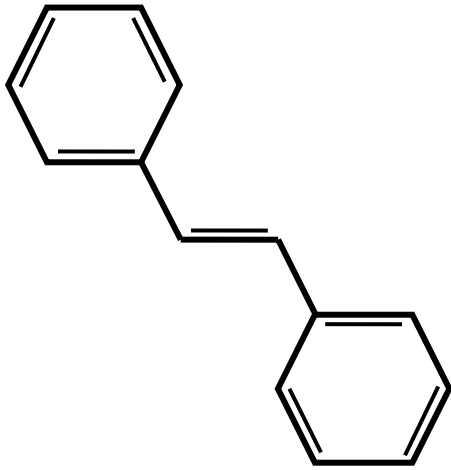


Effective dark matter “wind” from relative motion

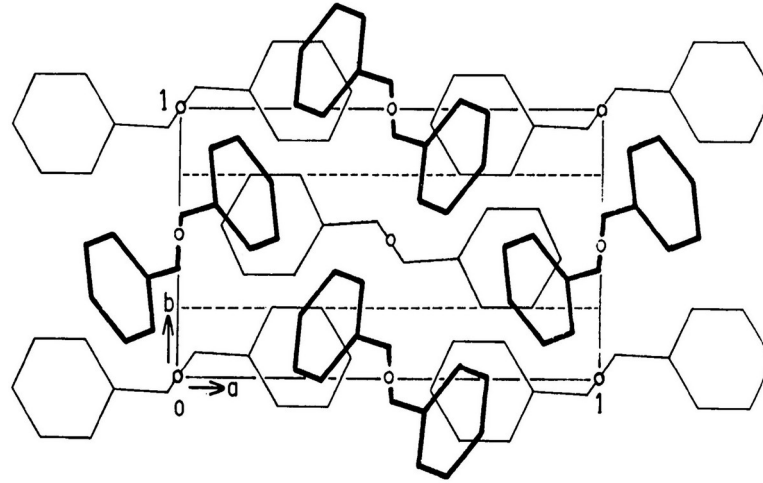


Change in relative orientation between detector and dark matter wind leads to *daily* modulation

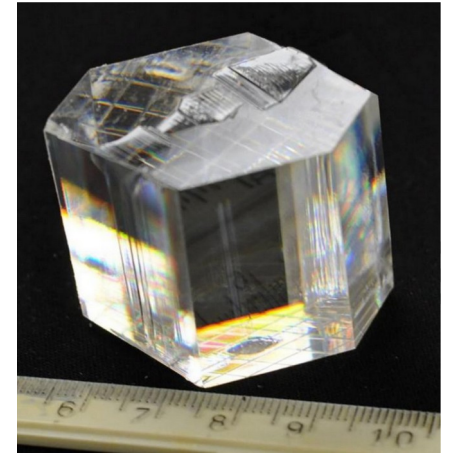
# Trans-Stilbene



Delocalized and planar network  
of double bonds



Molecular planes oriented in  
crystal lattice

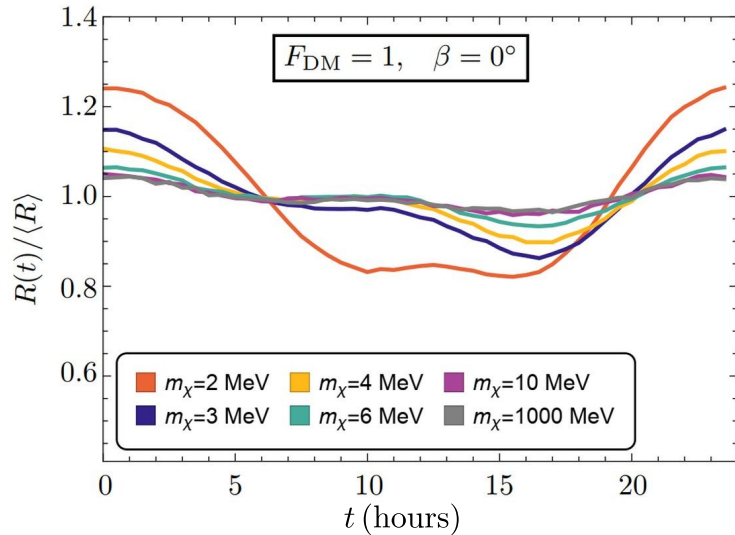


Carman, et.al. '18 (J. of Crystal Growth)

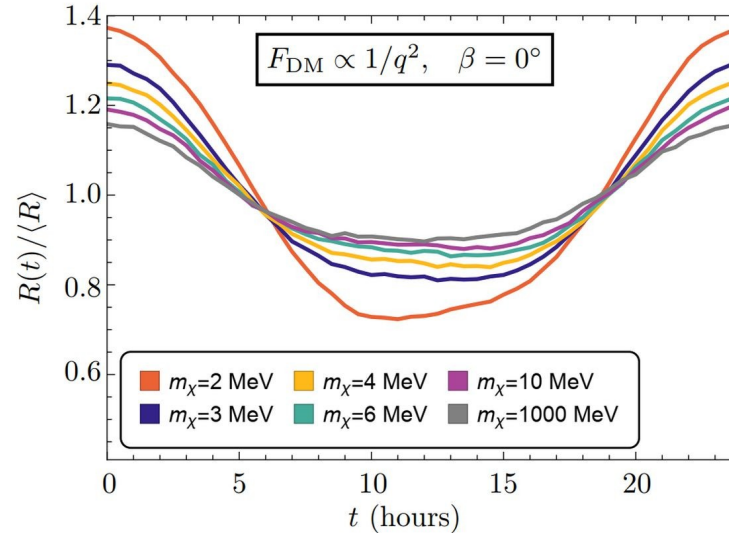
Large optical-quality crystals

# Daily Modulation

(Contact interaction)



(Long-range interaction)



} Varies up to 70%  
Verifiable signal!

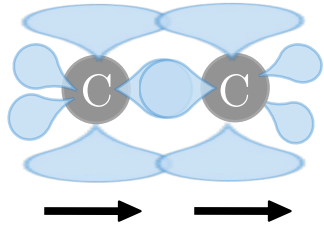
[CB, Kahn, Lillard, McDermott: 2103.08601]

Modulation amplitude remains as high as 10% even at the highest masses due to the fundamental anisotropy of the molecular form factor.

# The Molecular Migdal Effect(s)

Center of mass recoil (CMR)

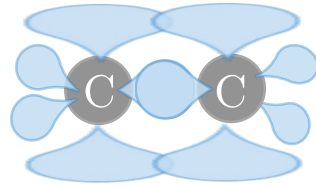
Cause by center of mass motion



Analogous to atomic Migdal effect

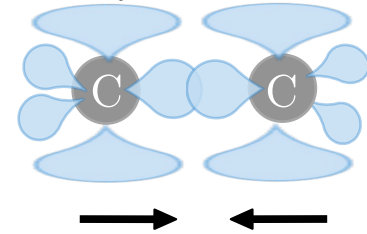
$$P_{CMR} \sim \frac{m_e}{M_{mol}}$$

Moving whole molecule  $\rightarrow$  BIG penalty



Non-adiabatic coupling (NAC)

Caused by relative motion



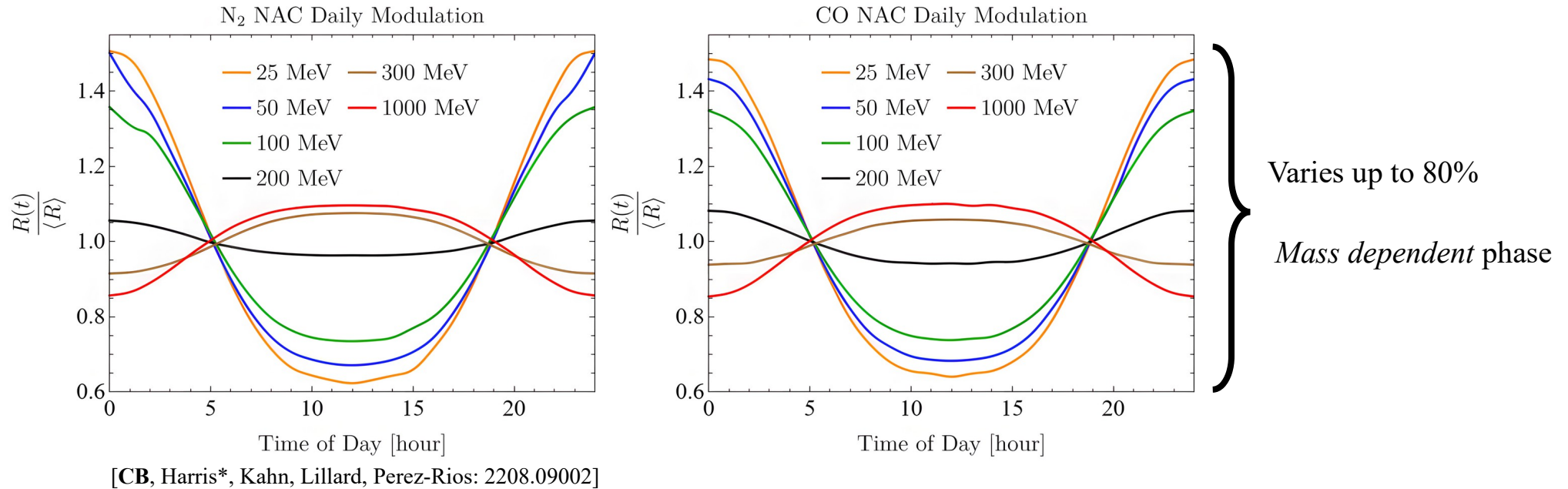
Effect beyond Born-Oppenheimer

$$P_{NAC} \sim \frac{m_e}{M_N}$$

Crumpling molecule  $\rightarrow$  small Penalty

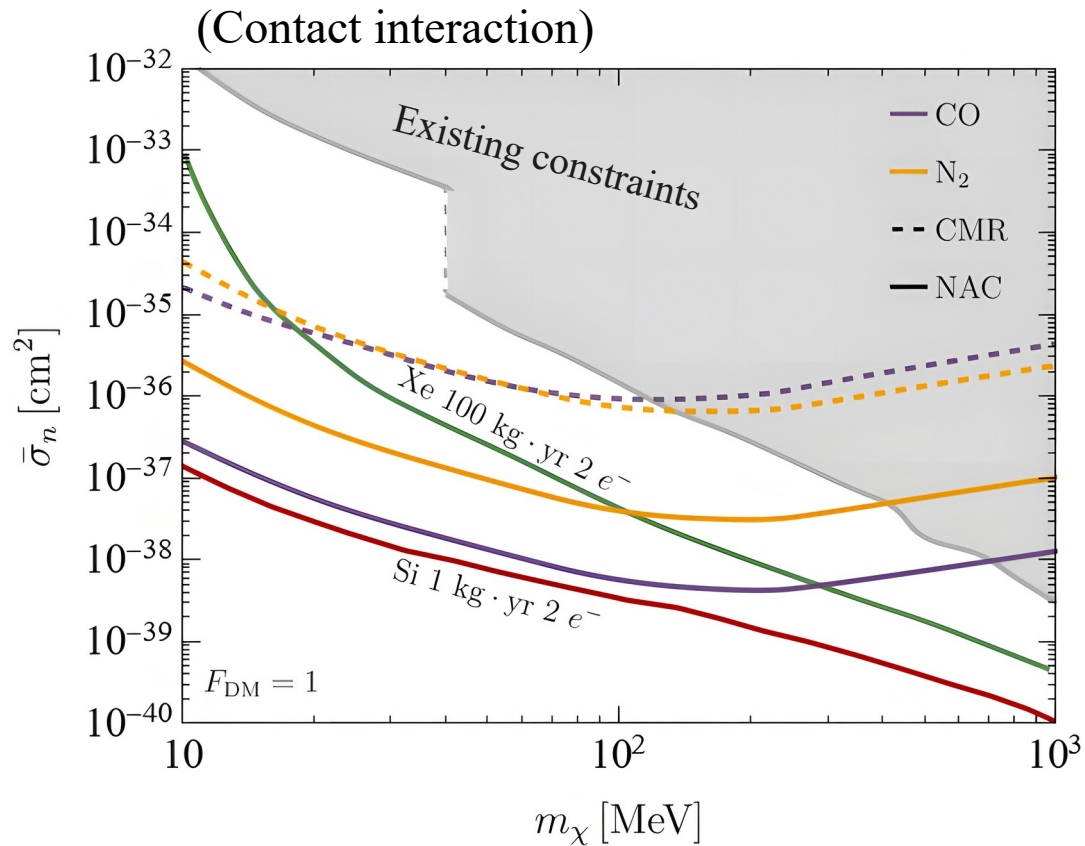


# Directional Molecular Migdal Effect



Molecular alignment → Directional electronic excitation → Directional molecular Migdal effect(s)

# Molecular Migdal Effect(s)



[CB, Harris\*, Kahn, Lillard, Perez-Rios: 2208.09002]

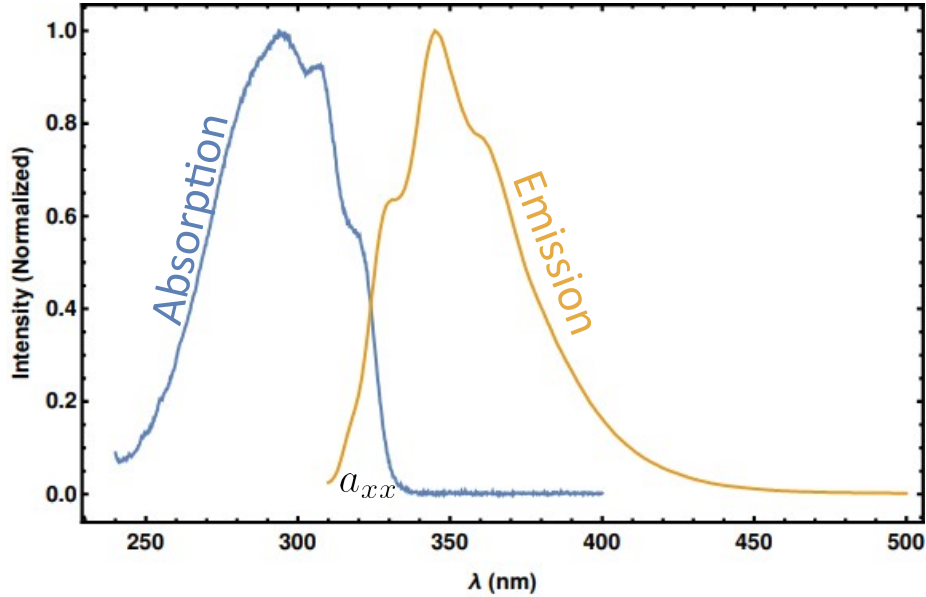
Si rate is calculated using the  
*CMR*-equivalent Migdal effect.  
 Is there an *NAC*-equivalent in Si?

← *enter of mass recoil*  
 Subdominant at all masses.

← *Non-adiabatic coupling*  
 Favorable kinematic factor.

Simplest molecular models already competitive.

# Fluorescence with DM



Next Step

## Option 1

Reduce background in the excitation.

### *Molecular crystals*

Anisotropic excitation  $\rightarrow$  Time-varying DM signal

## Option 2

Reduce background in the emission.

### *Quantum dots*

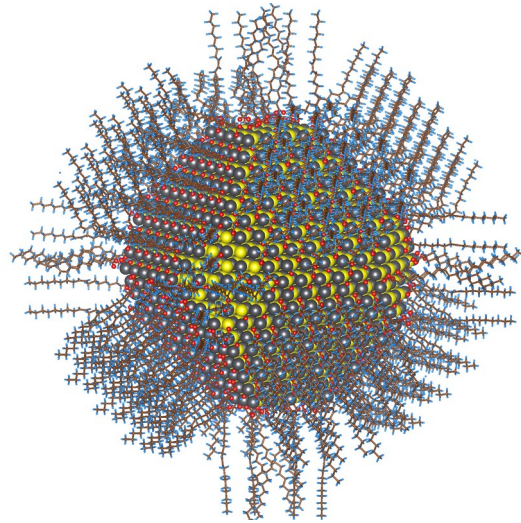
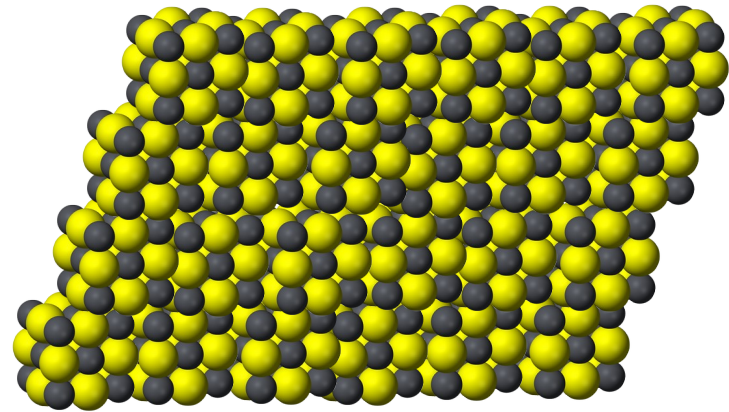
Multiple excitons  $\rightarrow$  Time-coincident DM signal

# Nanocrystals: Quantum Dots

Quantum confinement affects long-wavelength physics

Quantum confinement

Example: PbS



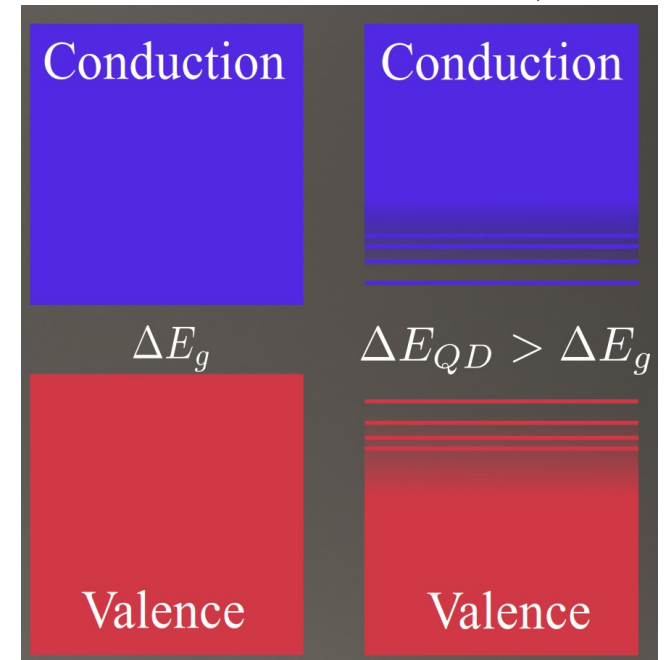
Zherebetsky et al., Science 344, 1380 (2014)

$$R \rightarrow \infty$$

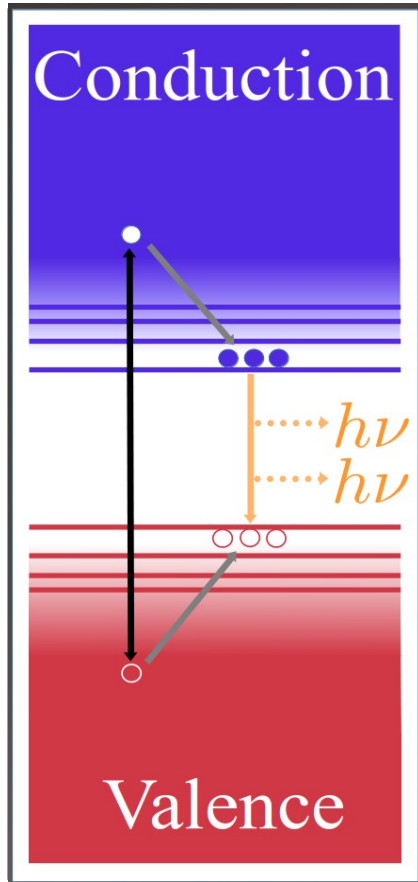
$$R \sim \mathcal{O}(\text{few nm})$$

$$|\psi_i\rangle \sim u_{\text{Bloch}}(r)e^{ik \cdot r}$$

$$|\psi_i\rangle \sim u_{\text{Bloch}}(r)\psi_{\text{bound}}(r)$$



# Quantum Dots: Coincident Signal



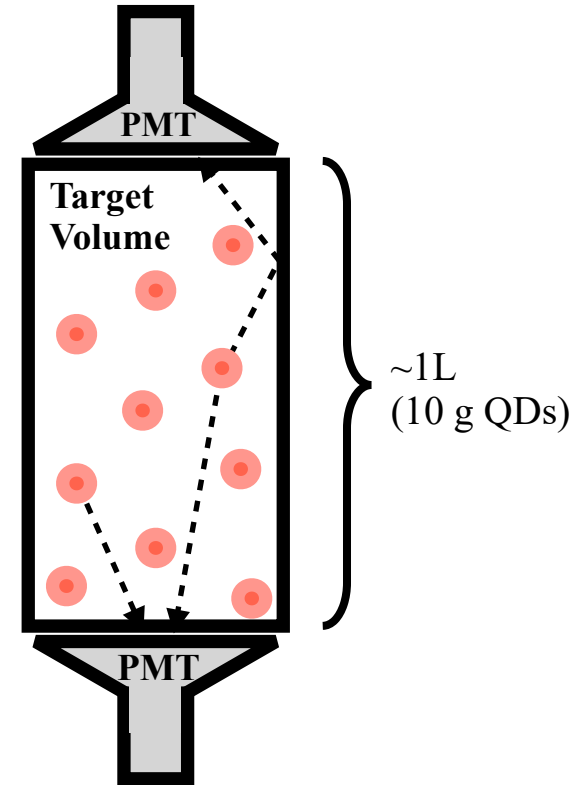
Absorption → *Very energetic* exciton

Multi-exciton generation → several excitons

If energy is greater than twice the band gap

Radiative recombination → coincident photons

Band-edge excitons produce light



# Deployment: DarkDot & QUADRA

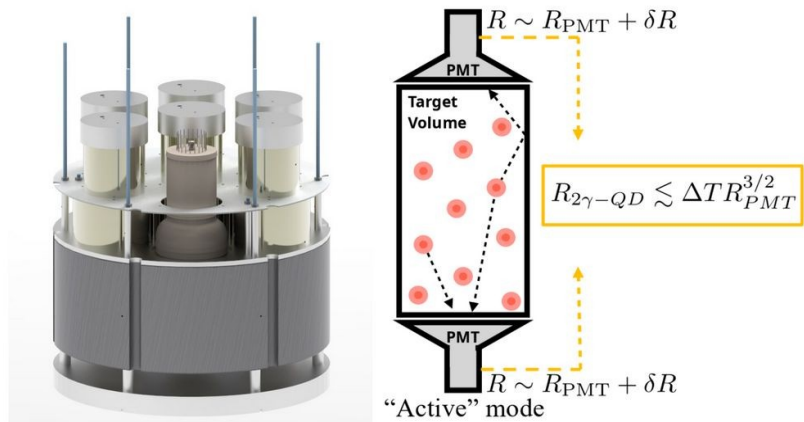
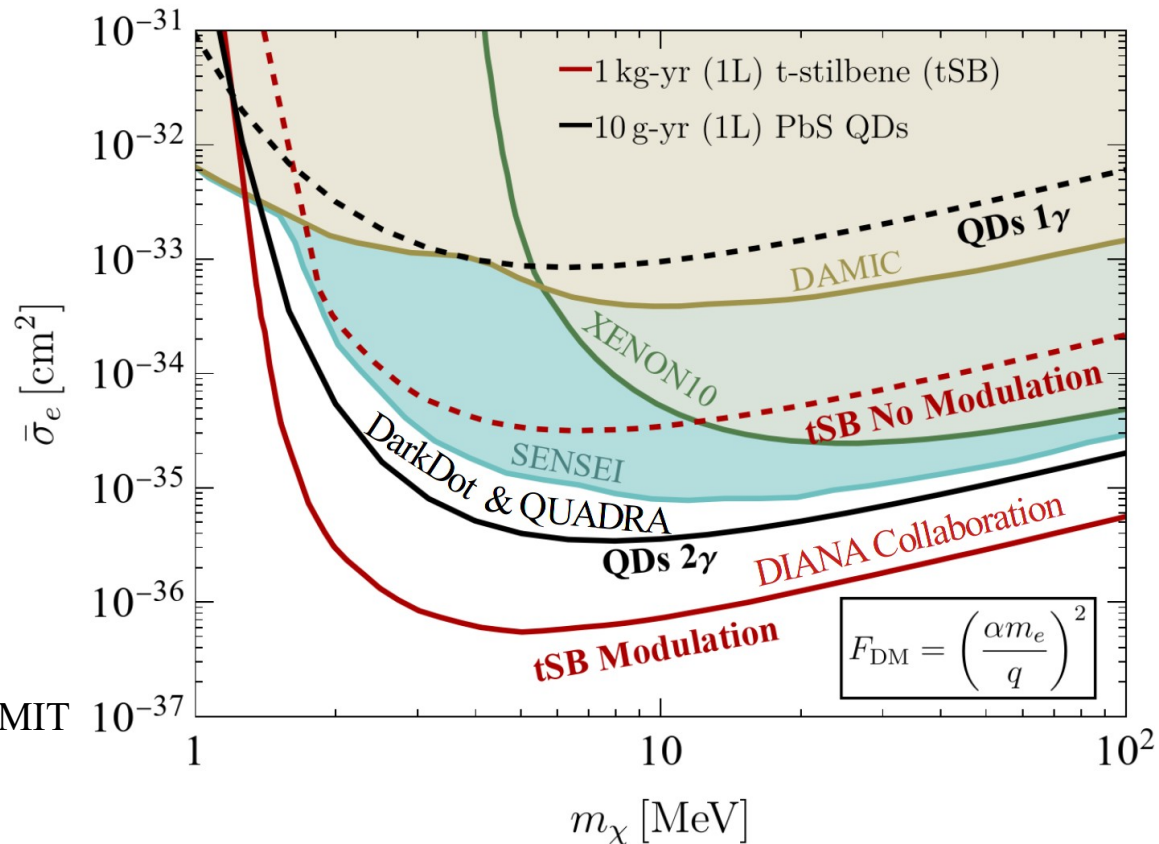


Figure 2: (Left:) Inner structure of the current SUXESs facility. (Right:) A diagram of a single module for the proposed detector.

## Experimental Collaborations

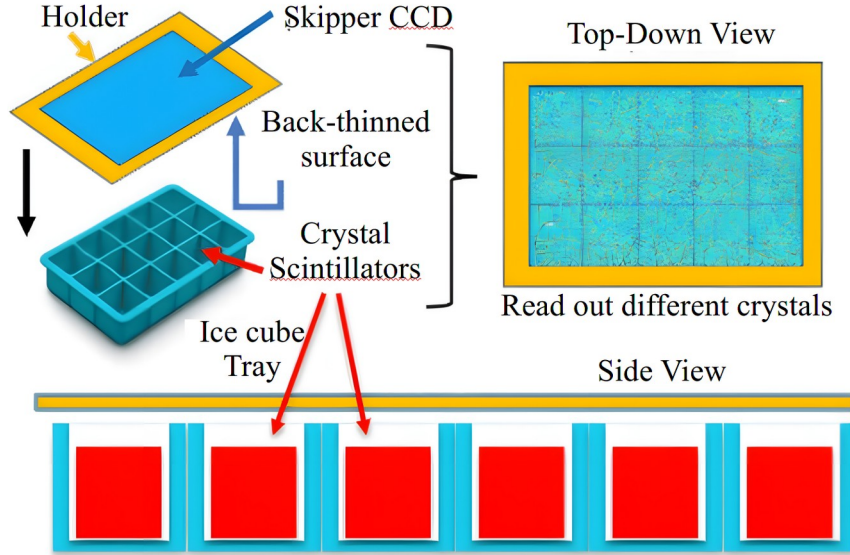
**DarkDot (SNSPDs):** MIT (Host), Stockholm U.

**QUADRA (Abalone PMTs):** Stockholm U. (Host), MIT

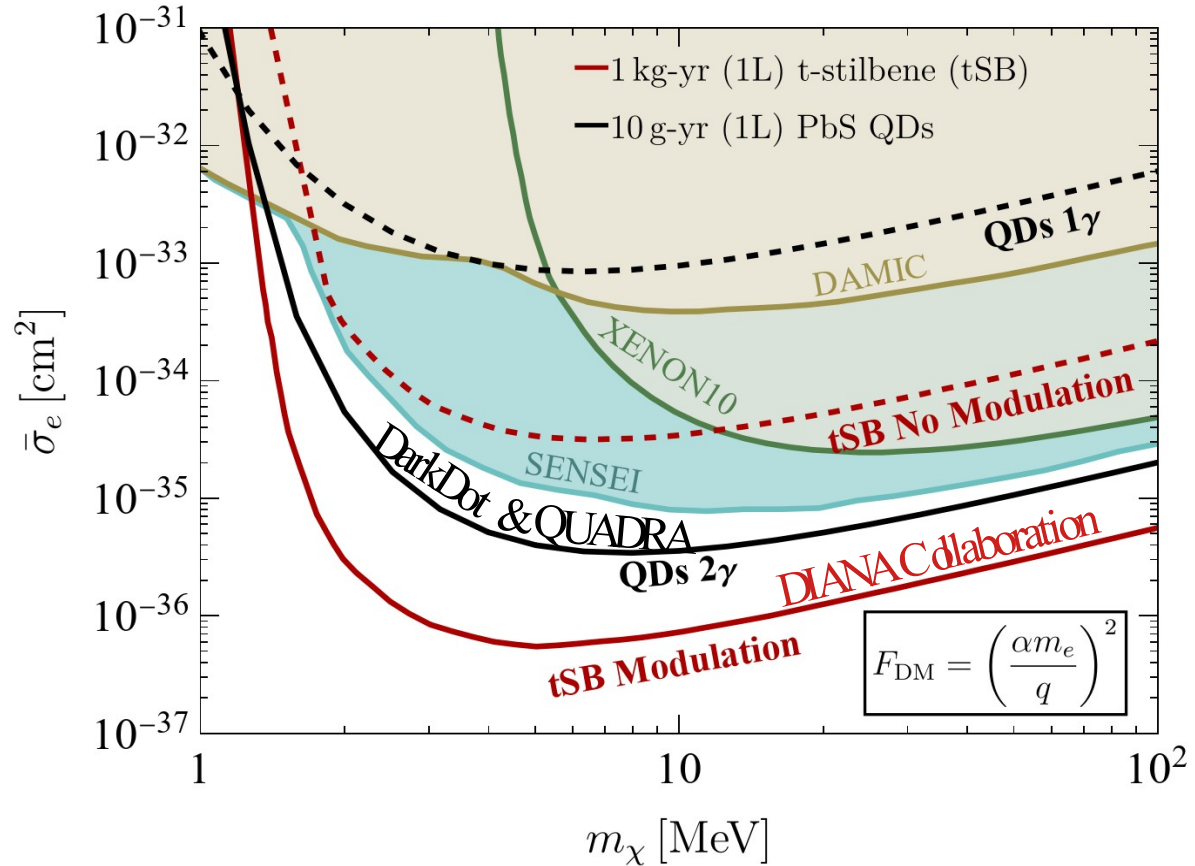


# Deployment: DIANA\*

## Daily Modulation in an Intrinsically ANisotropic Array



Collaboration: FermiLab, MIT, UIUC, U. Oregon



# Deployment: t-stilbene

Detecting fluorescence with Si

Skipper CCD Measurements of trans-Stilbene



Daniel Baxter, Alex Drlica-Wagner,  
Edgar Marrufo, Brandon Roach



# Deployment: t-stilbene

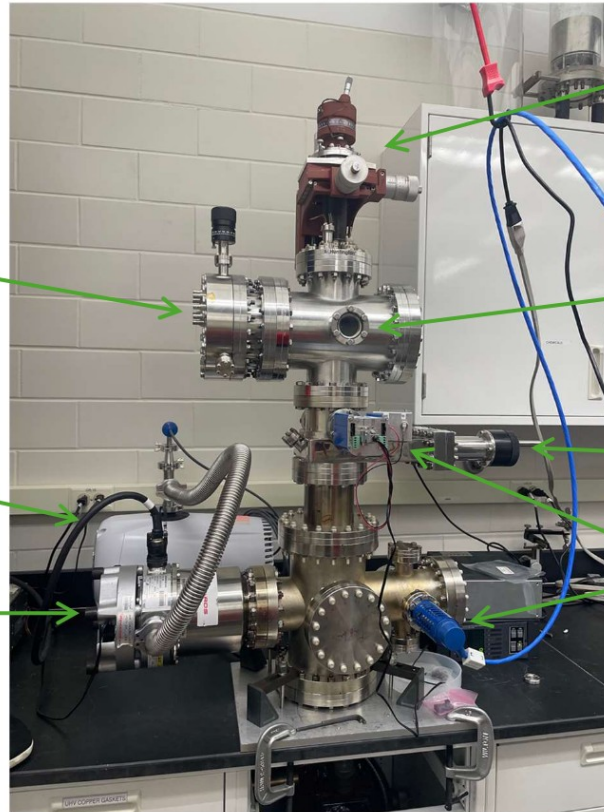
## Calibration of anisotropy

Angle-dependent stilbene fluorescence gizmo (ADSFG)

Electron gun (10-250 eV)  
+ phosphor screen for LEED

Scroll pump

Turbo, 500 l/s for  
“overpumping”



Sample manipulator  
(x,y,z, $\theta$ ) for rotating  
stilbene crystal

Viewport for  
PMT (not shown  
but working)

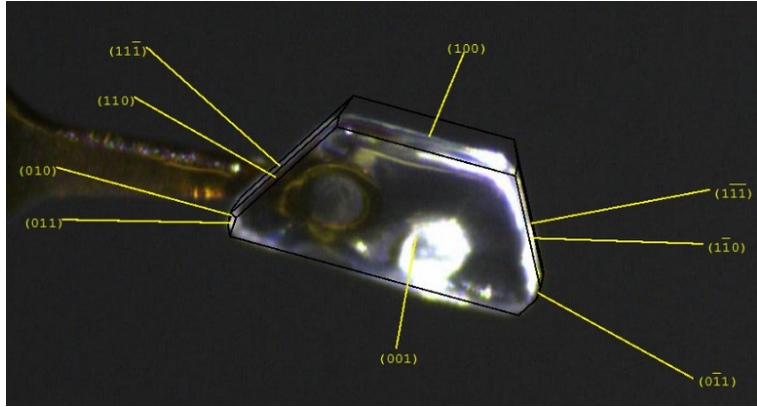
Gate valve

Ion Gauges

Abbamonte Group  
(UIUC)

# Deployment: t-stilbene

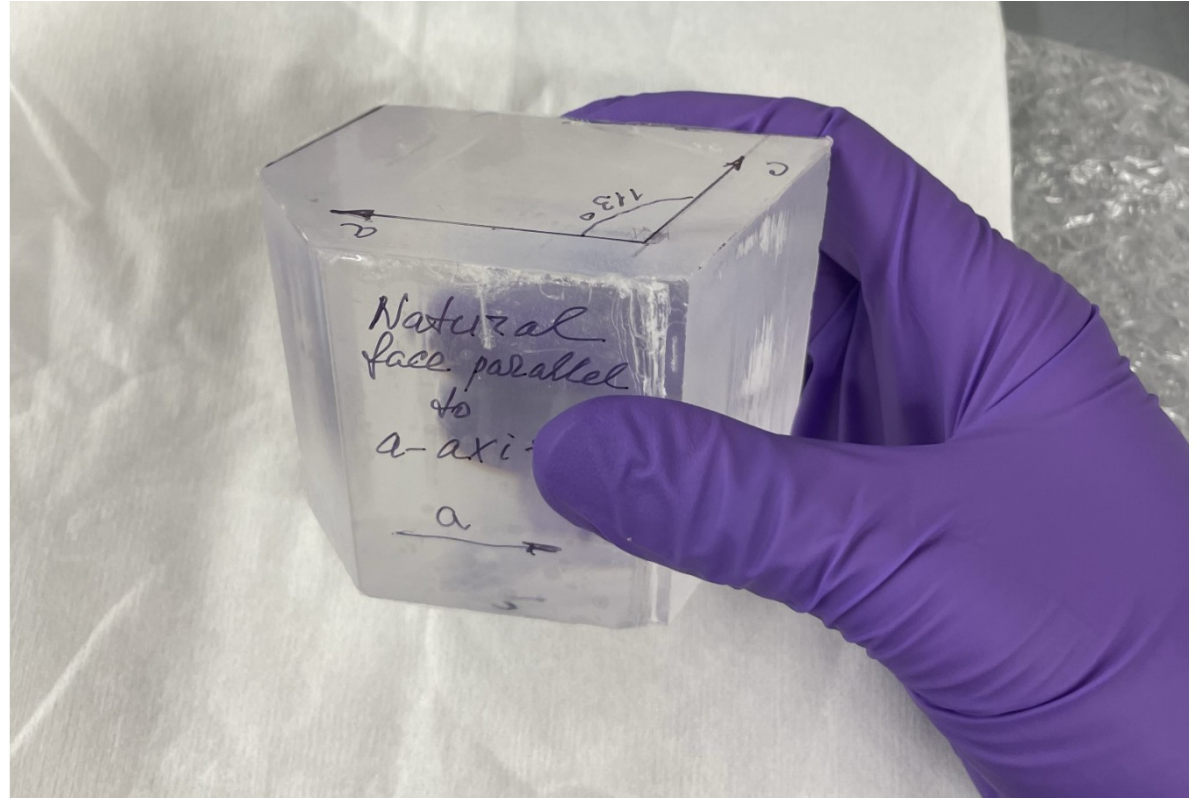
Calibrate at mg-scale



Credit: Dane Johnson  
Freedman Group (MIT)

Natalia Zaitseva (LLNL)

Scale to kg-scale



# Finding Optimal Targets

Problem: Chemical space is unreasonably large

How many molecules possible with  
C, O, N, F, H?

< 9 atoms: 100s of Thousands (DFT Computable)

< 30 atoms: 100s of Billions (Intractable)

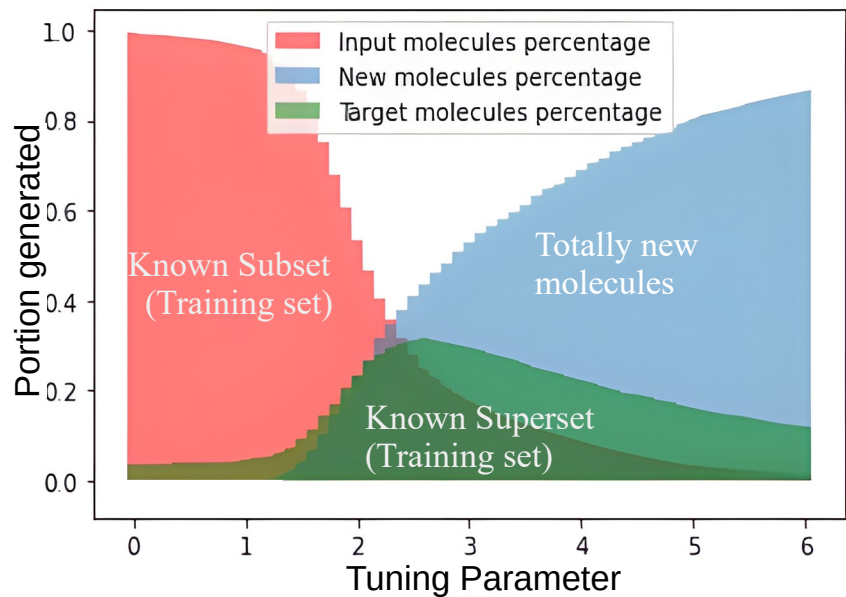
...toluene has 15, xylene has 18, t-stilbene has 26

## Method

1. Look for known favorable properties - *cheminformatics*
2. Extra(intra)polate ont

# ML for DM Direct Detection

Property prediction   Molecular Generation



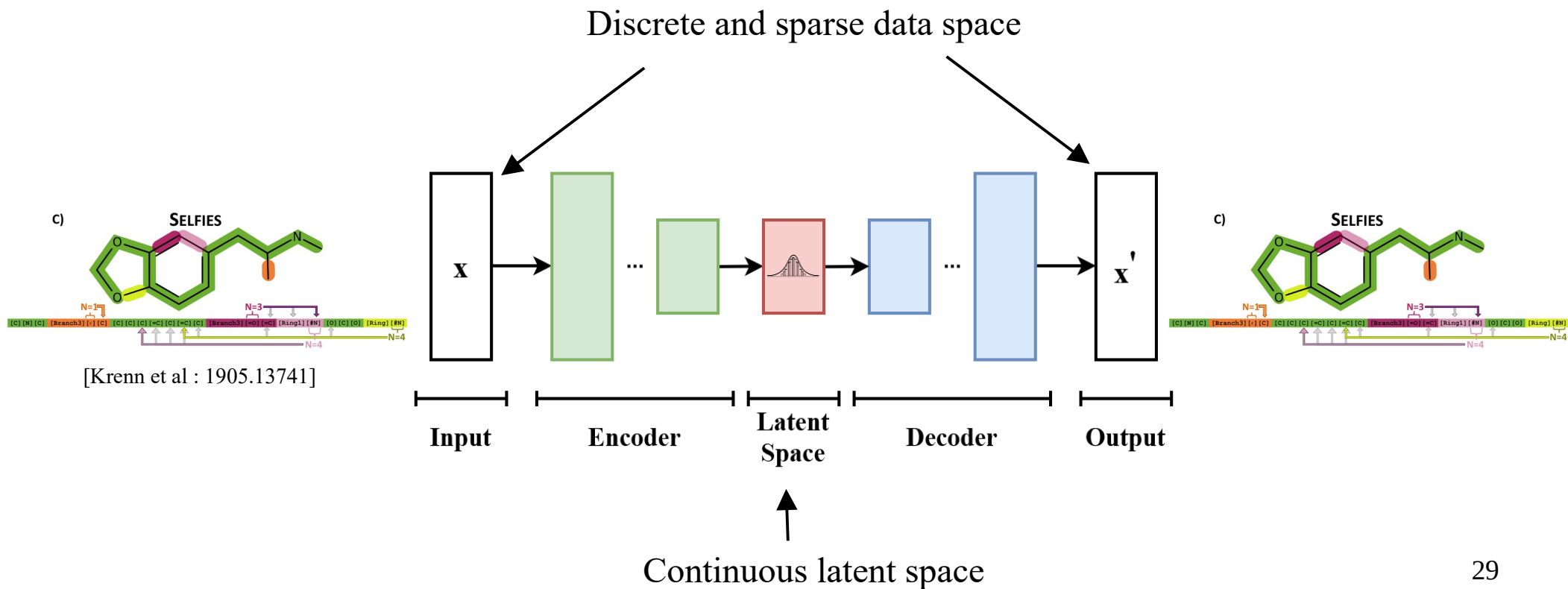
[CB, Cook\*, Smirnov: 2411.xxxxx]

Using exhaustive database ( $< 9$  atoms)  
Characterize neural nets  
→ Possible to learn from small subsample

Next: Large but sparse dataset up to 10s of atoms

# ML for DM Direct Detection

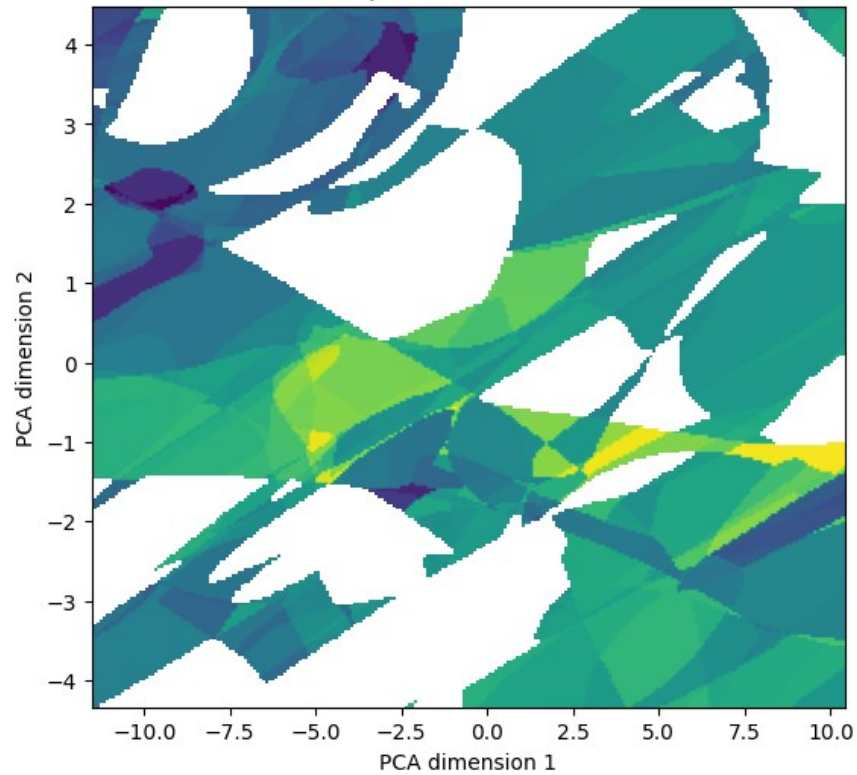
Molecular Space: Variational Autoencoders



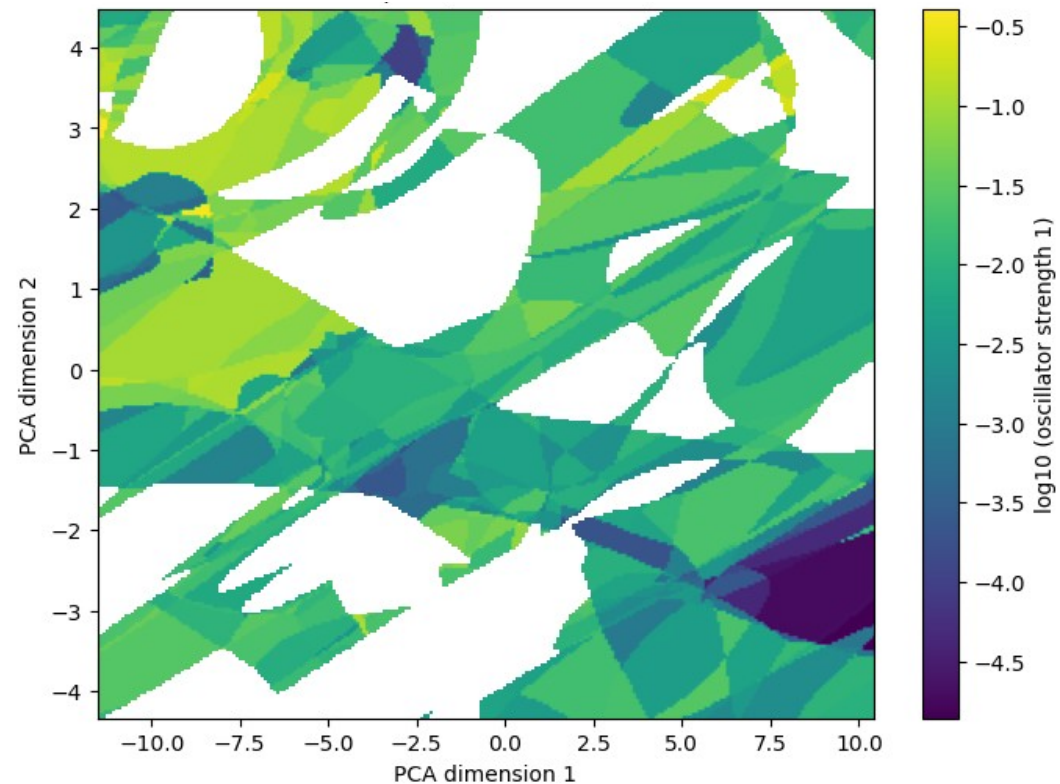
# ML for DM Direct Detection

Property prediction: Molecular latent space

$\Delta E$



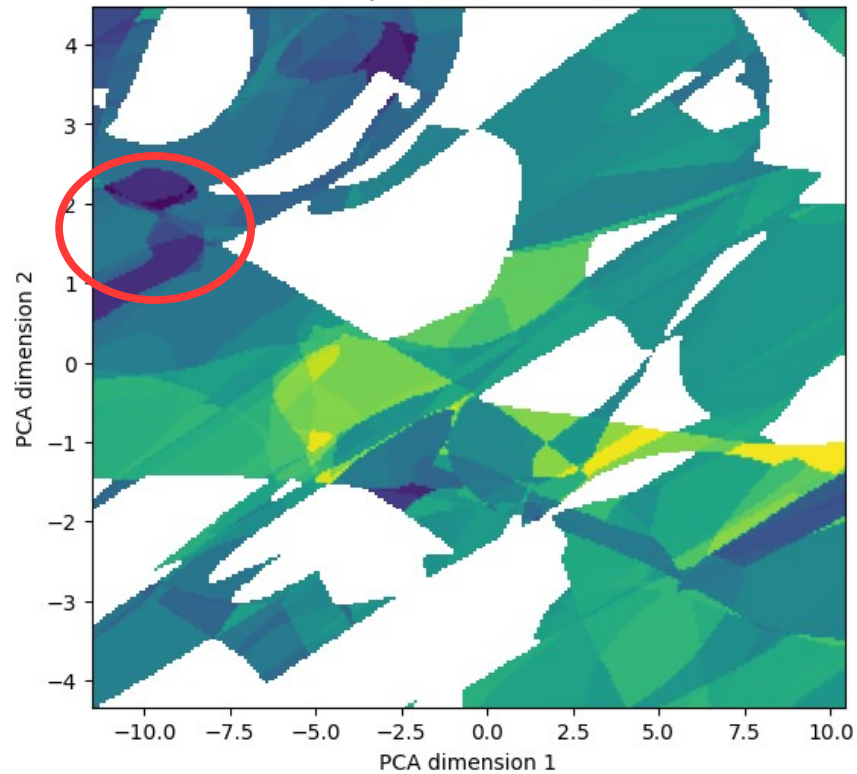
$\langle r_{01} \rangle \approx |f_{0,1}|/q$



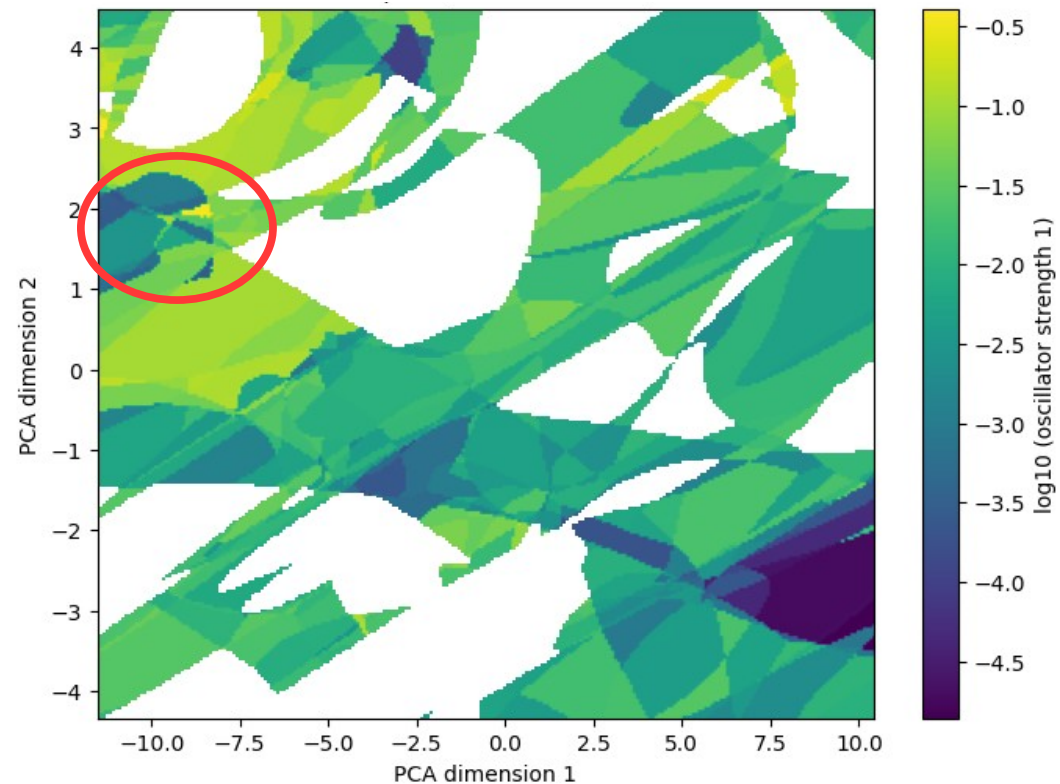
# ML for DM Direct Detection

Property prediction: Molecular latent space

$\Delta E$



$\langle r_{01} \rangle \approx |f_{0,1}|/q$



# ML Proposed Molecules

Property prediction   Molecular Generation



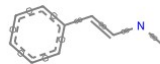
Clst 3 (43 mols)



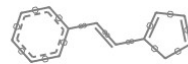
Clst 4 (35 mols)



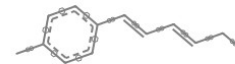
Clst 5 (32 mols)



Clst 6 (28 mols)



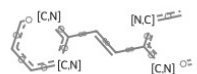
Clst 7 (28 mols)



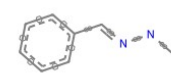
Clst 8 (26 mols)



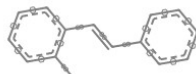
Clst 9 (25 mols)



Clst 13 (21 mols)



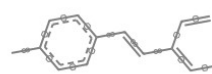
Clst 14 (20 mols)



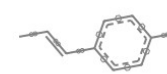
Clst 15 (18 mols)



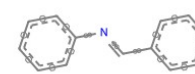
Clst 16 (18 mols)



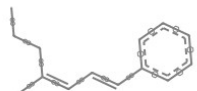
Clst 17 (17 mols)



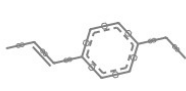
Clst 18 (16 mols)



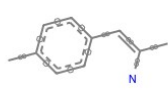
Clst 19 (16 mols)



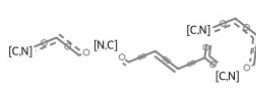
Clst 23 (14 mols)



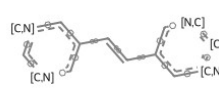
Clst 24 (13 mols)



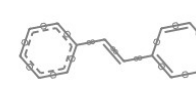
Clst 25 (13 mols)



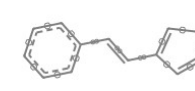
Clst 26 (13 mols)



Clst 27 (13 mols)



Clst 28 (13 mols)



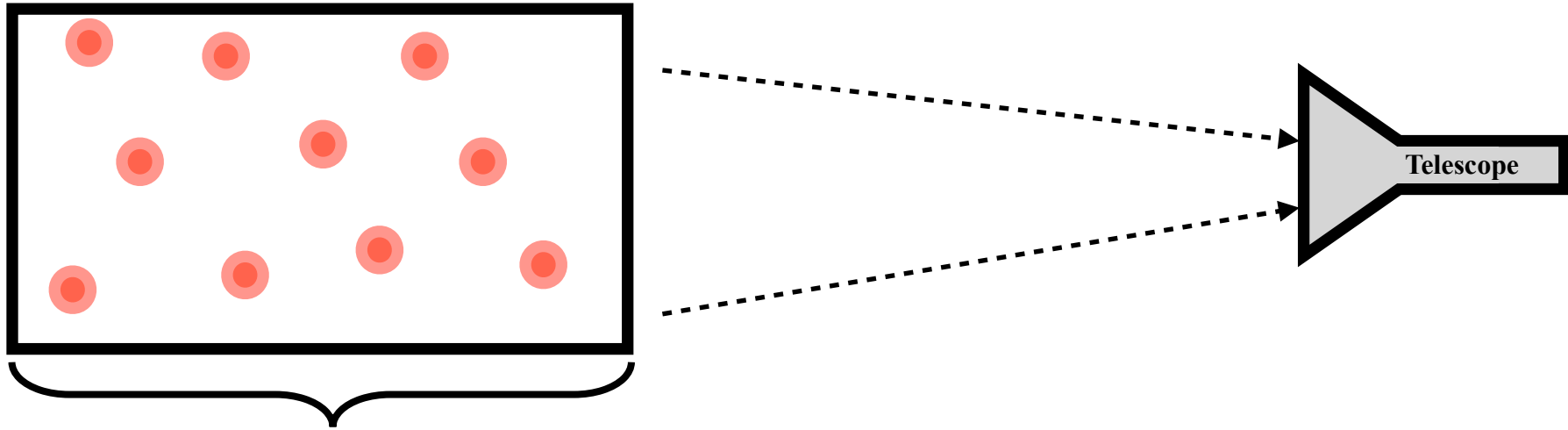
Clst 29 (13 mols)

## Molecular generation

- Generate candidate molecule *shortlist*:  $\text{Min}(\Delta E \sim 2\text{eV})$ ,  $\text{Max}(\langle r^2 \rangle)$ , filter for synthetic feasibility.
- Cluster by molecular similarity
- Extract largest common substructure



# Beyond direct detection



Astrophysical volume of molecules

Same theoretical techniques → Predict rates in *astrophysical* objects

# Dark Matter can be captured inside celestial objects.

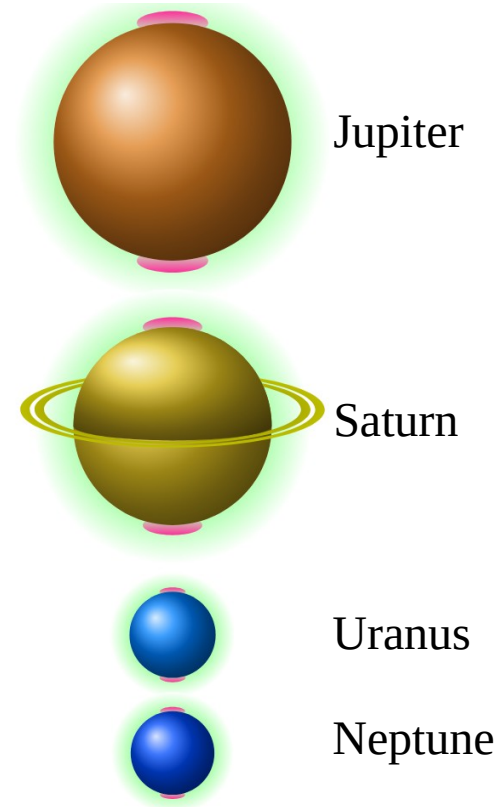
Assuming equilibrium is reached:

$$\Gamma_{\text{ann}} = f_{\text{cap}} \times \pi R^2 \rho_{\chi} v_{\chi} \sqrt{\frac{8}{3\pi}} \left( 1 + \frac{3}{2} \frac{v_{\text{esc}}^2}{v_{\chi}^2} \right)$$

Geometric scaling

Enhancement due to gravity

$$P_{\text{ion}}^{\text{DM}} = \frac{\Gamma_{\text{ann}} \times f_{\text{iono}}}{4\pi R^2}$$



# Equilibrium Radial Distribution

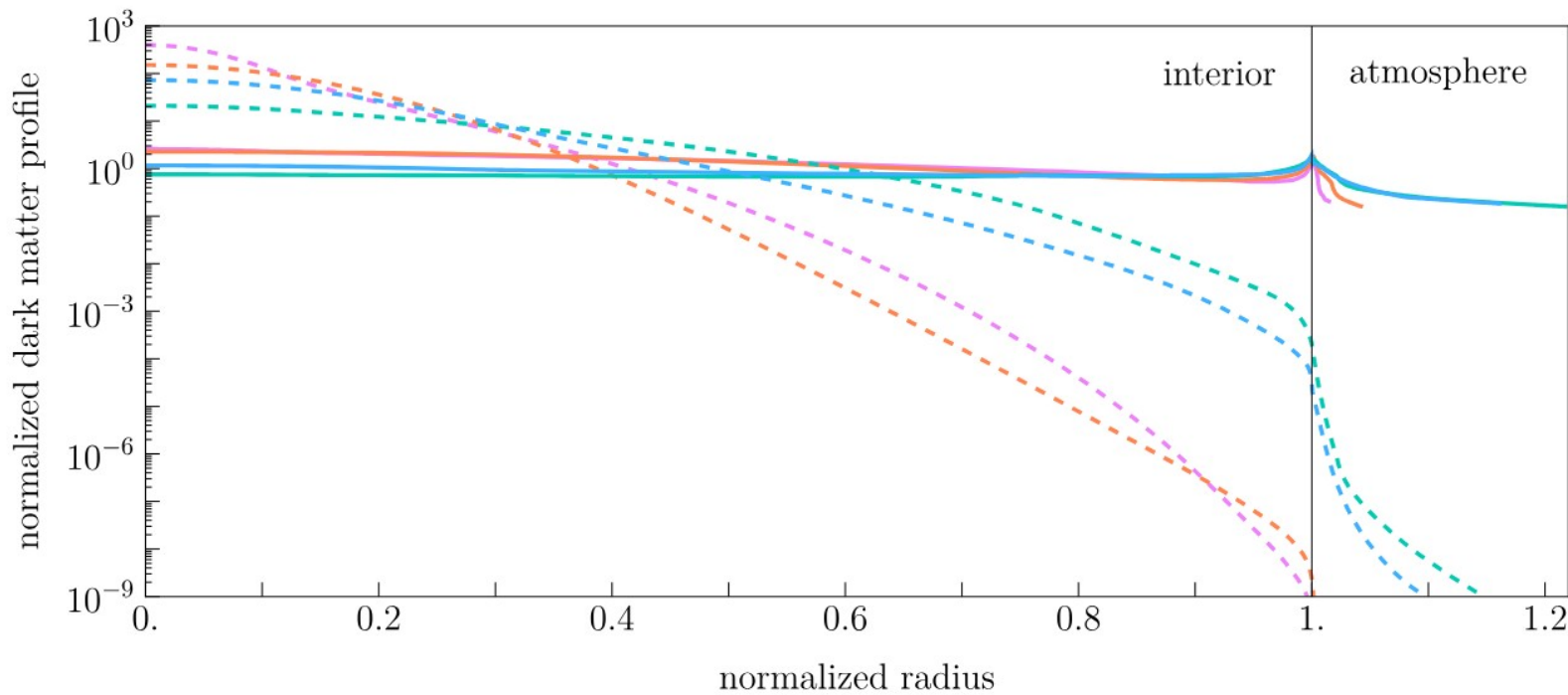
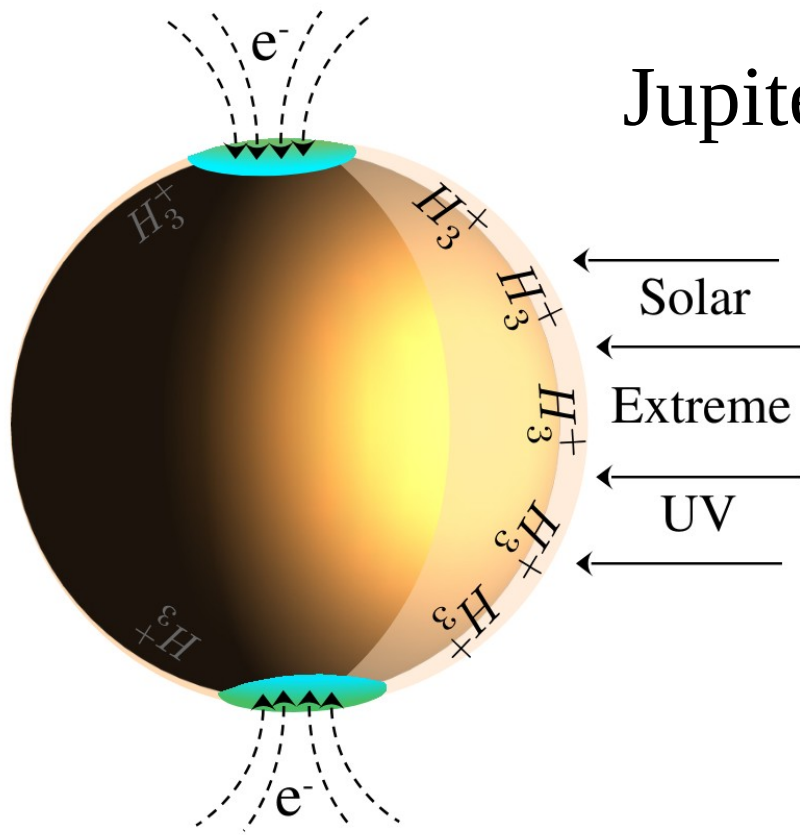
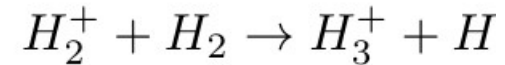
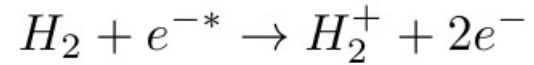


Figure 3. Dark matter radial profile in the Solar System gas and ice giants as labelled, for a 0.1 GeV (full) and a 1 GeV dark matter particle. The distance to the planet cores are normalized by each planetary radius; values above 1 indicate the atmosphere.

# Jupiter: the IR signal



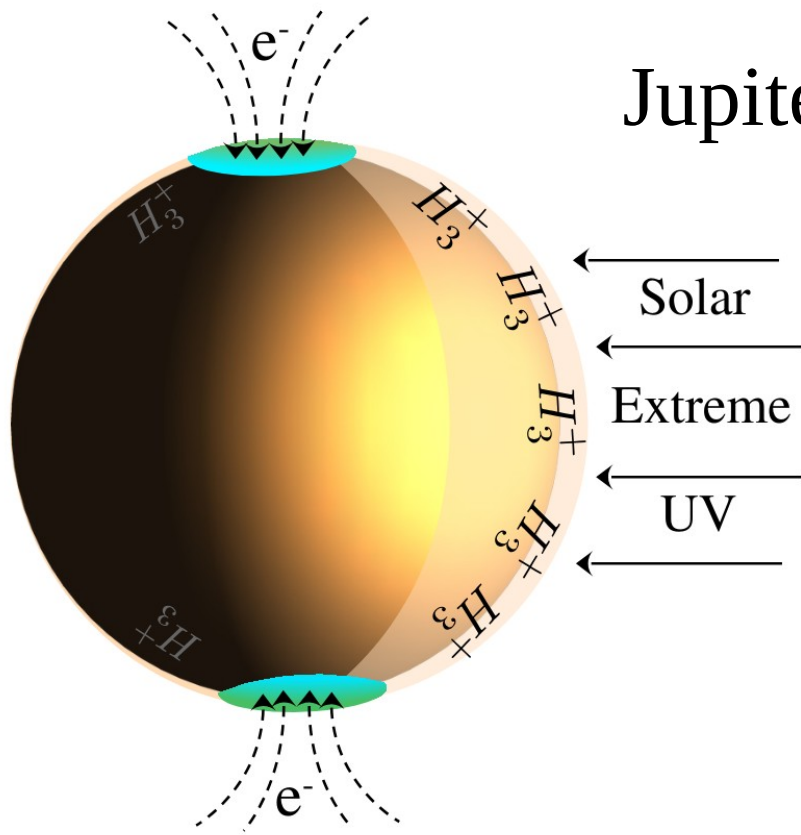
Ionizing radiation generates  $H_3^+$



$H_3^+$  radiates extremely efficiently in the IR. It is Jupiter's *thermostat*.

Figure 1. Schematic of  $H_3^+$  production in Jupiter. Auroral  $H_3^+$  emission near the magnetic poles is sourced by precipitating electrons, and solar extreme UV irradiates the day side and dominates  $H_3^+$  production near the equator. No  $H_3^+$  is expected from known processes at low latitudes on the night side, making it an ideal dark matter signal region.

# Jupiter: the IR signal



$\text{H}_3^+$  IR emission intensity is linear in ionizing power

$$I^{\text{H}_3^+} = \alpha \times \beta \times P_{\text{ion}} E_{\text{mol}}(T)$$

Efficiency and geometric factors

Molar emissivity

Night-side Intensity

$$I^{\text{H}_3^+} < 0.03 \text{ mW/m}^2 / \mu\text{m/sr}$$

Day-side Intensity

$$I_{\text{day}}^{\text{H}_3^+} = 0.09 \text{ mW/m}^2 / \mu\text{m/sr}$$

Day-side Ionizing Power

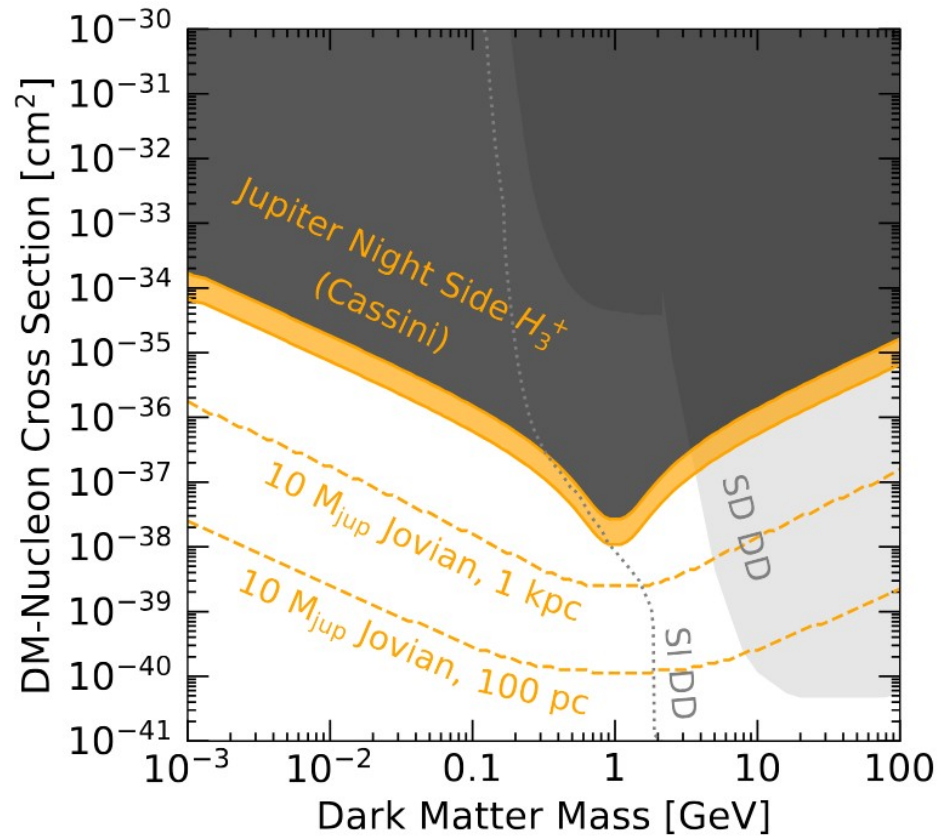
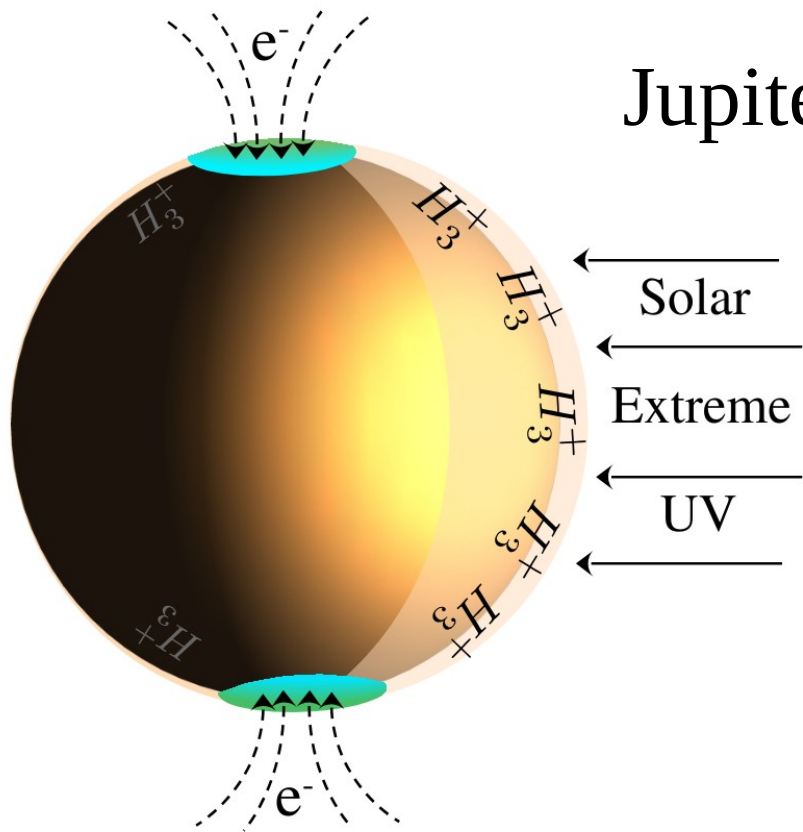
$$P_{\text{ion}}^{\text{EUV}} = 62 \mu\text{W/m}^2$$

Night-side Ionizing Power

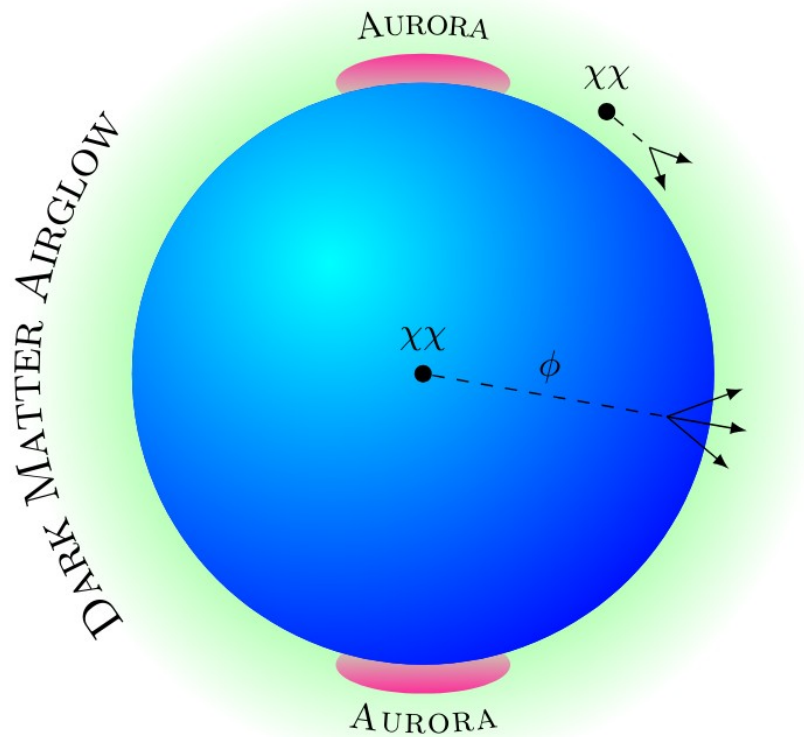
$$P_{\text{ion}}^{\text{night}} < \frac{I_{\text{max}}^{\text{H}_3^+}}{I_{\text{day}}^{\text{H}_3^+}} \times \frac{E_{\text{mol}}(T_{\text{day}})}{E_{\text{mol}}(T_{\text{night}})} \times P_{\text{ion}}^{\text{EUV}} \times 1.5$$

$$P_{\text{ion}}^{\text{night}} < 40 \pm 17 \mu\text{W/m}^2$$

# Jupiter: the IR signal



# Jupiter: the UV signal



Ionizing radiation generates excited  $\text{H}_2$

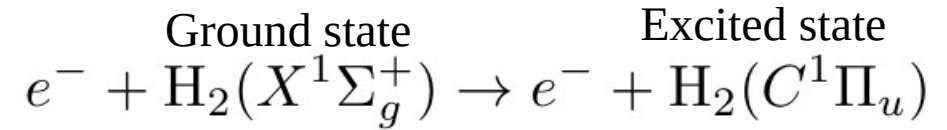
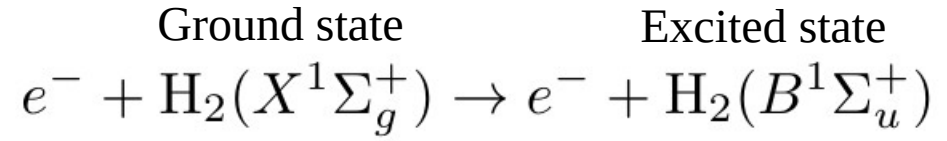
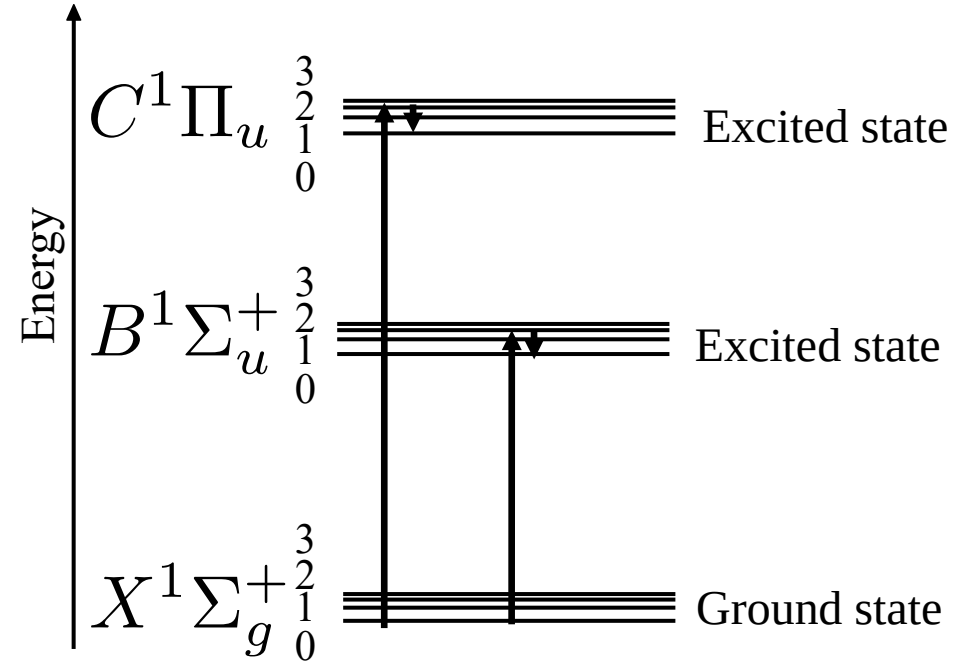
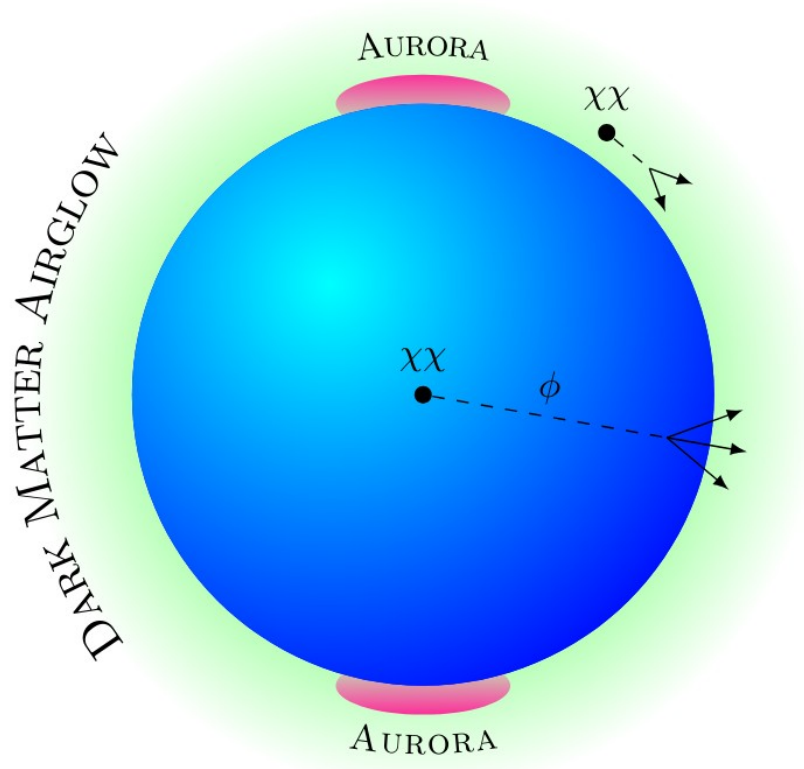


Figure 1. A planet with aurorae (at the poles, magenta), and a dark matter induced ultraviolet airglow (isotropic, green).

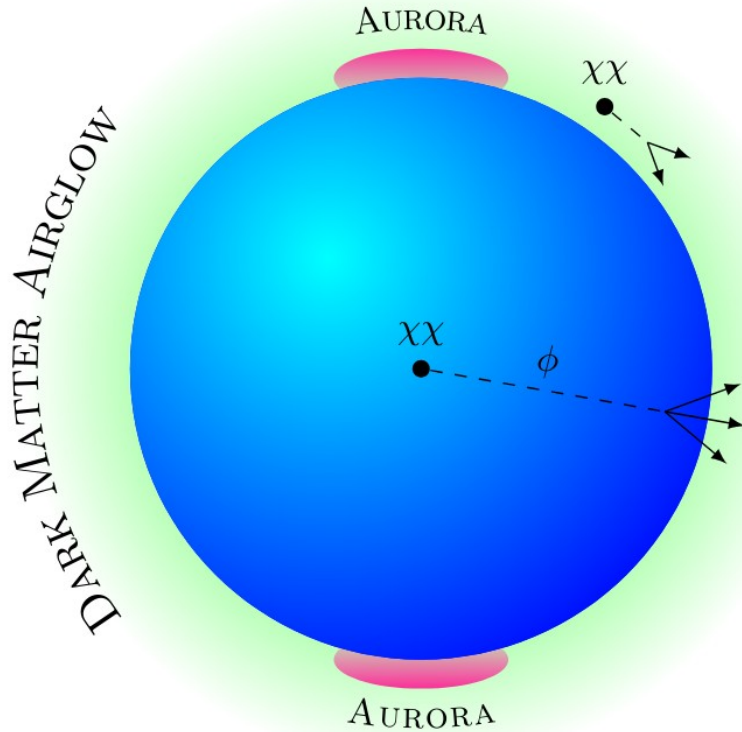
# Jupiter: the UV signal



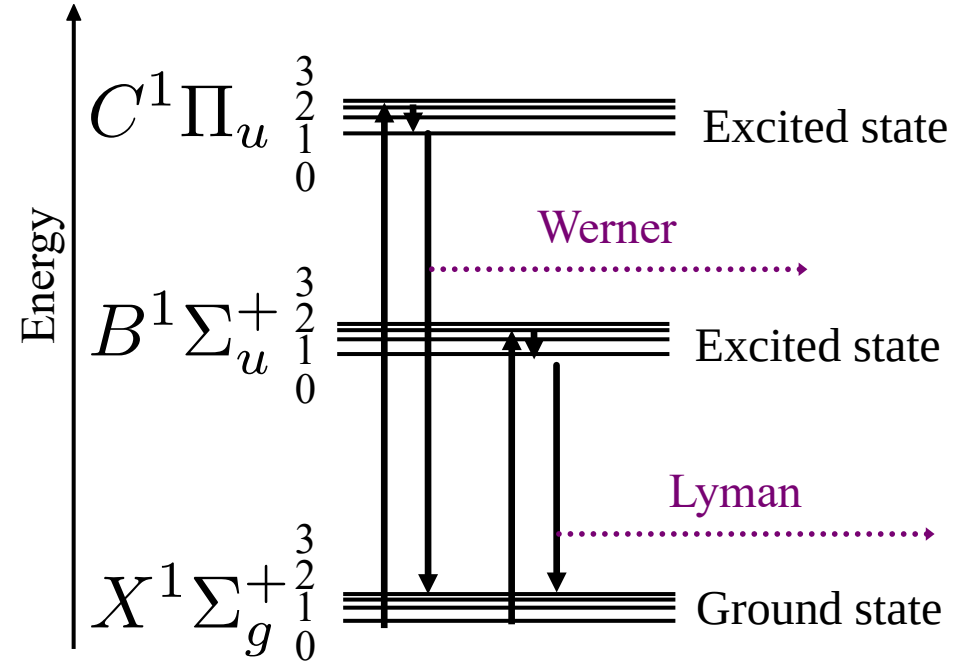
Collisions with charged particles drive excitations



# Jupiter: the UV signal



Singlet excited states have short lifetimes  $\tau \sim \text{ns}$ .



Radiative UV deexcitation  
(Following collisional excitation)

# Jupiter: the UV signal

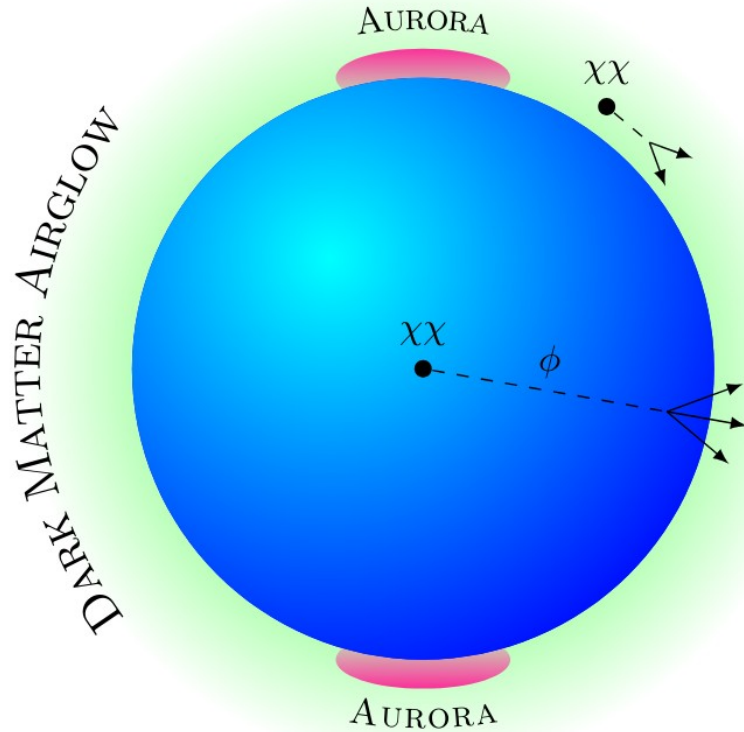
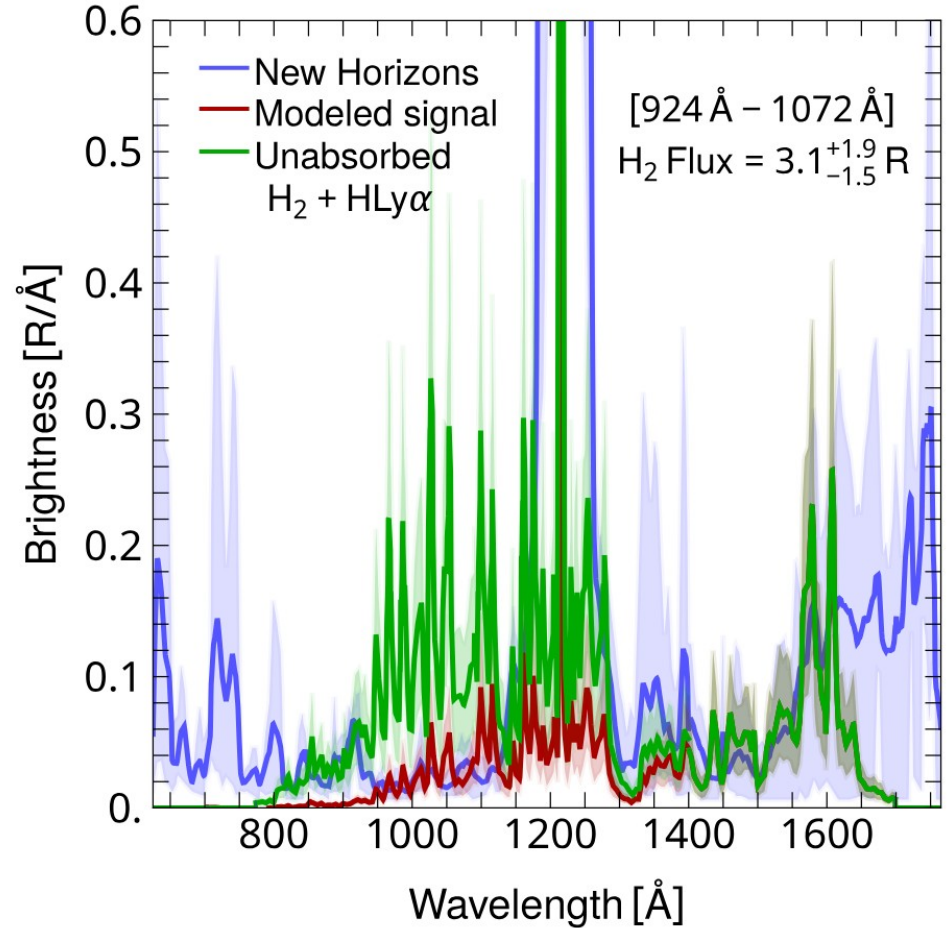
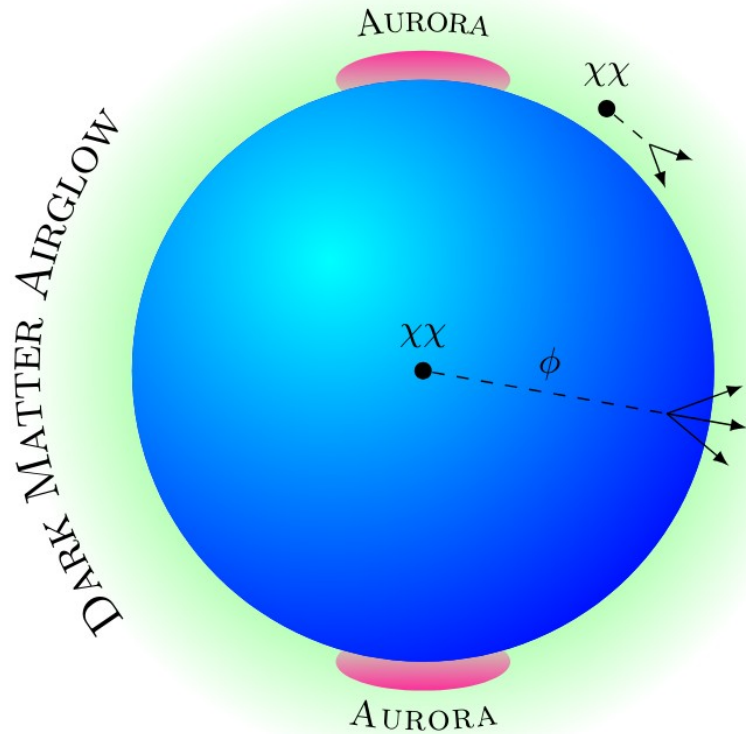


Table I. Measurements of the H<sub>2</sub> Lyman and Werner flux from the nightside equator region of the Solar System's giant planets, and corresponding input power into the planet's upper atmosphere. Uranus and Neptune's flux includes the Rydberg bands from H<sub>2</sub> and may originate from the dayside.

Planet	Flux [R]	Power [ $\mu\text{W}/\text{m}^2$ ]	Space Probe	Ref.
Jupiter	$3.1^{+1.9}_{-1.5}$	$0.31^{+0.19}_{-0.15}$	New Horizons	[22]
Saturn	< 10	< 1	Voyager 1	[13]
Uranus	46	4.6	Voyager 2	[24]
Neptune	$19 \pm 3$	$1.9 \pm 0.3$	Voyager 2	[24]

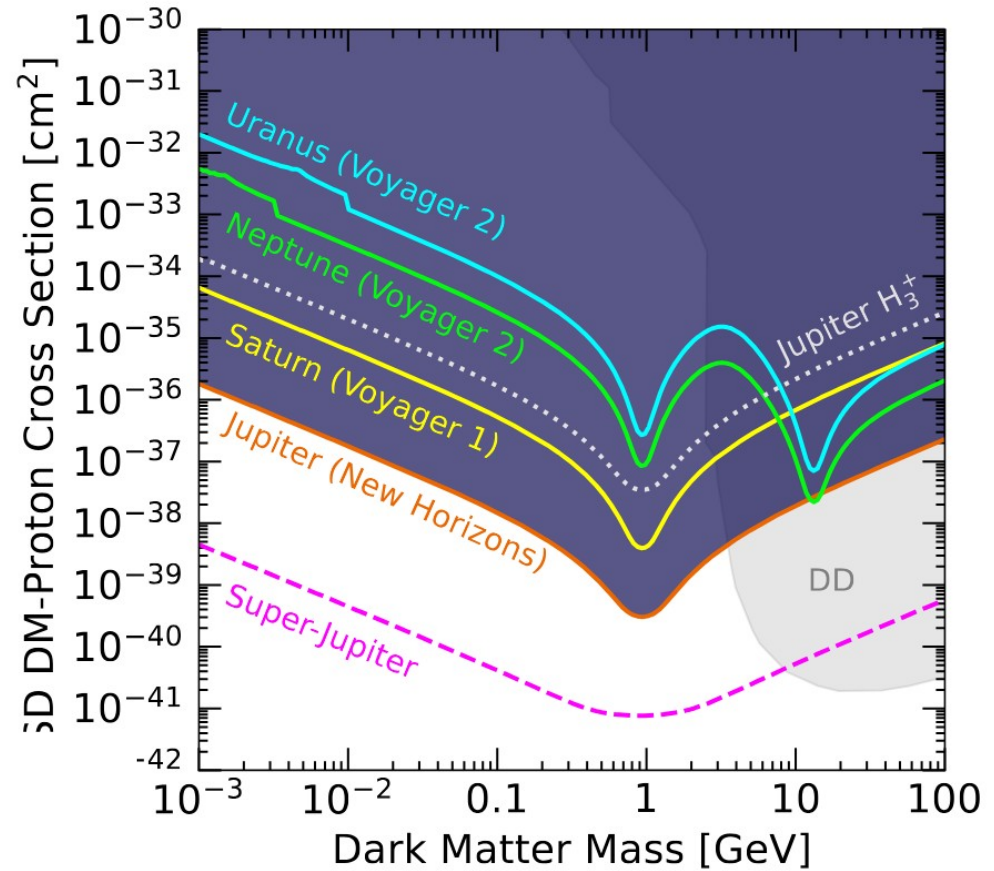
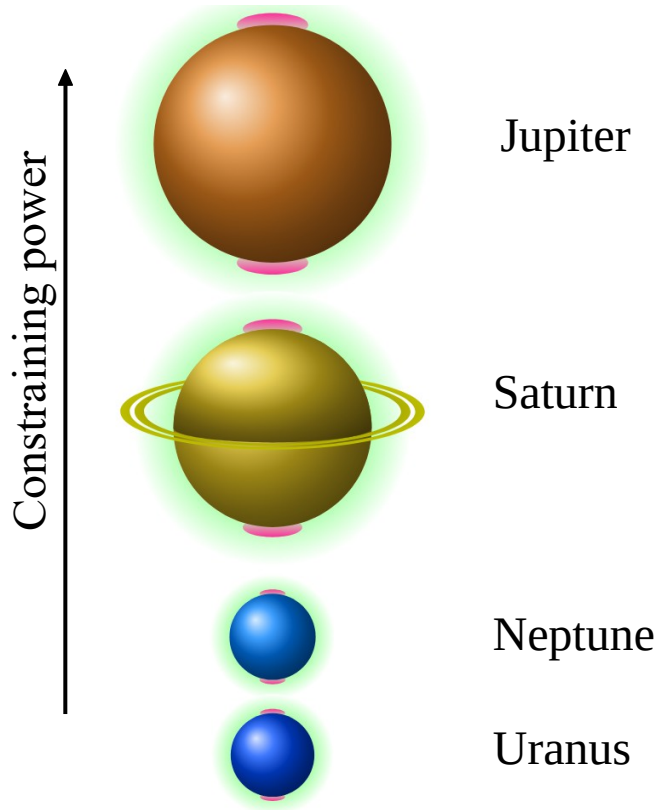
10 R in H<sub>2</sub> UV surface brightness corresponds to about  $1\mu\text{W}/\text{m}^2$  energy in precipitating electrons.

# Jupiter: the UV signal

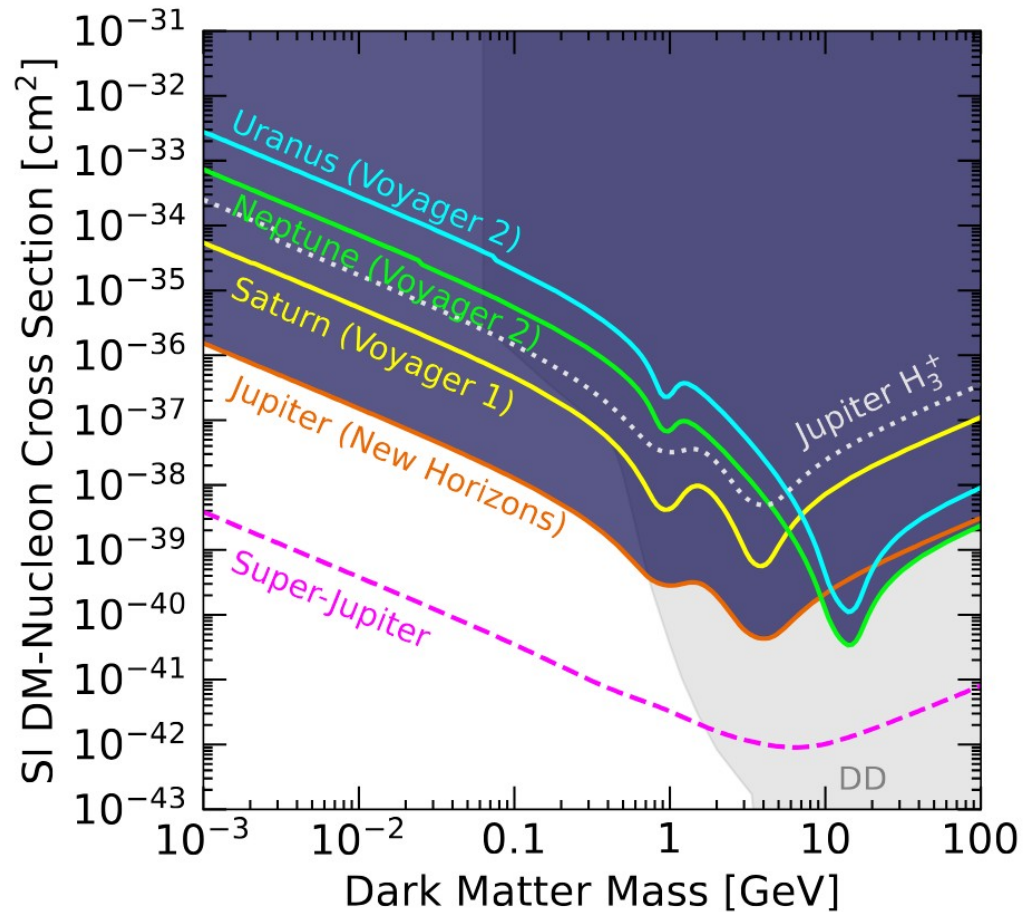
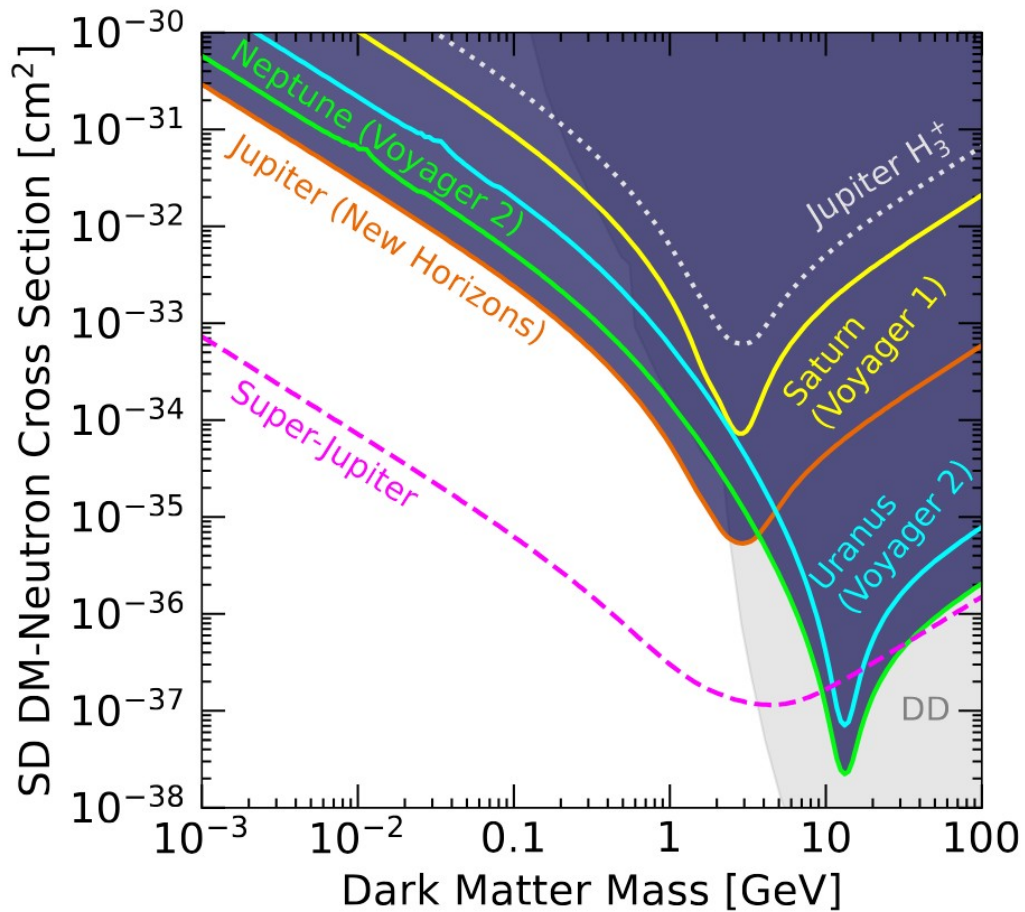


$$\text{Color Ratio} = \frac{I(1550 - 1620 \text{ \AA})}{I(1230 - 1300 \text{ \AA})} = 2.5$$

# Constraining Spin-dependent Scattering



# The UV signal



# The UV signal - Teaser

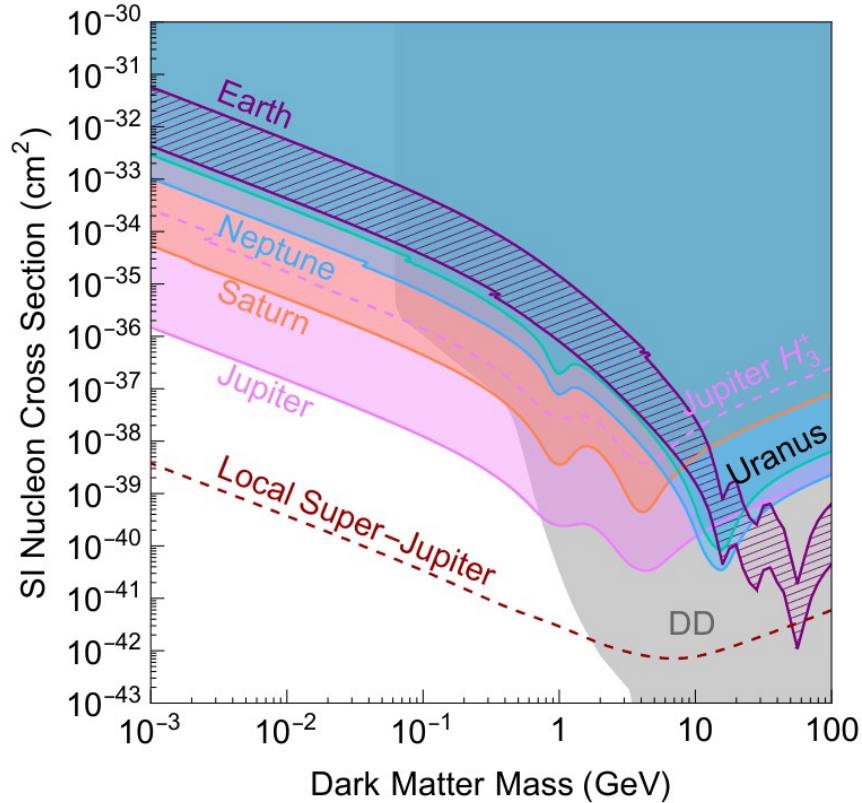


Figure 5. UV airglow limits on spin-independent dark matter interactions for the Solar System's giant planets, Earth, and a potential local Super-Jupiter, as labeled.